Radiative corrections in e⁺e⁻ collisions with the **BabaYaga** event generator

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International Symposium Advances in Dark Matter and Particle Physics

University of Messina Messina, October 24-27, 2016

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- ⋆ Motivations for precise luminometry
- ⋆ QED processes & radiative corrections
- * The **BabaYaga** and **BabaYaga@NLO** event generators
 - theoretical framework
 - improving theoretical accuracy: QED Parton Shower and matching with NLO corrections
- ⋆ Results, tuned comparisons, theoretical accuracy
- ⋆ Dark-photon searches through radiative return method
- Conclusions

Relevant references

★ Website

http://www.pv.infn.it/hepcomplex/babayaga.html

- *** BabaYaga** main references:
 - Barzè et al., Eur. Phys. J. C 71 (2011) 1680BabaYaga with dark photon• Balossini et al., Phys. Lett. 663 (2008) 209BabaYaga@NLO for $e^+e^- \rightarrow \gamma\gamma$ Balossini et al., Nucl. Phys. B758 (2006) 227BabaYaga@NLO for Bhabha• C.M.C.C. et al., Nucl. Phys. Proc. Suppl. 131 (2004) 48BabaYaga@NLO• C.M.C.C., Phys. Lett. B 520 (2001) 16improved PS BabaYaga• C.M.C.C. et al., Nucl. Phys. B 584 (2000) 459BabaYaga
- ★ Related work:
 - · S. Actis et al.

"Quest for precision in hadronic cross sections at low energy: Monte Carlo tools vs. experimental data", Eur. Phys. J. C **66** (2010) 585 Report of the Working Group on Radiative Corrections and Monte Carlo Generators for Low Energies

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 C.M.C.C. et al., JHEP 1107 (2011) 126 NNLO massive pair corrections

Why precision luminosity generators?

- Precision measurements require a precise knowledge of the machine luminosity
- *e.g.*, the measurement of the R(s) ratio is a key ingredient for the predictions of $a_{\mu} = (g_{\mu} 2)/2$ and $\Delta \alpha_{had}(M_Z)$ and in turn for SM precision tests



$$a_{\mu} = \frac{\alpha^2}{3\pi^2} \int_{m_{\pi}^2}^{\infty} \mathrm{d}s \, K(s) \frac{R(s)}{s} \qquad \Delta \alpha_{\mathrm{had}}^{(5)}(M_Z^2) = -\frac{\alpha M_Z^2}{3\pi} \mathsf{Re} \int_{m_{\pi}^2}^{\infty} \frac{R(s) \mathrm{d}s}{s(s - M_Z^2 - i\epsilon)}$$

Reference processes for luminosity

 Instead of getting the luminosity from machine parameters, it's more effective to exploit the relation

$$\sigma = \frac{N}{L} \quad \rightarrow \quad L = \frac{N_{\text{ref}}}{\sigma_{\text{theory}}} \qquad \frac{\delta L}{L} = \frac{\delta N_{\text{ref}}}{N_{\text{ref}}} \oplus \frac{\delta \sigma_{\text{theory}}}{\sigma_{\text{theory}}}$$

- Normalization processes are required to have a clean topology, high statistics and be calculable with high theoretical accuracy
- ★ Large-angle QED processes as $e^+e^- \rightarrow e^+e^-$ (Bhabha), $e^+e^- \rightarrow \gamma\gamma$, $e^+e^- \rightarrow \mu^+\mu^-$ are golden processes at flavour factories to achieve a typical precision at the level of $1 \div 0.1\%$

→ QED RC corrections are mandatory

→ BabaYaga has been developed for high-precision simulation of QED processes at flavour factories (primarily for luminosity determination)

Theory of QED corrections into MC generators

* The most precise MC generators include exact $\mathcal{O}(\alpha)$ (NLO) photonic corrections matched with higher–order leading logarithmic contributions

[+ vacuum polarization, using a data driven routine for the calculation of the non–perturbative $\Delta \alpha_{\rm had}^{(5)}(q^2)$ hadronic contribution]

- Common methods used to account for multiple photon corrections are the analytical collinear QED Structure Functions (SF), YFS exponentiation and QED Parton Shower (PS)
- The QED PS [implemented in BabaYaga/BabaYaga(NLO] is an exact MC solution of the QED DGLAP equation for the electron SF $D(x,Q^2)$

$$Q^2 \frac{\partial}{\partial Q^2} D(x, Q^2) = \frac{\alpha}{2\pi} \int_x^1 \frac{dt}{t} P_+(t) D(\frac{x}{t}, Q^2)$$

- The PS solution can be cast into the form
 - $D(x,Q^2) = \Pi(Q^2) \sum_{n=0}^{\infty} \int \frac{\delta(x-x_1 \cdots x_n)}{n!} \prod_{i=0}^{n} \left[\frac{\alpha}{2\pi} P(x_i) \ L \ dx_i \right]$
 - $\rightarrow \Pi(Q^2) \equiv e^{-\frac{\alpha}{2\pi}LI_+} \text{ Sudakov form factor, } I_+ \equiv \int_0^{1-\epsilon} P(x)dx, L \equiv \ln Q^2/m^2 \text{ collinear log,} \\ \epsilon \text{ soft-hard separator and } Q^2 \text{ virtuality scale}$
 - \rightarrow the kinematics of the photon emissions can be recovered \rightarrow exclusive photons generation
- · The accuracy is improved by matching exact NLO with higher-order leading log corrections
 - * theoretical error starts at $\mathcal{O}(\alpha^2)$ (NNLO) QED corrections, for all QED channels [Bhabha, $\gamma\gamma$ and $\mu^+\mu^-$]

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Pictorially, the corrections to the LO cross section can be arranged as (collinear log $L\equiv\log\frac{s}{m_e^2})$

$$\begin{array}{c|c} \mathsf{LO} & \alpha^{0} \\ \mathsf{NLO} & \alpha L & \alpha \\ \mathsf{NNLO} & \frac{1}{2}\alpha^{2}L^{2} & \frac{1}{2}\alpha^{2}L & \frac{1}{2}\alpha^{2} \\ \mathsf{h.o.} & \sum_{n=3}^{\infty} \frac{\alpha^{n}}{n!}L^{n} & \sum_{n=3}^{\infty} \frac{\alpha^{n}}{n!}L^{n-1} & \cdots \end{array}$$

Blue: Leading-Log PS, Leading-Log YFS, SF

Pictorially, the corrections to the LO cross section can be arranged as (collinear log $L\equiv\log\frac{s}{m_e^2}$)

Red: matched PS, YFS, SF + NLO

Matching NLO and PS in BabaYaga@NLO

Exact $\mathcal{O}(\alpha)$ (NLO) soft+virtual (*SV*) corrections and hard bremsstrahlung (*H*) matrix elements can be combined with QED PS *via* a matching procedure

•
$$d\sigma_{\underline{LL}}^{\infty} = \Pi(Q^2, \varepsilon) \sum_{n=0}^{\infty} \frac{1}{n!} |\mathcal{M}_{n,LL}|^2 d\Phi_n$$

- $d\sigma_{LL}^{\alpha} = [1 + C_{\alpha,LL}] |\mathcal{M}_0|^2 d\Phi_0 + |\mathcal{M}_{1,LL}|^2 d\Phi_1 \equiv d\sigma_{LL}^{SV}(\varepsilon) + d\sigma_{LL}^H(\varepsilon)$
- $d\sigma_{\mathrm{NLO}}^{\alpha} = [1 + C_{\alpha}] |\mathcal{M}_0|^2 d\Phi_0 + |\mathcal{M}_1|^2 d\Phi_1 \equiv d\sigma_{\mathrm{NLO}}^{SV}(\varepsilon) + d\sigma_{\mathrm{NLO}}^H(\varepsilon)$

•
$$F_{SV} = 1 + (C_{\alpha} - C_{\alpha,LL})$$
 $F_H = 1 + \frac{|\mathcal{M}_1|^2 - |\mathcal{M}_{1,LL}|^2}{|\mathcal{M}_{1,LL}|^2}$

$$d\sigma_{\text{matched}}^{\infty} = F_{SV} \Pi(Q^2, \varepsilon) \sum_{n=0}^{\infty} \frac{1}{n!} \left(\prod_{i=0}^{n} F_{H,i} \right) |\mathcal{M}_{n,LL}|^2 d\Phi_n$$

 $d\Phi_n$ is the exact phase space for n+2 final-state particles

Matching NLO and PS in BabaYaga@NLO

- F_{SV} and $F_{H,i}$ are infrared/collinear safe and account for missing $O(\alpha)$ non-logs, avoiding double counting of LL
- $\left[\sigma_{matched}^{\infty}\right]_{\mathcal{O}(\alpha)} = \sigma_{\mathrm{NLO}}^{\alpha}$
- resummation of higher orders LL contributions is preserved
- the cross section is still fully differential in the momenta of the final state particles (e^+ , e^- and $n\gamma$)
- as a by-product, part of photonic $\alpha^2 L$ included by means of terms of the type $F_{SV \mid H,i} \times LL$

G. Montagna et al., PLB 385 (1996)

• the theoretical error is shifted to $\mathcal{O}(\alpha^2)$ (NNLO, 2 loop) not infrared terms: very naively and roughly (for photonic corrections)

$$\frac{1}{2}\alpha^2 L \equiv \frac{1}{2}\alpha^2 \log \frac{s}{m_e^2} \sim 5 \times 10^{-4}$$

 to show the typical size of RC, the following setups and definitions are used (for Bhabha)

a
$$\sqrt{s} = 1.02 \text{ GeV}, E_{min} = 0.408 \text{ GeV}, 20^{\circ} < \theta_{\pm} < 160^{\circ}, \xi_{max} = 10^{\circ}$$

b $\sqrt{s} = 1.02 \text{ GeV}, E_{min} = 0.408 \text{ GeV}, 55^{\circ} < \theta_{\pm} < 125^{\circ}, \xi_{max} = 10^{\circ}$
c $\sqrt{s} = 10 \text{ GeV}, E_{min} = 4 \text{ GeV}, 20^{\circ} < \theta_{\pm} < 160^{\circ}, \xi_{max} = 10^{\circ}$
d $\sqrt{s} = 10 \text{ GeV}, E_{min} = 4 \text{ GeV}, 55^{\circ} < \theta_{\pm} < 125^{\circ}, \xi_{max} = 10^{\circ}$

$$\delta_{VP} \equiv \frac{\sigma_{0,VP} - \sigma_{0}}{\sigma_{0}} \qquad \qquad \delta_{\alpha} \equiv \frac{\sigma_{\alpha}^{NLO} - \sigma_{0}}{\sigma_{0}}$$
$$\delta_{HO} \equiv \frac{\sigma_{matched}^{PS} - \sigma_{\alpha}^{NLO}}{\sigma_{0}} \qquad \qquad \delta_{HO}^{PS} \equiv \frac{\sigma_{\alpha}^{PS} - \sigma_{\alpha}^{PS}}{\sigma_{0}}$$
$$\delta_{\alpha}^{non-log} \equiv \frac{\sigma_{\alpha}^{NLO} - \sigma_{\alpha}^{PS}}{\sigma_{0}} \qquad \qquad \delta_{\infty}^{non-log} \equiv \frac{\sigma_{matched}^{PS} - \sigma_{\alpha}^{PS}}{\sigma_{0}}$$

setup	(a)	(b)	(C)	(d)
δ_{VP}	1.76	2.49	4.81	6.41
δ_{lpha}	-11.61	-14.72	-16.03	-19.57
δ_{HO}	0.39	0.82	0.73	1.44
δ^{PS}_{HO}	0.35	0.74	0.68	1.34
$\delta^{non-log}_{\alpha}$	-0.34	-0.56	-0.34	-0.56
$\delta^{non-log}_\infty$	-0.30	-0.49	-0.29	-0.46

Table: Relative corrections (in per cent) to the Bhabha cross section for the four setups

- * in short, the fact that $\delta_{\alpha}^{non-log} \simeq \delta_{\infty}^{non-log}$ and $\delta_{HO} \simeq \delta_{HO}^{PS}$ means that the matching algorithm preserves both the advantages of exact NLO calculation and PS approach:
 - \rightarrow it includes the missing NLO RC to the PS
 - \rightarrow it adds the missing higher-order RC to the NLO

Estimating the theoretical accuracy

S. Actis et al. Eur. Phys. J. C 66 (2010) 585

- It is extremely important to compare independent calculations/implementations/codes, in order to
 - * asses the technical precision, spot bugs (with the same th. ingredients)
 - estimate the theoretical "error" when including partial/incomplete higher-order corrections
- A number of generators are available, some of them including QED h.o. and NLO corrections according to different approaches (collinear SF + NLO, YFS exponentiation,...)

Generator	Processes	Theory Accuracy		Web address
BHAGENF/BKQED	$e^+e^-/\gamma\gamma, \mu^+\mu^-$	$\mathcal{O}(\alpha)$	1%	www.lnf.infn.it/~graziano/bhagenf/bhabha.html
BabaYaga v3.5	$e^+e^-, \gamma\gamma, \mu^+\mu^-$	Parton Shower	~ 0.5%	www.pv.infn.it/~hepcomplex/babayaga.html
BabaYaga@NLO	$e^+e^-, \gamma\gamma, \mu^+\mu^-$	$\mathcal{O}(\alpha) + PS$	$\sim 0.1\%$	www.pv.infn.it/~hepcomplex/babayaga.html
BHWIDE	e^+e^-	$\mathcal{O}(\alpha)$ YFS	0.5%(LEP1)	placzek.home.cern.ch/placzek/bhwide
MCGP J	$e^+e^-, \gamma\gamma, \mu^+\mu^-$	$\mathcal{O}(\alpha) + SF$	< 0.2%	cmd.inp.nsk.su/~sibid

Large angle Bhabha: tuned comparisons & technical precision

Without vacuum polarization, to compare QED corrections consistently

At the Φ and τ -charm factories (cross sections in nb)

By BabaYaga group, Ping Wang and A. Sibidanov

setup	BabaYaga@NLO	BHWIDE	MCGPJ	$\delta(\%)$
$\sqrt{s} = 1.02 \text{ GeV}, 20^{\circ} \le \vartheta_{\mp} \le 160^{\circ}$	6086.6(1)	6086.3(2)	—	0.005
$\sqrt{s} = 1.02 \text{ GeV}, 55^{\circ} \le \vartheta_{\mp} \le 125^{\circ}$	455.85(1)	455.73(1)	—	0.030
$\sqrt{s} = 3.5 \text{ GeV}, \vartheta_+ + \vartheta \pi \le 0.25 \text{ rad}$	35.20(2)		35.181(5)	0.050

 \star Agreement well below 0.1%! \star

At BaBar (cross sections in nb)

By A. Hafner and A. Denig

angular acceptance cuts	BabaYaga@NLO	BHWIDE	$\delta(\%)$
$15^{\circ} \div 165^{\circ}$	119.5(1)	119.53(8)	0.025
$40^{\circ} \div 140^{\circ}$	11.67(3)	11.660(8)	0.086
$50^{\circ} \div 130^{\circ}$	6.31(3)	6.289(4)	0.332
$60^{\circ} \div 120^{\circ}$	3.554(6)	3.549(3)	0.141

 \star Agreement at the \sim 0.1% level! \star

Theoretical accuracy, comparisons with NNLO

- NLO RC being included, the theoretical error starts at $\mathcal{O}(\alpha^2)$ (NNLO) \hookrightarrow anyway large NNLO RC already included by h.o. exponentiation and
 - by $\mathcal{O}(\alpha)$ LL × finite-NLO
- The full set of NNLO QED corrections to Bhabha scattering has been calculated in the last years
- BabaYaga formulae can be truncated at $\mathcal{O}(\alpha^2)$ to be consistently and systematically compared with all the classes of NNLO corrections

$$\sigma^{\alpha^2} = \sigma^{\alpha^2}_{\rm SV} + \sigma^{\alpha^2}_{\rm SV,H} + \sigma^{\alpha^2}_{\rm HH}$$

- $\sigma_{\rm SV}^{\alpha^2}$: soft+virtual photonic corrections up to ${\cal O}(\alpha^2)$
 - \mapsto compared with the corresponding available NNLO QED calculation
- $\sigma_{SV,H}^{\alpha^2}$: one-loop soft+virtual corrections to single hard bremsstrahlung \mapsto presently estimated relying on existing (partial) results
- $\sigma_{\rm HH}^{\alpha^2}$: double hard bremsstrahlung
 - \mapsto compared with the exact $e^+e^- \rightarrow e^+e^-\gamma\gamma$ cross section, to register really negligible differences (at the 1×10^{-5} level)

NNLO Bhabha calculations

• Photonic corrections A. Penin, PRL 95 (2005) 010408 & Nucl. Phys. B734 (2006) 185



Electron loop corrections

R. Bonciani *et al.*, Nucl. Phys. **B701** (2004) 121 & Nucl. Phys. **B716** (2005) 280
 S. Actis, M. Czakon, J. Gluza and T. Riemann, Nucl. Phys. **B786** (2007) 26



NNLO Bhabha calculations

· Heavy fermion and hadronic loops

R. Bonciani, A. Ferroglia and A. Penin, PRL 100 (2008) 131601 S. Actis, M. Czakon, J. Gluza and T. Riemann, PRL 100 (2008) 131602

J.H. Kühn and S. Uccirati, Nucl. Phys. B806 (2009) 300



One-loop soft+virtual corrections to single hard bremsstrahlung

S. Actis, P. Mastrolia and G. Ossola, Phys. Lett. B682 (2010) 419

Comparison with NNLO calculation for $\sigma_{SV}^{\alpha^2}$

Using realistic cuts for luminosity @ KLOE

Comparison of $\sigma_{SV}^{\alpha^2}$ calculation of BabaYaga@NLO with

 Penin (photonic): function of the logarithm of the soft photon cut-off (left plot) and a fictitious electron mass (right plot)



- ★ differences are infrared safe, as expected
- $\star~\delta\sigma(\mathrm{photonic})/\sigma_0\propto lpha^2 L$, as expected
- Numerically, for various selection criteria at the Φ and B factories

$$\sigma^{lpha^2}_{
m SV}({
m Penin}) - \sigma^{lpha^2}_{
m SV}({
m BabaYaga@NLO}) \, < \, {
m 0.02\% imes \sigma_0}$$

Lepton and hadron loops & pairs at NNLO

- The exact NNLO soft+virtual corrections and $2 \rightarrow 4$ matrix elements $e^+e^- \rightarrow e^+e^-(l^+l^-)$ $[l = e, \mu, \tau], e^+e^- \rightarrow e^+e^-(\pi^+\pi^-)$ are available
- Compared to the *approximation* in BabaYaga@NLO, using realistic luminosity cuts ($S_i \equiv \sigma_i^{\text{NNLO}} / \sigma_{BY}$)

	\sqrt{s}		$\sigma_{ m BY}$	$S_{e^+e^-}$ [‰]	S_{lep} [‰]	S_{had} [‰]	S_{tot} [‰]
KLOE	1.020	NNLO		-3.935(4)	-4.472(4)	1.02(2)	-3.45(2)
		BB@NLO	455.71	-3.445(2)	-4.001(2)	0.876(5)	-3.126(5)
BES	3.650	NNLO		-1.469(9)	-1.913(9)	-1.3(1)	-3.2(1)
		BB@NLO	116.41	-1.521(4)	-1.971(4)	-1.071(4)	-3.042(5)
BaBar	10.56	NNLO		-1.48(2)	-2.17(2)	-1.69(8)	-3.86(8)
		BB@NLO	5.195	-1.40(1)	-2.09(1)	-1.49(1)	-3.58(2)
Belle	10.58	NNLO		-4.93(2)	-6.84(2)	-4.1(1)	-10.9(1)
		BB@NLO	5.501	-4.42(1)	-6.38(1)	-3.86(1)	-10.24(2)

 $\star\,$ The uncertainty due to lepton and hadron pair NNLO corrections is at the level of a few units in $10^{-4}\,$

Carloni, Czyz, Gluza, Gunia, Montagna, Nicrosini, Piccinini, Riemann et al., JHEP 1107 (2011) 126

Error budget for Bhabha luminometry

main conclusion of the Luminosity Section of the WG Report

Putting the sources of uncertainties (in large–angle Bhabha) all together:

Source of error (%)	$\Phi-factories$	\sqrt{s} = 3.5 GeV	<i>B</i> -factories
$\left \delta_{\mathrm{VP}}^{\mathrm{err}} ight $ [Jegerlehner]	0.00	0.01	0.03
$\left \delta_{ m VP}^{ m err} ight $ [HMNT]	0.02	0.01	0.02
$ \delta^{ m err}_{ m SV,lpha^2} $	0.02	0.02	0.02
$\left \delta_{\mathrm{HH},\alpha^{2}}^{\mathrm{err}}\right $	0.00	0.00	0.00
$ \delta_{\rm SV,H,\alpha^2}^{\rm err} $	0.05	0.05	0.05
$ \delta_{ m pairs}^{ m err} $	0.03	0.016	0.03
$ \delta_{ ext{total}}^{ ext{err}} $ linearly	0.12	0.1	0.13
$ \delta_{ m total}^{ m err} $ in quadrature	0.07	0.06	0.06

- For the experiments on top of and closely around the narrow resonances $(J/\psi, \Upsilon, ...)$, the accuracy deteriorates, because of the differences between the predictions of independent $\Delta \alpha_{had}^{(5)}(q^2)$ parameterizations
- The present error estimate appears to be rather robust and sufficient for high-precision luminosity measurements. It is comparable with that achieved for small-angle Bhabha luminosity monitoring at LEP/SLC

BabaYaga for dark photon searches at low-enegies

• From normalization to "discovery" tool \rightarrow

 $e^+e^-
ightarrow \gamma + \gamma^{
m dark}
ightarrow \ell^+\ell^-\gamma \; (n\gamma)$

- dark photon (U-boson) production via radiative return, including ISR (with LL collinear structure-functions)
- Implemented model:

"secluded" $U(1)_S$ symmetry with a light vector gauge field, heavy DM, complex Higgs field for $U(1)_S$ SSB

 \hookrightarrow coupling to SM fields through $\gamma^{\rm dark}/\gamma$ mixing

$$\mathcal{L}_{mix} = \frac{\epsilon}{2} F^{\mu\nu}_{\gamma^{\mathrm{dark}}} F^{\gamma}_{\mu\nu}$$

- extremely weak signal \rightarrow control of background mandatory



KLOE-2 results

 BabaYaga with dark-photon production used in KLOE-2 analyses for exclusion plots



A. Anastasi et al., Phys. Lett. B 750 (2015) 633

D. Babusci et al. Phys. Lett. B 736 (2014) 459

Conclusions

- In the last 15 years BabaYaga/BabaYaga@NLO has been developed for high-precision luminometry at flavour factories
- ⋆ It simulates QED processes

$$\begin{array}{l} \hookrightarrow e^+e^- \to e^+e^- \ (+n\gamma) \\ \hookrightarrow e^+e^- \to \mu^+\mu^- \ (+n\gamma) \\ \hookrightarrow e^+e^- \to \gamma\gamma \ (+n\gamma) \end{array}$$

with multiple-photon emission in a QED Parton Shower framework, matched with exact NLO matrix elements

- ★ A theoretical precision at the 0.5×10^{-3} level is achieved (at least for Bhabha), with a systematic comparison to independent calculations/codes and missing higher-order corrections
- * It has been extended to simulate dark-photon production via radiative-return method for search at low-energy e^+e^- colliders
- Improving the accuracy of QED processes would imply the inclusion of exact full 2-loop corrections, which is (in principle) feasible if needed by experiments