

# Measurement of <sup>7</sup>Be(n,α) and <sup>7</sup>Be(n,p) cross sections for the Cosmological Li problem in EAR2@n\_TOF

#### Request for a test beam at n\_TOF and sample preparation at ISOLDE

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## Outline

- Physics Case
- •<sup>7</sup>Be(n,p) measurement
- •<sup>7</sup>Be( $n,\alpha$ ) measurement

Step 2 (addendum to this proposal)

• <u>Test measurement and sample preparation</u> - Step 1(this proposal)

## **Big Bang Nucleosynthesis**

**Big Bang Nucleosynthesis (BBN)**, together with Hubble expansion and Cosmic Microwave Background Radiation is one of the cornerstones for Big Bang Theory.

BBN gives the sequence of nuclear reactions leading to the synthesis of light elements up to Na\* in the early stage of Universe (0.01-1000 sec)

At his first formulation, it depended on 3 parameters:

-the baryon-to-photon ratio η,
-the number of species of neutrino ν,
-the lifetime of neutron τ.

Nowadays **BBN is a parameter free theory**\*\*, being the **cross-sections** of reactions involved the only input to the theory.

\* A.Coc et al., The Astrophysical Journal, 744:158 (2012) \*\*D.N. Schramm and T.S Turner, Rev. Mod. Phys 70 (1998) 303



## The Cosmological Lithium Problem

Х

D/H

0.26

0.25

0.24

0.23

 $10^{-3}$ 

**BBN** successfully predicts the abundances of primordial elements such as <sup>4</sup>He, D and <sup>3</sup>He.

A serious discrepancy (factor 2-4) between the predicted abundance of <sup>7</sup>Li and the value inferred by measurements (Spite et al, many others.)



baryon-to-photon ratio  $\eta$ 

baryon density  $\Omega_{\rm b}h^2$  $10^{-2}$ 

\*

\* R.H.Cyburt et al., Journal of Cosmology and Astroparticle Physics 11 (2008) 012

\*\* A.Coc et al., The Astrophysical Journal, 744:158 (2012)

### Solution to CLiP?

Approximately 95% of primordial <sup>7</sup>Li is produced from the electron capture decay of <sup>7</sup>Be ( $T_{1/2}$ =53.2 d).

A higher destruction rate of <sup>7</sup>Be can solve or at least partially explain the CLiP.





# Status of evaluations and experimental data: <sup>7</sup>Be(n,p)

**Only one** direct measurement (Koehler et al., 1988, <u>0.025 eV - 13.5 keV</u>), covering partially the range of BBN interest (few keV-few hundreds of keV).



BBN calculations rely on **extrapolation to higher energies** of Koehler data or on **indirect measurements** 

### Status of evaluations and experimental data

**Only one** direct measurement (P. Bassi et al., 1963, @ <u>0.025 eV</u>) giving an **upper limit** for the cross section at thermal energy (<**0.1 mb**)



# The (n,p) measurement in EAR2 @n\_TOF



<sup>7</sup>Be(n,p)

Detection of protons (E~ 1 MeV or 1.4 MeV)

**Purification of sample needed: ISOLDE** 



#### Sample:

- 100 ng of <sup>7</sup>Be (material from water cooling of SINQ spallation source at **PSI**)
- Offline mass separation required at ISOLDE ≤1 day (starting from 100 GBq)

#### Implantation on backing





If successfully, the test will provide a new possibility to measure (n,p) and  $(n,\alpha)$  reactions.

## Conclusions

• Uncertainties in nuclear data strongly affect the Big Bang Nucleosynthesis calculations for the abundance of <sup>7</sup>Li and could probably explain (at least partially) the Cosmological Lithium **P**roblem.

• We plan to measure at **n\_TOF EAR2 for the first time** the <sup>7</sup>Be(n,p) and <sup>7</sup>Be(n, $\alpha$ ) cross sections in the whole range of interest for BBN, with the aim of reducing uncertainties in nuclear data used in calculations, thus setting stronger constraints to BBN theory and on CLiP.

#### • The needed <sup>7</sup>Be is **available at PSI**:

- For the (n,p) measurement, isotopic purity will be done at ISOLDE (100 ng)
   (3 shifts of offline ISOLDE mass separation no protons requested)
- For the  $(n,\alpha)$  measurement, no need for isotope separation (8 µg)
- 1.5x10<sup>18</sup> protons requested at n\_TOF EAR2 to test effect of direct neutrons on silicon detectors
  - check background induced on Si-detectors
  - <u>check degradation related to n and γ dose</u>
- Following to the results of the test, an addendum with the protons request for the (n,p) and  $(n,\alpha)$  measurements will be submitted.

Thanks for your attention

# Back-up slides

## <sup>7</sup>Be destruction



P D Serpico<sup>1</sup> Journal of Cosmology and Astroparticle Physics 12 (2004) 010

Figure B.7. Leading processes for production and destruction of <sup>7</sup>Be.

### **Background sources**



### **Energy deposition of 477 keV gamma**

 $\gamma$  477 keV - Be sample - 5  $\mu m$  AI backing - 5% energy resolution



# Energy loss of alpha from $n(^7Be,\alpha)\alpha$

 $\alpha$  8.9 MeV - Be sample - 5  $\mu m$  Al backing - 5% energy resolution



### **Expected counts**



### **Expected counts**

