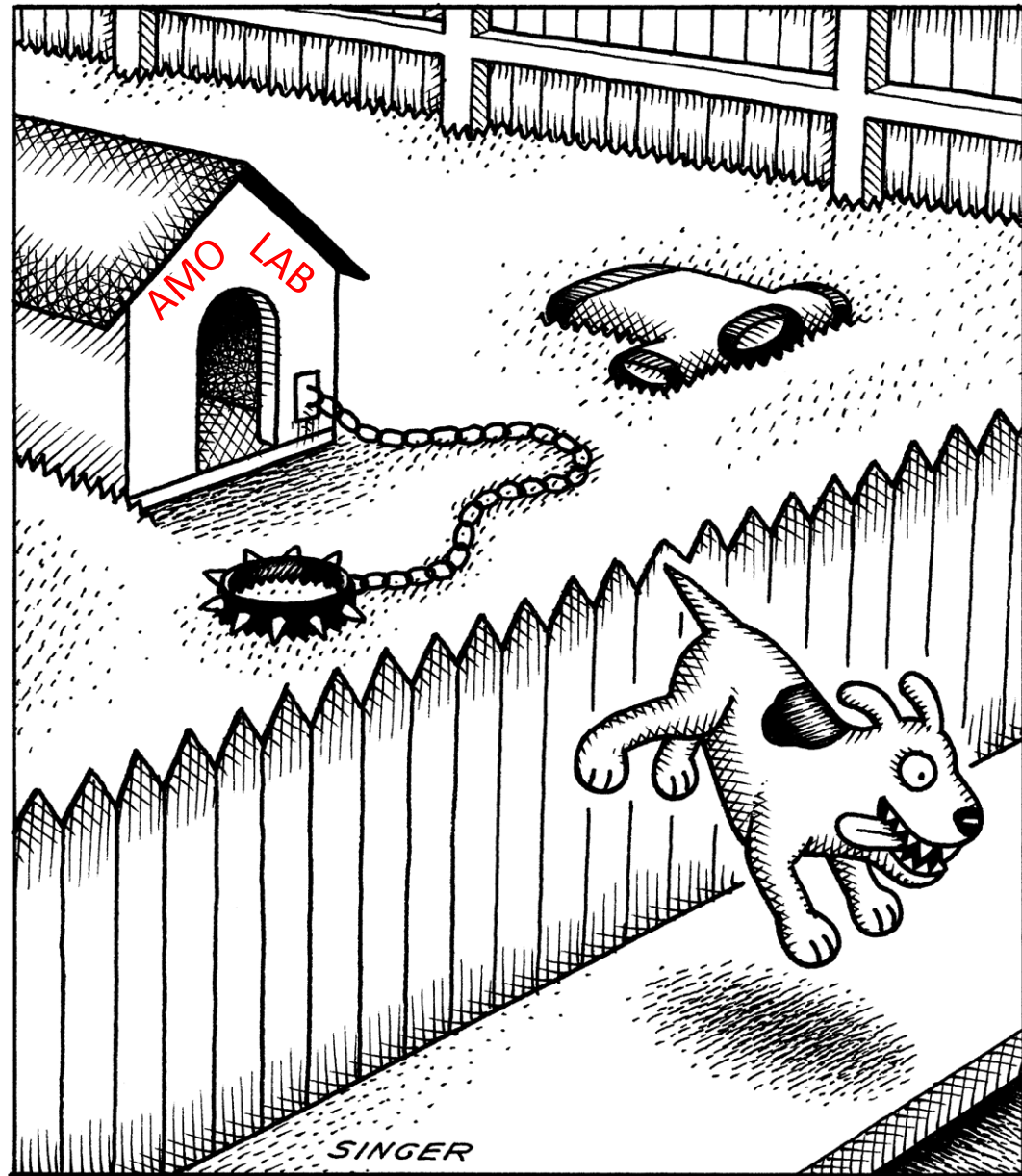
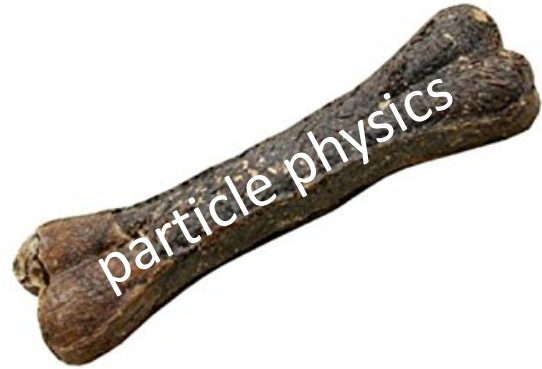
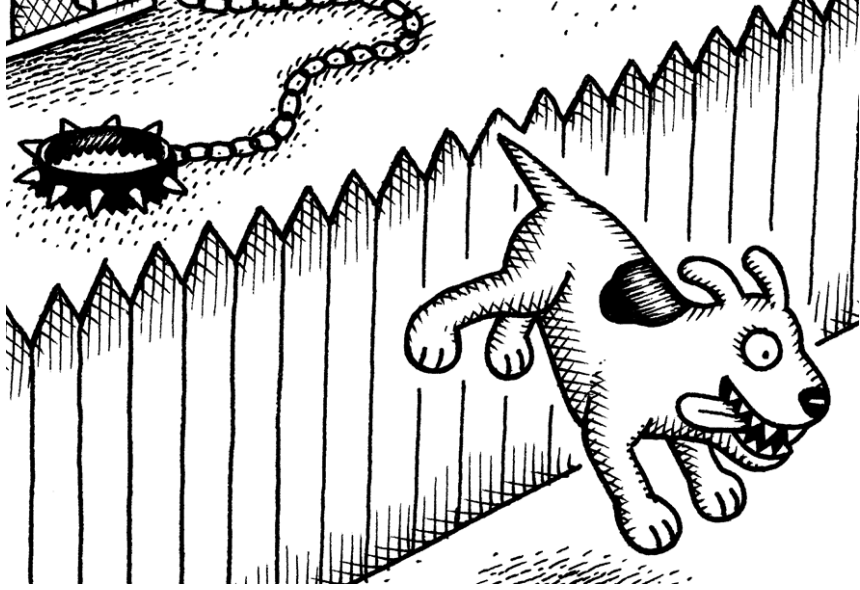


When Atomic
Physicists **Escape**:
Baryogenesis;
Materials; Liquids

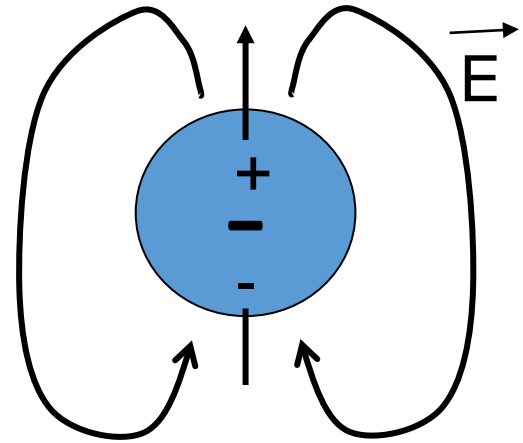


artist: Andy Singer Cagle.com



As of 2013, the electron
electric dipole moment eEDM
was already known to be small
 $d_e < 10^{-27}$ e-cm
 $< 10^{-14} e r_{\text{classical}}$ (Berkeley, Imperial)

Why do better?

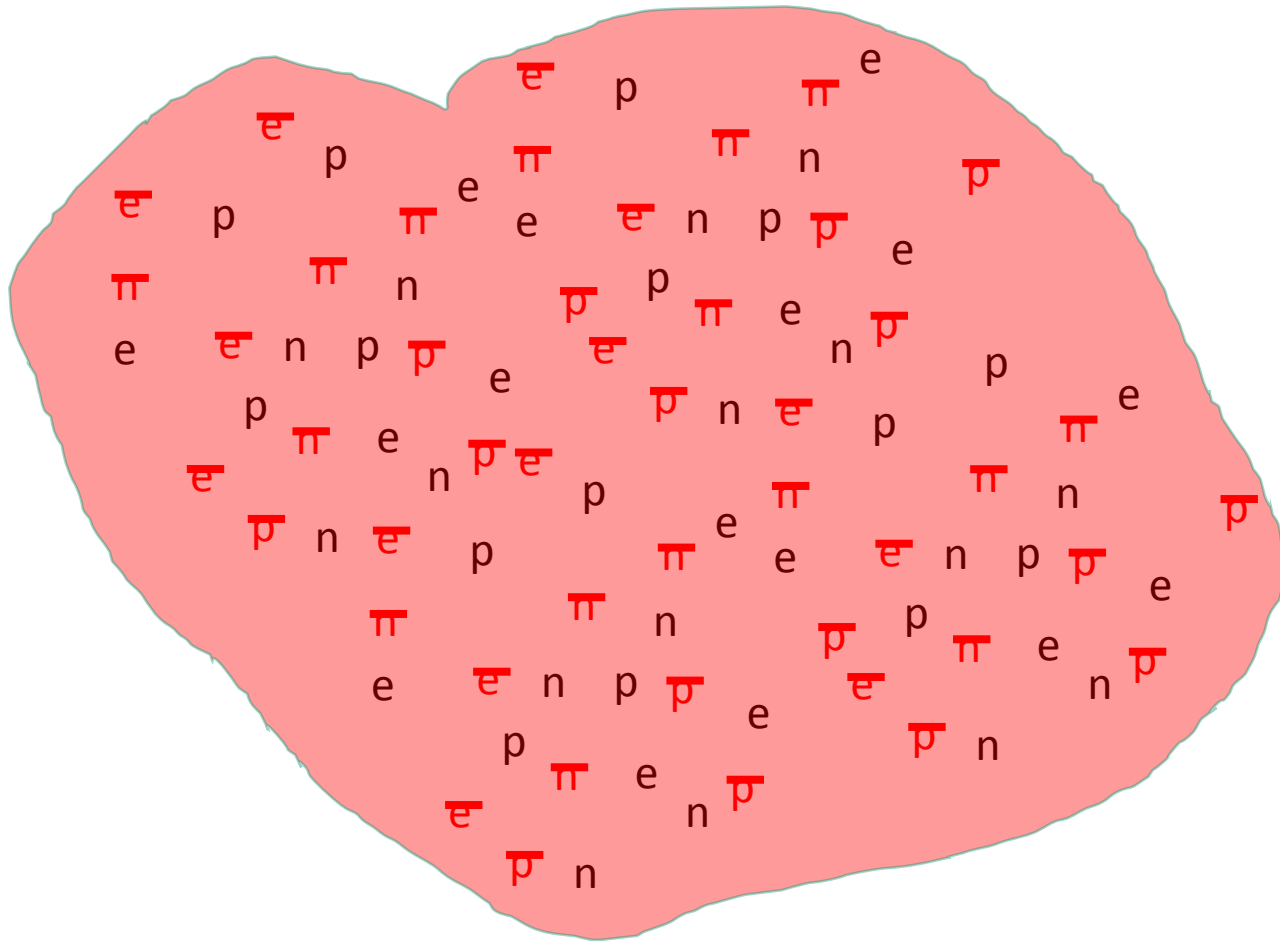


The situation, approximately
14 billion years before right now:

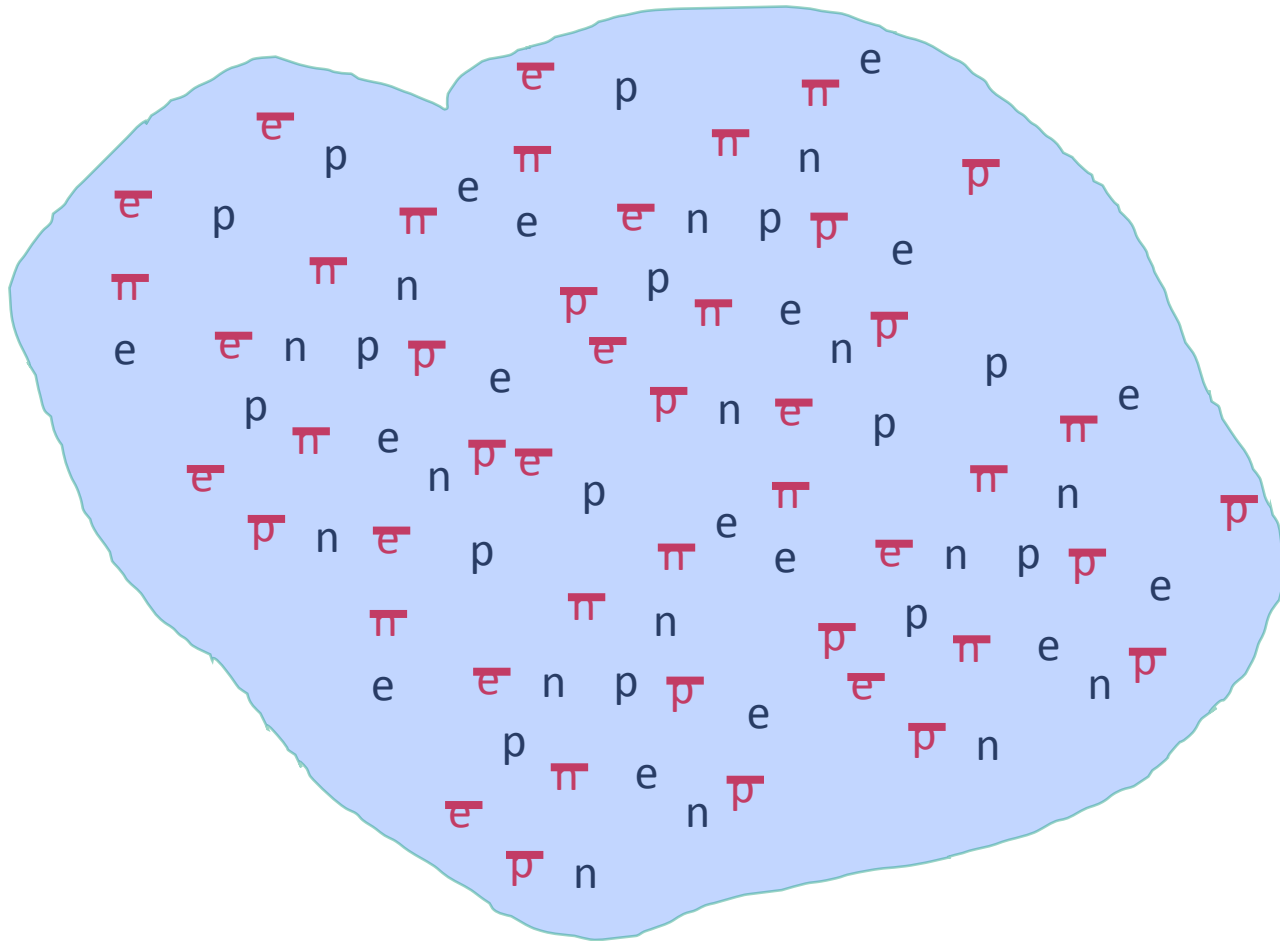


BANG

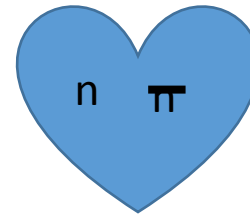
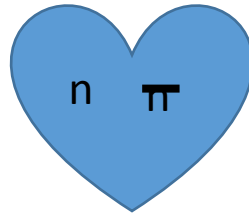
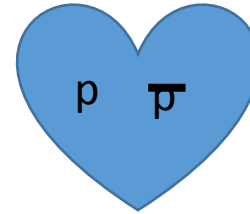
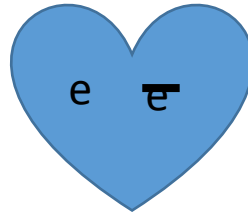
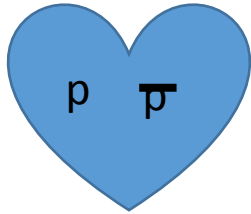
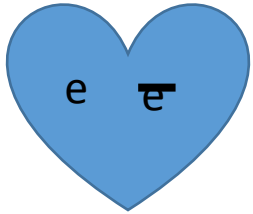
Then, shortly thereafter:

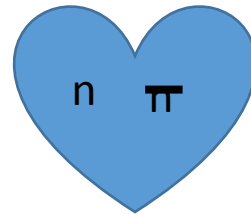
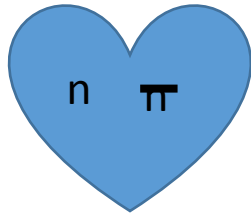
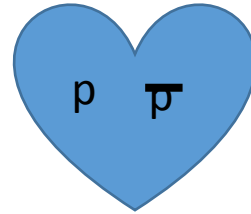
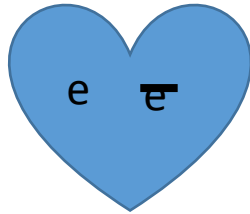
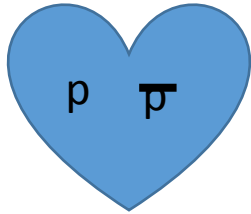


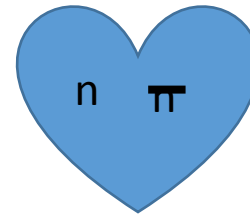
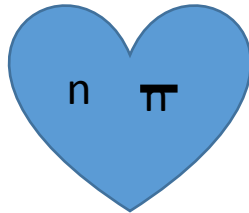
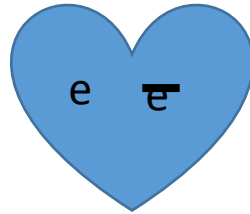
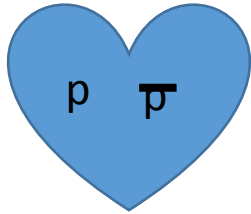
Then, the universe expanded and cooled:



Then, true love!:









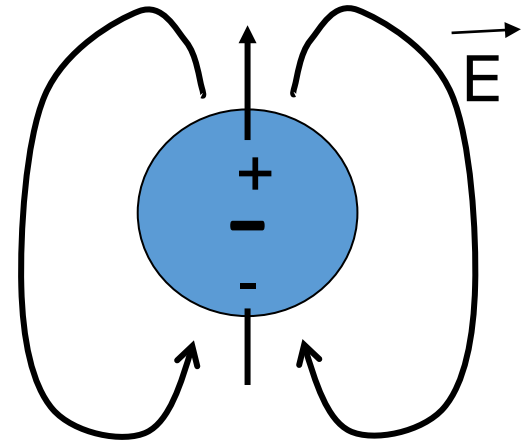
e



p

n

As of 2013, the electron
electric dipole moment eEDM
was already known to be small
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 $< 10^{-14} e r_{\text{classical}}$ (Berkeley, Imperial)

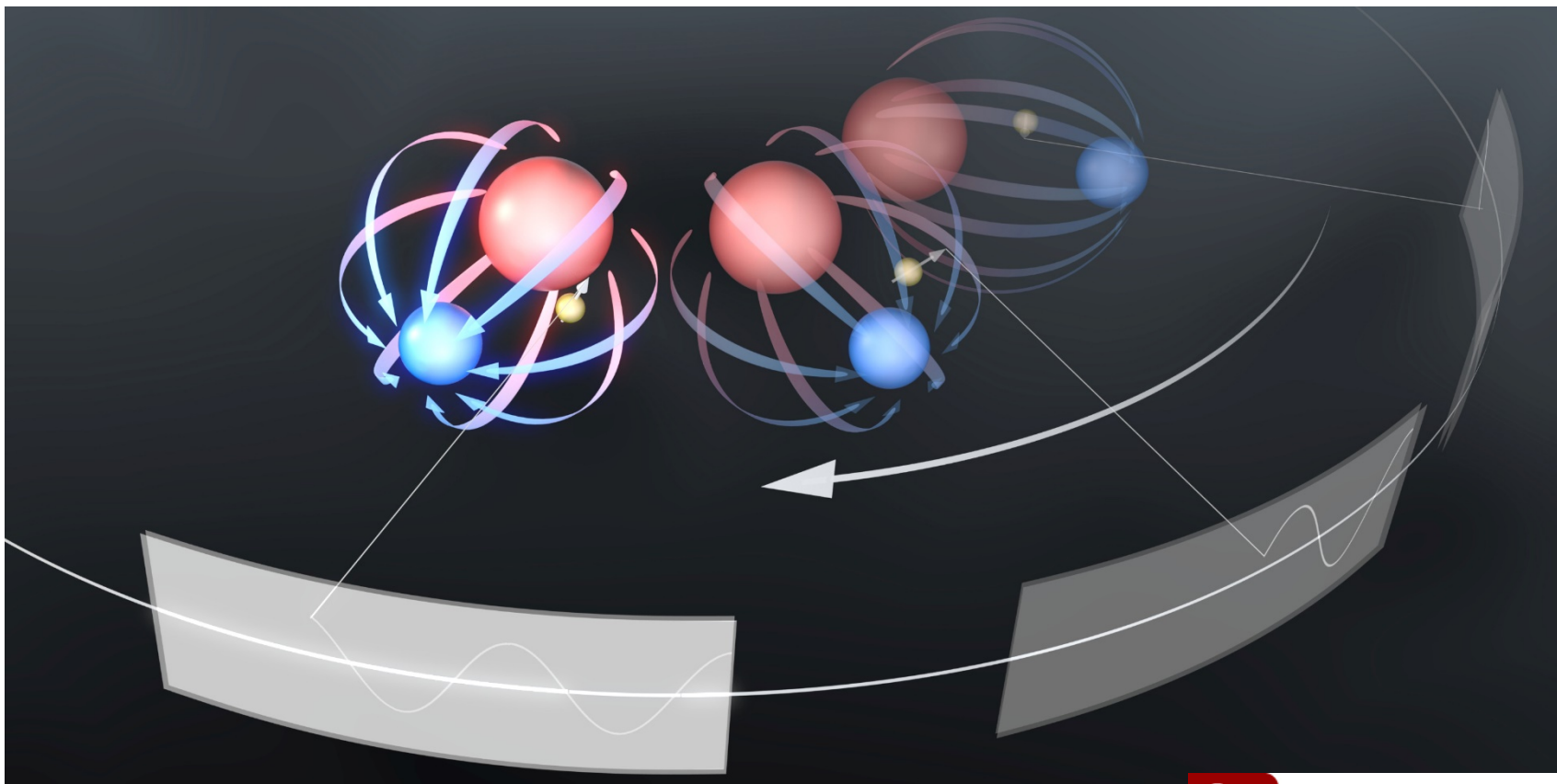


Why do better?

After the big bang cooled, most matter and antimatter annihilated,
but *tiny* bit of matter left over, hooray! Why?

There is CP violation in known present-day particle physics
but not enough. Where is it hiding?

Measuring electron EDM using molecular ions



JILA eEDM collaboration



JILA
NIST/UMD



- *Dr. Yan Zhou*
- *Dr. Yuval Shagam*
- *Kia Boon Ng*
- *Will Cairncross*
- *Dan Gresh*
- *Tanya Roussy*
- *Fatemeh Abbasi-Razgaleh*
- *Jeff Meyers, Kevin Boyce*
- *Jun Ye*
- *Eric Cornell*

Past Group Members

- Laura Sinclair
- Kang-Kuen Ni
- Kevin Cossel
- Russ Stutz
- Aaron Leanhardt
- Yiqi Ni
- Huanqian Loh
- Matt Grau

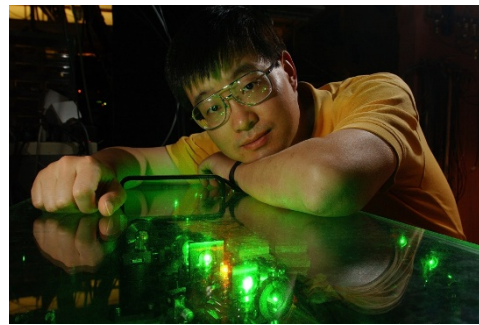
Local theory: John Bohn

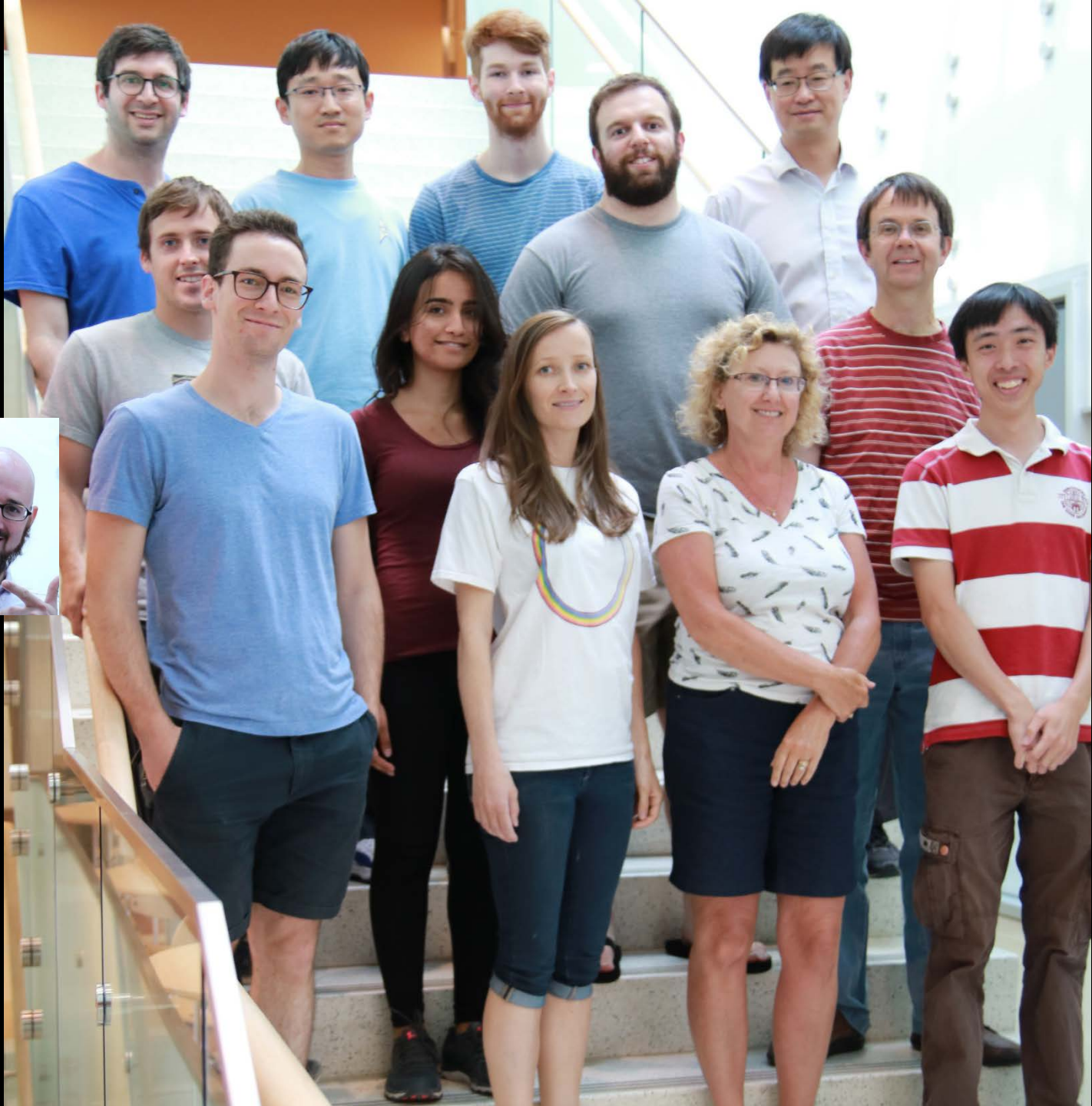
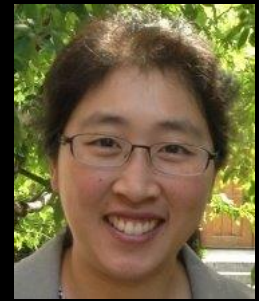
Non local Theory: Bob Field

Still Less Local Theory

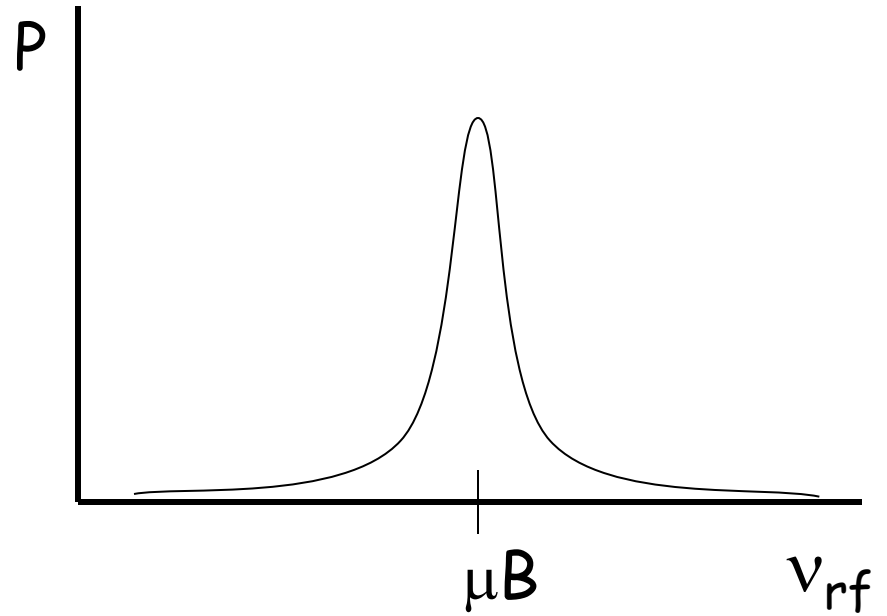
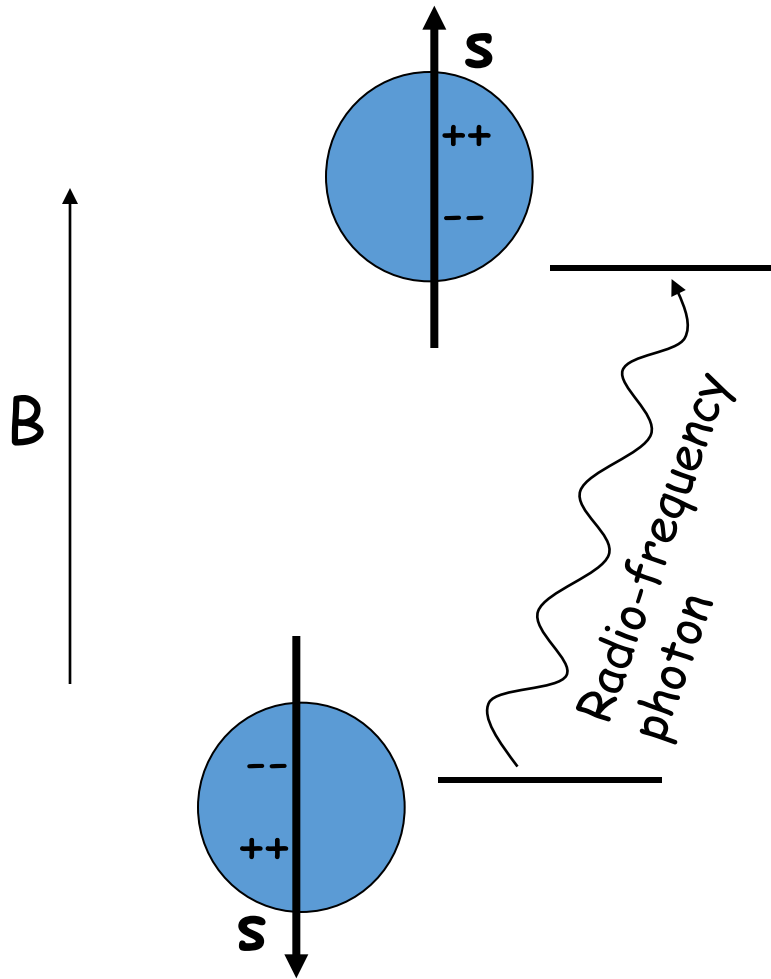
St. Petersburg quantum chemistry group

Thanks: NSF/PFC, NIST,
and Marsico Foundati

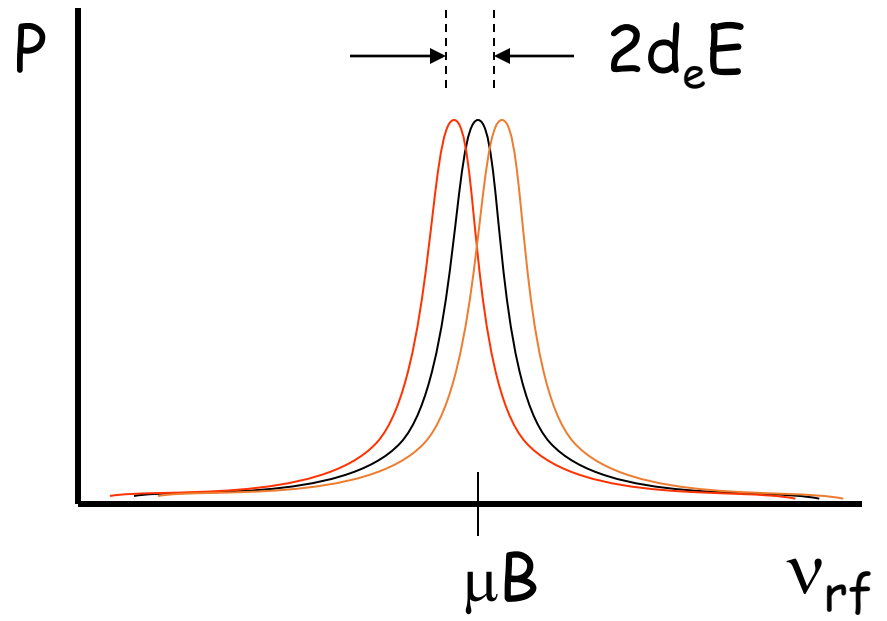
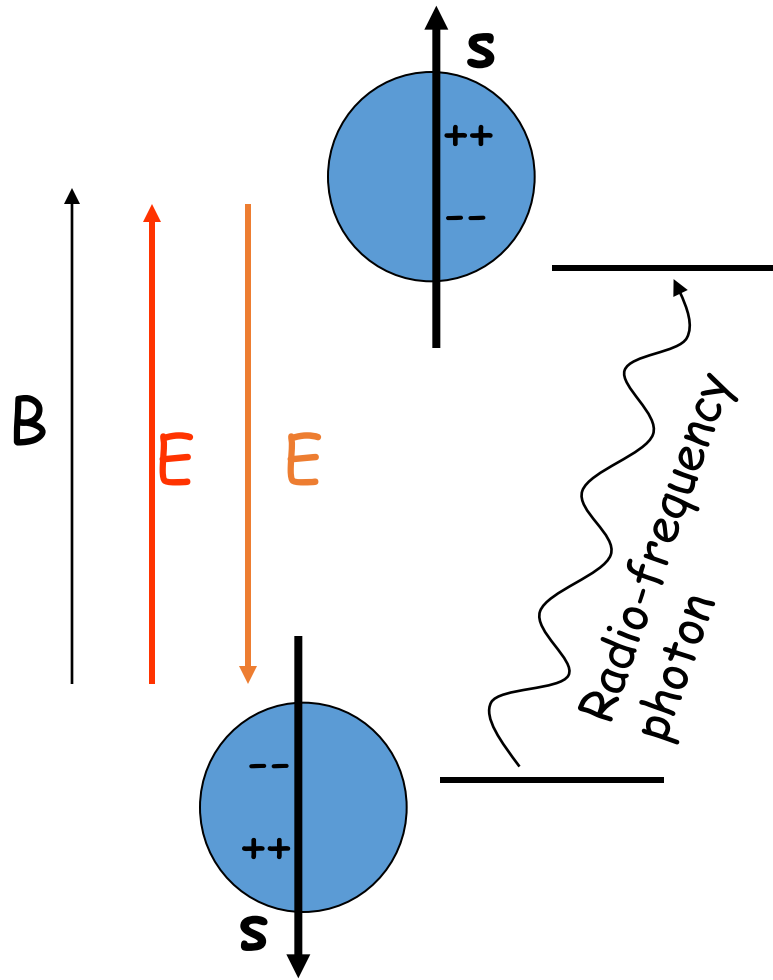




How to measure $eEDM$? First, how do we measure $eMDM$?



How to measure $eEDM$?



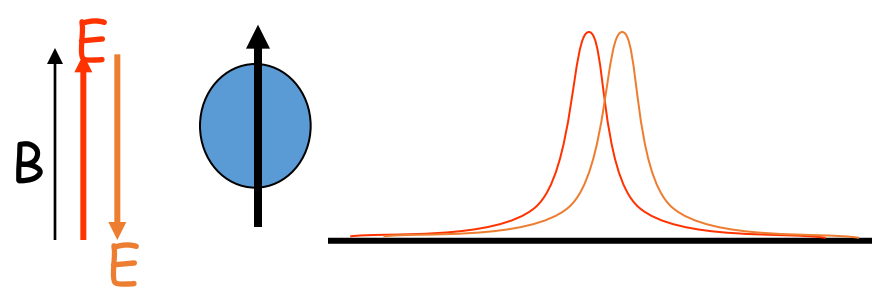


Figure-of-merit:
What makes a good EDM
experiment?

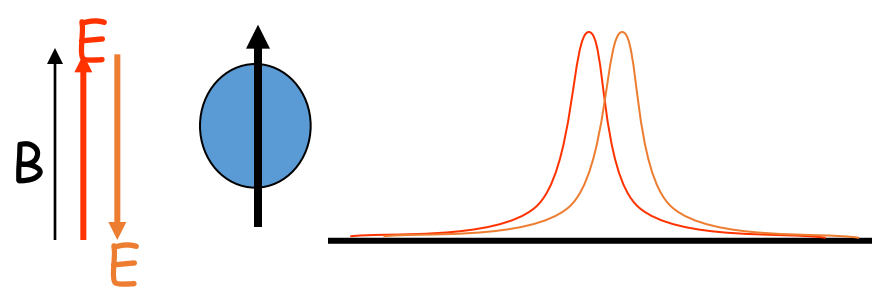
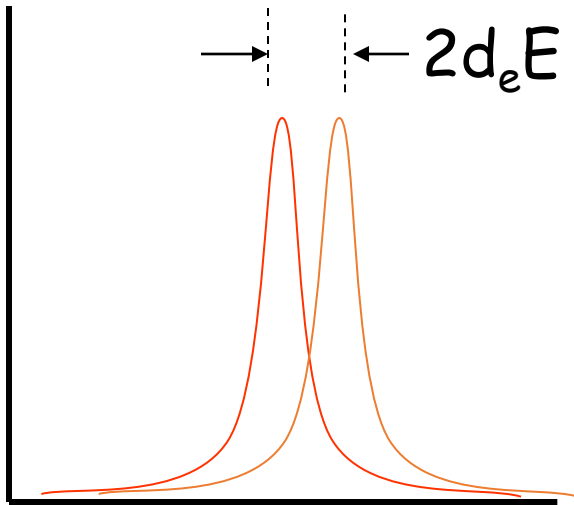
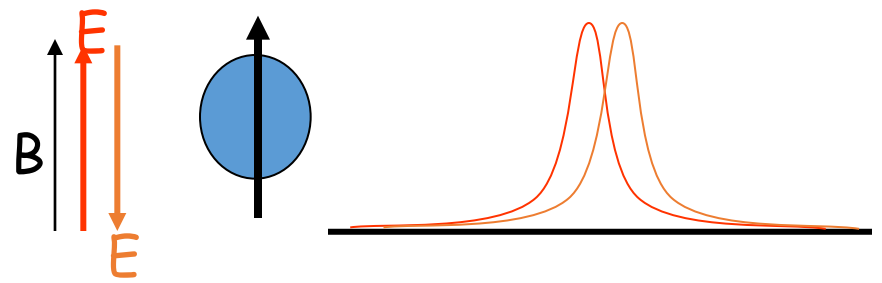
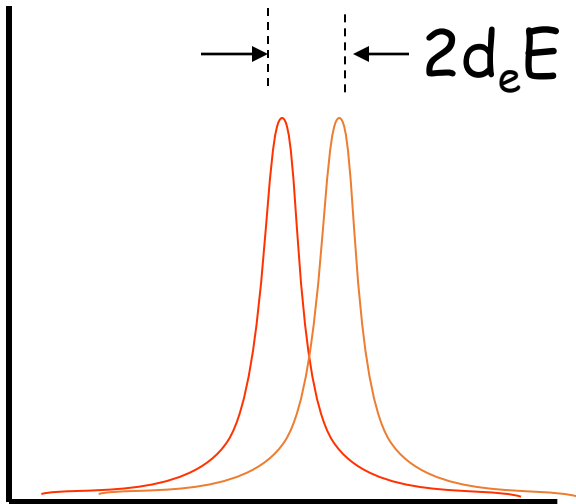
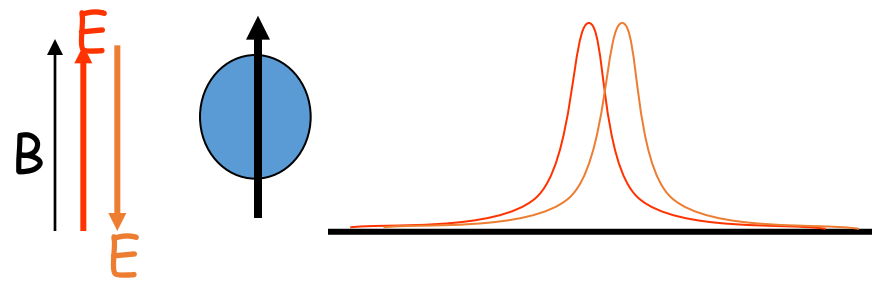


Figure-of-merit:
What makes a good EDM
experiment?

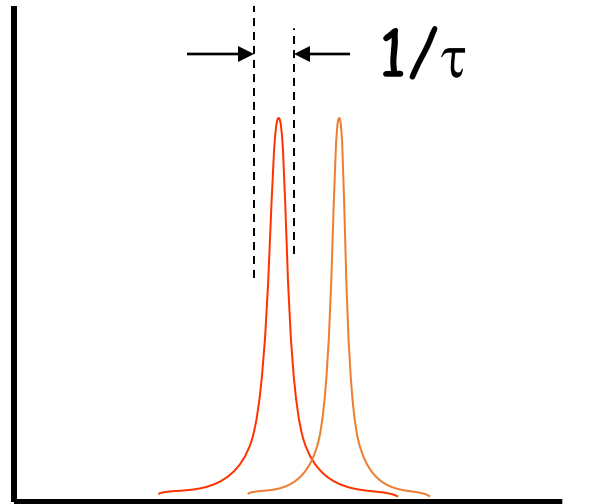


Big Electric
Field!

Figure-of-merit:
What makes a good EDM
experiment?



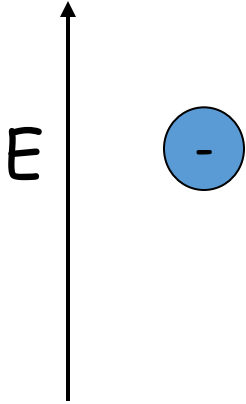
Big Electric
Field!



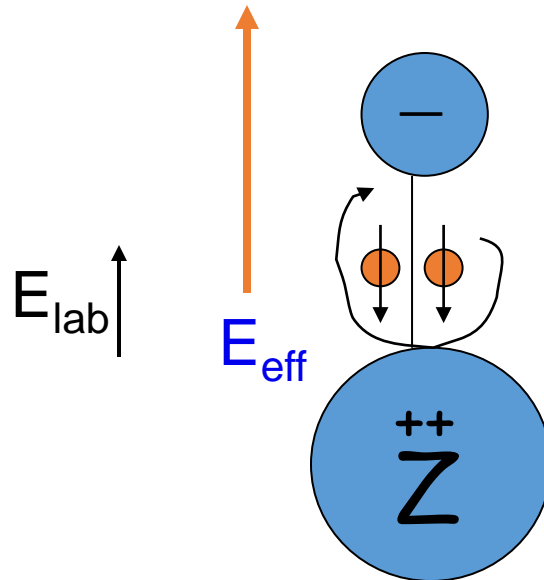
Long Coherence
Time (narrow
resonances)!

Problem:

Big E , long τ . Electron accelerates quickly, and is gone????



Our approach. 1. Use molecule for big E_{eff}
(we follow Hinds and Demille in this)

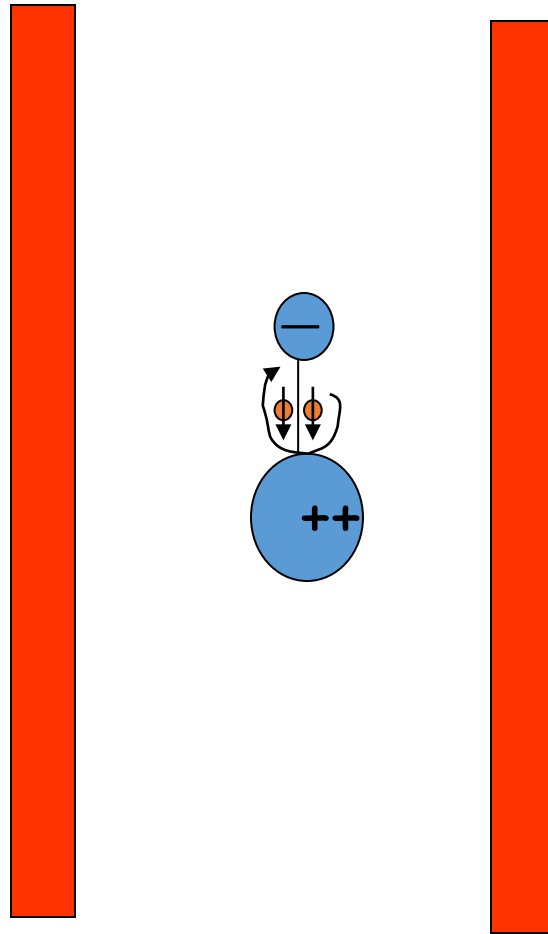


$$E_{\text{lab}} = 10 \text{ V/cm}$$

$$E_{\text{eff}} > 10^{10} \text{ V/cm}$$

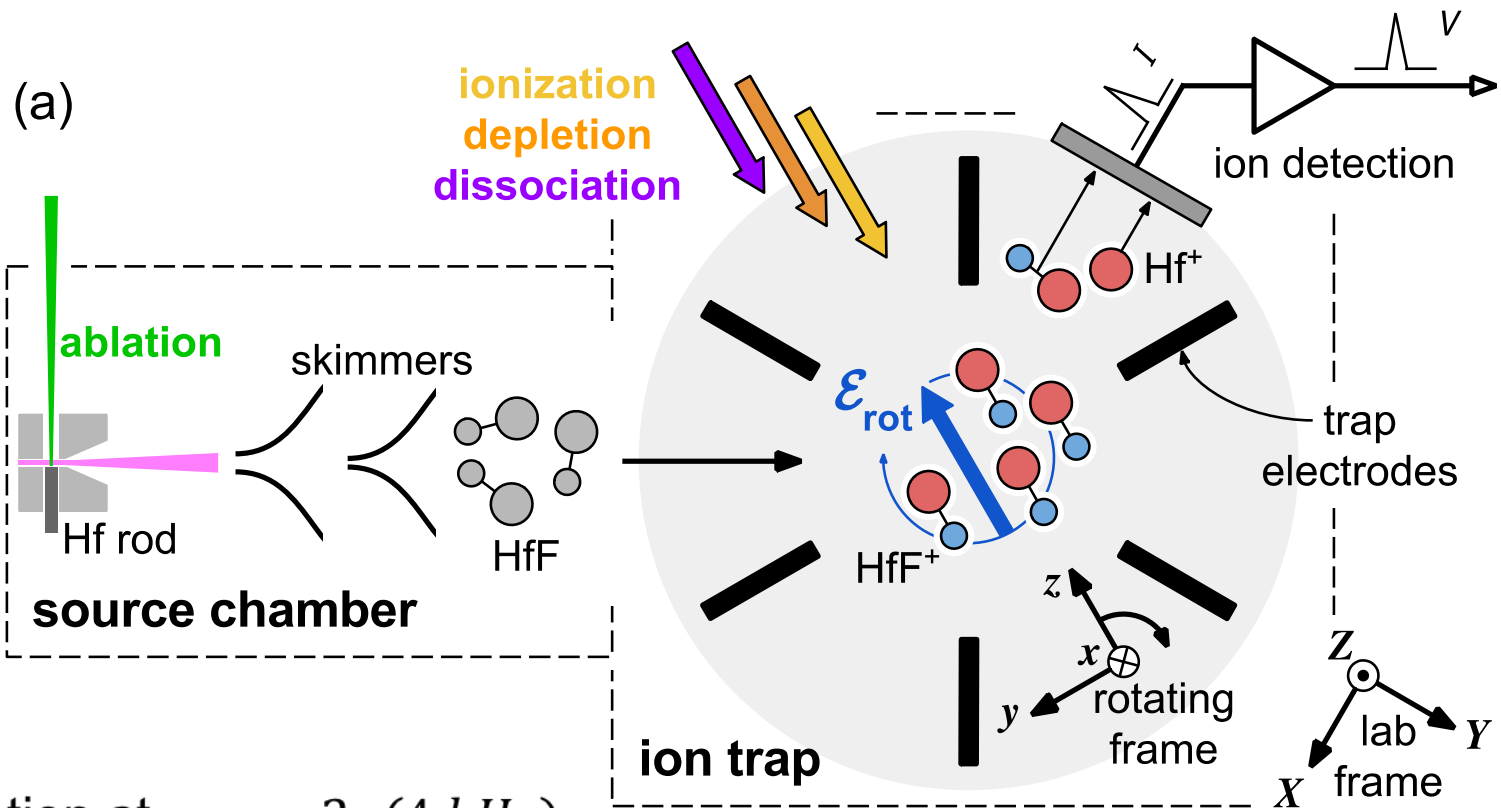
Our approach. 2. Use trapped ion for long τ

(atomic spectroscopy in ion traps sees many seconds)



We will work in
an ion trap.

Apparatus

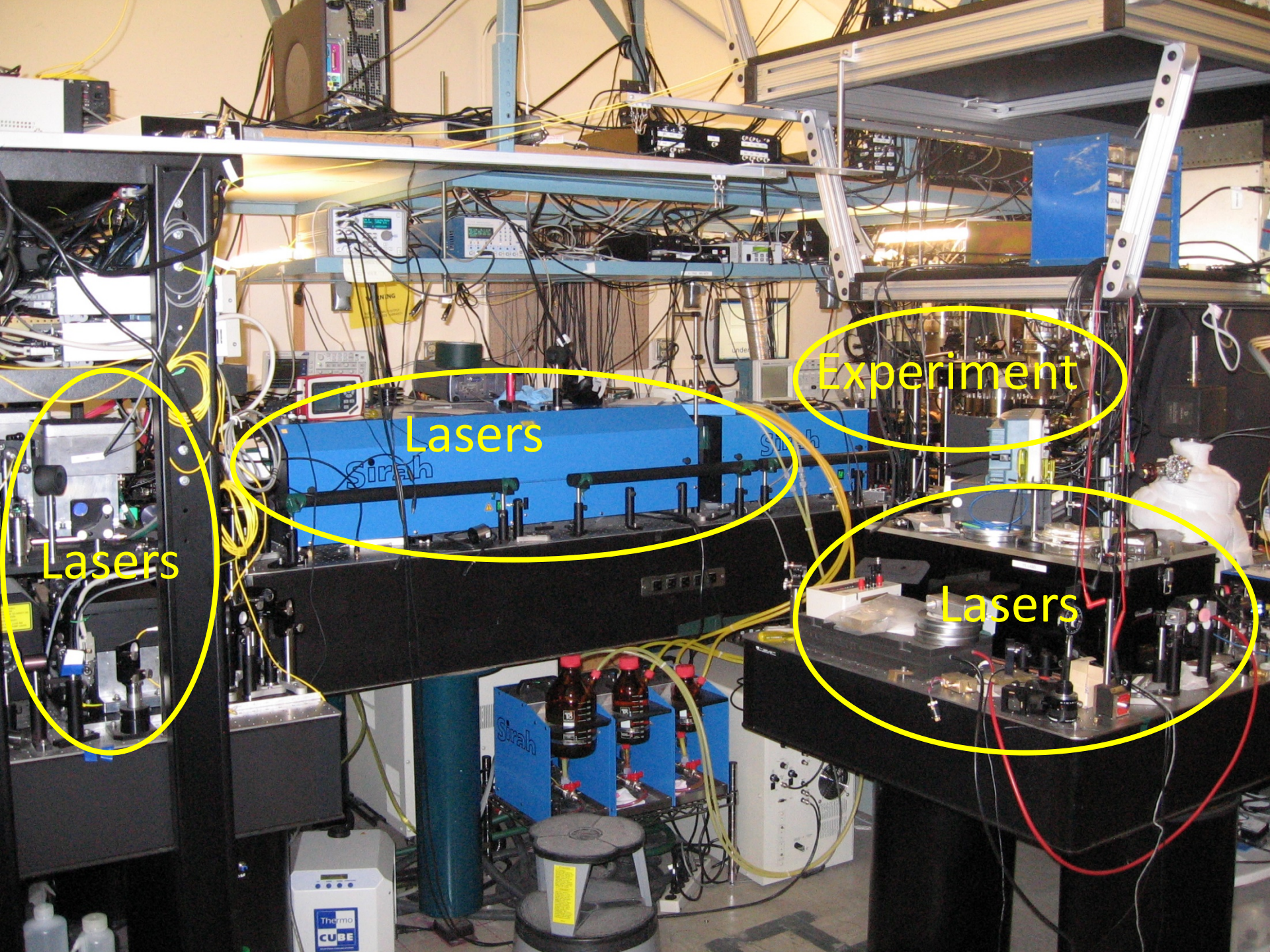


Secular trap motion at $\omega_{sec} \sim 2\pi(4 \text{ kHz})$

“RF” micromotion at $\omega_{rf} = 2\pi(50 \text{ kHz})$

Rotational micromotion at $\omega_{rot} = 2\pi(250 \text{ kHz})$

Rotating magnetic field: not sensitive to DC fields



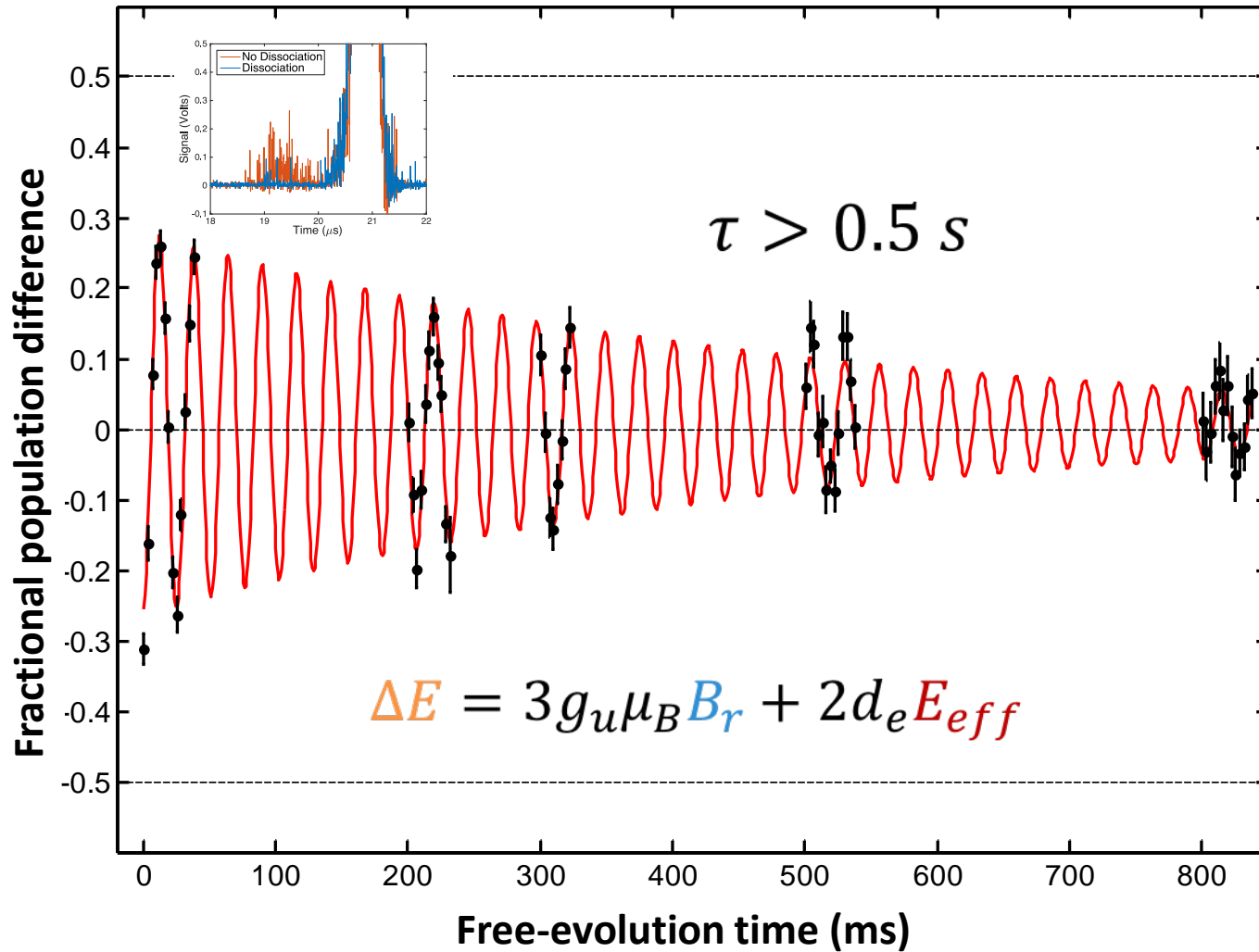
Lasers

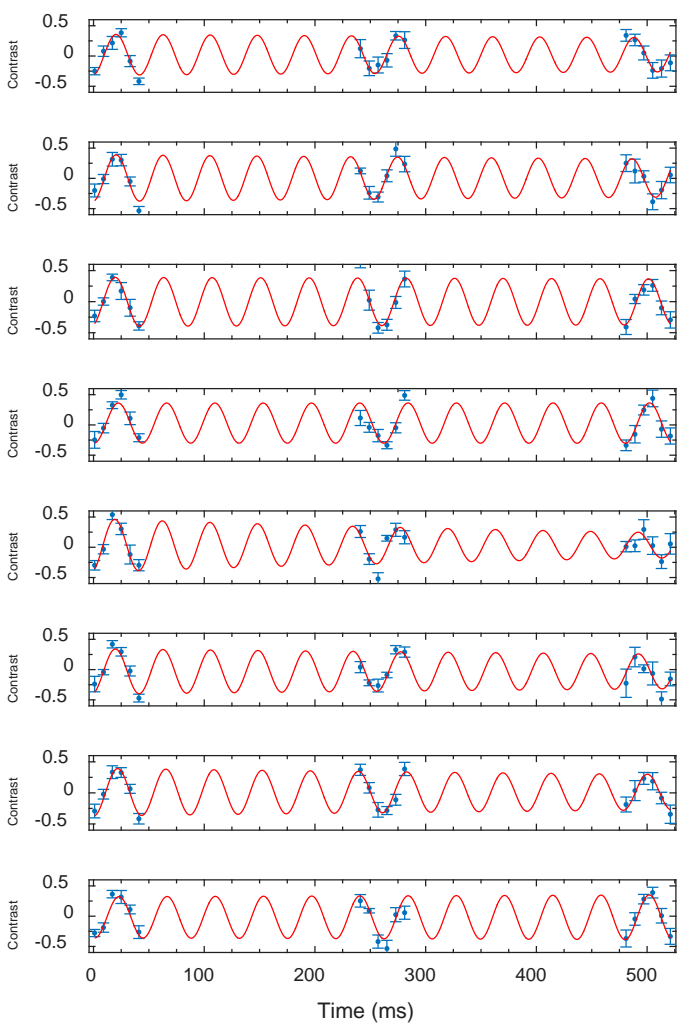
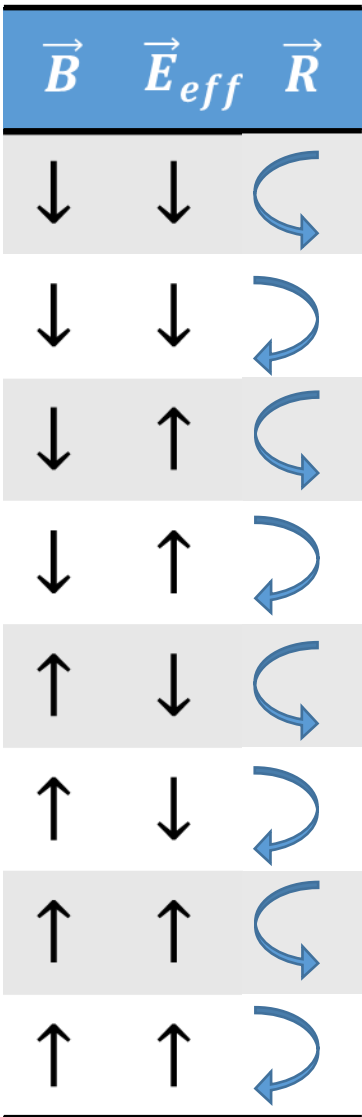
Experiment

Lasers

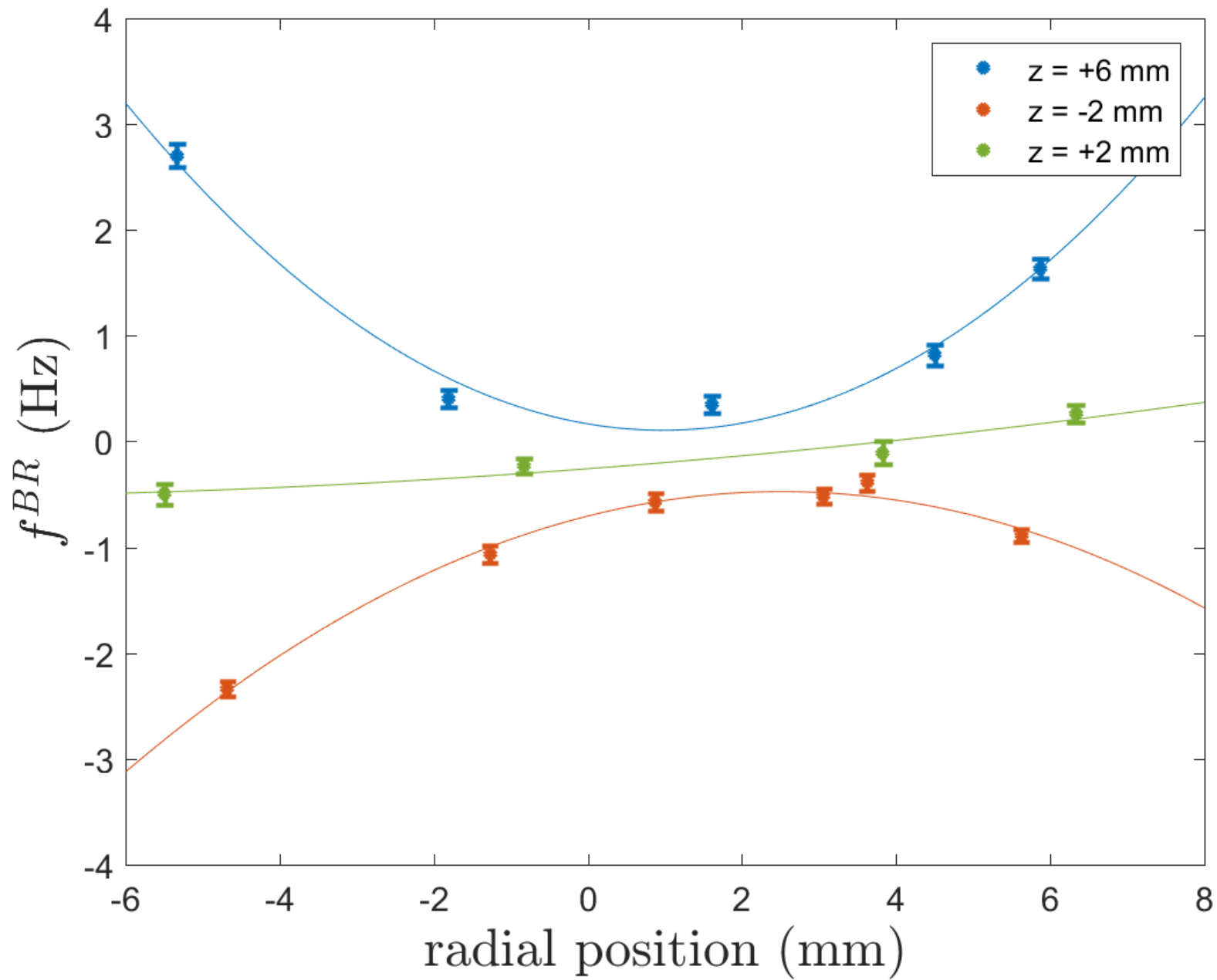
Lasers

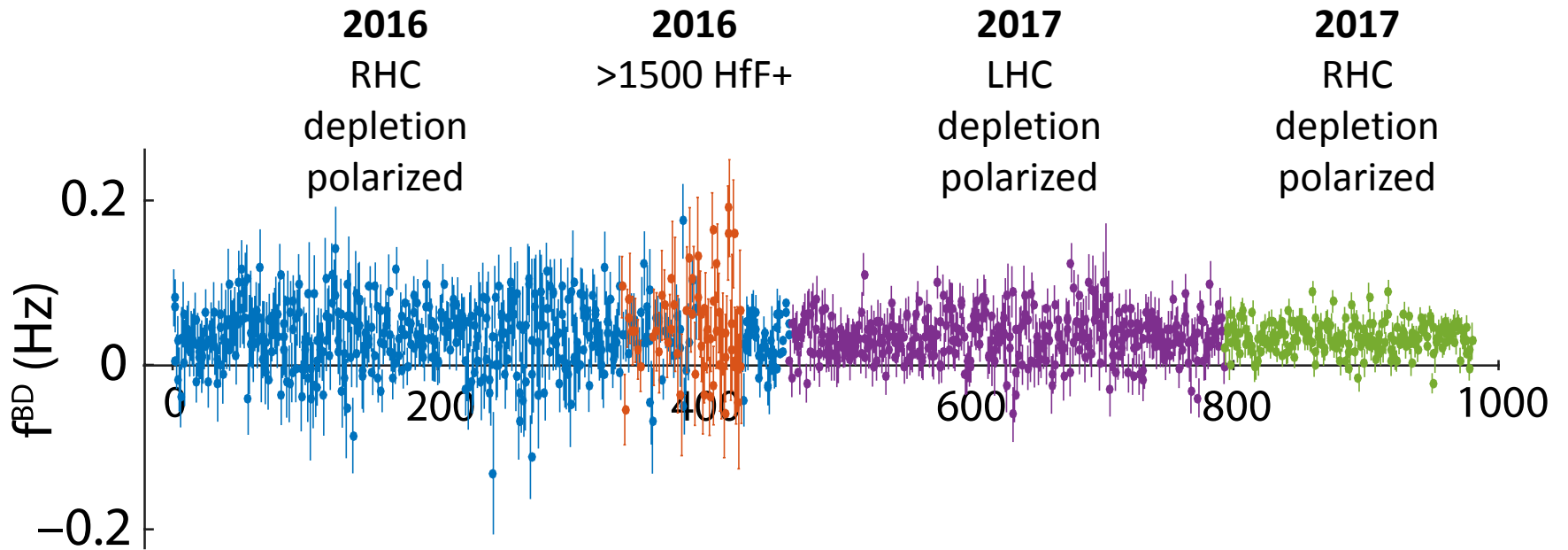
Ramsey Fringe – Electron Spin Resonance





+	-	-	-	+	+	+	-
+	-	-	+	+	-	-	+
+	-	+	-	-	+	-	+
+	-	+	+	-	-	+	-
+	+	-	-	-	-	+	+
+	+	-	+	-	+	-	-
+	+	+	-	+	-	-	-
+	+	+	+	+	+	+	+
f^0	f^B	f^D	f^R	f^{BD}	f^{BR}	f^{DR}	f^{BDR}





Frequency channel	All data	2017 only
f^{R}	2.6(9) mHz	3(1) mHz
f^{DR}	-0.6(8) mHz	-1(1) mHz
f^{BD}	34.5(8) mHz	34.4(1.0) mHz
f^{BDR}	0.4(9) mHz	-0.3(1.0) mHz

Users/willcaincross / Documents / eedm / paper_2016 / data

Figures - Figure 1

Expression	Correction	Uncertainty	Total
1. f_{BD}^{stat}	-1.137e-08	5.130e-08	5.255e-08
2. f_{BD}^{sys}	-2.0157e-08	2.2663e-08	3.0009e-08
3. $f_{BD}^{stat+sys}$	7.2670e-07	1.2209e-08	1.5125e-08
4. $f_{BD}^{stat+sys}$	3.1705e-07	1.8875e-08	3.1765e-07
5. $f_{BD}^{stat+sys}$	5.6240e-08	1.2731e-05	1.2829e-05
6. $f_{BD}^{stat+sys}$	-2.4075e-05	4.1263e-05	4.8225e-05
7. $f_{BD}^{stat+sys}$	-1.0265e-04	1.0265e-04	1.9465e-04

Systematics Parametric plot Histogram Normal Probability Plot

Copy Fig

$f_{BD} = (0.10 \pm 0.87_{stat} \pm 0.20_{sys}) \text{ mHz}$

$d_e = (0.09 \pm 0.77_{stat} \pm 0.18_{sys}) * 1e-28 \text{ e.cm}$

$|d_e| < 1.4 \text{ mHz}$

$|d_e| < 1.3e-28 \text{ e.cm}$

Editor - /Users/willcaincross/Documents/eedm/paper_2016/data/matlab/pr...

```
plot_edm.m process.m
379 -
380 -
381 - end
382 - known_switch_symbols = {'B', 'D', 'R', 'P', 'S'};
383 - [~,subs] = ismember(switch_labels,known_switch_labels);
384 - switch_symbols = known_switch_symbols(subs);
385 -
386 - %% Process data
387 - % Choose a column to average over identical entries. We
388 - % combine all the data rows that have the same block index
389 - % block = unique(block);
390 - % ublock = unique(block);
391 - % nblock = numel(ublock);
392 -
393 - % loop through unique blocks
394 - lcon_table = struct();
395 - for b = 1:nblock
396 -     lcon_table = struct2table(lcon_table);
397 -
398 - %% Find the EDM channel and apply the blind.
399 - % if any(ismember(lcon_table.Properties.VariableNames,'f_B
400 - % if exist('B Blind/B Blind', 'file')
401 - % % if the blind exists, load it into memory and a
402 - % % we apply one blind for all data taken before 2
403 - % blind_set1 = lcon_table.time_stamp
404 - % lcon_table.f_B(blind_set1) =
405 - %
406 - % apply a second, larger bl
407 - % an uncertainty of 5*10^-2
408 - % blind_set2 = lcon_table.time
409 - % lcon_table.f_B(blind_set2)
410 - else
411 - % if the blind doesn't exist
412 - % freakishly large blind (4e
413 - % 2*(10^-27 cm) = 23.3e9 et
414 - % = 0.012678304;
415 - % lcon_table.f_BD = lcon_
416 - % fprintf('BLIND not found
417 - %
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$$f_{\text{BD}} = (0.10 \pm 0.87_{\text{stat}} \pm 0.20_{\text{syst}}) \text{ mHz}$$

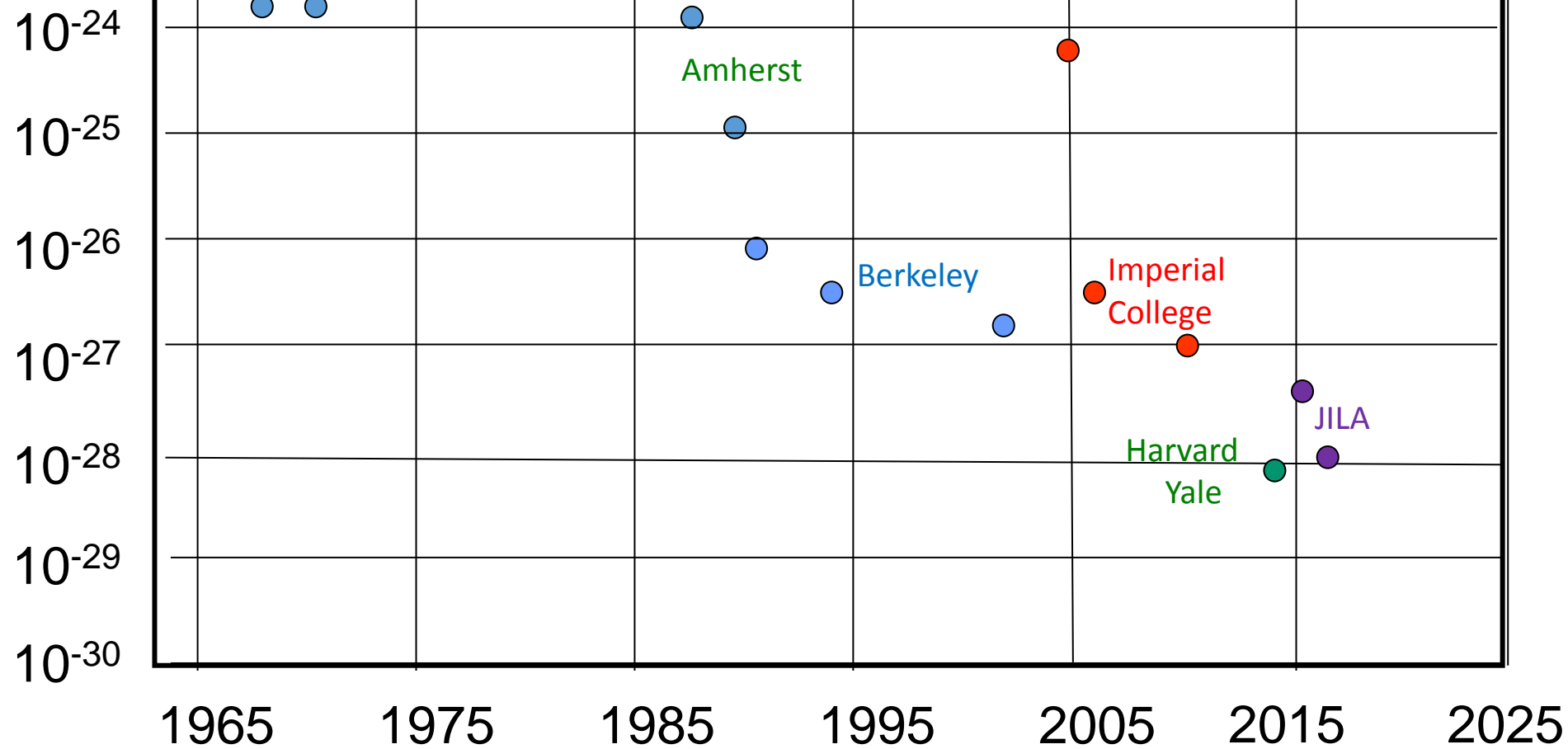
$$d_e = (0.09 \pm 0.77_{\text{stat}} \pm 0.18_{\text{syst}}) \times 10^{-28} \text{ e.cm}$$

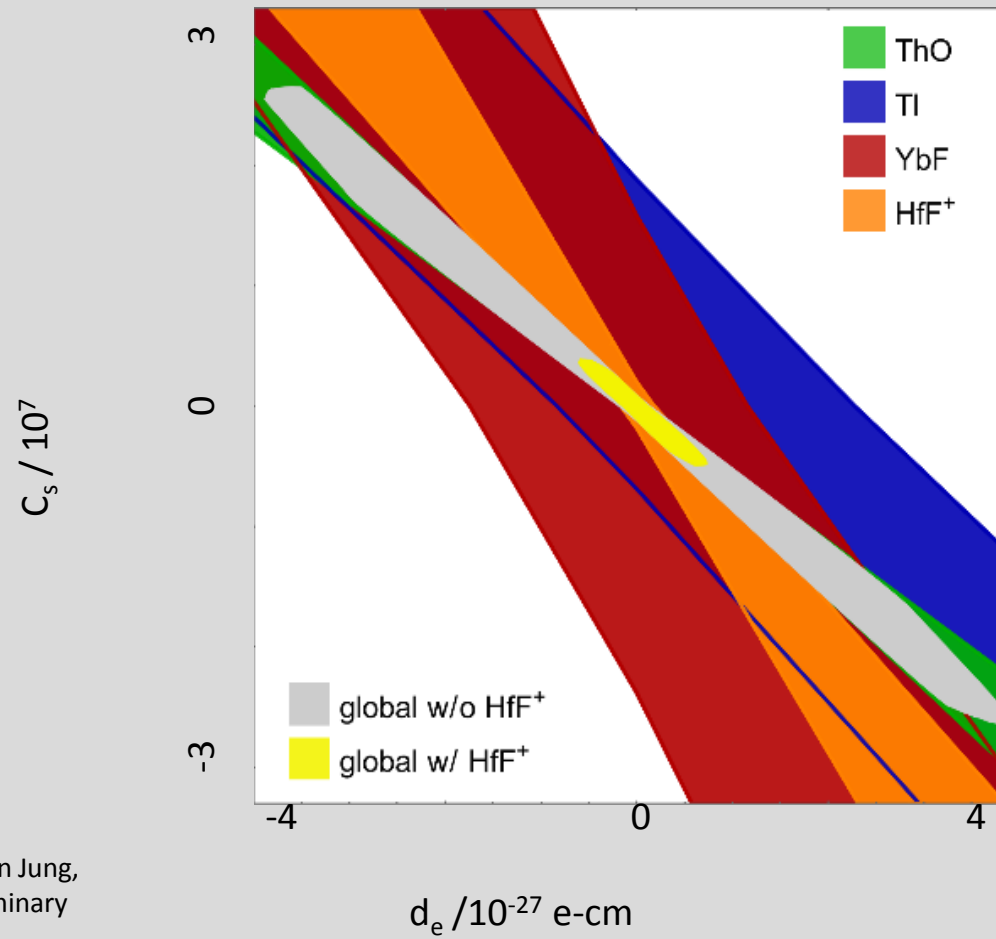
$$|d_e| < 1.4 \text{ mHz}$$

$$|d_e| < 1.3 \times 10^{-28} \text{ e.cm}$$



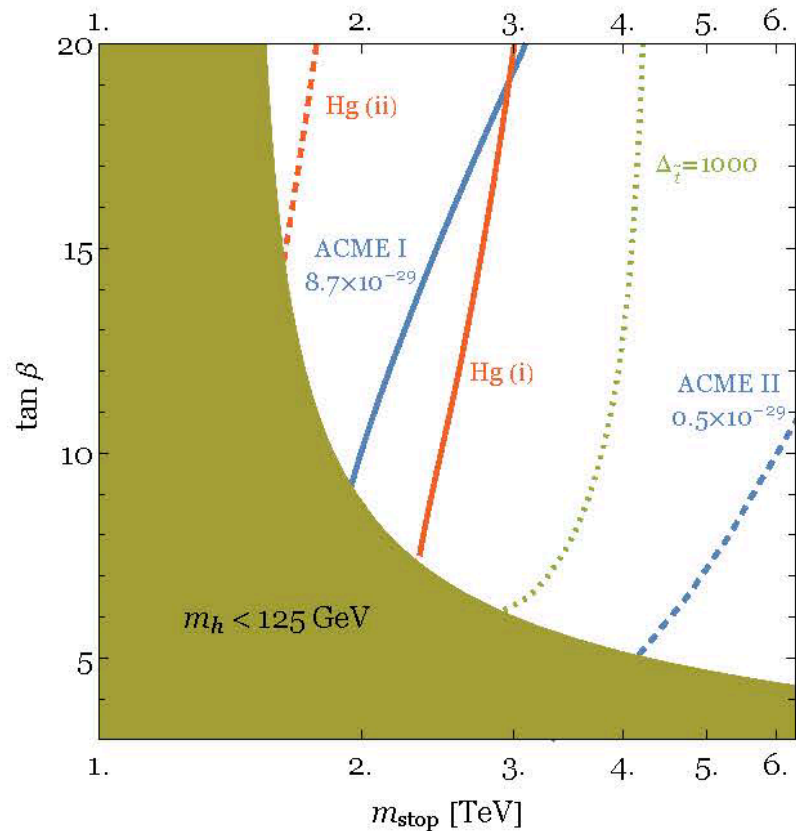
Limit on
eEDM
(e-cm)





Martin Jung,
Preliminary

$$d_{\text{ThO,Hg}}/e \text{ [cm]}, \arg(A_t\mu)=\pi/2, m_A=400 \text{ [GeV]}$$



$$d_{\text{ThO,Hg}}/e \text{ [cm]}, \arg(A_t\mu)=\pi/2, \tan\beta = 10$$

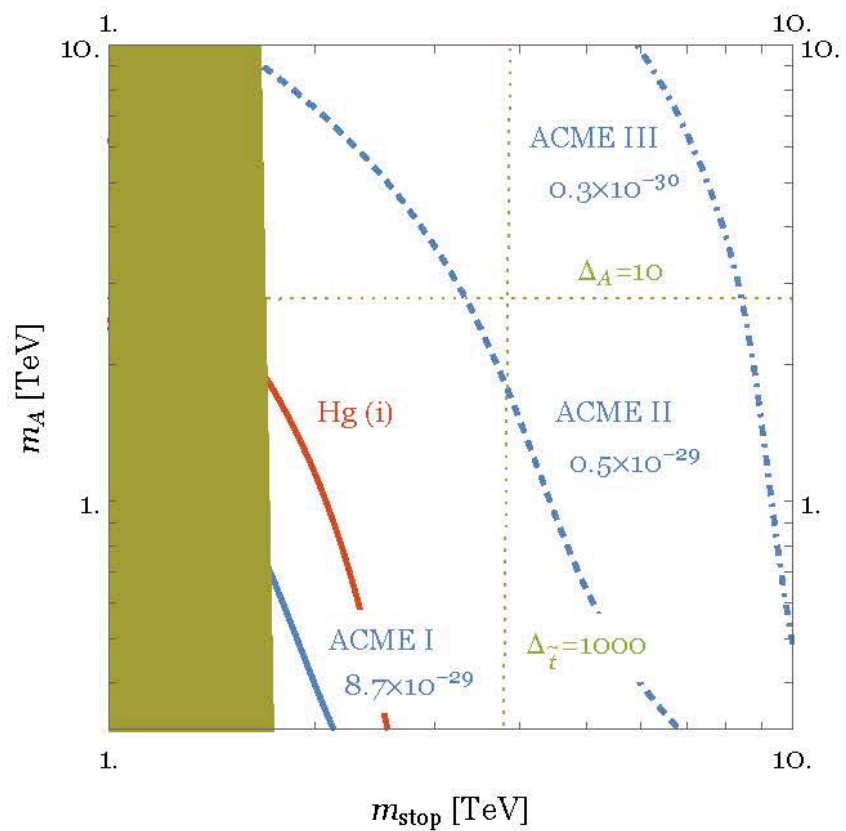
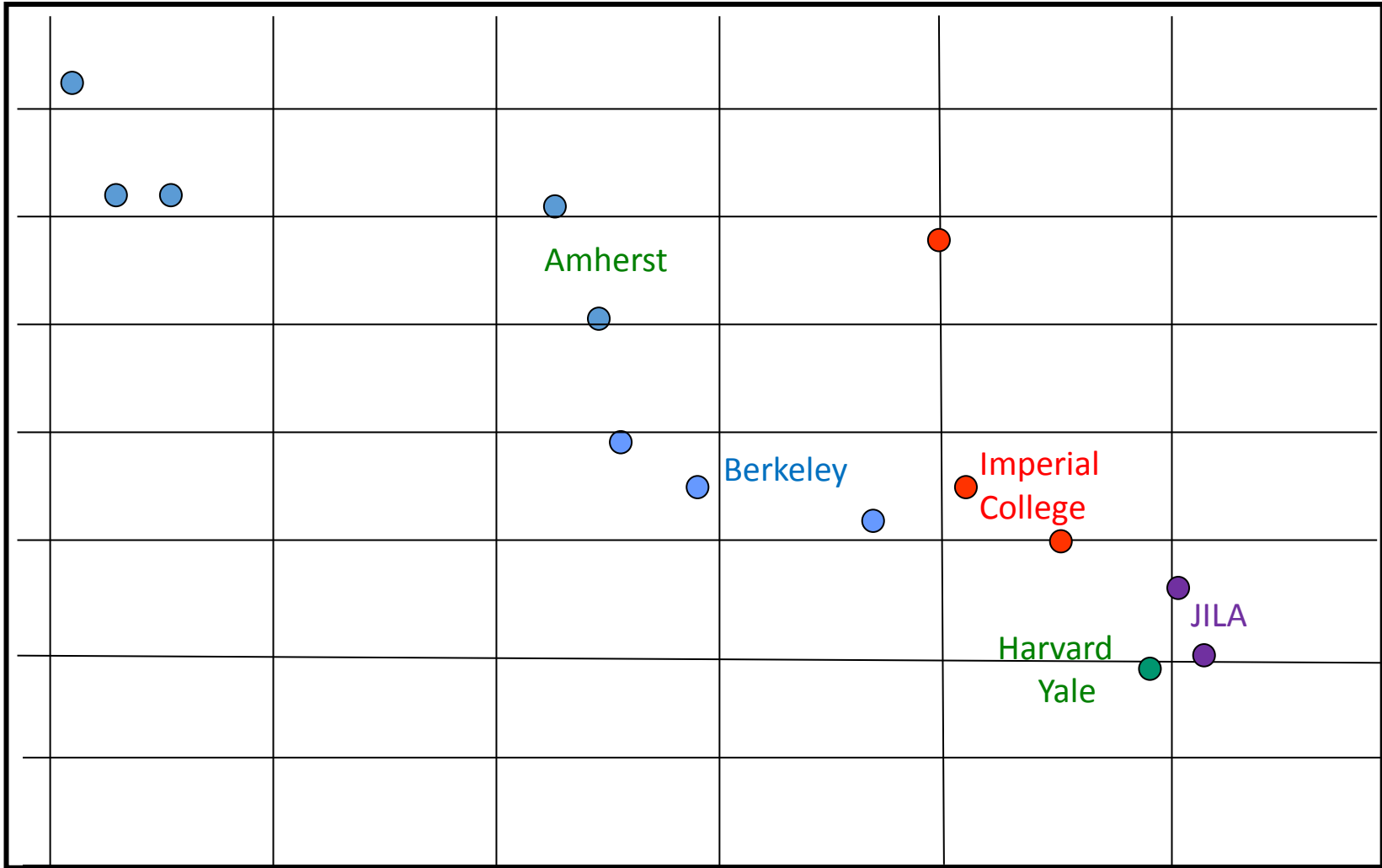


Figure 9: EDM constraints on the stop parameter space in the MSSM, where stop loops with large A -term lift the Higgs mass to 125 GeV. The horizontal axis shows the common stop soft mass $m_{\text{stop}} = \tilde{m}_{Q_3} = \tilde{m}_{u_3}$. At left we fix $m_A = 400$ GeV and vary $\tan\beta$ on the vertical axis; at right we fix $\tan\beta = 10$ and vary m_A on the vertical axis. In the brown/green shaded region, no choice of A_t is sufficient to achieve the correct Higgs mass. In the rest of the parameter space, at each point we choose A_t to achieve $m_h = 125$ GeV. Regions of parameter space to the left of the solid blue contours are excluded by measurements of ThO. Red solid and dashed contours denote the mercury EDM constraints for the cases (i) and (ii) discussed in Appendix C), respectively. The blue dashed and dot-dashed contours (“ACME II” and “ACME III”) are future projections. The dotted green lines display the stop tree-level Higgs fine tuning (26). We have fixed $|\mu| = 350$ GeV in these figures.

Limit on
eEDM
(e-cm)

10^{-24}
 10^{-25}
 10^{-26}
 10^{-27}
 10^{-28}
 10^{-29}
 10^{-30}



1965

1975

1985

1995

2005

2015

2025

Amherst

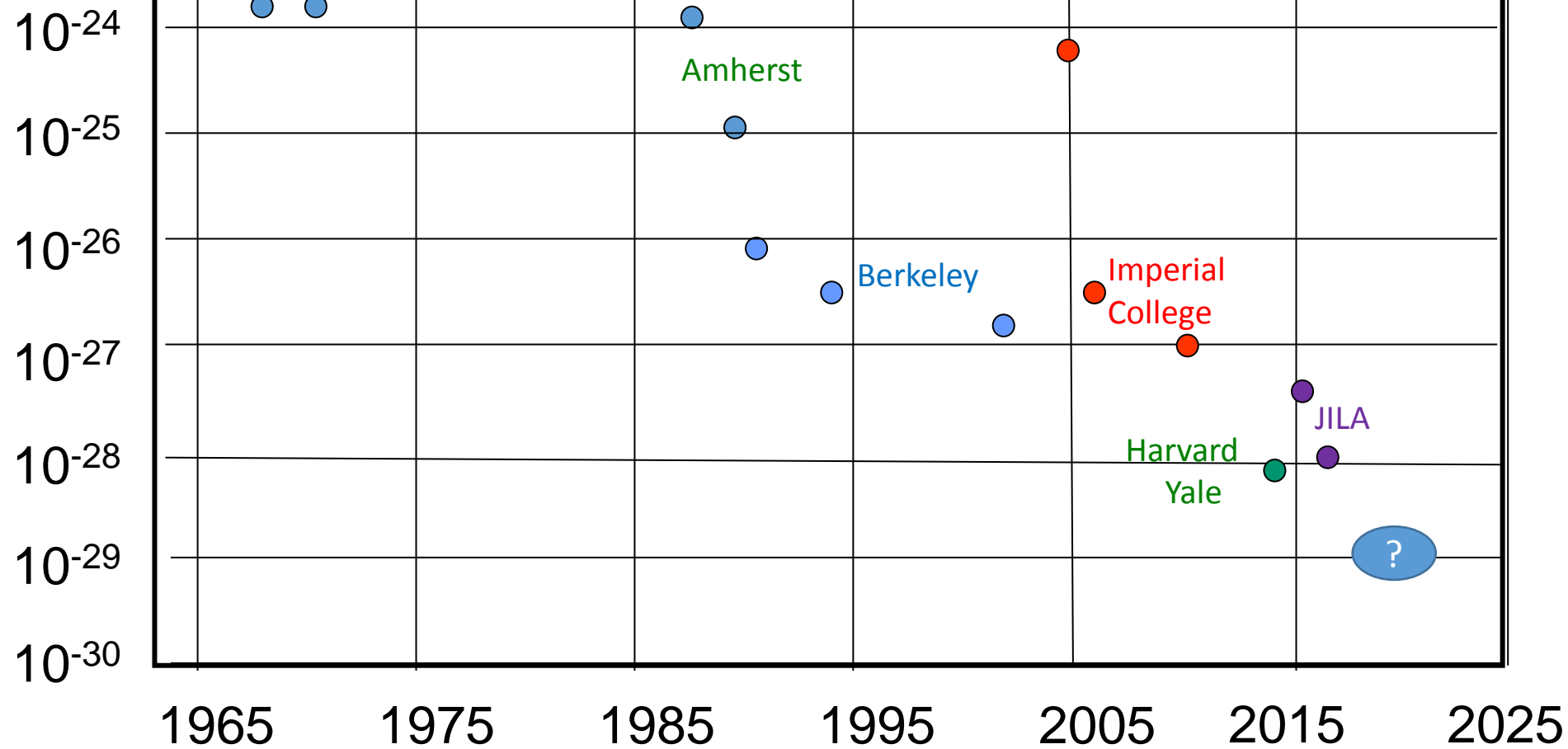
Berkeley

Imperial
College

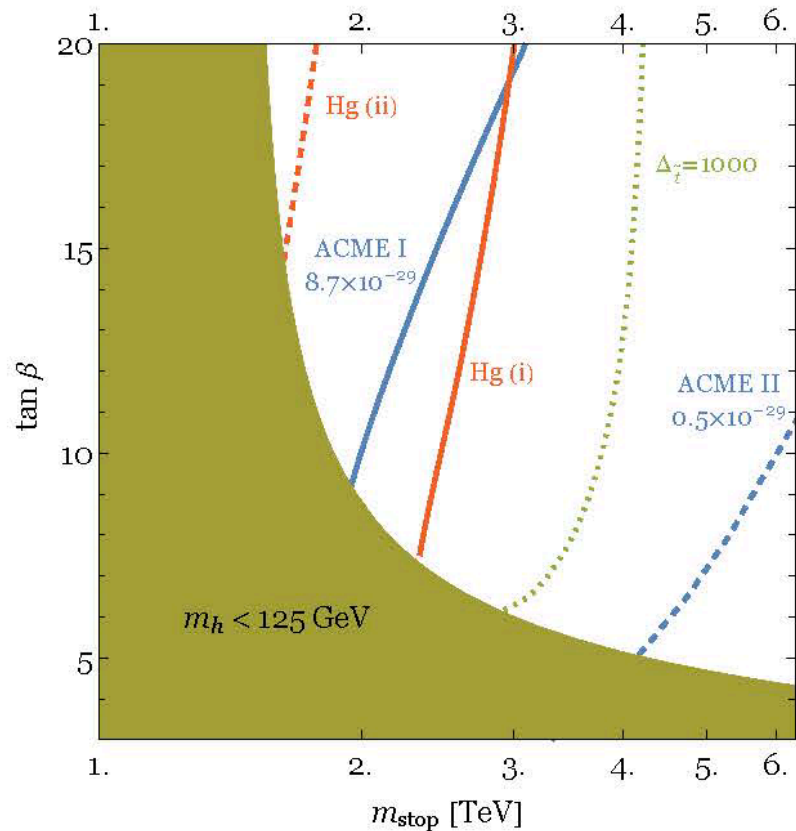
Harvard
Yale

JILA

Limit on
eEDM
(e-cm)



$$d_{\text{ThO,Hg}}/e \text{ [cm]}, \arg(A_t\mu)=\pi/2, m_A=400 \text{ [GeV]}$$



$$d_{\text{ThO,Hg}}/e \text{ [cm]}, \arg(A_t\mu)=\pi/2, \tan\beta = 10$$

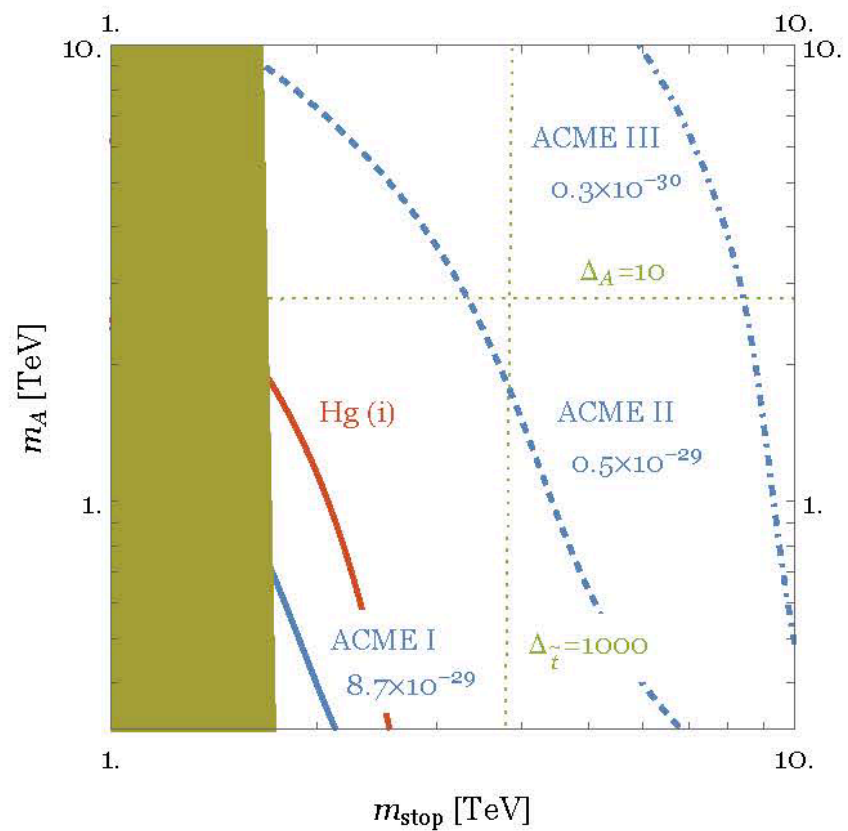


Figure 9: EDM constraints on the stop parameter space in the MSSM, where stop loops with large A -term lift the Higgs mass to 125 GeV. The horizontal axis shows the common stop soft mass $m_{\text{stop}} = \tilde{m}_{Q_3} = \tilde{m}_{u_3}$. At left we fix $m_A = 400$ GeV and vary $\tan\beta$ on the vertical axis; at right we fix $\tan\beta = 10$ and vary m_A on the vertical axis. In the brown/green shaded region, no choice of A_t is sufficient to achieve the correct Higgs mass. In the rest of the parameter space, at each point we choose A_t to achieve $m_h = 125$ GeV. Regions of parameter space to the left of the solid blue contours are excluded by measurements of ThO. Red solid and dashed contours denote the mercury EDM constraints for the cases (i) and (ii) discussed in Appendix C), respectively. The blue dashed and dot-dashed contours (“ACME II” and “ACME III”) are future projections. The dotted green lines display the stop tree-level Higgs fine tuning (26). We have fixed $|\mu| = 350$ GeV in these figures.

Many other “AMO type” particle physics work. **H. Muller**

other electric dipole moment experiments in Harvard/Yale, Penn State

electric dipole moments of mercury atom, or radium atom, of neutron.

anomalies in *magnetic* moments

time-varying physical “constants”.

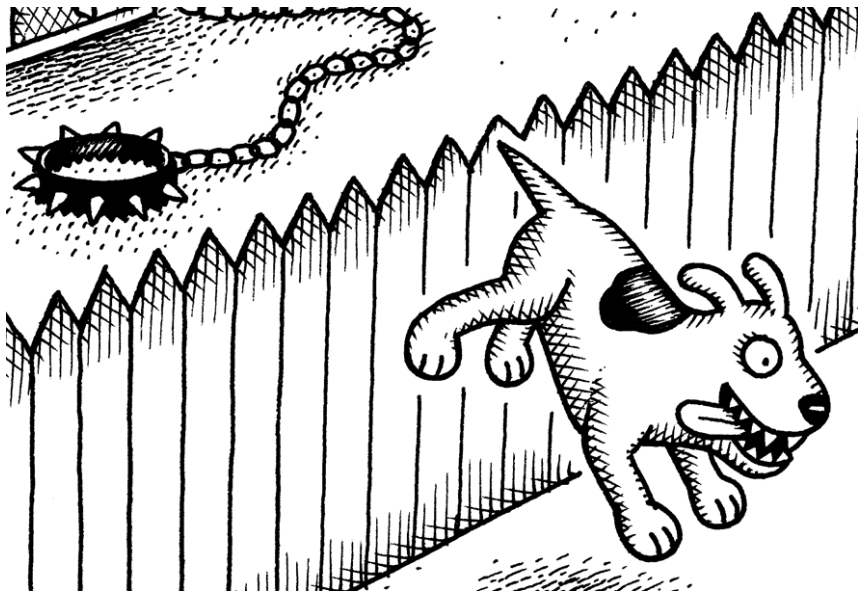
CPT violation? Gravity and antimatter? **F. Sorrentino**

search for axions

tests of gravitational theory. **I Carusotto. G. Rosi. V. Fleurov**

And let us recall the “O” in AMO. Optical detection of gravitational waves @ LIGO, VIRGO, maybe LISA? Optical gyroscopes. **B Patricelli. D. Virgilio.**

...



cold-atom capabilities:

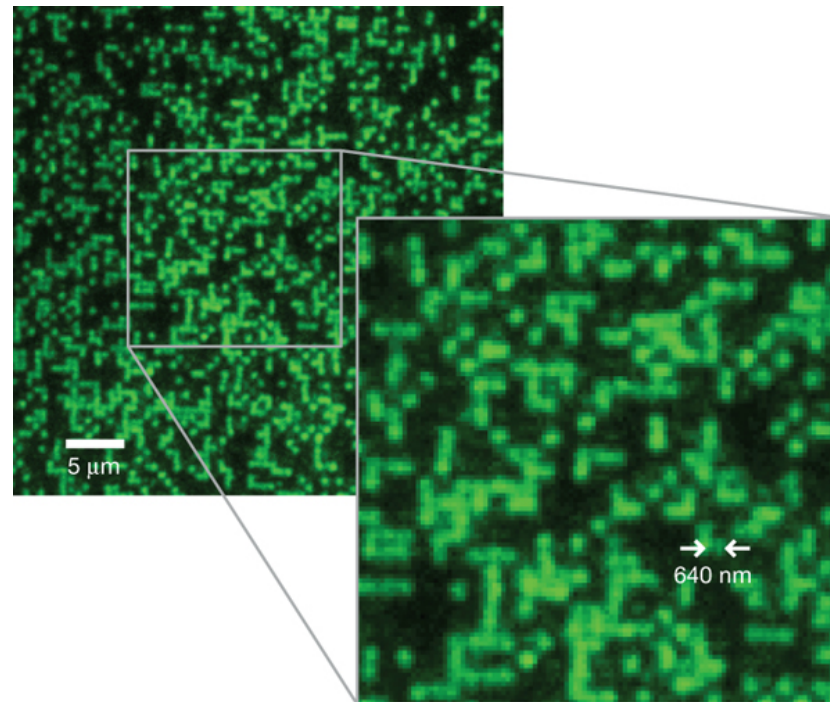
methods: lasers, evap cooling; individual atom detection
confining potentials : magnetic, or optical. harmonic or
square or lattice.

bosons and fermions. tunable interactions. reduced
dimensions.

*many of these things changeable in real time,
dynamically*

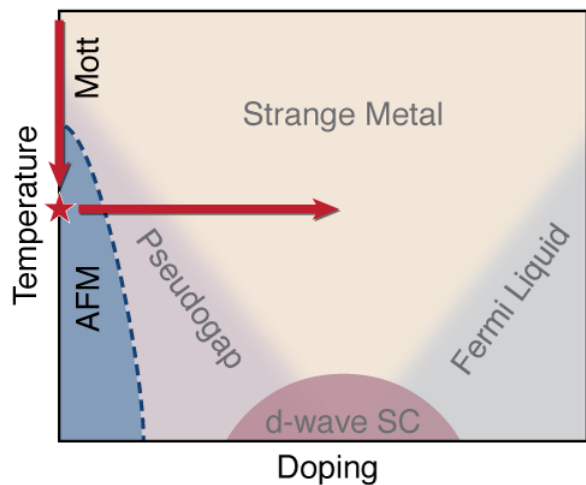
more recently, synthetic magnetic field.

Site-resolved imaging of single atoms on a 640-nm-period optical lattice, loaded with a high density Bose–Einstein condensate.

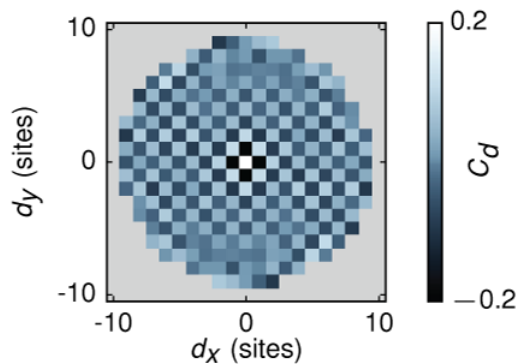


WS Bakr *et al.* *Nature* **462**, 74-77 (2009) doi:10.1038/nature08482

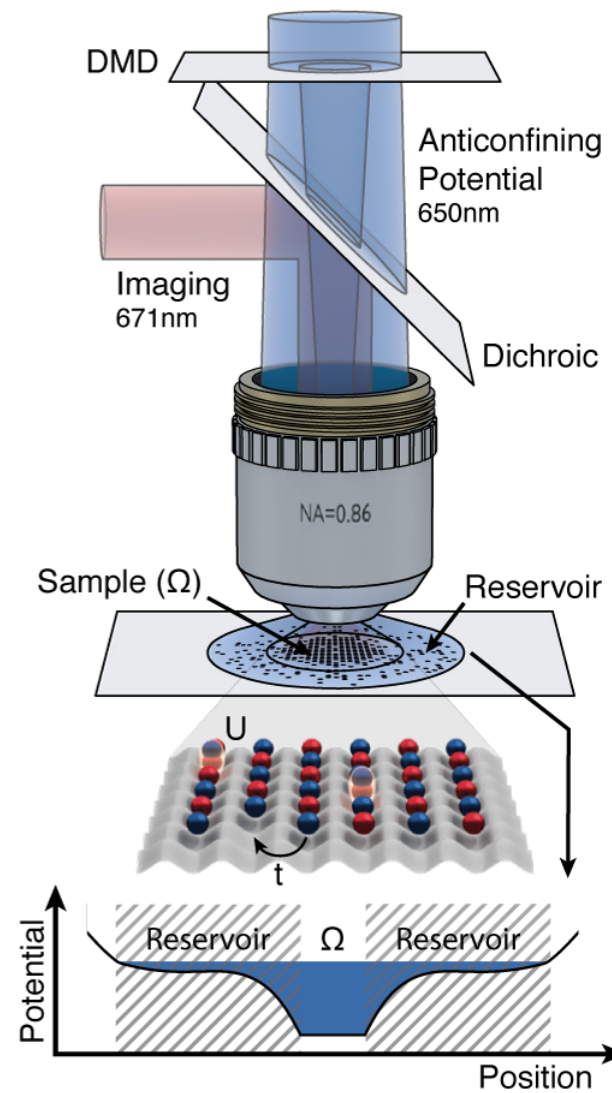
The Hubbard Model



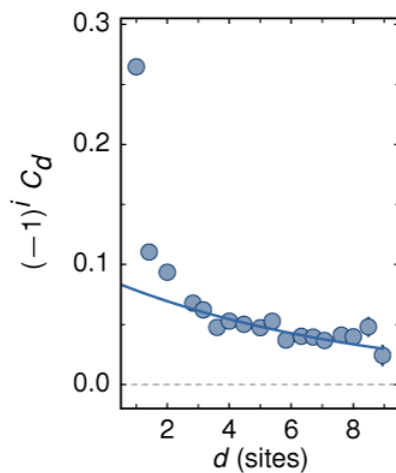
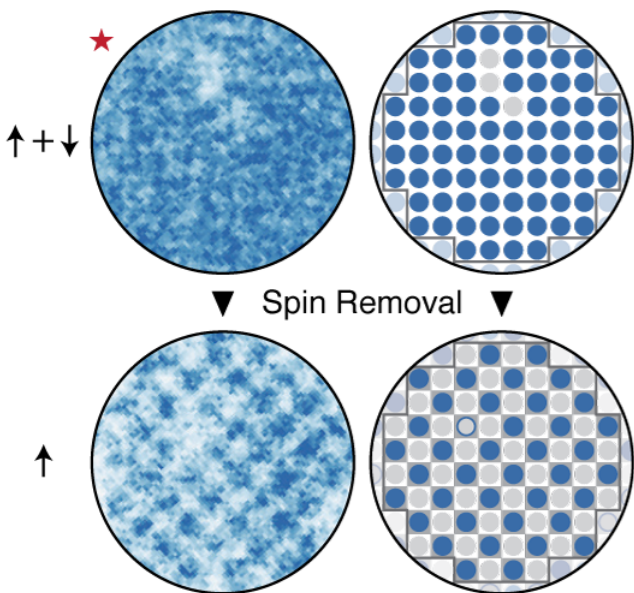
Spin Correlation Function



Entropy Redistribution



Long-range Antiferromagnet



cold-atom capabilities:

methods: lasers, evaporative cooling.; individual atom detection

containers : magnetic, or optical. harmonic or square *or lattice*.

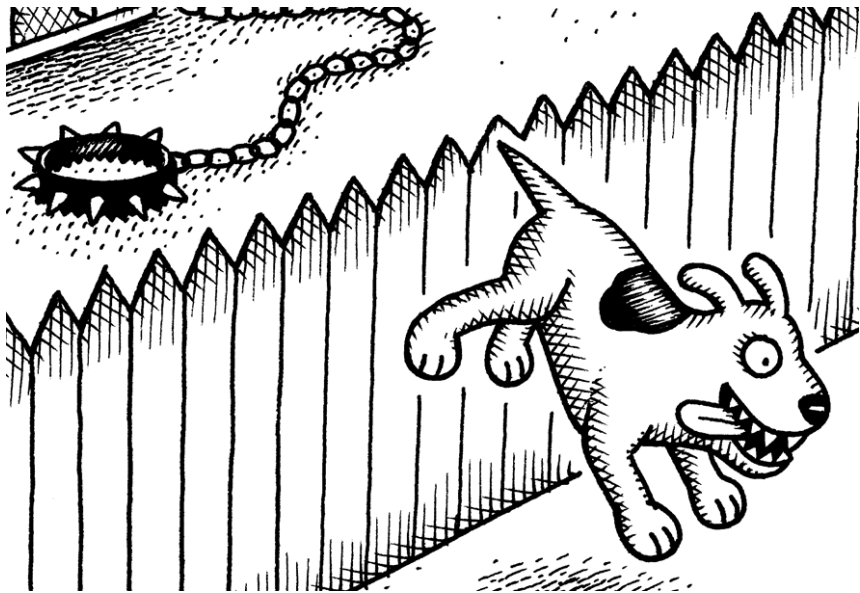
bosons and fermions. tunable interactions. reduced dimensions.

many of these things changeable in real time, dynamically

more recently, synthetic magnetic field.

Few body, many body.

See talks from: M Koehl. G. Rossini. V. Fleurov



liquids:

Where most
chemistry
happens.

Where all life
happens.



liquids:

Where most
chemistry
happens.

Where all life
happens.

macroscopic
definition:
has a free
surface.



liquids:

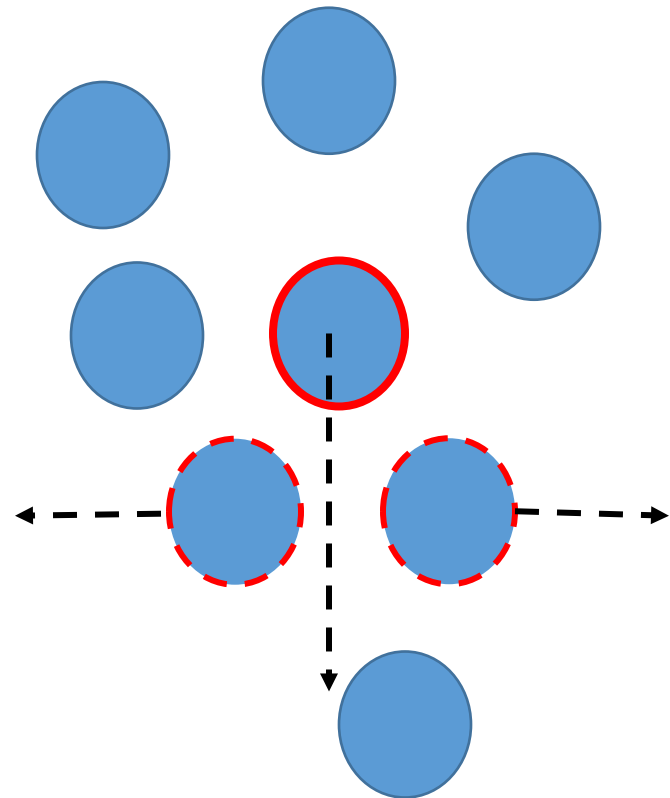
Where most
chemistry
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has a free
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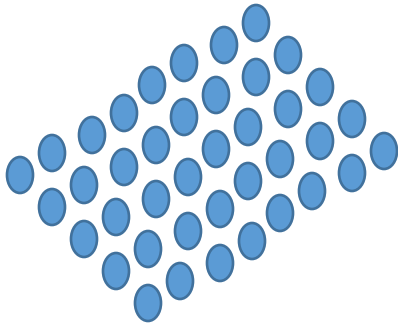


microscopically,
a liquid is
where
one atom can
move only if
some other
atoms get out of
the way.



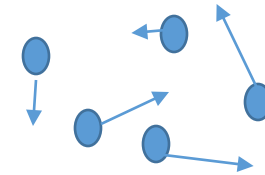
As physicists, we often like to approach a problem perturbatively. We start with...

A solid:
lattice at
 $T=0$.



perturbatively
add carriers,
phonons, etc

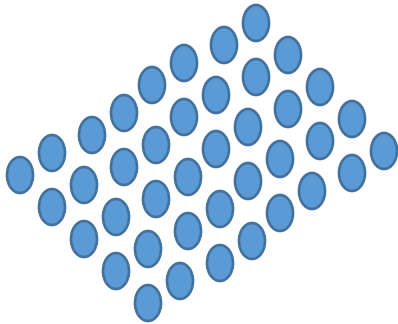
A gas:
noninteracting
molecules with good \mathbf{k} .



perturbatively add
rare \mathbf{k} -changing
collisions, mean field,
etc

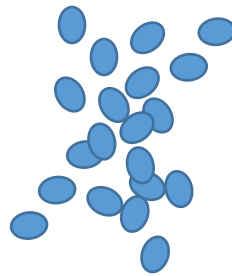
As physicists, we often like to approach a problem perturbatively. We start with...

A solid:
lattice at
 $T=0$.

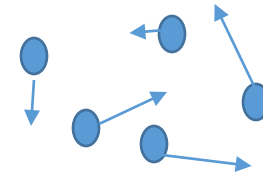


perturbatively
add carriers,
phonons, etc

A liquid:
??? Not $T=0$. No
single-atom \mathbf{k}

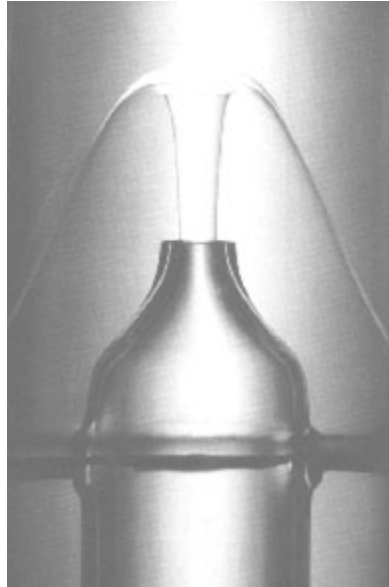


A gas:
noninteracting
molecules with good \mathbf{k} .



perturbatively add
rare \mathbf{k} -changing
collisions, mean field,
etc

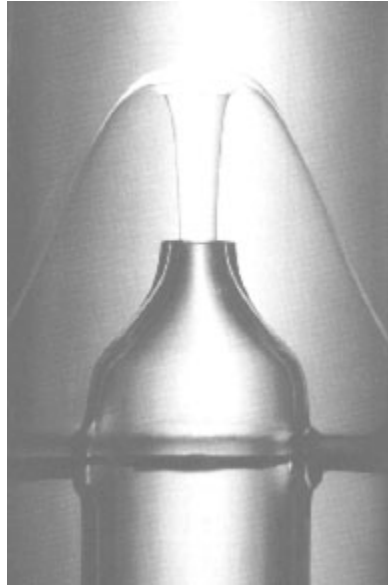
Except, superfluid liquid! Liquid at $T=0$!
Macroscopic physics very exotic.



Except, superfluid liquid! Liquid at $T=0$!

Macroscopic physics very exotic.

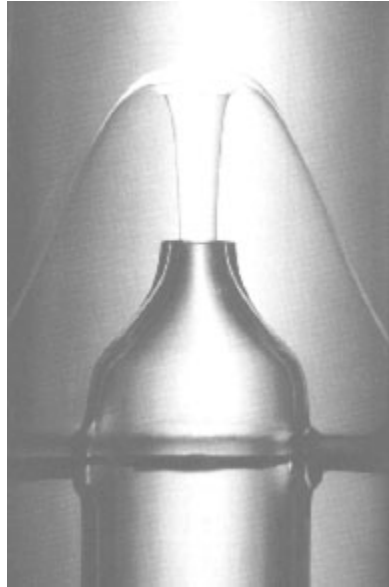
Microscopically it's actually easier to understand than a regular liquid.



Except, superfluid liquid! Liquid at $T=0$!

Macroscopic physics very exotic.

Microscopically it's actually easier to understand than a regular liquid.



But comes only as liquid helium.

Microscopic probes are difficult.

Microscopic probes in Liquid Helium are tricky.

LHe $T = 0.35$ K

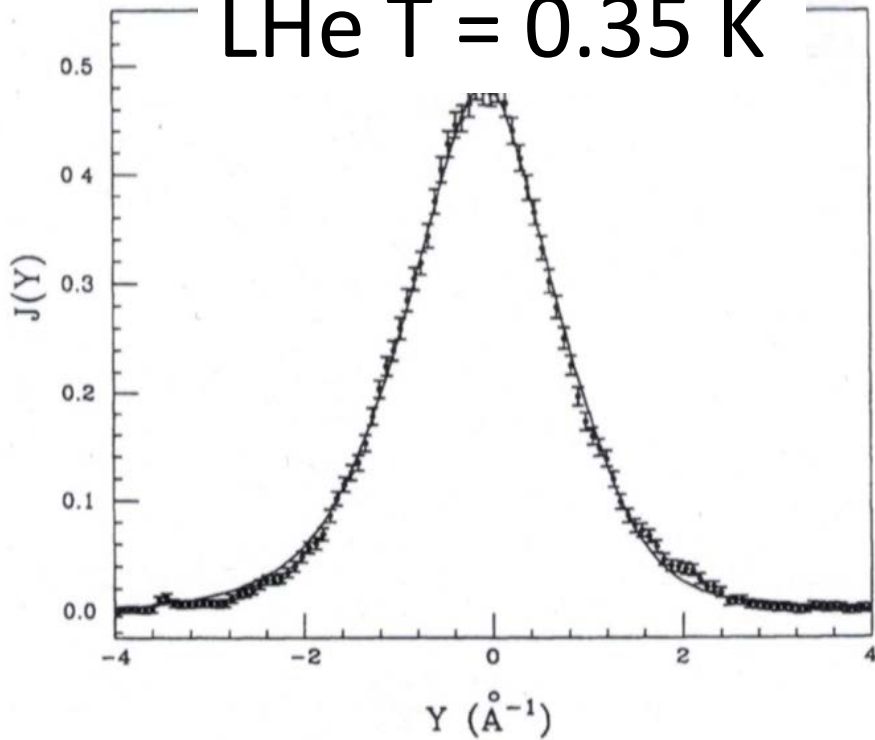


FIG. 8. Observed scattering at $T=0.35$ K. The dashed line is the GFMC prediction of Whitlock and Panoff broadened by the instrumental resolution function and by Silver's FSE broadening function.

LHe $T = 3.5$ K

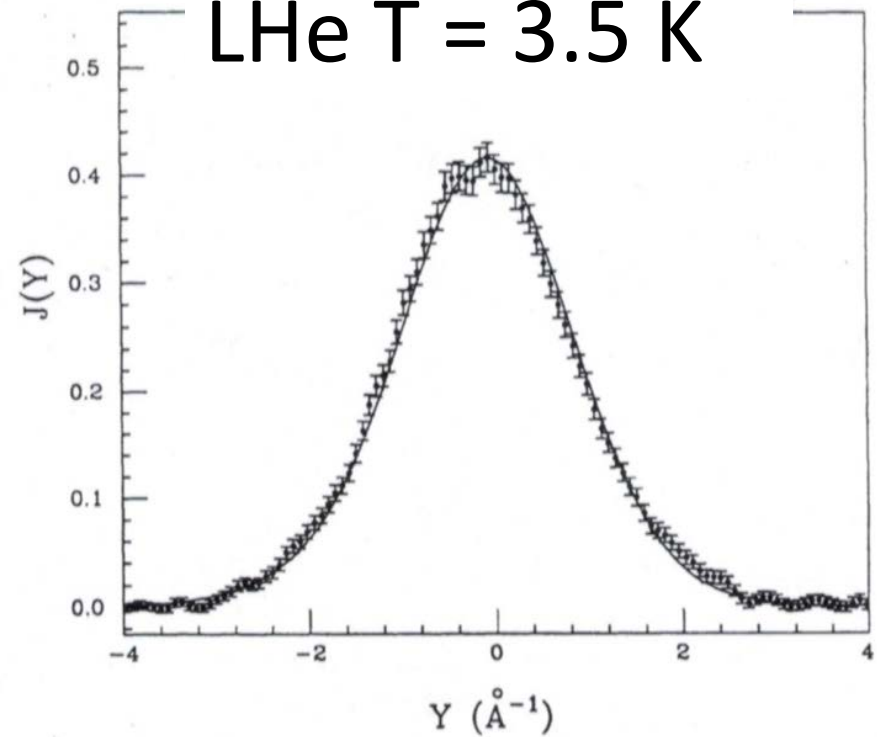


FIG. 7. Observed scattering at $T=3.5$ K. The dashed line is the PIMC prediction of Ceperley and Pollock broadened by the instrumental resolution function and by Silver's FSE broadening function.

Data from Sokol et al 1990.

Idea:

Ultracold atomic gases!

Many wonderful probes.

Tuneable interactions,

Tuneable dimensionality,

Lots more experimental tools than LHelium

Q: Can we make them liquid like, and in that way learn something about liquids?

A: Maybe.

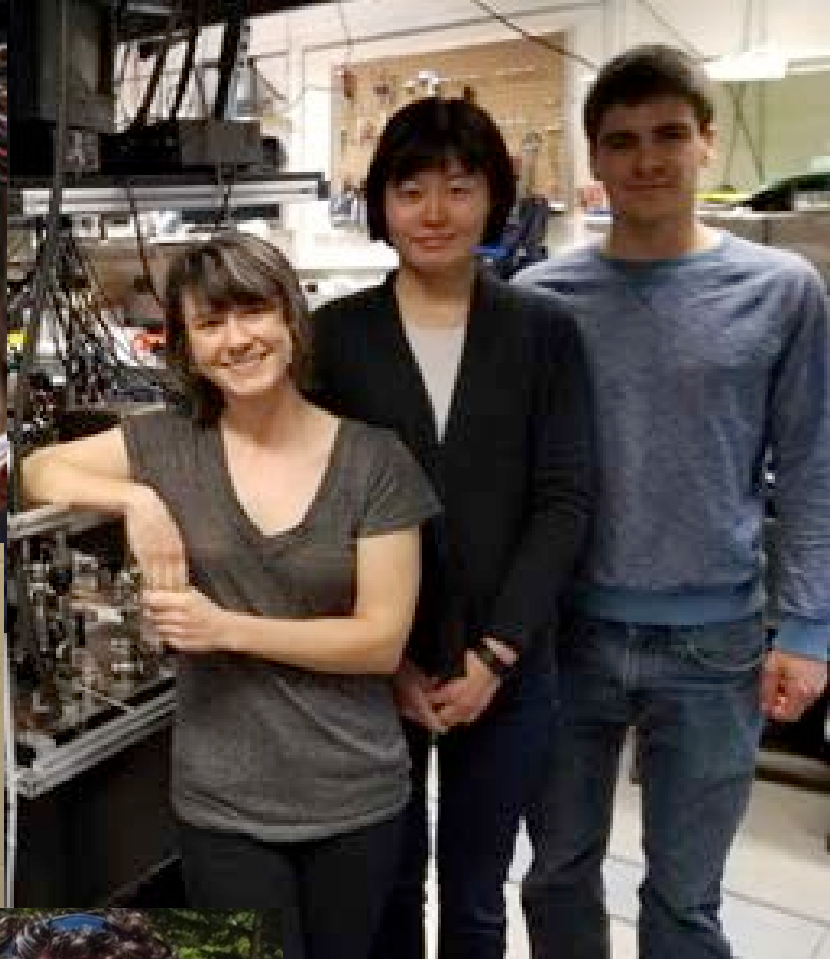
Work in progress!

Strongly interacting
degenerate bose
fluid:
A few-body approach
to a many-body
challenge

Cathy Klauss, Xin Xie
Carlos Lopez-Abadia

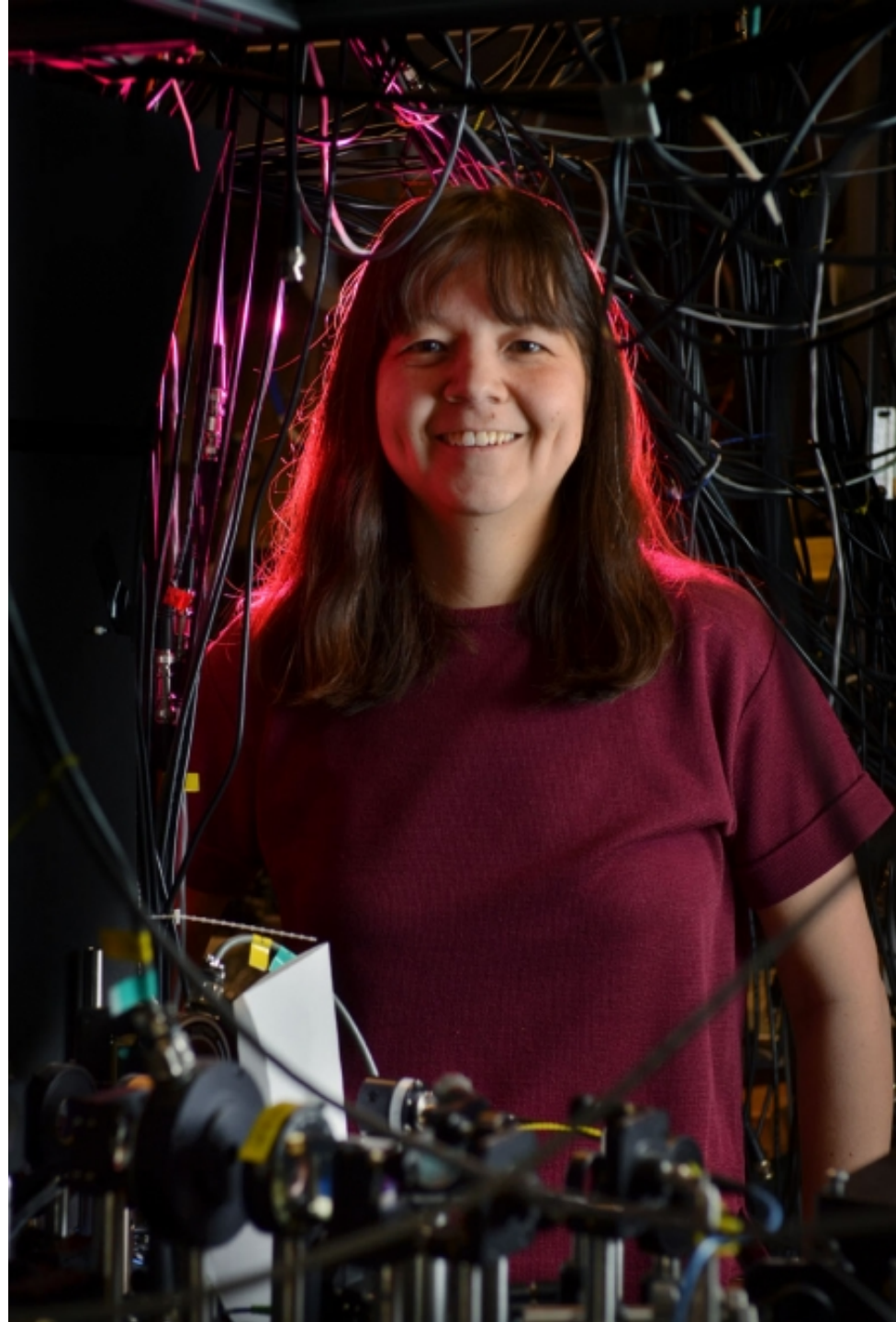
Debbie Jin, E.C
Jose d'Incao
visiting scientist:
Zoran Hadzibabic

former group members
Rob Wild, Phil Makotyn



Thanks: NSF, NIST, NASA

This portion of the
talk is dedicated to the
memory of
Deborah S. Jin
(1968-2016).



Idea:

Ultracold atomic gases!

Many wonderful probes.

Tuneable interactions,

Tuneable dimensionality,

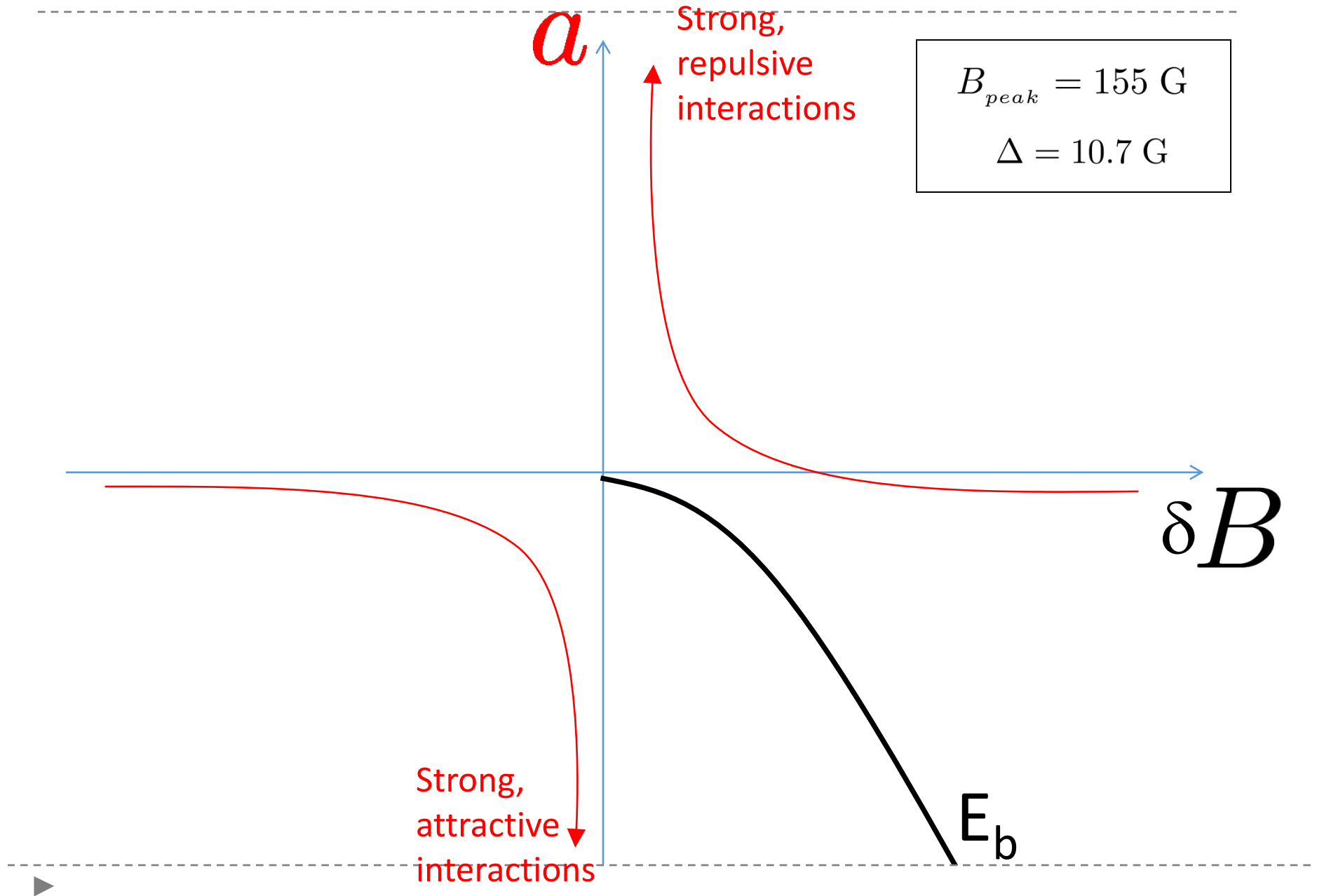
Lots more experimental tools than LHelium

Q: **Can we make them liquid like**, and in that way learn something about liquids?

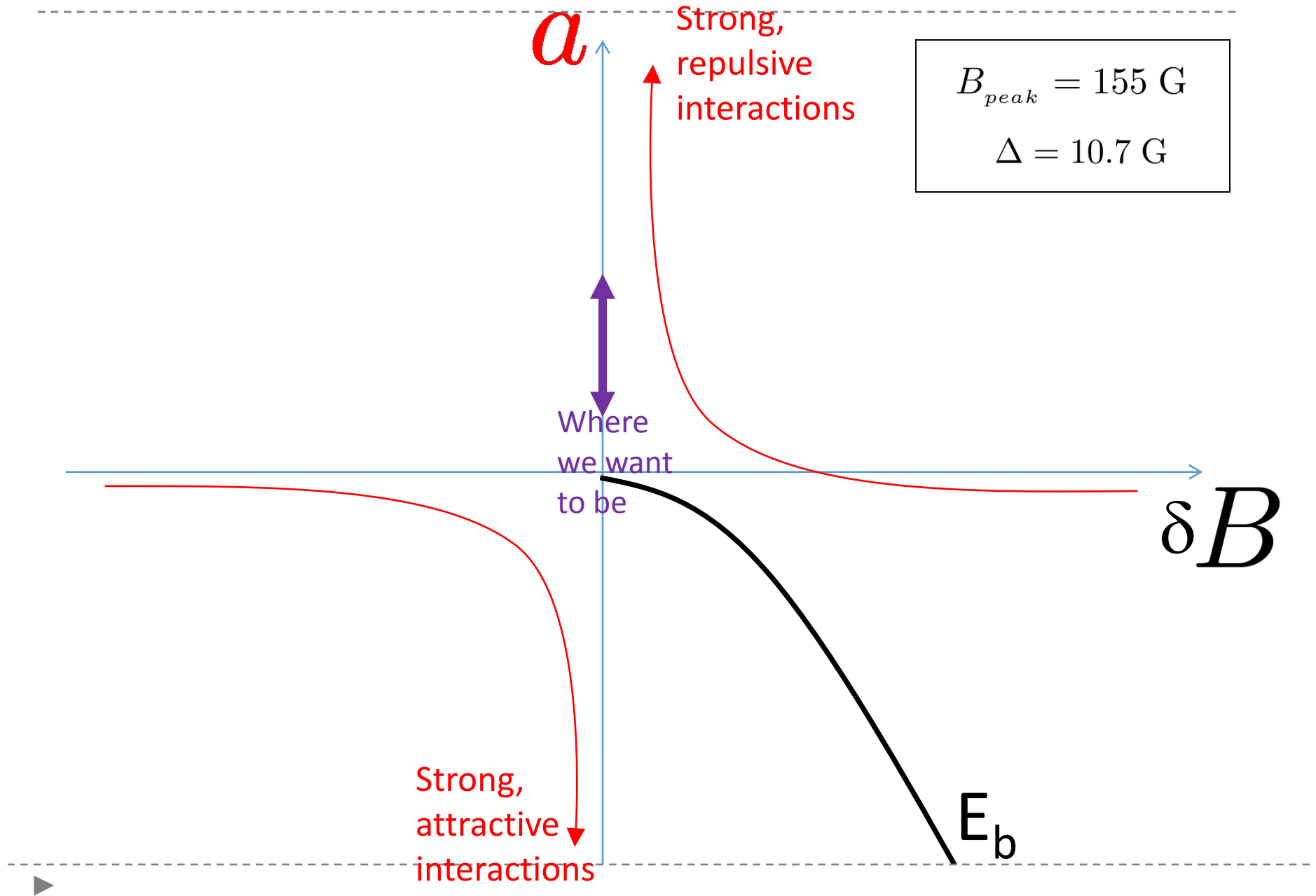
A: Maybe.

Work in progress!

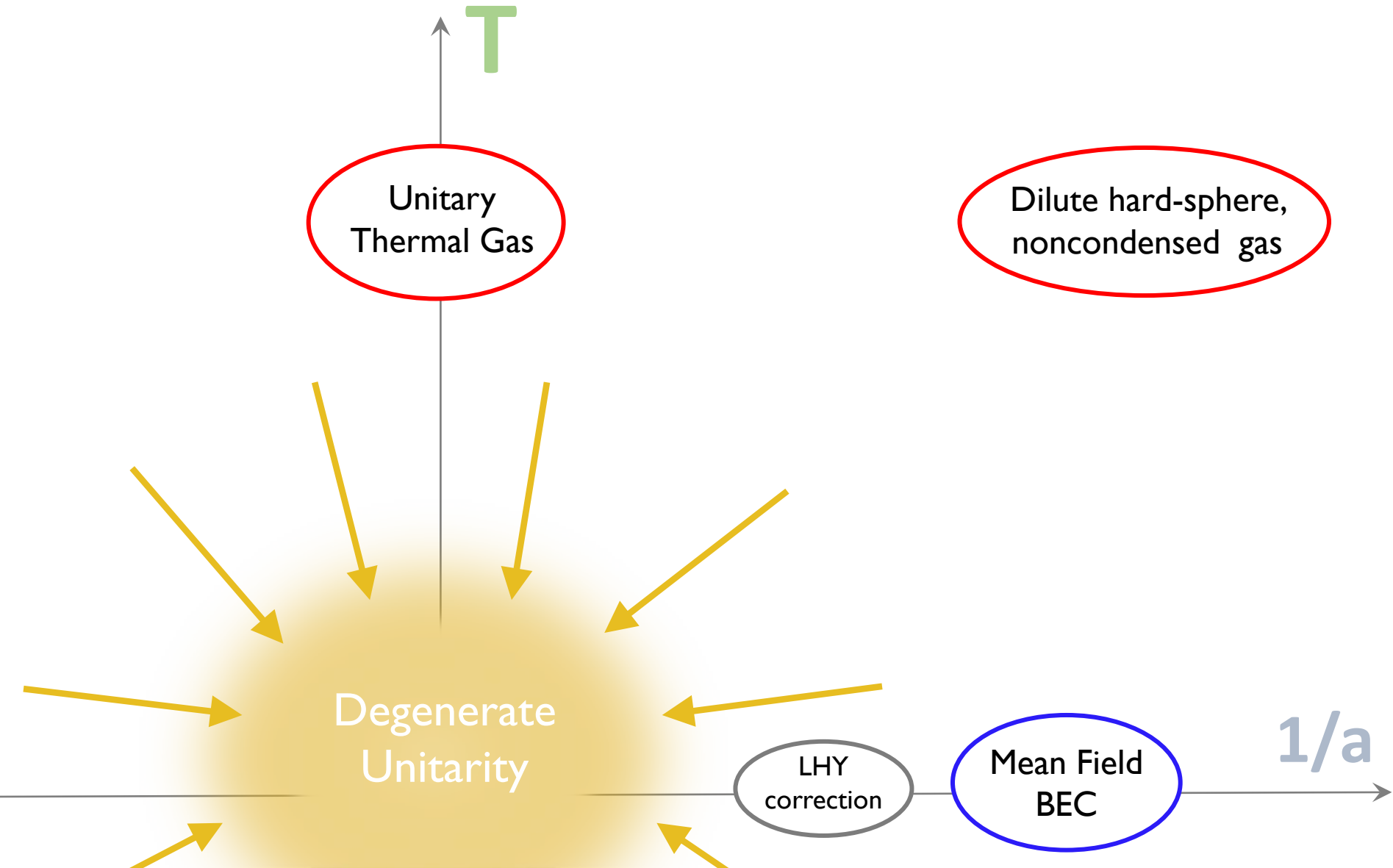
Feshbach Resonance



Feshbach Resonance

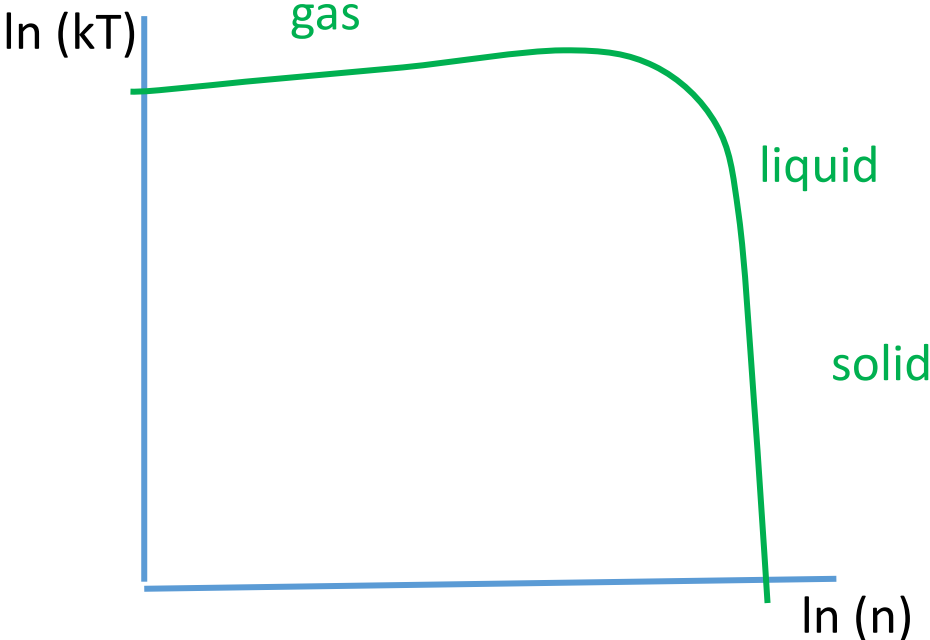


Cold Bosons: The Regimes



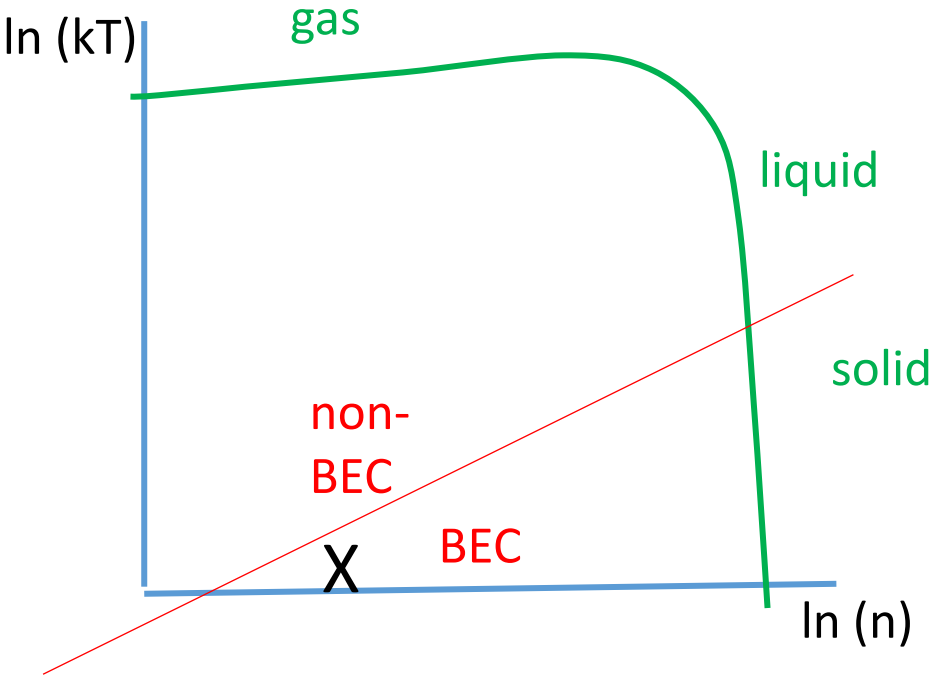
The 3-body recombination catastrophe.

Generic phase diagram



The 3-body recombination catastrophe.

Generic phase diagram



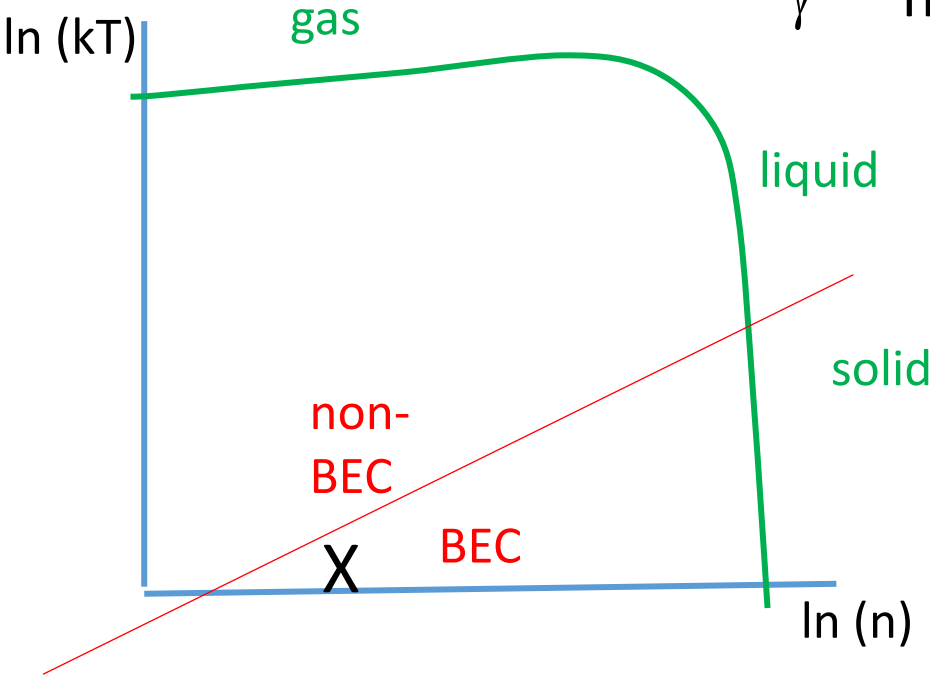
The 3-body recombination catastrophe.

Thermalization – two-body collisions

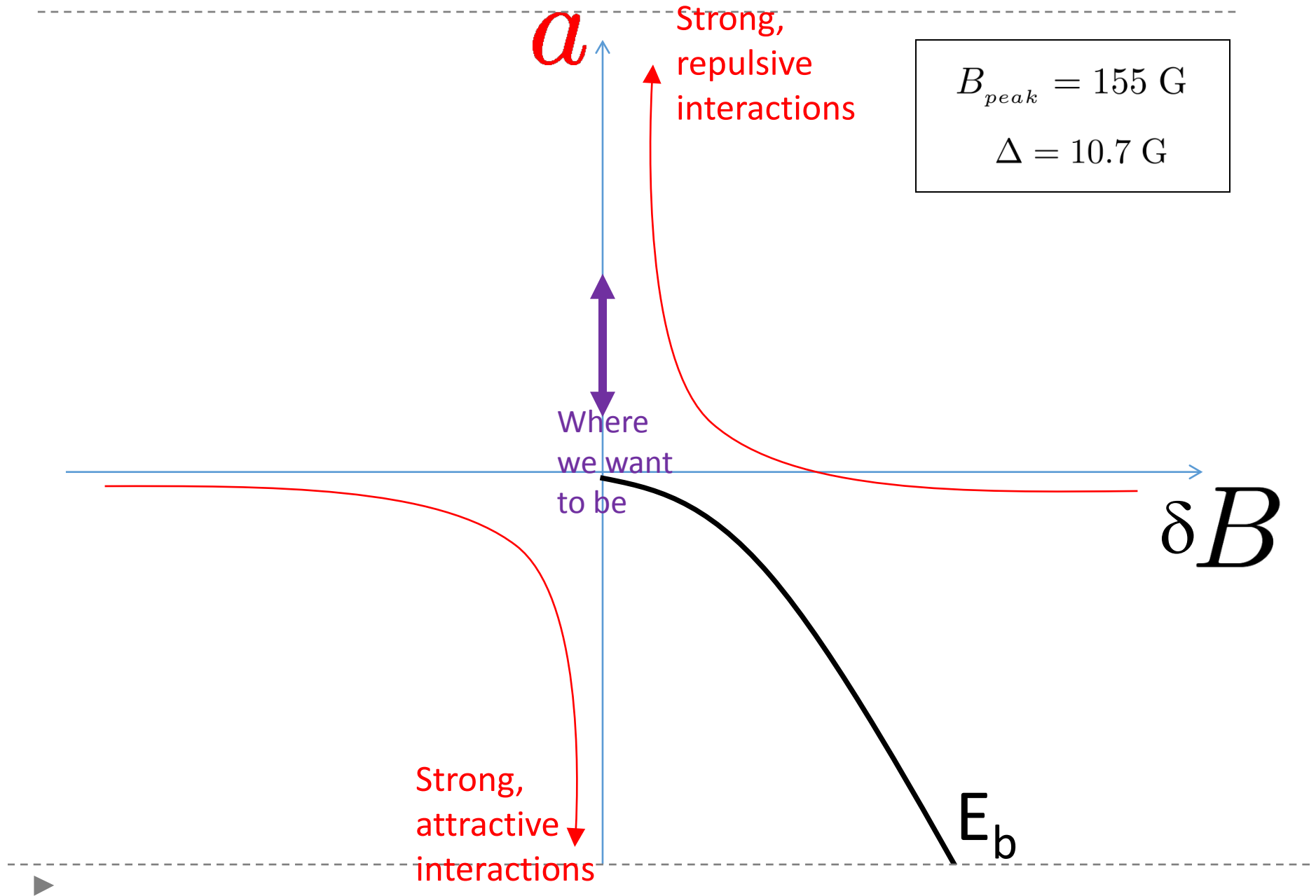
$$\gamma \sim n a^2$$

Decay – 3-body recombination

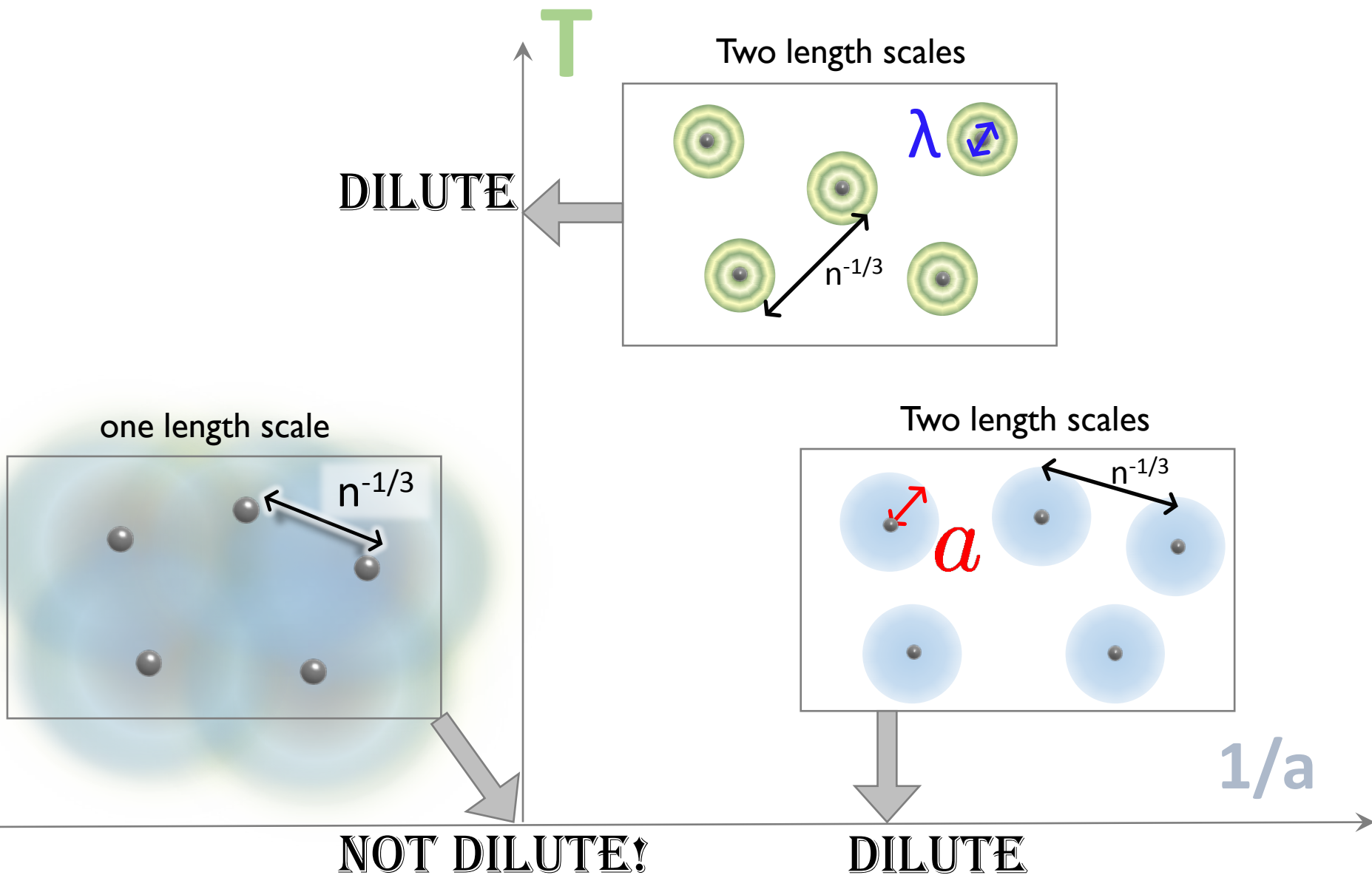
$$\gamma \sim n^2 a^4$$



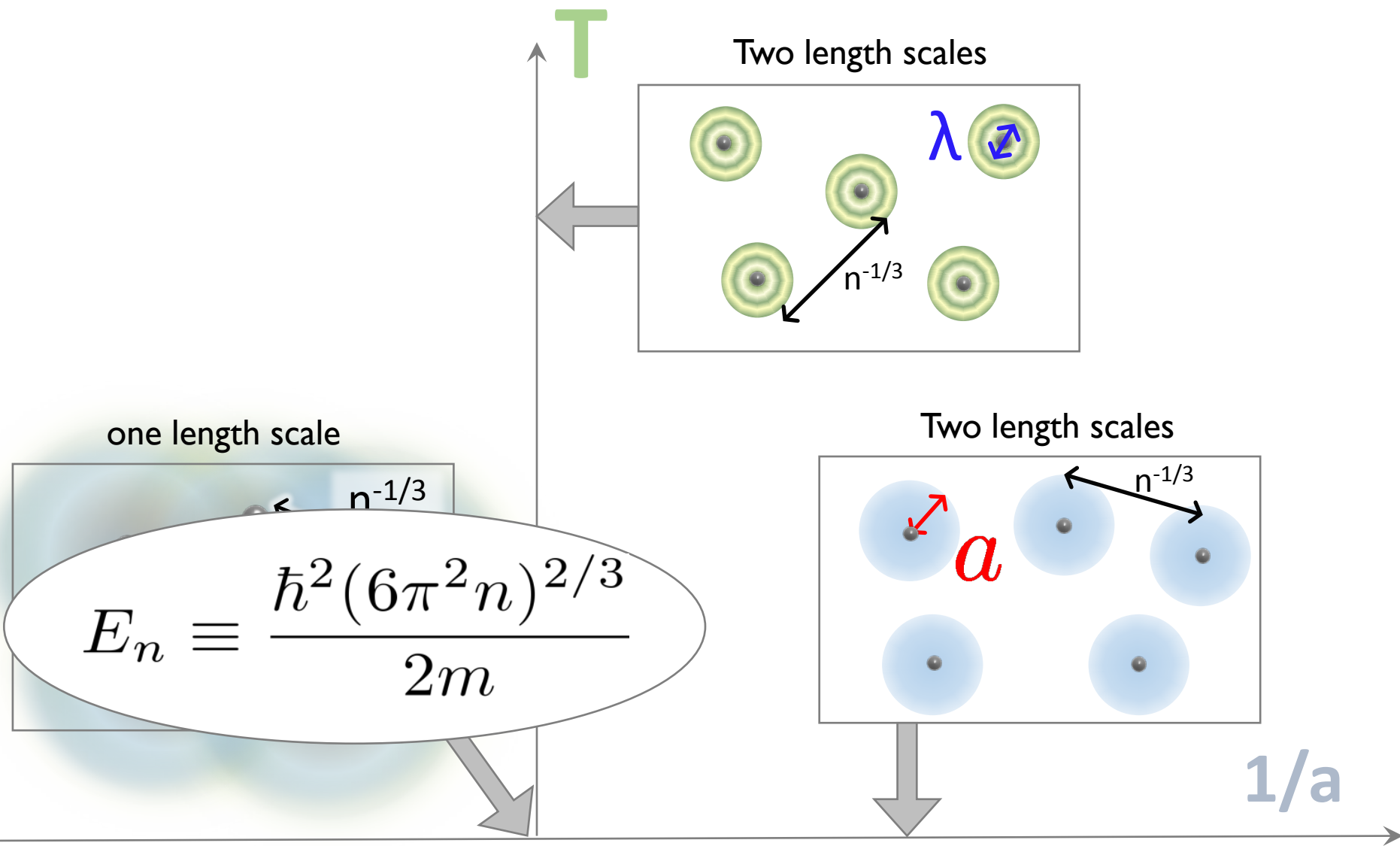
Feshbach Resonance



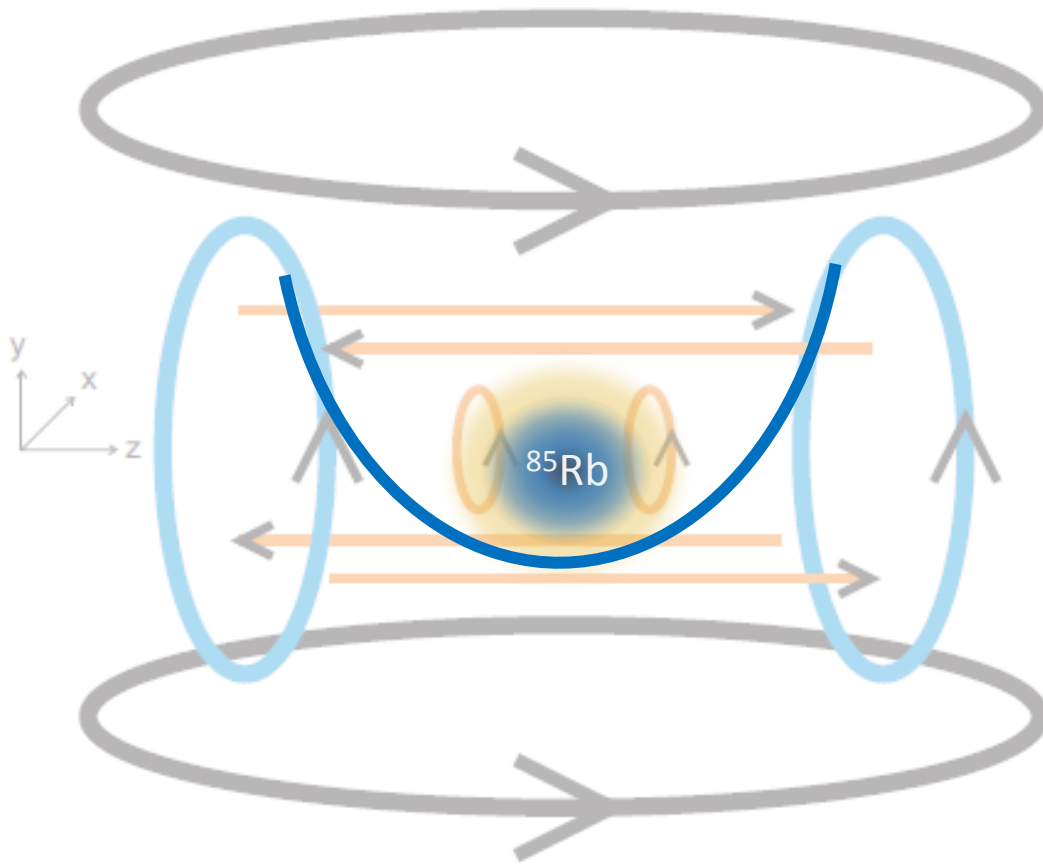
Degenerate Unitarity



Degenerate Unitarity



Basic spherical BEC setup



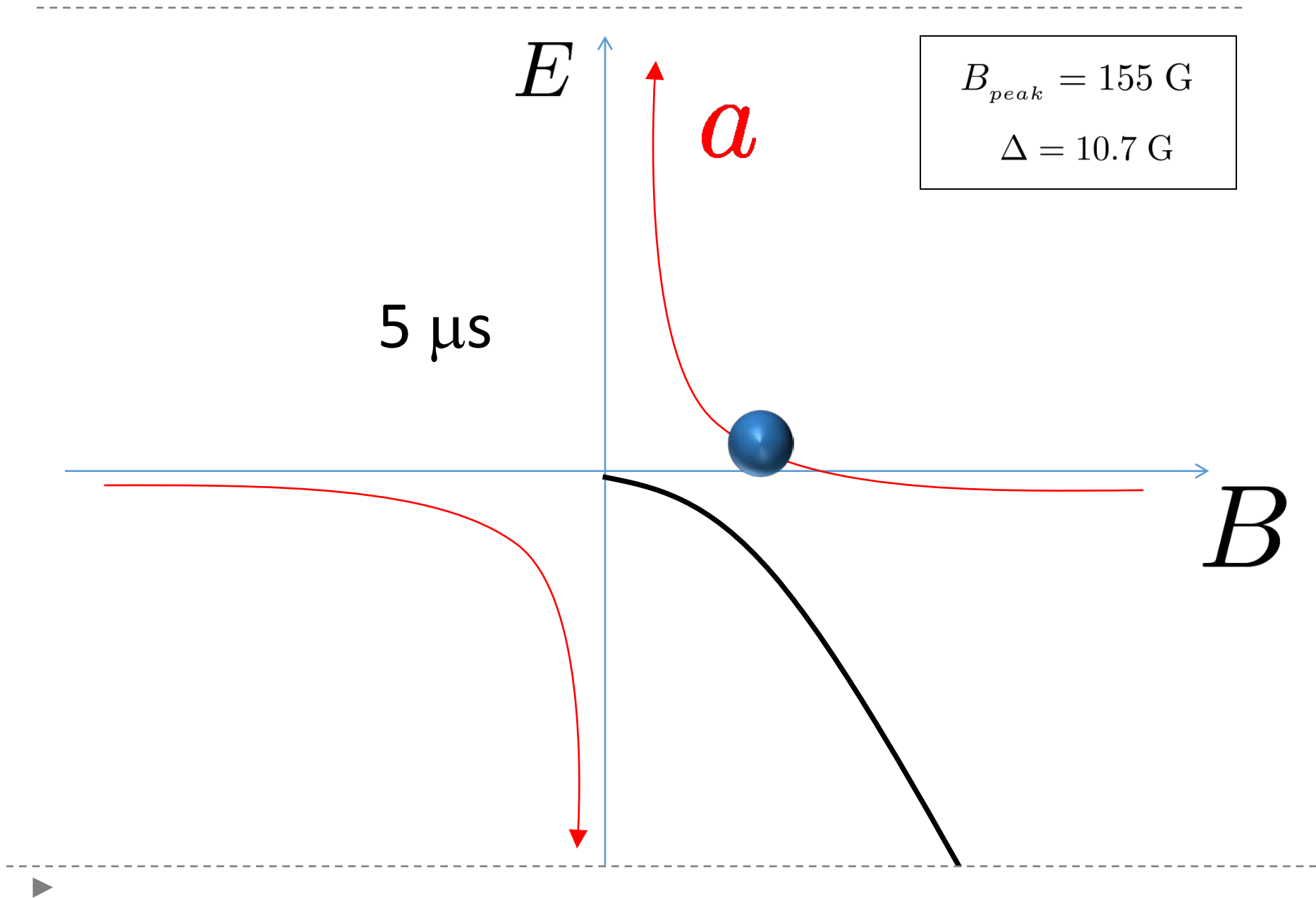
$$N_{BEC} = 6 \cdot 10^4$$

$$\omega = 2\pi 10 \text{ Hz}$$

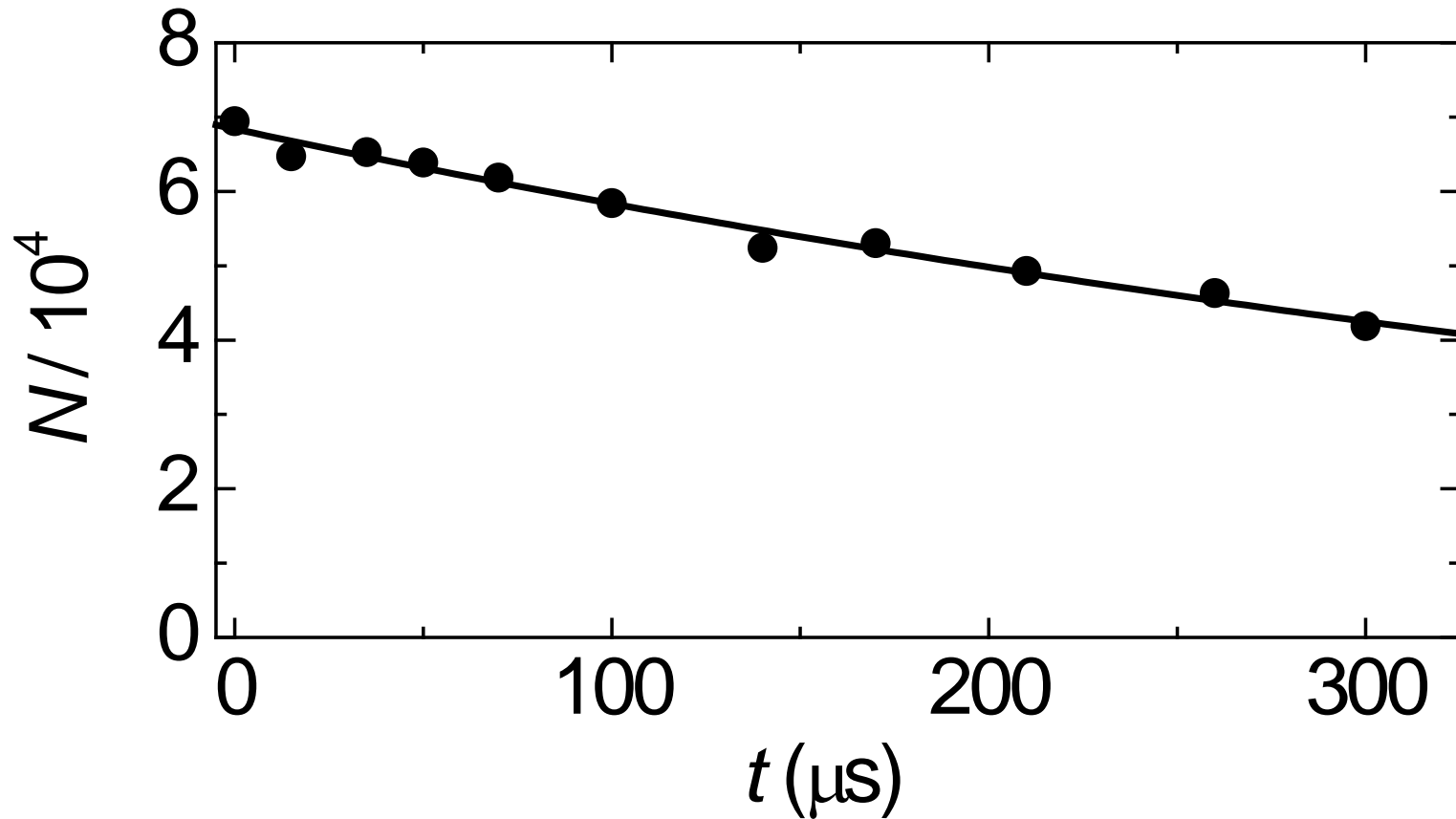
$$\langle n \rangle = 5 \cdot 10^{12} \text{ cm}^{-3}$$

$$T < 10 \text{ nK}$$

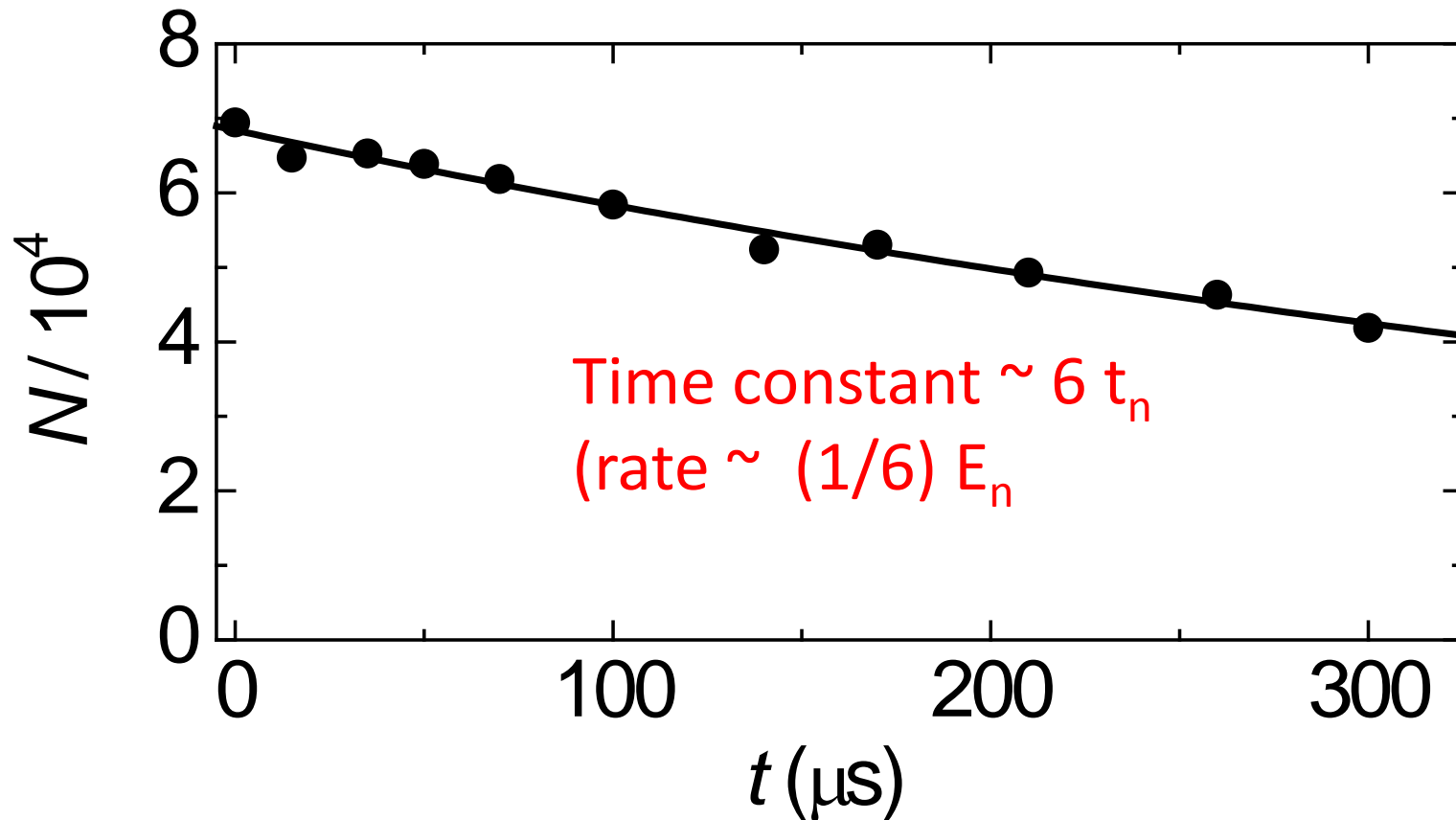
Feshbach Resonance



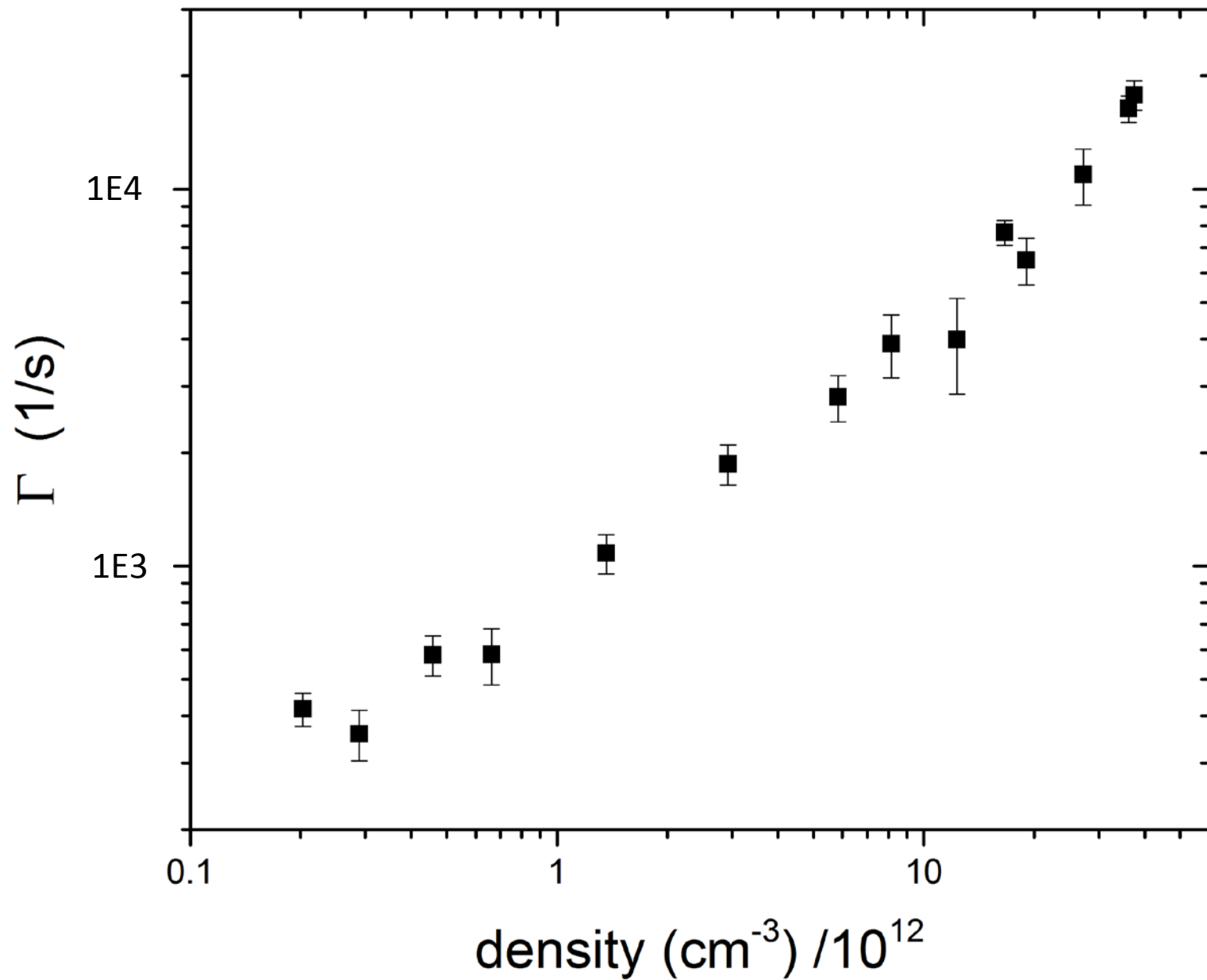
Total atoms remaining v. time at unitarity



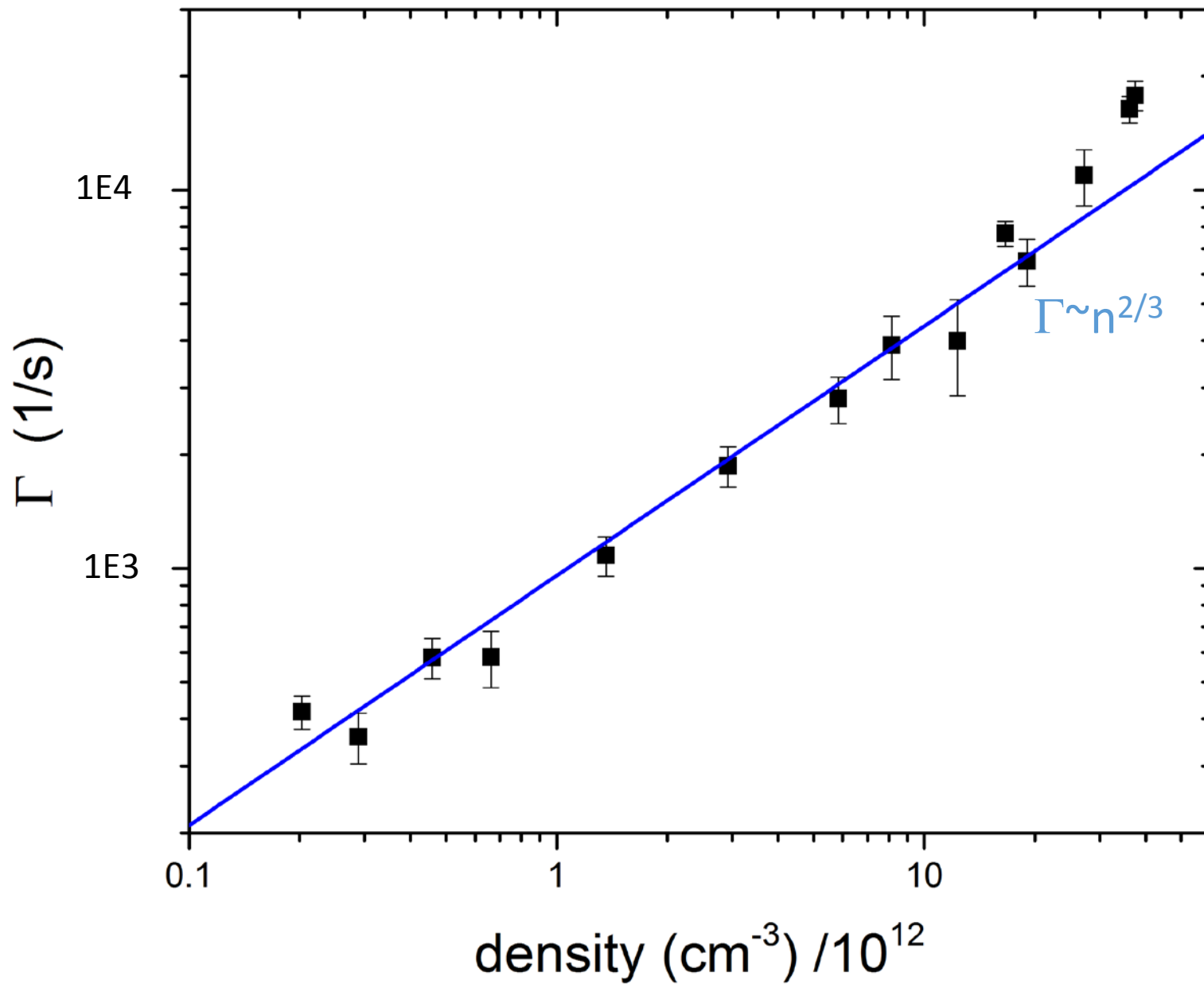
Total atoms remaining v. time at unitarity



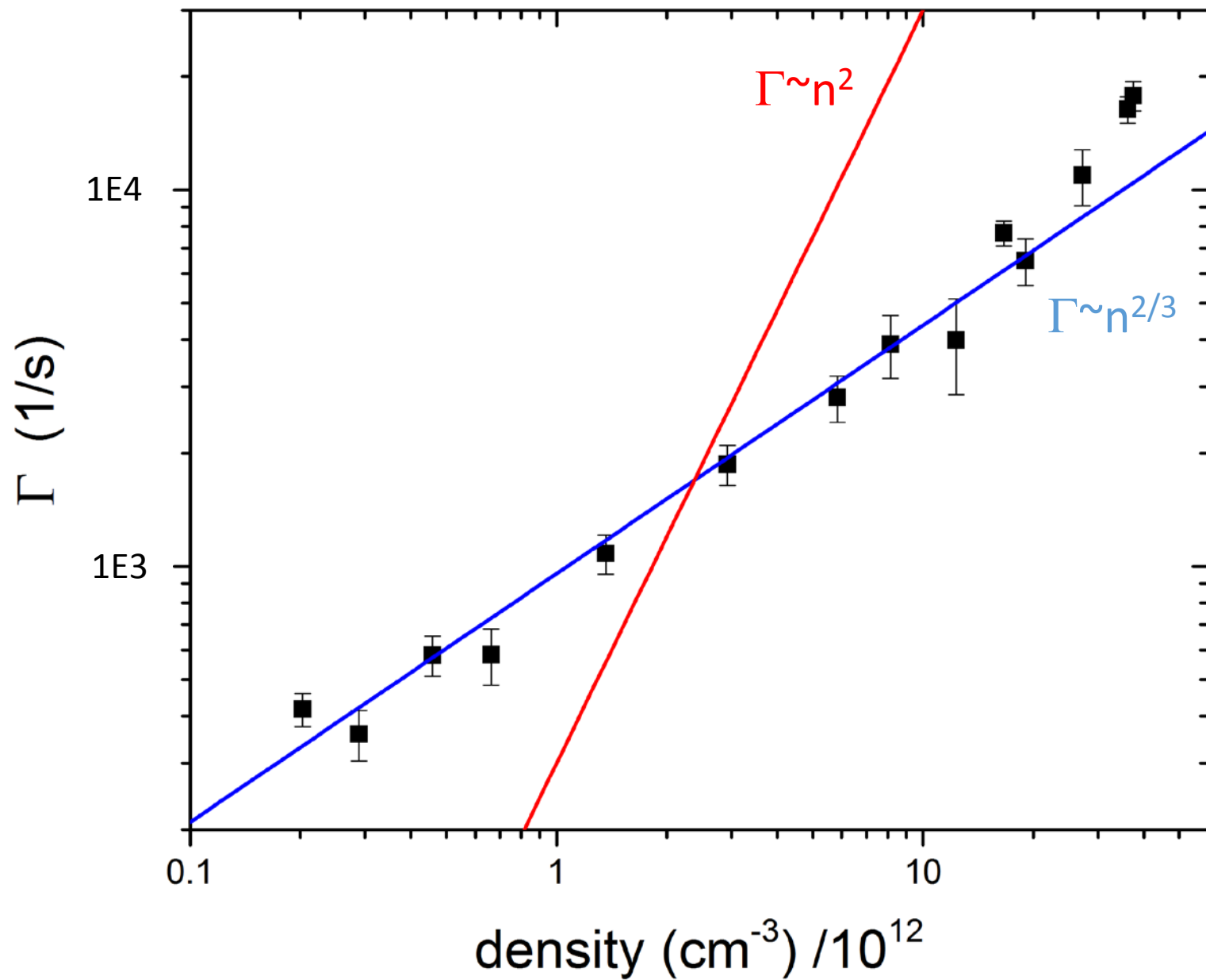
Loss Rate at Unitarity



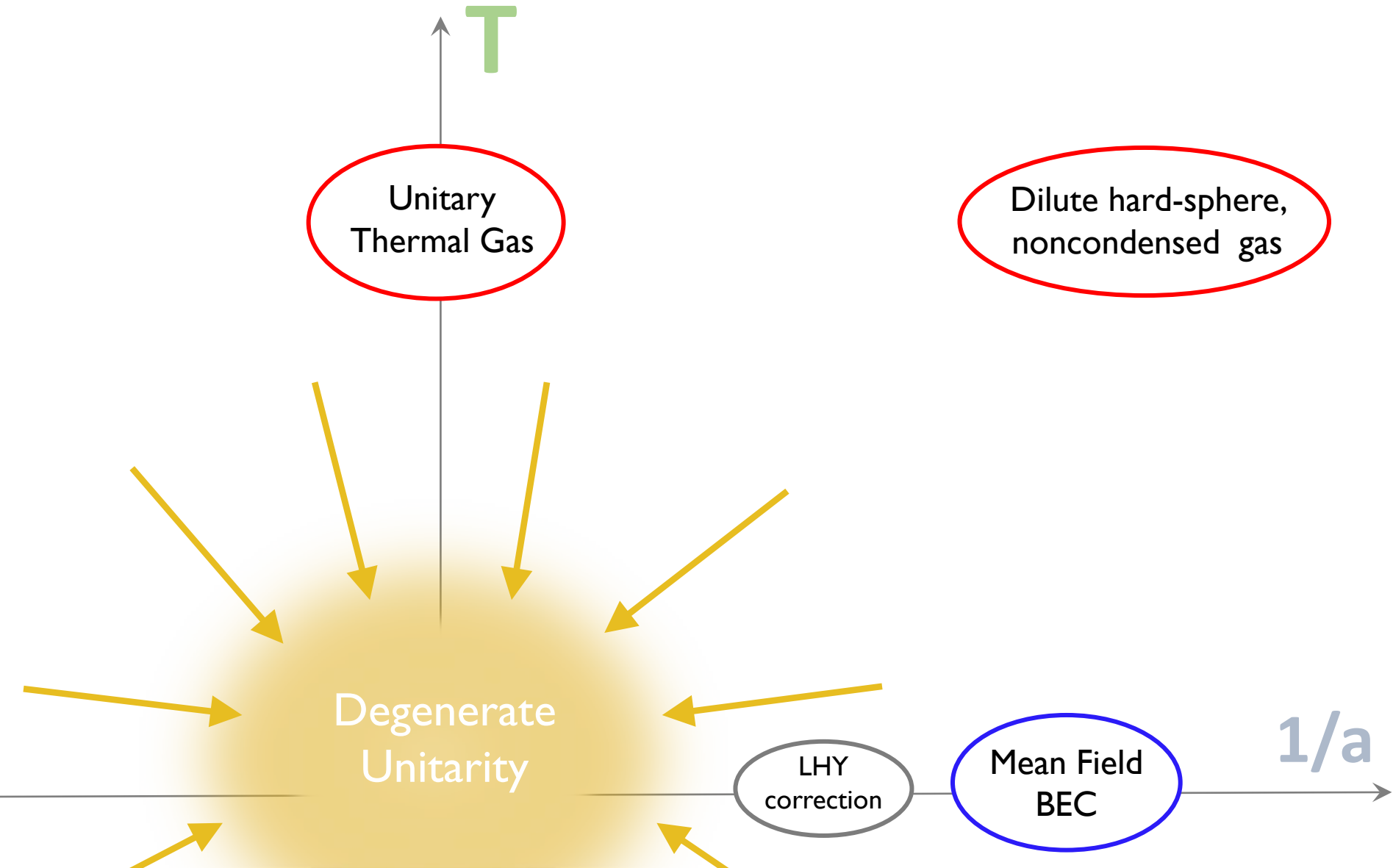
Loss Rate at Unitarity



Loss Rate at Unitarity



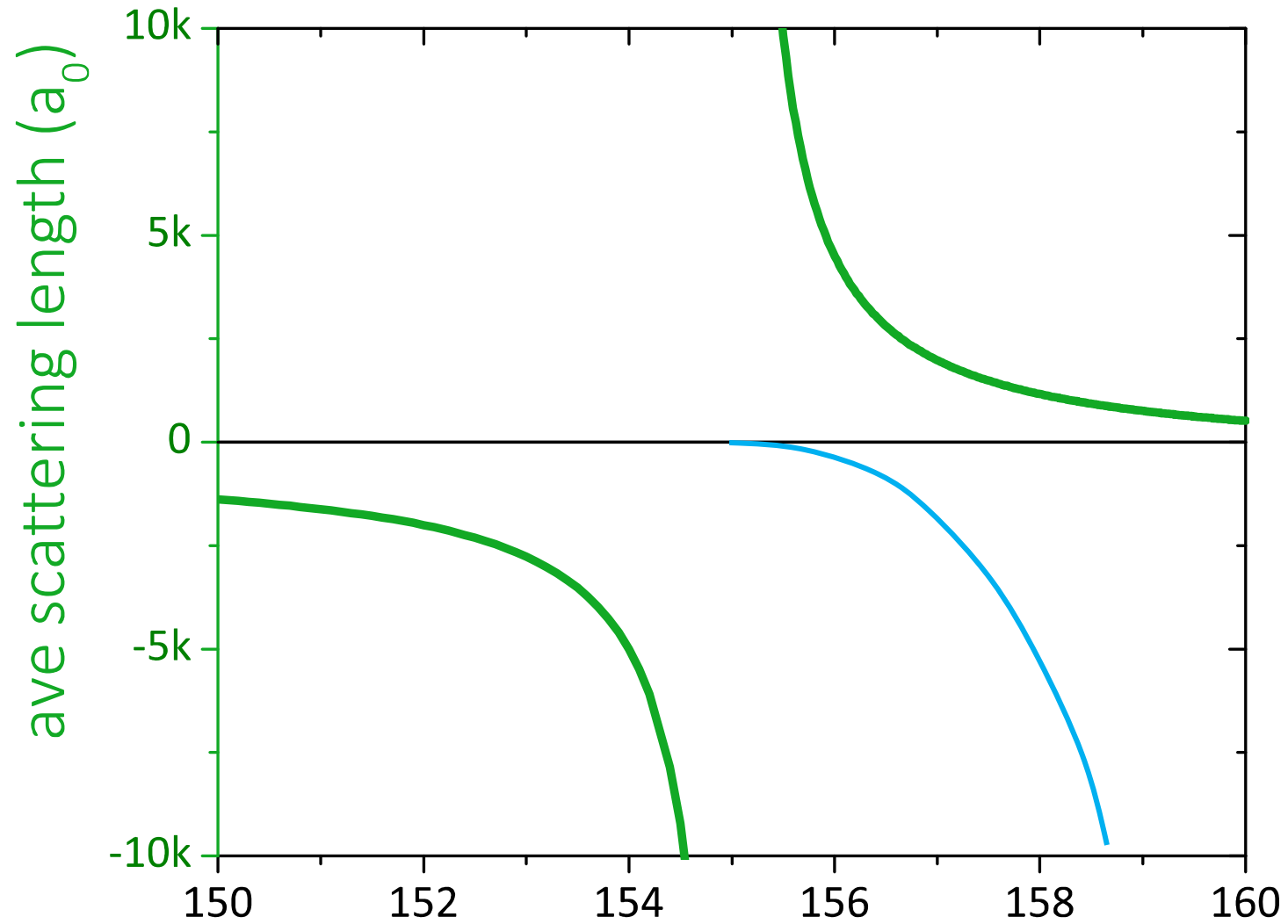
Cold Bosons: The Regimes

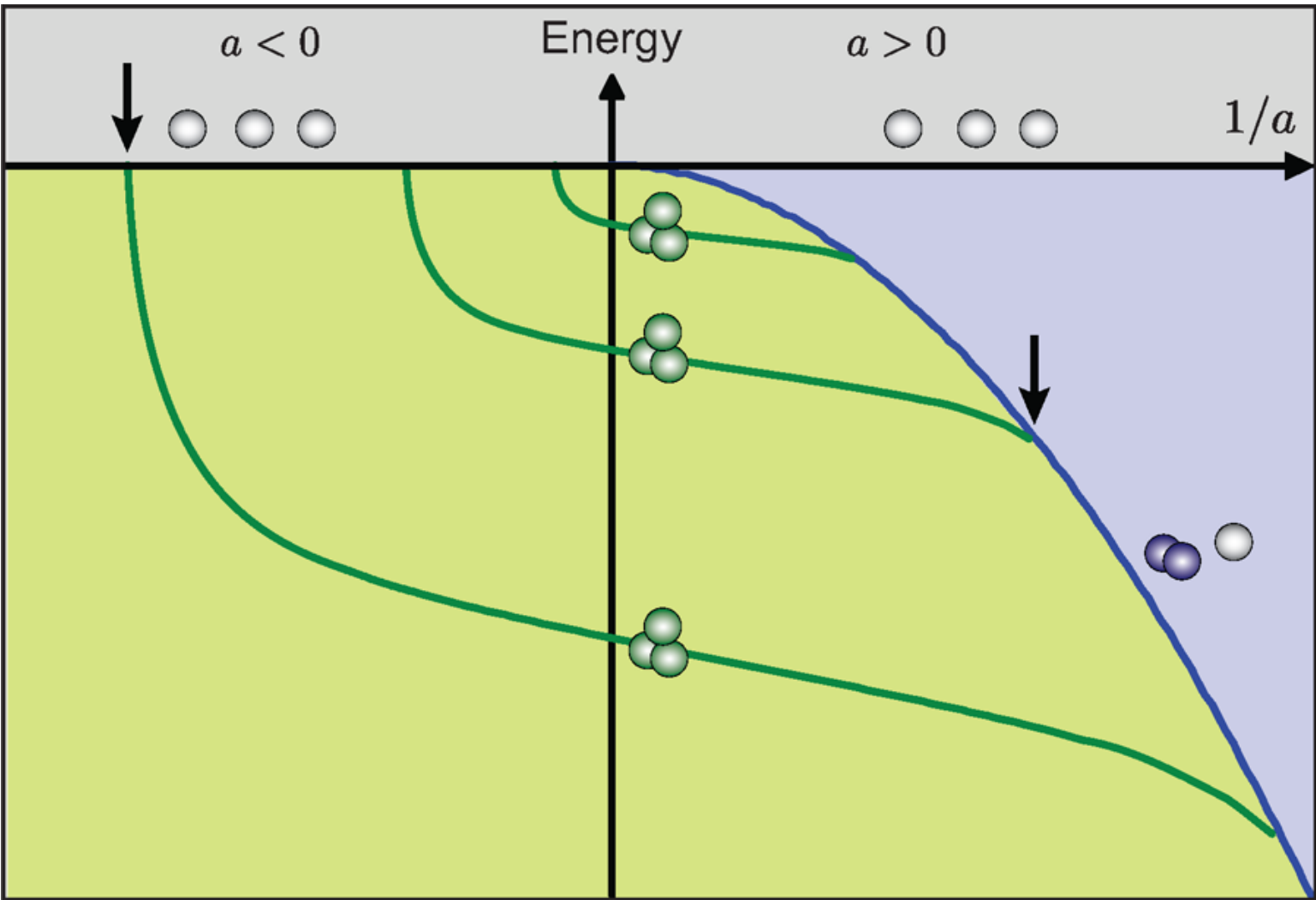


How can we learn more about
this short-lived (but not infinitely short-lived)
exotic, liquid-like *many-body* state?

First, let's review underlying *few-body* physics.

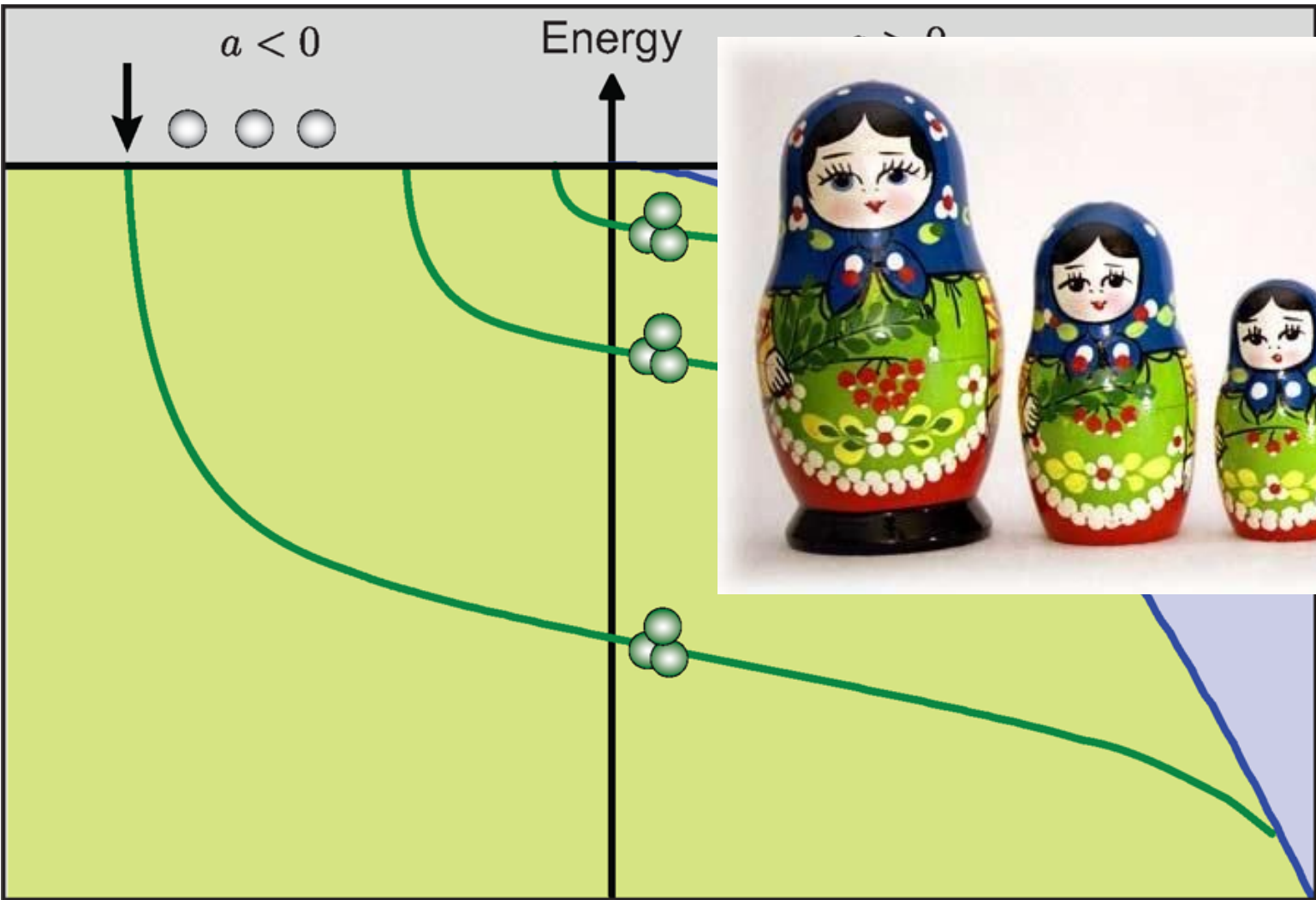
Feshbach Resonance: evidence for existence of... a two-body boundstate!





From Ferlaino and Grimm (Phys. Today, 2010)

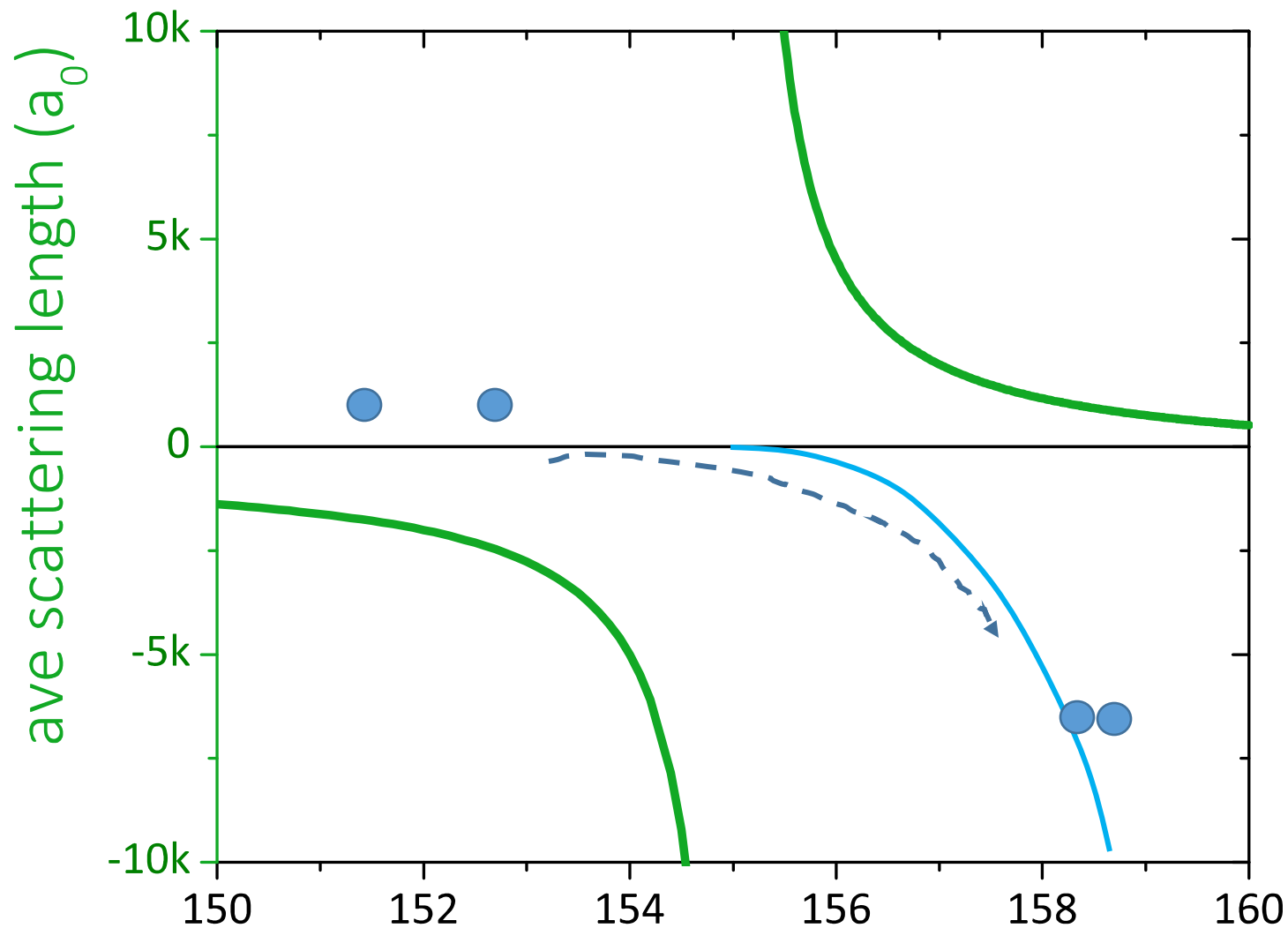
Three-body Efimov states



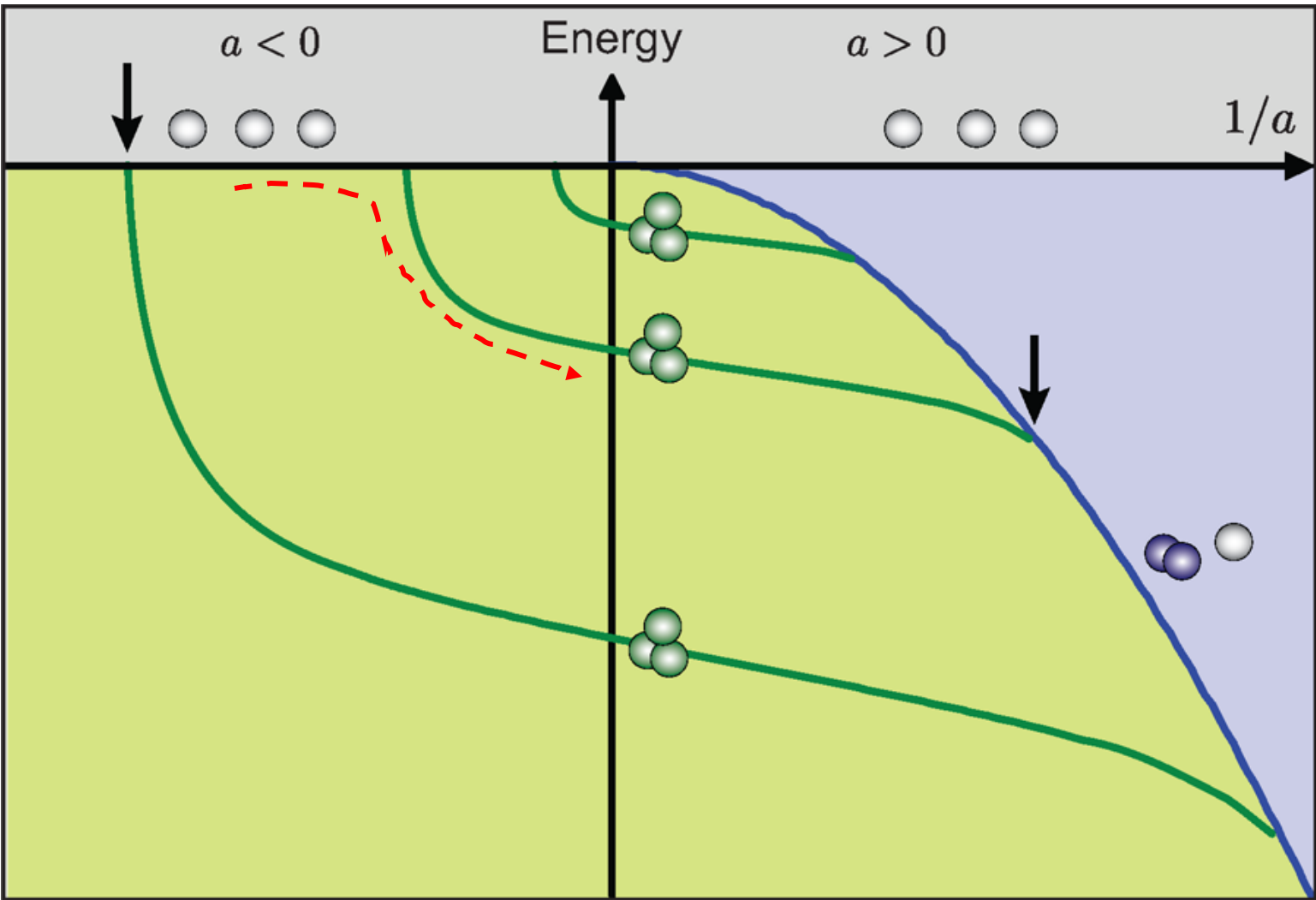
From Ferlaino and Grimm (Phys. Today, 2010)

Three-body Efimov states

Feshbach Resonance: evidence for existence of... a two-body boundstate!



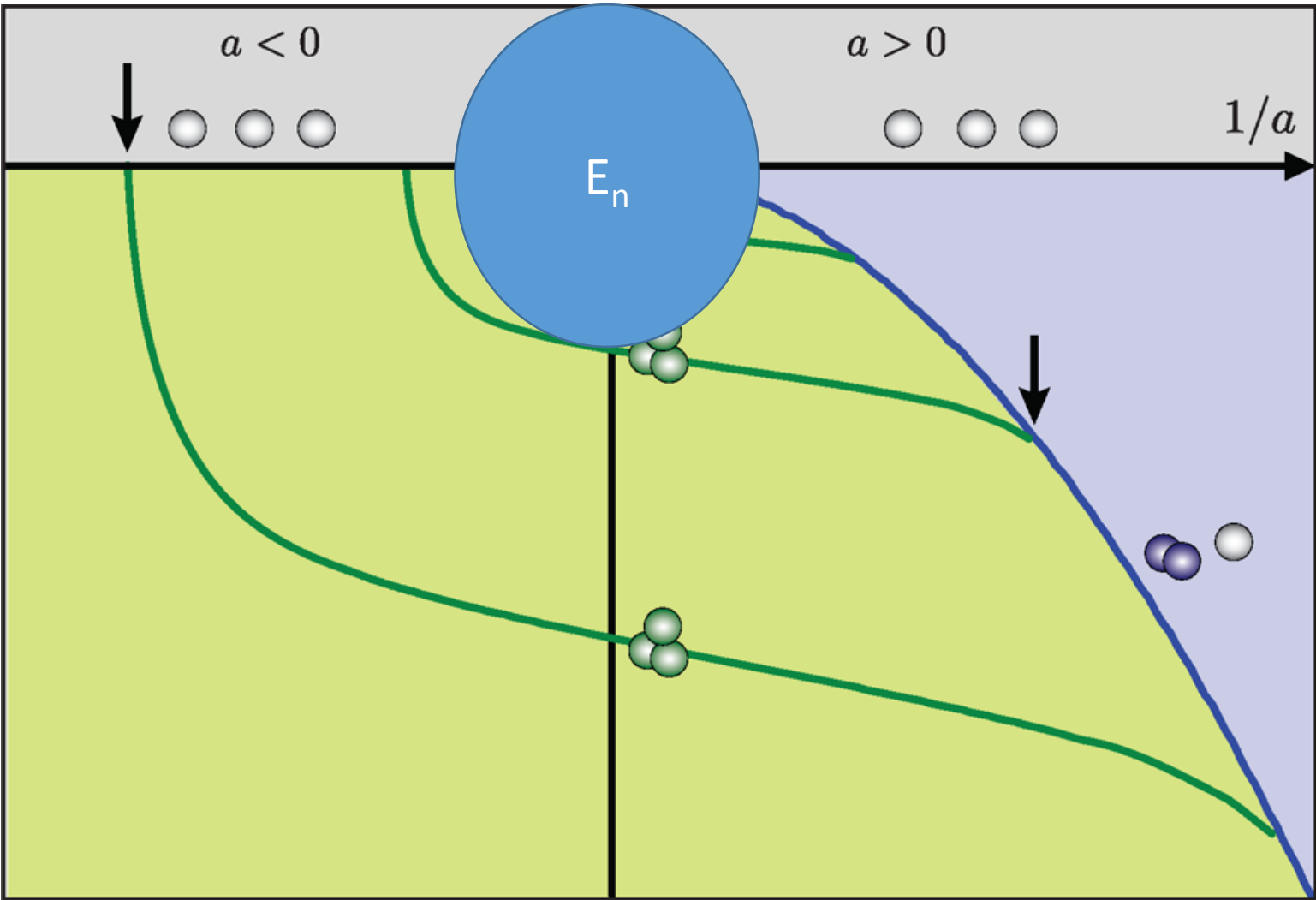
Various ways
to create a
population
in the
molecule
state.



Enhanced loss (changes in imaginary part of scattering length) identifies curve crossing in bound states

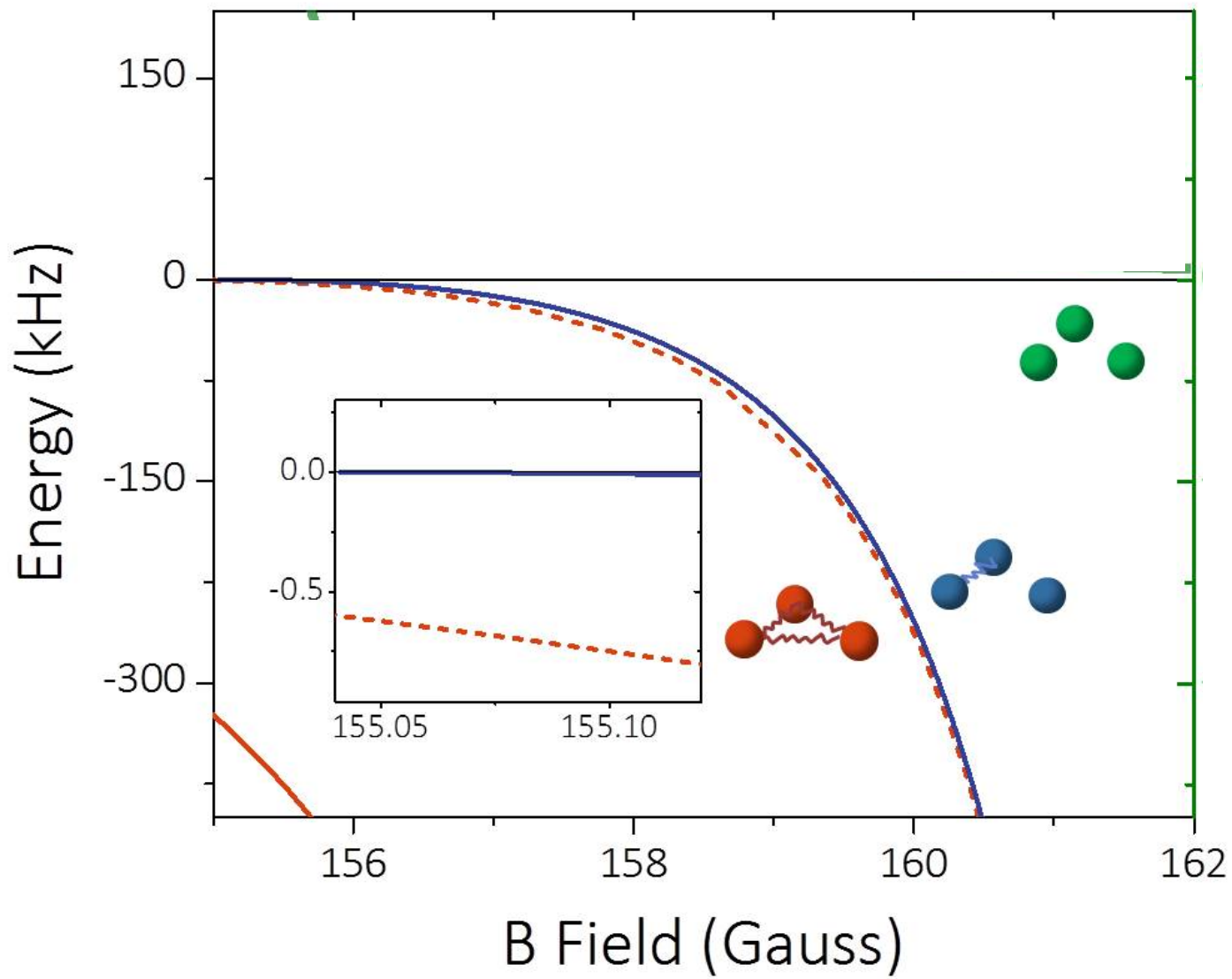
Can we, instead, create a "population"?

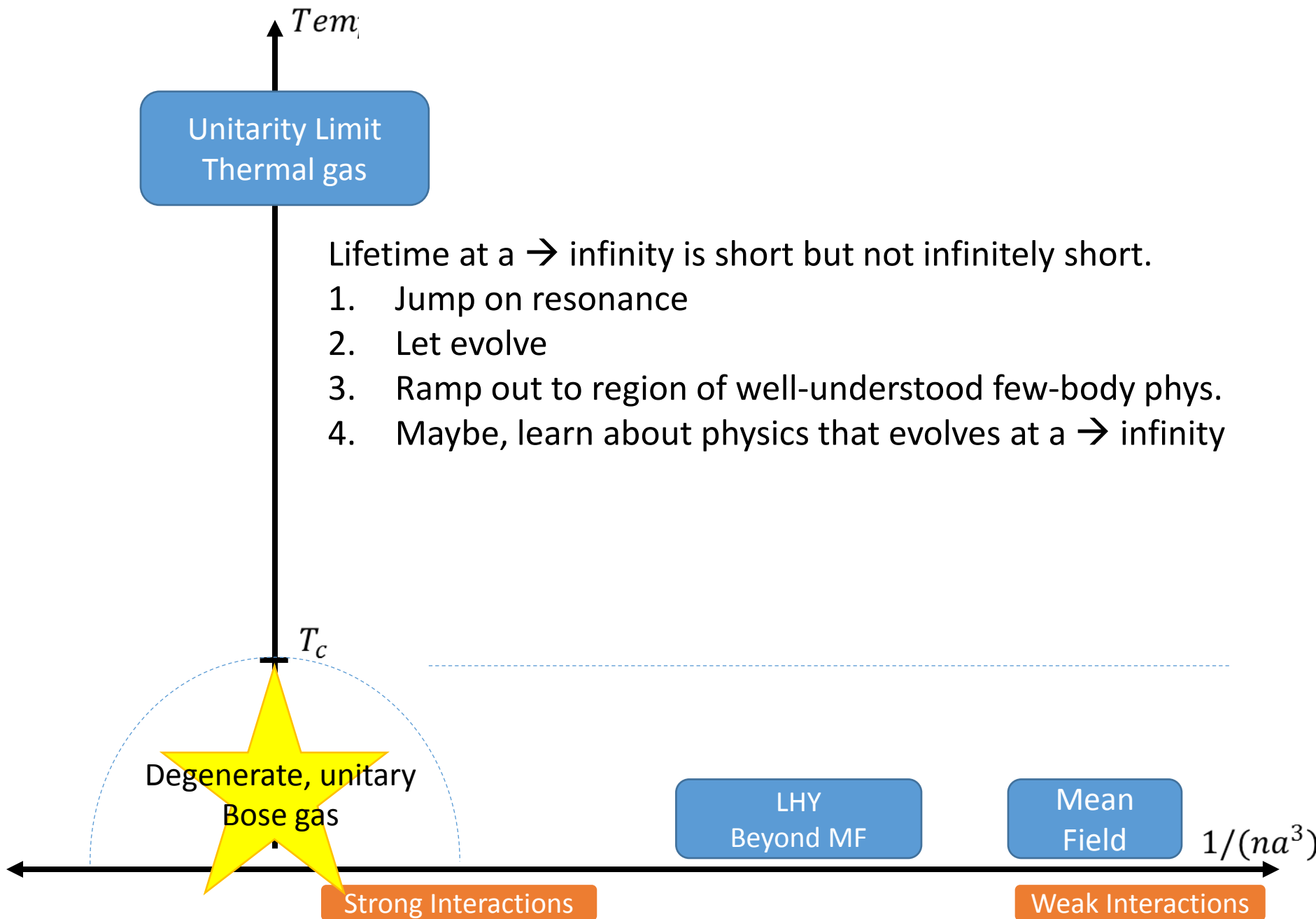
From Ferlaino and Grimm (Phys. Today, 2010)



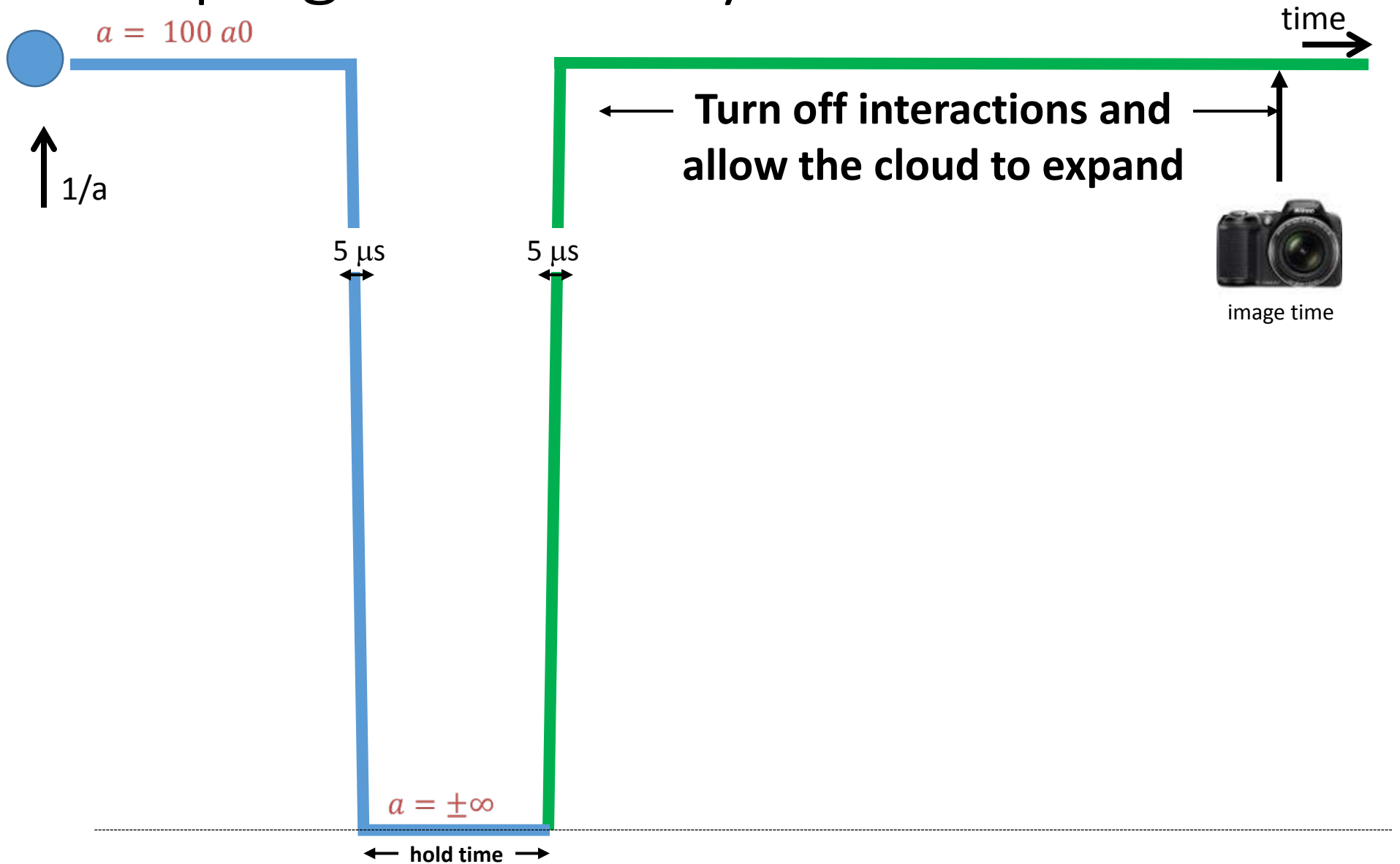
From Ferlaino and Grimm (Phys. Today, 2010)

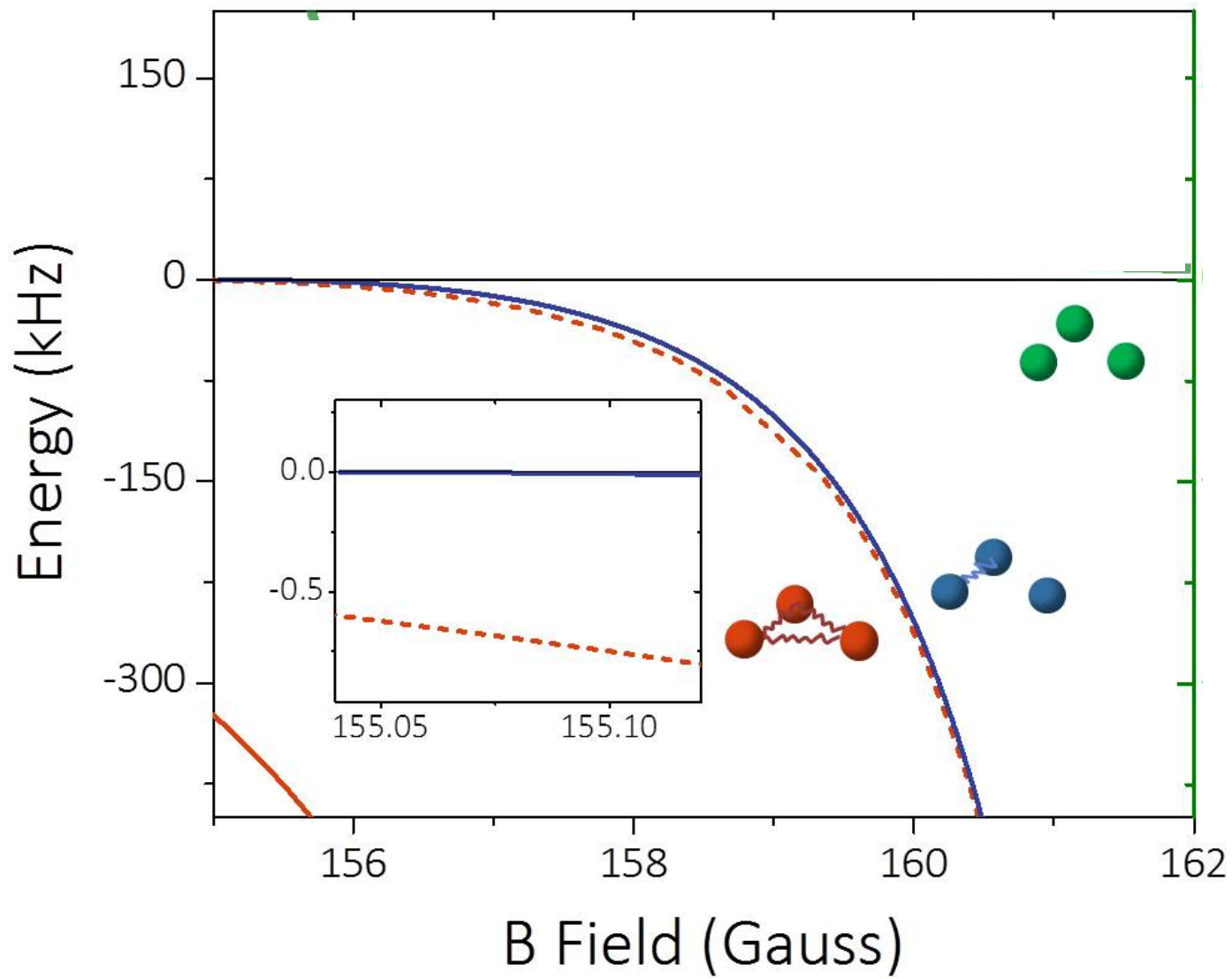
Three-body Efimov states

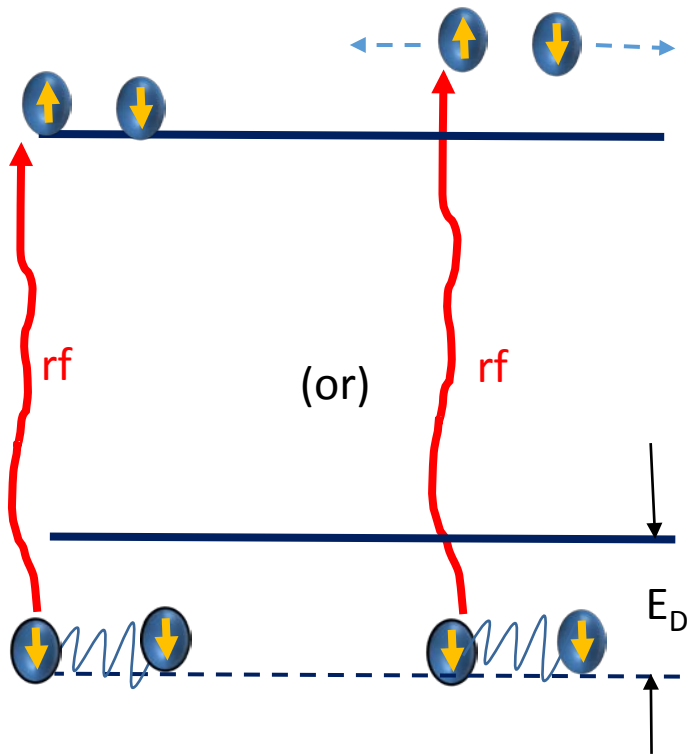




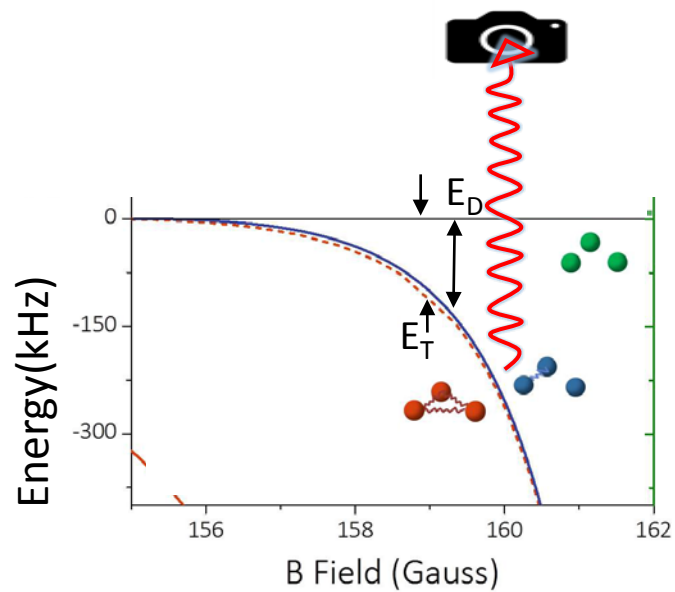
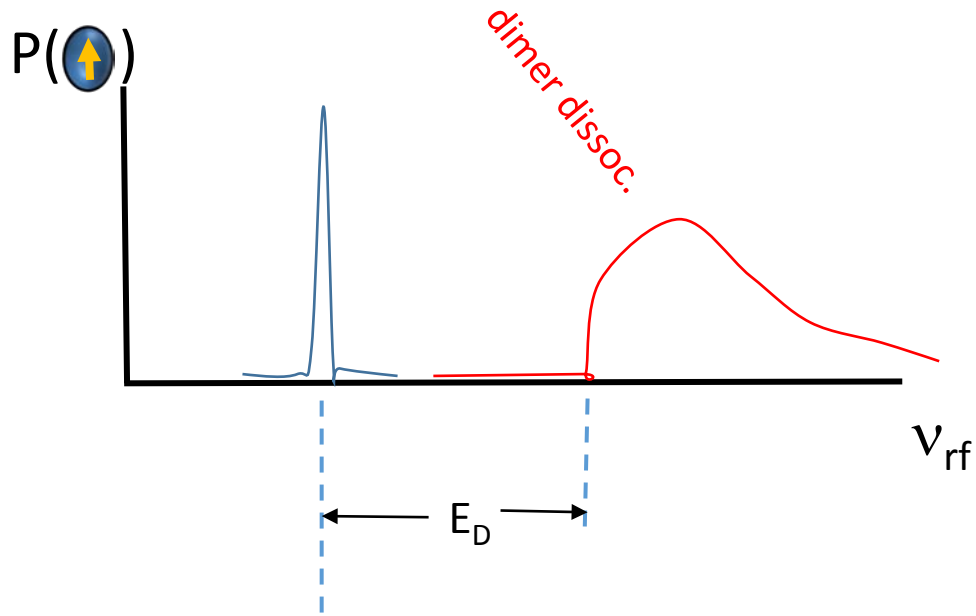
Jumping to Unitarity

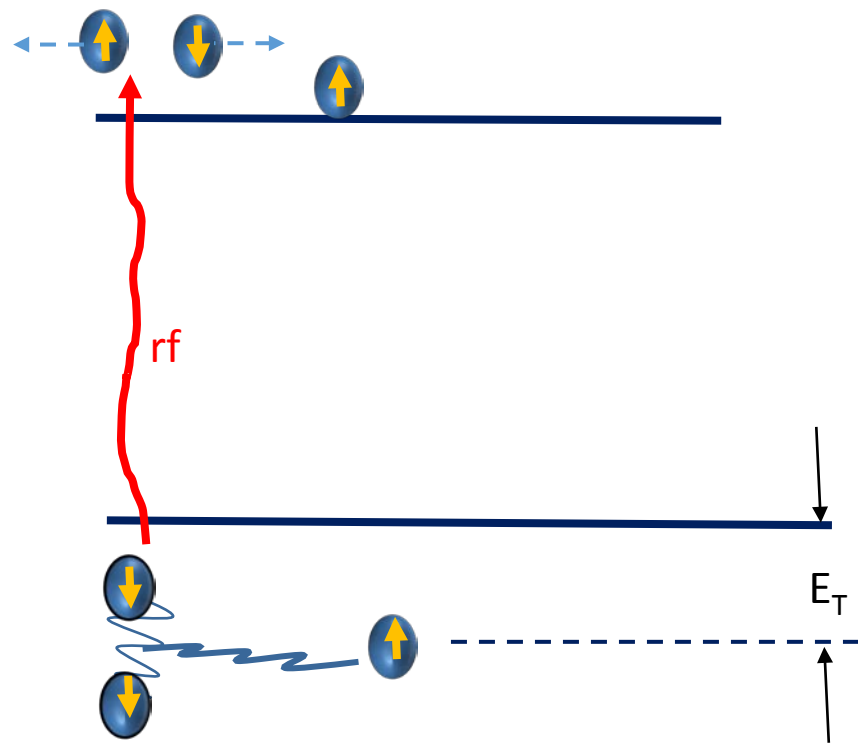




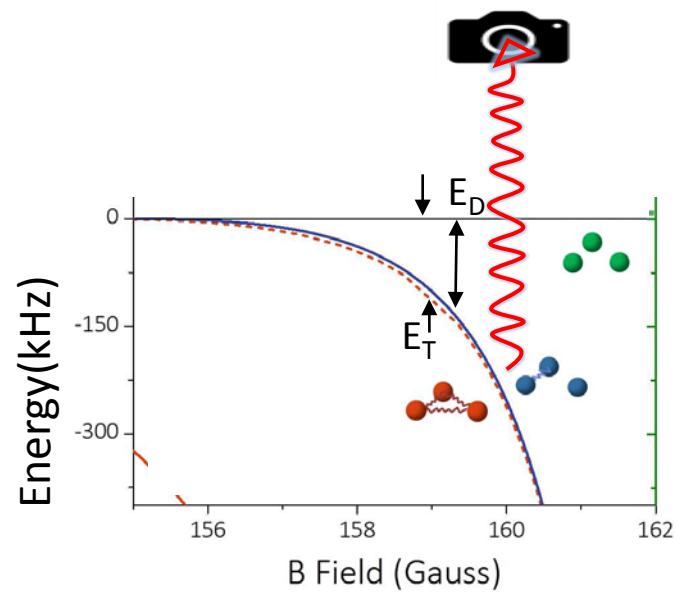
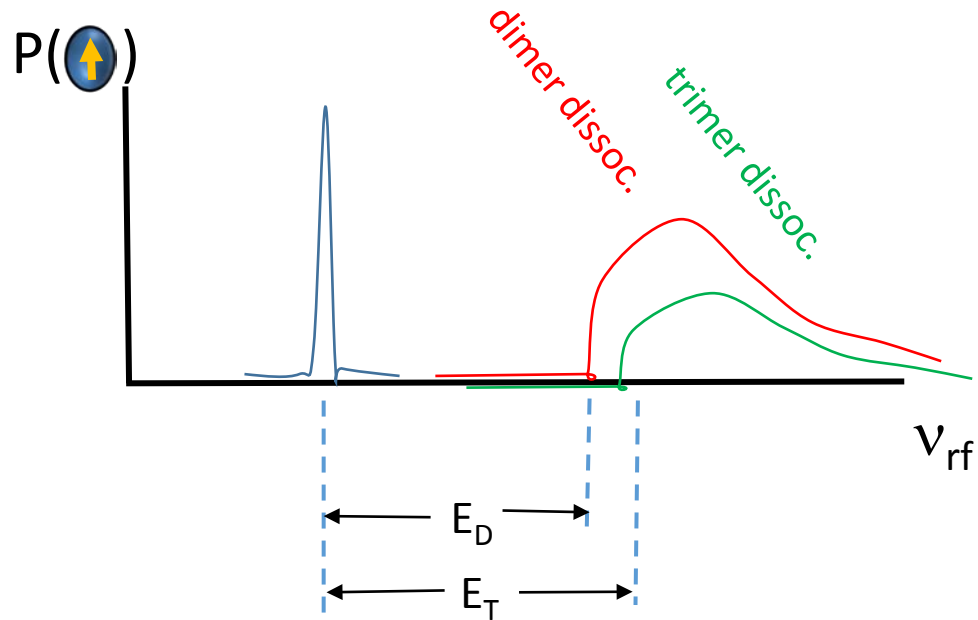


Dimer Dissociation

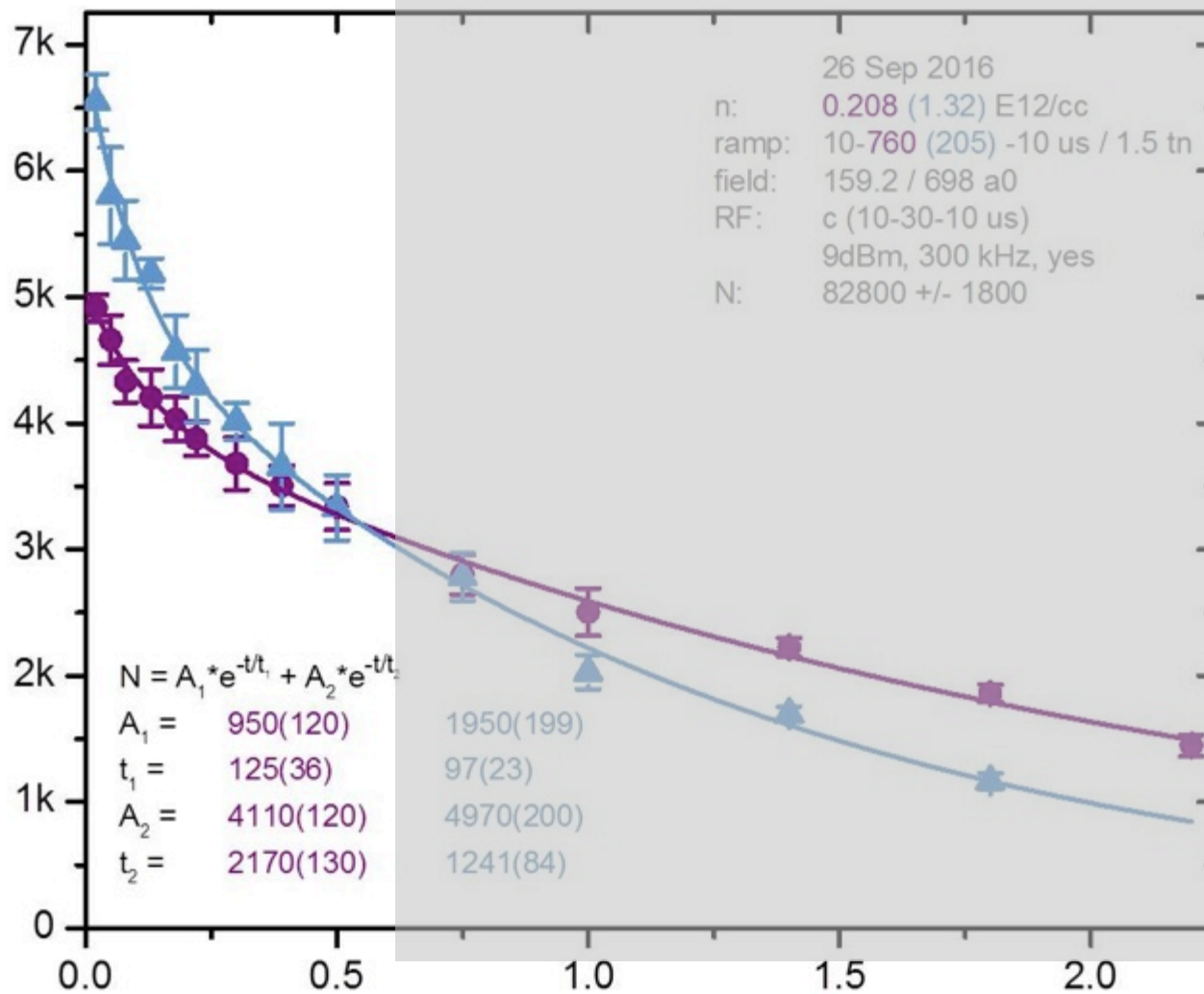




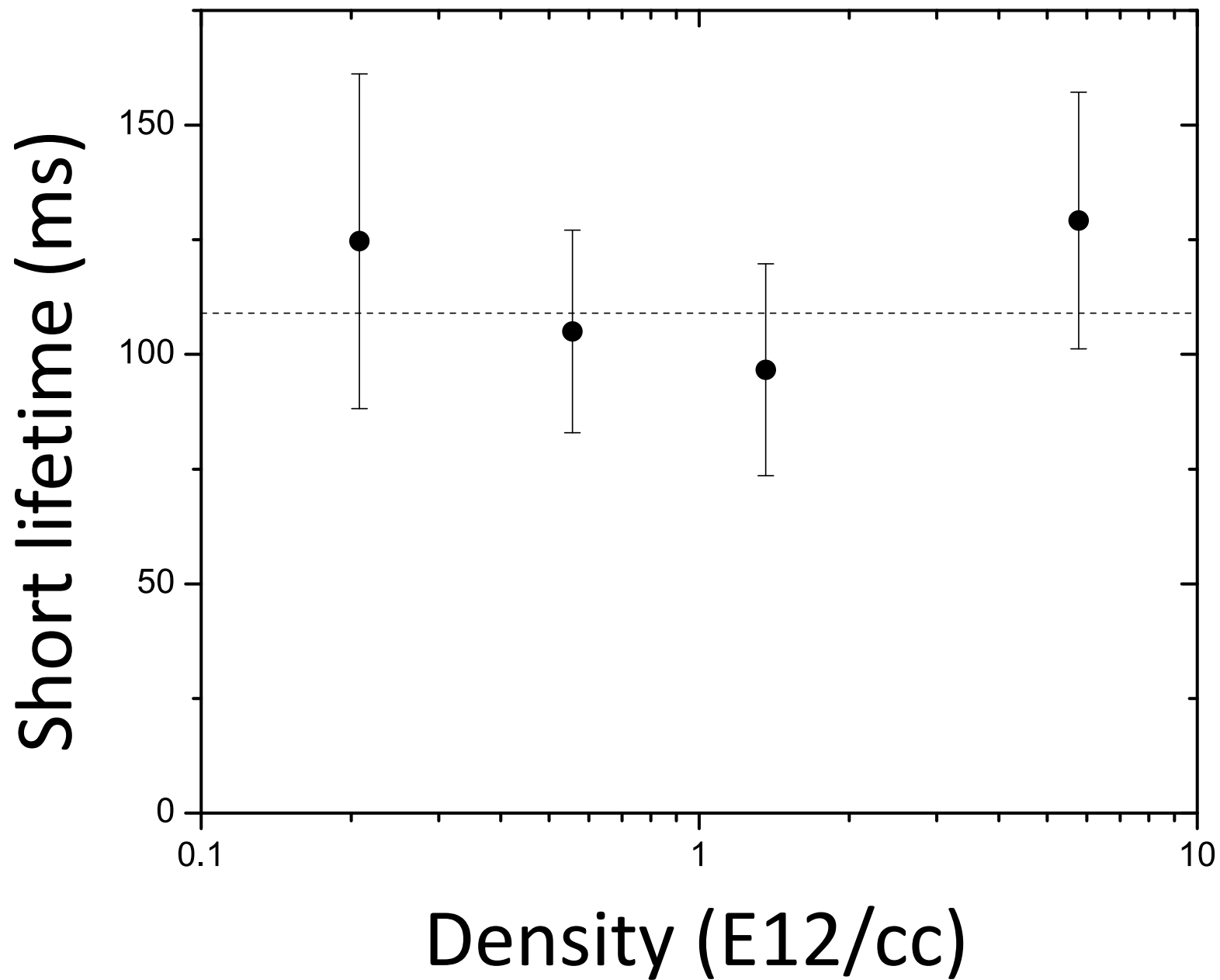
Trimer Dissociation



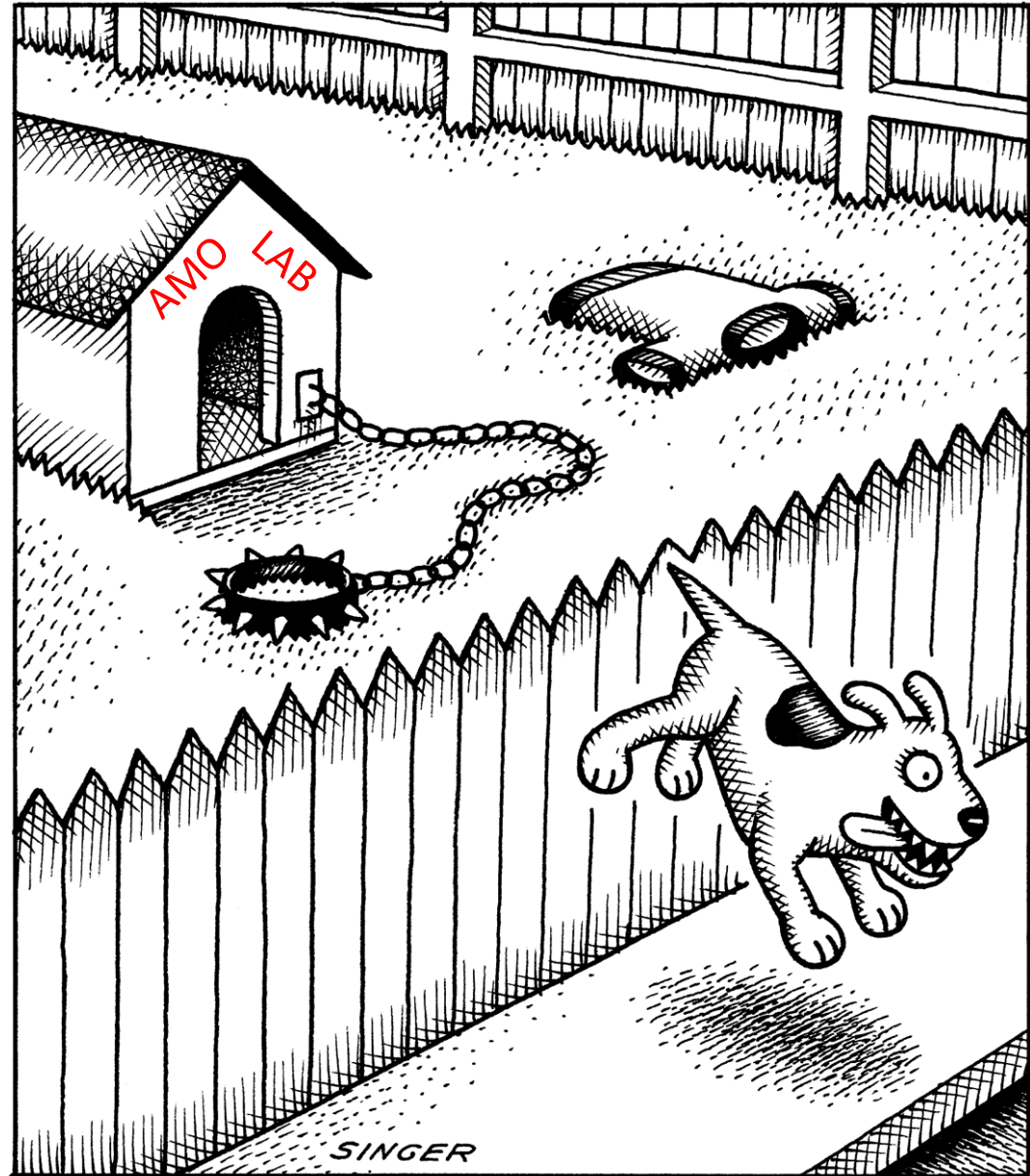
molecule number

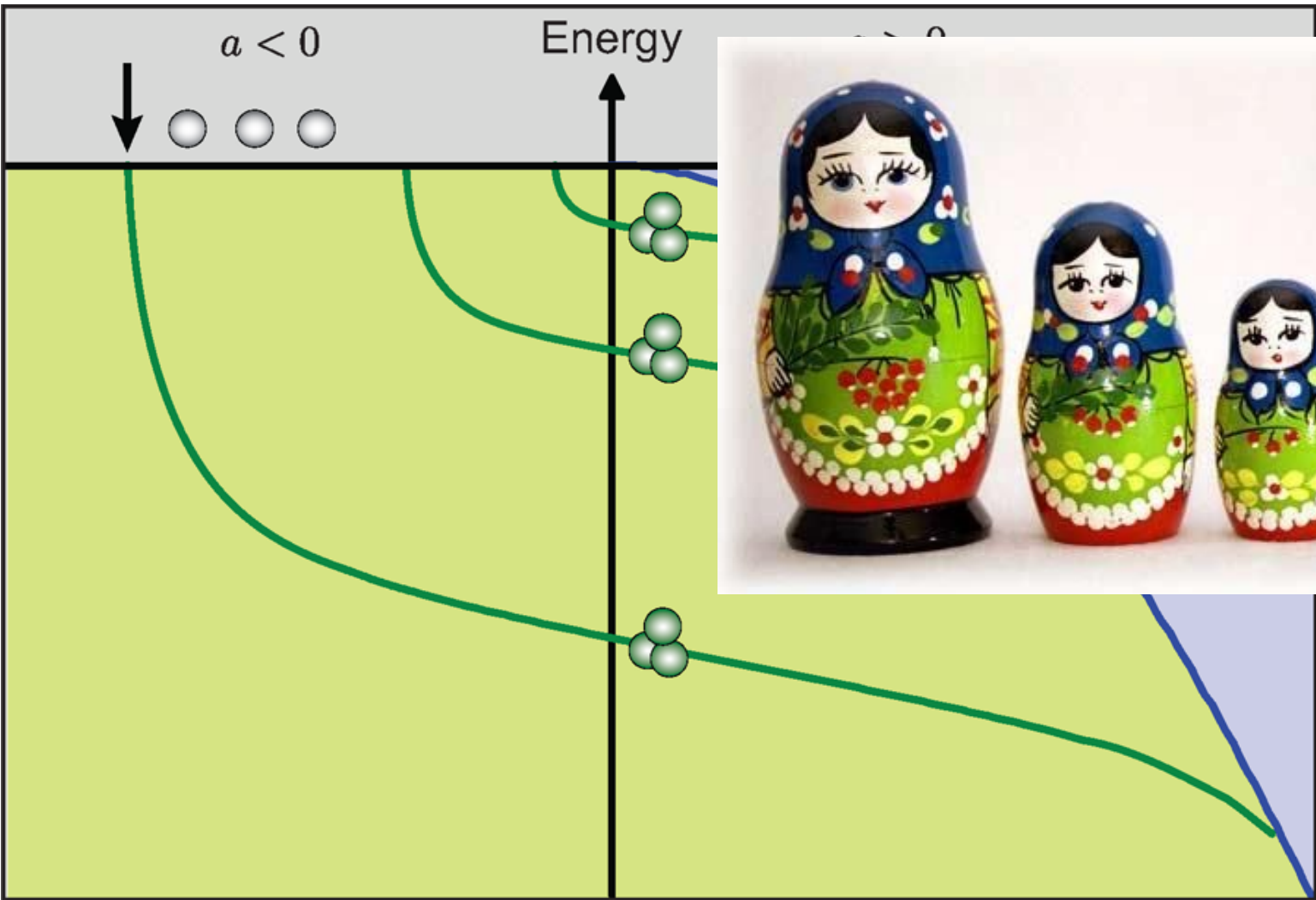


hold time (μs)



When Atomic
Physicists **Escape**:
Baryogenesis;
Materials; Liquids





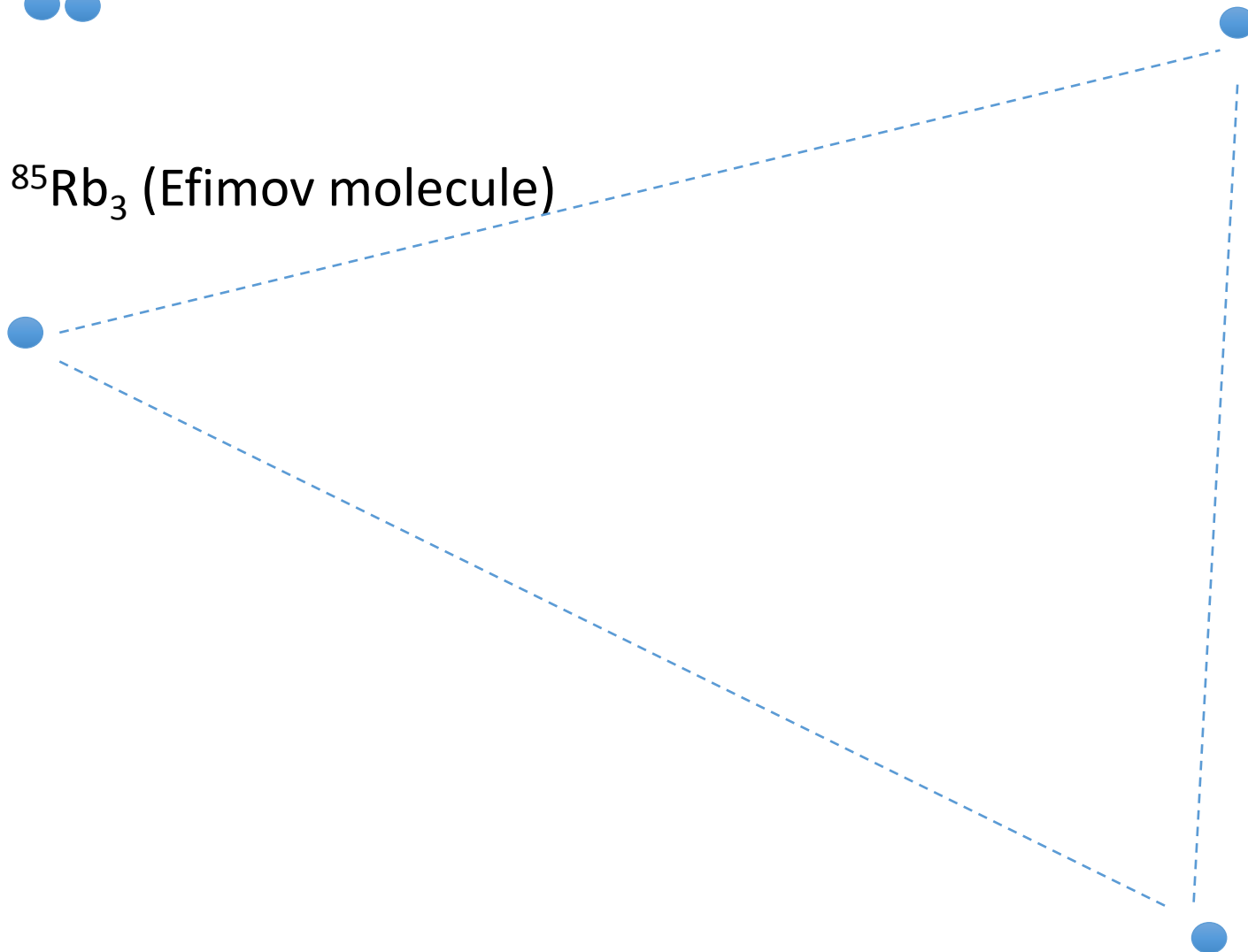
From Ferlaino and Grimm (Phys. Today, 2010)

Three-body Efimov states

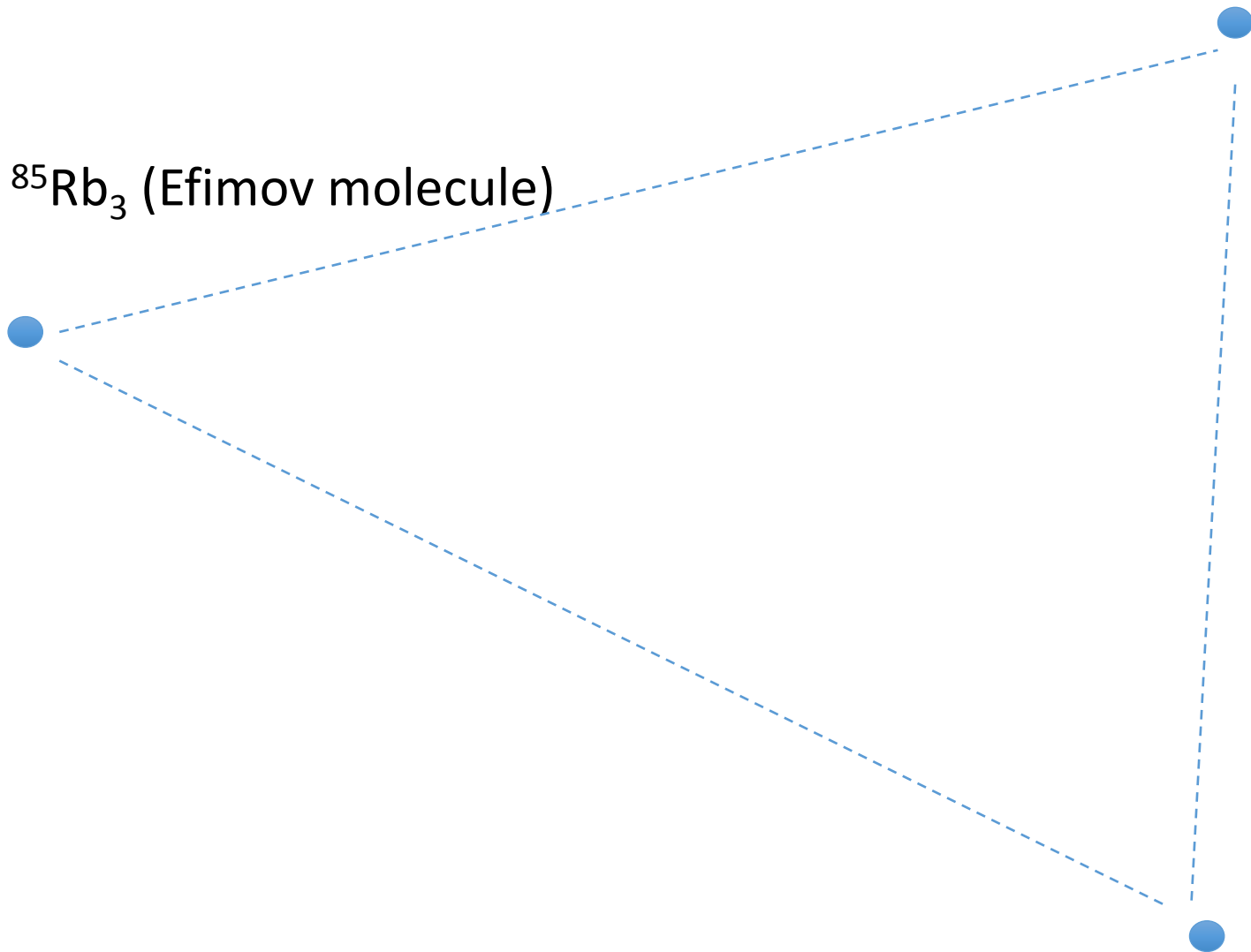
O_2 (traditional two-atom molecule)



$^{85}Rb_3$ (Efimov molecule)



$^{85}\text{Rb}_3$ (Efimov molecule)



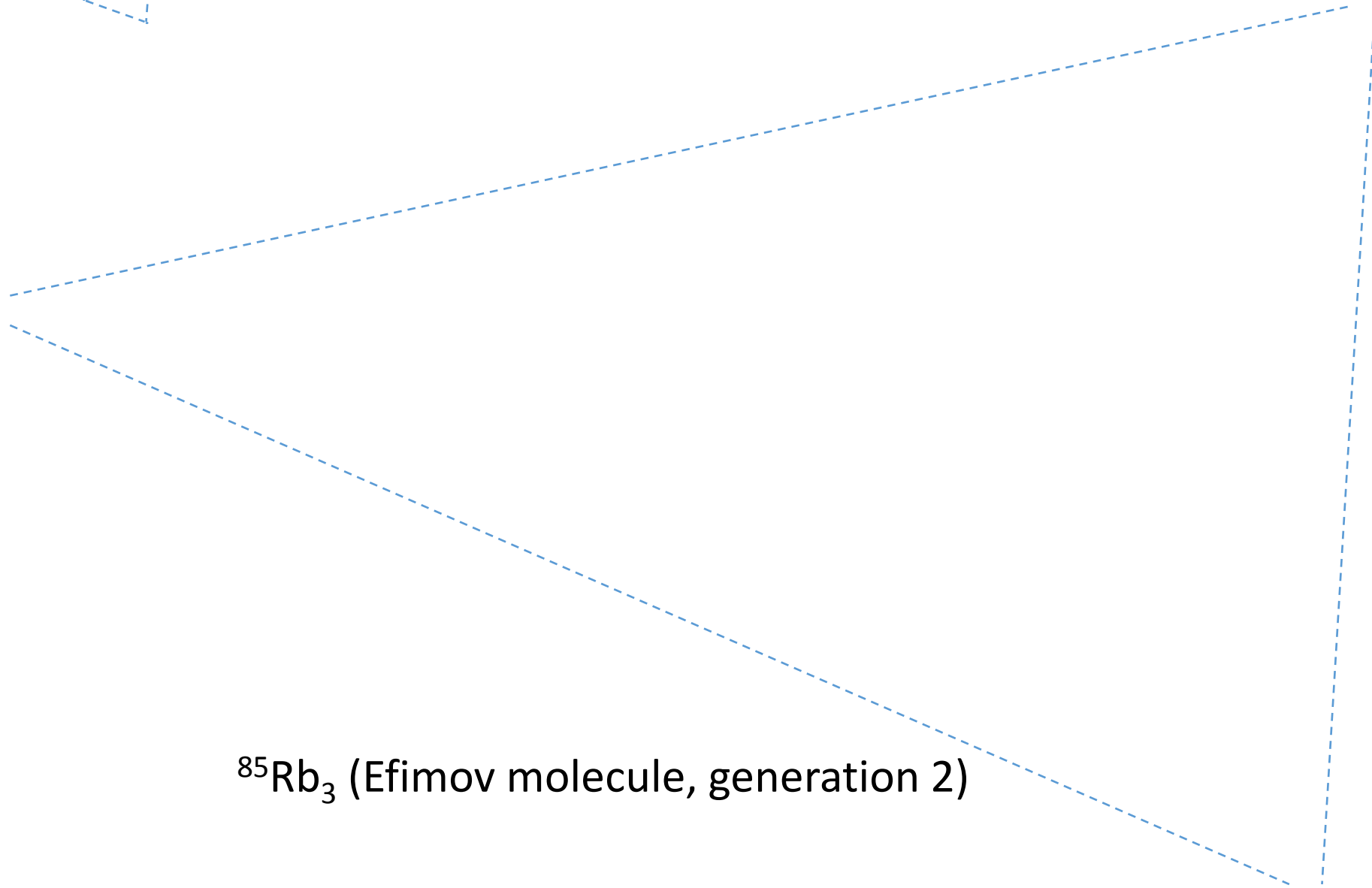
$^{85}\text{Rb}_3$ (Efimov molecule, generation 1)



$^{85}\text{Rb}_3$ (Efimov molecule, generation 1)

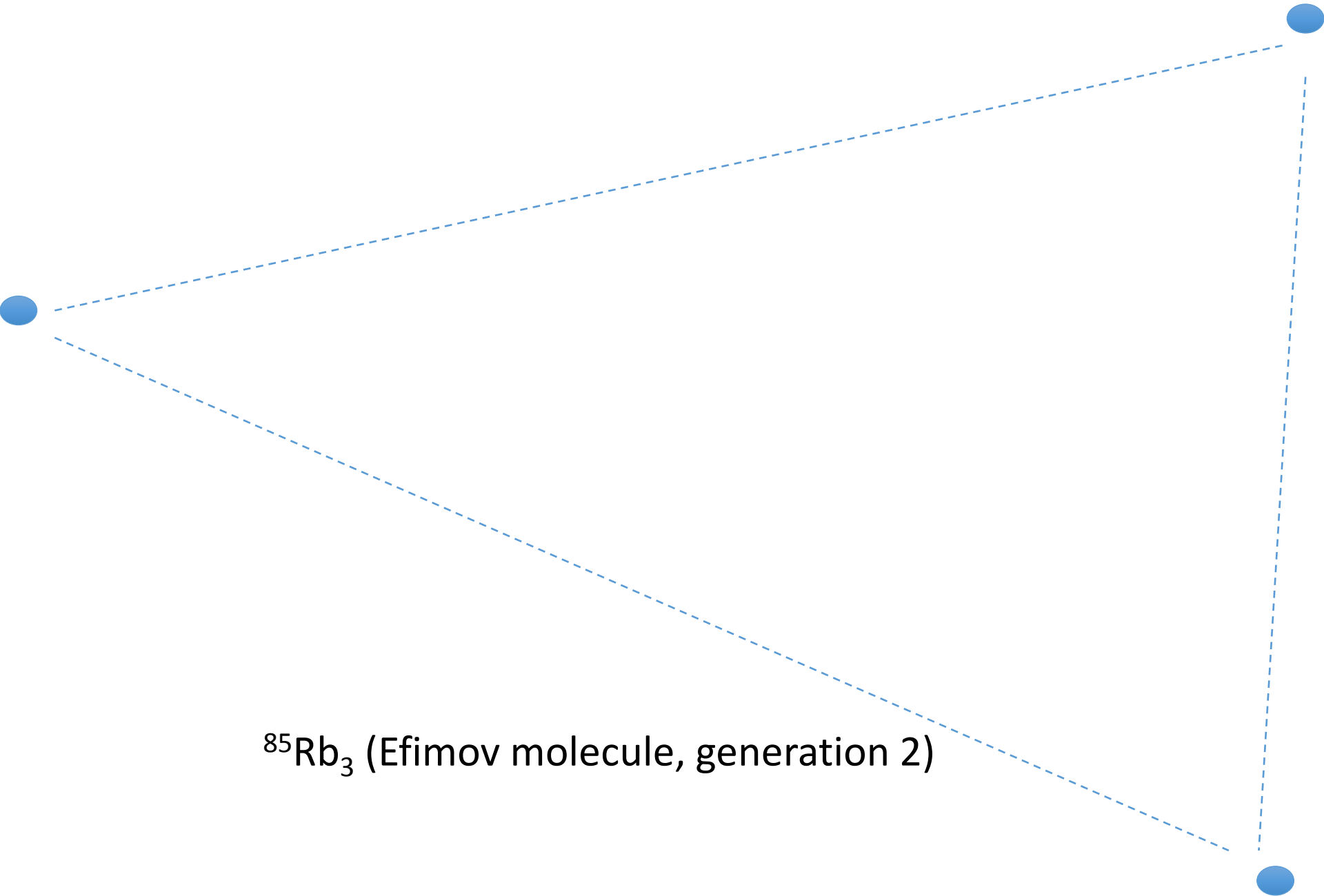


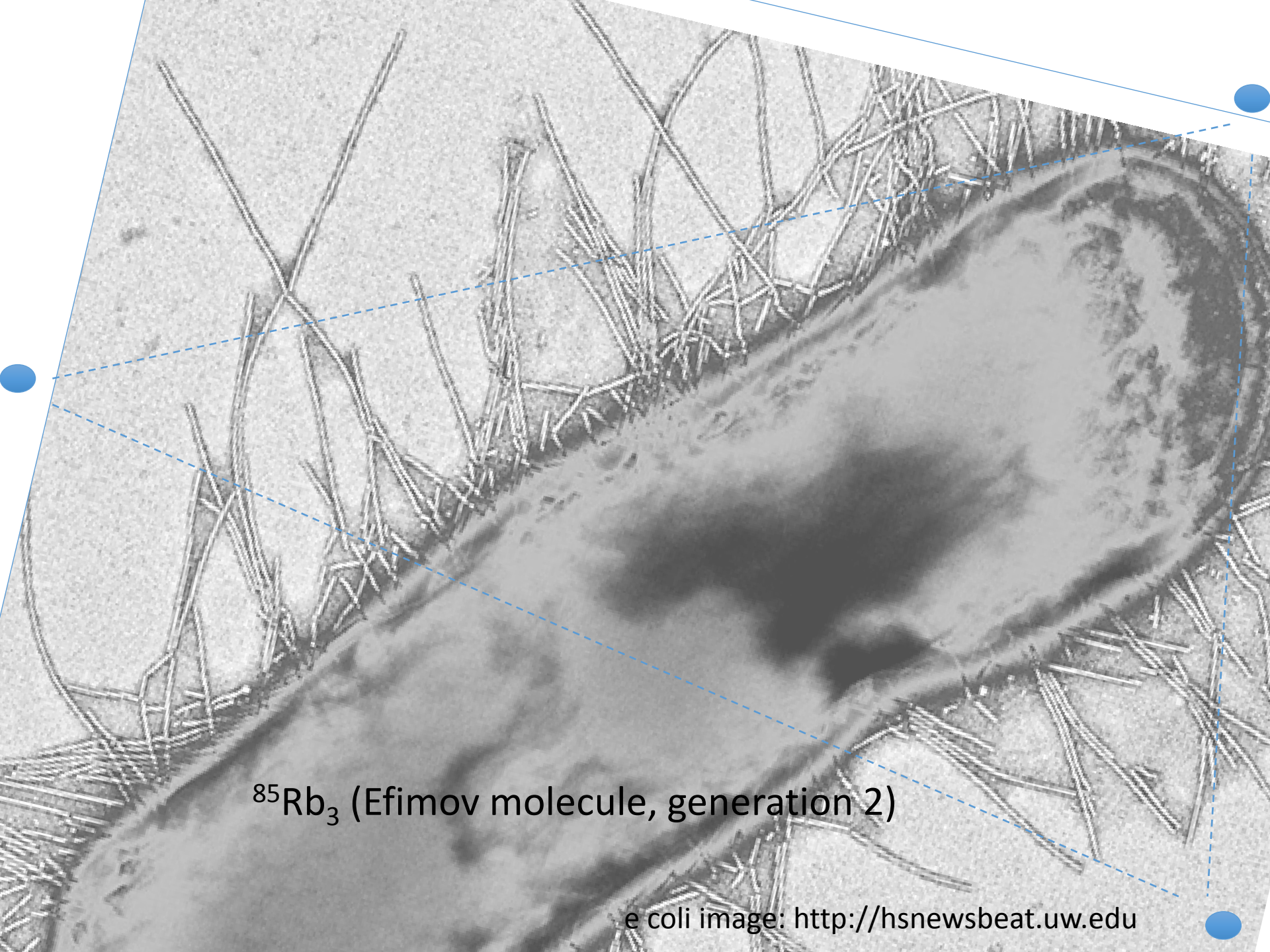
$^{85}\text{Rb}_3$ (Efimov molecule, generation 2)



From one “generation” to next,
linear size increases $\times 22.7$
binding energy decreases $\times 22.7^2 = 500$
volume increases $\times 22.7^3 = 12000$

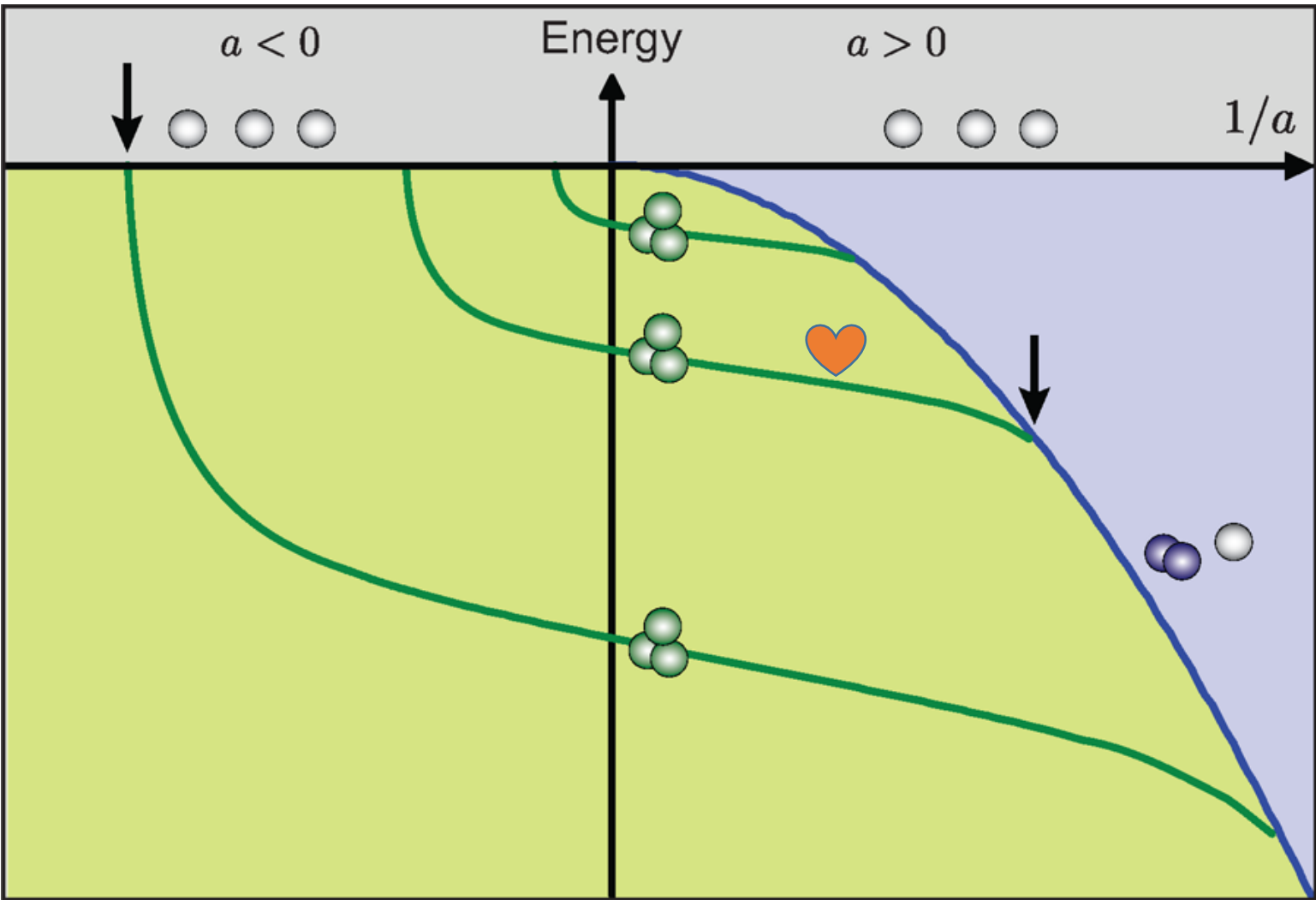






$^{85}\text{Rb}_3$ (Efimov molecule, generation 2)

e coli image: <http://hsnewsbeat.uw.edu>



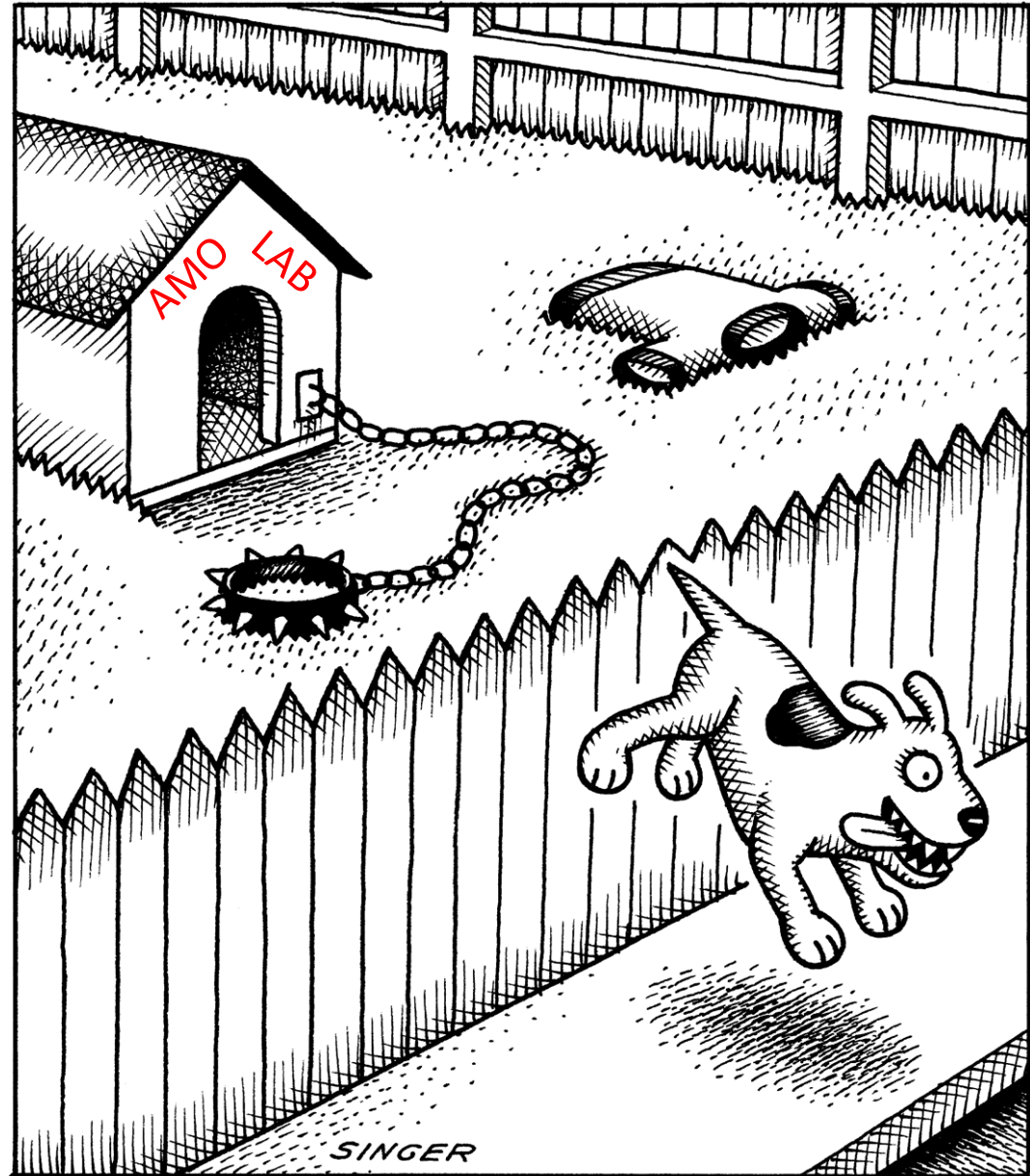
From Ferlaino and Grimm (Phys. Today, 2010)

From one “generation” to next,
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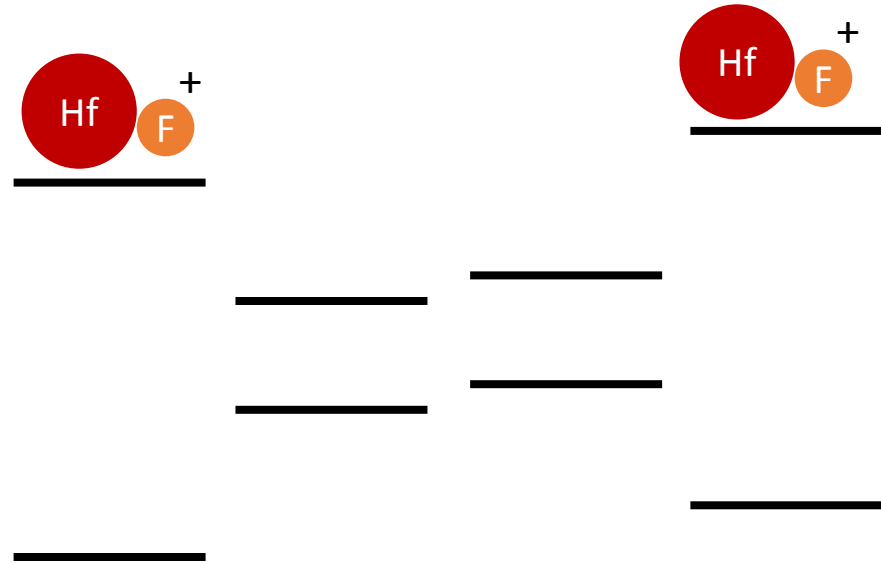


When Atomic
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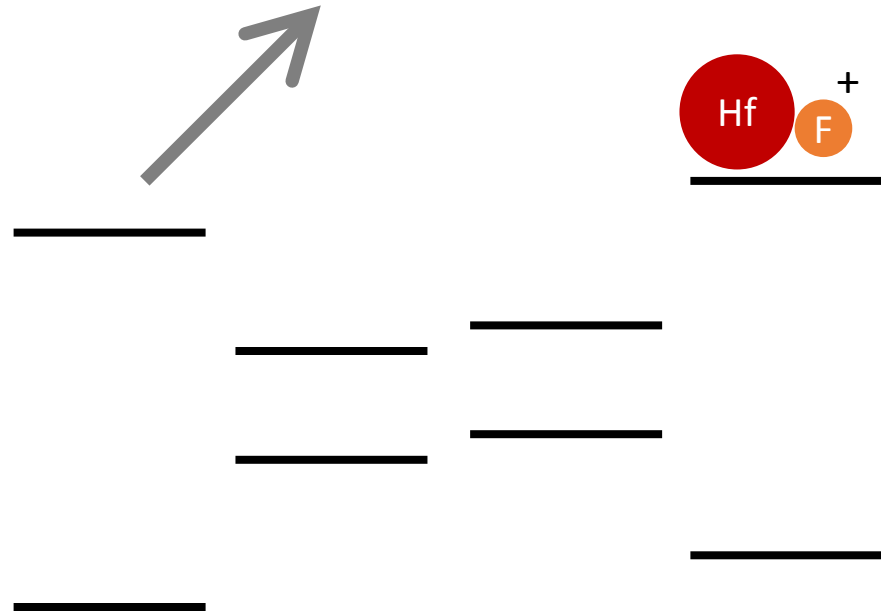
Ramsey Sequence

Transfer lasers
prepare population
in a single pair of
Stark states



Ramsey Sequence

Optically deplete the population of one m_F level using strobed circularly polarized light



Transfer

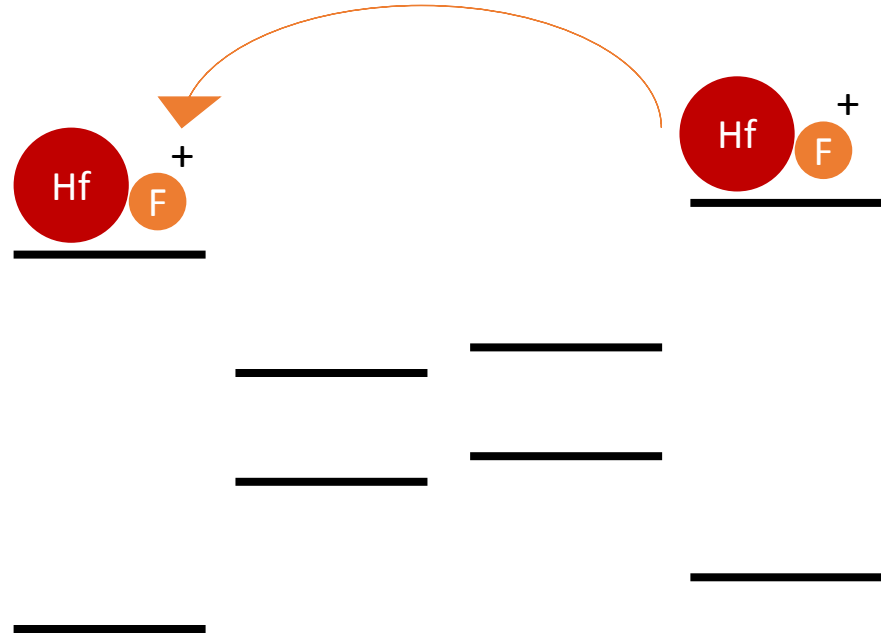
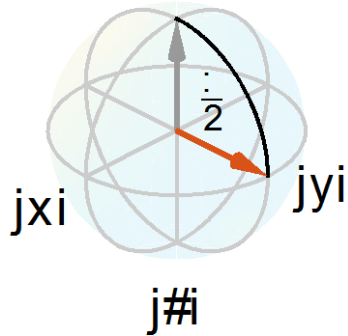


Ramsey Sequence

$\pi/2$ pulse puts system into the superposition

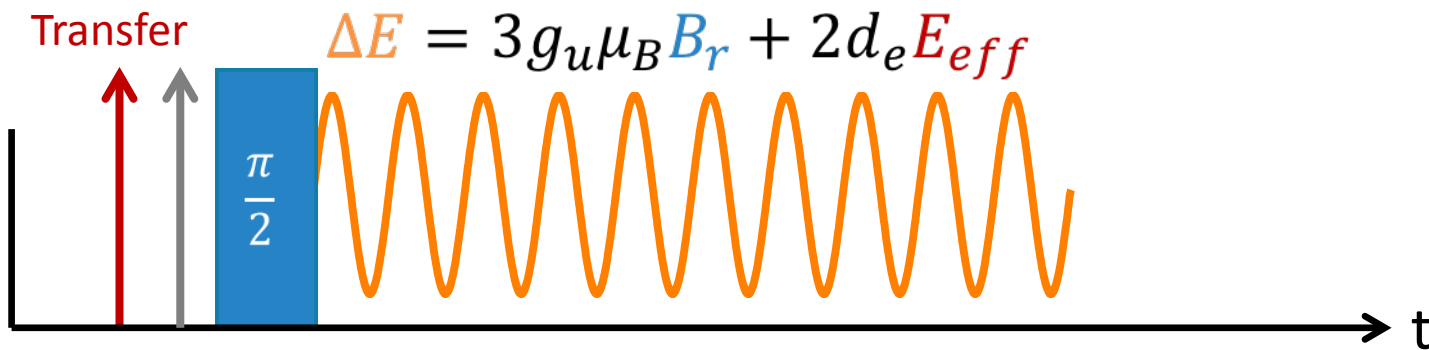
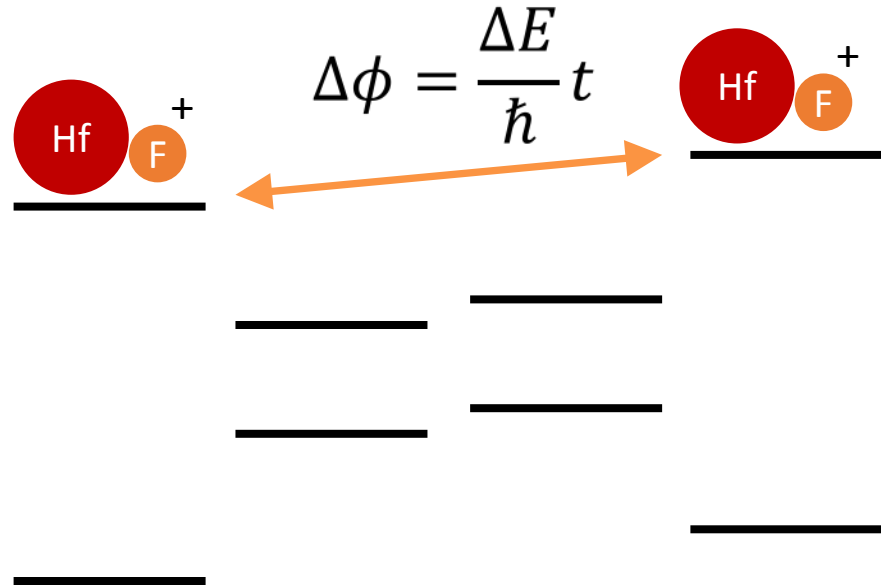
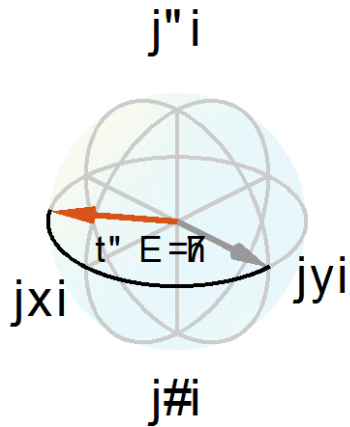
$$\left| m_F = -\frac{3}{2} \right\rangle + \left| m_F = +\frac{3}{2} \right\rangle$$

$j''i$



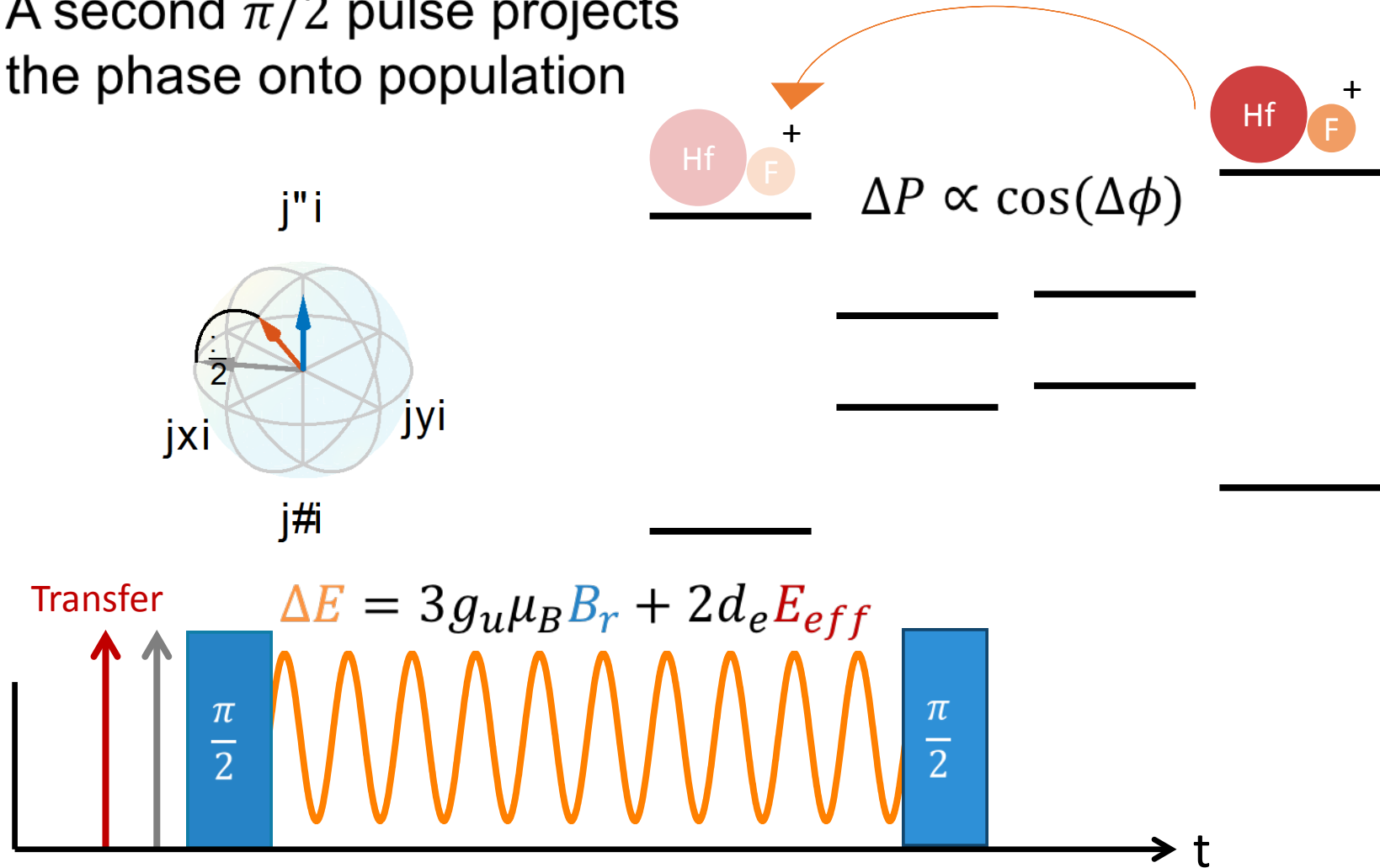
Ramsey Sequence

Free evolution



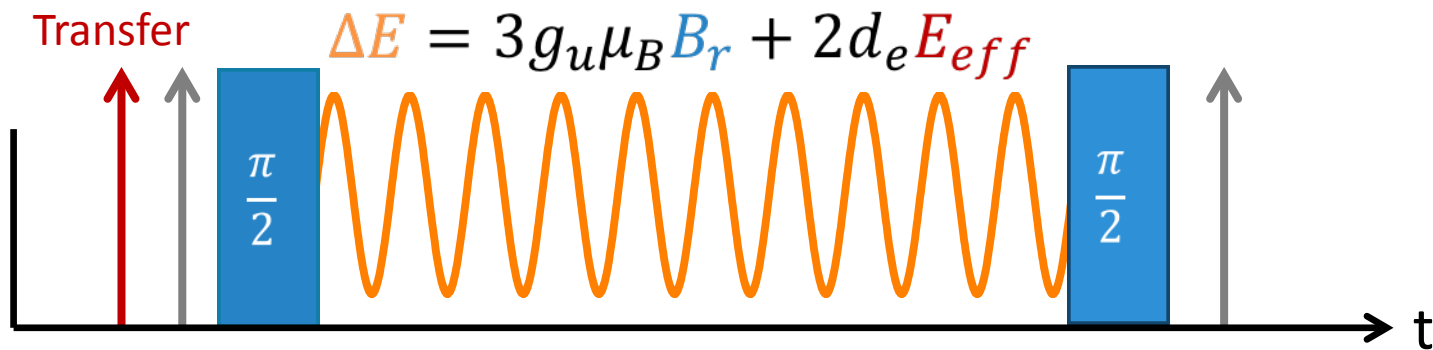
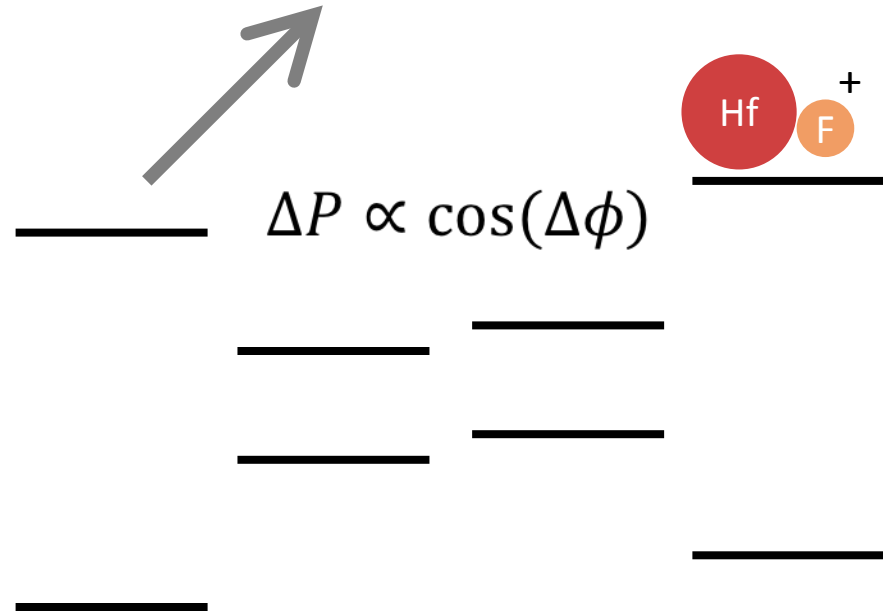
Ramsey Sequence

A second $\pi/2$ pulse projects the phase onto population



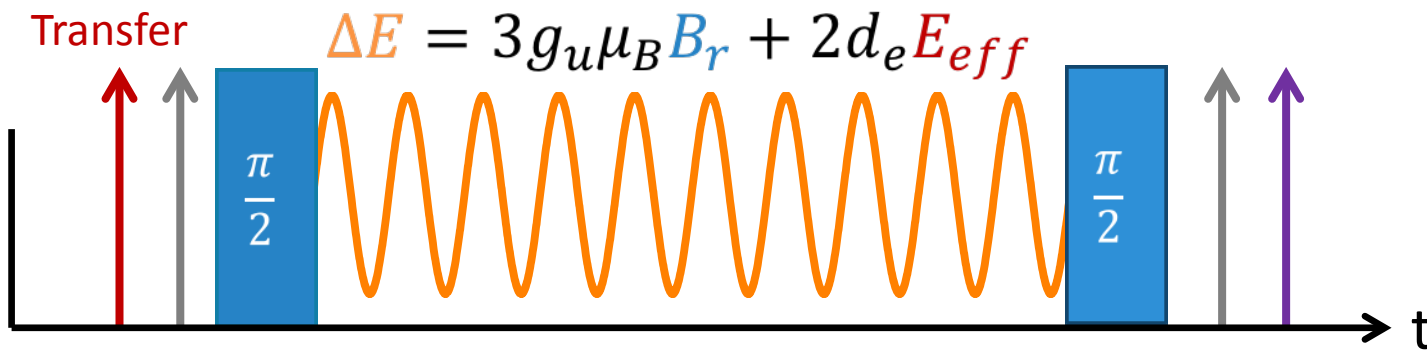
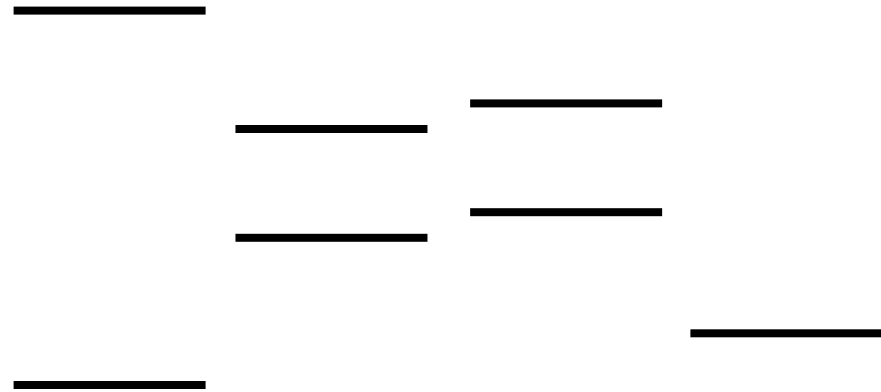
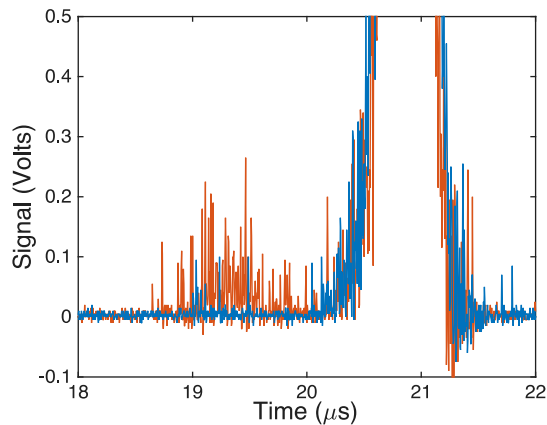
Ramsey Sequence

Optically deplete population out of one of the m_F levels

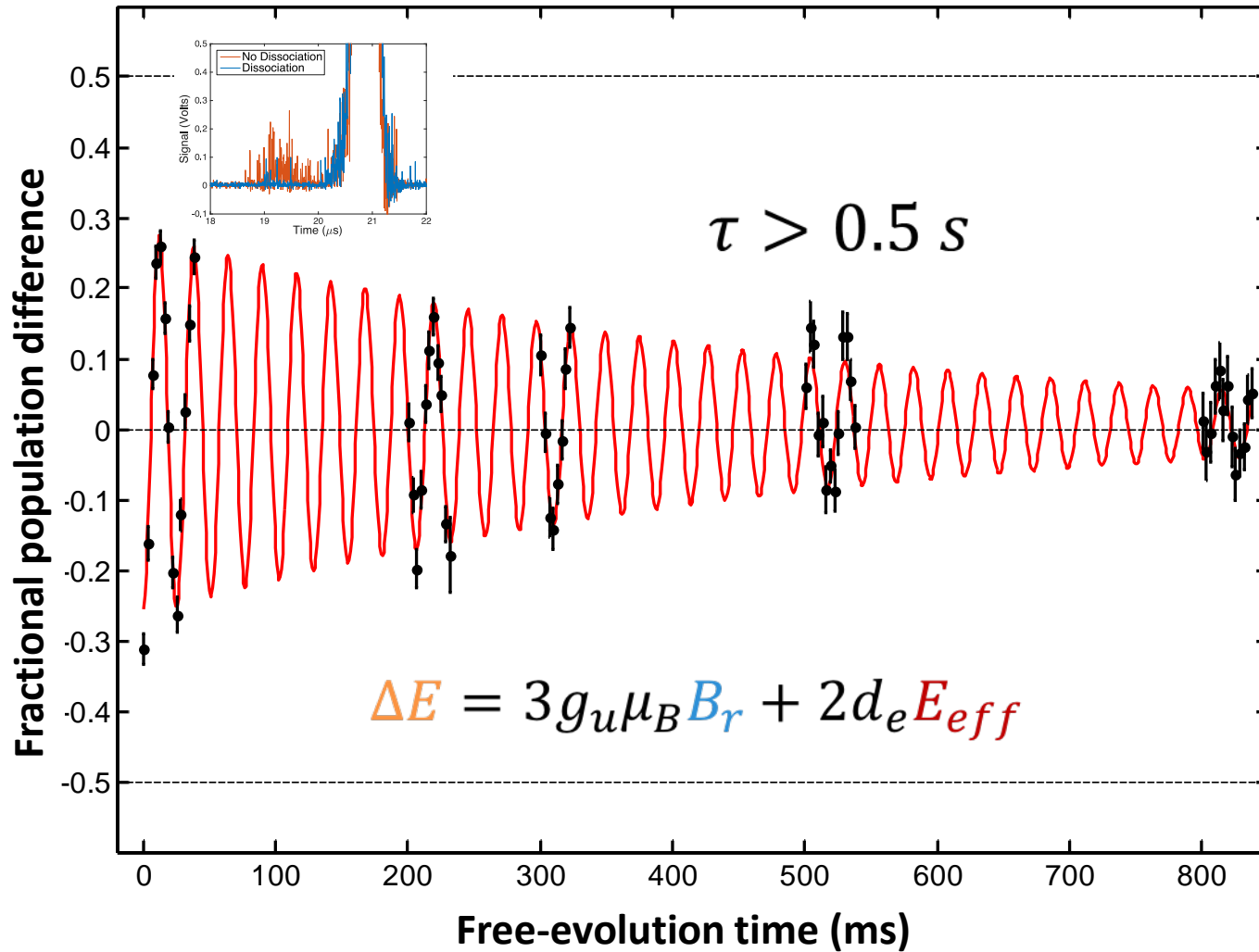


Ramsey Sequence

Dissociate all of the ions in the $J = 1$ level, and count Hf^+ ions in the trap



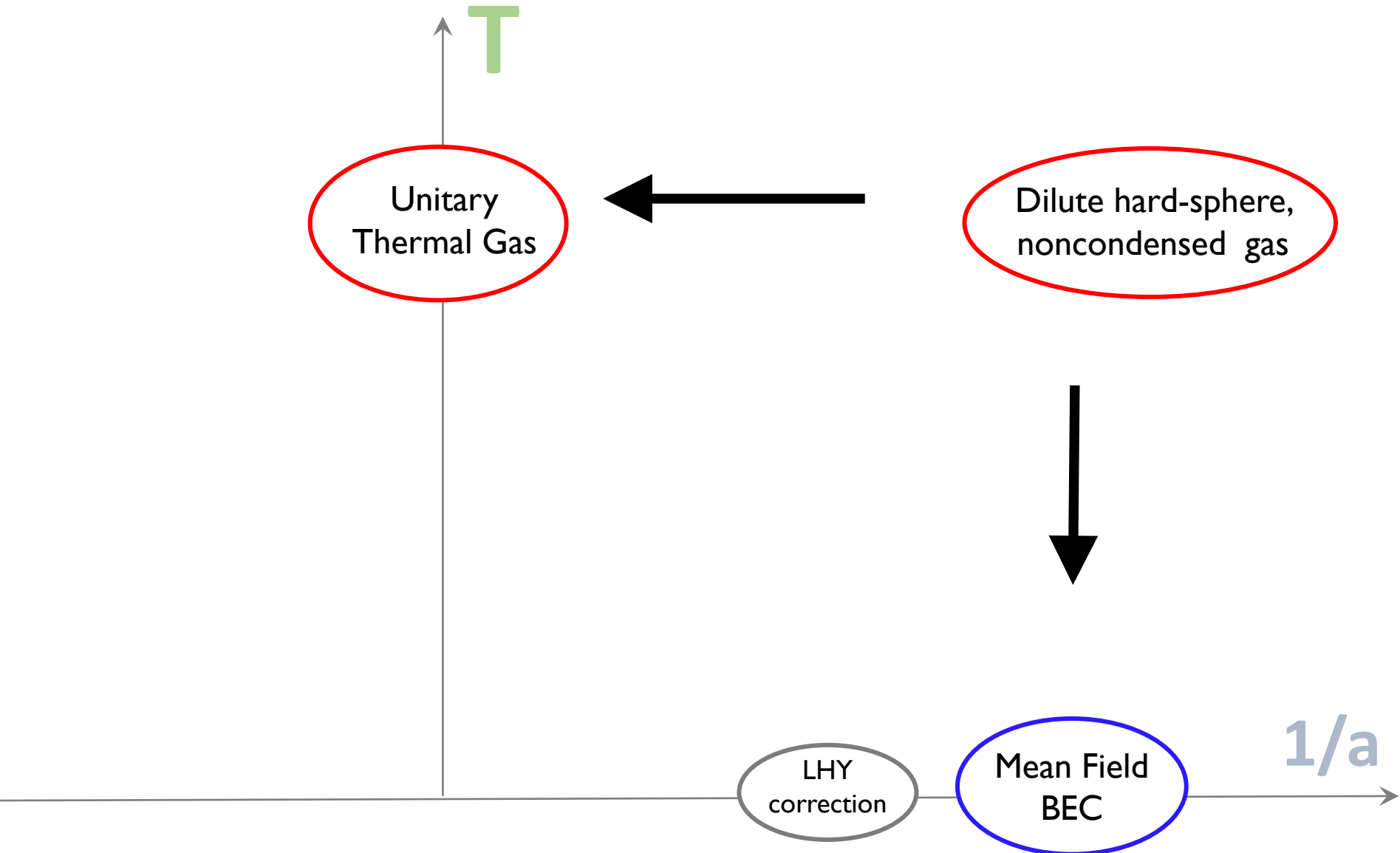
Ramsey Fringe



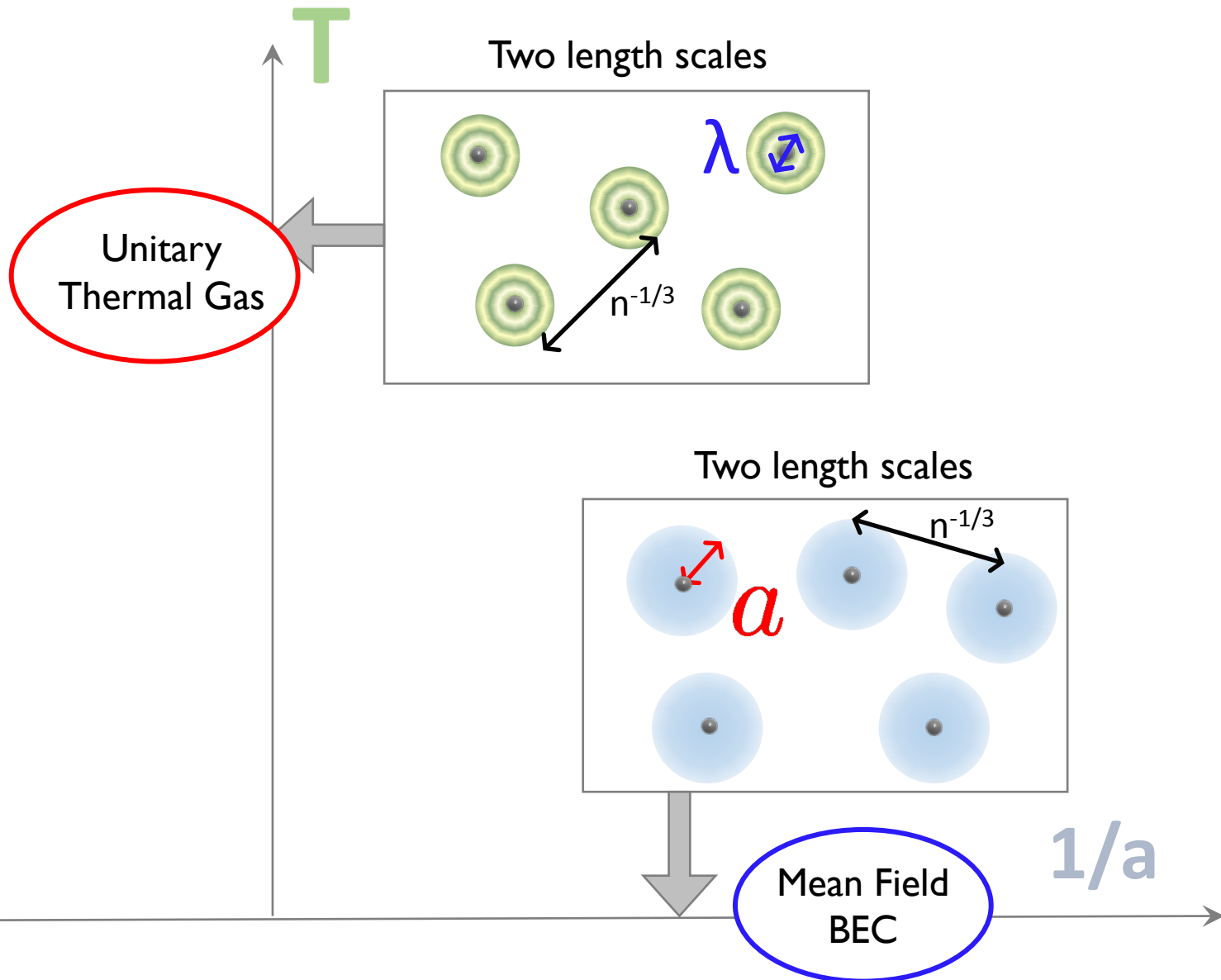
Cold Bosons: The Regimes



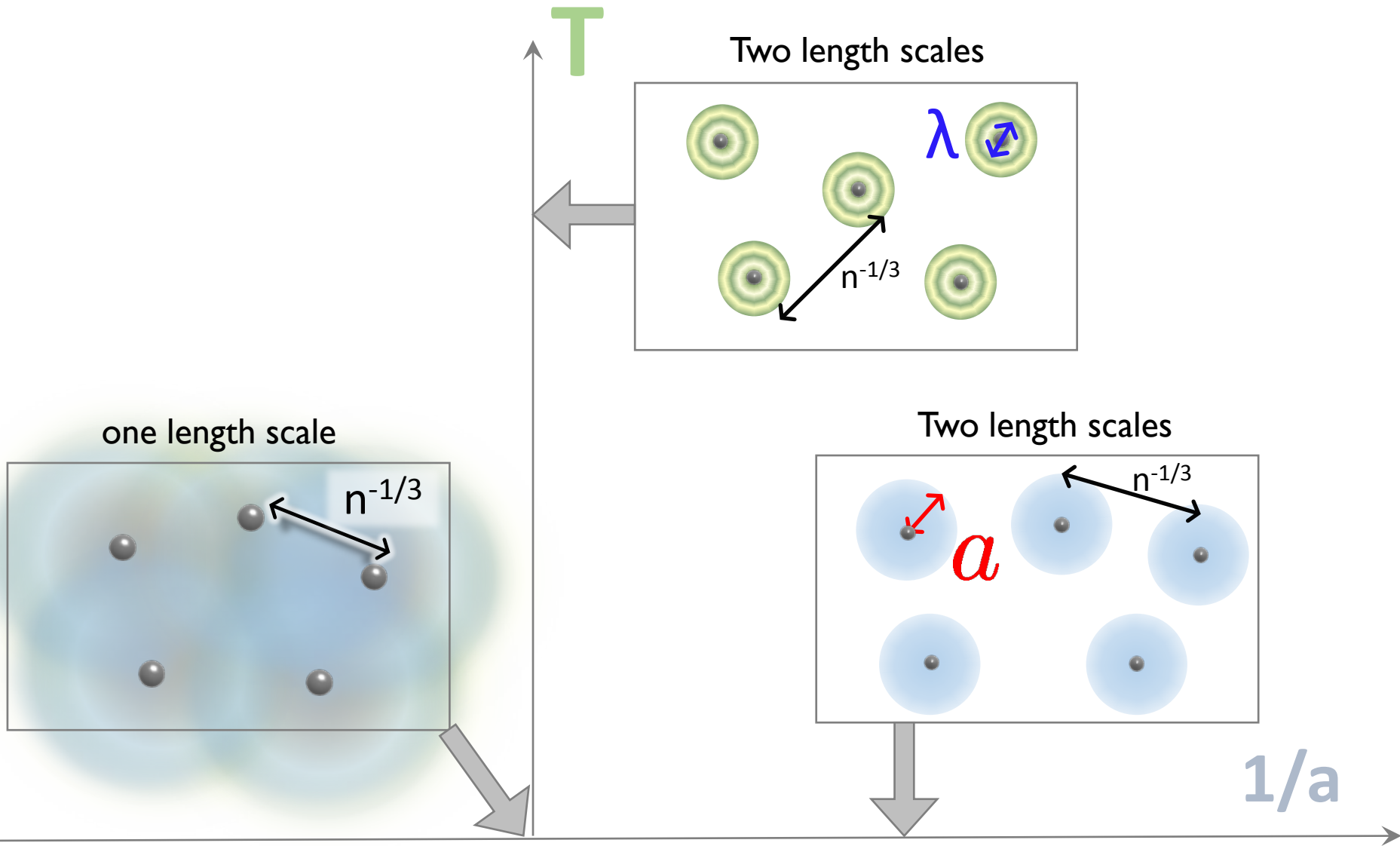
Cold Bosons: The Regimes



Degenerate Unitarity



Degenerate Unitarity: Why is it interesting?



Three boson inelastic loss: The Regimes

T

confirmed by
Paris group!

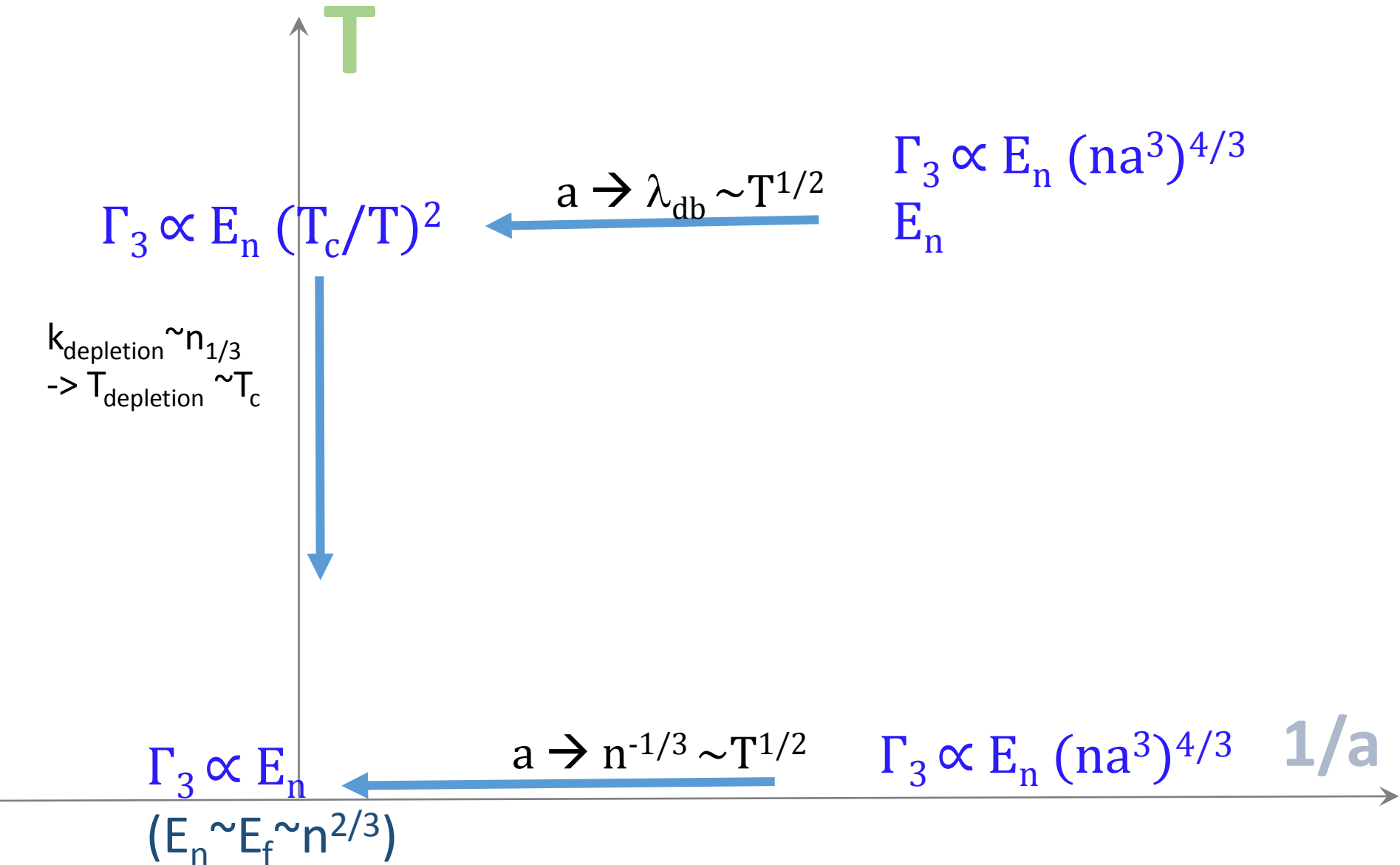
$$\Gamma_3 \propto E_n (T_c/T)^2$$

$a \rightarrow \lambda_{db} \sim T^{1/2}$

$$\Gamma_3 \propto E_n (na^3)^{4/3}$$

$$\Gamma_3 \propto E_n (na^3)^{4/3} \quad \mathbf{1/a}$$

Three boson inelastic loss: The Regimes



Three boson inelastic loss: The Regimes

T



$$\Gamma_3 \propto n^2 a^4$$

$$\Gamma_3 \propto n^{2/3} n^{4/3} a^4$$

$$\Gamma_3 \propto E_n (na^3)^{4/3}$$

1/a