

The direct neutrino mass search with KATRIN



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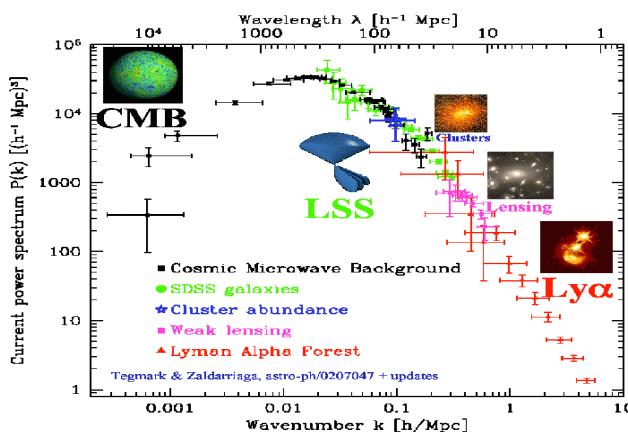


Introduction
The KArlsruhe TRitium Neutrino experiment
KATRIN
Commissioning data
Neutrino mass sensitivity
Conclusions

Three complementary ways to the absolute neutrino mass scale

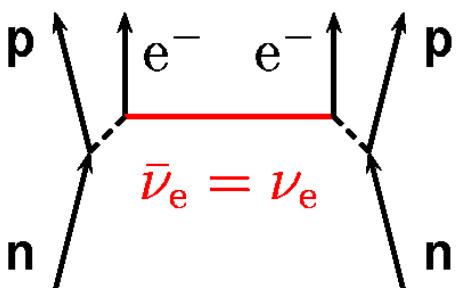
1) Cosmology

very sensitive, but model dependent
 compares power at different scales
 current sensitivity: $\sum m(\nu_i) \approx 0.23 \text{ eV}$



2) Search for $0\nu\beta\beta$

Sensitive to Majorana neutrinos
 Upper limits by EXO-200, KamLAND-Zen, GERDA, CUORE



3) Direct neutrino mass determination:

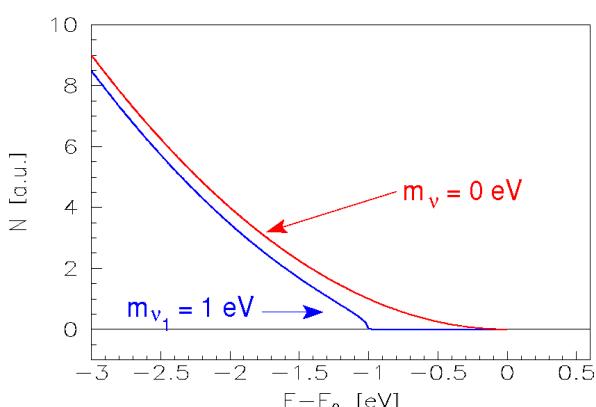
No further assumptions needed, use $E^2 = p^2c^2 + m^2c^4 \Rightarrow m^2(\nu) \text{ is observable mostly}$
Time-of-flight measurements (ν from supernova)

SN1987a (large Magellan cloud) $\Rightarrow m(\nu_e) < 5.7 \text{ eV}$

Kinematics of weak decays / beta decays

measure charged decay prod., E-, p-conservation

- β -decay searches for $m(\nu_e)$
 - tritium, ^{187}Re β -spectrum
 - ^{163}Ho electron capture (EC)



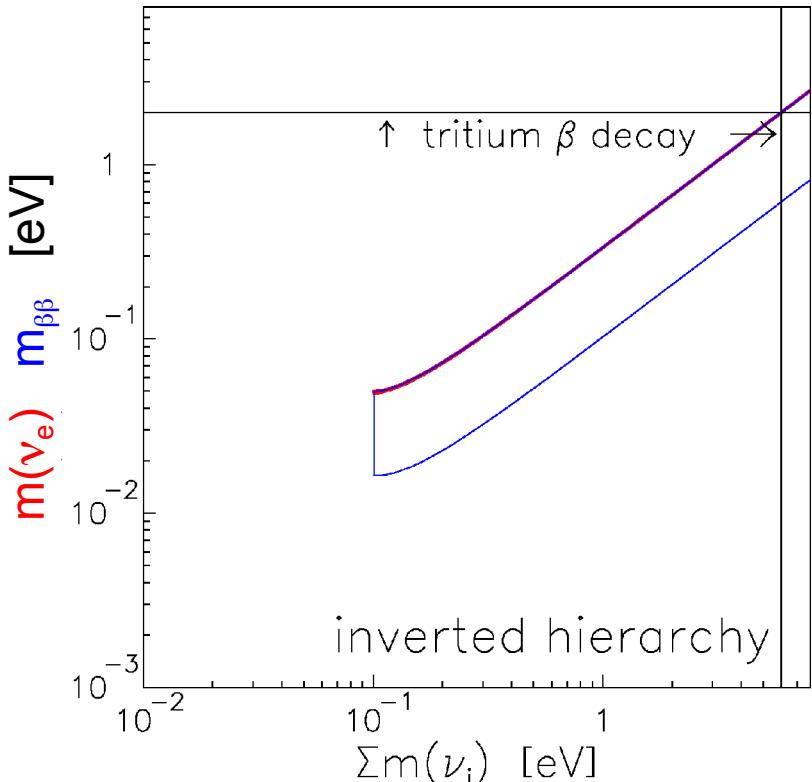
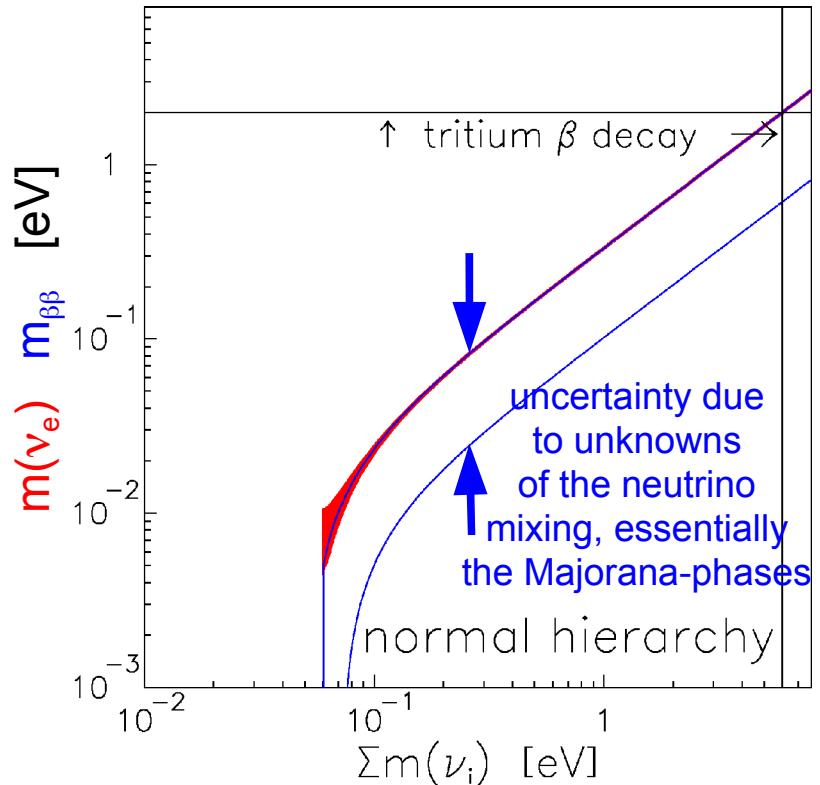
Comparison of the different approaches to the neutrino mass

Direct kinematic measurement: $m^2(\nu_e) = \sum |U_{ei}|^2 m^2(\nu_i)$ (incoherent)

Neutrinoless double β decay: $m_{\beta\beta}(\nu) = |\sum |U_{ei}|^2 e^{i\alpha(i)} m(\nu_i)|$ (coherent)

if no other particle is exchanged (e.g. R-violating SUSY)

without additional uncertainties of nuclear matrix elements M and quenching factor g_A



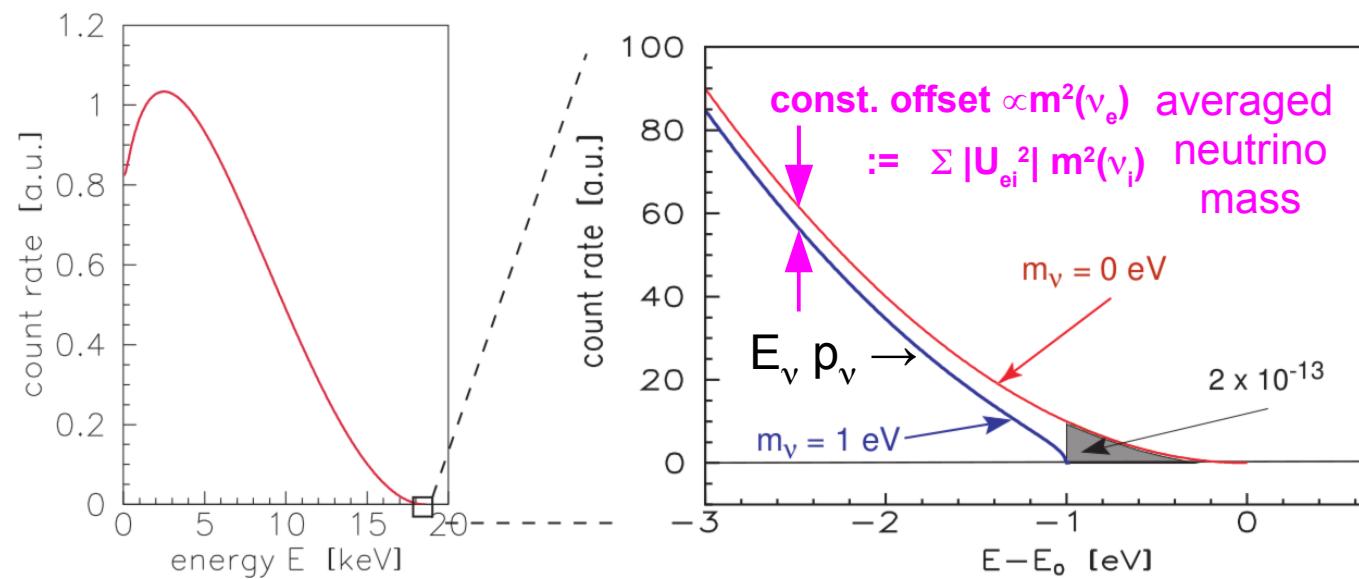
⇒ absolute scale/cosmological relevant neutrino mass in the lab by single β decay

Direct determination of $m(\nu_e)$ from β -decay (and EC)

$$\beta: \frac{dN}{dE} = K F(E, Z) \underbrace{p}_{p_e} \underbrace{E_{\text{tot}}}_{E_e} \underbrace{(E_0 - E_e)}_{E_\nu} \underbrace{\sum |U_{ei}|^2 \sqrt{(E_0 - E_e)^2 - m(\nu_i)^2}}_{p_\nu}$$

essentially phase space:

with “electron neutrino mass”: $m(\nu_e)^2 := \sum |U_{ei}|^2 m(\nu_i)^2$, complementary to $0\nu\beta\beta$ & cosmology
(modified by electronic final states, recoil corrections, radiative corrections)



$m(\nu) < 2$ eV (Mainz, Troitsk)
 ν do not solve DM problem

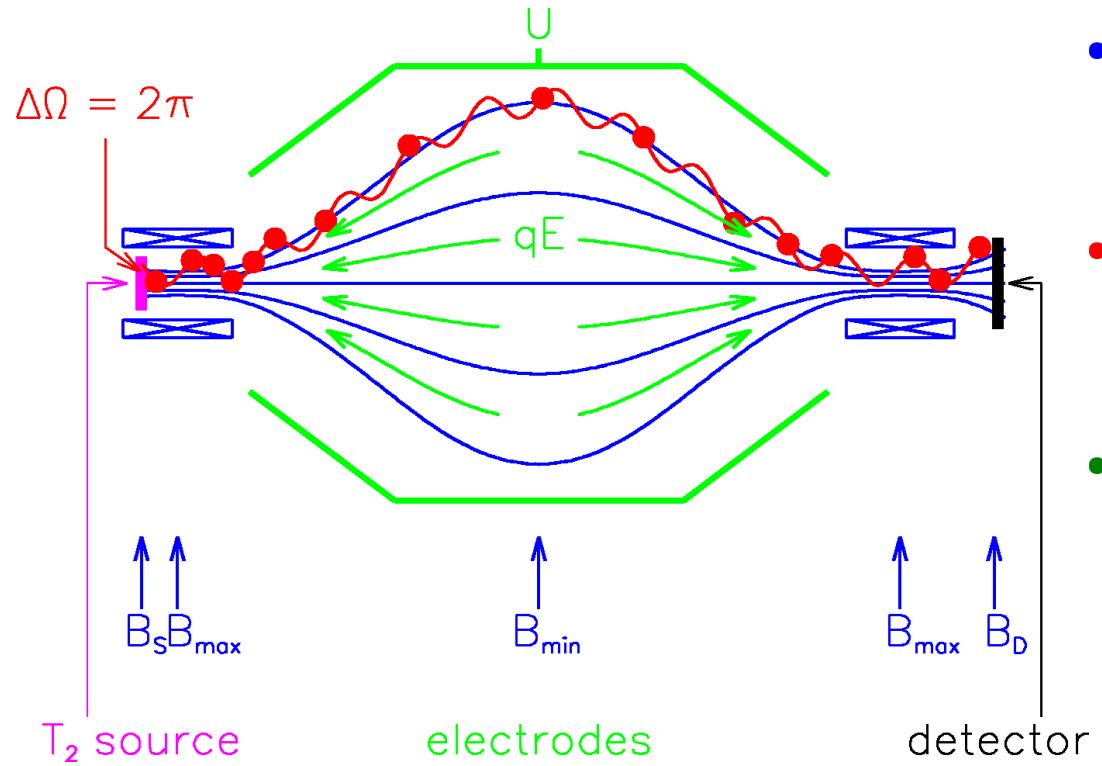


Need: **low endpoint energy**
very high energy resolution &
very high luminosity &
very low background

\Rightarrow **Tritium ${}^3\text{H}$ (${}^{187}\text{Re}$, ${}^{163}\text{Ho}$)**

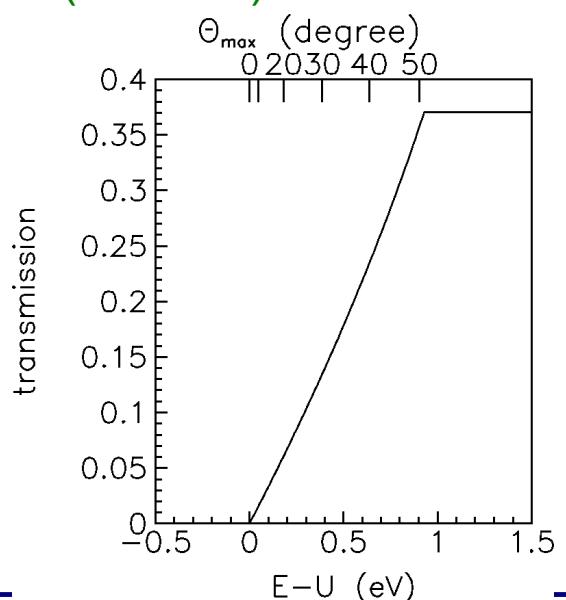
\Rightarrow **MAC-E-Filter**
(or bolometer for ${}^{187}\text{Re}$, ${}^{163}\text{Ho}$)

The classical way: Tritium β -spectroscopy with a MAC-E-Filter

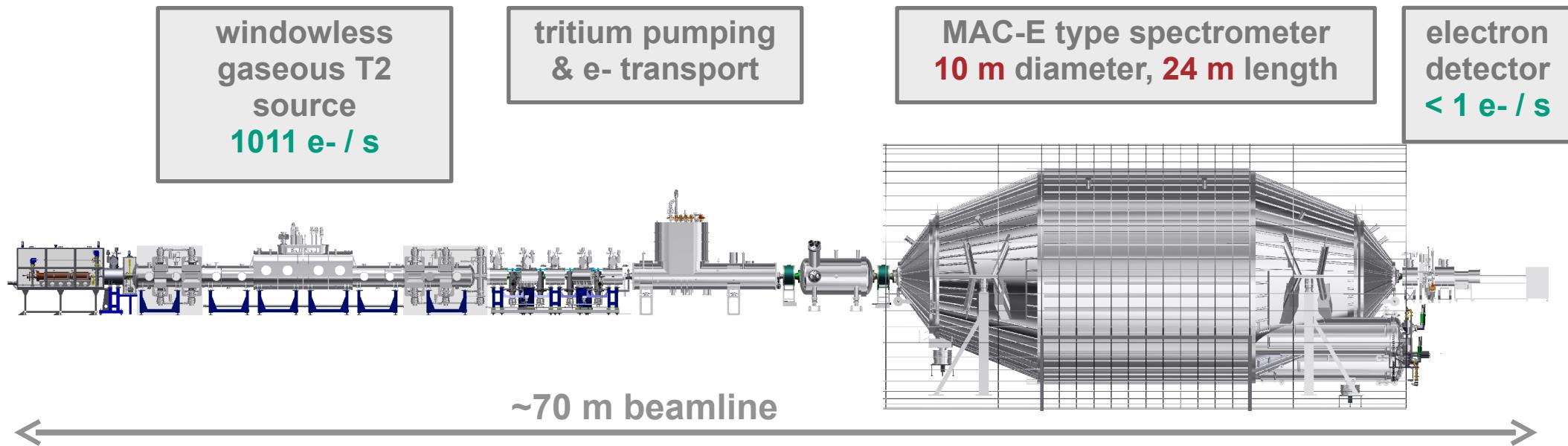


⇒ sharp integrating transmission function without tails →

Magnetic Adiabatic Collimation + Electrostatic Filter
(A. Picard et al., Nucl. Instr. Meth. 63 (1992) 345)



The KATRIN experiment



Sensitivity on $m(\nu_e)$:
 $2 \text{ eV} \rightarrow 200 \text{ meV}$



KATRIN at Karlsruhe Institute for Technology
Int. Collaboration: 20 institutions from 6 countries

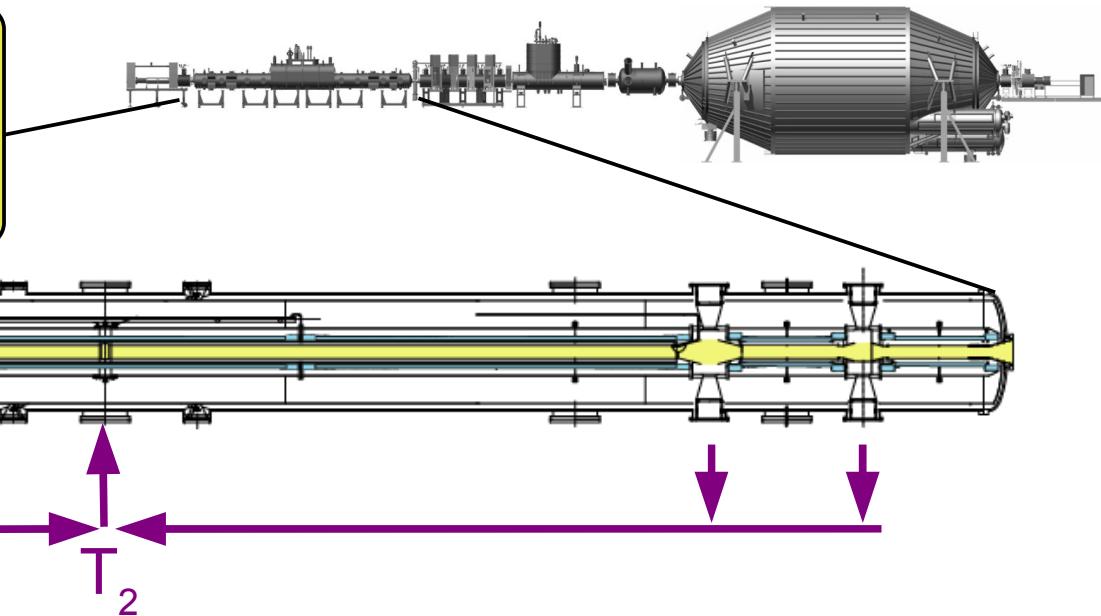
Molecular Windowless Gaseous Tritium Source WGTS



per mill stability source strength request:

$$dN/dt \sim f_T \cdot N / \tau \sim n = f_T \cdot p V / RT$$

tritium fraction f_T & ideal gas law



WGTS: tube in long superconducting solenoids
 \varnothing 9cm, length: 10m, $T = 30$ K

Tritium recirculation (and purification)

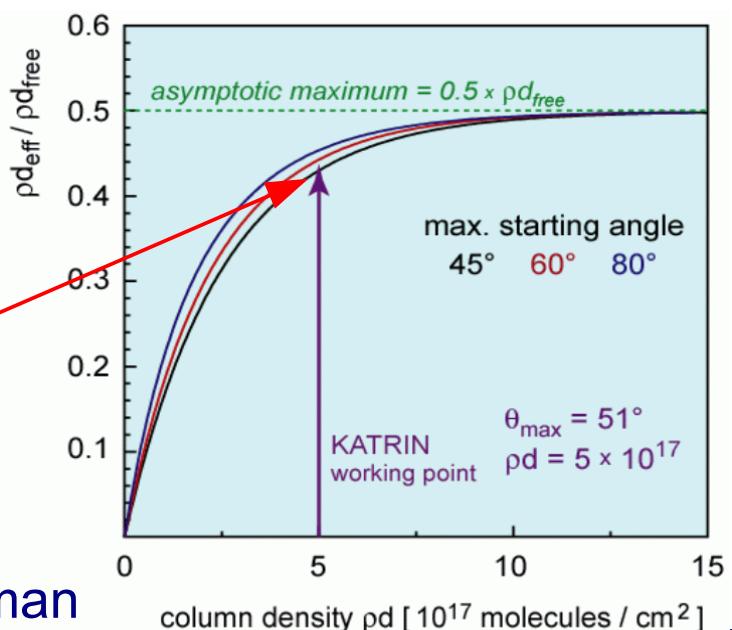
$$p_{\text{inj}} = 0.003 \text{ mbar}, q_{\text{inj}} = 4.7 \text{ Ci/s}$$

allows to measure with near to maximum count rate using

$$pd = 5 \cdot 10^{17} / \text{cm}^2$$

with small systematics

check column density by e-gun, T_2 purity by laser Raman



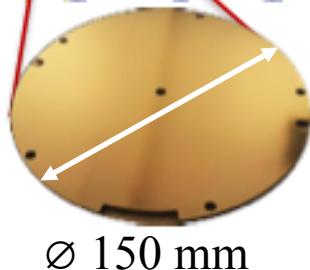
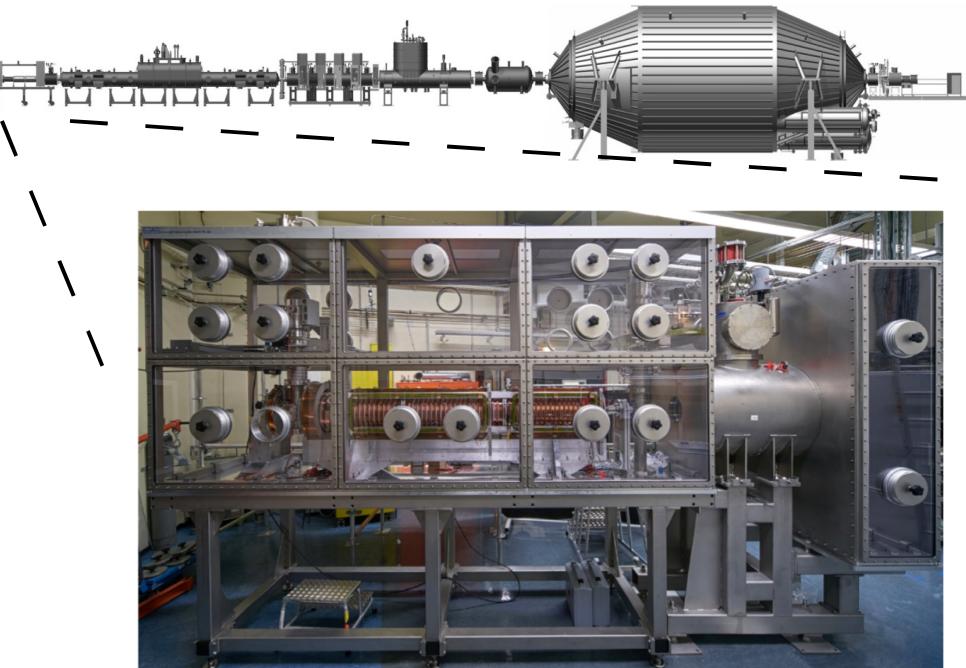
Molecular Windowless Gaseous Tritium Source WGTS



Calibration and monitoring rear system: controlling and studying systematics

Essential for diagnostics of tritium source
& spectrometer transmission

- **photo-electron gun:**
spectrometer transmission
column density & energy losses in source
- **rear wall:** definition of source potential,
neutralization of tritium plasma



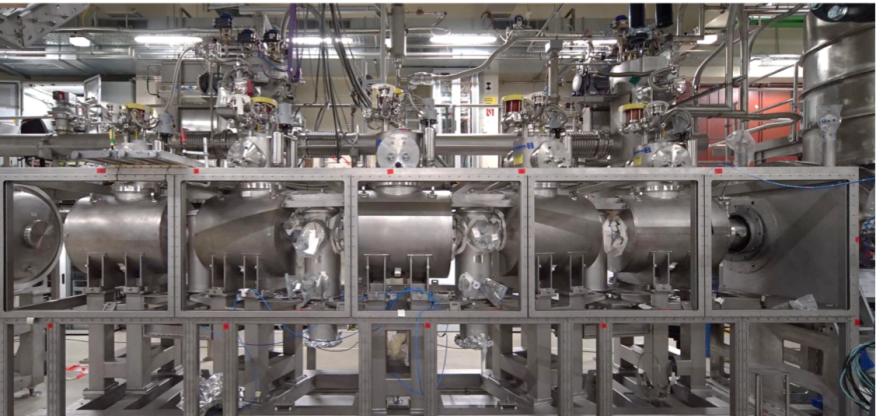
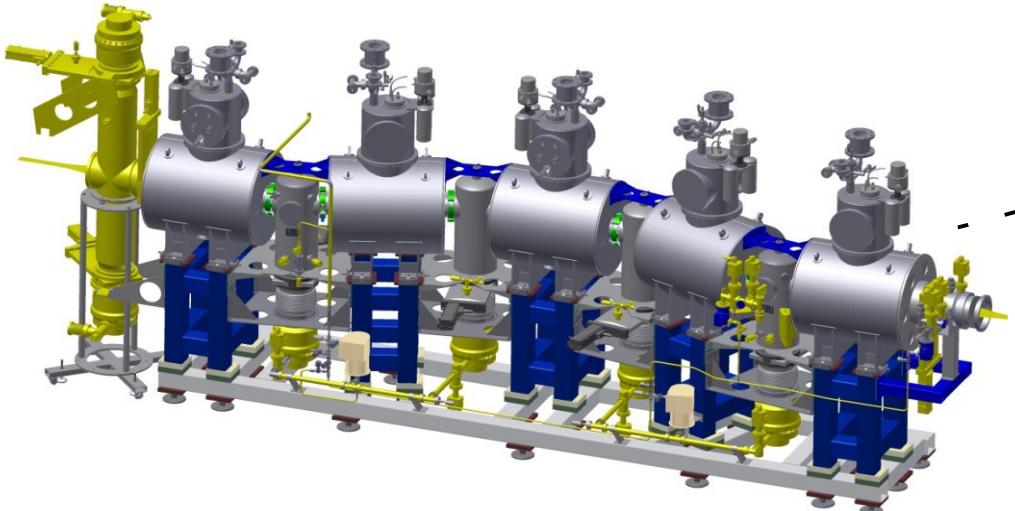
Rear Wall: Au surface creates stable and homogeneous electrostatic potential ($\sim 10\text{-}20 \text{ mV}$) in the source plasma



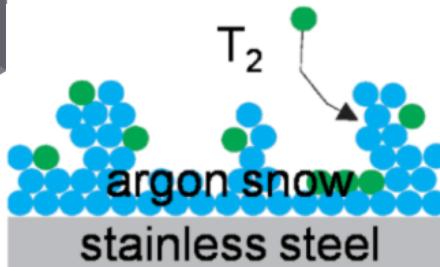
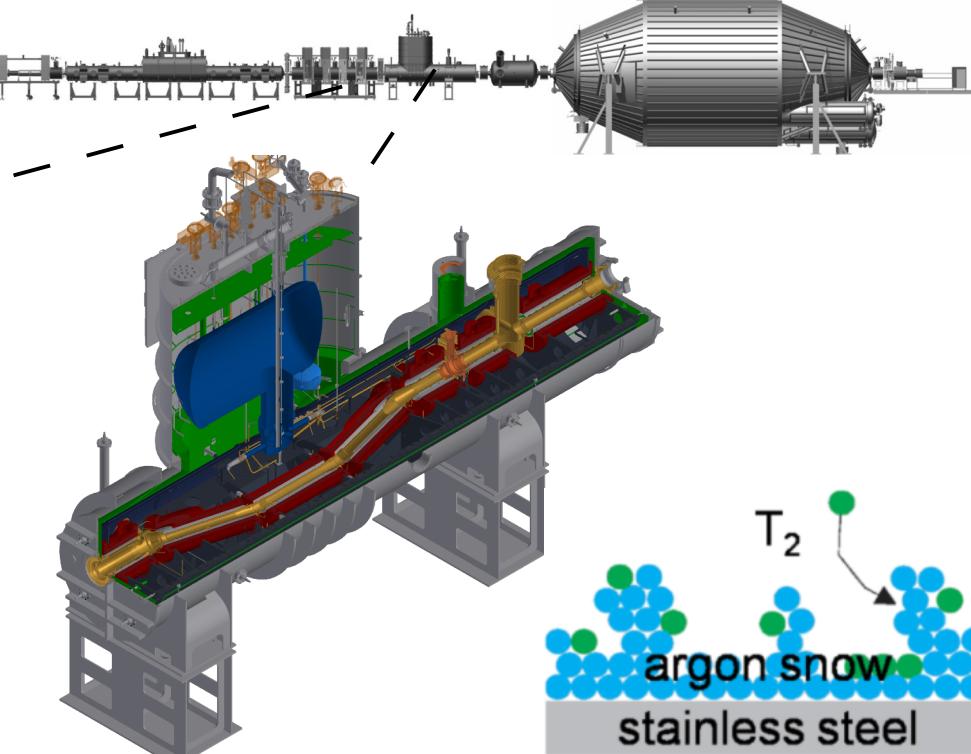
- X-ray detectors:

online monitoring of tritium β-decay activity via X-rays (BIXS)

Differential and cryo pumping sections: supression of T_2 by 10^{14} (incl. WGTS)



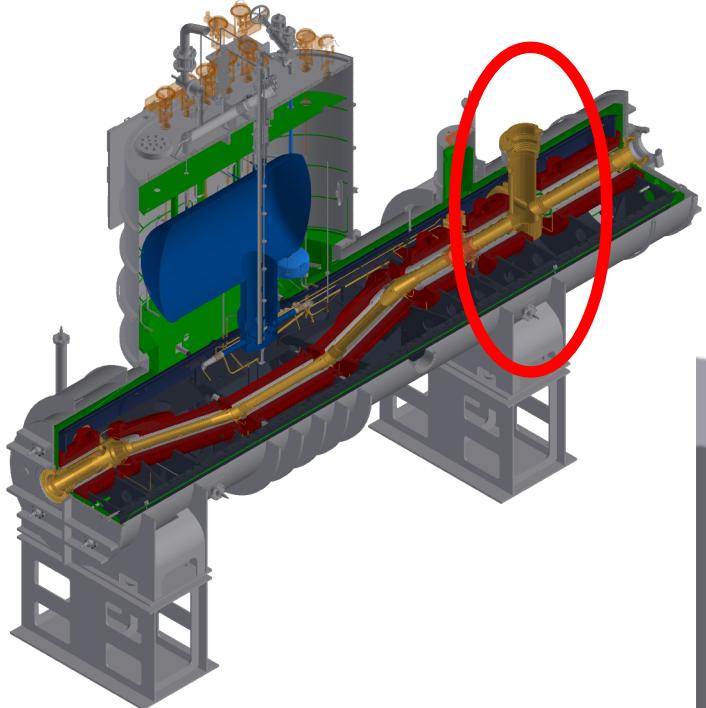
- active pumping: 4 TMPs
- Tritium retention: 10^5
- magnetic field: 5.6 T
- **Ion monitoring by FTICR and ion manipulation by dipole and monopole electrodes inside**



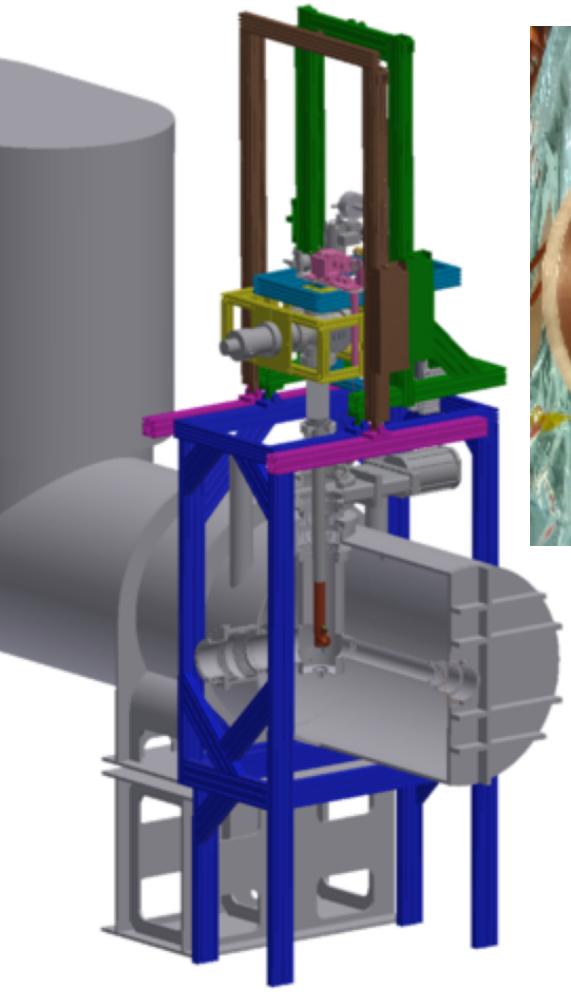
- based on cryo-sorption at Ar snow at 3-4 K
- Tritium retention: $>10^7$
- magnetic field: 5.6 T



Monitoring and calibration instrumentation of the CPS



Electron rate monitor
scanning small SD or PIN diode

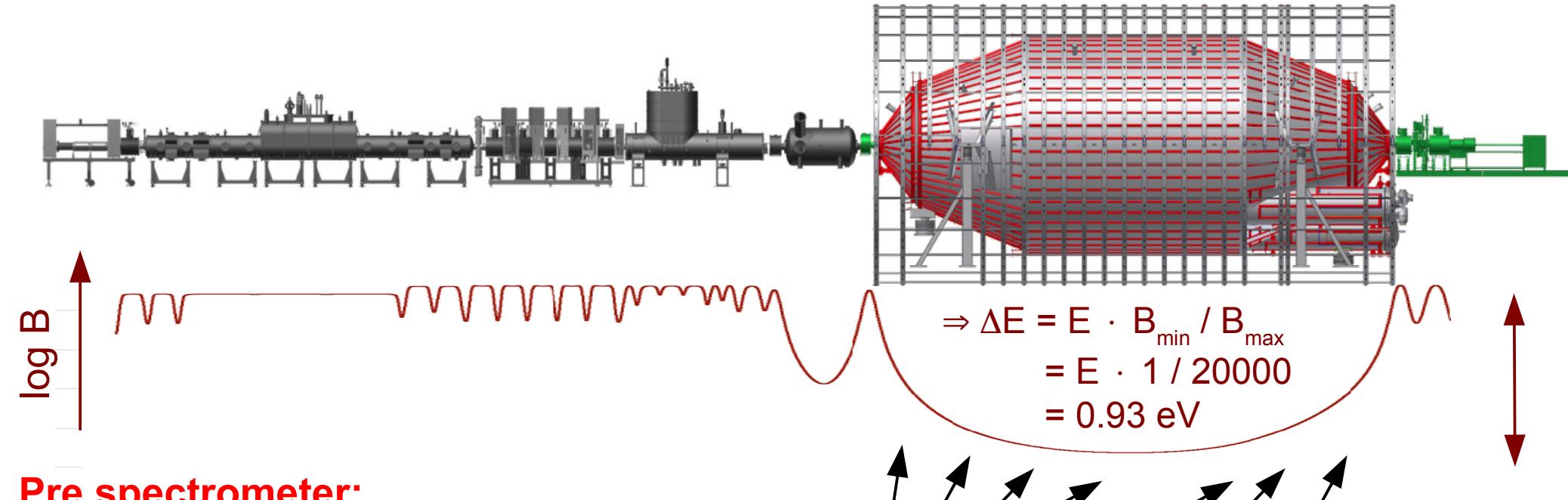


Condensed $^{83\text{m}}\text{Kr}$ conversion electron source

for energy calibration and studies of transmission properties
 HOPG @ $T=25\text{K}$, UHV, on HV, can scan full flux tube
 surface control: heating & laser ablation, laser ellipsometry



KATRIN spectrometers of MAC-E-Filter type



Pre spectrometer:

- successful tests & developments of new concepts

Main spectrometer:

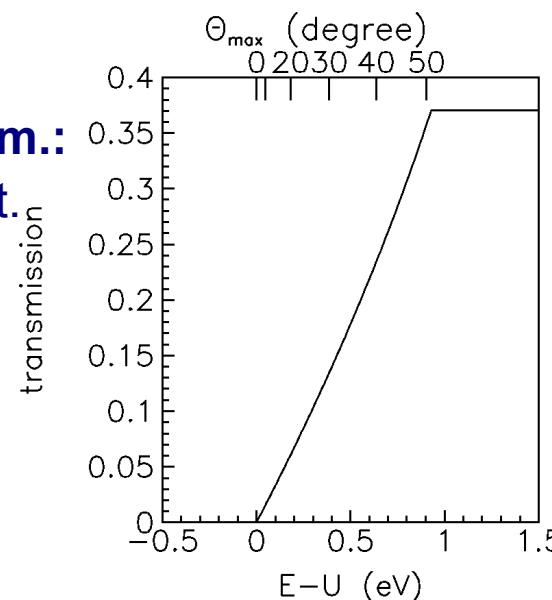
- huge size: 10m diameter, 24m length
1240 m³ volume, 690 m² inner surface
- ultra-high vacuum: $p = O(10^{-11} \text{ mbar})$
- ultra-high energy resolution: $\Delta E = 0.93 \text{ eV}$
- vacuum vessel on precise high voltage (ppm precision)

adiabatic transform.:

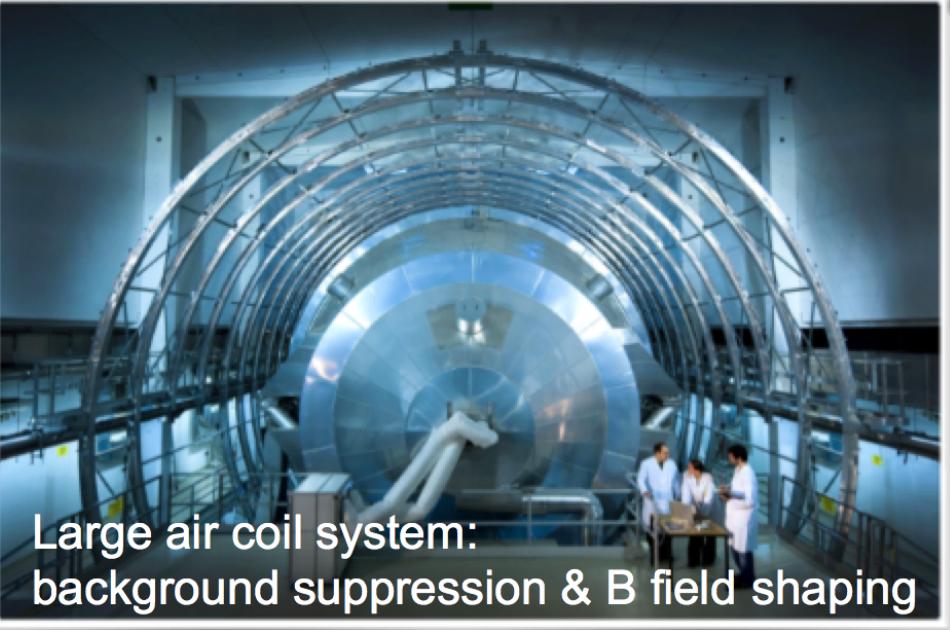
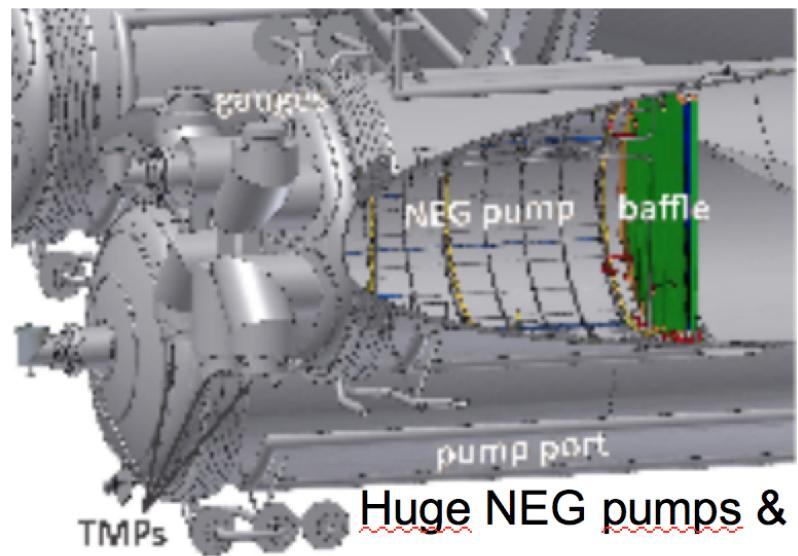
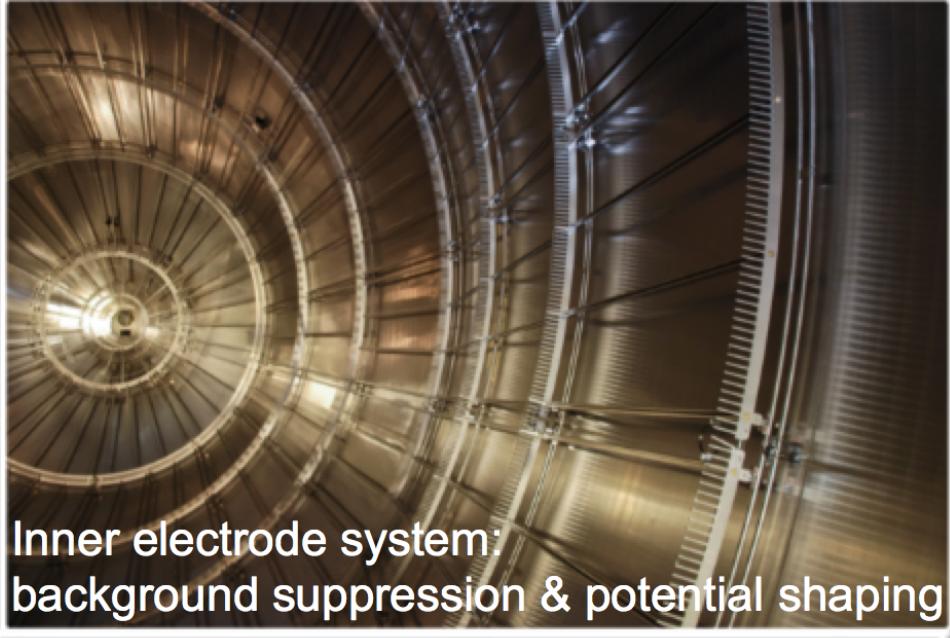
$$\mu = E_{\perp} / B = \text{const.}$$

\Rightarrow parallel e^- beam

$$\Delta E/E = B_{\min} / B_{\max}$$



KATRIN main spectrometer



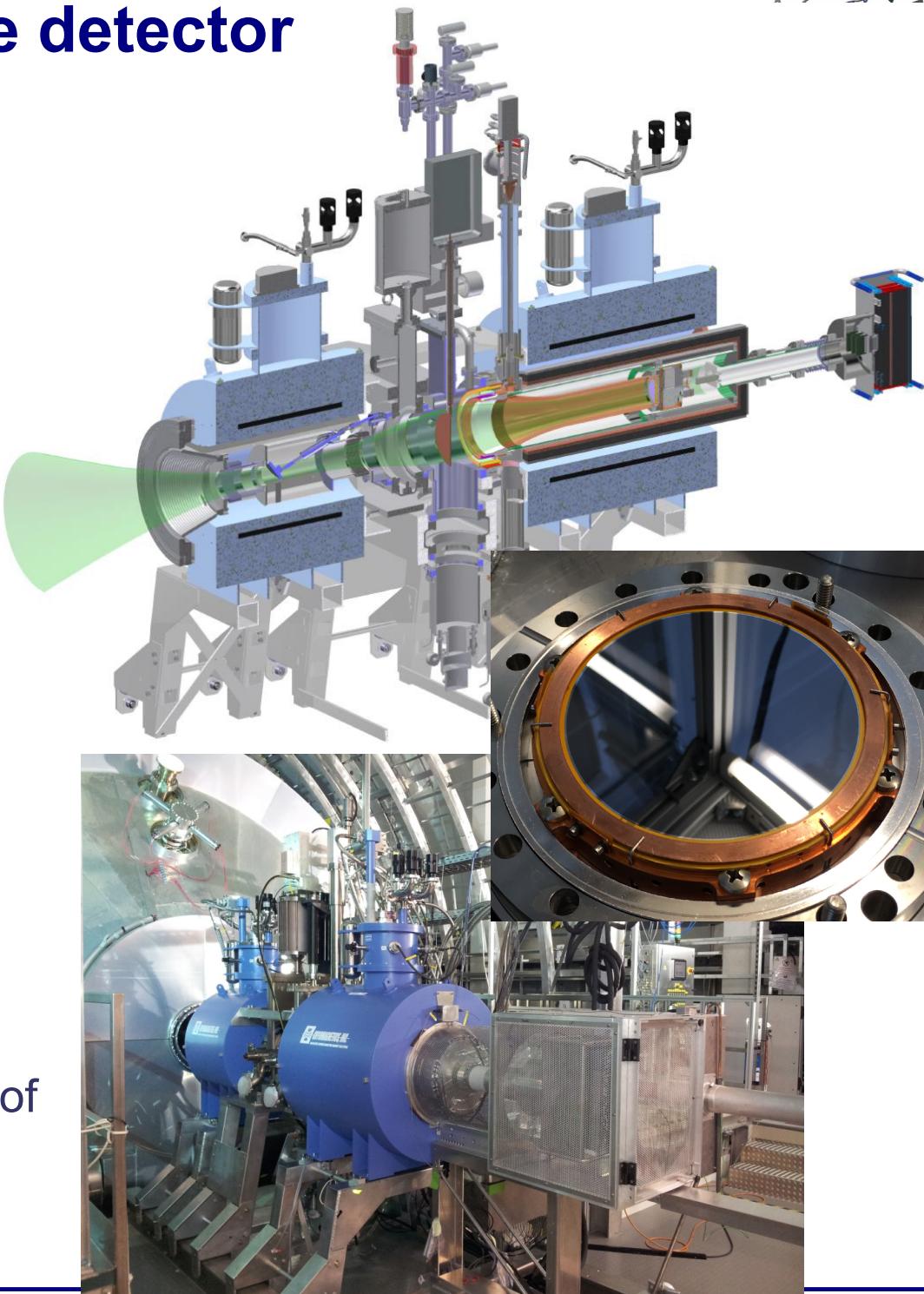
The detector

Requirements

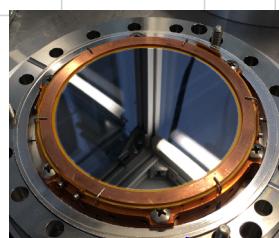
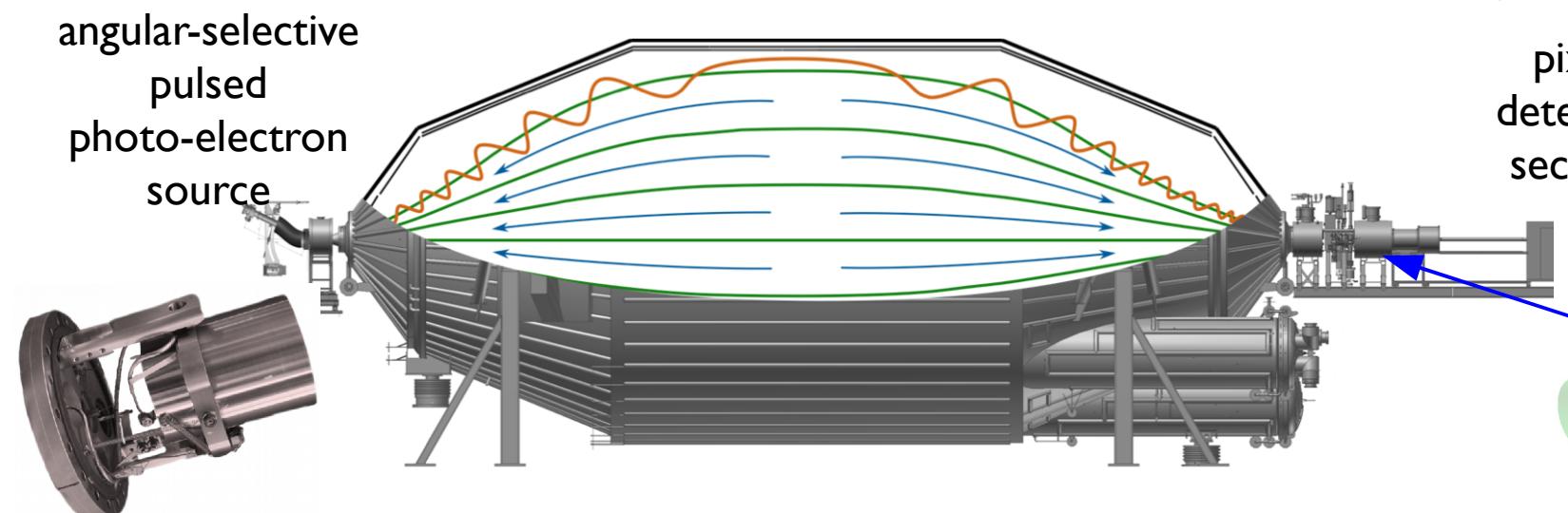
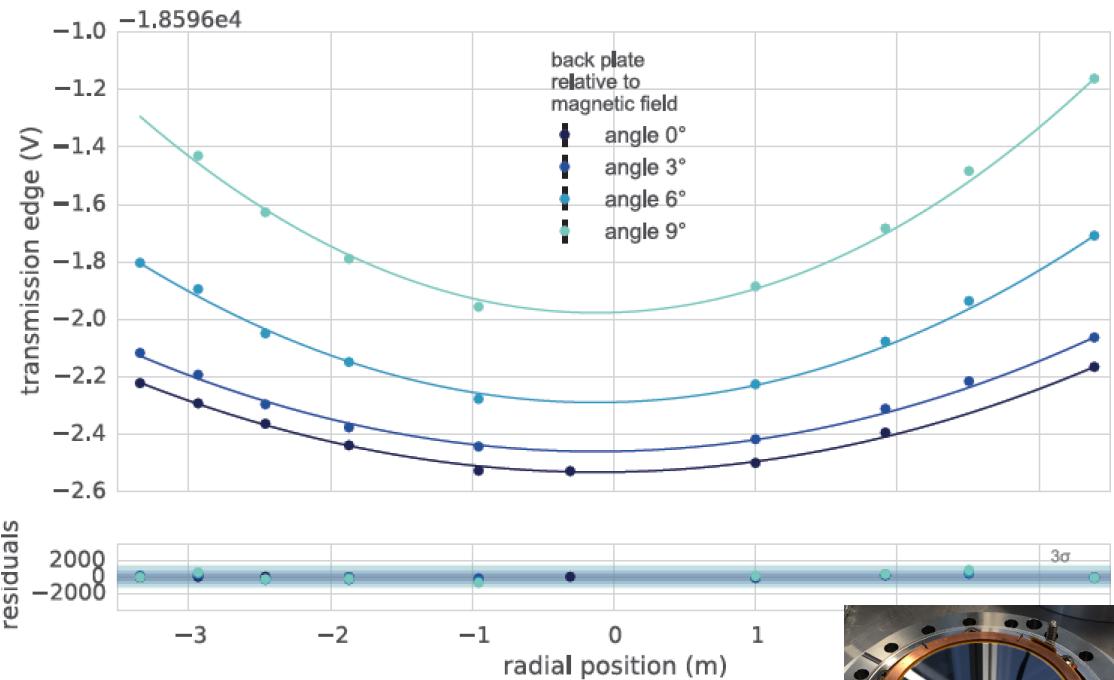
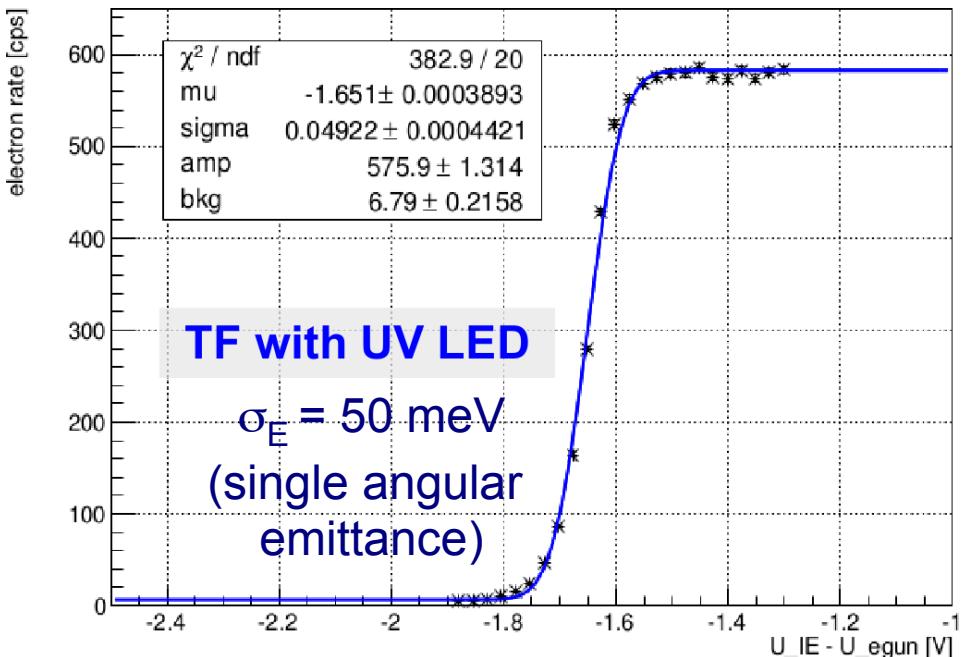
- detection of β -electrons (mHz to kHz)
- high efficiency (> 90%)
- low background (< 1 mHz)
(passive and active shielding)
- good energy resolution (< 1 keV)

Properties

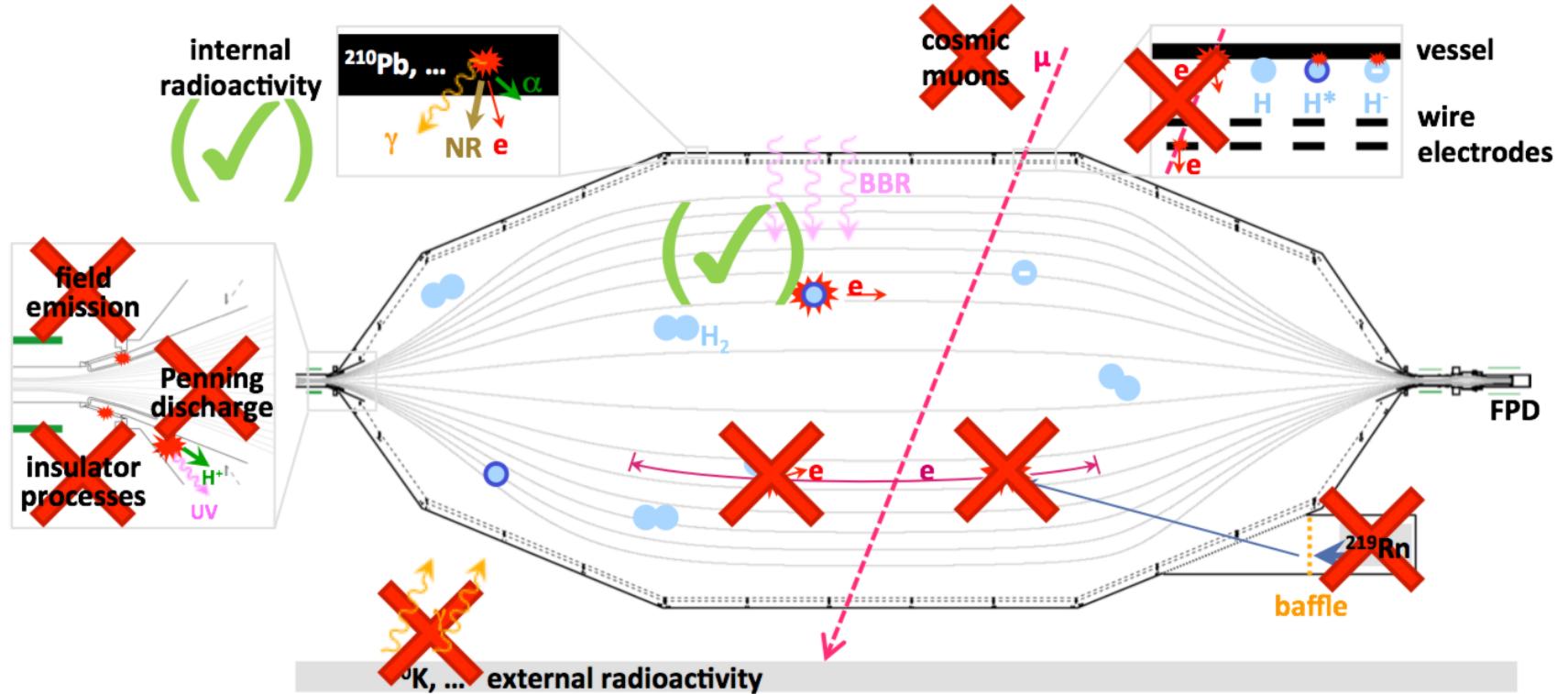
- 90 mm Ø Si PIN diode
- thin entry window (50nm)
- detector magnet 3 - 6 T
- post acceleration (30kV)
(to lower background in signal region)
- segmented wafer (148 pixels)
 - record azimuthal and radial profile of the flux tube
 - investigate systematic effects
 - compensate field inhomogeneities



Commissioning of main spectrometer ($\Delta E = 0.93$ eV) and detector



Background sources at KATRIN: detailed understanding, but ...



- 8 sources of background investigated and understood
- 7 out of 8 avoided or actively eliminated by
 - fine-shaping of special electrodes
 - symmetric magnetic fields
 - LN₂-cooled baffles (cold traps)
 - wire electrode grids

- 1 out of 8 remaining:

caused by ²¹⁰Pb on spectrometer walls (neutral H* atoms ionised by black-body radiation in spectrometer)

Background due to ionization of Rydberg atoms sputtered off by α decays

H* Rydberg atoms:

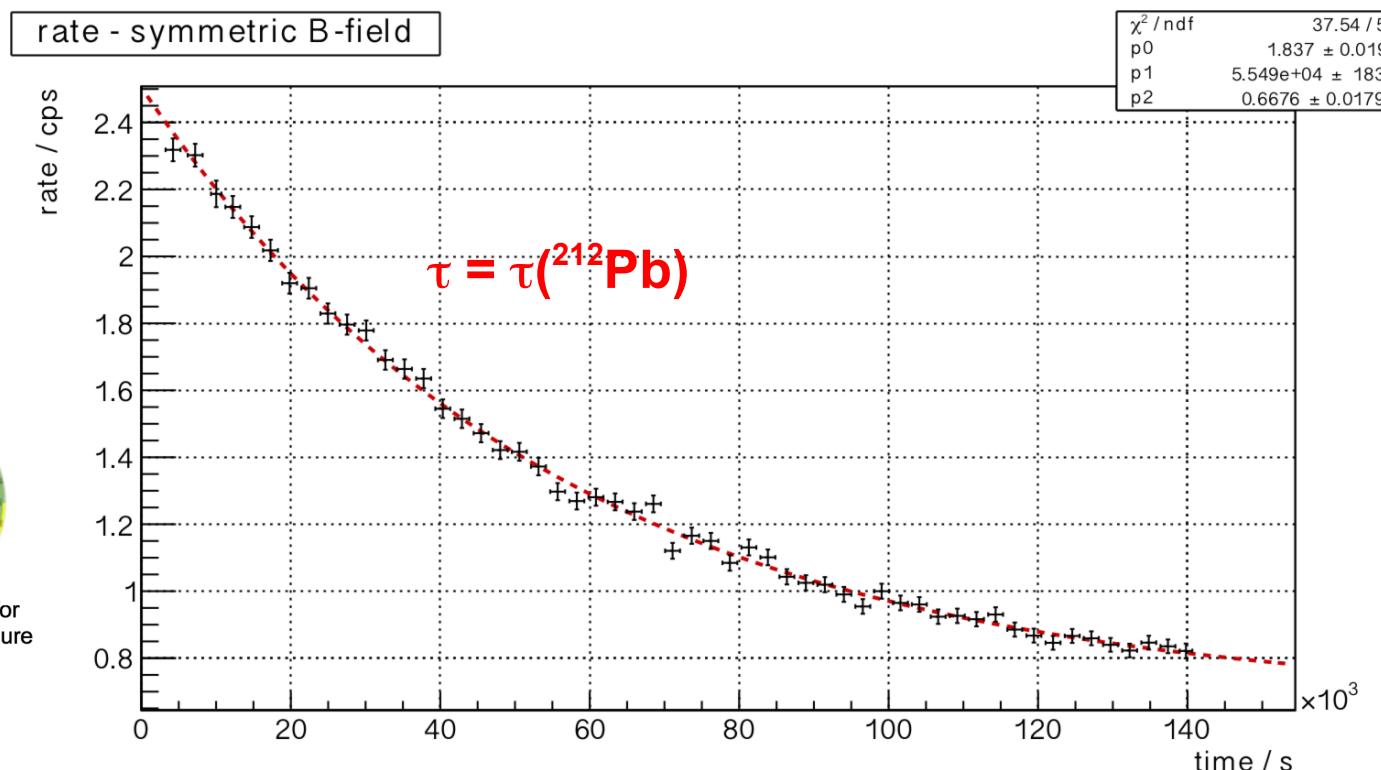
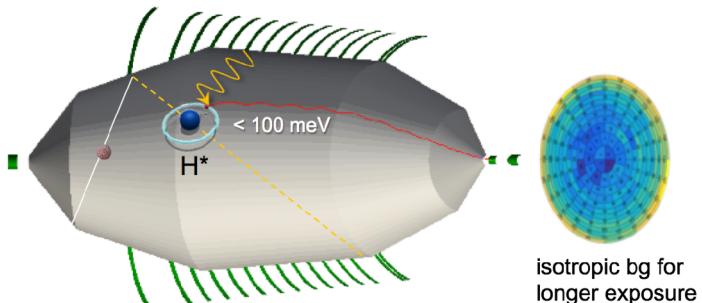
- desorbed from walls due to ^{206}Pb recoil ions from ^{210}Po decays
- non-trapped electrons on meV-scale
- bg-rate: ~ 0.5 cps

counter measures:

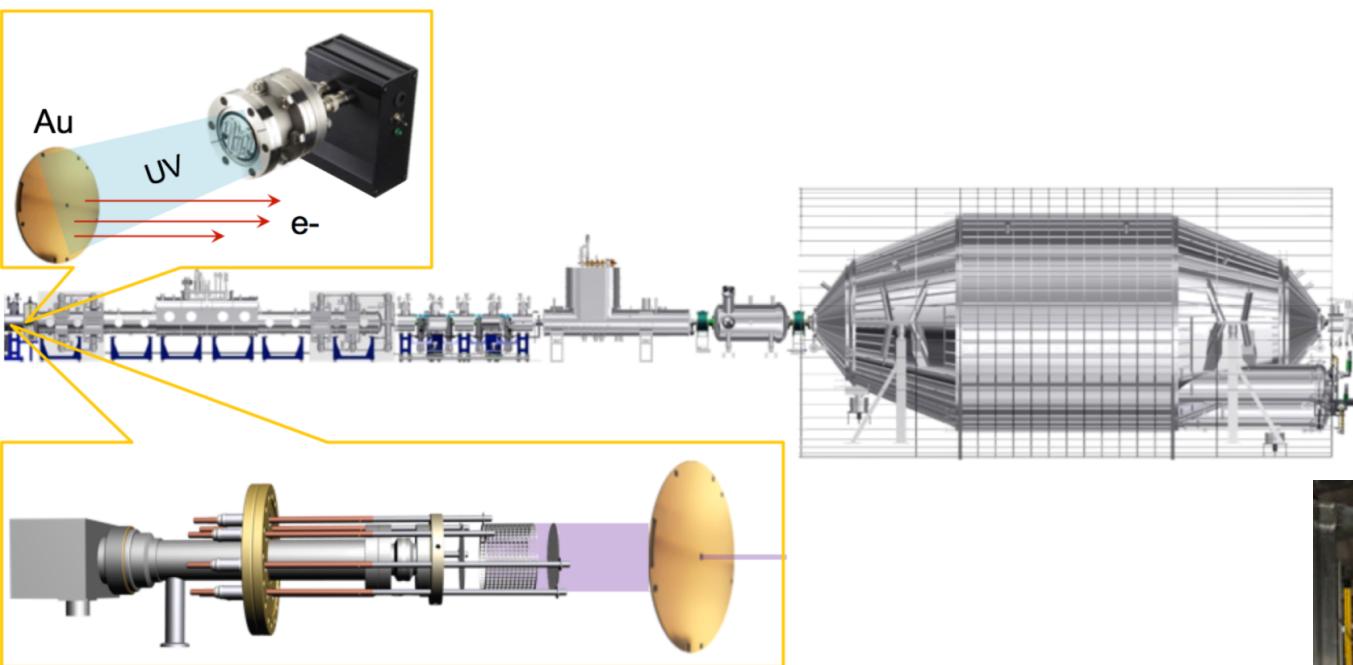
- reduce H-atom surface coverage:
 - a) extended bake-out phase: done
 - b) strong UV illumination source

Testing this hypothesis:

artificially contaminating the spectrometer with implanted short-living daughters of ^{220}Rn



Technical start of KATRIN: „1st light“, photo-electrons from rear wall & and ions



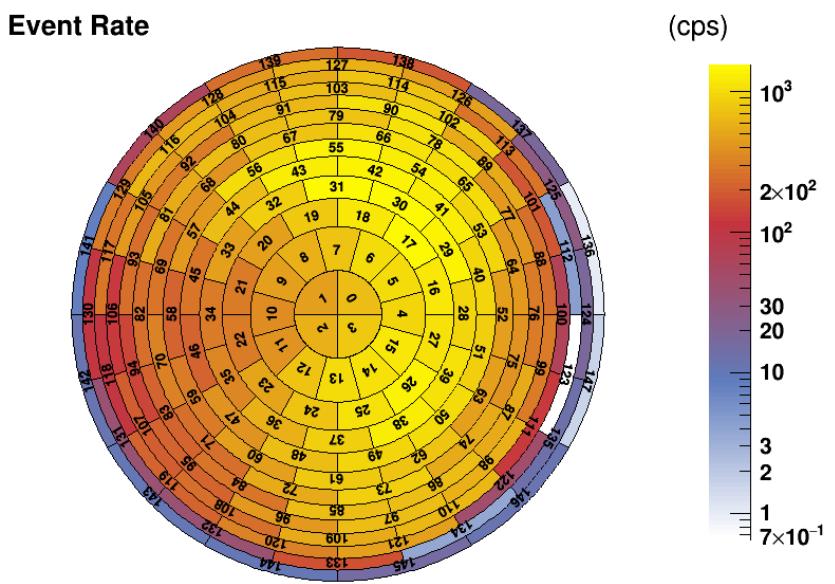
Testing whole 70m long beamline with electrons:

- alignment
- magn. steering of pencil beam

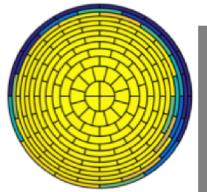
With ions:

- ion removal

no tritium yet

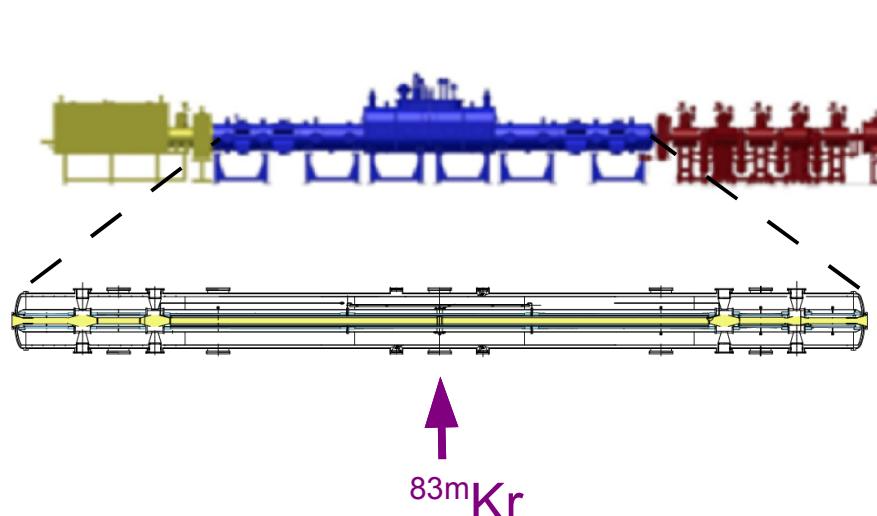
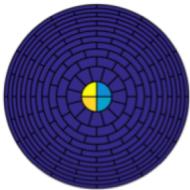


July 2017: calibration and commissioning campaign with all 3 ^{83m}Kr sources

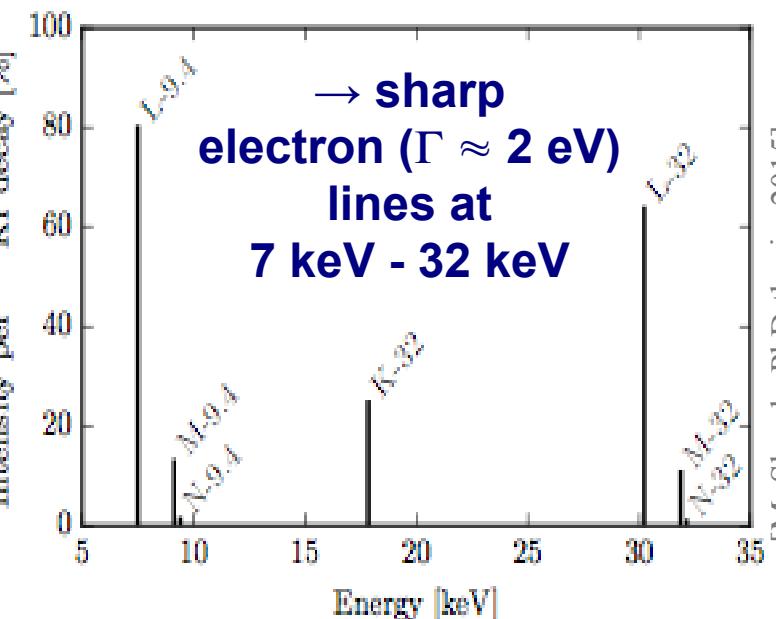
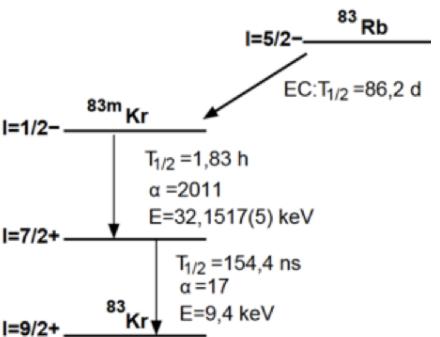


gaseous ^{83m}Kr source
decaying ^{83m}Kr atoms fill
whole WGTS (at 100 K)

condensed ^{83m}Kr source
point-like source
full flux tube scanable

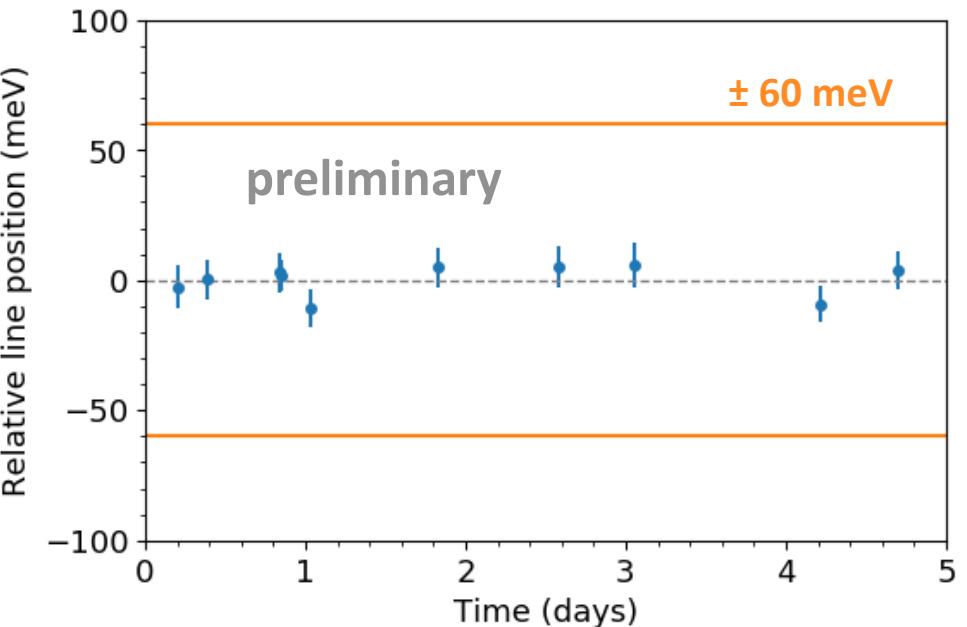
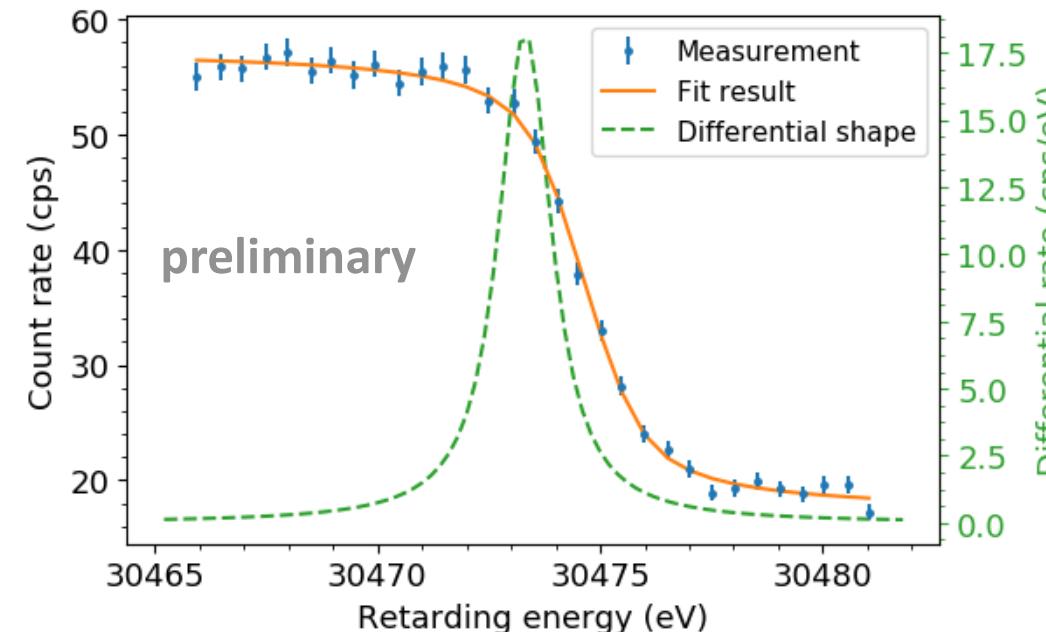


from 1 GBq ^{83}Rb source



Line scan & stability (gaseous Kr source GKrs)

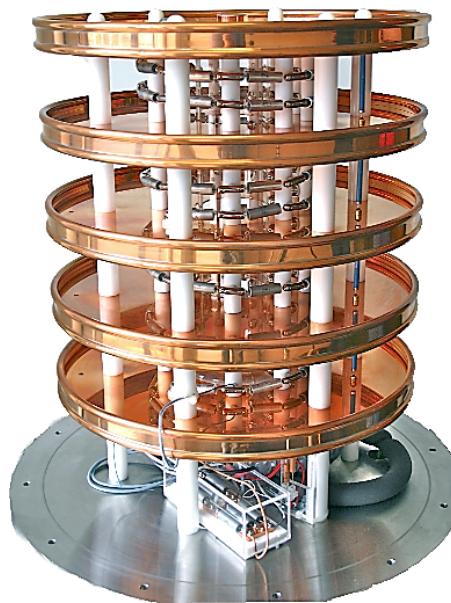
L3-32 line (30.47 keV, $\Gamma \sim 1.4$ eV)



- Just one example (one out of many line scans)
- Only central detector ring shown (x30 more statistics available)
- High-resolution scans of narrow N2,3-32 doublet (670 meV hyperfine splitting, sub-eV natural widths, background-free at 32 keV) currently being analyzed

Absolute energy scale calibration by difference of electron conversion lines

Measure retarding voltage
with ultra-high precision
HV divider:

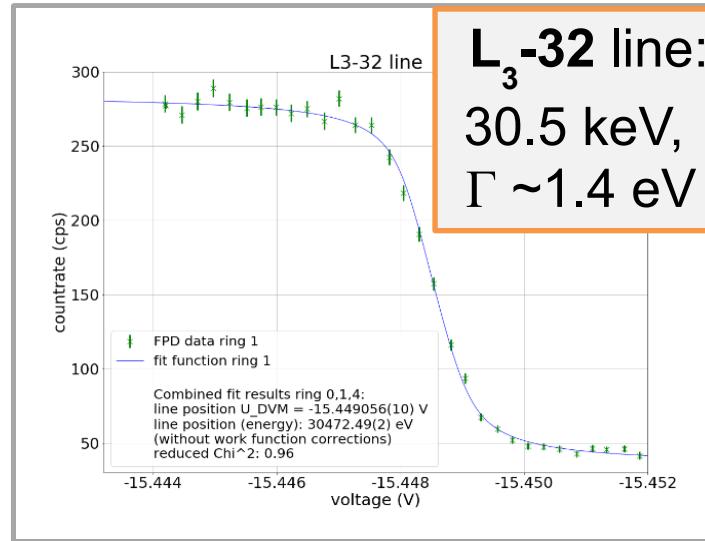


In cooperation with
German national
metrology institute

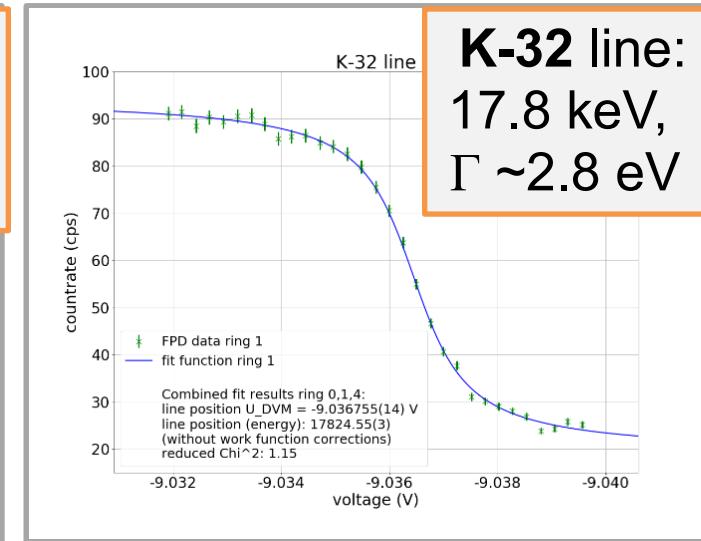


Last calibration at
PTB in 2013:
 $M = 1972.4531(20)$

Determine difference of conversion electron line positions:



L_3 -32 line:
30.5 keV,
 $\Gamma \sim 1.4$ eV



K-32 line:
17.8 keV,
 $\Gamma \sim 2.8$ eV

Energy of electrons:

$$E_{\text{kin}} = E_{\gamma} - E_{\text{binding}} + E_{\gamma}^{\text{rec}} - E_{\text{ion}}^{\text{rec}}$$

Energy of
 γ -transition

Binding
energy

Recoil energies

Transmission condition:

$$E_{\text{kin}} = -\Delta\Phi - q\Delta U + qU_{\text{spec}}$$

Work function difference
source - spectrometer

Potential decrease in
analyzing plane (≈ 2 V)

→ systematic effects ($\Delta\Phi, E_{\gamma}$) cancel out

$$q(U_{\text{spec}}^{L3-32} - U_{\text{spec}}^{K-32}) - q(\Delta U_{L3-32} - \Delta U_{K-32}) = E_{\text{binding}}^{K-32} - E_{\text{binding}}^{L3-32} + E_{\text{rec}}^{K-32} - E_{\text{rec}}^{L3-32}$$

$$M \cdot (U_{\text{DVM}}^{L3-32} - U_{\text{DVM}}^{K-32})$$

→ GKRS 2017:

$$M = 1972.453(10)$$

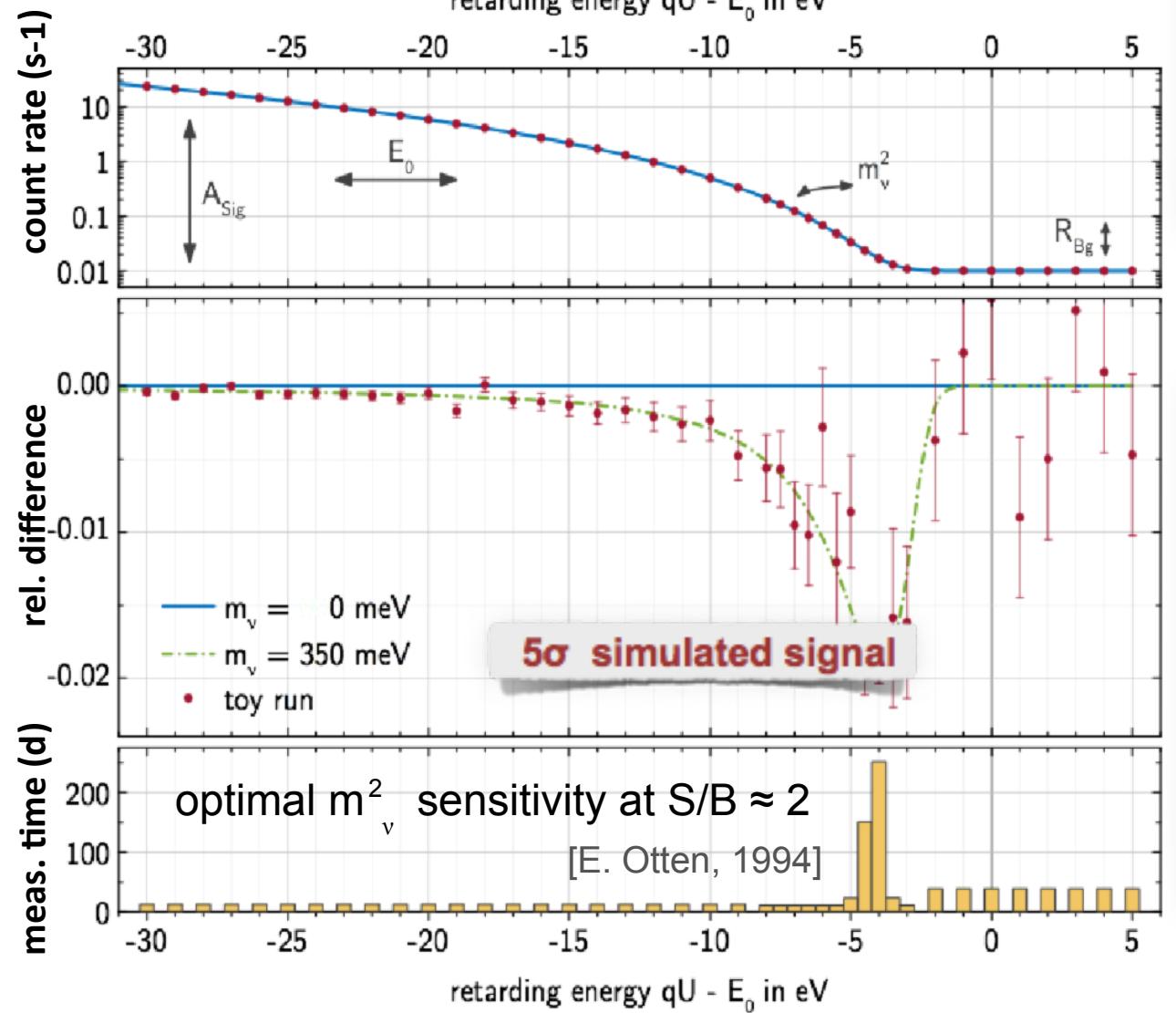
K-32 line
 $E(\text{ce}) = 17824.23(50)$ eV
 $E_{\text{binding}}(\text{ce}) = 14327.26(4)$ eV
 $E_{\text{rec}}(\text{ce}) = 0.120$ eV

L3-32 line
 $E(\text{ce}) = 30472.19(50)$ eV
 $E_{\text{binding}}(\text{ce}) = 1679.21(3)$ eV
 $E_{\text{rec}}(\text{ce}) = 0.207$ eV

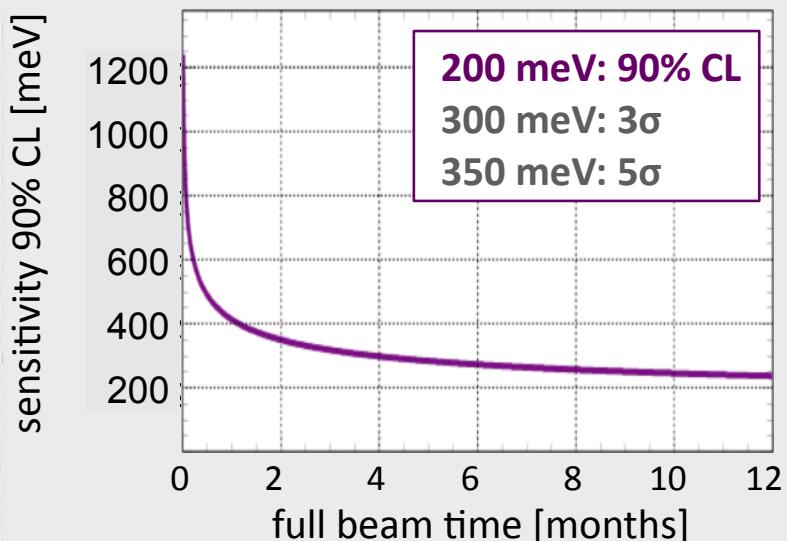
→ both values agree perfectly!

HV divider scale factor stays constant over 4 years (5 ppm uncertainty)

Neutrino mass analysis & sensitivity



- Relative **shape** measurement of **integrated β spectrum**
- 4 fit parameters:
 $m^2(\nu)$, E_0 , amp , bg



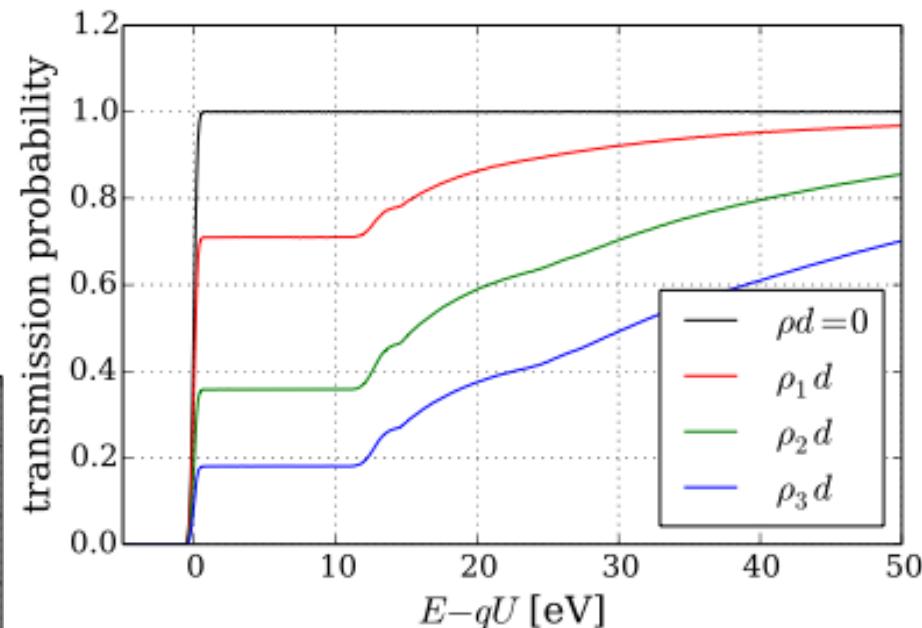
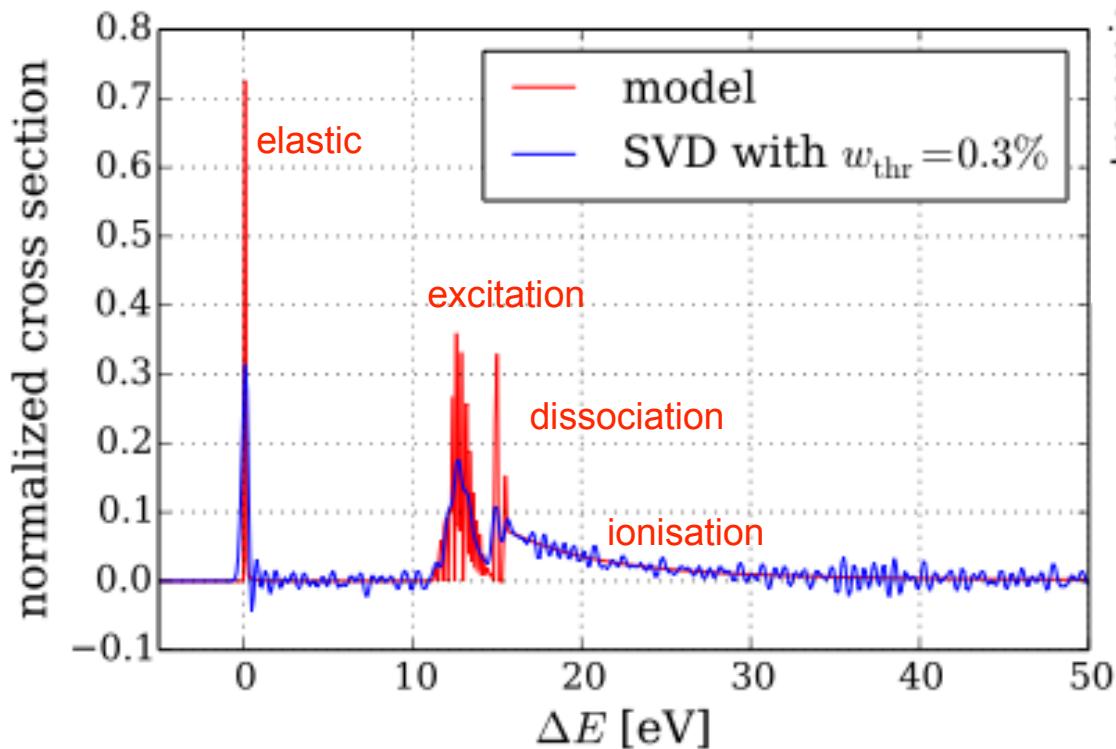
Determining the energy loss function



Measurement of electron transmission at different gas column densities

→ **determine energy loss function through deconvolution**

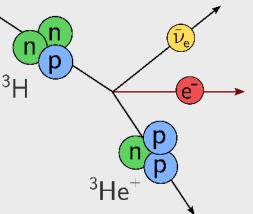
(here: singular value decomposition, SVD)



V. Hannen et al., Astropart. Phys. 89 (2017) 30

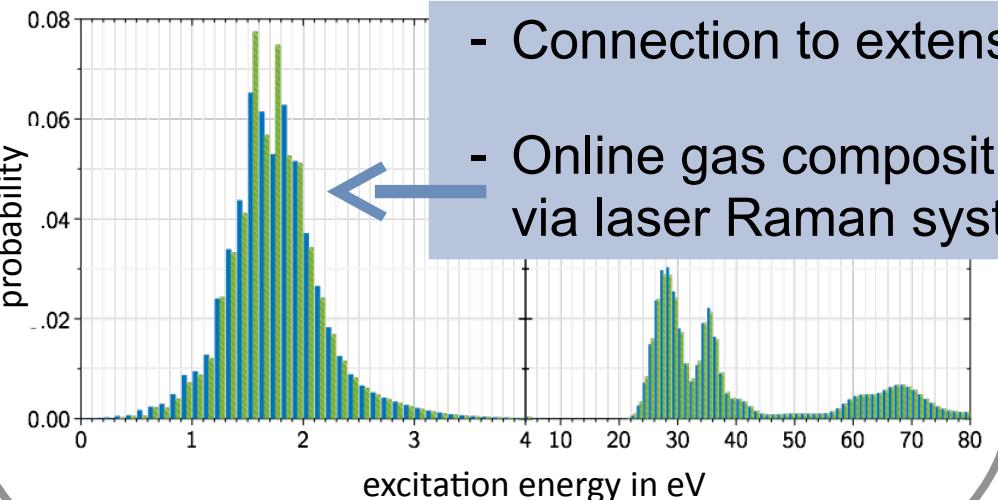
Modelling of the β spectrum

Reference spectrum generator



State-of-the-art nuclear & molecular theory

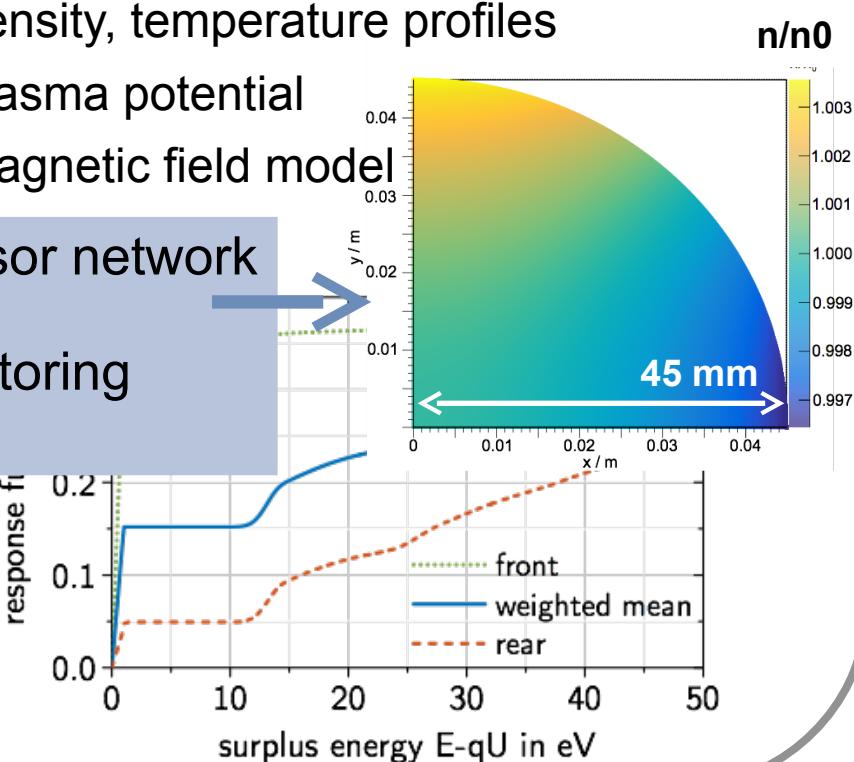
- electronic final states, radiative corrections, Doppler effect
- minor: relativistic Fermi function & recoil, screening, finite nuclear extension, ...



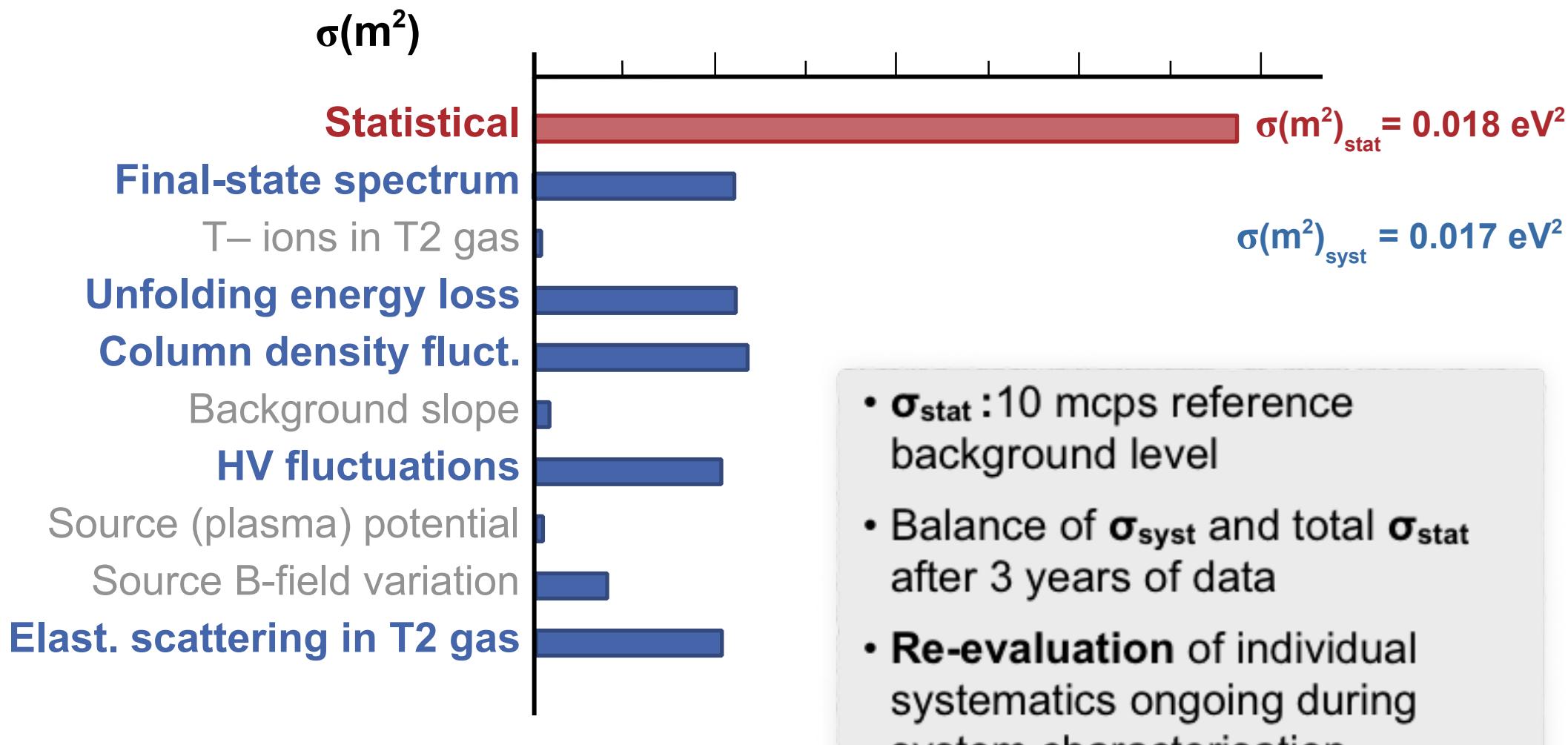
- Connection to extensive sensor network
- Online gas composition monitoring via laser Raman system

3d gas-dynamics modelling

- longitudinal and radial pressure, density, temperature profiles
- plasma potential
- magnetic field model



KATRIN's uncertainty budget (design sensitivity, ~2004):



KATRIN's uncertainty budget (design sensitivity, ~2004):

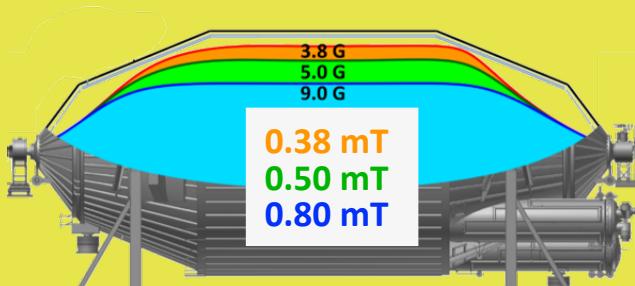
3 yr of data taking

sensitivity on the neutrino mass (stat.+sys. uncertainties):

→ 200 meV (design value)

Higher (Rydberg) background rate

→ using larger data range (E_0 -60 eV) and a bit less energy res.:



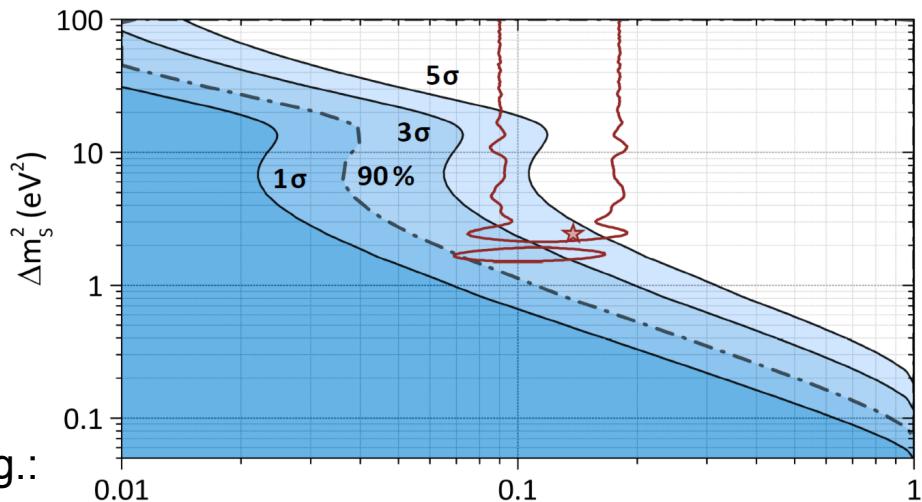
→ 240 meV (without further mitigation of the Rydberg background)

system characterisation

Sterile neutrinos

$$dN/dE = K \cdot F(E, Z) \cdot p \cdot E_{\text{tot}} \cdot (E_0 - E_e) \left(\cos^2(\theta) \sqrt{(E_0 - E_e)^2 - m(\nu_{1,2,3})^2} + \sin^2(\theta) \sqrt{(E_0 - E_e)^2 - m(\nu_4)^2} \right)$$

eV ν :



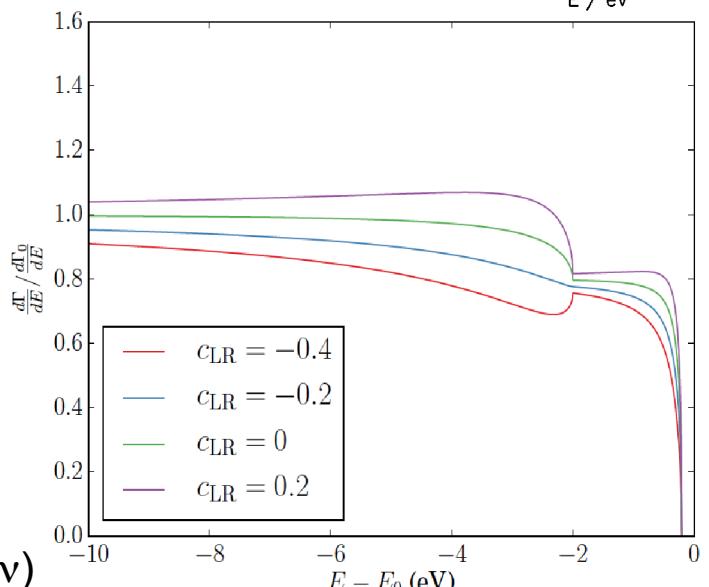
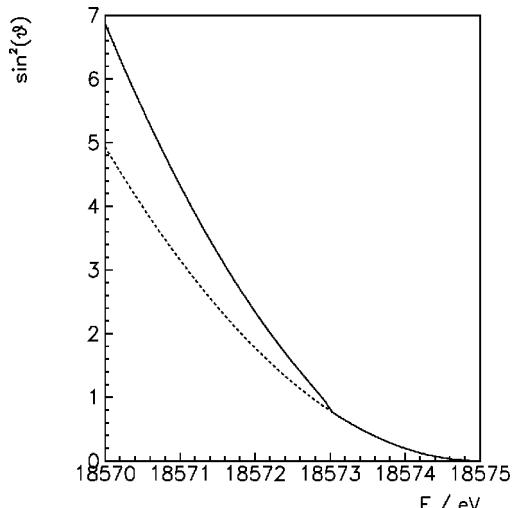
see e.g.:

J. A. Formaggio, J. Barret, PLB _{$\sin 2\theta_s$} 706 (2011) 68

A. Sejersen Riis, S. Hannestad, JCAP02 (2011) 011

A. Esmaili, O.L.G. Peres, arXiv:1203.2632

M.Kleesiek,
PhD thesis,
KIT (2014)



keV ν :

see e.g.

S. Mertens et al., JCAP 02 (2015) 020

M. Drewes et al. JCAP 01 (2017) 025

non SM currents, ...

see e.g.: N. Steinbrink et al., JCAP 6 (2017) 15 (RH currents & sterile ν)

Conclusions

Direct neutrino mass experiments:

complementary to cosmological analyses and $0\nu\beta\beta$
can look also for sterile neutrinos (eV, keV) and other BSM

KATRIN: the direct neutrino mass experiment with 200 meV sensitivity
System is complete (except tritium loops and rear wall and calibration system):

- 1st light in October 2016
- ^{83m}Kr calibration measurements in July 2017 successful
- **tritium data taking will start in 2018**

**KATRIN inauguration ceremony: June 11, 2018
(after Neutrino 2018 at Heidelberg)**



Thank you for your attention !