

# Chasing the Light Sterile Neutrino: Status of the STEREO Experiment

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on Behalf of the STEREO Collaboration



CNNP 2017 - Catania

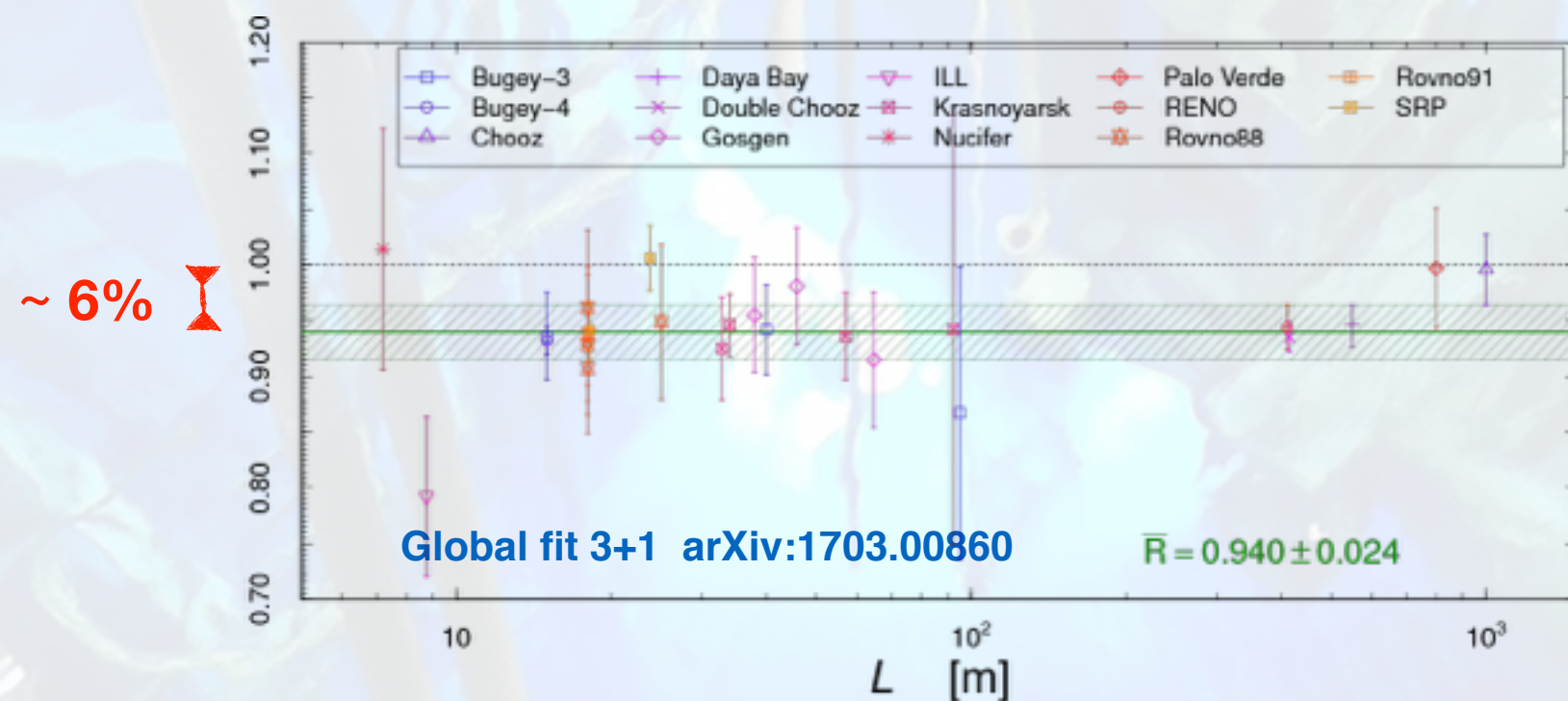




# Reactor Antineutrino Anomaly

- Estimation of reactor  $\bar{\nu}_e$  spectrum  $\rightarrow$   **$\sim 2.6 \sigma$  rate excess** wrt short baseline measures

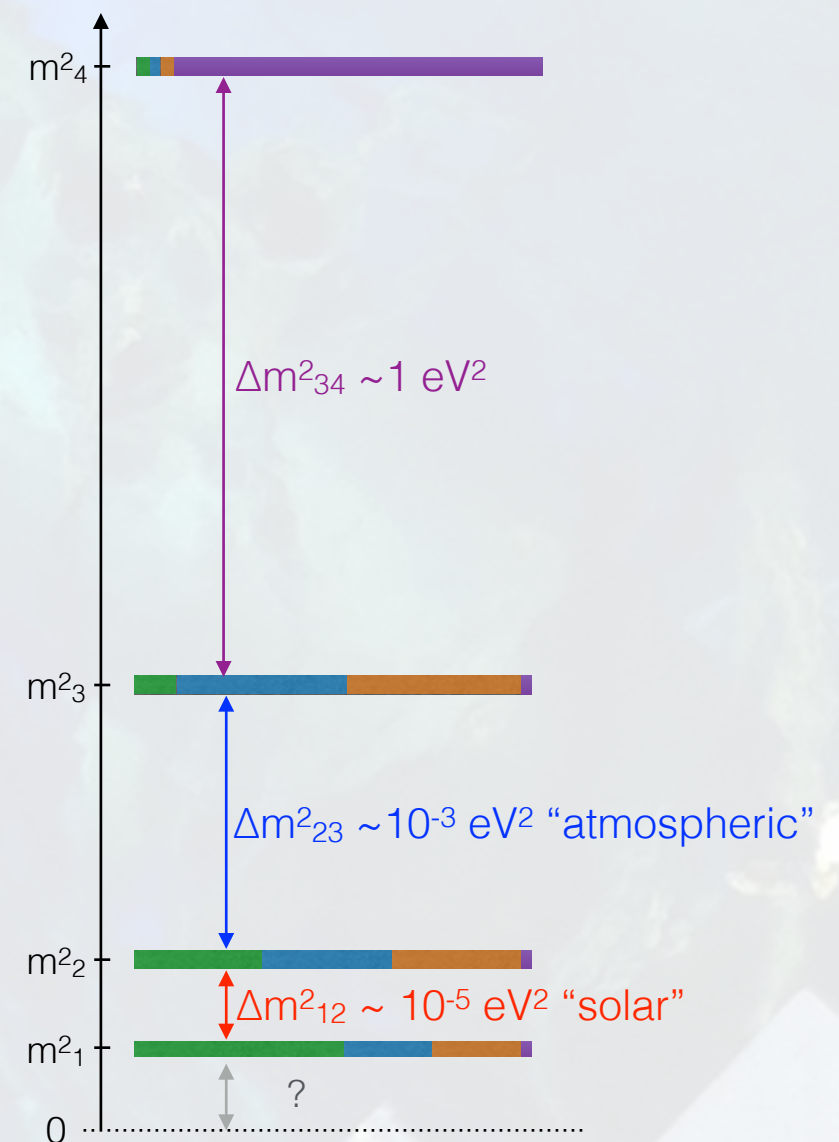
Phys.Rev.D83, 073006 (2011)



- Explained by **oscillation involving extra sterile neutrino(s)** with  $\Delta m^2 \sim 0.1-1 \text{ eV}^2$

$$P_{\bar{\nu}_e \rightarrow \bar{\nu}_e}(L \lesssim 10^2 m) \simeq 1 - \sin^2(2\theta_{ee}) \sin^2(1.27 \Delta m_{14}^2 L / E_{\bar{\nu}_e})$$

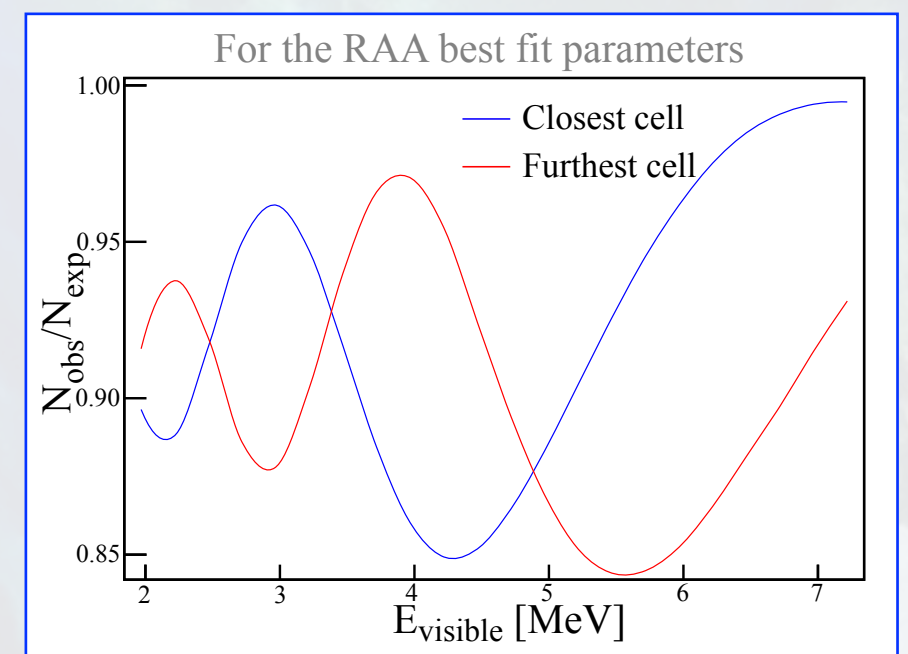
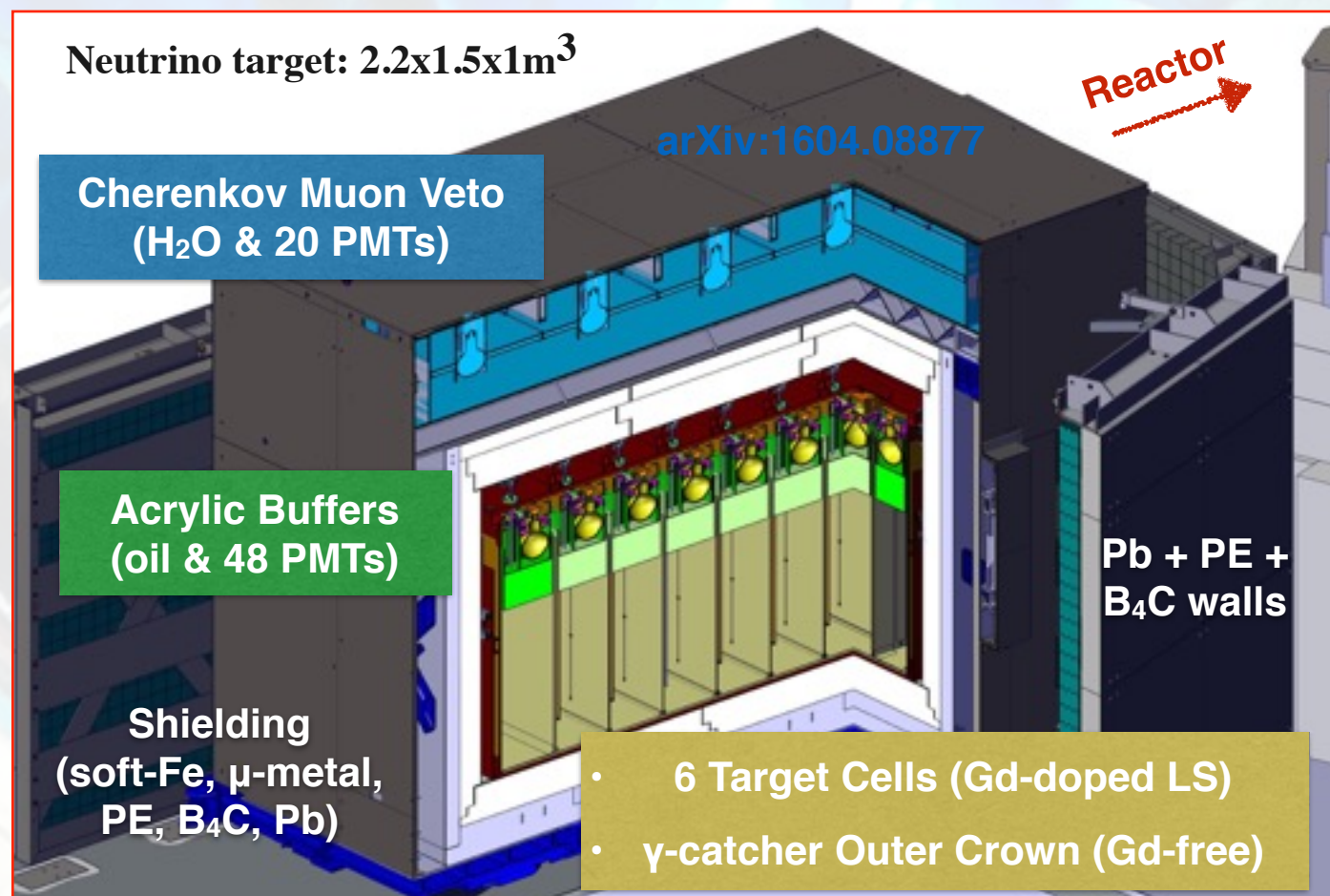
new parameters





# The STEREO experiment

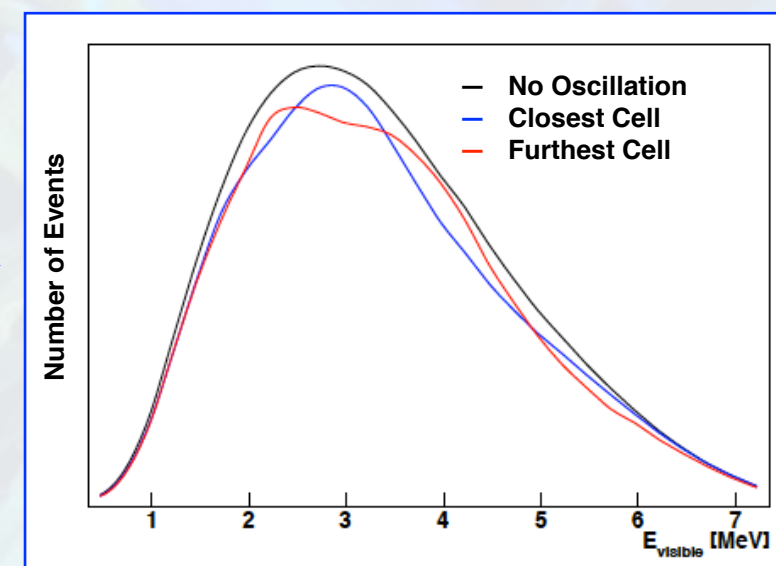
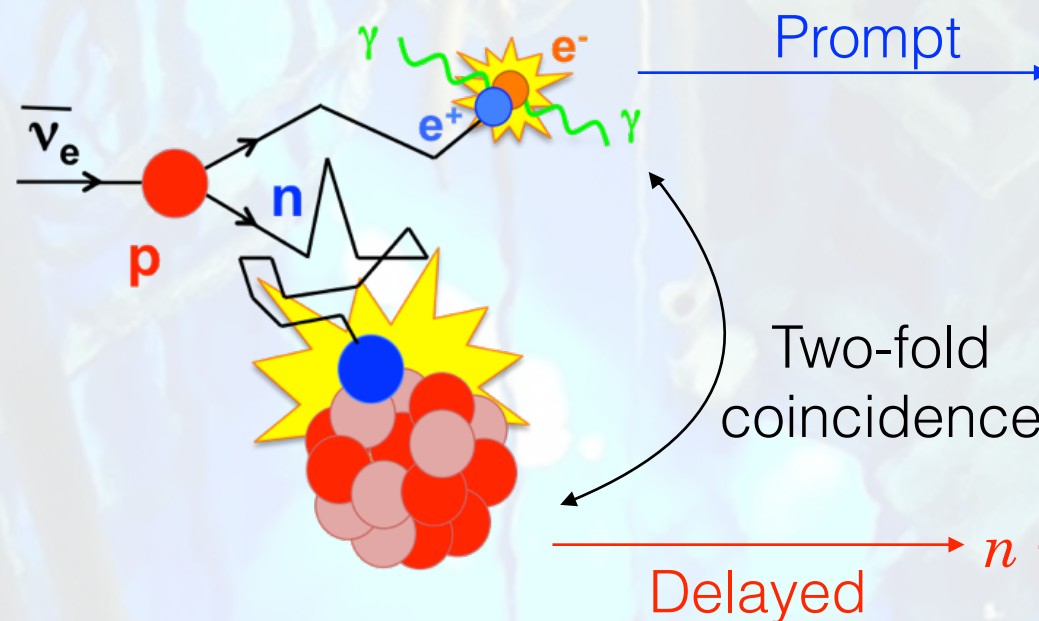
- STEREO detector: **9 m from ILL research reactor core** ( $\varnothing = 37$  cm, 93%  $^{235}\text{U}$ , 58.3 MW<sub>th</sub>)
- **Compare 6 target cells** to measure oscillation-driven **distortion in the  $E_{\bar{\nu}_e}$  spectrum**
- Main challenges: surface level, reactor noise ...and competition





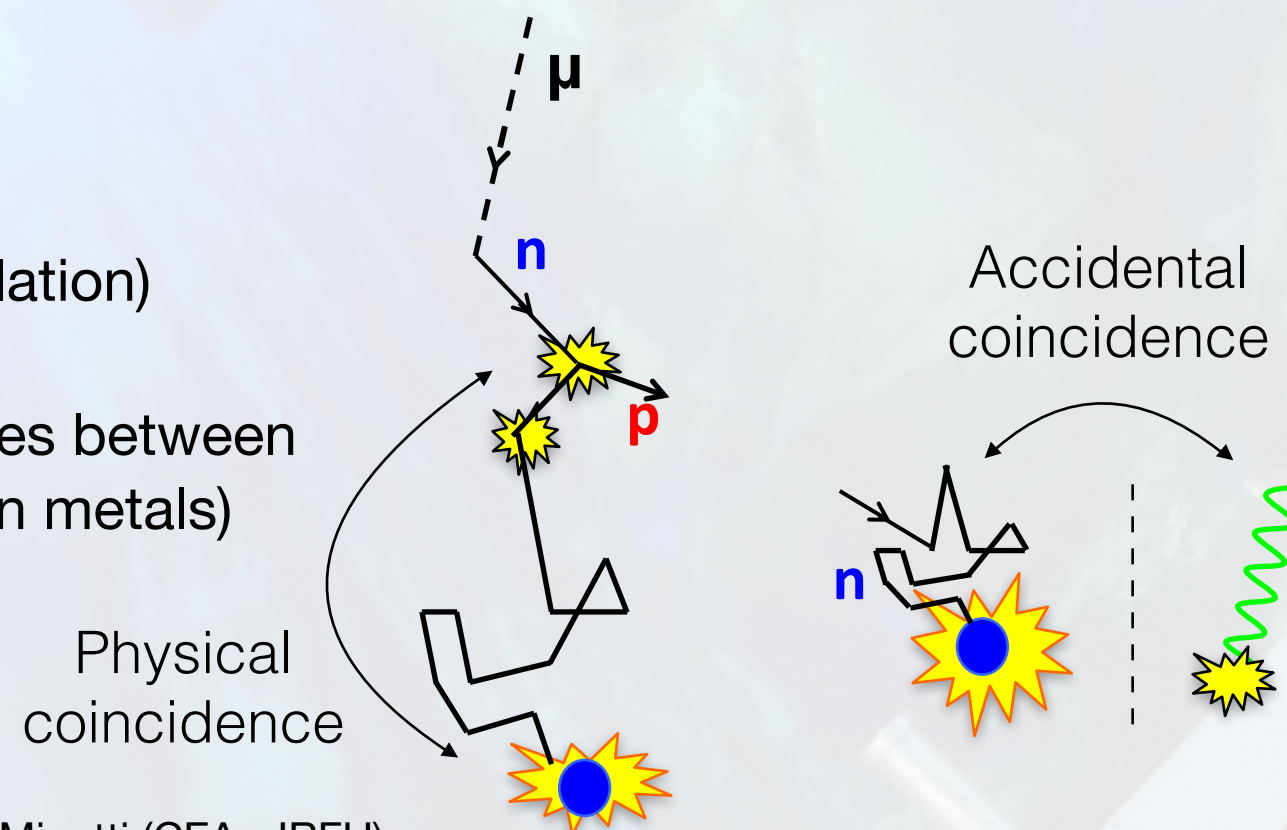
# Signal and Background

- Signal: **In**verse **B**eta **D**ecay  $\bar{\nu}_e p \rightarrow e^+ n$



$$E_\nu \approx E_{e^+} + 0.8 \text{ MeV}$$

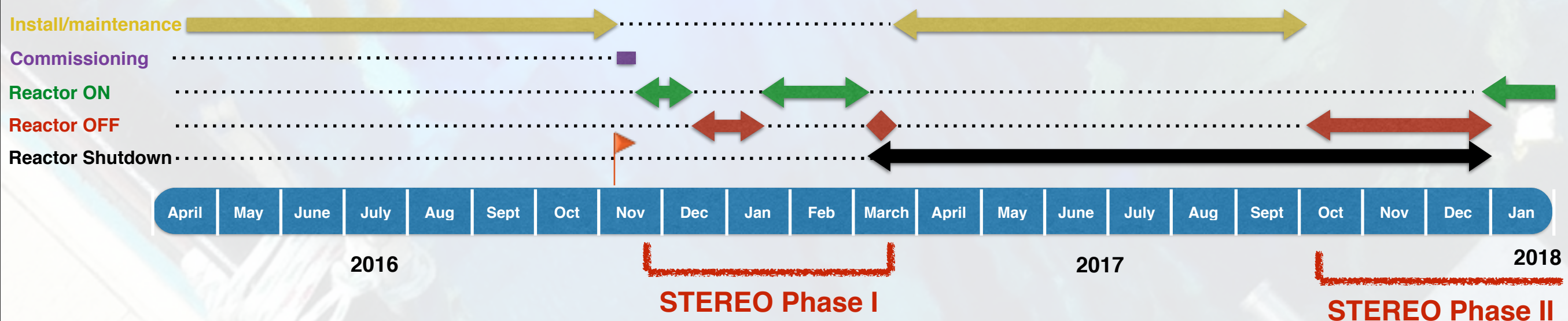
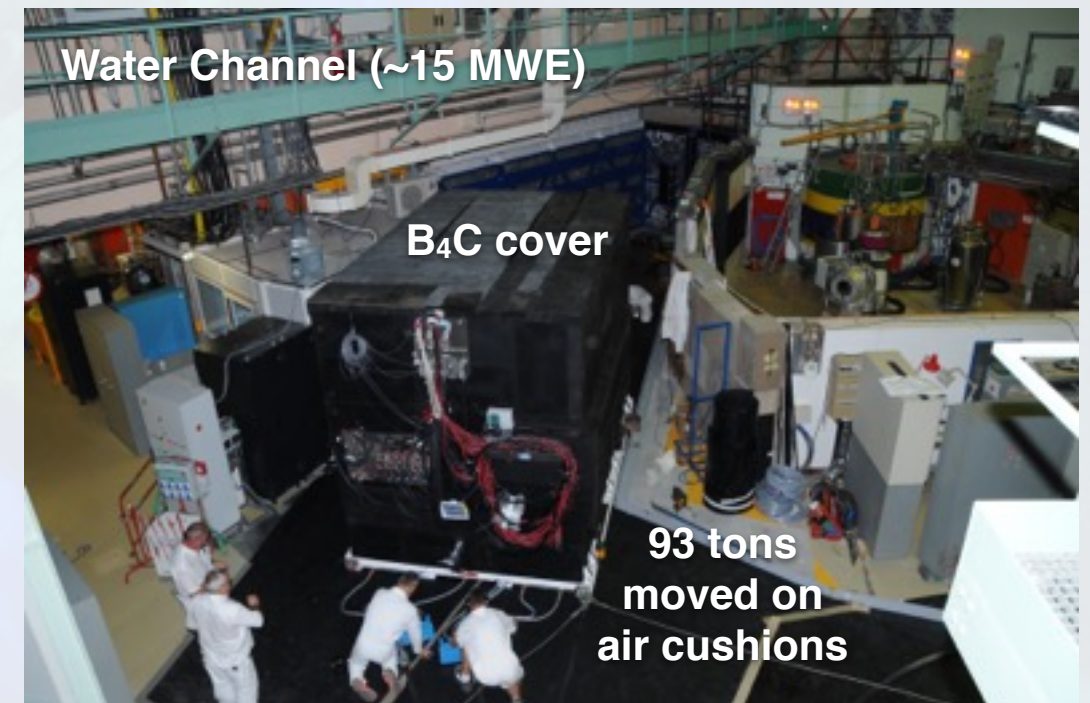
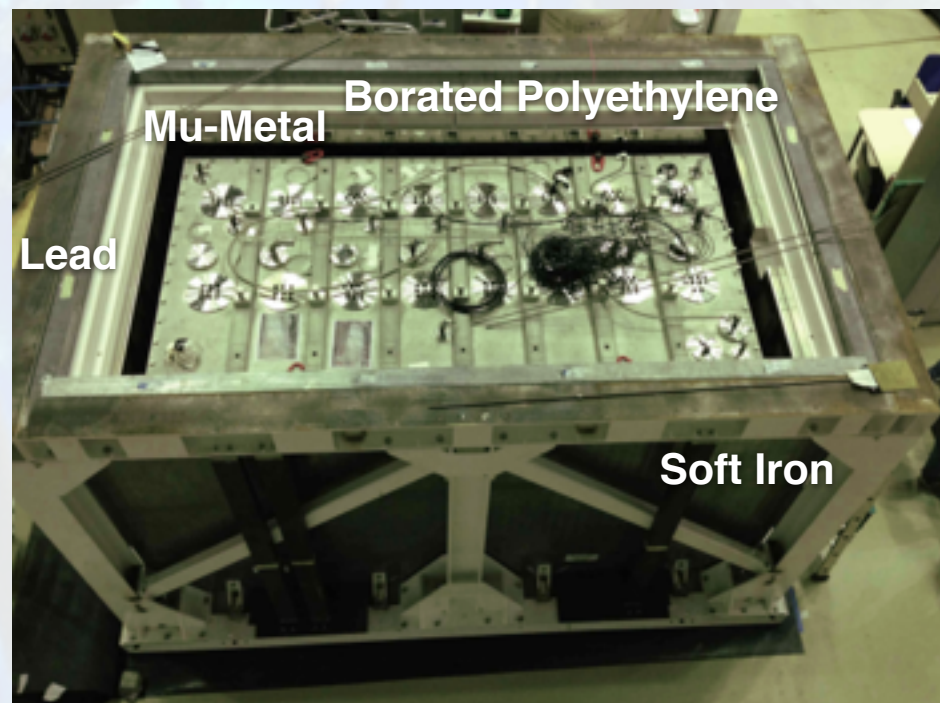
- Background
  - Cosmic induced (fast  $n$ ,  $n$ - $\gamma$  from  $\mu$  spallation)
  - Reactor-induced (accidental coincidences between  $n_{\text{th}}$  leakage, high-E  $\gamma$ 's from  $n$ -capture on metals)





# STEREO Timeline and Data Taking

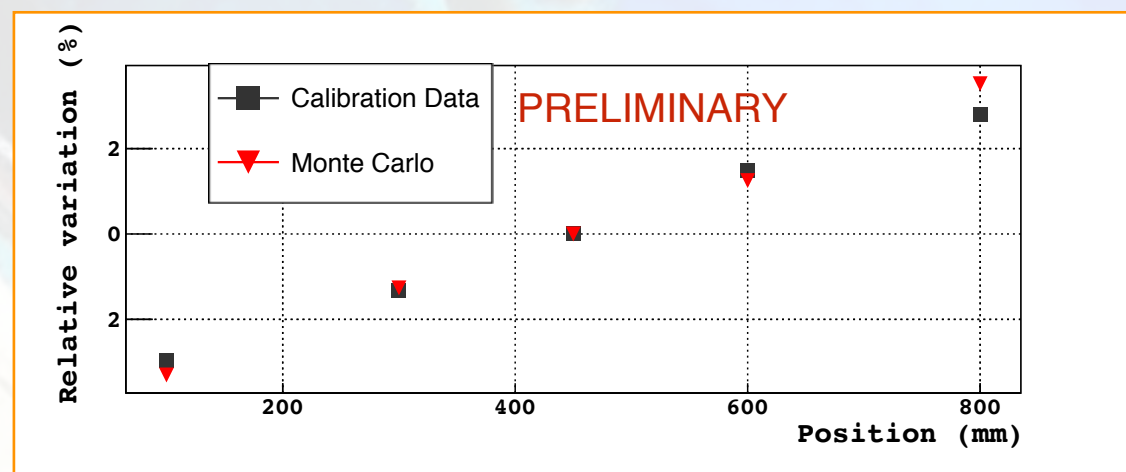
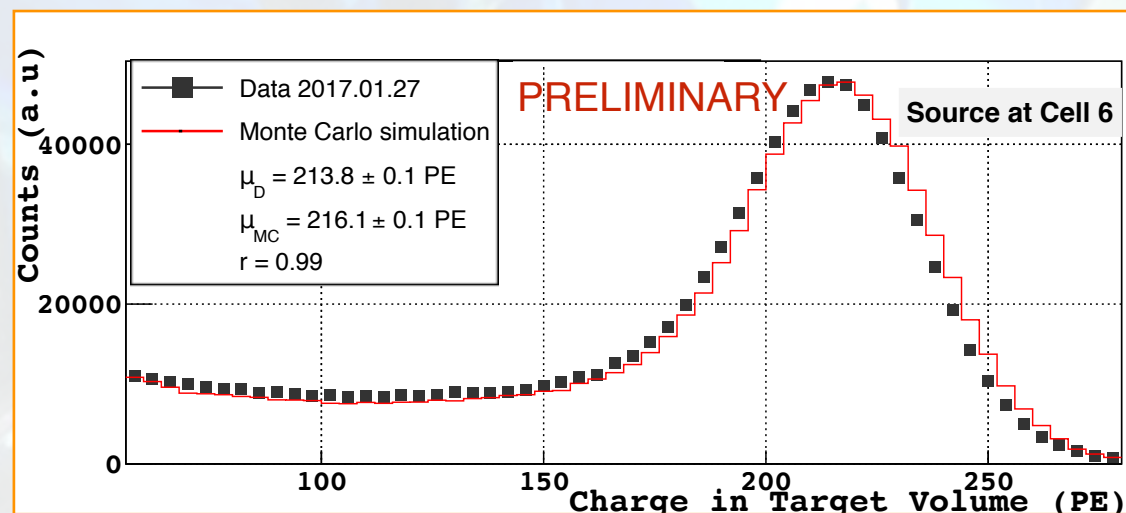
- Phase I (from Nov 2016): **70 days reactor ON** (~1.5 cycles), **25 days OFF**
- Phase II (from 4 Oct 2017): expecting 7 additional cycles in 2018-19





# Calibration: $\gamma$ Peaks

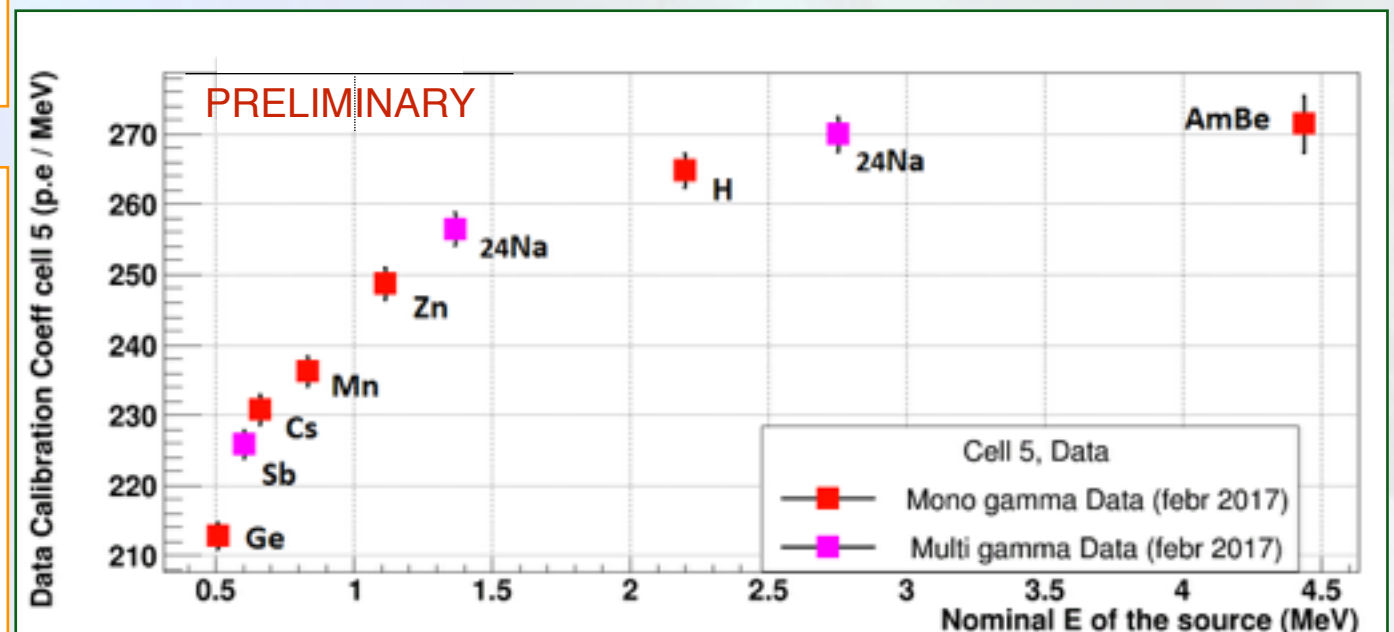
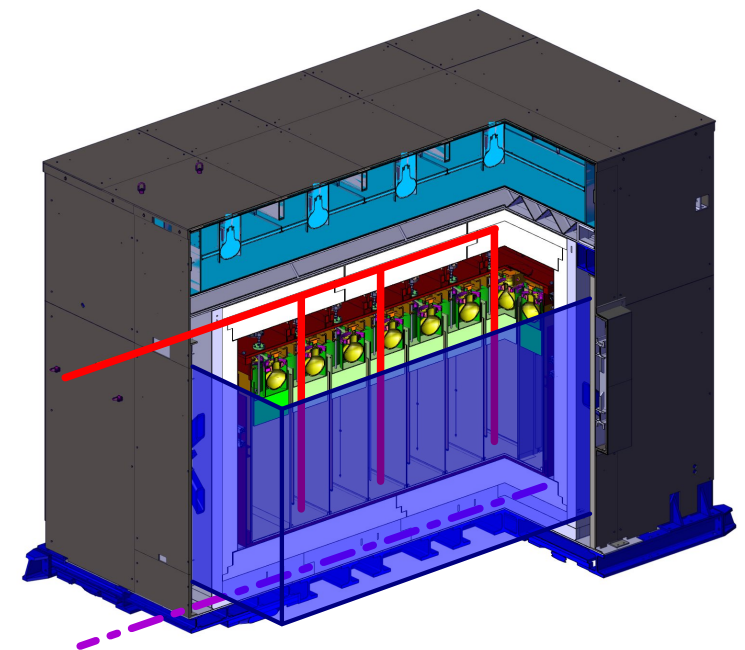
- STEREO detector **calibrated** regularly with  $\gamma$  and  $n$  sources using **3 independent systems**
- Energy peaks: low z-dependence, reproduced by MC



internal  
calibration  
(cell 1,4,6)

external  
calibration  
(2D, inside  
shielding)

underneath  
calibration





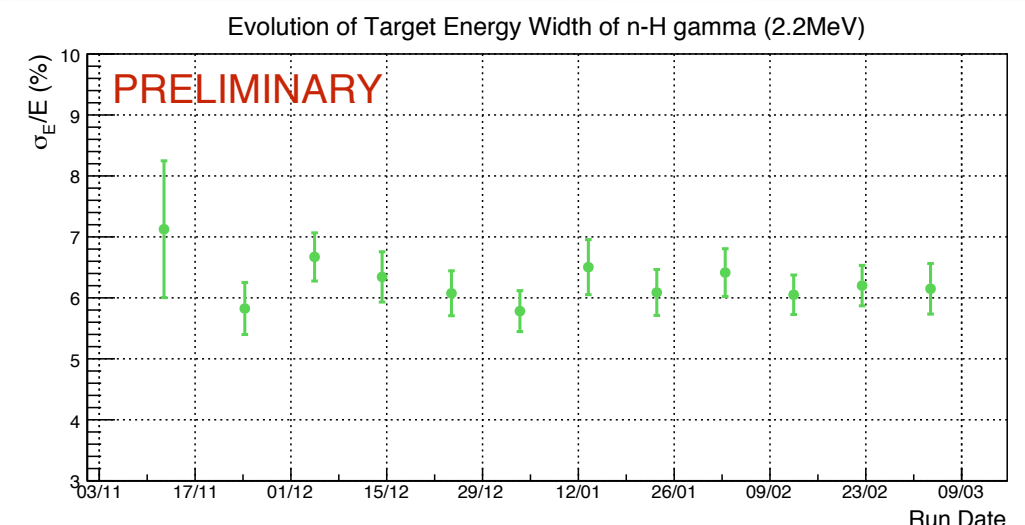
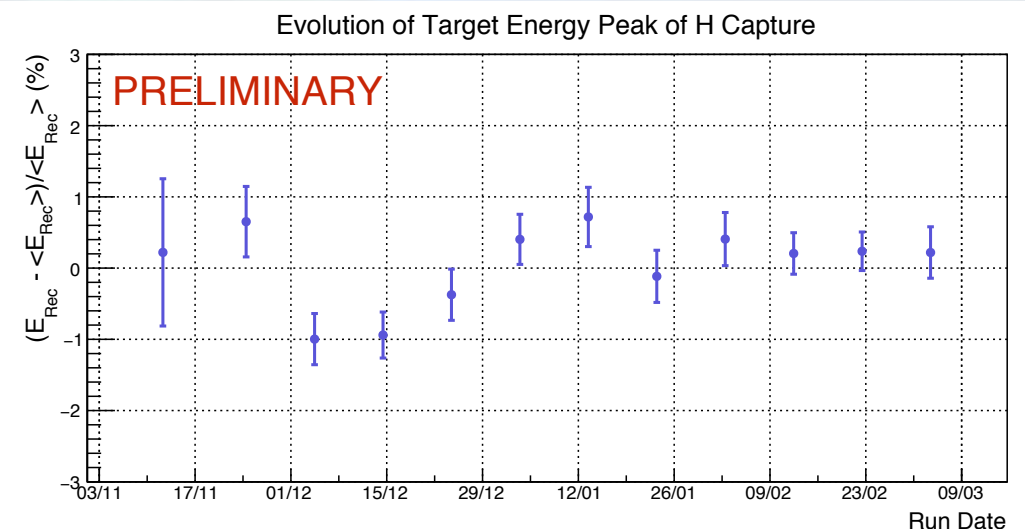
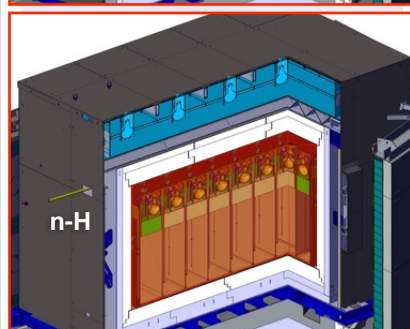
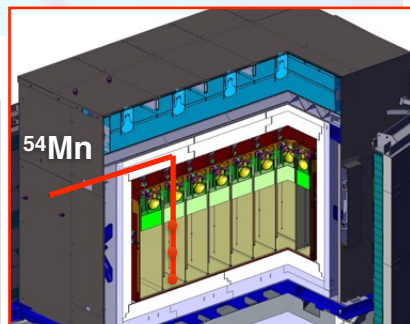
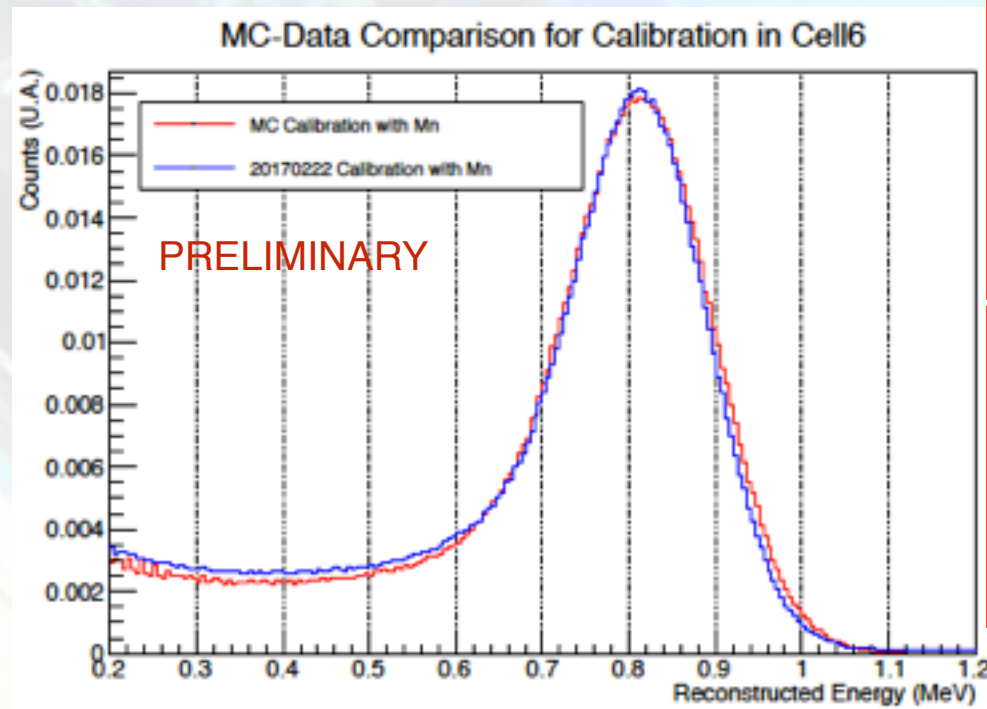
# Energy Reconstruction

- Deposited energy is obtained from charges collected in each cell (matrix inversion)

$$Q_i = \alpha_i^{\text{geom}} \sum_{j=\text{cells}} E_j^{\text{dep}} c_j L_{ji}$$

Charge collected cell  $i$       Photon acceptance cell  $i$       Energy deposit cell  $j$       Photons/MeV in cell  $i$  (from  $^{54}\text{Mn}$  calibration)      Cross-talk cells  $i$  and  $j$  (from  $^{54}\text{Mn}$ ,  $\mu$ 's)

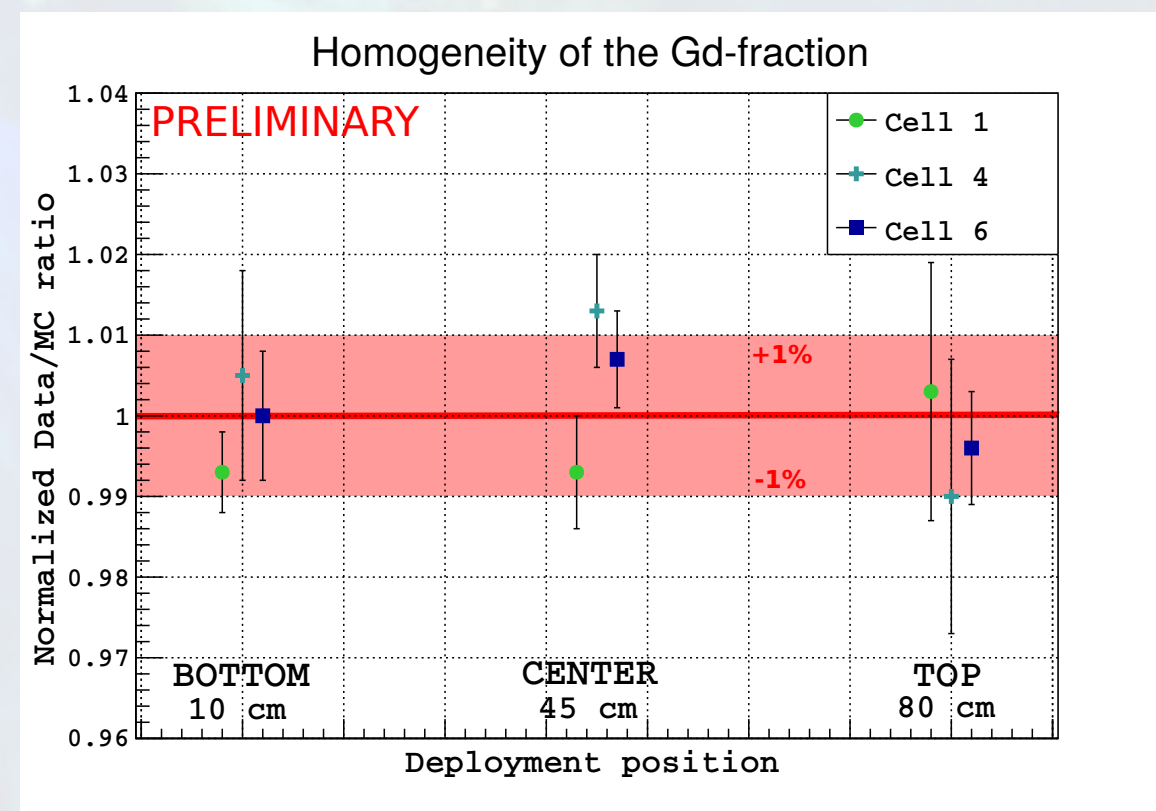
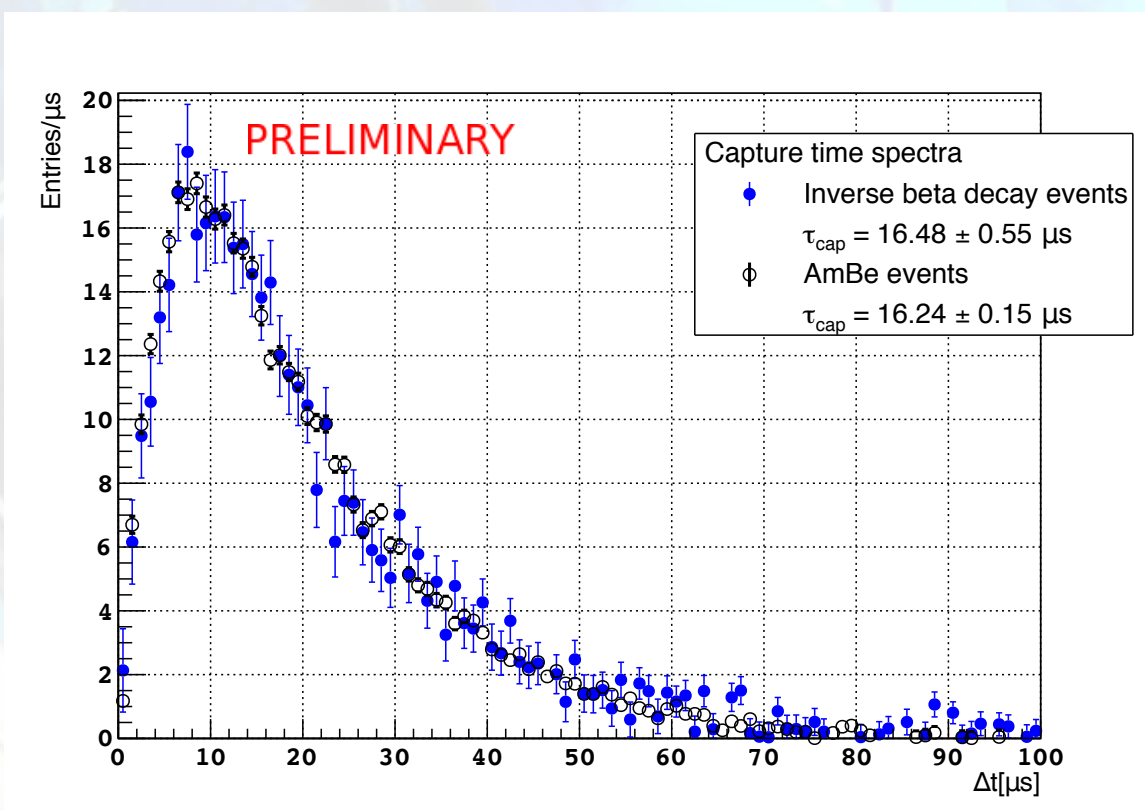
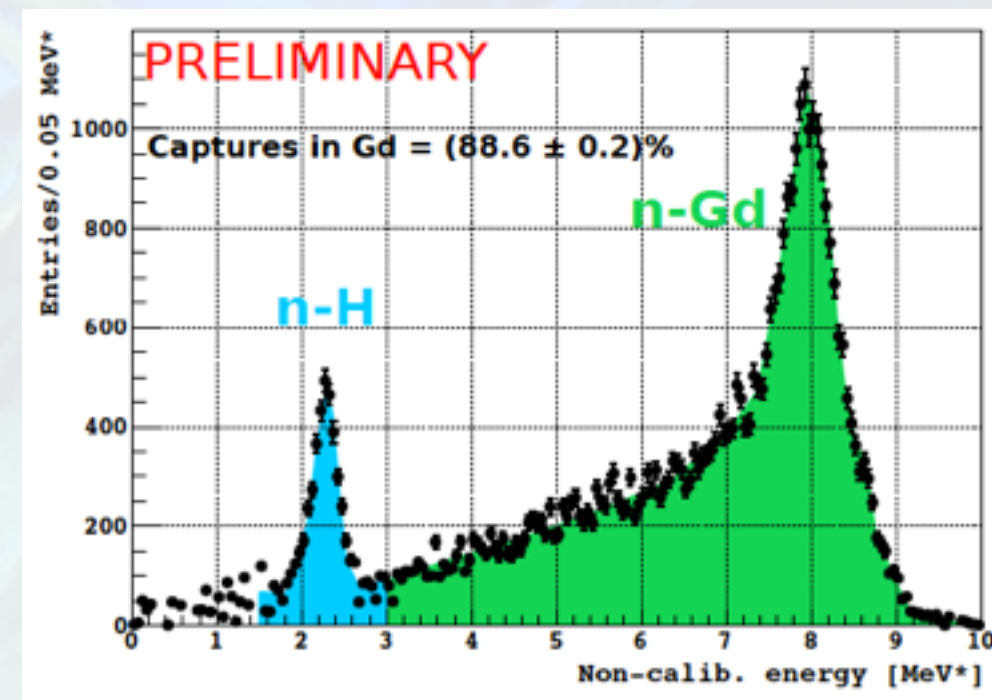
- Very good data-MC agreement on  $^{54}\text{Mn}$  peak
- Validation with  $n$ -H capture peak stability





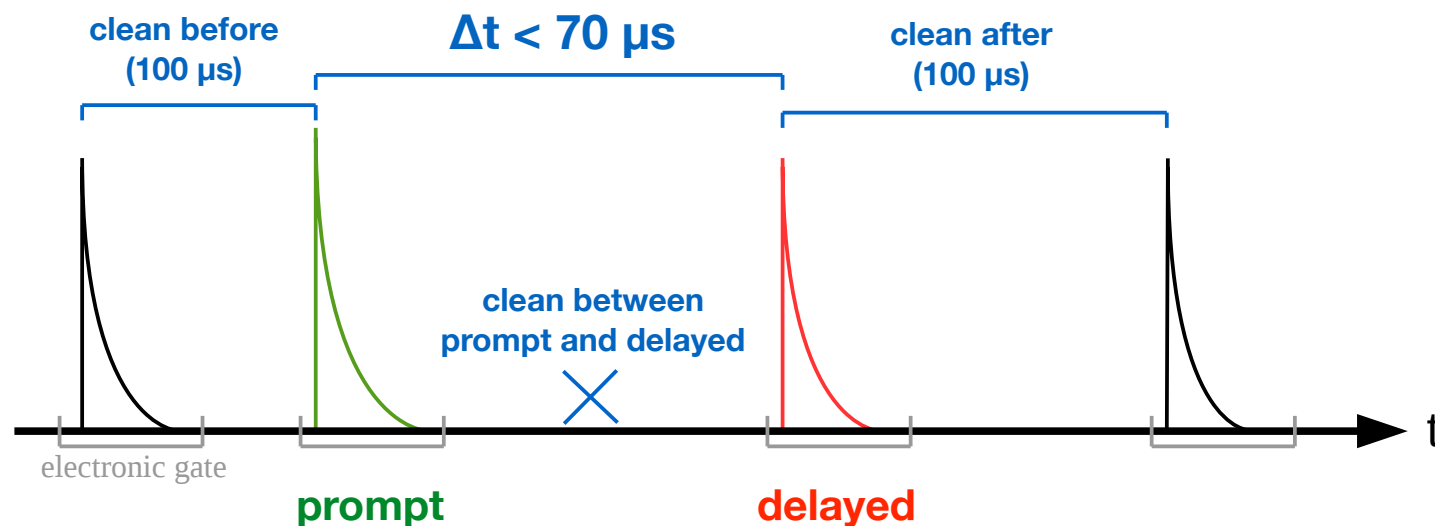
# Calibration: Detection Efficiency

- Neutron calibration using AmBe coincidences
- n-capture time in agreement with IBD candidates
- Correction factor for Gd-capture fraction in MC  
→ **MC validation** for global **n-capture efficiency** calculation



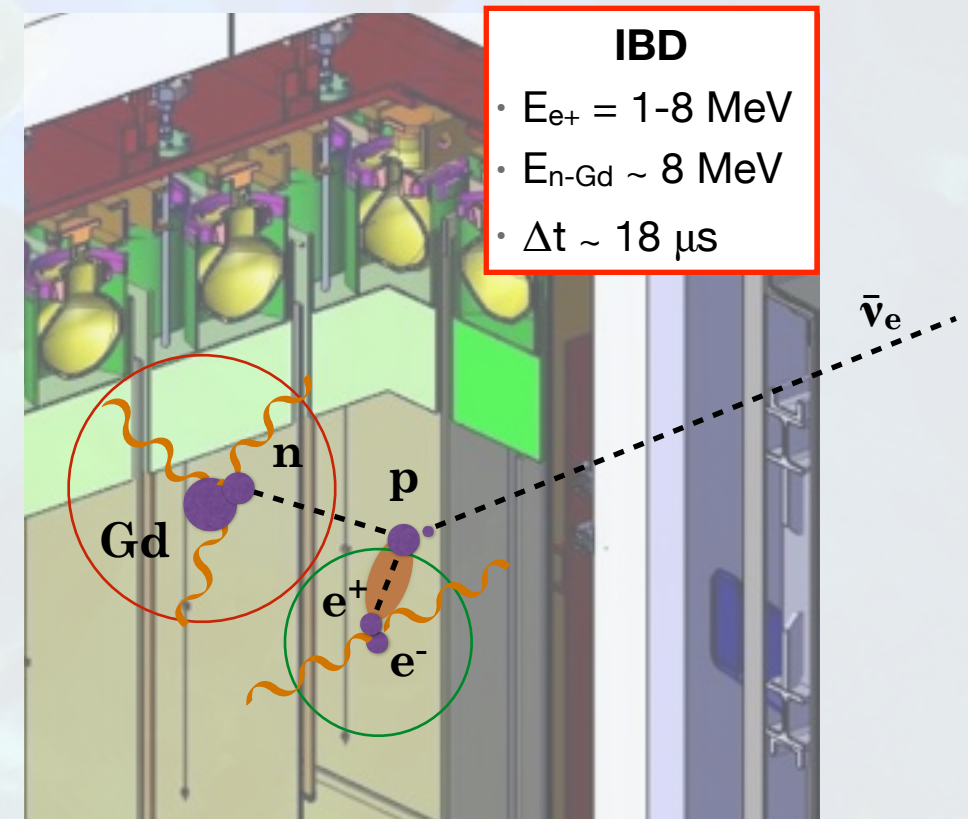


# Neutrino Selection and Cuts

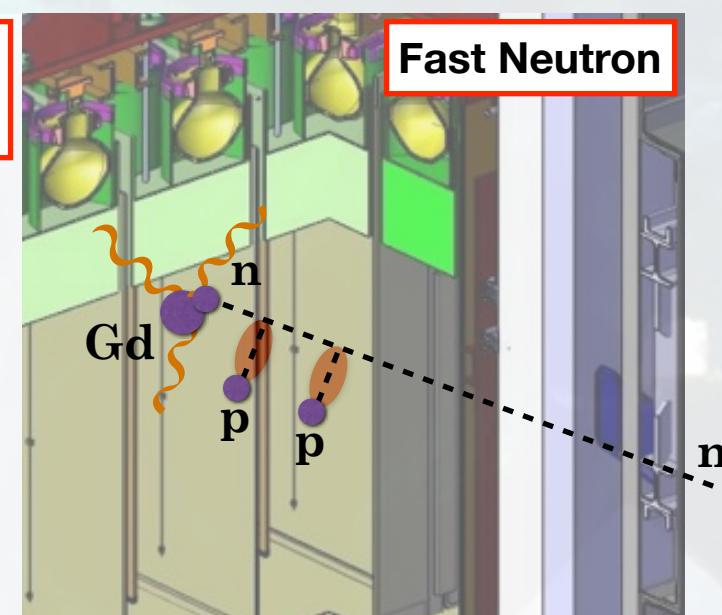
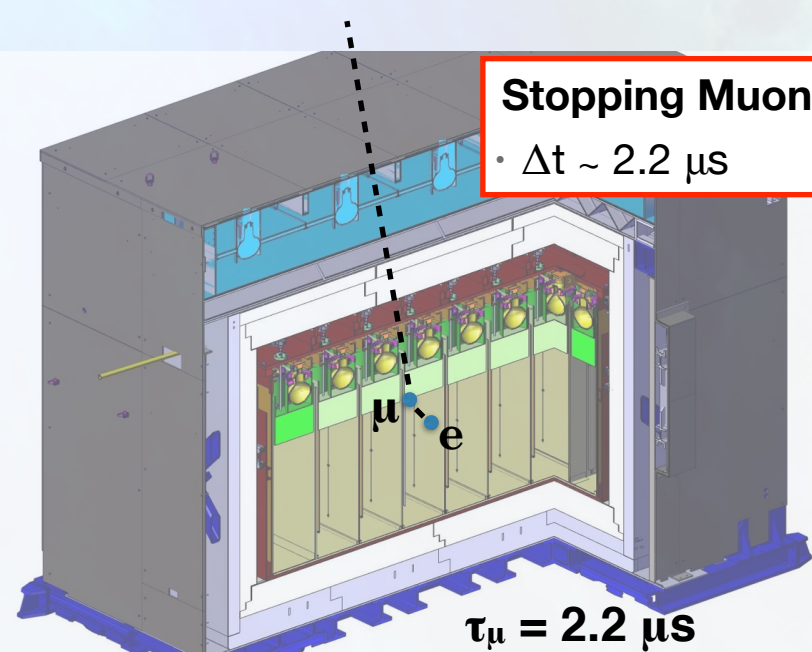


$$2 \text{ MeV} < E_{\text{det}}^{\text{dep}} < 8 \text{ MeV} \\ \& E_{\text{GC}}^{\text{dep}} < 1.1 \text{ MeV}$$

$$5 \text{ MeV} < E_{\text{det}}^{\text{dep}} < 10 \text{ MeV} \\ \& E_{\text{Tg}}^{\text{dep}} > 1 \text{ MeV}$$



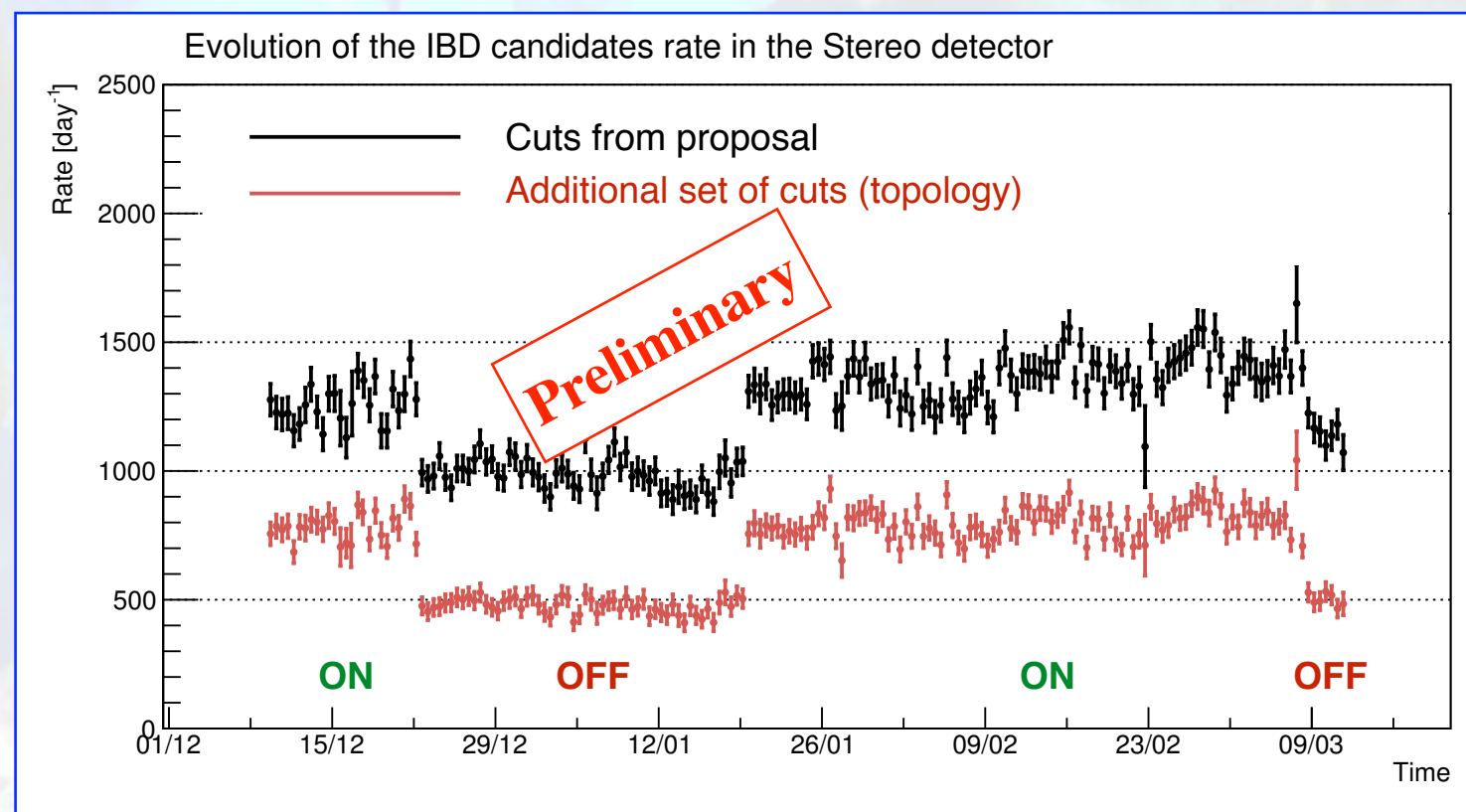
Cuts	Fast Neutrons	Stopping Muons	Neutrino efficiency
100 $\mu s$ after- $\mu$ veto	✗		$\sim 6.5 \%$ deadtime
PMT charge homogeneity		✗	98 %
PSD ( $Q_{\text{tail}}/Q_{\text{tot}}$ )	✗	✗	98 % with $2\sigma$ cut
Cell multiplicity	✗		$> 99 \%$





# Neutrino Rate

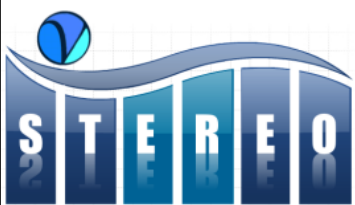
- Accidental coincidences measured online using multiple off-time windows and subtracted
- After all cuts applied and atmospheric pressure correction: neutrino efficiency  $>95\%$ , **neutrino rate  $\sim 300$  /day**
- Next analysis steps: refine cuts, define systematics for reactor-OFF subtraction, finalise energy reconstruction for spectral analysis







# Thank you for your attention



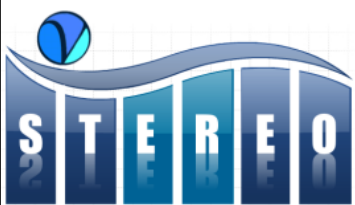
is a collaboration of







**Questions?**



*is a collaboration of*



*supported by*





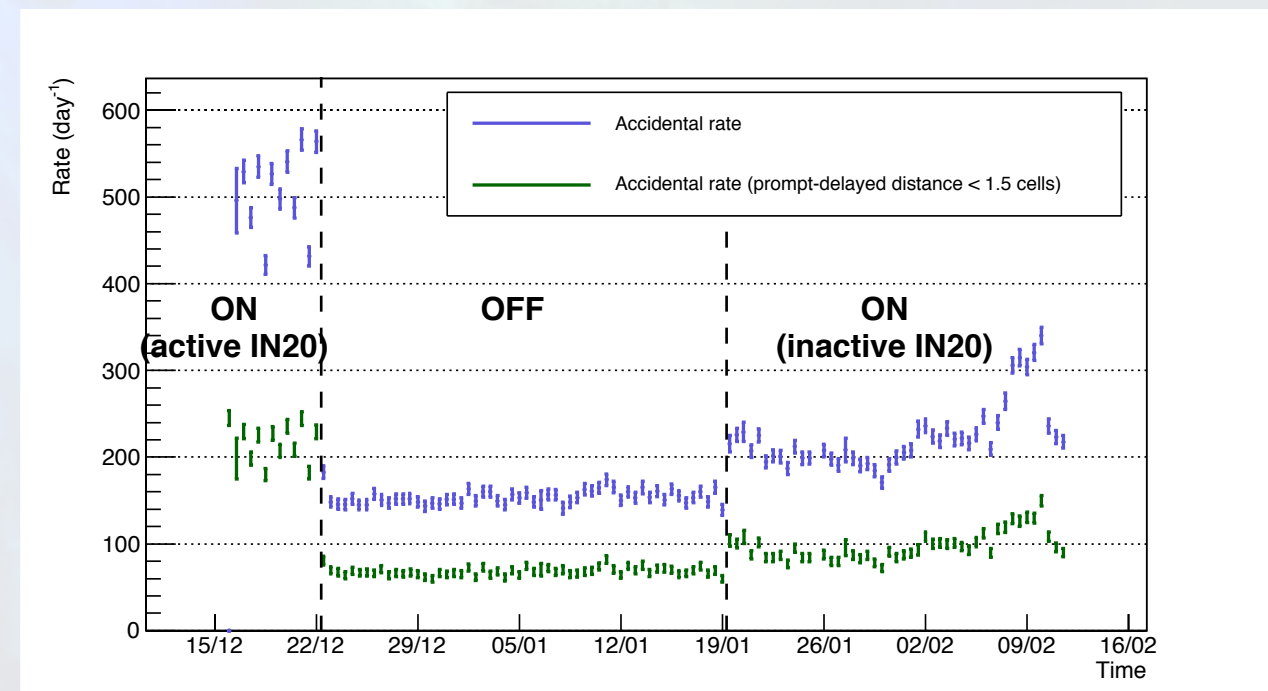
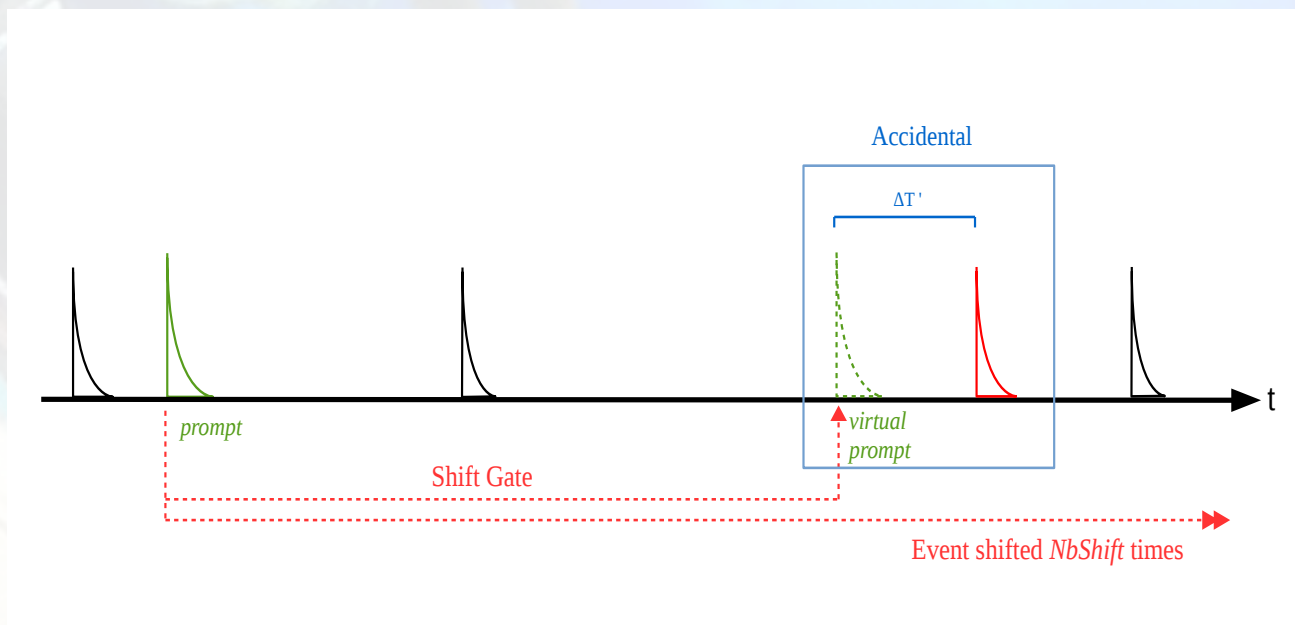
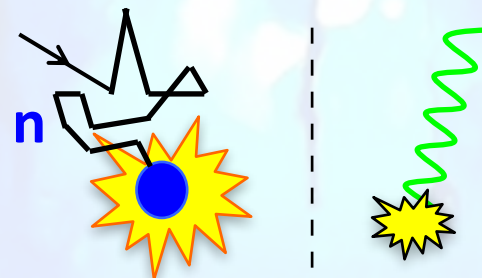
# Backup





# Accidental Coincidences

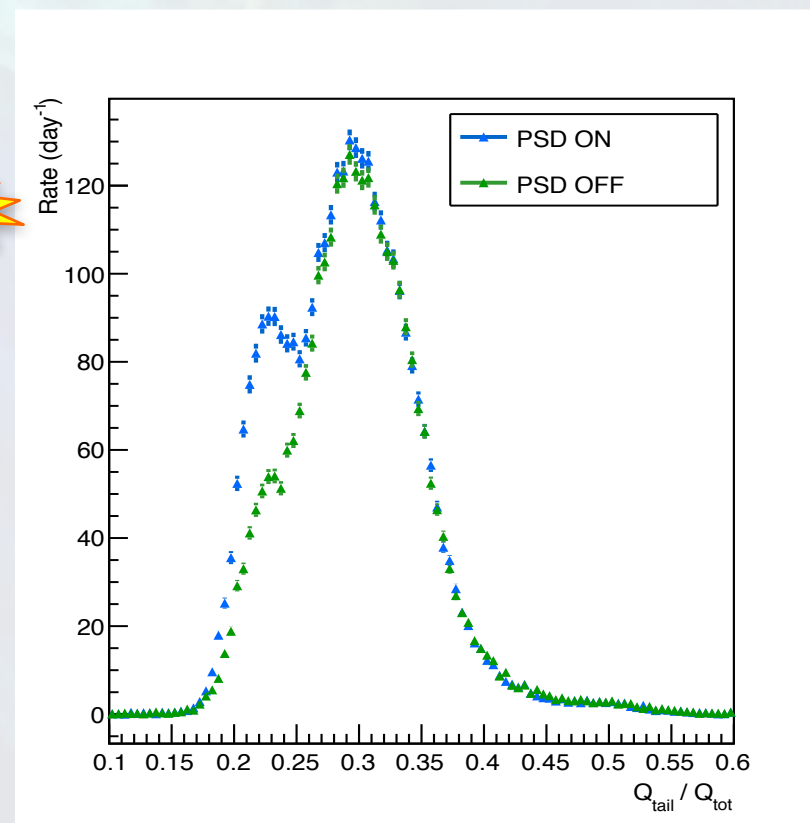
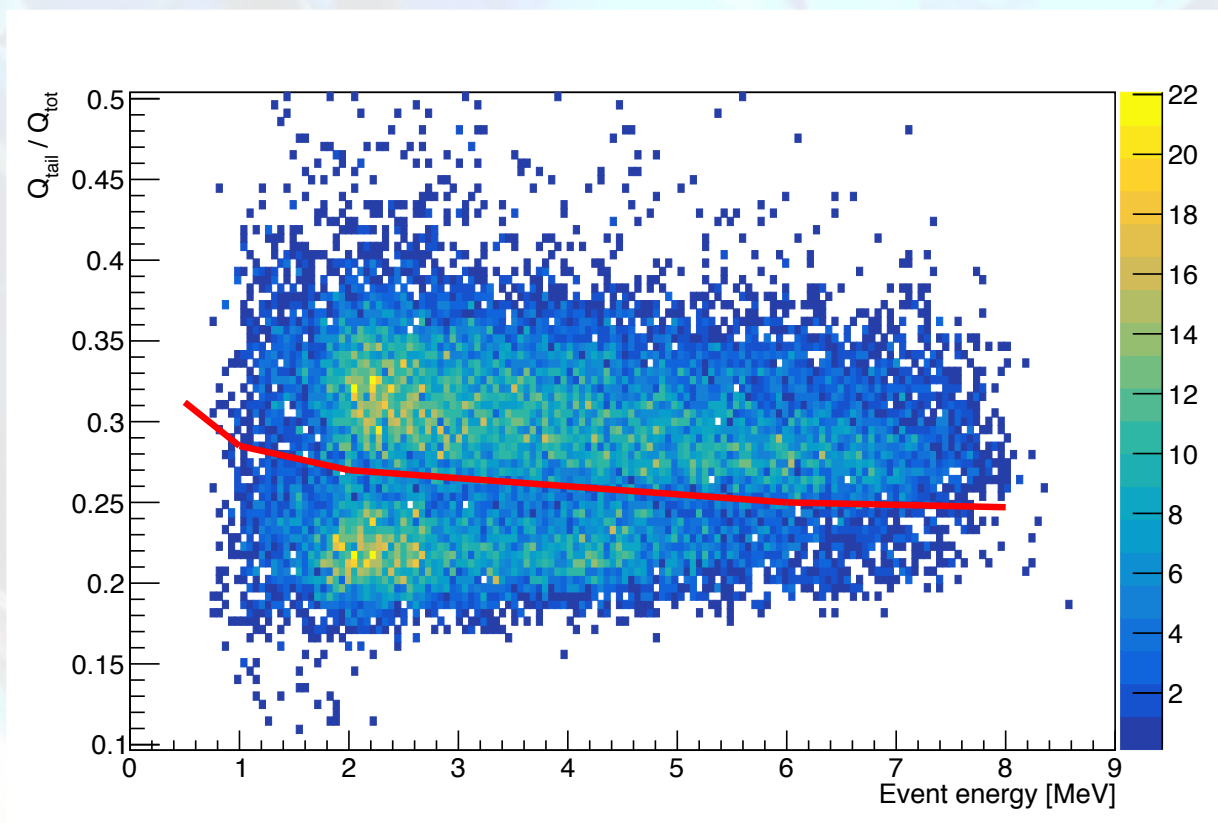
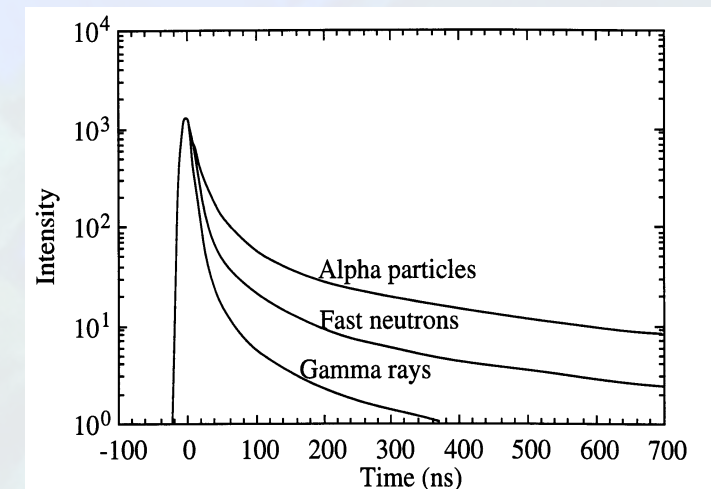
- Accidental  $\gamma$ 's/ $n$ 's coincidences (reactor-induced, influenced by neighbouring experiments)
- Remaining accidentals **measured online** & **subtracted statistically** with off-time windows
- Prompt-delayed vertices strongly correlated for IBD  $\rightarrow$  cut on spatial distance





# Pulse Shape Discrimination

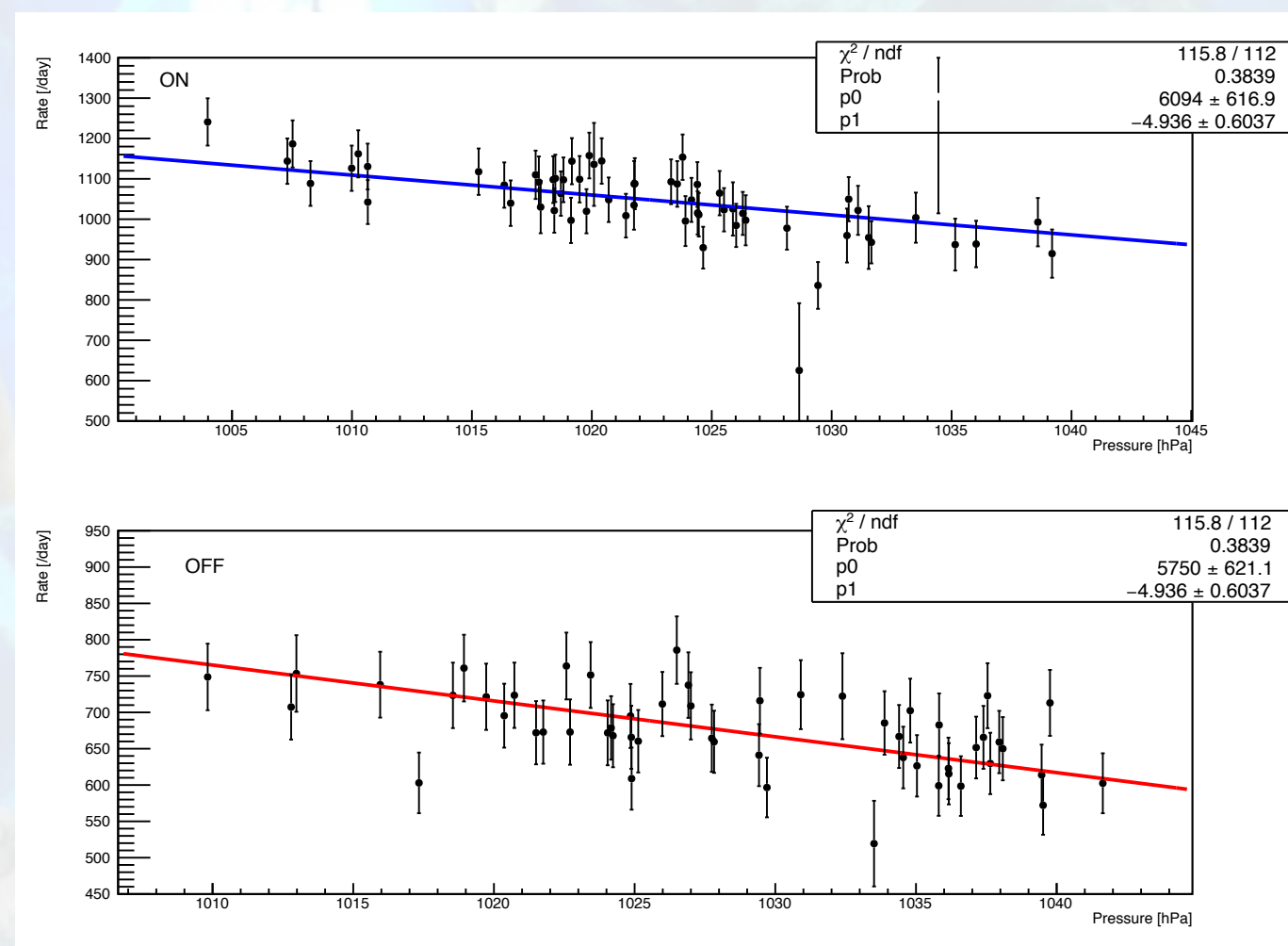
- **Pulse Shape Discrimination:** disentangle particles with large  $dE/dx$  (recoil protons, low-energy decaying  $\mu$ 's)
- Intrinsic property of LS
- Use  $Q_{\text{tail}}/Q_{\text{tot}}$  ratio (@ read-out)





# Correlated Rate Variation with Atmospheric Pressure

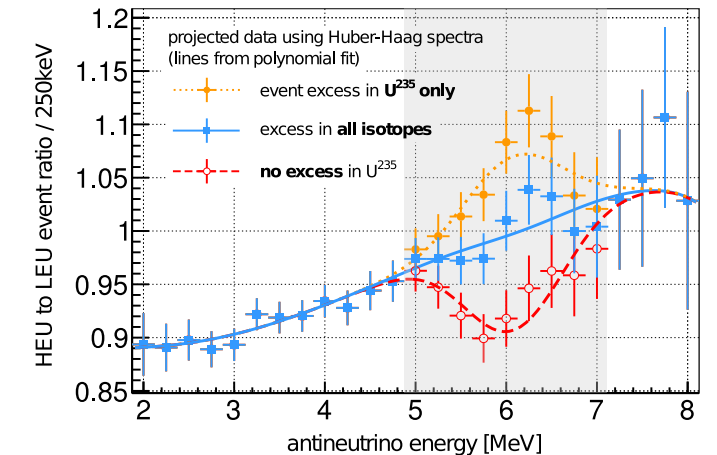
- **Residual cosmic background** is measured during **reactor off**
- Subtraction must take into account differences in  $\mu$  rates between on and off periods
- **Correction for  $\mu$  rate variation** with atmospheric pressure  $p_{\text{atm}}$  is made by extracting the dependency (on-off combined fit) and **re-normalising to a  $p_{\text{ref}}$**



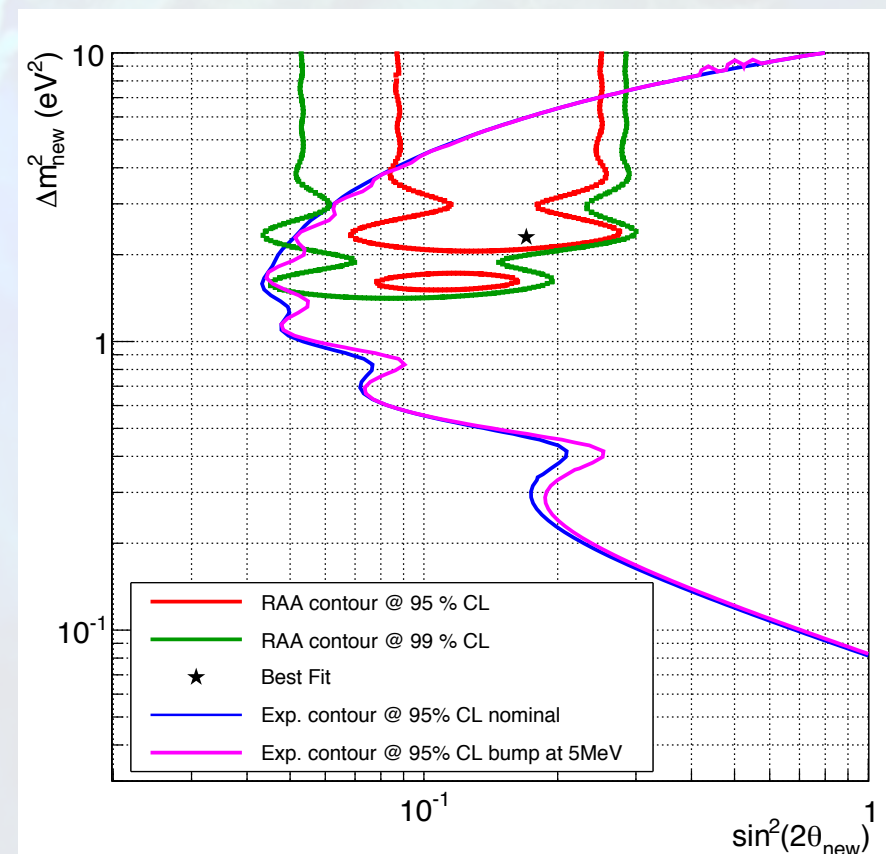
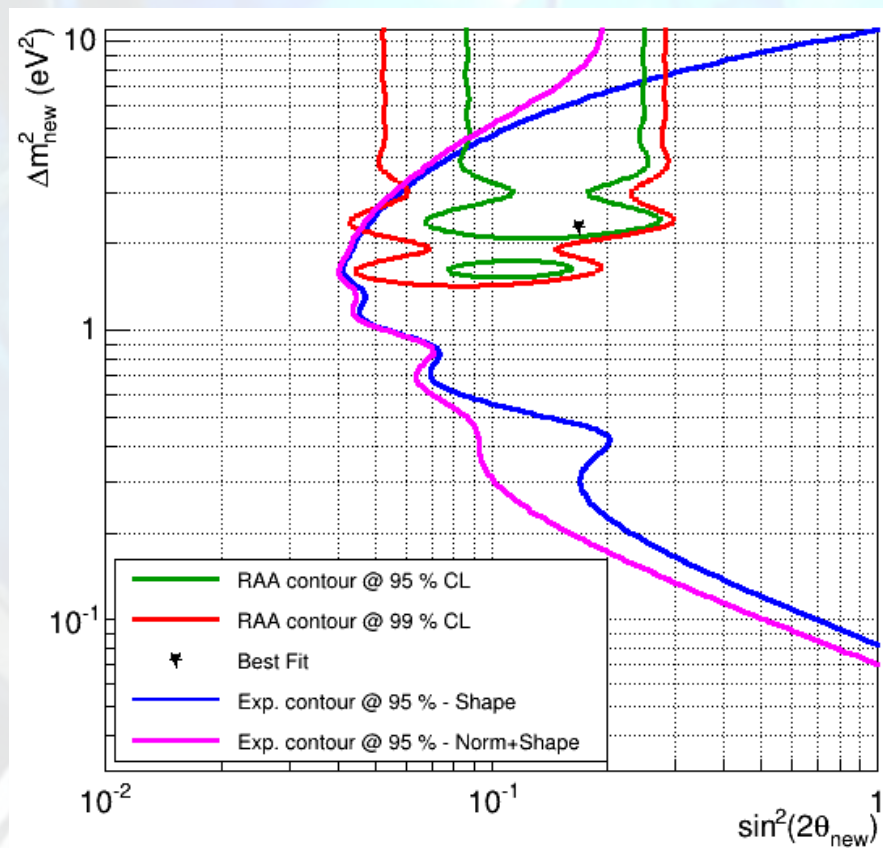


# STEREO Sensitivity & the 5 MeV Distortion

- STEREO will provide a new reference measurement of pure  $^{235}\text{U}$   $\bar{\nu}$  spectrum  $\Rightarrow$  isolate source of the distortion, constrain models
- Sensitivity after 300 days reactor ON, not reduced by distortion
  - 400  $\bar{\nu}$ /day in 300 days - **S/B = 1.5**
  - $E_{e^+} > 2 \text{ MeV}$ ,  $E_n > 5 \text{ MeV}$  -  $\varepsilon_{\text{det}} = 60\%$
  - $\delta E_{\text{scale}} = 2\%$ ,  $\delta \Phi_{\bar{\nu}} = 4\%$



arXiv:1512:xxx  
C. Buck, A.P. Collin, J. Haser, M. Lindner





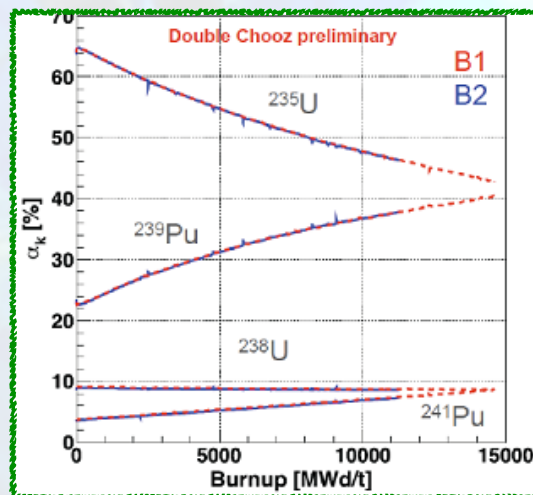
# Antineutrino Spectrum Estimation

- In LEU facilities the fuel evolution (**burnup**) involves mostly  $^{235}\text{U}$  and  $^{239}\text{Pu}$

$$N_{\text{IBD}}^{\text{exp}}(E_{\bar{\nu}}, t) = \frac{N_p \epsilon}{4\pi L^2} \times \frac{P_{th}(t)}{\langle E_f \rangle(t)} \times \langle \sigma_f \rangle(E_{\bar{\nu}}, t)$$

Average E released per fission

$$\langle E_f \rangle = \sum_k \alpha_k \langle E_F \rangle$$



Average Cross Section per fission

$$\langle \sigma_f \rangle_k = \int dE S_k(E) \sigma_{\text{IBD}}(E)$$

Neutrinos per fission per MeV (prediction)

CS (theoretical + corrections)

$k = \text{isotope } (^{235}\text{U}, ^{238}\text{U}, ^{239}\text{Pu}, ^{241}\text{Pu})$

- Single  $\bar{\nu}$  spectra  $S_k(E)$  not available, obtained from  $\beta$  spectrum

Mueller et al., *Phys. Rev. C* 83.5 (2011): 054615

- Model currently most used: Mueller ( $^{238}\text{U}$ ) - Huber ( $^{235}\text{U}$ , Pu) [Huber P., \*Phys. Rev. C\* 84.2 \(2011\): 024617](#)



# Light Sterile Search @ Reactors

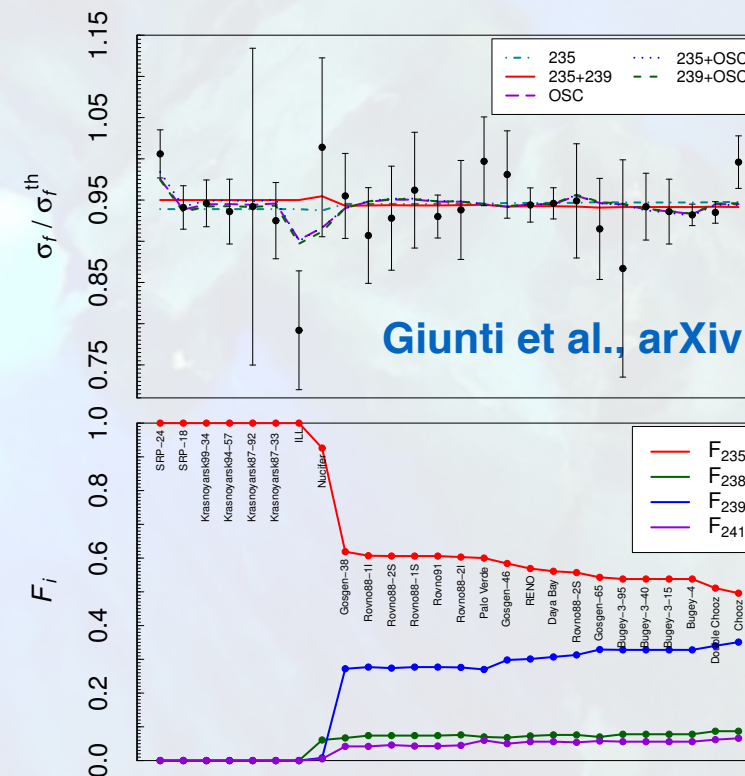
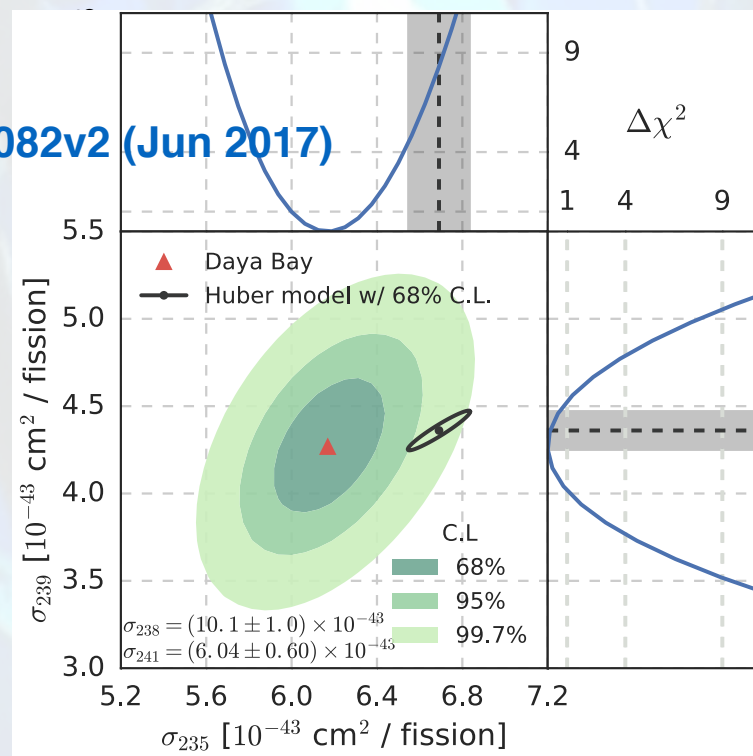
	Core $P_{Th}$	Core Size	Overburden	Segmentation	Baseline	Material
<b>Chandler</b>	72 MW ( $^{235}U$ )	$\varnothing = 50$ cm	$\sim 10$ mwe	6.2 cm (3D)	5.5 m	PS + Li layer
<b>DANSS</b>	3 GW (LEU)	$h = 3.6$ m $\varnothing = 3.1$ m	$\sim 50$ mwe	5 cm (2D)	10.7-12.7 m	Gd-doped PS
<b>NEOS</b>	2.8 GW (LEU)	$h = 3.7$ m $\varnothing = 3.1$ m	$\sim 20$ mwe	-	23.7 m	Gd-doped LS
<b>Neutrino4</b>	90 MW ( $^{235}U$ )	35x42x42 cm <sup>3</sup>	few mwe	22.5 cm (2D)	6-12 m	Gd-doped LS
<b>NuLat</b>	40/1790 MW ( $^{235}U$ /LEU)		few mwe	6.35 cm (3D)	4.7/24 m	Li-doped PS
<b>Prospect</b>	85 MW ( $^{235}U$ )	$h = 0.5$ m $\varnothing = 0.2$ m	few mwe	15 cm (2D)	7 m	Li-doped LS
<b>SoLiδ</b>	72 MW ( $^{235}U$ )	$\varnothing = 0.5$ m	$\sim 10$ mwe	5 cm (3D)	5.5 m	PS + Li layer
<b>Stereo</b>	58 MW ( $^{235}U$ )	$\varnothing = 37$ cm	$\sim 15$ mwe	25 cm (1D)	8.9 m	Gd-doped LS



# Sterile Neutrino vs Flux Normalisation

- Daya Bay slightly favours  $^{235}\text{U}$   $\bar{\nu}_e$  flux suppression over sterile hypothesis
- Situation is reversed for other experiments (flux deficit independent from fuel composition)
- $^{235}\text{U}$  and  $^{239}\text{Pu}$  fluxes are normalised on separate “vintage”  $\beta$ -spectrum measurements @ ILL

arXiv:1704.01082v2 (Jun 2017)



Giunti et al., arXiv:1708.01133v1 (Aug 2017)

- Daya Bay (but also DANSS and NEOS) are LEU experiments (mixed fuel)
- **Need results from short baseline HEU experiments to solve the conundrum**



# The Global Picture

- If we include NEOS and DANSS recent results in the game, the sterile hypothesis cannot still be rejected
- This is mainly due to (matching) features in the experimental spectra, reflected in the exclusion plot, that prefer oscillation rather than scaling
- **A global analysis from multiple experiments can discriminate whether these are fluctuations or there is an oscillatory pattern**

Gentler et al., arXiv:1709.04294v1 (Sep 2017)

