New dark forces hidden in low-energy QCD

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References:

ST. PRD [arXiv:1404.4370]

Kaplinghat, ST, Yu. PRL [arXiv:1508.03339]



Outline

Motivations from the dark sector
 Astrophysical hints from small scale structure
 Self-interacting dark matter

• Blind spot for new force searches: Leptophobic gauge bosons

Searching for new forces

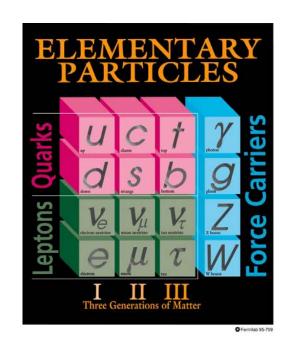
SM based on $SU(3)_C \times SU(2)_L \times U(1)_Y$ gauge symmetry.

Are there any additional gauge symmetries? Look for new gauge bosons.

Motivations:

- 1. Grand unified theories: Generically have additional gauge bosons, but typically very heavy (10¹⁶ GeV).
- 2. Dark matter: Stability of dark matter related to new gauge symmetry?

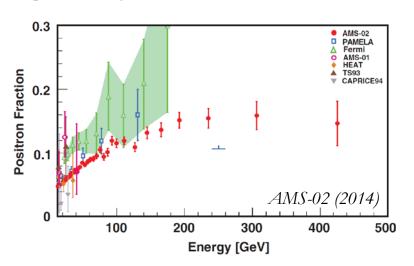
 Can also give the right relic density.

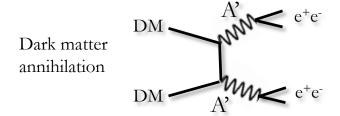


Motivations for new GeV-scale forces

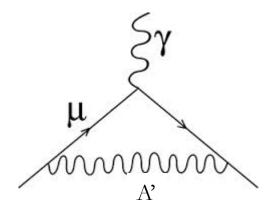
Dark matter indirect detection anomalies e.g. Pamela/AMS-02 positron excess

Pospelov & Ritz (2008); Arkani-Hamed et al (2008)





 $(g-2)_{\mu}$ anomaly Pospelov (2008)



Dark matter and structure of galaxies







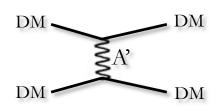
Core/cusp problem: Galaxies and clusters are less dense than cold dark matter (WIMPs) predictions

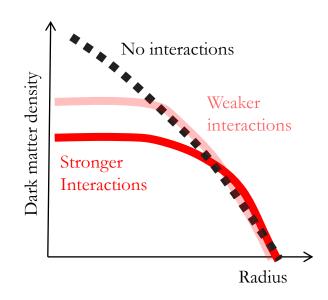
Moore (1994), Flores & Primack (1994)

Self-interacting dark matter

Spergel & Steinhardt (2000)

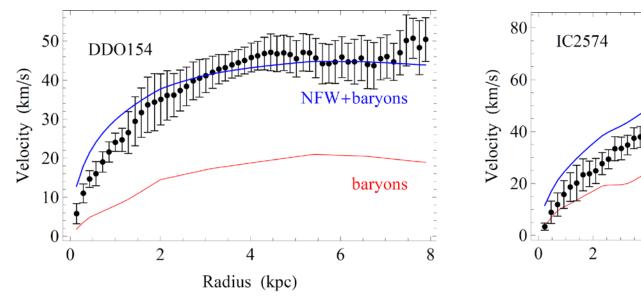
MeV–GeV scale dark force ST, Yu, Zurek (2013)

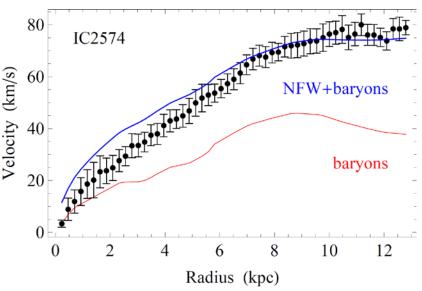




Cores in field galaxies

THINGS (dwarf galaxy survey) - Oh et al. (2011)

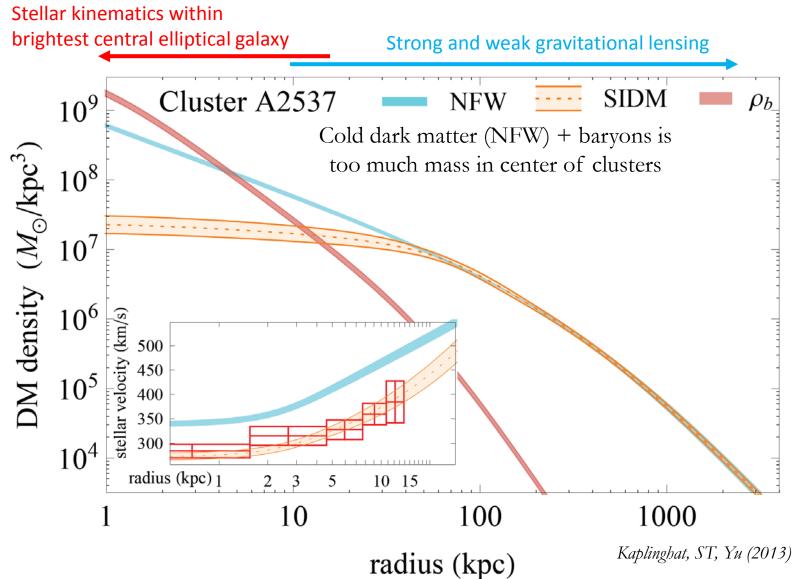




Core/cusp problem:

Cold dark matter halo (NFW profile) + baryons predict too much mass in centers of galaxies

Cores in galaxy clusters



Self-interacting Dark Matter

Unknown if core/cusp is solved by DM physics or baryonic astrophysics If DM is responsible, what are the particle physics implications?

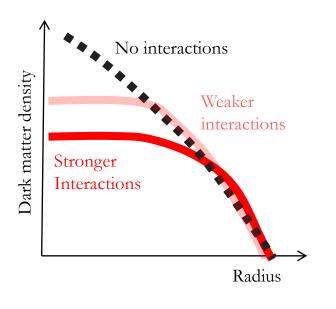


Figure of merit:
Velocity-weighted cross
section per DM mass

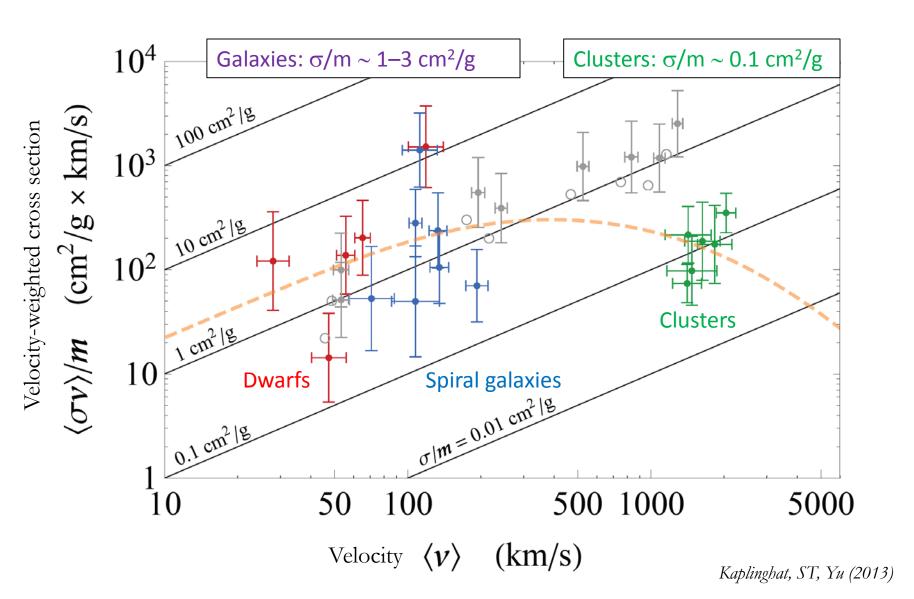
$$rate = \frac{\langle \sigma v \rangle}{m} \, \rho_{\rm dm}$$

Dwarf and spiral galaxies: velocity ~ 30-200 km/s

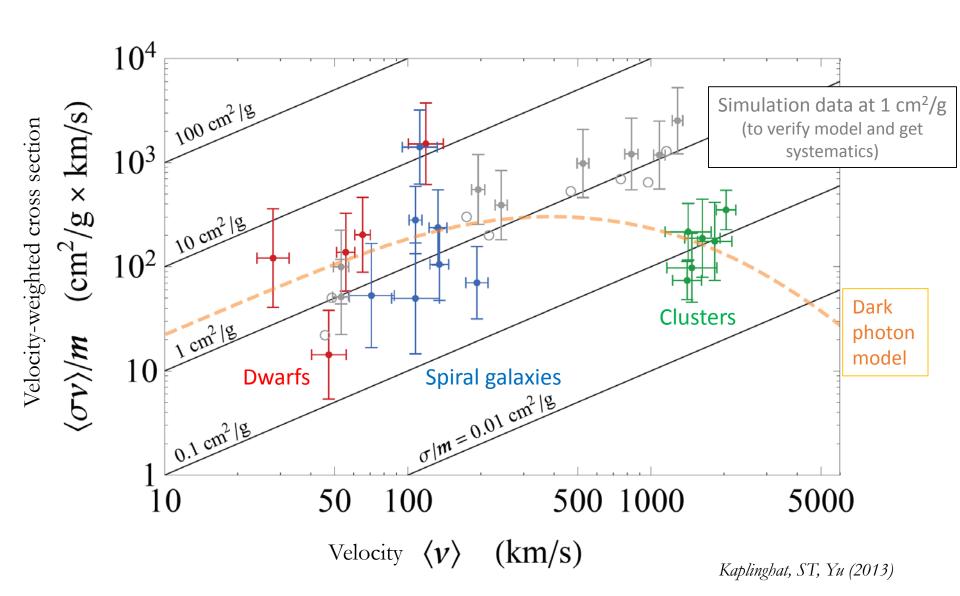
Clusters: velocity $\sim 1500 \text{ km/s}$

Can all observations be fit by a simple particle physics model?

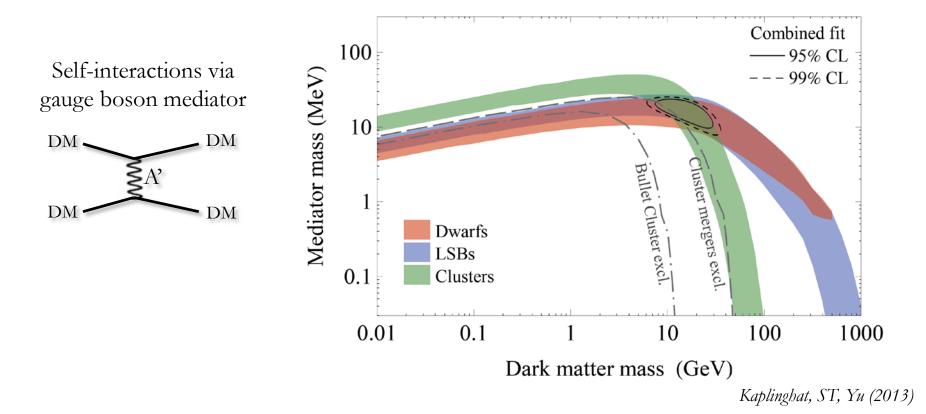
Fits to dwarfs, LSBs, and clusters



Fits to dwarfs, LSBs, and clusters



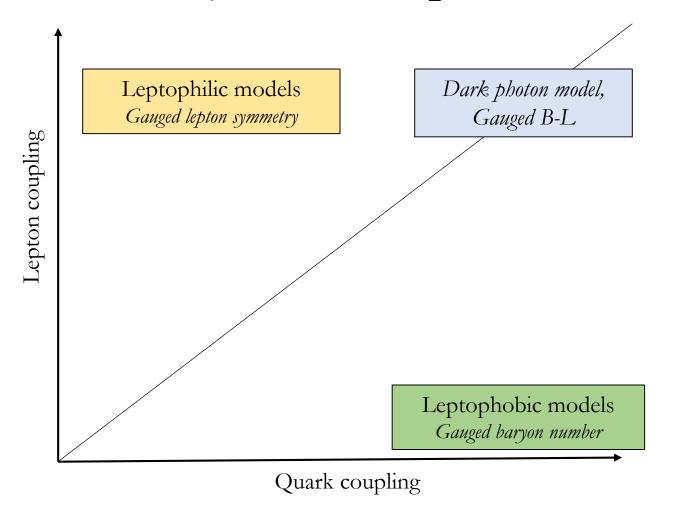
Dark matter with dark photon



Model-dependent: Dark sector parameters can be fit from astrophysical data. Not fixed how dark photon couples to Standard Model (kinetic mixing unknown)

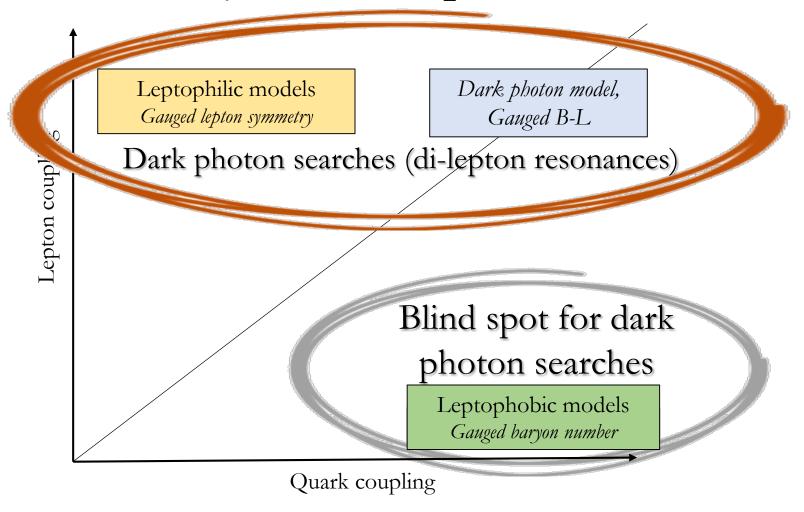
Model-independent: Dark sector particles below GeV scale to get large enough cross section

Beyond dark photons



Also a third axis: decays to invisible states (neutrinos, light dark matter) Davoudiasl et al (2012), Batell et al (2009), deNiverville et al (2011,2012)

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New force coupling to quarks

Most dark photon searches are for A' coupling to leptons (or invisible states)

What if a new force couples mainly to quarks?

Old idea: Radjoot (1989), Foot et al (1989), Nelson & Tetradis (1989), He & Rajpoot (1990), Carone & Murayama (1995), Bailey & Davidson (1995), Aranda & Carone (1998), Fileviez Perez & Wise (2010), Graesser et al (2011), Dobrescu & Frugiule (2014), Batell et al (2014), ...

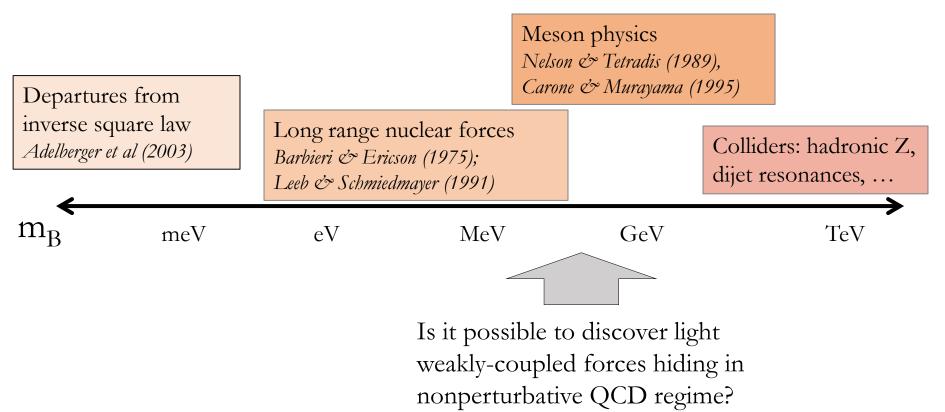
Simplest model: U(1)_B gauge boson coupled to baryon number

$$\mathscr{L} = \frac{g_B}{3} \, \bar{q} \gamma^\mu q B_\mu$$
 Flavor-universal charge g_B coupling to all quarks

Also known as: "leptophobic Z" or "baryonic photon γ_B " or " Z_B " or "B boson"

New force coupling to quarks

B boson = gauge boson coupled to baryon number Discovery signals depend on the B mass



Theoretical constraints from anomalies

• U(1)_B gauge symmetry is anomalous. Requires new fermions with electroweak quantum numbers.

• Absence of new fermions at colliders (mass > 100 GeV) implies new baryonic force must have very small coupling

$$g_B \lesssim 10^{-2} \times (m_B/100 \text{ MeV})$$

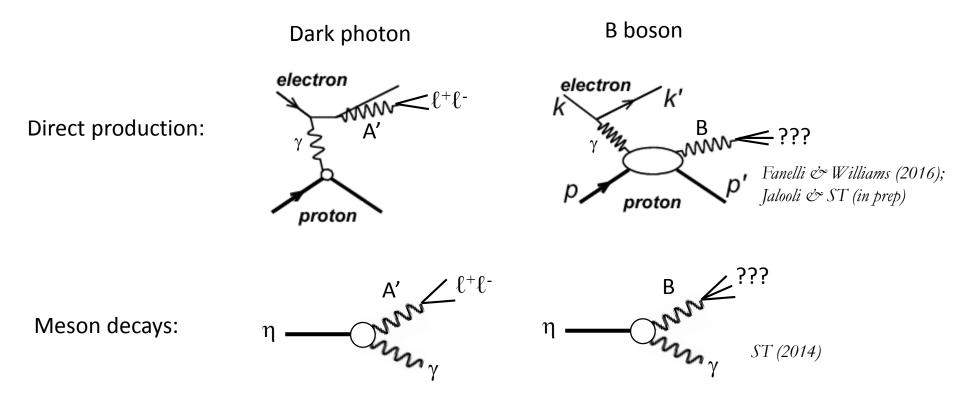
$$\alpha_B = \frac{g_B^2}{4\pi} \lesssim 10^{-5} \times (m_B/100 \,\text{MeV})^2$$

Detecting the B boson

- Can a weakly-coupling force ($g_B << 1$) be detected in the nonperturbative regime of QCD?
- B boson preserves the symmetries of QCD
 - Charge conjugation, parity, and isospin or SU(3)_{flavor}
- Previous lore: Nelson & Tetradis (1989)
 - Above $2m_{\pi}$, decay dominated by B $\rightarrow \pi\pi$
 - B boson buried under huge $\rho \rightarrow \pi\pi$ background

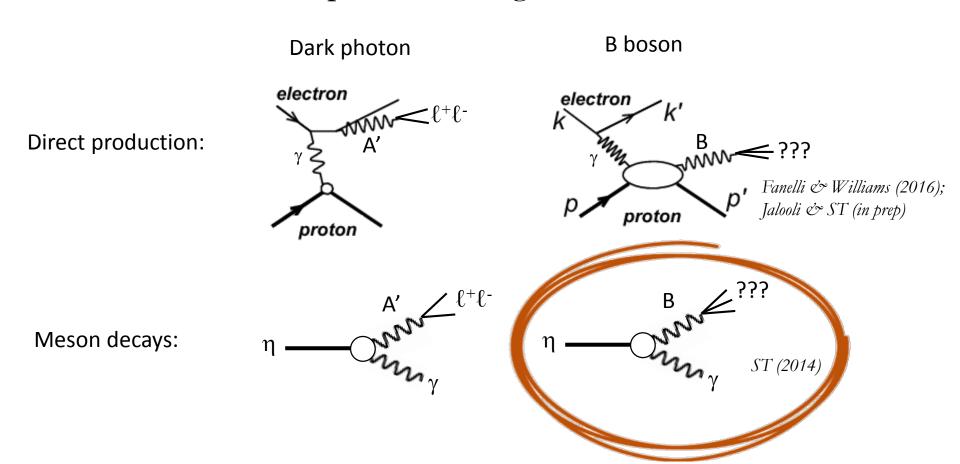
Baryonic force at the QCD scale

- How are the gauge bosons produced?
- What are the experimental signatures?



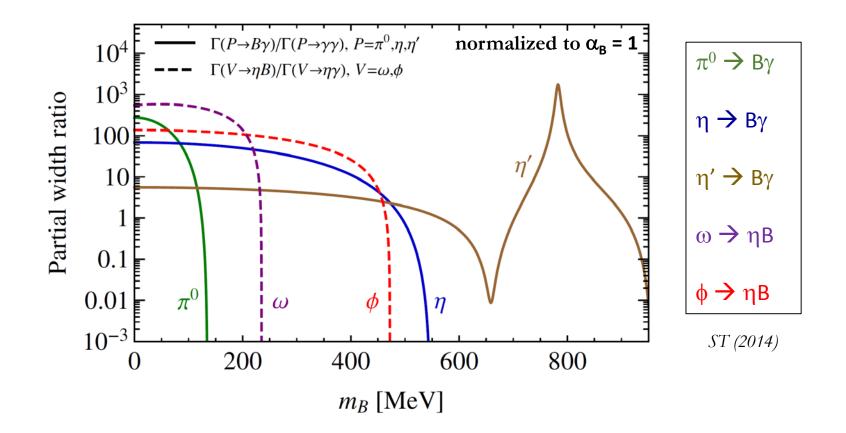
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B bosons production

Meson decays: $\pi^0 \to B\gamma$, $\eta \to B\gamma$, $\eta' \to B\gamma$, $\omega \to \eta B$, $\phi \to \eta B$ Like SM decays: $\pi^0 \to \gamma\gamma$, $\eta \to \gamma\gamma$, $\eta' \to \gamma\gamma$, $\omega \to \eta\gamma$, $\phi \to \eta\gamma$



B boson decay

How does B decay? Worry: B $\rightarrow \pi\pi$ is hopeless.

Recall the original Lagrangian:
$$\mathscr{L} = \frac{g_B}{3} \bar{q} \gamma^{\mu} q B_{\mu}$$

The quantum numbers for B:

- J = 1
- *P* = *C* = –
- I = 0
- G = -

B boson decay

B has same quantum numbers as the ω meson

Particle Data Book

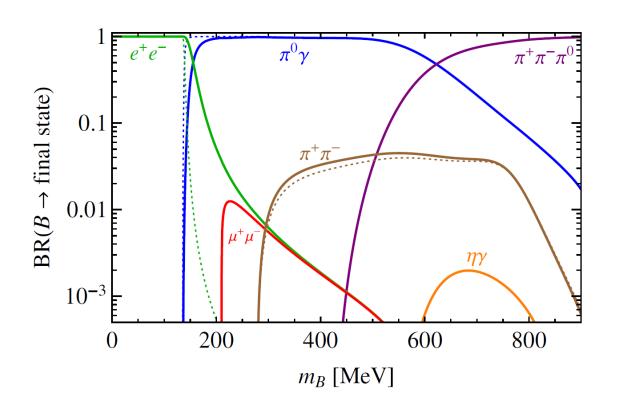
Scale factor/

$$I^G(J^{PC}) = 0^-(1^{-})$$

ω (782) DECAY MODES

	Mode		Fraction (Γ_i/Γ)	Confidence level
_	$\pi^+\pi^-\pi^-\pi^0\gamma$	0	(89.2 ±0.7) % (8.28±0.28) %	S=2.1
Γ ₃	$\pi^+\pi^-$	$\omega \rightarrow \pi\pi$ forbidden by G-parity (Isospin-violating ρ – ω mixing)	$(1.53^{+0.11}_{-0.13})\%$	S=1.2
Γ_9	e^+e^-		$(7.28\pm0.14)\times10$	o ⁻⁵ S=1.3
Γ ₁₅ Γ ₁₆	$\mu^+\mu^ 3\gamma$		$(9.0 \pm 3.1) \times 10$ $< 1.9 \times 10$	

B boson decay branching ratios



Computed using vector meson dominance

Subleading lepton couplings arise by $B-\gamma$ mixing

Solid: $\varepsilon = eg_B/16\pi^2$ Dotted: $\varepsilon = 0.1 eg_B/16\pi^2$

New signatures not covered in dark photon searches: $B \rightarrow \pi^0 \gamma$, $\pi^+ \pi^- \pi^0$

B boson signal channels

Decay →	$B \rightarrow e^+e^-$	$B \to \pi^0 \gamma$	$B \to \pi^+\pi^-\pi^0$	
Production ↓	$m_B \sim 1 - 140 \text{ MeV}$	140–620 MeV	620-1000 MeV	$B \to \eta \gamma$
$\pi^0 \to B\gamma$	$\pi^0 ightarrow e^+ e^- \gamma$			
$\eta \to B\gamma$	$\eta ightarrow e^+ e^- \gamma$	$\eta o\pi^0\gamma\gamma$	• • •	
$\eta' \to B\gamma$	$\eta' ightarrow e^+ e^- \gamma$	$\eta' o \pi^0 \gamma \gamma$	$\eta^\prime o \pi^+\pi^-\pi^0\gamma$	$\eta' o \eta \gamma \gamma$
$\omega \to \eta B$	$\omega ightarrow \eta e^+ e^-$	$\omega o \eta \pi^0 \gamma$	• • •	• • •
$\phi \to \eta B$	$\phi \to \eta e^+ e^-$	$\phi o \eta \pi^0 \gamma$		

Covered by dark photon searches Limits are more model dependent New signals not being covered in dark photon searches

A new type of signature for meson factories:

 $\pi^0 \gamma$ resonances in rare decays

B boson signal channels

Decay →	$B \rightarrow e^+ e^-$	$B o \pi^0 \gamma$	$B o \pi^+\pi^-\pi^0$	
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$\pi^0 \to B\gamma$	$\pi^0 o e^+ e^- \gamma$			
$\eta \to B\gamma$	$\eta ightarrow e^+e^-\gamma$	$\eta \to \pi^0 \gamma \gamma$		
$\eta' \to B\gamma$	$\eta' ightarrow e^+ e^- \gamma$	$\eta' o \pi^0 \gamma \gamma$	$\eta^\prime o \pi^+\pi^-\pi^0\gamma$	$\eta' o \eta \gamma \gamma$
$\omega \to \eta B$	$\omega ightarrow \eta e^+ e^-$	$\omega o \eta \pi^0 \gamma$	• • •	• • •
$\phi \to \eta B$	$\phi ightarrow \eta e^+ e^-$	$\phi \to \eta \pi^0 \gamma$	• • •	
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Covered by dark photon searches Limits are more model dependent New signals not being covered in dark photon searches

A new type of signature for meson factories:

 $\pi^0 \gamma$ resonances in rare decays

$\eta \rightarrow \pi^0 \gamma \gamma$

Particle Data Book

η		$I^{G}(J^{PC}) = 0^{+}(0^{-}$	+)			
η DECAY MODES						
	Mode	Fraction (Γ_i/Γ)	Scale factor/ Confidence level			
		Neutral modes	_			
Γ_1	neutral modes	(72.12±0.34) %	S=1.2			
Γ_2	$\frac{2\gamma}{3\pi^0}$	(39.41 ± 0.20) %	S=1.1			
Γ ₃	$3\pi^0$	(32.68 ± 0.23) %	S=1.1			
Γ ₄	$\pi^0 2\gamma$	(2.7 ± 0.5) \times 10	o ⁻⁴ S=1.1			

B boson signature: $\eta \rightarrow B\gamma \rightarrow \pi^0\gamma\gamma$

Mimics the rare SM decay $\eta \rightarrow \pi^0 \gamma \gamma$

Nelson & Tetradis (1989)

Total rate constraint:
$$\frac{\Gamma(\eta \to B\gamma)}{\Gamma(\eta \to \gamma\gamma)} = 2\frac{\alpha_B}{\alpha_{\rm em}} \left(1 - \frac{m_B^2}{m_\eta^2}\right)^3 \times O(1) < \frac{\Gamma(\eta \to \pi^0 \gamma\gamma)}{\Gamma(\eta \to \gamma\gamma)} \sim 10^{-3}$$

Requires $\alpha_{\rm B} < 10^{-5} << \alpha_{\rm em}$

Boost sensitivity by searching for $\pi^0 \gamma$ resonance in $\eta \rightarrow \pi^0 \gamma \gamma$ Proposal by Jefferson Eta Factory (JLab)

$\phi \rightarrow \eta \pi^0 \gamma$

B boson signature: $\phi \rightarrow \eta B \rightarrow \eta \pi^0 \gamma$

Mimics the rare SM decay $\phi \rightarrow \eta \pi^0 \gamma$

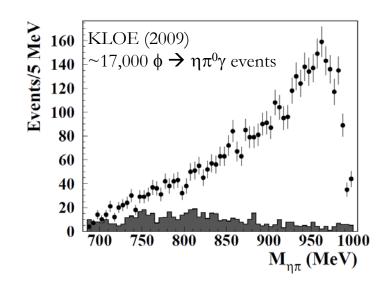
SM decay target for understanding scalar resonance in QCD

$$\phi \rightarrow a_0(980) * \gamma \rightarrow \eta \pi^0 \gamma$$

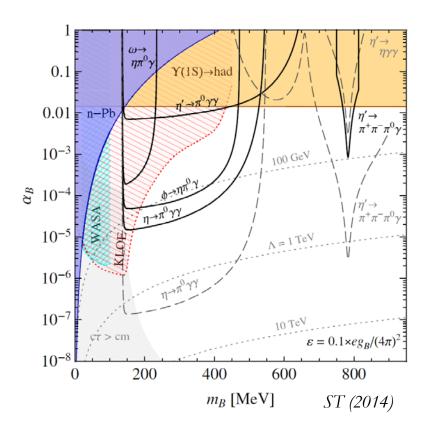
Achasov & Ivanchenko (1989)

B boson signature: $\pi^0 \gamma$ invariant mass peak

Total rate constraint: $\alpha_{\rm B} < 5 {\rm x} 10^{-5} << \alpha_{\rm em}$



Constraints on B boson



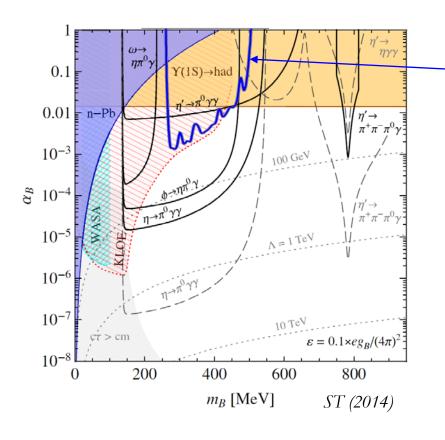
Black lines = rare meson decays (based on total rate)

Blue = low-energy neutron scattering

Orange = $U \rightarrow$ hadrons

Red/Blue hatched = Dark photon searches (model-dependent)

Constraints on B boson



First official B boson search Won et al |Belle| (2016)

Search for $\pi^+\pi^-$ resonance in $\eta \rightarrow B\gamma \rightarrow \pi^+\pi^-\gamma$

Conclusions

- New forces beyond the Standard Model:
 - Motivated by dark matter
 - Would be a game-changing particle physics discovery
- GeV-scale leptophobic forces
 - Blind spot to dark photon searches
 - Even very small couplings can be discovered (10⁵x smaller than EM)
 - Smoking gun signature: a $\pi^0 \gamma$ resonance in rare meson decays.
 - No bump hunt done (with potential for discovery)