

# Astromaterials in Accreting Neutron Stars

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June 15, 2017



U.S. DEPARTMENT OF  
**ENERGY**

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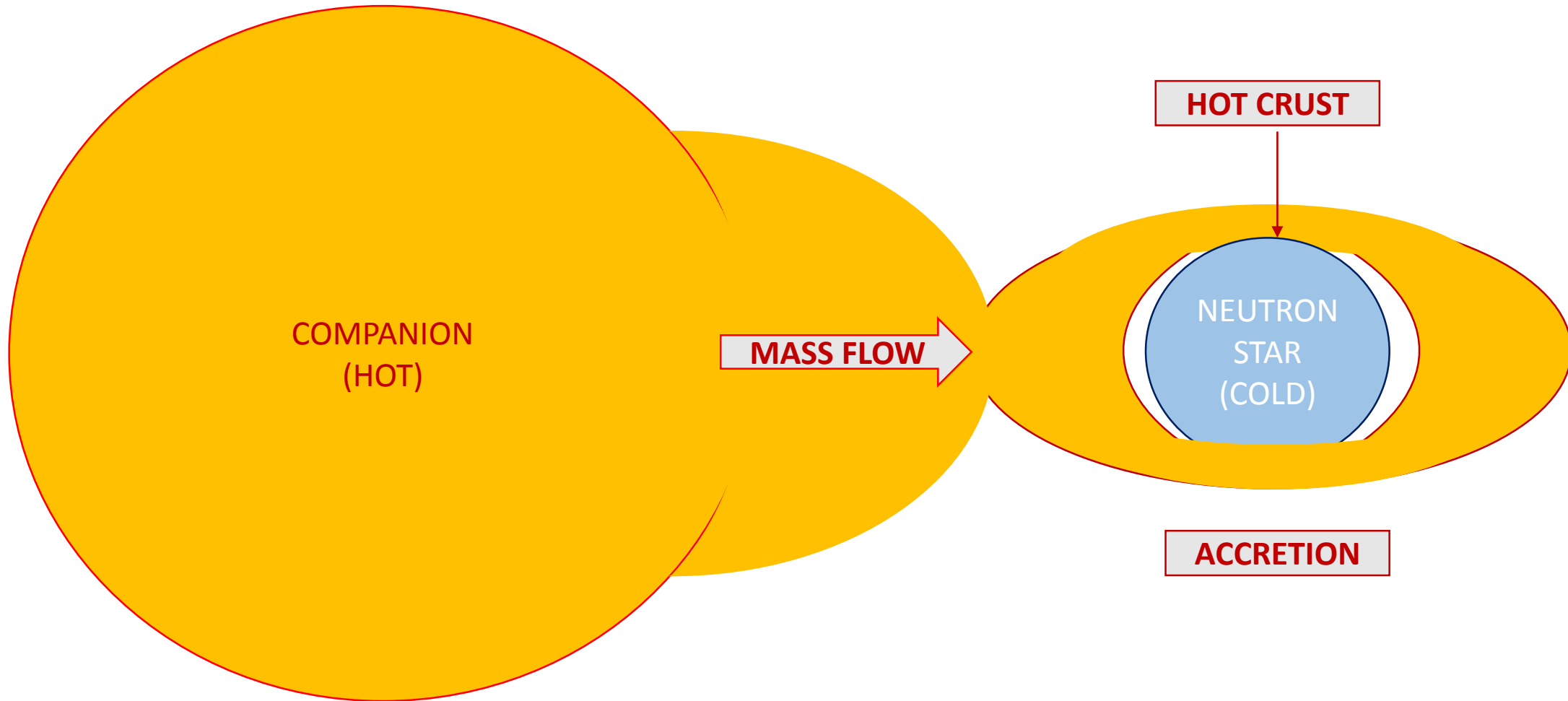
**NUCLEI**  
Nuclear Computational Low-Energy Initiative



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# X-Ray Binaries

# Low Mass X-Ray Binaries



## Quiz 1



Hydrogen on the surface of a neutron star is being heated and compressed.

What happens now?

## Quiz 1

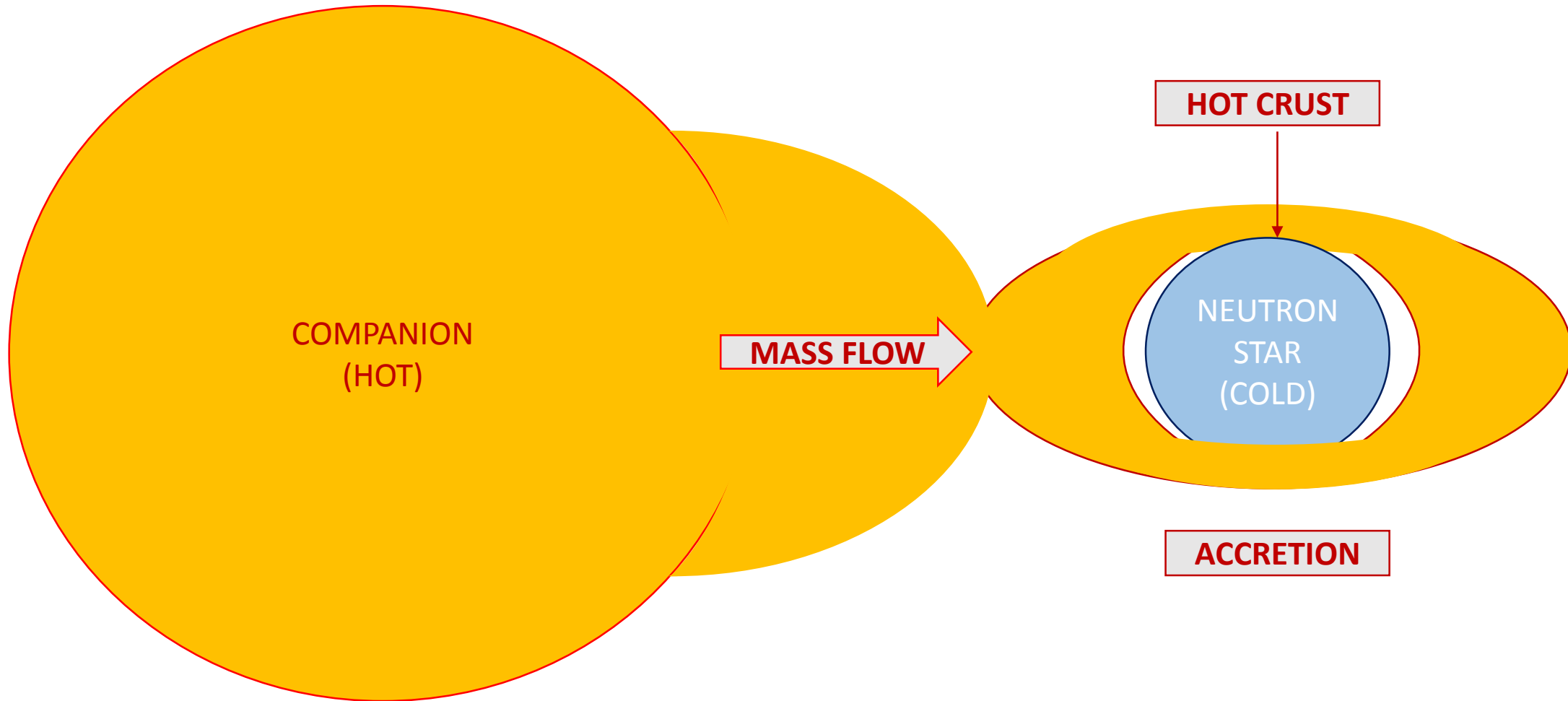


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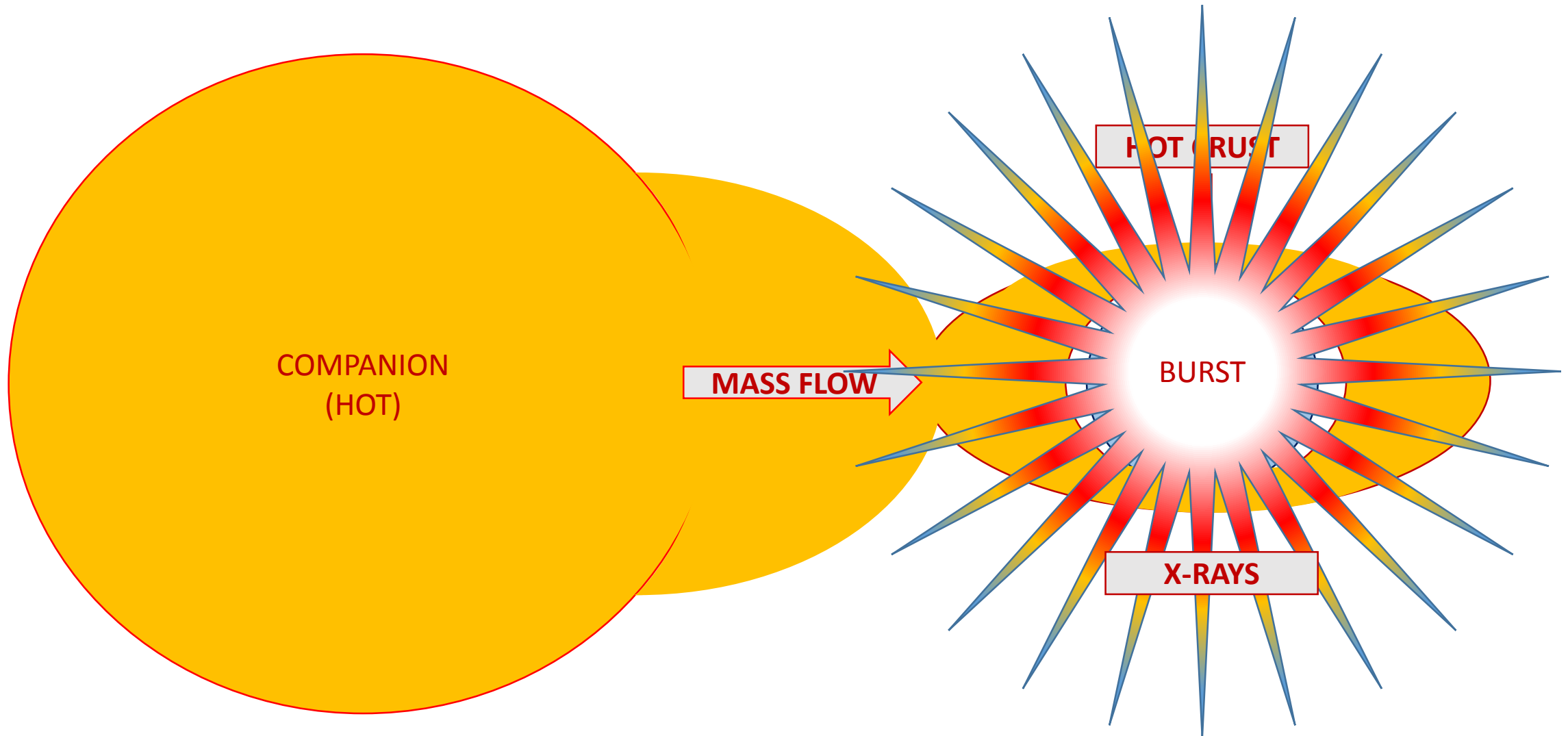
What happens now?

**A: Explosive nuclear burning.**  
(Similar to a WD nova)

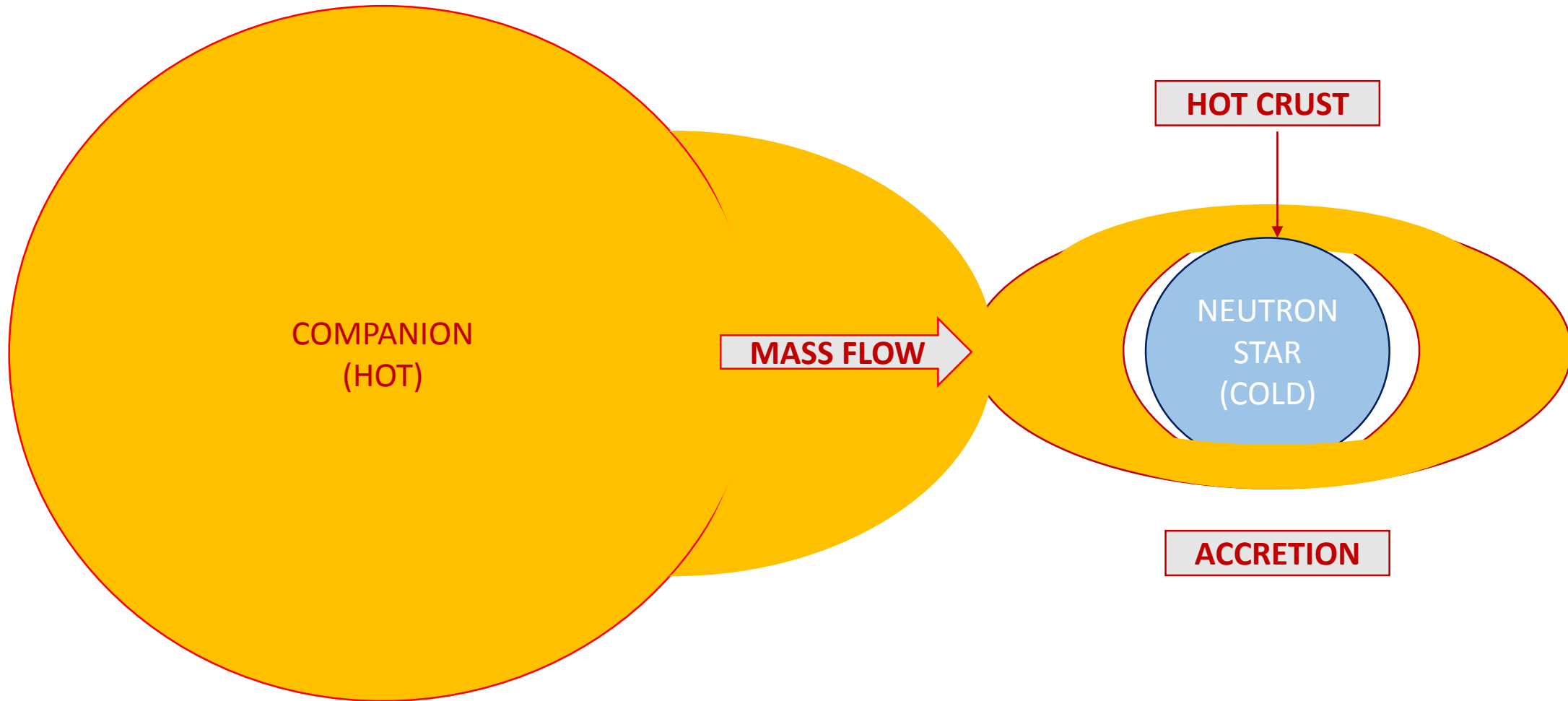
# Low Mass X-Ray Binaries



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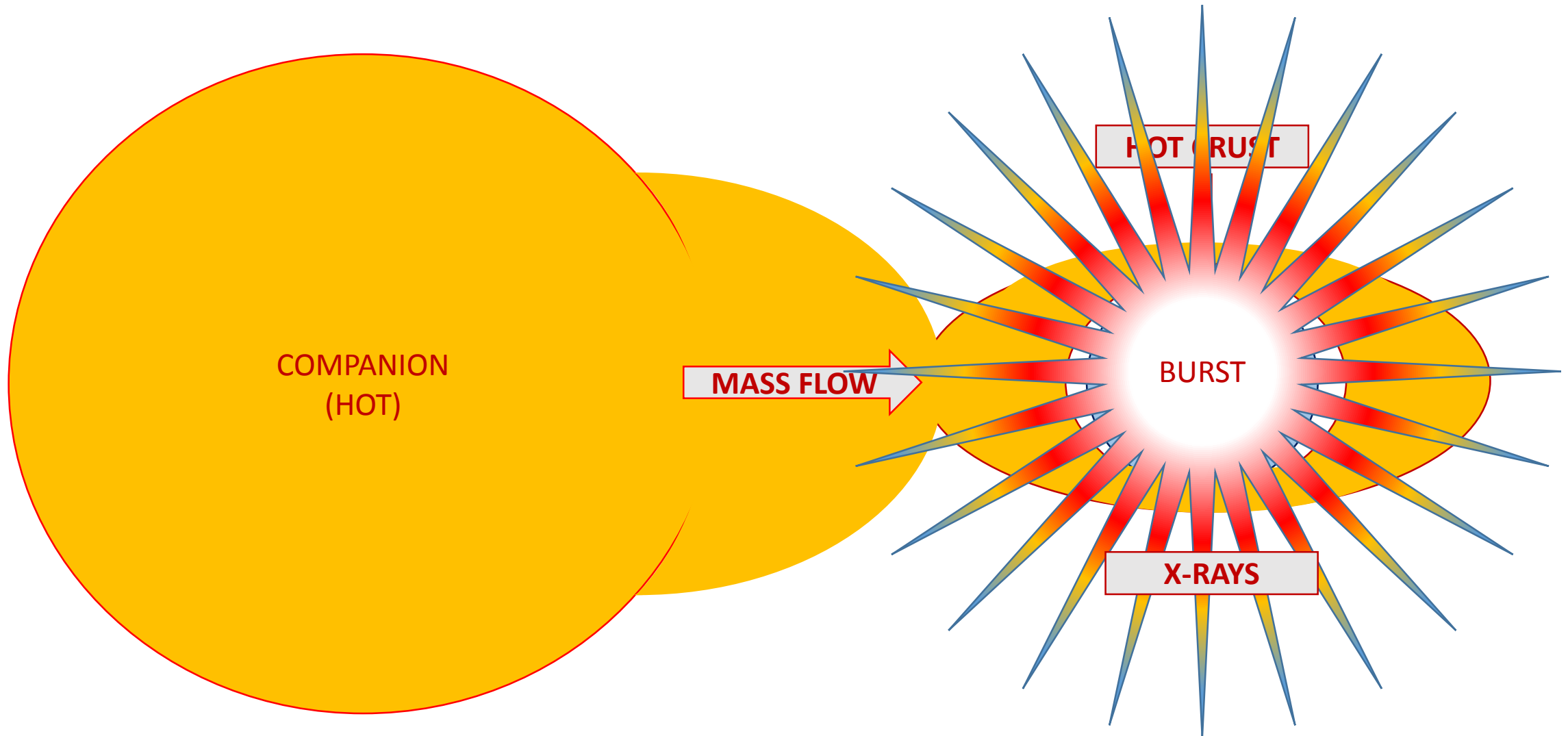


# Low Mass X-Ray Binaries





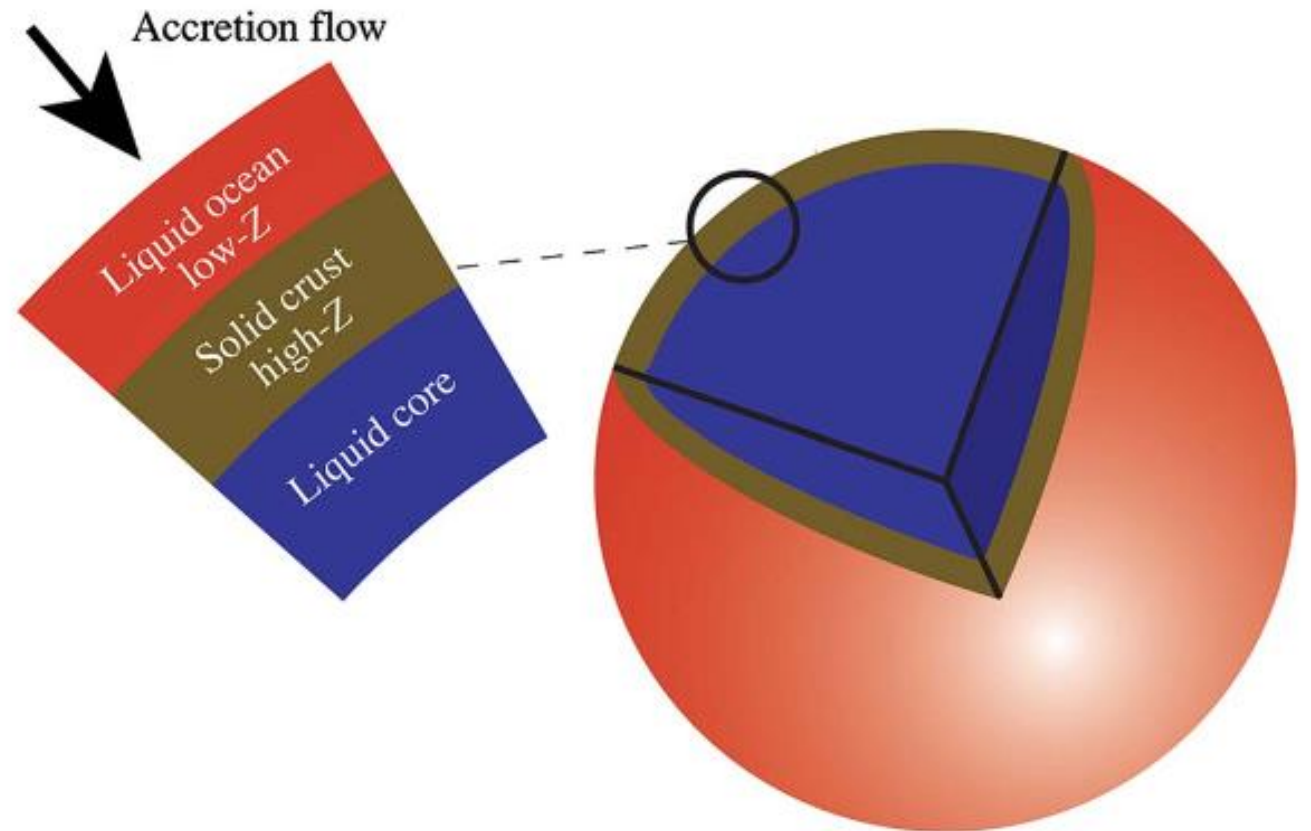
# Low Mass X-Ray Binaries



# X-ray bursts



- As matter accretes, it is compressed, buried, and heated
- Explosive nuclear burning produces a mix of heavy nuclei (rp-process), and an X-ray burst
- Ash is buried, and crystallizes
- This matter forms the crust



## Quiz 2



What is a material?

## Quiz 2



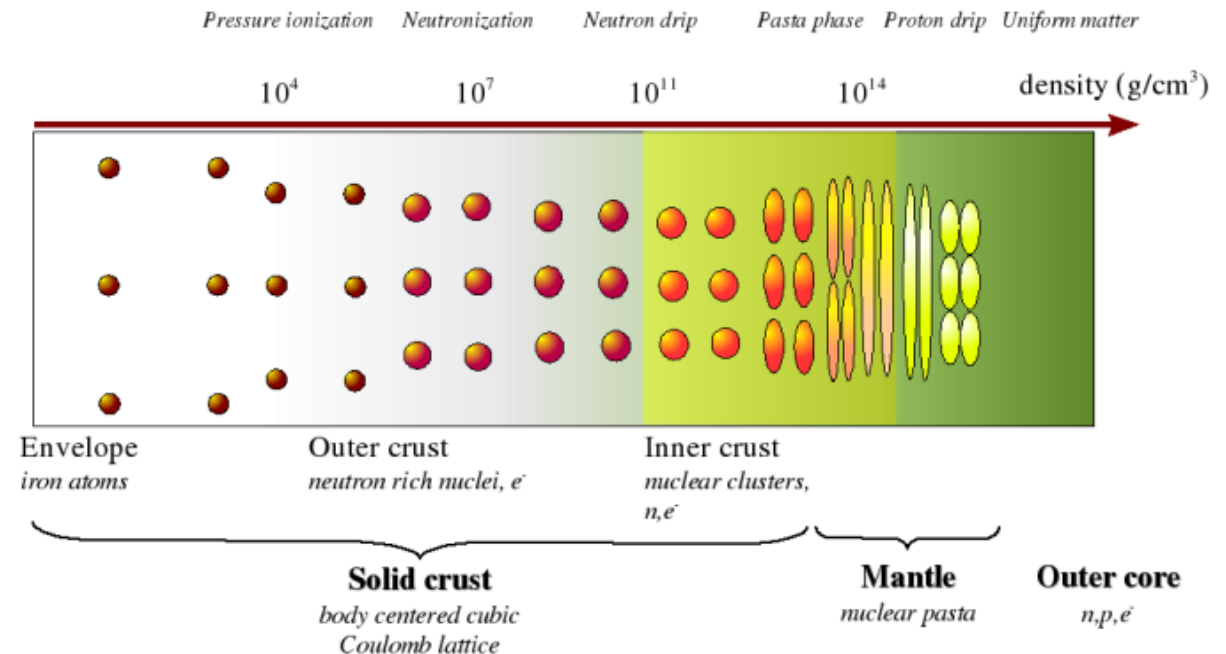
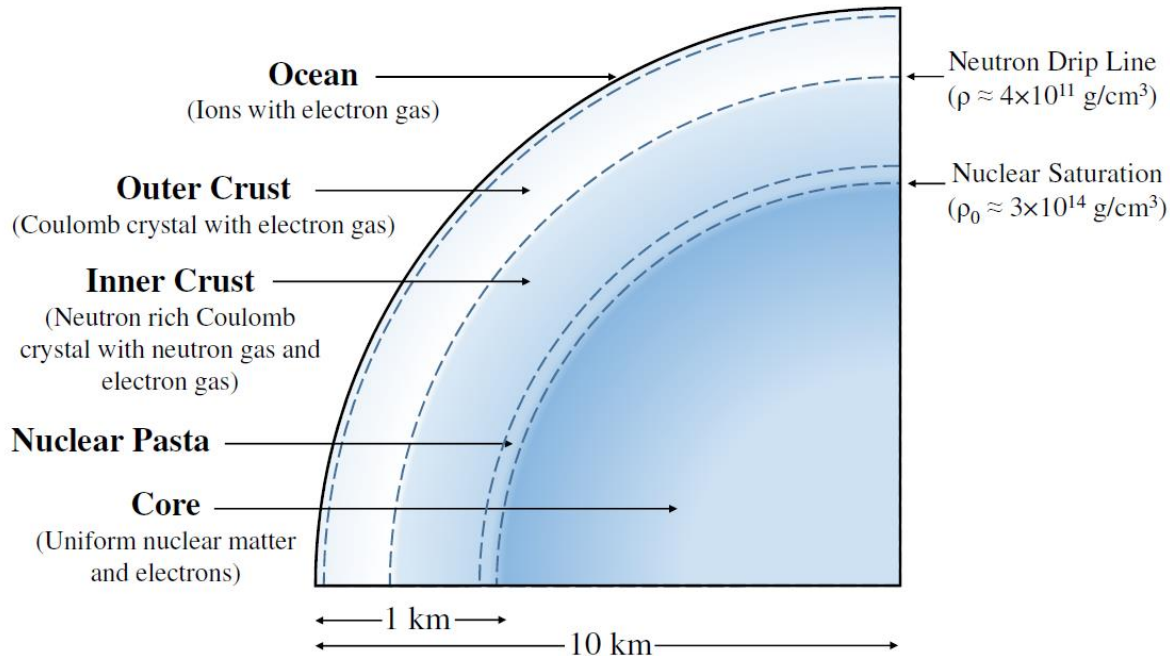
What is a material?

A: Hard to define. But, at least,  
a material is not a fluid.

# Neutron Star Structure



- What's inside a neutron star? Crust and core, similar to earth...

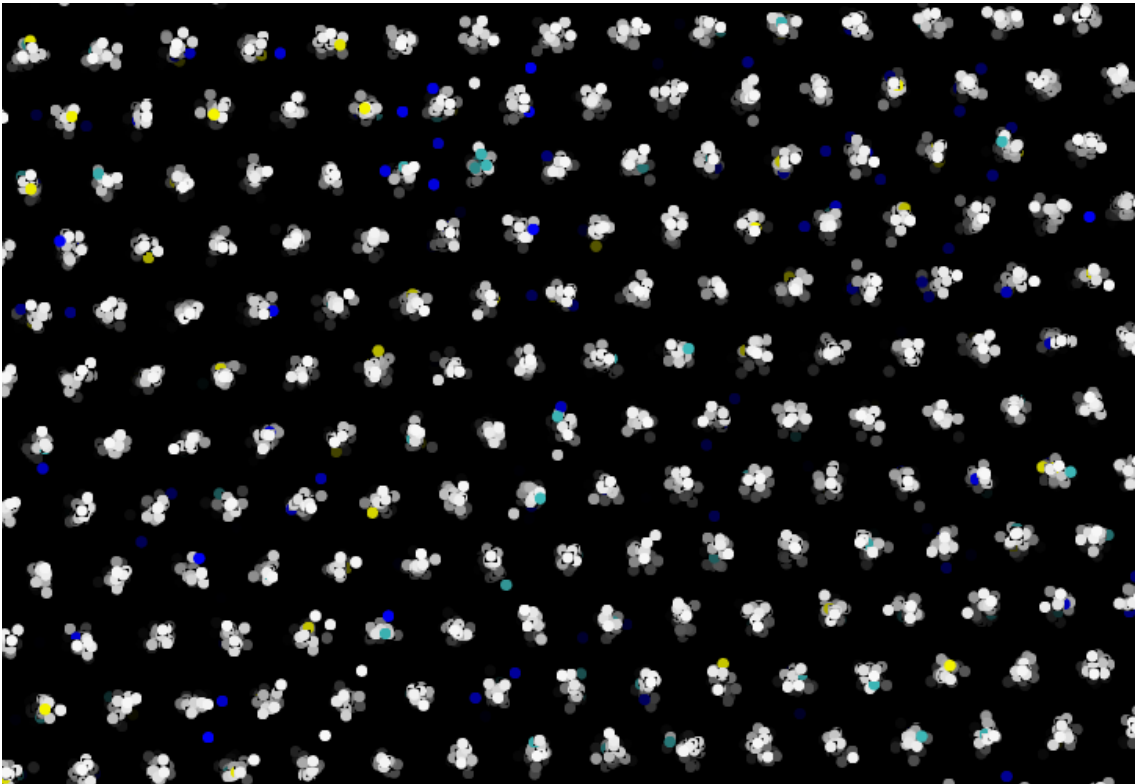


- Not just a “giant nucleus in space!”

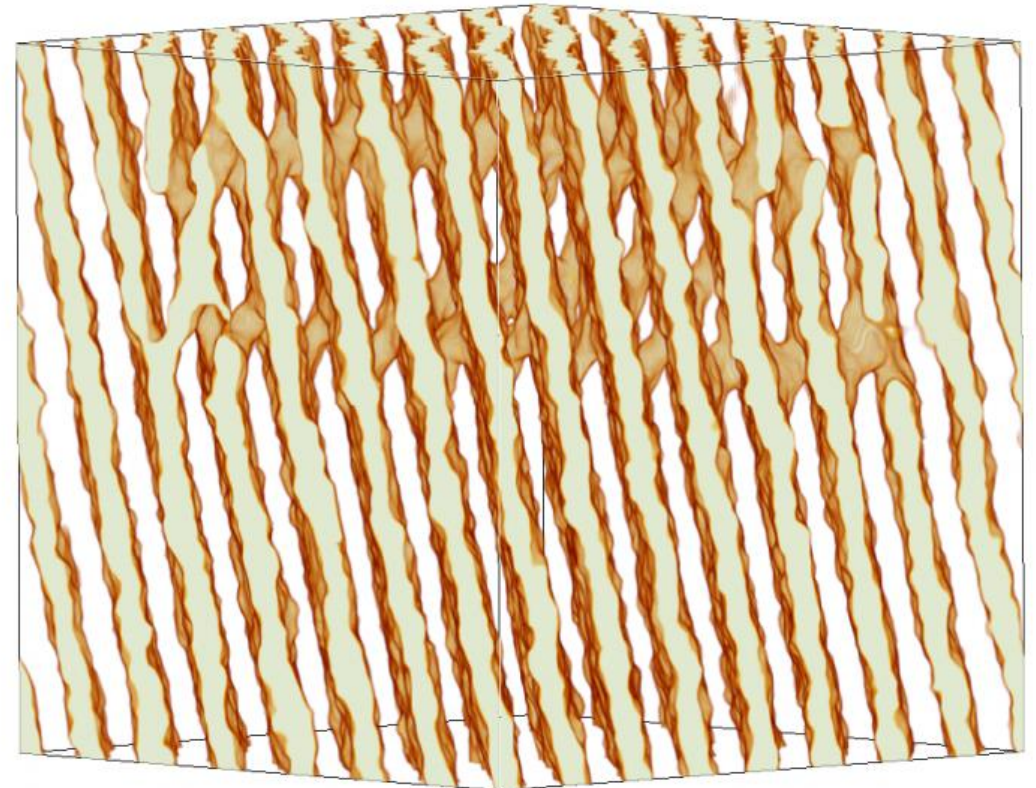
# Astromaterials



- The neutron star crust contains *materials*, like a crystalline outer crust, and a ‘goosey’ inner crust of nuclear matter, like a liquid crystal.

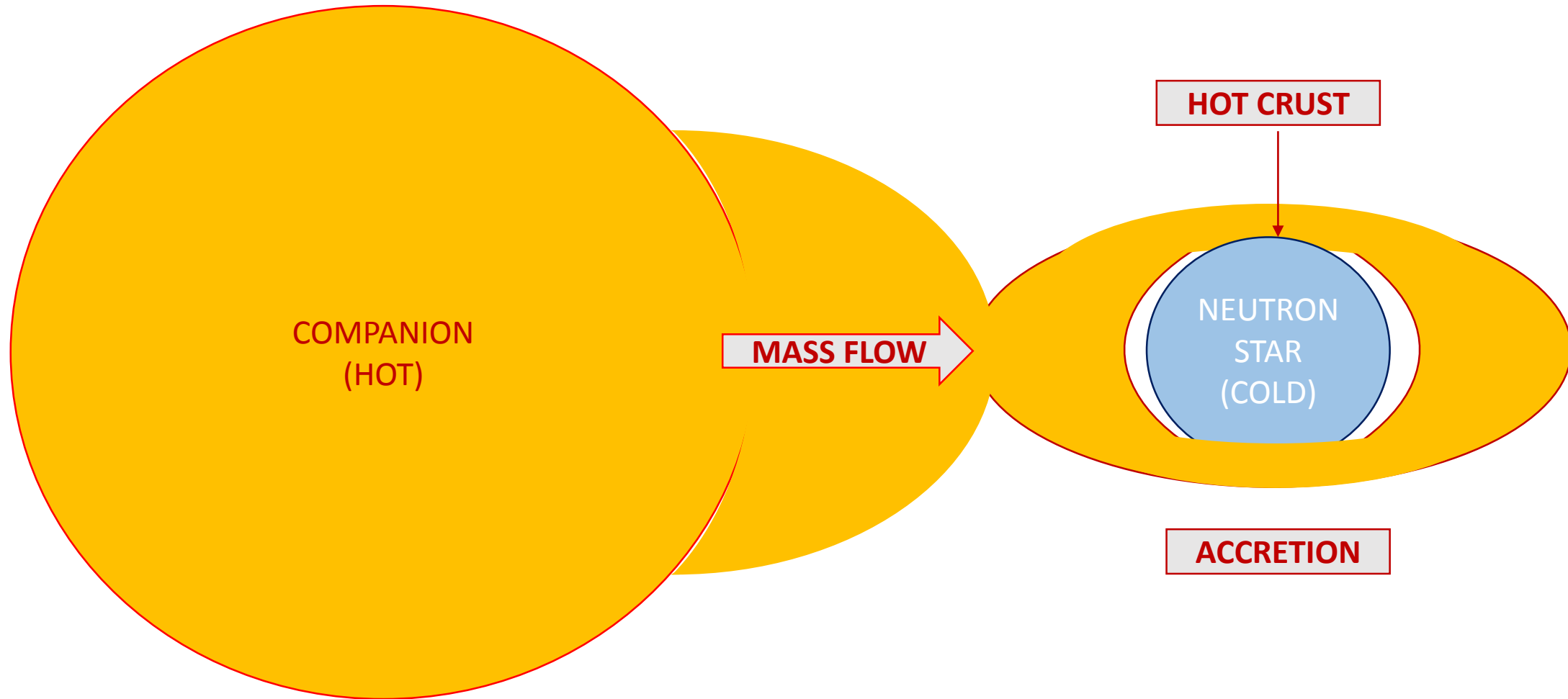


**HARD (Outer Crust)**

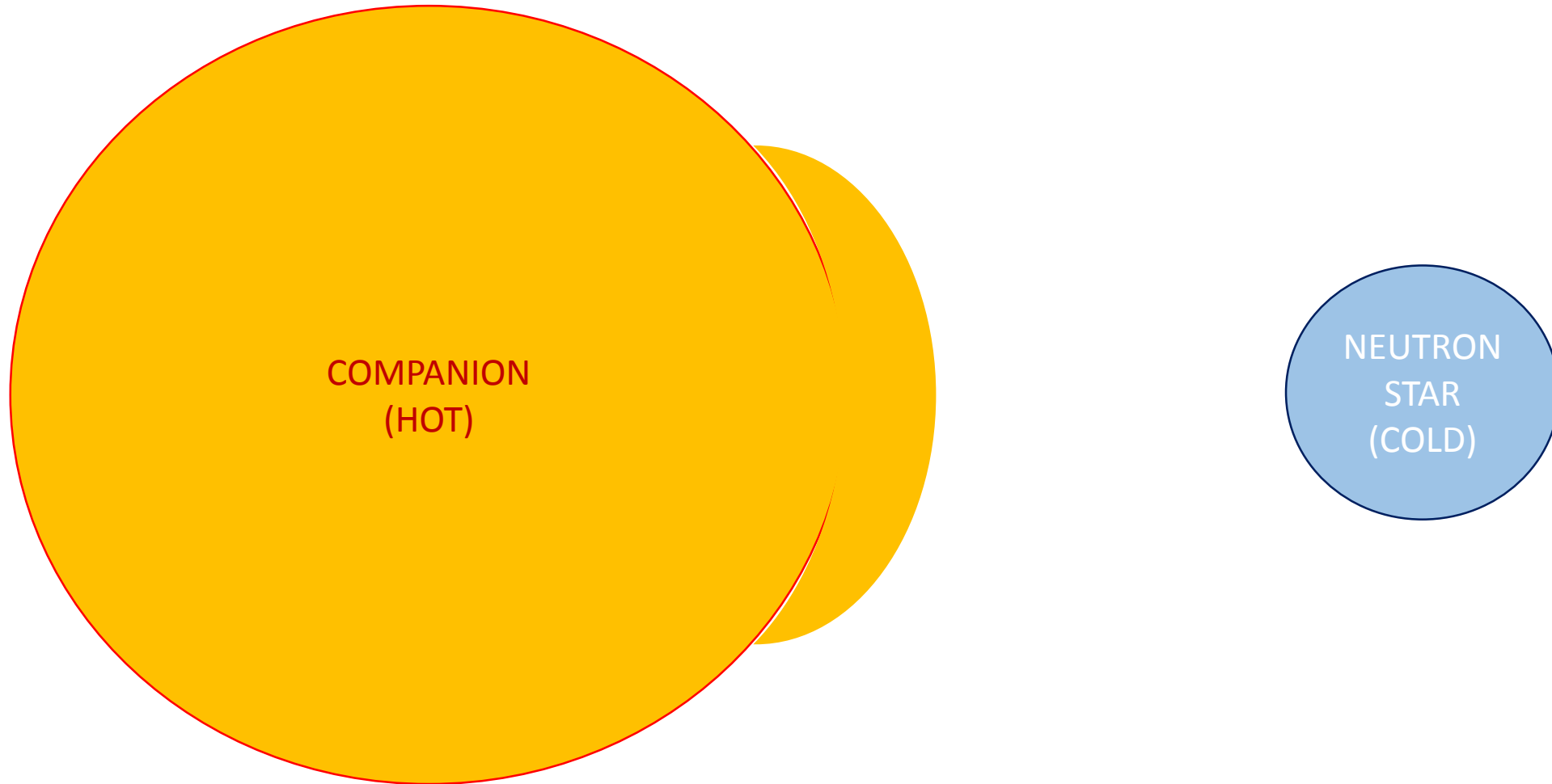


**SOFT (Inner Crust)**

# Low Mass X-Ray Binaries



# Low Mass X-Ray Binaries





## Quiz 3

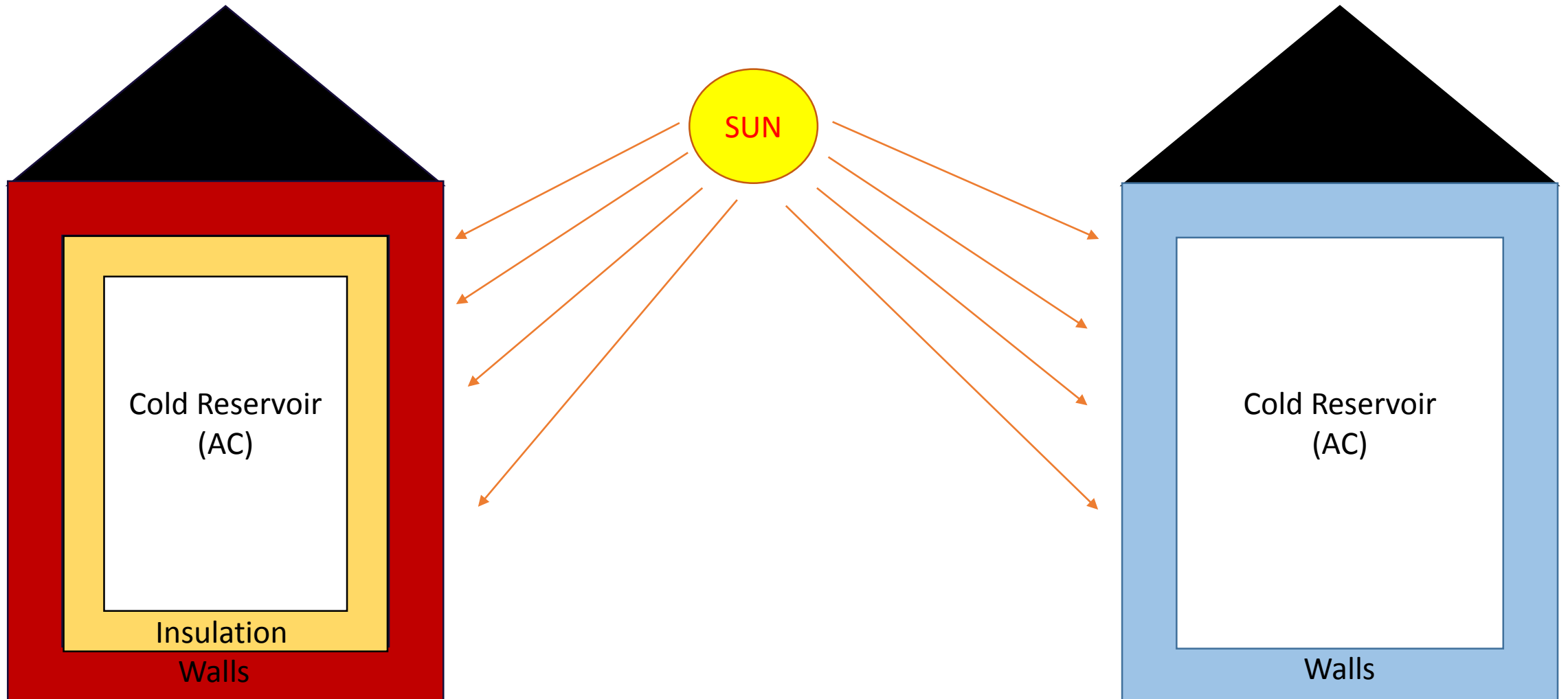


Suppose the accretion stops.

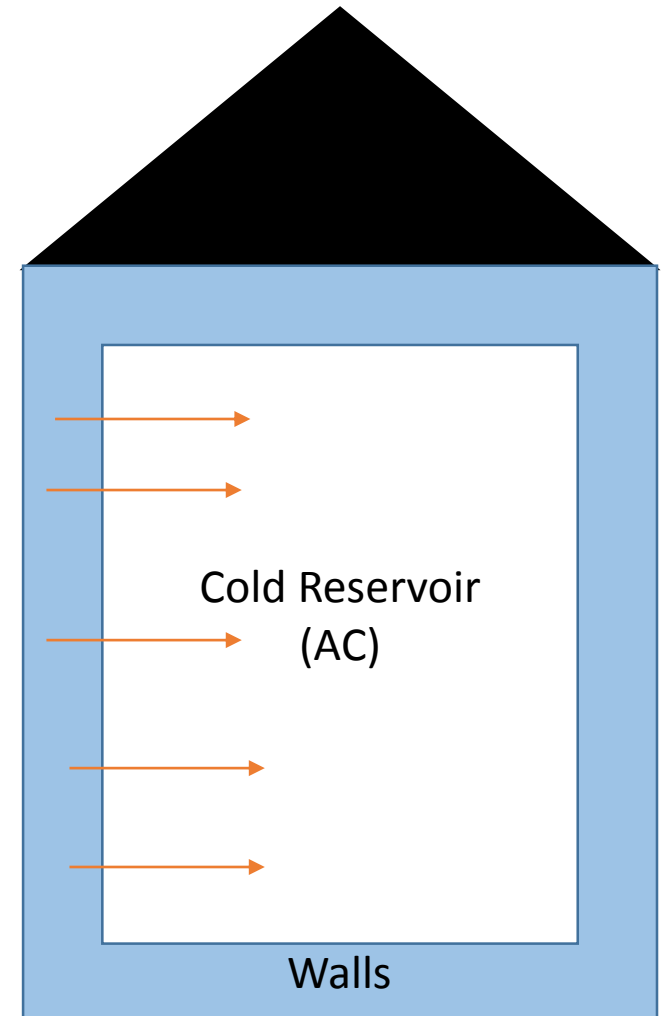
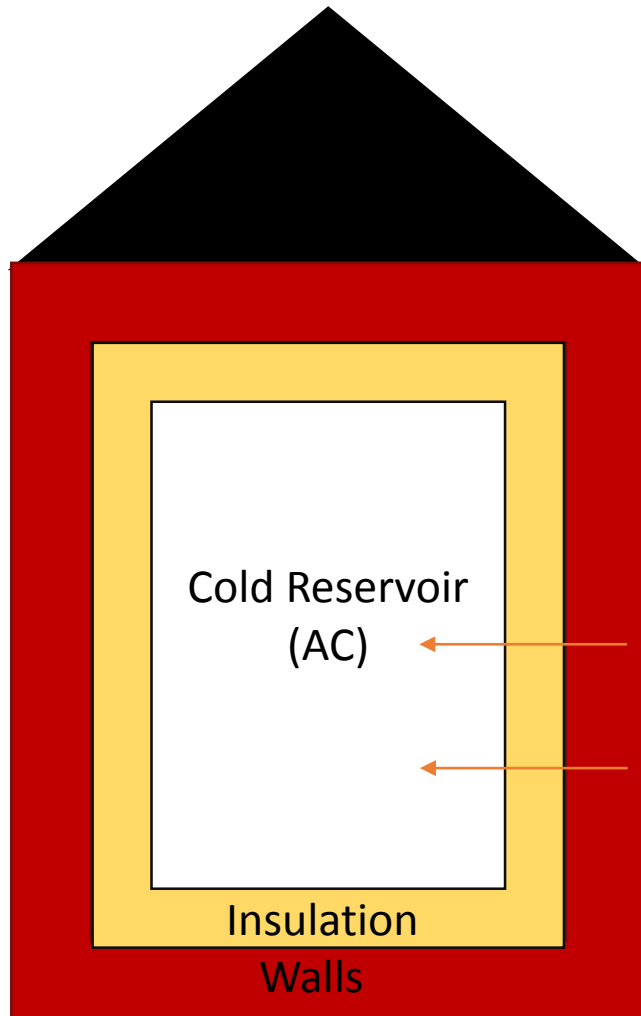
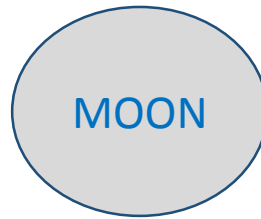
What happens now?

A: The neutron star crust cools.

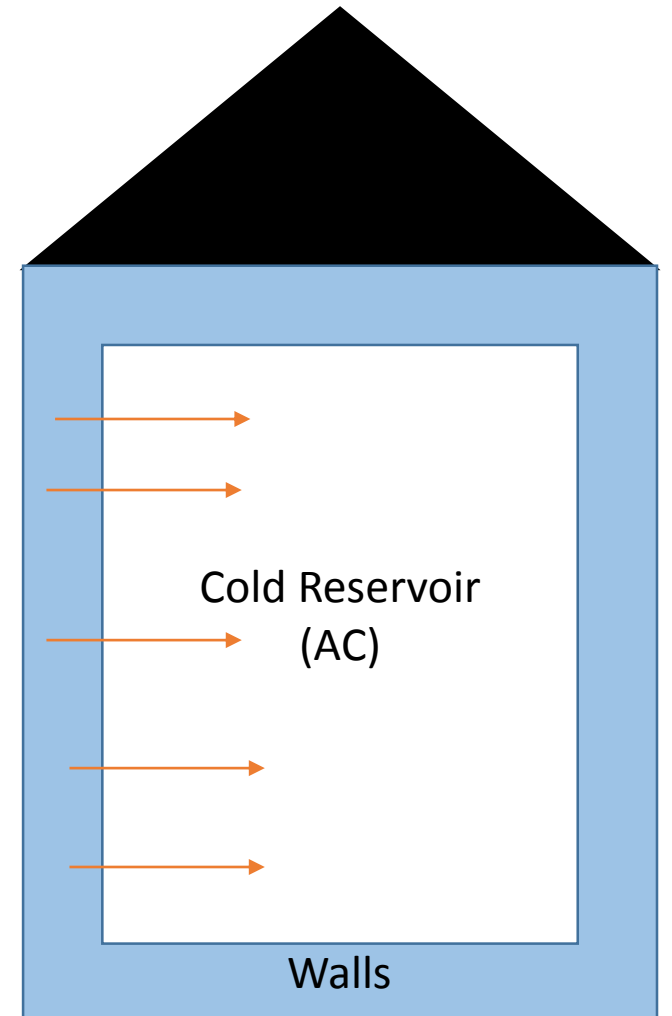
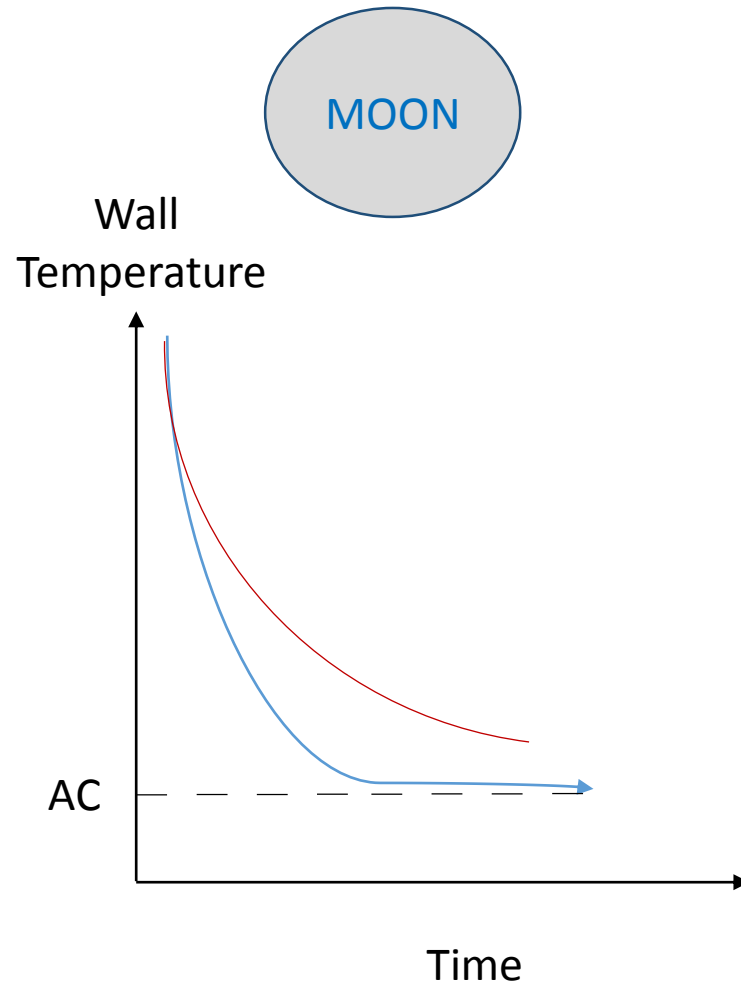
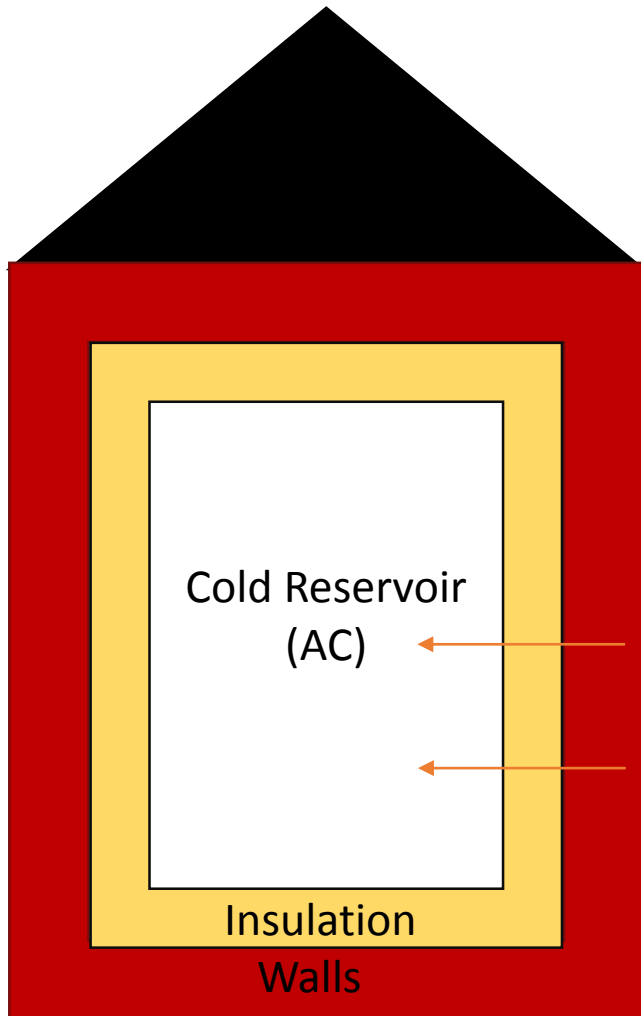
# An Example



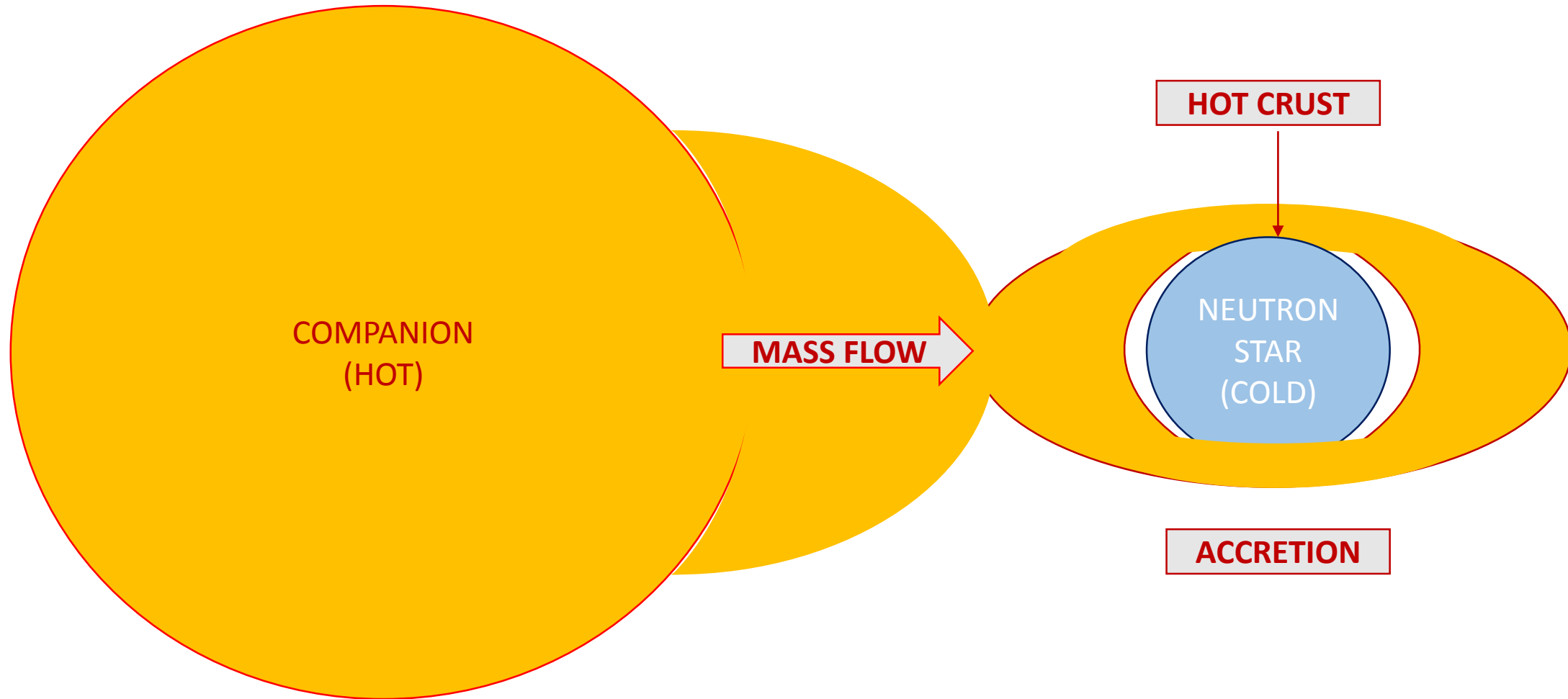
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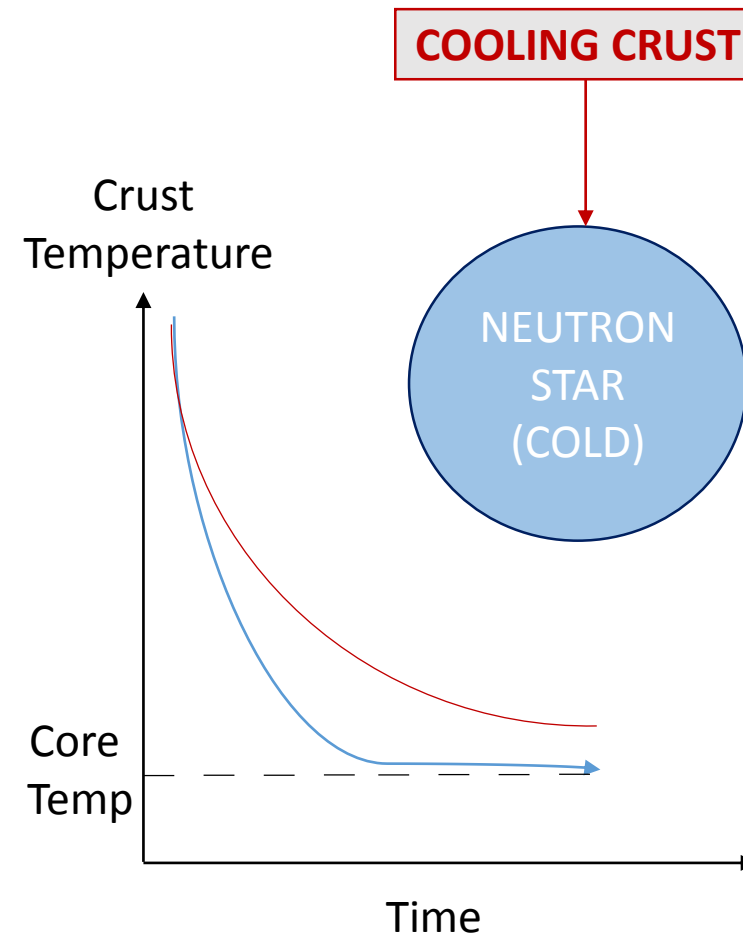
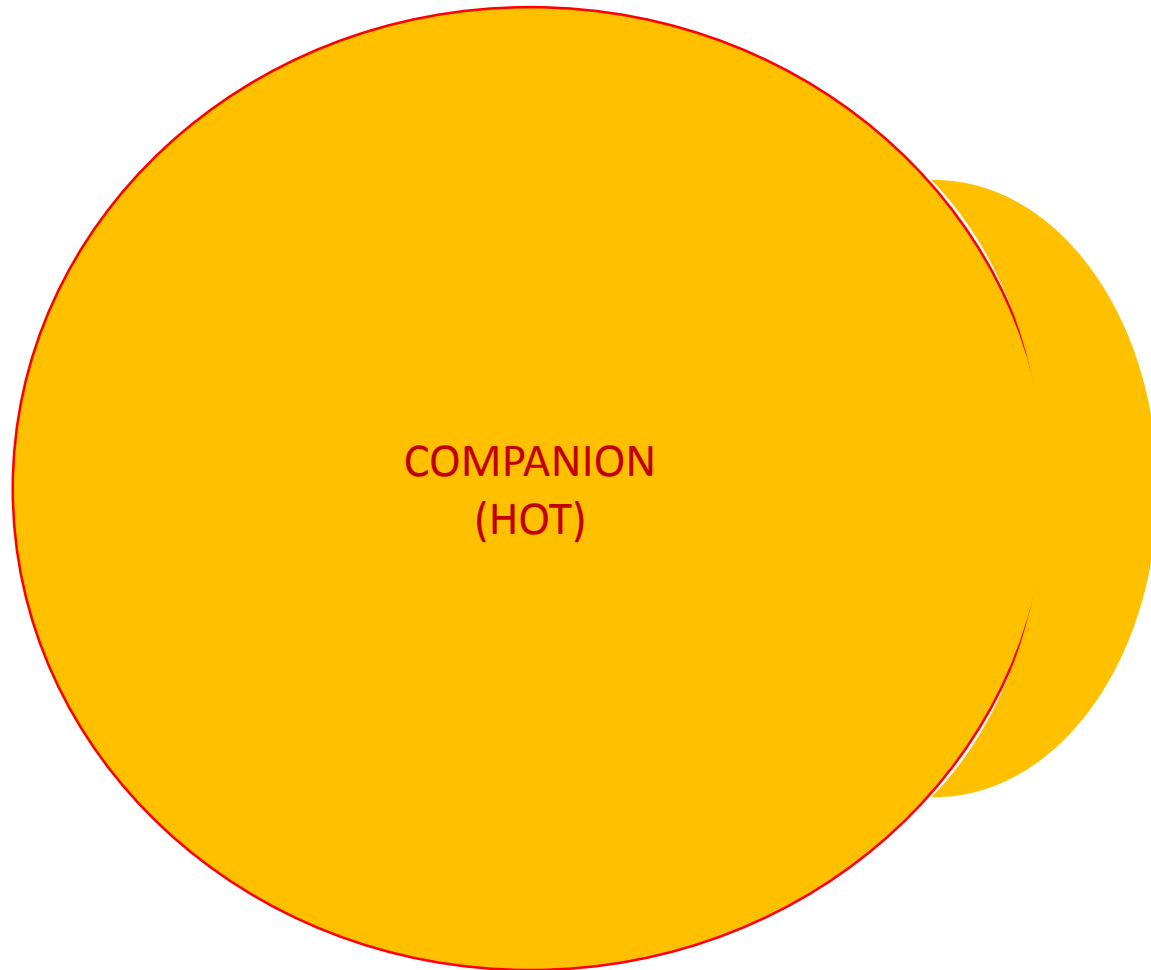
# An Example



# Low Mass X-Ray Binaries



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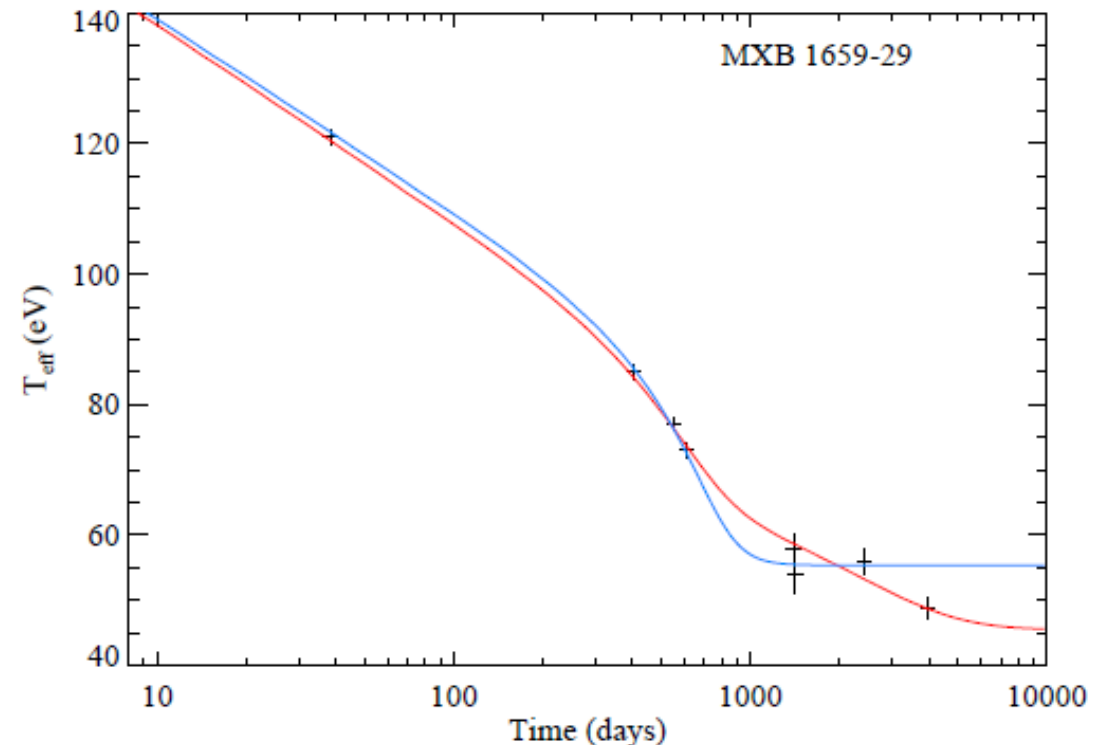
# Observables – Thermal Properties



- Use our models to study neutron star interiors from their cooling
- Cooling curves: low mass X-ray binary MXB 1659-29

$$Q_{\text{imp}} \equiv n_{\text{ion}}^{-1} \sum_i n_i (Z_i - \langle Z \rangle)^2$$

- **Blue**: Model with only a conductive crust
- **Red**: Model with conductive crust, a deep insulating layer



## Quiz 4



What should we do next?



## Quiz 4



What should we do next?

**A: Find more sources, make more observations.**

# Summary



To interpret observations of neutron stars, we must first develop microscopic models of their interiors. By simulating the kinds of matter we expect to find in the crust we can calculate properties of the star, allowing us to interpret observations.