



cherenkov
telescope
array

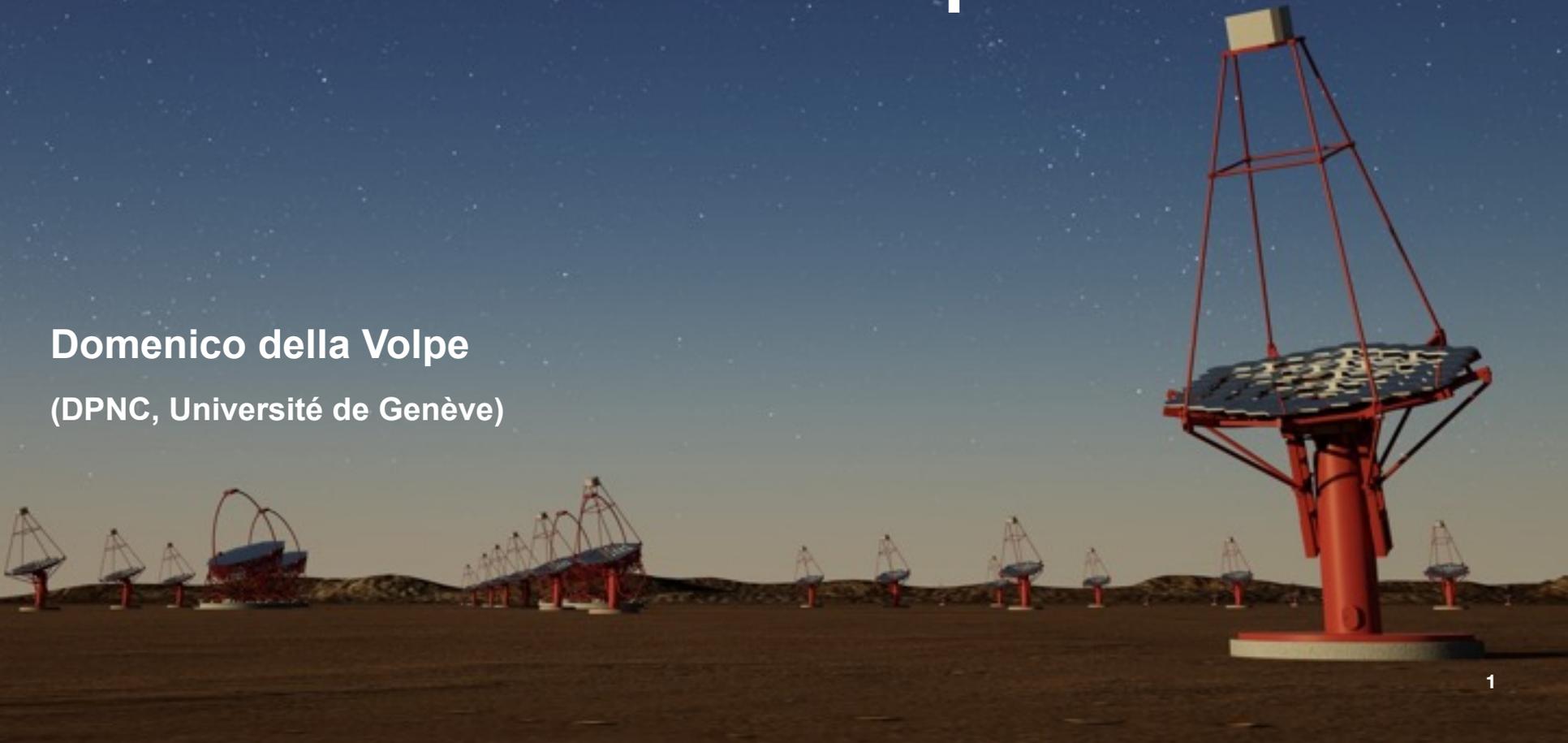


UNIVERSITÉ
DE GENÈVE

FACULTÉ DES SCIENCES

Innovative design of a camera for IACT telescopes

Domenico della Volpe
(DPNC, Université de Genève)



The context

- The Cherenkov Telescope Array will be composed of 3 different sizes of telescopes: Large (LST), Medium (MST) and Small (SST)
- Different optics and a camera designs are proposed for these telescopes
- The Geneva group is the leader of one of the proposed SST telescopes using a Davies-Cotton single mirror design and proposing an innovative fully digital camera based on SiPM
- The camera presented here is the design proposed for this Single mirror Small Size telescope (SST-1M)



Why SiPMs

Large scale mass production

No Ageing

Lightweight

High yield

Low Op. Voltage

Higher PDE over larger spectral range

Robust

Insensitive to Magnetic Field

Compact

Very good Single Photon Response

Small variation sample-by sample

Why SiPM

- High Quantum Efficiency
- Single Photon sensitivity
- Fast pulses
- Low noise
- High Fill factor
- Robustness
- Uniformity
- High dynamic range
- Large area to be covered
- Linear response

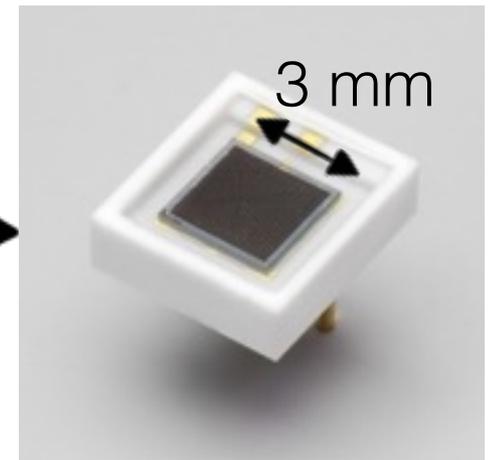
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Improve

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Worse

40 mm



Why SiPM

Improve

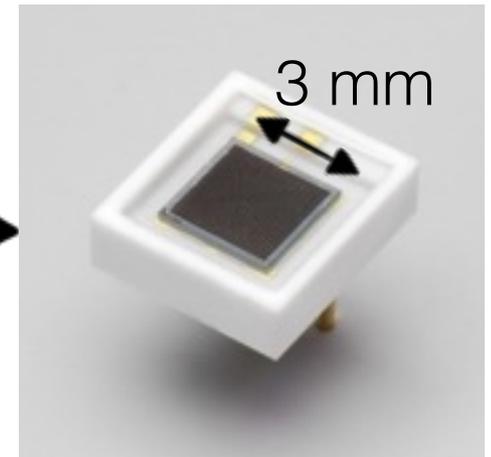
- High Quantum Efficiency
 - Single Photon sensitivity
 - Fast pulses
 - Low noise (dark count, afterpulses)
 - High Fill factor
 - Robustness
 - Uniformity
 - High dynamic range
 - Large area to be covered
 - Linear response (optical cross talk, pile up)
-
- Lower voltage and easier cooling
 - Lightweight
 - High potential for performance improvement and cost decrease
 - Characteristics depend on operation temperature

Worse

40 mm



3 mm



Why SiPM

Improve

Worse

- Fast pulses
- Low noise (dark count, afterpulses)

- Large area to be covered
- Linear response (optical cross talk, pile up)

- Characteristics depend on operation temperature

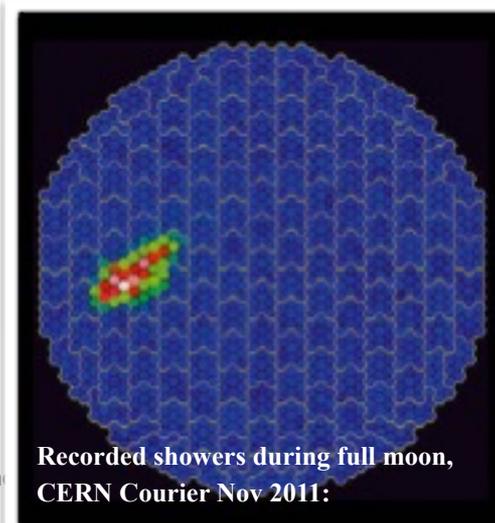
Are these parameters worrisome for gamma ray astronomy ?

Why SiPM in gamma-ray Astronomy

- Excellent single PE sensitivity
- Lightweight and robust cameras
- No evidence of ageing after 18 months
- Night Sky Background (NSB) rate dominates wrt Dark noise (MHz)
- Current Photo-Detection Efficiency > 40%.
- Operation during Moonlight: ~30% larger duty cycle
- As demonstrated by FACT, SiPM work on the field and with moonlight!



New approach, use fully digital SiPM-based camera on a Davies-Cotton telescope.



Telescope Design Drivers

Conditions:

Dish = 4 m



energy threshold

FoV = 9°



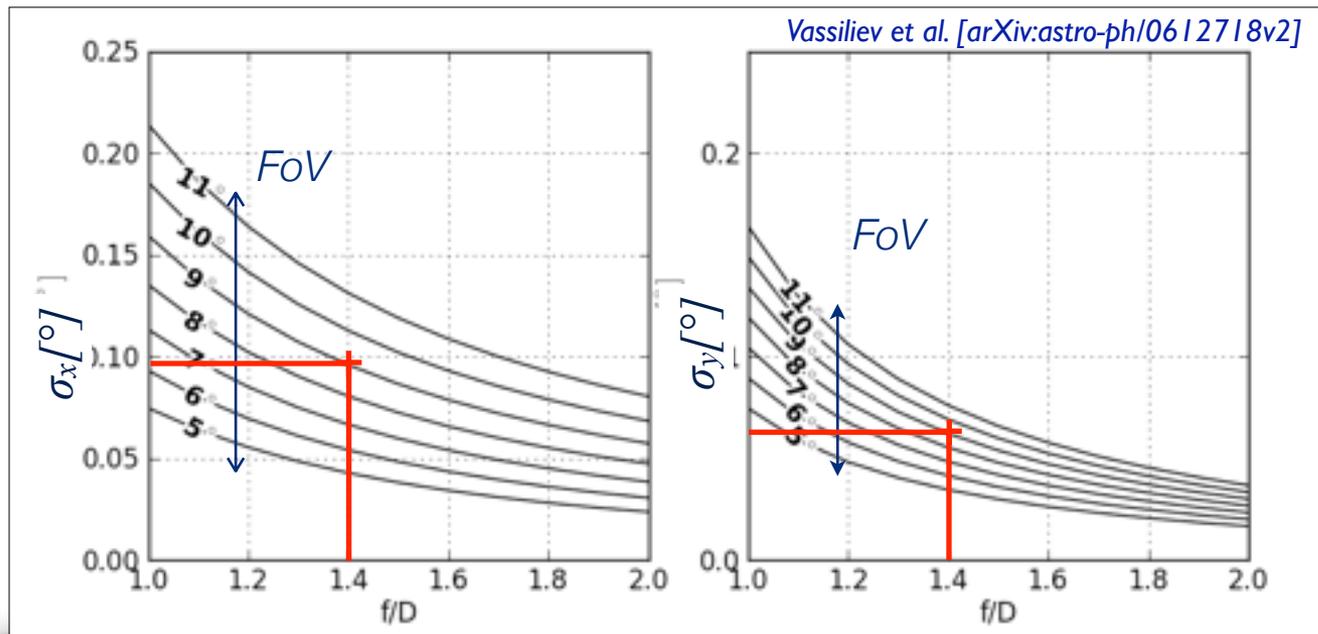
explore the galactic center, map extended source

f/D = 1.4



suitable angular resolution

PSF OF A 4M DAVIES-COTTON



Telescope Design Drivers

Too big for a SiPM!!

Conditions:

Dish = 4 m

FoV = 9°

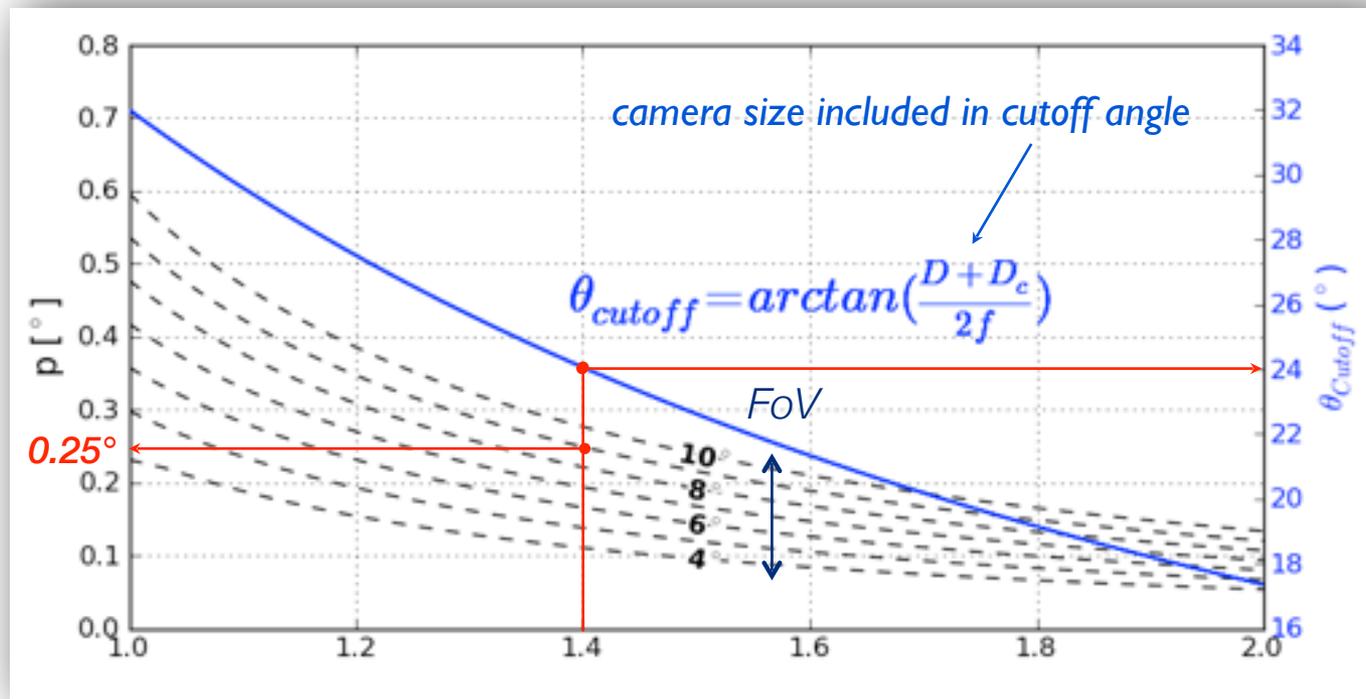
f/D = 1.4

pixel size = $4 \cdot \min(\sigma_x, \sigma_y) = 0.25^\circ$

pixel size (linear) = 2.44 cm

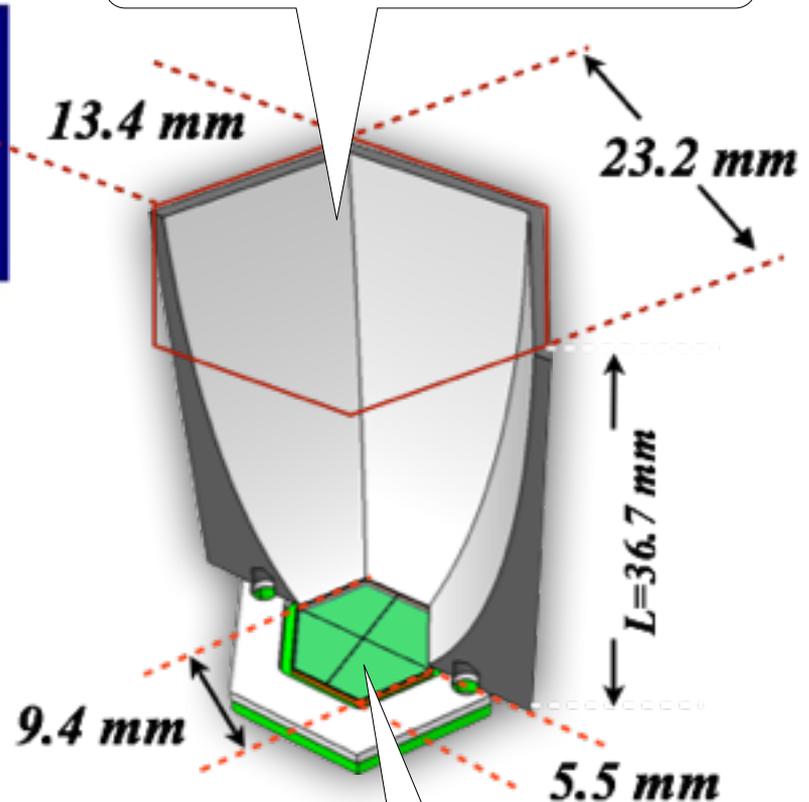
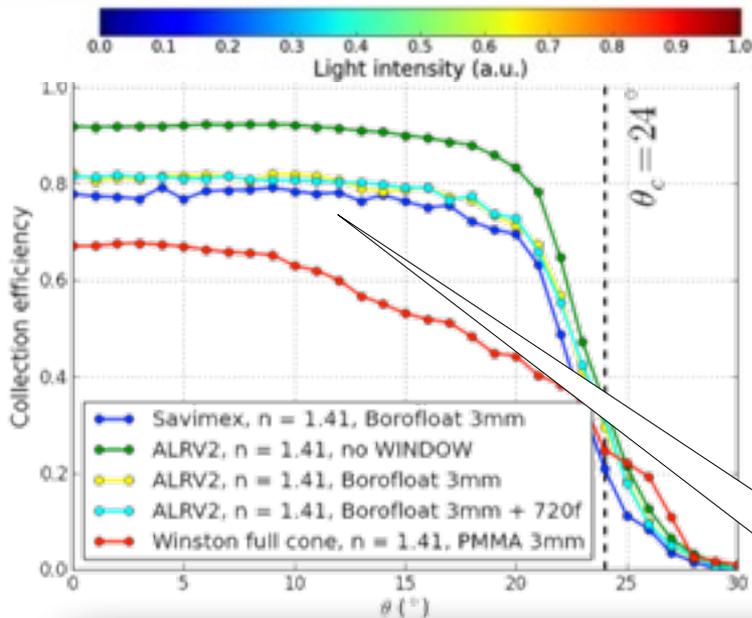
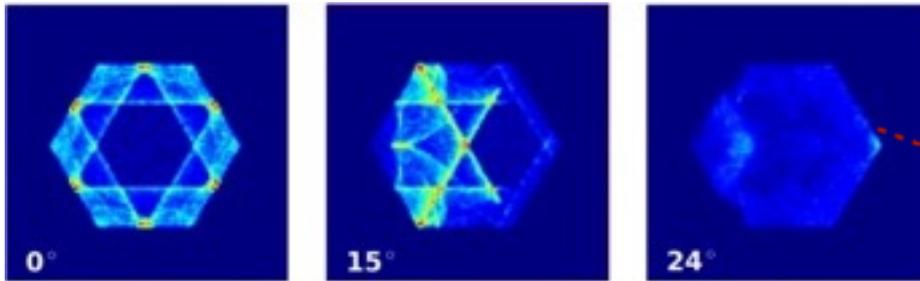
$n_p = 1296$ pixels

Camera size (D_c) = 88 cm



Winston cones

We need an Hollow cone to improve the response in the UV

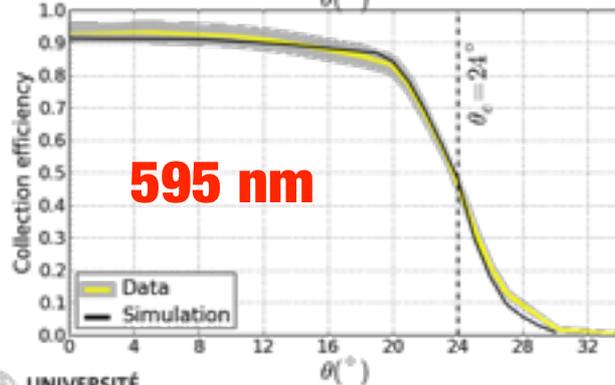
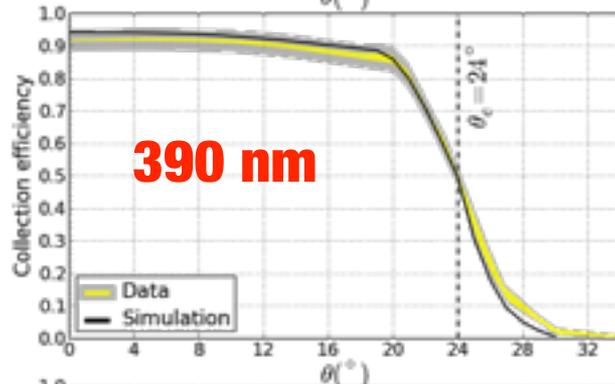
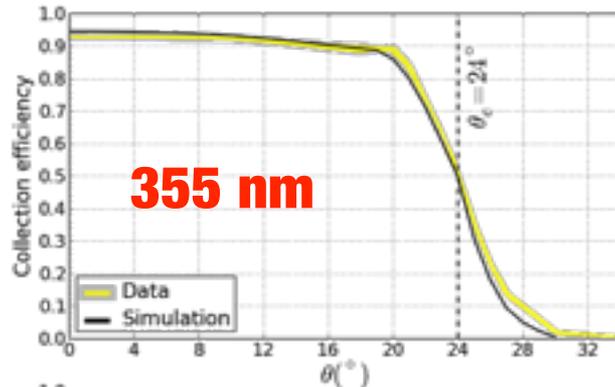


It's all about coating!

A new type of sensor!

Point Spread Function	➡	angular pixel size: 0.25°	➡	Top physical size
f/D and Camera diameter	➡	Cutoff angle: 24°	➡	Cone Height

Prototypes and first measurements

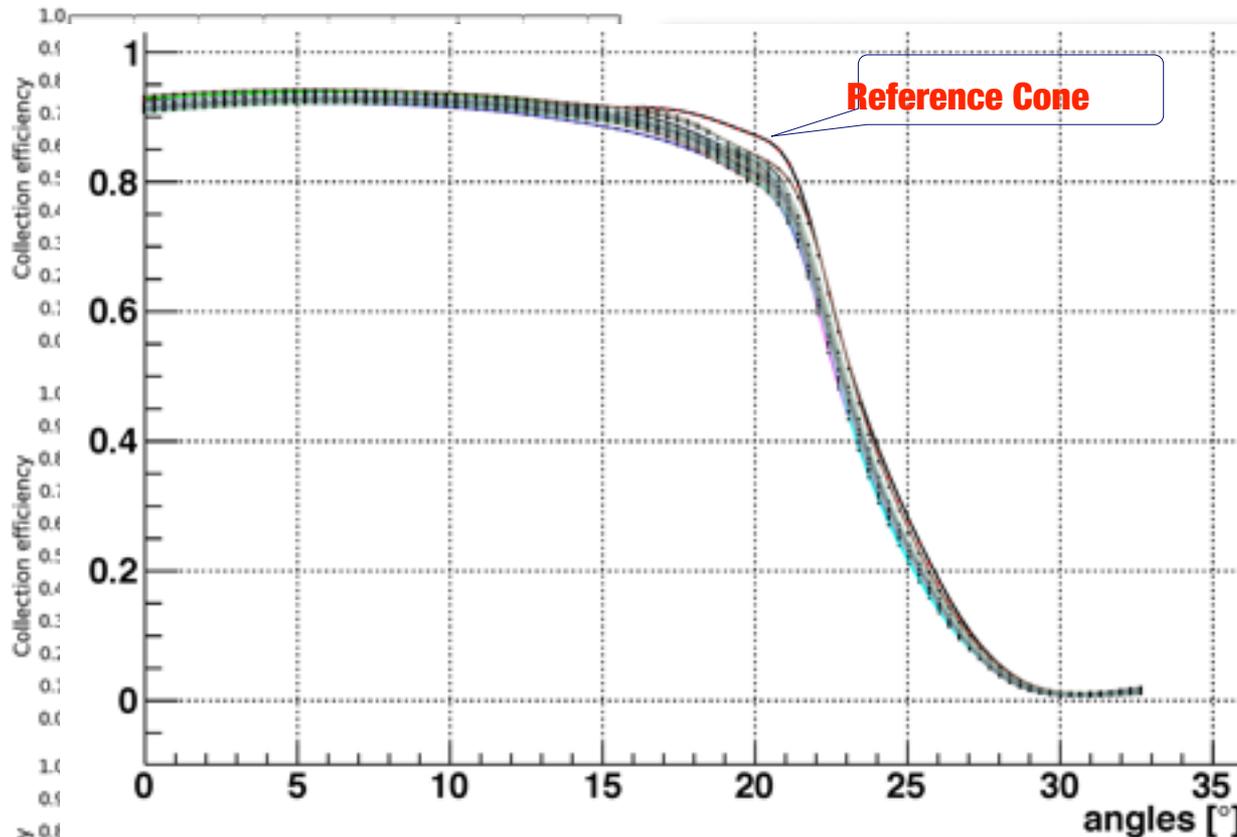


- ➔ Measurement done for different wavelengths
- ➔ Simulation of the set-up to validate the coating
- ➔ Very good agreement between simulation and measurement
- ➔ Coating qualified also with 40 thermal cycles (from -15° to $+30^{\circ}$) - no measurable effect
- ➔ The collection efficiency shown here does not take into account the effect of the entrance window

Related paper:

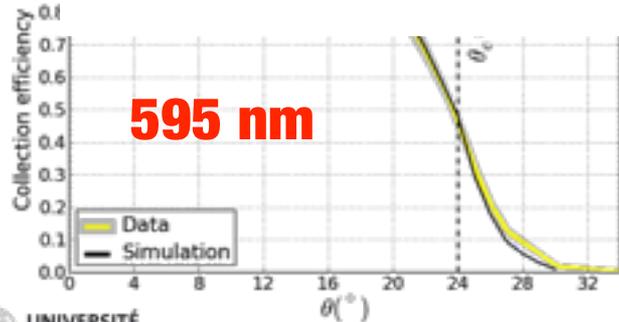
J.A. Aguilar et al., *Design, optimization and characterization of the light concentrators of the single mirror small size telescopes of the Cherenkov Telescope Array*, *Astroparticle Physics*, doi:10.1016/j.astropartphys.2014.05.010

Prototypes and first measurements



films Zurich
Viaoptic GmbH

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validate the coating
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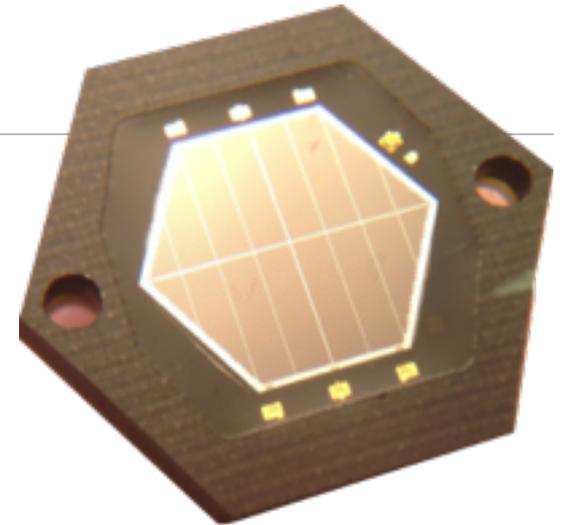
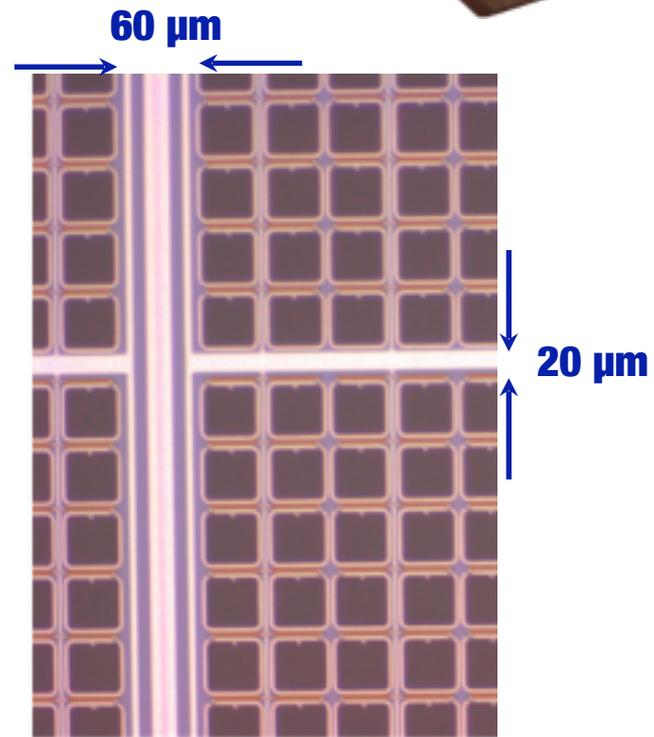
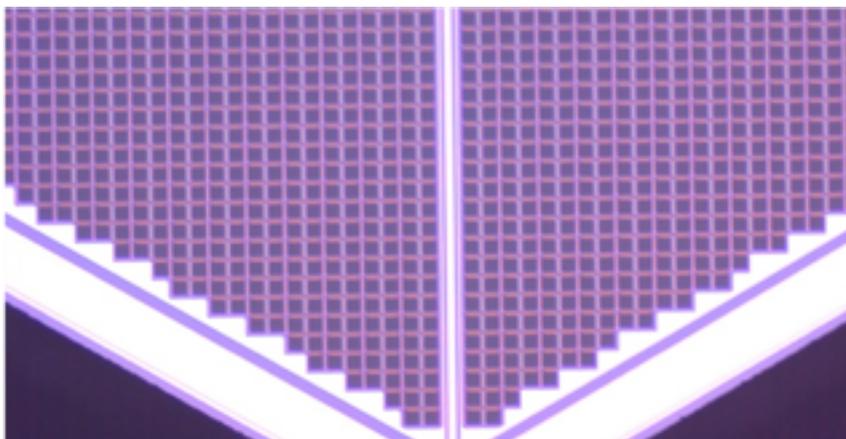
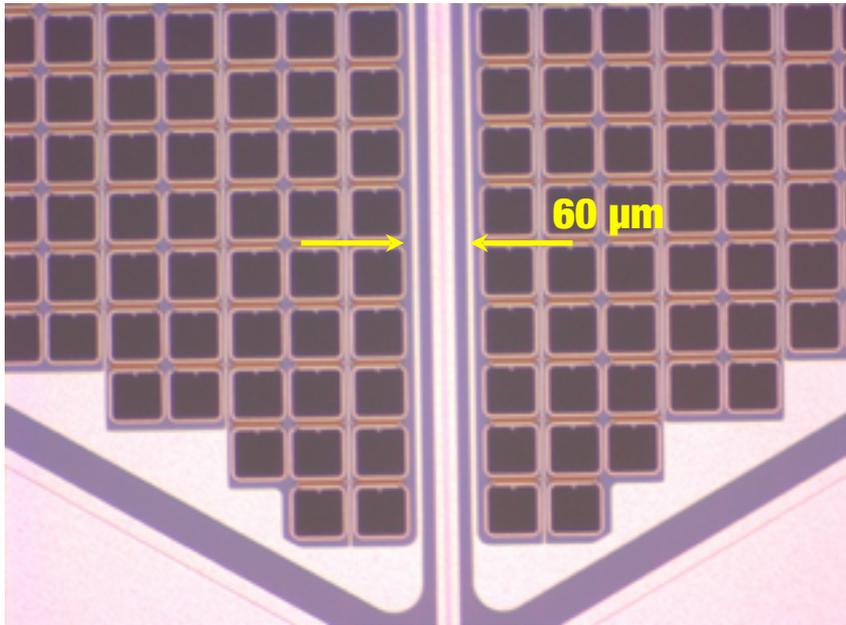


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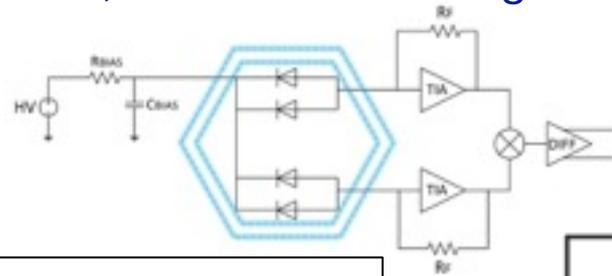
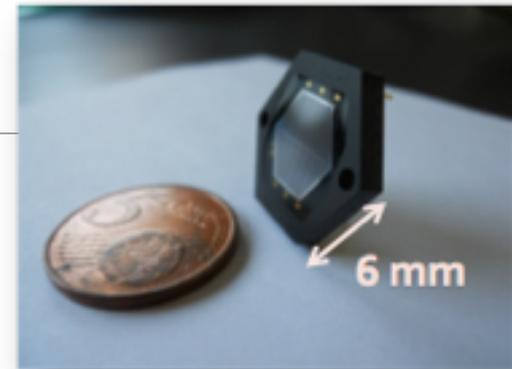
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The S12516-050 sensor

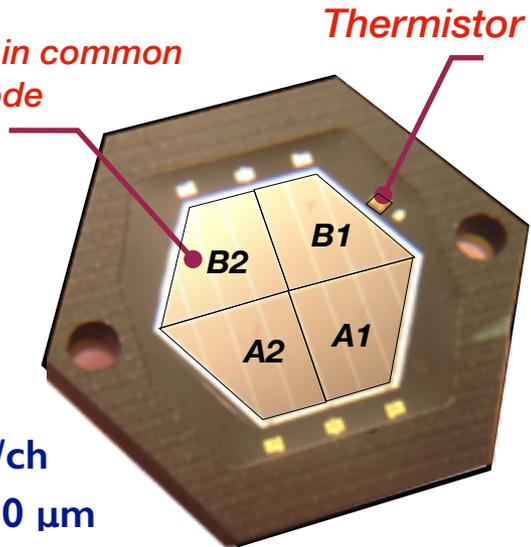
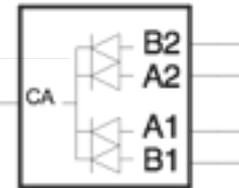
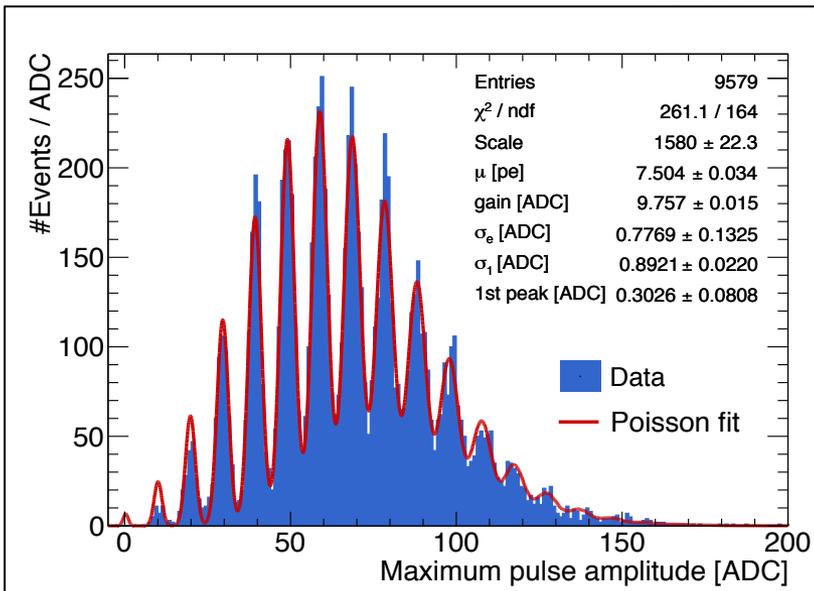


The Hexagonal Sensor

- Despite the use of the light concentrators, the pixel size remains large compared to common devices,
- Result of a collaboration between DPNC University of Geneva and Hamamatsu, the sensors are large hexagonal arrays of G-APD.



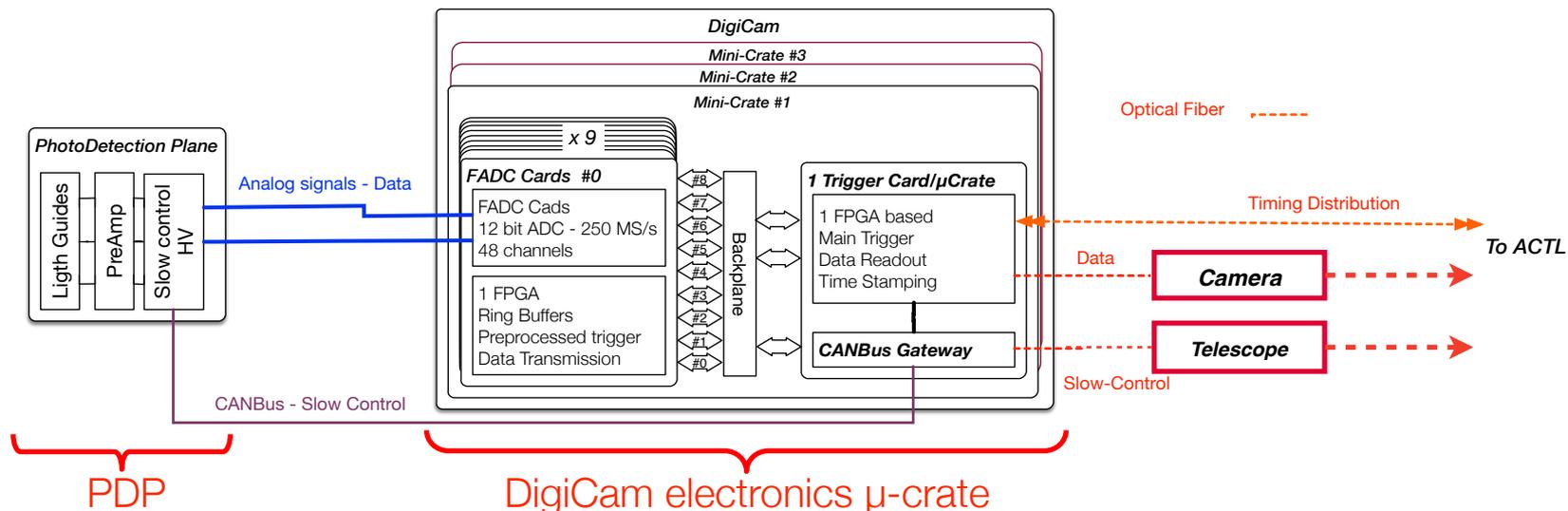
4 channels in common cathode mode



- Channels:** 4
- Area:** 23.8 mm²/ch
- Cell size:** 50 μ m x 50 μ m
- Fill factor:** 68%
- N. pixels:** 9210 SPADs/
- Capacitance:** 840 pF/ch
- DC rate:** 2-3 MHz/ch

- ➔ 93.6 mm² surface
- ➔ 10.4 mm flat-to-flat

The Camera concept



- Separation of PDP and Digital Readout
 - Separate mechanics and power supplies
 - Analogue signals over CAT6/RJ45
 - DC coupling for NSB monitoring
- Window and chassis sealed, IP65
- Water cooled - Heat pipes on Digital board
- Compact, robust, lightweight and self-contained

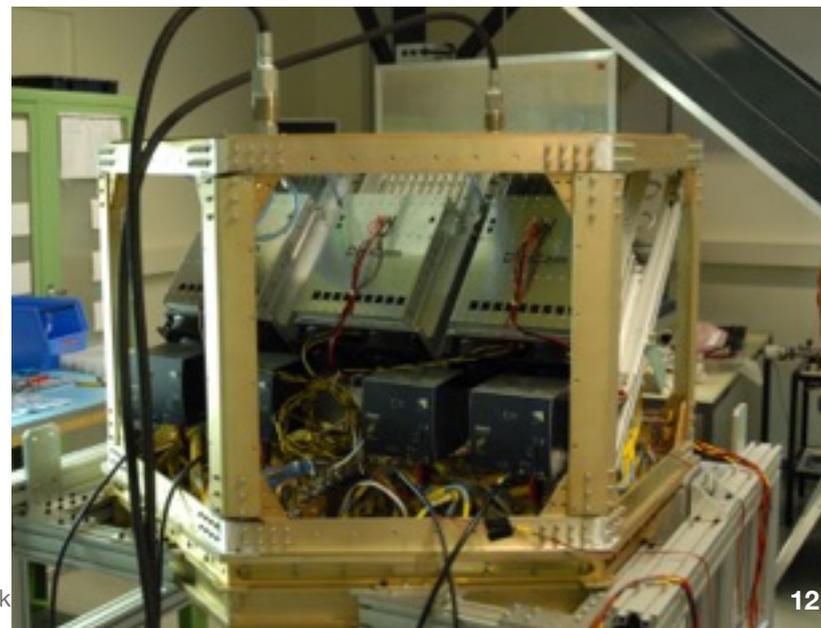


Photo Detection Plane

- 1296 pixels, 108 modules (12 pixels each)
- Power consumption ~ 500 W
- Total weight 35 kg
- Borofloat entrance window 3.3 mm coated with AR filter (Cut-off at 540 nm)
- Aluminum backplate (6 mm) as backbone and cold plate for Cooling
- Sensor bias automatically adjusted according to temperature (reso. of 0.17 deg)

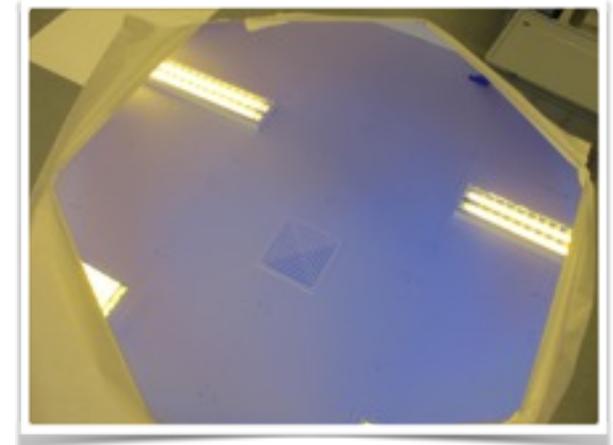
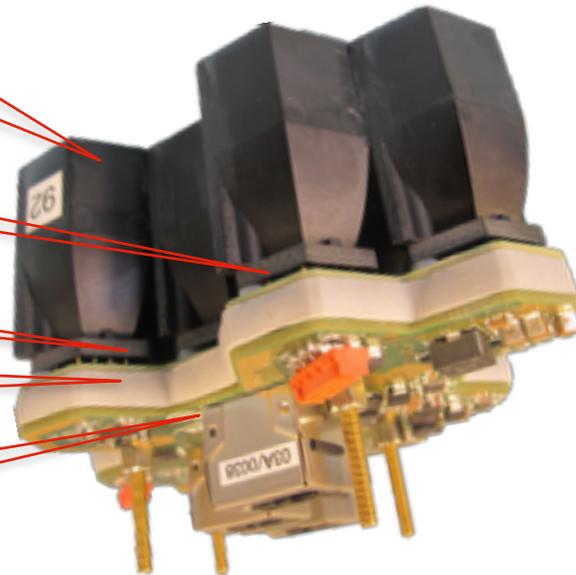
Hollow light guides

Sensors

PreAmp board

Thermal gap filler

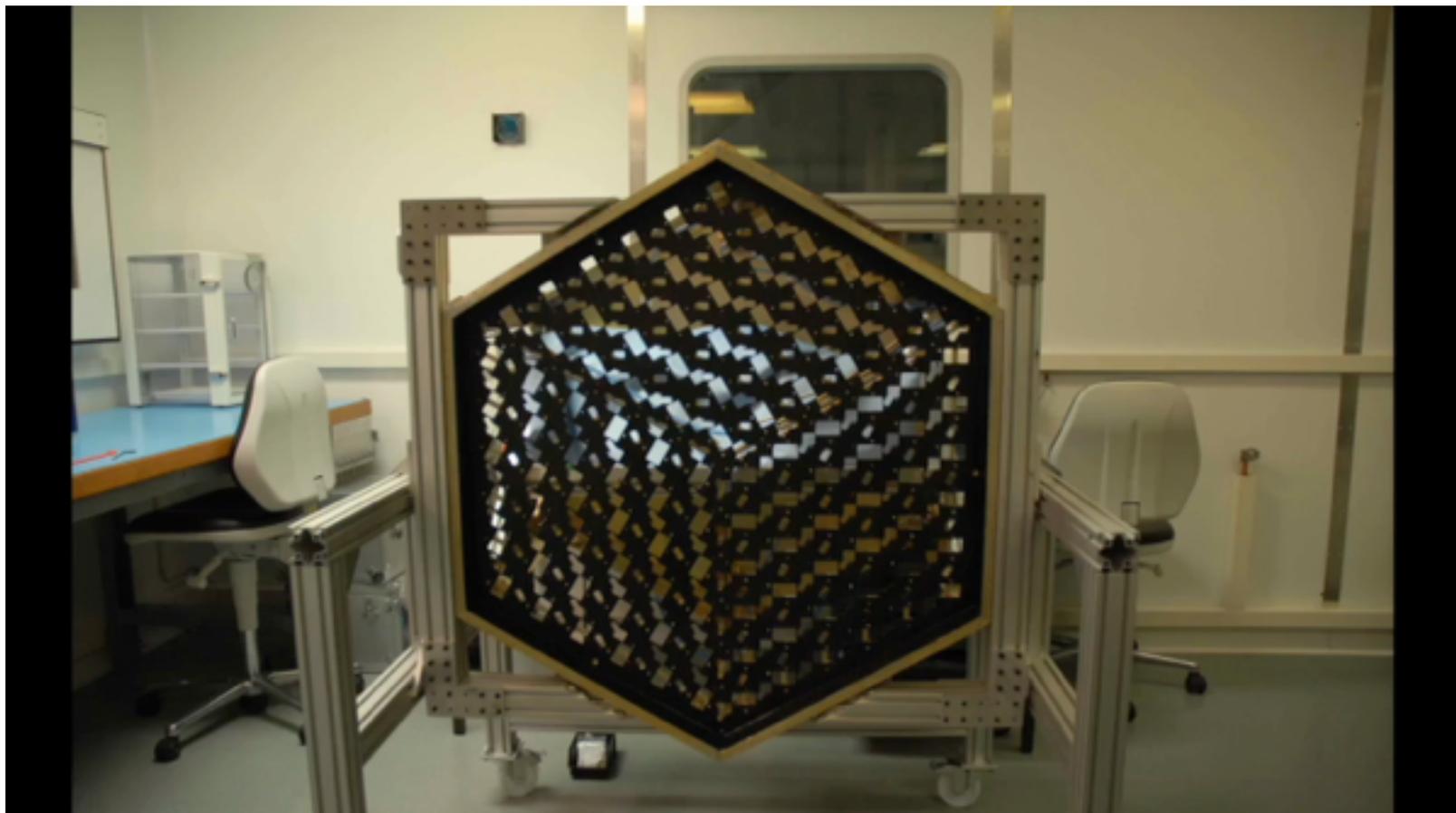
Slow control board



Borofloat window

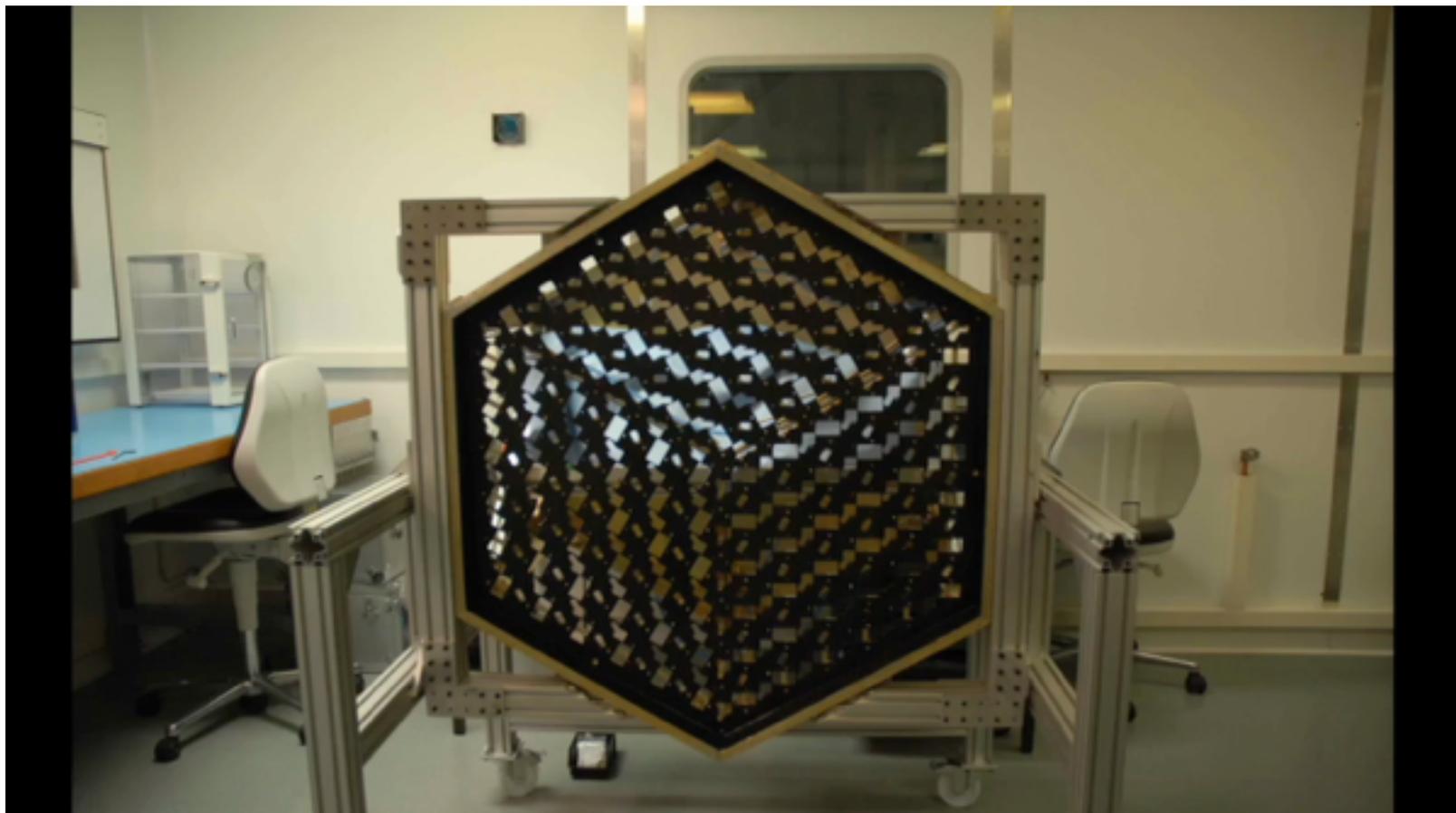
Assembly - Photo Detection Plane

1 day of work to assemble full PDP



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1 day of work to assemble full PDP



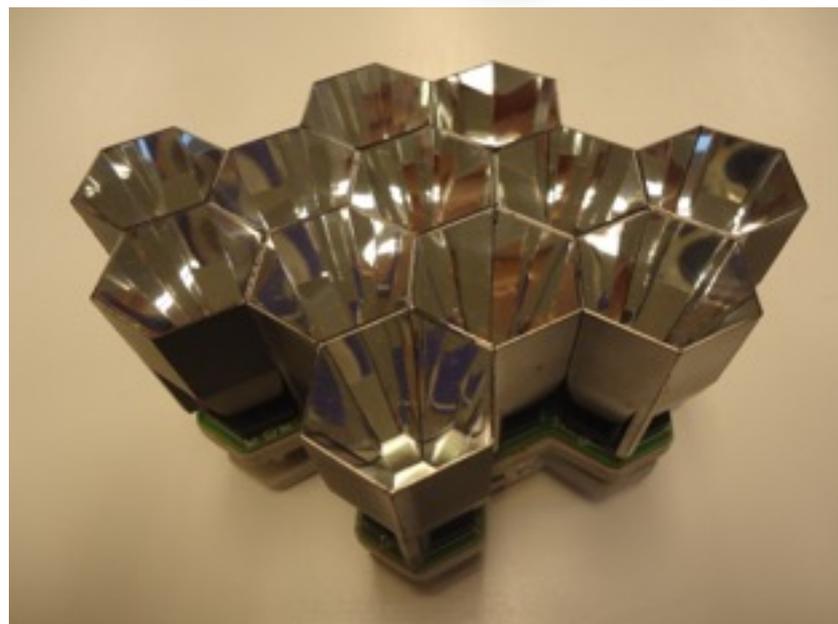
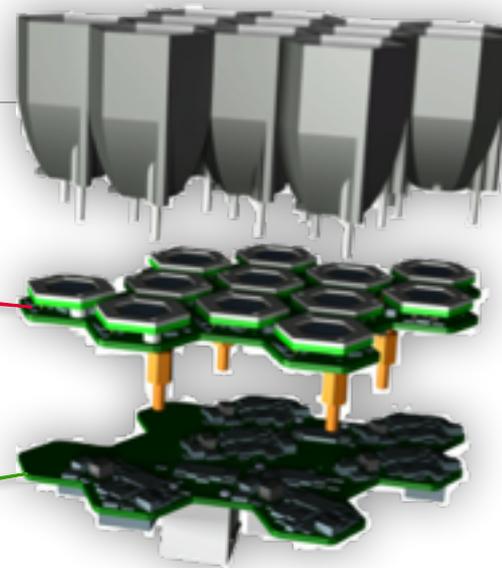
The Front-end electronics

• The pre-amp boards

- read and pre-amplify the analog signals
- routes out signals, the HV and the output of the NTC probe, present on the sensor package.

• SlowControl Board

- route pixel signals to Digicam via the RJ45 connector
- distributes the power and the HV
- regulate HV for each pixel
- Use a micro controller to change HV of each pixel to compensate temperature variations measured by the NTC probe in sensor package
- Board accessed via CAN-bus



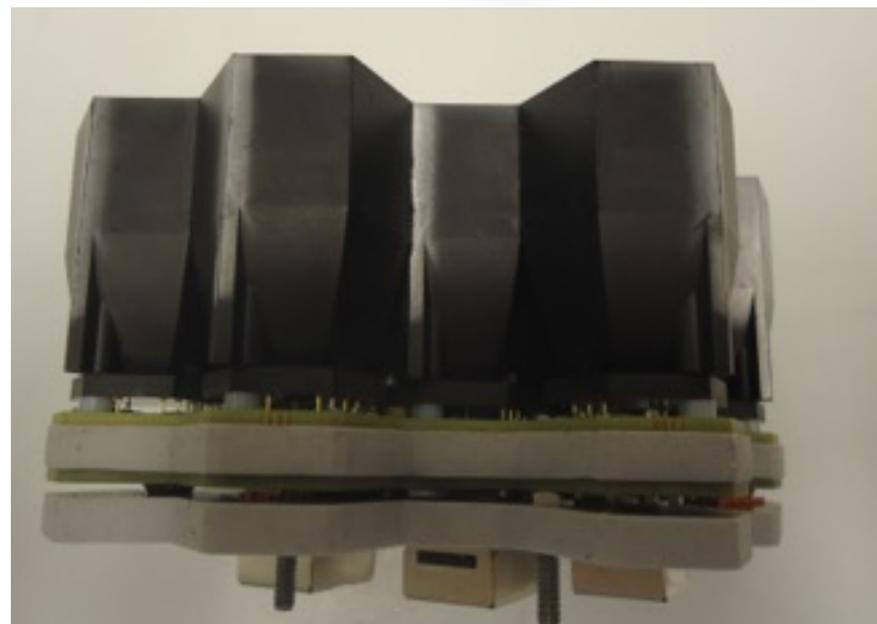
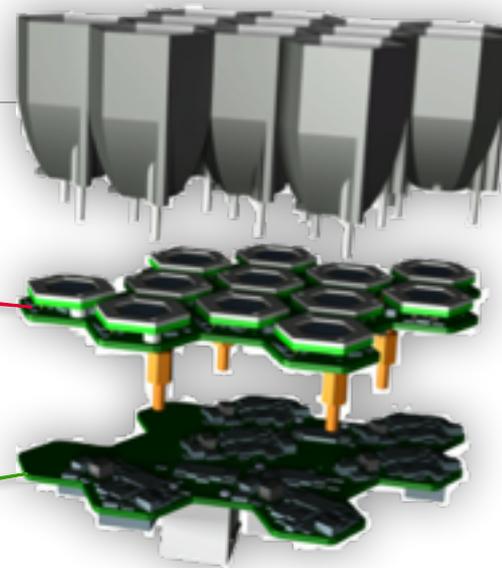
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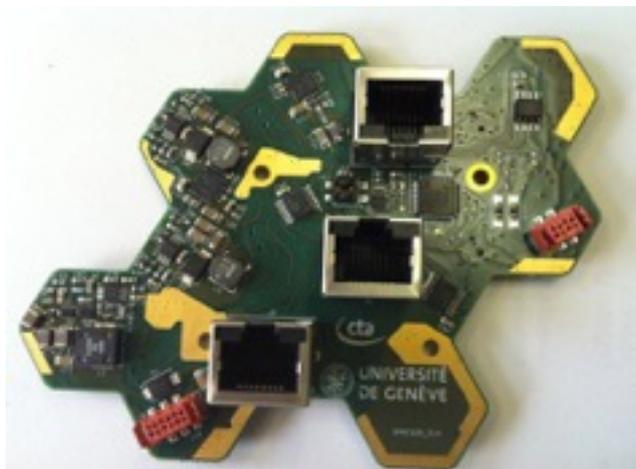
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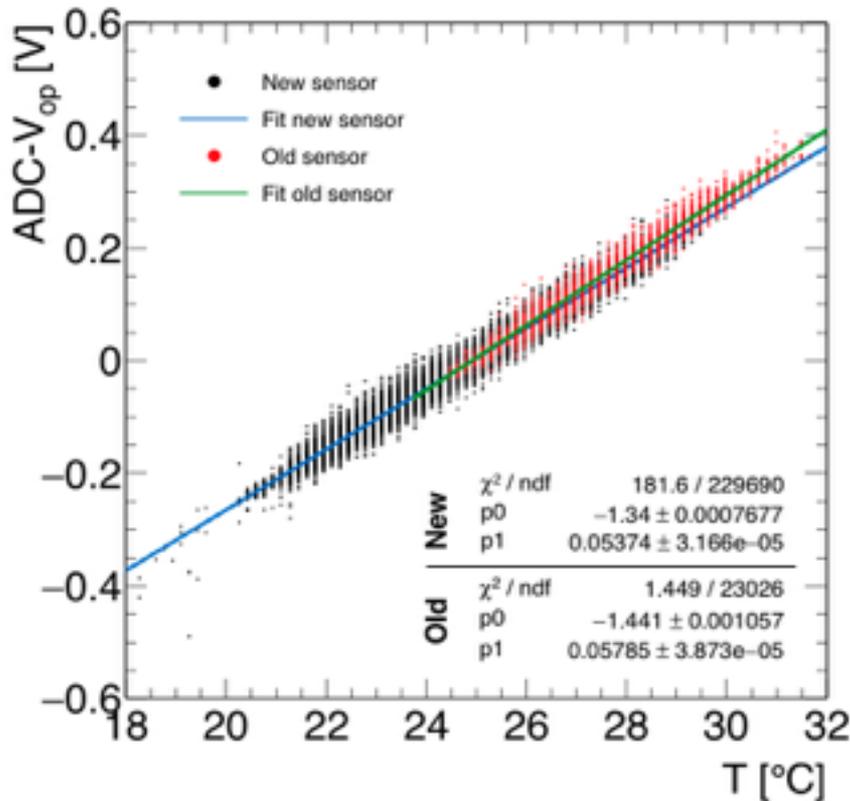
Slow control Board



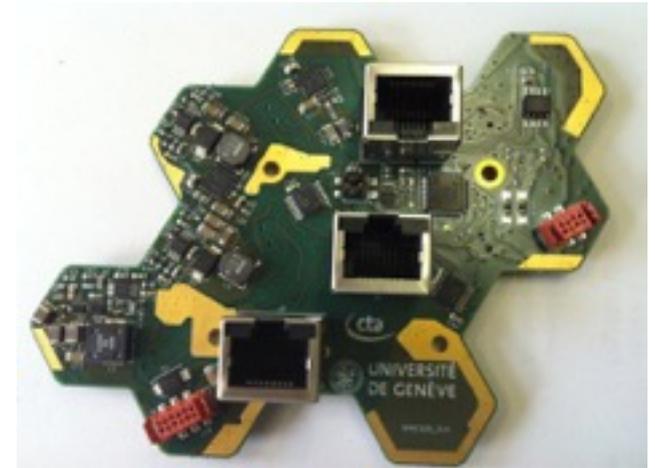
- The slow control board has the following task
 - route pixel signals to Digicam via the RJ45 connector
 - provide the bias and
 - change the bias to compensate for temperature variations
- a micro controller supervises the temperature compensation
 - look-up table is loaded with correction
 - each pixel has its own correction
 - temperature reading
- Board accessed via CAN-bus



Commissioning - Design validation



✓ Slow control and communication validated



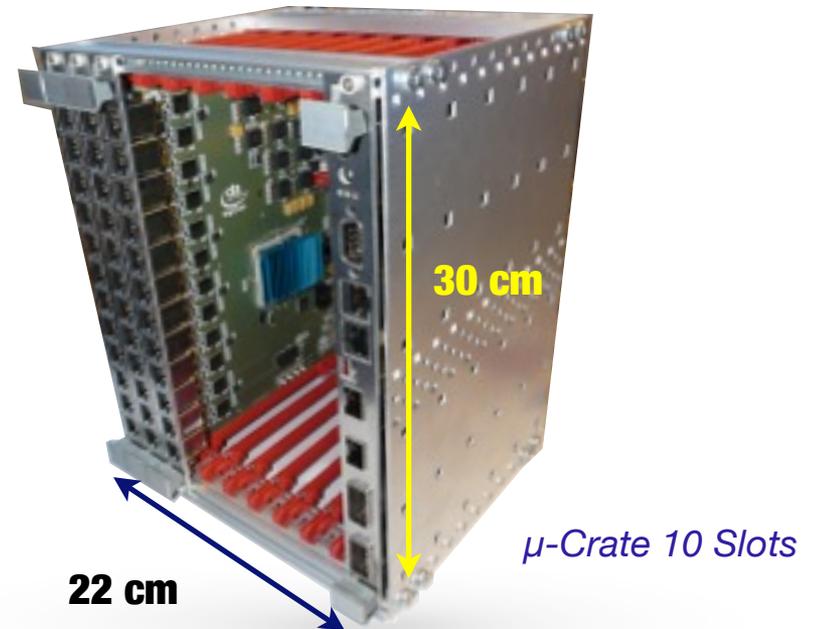
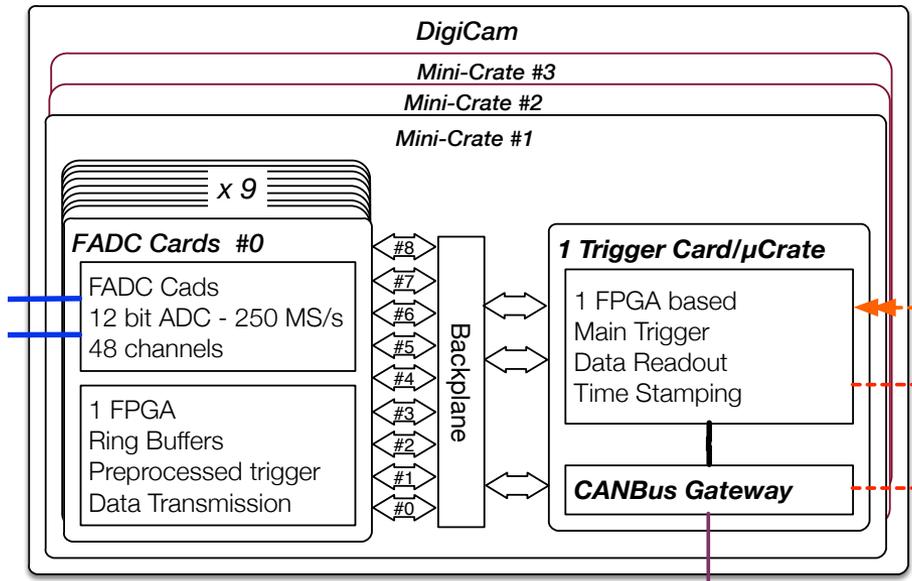
Control:

- ✓ Enable boost for bias
- ✓ Activate bias voltage
- ✓ Reference voltage and temperature setting
- ✓ Temperature variation factor

Monitoring of:

- ✓ Temperature and bias voltage
- ✓ Compensation loop

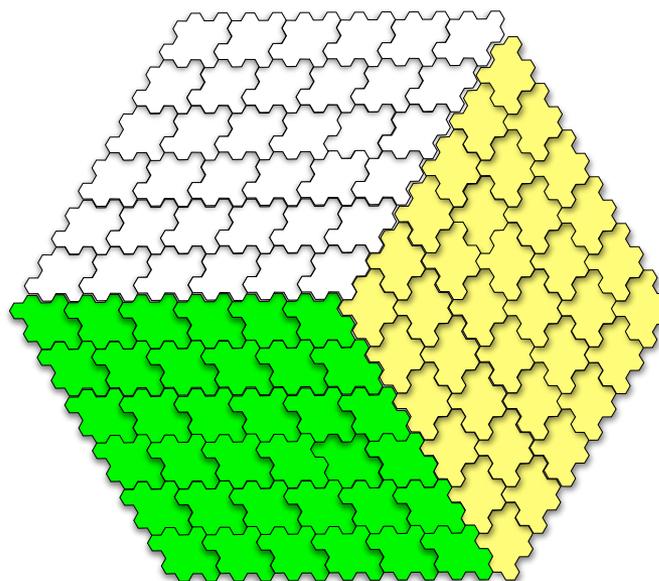
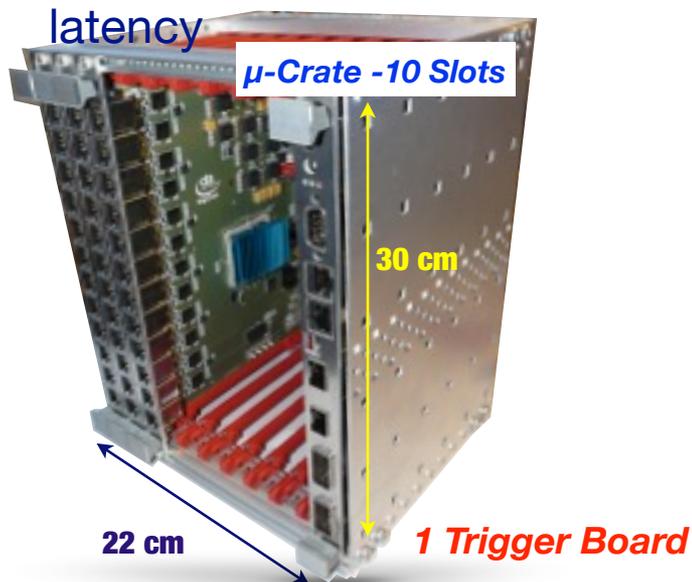
Digital readout and trigger / DigiCam



- Sampling rate 250 MHz
- Readout rate: 32 kHz @ 80 ns readout window, no dead-time
- Fully digital trigger and readout (High-speed/High-throughput)
- Serial architecture based on multi-Gigabit links (trigger and ADC readout)
- Trigger path with reconfigurable algorithms and signal preprocessing

DigiCam Electronics

- Based on last generation of High performance Xilinx FPGA \Rightarrow Flexibility, High Speed, Low power consumption
 - \rightarrow up to **56 GTX** transceivers, max 12.5Gbit/s
 - \rightarrow up to **96 GTH** transceivers, max 13.1Gbit/s
- Using DDR expandable RAM for data buffering \Rightarrow Upgradable, handle high latency

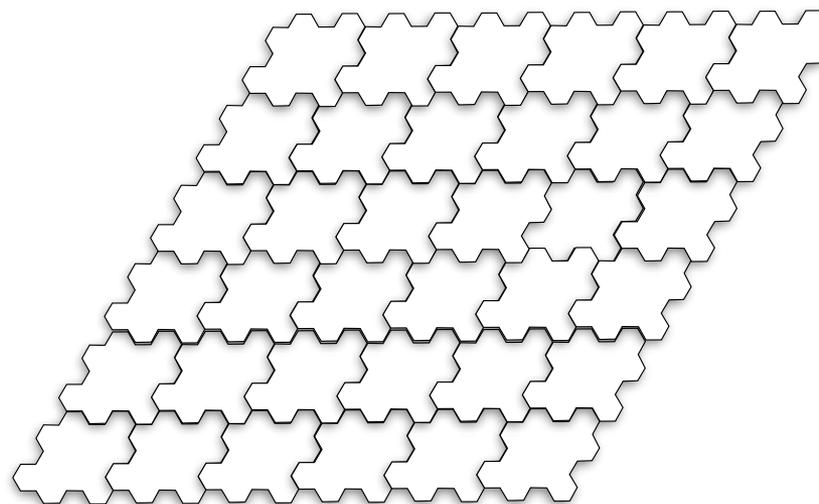
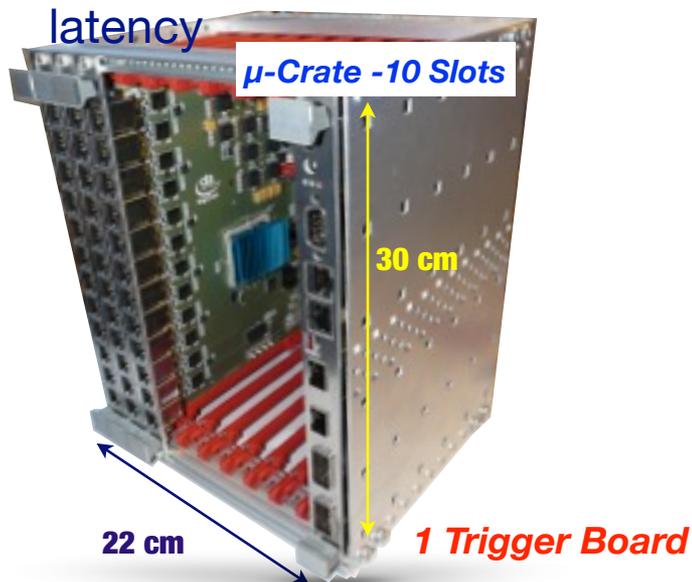


BackPlane with 54 differential pairs

- 36 - high speed 8Gbit/s for crate trigger data
- 18 - low speed 1Gbit/s for readout and slow control

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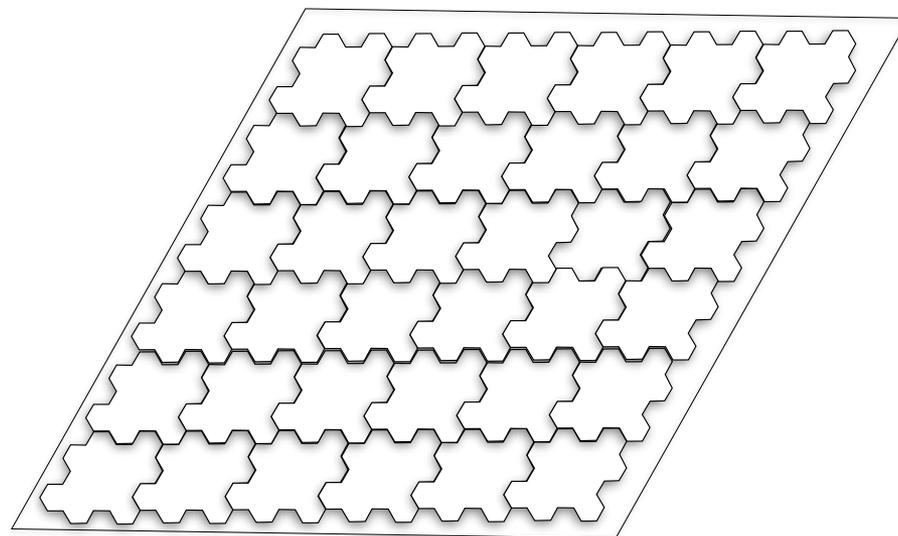
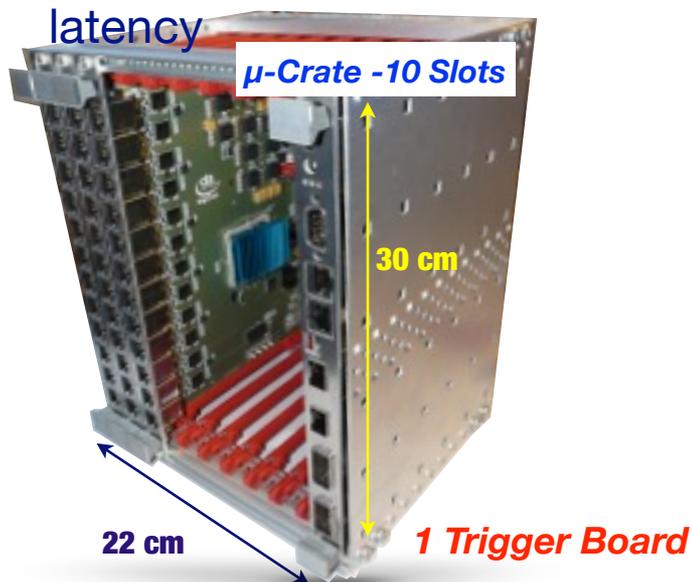


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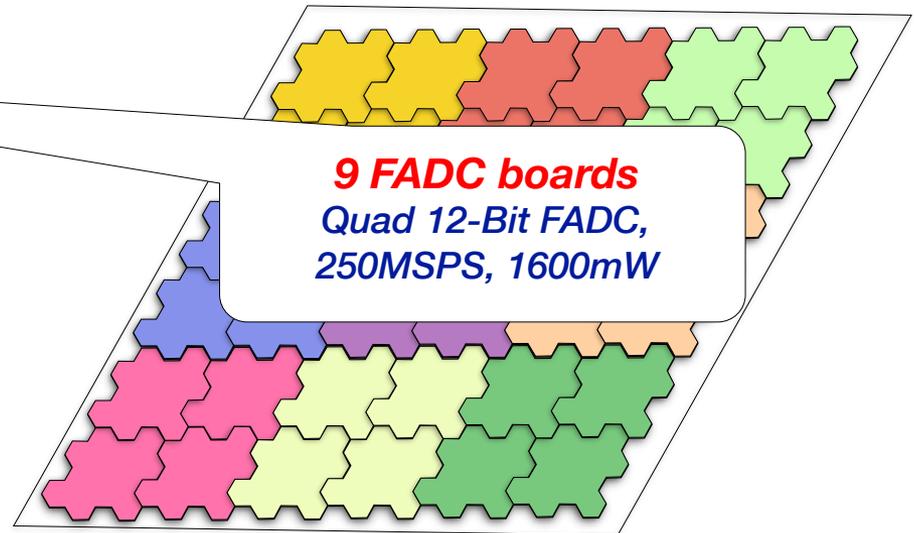
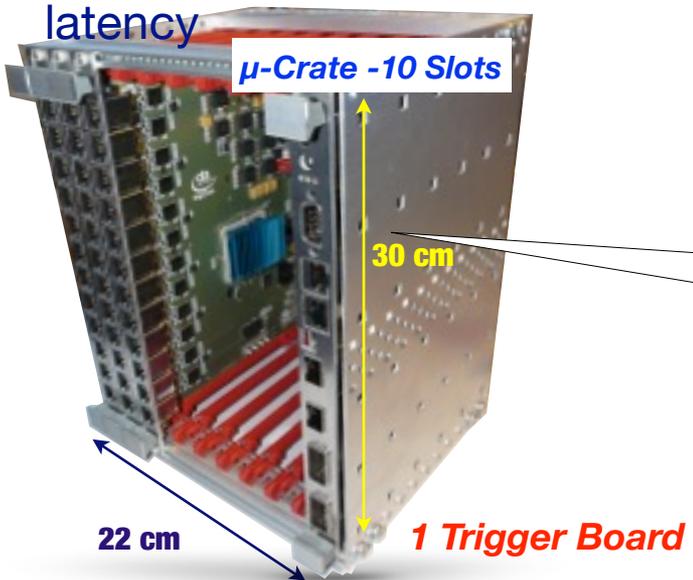


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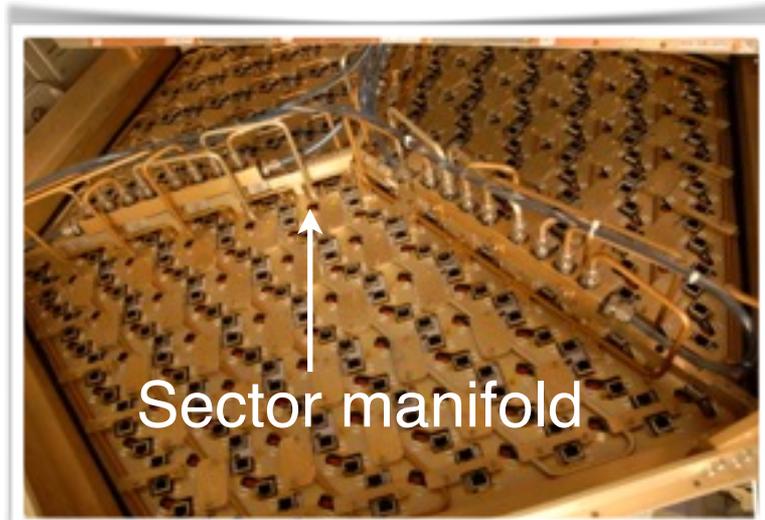
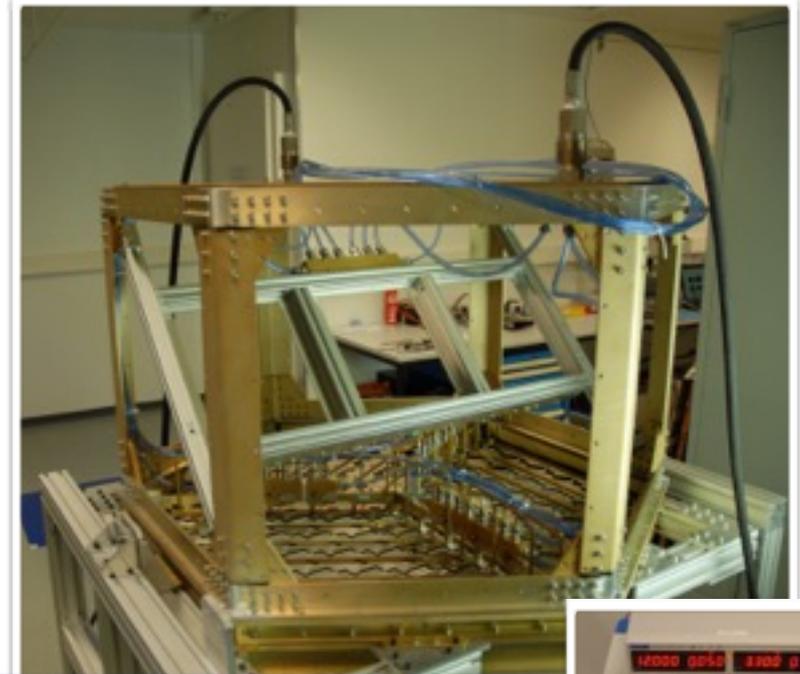
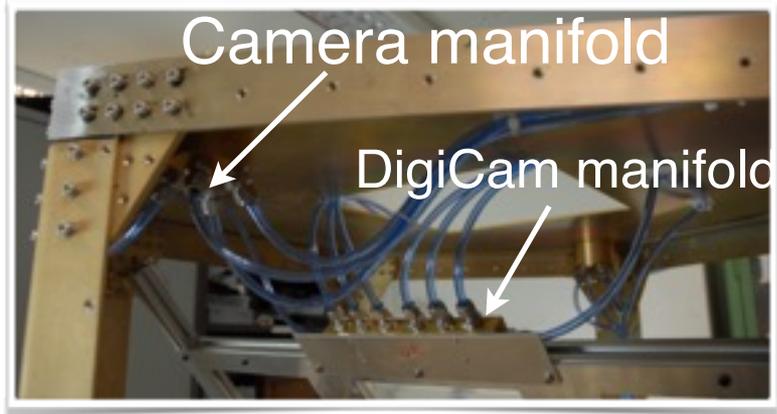
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Assembly - Water Cooling system

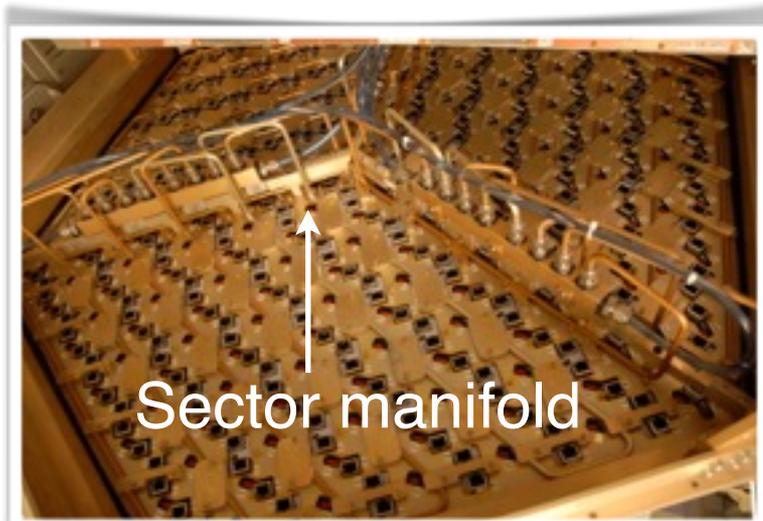
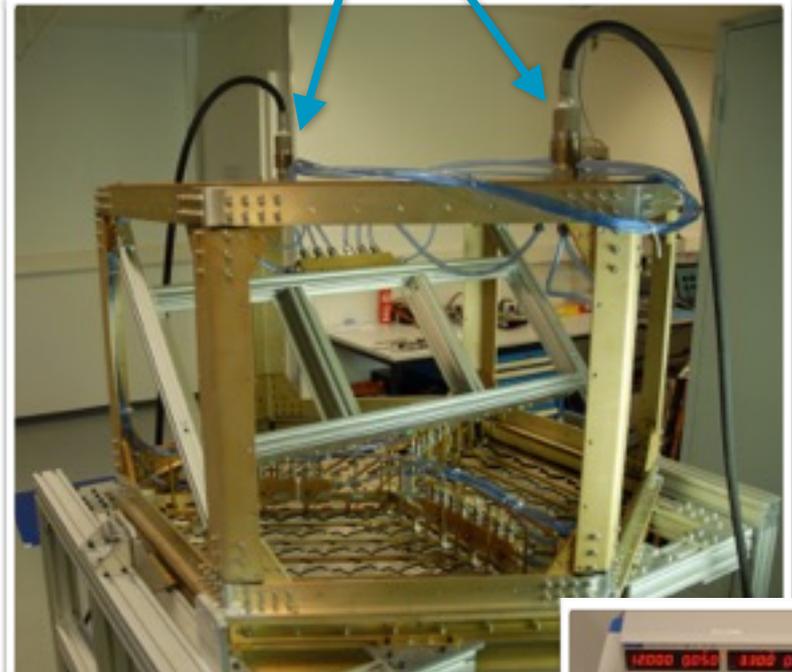
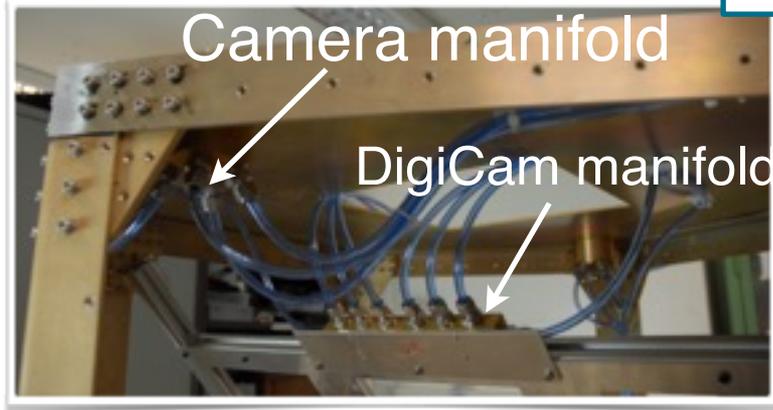


- Cooling pipes installed prior to any electrical system
- Water tightness checked
- Epoxy resin applied on joints for safety



Assembly - Water Cooling system

fast connectors for inlet/outlet camera



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Assembly - Water Cooling system

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Camera manifold



Sector manifold

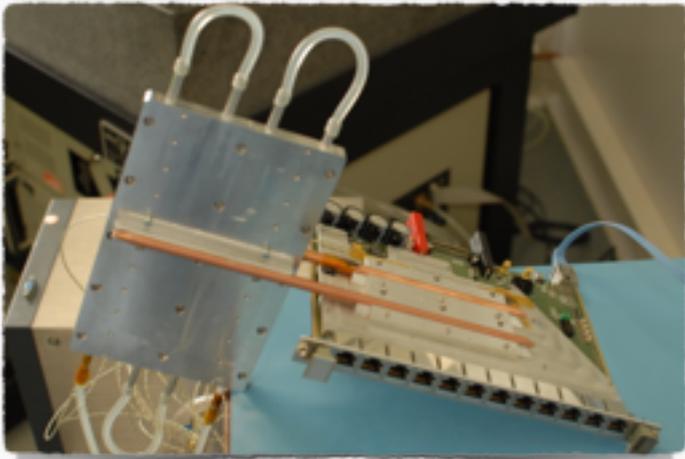


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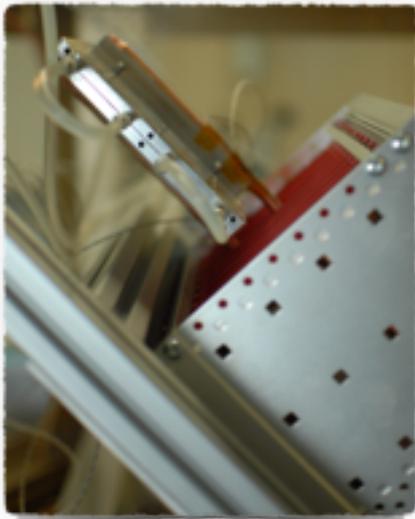


Chiller

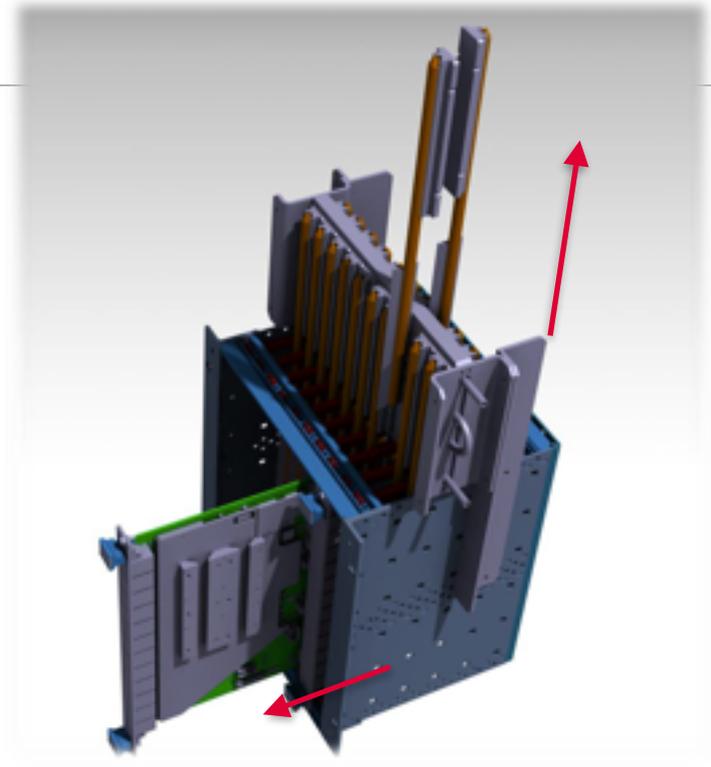
Digicam Cooling



2 Heat pipes (25W each) per board coupled to a water cooled plate

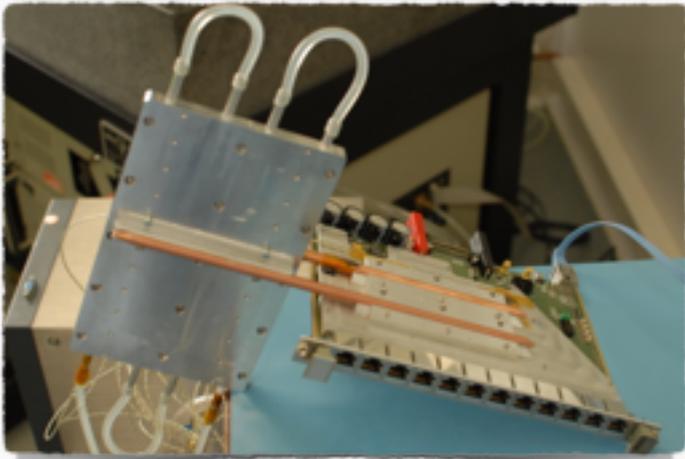


Micro-crate mounted at 45° to have always optimal heat-pipes efficiency

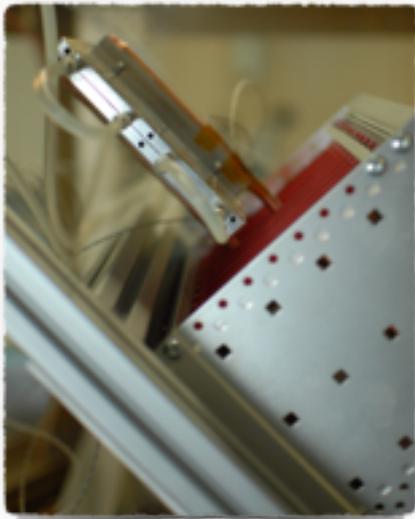


- Board can be removed from the crate with limited intervention on the cooling system

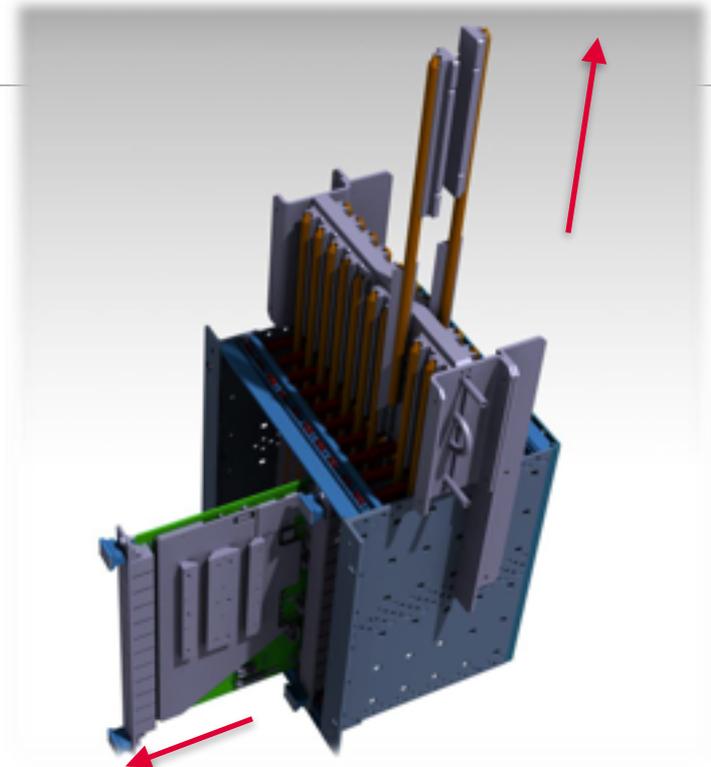
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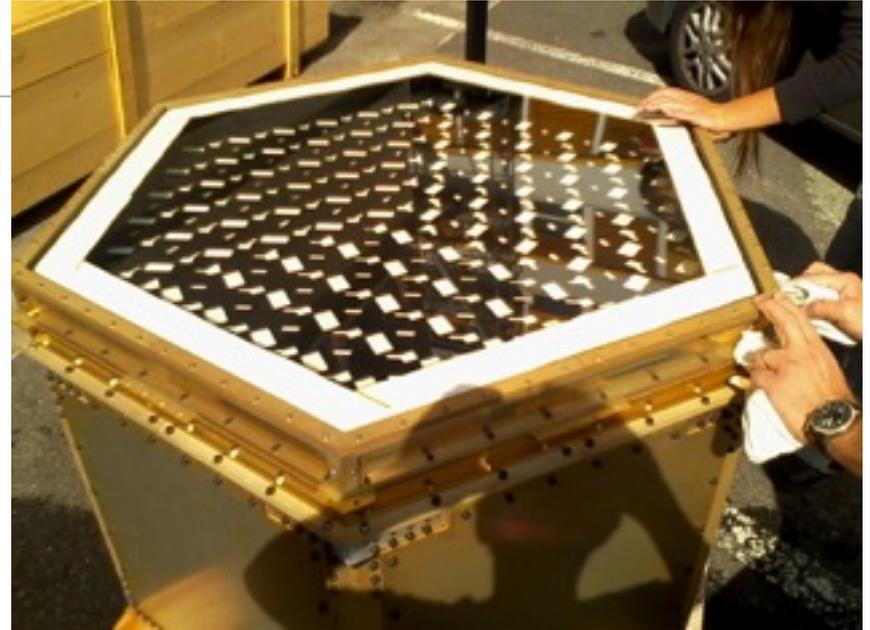
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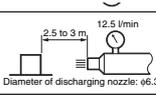


- Board can be removed from the crate with limited intervention on the cooling system

Camera Chassis IP65 Test

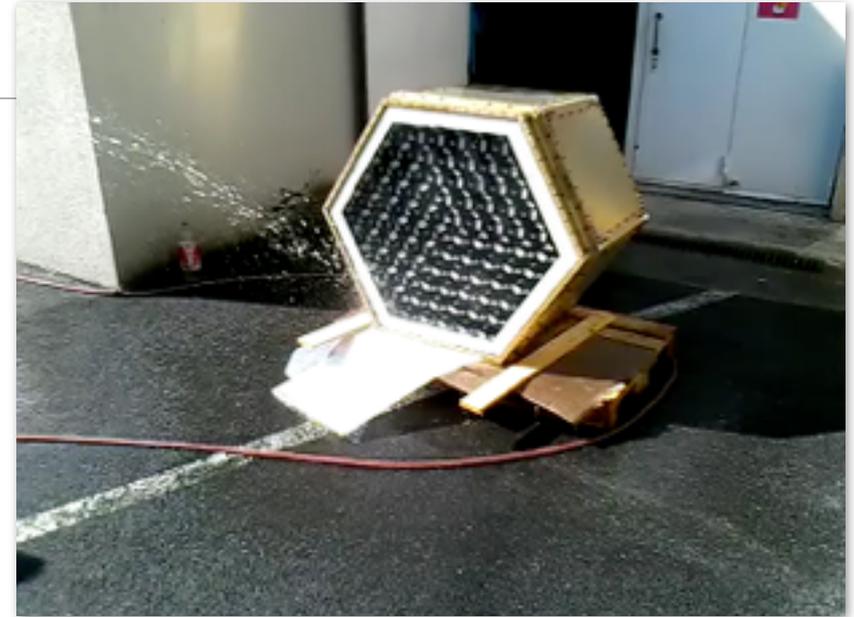
- Camera tested against IP65 Specs
 - to check even tiny leaks the internal surface of the camera was covered with a special paper which change color in case it comes in contact to water
- No water leak but on 2 points on the patch panel
 - Due to a wrong machining of the holes
 - cured with a small joint
- Another test is scheduled next week to qualify the new solution



5	 <p>Protection from water jets</p>	No harmful effect of water splash from all direction	<p>By using tool as described in right picture, Jet the water from all direction to the object surface for 1 m²/min, at least for 3 min in total.</p>  <p>2.5 to 3 m 12.5 l/min Diameter of discharging nozzle: ø6.3</p>
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Camera Chassis IP65 Test

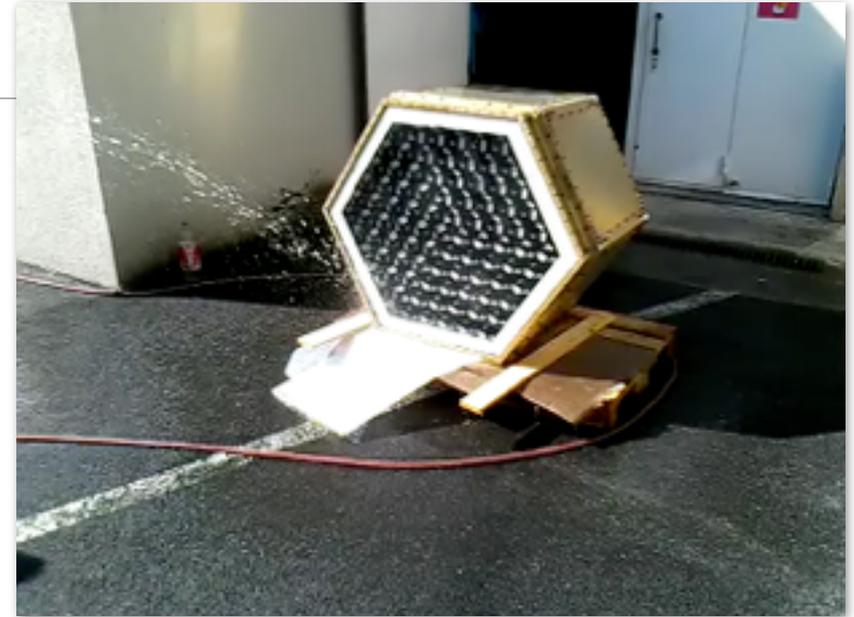
- Camera tested against IP65 Specs
 - to check even tiny leaks the internal surface of the camera was covered with a special paper which change color in case it comes in contact to water
- No water leak but on 2 points on the patch panel
 - Due to a wrong machining of the holes
 - cured with a small joint
- Another test is scheduled next week to qualify the new solution

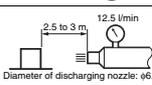


5	<p>Protection from water jets</p>	<p>No harmful effect of water splash from all direction</p>	<p>By using tool as described in right picture, Jet the water from all direction to the object surface for 1 m²/min, at least for 3 min in total.</p> <p>2.5 to 3 m 12.5 l/min Diameter of discharging nozzle: 6.3</p>
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Camera Chassis IPX5 Test

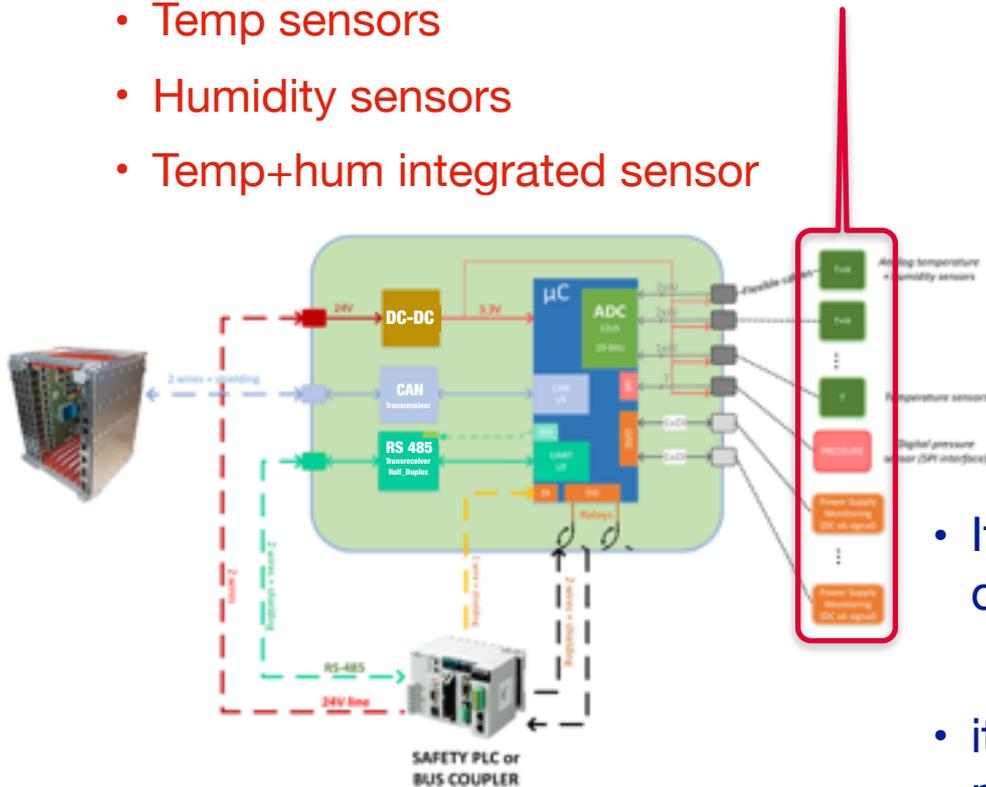
- Camera tested against IP65 Specs
 - to check even tiny leaks the internal surface of the camera was covered with a special paper which change color in case it comes in contact to water
- No water leak but on 2 points on the patch panel
 - Due to a wrong machining of the holes
 - cured with a small joint
- Another test is scheduled next week to qualify the new solution



5	Protection from water jets 	No harmful effect of water splash from all direction	By using tool as described in right picture, Jet the water from all direction to the object surface for 1 m ² /min, at least for 3 min in total. 
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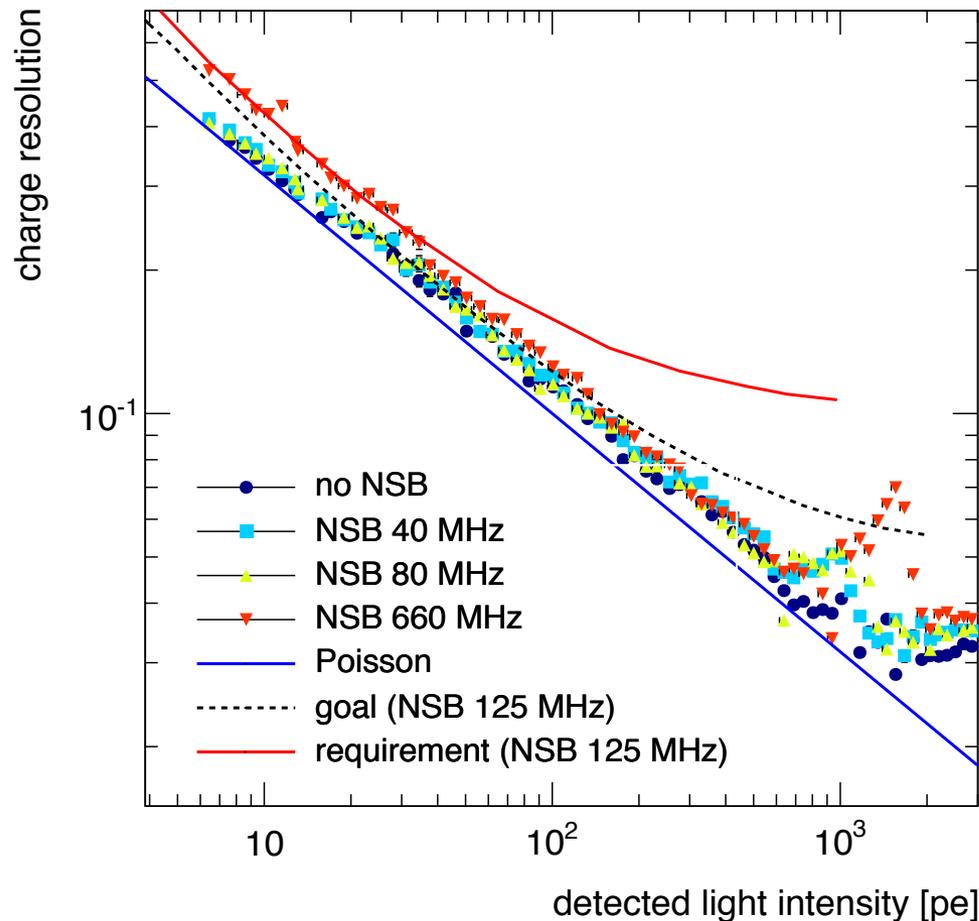
Camera Safety - Housekeeping Board

- to have as many possible probes but not having lots of cables from camera to the PLC we designed an Housekeeping board
- It can host 3 different type of sensors
 - Temp sensors
 - Humidity sensors
 - Temp+hum integrated sensor



- It can be connected to the PLC with 2 different protocols:
 - RS485 or CanBus
- it can trigger alarm to PLC in case of problems
- it will also monitor the power supply

Performance

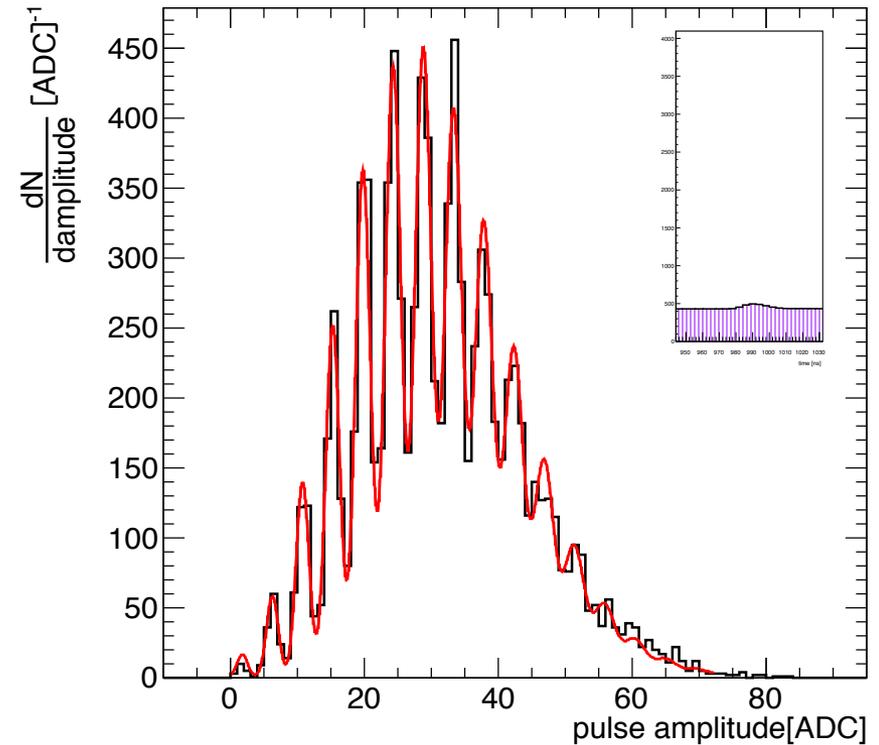
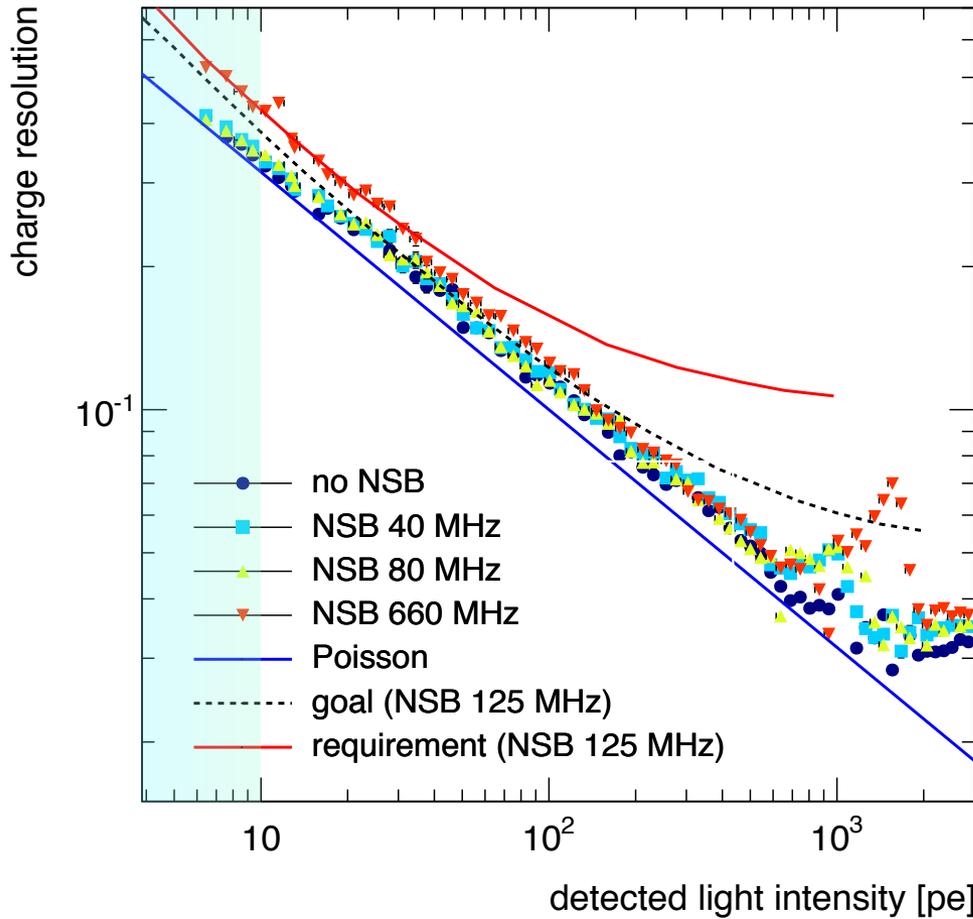


Charge resolution measurement:

- Key parameter to assess quality of image reconstruction (Hillas parameters)
- Measurement performed injecting both pulsed and continuous light
- Different NSB level from dark night to half-moon (45° off-axis)

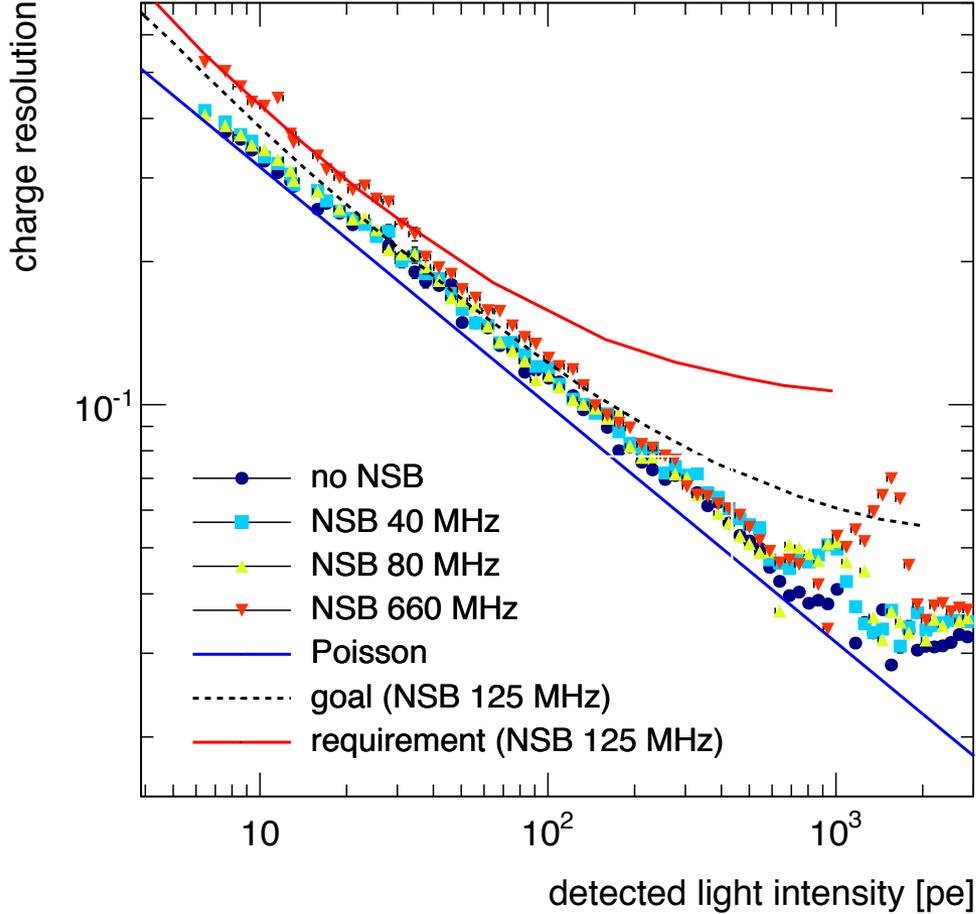
Performance

Linear
low light



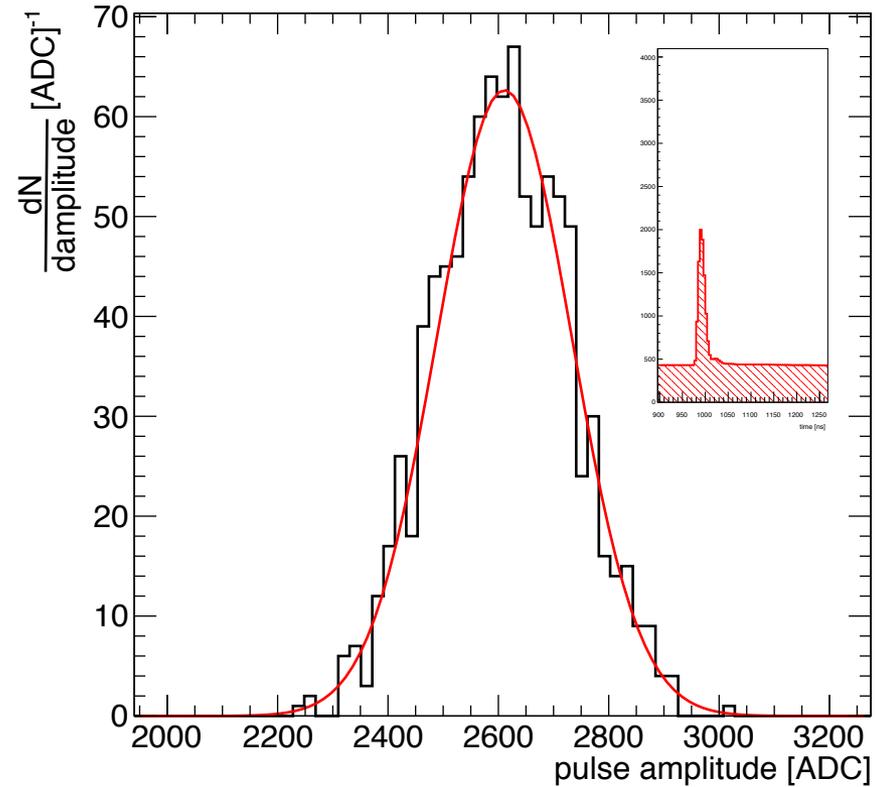
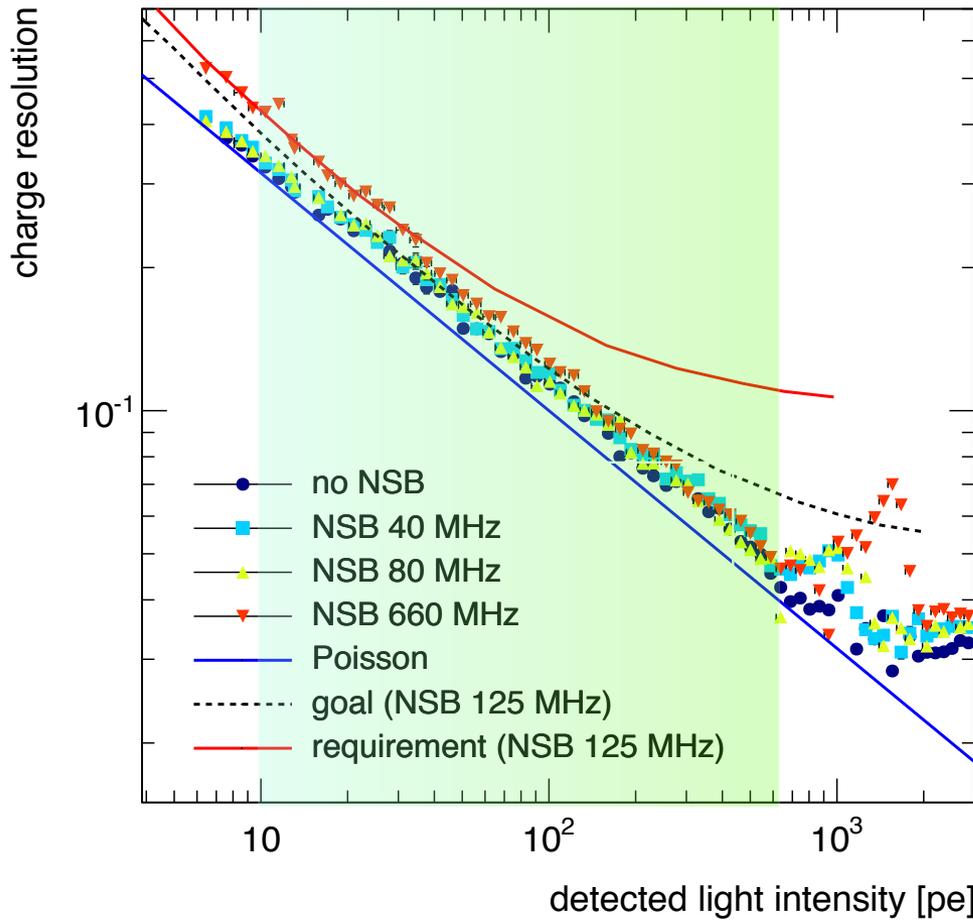
Performance

Linear
medium light



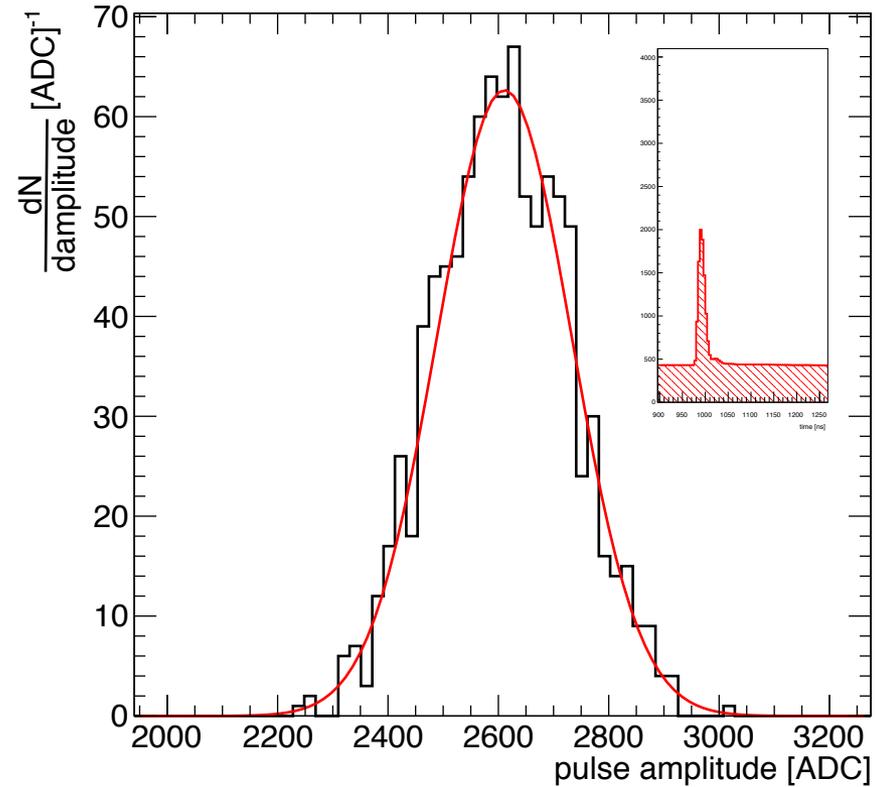
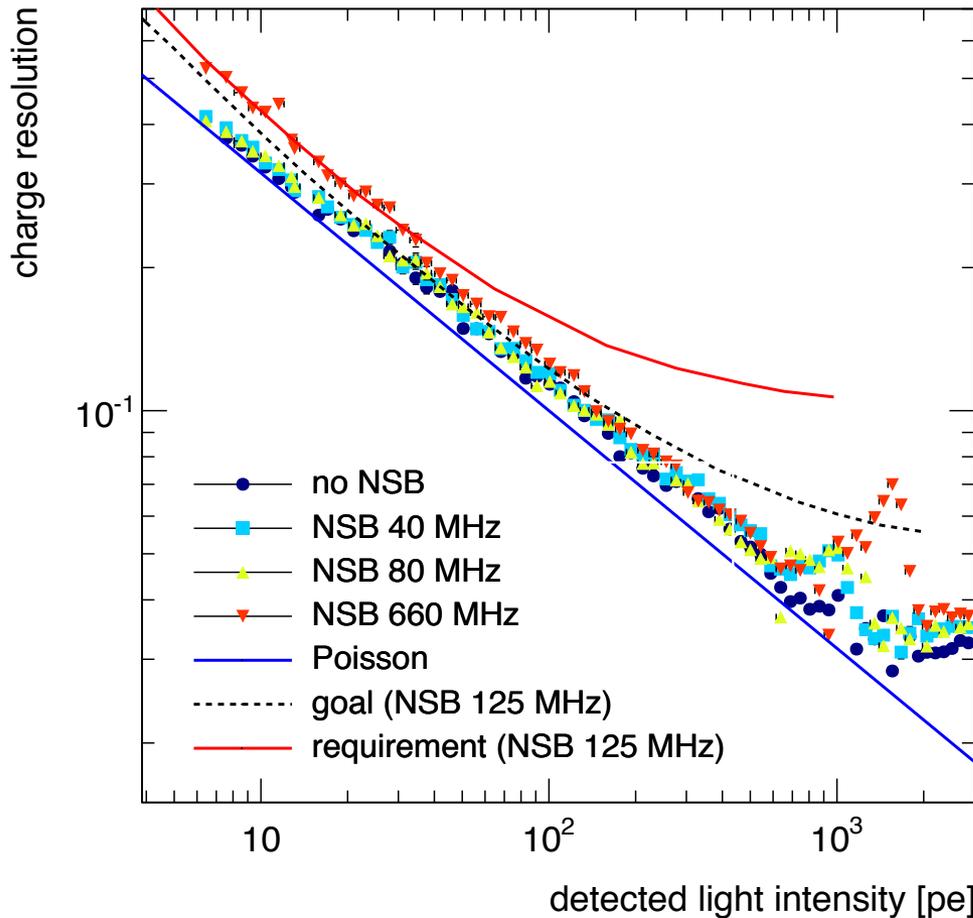
Performance

Linear
medium light



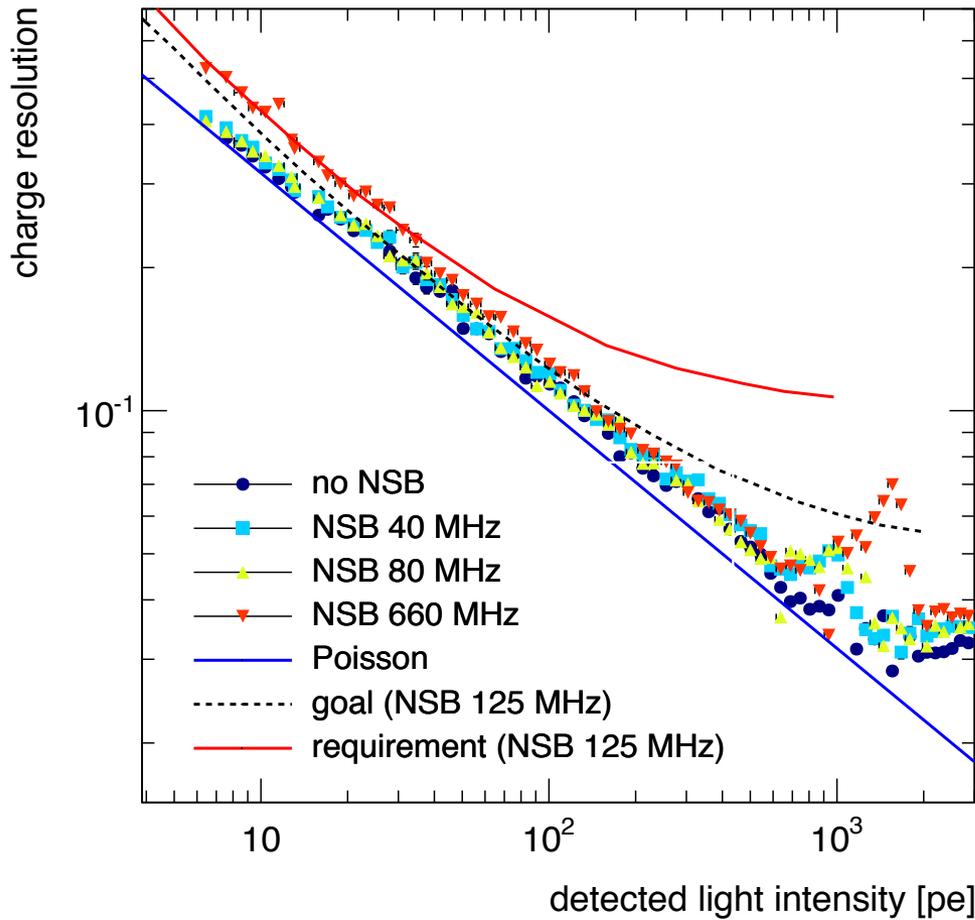
Performance

Linear
medium light

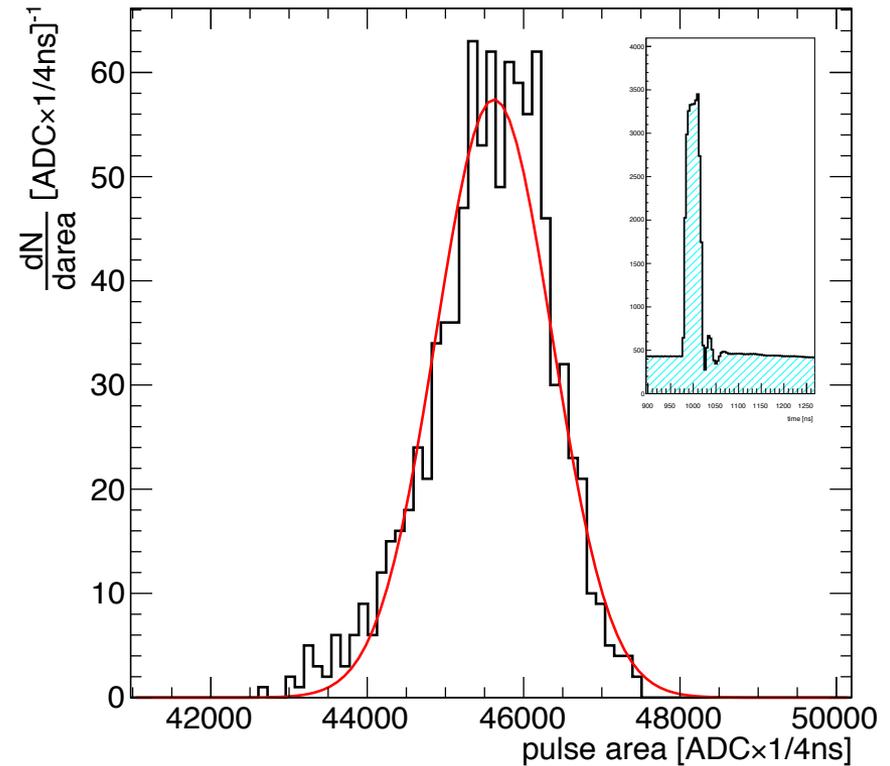
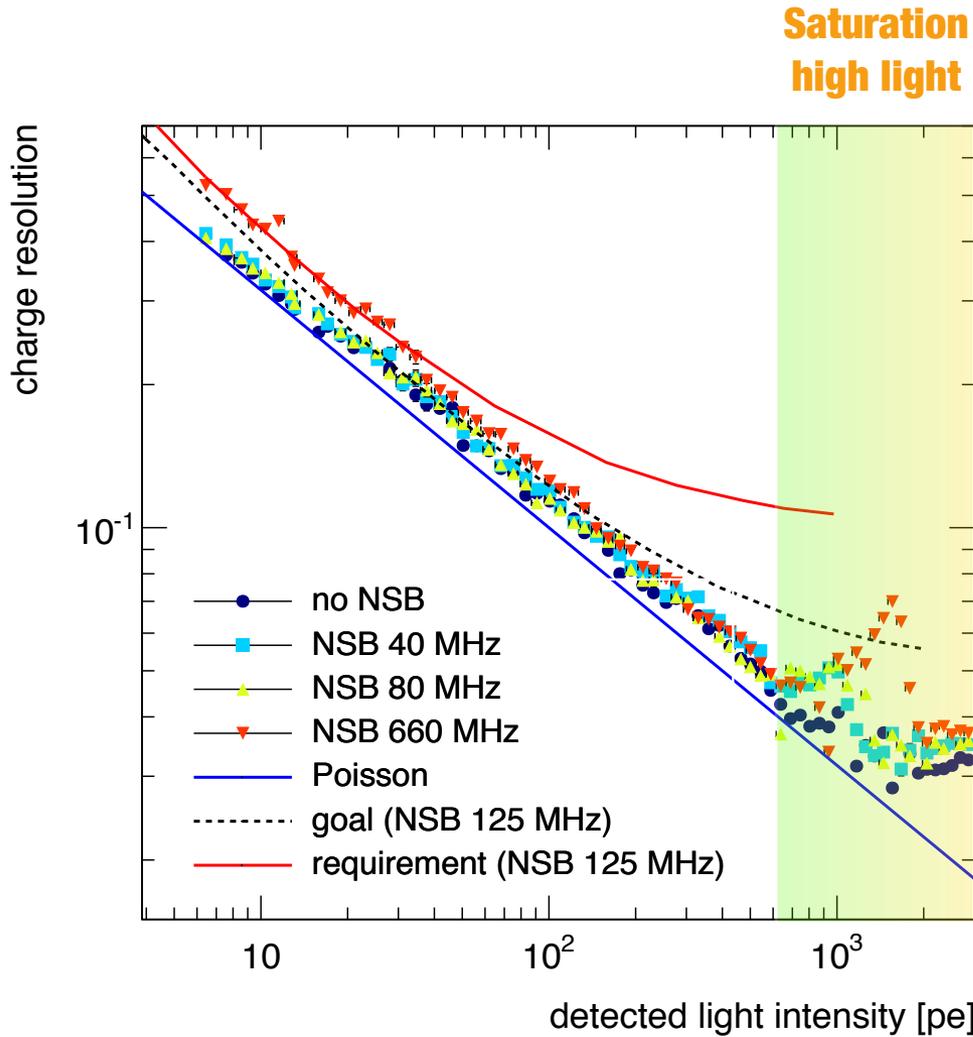


Performance

Saturation
high light

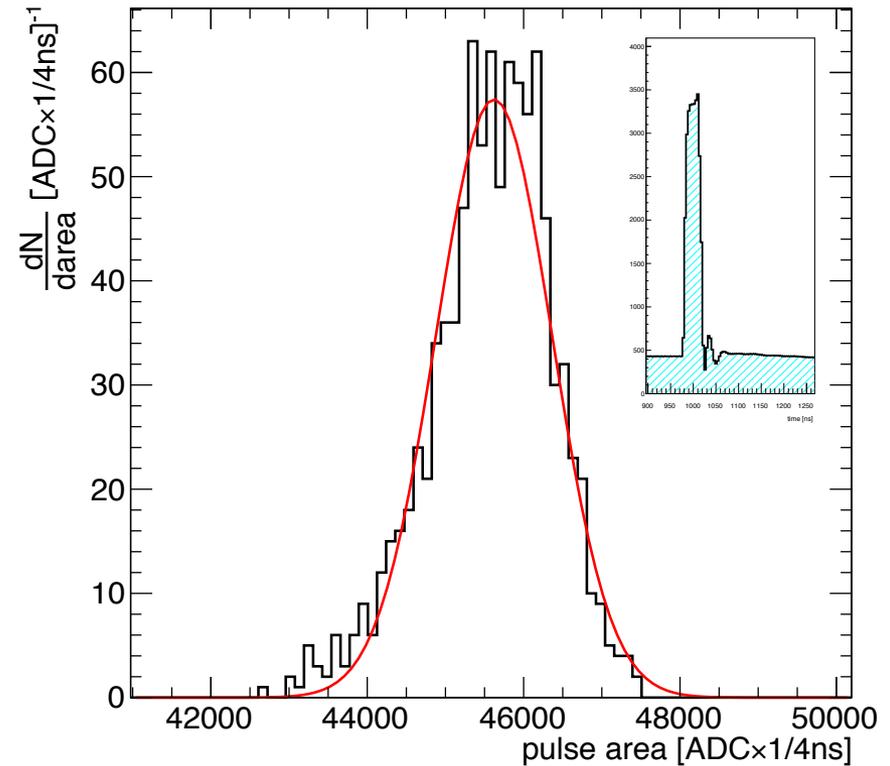
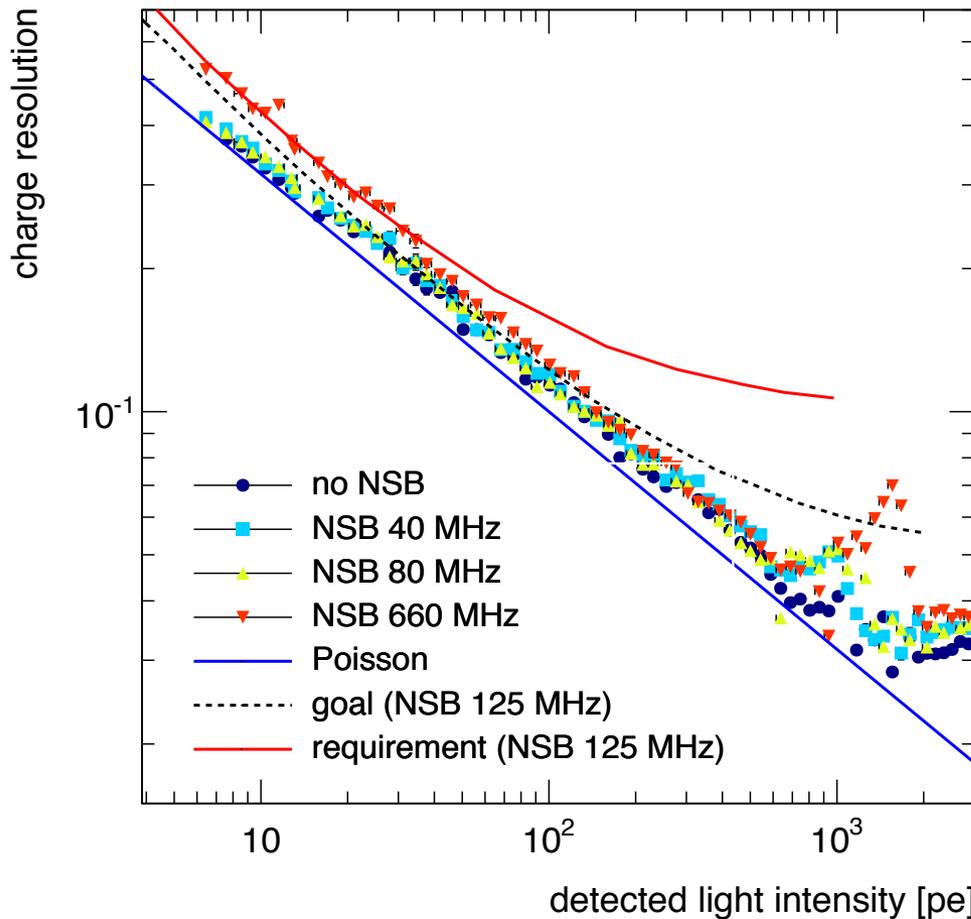


Performance

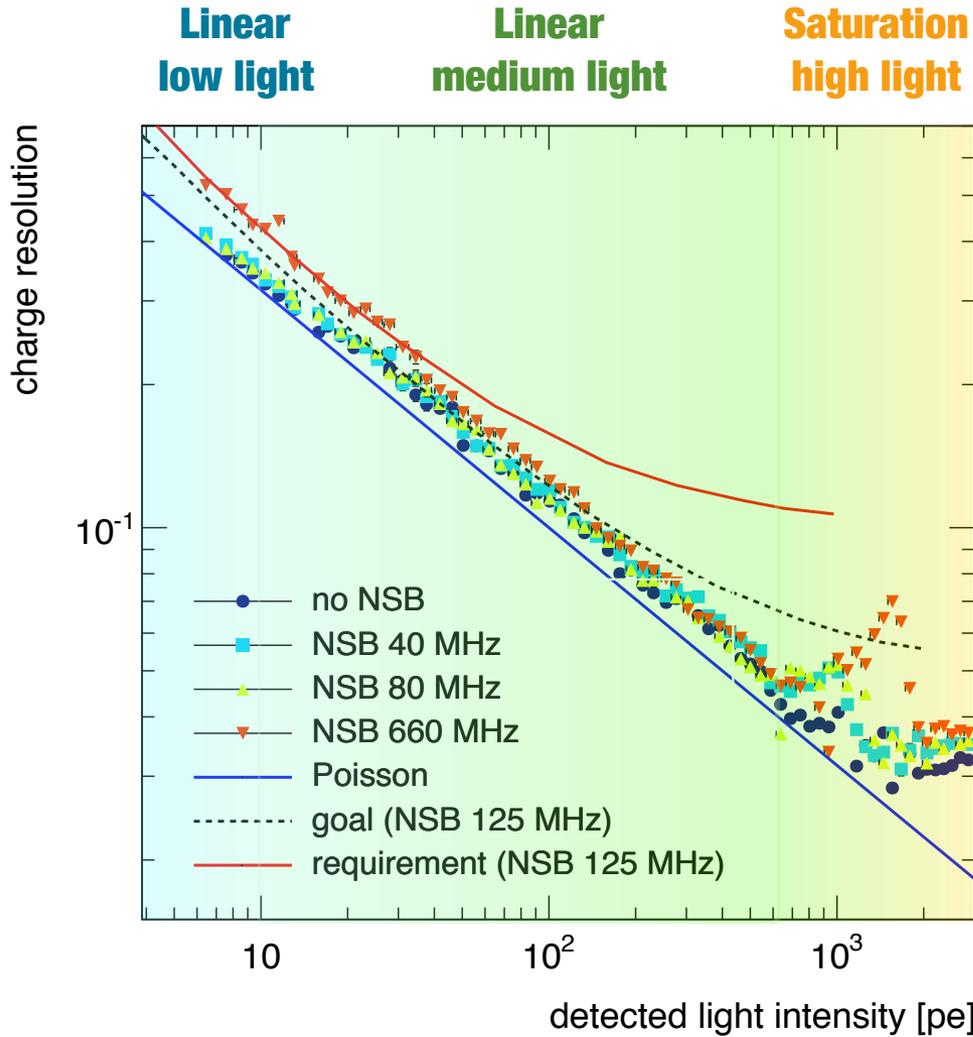


Performance

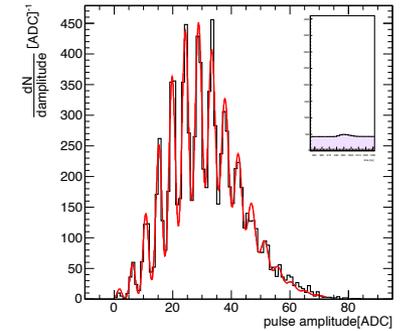
Saturation
high light



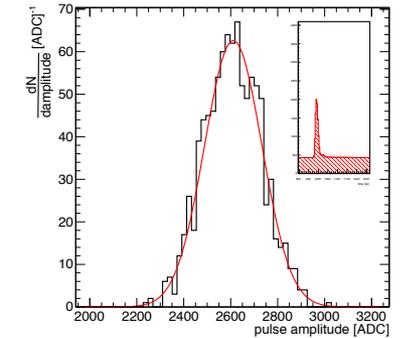
Performance



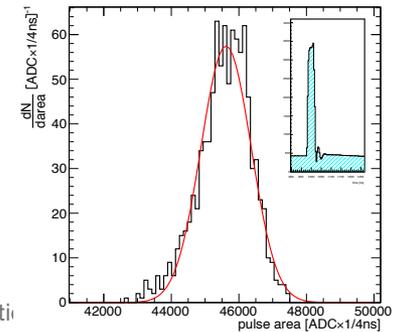
Linear low light



Linear medium light

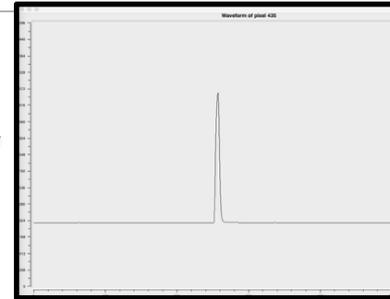
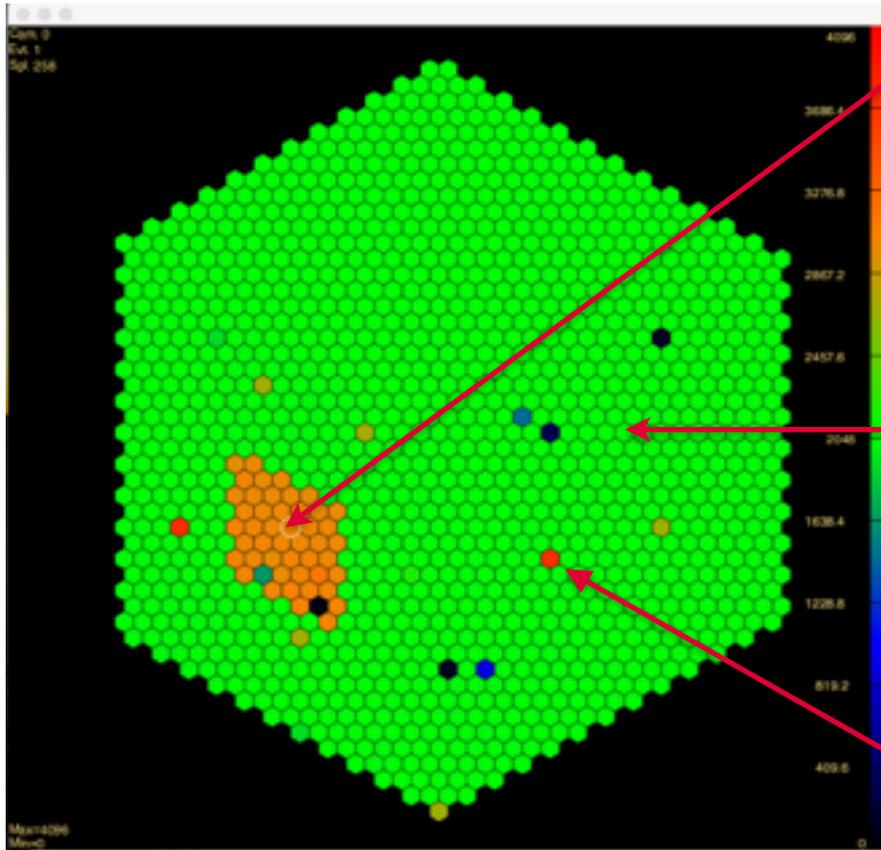


Saturation high light



First operations of the camera

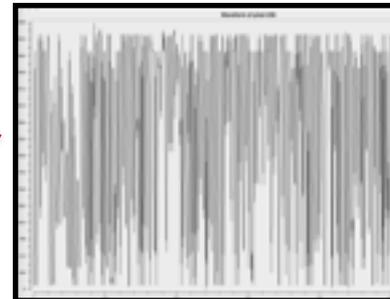
- Current Setup (See Vitalii's talk)



Normal pixel with injected signal



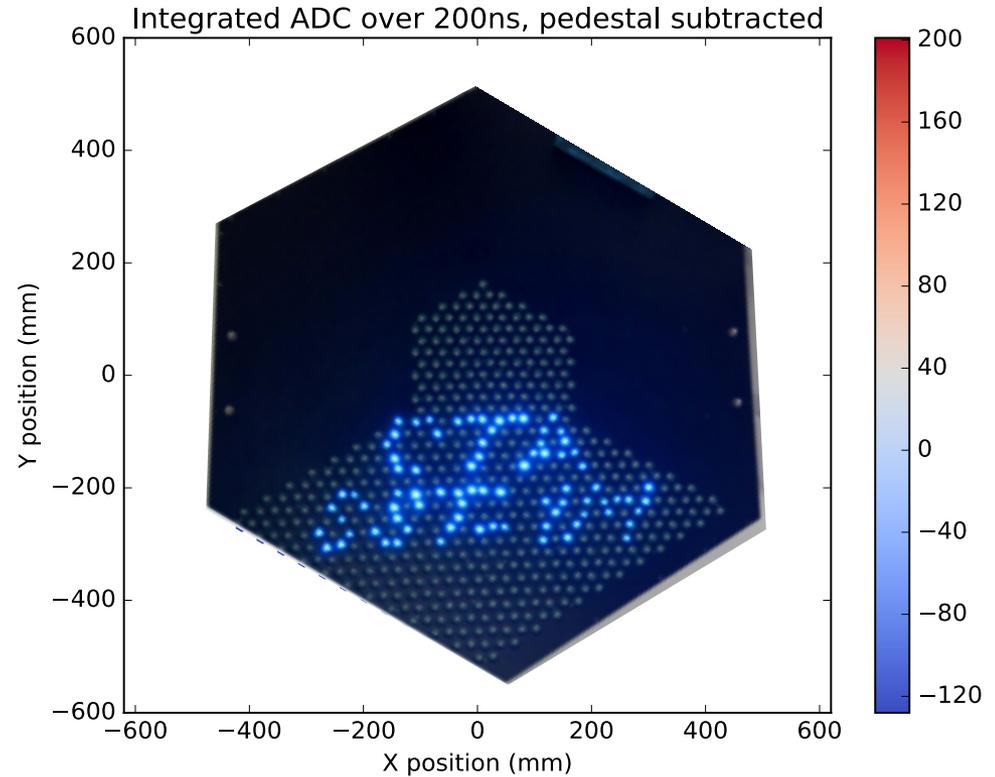
Normal pixel w/o signal



Unsynchronised pixel

First operations of the camera

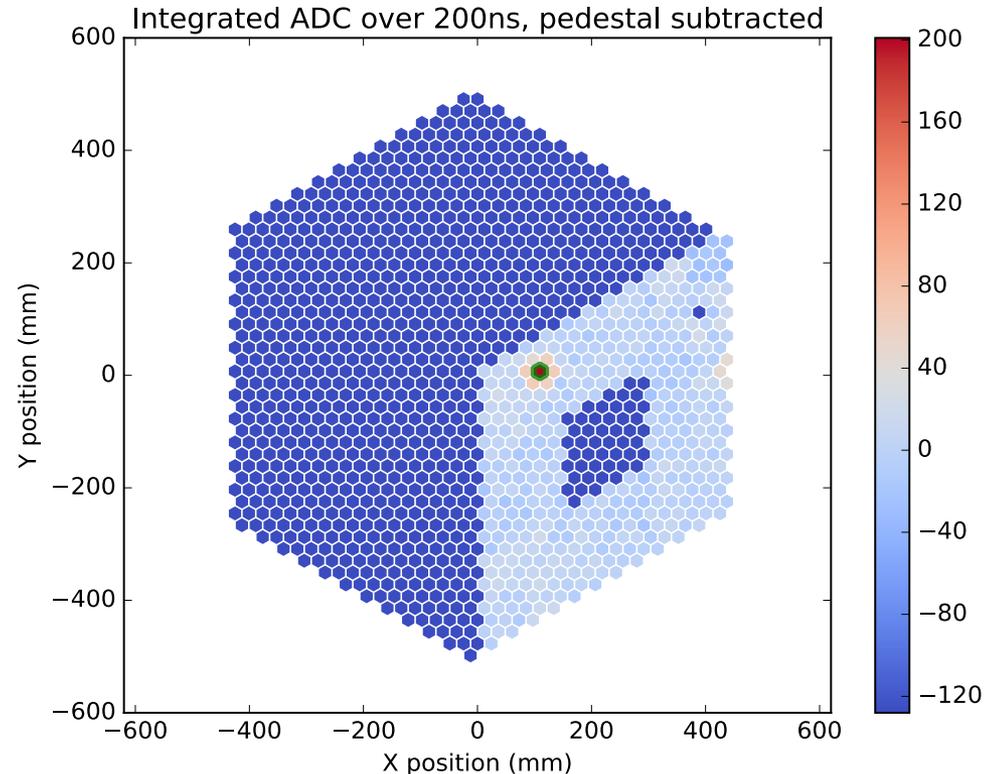
- A camera test set-up has been realised.
- 1 sector (1/3 of the camera) + the centre of the camera can be tested at the same time
- For each pixel, 2 LEDs are present: one flashing (AC) to emulate Cherenkov flashes and one DC to emulate NSB.
- Cherenkov shower-like shapes can be produced to verify and validate trigger algs and efficiency
- The set-up is very flexible and allow to check the cabling/ mapping, the working pixels, the flat fielding, charge resolution for each individual pixel, calibrations, trigger algs



Pixels scan to check mapping and working pixels

First operations of the camera

- A camera test set-up has been realised.
- 1 sector (1/3 of the camera) + the centre of the camera can be tested at the same time
- For each pixel, 2 LEDs are present: one flashing (AC) to emulate Cherenkov flashes and one DC to emulate NSB.
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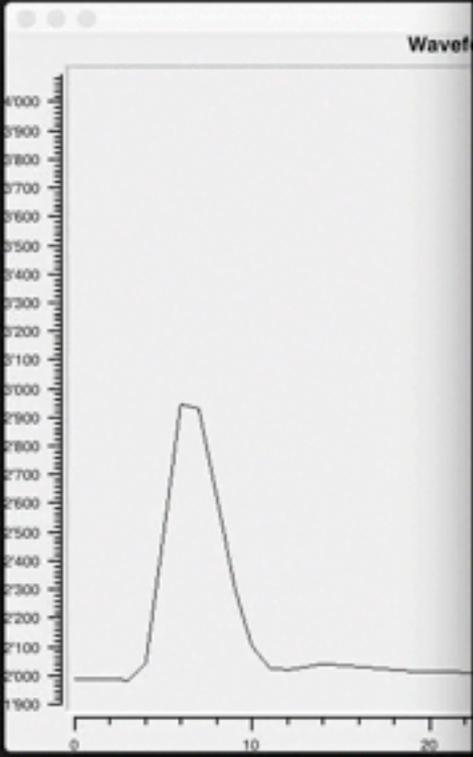
Pixels scan to check mapping and working pixels

Conclusions

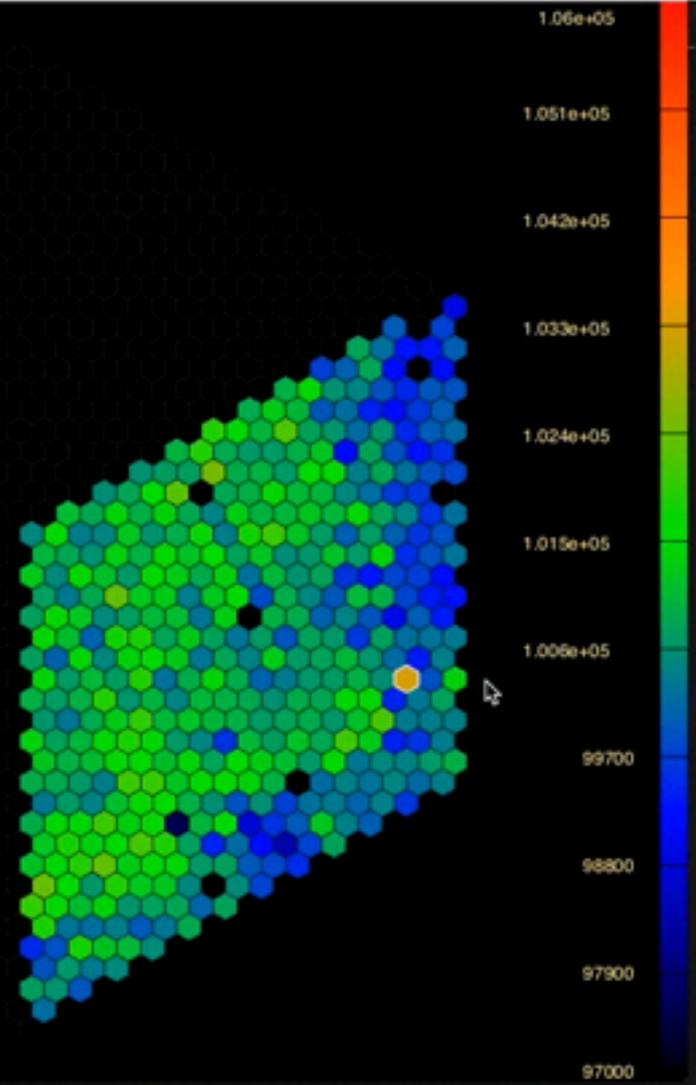
- SiPM use is spreading in many fields given their advantages.
- SiPM are particularly fit for gamma-ray astronomy,
 - Operation during Moonlight ~ 30% larger duty cycle
 - No evidence of ageing
 - Lightweight and robust cameras
 - Excellent single PE sensitivity
 - High Photo-Detection Efficiency at ~ 40%
- SST-1M camera goes in this direction but tried to open a new road towards large area devices.
 - Custom designed hexagonal device in collaboration with Hamamatsu
 - Large Area devices are complicated to handle but can be done!
- Many Lessons learned
 - SiPM parameter spread verified and validated
 - Large capacitance can be mastered
- The camera is under test . Very soon result on the field with real data.... stay tuned
- The solution worked out for CTA can be exploit without major modification both in larger CTA telescope and LHAASO WFCTA.
- Now under evaluation last generation of FBK sensor to see if they can work in our approach at room temperature.

```
Dysnomia:/local/etienne/workspace/CamerasToACTL
17:44%Build.MacOS/bin/viewer --input /local/etienne/Ca
pixels 89,239,525,739,749,964,1036,1072 --first_event Spl.All
Info: There are 15924 events in this table
Assuming dead pixels list: 89 239 525 739 749 964 1036
Evt. 9234]
```

Cam 1
Evt. 160912



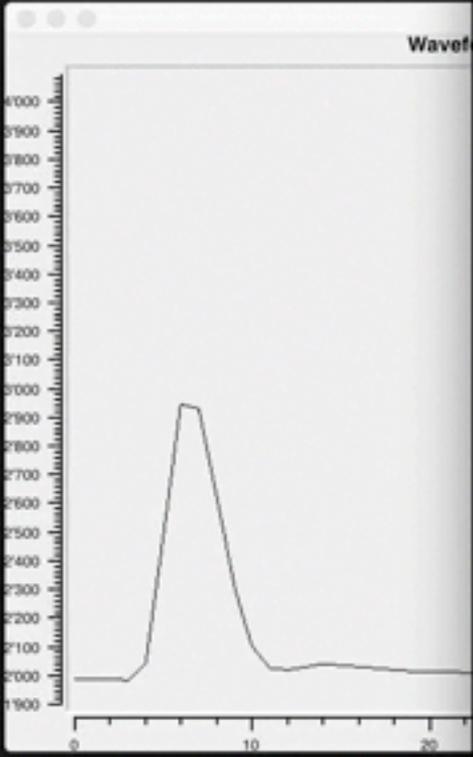
Max=1.06e+05
Min=97000



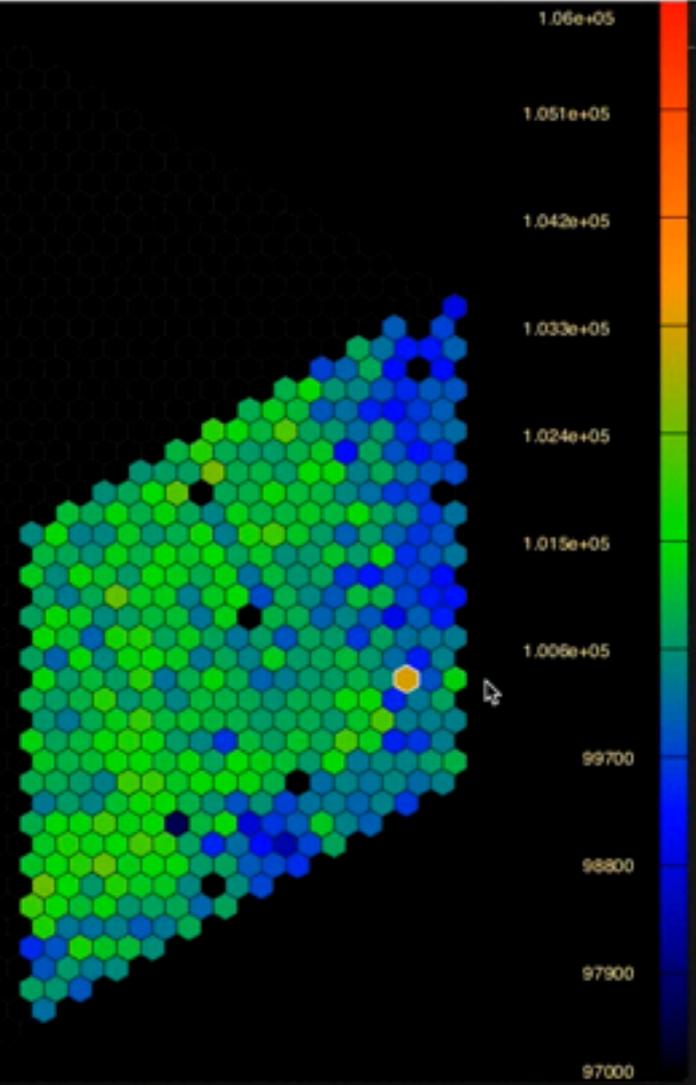
Looking for seeing real shower!!!

```
Dysnomia:/local/etienne/workspace/CamerasToACTL
17:44%Build.MacOS/bin/viewer --input /local/etienne/Ca
pixels 89,239,525,739,749,964,1036,1072 --first_event Spl.All
Info: There are 15924 events in this table
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Evt. 9234]
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Cam 1
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Max=1.06e+05
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Looking for seeing real shower!!!