

Hadronic Interaction Models

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**7th Workshop on Air Shower Detection at High Altitude,
Turin, Italy**

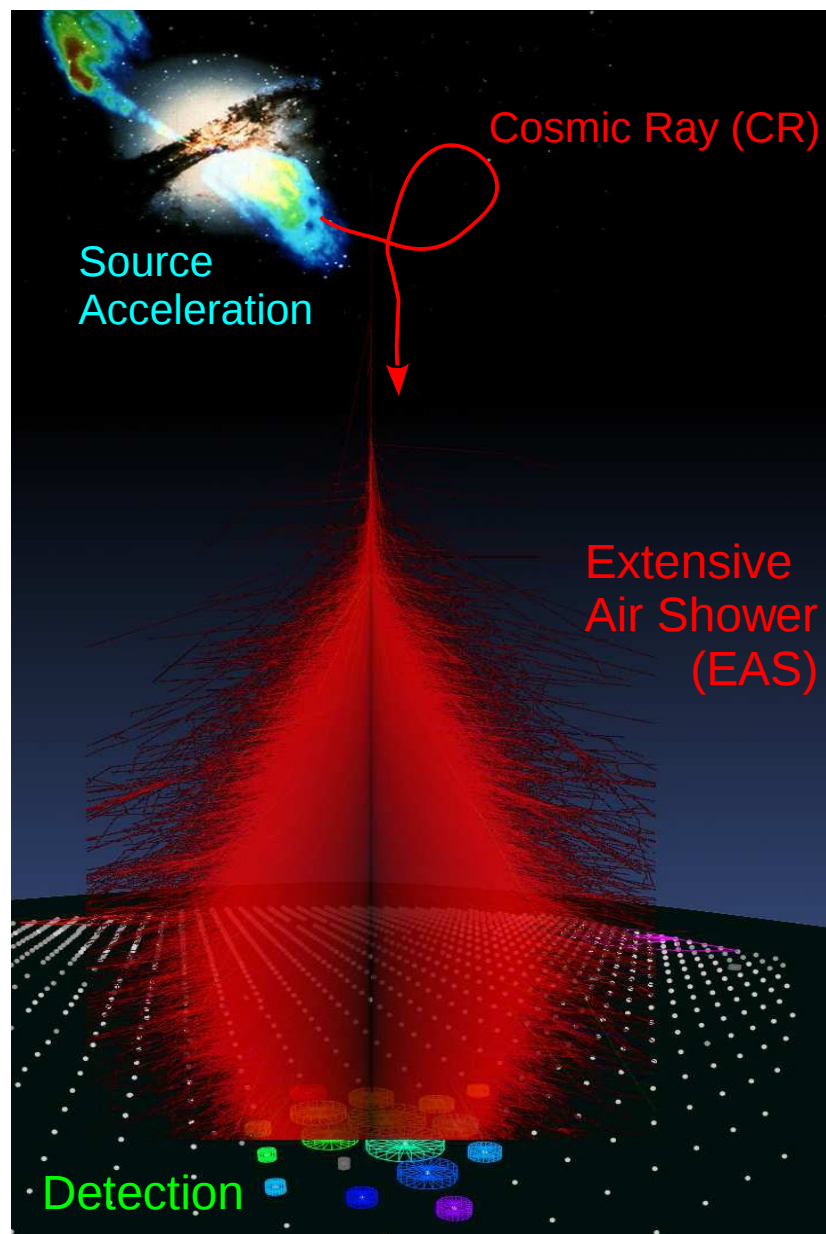
November the 30th 2016

Outline

- Monte-carlo for Cosmic Ray analysis
- MC comparison to accelerator data
- Electromagnetic (EM) signal in extended air showers
 - ➔ source of uncertainties
- Muon signal

LHC data reduced the model uncertainties and **exclude old models** for mass composition of cosmic rays. **Remaining uncertainties** can be further reduced taking into account **forward measurements** AND using (light) **nuclear target**.

Preamble



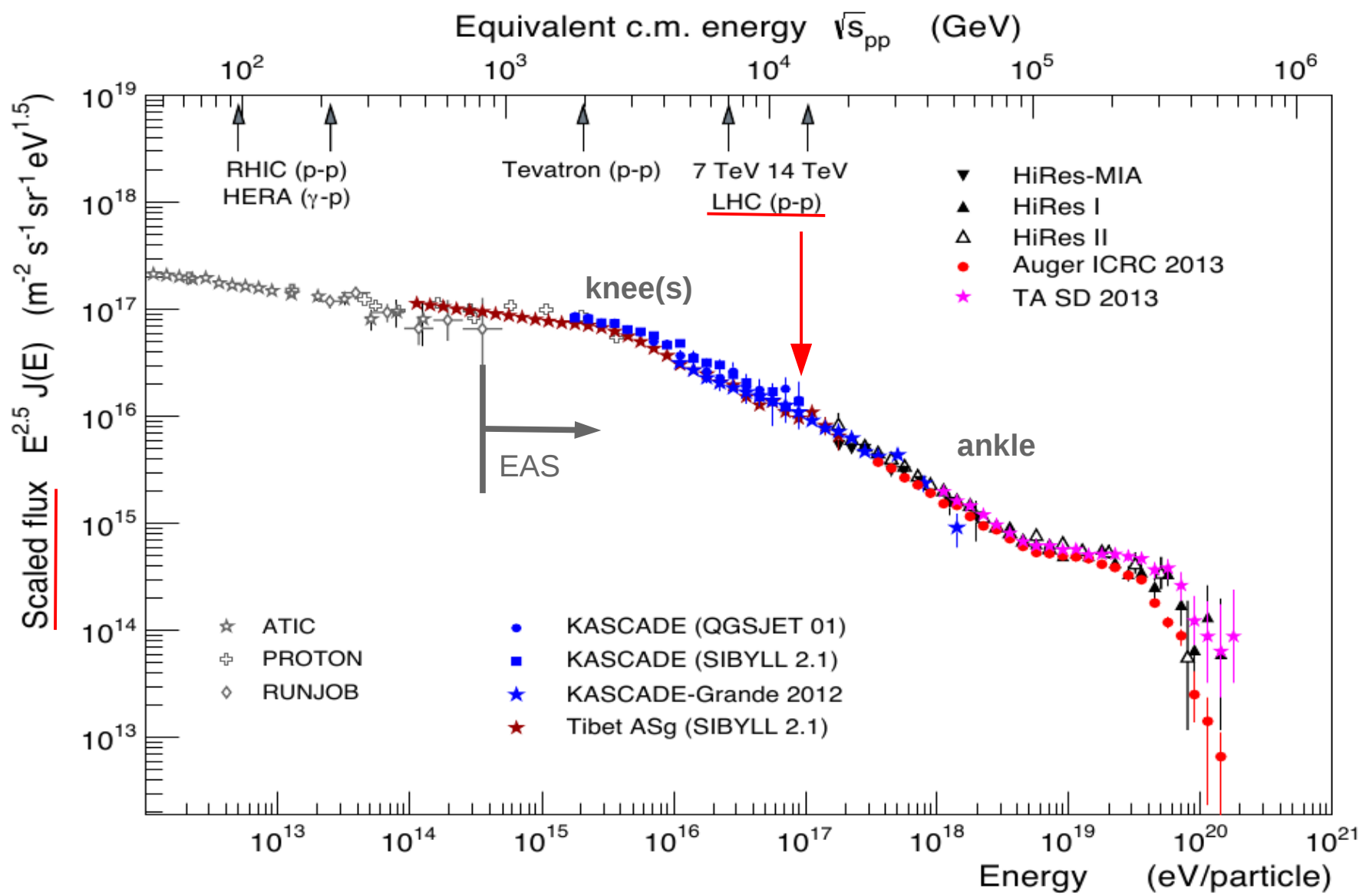
From R. Ulrich (KIT)

- **Goal of Astroparticle Physics :**
 - ➔ astronomy with high energy particles

- **How to test hadronic interactions ?**
 - ➔ if the source mechanism is well understood we could have a known beam at ultra-high energy (10^6 GeV and more)
 - ➔ improving but not very precise
 - ➔ reasonable minimum limits from CR abundance :
 - ◆ low = hydrogen (proton)
 - ◆ high = iron ($A=56$)
 - ➔ test of hadronic interactions in EAS via correlations between observables.

mass measurements should be consistent
and lying between proton and iron
simulated showers if physics is correct

Cosmic Ray Spectrum



R. Engel (KIT)

Hadronic Interaction Models in CORSIKA

(HDPM)

Old generation : (SIBYLL 2.1 QGSJET01 DPMJET 2.55 VENUS) (<2001)

All Glauber based
But differences in hard, remnants, diffraction ...

New (!) generation :

LHC tuned :

LHC inspired : SIBYLL 2.3

Motivation :
- update with latest LHC results in simple model

Engel et al.

semi-hard

(QGSJET II-03)

QGSJET II-04

Ostapchenko

QGSJET III (?)

Motivation :
- Hard Pomeron-Pomeron connexion

NEXUS 3.97

soft

Attempt to get everything described in a consistent way (energy sharing)

(EPOS 1.99)

EPOS LHC

Pierog & Werner

Fedinitch & Engel

DPMJET III

Motivation :
- update with LHC results
- fix high energy

EPOS 3

Motivation :
- binary scaling in hard probes
- better diffraction

Riehn & Engel

(2005-2012)

(2013-)

(2017-)

Cosmic Ray Hadronic Interaction Models

● Theoretical basis :

- pQCD (large p_t)
- Gribov-Regge (cross section with multiple scattering)
- energy conservation

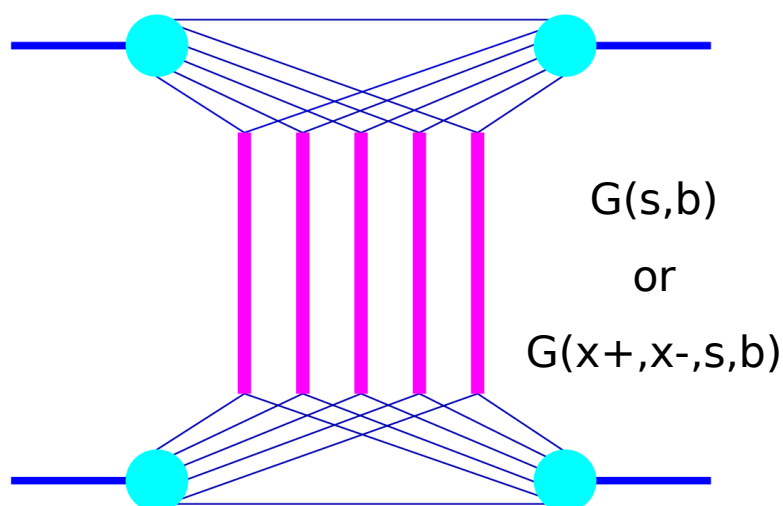
● Phenomenology (models) :

- hadronization
 - string fragmentation
 - EPOS : high density effects (statistical hadronization and flow)
- diffraction (Good-Walker, ...)
- higher order effects (multi-Pomeron interactions)
- remnants

● Comparison with data to fix parameters

Better predictive power than HEP models thanks to link between total cross section and particle production (GRT) tested on a broad energy range (including EAS)

Cross Section and Multiplicity in Models



● Gribov-Regge and optical theorem

- ➔ Basis of all models (multiple scattering) but
 - ◆ Classical approach for QGSJET, SIBYLL and DPMJET (no energy conservation for cross section calculation)
 - ◆ Parton based Gribov-Regge theory for EPOS (**energy conservation at amplitude level**)

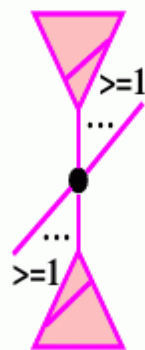
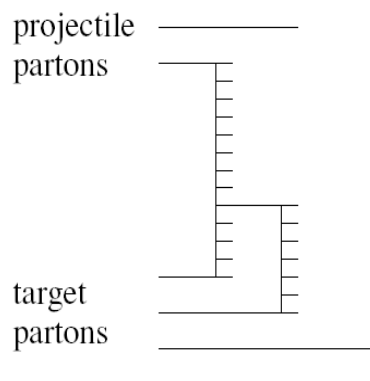
● pQCD

- ➔ Minijets with cutoff in SIBYLL and DPMJET
- ➔ Same hard Pomeron (DGLAP convoluted with soft part : no cutoff) in QGSJET and EPOS but
 - ◆ Generalized enhanced diagram in QGSJET-II
 - ◆ Simplified non linear effect in EPOS

- Phenomenological approach

EPOS

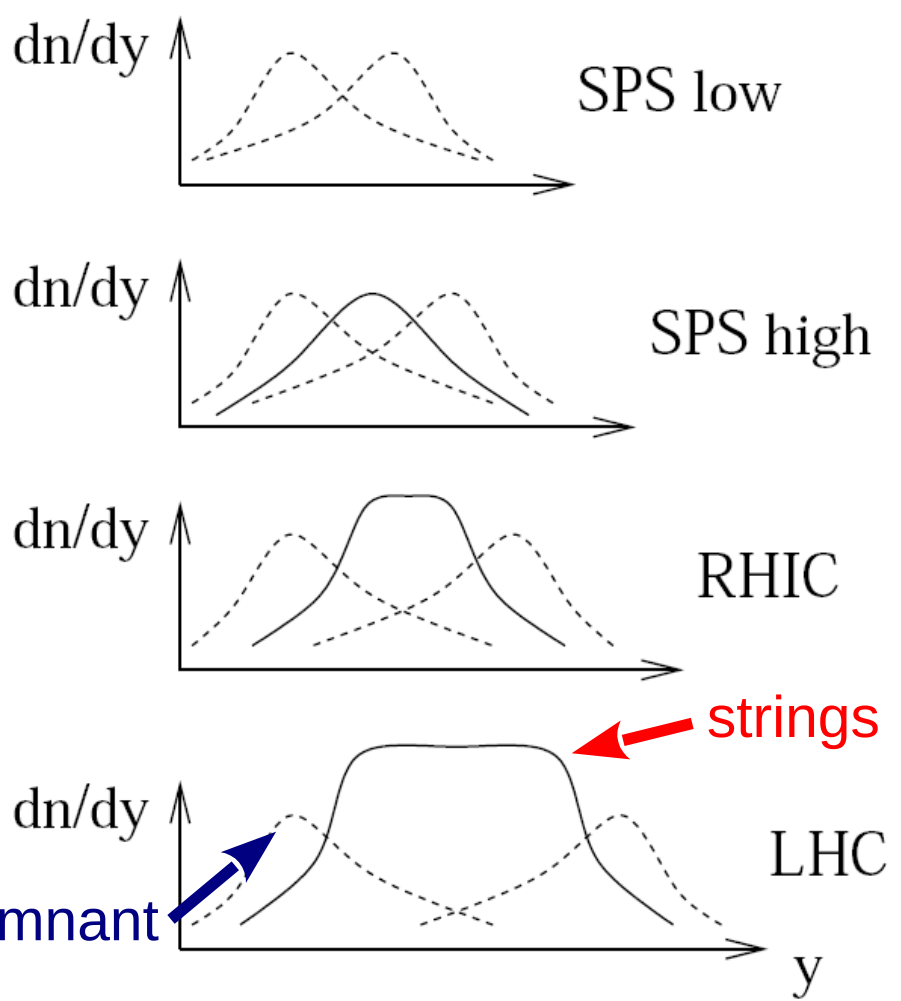
QGSJET II



Remnants

Forward particles mainly from projectile remnant

The (in)elasticity is closely related to diffraction and forward spectra

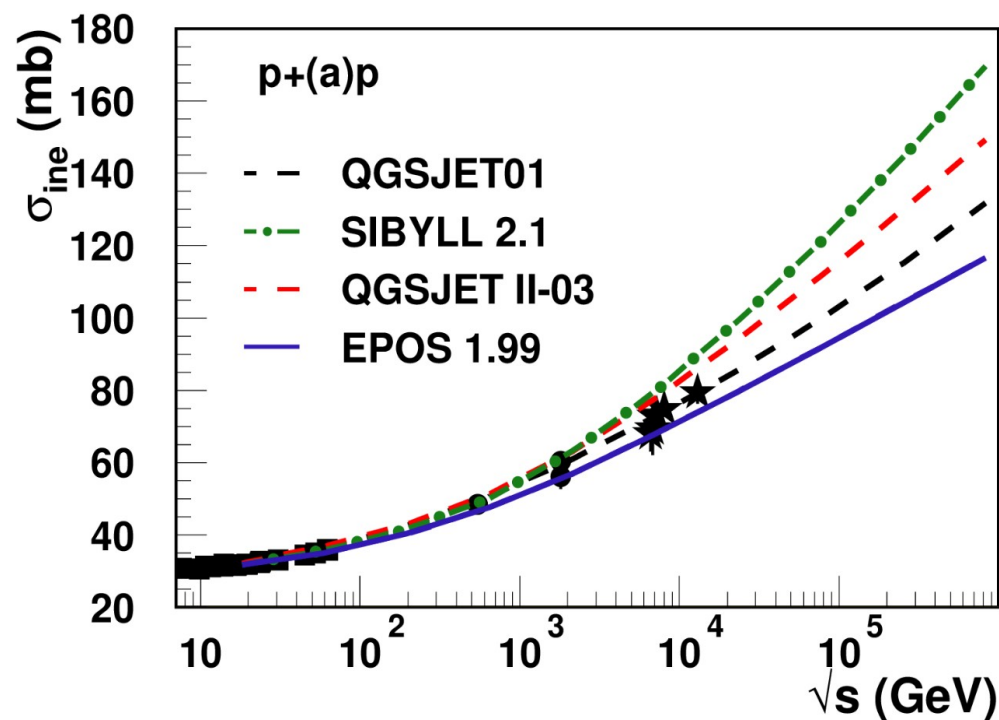


- ➔ SIBYLL
 - ◆ No remnant except for diffraction
 - Leading particle from string ends
- ➔ QGSJET
 - ◆ Low mass remnants
 - Leading particle similar to proj.
- ➔ EPOS
 - ◆ Low and high mass remnants
 - Any type of leading particle
 - from resonance
 - from string
 - from statistical decay

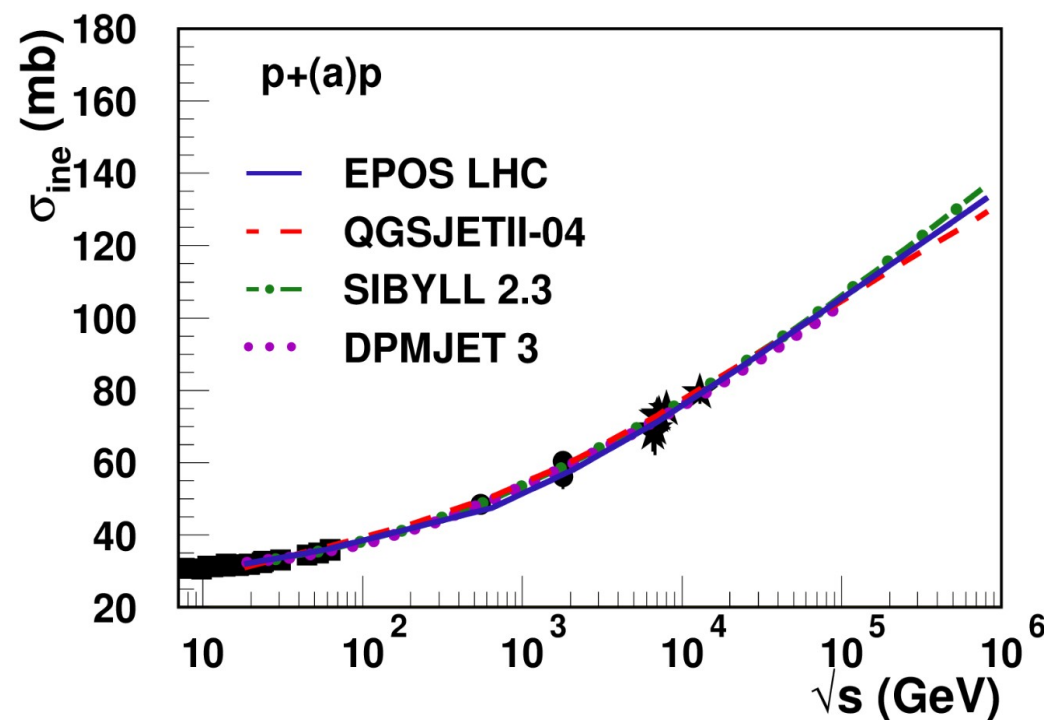
Cross Sections

- ➔ Same cross section prediction at pp level and low energy (data for tuning)
- ➔ extrapolation to high energy looks settled
 - ◆ different amplitude and scheme
 - ➔ same extrapolations

Pre - LHC



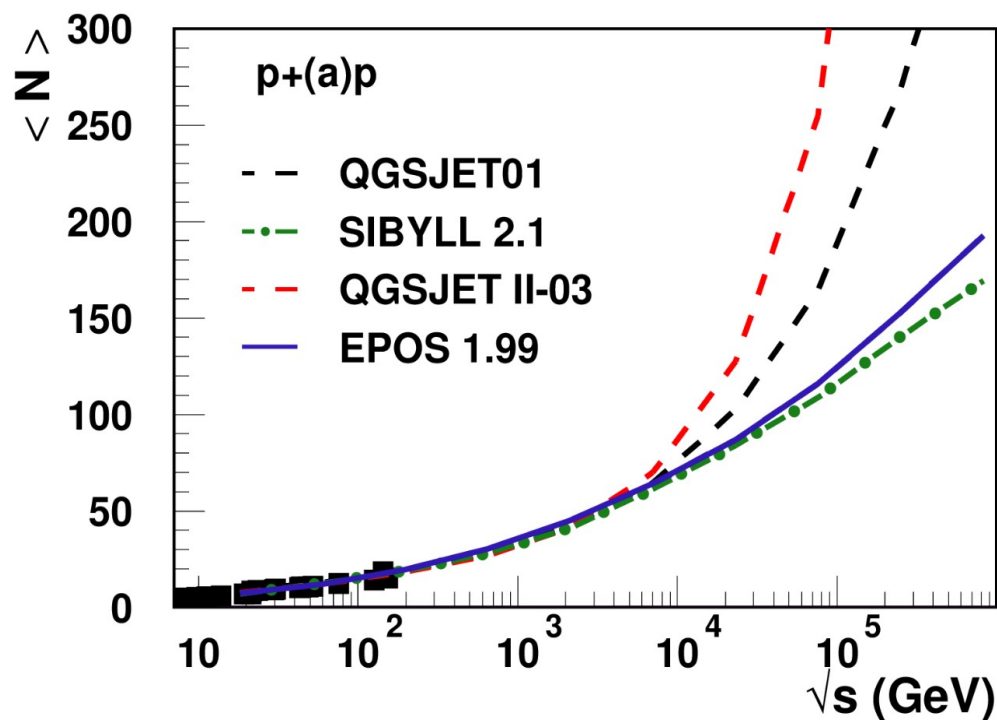
Post - LHC



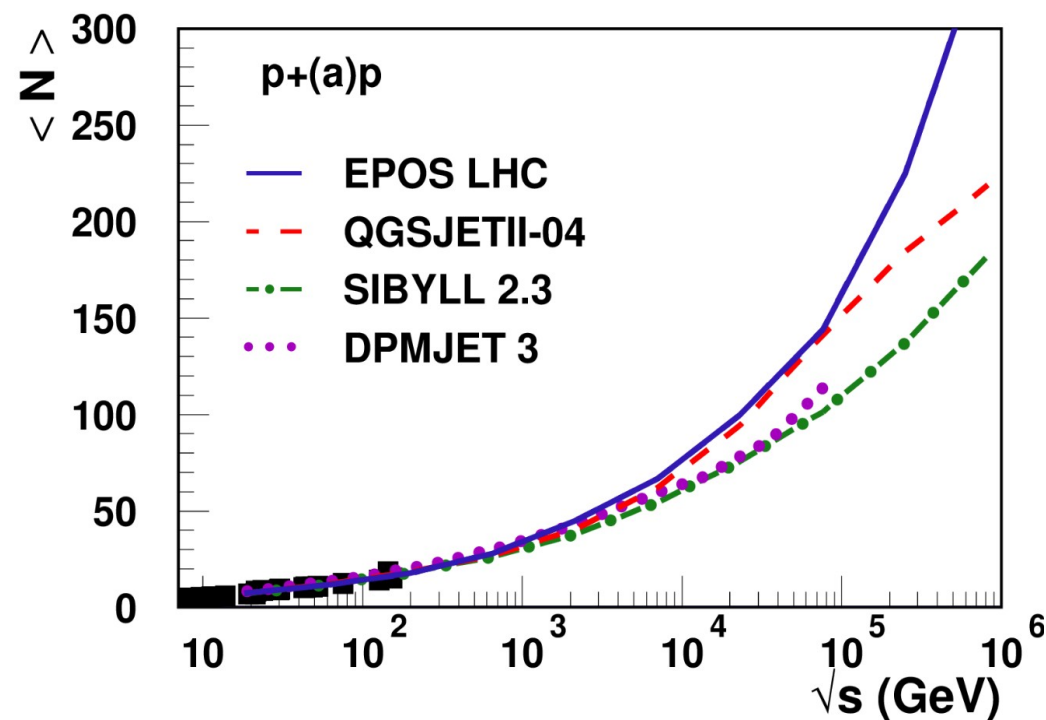
Multiplicity

- Multiplicity fixed by data up to 900 GeV
- extrapolation to high energy is still model dependent ?

Pre - LHC



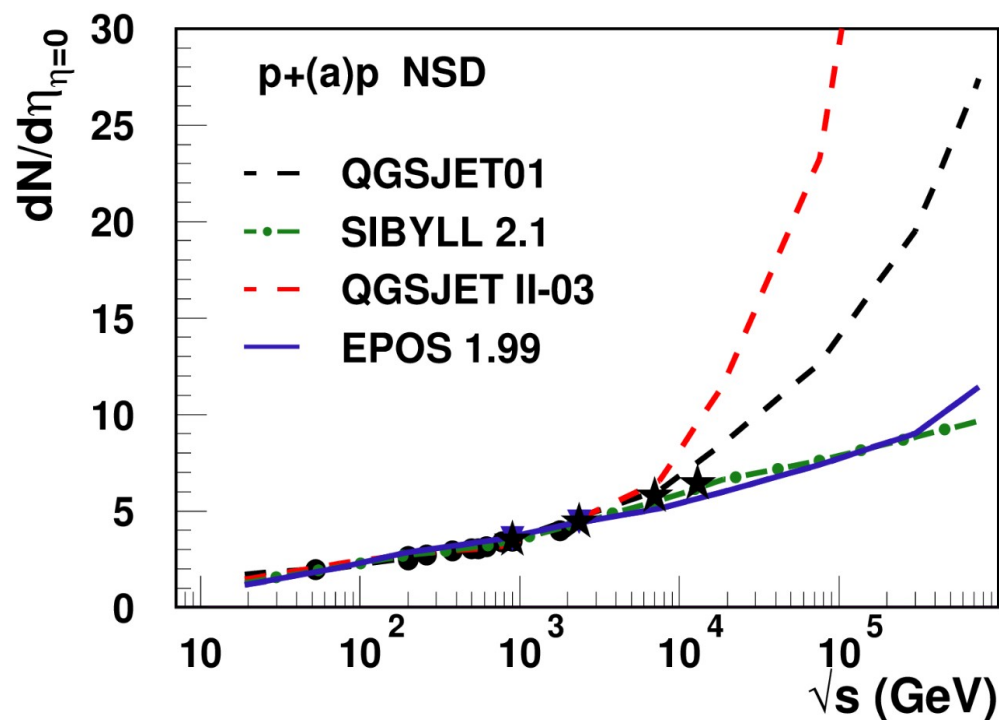
Post - LHC



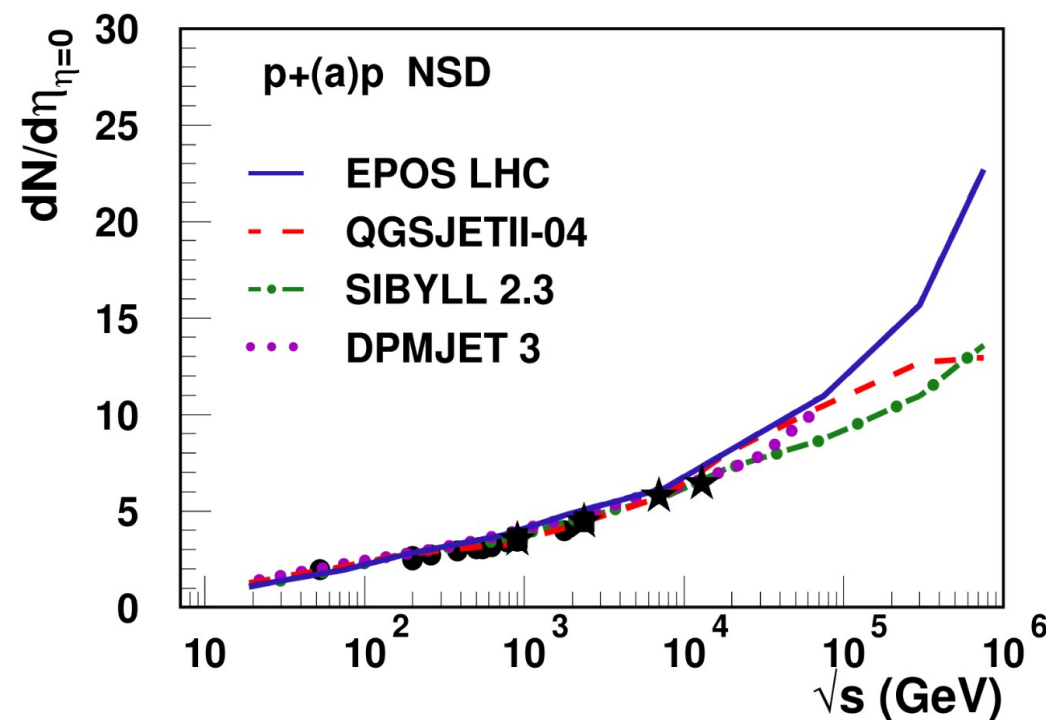
Multiplicity at mid-rapidity

- Multiplicity fixed by data up to 13 TeV
- extrapolation to high energy less model dependent after LHC
- QGSJET01 and QGSJETII-03 extrapolation excluded

Pre - LHC



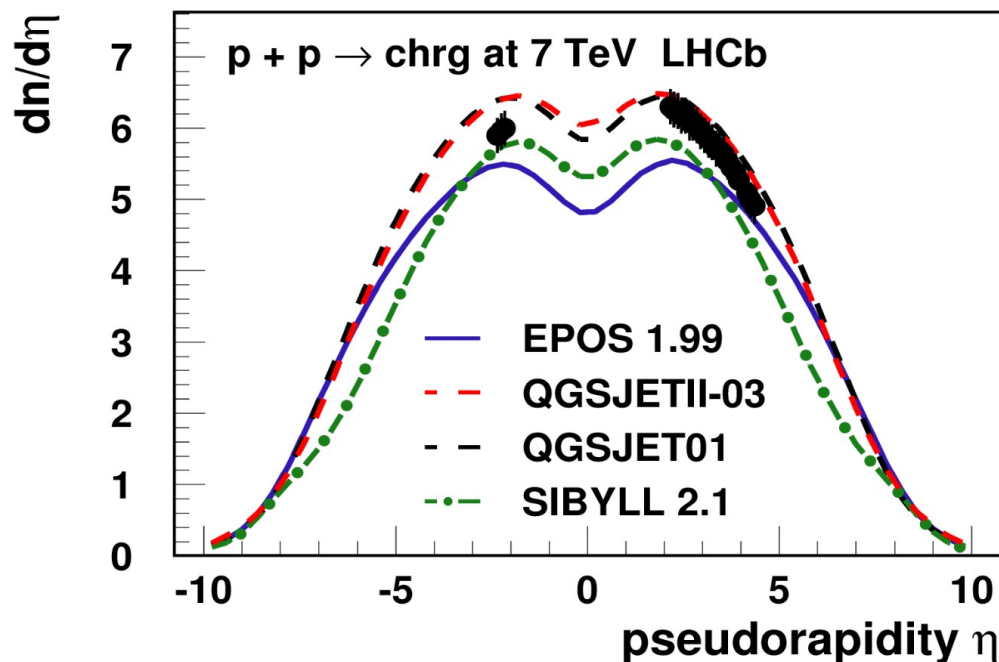
Post - LHC



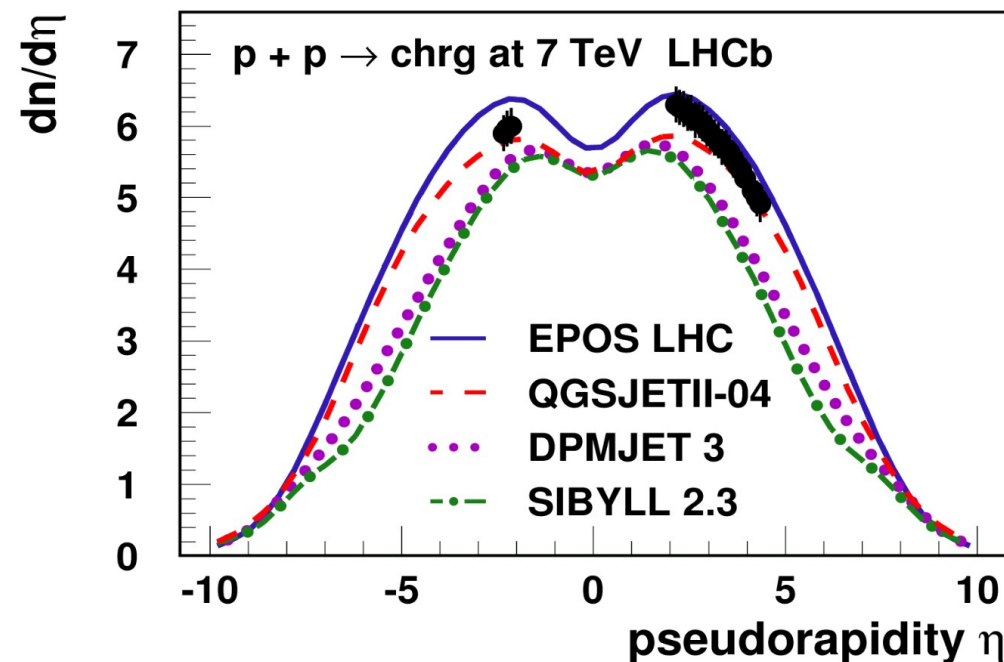
Pseudorapidity

- ➔ Difference between mid-rapidity and full multiplicity coming from the width of the pseudorapidity distributions
- ➔ From LHC data
 - DPMJET 3 and SIBYLL 2.3 too narrow
 - QGSJETII-04 ~ OK
 - EPOS LHC a bit too large

Pre - LHC



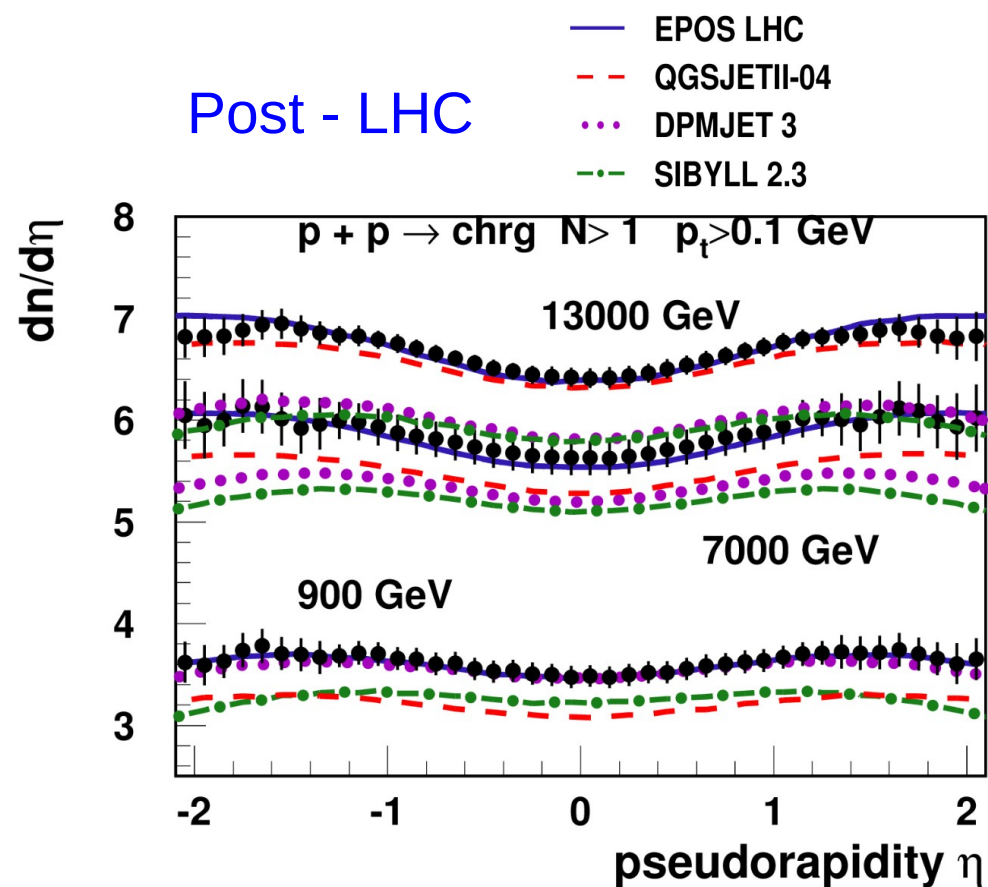
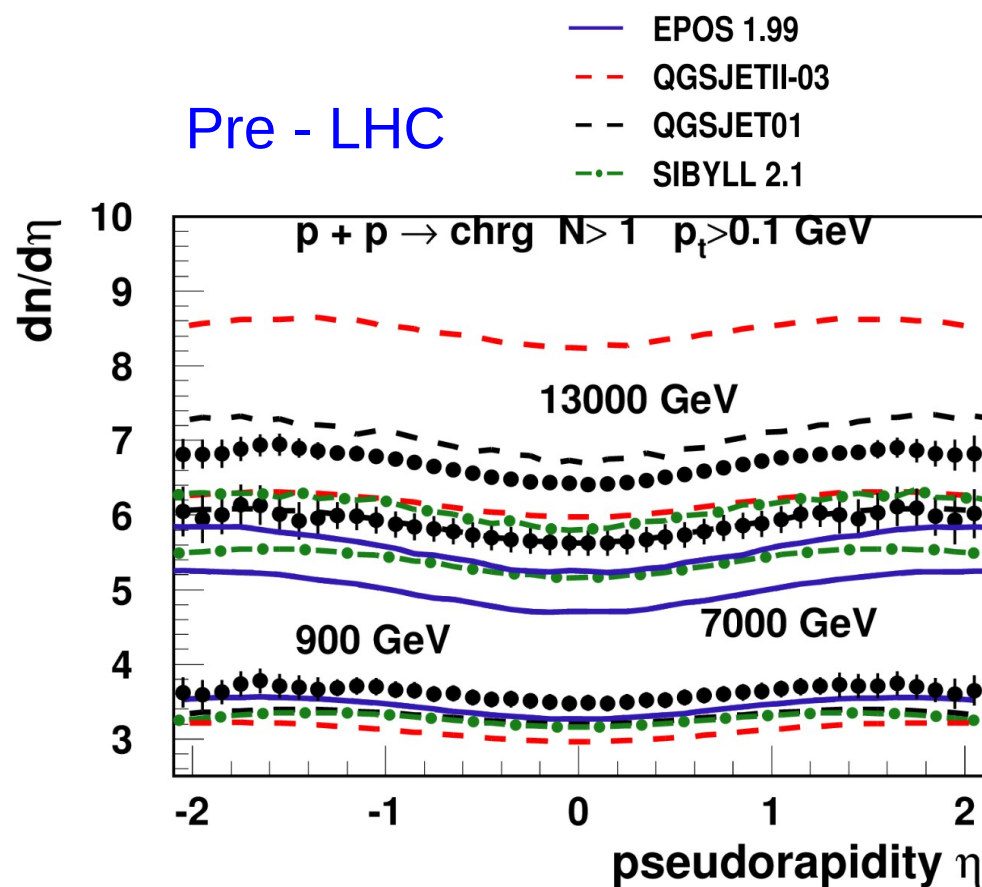
Post - LHC



Test of Models vs Accelerator Data

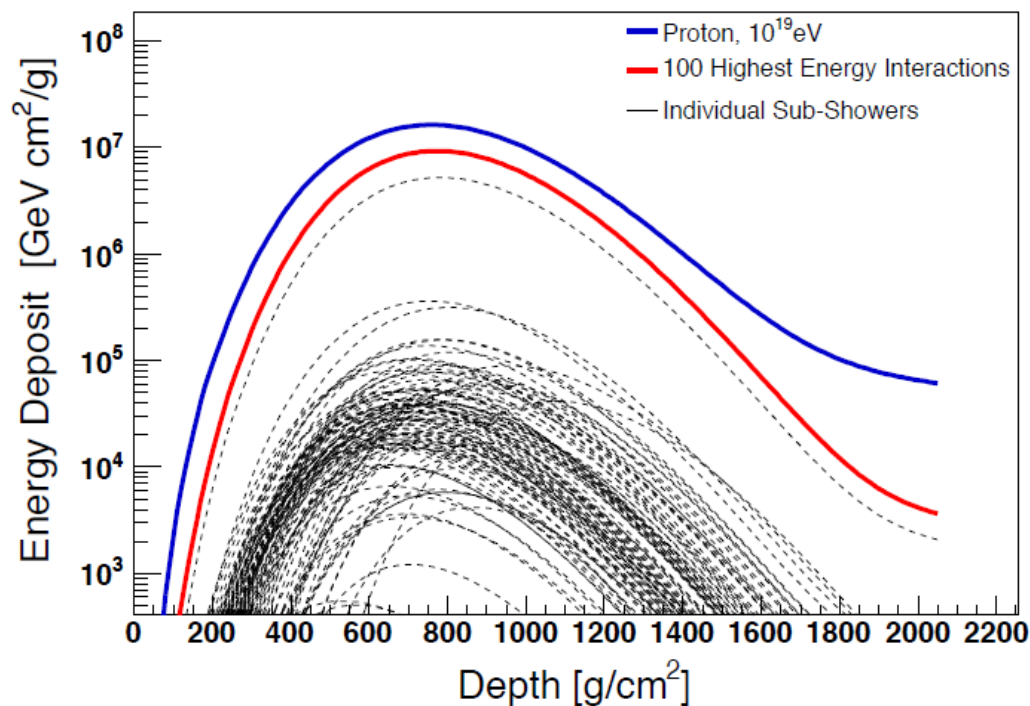
➔ From LHC data

- All pre-LHC models extrapolation excluded
- DPMJET 3 and SIBYLL 2.3 underestimate multiplicity
- QGSJETII-04 and EPOS LHC ~ OK (and similar to Pythia 8)



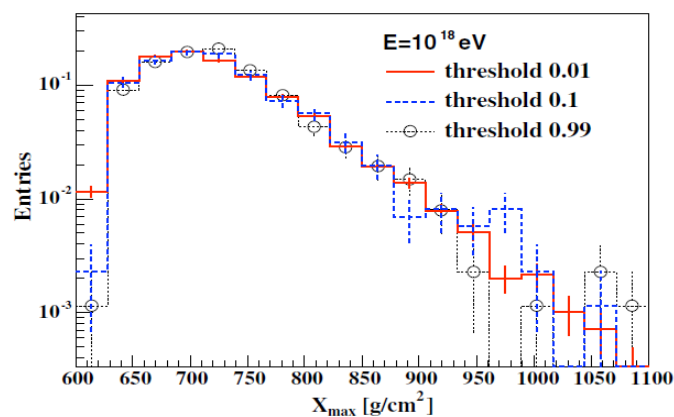
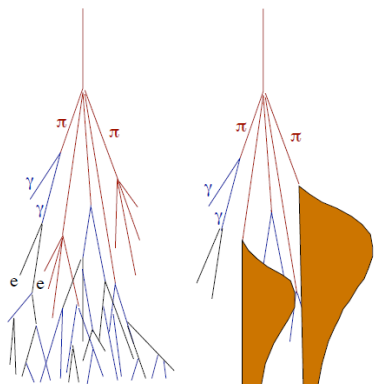
Energy of hadronic interactions for X_{\max}

Electrons



Shower particles produced in 100 interactions of highest energy

X_{\max} dominated by first (high energy) interaction(s) : proton (nucleus)-Air



Fluctuations mainly coming from the first hadronic interaction.

Simplified Shower Development

Using generalized Heitler model and superposition model :

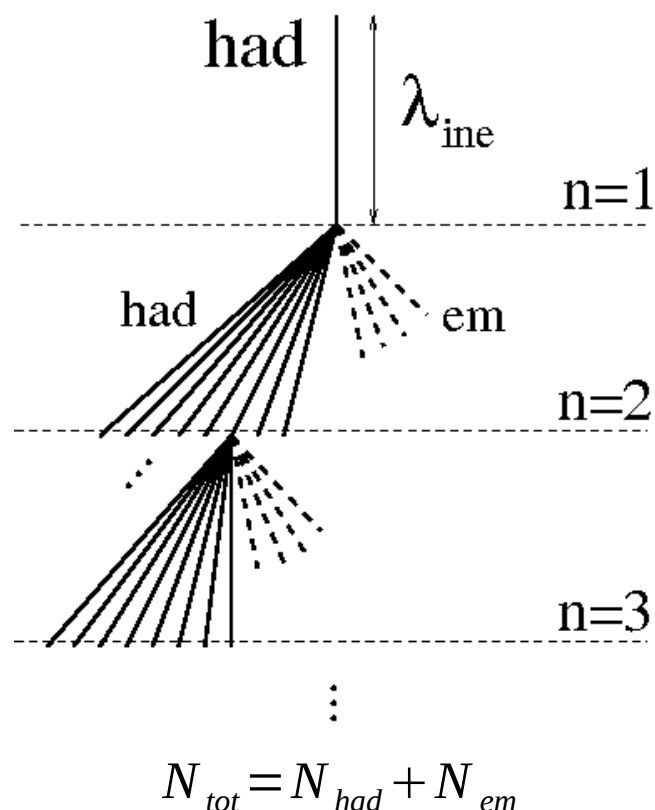
$$X_{max} \sim \lambda_e \ln \left((1-k) \cdot E_0 / (2 \cdot N_{tot} \cdot A) \right) + \lambda_{ine}$$

→ Model independent parameters :

- E_0 = primary energy
- A = primary mass
- λ_e = electromagnetic mean free path

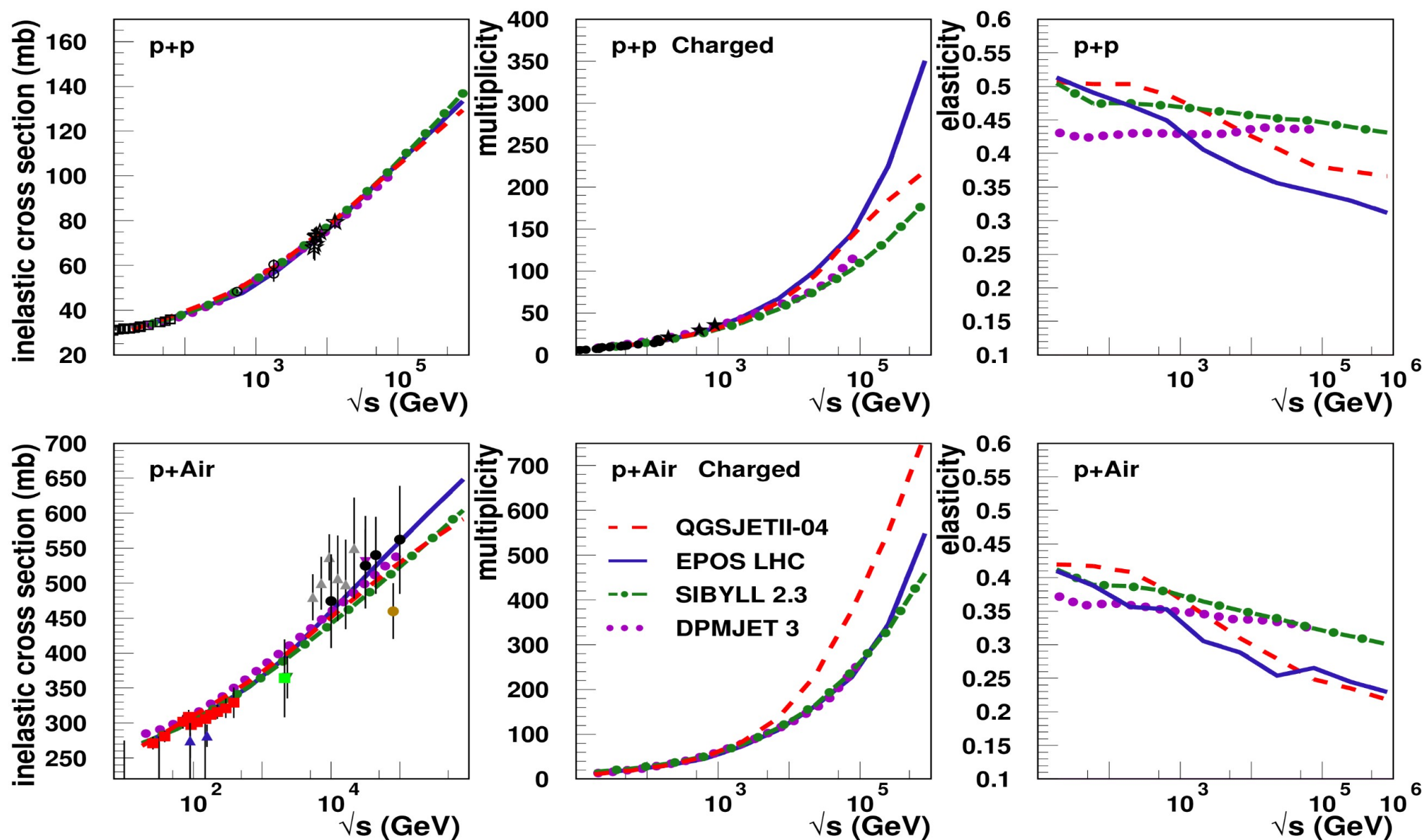
→ Model dependent parameters :

- k = elasticity
- N_{tot} = total multiplicity
- λ_{ine} = hadronic mean free path (cross section)

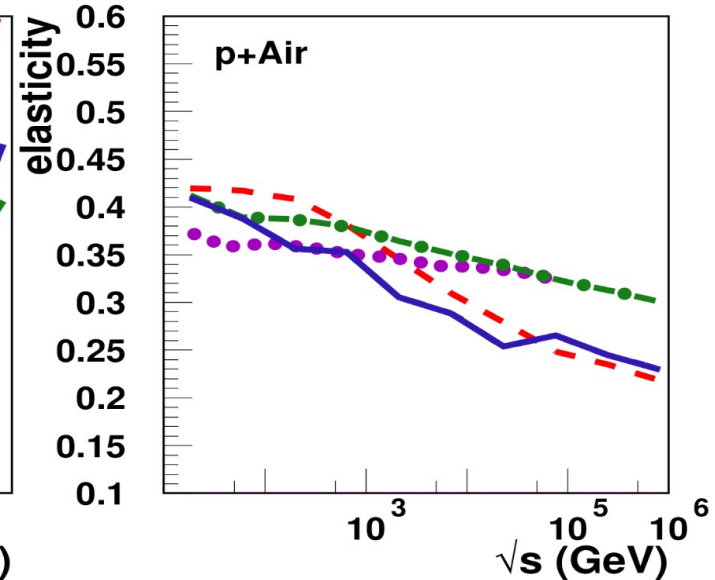
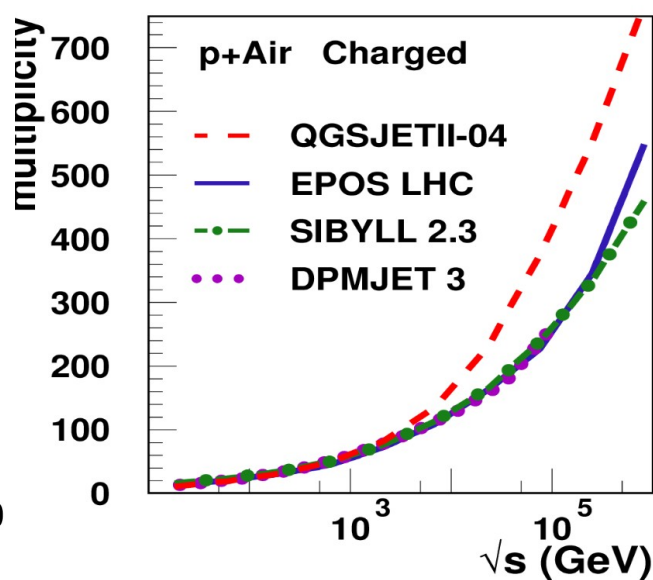
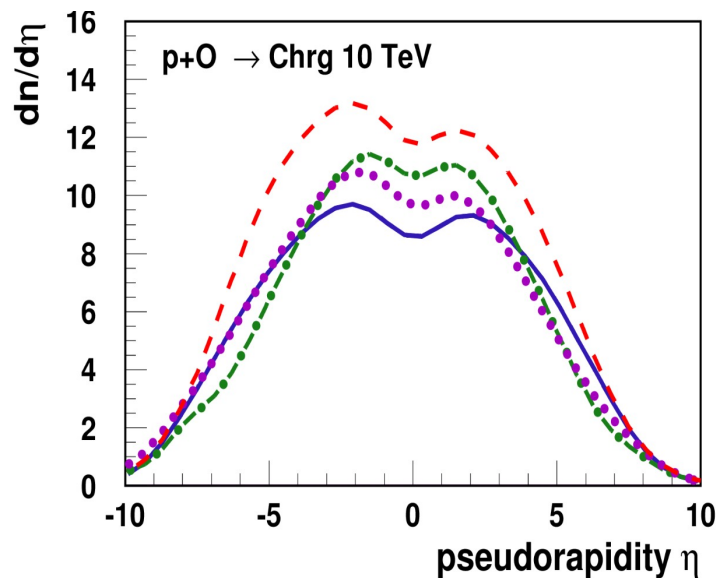
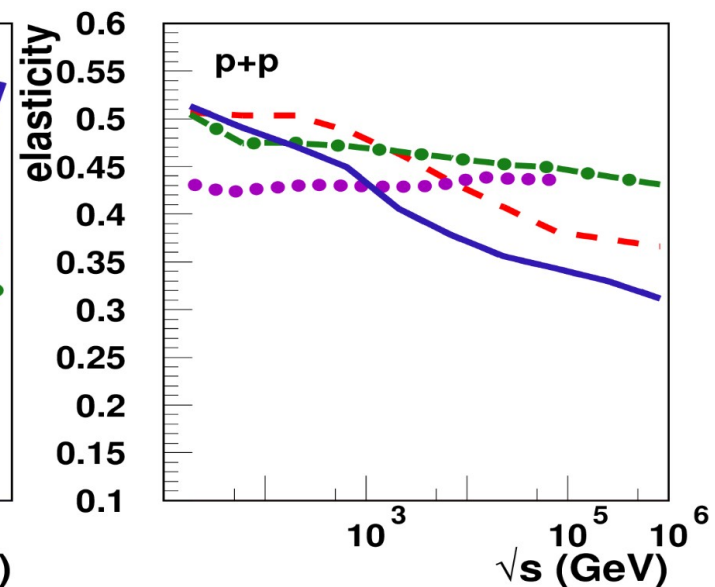
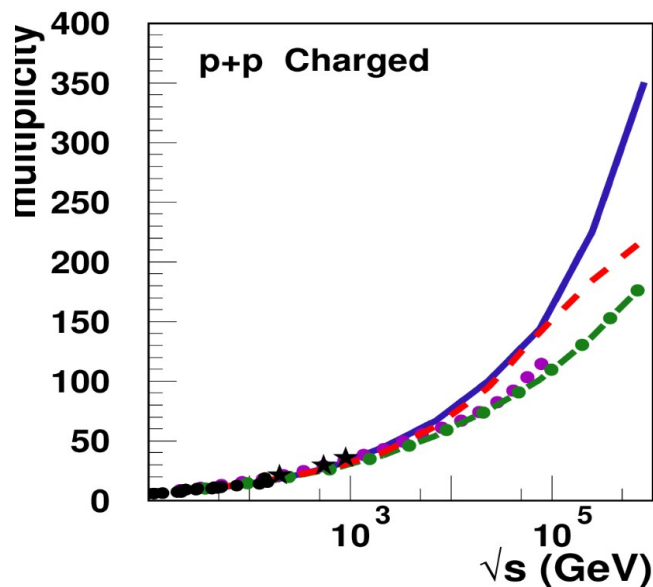
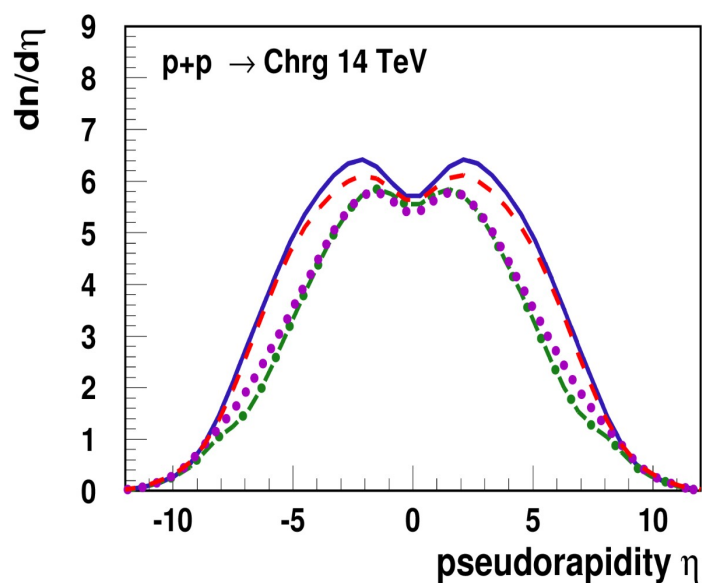


J. Matthews, Astropart.Phys. 22
(2005) 387-397

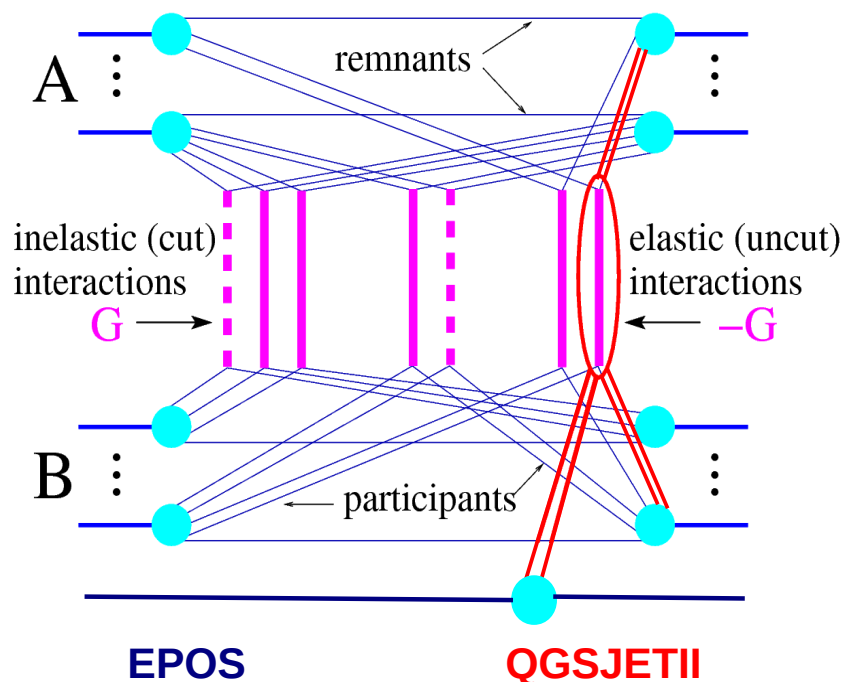
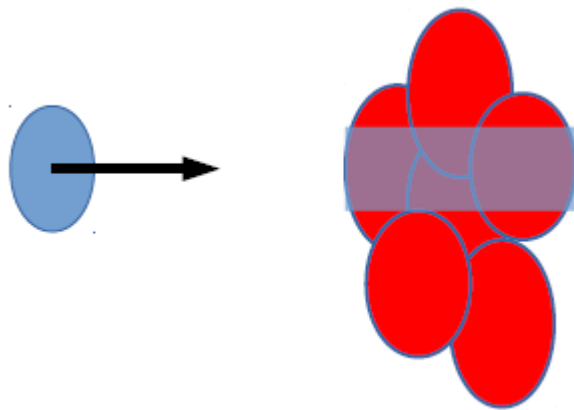
Ultra-High Energy Hadronic Model Predictions p-Air



Ultra-High Energy Hadronic Model Predictions p-Air



Nuclear Interactions



● Sibyll

→ Glauber for pA

■ with inelastic screening for diffraction in new Sibyll 2.3 (only nuclear effect)

→ superposition model for AA ($A \times pA$)

● QGSJETII

→ Pomeron configuration based on A projectiles and A targets

→ Nuclear effect due to multi-leg Pomerons

● EPOS

→ Pomeron configuration based on A projectiles and A targets

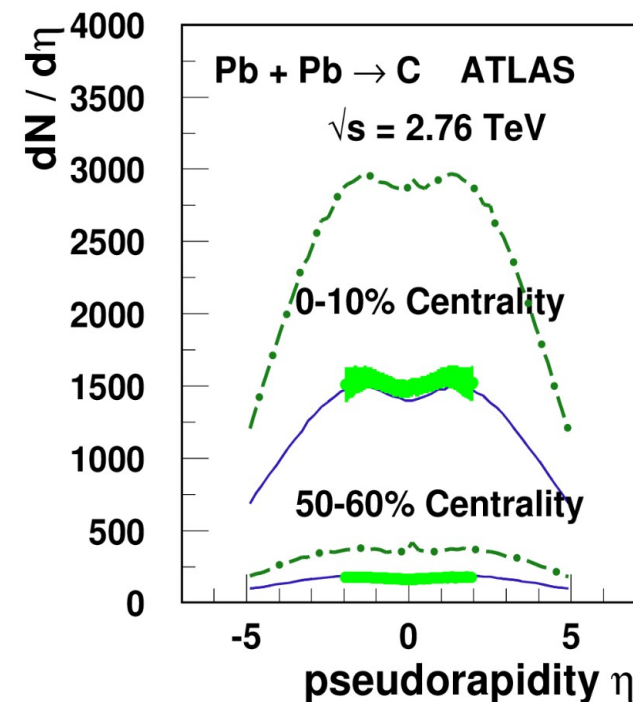
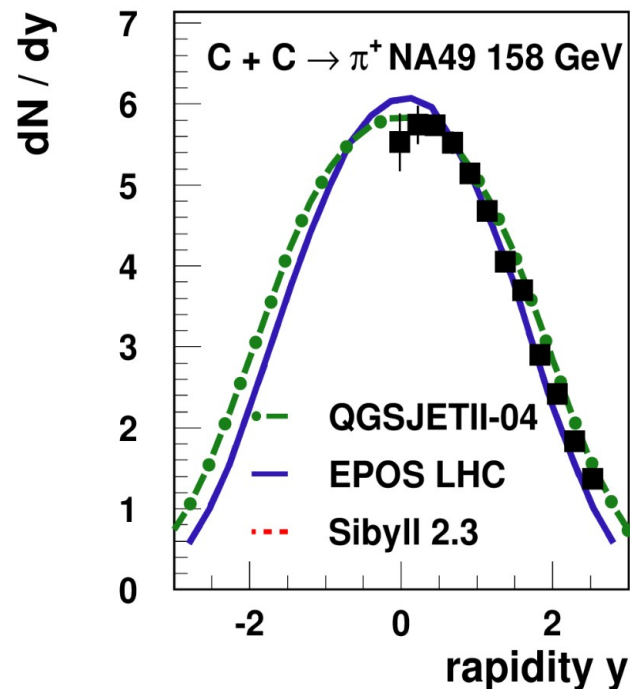
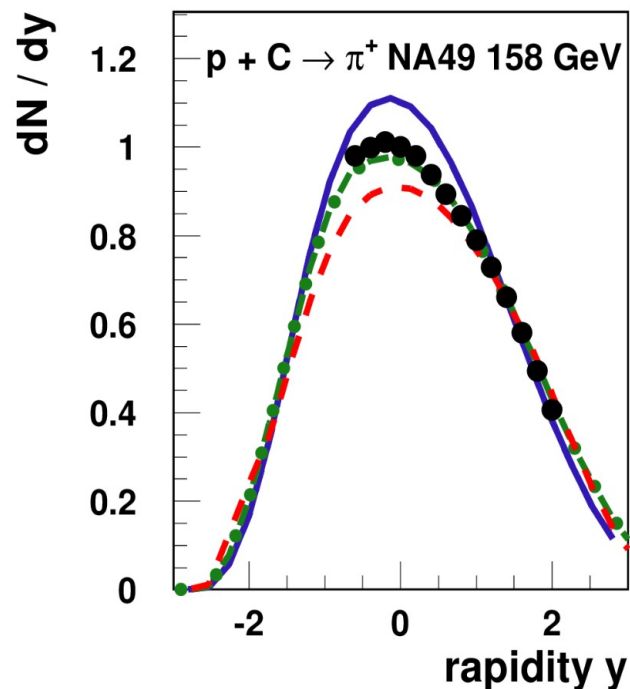
→ screening corrections depend on nuclei

→ final state interactions (core-corona approach and collective hadronization with flow for core)

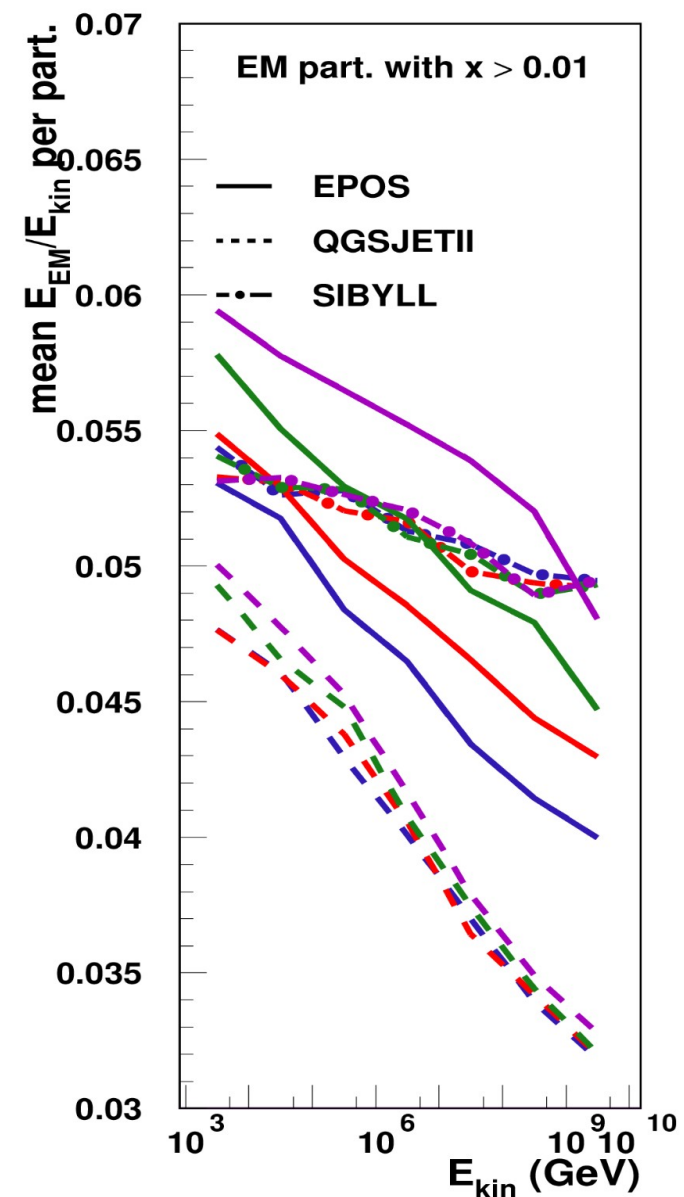
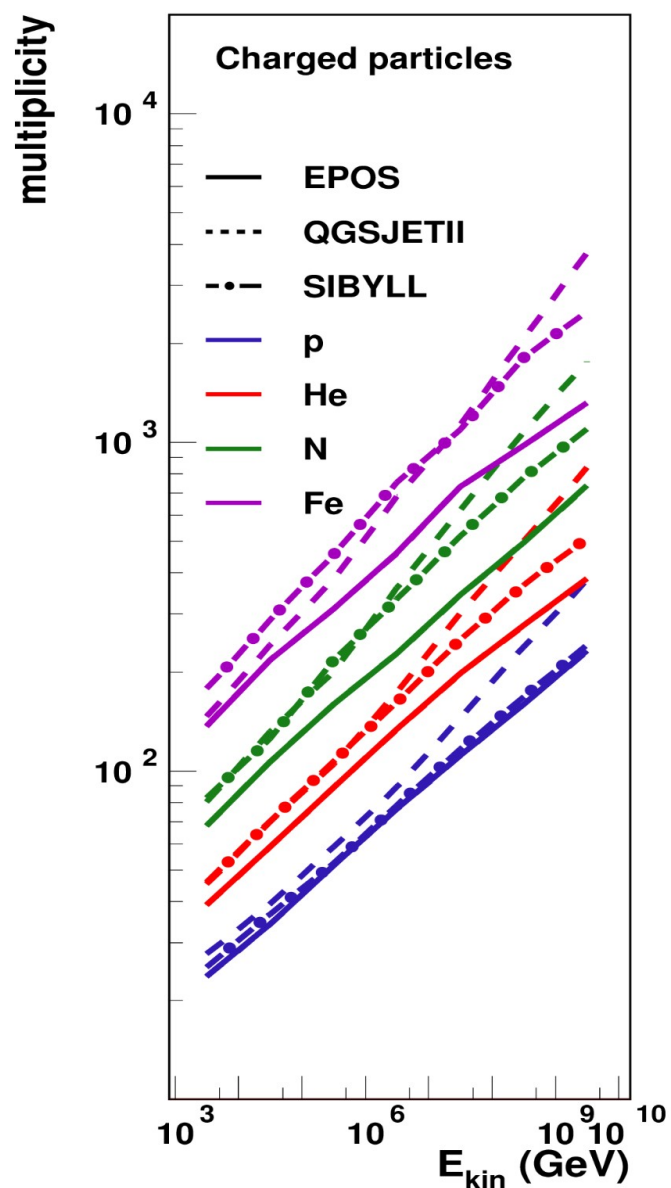
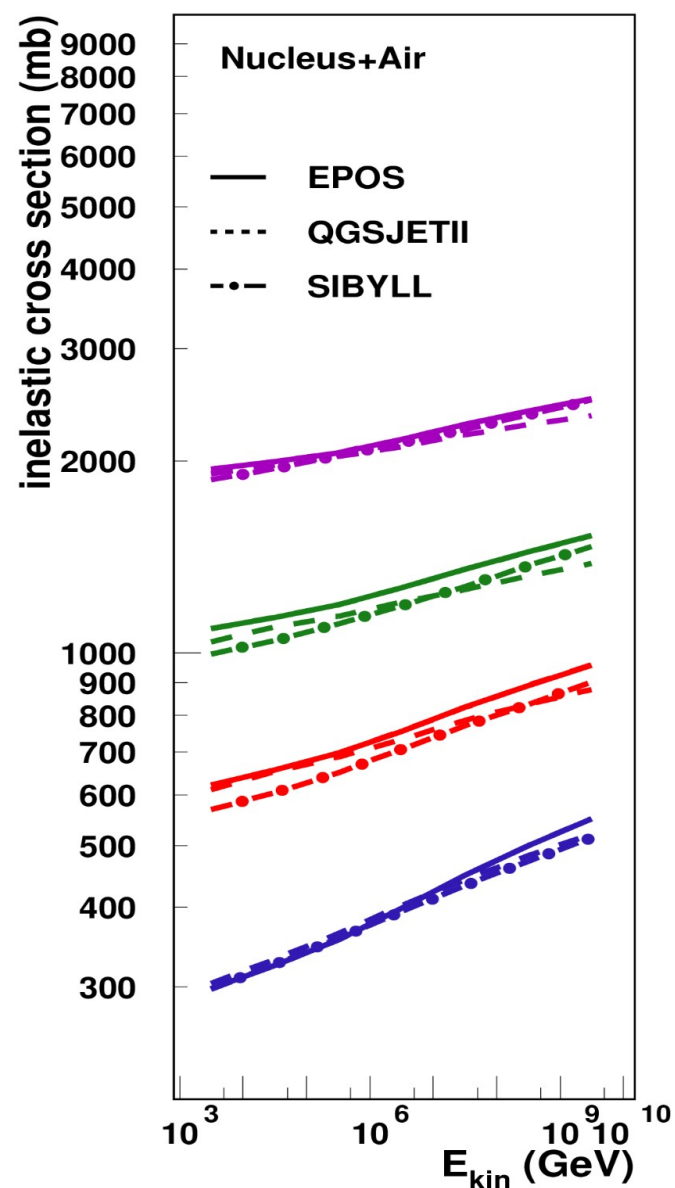
Light Ion Data

Very few data to compare with all CR models :

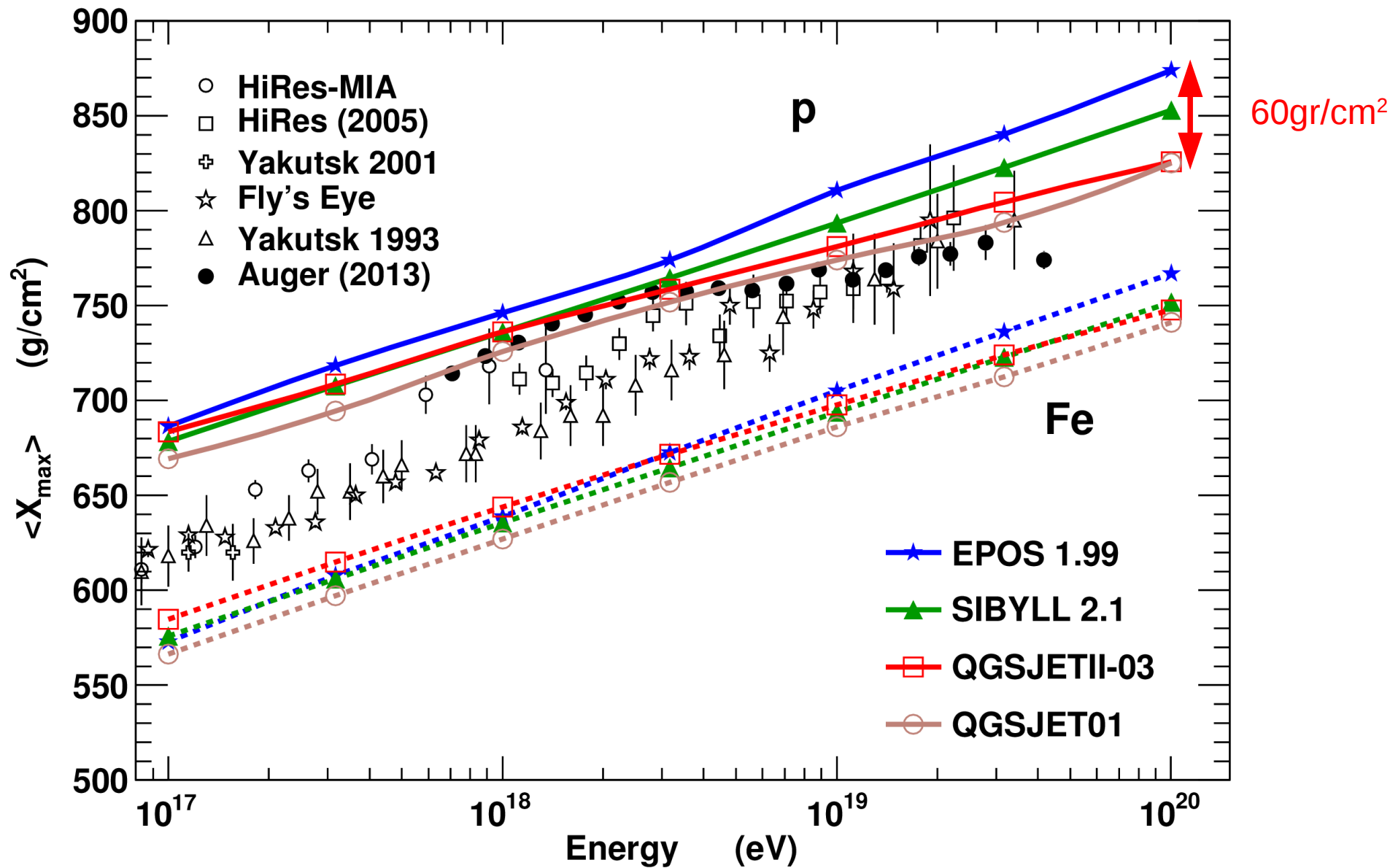
- ➔ strong limitations in Sibyll (projectile up to Fe only and target up to O !)
- ➔ no final state interactions exclude heavy nuclei for QGSJETII
- ➔ no light ion data at high energy



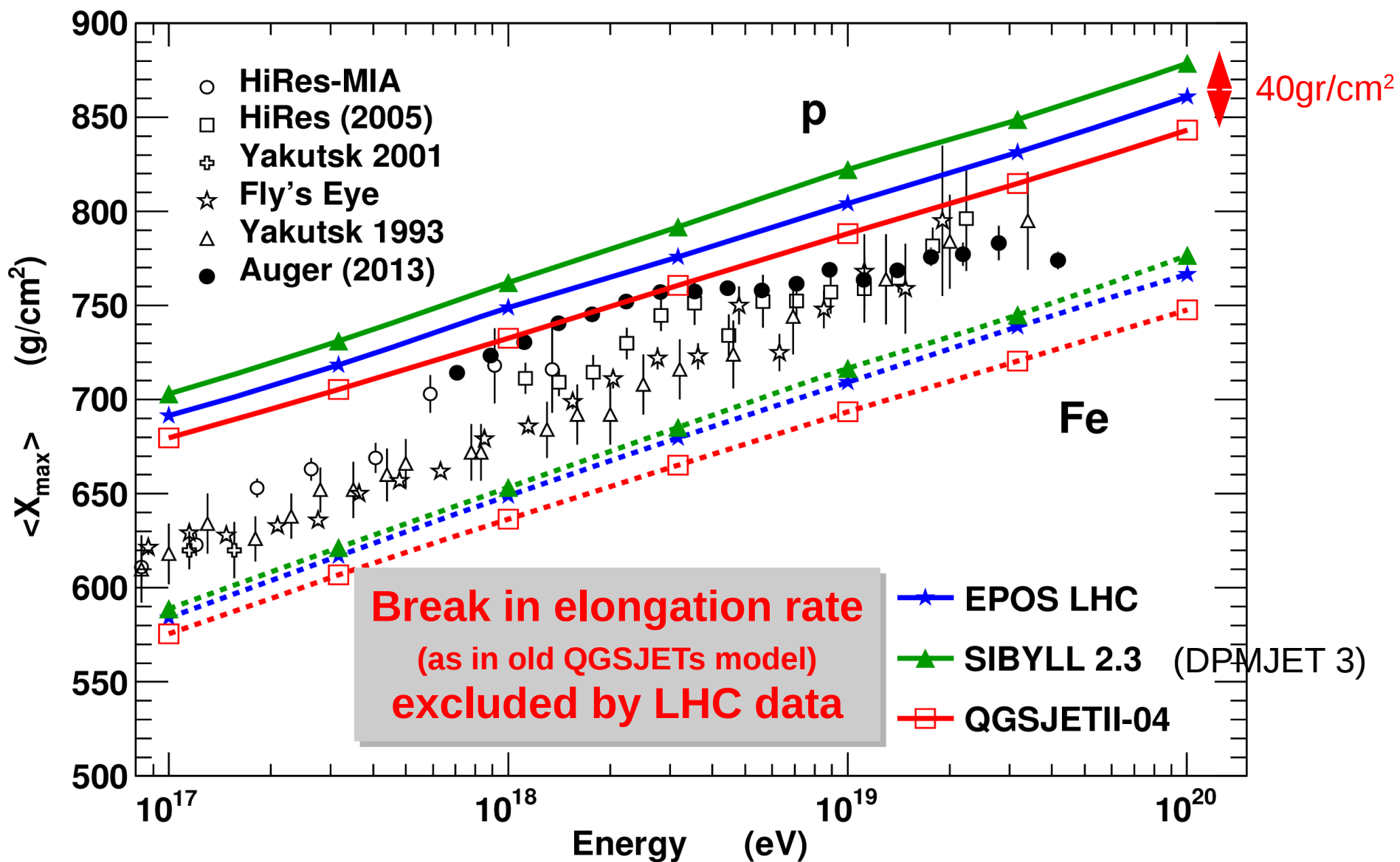
Ultra-High Energy Hadronic Model Predictions A-Air



EAS with Old CR Models : X_{\max}



EAS with Re-tuned CR Models : X_{\max}



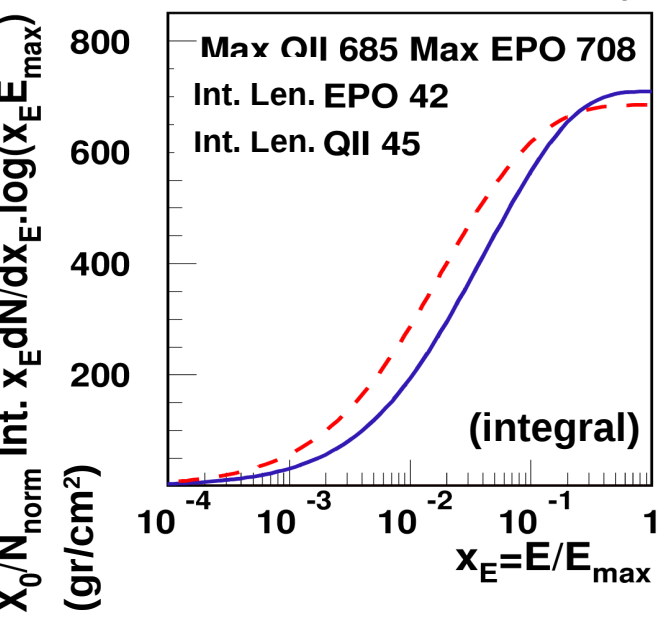
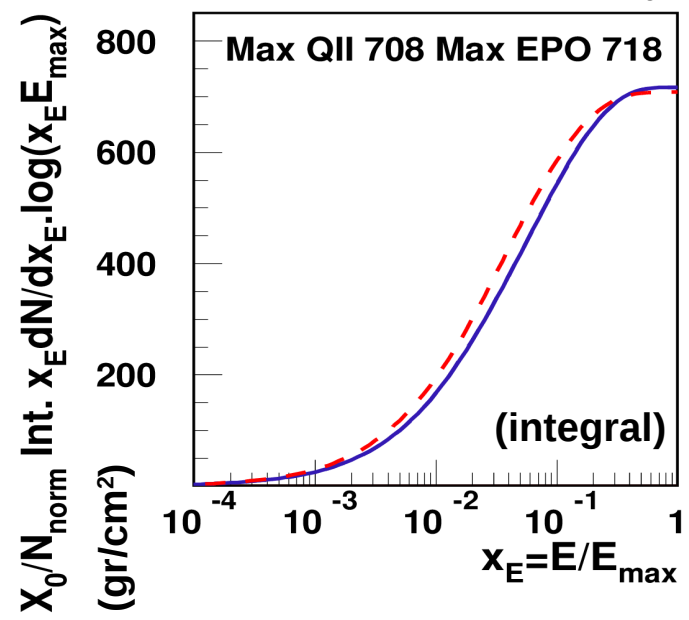
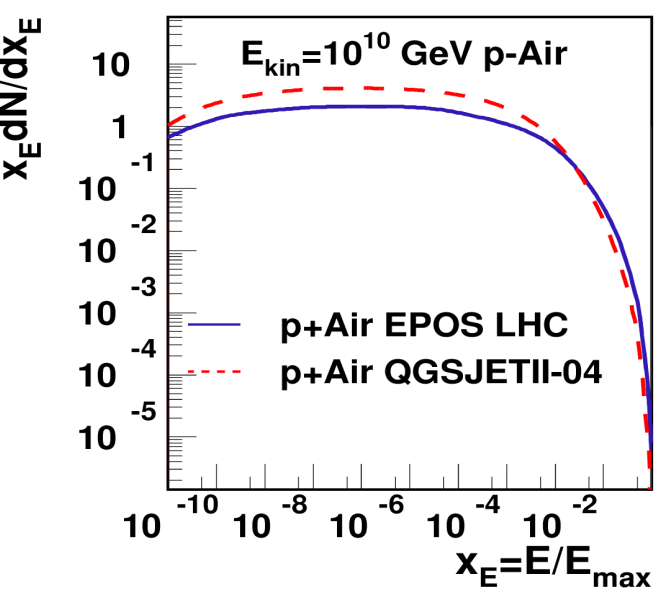
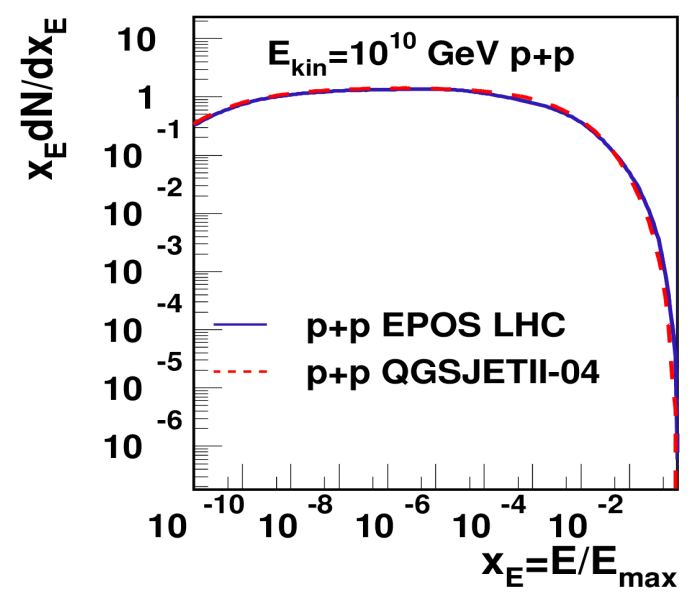
Photon Energy Spectra

● **Uncertainties in X_{max}**

- ➔ photon energy spectra
- ➔ elasticity (for 2^d interaction)
- ➔ **extrapolation to nuclear interactions**

● **Use directly energy spectra from first interaction**

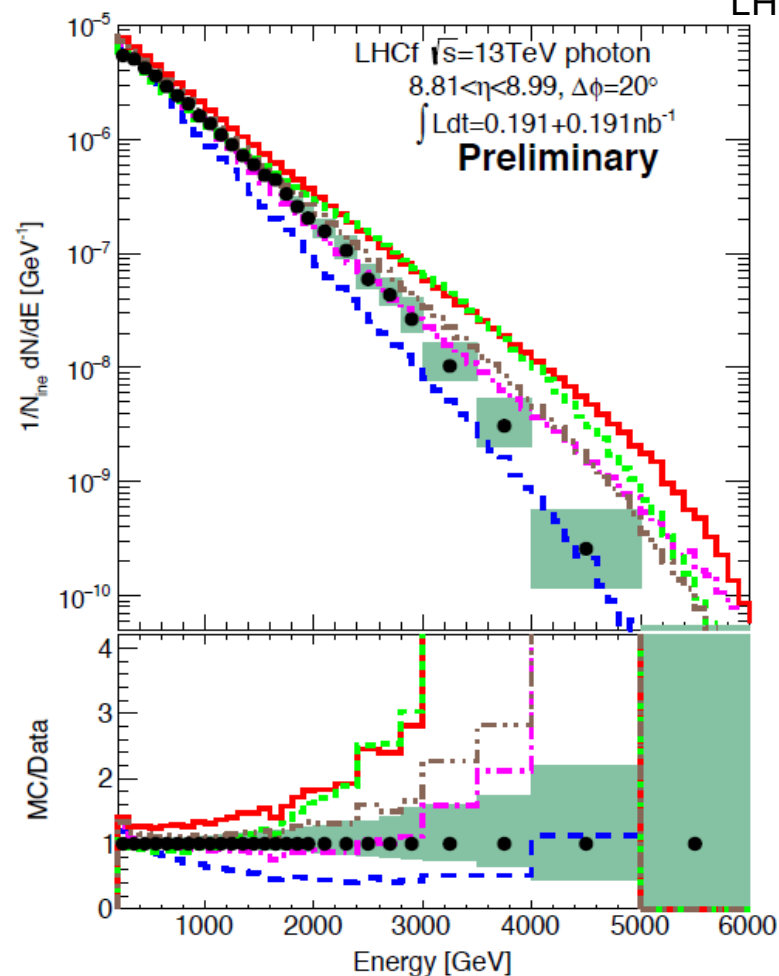
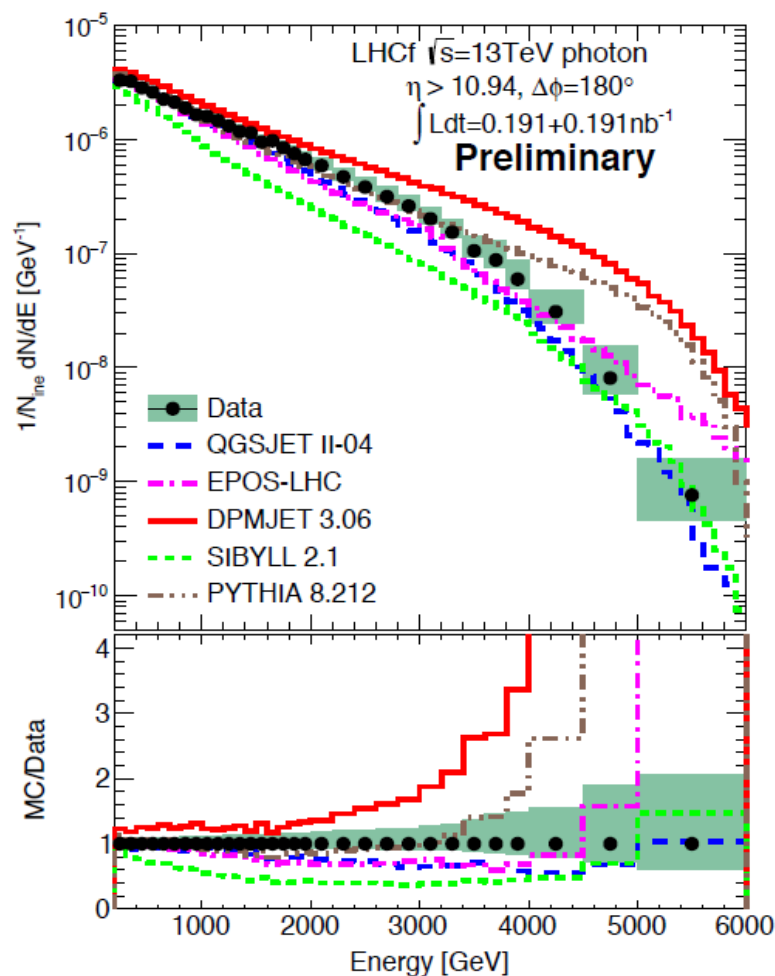
- ➔ which energy is important ?



Comparison with LHCf

- ➔ LHCf favor not too soft photon spectra (EPOS LHC, SIBYLL 2.3) : deep X_{\max}
- ➔ No model compatible with all LHCf measurements : room for improvements !
- ➔ Can p-Pb data be used to mimic light ion (Air) interactions ?

T.Sako for the
LHCf collaboration

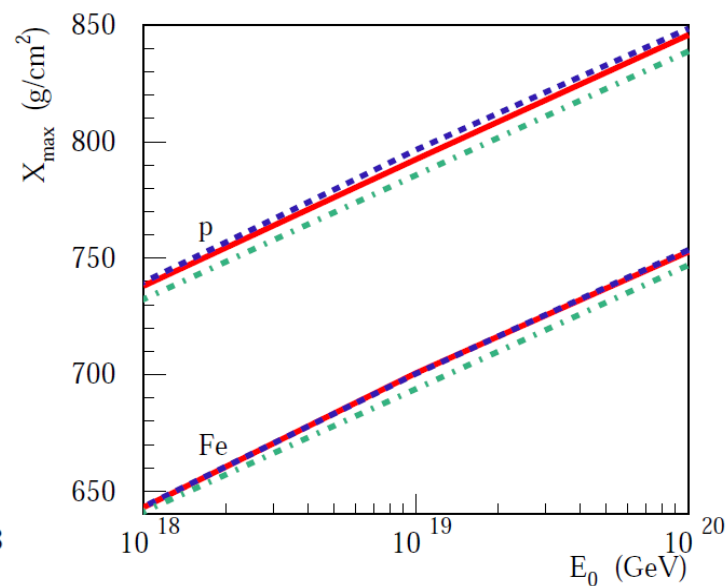
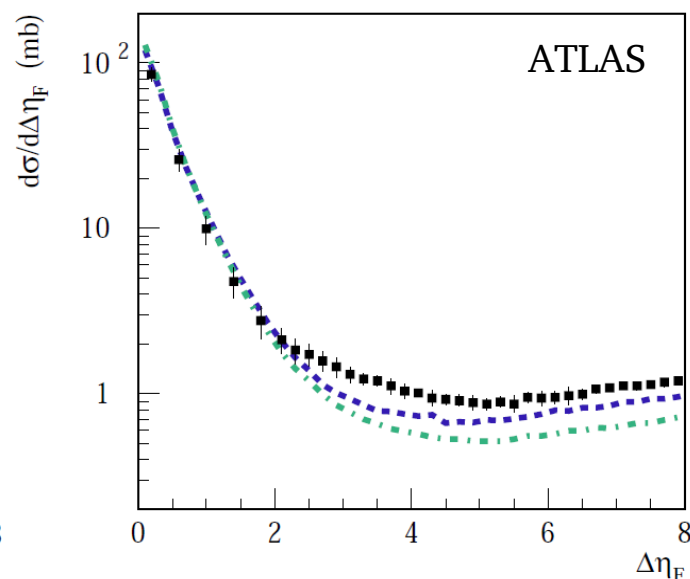
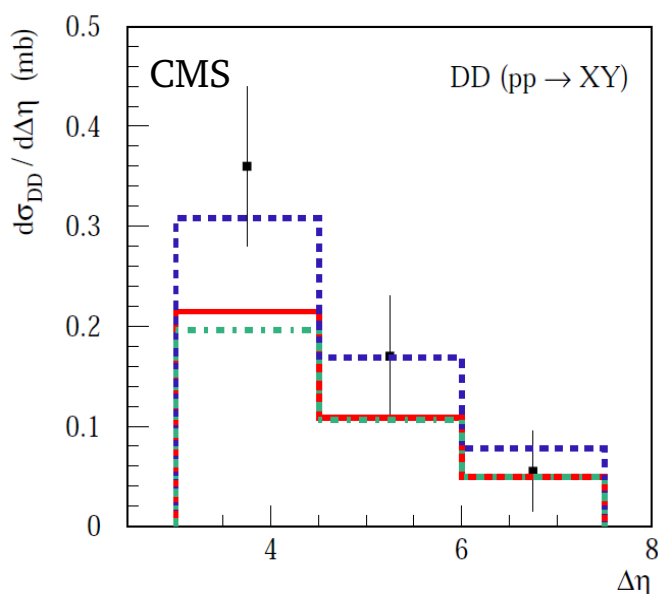


Diffraction measurements

- TOTEM and CMS diffraction measurement not fully consistent
- Tests by S. Ostapchenko using QGSJETII-04 (PRD89 (2014) no.7, 074009)
 - ➔ SD+ option compatible with CMS
 - ➔ SD- option compatible with TOTEM

M_X range	< 3.4 GeV	$3.4 - 1100$ GeV	$3.4 - 7$ GeV	$7 - 350$ GeV	$350 - 1100$ GeV
TOTEM [13, 24]	2.62 ± 2.17	6.5 ± 1.3	$\simeq 1.8$	$\simeq 3.3$	$\simeq 1.4$
QGSJET-II-04	3.9	7.2	1.9	3.9	1.5
option SD+	3.2	8.2	1.8	4.7	1.7
option SD-	2.6	7.2	1.6	3.9	1.7

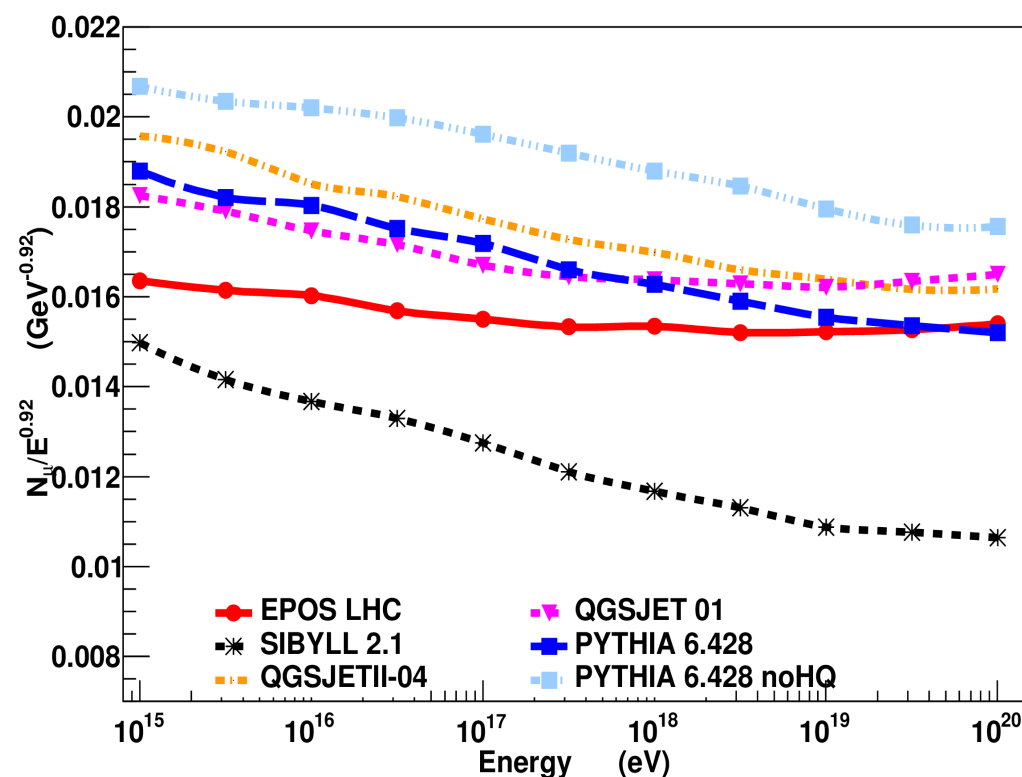
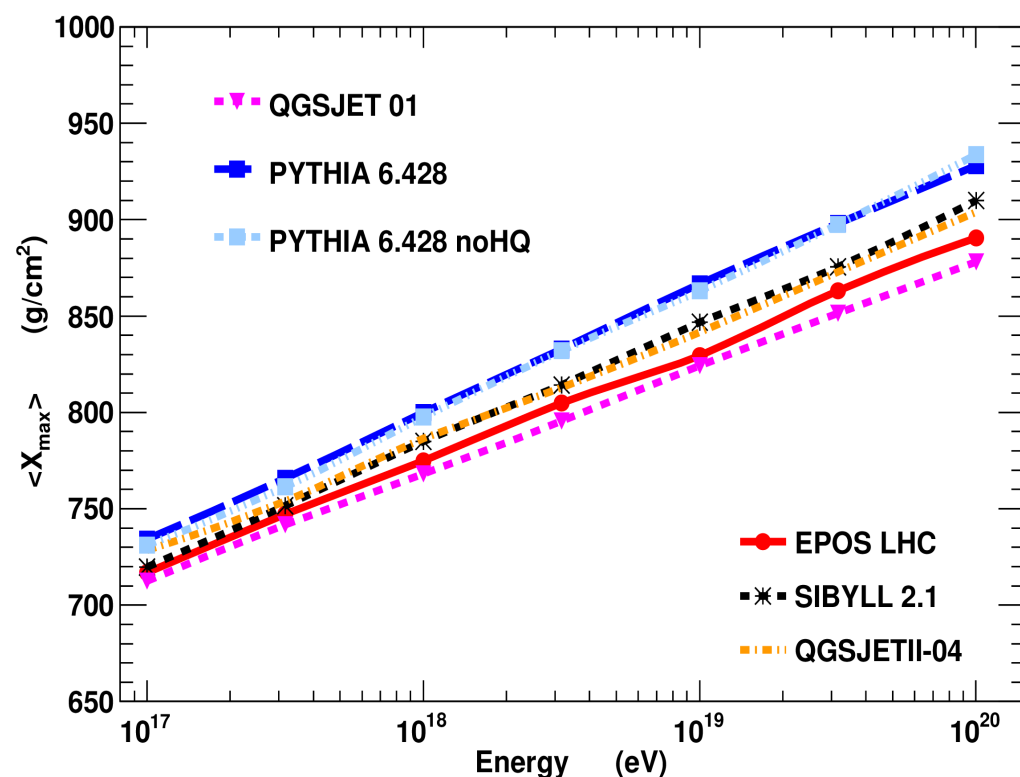
➔ difference of ~ 10 gr/cm² between the 2 options



Tests using hydrogen atmosphere

- Work done with David D'Enterria (CERN) and Sun Guan hao
 - ➔ test of Pythia event generator
- Modified air shower simulations with air target replaced by hydrogen
 - ➔ for interactions only (no change in density) : no nuclear effect
 - ➔ Without nuclear effect relative order of the models is changed

➔ Nuclear effects are the main remaining source of uncertainty

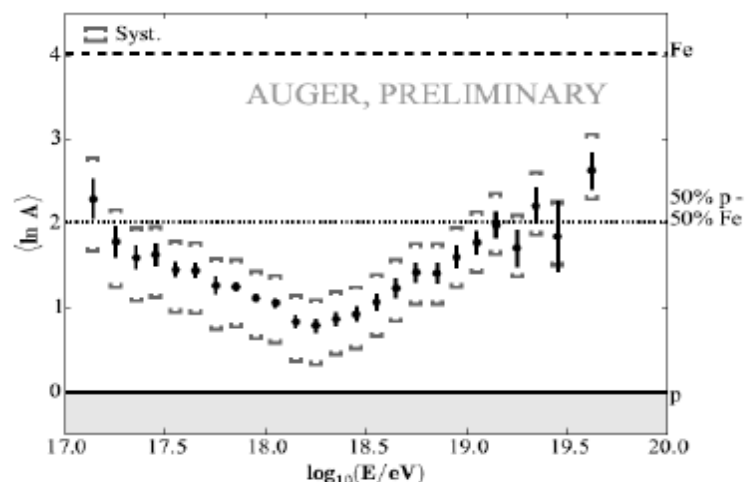


Model Consistency using Electromagnetic Component

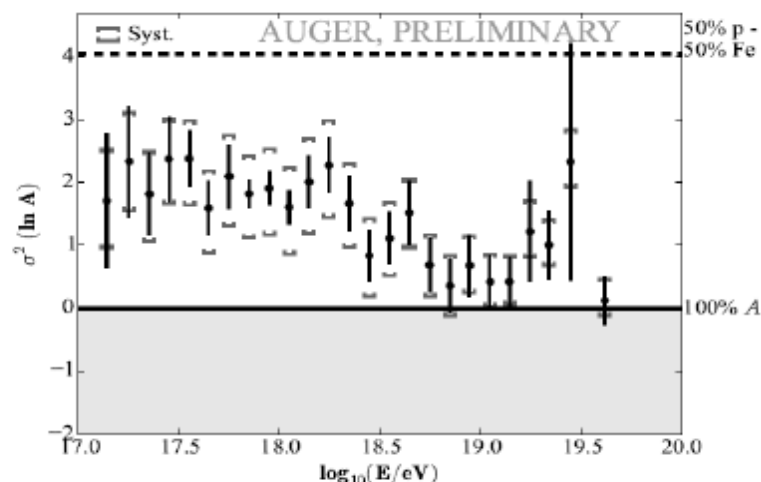
Study by Pierre Auger Collaboration

→ std deviation of $\ln A$ allows to test model consistency.

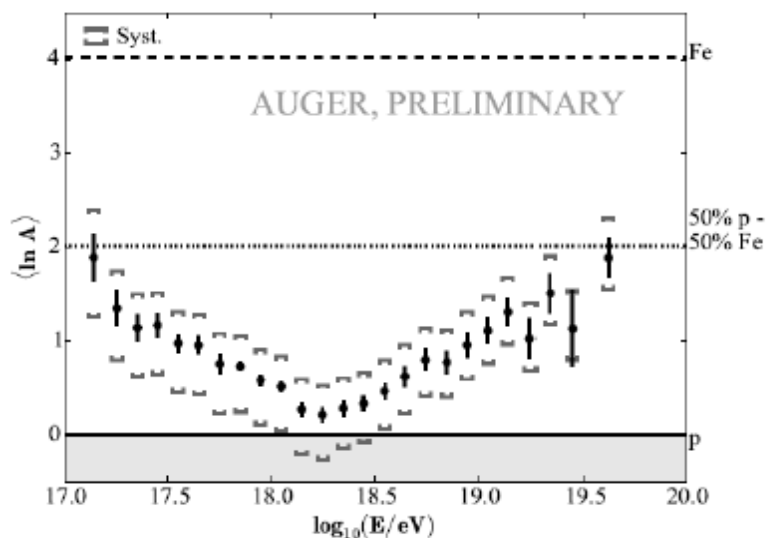
EPOS-LHC (Mean of $\ln A$)



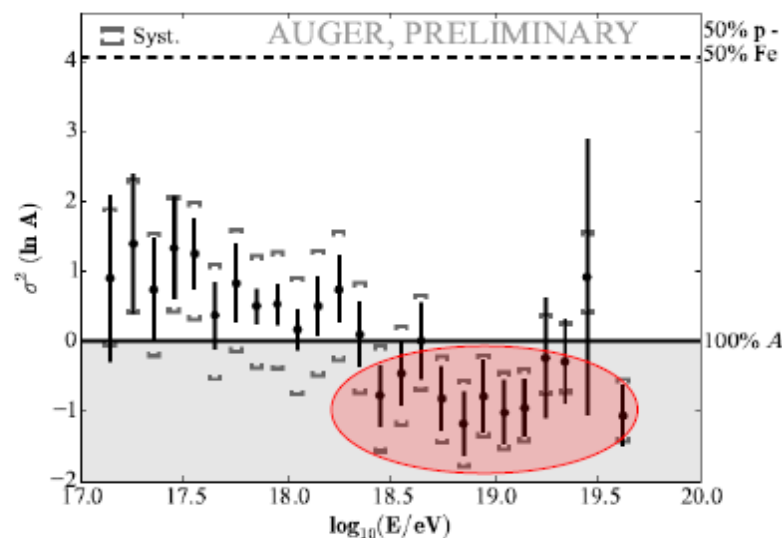
EPOS-LHC (Std. Deviation of $\ln A$)



QGSJetII-04 (Mean of $\ln A$)



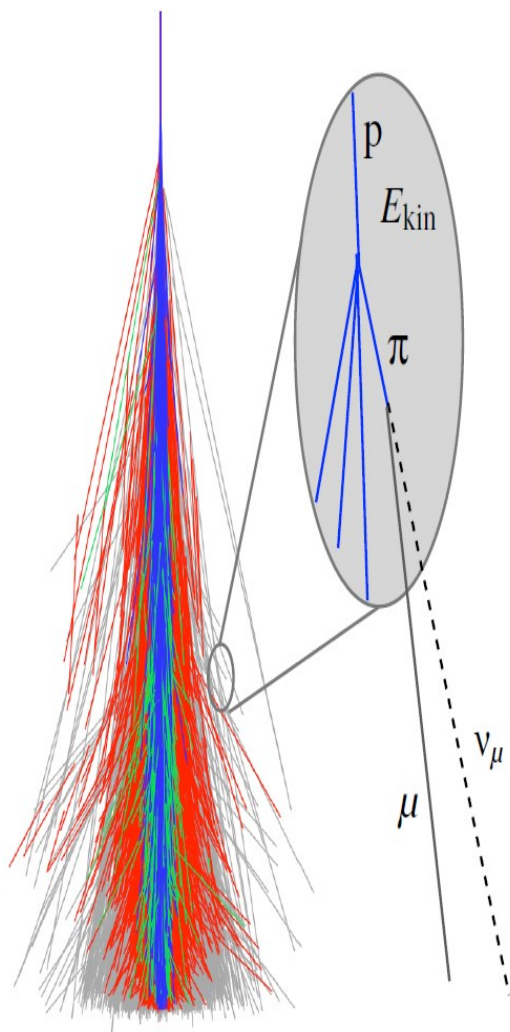
QGSJetII-04 (Std. Deviation of $\ln A$)



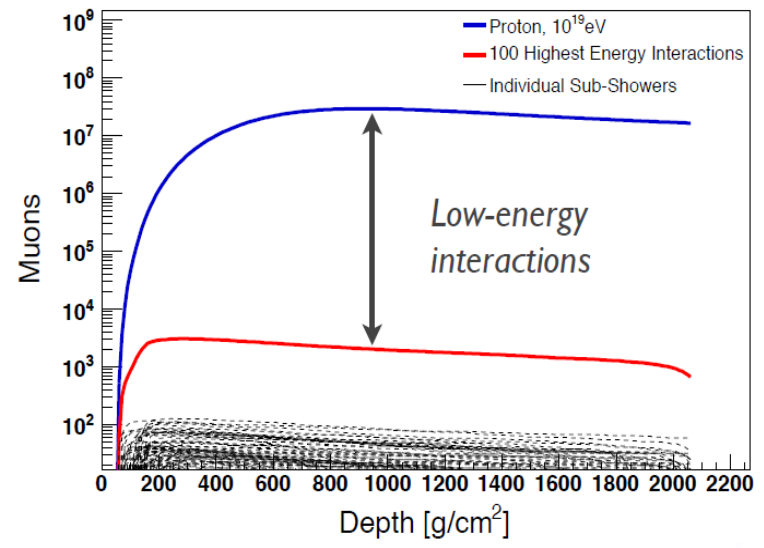
tensions if $\langle X_{\max} \rangle$
too small

QGSJETII-04 is a
lower limit for X_{\max}

Muon production by low energy interactions

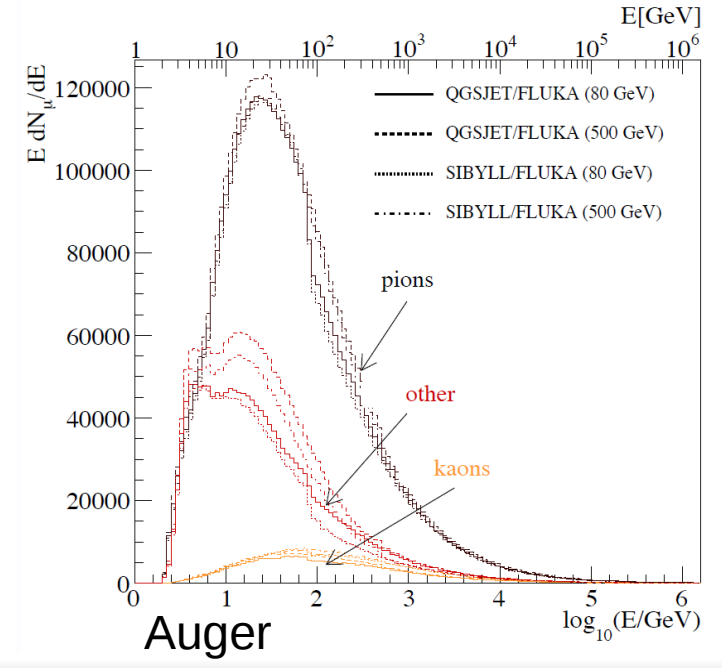
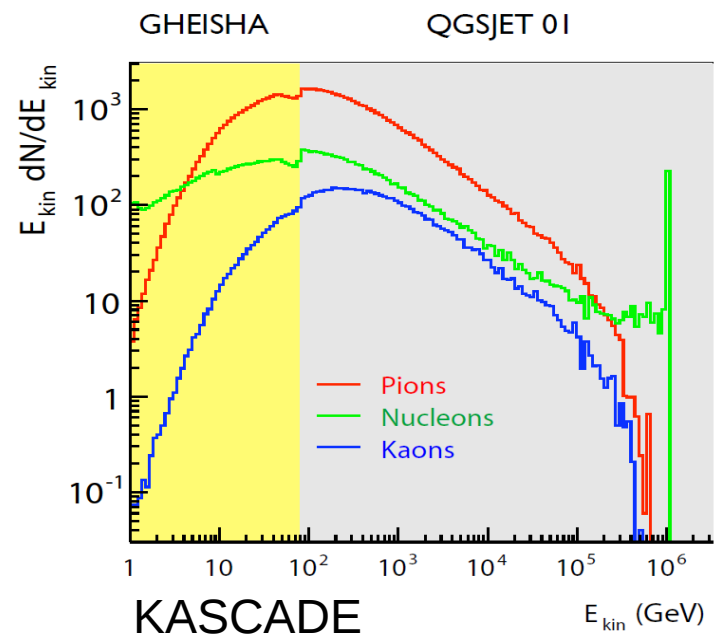


Muons



Energy of Last generation:
 ~ 100 GeV for KASCADE
 ~ 30 GeV for Auger

N_{mu} generated by all
 (low energy)
 interactions :
 pion-Air

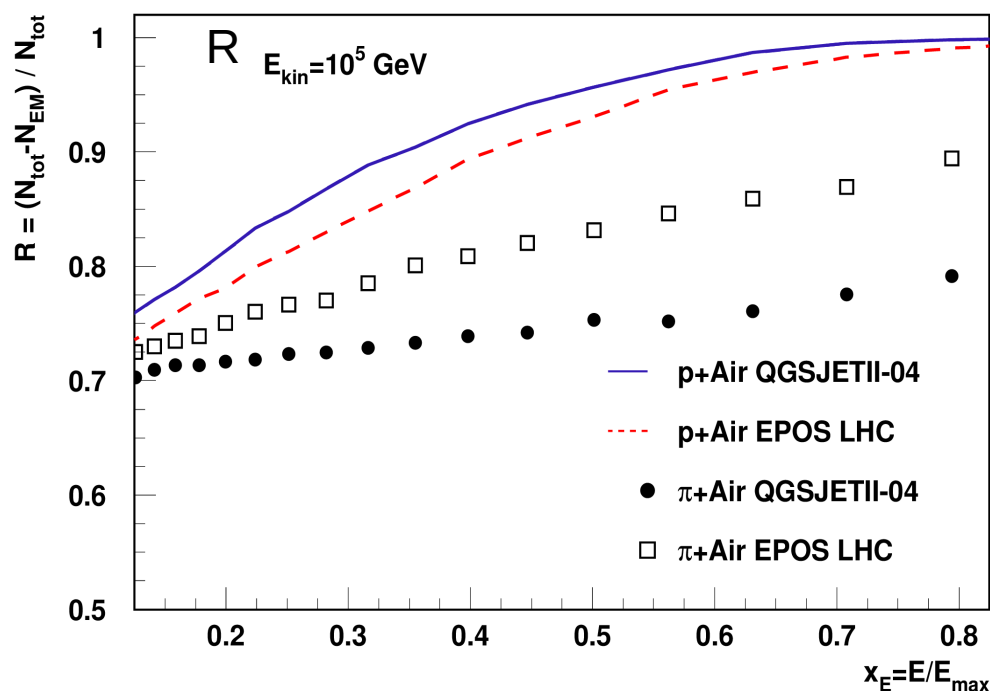


Muon Number

From Heitler

$$N_{\mu} = \left(\frac{E_0}{E_{dec}} \right)^{\alpha}, \quad \alpha = \frac{\ln N_{had}}{\ln (N_{had} + N_{em})}$$

➔ In real shower, not only pions : Kaons, (anti)Baryons and resonances



$$\alpha = \frac{\ln (N_{had})}{\ln (N_{tot})} = 1 + \frac{\ln (R)}{\ln (N_{tot})}$$

$$R = \frac{N_{had}}{N_{tot}} \approx \frac{N_{\pi^{ch}} + N_B}{N_{\pi^{ch}} + N_B + N_{\pi^0}} < 1$$

Very important :
in (a)Baryon-Air interactions, no
leading neutral pion !
 $R \sim 1$

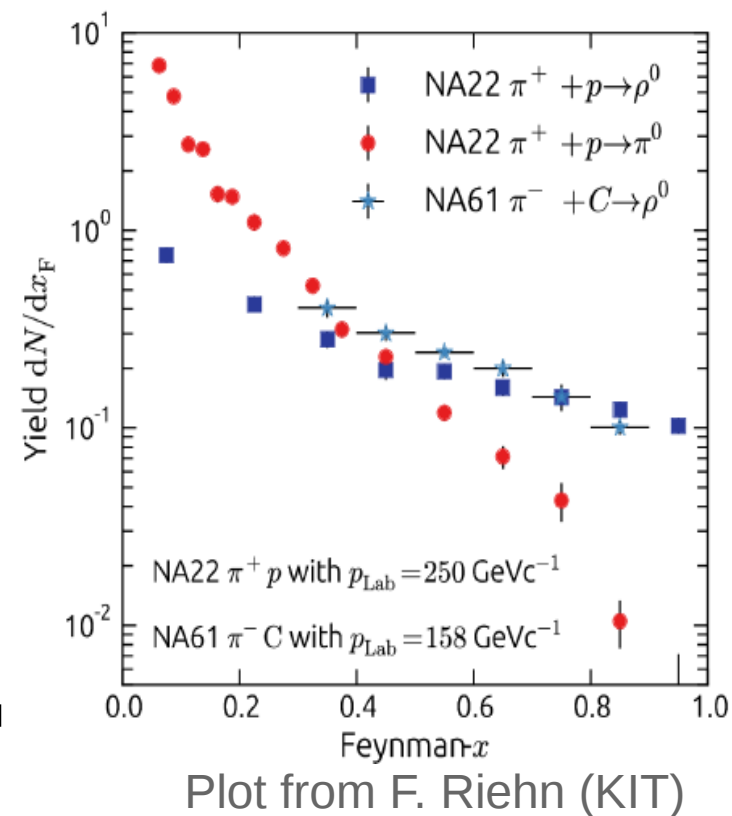
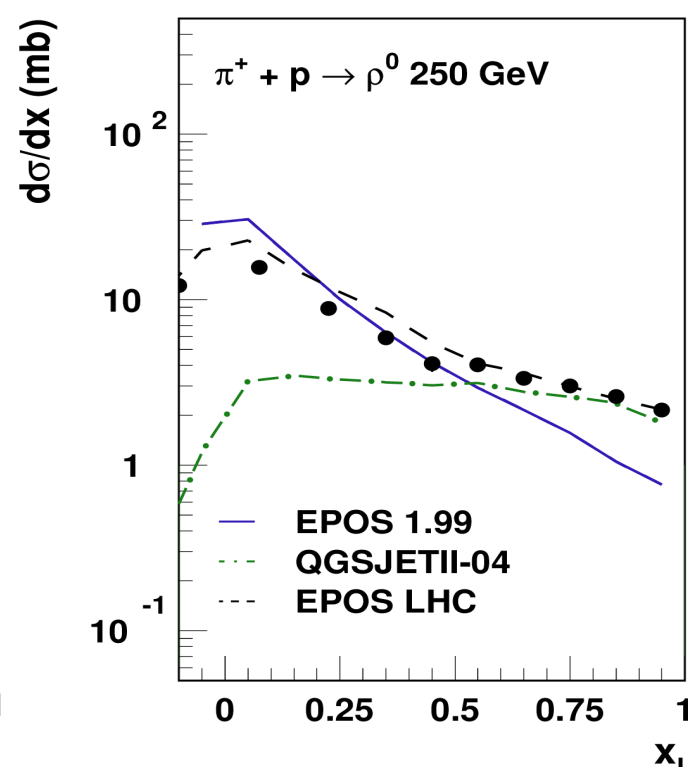
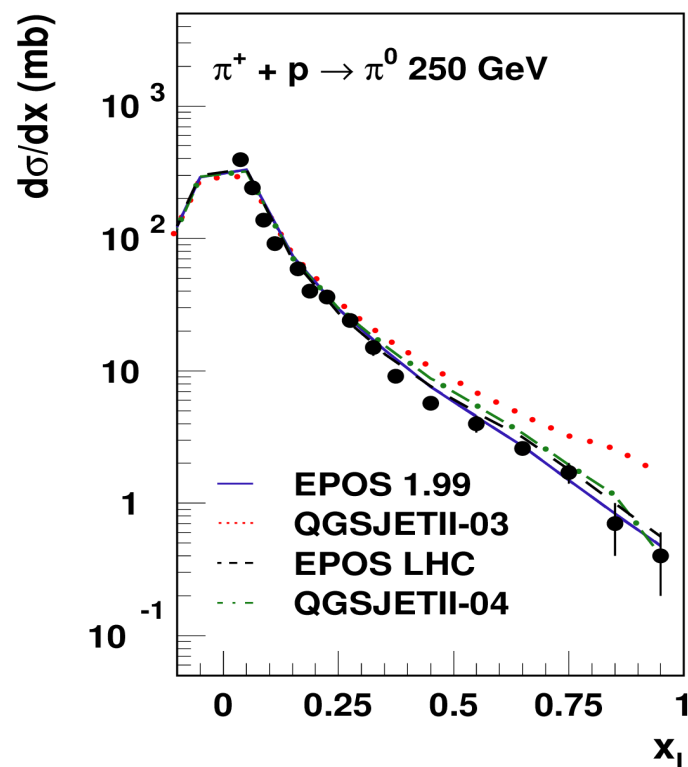
R depends on the number of (anti)B and ρ^0 in p- or π -Air interactions

More fast (anti)baryons or ρ^0 or larger $N_{tot} = \alpha \rightarrow 1$ = more muons

Pion Leading Particle Effect

Rho meson production added in QGSJETII (and Sibyll 2.3) to take into account leading particle effect in pion-Air interaction

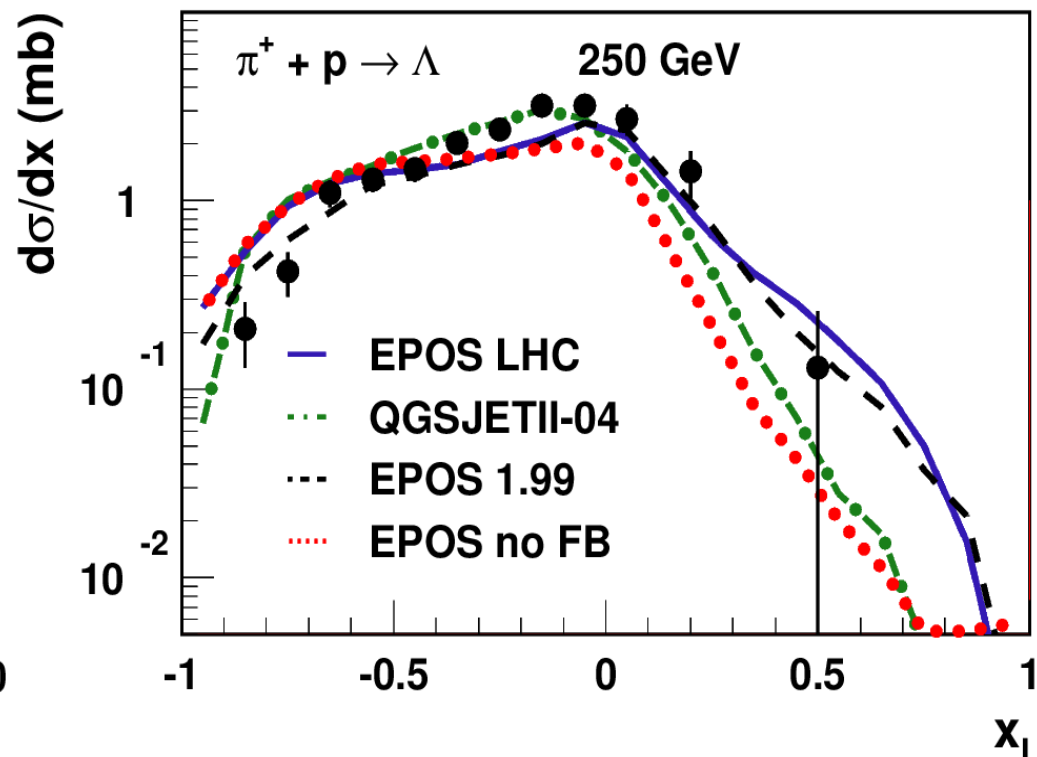
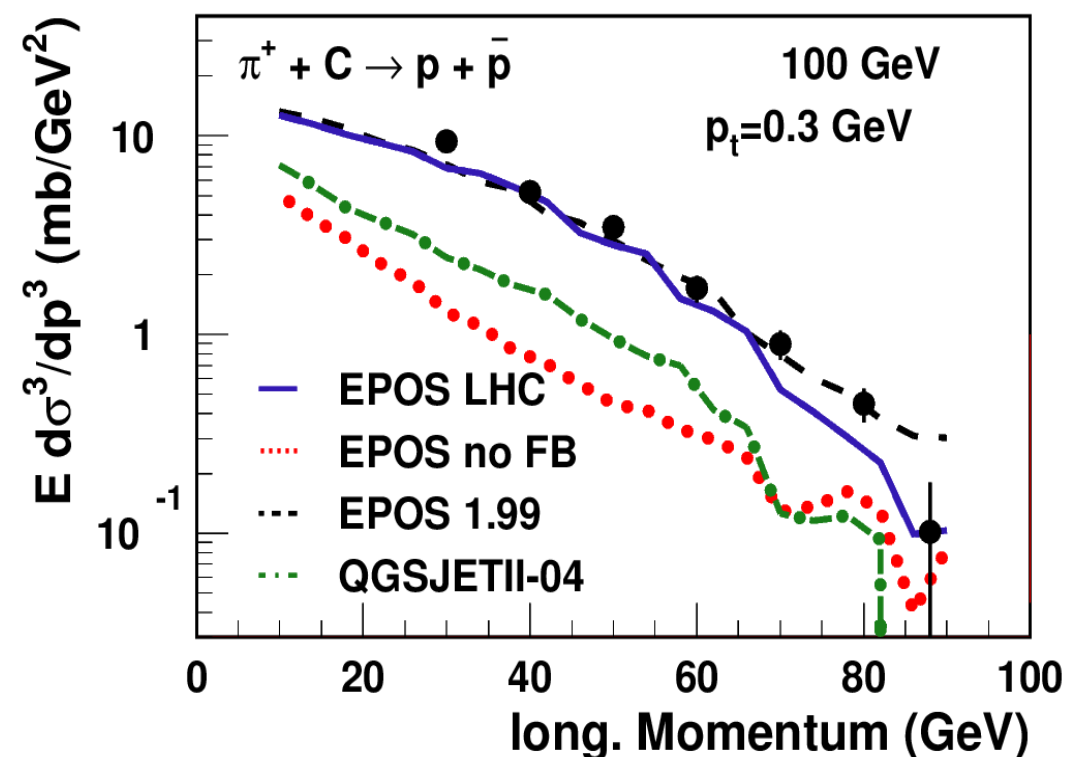
- ➔ same effect as baryon production : forward π^0 replaced by charged pions (reduced leading π^0)
- ➔ increase muon production
- ➔ higher minimum muon energy (less generations) compared to baryons



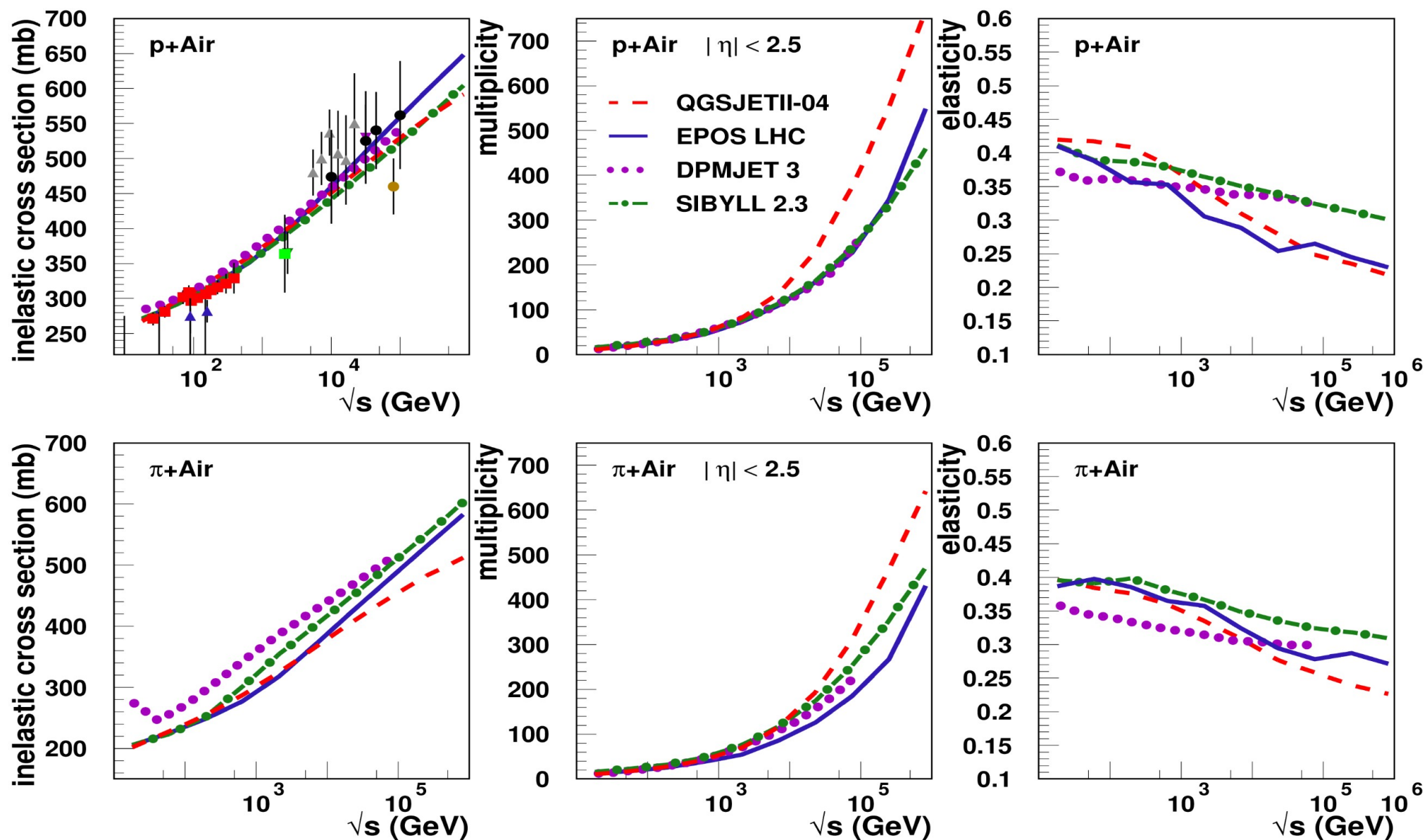
Baryons in Pion-Carbon

Very few data for baryon production from meson projectile, but for all :

- ➔ strong baryon acceleration (probability $\sim 20\%$ per string end)
- ➔ proton/antiproton asymmetry (valence quark effect)
- ➔ target mass dependence

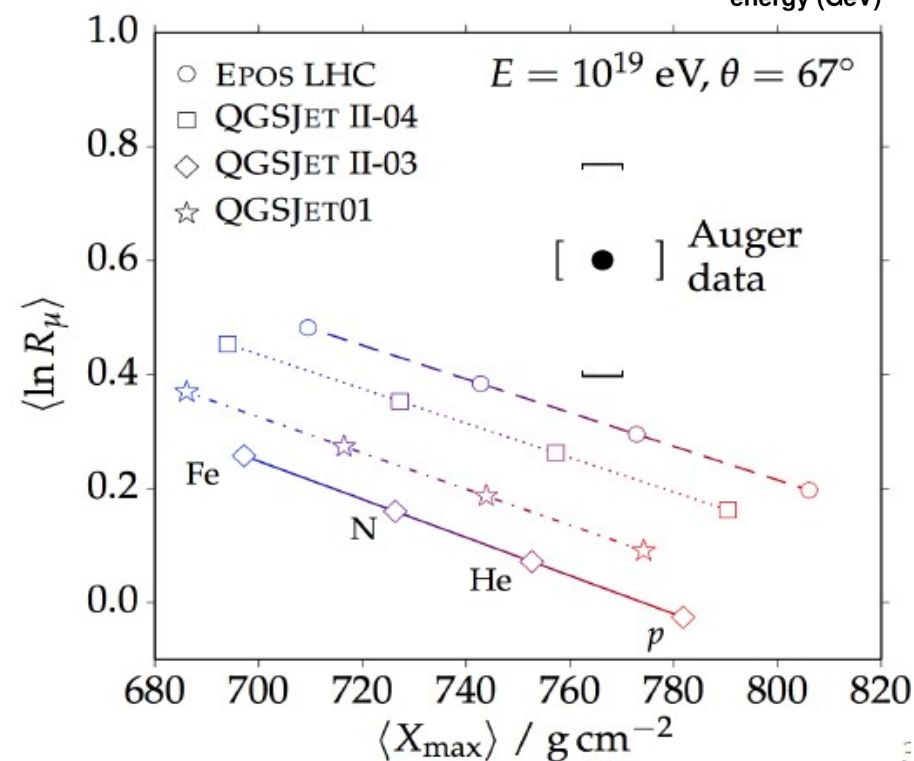
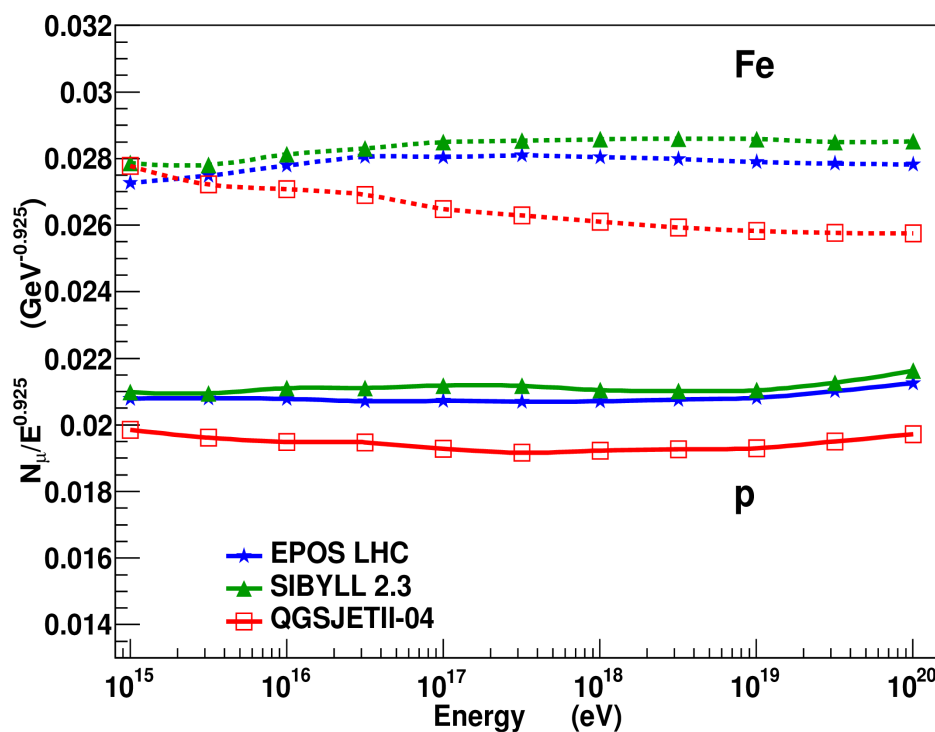
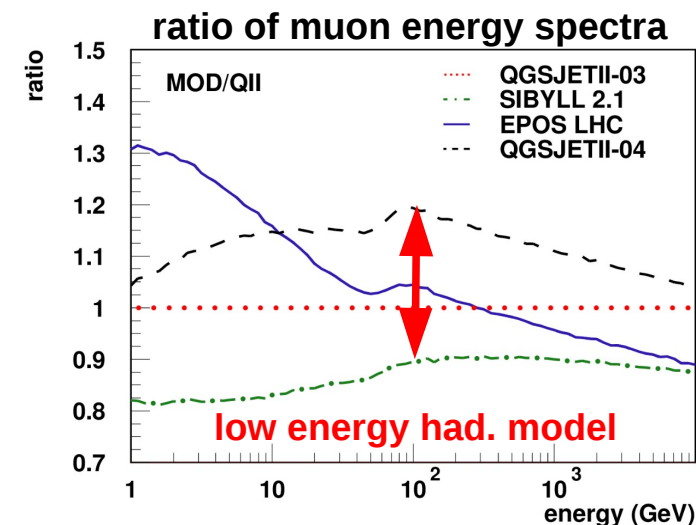


Ultra-High Energy Hadronic Model Predictions π -Air



Muons at Ground

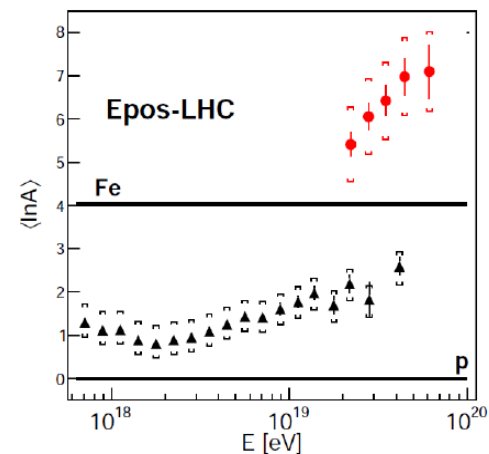
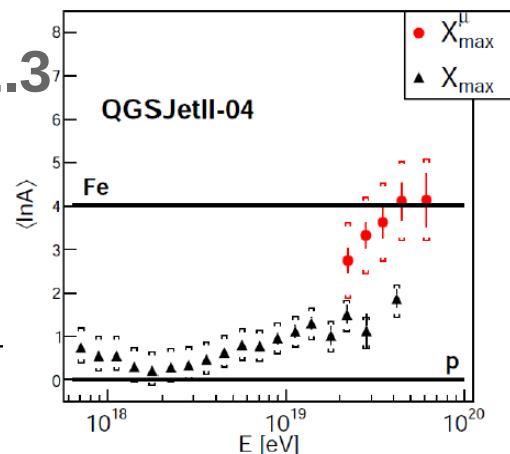
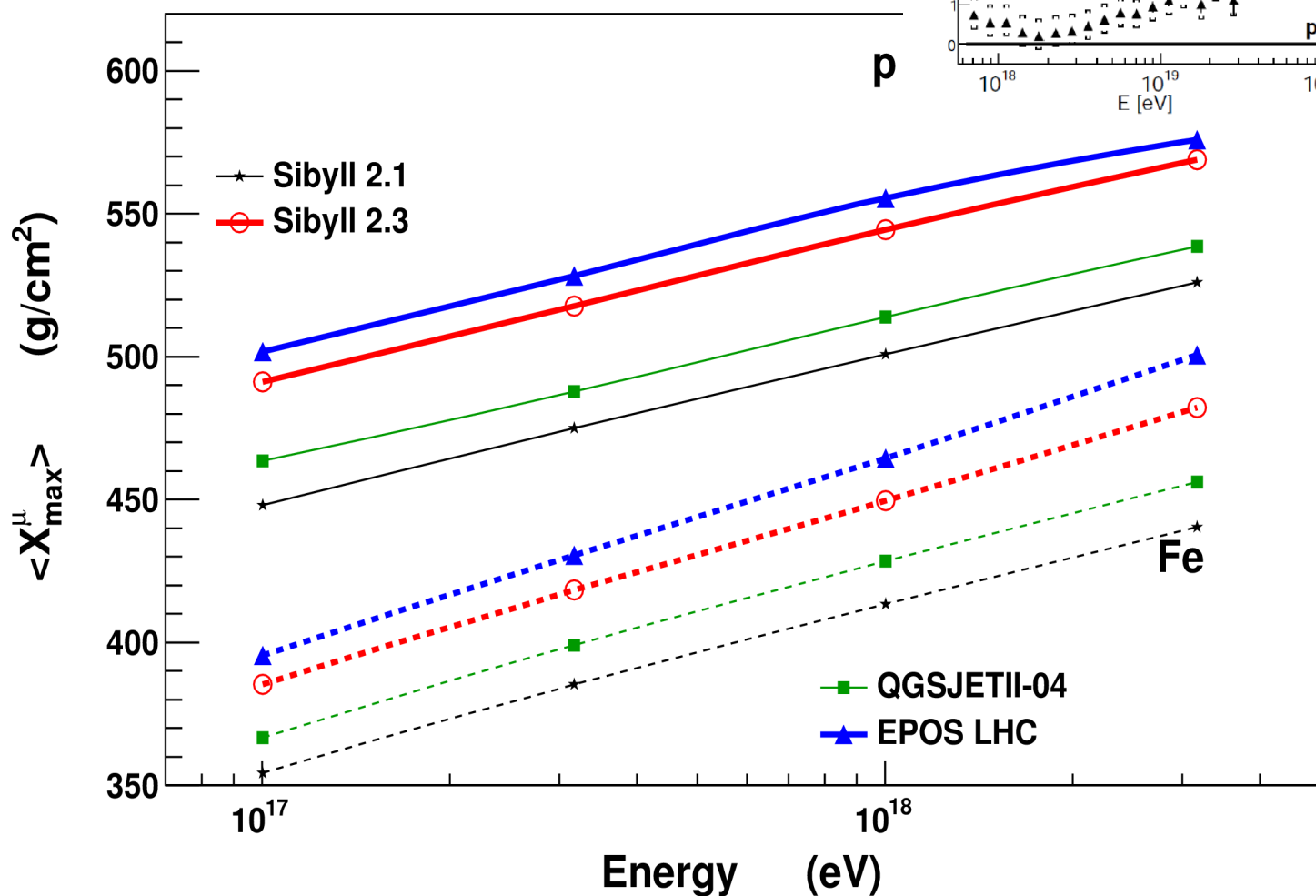
- ➔ Muon production depends on all int. energies
- ➔ Muon production dominated by pion interactions (LHC indirectly important)
- ➔ Resonance and baryon production important
- ➔ Post-LHC Models ~ agrees on numbers but with different production height and spectra



Muon Production Depth

Same for EPOS LHC and SIBYLL 2.3

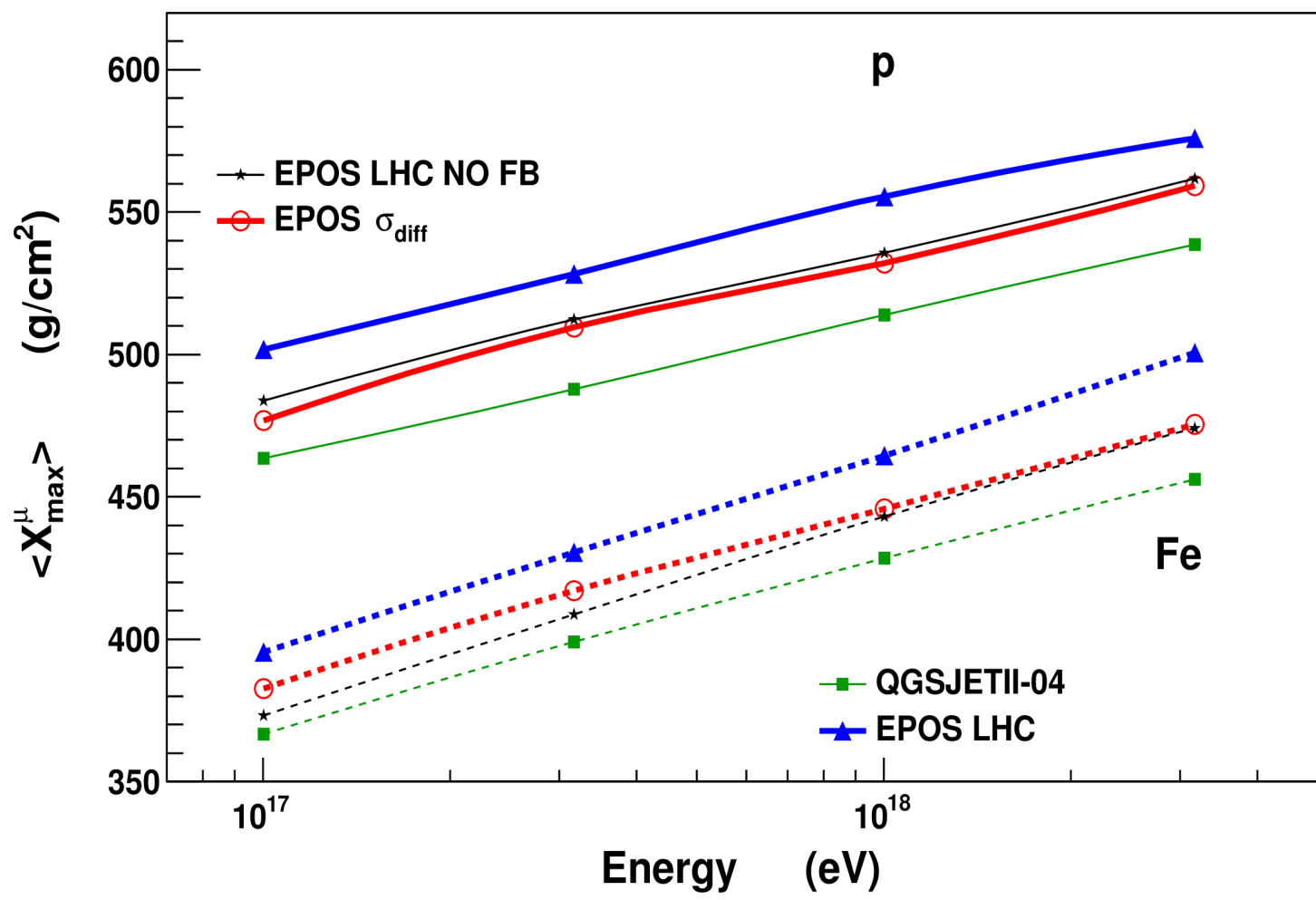
- ➔ low pion-air elasticity: deeper X_{\max}^{μ}
- ➔ more forward baryons: deeper X_{\max}^{μ}



$\langle X_{\max}^{\mu} \rangle$ with modified EPOS LHC

EPOS LHC without forward baryons or more inelastic pion int.

- ➔ softer meson spectra (lower elasticity) : smaller X_{\max}^{μ}
- ➔ less forward baryons: smaller X_{\max}^{μ}



-25 g/cm² for diff
 -20 g/cm² for baryons

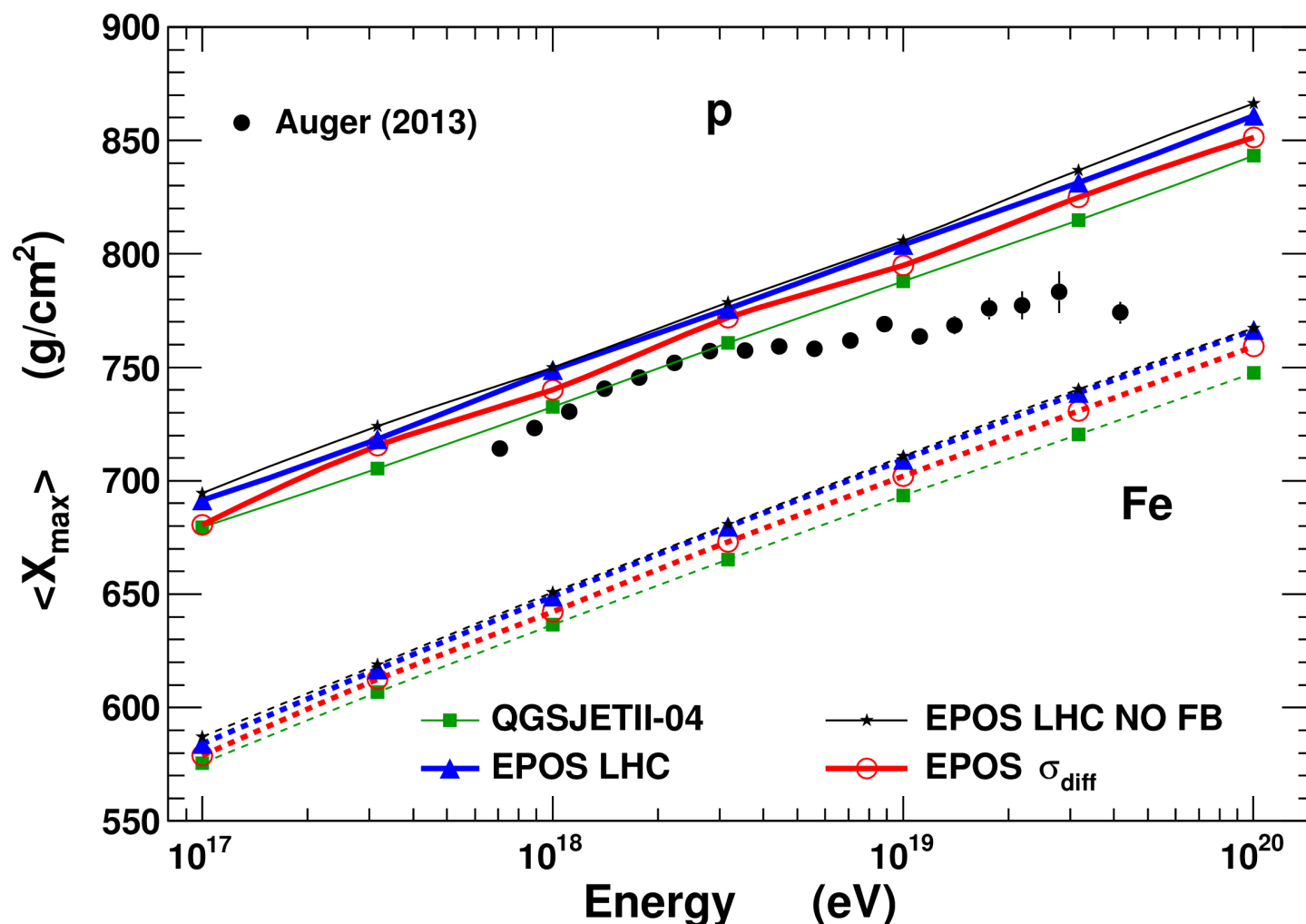
MPDs sensitive to baryon (less generation) and meson spectra in pion interactions

$\langle X_{\max} \rangle$ with Modified EPOS LHC

EPOS LHC without forward baryons or more inelastic pion int.

→ softer meson spectra: smaller X_{\max}

→ forward baryons: negligible effect



-10 g/cm² for diff

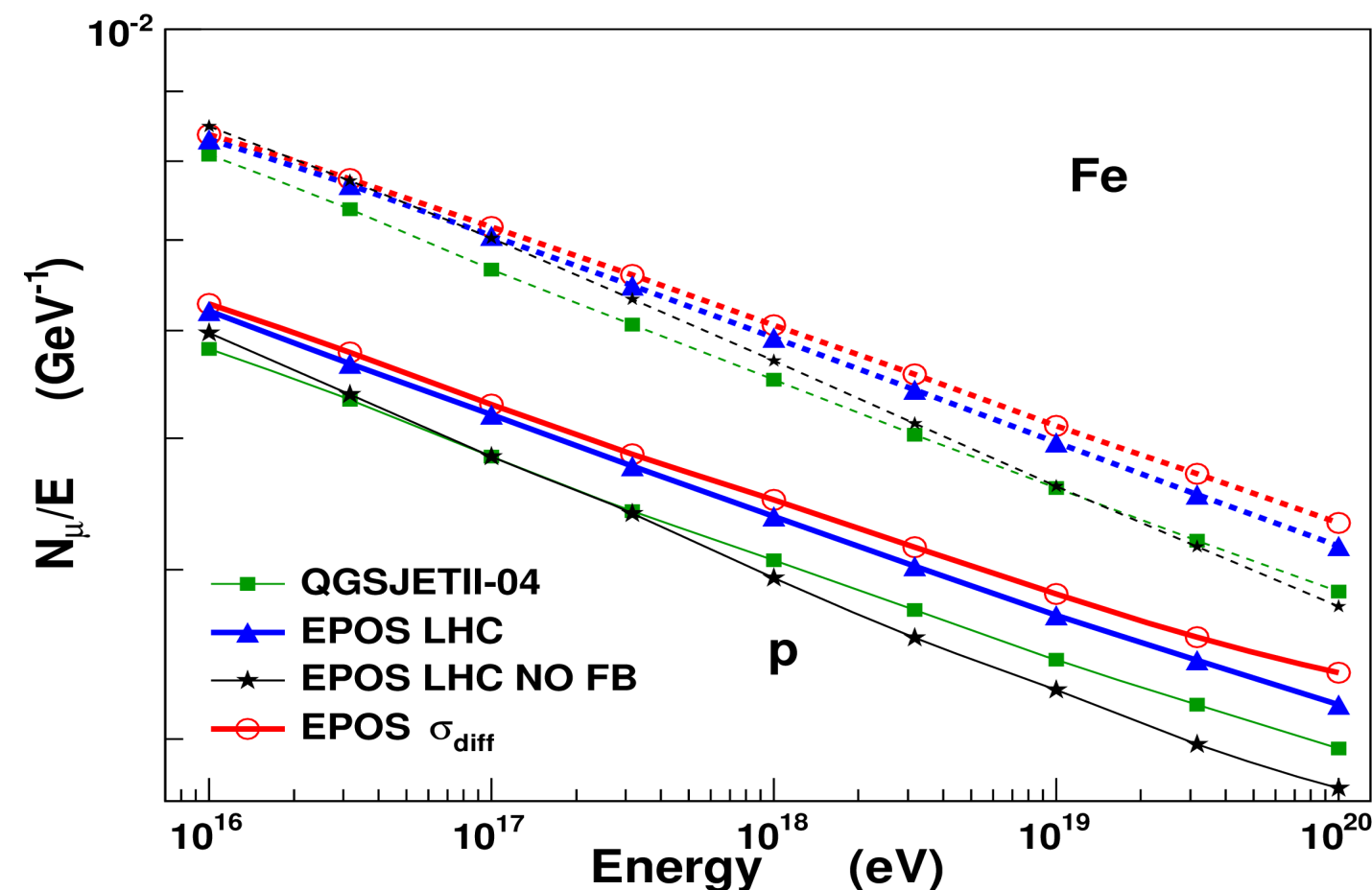
~0 g/cm² for baryons

X_{\max} less sensitive to baryon spectra than to pion spectra in pion interactions

N_μ with Modified EPOS LHC

Number of muons depends on the same parameters

- softer meson spectra: larger N_μ
- forward baryons: lower N_μ but could be compensated by ρ^0 (keep energy to produce muons but doesn't change the number of generations: lower MPD)



N_μ sensitive to baryon (less generation) and meson spectra in pion interactions

+5% for diff
-15% without forward baryons

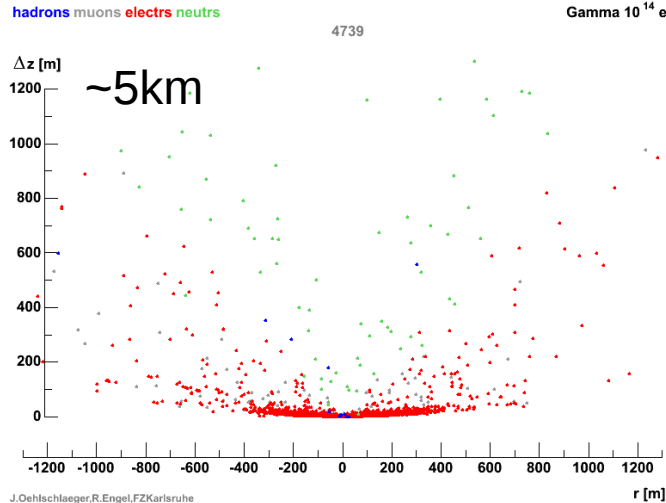
New Baryon Measurement

- **Very few data for baryon production from meson projectile, but for all :**
 - ➔ strong baryon acceleration (probability $\sim 20\%$ per string end)
 - ➔ proton/antiproton asymmetry (valence quark effect)
 - ➔ target mass dependence
- **New data set from NA49 (G. Veres' PhD)**
 - ➔ test π^+ and π^- interactions and productions at 158 GeV with C and Pb target
 - ➔ confirm large forward proton production in π^+ and π^- interactions but not for anti-protons
 - ◆ forward protons in pion interactions are due to strong baryon stopping (nucleons from the target are accelerated in projectile direction)
 - ◆ strong effect only at low energy
 - ➔ EPOS overestimate forward baryon production at high energy
 - ➔ Source of discrepancy understood and will be corrected in EPOS 3.
- **New measurements by NA61 can be used to confirm this result**

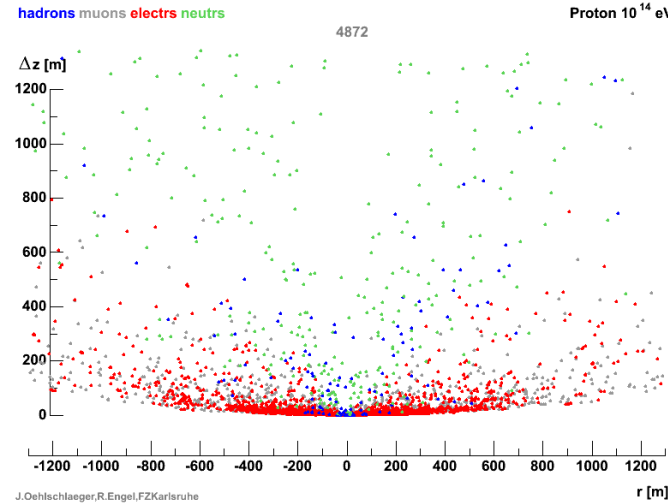
Air Showers at High Altitude

Thick shower front (close to maximum) : \rightarrow more particles and less fluctuations
 \rightarrow sensitive to details of nuclear int. ? \rightarrow shell structure detection ?

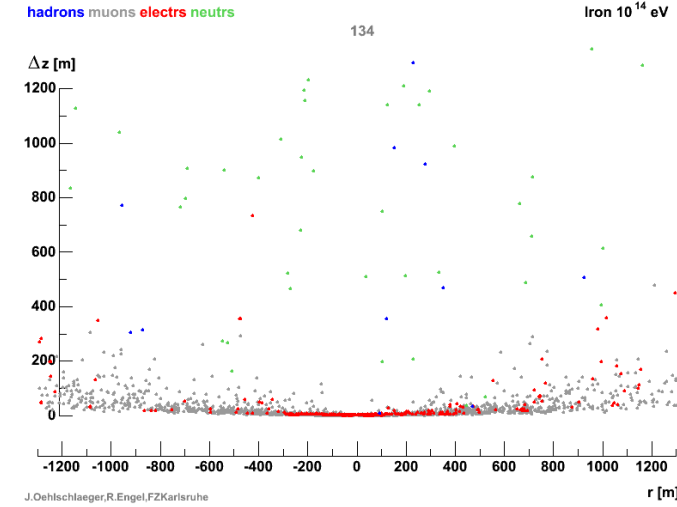
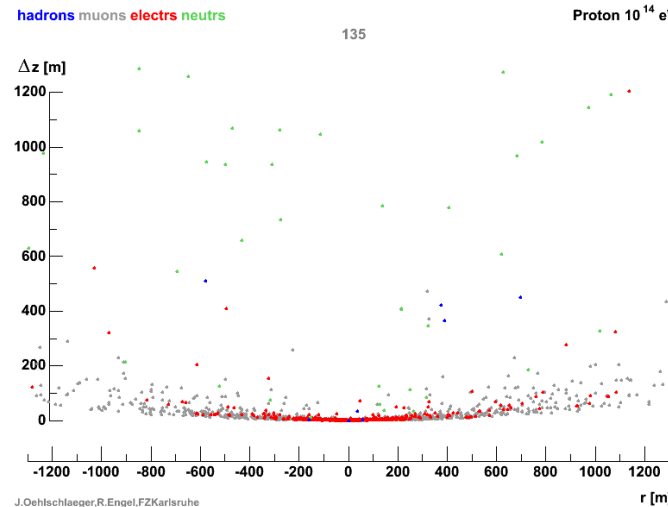
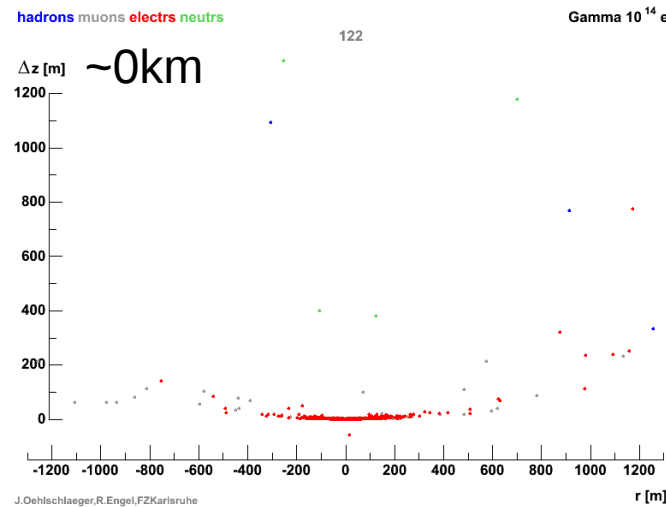
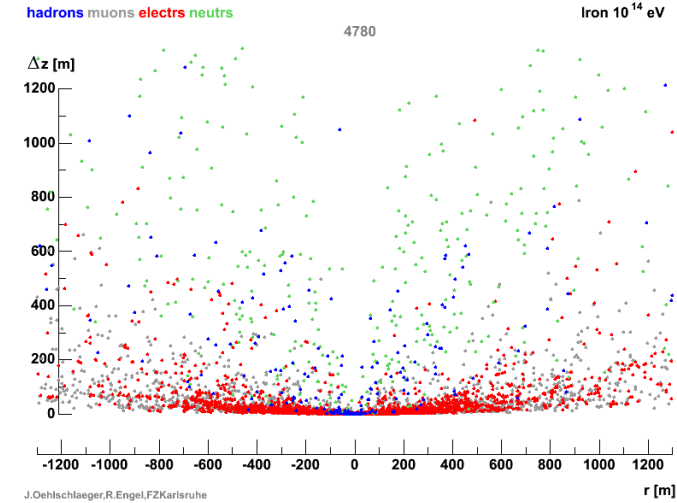
Gamma



Proton



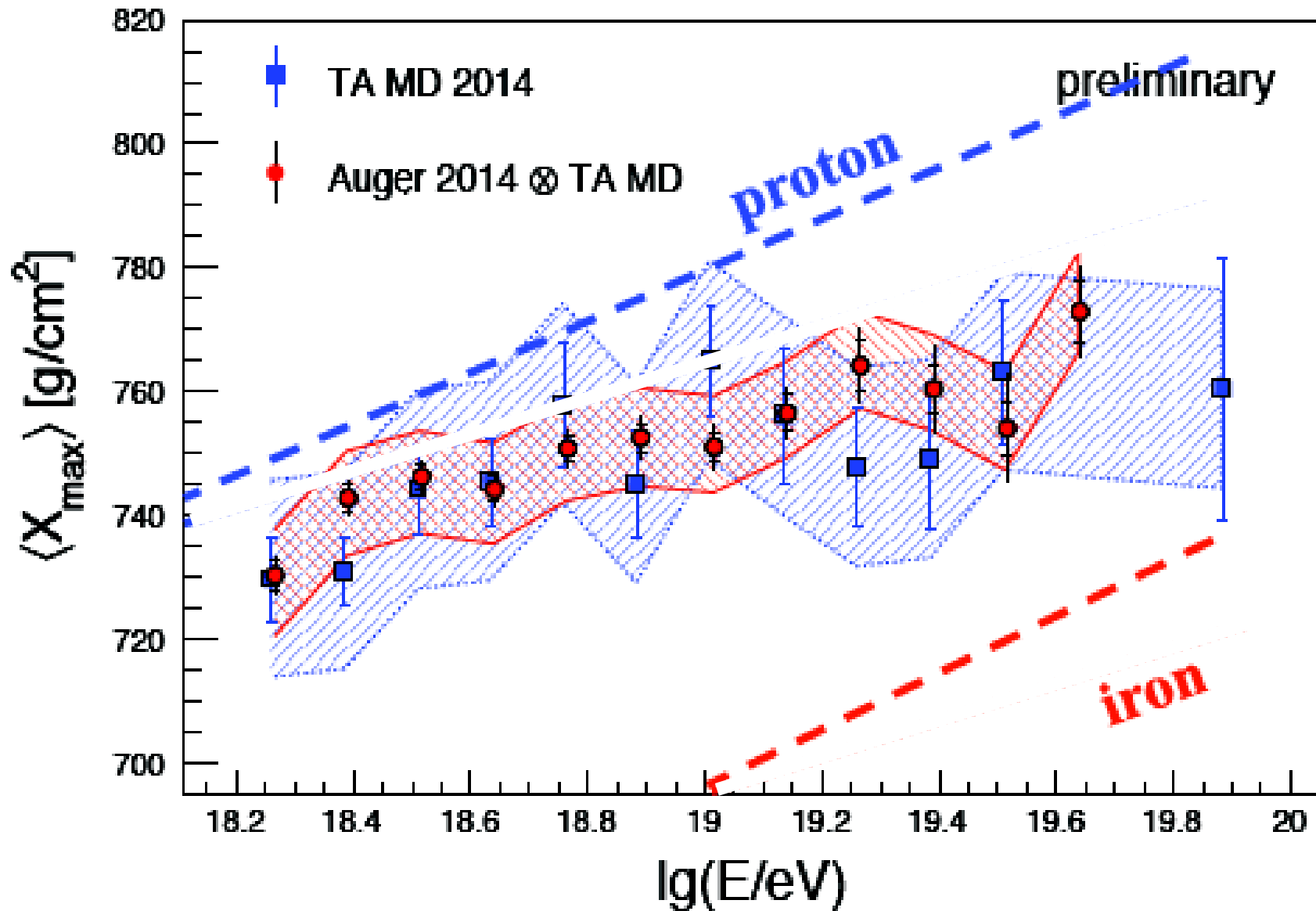
Iron



Summary

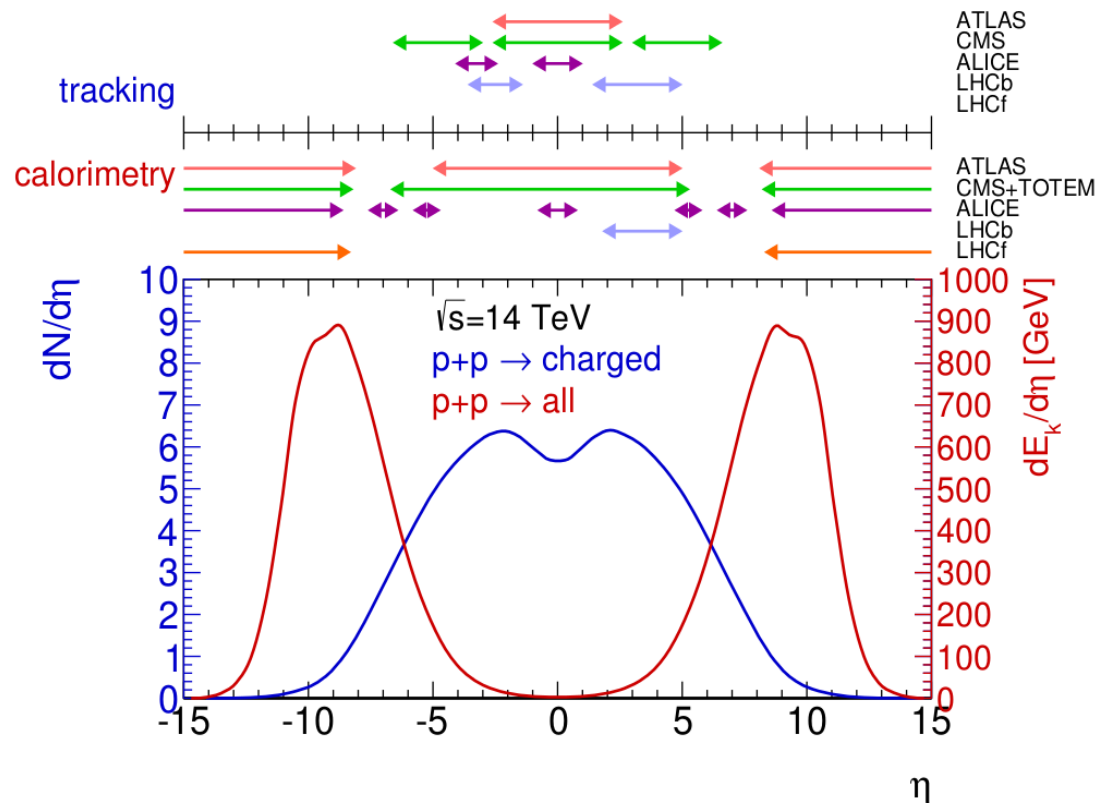
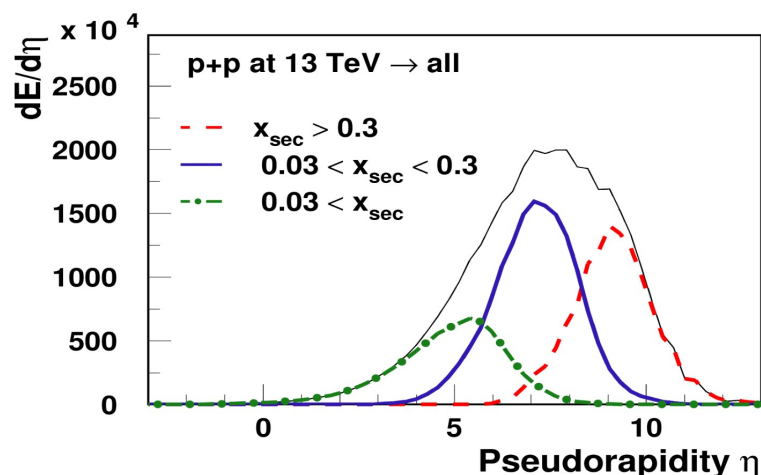
- Auger data (and other low energy cosmic ray experiments) not consistently described by hadronic interaction models (even post LHC)
 - ◆ $\langle X_{\max} \rangle$ and fluctuations, number of muons and muon production depth ...
 - ➔ but it has never been so good ! only 1 to 2 sigma difference in most of the cases
- Central particle production at LHC reduced model uncertainties in slope of X_{\max}
 - ➔ same energy evolution in models important for mass of primary cosmic rays
 - ➔ all pre-LHC models in contradiction with LHC data (central and forward prod.)
 - ➔ using latest model version reduce uncertainties and avoid unphysical behavior
- Improvements to come (EPOS 3 for ICRC 2017, others ?)
 - ➔ forward physics: photon and neutron spectra and diffraction measured at LHC, and baryon stopping and resonance production at SPS
 - ➔ effect of extrapolation to p-Air interaction: p-Pb measurements can be used to constrain nuclear effects (p-O would be the best check).
 - ➔ effect of (very) low energy: extension to very low energy (few GeV) to have a better control on the muon production.

TA ...



From Roberto Aloiso talk (2015 working group)

LHC acceptance



- p-p data of central detectors used to reduce uncertainty by factor ~2
 - ➔ p-Pb difficult to compare to CR models (only EPOS)
 - ➔ special centrality selection
 - ➔ pO ?
- Direct photon energy spectra from LHCf
 - ➔ small phase space but relevant for X_{max}
 - ➔ p-Pb (O) and correlation with ATLAS
- Average elasticity/inelasticity (energy fraction of the leading particle)
 - ➔ all diffraction measurement to be taken into account