

Struttura del workshop

- **Session 1**: Theorists motivations, ideas and wishes
 - ... direct and indirect searches for new particles (axions, dark photons,
 HNL, etc...), as well as testing symmetries and interactions...
- Session 2: Accelerator and infrastructure opportunities at CERN
 - ... upgrades of the CERN complex. Anticipated topics include proton yields in the LINAC4 era, beam dump facilities, fixed target options, and non-accelerator options. Novel ideas such as proton drivers, muon rings, EDM rings will be explored.
- **Session 3**: Potential future of existing programmes
 - The main experimental collaborations other than LHC currently operating at CERN will present their views for their mid-term and longterm future.
- **Session 4**: New experimental ideas
 - This session will be the focus of the workshop and aims at collecting ideas from the community for new experimental programmes exploiting the unique features of CERN.

Questo report: approccio "sperimentale"

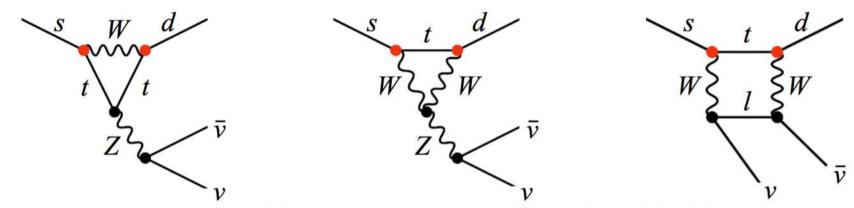
- 1. k->pnn + neutrini right handed
- 2. Dark sector with missing energy
- 3. QCD, misure di sezione d'urto
- 4. EDM + (g-2)
- 5. Antimatter experiments
- 6. Beam facilities: gamma factories + Awake
- 7. Esperimenti non agli acceleratori: axions, Kwisp, DarkSide

$1 - k > \pi \nu \nu$ e dark sector

(ma perche' sono messi insieme?)

$K \to \pi \nu \bar{\nu}$ in the Standard Model

FCNC processes dominated by Z-penguin and box amplitudes:



Extremely rare decays with rates very precisely predicted in SM:

- Hard GIM mechanism + pattern of CKM suppression (V_{ts}*V_{td})
- No long-distance contributions from amplitudes with intermediate photons
- Hadronic matrix element obtained from $BR(K_{e3})$ via isospin rotation

	SM predicted rates Buras et al, JHEP 1511*	Experimental status
$K^+ \rightarrow \pi^+ \nu \overline{\nu}$	BR = $(8.4 \pm 1.0) \times 10^{-11}$	BR = $(17.3^{+11.5}_{-10.5}) \times 10^{-11}$ Stopped K^+ , 7 events observed BNL 787/949, PRD79 (2009)
$K_L ightarrow \pi^0 u \overline{ u}$	BR = $(3.4 \pm 0.6) \times 10^{-11}$	BR < 2600 × 10 ⁻¹¹ 90%CL KEK 391a, PRD81 (2010)

^{*} Tree-level determinations of CKM matrix elements

$K \rightarrow \pi \nu \bar{\nu}$ and the unitarity triangle

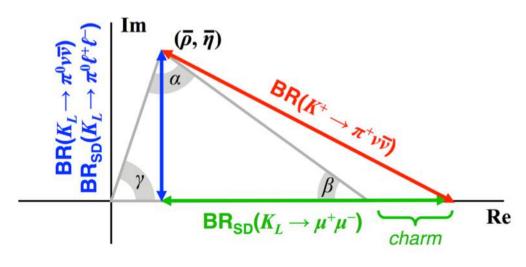
Dominant uncertainties for SM BRs are from CKM matrix elements

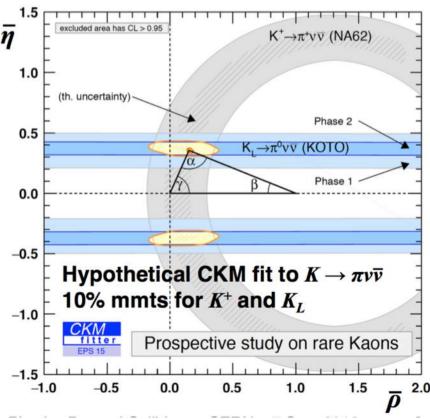
$$BR(K^+ \to \pi^+ \nu \bar{\nu}) = (8.39 \pm 0.30) \times 10^{-11} \cdot \left[\frac{|V_{cb}|}{0.0407} \right]^{2.8} \cdot \left[\frac{\gamma}{73.2^{\circ}} \right]^{0.74}$$

Buras et al., JHEP 1511

$$BR(K_L \to \pi^0 \nu \bar{\nu}) = (3.36 \pm 0.05) \times 10^{-11} \cdot \left[\frac{|V_{ub}|}{3.88 \times 10^{-3}} \right]^2 \cdot \left[\frac{|V_{cb}|}{0.0407} \right]^2 \cdot \left[\frac{\sin \gamma}{\sin 73.2^{\circ}} \right]^2$$

- Intrinsic theory uncertainties ~ few percent
- BR measurements for both K⁺ and K_L
 determine the unitarity triangle
 independently from B inputs
 - Overconstrain CKM matrix → reveal NP?





Decadimento carico: NA62

Run in 2015:

Commissioning of L0 trigger Run up to nominal intensity, 3.3 10¹² POT/spill, 3.5-s effective-length spill

Running in 2016:

Stable running at 20% of the nominal beam intensity

Data already collected: sensitivity to BR(K $\rightarrow \pi \nu \nu$) up to 10⁻⁹

End of 2016: reach SM-expectation sensitivity, O(10⁻¹⁰)

End of 2017 run: improve (by much) on present state of the art (BNL measurement)

End of 2018 run: measurement of BR at 10%

$K_L \to \pi^0 \nu \bar{\nu}$: experimental considerations

Essential signature: 2γ with unbalanced p_{\perp} + nothing else!

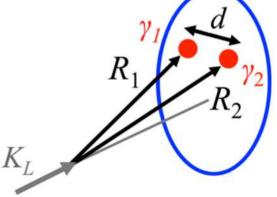
All other K_L decays have ≥ 2 extra γ s or ≥ 2 tracks to veto (Exception: $K_L \rightarrow \gamma \gamma$, not a big problem since $p_{\perp} = 0$)

 $M(\gamma\gamma) = m(\pi^0)$ is the only sharp kinematic constraint

Generally used to reconstruct vertex position

$$m_{\pi^0}^2 = 2E_1 E_2 (1 - \cos \theta)$$

$$R_1 \approx R_2 \equiv R = \frac{d\sqrt{E_1 E_2}}{m_{\pi^0}}$$



Main backgrounds:

$$K_L \rightarrow \pi^0 \pi^0$$
 with 2 lost γ s BR = 8.65 × 10⁻⁴

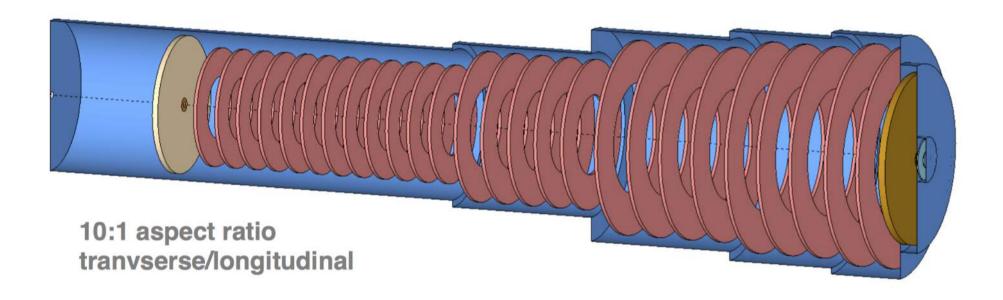


Hermetic veto, including beam exit

$$n + gas \rightarrow X\pi^0, X\eta$$



High vacuum decay region

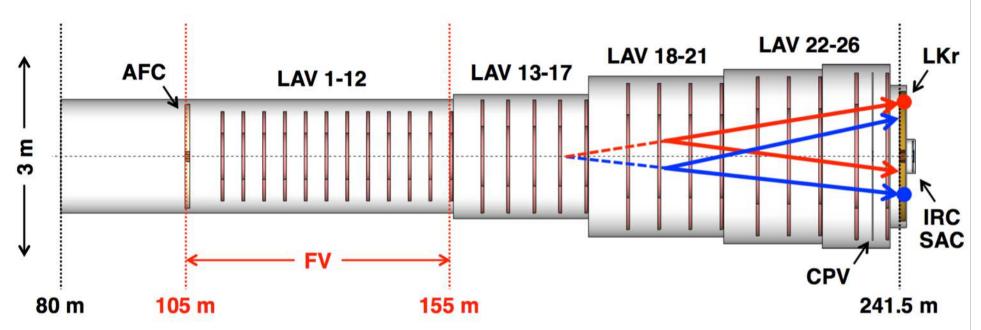


Roughly same vacuum tank layout and fiducial volume as NA62

- Upstream edge of vacuum tank 20 m in front of AFC
- FV starts ~105 m downstream of target decay away A, K_S
- About 2.2% of K_L in beam decay in FV

Possible to re-use NA48 LKr calorimeter (?)

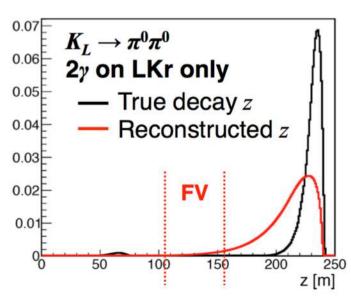
26 new large-angle photon veto stations (LAV), coverage to 100 mrad New small-angle photon veto detectors (IRC/SAC)

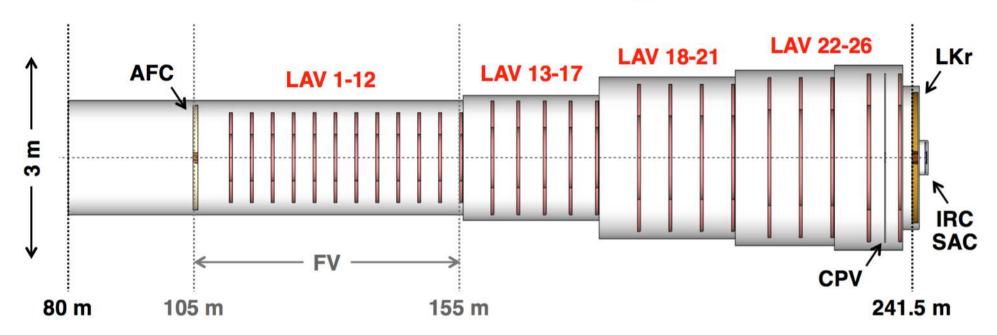


Vacuum tank layout and FV similar to NA62

90-m distance from FV to LKr significantly helps background rejection

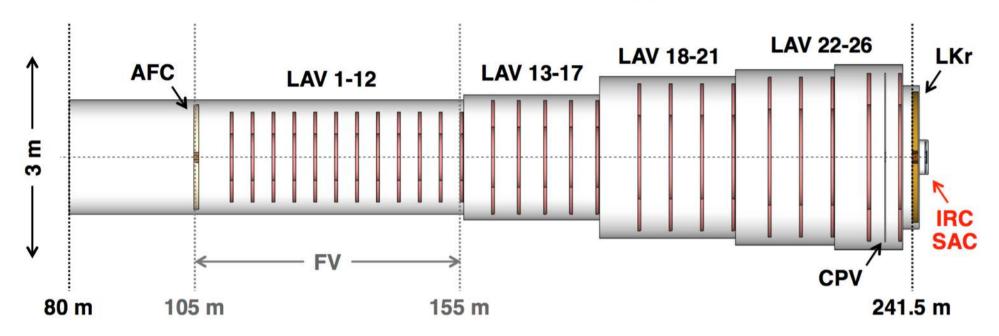
- Most $K_L \to \pi^0 \pi^0$ decays with lost photons occur just upstream of the LKr
- " π^0 s" from mispaired γ s are mainly reconstructed downstream of FV





26 new large-angle photon veto stations (LAV)

- 5 sizes, sensitive radius 0.9 to 1.6 m, at intervals of 4 to 6 m
- Hermetic coverage out to 100 mrad for E_{γ} down to ~100 MeV
- Baseline technology: Lead/scintillator tile with WLS readout Based on design of CKM VVS
 Assumed efficiency based on E949 and CKM VVS experience



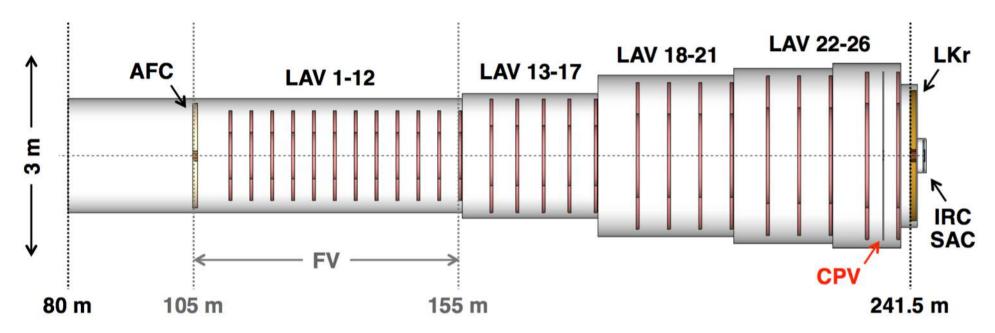
Small-angle photon veto systems (IRC, SAC)

- Reject high-energy γ s from $K_L \to \pi^0 \pi^0$ escaping through beam hole
- Must be insensitive as possible to 3 GHz of beam neutrons

	Rate (MHz)	Req. 1 – ε
$\gamma, E > 5 \text{ GeV}$	230	10-2
γ , $E > 30 \text{ GeV}$	20	10-4
n	3000	_

Baseline solution:

Tungsten/silicon-pad sampling calorimeter with crystal metal absorber



Charged particle rejection

- Charged particle vetoes (CPV): Scintillating tiles, 5 m upstream of LKr
- Re-use NA62 hadronic calorimeters (MUV1/2, not shown)
 Ratio of hadronic/total energy effective to identify π showers
- LKr shower profile: Use cluster RMS to identify and reject π

$K_L \rightarrow \pi^0 \nu \bar{\nu}$ sensitivity summary

Channel	Simulated statistics	Events found	Expected in 5 yrs*
$K_L \! o \pi^0 u \overline{ u}$	100k yr	1.94M	97
$K_L \! o \pi^0 \pi^0$	5 yr	111	111
$K_L ightarrow \pi^0 \pi^0 \pi^0$ All bkg evts from cluster fusion Upstream decays not yet included	1 yr	3	15
$K_L \! o \gamma \gamma \ _{p_{\perp}}$ cut very effective	3 yr	0	0
$K_L o$ charged	though	nt to be redu	cible

^{*}Must subtract 35% for K_L losses in dump γ converter

~ 60 SM $K_L \rightarrow \pi^0 v \bar{v}$ in 5 years with S/B ~ 1

Background study incomplete!

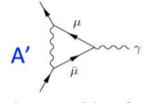
 π^0 from interactions of halo neutrons on residual gas, detector materials Radiative K_L decays, K_S /hyperon decays

Hidden sector at NA62: motivations

If DM is a thermal relic from hot early universe, can hunt for it in particle-physics: search for non-gravitational interactions DM-SM

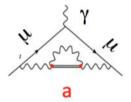
A mediator of a hidden sector might exist, inducing DM-SM field (feeble) interactions many possible dynamics: vector (A', aka dark photon), neutrino (HNL), axial (ALP a), scalar...

Various experimental hints for hidden sector at MeV-GeV, e.g., a_{μ} 3.5- σ discrepancy:



Might be due to a dark photon A' ...

A' μ enhancing light-by-light?



[Okun, Holdom]

[Marciano, et al. arXiv:1607.01022]

Model dependence: experimentally driven approach

Feeble interaction: ultra-suppressed production rate, very long-lived states E.g.: 1-GeV mass HNL, $\tau \sim 10^{-5}$ --10⁻² s, decay length ~ 10 --10000 Km at SPS energies, suppression at production 10⁻⁷--10⁻¹⁰



New Physics below the Fermi Scale

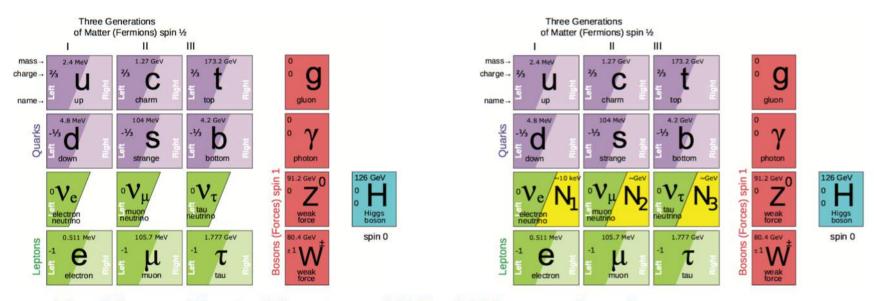
Mikhail Shaposhnikov



Superweakly interacting new physics below the Fermi scale

- Light messengers between the SM and heavy sectors such as SUSY, extra dimensions, new strong dynamics.
- Light Dark matter and/or light particles in the hidden sector
- Absence of new particle thresholds between the Fermi and Planck scales? Then all the experimental BSM problems should be explained by light particles.

An example: ν MSM



N = Heavy Neutral Lepton - HNL, Majorana fermion

Role of N_1 with mass in keV region: dark matter

Role of N_2 , N_3 with mass in 100 MeV – 100 GeV region: "give" masses to neutrinos and produce baryon asymmetry of the Universe Role of the Higgs: give masses to quarks, leptons, Z and W and inflate the Universe.

For comprehensive study of possibilities at CERN SPS see:

IOP Publishing

Reports on Progress in Physics

Rep. Prog. Phys. 00 (2016) 000000 (137pp)

UNCORRECTED PROOF

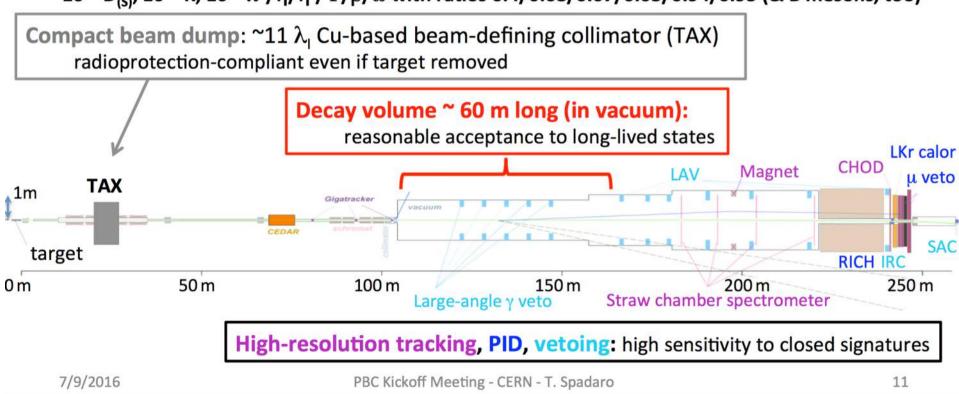
Report on Progress

A facility to search for hidden particles at the CERN SPS: the SHiP physics case

```
Sergey Alekhin<sup>1,2</sup>, Wolfgang Altmannshofer<sup>3</sup>, Takehiko Asaka<sup>4</sup>,
Brian Batell<sup>5</sup>, Fedor Bezrukov<sup>6,7</sup>, Kyrylo Bondarenko<sup>8</sup>, Alexey Boyarsky<sup>8,68</sup>,
Ki-Young Choi9, Cristóbal Corral10, Nathaniel Craig11, David Curtin12,
Sacha Davidson<sup>13,14</sup>, André de Gouvêa<sup>15</sup>, Stefano Dell'Oro<sup>16</sup>,
Patrick deNiverville<sup>17</sup>, P S Bhupal Dev<sup>18</sup>, Herbi Dreiner<sup>19</sup>,
Marco Drewes<sup>20</sup>, Shintaro Eijima<sup>21</sup>, Rouven Essig<sup>22</sup>, Anthony Fradette<sup>17</sup>,
Björn Garbrecht<sup>20</sup>, Belen Gavela<sup>23</sup>, Gian F Giudice<sup>5</sup>, Mark D Goodsell<sup>24,25</sup>,
Dmitry Gorbunov<sup>26,27</sup>, Stefania Gori<sup>3</sup>, Christophe Grojean<sup>28,29,69</sup>,
Alberto Guffanti<sup>30</sup>, Thomas Hambye<sup>31</sup>, Steen H Hansen<sup>32</sup>,
Juan Carlos Helo<sup>10,33</sup>, Pilar Hernandez<sup>34</sup>, Alejandro Ibarra<sup>20</sup>,
Artem Ivashko<sup>8,35</sup>, Eder Izaguirre<sup>3</sup>, Joerg Jaeckel<sup>36,69</sup>, Yu Seon Jeong<sup>37</sup>,
Felix Kahlhoefer<sup>29</sup>, Yonatan Kahn<sup>38</sup>, Andrey Katz<sup>5,39,40</sup>,
Choong Sun Kim<sup>37</sup>, Sergey Kovalenko<sup>10</sup>, Gordan Krnjaic<sup>3</sup>,
Valery E Lyubovitskij<sup>41,42,43</sup>, Simone Marcocci<sup>16</sup>, Matthew Mccullough<sup>5</sup>,
David McKeen<sup>44</sup>, Guenakh Mitselmakher<sup>45</sup>, Sven-Olaf Moch<sup>46</sup>,
Rabindra N Mohapatra<sup>47</sup>, David E Morrissey<sup>48</sup>, Maksym Ovchynnikov<sup>35</sup>,
Emmanuel Paschos<sup>49</sup>, Apostolos Pilaftsis<sup>18</sup>, Maxim Pospelov<sup>3,17,69</sup>,
Mary Hall Reno<sup>50</sup>, Andreas Ringwald<sup>29</sup>, Adam Ritz<sup>17</sup>, Leszek Roszkowski<sup>51</sup>,
Valery Rubakov<sup>26</sup>, Oleg Ruchayskiy<sup>21,52,68</sup>, Ingo Schienbein<sup>53</sup>,
Daniel Schmeier<sup>19</sup>, Kai Schmidt-Hoberg<sup>29</sup>, Pedro Schwaller<sup>5</sup>,
Goran Senjanovic<sup>54,55</sup>, Osamu Seto<sup>56</sup>, Mikhail Shaposhnikov<sup>21,68,69</sup>,
Lesya Shchutska<sup>45,69</sup>, Jessie Shelton<sup>57</sup>, Robert Shrock<sup>58</sup>, Brian Shuve<sup>3</sup>,
Michael Spannowsky<sup>59</sup>, Andy Spray<sup>60</sup>, Florian Staub<sup>5</sup>, Daniel Stolarski<sup>5</sup>,
Matt Strassler<sup>40</sup>, Vladimir Tello<sup>54</sup>, Francesco Tramontano<sup>61,62,69</sup>,
Anurag Tripathi<sup>61</sup>, Sean Tulin<sup>63</sup>, Francesco Vissani<sup>16,64</sup>, Martin W Winkler<sup>65</sup>
and Kathryn M Zurek<sup>66,67</sup>
```

NA62 perfectly suited for hidden sector searches

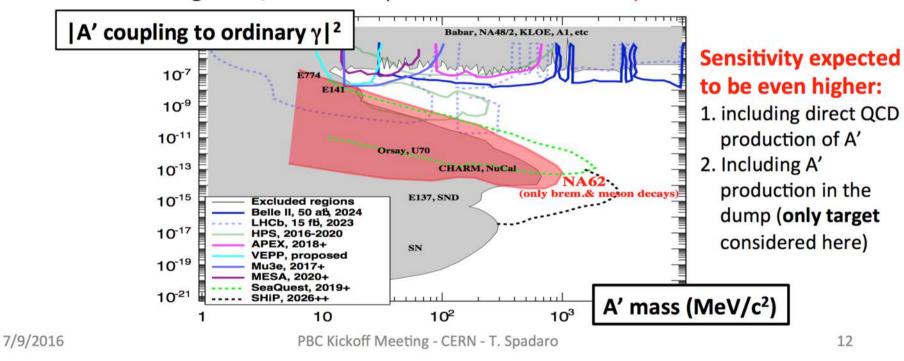
High-intensity 400-GeV proton beam \rightarrow boost charm/beauty, other meson production 10¹⁸ POT / nominal year: 10¹² POT/sec on spill, 3.5-s/16.8 s, 100 days/year, 60% run efficiency 10¹⁵ D_(s), 10¹⁴ K, 10¹⁸ $\pi^0/\eta/\eta'/\Phi/\rho/\omega$ with ratios 6.4/0.68/0.07/0.03/0.94/0.95 (& B mesons, too)



Search for visible decays of long-lived A'

Assume 2x10¹⁸ 400-GeV POT:

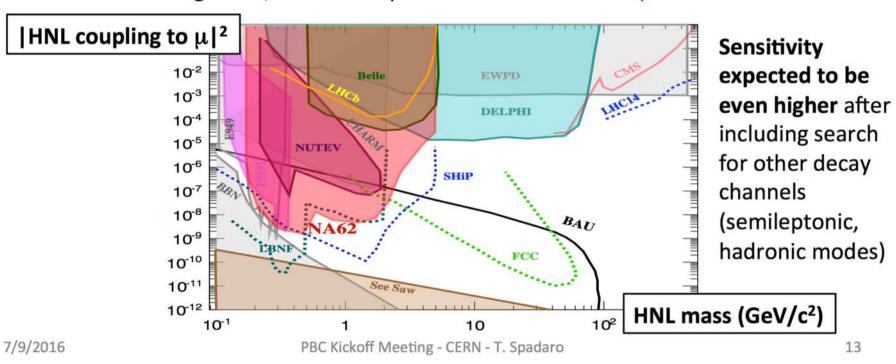
search for displaced, dilepton decays of dark photons, A' \rightarrow ee, $\mu\mu$ include trigger/acceptance/selection efficiency assume zero-background, evaluate expected 90%-CL exclusion plot



Search for visible decays of heavy neutral leptons

Assume 2 1018 400-GeV POT:

search for displaced, leptonic decays HNL $\rightarrow \pi e$, $\pi \mu$ include trigger/acceptance/selection efficiency assume zero-background, evaluate expected 90%-CL exclusion plot



commento

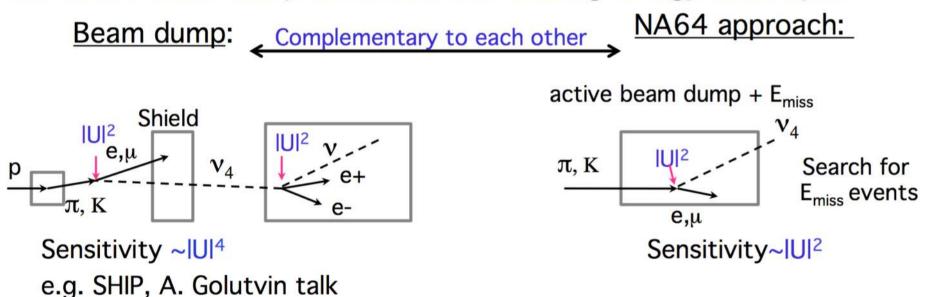
- misure di SHIP (molto) migliori di quelli di NA62 sia come statistica che come controllo del fondo
- ... ma quelle di NA62 possono essere fatte 10 anni prima ad un costo inferiore di ~100M!
- naturalmente l'utilizzo di NA62 come rivelatore per "hidden sector" ritarda lo sviluppo del progetto sui K neutri
- (penso che) una decisione sara' presa nel 2018 anche alla luce della quantita' di dati raccolti nel canale carico

2 – Dark sector with missing energy

NA64

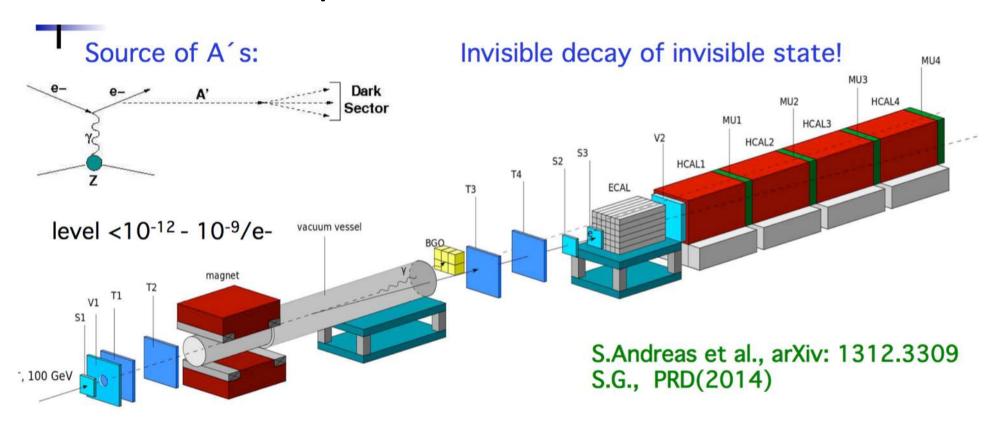
NA64 approach to search for new physics

The NA64 is a new fixed-target experiment aiming to search for dark photon A´->invisible decays at the CERN SPS with a new approach: the active beam dump combined with missing energy technique.



Goal of the talk: to show that this approach allows a sensitive probe of "light new physics" (dark sector physics, new symmetries, new WI sub-GeV particles coupled to e, μ , q's) by using e, μ , π , K, and p beams from existing facilities at CERN.

esempio con fasci di elettroni



Main components:

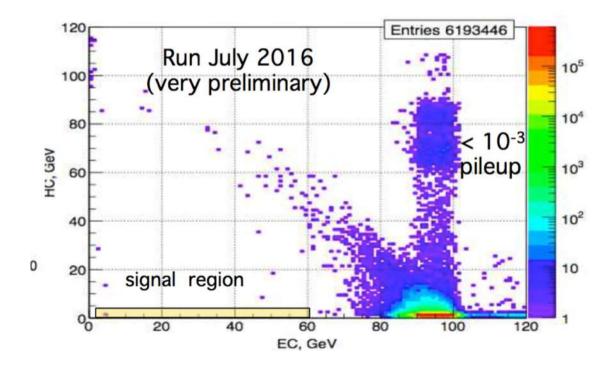
- clean, mono-energ. 100 GeV e- beam
- e- tagging system: MM/GEM/Straw tracker + SRD (BGO, LYSO; Pb-SC)
- 4π fully hermetic ECAL+ HCAL

Signature:

- in: 100 GeV e- track
- out: < 50 GeV e-m shower in ECAL
- no energy in the Veto and HCAL
- Sensitivity ~ ε²
- same signature for milli-q, APLs,

Esperimento NA64

- la qualita' e l'intensita' del fascio di ingresso sono fondamentali -> vedi in seguito proposta di AWAKE
- al momento si usa il fascio della North Area
- plot di energia rilasciata nel calorimetro adronico (HC) vs calorimetro elettromagnetico (EC)

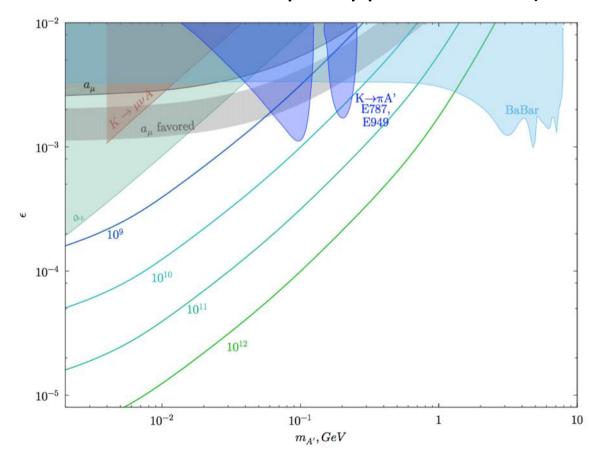


Summary of background sources for A´-> invisible

Source	Expected level	Comment
Beam contamination		
$-\pi$, p , μ reactions and punchthroughs,	< 10 ⁻¹³ -10 ⁻¹²	Impurity < 1% high precision MM
- e ⁻ low energy tail due to bremss., π , μ -decays in flight	< 10 ⁻¹²	tracker + e ⁻ SR photon tag
Detector		
ECAL+HCAL energy resolution, transverse hermeticity, holes, dead material, cracks	<10 ⁻¹³	Full upstream coverage
Physical		
- hadron electroproduction, e.g. e-A->e-A* + n,π,ρ,J/ψ-	< 10 ⁻¹³	HERA ep-data
 n punchthrough, μ inefficiency WI process: e⁻ Z->e⁻ Zvv 	< 10 ⁻¹³	(H1 Collaboration) WI σ estimated.
Total	< 10 ⁻¹²	

S.N. Gninenko - Search for dark sector physics - PBC kickoff workshop, CERN, Sept 6-7, 2016

- il regime di esclusione dipende fondamentalmente dal numero di "electrons on target"
- nel caso ottimistico di "zero background"!
- (non ho studiato in maniera approfondita l'esperimento, ma le stime fatte sembrano un po' approssimative)



3 – QCD, misure di sezione d'urto

COMPASS (m, e, hadron beams)

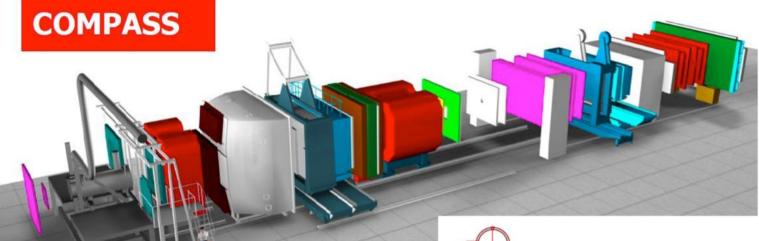
NA61/SHINE (hadron and nuclear beams)

Fixed target @ LHC



COMPASS Spectrometer at SPS M2 beam line (CERN)





Universal and flexible apparatus.

Most important features of the two-stage COMPASS Spectrometer:

- 1. Muon, electron or hadron beams with the momentum range 20-250 GeV and intensities up to 108 particles per second
- 2. Solid state polarised targets (NH₃ or ⁶LiD) as well as liquid hydrogen target and nuclear targets
- 3. Adv ~240 physicists (~60 PhD, ~25 Master&Diploma syst students), 12 countries + CERN, 24 institutions 11



Beyond 2020 dedicated Workshop



This week - regular annual COMPASS Workshop (IWHSS'16 Kloster Seeon)

COMPASS beyond 2020 Workshop

2-

- 21 Mar 2016, 08:05 → 22 Mar 2016, 17:10 Europe/Zurich
- 222-R-001 (CERN)

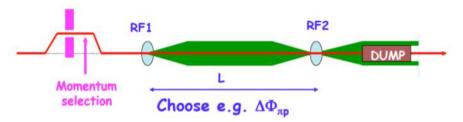
Description The goal of the workshop is to explore hadron physics opportunities for fixed-target COMPASS-like experiments at CERN beyond 2020 (CERN Long Shutdown 2 2019-2020). The programme comprises

- Reviews of the various physics domains: TMDs, GPDs, FFs, spectroscopy, exotics, tests of ChPT, astrophysics
- Reviews of physics results expected in the next 10 years from major labs around the world
- Good attendance (>100 physicists), large interest
- 11 "outside" review talks Jefferson Lab, RHIC, Fermilab, KEK (Japan) BEPC II (IHEP, Beijing), NICA (JINR, Dubna), CERN (After, LHCb), GSI (Panda), J-PARC (Japan), EIC China;
- 7 COMPASS talks (chronol.) SIDIS, GPDs, Chiral Dynamics, astrophysics (dark matter), Drell-Yan, hadron spectroscopy;
- 2 "round-table"-like discussions on possible future with hadron and muon beams;
- Outcome of the Workshop:
 - RF Separated antiproton/kaon beam would provide a unique opportunity for future fixed target COMPASS-like program at CERN
 - Existing muon and hadron beam allows to extend current COMPASS program by doing unique or first class measurements of exclusive processes, SIDIS and Drell-Yan



RF separated antiproton/kaon beam





$$\Delta\Phi = 2\pi (L f / c) (\beta_1^{-1} - \beta_2^{-1})$$
 with $\beta_1^{-1} - \beta_2^{-1} = (m_1^2 - m_2^2)/2p^2$

"Normal" h beam composition: ~97% (π) ~2.5%(K) ~0.5% (pbar)

Assumptions:

8 x 10⁷ antiprotons for 10¹³ ppp (10 seconds) (optimistic estimate by Lau Gatignon);
 we assume here 4 x 10¹³ protons.

Antiprotons RF separated beam: 3.2 x 10^7 /s - Gain is a factor of 50 compared to the standard harmonic beam for Drell-Yan experiment (~1% of harmonic beam 6x10⁷ /s dominated by π -)

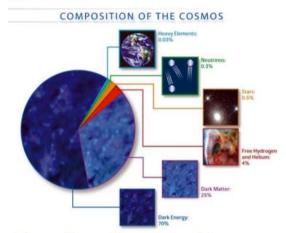
Using the same assumption for RF separated kaon beam, possible kaon beam intensity is 8 x 10⁶ /s - Gain is a factor of 80 compared to to the standard "spectroscopy" h beam

High intensity RF separated beam will provide unique opportunities for Hadron Spectroscopy and Drell-Yan physics



Astrophysics – search for dark matter,

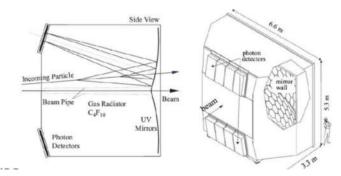


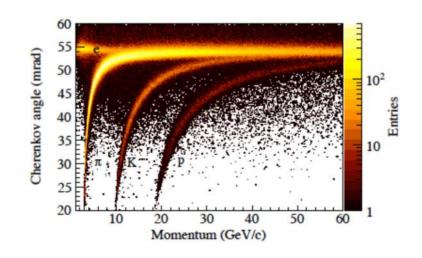


- New AMS(2) data the antiparticle flux is well known now (few % pres.);
- Two type of processes contribute SM interactions (proton on the ISM with the production for example antiprotons in the f.s.) and contribution from dark particle antiparticle annihilation;
- In order to detect a possible excess in the antiparticles flux a good knowledge of inclusive cross sections of p-He interaction with antiparticles in the f.s. is a must, currently the typical precision is of 30-50%.

Thus the primary goal is to measure inclusive antiproton (positron, gamma) production cross section in a wide kinematical range with the precision <10%. Compared to NA49 COMPASS have factor ~1000 as luminosity. COMPASS:

- Proton beam energy range 50-250 GeV
- Secondary particles identification:
 - Antiprotons (RICH)
 - Positrons and Gamma (ECals)

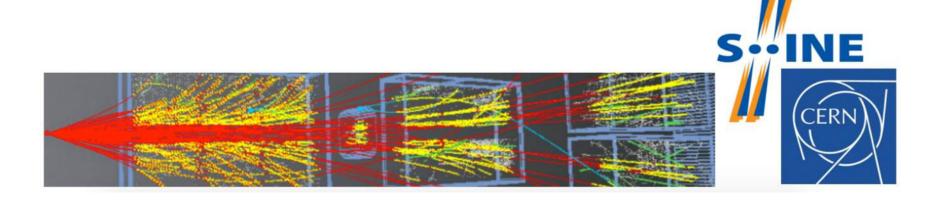




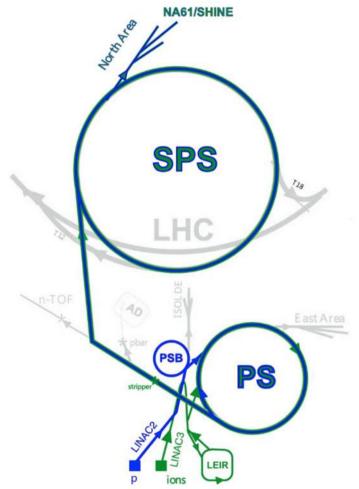
NACI/SHINE; PHYSICS AND FACILITY BEYOND 2020

M. GAZDZICKI FOR NAGI/SHINE

- MAGI/SHINE: PHYSICS AND FACILITY
- 🛅 🗎 IPEAS FOR MEASUREMENTS BEYOND 2020







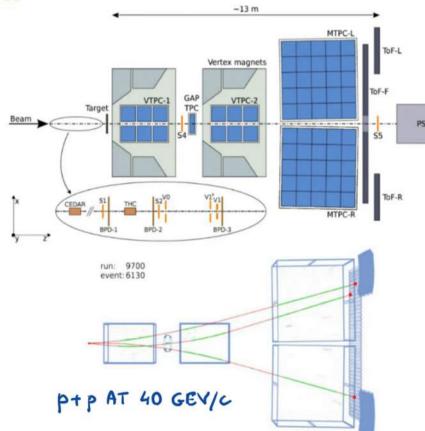
PRIMARY BEAMS:

- PROTONS AT 400 GEV/C
- 10NS (Ar, Xe, Pt) AT 13A-150A GEV/c

SECONDARY BEAMS!

- HADRONS (11- K+- P/p)
 AT 13-400 GEV/C
- IONS (Be) AT 13A-150A GEV/C

S.INE FACILITY: DETECTOR



JINST 9 (2014) P06005

- A LARGE ACCEPTANCE (2 50%)
 HADRON SPECTROMETER
- BEAM PARTICLES MEASURED BY COUNTERS AND MWPCS
- CHARGE PARTICLES MEASURED BY 5 (+2) TPCS
- PID VIA dE/dx IN TPCS AND tof IN 3 TOF DETECTORS
- ENERGY OF PROJECTILE SPECTATORS MEASURED IN PSD
- PRECISE VERTEXING VIA SMALL ACCEPTANCE VERTEX PETECTOR



IDEAS FOR MEASURMENTS BEYOND 2020

KEY QUESTIONS/ REQUESTED MEASUREMENTS:

- STRONG INTERACTIONS:

 STATISTICAL US DYNAMICAL MODELS

 (NON-LOCAL VS LOCAL PARTICLE PRODUCTION)

 OPEN CHARM, CUMULATIVE HADRONS,

 FLUCTUATIONS IN & FULL ACCEPTANCE
 - NEUTRINGS:
 INITIAL NEUTRING FLUX FOR DUNE
 / HADRON YIELDS FROM THE SURFACE OF
 THE DUNE REPLICA TARGET

COMPASS + NA61/SHINE

- in generale si nota un interesse nell'estendere le misure di "precisione" di sezioni d'urto adroniche
- questo e' spinto dalla richiesta di varie comunita':
 - astroparticle
 - neutrini
 - interazioni di QCD
 - sviluppo di MC

– ...

The fixed-target mode with TeV beams: why and what for?

4 decisive features

accessing the high x frontier

$$[|x_F| \equiv \frac{|p_z|}{p_{z\,\text{max}}} \to 1]$$

- achieving high luminosities,
- varying the atomic mass of the target almost at will,
- polarising the target.

3 physics cases

- High-x gluon, antiquark and heavy-quark content in the nucleon & nucleus
- Transverse dynamics and spin of gluons inside (un)polarised nucleons
- Heavy-ion physics between SPS & RHIC energies towards large rapidities

All this can be realised at CERN in a parasitic mode with the most energetic beams ever!

Nota: all (past) colliders with $E_p \ge 100$ GeV have had a fixed-target program (Tevatron, HERA, SPS, RHIC)

$$4 - EDM + (g-2)$$

EDM storage ring misura della sezione d'urto adronica per g-2

EDMs and precision g-2 as a window into New Physics

Maxim Pospelov Perimeter Institute/U of Victoria

Purcell and Ramsey (1949) ("How do we know that strong interactions conserve parity?" $\longrightarrow |d_n| < 3 \times 10^{-18} ecm.$)

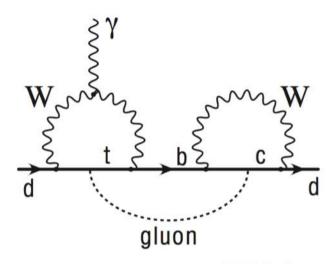
$$H = -\mu \mathbf{B} \cdot \frac{\mathbf{S}}{S} - d\mathbf{E} \cdot \frac{\mathbf{S}}{S}$$

 $d \neq 0$ means that both P and T are broken. If CPT holds then CP is broken as well.

CPT is based on locality, Lorentz invariance and spin-statistics = very safe assumption.

search for EDM = search for CP violation, if CPT holds

EDMs from CKM



CKM phase generates tiny EDMs:

$$d_d \sim \text{Im}(V_{tb}V_{td}^*V_{cd}V_{cb}^*)\alpha_s m_d G_F^2 m_c^2 \times \text{loop suppression}$$

 $< 10^{-33} e \text{cm}$

Direct quark EDMs identically vanish at 1 and 2 loop levels (Shabalin, 1981). 3-loop EDMs are calculated by Khriplovich.

d_e first appears at 4 loops (Khriplovich, MP, 1991) < 10⁻³⁷ cm

d_n is dominated by long distance effects but does not exceed 10⁻³¹.

CP violation from the Theta term

■ If CKM gave too small an EDM, there is a much bigger source of CP violation in the flavor conserving channel - theta term

Energy of QCD vacuum depends on θ -angle:

$$E(\bar{\theta}) = -\frac{1}{2}\bar{\theta}^2 m_* \langle \bar{q}q \rangle + \mathcal{O}(\bar{\theta}^4, m_*^2)$$

where $\langle \overline{q}q \rangle$ is the quark vacuum condensate and m_* is the reduced quark mass, $m_* = \frac{m_u m_d}{m_u + m_d}$. In CP-odd channel,

$$d_n \sim e \frac{\bar{\theta} m_*}{\Lambda_{\text{had}}^2} \sim \bar{\theta} \cdot (6 \times 10^{-17}) \ e \text{ cm}$$

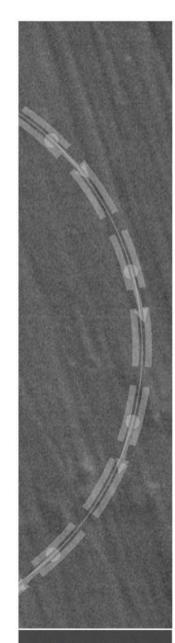
Strong CP problem = naturalness problem = Why $|\bar{\theta}| < 10^{-9}$ when it could have been $\bar{\theta} \sim O(1)$? $\bar{\theta}$ can keep "memory" of CP violation at Planck scale and beyond. Suggested solutions

Axions or clever symmetry for keeping theta=0; m_u =0...

Come misurare un EDM

- neutrone: precessione di Larmor dello spin in presenza di campo elettrico+magnetico parallelo e antiparallelo; adesso si usano Ultra Cold Neutrons (UCN) ed un campo magnetico oscillante oltre ai due E+B paralleli e antiparalleli
- nuclei (quark): nuclei pesanti (Tl, PbF, Hg, ...) vengono polarizzati, intrappolati (e.g. Penning trap) e sottoposti ad un campo elettrico
- protone, deuterio, (eventualmente) muone: intrappolare la particella in uno storage ring e combinare effetti di campo magnetico ed elettrico sui momenti magnetici ed elettrici della particella per osservare la precessione dello spin rispetto all'impulso p

$$H = -\mu \vec{B} \cdot \vec{s} - d\vec{E} \cdot \vec{s}$$



Technique $\vec{\omega}_{a} = \frac{e}{m} \left[G\vec{B} - \left(G - \frac{1}{\gamma^{2} - 1} \right) \frac{\vec{\beta} \times \vec{E}}{c} \right]$ $\vec{\omega}_{e} = \frac{\eta e}{2m} \left(\frac{\vec{E}}{c} - \vec{\beta} \times \vec{B} \right)$ $\vec{B} = 0 \quad \& \quad \gamma = \sqrt{1 + 1/G}$ Frozen Spin Teanique 8=1,248 p=0,7007 (ev/c

Physics Beyond Colliders

10



- Two technical reviews have been performed BNL: Dec 2009, March 2011
- Fermilab review. Lebedev "concept sound"
- First all-electric ring: AGS-analogue ('53-'57)

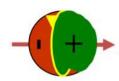
There were electrostatic quadrupoles and sextupoles, a single gap rf system, a pickup electrode system and the other appurtenances necessary to an accelerator. The Electron Analogue is the only electrostatic alternating gradient, strong focusing, synchrotron ever built and, as such, occupies a unique place in particle accelerator history. M. Plotkin '91

Ring radius 4.7m Proposed-built 1953-57

 Heidelberg Cryogenic Storage Ring: (expertise in collab.)

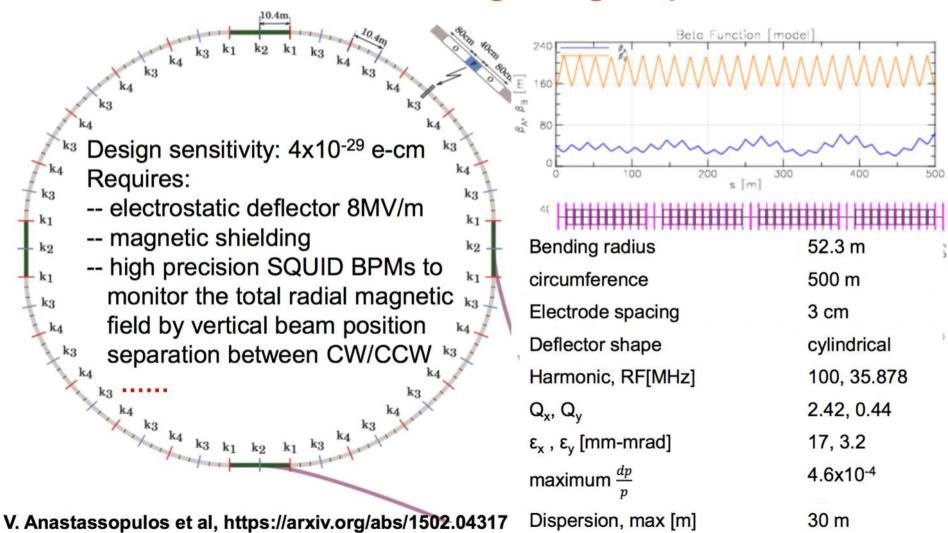


pEDM Storage Ring Mei Bai – JULICH, Aachen





Pure Electrostatic Storage Ring for proton EDM



FORSCHUNGSZENTRUM

Summary

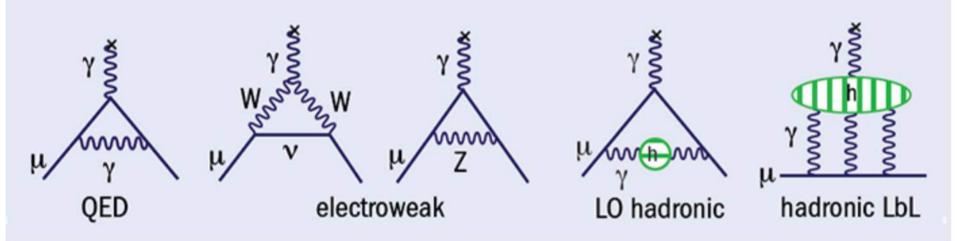




- Storage ring based EDM search offers fantastic physics
- Significant effort and progress are made worldwide
 - Experimental demonstration of long spin coherence time
 - The new low magnetic field shielding
 - High efficient polarimeter for deuteron beam that enabled key spin manipulations
 - Many others ...
- For the implementation of the EDM ring, things to consider
 - the progress of the ongoing R&D efforts worldwide
 - scenarios that allow the search of EDM of multiple ion species for fully understand the CP violation
 - and with this in mind, a staged approach for reaching the holy grail of 10⁻²⁹ e-cm or better sensitivity may benefit the community, both physics and accelerator, in the long run



Theoretical status of muon g-2, SM



The history of theoretical calculations goes very far in the past. Back to Schwinger's result, $a_u^{1-\text{loop QED}} = \alpha / (2 \pi)$

Currently, the QED calculations are carried out to four-loop order,

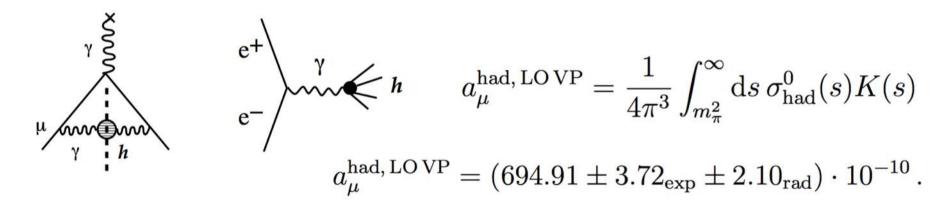
$$a_{\mu}^{\text{QED}} = 116584718.08(15) \cdot 10^{-11}$$

Electroweak calculations are also under control,

$$a_{\mu}^{\text{EW}} = (154 \pm 2) \cdot 10^{-11}$$

Hadronic contributions: A. vacuum polarization B. "light-by-light"

For A. unitarity + analyticity comes to rescue:

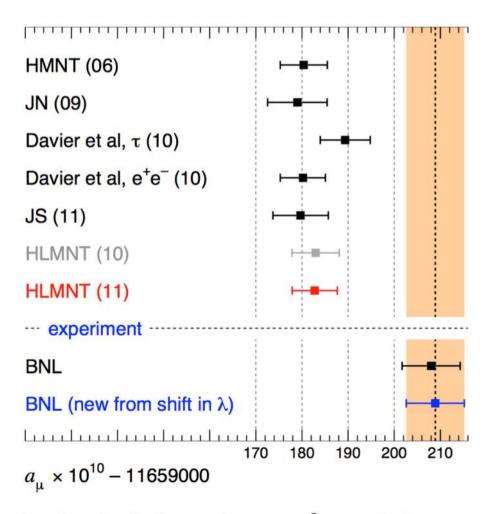


For B. some hadronic model calculations are necessary at this points (loops of heavy quarks + loops of light mesons, e.g. pions)

Recent evaluations give
$$a_{\mu}^{\rm had,\,l-by-l} = (10.5 \pm 2.6) \cdot 10^{-10} \\ a_{\mu}^{\rm had,\,l-by-l} = (11.6 \pm 4.0) \cdot 10^{-10}$$

Summary of the current discrepancy

From Hagiwara et al. (2011)



There is a theoretical "deficit" of 3×10^{-9} , and the tension is $\sim 3.5~\sigma$

Future: new experiment at Fermilab



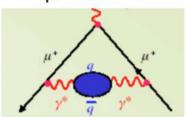
Repeat of the BNL experiment with the same storage ring, but much enhanced intensities, and improved monitoring of the magnetic field.

Goal: *shrink the size of the experimental errors by a factor of* ~4.

a_{μ}^{HLO} calculation, traditional way: time-like data

$$a_{\mu}^{HLO} = \frac{1}{4\pi^3} \int_{4m_{\pi}^2}^{\infty} \sigma_{e^+e^- \to hadr}(s) K(s) ds$$

$$a_{\mu} = (g-2)/2$$



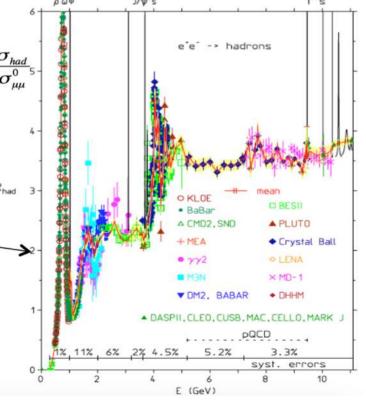
$$K(s) = \int_{0}^{1} dx \frac{x^{2}(1-x)}{x^{2} + (1-x)(s/m^{2})} \sim \frac{1}{s} \qquad \sigma_{e^{+}e^{-} \to hadr}(s) = \frac{4\pi}{s} \text{Im} \,\Pi_{had}(s)$$

 $2 \operatorname{Im} \sim = \left| \sim < \right|^2$

Traditional way: based on precise experimental (time-like) data:

$$a_{\mu}^{HLO} = (692.3 \pm 4.2)10^{-10} (DHMZ)$$

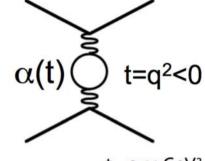
- Main contribution in the low energy region (highly fluctuating!)
- Current precision at o.6% → needs to be reduced by a factor ~2 to be competitive with the new g-2 experiments



Alternative approach: a_{μ}^{HLO} from space-like region

[C.M. Carloni Calame, M. Passera, L. Trentadue, G. Venanzoni Phys.Lett. B746 (2015) 325-32]

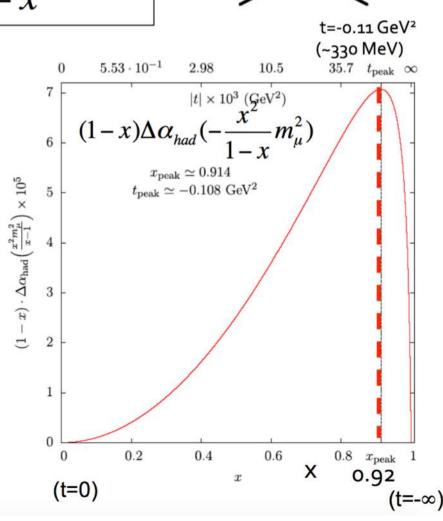
$$a_{\mu}^{HLO} = -\frac{\alpha}{\pi} \int_{0}^{1} (1-x) \Delta \alpha_{had} \left(-\frac{x^{2}}{1-x} m_{\mu}^{2}\right) dx$$



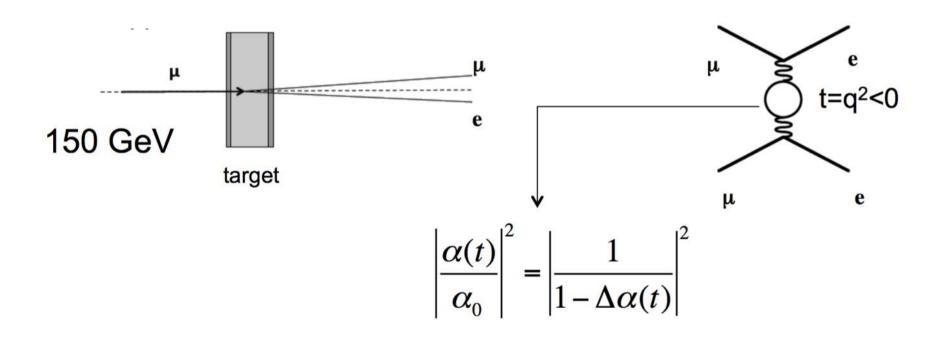
$$t = \frac{x^2 m_{\mu}^2}{x - 1} \quad 0 \le -t < +\infty$$

$$x = \frac{t}{2m_{\mu}^2} (1 - \sqrt{1 - \frac{4m_{\mu}^2}{t}}); \quad 0 \le x < 1;$$

- a_μ^{HLO} is given by the integral of the curve (smooth behaviour)
- It requires a measurement of the hadronic contribution to the effective electromagnetic coupling in the space-like region $\Delta\alpha_{had}(t)$ (t=q²<0)
- It enhances the contribution from low q² region (below 0.11 GeV²)
- Its precision is determined by the uncertainty on $\Delta\alpha_{had}$ (t) in this region

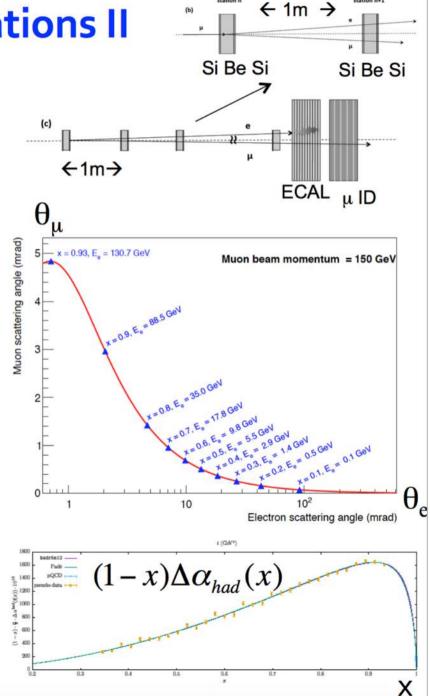


$\Delta\alpha_{had}(t)$ can be measured with high energy muon beam (E~150 GeV) on electron low-Z target, through the elastic scattering $\mu e \rightarrow \mu e$



Detector considerations II

- Modular apparatus: 20 layers of 3 cm Be (target), each coupled to 1 m distant Si (0.3 mm) planes. It provides a 0.02 mrad resolution on the scattering angle
- The t=q² <0 of the interaction is determined by the electron (or muon) scattering angle (a` la NA7)
- ECAL and μ Detector located downstream to solve PID ambiguity below 5 mrad. Above that, angular measurement gives correct PID
- It provides uniform full acceptance, with the potential to keep the systematic errors at 10⁻⁵ (main effect is the multiple scattering for normalization which can be studied by data)
- Statistical considerations show that a 0.3% error can be achieved on a_{μ}^{HLO} in 2 years of data taking with 2x10⁷ μ/s



- la misura puo' essere fatta al fascio di μ del Cern (M2) che fornisce $2*10^7 \,\mu/\text{sec}$
 - con il fascio dedicato, l'errore statistico della misura e'
 0.3% in 2 anni di presa dati
 - piccolo problema: sullo stesso fascio c'e' Compass!
- La parte piu' delicata e' la targhetta:
 - alto Z aumenta interazioni (=statistica) ma anche multiple scattering e background
 - 20 bersagli di Berillio (2cm) equipaggiato con Silicio per il tracking
 - puo' essere fatta una proof of concept con una stazione per verificare la fattibilita'

5 – antimateria, antiidrogeno







Probing Fundamental Physics with Antimatter

Stefan Ulmer (BASE) for the AD-Collaborations













Strategy and Physics Motivation

- Precise comparisons of the fundamental properties of simple baryonic matter/antimatter conjugates at low energy and with high precision.
- Such comparisons provide stringent tests of CPT invariance.
- Simple systems can be understood, and understanding provides sensitivity with respect to deviations.

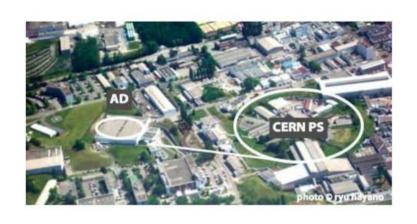
matter sector				
proton lifetime (direct)	>1.67 e34 y			
proton m	90 p.p.t.			
proton magn. moment	3.3 p.p.b.			
hydrogen 1S/2S	0.004 p.p.t.			
hydrogen GSHFS	0.7 p.p.t.			

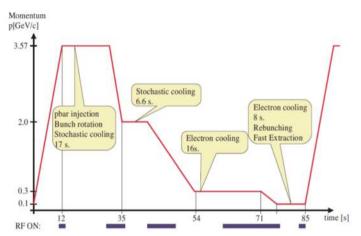
antimatter sector					
antiproton lifetime	>1.2 y				
antiproton m	90 p.p.t.				
antiproton m. moment	4.4 p.p.m.				
antihydrogen 1S/2S	?				
antihydrogen GSHFS	?				



ATRAP

Methods – the antiproton decelerator





AD provides: 30.000.000 antiprotons at 5.3 MeV each 120s

- Degrading / RFQ-D
- Catching
- Electron Cooling and kickout

work pioneered by the TRAP collaboration

Compression

resistive cooling -> particles at 8ueV

bridge 6 orders of magnitude within a few 100s

Catching: G. Gabrielse et al, PRL 57, 2504 (1986) Cooling: G. Gabrielse et al, PRL 63, 1360 (1989) Measurement: G. Gabrielse et al, PRL 65, 1317 (1990)



BSE

ATRAP

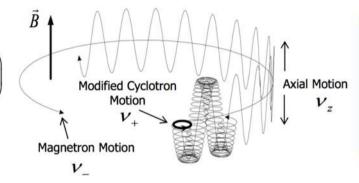
plasma di anti-protoni e positroni

Methods: Workhorse Penning Trap

radial confinement: $\vec{B} = B_0 \hat{z}$

axial confinement: $\Phi(\rho, z) = V_0 c_2 \left(z^2 - \frac{\rho^2}{2} \right)$





Measurements

 $\omega_c = \frac{e}{m_p} B$

charge-tomass ratios

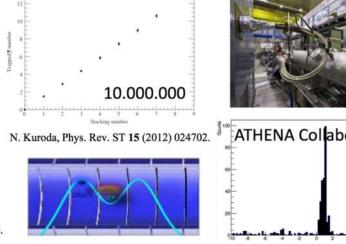
 $\omega_L = g \frac{e}{2m_p} B$

magnetic moments

 accumulation and compression of antiproton / positron plasmas

antihydrogen formation using a nested trap scheme.

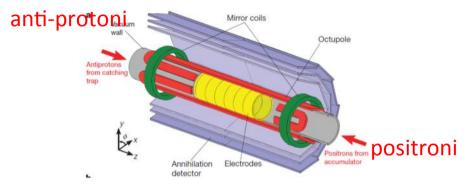
Proposed: G. Gabrielse, Phys. Lett. A **129** (1988) 38. First Demonstration: M. Amoretti, Nature **419** (2002) 456.





produzione di anti-idrogeno

Physics Measurements



- Antihydrogen 1S/2S laser spectroscopy in magnetic trap.
- Hydrogen in trap: 5 p.p.t.
- · Trapping via atomic magnetic moment, Challenge: shallow trap (0.5 K)

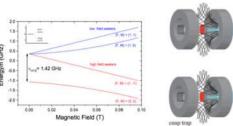
C. L. Cesar, Phys. Rev. Lett. 77 (1996) 255.

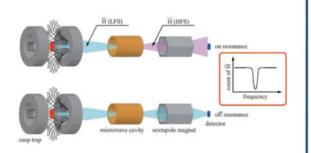
Antihydrogen beam spectroscopy (ASACUSA Scheme)

$$v_{\rm HF} = \frac{16}{3} \cdot \text{Ry} \cdot \alpha^2 c \cdot \left(\frac{1}{1 + \frac{m_e}{m_p}}\right)^3 \cdot \frac{m_e}{m_p} \cdot \frac{\mu_e}{\mu_B} \underbrace{\frac{\mu_p}{\mu_N}}_{\text{UN}} \cdot (1 + \underbrace{\delta_{str}}) + \delta_{QED})$$
<1ppb

- · extract antiproton magnetic moment
 - probe antiproton substructure
 - Test SME coefficients

Achieved 3.3 p.p.b. using hydrogen beam (Widmann Group in ASACUSA)





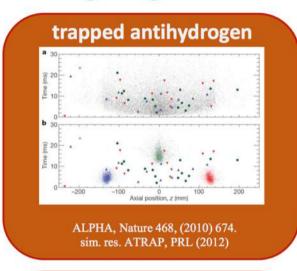
BSE

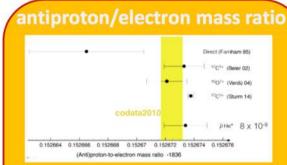
ATRAP

A. Mohri and Y. Yamazaki, Europhys. Lett. 63 (2003) 207.

risultati gia' prodotti

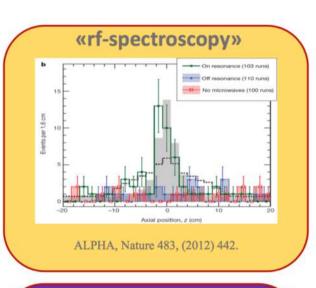
Highlights – CPT Sector

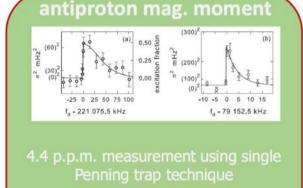




2-photon spectroscopy of anti-protonic helium / 8 p.p.b.

ASACUSA, Nature 484 (2011) 475





Antihydrogen Beam

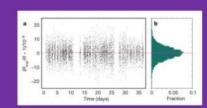
Table 1 | Summary of antihydrogen events detected by the antihydrogen detector.

	Scheme 1	Scheme 2	Background
Measurement time (s)	4,950	2,100	1,550
Double coincidence events, N _t	1,149	487	352
Events above the threshold			
(40 MeV), N _{>40}	99	29	6
Z-value (profile likelihood ratio) (σ)	5.0	3.2	_
Z-value (ratio of Poisson means) (σ)	4.8	3.0	-

Major step towards planned antihydrogen spectroscopy.

ASACUSA, Nature Comm. 3 (2014) 475.

antiproton/proton Q/M ratio



69 p.p.t. measurement using two particle Penning trap technique

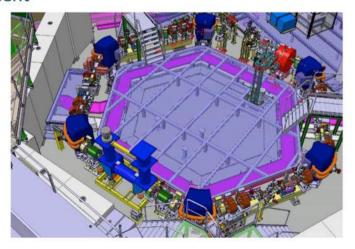
BASE, Nature 493 (2014) 502.

Per il futuro: modifica del sistema di rallentamento dei positroni e degli anti-protoni -> citato ma non descritto!

ELENA

- Antiprotons are caught in Penning traps using degraders – 99.9% of particles are lost.
- ELENA provides antiprotons decelerated to 100keV

 compared to the AD at improved beam
 emittance.
- Degrading at low particle energies is much more efficient



Experiment	ELENA Gain Factor
ALPHA	100
ATRAP	100
ASACUSA	10
AEgIS	100

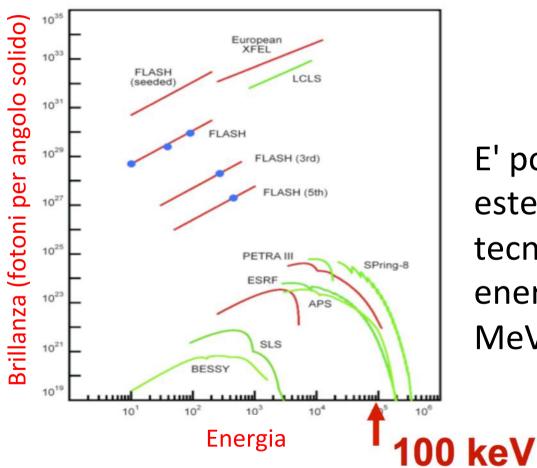
 ELENA will be able to deliver beams almost simultaneously to all experiments resulting in an essential gain in total beam time for each experiment. This also opens up the possibility to accommodate an extra experimental zone

Provides bright future perspective for antiproton-physics at CERN

6 – beam facilities

- gamma factory
 - Awake

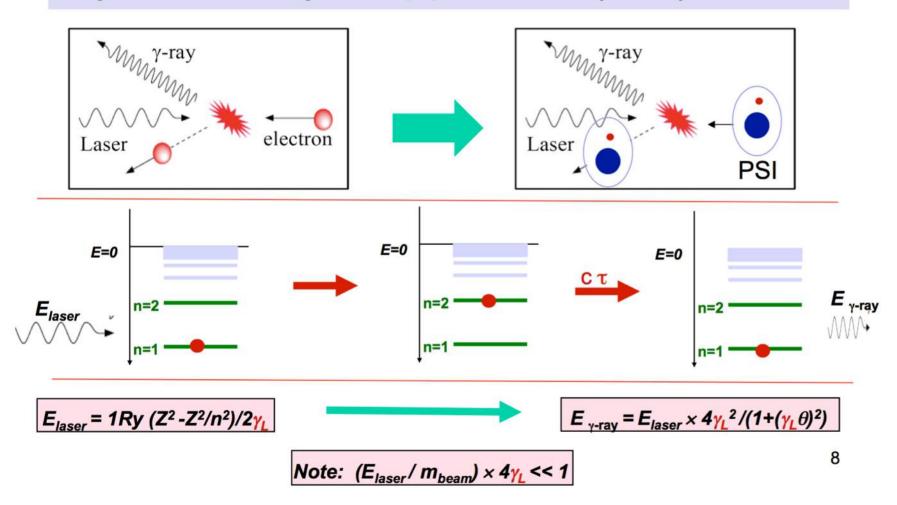
X-ray factories



E' possibile estendere questa tecnologia a energie di 1-100 MeV?

Energies up to ~ 100 keV

Simple Idea: replace an Electron beam by a Partially Stripped Ion (PSI) beam



"Tuning" dell'energia

$$v_i^{\text{max}} \longrightarrow (4 \gamma_L^2) v_i$$

 $\gamma_1 = E/M$ - Lorentz factor for the ion beam

The tuning of the PSI beam energy (SPS or LHC), the choice of the ion type, and the number of left electrons and of the laser type allows to tune the γ -ray energy in the requisite energy domain of 100 keV – 400 MeV.

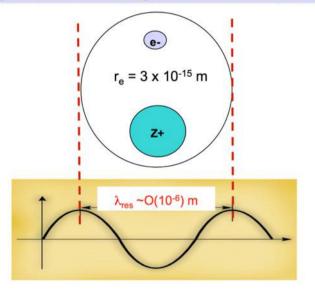
Example (maximal energy):

LHC, Pb⁸⁰⁺ ion, $\gamma_L = 2887$, n=1 \rightarrow 2, $\lambda_{laser} = 104.4$ nm, E_{γ} (max) = 396 MeV

Intensita' del fascio

The origin of the γ-beam intensity jump

(matching the size of the target to the laser wavelength)



Sezioni d'urto:

- elettroni

$$\sigma_e \approx r_e^2 \approx 10^{-25} cm^2$$

- PSI

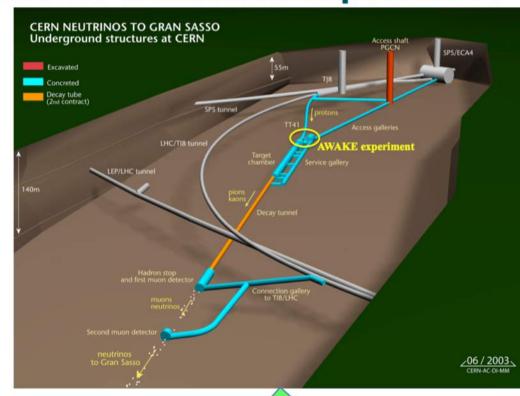
$$\sigma_{PSI} \approx \lambda_{res}^2 \approx 10^{-18} - 10^{-16} cm^2$$

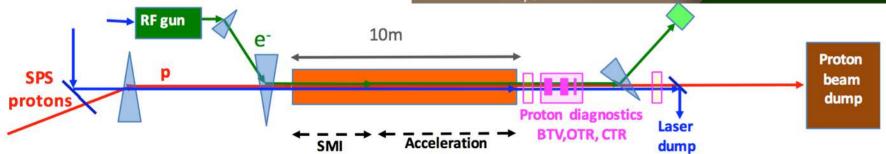
- brillanza aumentata di 7-9 ordini di grandezza

Fascio di elettroni O(50GeV) di alta purezza e intensita'

AWAKE: proton driven plasma wakefield experiment

- Demonstration experiment to show effect for first time and obtain GV/m gradients.
- Use 400 GeV SPS proton bunches with high charge.
- To start running this year and first phase to continue to LS2.
- Apply scheme to particle physics experiments leading to shorter or higher energy accelerators.





Physics behind AWAKE

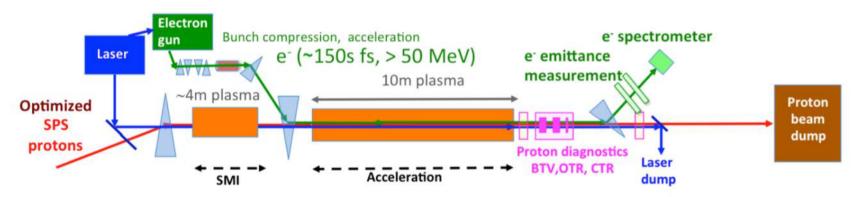
A plasma wakefield is a type of wave generated by particles travelling through a plasma. AWAKE will send proton beams through plasma cells to generate these fields. By harnessing wakefields, physicists may be able to produce accelerator gradients hundreds of times higher than those achieved in current radiofrequency cavities. This would allow future colliders to achieve higher energies over shorter distances than is possible today.

Usando protoni dell'SPS → elettroni di ~50 GeV Usando protoni di LHC → elettroni di ~2 TeV

Test attuali svolti con SPS

AWAKE Run 2

- Preparing AWAKE Run 2, after LS2 and before LS3.
 - Accelerate electron bunch to higher energies.
 - Demonstrate beam quality preservation.
 - Demonstrate scalability of plasma sources.



Preliminary Run 2 electron beam parameters

Parameter	Value	
Acc. gradient	>0.5 GV/m	
Energy gain	10 GeV	
Injection energy	$\gtrsim 50 \text{ MeV}$	
Bunch length, rms	40–60 μm (120–180 fs)	
Peak current	200–400 A	
Bunch charge	67–200 pC	
Final energy spread, rms	few %	
Final emittance	$\lesssim 10 \ \mu \text{m}$	

- Are there physics experiments that require an electron beam of up to O(50 GeV)?
- Use bunches from SPS with 3.5 × 10¹¹ protons every ~ 5 s.
- Using the LHC beam as a driver, TeV electron beams are possible.

E. Adli (AWAKE Collaboration), IPAC 2016 proceedings, p.2557 (WEPMY008).



Possible physics experiments

- Use of electron beam for test-beam campaigns.
 - Test-beam infrastructure for detector characterisation often over-subscribed
 - Also accelerator test facility.
 - Variation of energy.
 - Provide pure electron beam.
- Fixed-target experiments using electron beams, e.g. deep inelastic scattering.
 - Measurements at high x with higher statistics than previous experiments.
 - Polarised beams and spin structure of the nucleon.

Search for dark photons à la NA64

- Consider beam-dump and counting experiments.
- High energy electron-proton collider
 - A low-luminosity LHeC-type experiment.
 - A very high energy electron-proton collider.

This is not a definitive list, but a quick brainstorm and people are invited to think of other possible uses/applications/experiments.

7 – esperimenti non agli acceleratori

ALPS
IAXO (seguito di CAST)

KWISP

DARK-SIDE

Axion and axion-like-particle (ALP) couplings

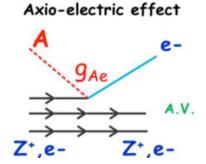
Axion and other Nambu-Goldstone bosons arising from spontaneous breakdown of global symmetries are theoretically well-motivated very weakly interacting slim (ultra-light) particles.

The coefficients are determined by specific ultraviolet extension of SM.

$$\mathcal{L} \supset -\frac{\alpha_s}{8\pi} \frac{C_{ag}}{f_a} a G^b_{\mu\nu} \tilde{G}^{b,\mu\nu} - \frac{\alpha}{8\pi} \frac{C_{a\gamma}}{f_a} a F_{\mu\nu} \tilde{F}^{\mu\nu} + \frac{1}{2} \frac{C_{af}}{f_a} \partial_{\mu} a \overline{\psi}_f \gamma^{\mu} \gamma_5 \psi_f$$

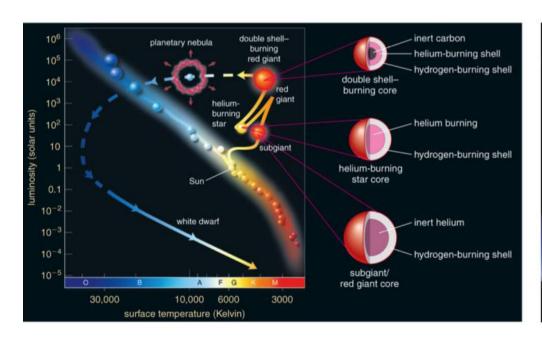
$$a - - - \mathcal{C}^{\mu\nu} g \qquad a - - - \mathcal{C}^{\mu\nu} \gamma_5 \psi_f$$

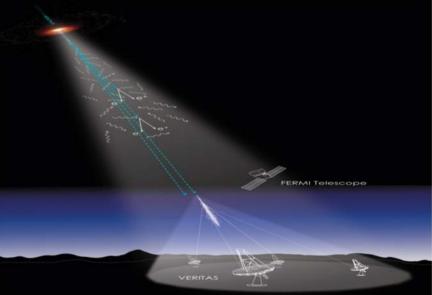




Astro-Hints on Axion/ALPs?

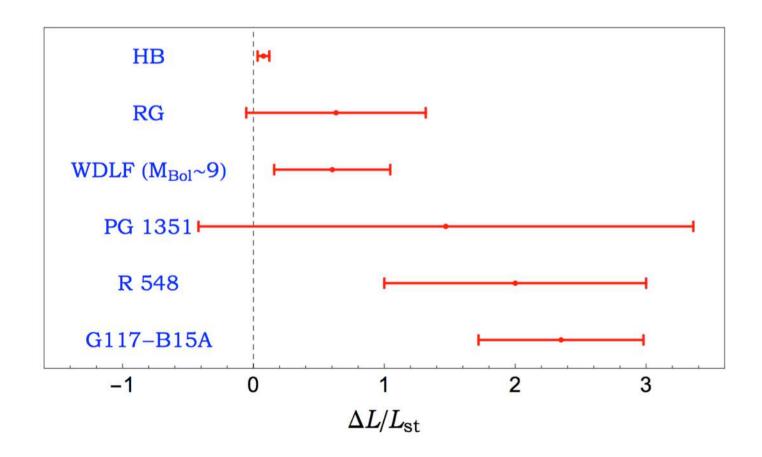
- Modest hints for
 - excessive energy losses of stars in various evolutionary stages,
 - photon-ALP conversion in astrophysical magnetic fields





Hints of Axion/ALP Energy Losses of Stars?

Excessive energy losses of HBs, RGs and WDs:

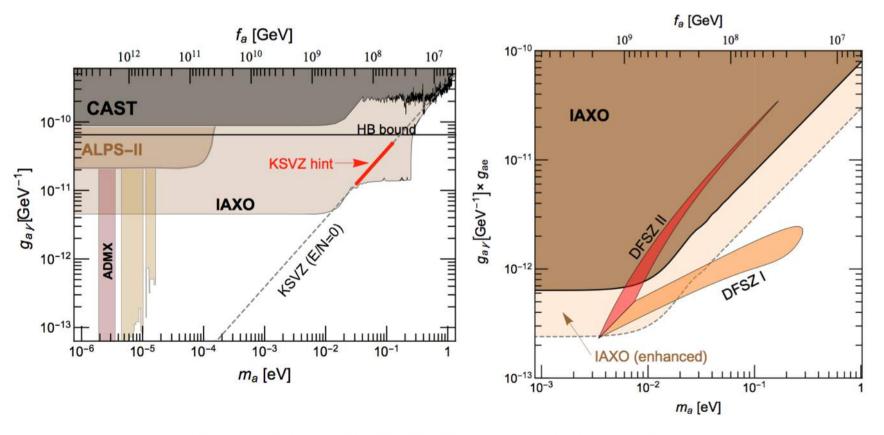


[Giannotti,Irastorza,Redondo,AR, arXiv:1512.08108]



Hints of Axion/ALP Energy Losses of Stars?

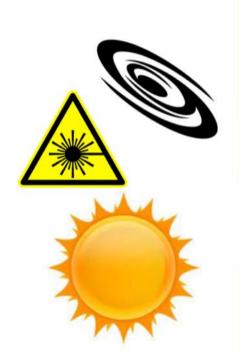
Excessive energy losses of HBs, RGs and WDs can all be explained at one stroke by production and emission of axion/ALP with coupling both to electrons and photons, e.g. in specific axion models:



[Giannotti,Irastorza,Redondo,AR, in preparation]



Tecniche di ricerca degli assioni

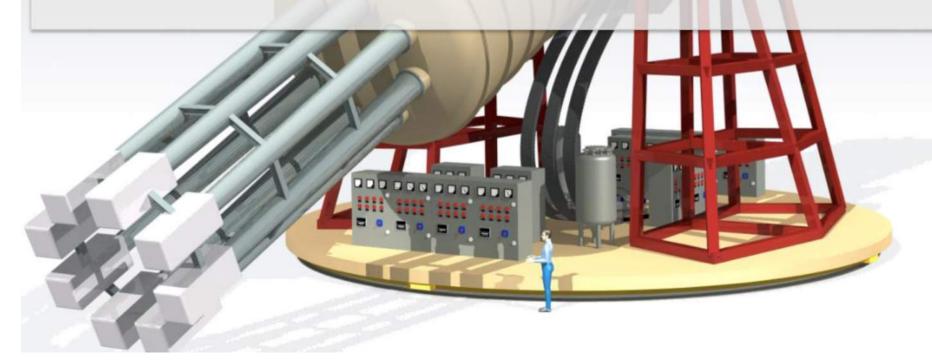


Source	Experiments	Model & Cosmology dependency	Technology
Relic axions	ADMX, ADMX-HF, Casper, CAPP,	High	New ideas emerging, Active R&D going on,
Lab axions	ALPS, OSQAR, fifth force exps,	Very low	
Solar axions	SUMICO, CAST,	Low	Ready for large scale experiment

- Helioscopes → do not rely on the axion being the dominant DM component. Solar axion emission robust prediction
- Helioscopes → No R&D needed. Technology mature enough for a large scale experiment (IAXO)
- Large complementarity with other detection strategies

Physics case, prospects and status of the International AXion Observatory IAXO

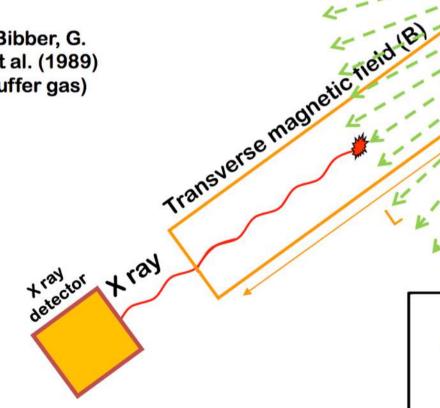
Igor G. Irastorza (U. Zaragoza) on behalf of the IAXO collaboration Beyond Colliders, CERN, September 7th 2016



Helioscopes

Axion helioscope concept **P. Sikivie, 1983**

+ K. van Bibber, G. Raffelt, et al. (1989) (use of buffer gas)



 $P_{a\gamma} = 2.6 \times 10^{-17} \left(\frac{B}{10 \text{ T}}\right)^2 \left(\frac{L}{10 \text{ m}}\right)^2$ $(g_{a\gamma} \times 10^{10} \text{ GeV})^2 \mathcal{F}$

CAST experiment @ CERN

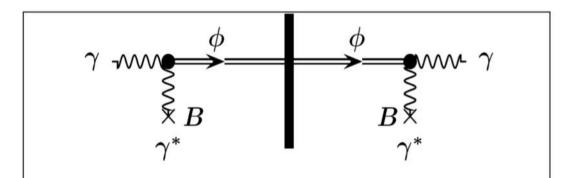
- Decommissioned LHC test magnet (L=10m, B=9 T)
- Moving platform ±8°V ±40°H (to allow up to 50 days / year of alignment)
- 4 magnet bores to look for X rays
- 2 X ray telescopes to increase signal/noise ratio.



How to look for ALPs ...

- ... by exploiting their coupling to photons:
- WISPs pass any barrier and could make

light-shining-through-a-wall.

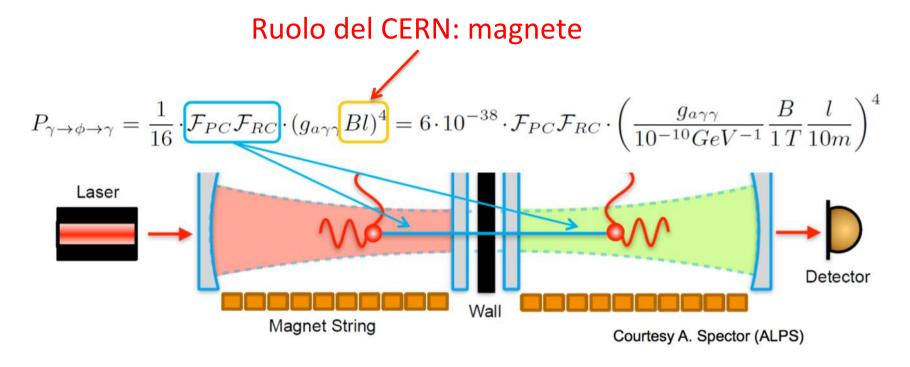


Okun 1982, Skivie 1983, Ansel'm 1985, Van Bibber et al. 1987

Axions, axion-like particles

$$P_{\gamma \to \phi}(B,\ell,q) = \frac{1}{4} \left(g \, B \, \ell \right)^2 \, F(q\ell) \qquad F(q\ell) = \left[\frac{\sin\left(\frac{1}{2}q\ell\right)}{\frac{1}{2}q\ell} \right]^2 \qquad \textit{q = p_{\gamma} - p_{\phi}}$$
 I: length of B field





- Use long strings of strong dipole magnets.
- Implement optical resonators to recycle the light before the wall and to boost the re-conversion of ALPs to photons behind the wall.
- The maximal length is given by the aperture of the magnets and the beam shape of the resonating mode in the optical resonators.
- The maximal circulating light power before the wall is given by the damage threshold of the mirrors and the beam spot size.

To-take-home

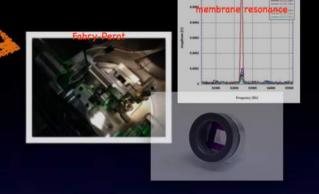
- Accelerator dipole magnets under development at CERN could be the basis of an "ultimate" light-shining-through-a-wall experiment mainly searching for axion-like particles.
- The costs of such an experiment would be dominated by the costs of the magnet strings.
 - The costs probably strongly correlate with decisions on future accelerator projects.
- Considering the development of the physics case for axion-like particles in the recent years, experiments have to move from the niche "make best use of existing equipment built for something else" to more dedicated installations:
 - IAXO and "ALPS III".

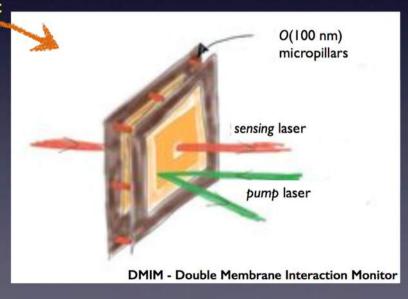
- studi odi forze a short range (~verifica di 1/r^2 a distanze ~110-100nm)
- non mi e' chiaro cosa c'entri il Cern



aKWISP in a nutshell

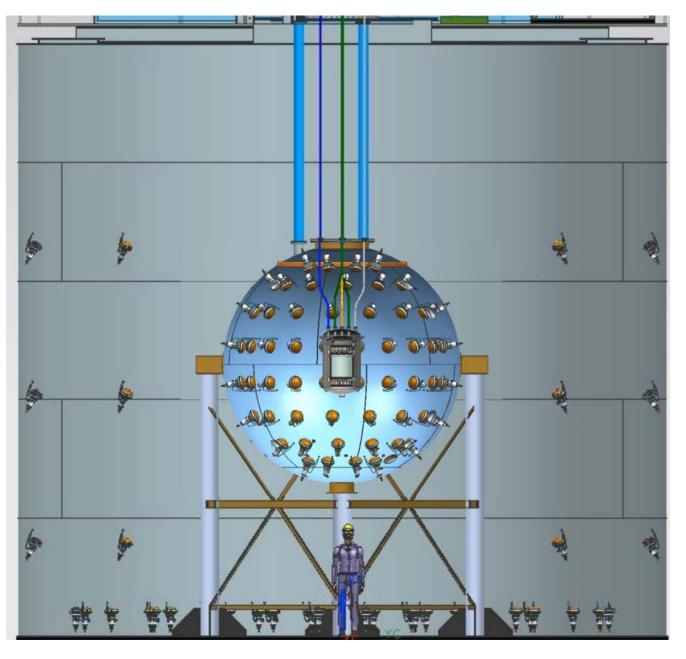
- Build on the KWISP (*) force sensor core apparatus
 - membrane-based optomechanical force sensor
 - sensitivity enhanced by the combined quality factors of two resonators: mechanical (membrane) and optical (FP)
 - sensitive to extremely tiny forces and subnuclear size displacements
- Introduce the double-membrane device concept
 - two membranes separated by O(10 nm)-size micropillars acting as sensing and source masses
 - different Q's and resonant frequencies
- Implement advanced technologies to achieve the ultimate sensitivity
 - homodyne detection
 - membrane optimisation
 - cryogenic cooling
- Study Short Range Interactions





10 m Height
11 m Diameter
1,000 Tonnes
Vater Cherenkov
Muon Veto

Ruolo del Cern: costruire presso il Cern un prototipo di 1 tonnellata



Kickoff - general points

- Four sessions
 - Theorists' motivations, ideas and wishes
 - Accelerator and infrastructure opportunities
 - Potential future of existing programs
 - New experiment ideas
- Avoided a top-down approach
- With the aim of getting ideas on the table
- No commitment at this stage! Goal is to provide input to European Strategy Update (ESU)

Conclusions

- Considerable potential for novel exploitation of existing complex
 - Cleary need to evaluate, prioritize, and commit resources where appropriate
- Some compelling new ideas at varying degrees of maturity
 - Evaluate, select, develop, report
 - Continue exploration of new ideas
- Possible collaboration with other labs to be explored

Foresee PBC follow-up meeting next year