

Neutron capture cross-section of ^{93}Zr

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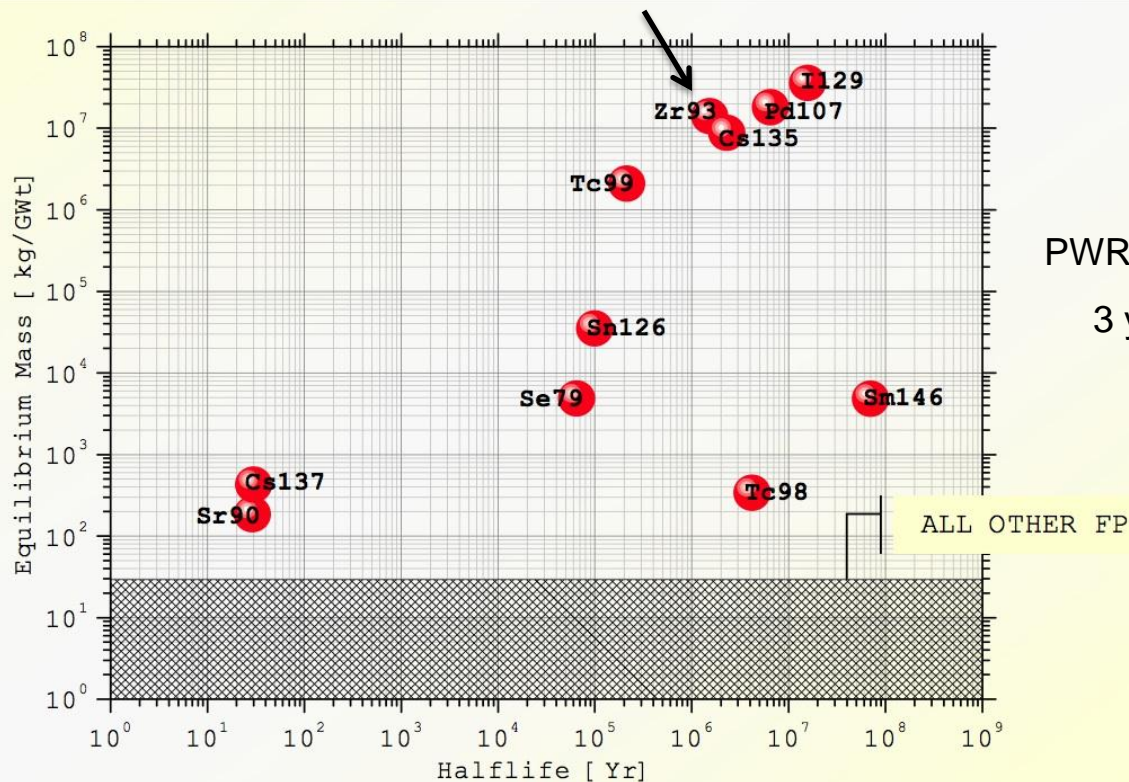
Technical coordinators: E. Berthomieux

Scientific Motivations

- Nuclear Technology
- Nuclear Astrophysics

Nuclear Technology

ACCUMULATION OF FISSION PRODUCTS



PWR 33GWtd THM⁻¹ burnup
3 yr cooling spent fuel

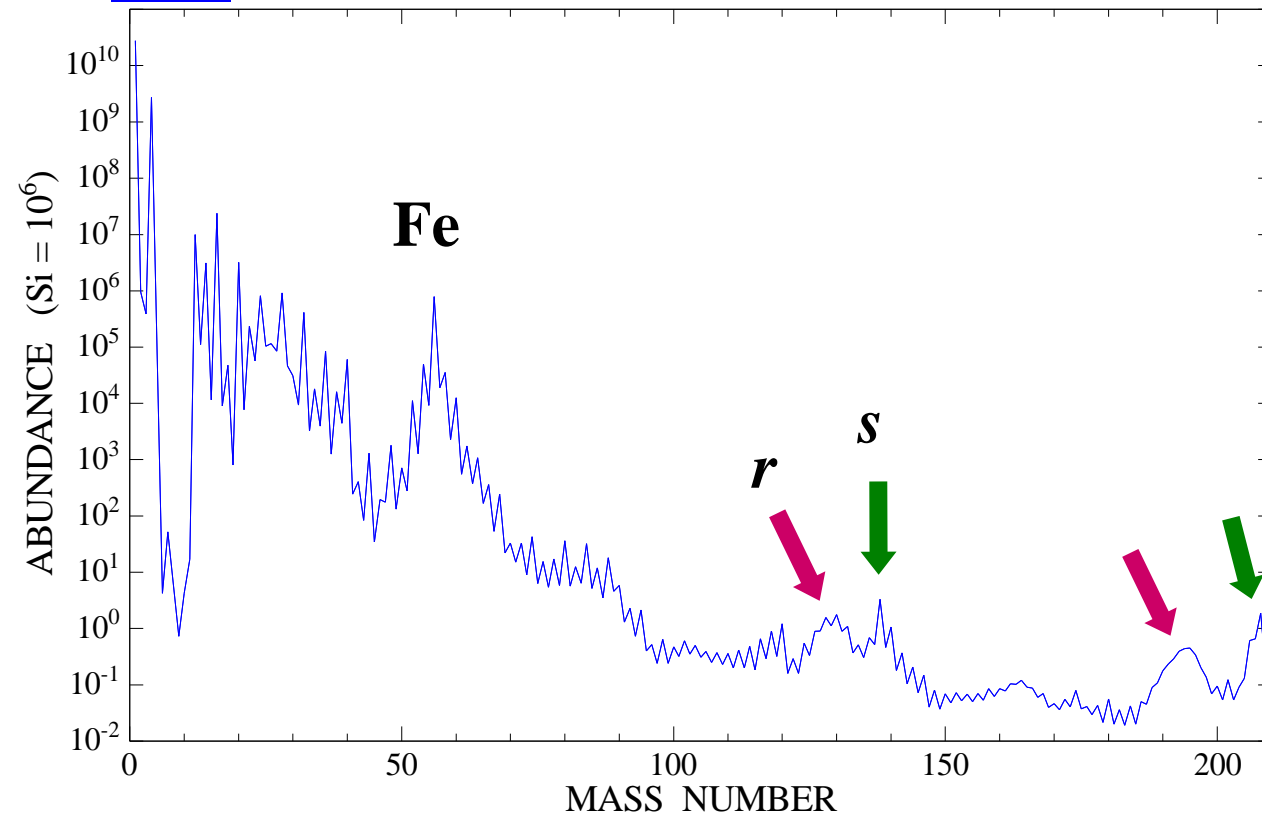
Nuclear Technology

- Partitioning and transmutation of nuclear wastes should make possible to reduce both the size of the repository for nuclear wastes and the long term risk. Fission products, such as ^{93}Zr , are considered as candidates for transmutation otherwise their very long half-lives necessitate storage in a repository for extremely long times.
- Neutron capture cross section of ^{93}Zr , as well as other long-lived fission products, is required for transmutation studies, as well as for nuclear reactor design purposes.
- Formal requests have been made for 5% uncertainties on capture cross sections of ^{93}Zr .

Nuclear Astrophysics

Abundances beyond Fe—ashes of stellar burning

BB **Fusion** **Neutrons**

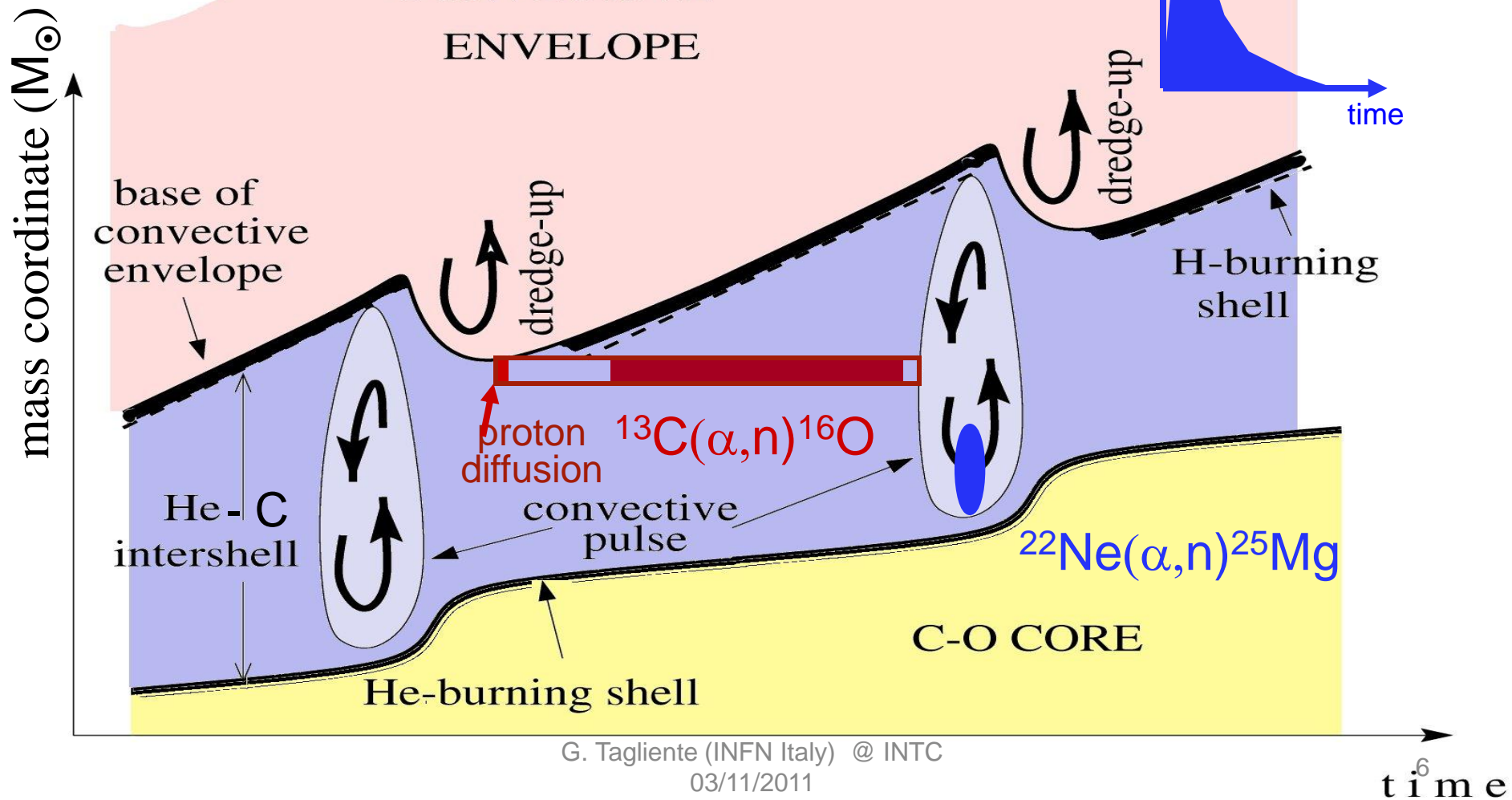


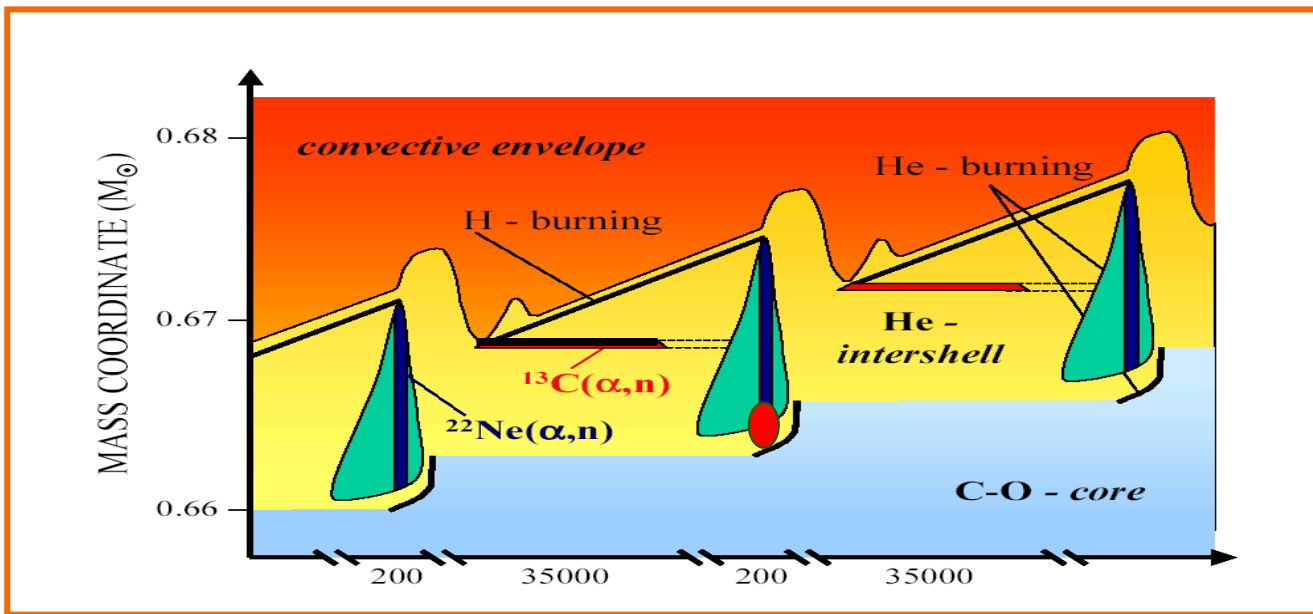
s-processes $n_n \sim 10^{10} \text{ n/cm}^3$
site: low mass AGB stars

r-processes $n_n \sim 10^{22} \text{ n/cm}^3$
site: supernovae explosions

AGB Stars

Thermal Pulse Stellar model





^{92}Mo 14.84	^{93}Mo 4.00 ka	^{94}Mo 9.25	^{95}Mo 15.82	^{96}Mo 16.68	^{97}Mo 9.55	^{98}Mo 24.13	^{99}Mo 2.75 d	^{100}Mo 9.63
^{91}Nb 680.04 a	^{92}Nb 34.70 Ma	^{93}Nb 100	^{94}Nb 20.30 ka	^{95}Nb 34.99 d	^{96}Nb 23.35 h	^{97}Nb 1.20 h	^{98}Nb 2.86 s	^{99}Nb 15.00 s
^{90}Zr 51.45	^{91}Zr 11.22	^{92}Zr 17.15	^{93}Zr 1.53 Ma	^{94}Zr 17.38	^{95}Zr 64.03 d	^{96}Zr 2.8	^{97}Zr 23.74 h	^{98}Zr 30.70 s
^{89}Y 100	^{90}Y 2.67 d	^{91}Y 58.51 d	^{92}Y 3.54 h	^{93}Y 10.18 h	^{94}Y 18.70 m	^{95}Y 10.30 m	^{96}Y 5.34 s	^{97}Y 3.75 s

G. Tagliente (INFN Italy) @ INTC
03/11/2011

The Thermal Pulse Stellar model: the Zr case

There is some inconsistency using the TP stellar model to calculate the N_s abundances with values of the Zr cross sections before n_TOF.

Nucleus	N_{\odot} Normalized to $N(\text{Si})=10^6$ atoms	N_s / N_{\odot}
^{90}Zr	5.546	0.789
^{91}Zr	1.21	1.066
^{92}Zr	1.848	1.052
^{94}Zr	1.873	1.217
^{96}Zr	0.302	0.842

The uncertainty on the N_{\odot} is 10%

The uncertainty on Zr cross sections

ranges from 5% to 20% (depending on the isotopes).

There are discrepancies up to 50% on the results of some measurements

New measurements with high accuracy needed !

Zr measurements:INTC 2002-034/P163

Zr isotope samples

Sample	Isotopic content (%)					
	^{90}Zr	^{91}Zr	^{92}Zr	^{93}Zr	^{94}Zr	^{96}Zr
^{90}Zr	97.7	0.87	0.6	-	0.67	0.16
^{91}Zr	5.43	89.9	2.68	-	1.75	0.24
^{92}Zr	4.65	1.62	91.4	-	2.03	0.3
$^{93}\text{Zr}^*$	1.5	19.0	20.0	20.0	20.0	19.0
^{94}Zr	4.05	1.18	1.93	-	91.8	1.04
^{96}Zr	19.41	5.21	8.2	-	8.68	58.5

Admixture: Hf, Na, Mg, Al ...

* Radio isotope ($T_{1/2} = 1.5 \times 10^6$ year)

Zr measurements:INTC 2002-034/P163

	^{90}Zr	^{91}Zr	^{92}Zr	^{93}Zr	^{94}Zr	^{96}Zr	^{197}Au	Pb
Mass (g)	2.717	1.404	1.349	4.88	2.015	3.398	1.871	3.895
Thickness (cm)	0,127	0,065	0,062	0,37	0,091	0,151	0.025	0.09
Chemical form	ZrO ₂	ZrO ₂	ZrO ₂	ZrO ₂	ZrO ₂	ZrO ₂	Metal	Metal
Enrichment (%)	97.7	89.9	91.4	20.0	91.8	58.5	100	Nat.

Samples 2.2 cm in diameter, 1 mm thick
Stable Zr isotopes encapsulated in 0.2 mm Al can
 ^{93}Zr isotope encapsulated in 0.2 mm Al + 0.2 mm Ti



Chemical form:ZrO₂

^{93}Zr isotope activity 92.5 MBq

Zr (INTC 2002-034/P163) publications

Neutron capture cross section of ^{90}Zr : bottleneck in the s-process reaction flow. G. Tagliente et al., PRC 77(2008)

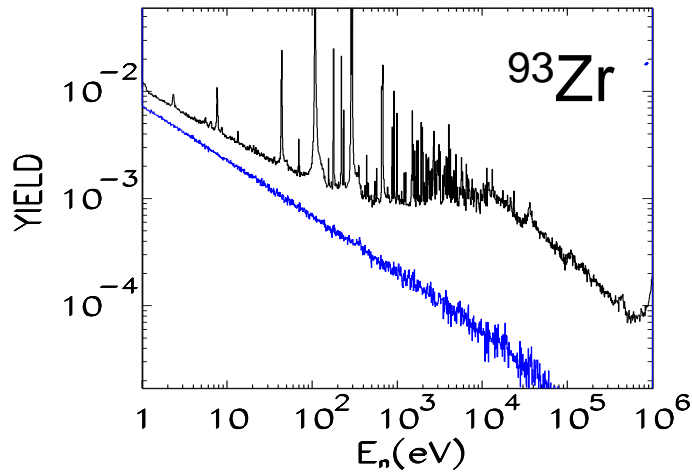
Study of the $^{91}\text{Zr}(n, \gamma)$ reaction up to 26 keV: G. Tagliente et al., PRC 78(2008)

The $^{92}\text{Zr}(n, \gamma)$ reaction and its implications on stellar nucleosynthesis: G. Tagliente et al., PRC 81(2010)

Neutron capture on ^{94}Zr : Resonance parameters and Maxwellian-averaged cross sections: G. Tagliente et al., PRC 84(2011)

$^{96}\text{Zr}(n, \gamma)$ measurement at the n_TOF facility at CERN: G. Tagliente et al., PRC accepted

Results - ^{93}Zr yield

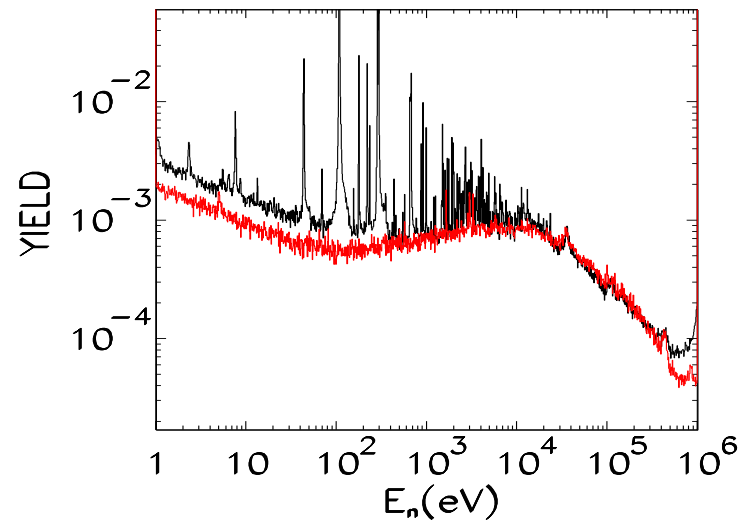


raw Yield

Natural radioactivity of the sample

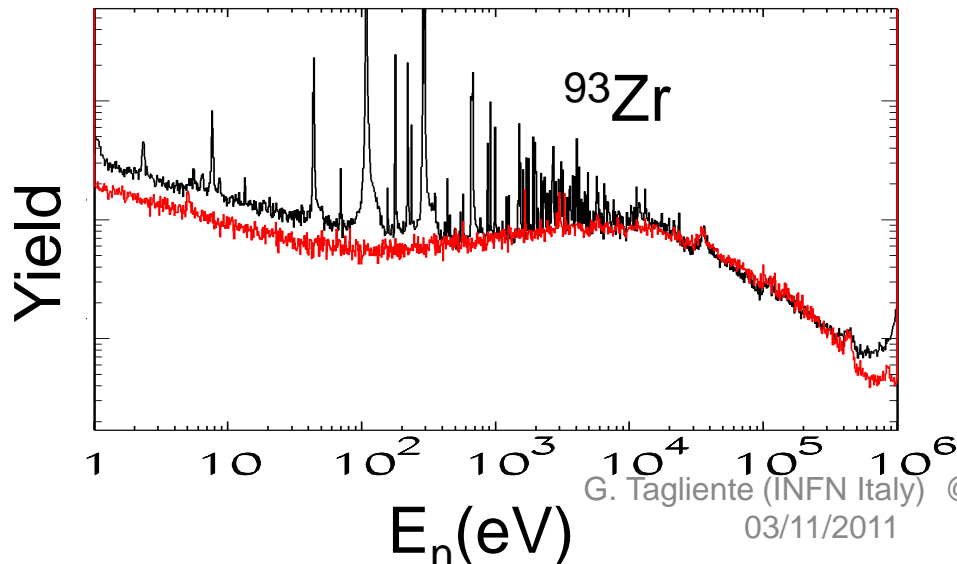
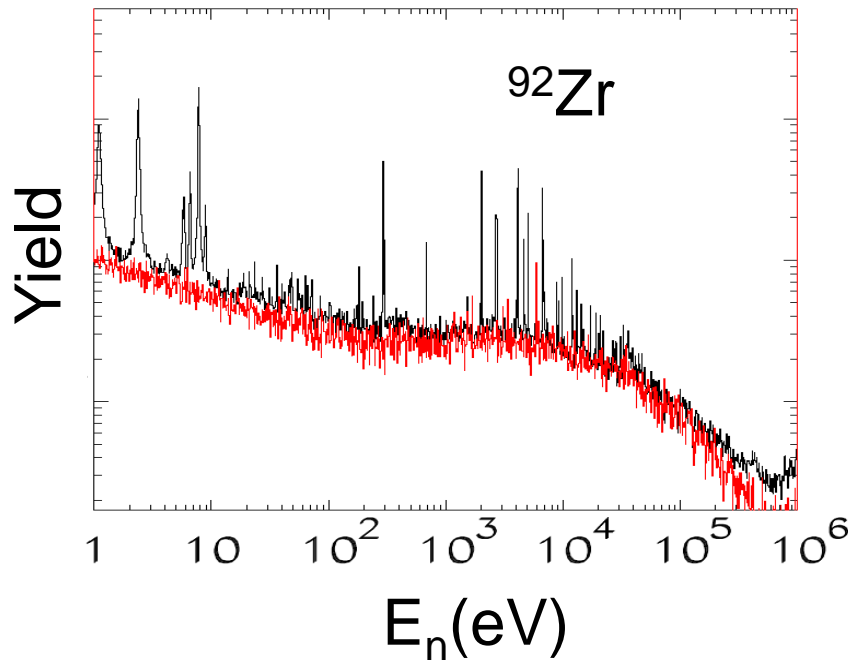
Yield – nat. radioactivity of the
sample

overall background



Results : MACS & experimental energy ranges

Maxwellian Averaged Cross Sections (MACS)

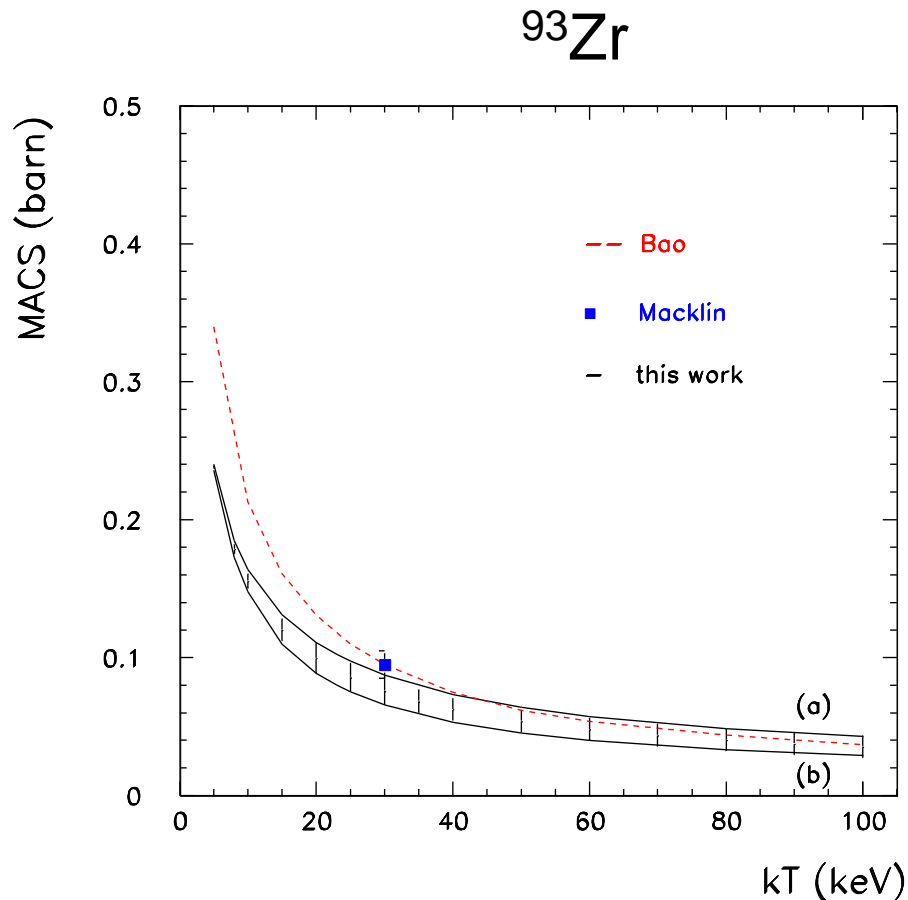


	Energy range (keV) n_TOF
^{90}Zr	0.01 - 66
^{91}Zr	0.01 - 26
^{92}Zr	0.01 - 40
^{93}Zr	0.01 - 8
^{94}Zr	0.01 - 60
^{96}Zr	0.01 - 37

The n_TOF data have to be complemented with a data library Jendl 3.3 and ENDF IV

MACS: results

The ^{93}Zr cross sections measured at n_TOF are 35% lower than the previous measurements

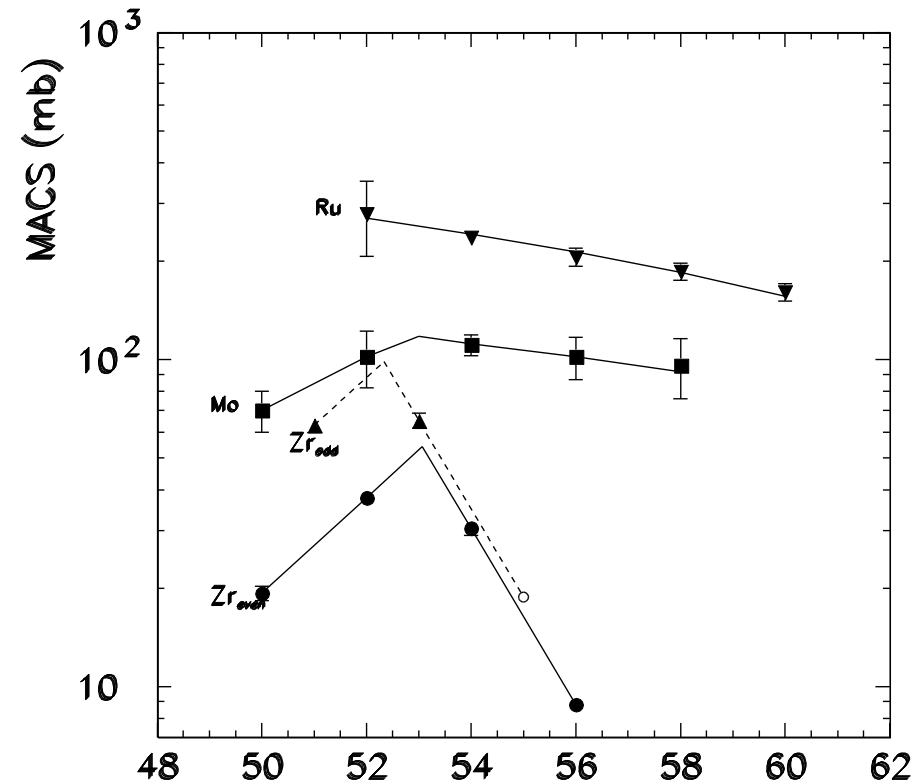


Preliminary

MACS: @ 30 keV

MACS in mbarn

Isotope	KADoNIS	n_TOF
^{90}Zr	21 ± 2	$19.3 \pm$
^{91}Zr	60 ± 8	$63^{0.9} \pm 4$
^{92}Zr	34 ± 6	38 ± 3
^{93}Zr	95 ± 10	$65.1 \pm$
^{94}Zr	26 ± 1	30.5 ± 2
^{95}Zr	79	18.9 ± 0.9
^{96}Zr	10.7 ± 0.5	8.9 ± 0.5



KADoNIS: Karlsruhe Astrophysical Database of Nucleosynthesis in Stars

Neutron Number

Astrophysical implication: Abundances

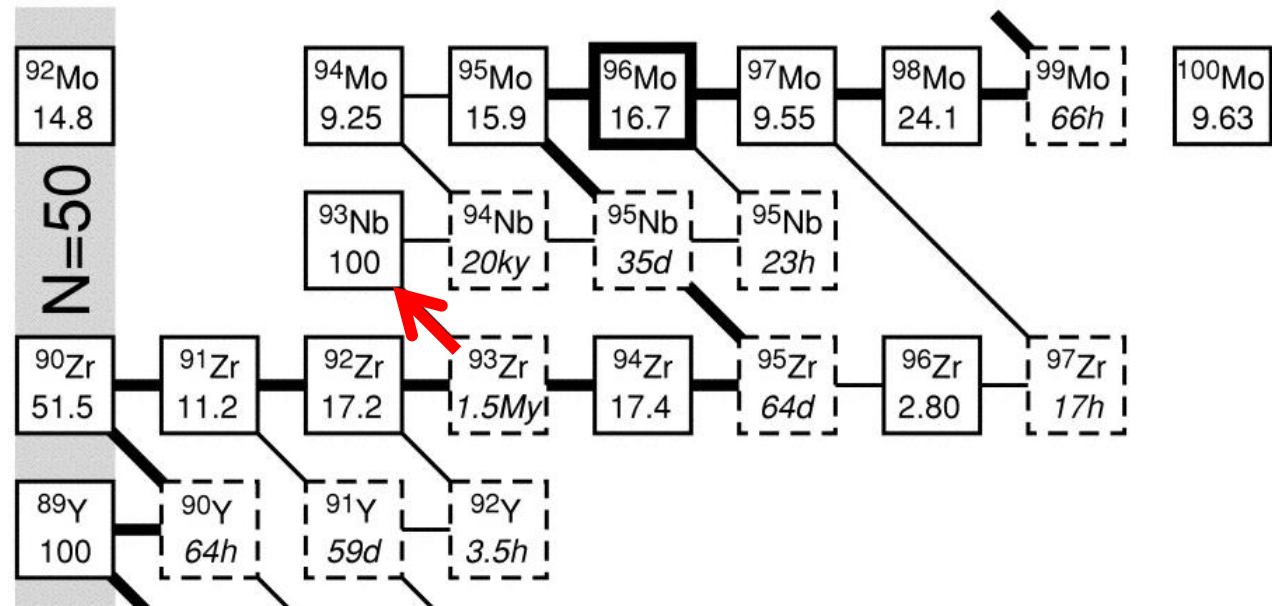
Nucleus	N_{\odot}	N_s / N_{\odot} %	N_s / N_{\odot} %
	Normalized to	Old MACS	n_TOF MACS
	$N(\text{Si})=10^6$ atoms		
^{90}Zr	5.546	0.789	0.844
^{91}Zr	1.21	1.066	1.024
^{92}Zr	1.848	1.052	0.981
^{94}Zr	1.873	1.217	1.152
^{96}Zr	0.302	0.842	0.321

Solar abundances, N_{\odot} , from Lodders 2009, accuracy 10%

The s-abundances, N_s , are calculated using the TP stellar model for low mass AGB star (1.5 - 3 M_{\odot}), accuracy 8%.

Old MACS are from the KADoNiS data base 2008. Since 2009 the databases has been update at the new n_TOF data, as the new data are relased.

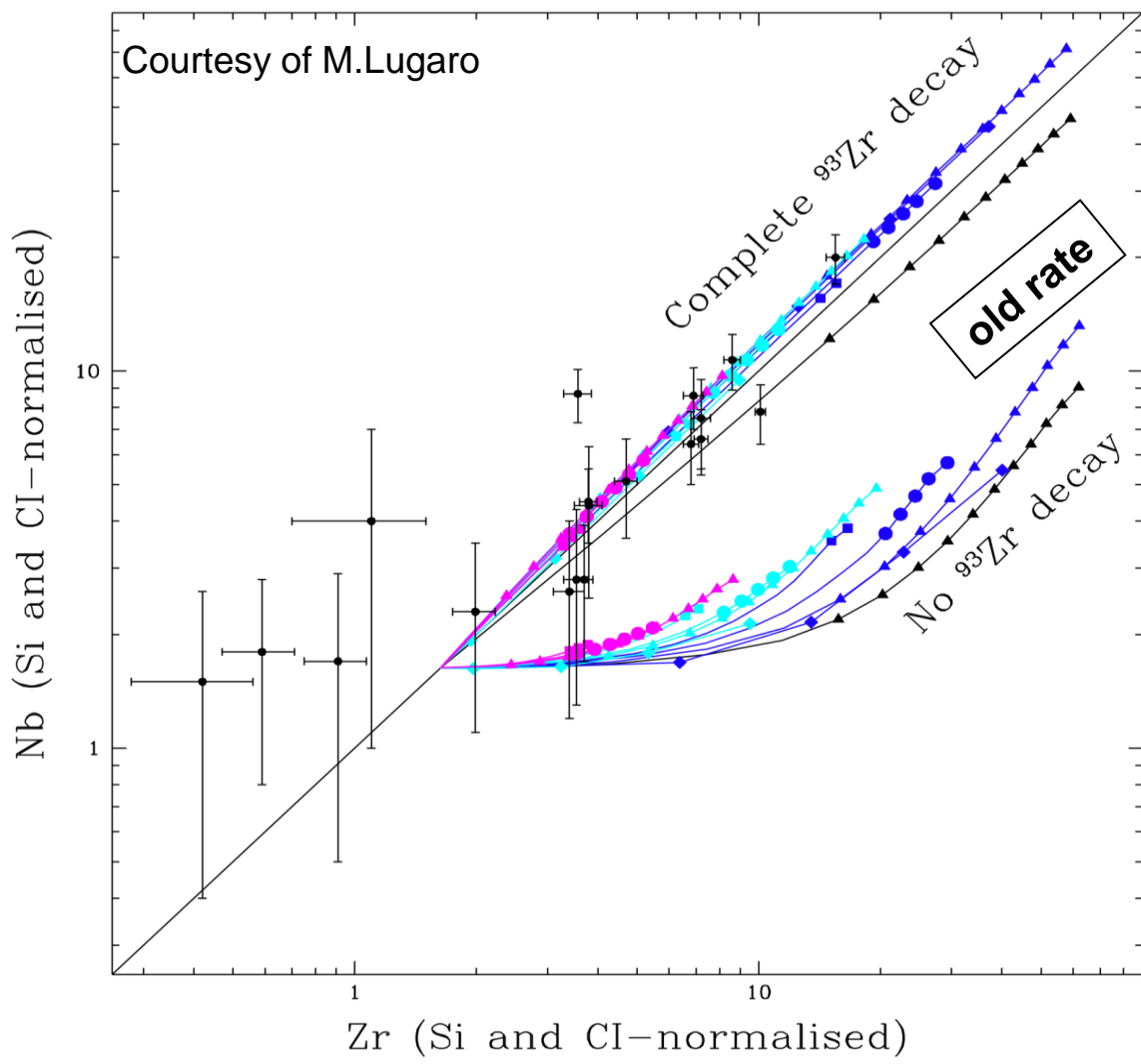
Astrophysical implication: Zr/Nb



A lower $^{93}\text{Zr}(n,\gamma)$ value means that more ^{93}Zr is produced. After **radiogenic decay** of ^{93}Zr more Nb will result.

The final result is **~50% more Nb!**

Elemental Nb and Zr abundances in SiC



- Mass, Metallicity
- ◇ 1.8, 0.01 ~ 0.7 Z_{\odot}
 - △ 3, 0.01 ~ 0.7 Z_{\odot}
 - 3, 0.02 ~ 1.5 Z_{\odot}
 - 3, 0.03 ~ 2 Z_{\odot}
- $M_{\text{mix}}(M_{\odot})$
- 0.002
 - 0.0005
 - 0.0002

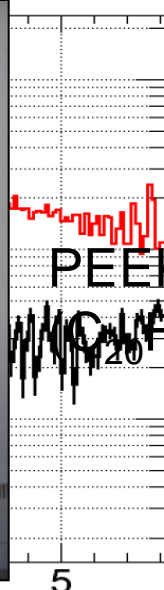
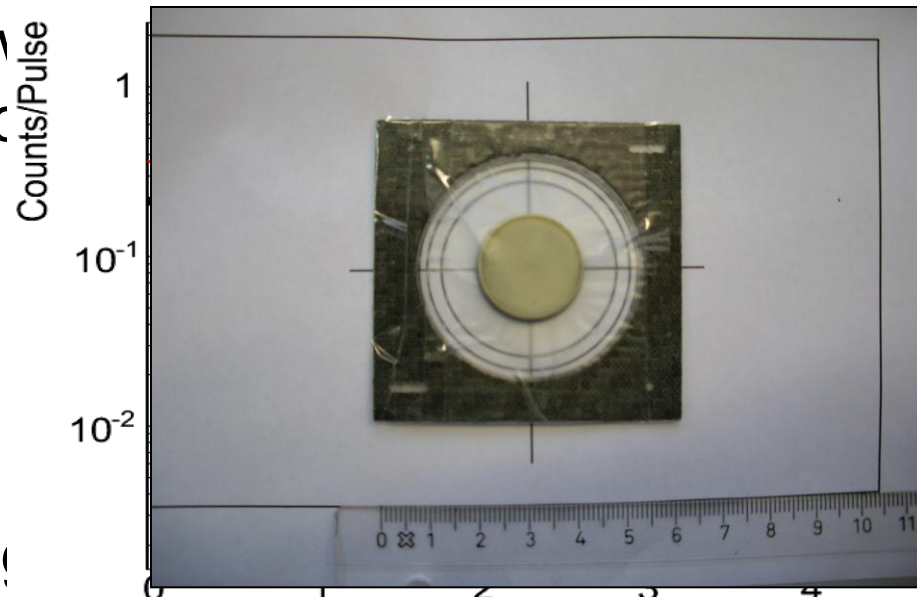
With the new $^{93}\text{Zr}(n,\gamma)$ cross section the problem is solved.

BUT!!!

The data, specially for the ^{93}Zr , have to be confirmed with new data in a wider neutron energy range

New ^{93}Zr capture measurement @ n_TOF

2009 new
lower background



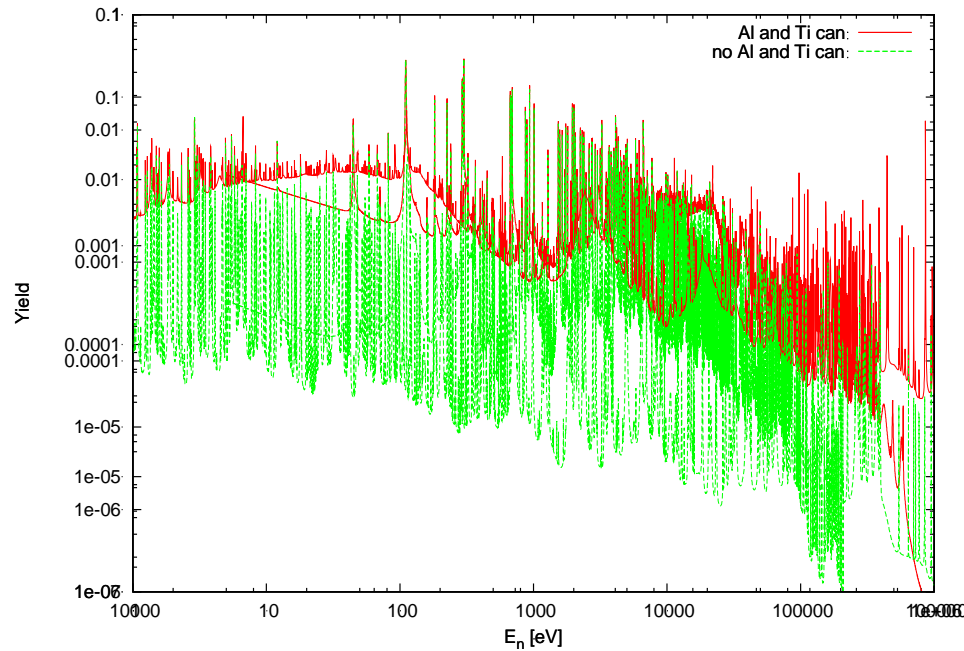
H_3BO_3):
PEEK container
(^{12}C , ^1H , ^{16}O)

2010 Upgrade
easier to

^{63}Ni measurement (CERN-INTC 2010-067/INTC-P-283)

of class A:

New ^{93}Zr capture measurement @ n_TOF



New ^{93}Zr capture measurement @ n_TOF

- Detectors: C_6D_6
- Samples
 - ^{93}Zr
 - Pb, empty ^{93}Zr container: background
 - Au, ^{92}Zr : normalization
- Goal: extend the measured cross section up to $E_n \approx 100$ keV with an accuracy $< 5\%$
- Number of protons required : 3×10^{18}