



ECL forward for the BelleII experiment at SuperKEKB

OUTLINE:

- Introduction
- R&D on pure CsI crystals + Front End readout
- future and plans





FORWARD Calorimeter of BelleII experiment: 1150 crystals of CsI(Tl) to be replaced with pure CsI:

- same dimensions of CsI(Tl) \rightarrow mechanics
- fast crystal (30 ns decay time) to avoid pile-up and occupancy
- rad-hardness good (report presented at the first JENNIFER General meeting)
- very low light yield

R&D on readout of the crystals is fundamental for the project to obtain a good energy resolution at relatively low energy with very low light output







- APD is: small good for the mechanics
 - low gain
 - S/N ratio deeply studied during R&D



- Pure CsI: low light yield @310nm where the Q.E. (Quantum Efficiency) of the APD is low (APD has max Q.E. around 420nm)
 - very fast prompt emission of the light (30ns) 🔶
 - slow component of the light up to lus

GOAL: eliminate the slow component of the light to avoid pile-up between physics events and background from machine → use a filter to select only the fast component → use a WLS to shift the fast emitted light in the 420nm region to match the APD window C. Cecchi 3







Resolution







INFN

August 2015: 6 pure CsI crystals with phototube readout have been installed around the Interaction Point (IP) for background studies.







Performance studies

First result: material budget Negligible effect from material budget on resolution

Second result: background impact Large impact, as expected, on resolution

Third result: ENE (Equivalent Noise Energy) Important impact on resolution mainly in the BKG configuration Fourth result: photostatistic fluctuations (low LY of the pure CsI)

Large impact on resolution





Summary of the different configurations

Study the resolution @100MeV

CsI(Tl) 12% actual calorimeter, this is the starting point.....

Nominal Background:

- 1st configuration (PESSIMISTIC)
- ENE 1.3 MeV
- Nphe/MeV (2 APD) 12.5

 $\sigma(E)/E(\%) = 6\%$

2nd configuration (REALISTIC)

- ENE 1.3 MeV
- Nphe/MeV (2APD) 25

 $\sigma(E)/E(\%) = 4.8\%$



Performance studies





PG: ENE 0.7 MeV and 25 ph-e/MeV LNF: ENE 1.3 MeV and 6 ph-e/MeV





Performance studies



-Nominal background (BG =1) degrade the resolution by a factor 10%

- With BG=1 ENE + photostatistic fluctuations degrade resolution of a factor 20% each

- With current simulation \rightarrow energy resolution @100 MeV about 5% (to be compared to 12% CsI(Tl))





Planning

SuperKEKB Commissioning Phases

Y. Funakoshi at June '16 B2GM

2016 Feb. ~ June

Phase 1 w/o QCS and Belle II

basic machine tuning vacuum scrubbing Optics tuning BKG study ~2017 Autumn (~ 5month)

Phase 2 w/QCS and Belle II w/o Vertex detector

DR commissioning BKG study Luminosity tuning Target luminosity: 1 x 10³⁴ cm⁻² s⁻¹ ~2018 Autumn

Phase 3 w/ full Belle II

Physics Run Luminosity tuning







Due to the delay of the SuperKEKB operations (about one year) the decision on the FWD EMC upgrade has also been delayed

- a task force has been settled (C. Cecchi co-chair) to finish the R&D and take the final decision

JENNIFER DELIVERABLES:

- FWD ECL TDR 10 months (Feb. 2016) \rightarrow R&D report March 2017, 24 months

- Commissioning report FWD ECL (48 months) \rightarrow TDR when Background studies will be available as they are crucial







• R&D is going on and is approaching the final steps

• Filter + WLS on pure CsI give good results \rightarrow ECL FWD detector is not affected by presence of BKG a slightly loss of signal is observed (74% of the light is retained after shaping)

• CsI(Tl) + APD shows good level of noise \rightarrow to be tested at more stable values of APD Gain

• Performance studies are ongoing \rightarrow pure CsI is better is presence of BKG than CsI(Tl) with the present level of BKG and with the actual clustering \rightarrow clustering is a key ingredient for these studies and to understand how CsI(Tl) can be improved

• Besides the delay non directly coming from the ECL group the ECL FWD Task Force will decide on the upgrade max in 1 year from now and there is still a slot of opportunity for the installation and commissioning of the detector which will have sensible impact on the physics results.