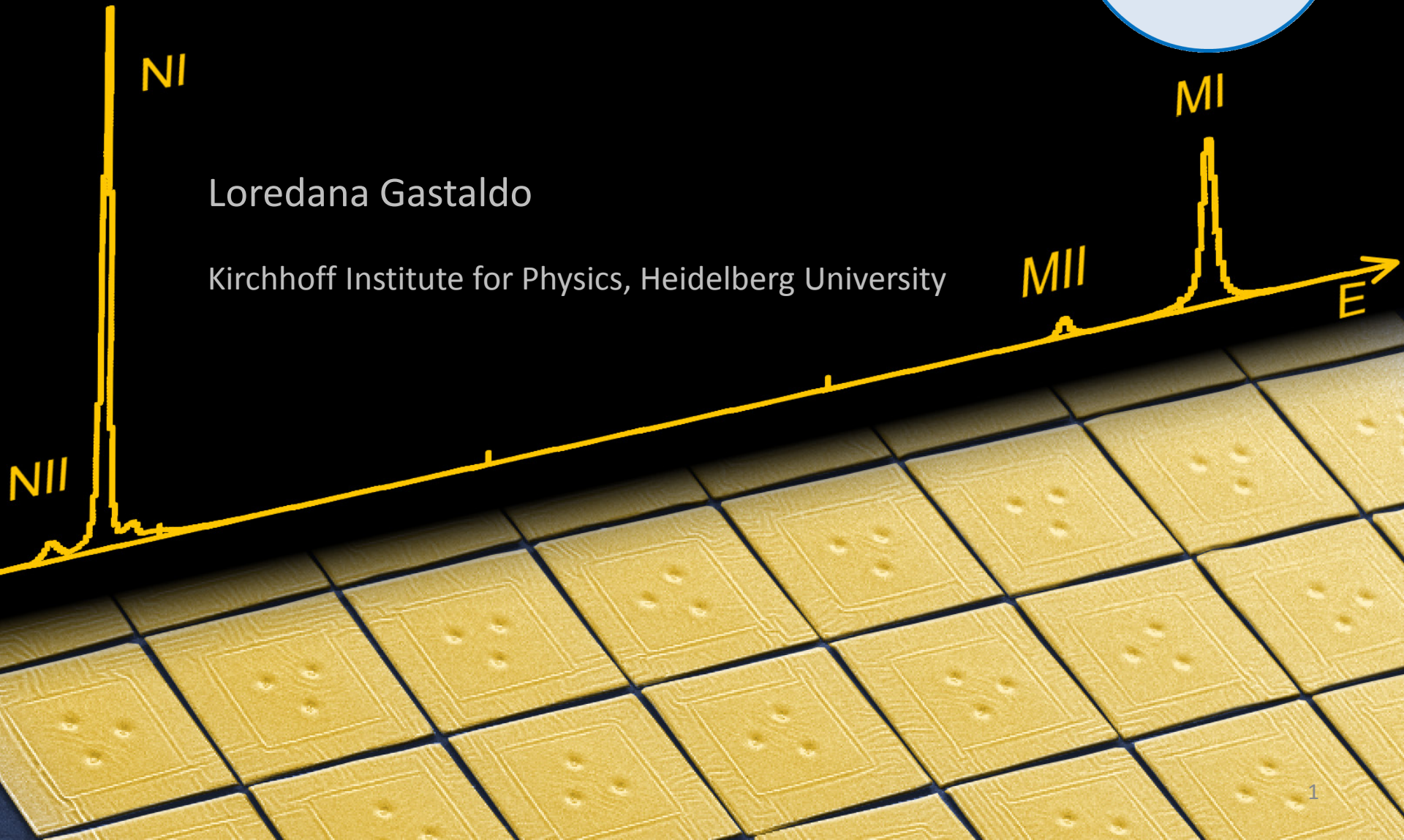


Electron Capture in ^{163}Ho experiment - ECHo



Loredana Gastaldo

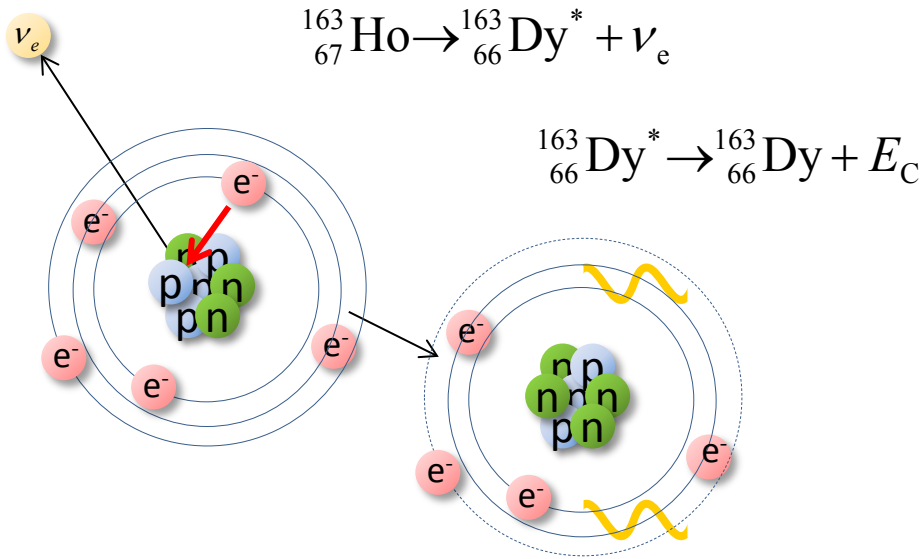
Kirchhoff Institute for Physics, Heidelberg University

Outline

- Electron capture in ^{163}Ho and neutrino mass
- Requirements to achieve sub-eV sensitivity on the electron neutrino mass
- The Electron Capture in ^{163}Ho experiment - ECHO
- Conclusions and outlook



Electron capture in ^{163}Ho



A non-zero neutrino mass affects the [de-excitation energy spectrum](#)

- $\tau_{1/2} \cong 4570$ years ($2 \cdot 10^{11}$ atoms for 1 Bq)

- $Q_{\text{EC}} = (2.833 \pm 0.030^{\text{stat}} \pm 0.015^{\text{syst}})$ keV

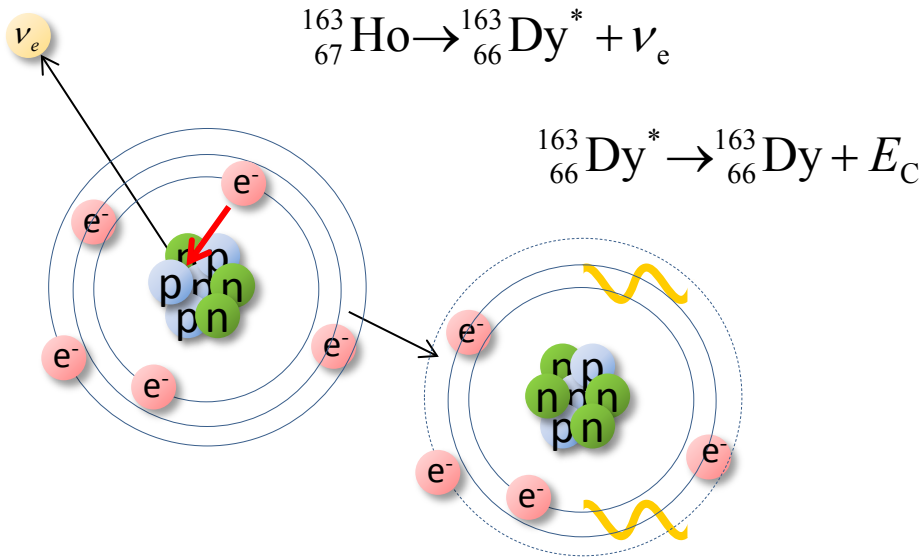
S. Eliseev et al., *Phys. Rev. Lett.*, 115, 062501 (2015)

- $Q_{\text{EC}} = (2.858 \pm 0.010^{\text{stat}} \pm 0.05^{\text{syst}})$ keV

P. C.-O. Ranitzsch et al., *Phys. Rev. Lett.*, 119, 122501 (2017)

- Low Q_{EC} -value allows capture only for:
 $3s, 3p_{1/2}, 4s, p_{1/2}, 5s, 5p_{1/2}, 6s$ electrons

Electron capture in ^{163}Ho



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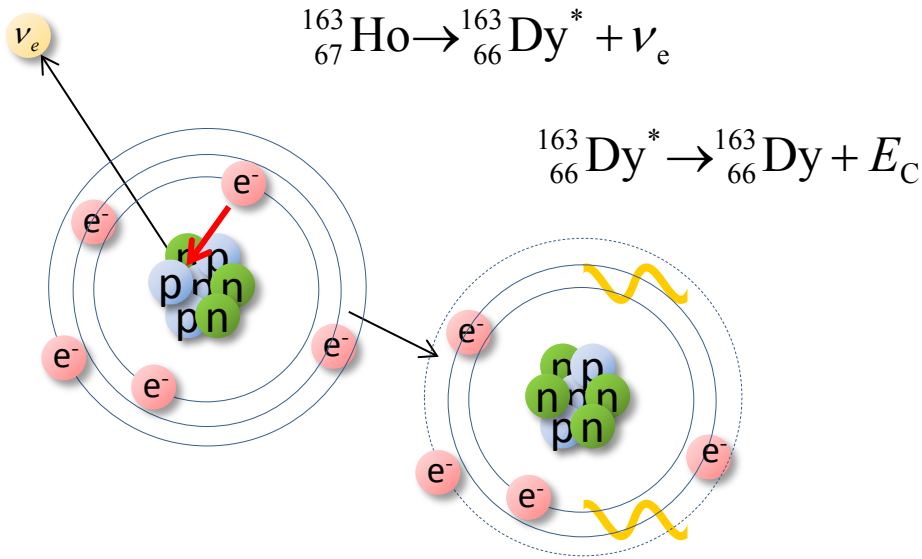
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Talk by Klaus Blaum on Wednesday

S. Eliseev et al., *Phys. Rev. Lett.*, 115, 062501 (2015)

P. C.-O. Ranitzsch et al., *Phys. Rev. Lett.*, 119, 122501 (2017)

Electron capture in ^{163}Ho



Atomic de-excitation:

- X-ray emission
- Auger electrons
- Coster-Kronig transitions

• $\tau_{1/2} \cong 4570$ years ($2 \cdot 10^{11}$ atoms for 1 Bq)

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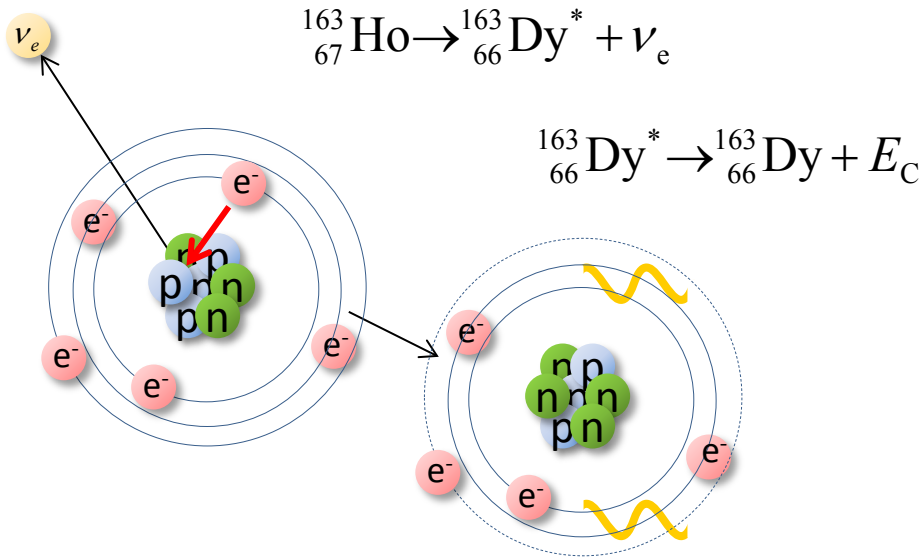
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Electron capture in ^{163}Ho

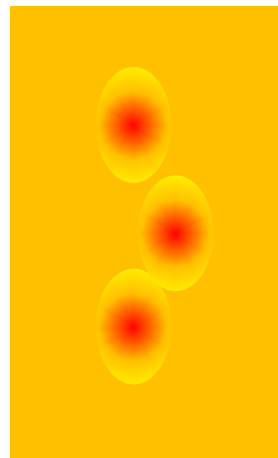


Atomic de-excitation:

- X-ray emission
- Auger electrons
- Coster-Kronig transitions

Calorimetric measurement

All the energy released in the electron capture process minus the one of the electron neutrino is measured by the detector



Source = Detector

ν_e

ν_e

ν_e

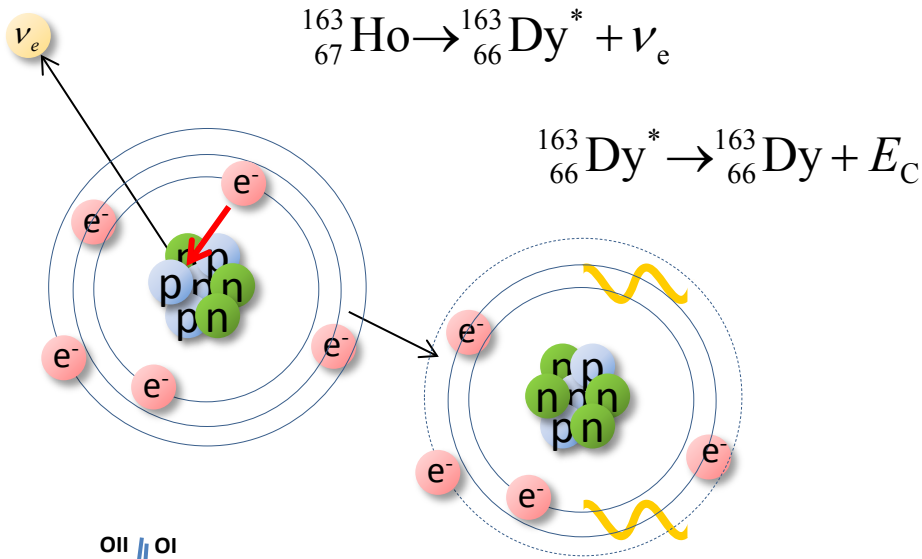


CALORIMETRIC MEASUREMENTS OF ^{163}Ho DECAY AS TOOLS TO DETERMINE THE ELECTRON NEUTRINO MASS

A. DE RÚJULA and M. LUSIGNOLI ¹
 CERN, Geneva, Switzerland

Phys. Lett. 118 B (1982) 118

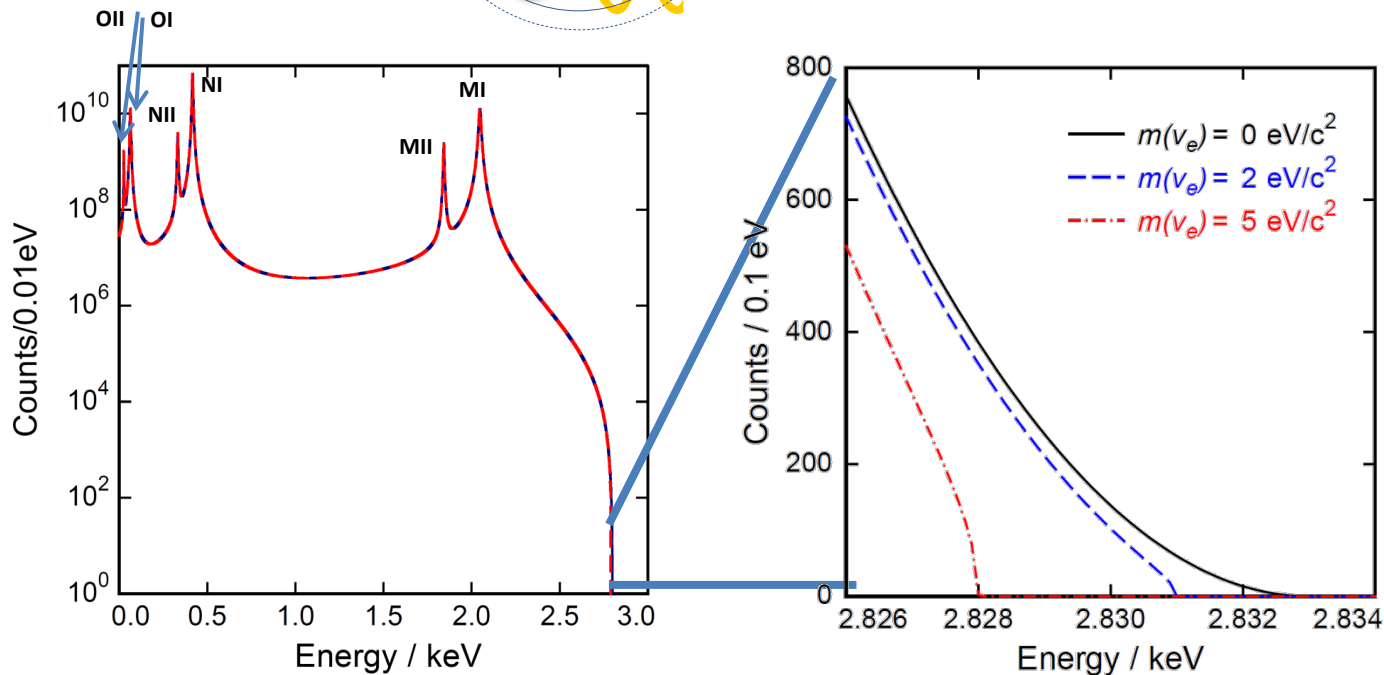
Electron capture in ^{163}Ho



Atomic de-excitation:

- X-ray emission
- Auger electrons
- Coster-Kronig transitions

Calorimetric measurement



Requirements for sub-eV sensitivity in ECHO

Statistics in the end point region

- $N_{\text{ev}} > 10^{14} \rightarrow A \approx 1 \text{ MBq}$

Unresolved pile-up ($f_{\text{pu}} \sim a \cdot \tau_r$)

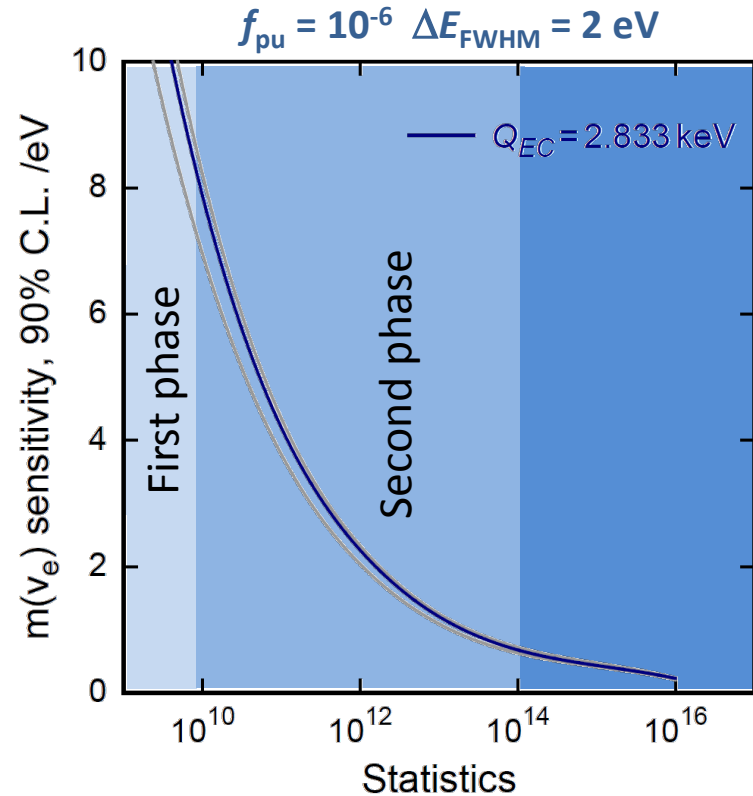
- $f_{\text{pu}} < 10^{-5}$
- $\tau_r < 1 \mu\text{s} \rightarrow a \sim 10 \text{ Bq}$
- 10^5 pixels \rightarrow multiplexing

Precision characterization of the endpoint region

- $\Delta E_{\text{FWHM}} < 3 \text{ eV}$

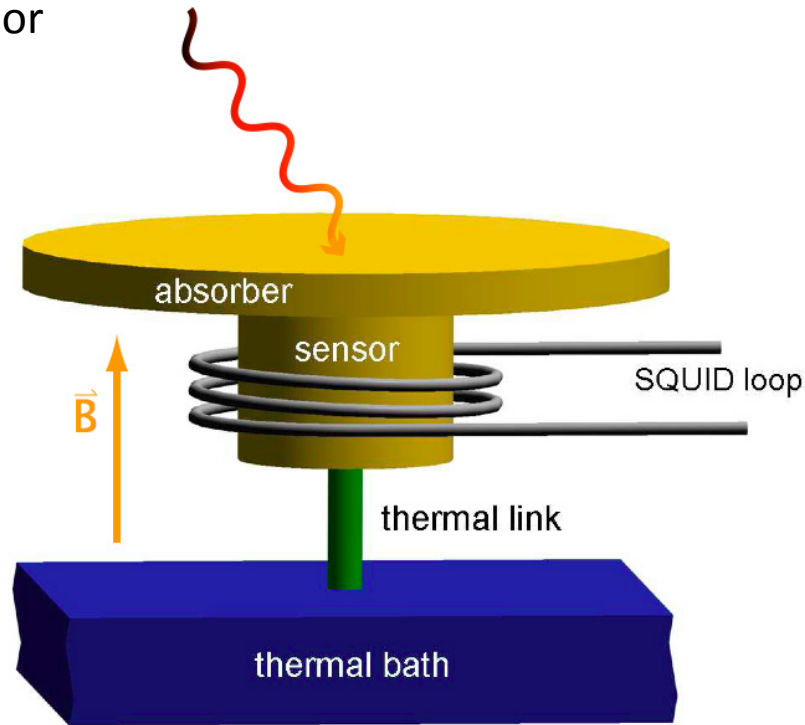
Background level

- $< 10^{-5} \text{ events/eV/det/day}$



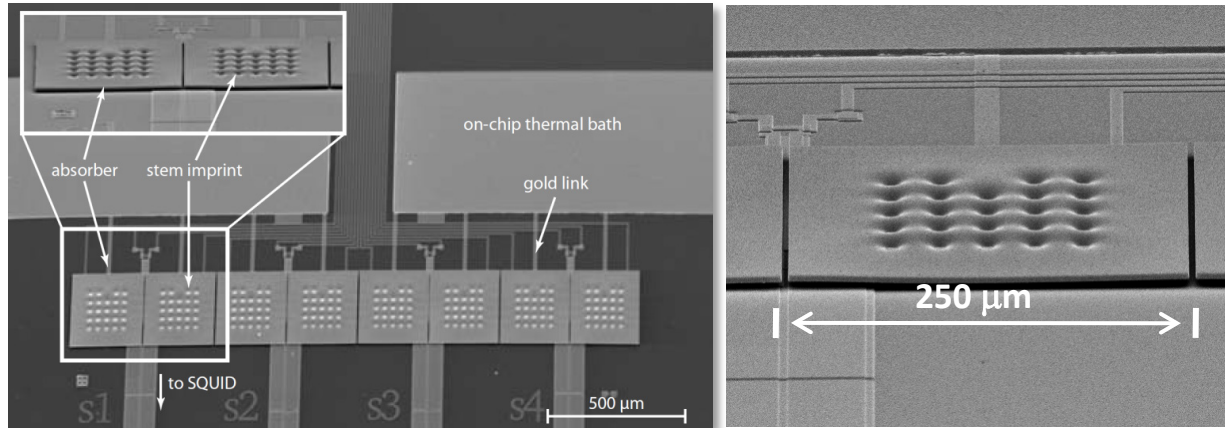
Metallic magnetic calorimeters (MMCs)

- Paramagnetic Ag:Er sensor



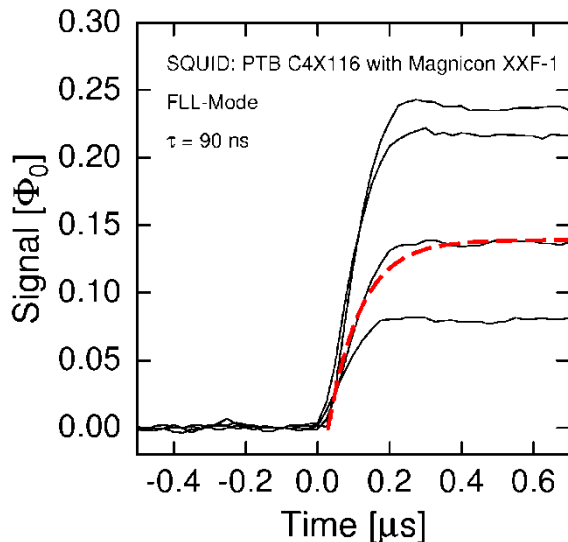
$$\Delta\Phi_s \propto \frac{\partial M}{\partial T} \Delta T \quad \rightarrow \quad \Delta\Phi_s \propto \frac{\partial M}{\partial T} \frac{E}{C_{\text{sens}} + C_{\text{abs}}}$$

MMCs: 1d-array for soft x-rays ($T=20$ mK)



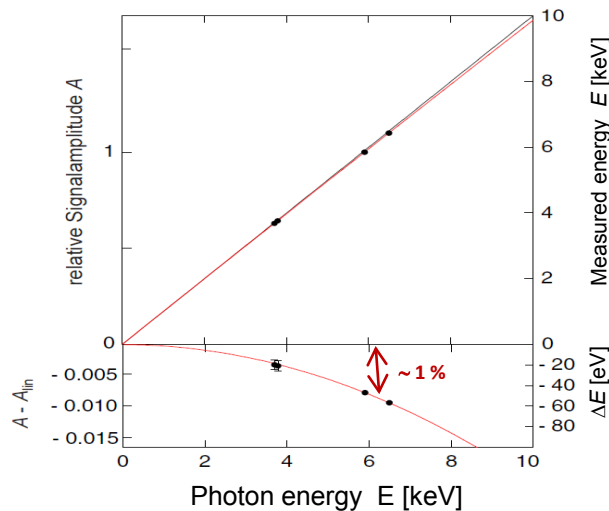
$$\Delta E_{FWHM} = 1.6 \text{ eV @ } 6 \text{ keV}$$

Rise Time: 90 ns

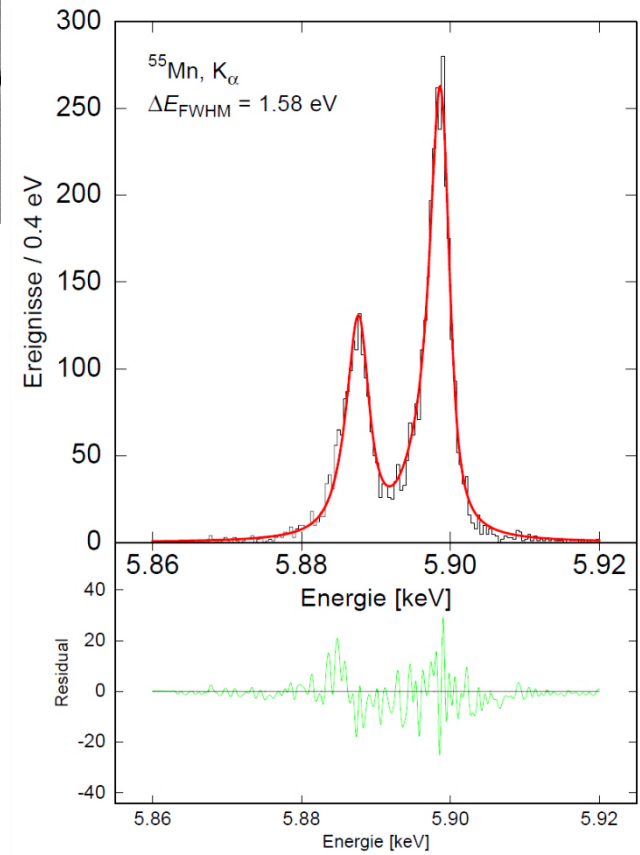


Reduction
un-resolved pile-up

Non-Linearity < 1% @6keV



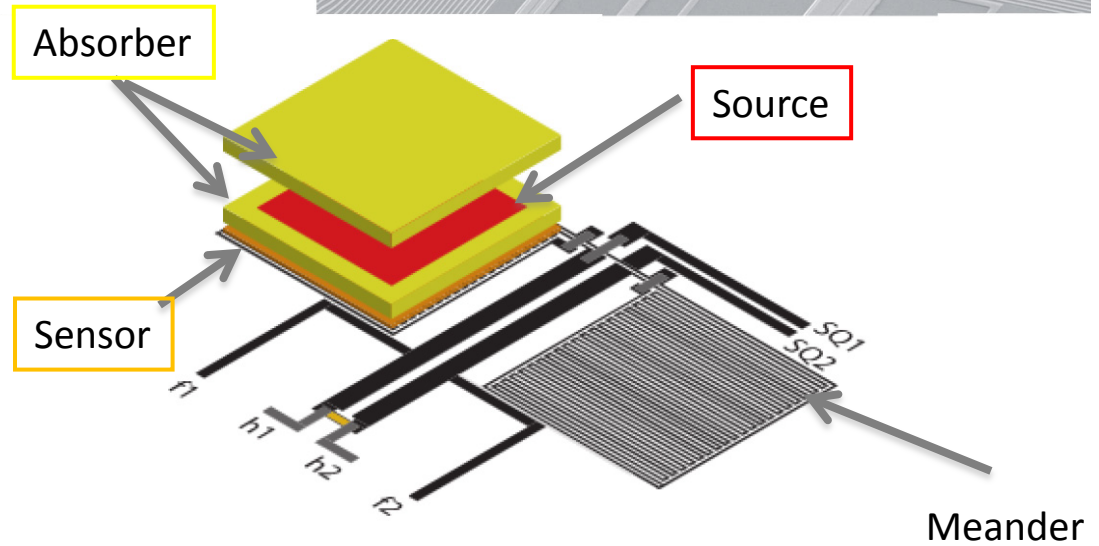
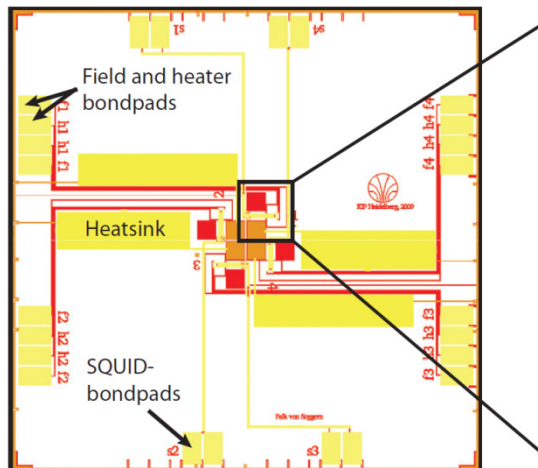
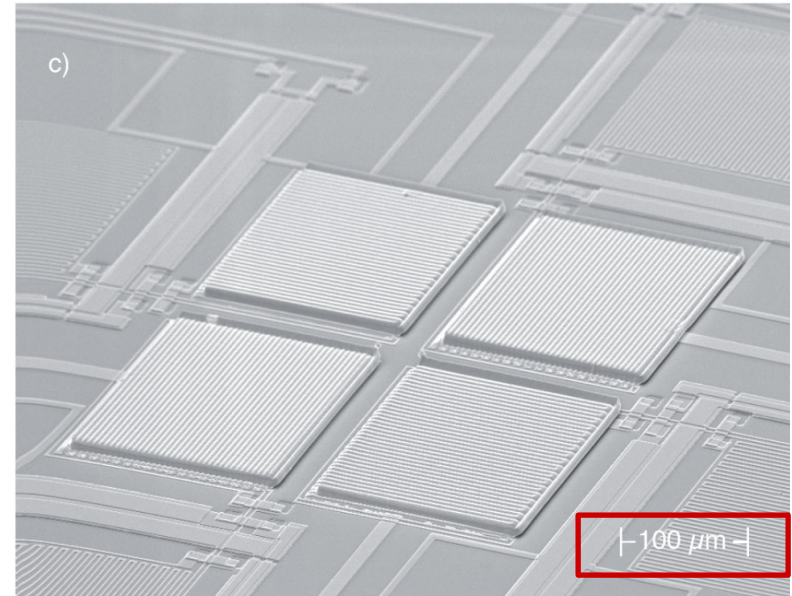
Definition
of the energy scale



Reduced smearing
in the end point region

First prototype of ^{163}Ho loaded MMC

- Absorber for metallic magnetic calorimeters
→ ion implantation @ ISOLDE-CERN in 2009
on-line process
- About 0.01 Bq per pixel
- Operated over more than 4 years

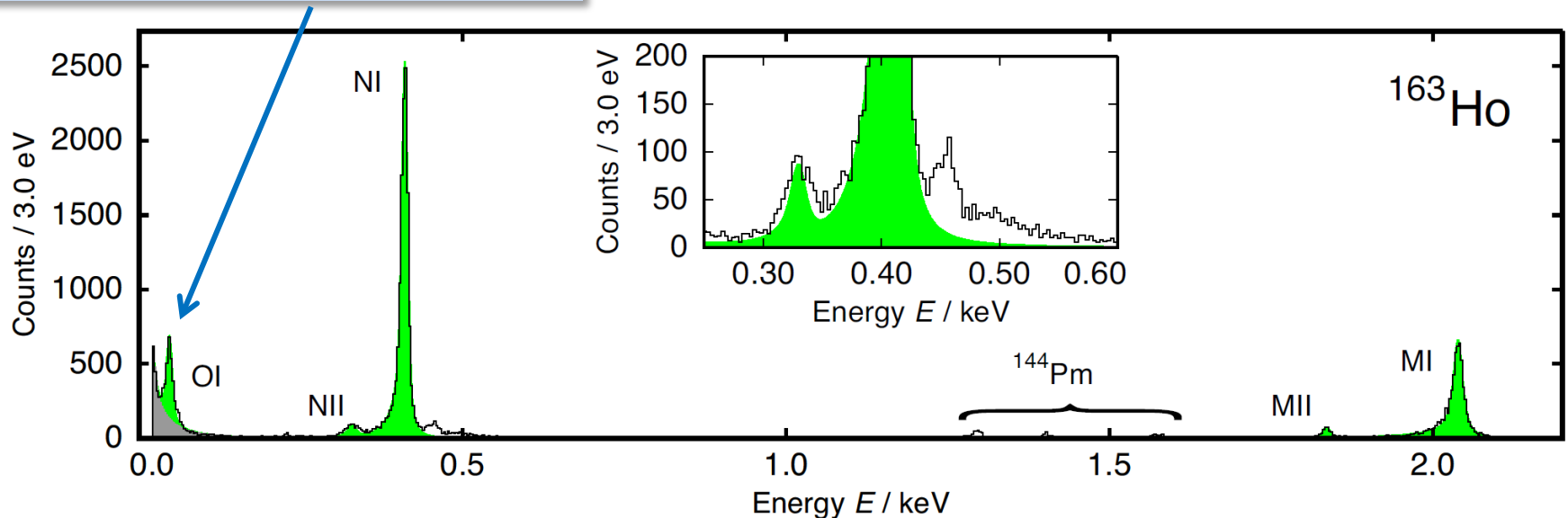


Calorimetric spectrum

- Rise Time ~ 130 ns
- $\Delta E_{\text{FWHM}} = 7.6$ eV @ 6 keV (2013)
- Non-Linearity $< 1\%$ @ 6keV

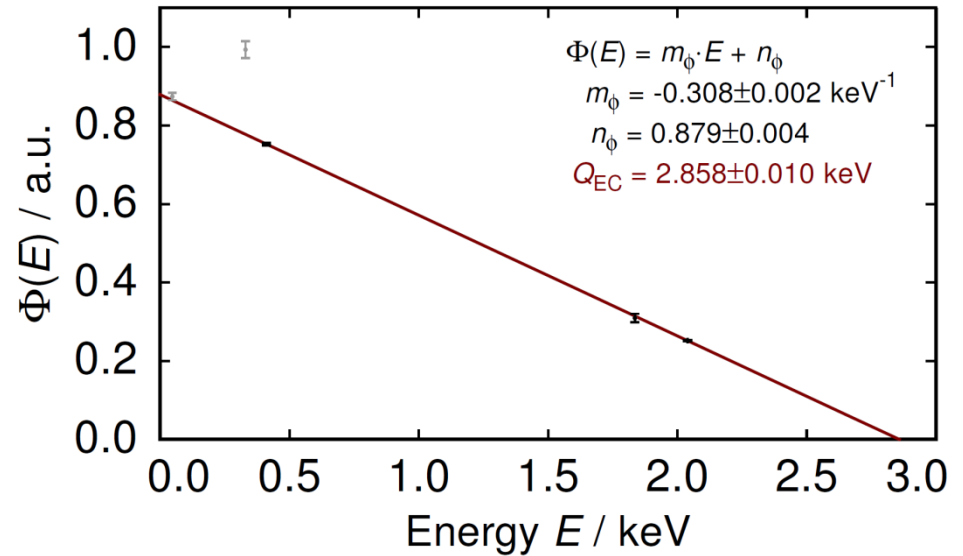
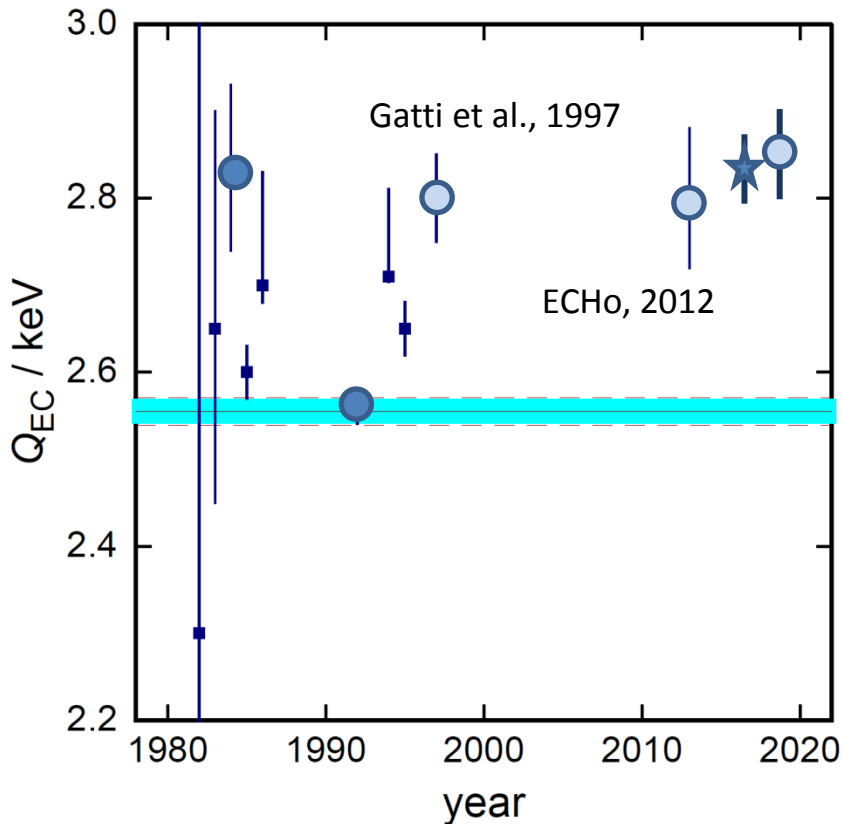
	E_{H} bind.	E_{H} exp.	Γ_{H} lit.	Γ_{H} exp
MI	2.047	2.040	13.2	13.7
MII	1.845	1.836	6.0	7.2
NI	0.420	0.411	5.4	5.3
NII	0.340	0.333	5.3	8.0
OI	0.050	0.048	5.0	4.3

First calorimetric measurement
of the OI-line



Q_{EC} determination

$$\Phi_H(E) = \sqrt{\frac{n_H}{\varphi_H^2(0)B_H}} \propto \sqrt{C}(Q_{EC} - E_H)$$



Our result:

$$Q_{EC} = (2.858 \pm 0.010^{\text{stat}} \pm 0.05^{\text{syst}}) \text{ keV}$$

P. C.-O. Ranitzsch et al., *Phys. Rev. Lett.*, 119, 122501 (2017)

Penning Trap Mass Spectrometry result:

$$Q_{EC} = (2.833 \pm 0.030^{\text{stat}} \pm 0.015^{\text{syst}}) \text{ keV}$$

S. Eliseev et al., *Phys. Rev. Lett.*, 115, 062501 (2015)



Scaling up

^{163}Ho high purity source

Required activity in the detectors: Final experiment $\rightarrow >10^6 \text{ Bq} \rightarrow >10^{17}$ atoms

- Neutron irradiation
(n, γ)-reaction on ^{162}Er

High cross-section 

Radioactive contaminants 

Er161 3.21 h 3/2- EC	Er162 0+ 0.14	Er163 75.0 m 5/2- EC	Er164 0+ 1.61	Er165 10.36 h 5/2- EC	Er166 0+ 33.6
Ho160 25.6 m 5+ EC *	Ho161 2.48 h 7/2- EC *	Ho162 15.0 m 1+ EC *	Ho163 4.70 y 2- EC *	Ho164 29 m 1+ EC, β^- *	Ho165 7/2- 100
Dy159 144.4 d 3/2- EC	Dy160 0+ 2.34	Dy161 5/2+ 18.9	Dy162 0+ 25.5	Dy163 5/2- 24.9	Dy164 0+ 28.2
Tb158 180 y 3- EC, β^- *	Tb159 3/2+ 100	Tb160 72.3 d 3- β^-	Tb161 6.88 d 3/2+ β^-	Tb162 7.60 m 1- β^-	Tb163 19.5 m 3/2+ β^-

^{163}Ho high purity source

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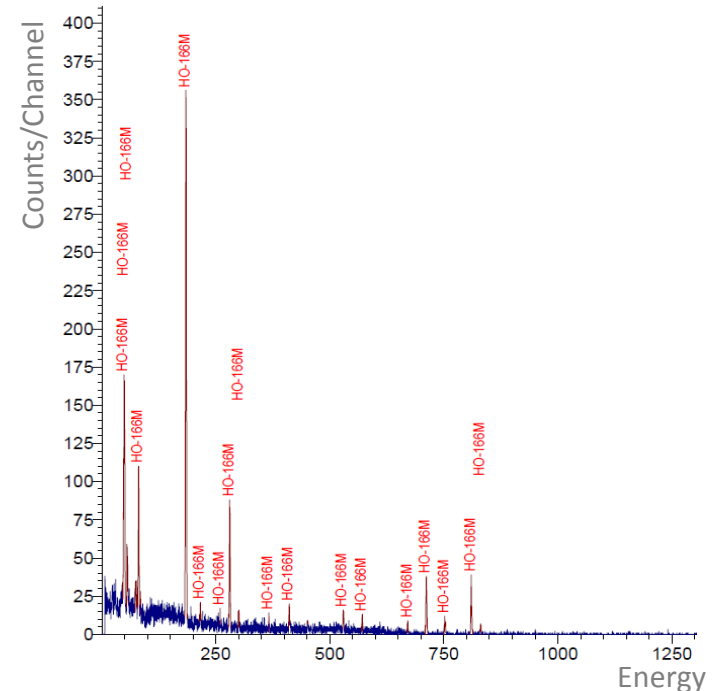


Radioactive contaminants

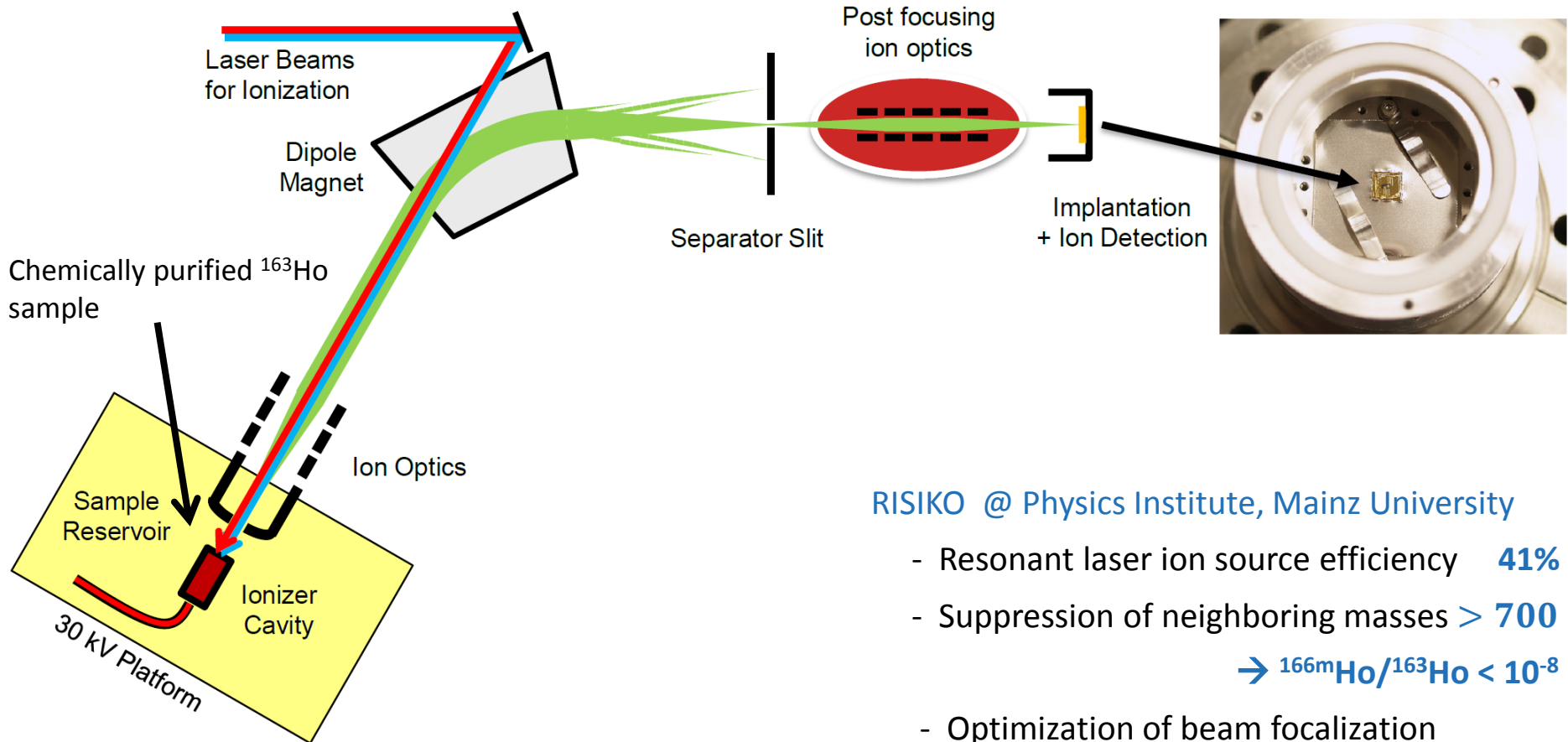


Excellent chemical separation

Er161 3.21 h 3/2- EC	Er162 0+ 0.14	Er163 75.0 m 5/2 EC	Er164 0+ 1.61	Er165 10.36 h 5/2- EC	Er166 0+ 33.6
Ho160 25.6 m 5+ EC *	Ho161 2.48 h 7/2- EC *	Ho162 15.0 m 1+ EC *	Ho163 1.70 y 2- EC *	Ho164 29 m 1+ EC, β^- *	Ho165 7/2- 100
Dy159 144.4 d 3/2- EC	Dy160 0+ 2.34	Dy161 5/2+ 18.9	Dy162 0+ 25.5	Dy163 5/2- 24.9	Dy164 0+ 28.2
Tb158 180 y 3- EC, β^- *	Tb159 3/2+ 100	Tb160 72.3 d 3- β^-	Tb161 6.88 d 3/2+ β^-	Tb162 7.60 m 1- β^-	Tb163 19.5 m 3/2+ β^-



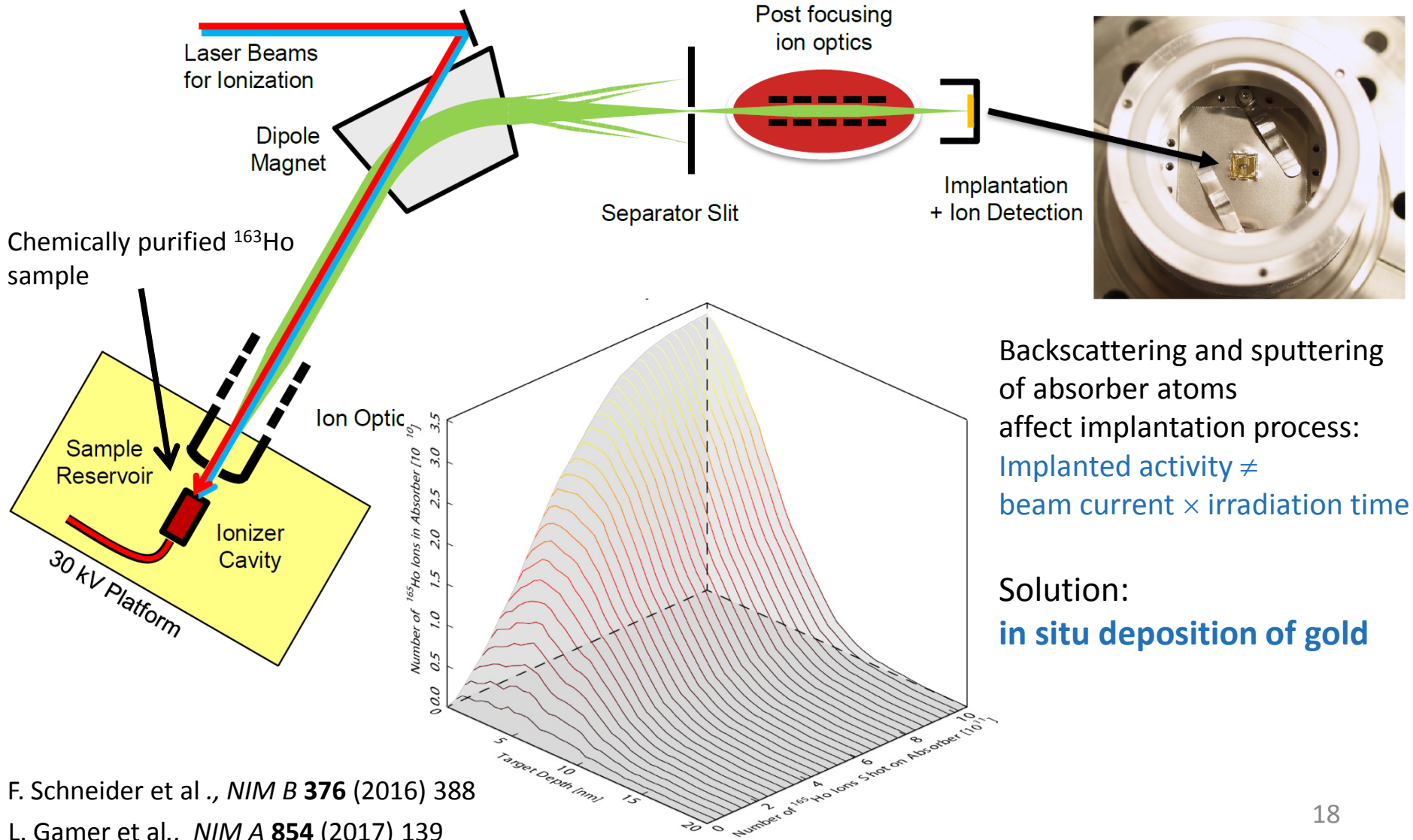
Mass separation and ^{163}Ho ion-implantation



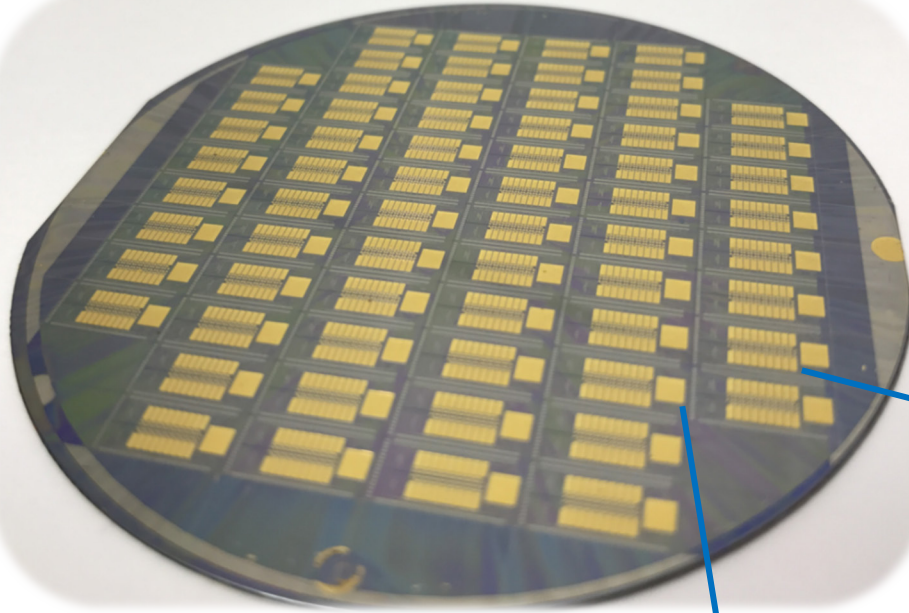
RISIKO @ Physics Institute, Mainz University

- Resonant laser ion source efficiency **41%**
- Suppression of neighboring masses **> 700**
→ $^{166}\text{mHo}/^{163}\text{Ho} < 10^{-8}$
- Optimization of beam focalization

Mass separation and ^{163}Ho ion-implantation



ECHO-1k array



3" wafer with 64 ECHO-1k chip

Suitable for
parallel and multiplexed readout

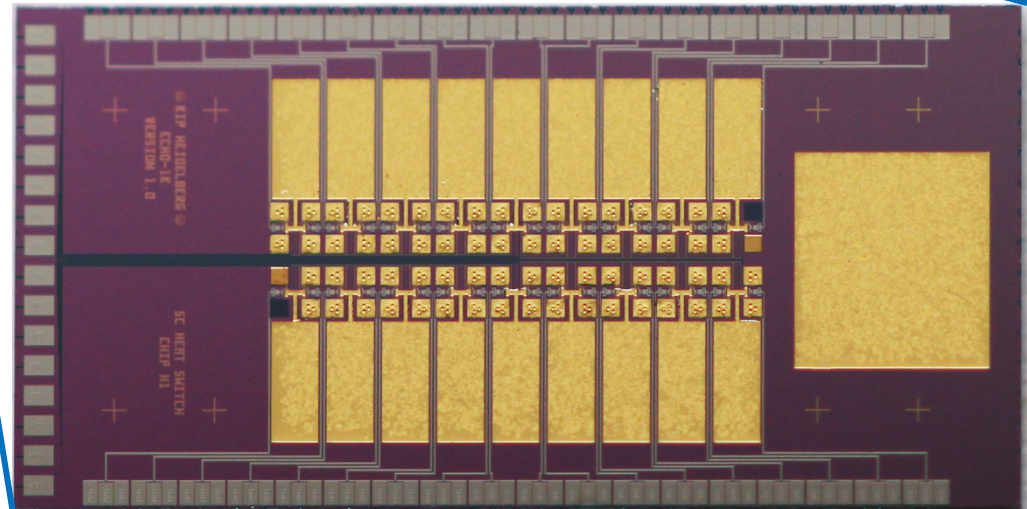
64 pixels which can be loaded with ^{163}Ho
+ 4 detectors for diagnostics

Design performance:

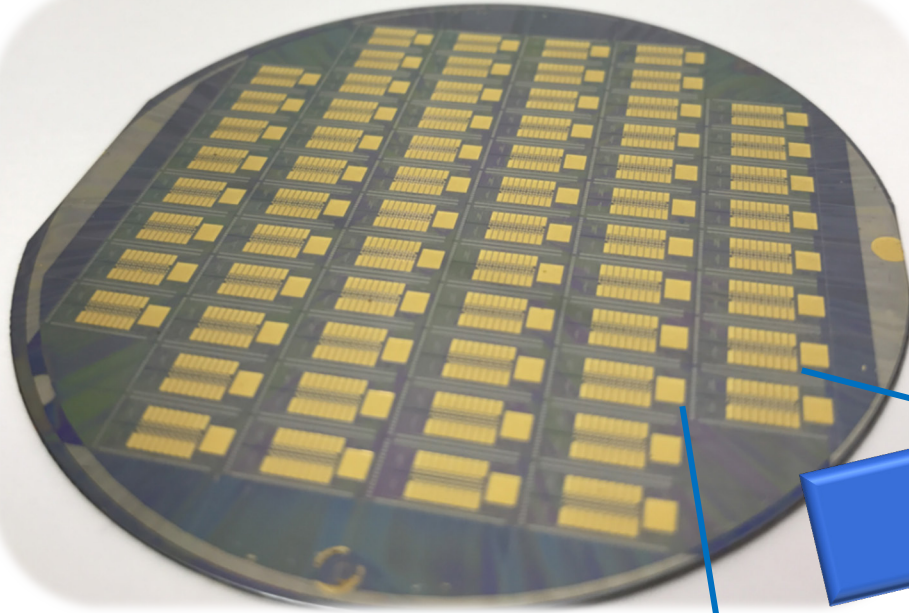
$$\Delta E_{\text{FWHM}} \sim 5 \text{ eV}$$

$$\tau_r \sim 90 \text{ ns (single channel readout)}$$

$$\tau_r \sim 300 \text{ ns (multiplexed read-out)}$$



ECHO-1k array



3" wafer with 64 ECHO-1k chip

Suitable for
parallel and multiplexed readout

4 chips implanted:
Expected activity between 1 Bq and 3 Bq

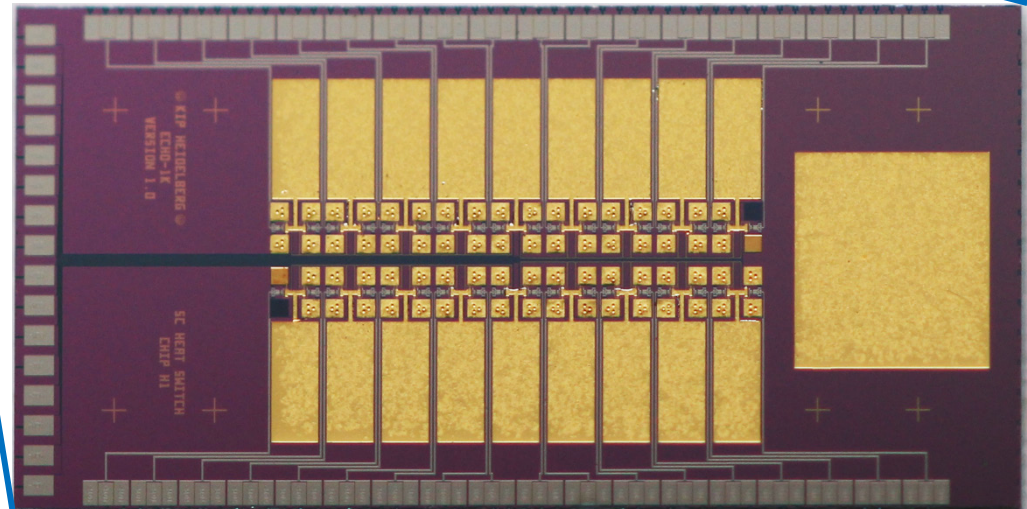
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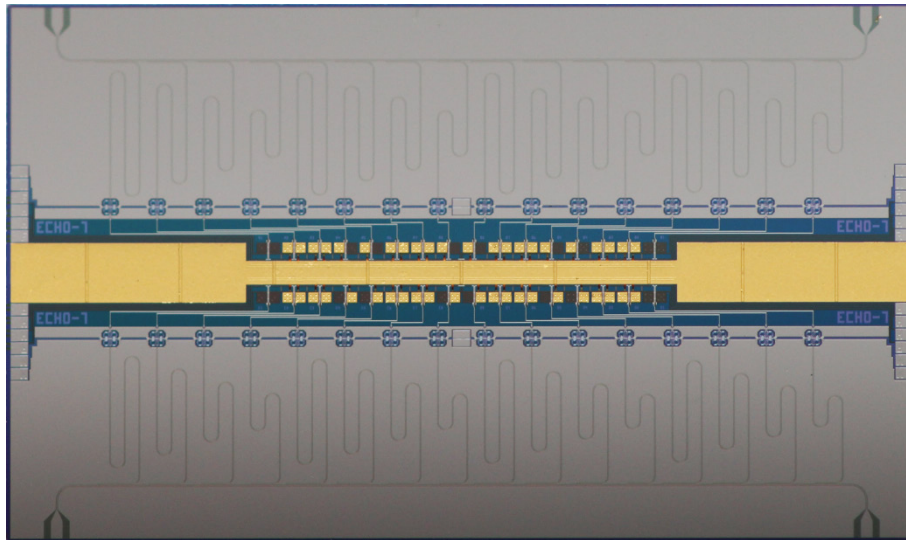
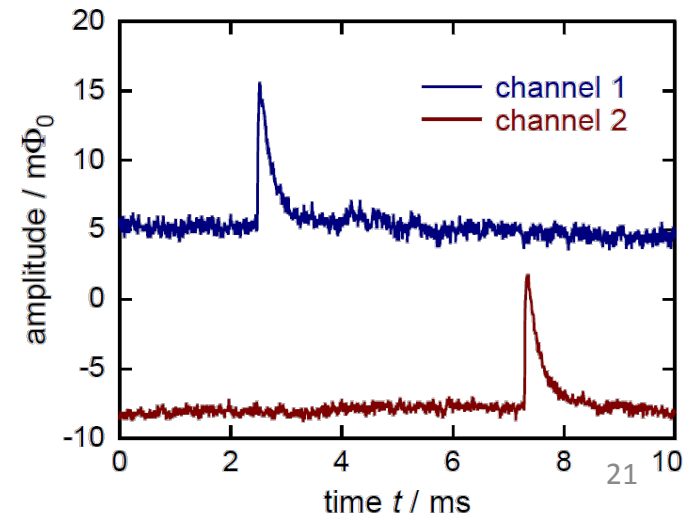
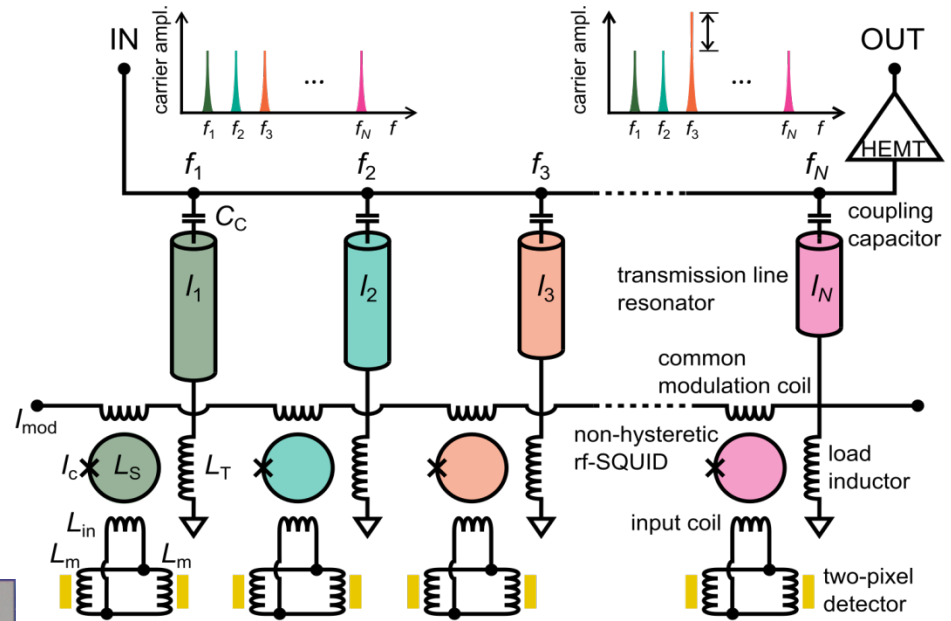


Multiplexing readout

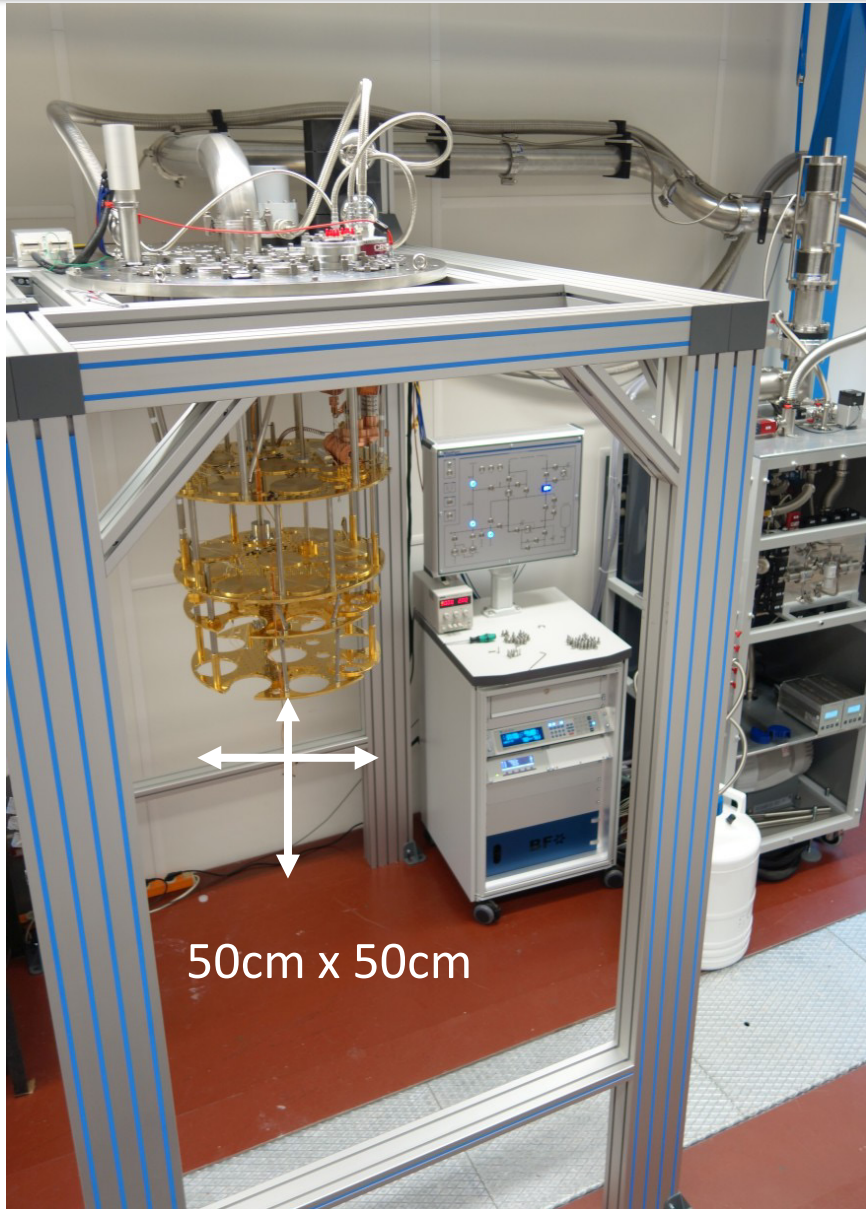
Microwave SQUID multiplexing

Single HEMT amplifier and 2 coaxes to read out **100 - 1000** detectors

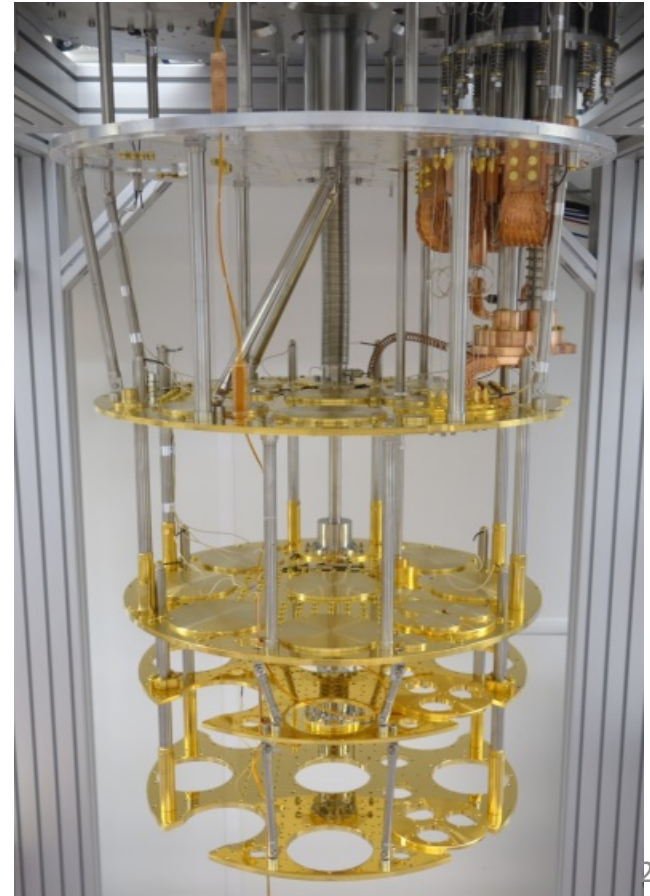
- Reliable fabrication of **64-pixel array**
- Successful characterization of first prototypes
→ **optimization of design parameters**



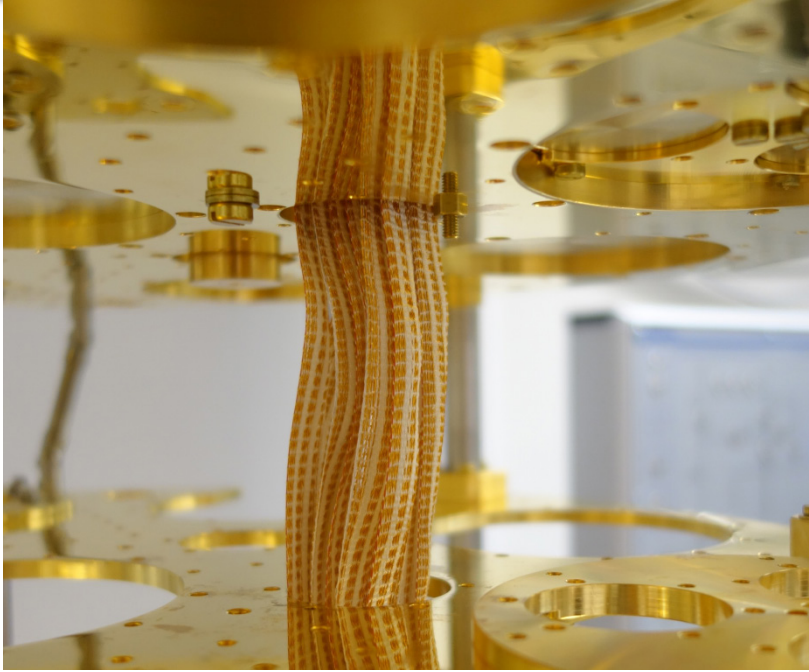
ECHO cryogenic platform



- Large space at MXC enough for several ECHO phases
- cooling power: $15\mu\text{W}$ @ 20 mK
- Possibility to load 200kg for passive shielding



ECHO cryogenic platform

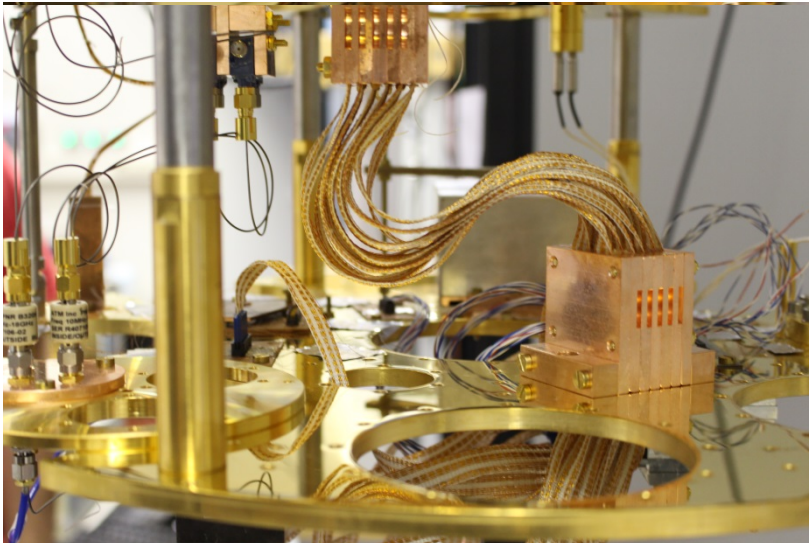


- Large space at MXC enough for several ECHO phases
- cooling power: $15\mu\text{W}$ @ 20 mK
- Possibility to load 200kg for passive shielding
- Presently equipped with:

2 RF lines for microwave multiplexing readout of 2 MMC arrays

12 ribbons each with 30 Cu98Ni2 0.2 mm, 1.56 Ohm/m, cables from RT to mK

→ allows for parallel readout of 36 two-stage SQUID set-up



ECHo-1k (2015 - 2018)

^{163}Ho activity: $A_t = 1 \text{ kBq}$

Detectors: **Metallic Magnetic Calorimeters**

→ Energy resolution $\Delta E_{\text{FWHM}} \leq 5 \text{ eV}$

→ Time resolution $\tau \leq 1 \mu\text{s}$

Unresolved pile-up fraction $f_{\text{pu}} \leq 10^{-5}$

→ activity per pixel: $A = 10 \text{ Bq}$

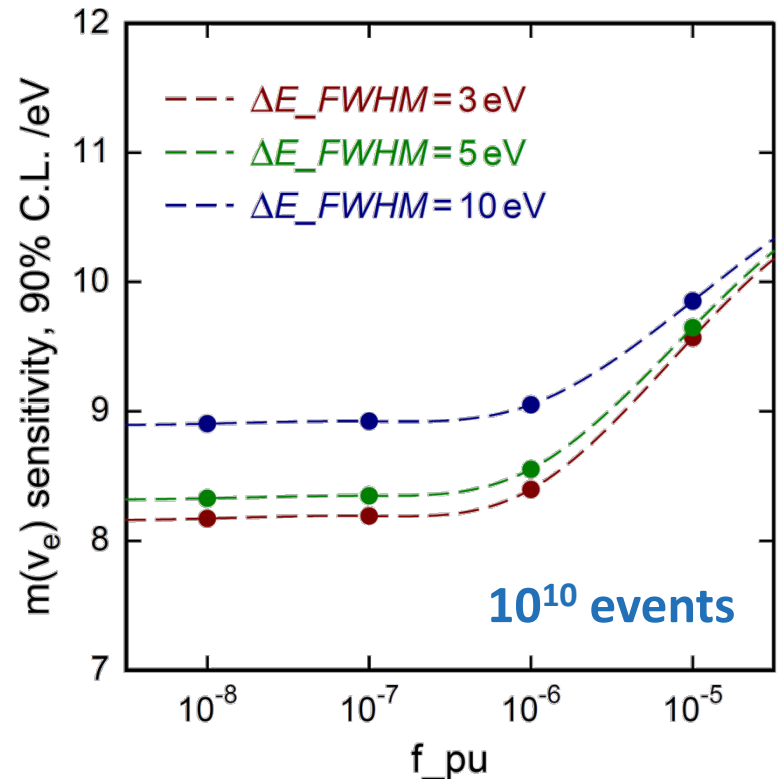
→ number of detectors $N = 100$

Read-out : **Microwave SQUID Multiplexing**

→ 2 arrays with ~ 50 single pixels

Background $b < 10^{-5} \text{ /eV/det/day}$

Measuring time $t = 1 \text{ year}$



$m(\nu_e) < 10 \text{ eV}$ 90% C.L.

ECHo-1M (next future)

^{163}Ho activity: $A_t = 1 \text{ MBq}$

Detectors: **Metallic Magnetic Calorimeters**

→ Energy resolution $\Delta E_{FWHM} \leq 3 \text{ eV}$

→ Time resolution $\tau \leq 0.1 \mu\text{s}$

Unresolved pile-up fraction $f_{pu} \leq 10^{-6}$

→ activity per pixel: $A = 10 \text{ Bq}$

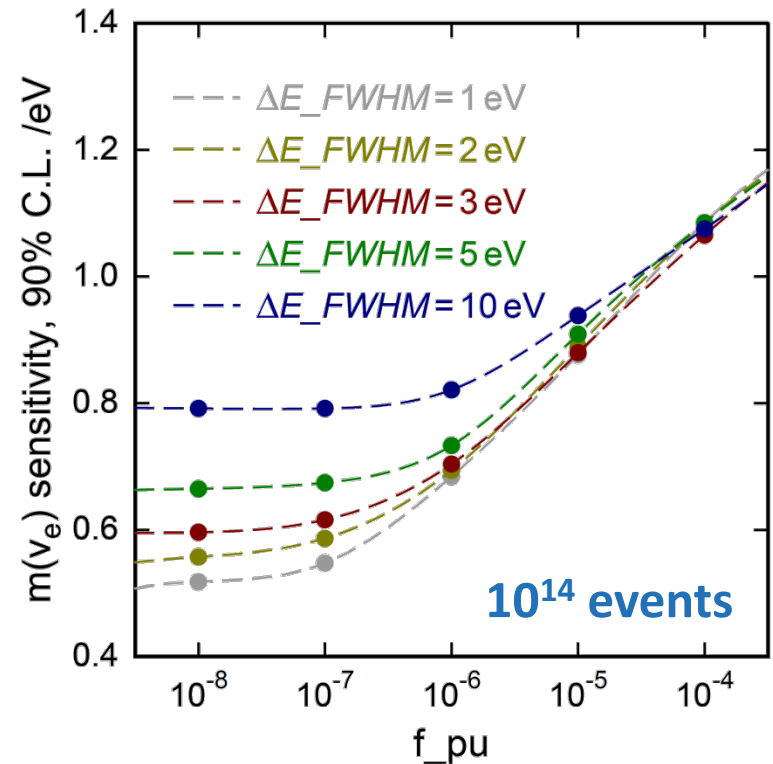
→ number of detectors $N = 10^5$

Read-out : **Microwave SQUID Multiplexing**

→ 100 arrays with ~ 1000 single pixels

Background $b < 10^{-6} \text{ /eV/det/day}$

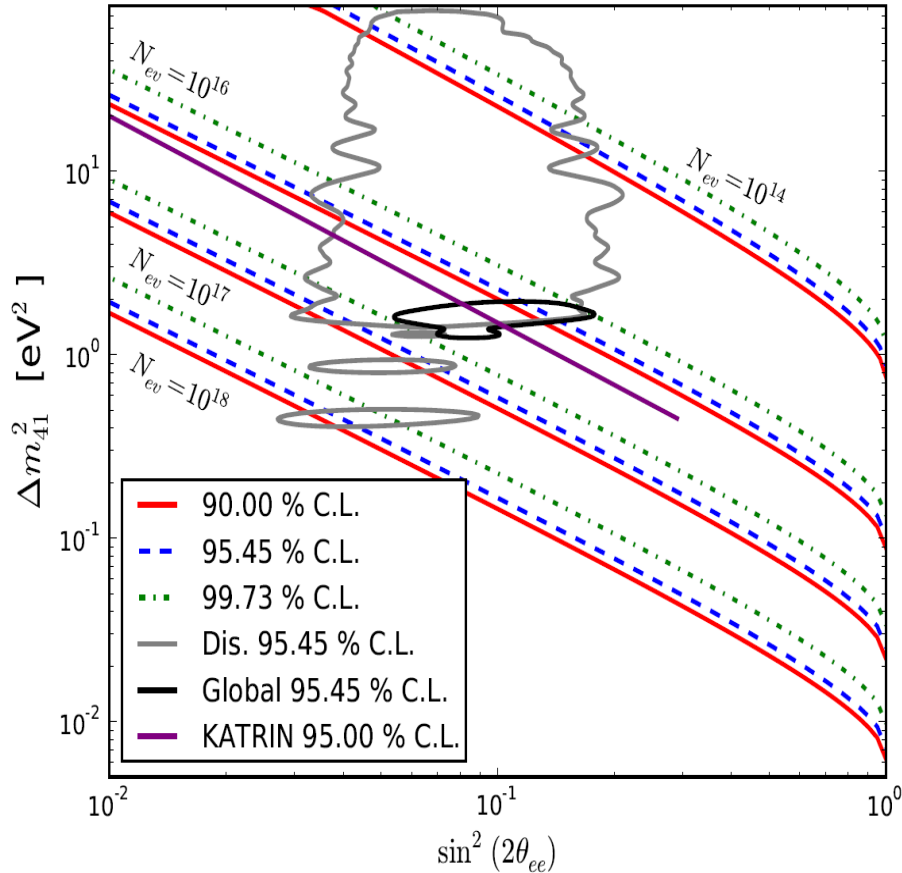
Measuring time $t = 1 - 3 \text{ year}$



$m(\nu_e) < 1 \text{ eV } 90\% \text{ C.L.}$

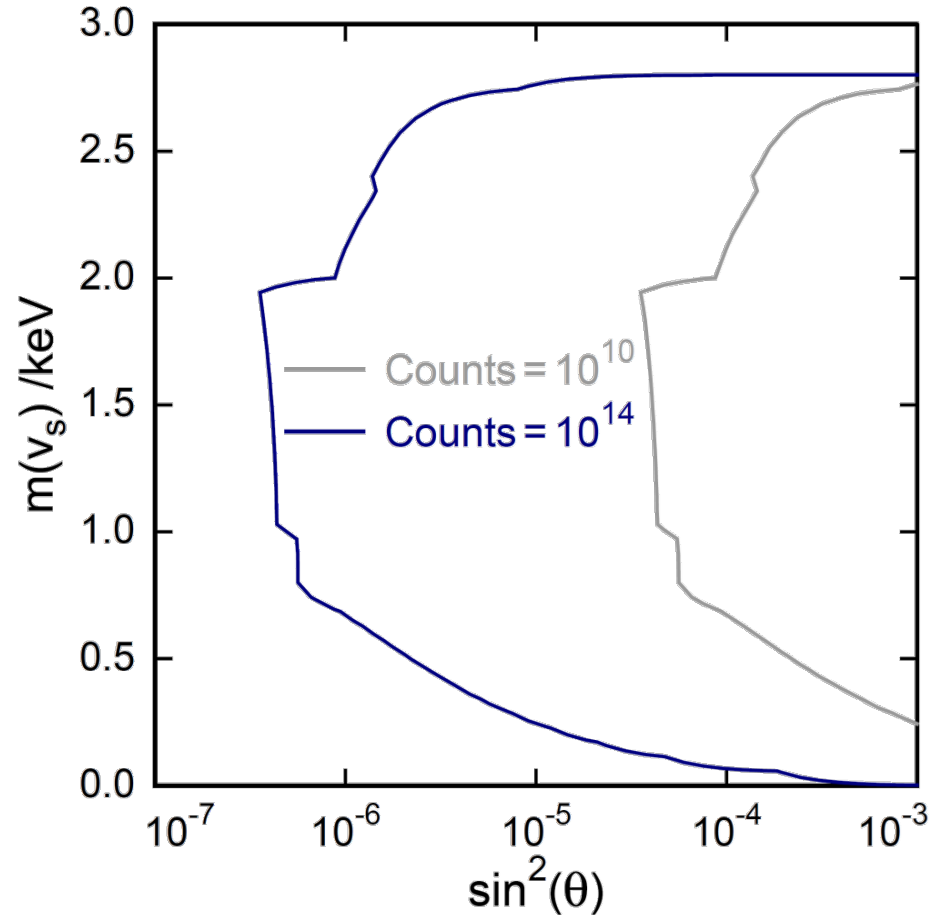
Sterile neutrinos

eV-scale sterile neutrinos



L. Gastaldo, C. Giunti, E. Zavanin.,
High Energ. Phys. **06** (2016) 61.

keV-scale sterile neutrinos



A White Paper on keV Sterile
 Neutrino Dark Matter, JCAP01(2017)025

Conclusions and outlook

The ECHO collaboration aims to reach sub-eV sensitivity on the electron neutrino mass analysing high statistics and high resolution ^{163}Ho spectra

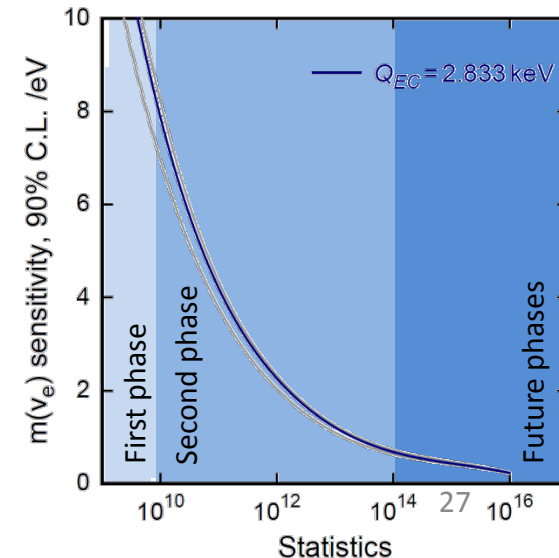
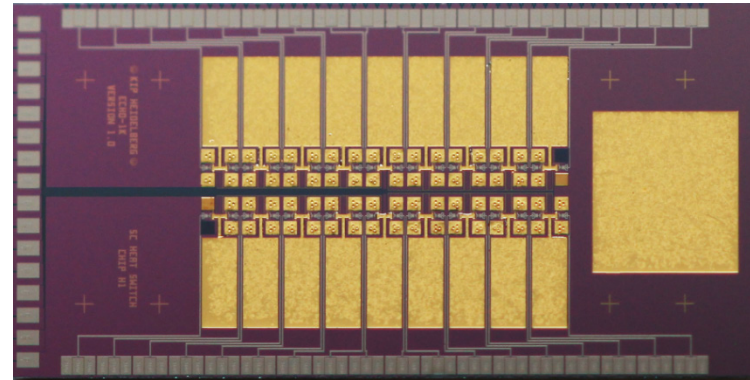
- Independent ^{163}Ho Q_{EC} measurement

$$Q_{\text{EC}} = (2.833 \pm 0.030^{\text{stat}} \pm 0.015^{\text{syst}}) \text{ keV}$$

$$Q_{\text{EC}} = (2.858 \pm 0.010^{\text{stat}} \pm 0.05^{\text{syst}}) \text{ keV}$$

- High purity ^{163}Ho sources have been produced
- ^{163}Ho ions can be successfully enclosed in microcalorimeter absorbers
- Large arrays have been tested and microwave SQUID multiplexing has been successfully proved
- **A new limit on the electron neutrino mass is approaching**

Er161 3.21 h 3/2-	Er162 0+	Er163 75.0 m 5/2-	Er164 0+	Er165 10.36 h 5/2-	Er166 0+
EC	0.14	EC	1.61	EC	33.6
Ho160 25.6 m 5+	Ho161 2.48 h 7/2- *	Ho162 15.0 m 1+ *	Ho163 4570 y 7/2- *	Ho164 29 m 1+ *	Ho165 7/2-
EC	EC	EC	EC	EC, β^-	100



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Fedor Simkovic

Department of Physics, Indian Institute of Technology Roorkee, India

Moumita Maiti

Goethe Universität Frankfurt am Main

Udo Kbschull, Panagiotis Neroutsos

Institute for Nuclear Chemistry, Johannes Gutenberg University Mainz

Christoph E. Düllmann, Klaus Eberhardt, Holger Dorrer, Fabian Schneider

Institute of Nuclear Research of the Hungarian Academy of Sciences

Zoltán Szúcs

Institute of Nuclear and Particle Physics, TU Dresden, Germany

Kai Zuber

Institute for Physics, Humboldt-Universität zu Berlin

Alejandro Saenz

Institute for Physics, Johannes Gutenberg-Universität

Klaus Wendt, Sven Junck, Tom Kieck

Institute for Theoretical Physics, Heidelberg University

Maurits Haverkort, Martin Brass

Institute for Theoretical Physics, University of Tübingen, Germany

Amand Fäßler

Institut Laue-Langevin, Grenoble, France

Ulli Köster

ISOLDE-CERN

Marsh Bruce, Day Goodacre Tom, Johnston Karl, Rothe Sebastian,

Stora Thierry, Veinhard Matthieu

Kirchhoff-Institute for Physics, Heidelberg University, Germany

Christian Enss, Dorothea Fonnesu, Loredana Gastaldo, Andreas Fleischmann,

Clemens Hassel, Federica Mantegazini, Sebastian Kempf, Mathias Wegner

Max-Planck Institute for Nuclear Physics Heidelberg, Germany

Klaus Blaum, Andreas Dörr, Sergey Eliseev, Mikhail Goncharov,

Yuri N. Novikov, Alexander Rischka, Rima Schüssler

Petersburg Nuclear Physics Institute, Russia

Yuri Novikov, Pavel Filianin

Physics Institute, University of Tübingen, Germany

Josef Jochum, Stephan Scholl

Saha Institute of Nuclear Physics, Kolkata, India

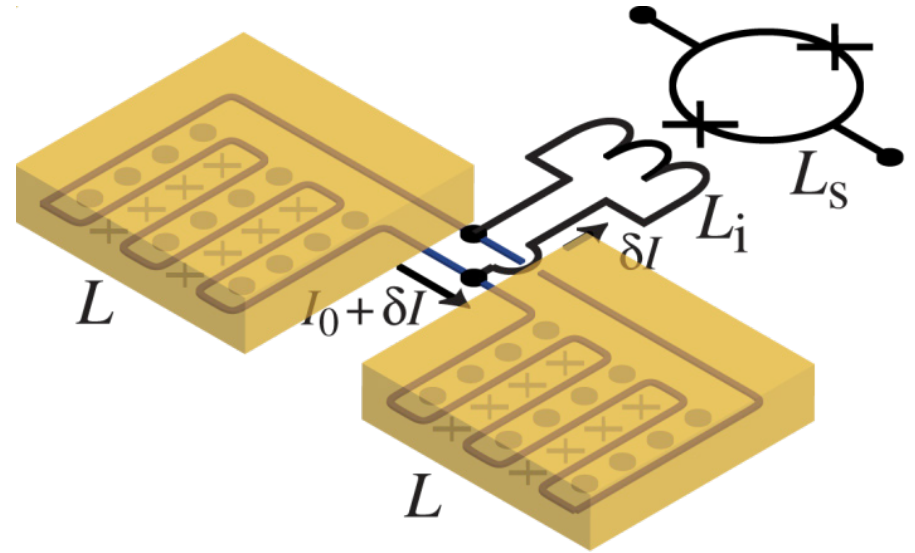
Susanta Lahiri



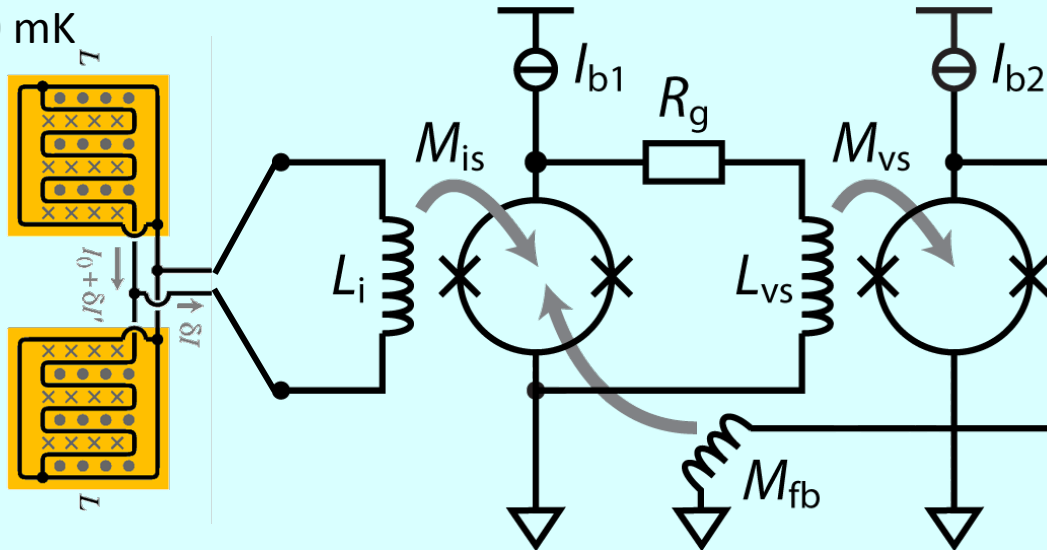
Thank you!

MMC geometry and read-out

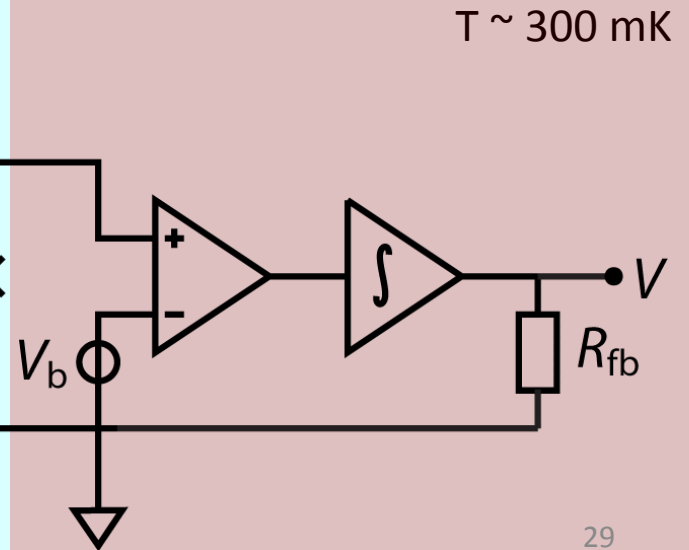
- Planar temperature sensor
 - B-field generated by persistent current
 - transformer coupled to SQUID
-
- Two-stage SQUID read-out



$T \sim 30 \text{ mK}$



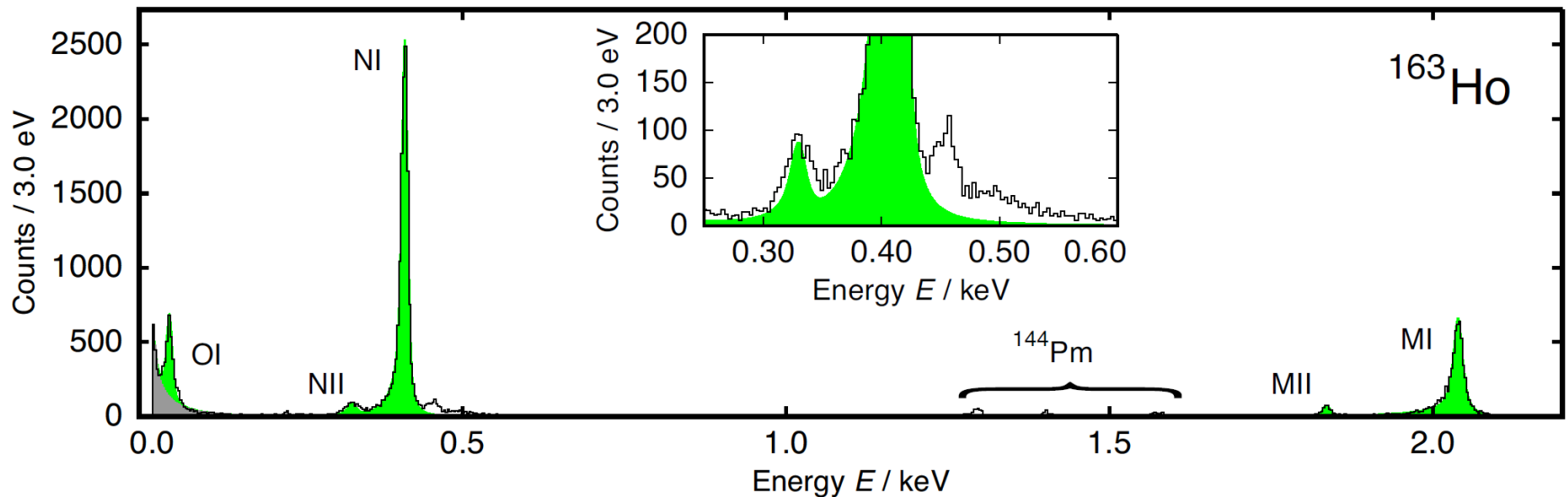
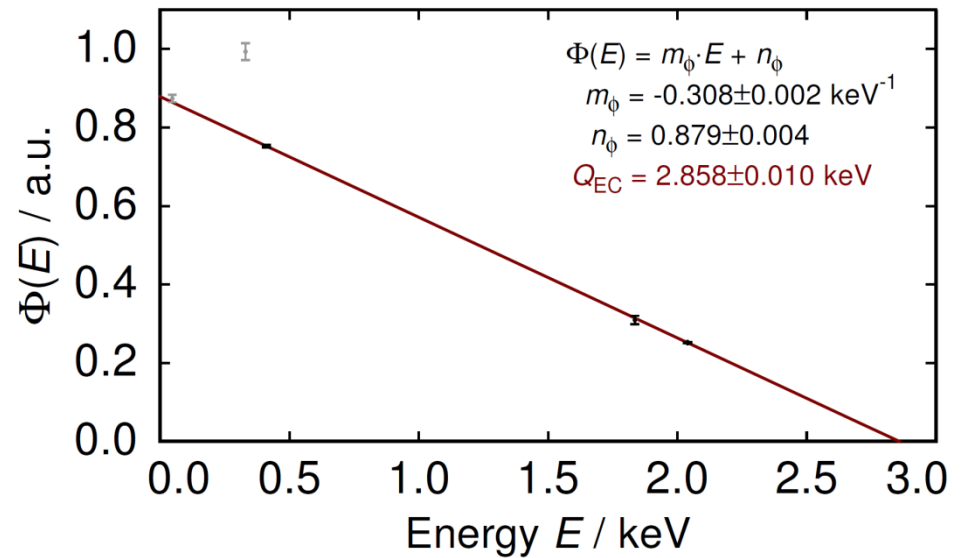
$T \sim 300 \text{ mK}$



Q_{EC} determination

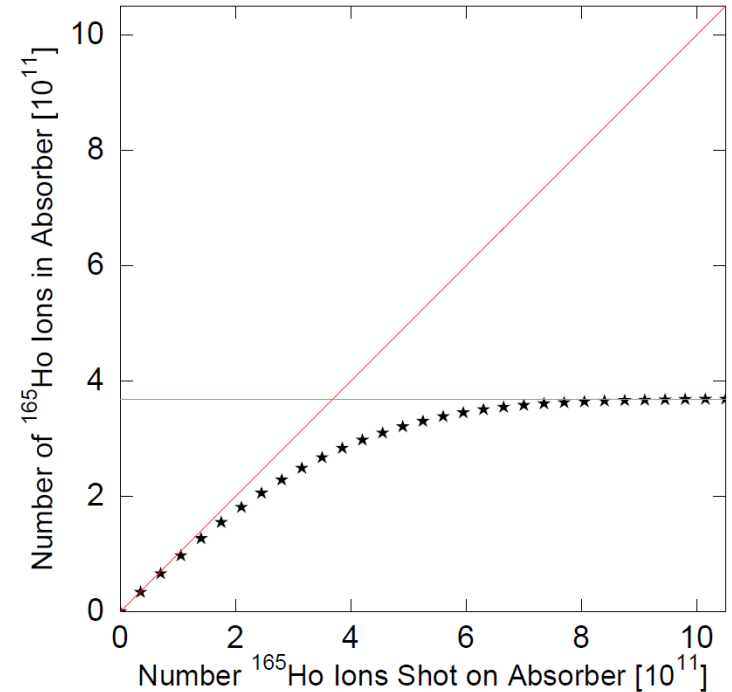
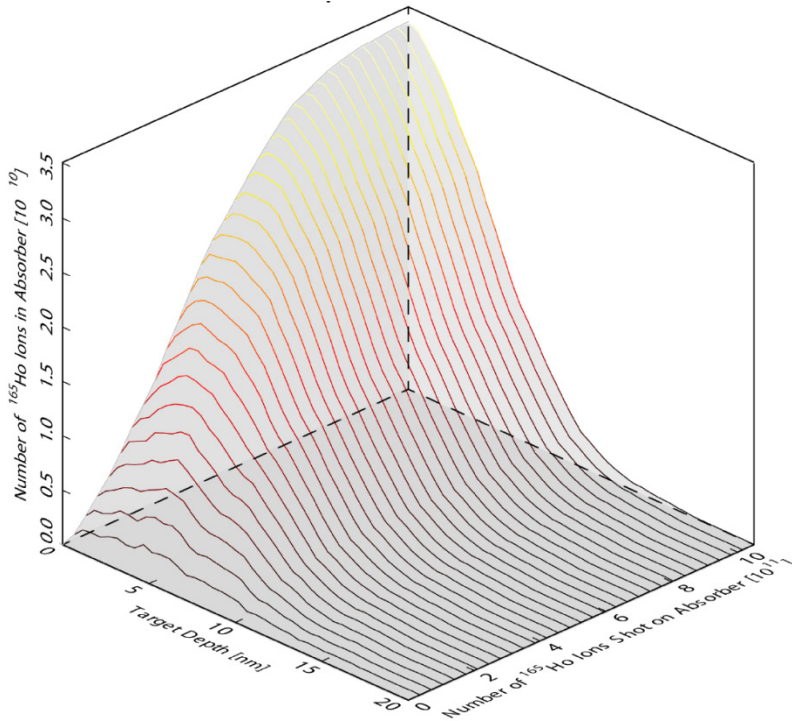
$$\Phi_H(E) = \sqrt{\frac{n_H}{\varphi_H^2(0)B_H}} \propto \sqrt{C}(Q_{EC} - E_H)$$

Line amplitudes are affected
by the phase space factor



Mass separation and ^{163}Ho ion-implantation

Implantation with 30 keV at RISIKO in Mainz in an area of $150\ \mu\text{m} \times 150\ \mu\text{m}$



Backscattering and sputtering of absorber atoms affect implantation process:

Implanted activity \neq beam current \times irradiation time

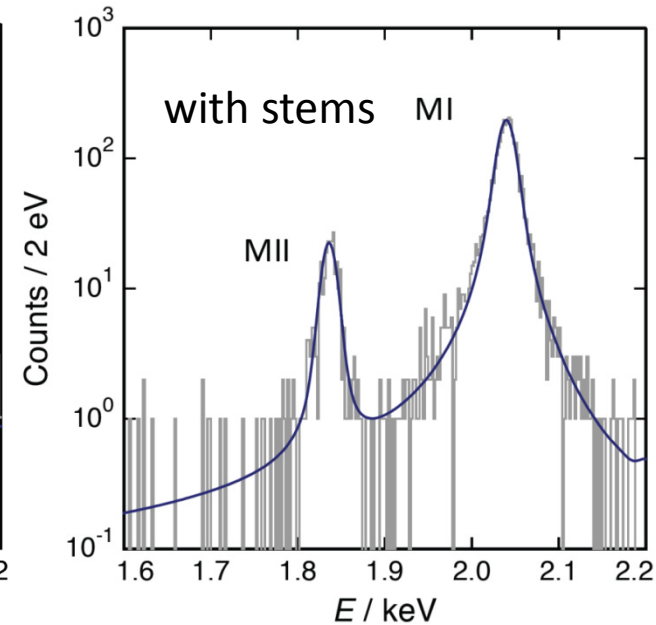
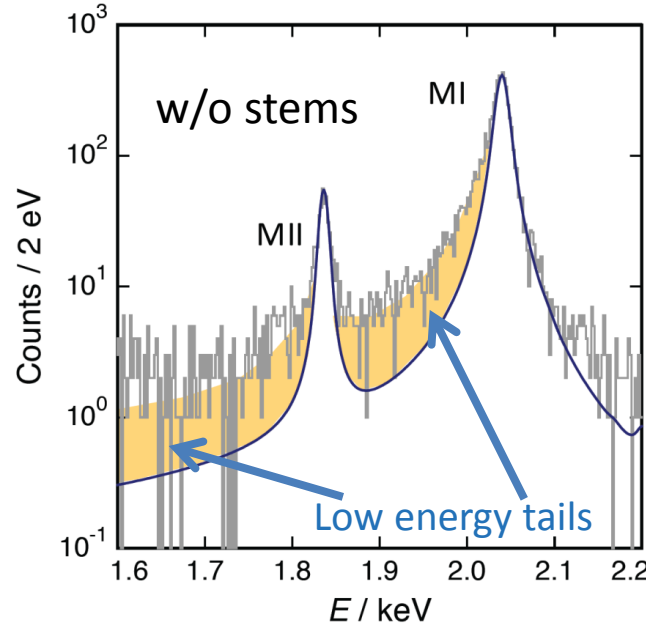
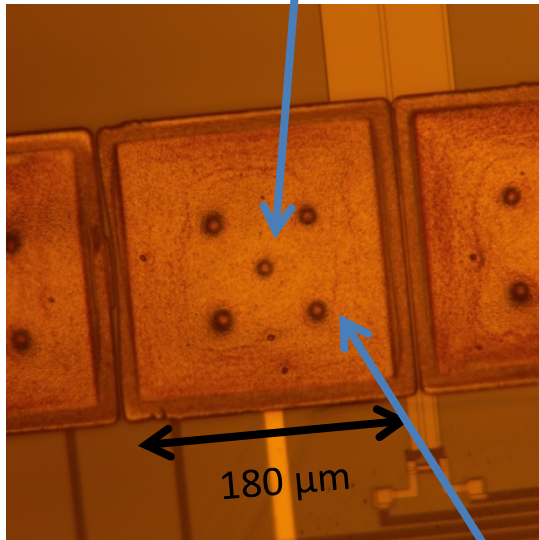


Implantation of Ho in gold with $E = 30\ \text{keV}$:
maximum number of Ho ions $\sim 3.6 \times 10^{11}$
corresponding to only $\sim 2\ \text{Bq}$

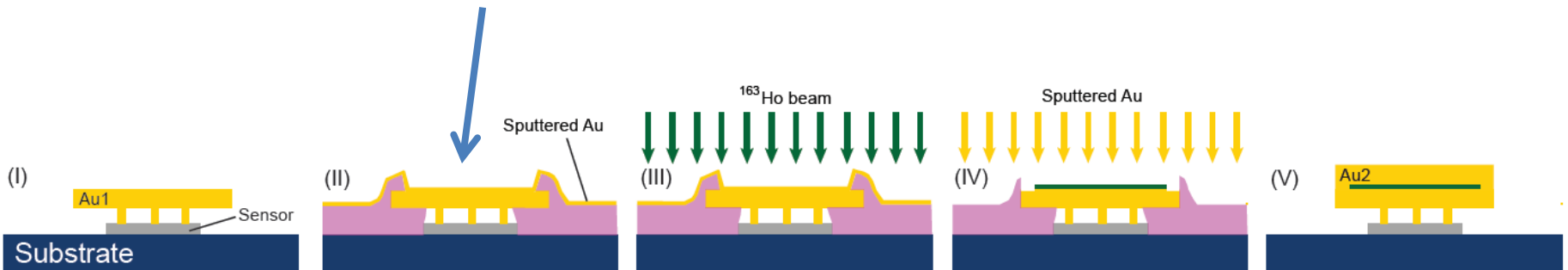
Solution: in situ deposition of gold

Fabrication 4π absorber

Stems between absorber and sensor prevent athermal phonon loss to the substrate



Definition of the implantation area by microstructuring a photoresist layer



^{163}Ho high purity source

Required activity in the detectors: Final experiment $\rightarrow >10^6 \text{ Bq} \rightarrow >10^{17}$ atoms

- Neutron irradiation
(n, γ)-reaction on ^{162}Er

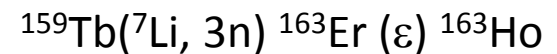
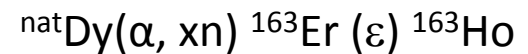
High cross-section 

Radioactive contaminants 



Er161 3.21 h 3/2- EC	Er162 0+ 0.14	Er163 75.0 m 5/2- EC	Er164 0+ 1.61	Er165 10.36 h 5/2- EC	Er166 0+ 33.6
Ho160 25.6 m 5+ EC *	Ho161 2.48 h 7/2- EC *	Ho162 15.0 m 1+ EC *	Ho163 4.70 y 2- EC	Ho164 29 m 1+ EC, β^-	Ho165 2.3 y 2- EC *
Dy159 144.4 d 3/2- EC	Dy160 0+ 2.34	Dy161 5/2+ 18.9	Dy162 0+ 25.5	Dy163 5/2- 24.9	Dy164 0+ 28.2
Tb158 180 y 3- EC, β^- *	Tb159 3/2+ 100	Tb160 72.3 d 3- β^-	Tb161 6.88 d 3/2+ β^-	Tb162 7.60 m 1- β^-	Tb163 19.5 m 3/2+ β^-

- Charged particle activation



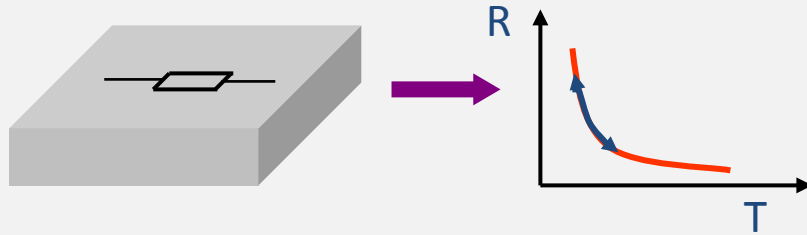
Small cross-section 

Few radioactive contaminants 

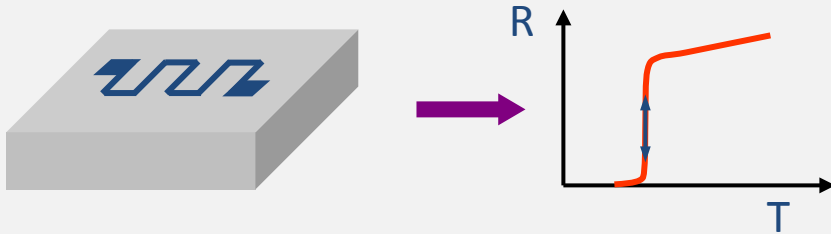
NuMECS

Temperature sensors

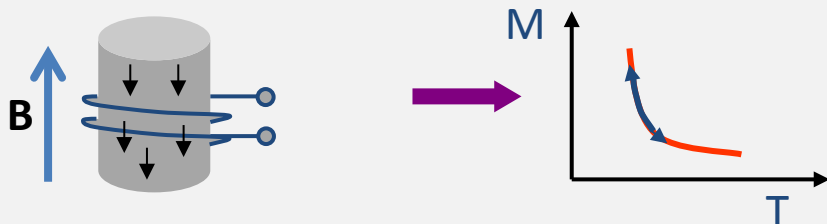
Resistance of highly doped semiconductors



Resistance at superconducting transition, TES

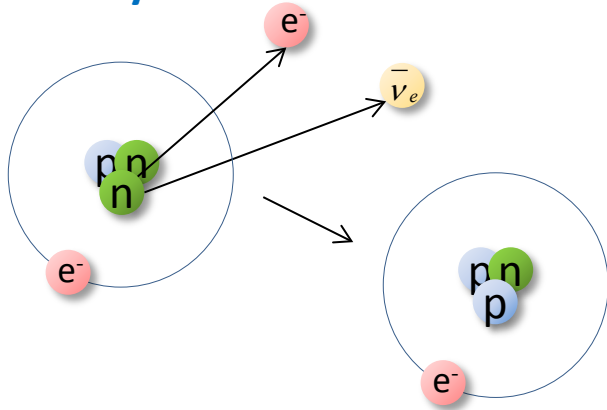


Magnetization of paramagnetic material, MMC

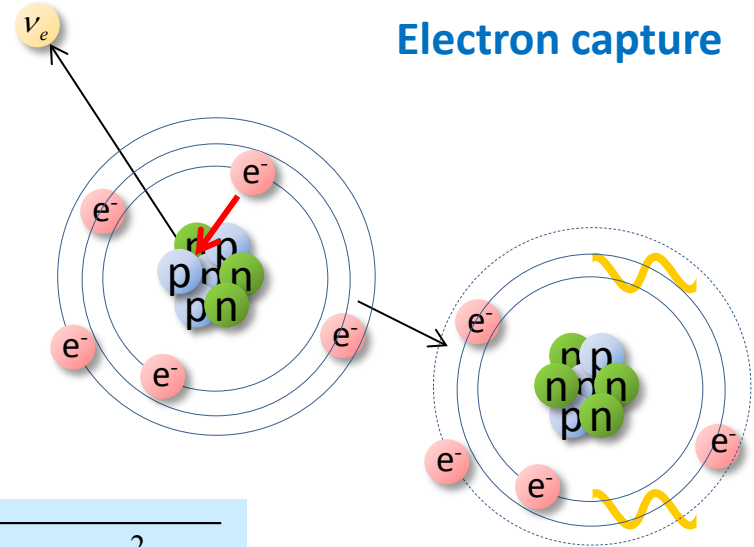


Kinematic approach

Beta decay



Electron capture

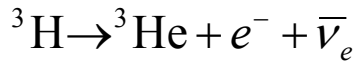
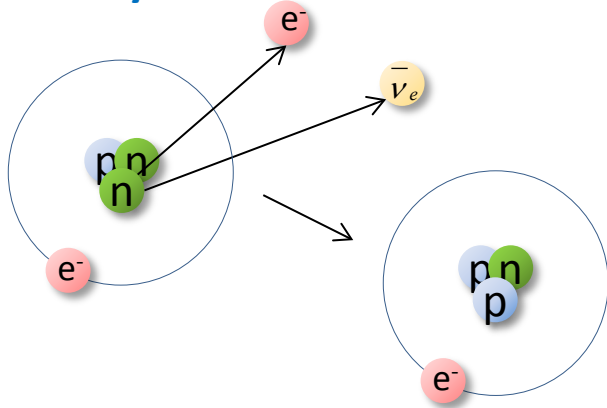


$$\frac{dW}{dE} \propto (Q - E)^2 \sqrt{1 - \frac{m_\nu^2}{(Q - E)^2}}$$

- A finite neutrino mass modify the spectrum in a small region close to the end-point
- Low Q-values enhance the fraction of events in the region of interest

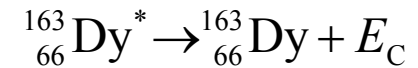
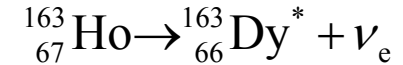
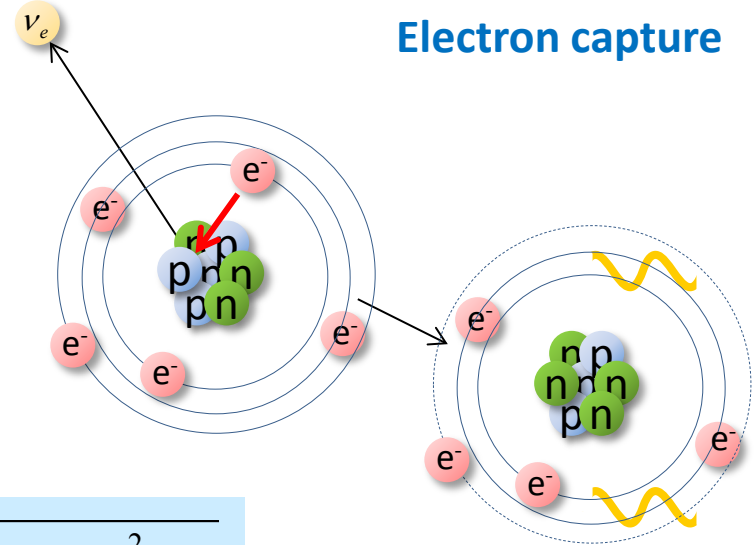
Kinematic approach

Beta decay



$$\frac{dW}{dE} \propto (Q - E)^2 \sqrt{1 - \frac{m_\nu^2}{(Q - E)^2}}$$

Electron capture



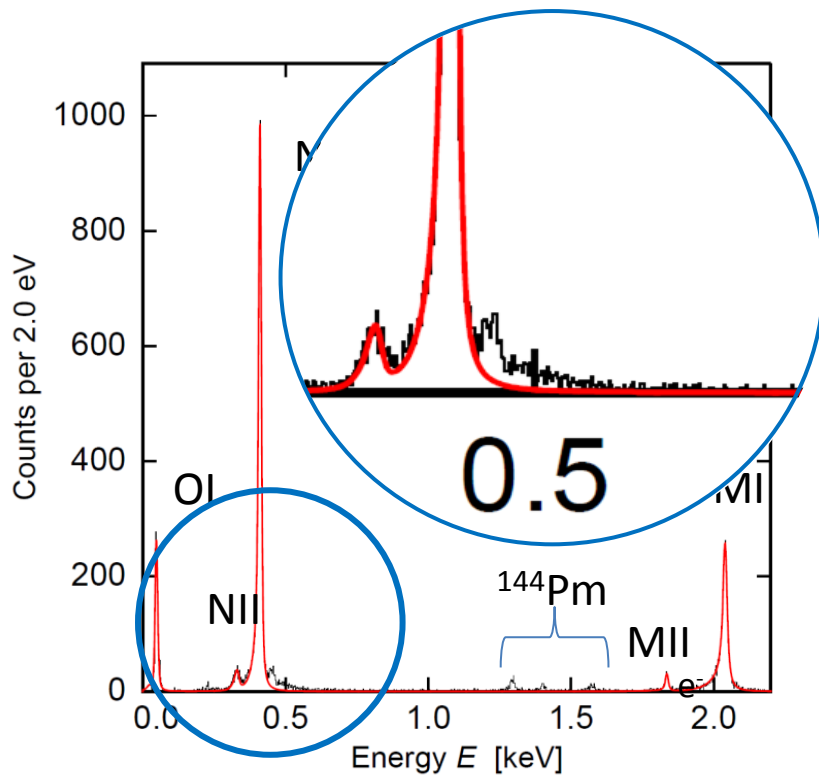
$$m(\bar{\nu}_e) < 2.2 \text{ eV} \quad (1)$$

$$m(\nu_e) < 225 \text{ eV} \quad (2)$$

(1) Ch. Kraus *et al.*, *Eur. Phys. J. C* **40** (2005) 447
 Ch. Weinheimer, *Prog. Part. Nucl. Phys.* **57** (2006) 22
 N. Aseev *et al.*, *Phys. Rev D* **84** (2011) 112003

(2) P. T. Springer, C. L. Bennett, and P. A. Baisden *Phys. Rev. A* **35** (1987) 679

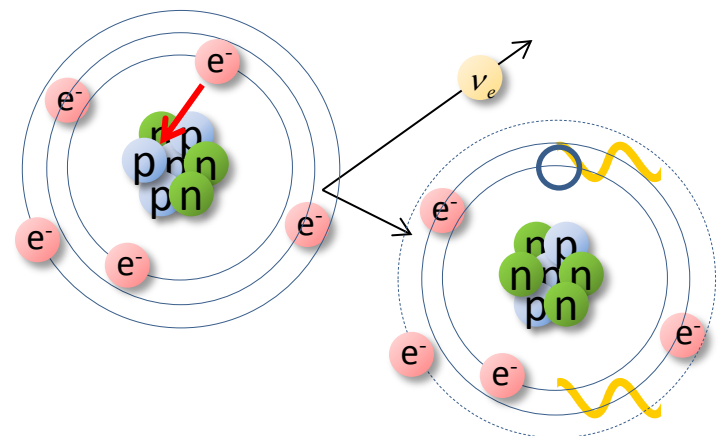
Characterisation of spectral shape



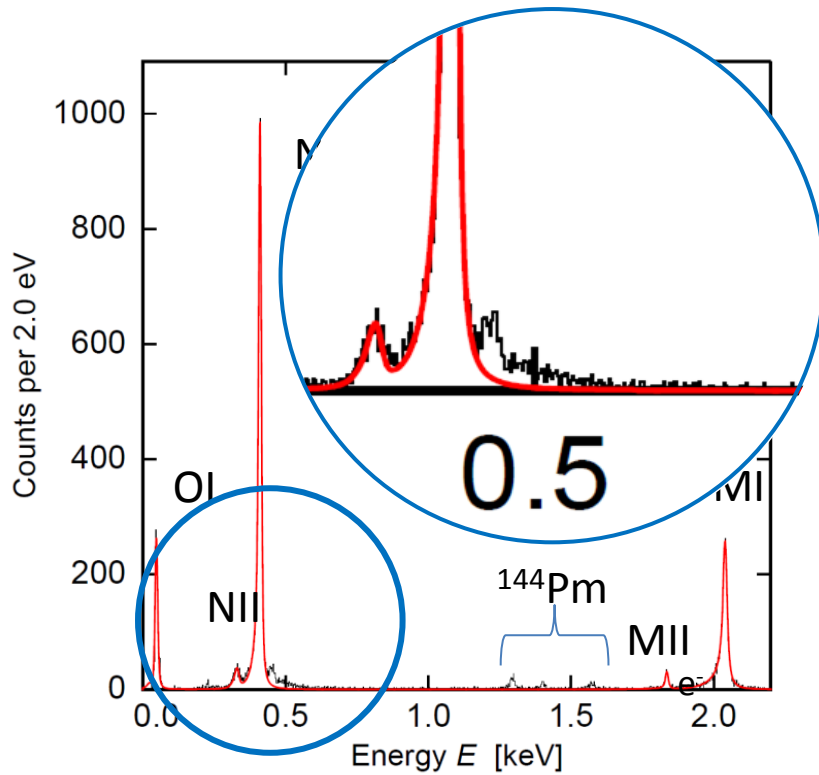
Estimate the effect of

- Higher order excitation in ^{163}Ho

- A. Faessler et al.
J. Phys. G **42** (2015) 015108
- R. G. H. Robertson
Phys. Rev. C **91**, 035504 (2015)
- A. Faessler et al.
Phys. Rev. C **91**, 045505 (2015)
- A. Faessler et al.
Phys. Rev. C **91**, 064302 (2015)
- A. De Rujula et al.
arXiv:1601.04990v1 [hep-ph] 19 Jan 2016

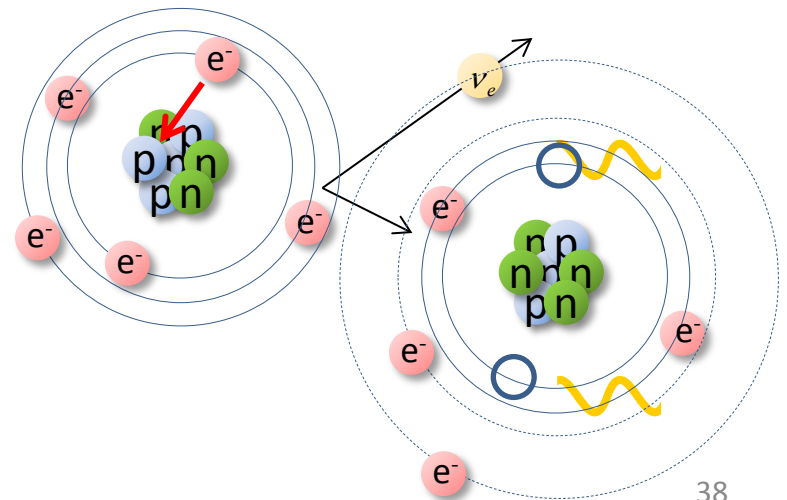


Characterisation of spectral shape

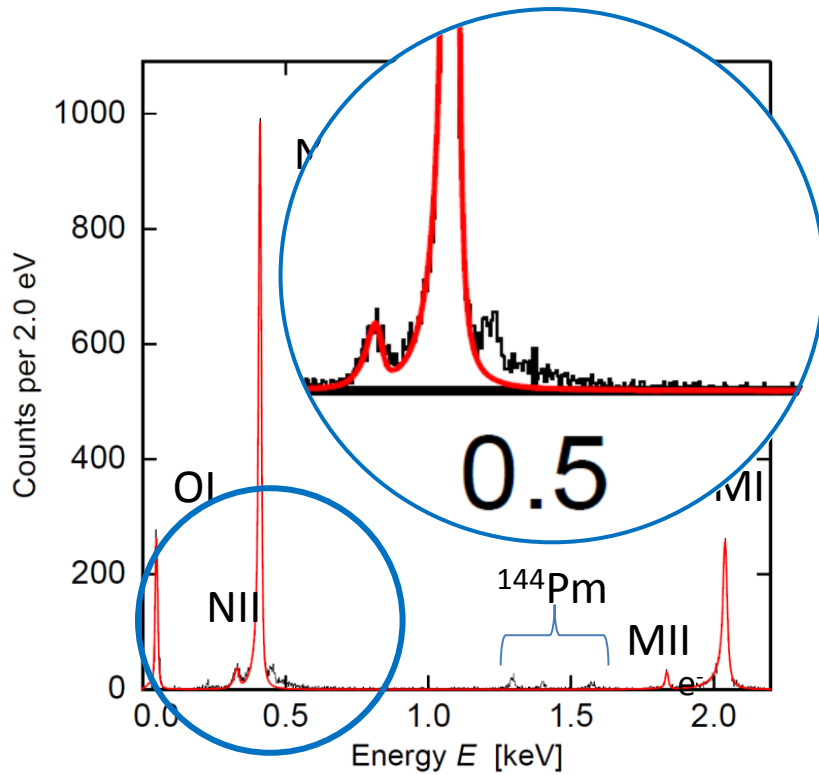


Two-holes excited states: shake-up

- A. Faessler et al.
J. Phys. G **42** (2015) 015108
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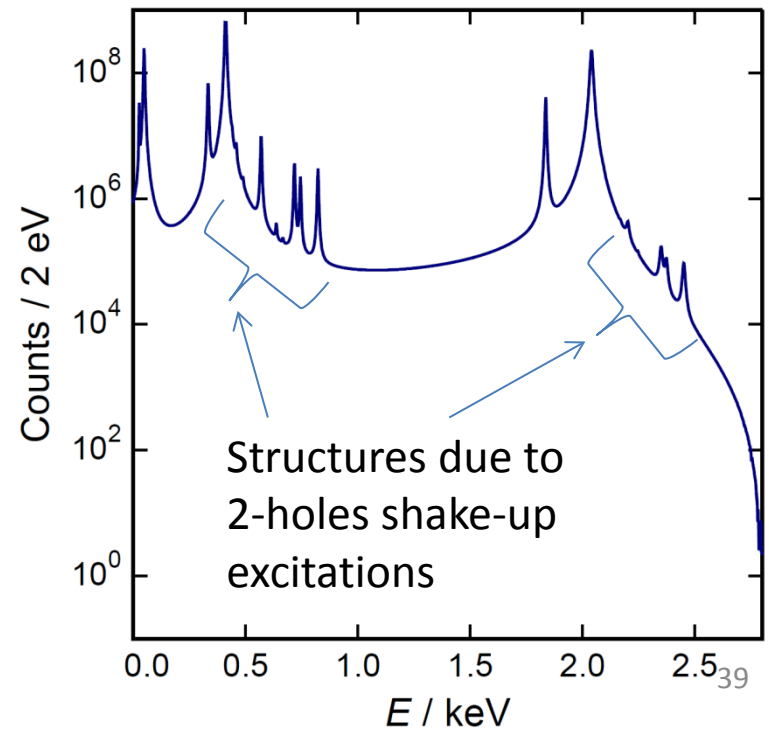


Characterisation of spectral shape

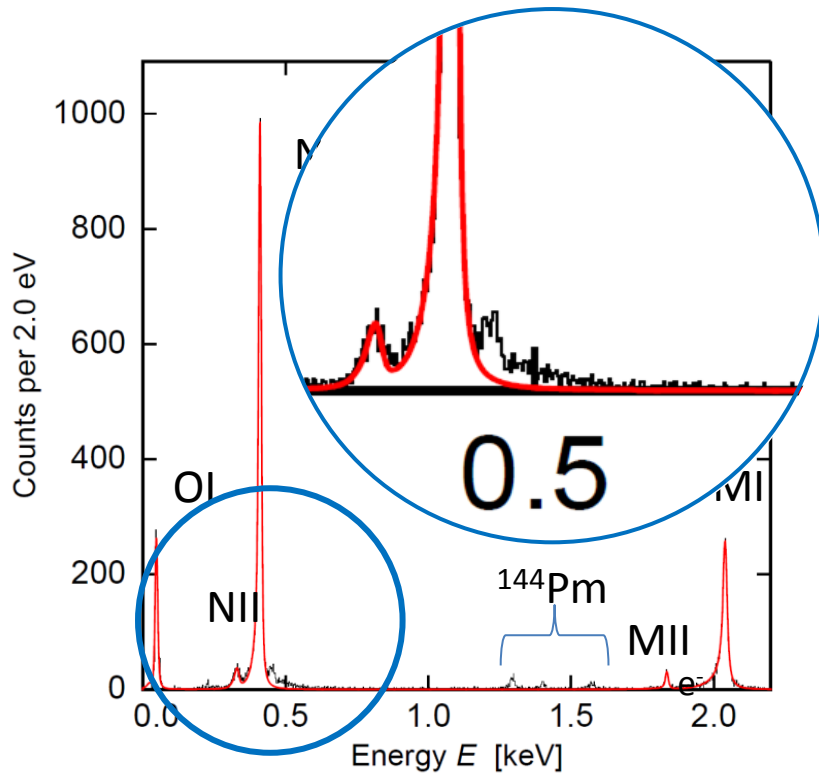


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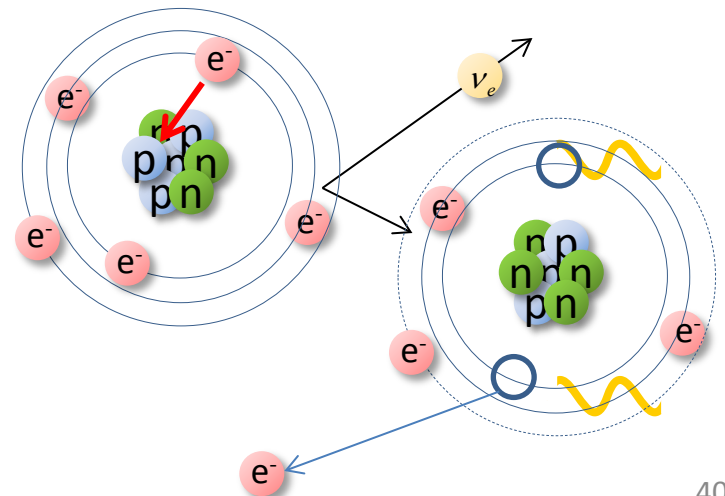


Characterisation of spectral shape



Two-holes excited states: shake-up
shake-off

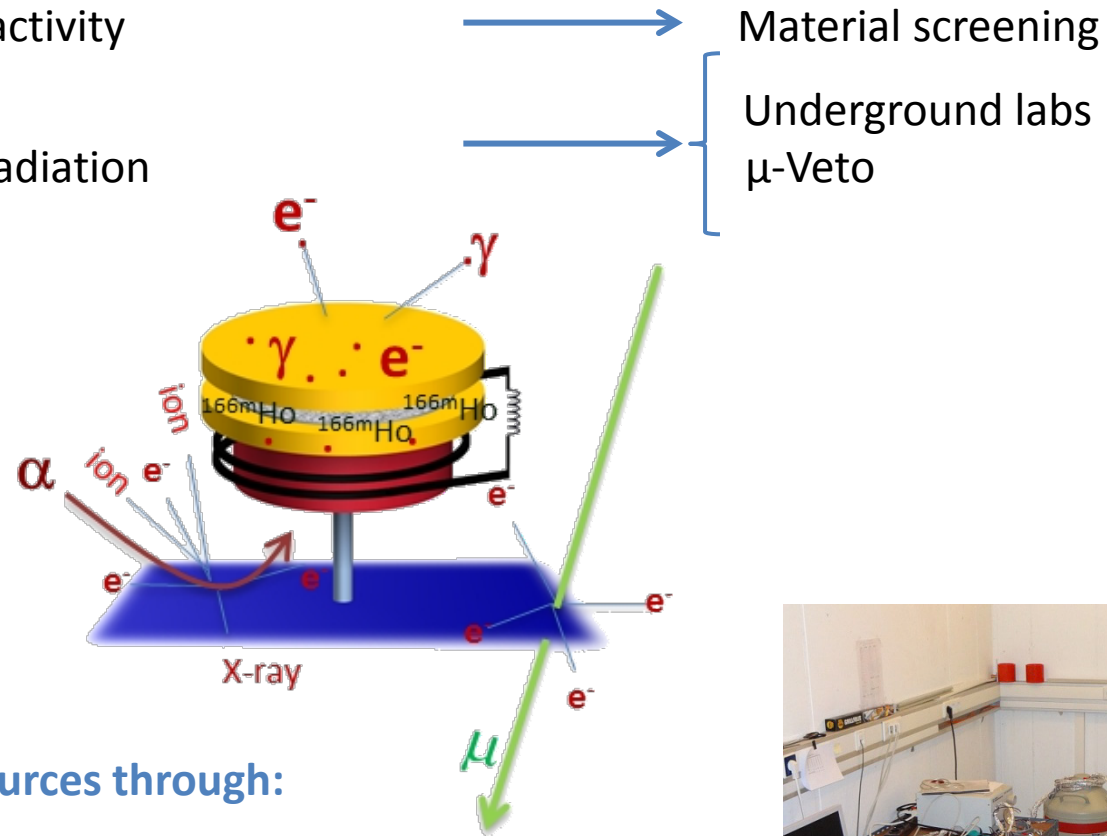
- A. Faessler et al.
J. Phys. G **42** (2015) 015108
- R. G. H. Robertson
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- A. Faessler et al.
Phys. Rev. C **91**, 045505 (2015)
- A. Faessler et al.
Phys. Rev. C **91**, 064302 (2015)
- A. De Rujula et al.
[arXiv:1601.04990v1 \[hep-ph\]](https://arxiv.org/abs/1601.04990v1) 19 Jan 2016



Background

Background sources:

- Radioactivity in the detector
- Environmental radioactivity
- Cosmic rays
Induced secondary radiation



Study of background sources through:

- Monte Carlo simulations
- Dedicated experiments

Screening facilities

- Uni-Tübingen
- Felsenkeller

