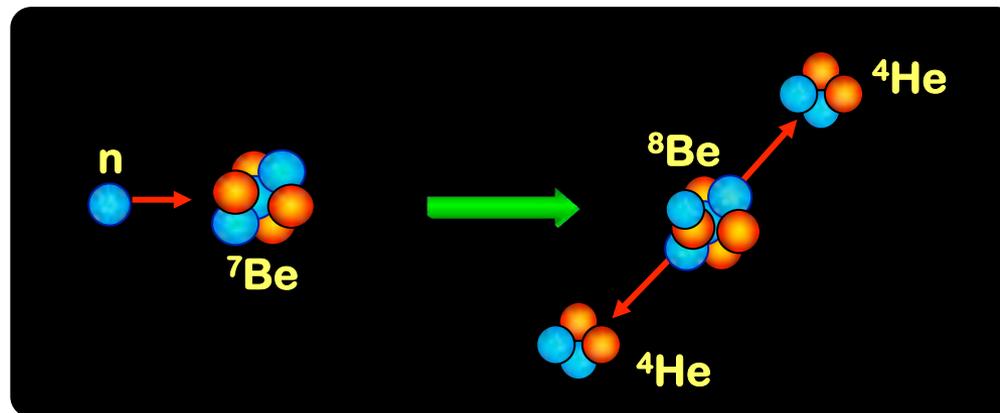


**Recent results at nTOF: ${}^7\text{Be}(n,\alpha)\alpha$ cross section measurement
does it solve the Cosmological Lithium Problem?
(not yet!)**

L. Cosentino, P. Finocchiaro,
A. Pappalardo,
A. Musumarra,
M. Barbagallo, N. Colonna,

INFN-LNS Catania
ELI-NP Bucharest
INFN-LNS & Dip. Fis. Astr. Catania
INFN Bari

and the n_TOF collaboration @CERN



...subtitle

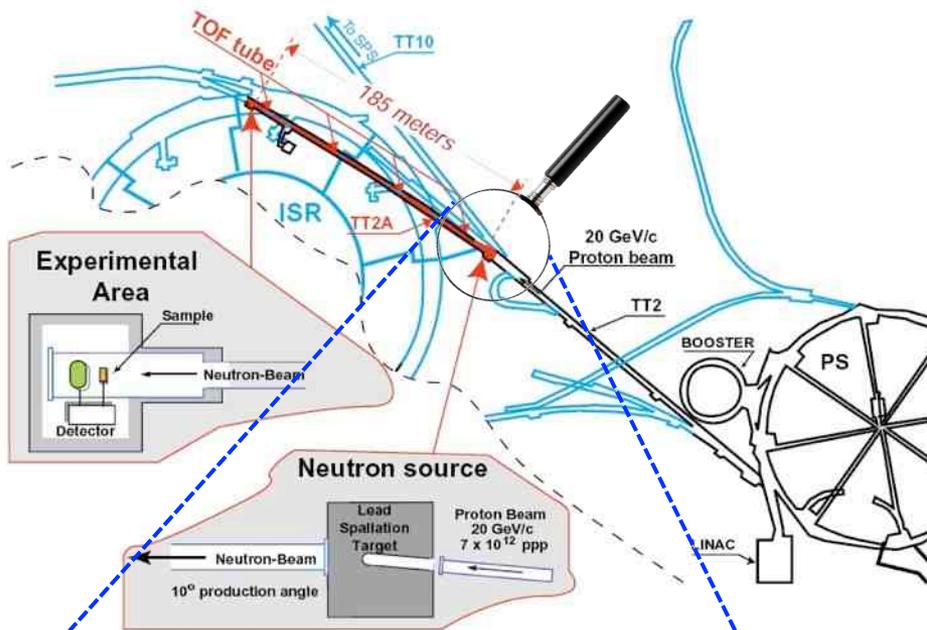
${}^7\text{Be}(n,\alpha)\alpha$: a 50-year wait for an “impossible” measurement

Summary

- few words and a short movie about n_TOF
- semiconductor detectors for neutrons (SiLiF)
- the challenging ${}^7\text{Be}(n,\alpha)\alpha$ reaction: preparation and experiment

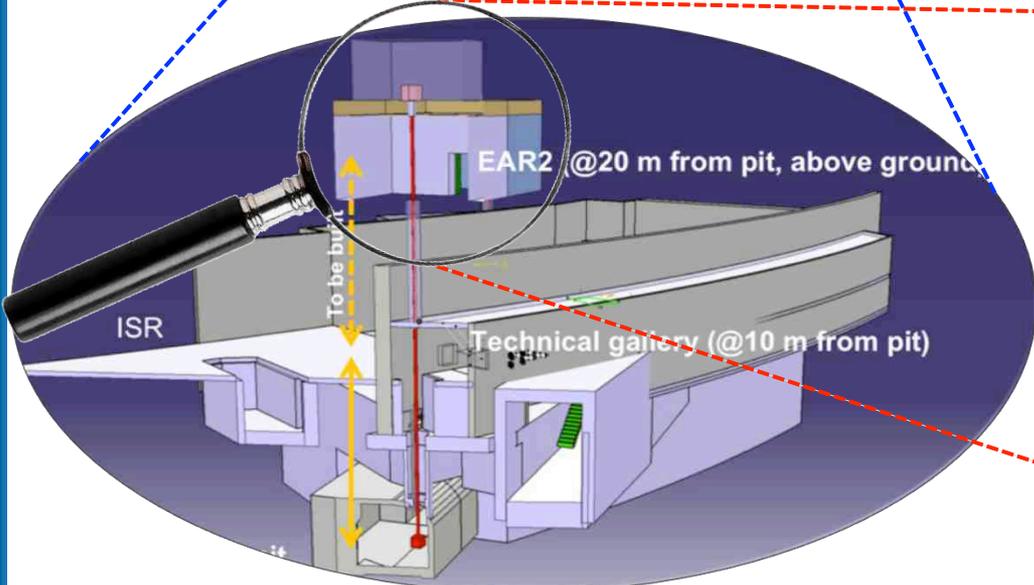
the n_TOF collaboration in Europe

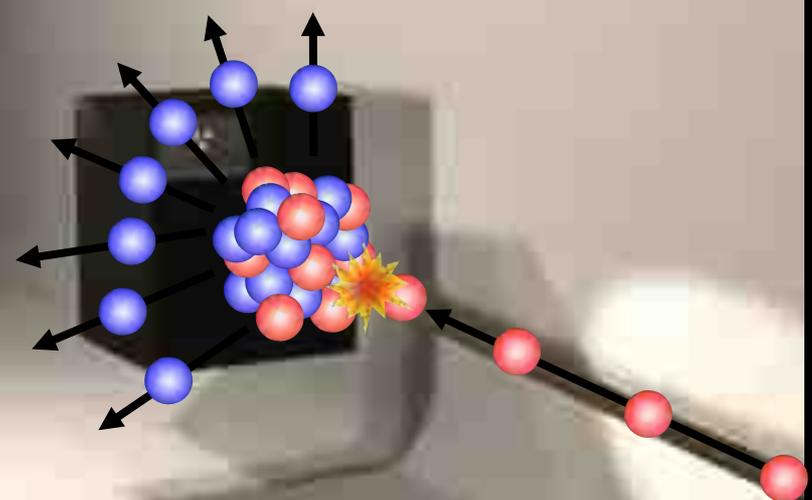




the n_TOF facility at CERN

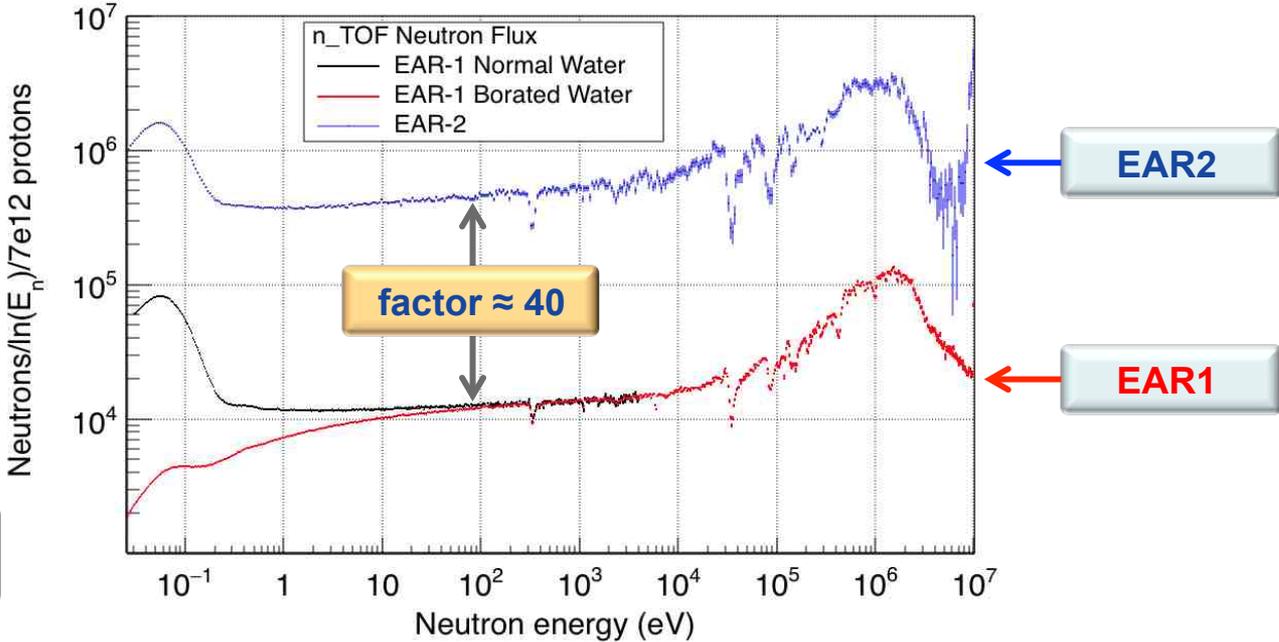
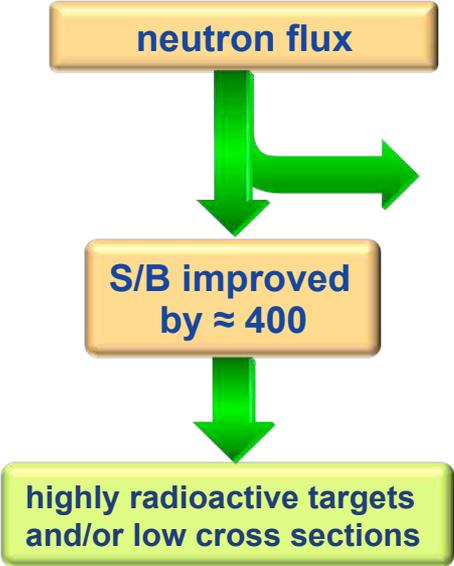
$\approx 10^7$ n/pulse



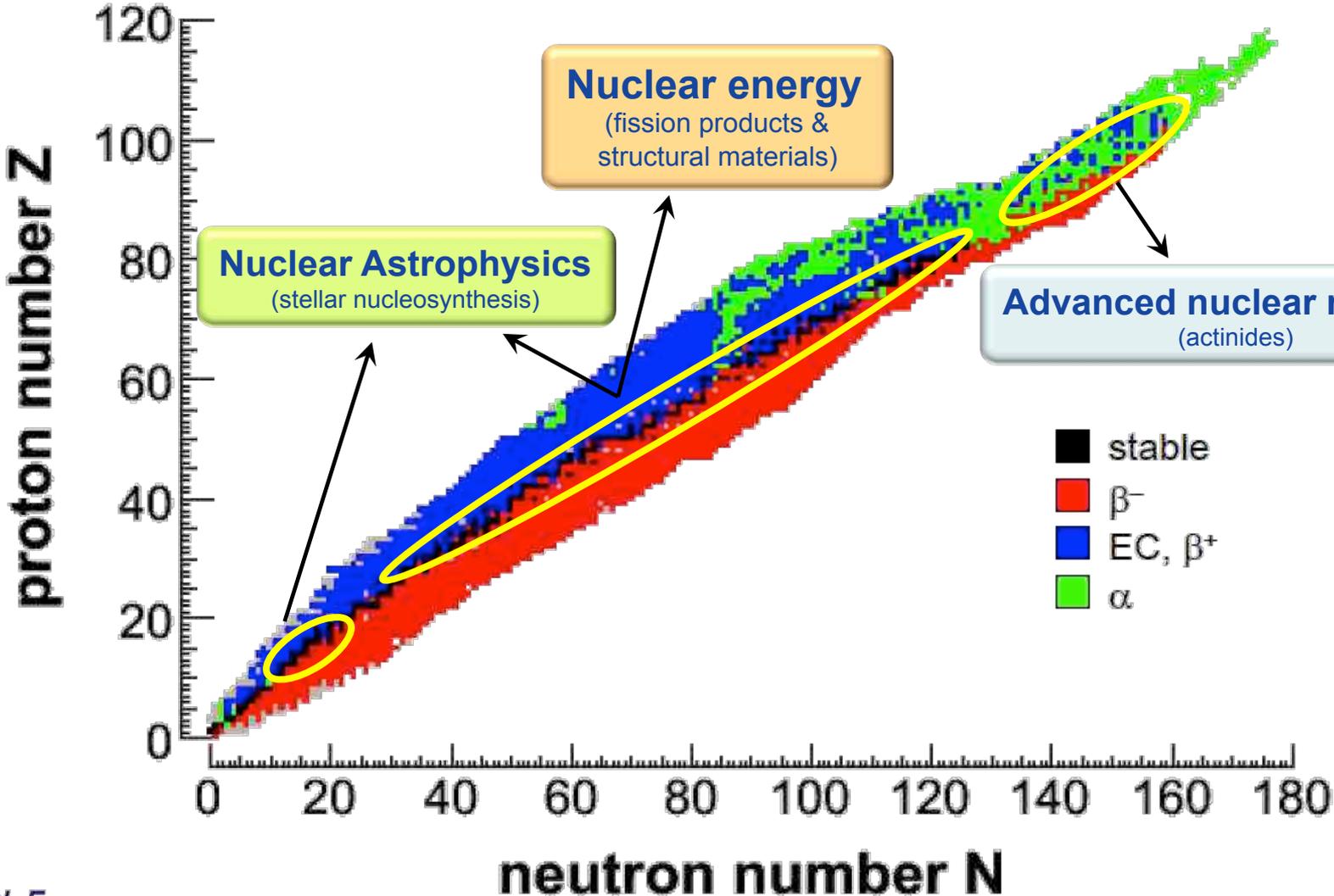


n_TOF main features

	EAR-1	EAR-2
Wide energy spectrum	25 meV < E _n < 1 GeV	25 meV < E _n < 300 MeV
High instant flux	10 ⁵ n/cm ² /pulse	10 ⁶ n/cm ² /pulse
Low repetition rate	< 0.8 Hz (1 pulse/2.4 s maximum)	
High energy resolution	ΔE/E = 10 ⁻⁴ (E _n < 10 keV)	ΔE/E = 10 ⁻³ (E _n < 10 keV)



**Neutron cross sections
are needed for...**



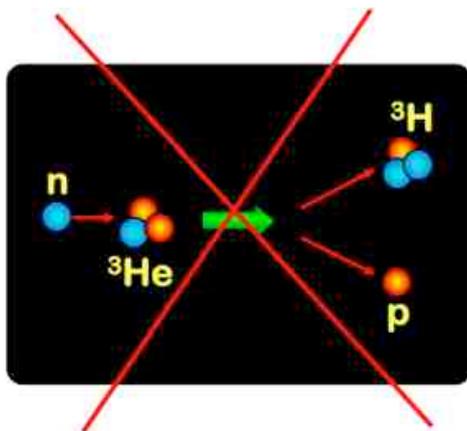
semiconductor detectors for neutrons: SiLiF

materials for thermal neutron conversion

³He

$\sigma(0.025)$
 ≈ 5330 b

available energy
0.76 MeV
no gamma rays



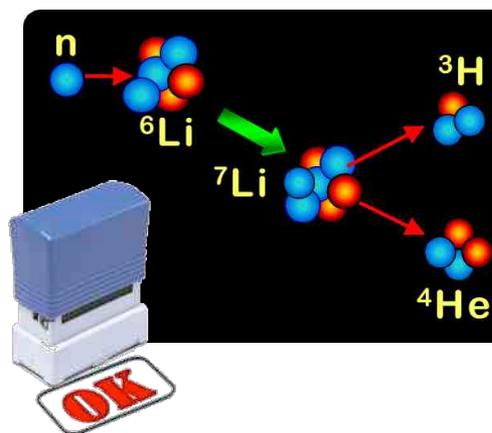
perfect gas detector but...
worldwide lack of ³He



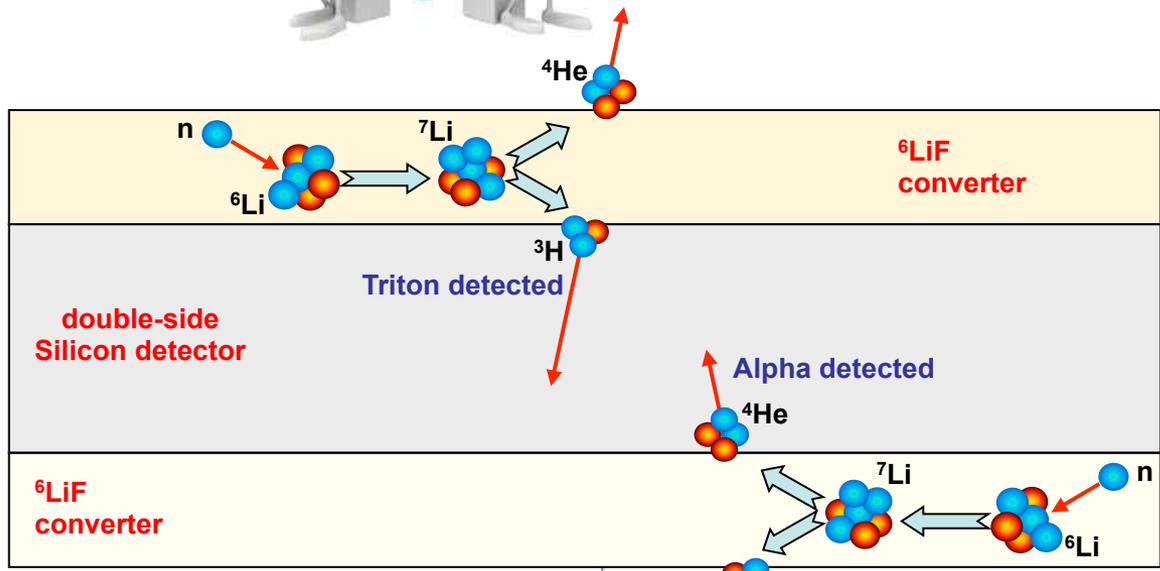
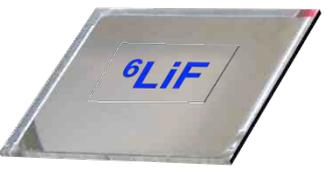
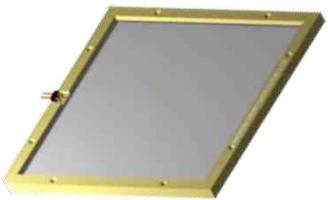
⁶Li

$\sigma(0.025)$
 ≈ 940 b

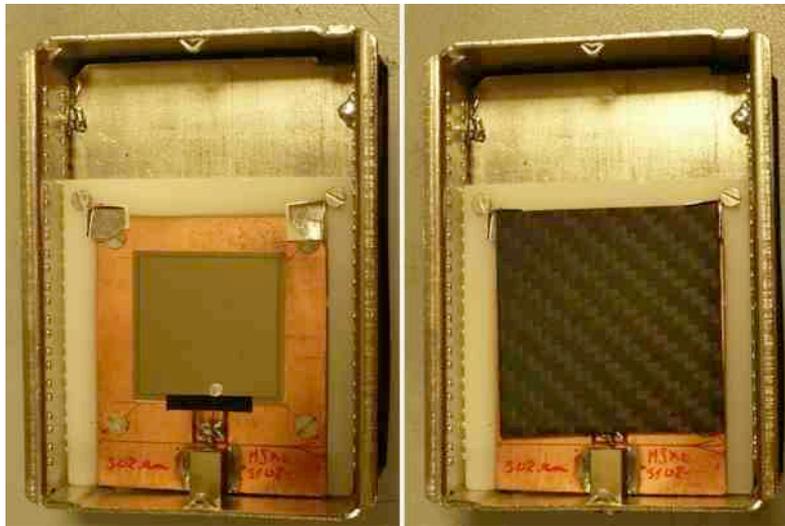
available E
4.78 MeV



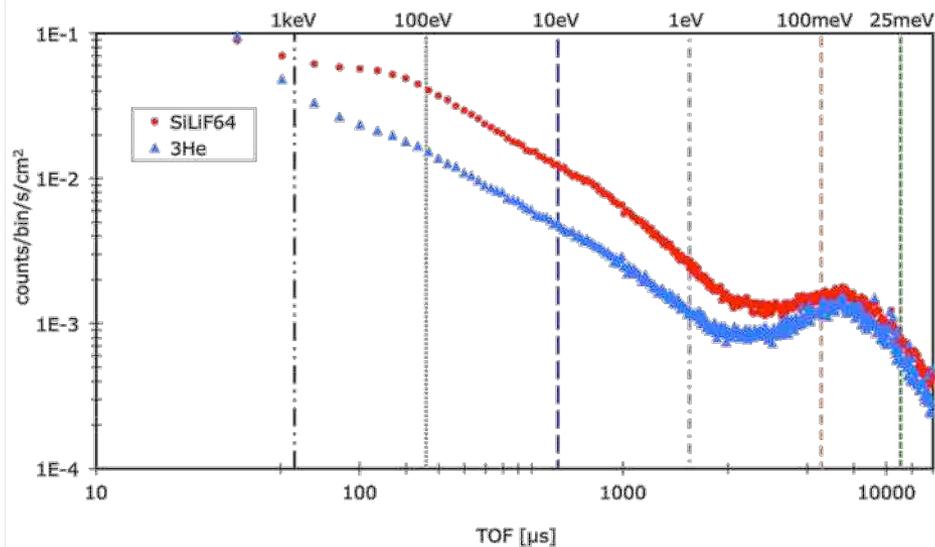
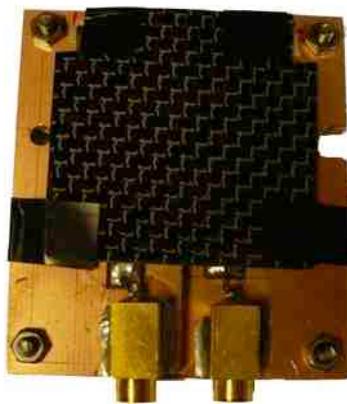
SiLiF: how?



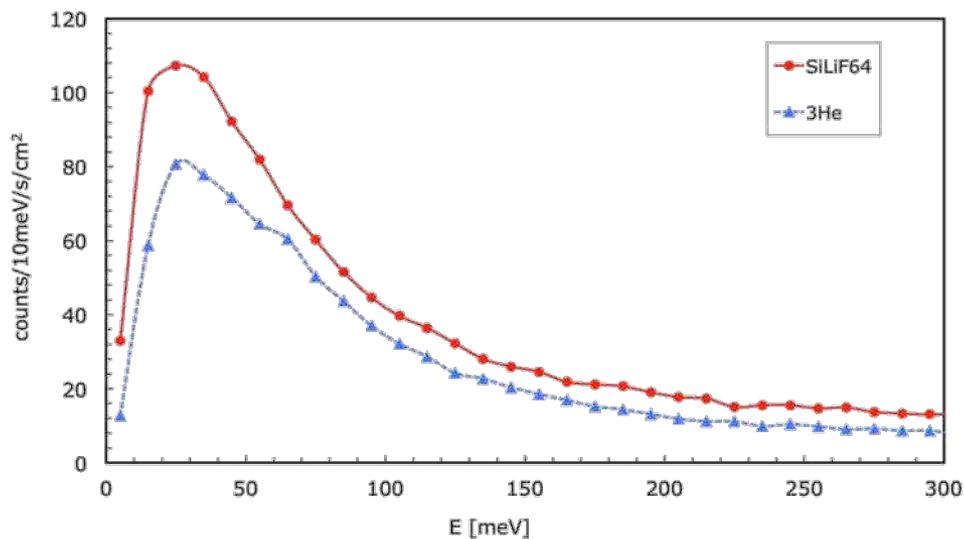
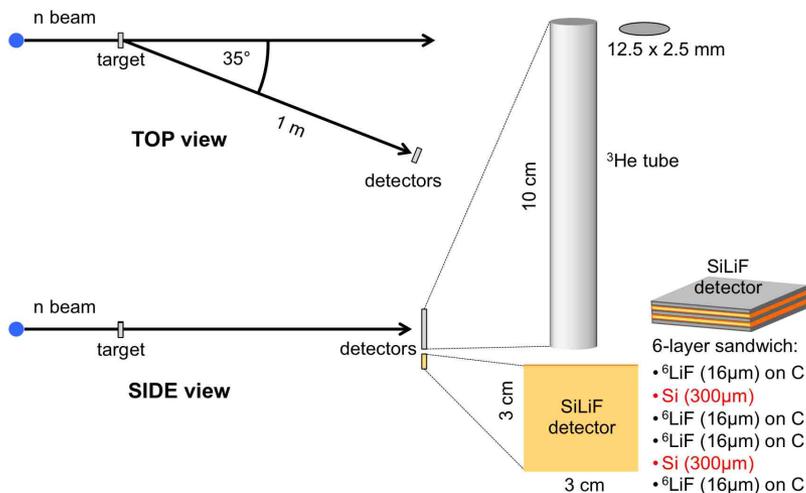
- low cost, technology cheaper than ^3He
- low voltage (20-30 V)
- flat detectors, compact, robust, simple to use, easily handled
- efficiency evaluated analytically
- fine position sensitivity (mm) easily achieved
- coarse position sensitivity (cm) with pads
- this sample 3cm x 3cm active area
- 10^{-6} neutron/gamma discrimination



SiLiF detectors

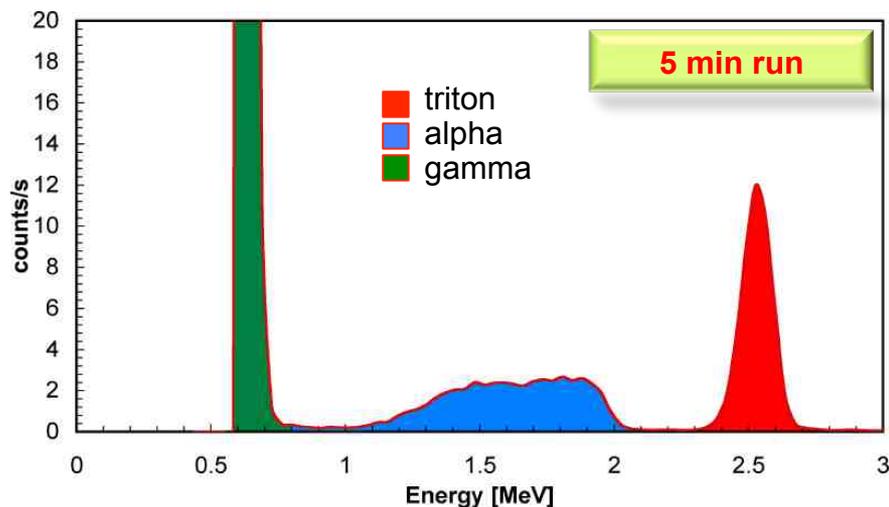
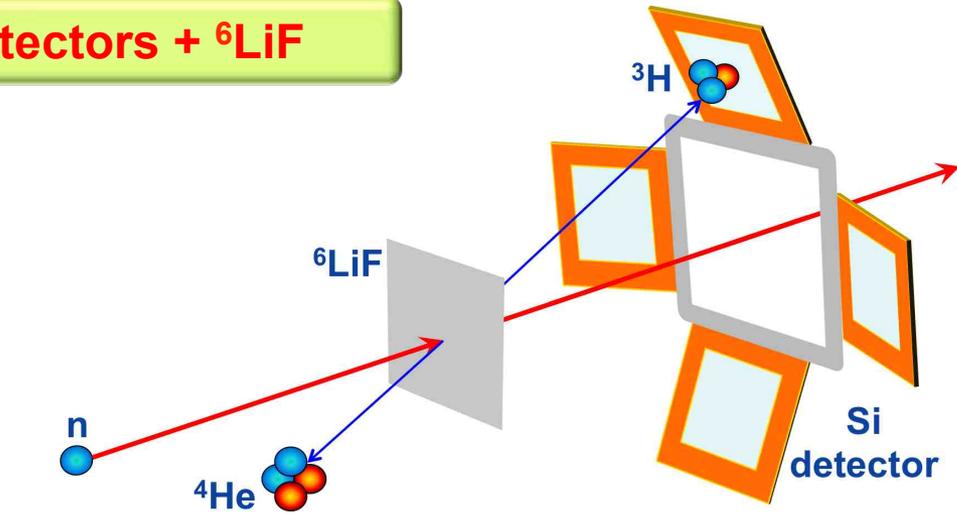
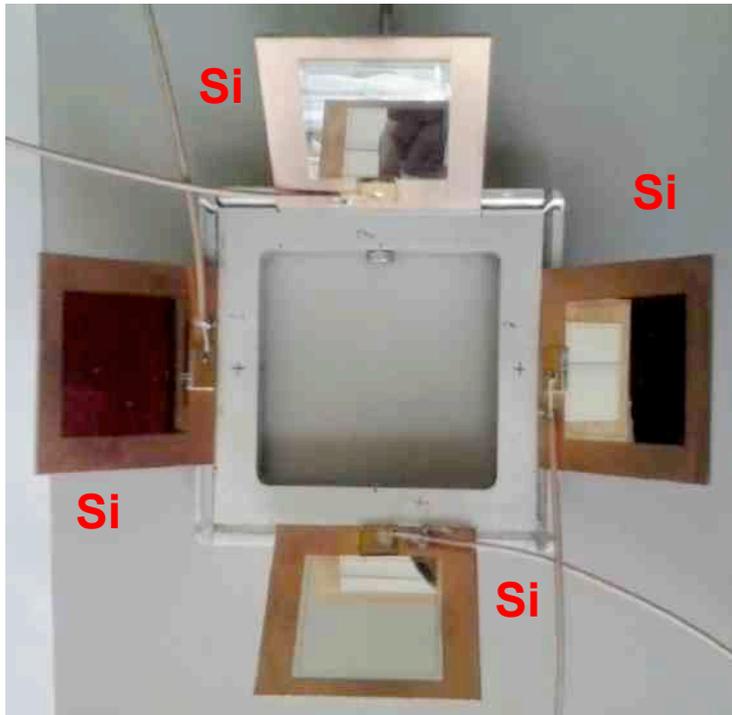


comparison with ³He tube at INES/ISIS

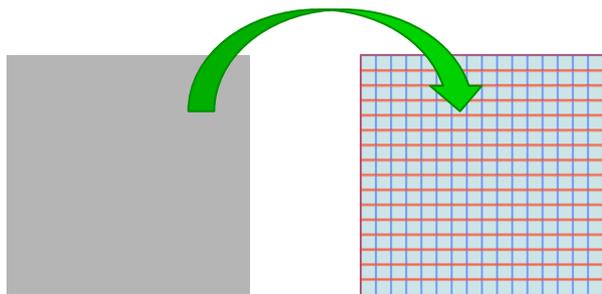


SiMon2 @ n_TOF neutron beam monitor for flux normalization

Si detectors + ^6LiF



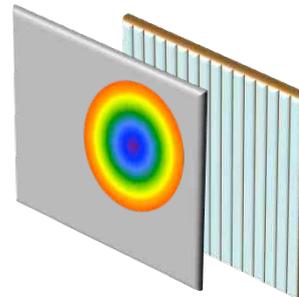
SiLiF



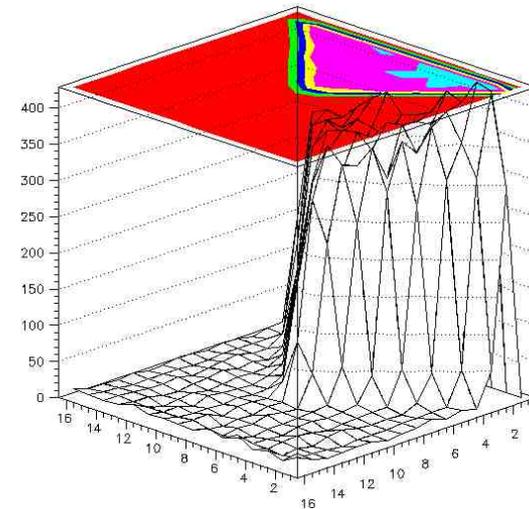
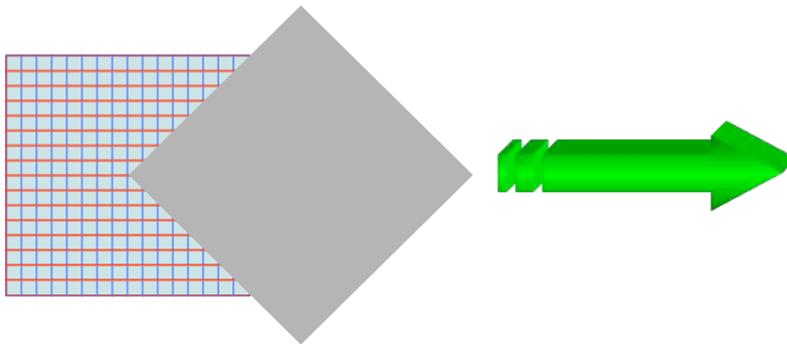
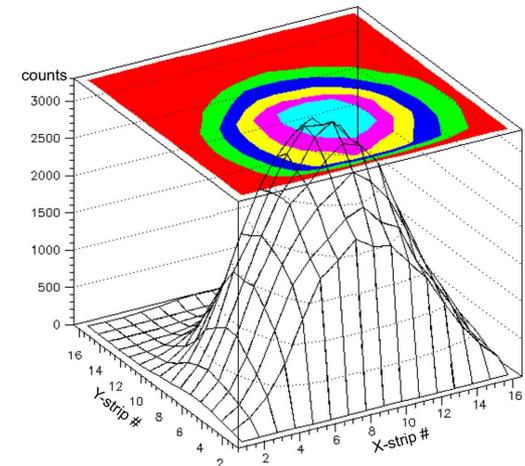
^6LiF foil

25+25 strip Si

the idea

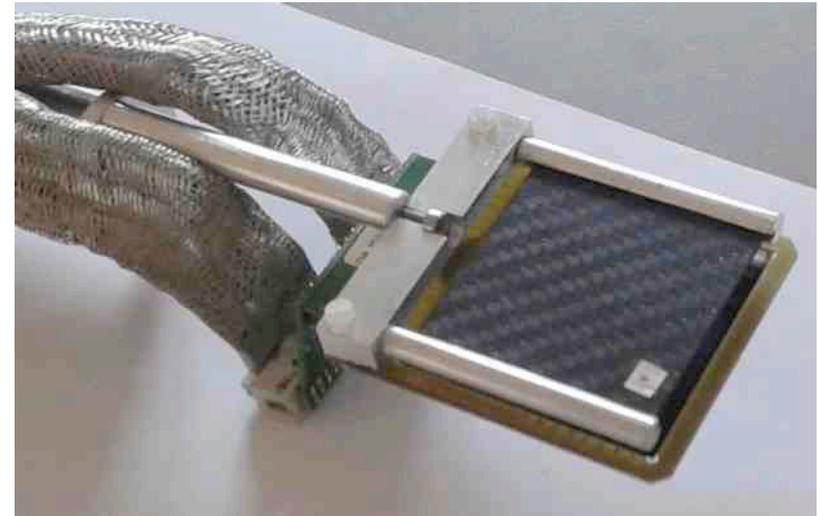
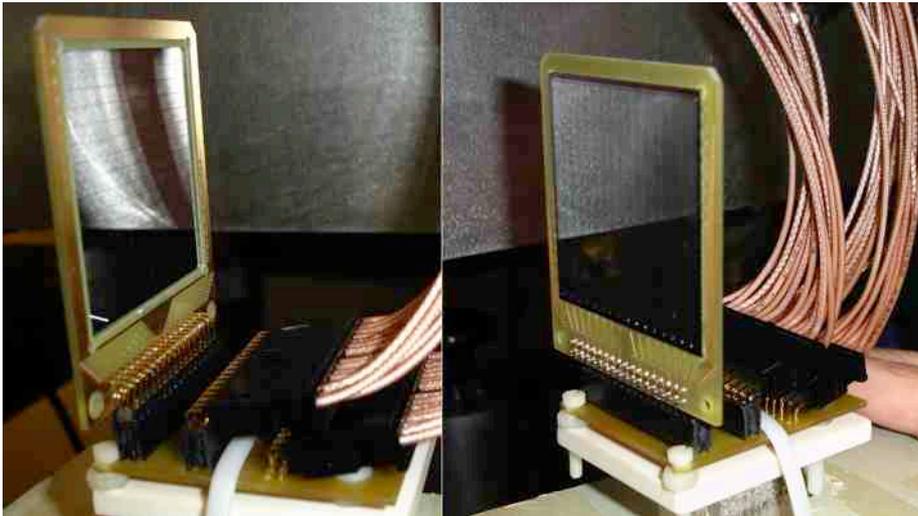
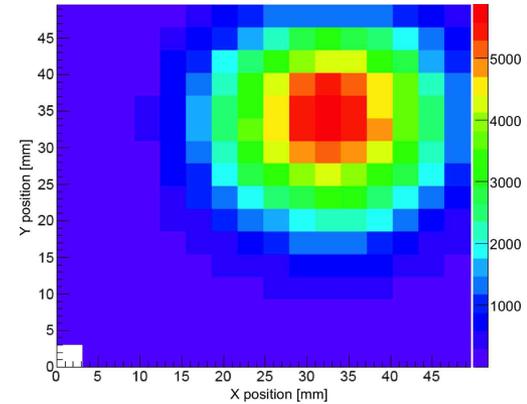


the test

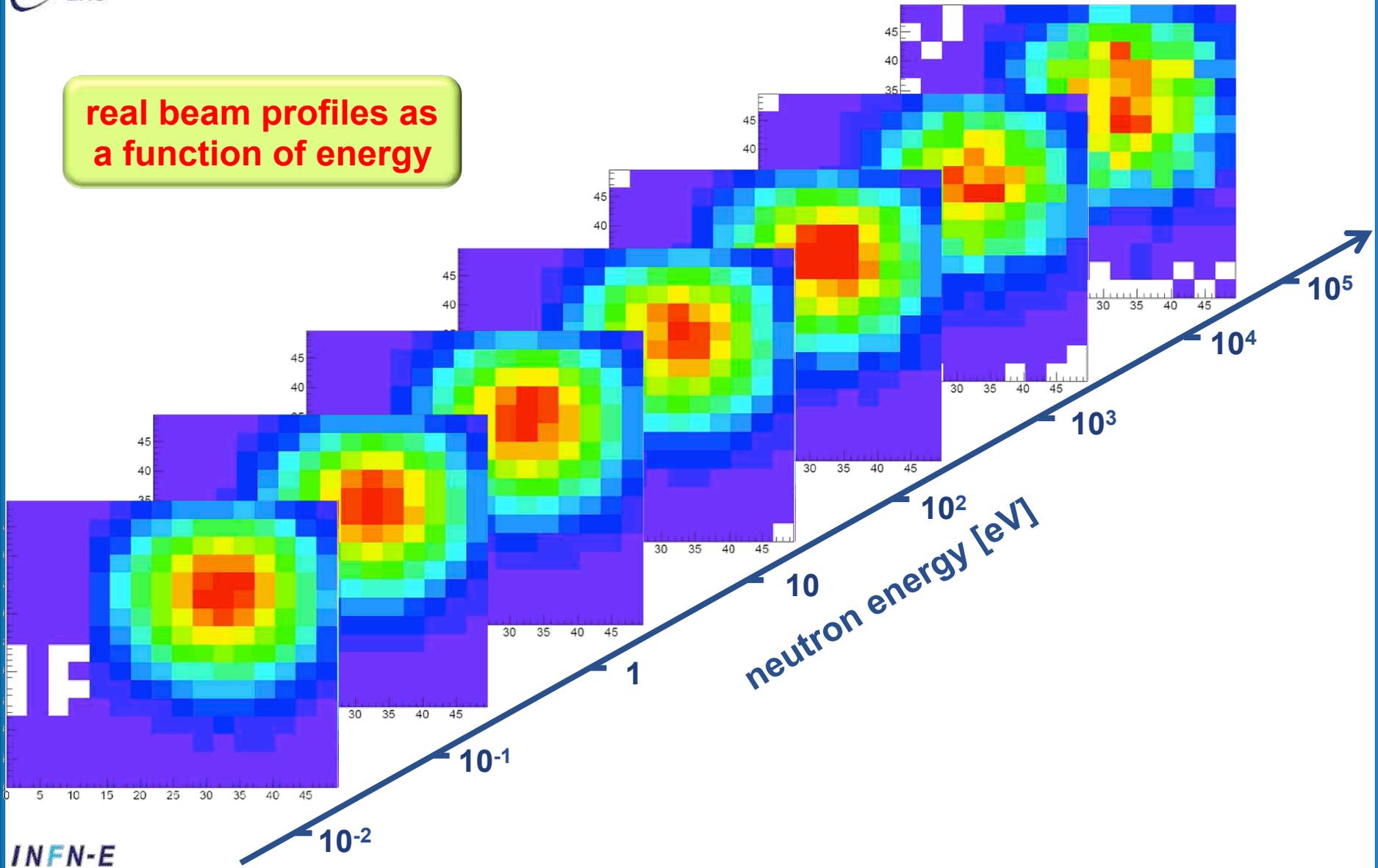


5cm x 5cm strip SiLiF detector
25 strips, 2mm x 5cm

real beam profile



real beam profiles as
a function of energy

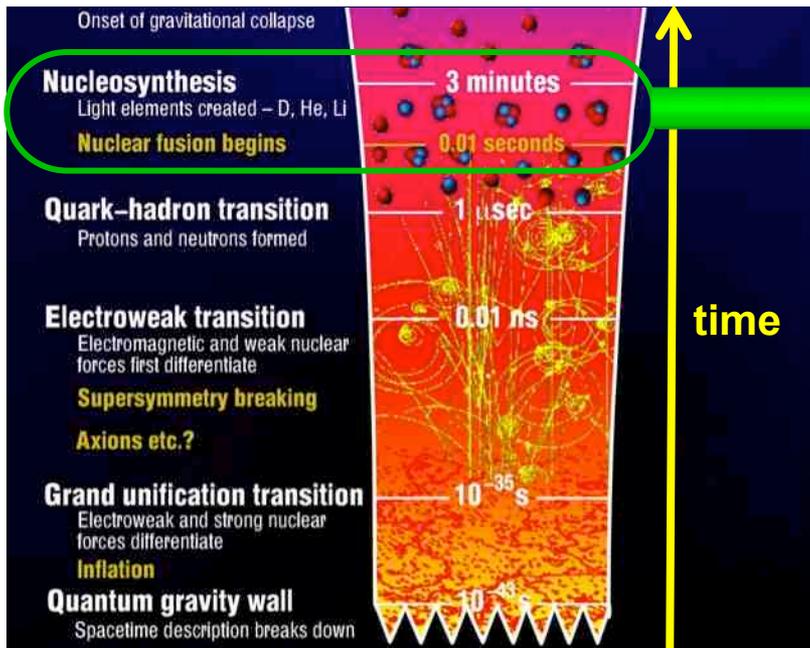


INFN-E

the challenging ${}^7\text{Be}(n,\alpha)$ reaction:
preparation and experiment

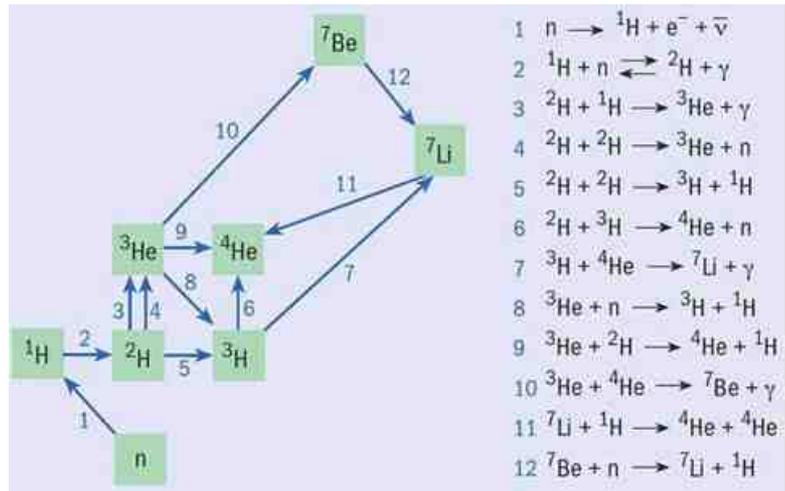
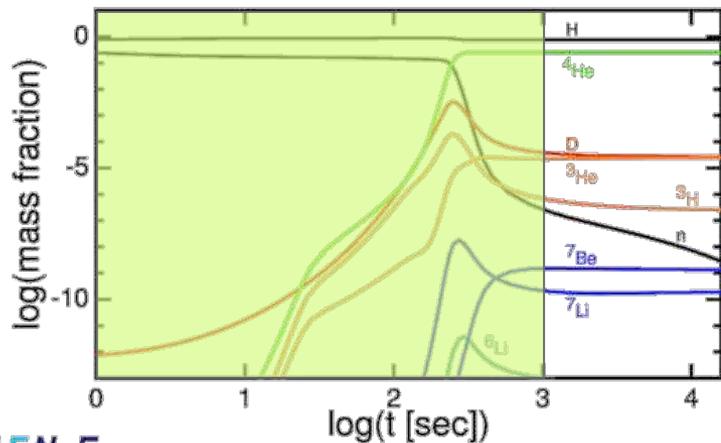


Big Bang Nucleosynthesis (BBN)



theory without free parameters
only depends on cross sections

sequence of nuclear reactions
leading to the synthesis of the
light elements (0.01 ÷ 1000 s)

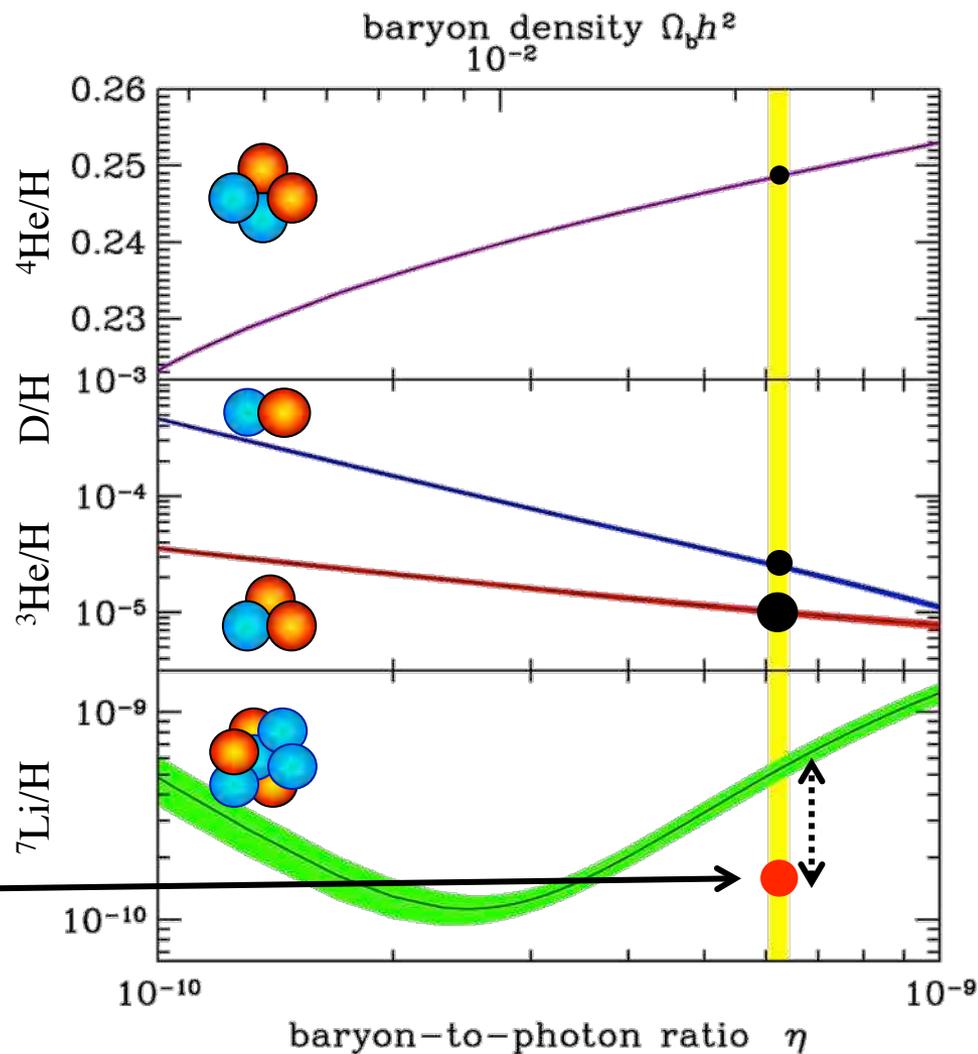


- 1 $n \rightarrow {}^1\text{H} + e^- + \bar{\nu}$
- 2 ${}^1\text{H} + n \rightleftharpoons {}^2\text{H} + \gamma$
- 3 ${}^2\text{H} + {}^1\text{H} \rightarrow {}^3\text{He} + \gamma$
- 4 ${}^2\text{H} + {}^2\text{H} \rightarrow {}^3\text{He} + n$
- 5 ${}^2\text{H} + {}^2\text{H} \rightarrow {}^3\text{H} + {}^1\text{H}$
- 6 ${}^2\text{H} + {}^3\text{H} \rightarrow {}^4\text{He} + n$
- 7 ${}^3\text{H} + {}^4\text{He} \rightarrow {}^7\text{Li} + \gamma$
- 8 ${}^3\text{He} + n \rightarrow {}^3\text{H} + {}^1\text{H}$
- 9 ${}^3\text{He} + {}^2\text{H} \rightarrow {}^4\text{He} + {}^1\text{H}$
- 10 ${}^3\text{He} + {}^4\text{He} \rightarrow {}^7\text{Be} + \gamma$
- 11 ${}^7\text{Li} + {}^1\text{H} \rightarrow {}^4\text{He} + {}^4\text{He}$
- 12 ${}^7\text{Be} + n \rightarrow {}^7\text{Li} + {}^1\text{H}$

the Cosmological Lithium Problem (CLIP)

BBN successfully predicts the abundance of primordial elements (^4He , D, ^3He)

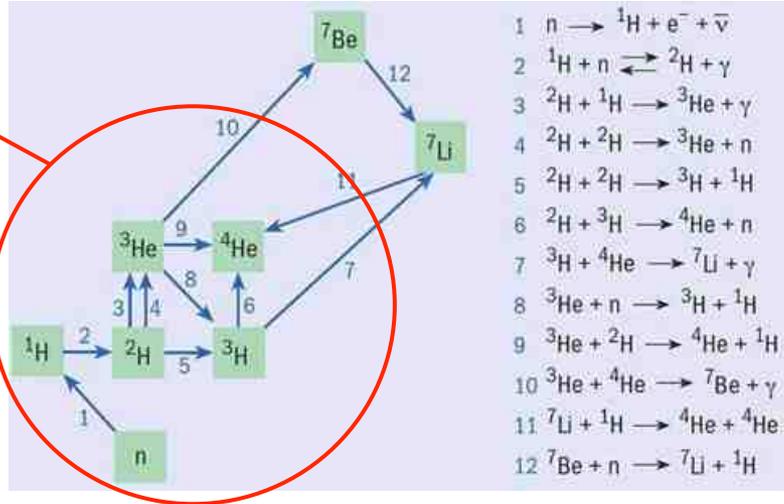
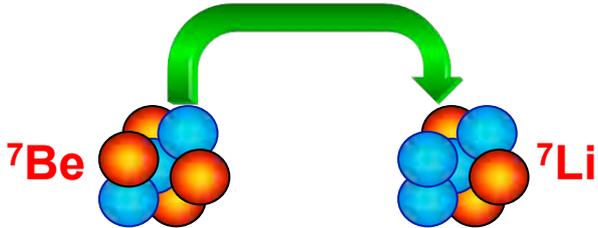
serious discrepancy for ^7Li theory/observation $\approx 2 \div 4$



the Cosmological Lithium Problem (CLIP)

a few minutes after the Big Bang
this cycle stops

≈ 95% of primordial ⁷Li
produced by EC of ⁷Be
 $t_{1/2} = 53.2d$

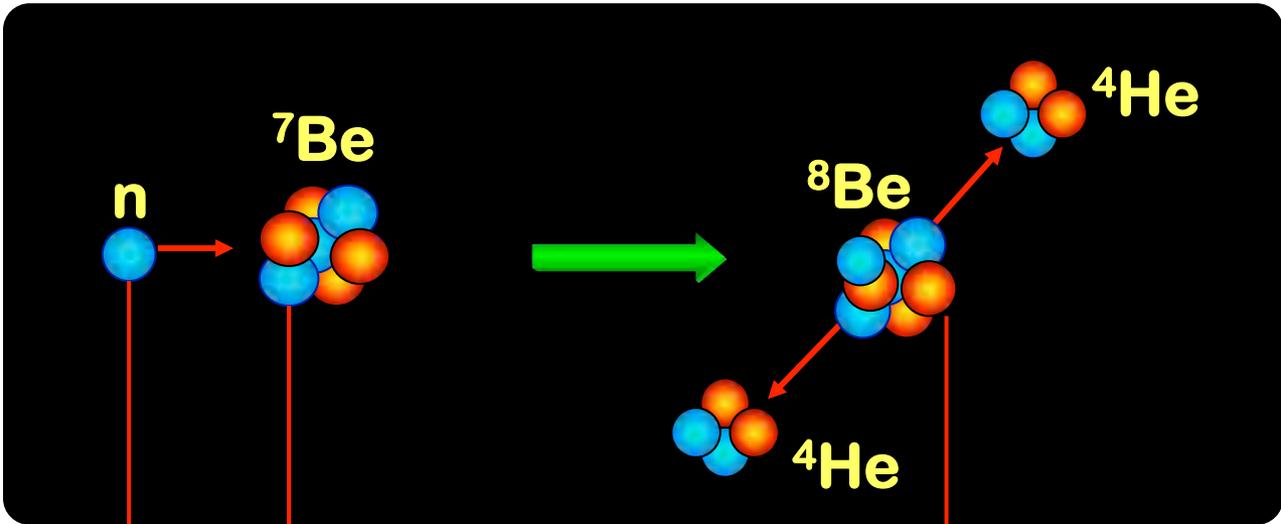


if ⁷Be has been somehow destroyed
less ⁷Li would be produced

solution to CLIP?

remaining nuclear processes to be investigated:
⁷Be(n,α)α ⁷Be(n,p)⁷Li

${}^7\text{Be}(n,\alpha)\alpha$: why impossible?



huge background
(n and gamma)

${}^7\text{Be}$ production $\ll 1 \mu\text{g}$

${}^7\text{Be}$ target highly radioactive (13 GBq/ μg),
478 keV gamma rays, $T_{1/2} \approx 53\text{d}$

intensity of available n-beams not enough

${}^7\text{Be}(n,\alpha)\alpha$ the only existing data is at 25 meV
(Bassi et al., 1963)

measured with thermal neutrons inside a reactor
(Ispra, Italy)

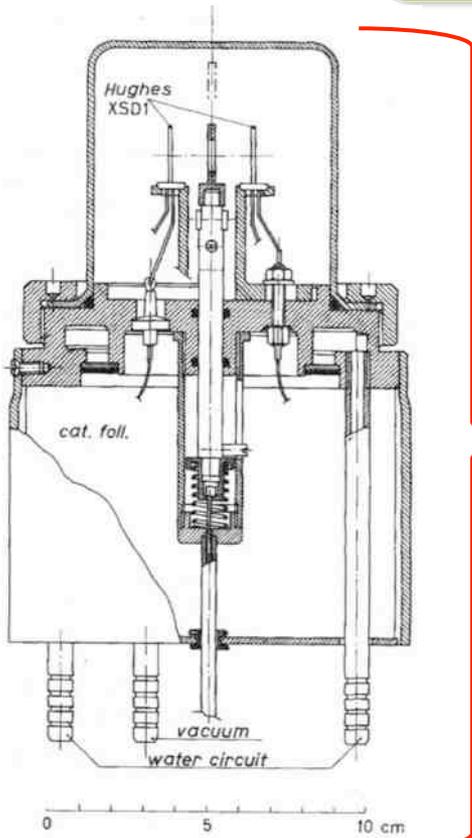
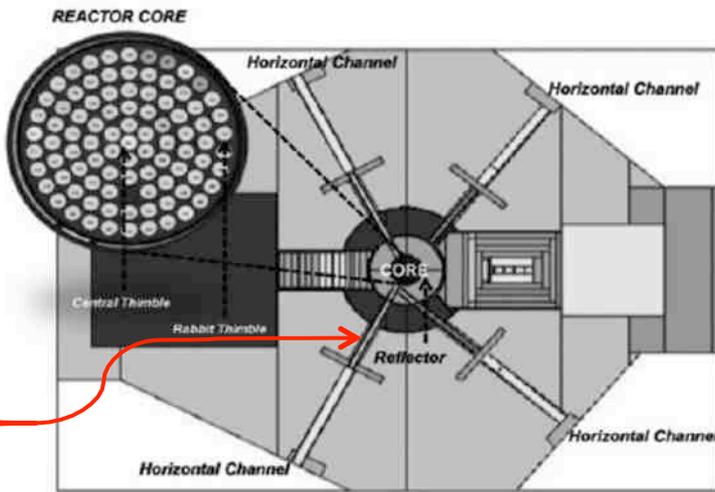
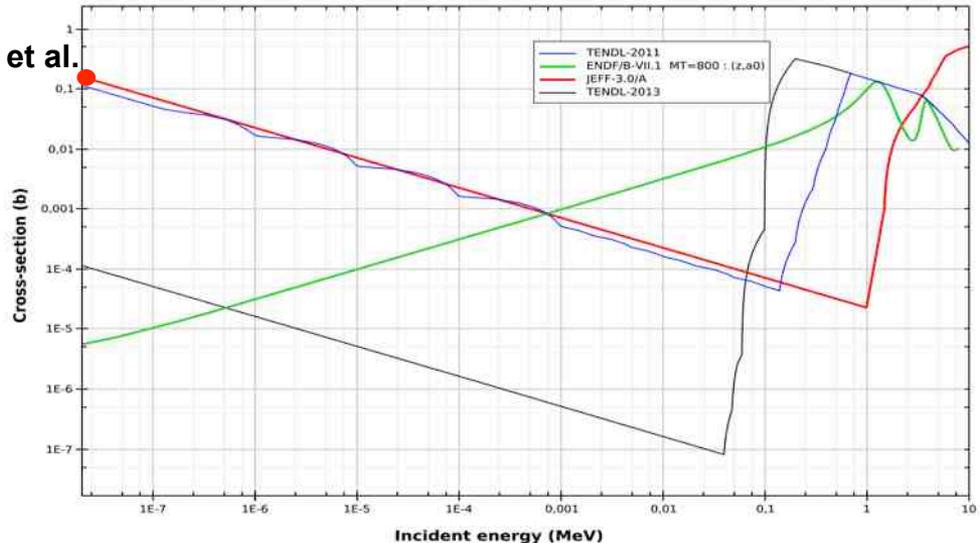


Fig. 1.



Bassi et al.



$^7\text{Be}(n,\alpha)\alpha$: how?

beam intensity not enough



n_TOF beam @ EAR2



^7Be production $\ll 1 \mu\text{g}$



PSI Zurich: extraction and purification of cooling water from SINQ



highly radioactive target



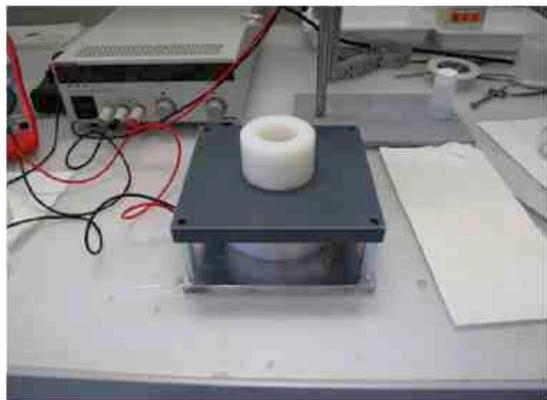
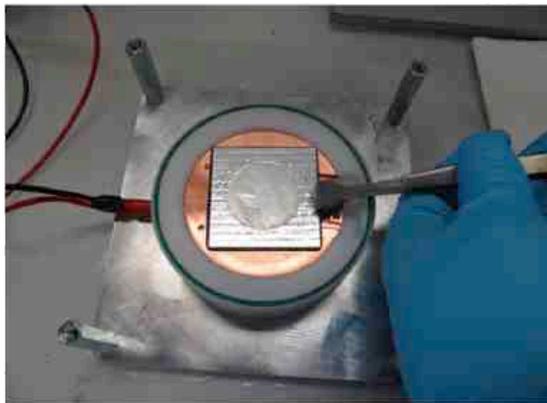

huge background



**2 targets produced
in a hot cell**

Deposition techniques

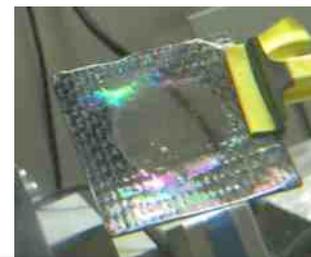
Molecular plating



Vaporization of droplets



**molecular plating
18 GBq**

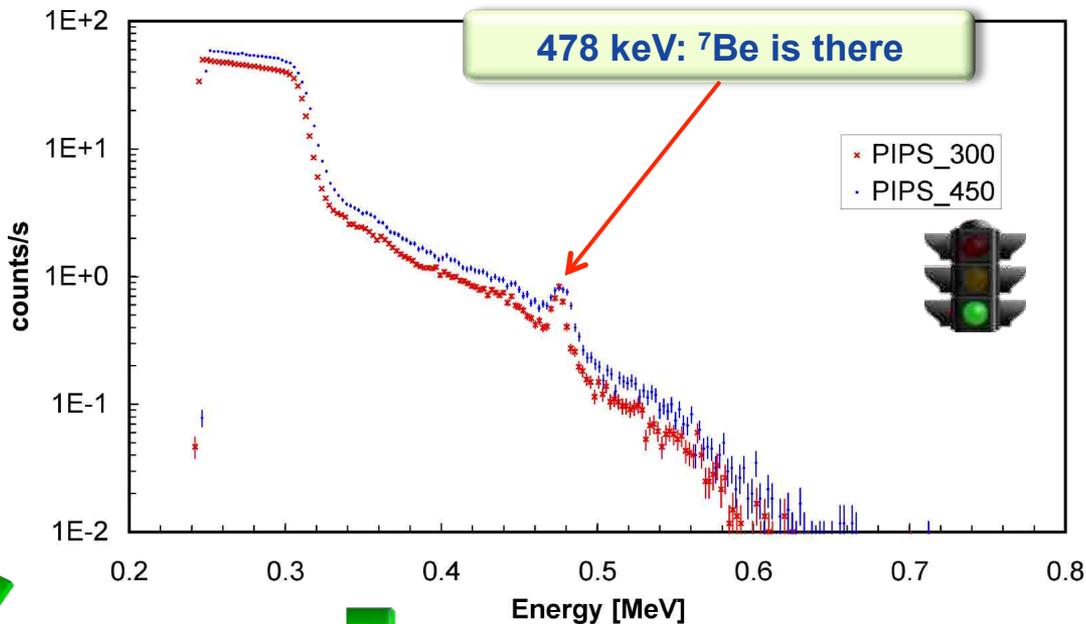
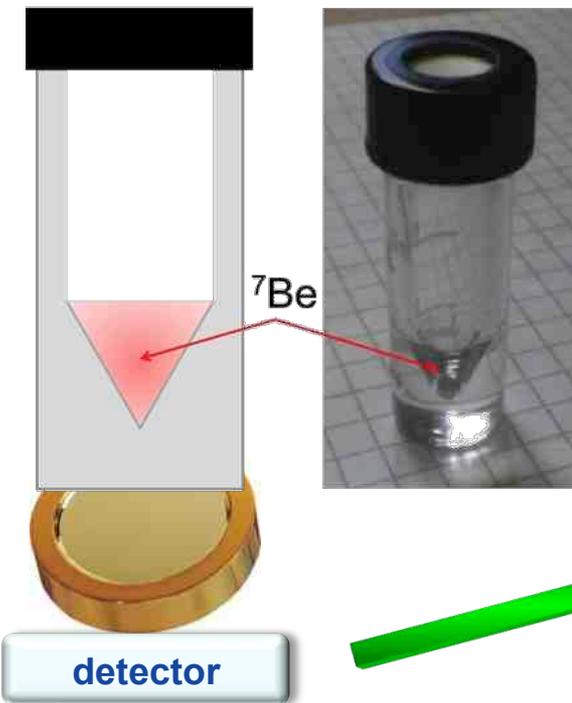


**vaporization
17 GBq**

PSI 23-25.11.2015 CHANDA – workshop on target preparation – the needs and the possibilities

previously tested and checked with (⁹Be + trace quantity of ⁷Be)

(⁹Be + trace quantity of ⁷Be): check ⁷Be presence



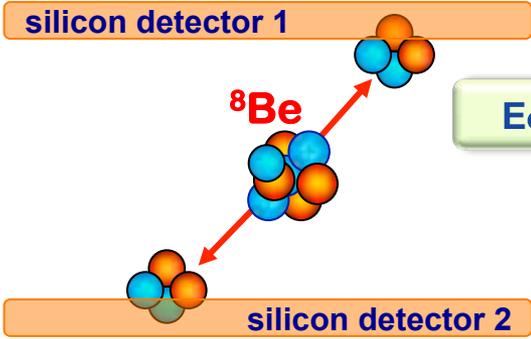
**produced a test target
⁹Be + trace quantity of ⁷Be**

**check the transferred Be: > 99.5%
(activity measurement)**

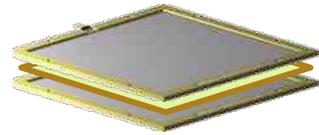
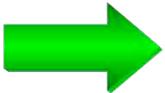
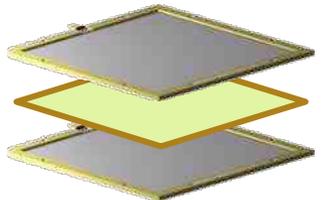


${}^7\text{Be}(n,\alpha)\alpha$: detection technique

α - α coincidence for background suppression (n and gamma)

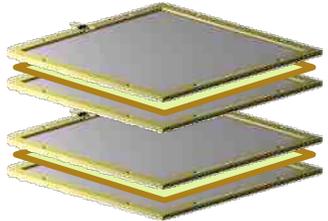


$E_\alpha \approx 8 \text{ MeV}$



large solid angle

redundancy (to evaluate systematic errors and spurious counts)





home-made detection system: no wizardry

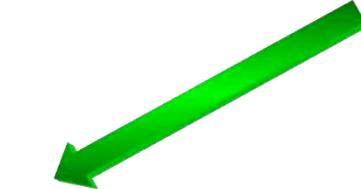
hands-on work



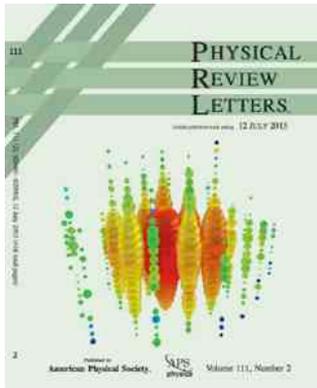
data acquisition



data analysis

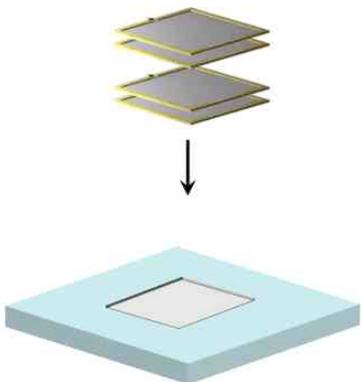


publication

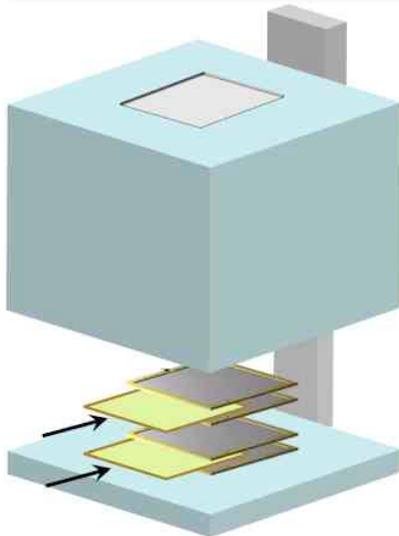


$^7\text{Be}(n,\alpha)\alpha$: shielding and mechanics

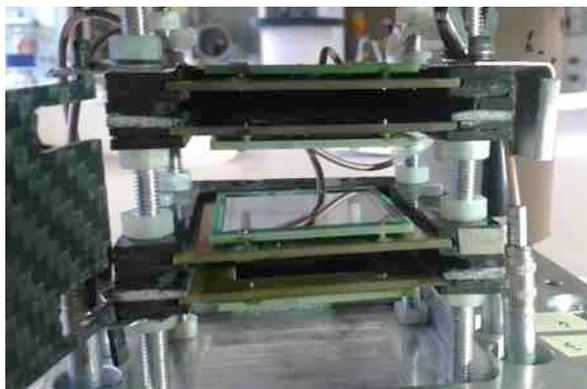
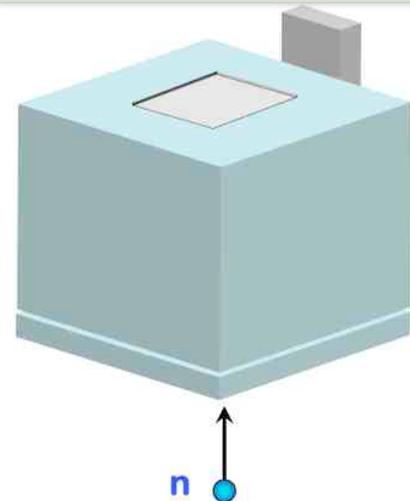
detector installation



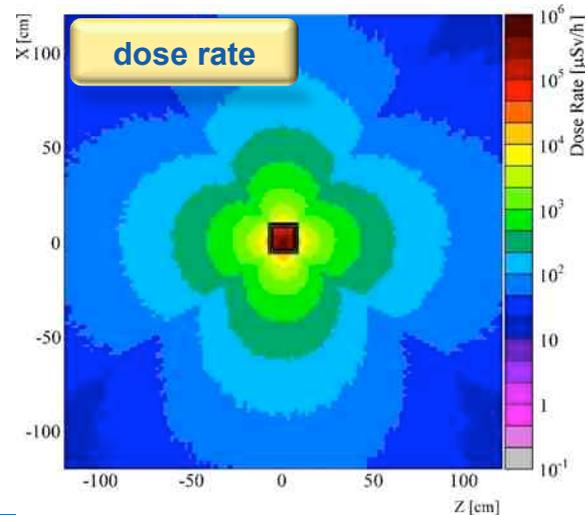
target insertion



pneumatic closure



INFN-E



target insertion (inside a hot cell)



chamber extraction



transport PSI - CERN



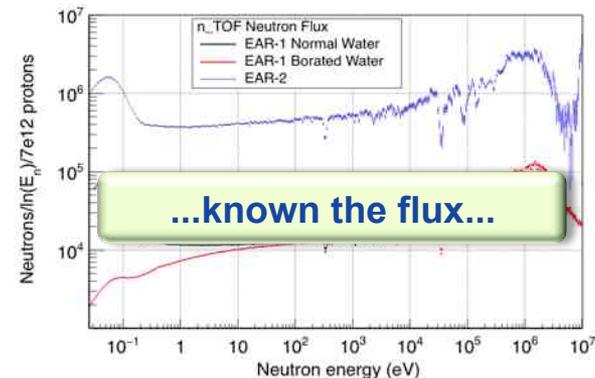
installation at EAR2 n_TOF

$^7\text{Be}(n,\alpha)\alpha$: the measurement?

1 “burst” of neutrons per second, duration $\approx 10\text{ms}$

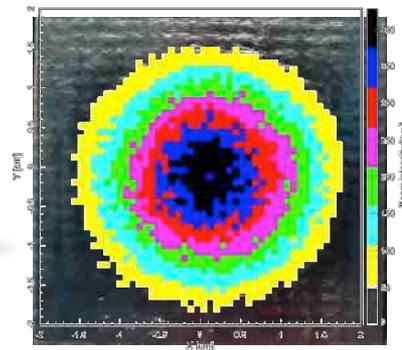
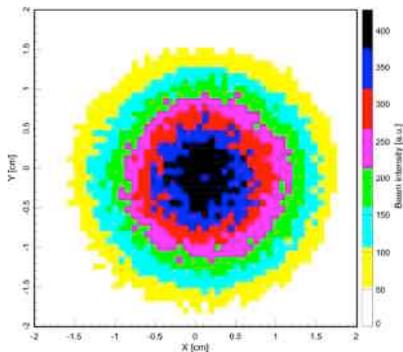
incident neutron energy for each α - α coincidence

from the number of coincidence events one can deduce $\sigma(E)$...



...the spatial distribution of the beam...

...and the target features



**incident neutron energy:
the Time-Of-Flight technique**
(invented long time ago: Orazi and Curiazi legend)

start



Duella fra Orazi e Curiazi



finish



run time (flight)
proportional to $1/v$

measured t (TOF)

$v = s/t$



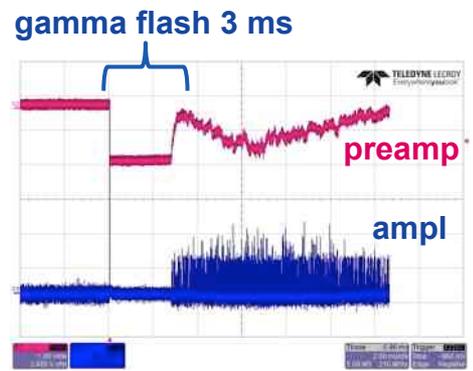
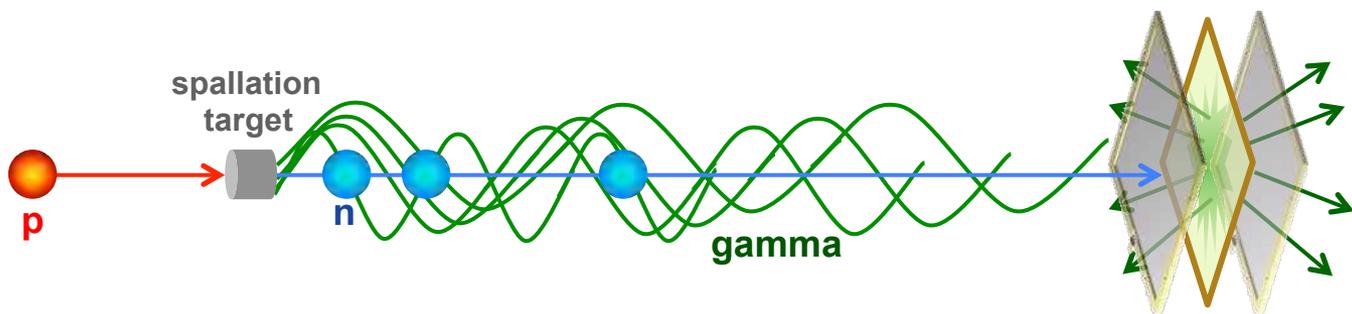
compute v

$E = \frac{1}{2} mv^2$



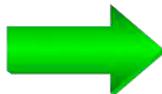
compute E

${}^7\text{Be}(n,\alpha)\alpha$: 4 crucial questions

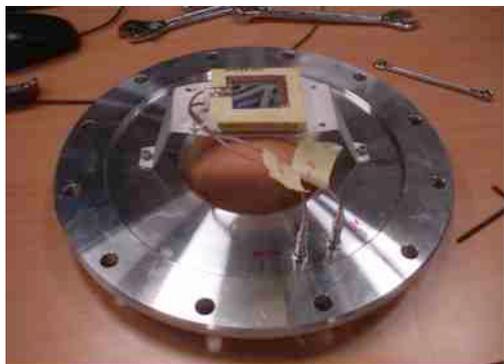


- | | | |
|---|---|--|
| Q1: damage from neutrons | ➔ | will the detectors survive? |
| Q2: energy resolution | ➔ | worsens, is it enough to select α 's? |
| Q3: gamma flash | ➔ | fast electronics to recover after blinding? |
| Q4: gammas from ${}^7\text{Be}$ | ➔ | will the system work under 1GHz gammas? |

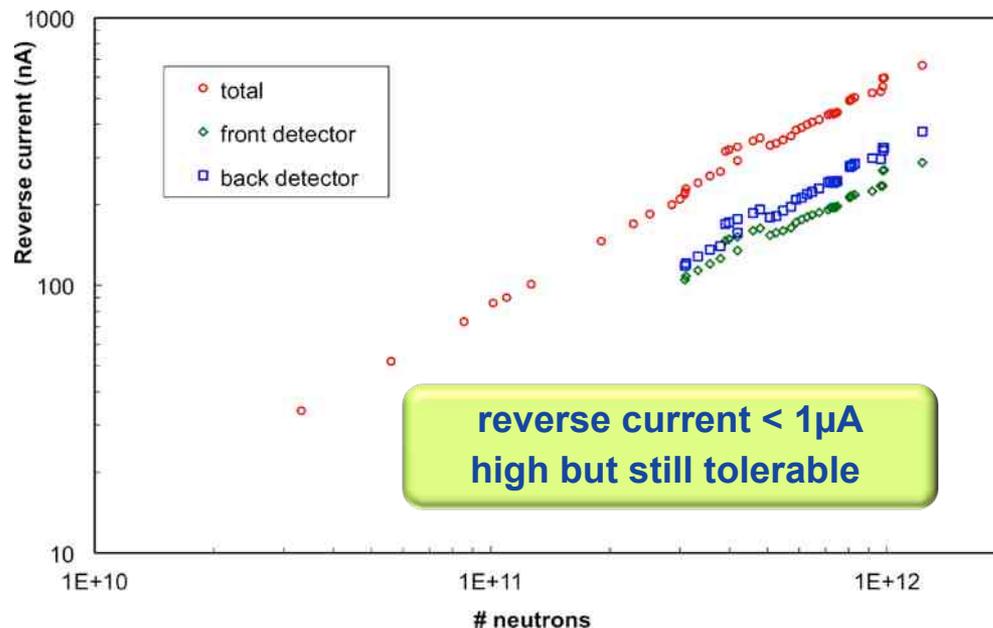
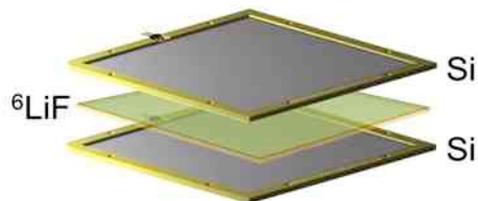
Q1: damage from neutrons



will the detectors survive?



produced a sandwich Si-⁶LiF-Si
installed in EAR2 near the beam dump
45 days irradiation, $\approx 10^{12}$ neutrons

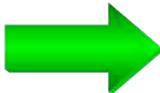


reverse current < 1 μ A
high but still tolerable

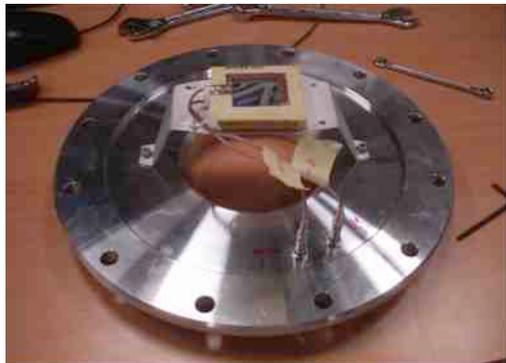
Q1: OK



Q2: energy resolution



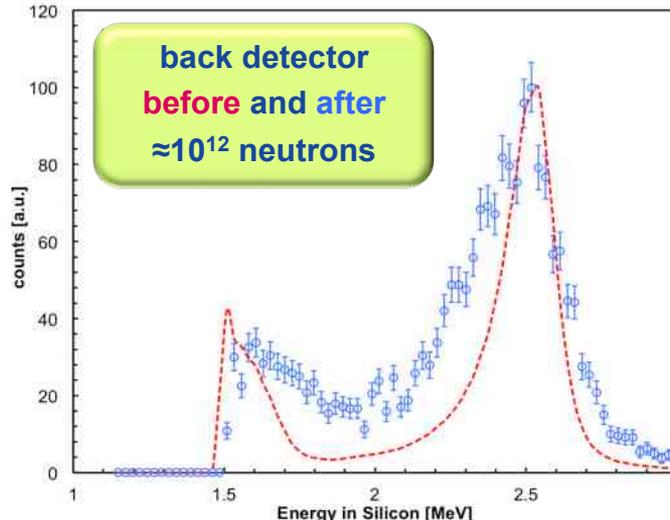
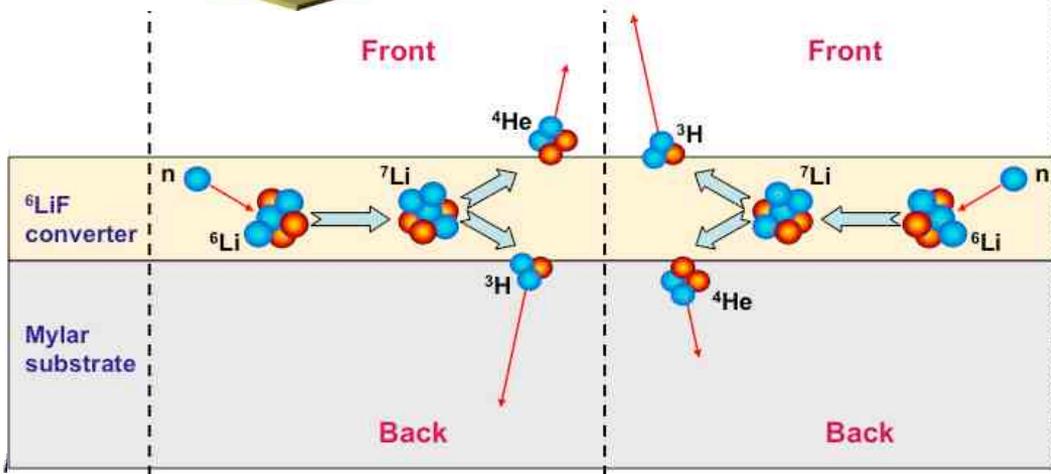
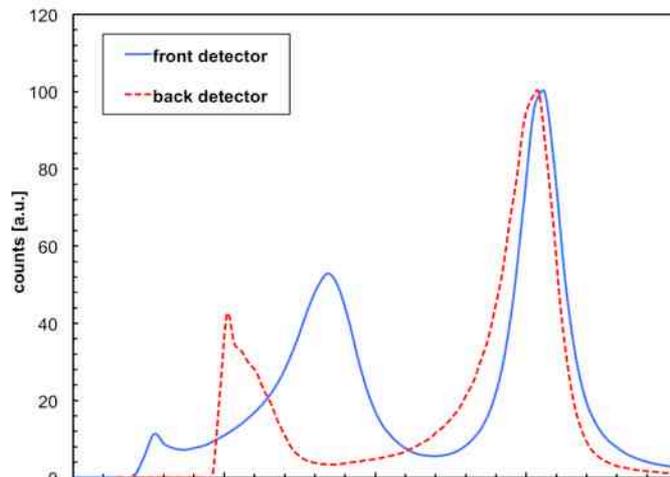
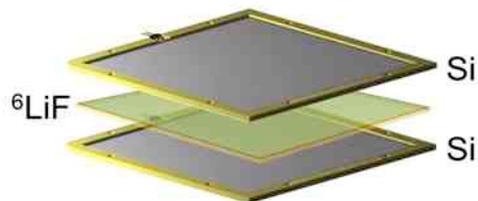
worsens, is it enough to select α 's?



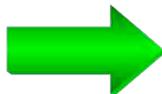
Q2: OK



${}^6\text{Li}(n,\alpha)t$

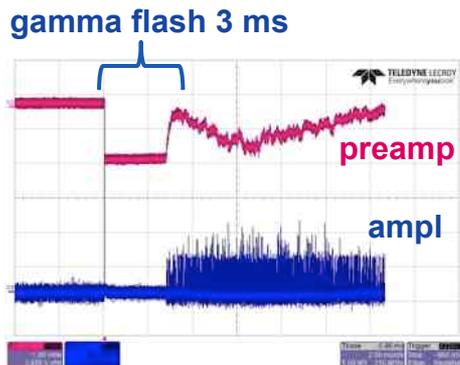
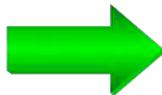


Q3: gamma flash



fast electronics to recover after blinding?

traditional preamp

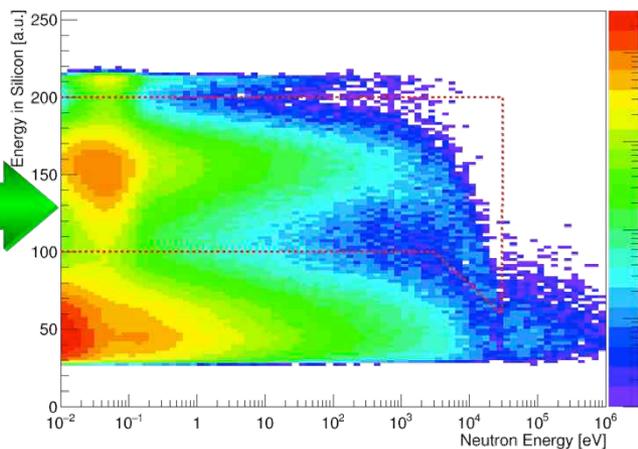


limit 0.1 ÷ 1 eV

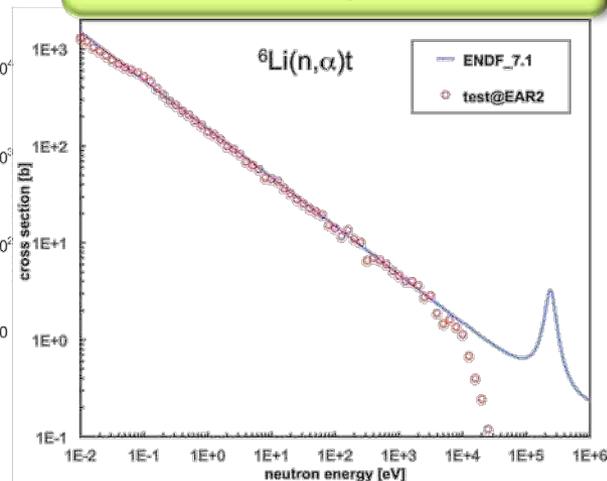
Mesytec MPR-16
lin-log preamp



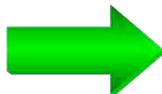
Q3: OK



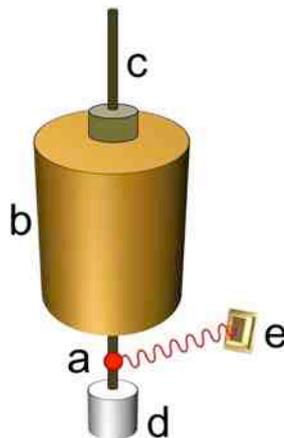
sigma ⁶Li(n,α)t
reconstructed up to 1 ÷ 10 keV



Q4: gammas from ^7Be



will the system work under 1GHz gammas?



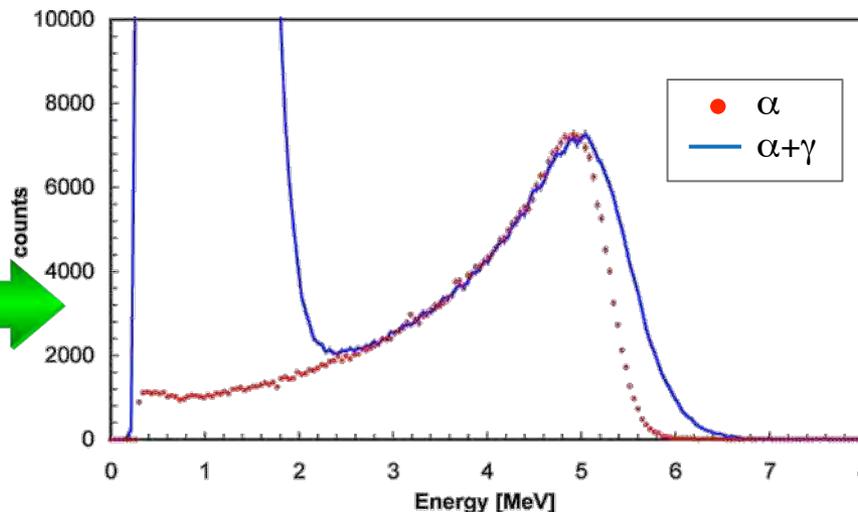
^{137}Cs gamma source at LNS (39 GBq)
reference Geiger: $10^9 \gamma/\text{s}$ on the detector
 $\rightarrow 10^6 \div 10^7$ signals/s from the detector

traditional
preamp



saturation:
no signal!

Mesytec MPR-16
lin-log preamp



Q4: OK

Conclusion 1

neutron beam intensity



EAR2 n_TOF @ CERN

production of ^7Be targets



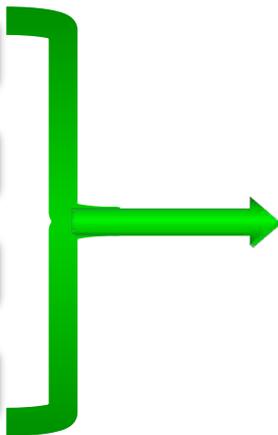
PSI @ Zurich

Q1: damage from neutrons

Q2: energy resolution

Q3: gamma flash

Q4: gammas from ^7Be



LNS @ Catania



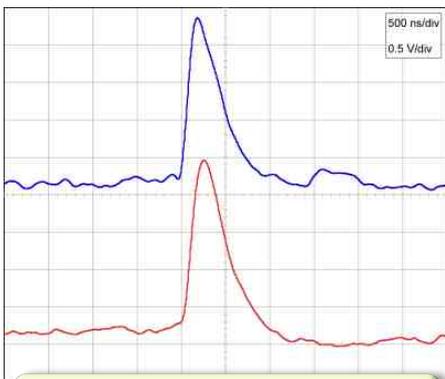
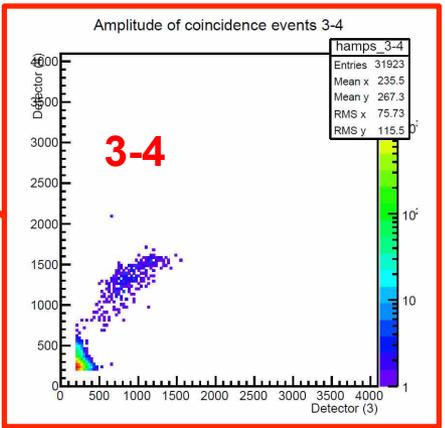
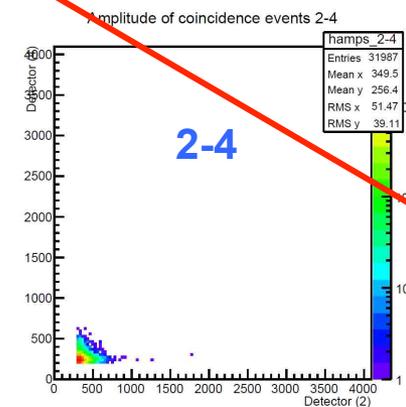
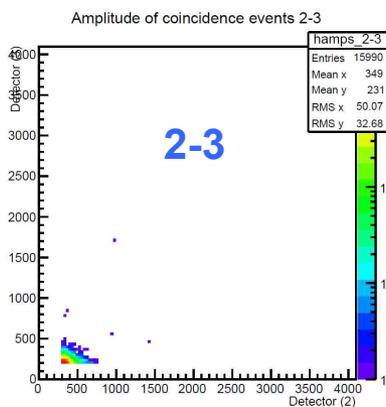
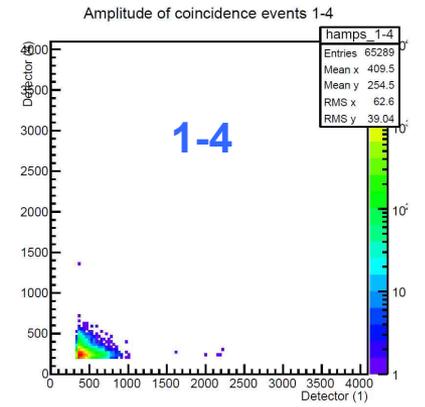
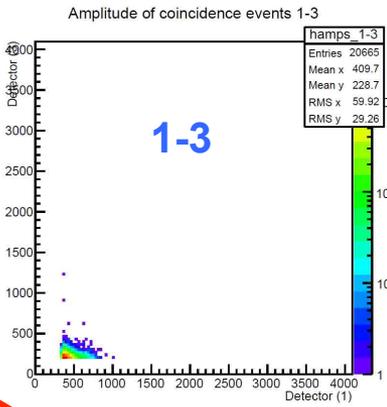
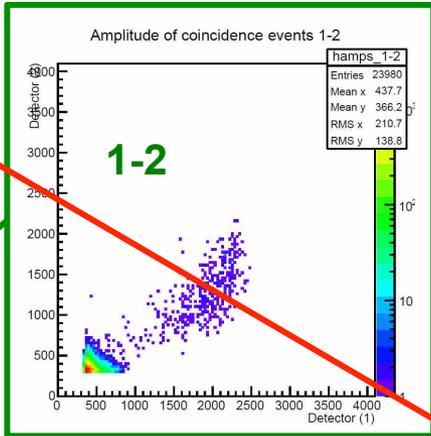
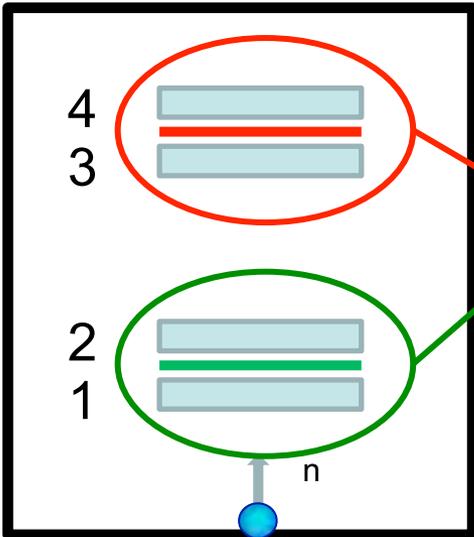
"impossibile" measurement?



possible measurement !!!



measurement results

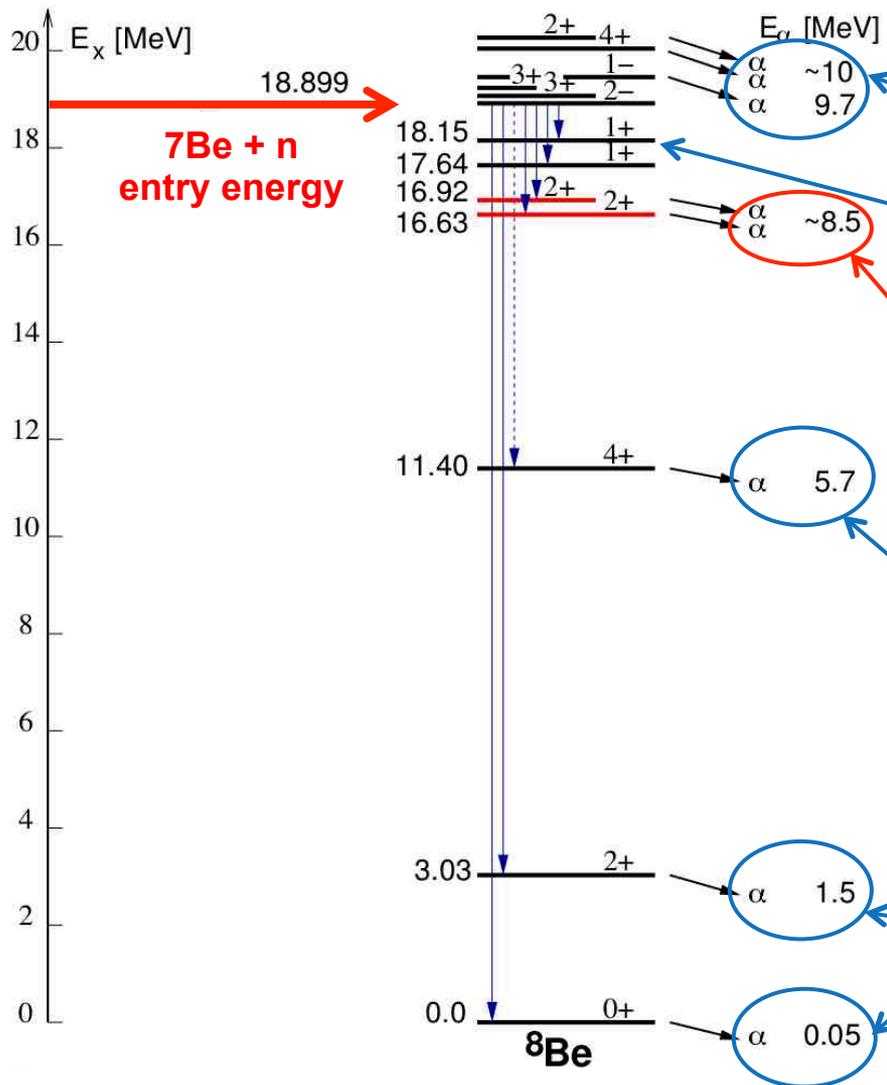


α - α coincidence

lower energy threshold at 2 MeV

⁸Be level scheme and decay channels

by A. Mengoni



Possible 2α channels not observed due to experimental n -energy cutoff

Forbidden (selection rules)

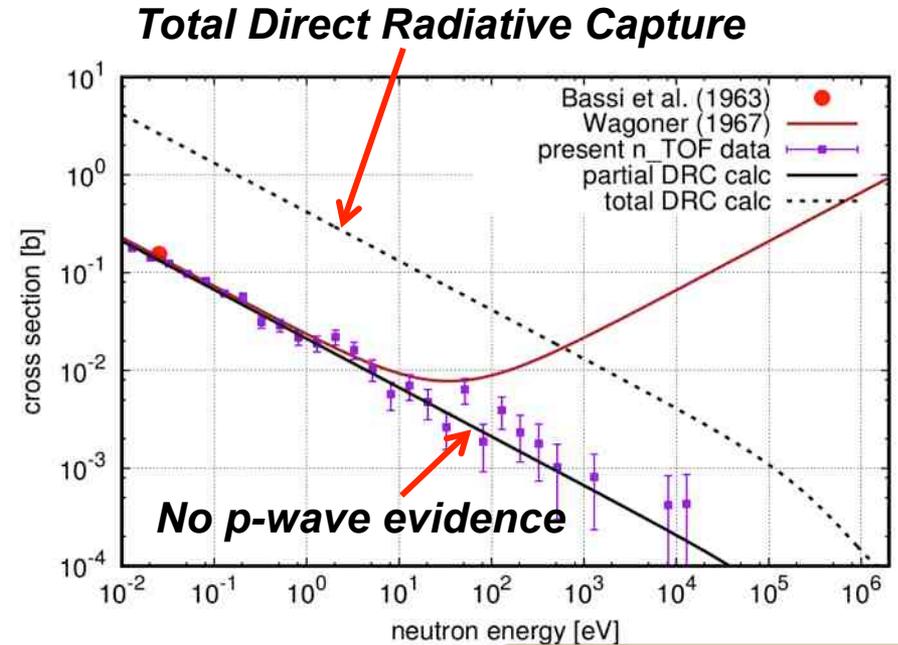
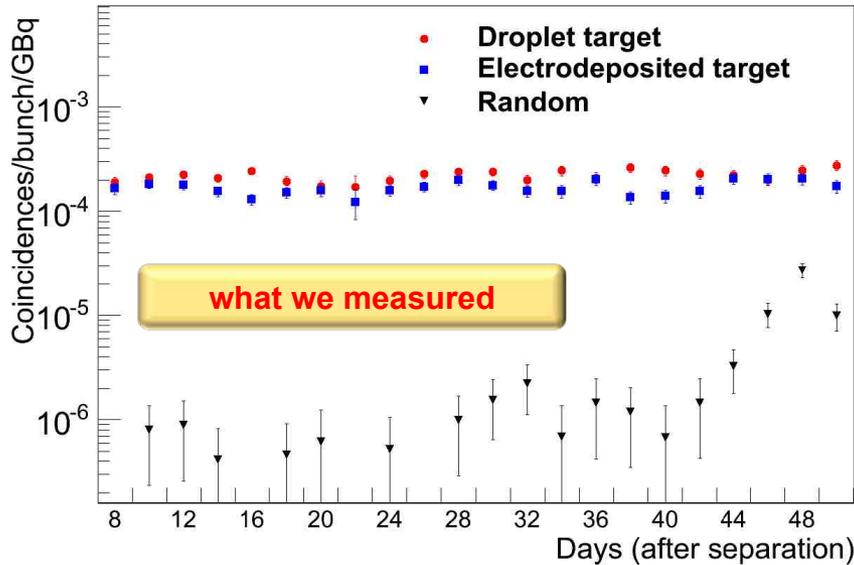
Observed γ - 2α channels

Suppressed γ - 2α channel (negligible, not observed)

Predicted relevant γ - 2α channels not observed due to experimental apparatus energy threshold



Impact on Cosmological Lithium Problem



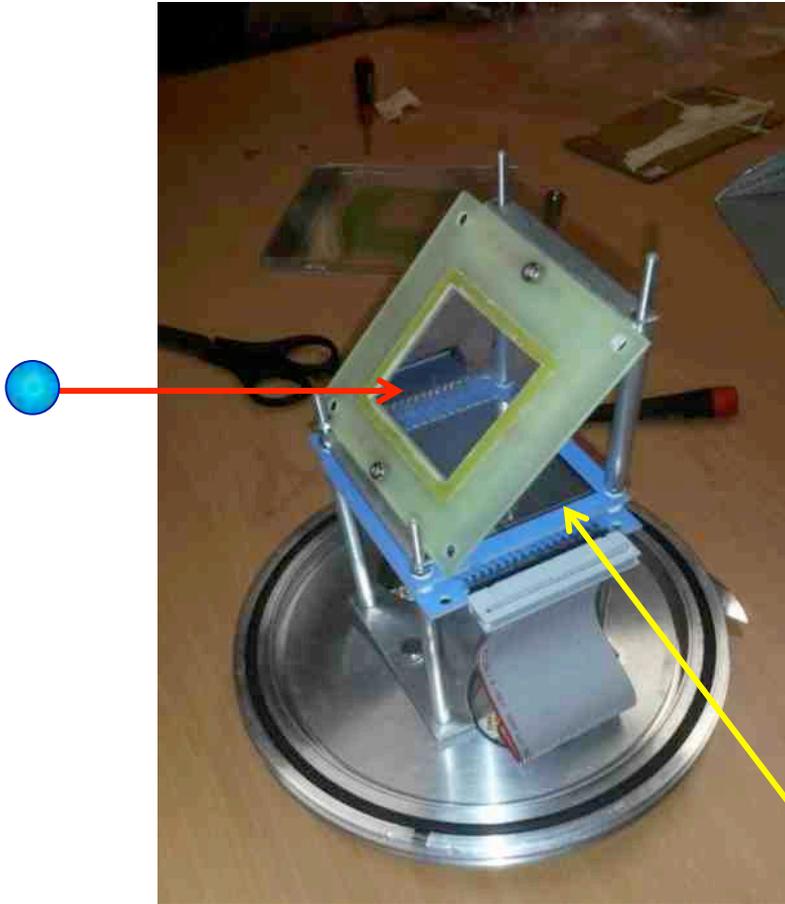
Article submitted to PRL

the n TOF results hint to a minor role of this reaction in BBN, leaving the long-standing Cosmological Lithium problem unsolved

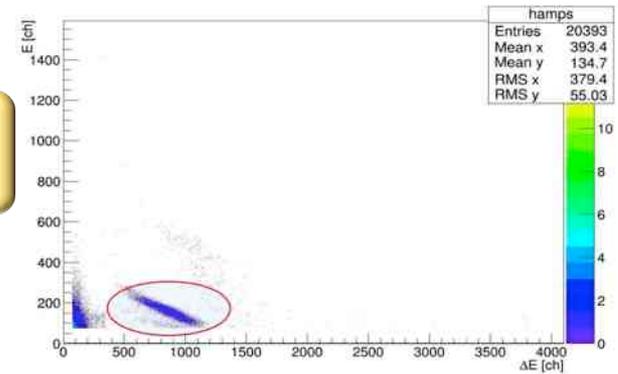
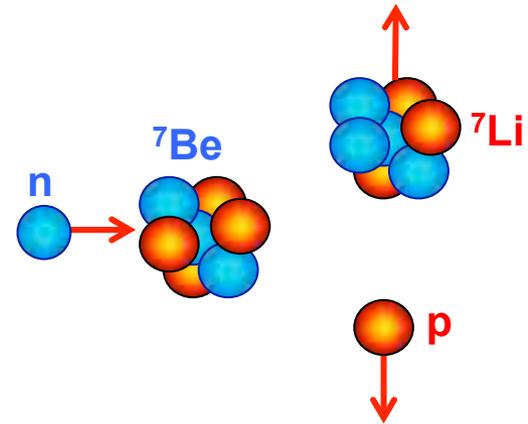
the reaction rate currently used in BBN calculations requires substantial revision

${}^7\text{Be}(n,p){}^7\text{Li}$

experiment just completed



silicon strip telescope



Conclusion 2

the past

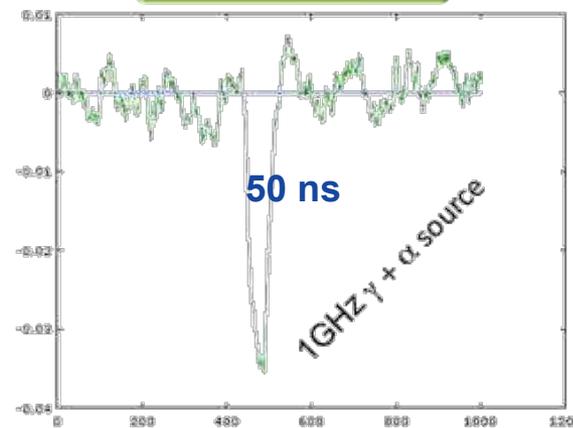
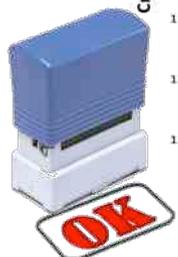
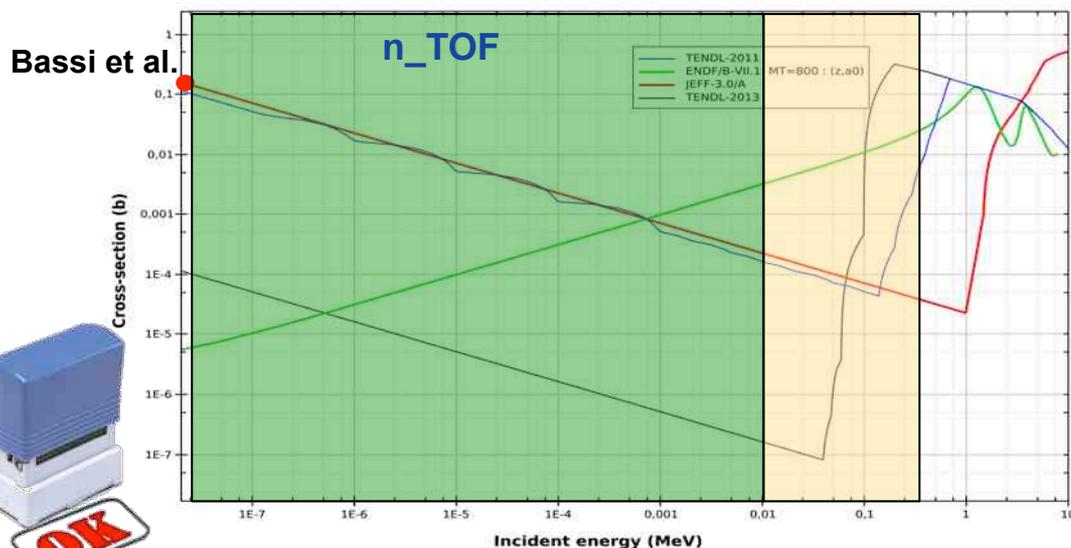
${}^7\text{Be}(n,\alpha)\alpha$ only existing data at 25 meV
(Bassi et al., 1963)

the present

August-October 2015: just measured ${}^7\text{Be}(n,\alpha)\alpha$

article submitted to Physical Review Letters

the future?



new preamp @ LNS ?

main characters



A.Musumarra

L.Cosentino



A.Pappalardo



N.Colonna



M.Barbagallo

without them the experiment would not have been possible

other characters

M.Piscopo, S.Russo, Lab. Bersagli LNS, Officina LNS, Elettronica LNS,
 E.Maugeri, D.Schumann, O.Aberle, B.Langhans, E.Chiaveri, F.Gunsing
and the whole n_TOF collaboration
 ...and I apologize if I forgot somebody !!!



THE END
thank you !!!

