# The Mu2e experiment at Fermilab: design and status



#### Raffaella Donghia LNF-INFN and Roma Tre university On behalf of the Mu2e collaboration

Les Rencontres de Physique de la Vallée d'Aoste La Thuile, March 5-11, 2017







# The Mu2e Collaboration



#### ~230 Scientists from 37 Institutions

Argonne National Laboratory, Boston University, Brookhaven National Laboratory, University of California Berkeley, University of California Irvine, California Institute of Technology, City University of New York, Joint Institute of Nuclear Research Dubna, Duke University, Fermi National Accelerator Laboratory, Laboratori Nazionali di Frascati, University of Houston, Helmholtz-Zentrum Dresden-Rossendorf, University of Illinois, INFN Genova, Lawrence Berkeley National Laboratory, INFN Lecce, University Marconi Rome, Institute for High Energy Physics Protvino, Kansas State University, Lewis University, University of Liverpool, University College London, University of Louisville, University of Manchester, University of Minnesota, Muons Inc., Northwestern University, Institute for Nuclear Research Moscow, Northern Illinois University, INFN Pisa, Purdue University, Novosibirsk State University/Budker Institute of Nuclear Physics, Rice University, University of South Alabama, University of Virginia, University of Washington, Yale University



# Talk overview

- Charged Lepton Flavor Violation (CLFV)
  - o BSM
  - o CLFV Muon sector
  - o History
- Muon Conversion
  - Measurement Technique
- Mu2e
  - o Goal
  - o Design
  - Detectors
- Summary





With neutrino mass, we know that lepton flavor is not conserved

The SM C strongly s

CLFV process would be  
suppressed:  
$$\mathcal{B}(\mu \to e\gamma) = \frac{3\alpha}{32\pi} \left| \sum_{i=2,3} U_{\mu i}^* U_{ei} \frac{\Delta m_{1i}^2}{M_W^2} \right|^2 < 10^{-54}$$

W

Any observation of CLFV would be new physics Beyond the Standard Model (BSM)!

CIFV

# Muon CLFV history





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e



Muon CLFV - BSM



also see Flavour physics of leptons and dipole moments, arXiv:0801.1826; Marciano, Mori, and Roney, Ann. Rev. Nucl. Sci. 58, doi:10.1146/annurev.nucl.58.110707.171126

# Probe mass scales $\,\lambda$ 2000~10000 TeV, significantly above the direct reach of LHC

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• Measure the ratio of  $\mu$  - e conversions to conventional muon captures

 $\mu$ -e conversion in the presence of a nucleus

$$R_{\mu e} = \frac{\mu^{-} + N(A, Z) \to e^{-} + N(A, Z)}{\mu^{-} + N(A, Z) \to \nu_{\mu} + N(A, Z - 1)}$$

Nuclear captures of muonic Al atoms

- And set an upper limit:
   R<sub>µe</sub> < 6 x 10<sup>-17</sup> (@ 90% CL, with ~ 10<sup>18</sup> stopped muons in 3 years of running)
- Discovery sensitivity: all  $R_{\mu e} > \text{few x } 10^{-16}$ Covers broad range of new physics theories



- 1. Generate low momentum  $\mu^{-}$  beam
- 2. Stop the muons in an AI target  $\rightarrow$  trapped in orbit around the nucleus
  - 3. Look for an excess around 104.97 MeV/c in the electron spectrum
- Main Backgrounds
- Muon decay in orbit (DIO)  $\mu^- + Al \rightarrow e^- + \overline{v_e} + v_\mu + Al$
- Radiative  $\mu/\pi$  capture  $\mu^- + Al \rightarrow \upsilon_{\mu} + \gamma + Mg$   $\pi^- N \rightarrow \gamma N^*, \gamma \rightarrow e^+ e^ \pi^- N \rightarrow e^+ e^- N^*$
- $\pi/\mu$  decay in flight
- Antiproton annihilation
- Electrons from beam, cosmic rays





# Mu2e design



#### **Production Solenoid / Target**

- Protons hitting target and producing mostly  $\pi$
- Solenoid reflects slow forward  $\mu/\pi$  and contains backward  $\mu/\pi$

#### **Transport Solenoid**

- Selects and transports low momentum  $\mu^{-}$ 



#### Detector Solenoid: stopping target and detectors

- Stops  $\mu^-$  on Al foils (decay time ~ 864 ns)
- Events reconstructed by detectors, optimized for 105 MeV momentum







- ~ 20000 straw drift tubes, divided in 18 stations, 2 planes/station
- Each straw is 5 mm diameter, with 25  $\mu m$  sense wire, 15  $\mu m$  thick mylar walls
- 3 m long, 1.4 m diameter, in a uniform 1 T magnetic field

#### Momentum resolution < 170 keV/c (@ 100 MeV/c) Timing resolution ~ 1 ns Spatial resolution ~ 100 µm

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Cosmic Rays are a major sources of background  $\rightarrow$  CRV required

- Composed of 4 layers of overlapping scintillators (a coincidence of 3 out of 4 is used)
- Placed around DS and part of TS area
- Required efficiency: 0.9999







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stopped and unstopped pions will arrive at the detector before observation window









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# Conclusion



- Mu2e is a discovery CLFV experiment, looking for NP BSM with high complementarity to other programs while increasing reach and diversification in model testing
- Mu2e will improve previous conversion experiment by 4 orders of magnitude and probe mass scales up to thousands of TeV
  - 8 years timeline for completion of first phase
- Mu2e has purchased its superconductors, will soon occupy its building
   Construction period 2017-2018
   Installation will begin in 2019
- Mu2e phase-2 being planned to increase (x 10) intensity and sensitivity!



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#### Probe SUSY through loops



If SUSY seen at LHC  $\rightarrow$  rate ~10^{-15}

Implies ~ 40-50 signal events with negligible background in Mu2e for many SUSY models.

#### SUSY GUT in an SO(10) framework



L. Calibbi et al., hep-ph/0605139

# Complementary with the LHC experiments while providing models' discrimination

Table 3: Comparison of various ratios of branching ratios in the LHT model (f = 1 TeV) and in the MSSM without [92,93] and with [96,97] significant Higgs contributions.

O(1)	7 1 (
Other	prediction

antino	ratio	LHT	MSSM (dipole)	MSSM (Higgs)	
C.Tare	${Br(\mu^-  ightarrow e^- e^+ e^-) \over Br(\mu  ightarrow e \gamma)}$	0.021	$\sim 6\cdot 10^{-3}$	$\sim 6\cdot 10^{-3}$	
$\begin{array}{c c} \hline \\ \hline $		0.04 0.4	$\sim 1\cdot 10^{-2}$	$\sim 1\cdot 10^{-2}$	~
S.Reck	$\frac{Br(\tau^-\!\!\rightarrow\!\!\mu^-\mu^+\mu^-)}{Br(\tau\!\rightarrow\!\!\mu\gamma)}$	0.04 0.4	$\sim 2\cdot 10^{-3}$	0.060.1	
uling, S h]	$\frac{Br(\tau^-{\rightarrow}e^-\mu^+\mu^-)}{Br(\tau{\rightarrow}e\gamma)}$	0.04 0.3	$\sim 2\cdot 10^{-3}$	0.020.04	• Relative rates
$\begin{array}{c c} \begin{array}{c} O \\ S \\ S \\ S \\ \end{array} \end{array} \left  \begin{array}{c} \frac{Br(\tau^- \rightarrow \mu^- e^+ e^-)}{Br(\tau \rightarrow \mu \gamma)} \end{array} \right $	0.04 0.3	$\sim 1\cdot 10^{-2}$	$\sim 1\cdot 10^{-2}$	are model dependent	
.J.Burc 454v2	$\frac{Br(\tau^-{\rightarrow}e^-e^+e^-)}{Br(\tau^-{\rightarrow}e^-\mu^+\mu^-)}$	0.82.0	$\sim 5$	0.30.5	
$ \begin{array}{c} \stackrel{\text{V.}}{} \\ \stackrel{\text{O}}{} \\ \stackrel{\text{O}}{} \\ \stackrel{\text{O}}{} \\ \stackrel{\text{O}}{} \\ \stackrel{\text{O}}{} \\ \frac{Br(\tau^- \rightarrow \mu^- \mu^+ \mu^-}{Br(\tau^- \rightarrow \mu^- e^+ e^-)} \end{array} $	$\frac{Br(\tau^-{\rightarrow}\mu^-\mu^+\mu^-)}{Br(\tau^-{\rightarrow}\mu^-e^+e^-)}$	0.71.6	$\sim 0.2$	510	<ul> <li>Measure ratios to pin-</li> </ul>
A.Blar arXiv:C	$\frac{R(\mu \mathrm{Ti} \rightarrow e \mathrm{Ti})}{Br(\mu \rightarrow e\gamma)}$	$10^{-3}\dots10^2$	$\sim 5\cdot 10^{-3}$	0.080.15	down theory details





## Points vs LHC



TABLE XII: LFV rates for points SPS 1a and SPS 1b in the CKM case and in the  $U_{e3} = 0$  PMNS case. The processes that are within reach of the future experiments (MEG, SuperKEKB) have been highlighted in boldface. Those within reach of post-LHC era planned/discussed experiments (PRISM/PRIME, Super Flavour factory) highlighted in italics.

	SPS	8 1a	SPS	8 1b	SP	S 2	$\mathbf{SP}$	S 3	Future
Process	CKM	$U_{e3} = 0$	CKM	$U_{e3}=0$	CKM	$U_{e3} = 0$	CKM	$U_{e3} = 0$	Sensitivity
$BR(\mu \rightarrow e \gamma)$	$3.2 \cdot 10^{-14}$	$3.8 \cdot 10^{-13}$	$4.0 \cdot 10^{-13}$	$1.2 \cdot 10^{-12}$	$1.3 \cdot 10^{-15}$	$8.6 \cdot 10^{-15}$	$1.4 \cdot 10^{-15}$	$1.2\cdot10^{-14}$	$O(10^{-14})$
$BR(\mu \rightarrow e e e)$	$2.3 \cdot 10^{-16}$	$2.7 \cdot 10^{-15}$	$2.9 \cdot 10^{-16}$	$8.6 \cdot 10^{-15}$	$9.4 \cdot 10^{-18}$	$6.2 \cdot 10^{-17}$	$1.0 \cdot 10^{-17}$	$8.9 \cdot 10^{-17}$	$O(10^{-14})$
$CR(\mu \rightarrow e \text{ in Ti})$	$2.0 \cdot 10^{-15}$	$2.4 \cdot 10^{-14}$	$2.6 \cdot 10^{-15}$	$7.6 \cdot 10^{-14}$	$1.0 \cdot 10^{-16}$	$6.7 \cdot 10^{-16}$	$1.0 \cdot 10^{-16}$	$8.4 \cdot 10^{-16}$	$O(10^{-18})$
$BR(\tau \rightarrow e \gamma)$	$2.3 \cdot 10^{-12}$	$6.0 \cdot 10^{-13}$	$3.5 \cdot 10^{-12}$	$1.7 \cdot 10^{-12}$	$1.4 \cdot 10^{-13}$	$4.8 \cdot 10^{-15}$	$1.2 \cdot 10^{-13}$	$4.1 \cdot 10^{-14}$	$O(10^{-8})$
$BR(\tau \rightarrow e e e)$	$2.7 \cdot 10^{-14}$	$7.1 \cdot 10^{-15}$	$4.2 \cdot 10^{-14}$	$2.0 \cdot 10^{-14}$	$1.7 \cdot 10^{-15}$	$5.7 \cdot 10^{-17}$	$1.5 \cdot 10^{-15}$	$4.9 \cdot 10^{-16}$	$O(10^{-8})$
$BR(\tau \rightarrow \mu \gamma)$	$5.0 \cdot 10^{-11}$	$1.1 \cdot 10^{-8}$	$7.3 \cdot 10^{-11}$	$1.3 \cdot 10^{-8}$	$2.9 \cdot 10^{-12}$	$7.8 \cdot 10^{-10}$	$2.7 \cdot 10^{-12}$	$6.0 \cdot 10^{-10}$	$O(10^{-9})$
$BR(\tau \rightarrow \mu  \mu  \mu)$	$1.6\cdot 10^{-13}$	$3.4 \cdot 10^{-11}$	$2.2\cdot 10^{-13}$	$3.9\cdot10^{-11}$	$8.9\cdot10^{-15}$	$2.4\cdot 10^{-12}$	$8.7 \cdot 10^{-15}$	$1.9\cdot 10^{-12}$	$\mathcal{O}(10^{-8})$

- These are SuSy benchmark points for which LHC has discovery sensitivity
- Some of these will be observable by MEG/Belle-2
- All of these will be observable by Mu2e





#### Leptoquarks

Presenza di leptoquarks alla scala del TeV potrebindurre processi CLFV con una costante di accoppiamento  $\lambda$ .

- Rosso: MEG-II
- Blu: Mu2e

• Raffaella Donghia - Roma Tre JC  $\sqrt{|Y_{\mu\ell}|^2 + |Y_{e\mu}|^2}$ 

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### Specific Example: Higgs triplet and LHT



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## A few more models



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- □ Search for tracking hits with time and azimuthal angle compatible with the calorimeter clusters ( $|\Delta T| < 50 \text{ ns}$ ) → simplification of pattern recognition
- Add search of an Helix passing through cluster and selected hits + use calorimeter time to calculate tracking Hit drift times.
  - → Reduce the wrong drift sign assignments i.e. smaller positive momentum tail







Process	TDR Yield	CD3 Yield	Comment	- Tracker digitization
DIO	$\textbf{0.199} \pm \textbf{0.092}$	$\textbf{0.19} \pm \textbf{0.08}$	Fixed by the procedure. Instead one can com-	performance
			for the fixed DIO back- ground: $2.6 \times 10^{-17}$ (TDR),	- New Csl calorimeter
	0.001		$2.69 \times 10^{-17}$ (CD3).	- Better cosmic
RMC	$0.000^{+0.004}_{-0.000}$	$\Leftarrow$ same	not updated	backaround estimate
Pion capture	$0.023\pm0.006$	0.030	assumes $10^{-10}$ extinction	(much higher statistics)
Muon DIF	< 0.003	<i>⇐</i> same	not updated	
Pion DIF	$0.001 \pm < 0.001$	<b>← same</b>	not updated	- Largest background in
Beam electrons	$0.003\pm0.001$	$< 5 \times 10^{-4}$	assumes 10 <sup>-10</sup> extinction	Mu2e:
Antiprotons	$\textbf{0.047} \pm \textbf{0.024}$	← same	consistent with TDR to 10%	.DIO, according to the
Cosmic rays	$0.082\pm0.018$	$\textbf{0.247} \pm \textbf{0.055}$	assumes CRV inefficiency 10 <sup>-4</sup>	TDR numbers .Cosmic, as we currently
Total	$0.36\pm0.10$	0.52±?		1 think

Assuming  $6 \times 10^{17}$  stopped muons in  $6 \times 10^7$  sec of beam tine

Less than 1 background event allows us to reach desired single event sensitivity



# Why muon conversion is unique?



Muon conversion is a unique probe for BSM:

- Broad discovery sensitivity across all models:
  - $\rightarrow$  Sensitivity to the same physics of MEG but with better mass reach
  - ightarrow Sensitivity to physics that MEG is not
  - → If MEG observes a signal, MU2E/COMET do it with improved statistics. Ratio of the BR allows to pin-down physics model
  - → If MEG does not observe a signal, MU2E/COMET have still a reach to do so
- Sensitivity to  $\lambda$  (mass scale) up to hundreds of TeV beyond any current existing accelerator





#### □ Use the fact that muonic atomic lifetime >> prompt background

Need a pulsed beam to wait for prompt background to reach acceptable levels → Fermilab provides the beam we need !

### OUT of time protons are also a problem. To keep associated background low we need proton extinction of 10<sup>-10</sup>: proton extinction (between pulses): # protons out of beam/# protons in pulse





#### The DIO background is the most difficult one

Electron energy distribution from the decay of bound muons is a (modified) Michel spectrum:

 → Presence of atomic nucleus and momentum transfer create a recoil tail with a fast falling slope close to the endpoint
 → To separate DIO

→ To separate DIO endpoint from CE line we need a high Resolution Spectrometer



Czarnecki et al., Phys. Rev. D 84, 013006 (2011) arXiv:1106.4756v2

# **Stopping Monitor**



Figure 7.18. Preliminary singles germanium spectrum from the AlCap experiment at PSI. When muons stop in aluminum, they capture on the nucleus 60% of the time. A fraction of the captures produce <sup>27</sup>Mg in the ground state, which has a half-life of 9.5 minutes. In the decay, an 844 keV gamma is produced 72% of the time.

- Need a high precise gamma detector (HpGe)
- Energy of gamma ray is unique to the detector
- Detecting the delayed gamma rays eliminate problems related to beam flash
- Proton beam structure is 0.5 s on followed by 0.8 s idle. Gamma spectrum wil be acquired during idle time.
- Hpge should view the target far from the source and beyond DS



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## MEG - Backgrounds

- Accidental coincidence of e<sup>+</sup> and γ:
- Proportional to  $I^2_{\mu}$  (while signal proportional to  $I_{\mu}$ )
- Compromise between high intensity and low background
- Radiative muon decay background
- Proportional to  $I_{\mu}$
- Note: e<sup>+</sup> and  $\gamma$  simultaneous as for signal





proton t



# Accelerator scheme and proton extinction

- Booster: batch of 4×10<sup>12</sup> protons every 1/15<sup>th</sup> second
- Booster "batch" is injected into the Recycler ring and re-bunched into 4 bunches
- These are extracted one at a time to the Delivery ring
- As a bunch circulates, protons are extracted to produce the desired beam structure →
   bunches of ~3x10<sup>7</sup> protons each, separated by 1.7 μs



#### Proton Extinction achieving 10-10 is hard; normally get 10-2 - 10-3

- Internal (momentum scraping) and bunch formation in Accumulator
- External: oscillating (AC) dipole

Calculations based on accelerator models That take into account collective effects Shows that this combination gets  $\sim 10^{-12}$ 





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# The Transport Solenoid

- The S shape eliminates photons and neutrons
- Curvature and collimators select momentum and sign of particles
- Negative gradient that accelerates particles from the PS through the DS

<sup>,</sup> Collimator

Antiproton absorber



Positive

# The Detector Solenoid



- Graded field "reflects" downstream a fraction of conversion electrons emitted upstream (isotropic process)
- For the sensitivity goal → ~ 6 x 10<sup>17</sup> stopped muons for 3 years run → 10<sup>10</sup> stopped muon/s (10 GHz)



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# Straw Tracker (2)

- Tracker is made of arrays of straw drift tubes
- 20000 tubes arranged in planes on stations, the tracker has 18 stations
- Tracking at high radius only ensures operability (beam flash produces a lot of low momentum particles + DIO background)



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# Straw tubes



- 5 mm diameter;
- wall thickness of 15 µm



- two layers of ~6 μm (25 gauge) Mylar, spiral wound, with a ~3 μm layer of adhesive between layers. The inner surface has 500 Å aluminum overlaid with 200 Å gold as the cathode layer. The outer surface has 500 Å of aluminum to act as additional electrostatic shielding and improve the leak rate The straws will be tensioned to 500 g
- sense wire is 25  $\mu m$  gold plated tungsten, centered in the straw

 $_{\odot}\,$  The wire will be tensioned to 80 g

- The drift gas is tentatively taken as 80:20 Argon:CO\_2 with an operating voltage of ~1500 V, with maximum drift time of ~50 ns and gain of ~3\*10<sup>4</sup>



<u>sl lest beam</u>



Orthogonal single crystal





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