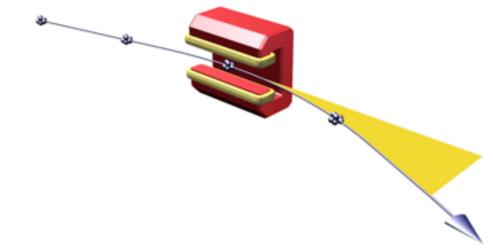




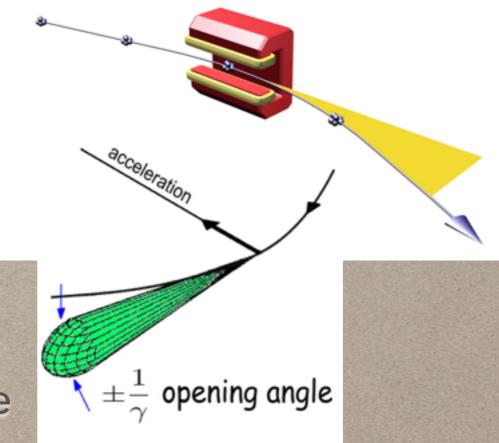
SYNCHROTRON RADIATION CONTRIBUTION IN THE INTERACTION REGION IN FCC-HH FRANCESCO COLLAMATI 24-05-2016

1

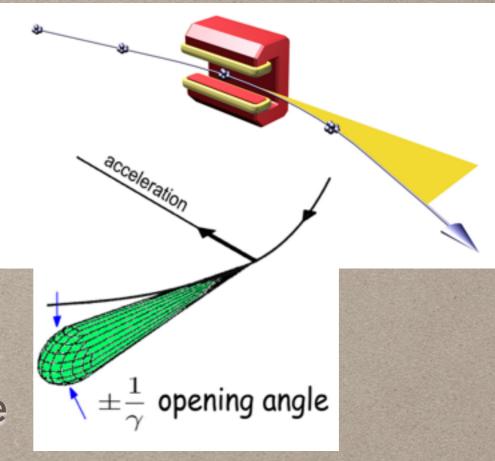
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- This radiation is distributed in a cone tangential to the moving direction of the emitting particle

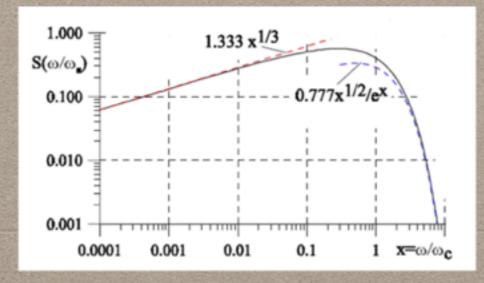


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 The emitted radiation is composed by several harmonics of the revolution frequency:

2



critical frequency (& energy): the frequency above which half of the power is emitted

$$v_c = \frac{3}{2}c\frac{\gamma^3}{\rho}$$

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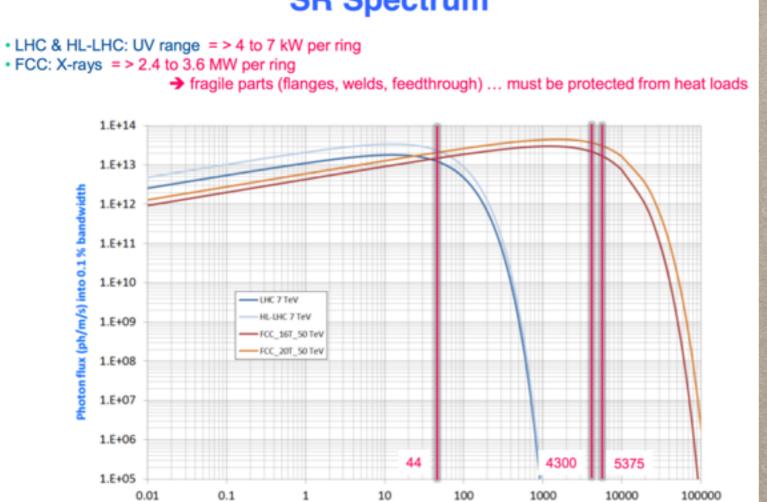
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- However, in Very High Energy p-p colliders the effect starts to be visible, and should be carefully evaluated
 - I.e. LHC, FCC-hh...

 Energy distribution of synchrotron radiation photons for LHC and FCC-hh:

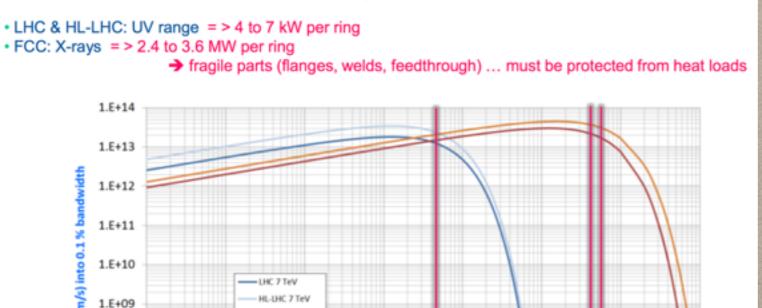
SR Spectrum



Photon energy (eV)

 Energy distribution of synchrotron radiation photons for LHC and FCC-hh:

SR Spectrum



44

Photon energy (eV)

100

10

FCC_16T_50 TeV FCC_20T_50 TeV

1

1.E+08

1.E+07

1.E+06

1.E+05

0.01

0.1

 $E_{_{FCC_{hh}}} \propto 7 \times E_{_{LHC}}$

5375

100000

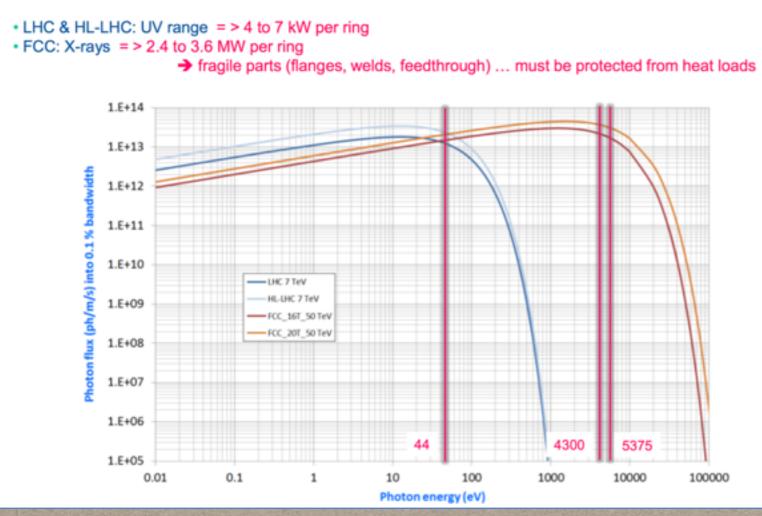
10000

4300

1000

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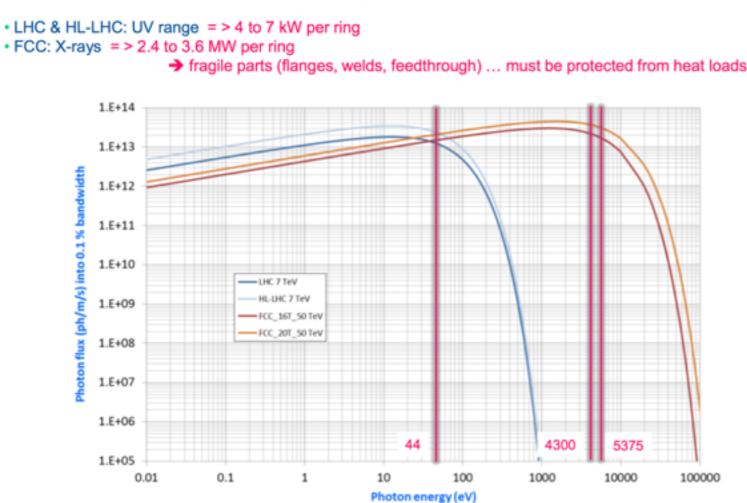
SR Spectrum



 $E_{FCC_{hh}} \propto 7 \times E_{LHC}$ $P_{SR_{FCC_{hh}}} \propto 170 \times P_{SR_{LHC}}$

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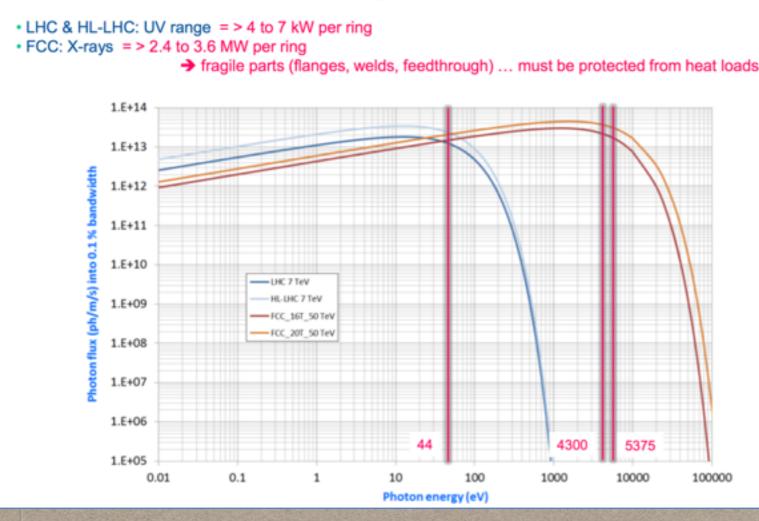
SR Spectrum



 $E_{FCC_{hh}} \propto 7 \times E_{LHC}$ $P_{SR_{FCC_{hh}}} \propto 170 \times P_{SR_{LHC}}$ $E_{Crit_{FCC_{hh}}} \propto 100 \times E_{Crit_{LHC}}$

 Energy distribution of synchrotron radiation photons for LHC and FCC-hh:

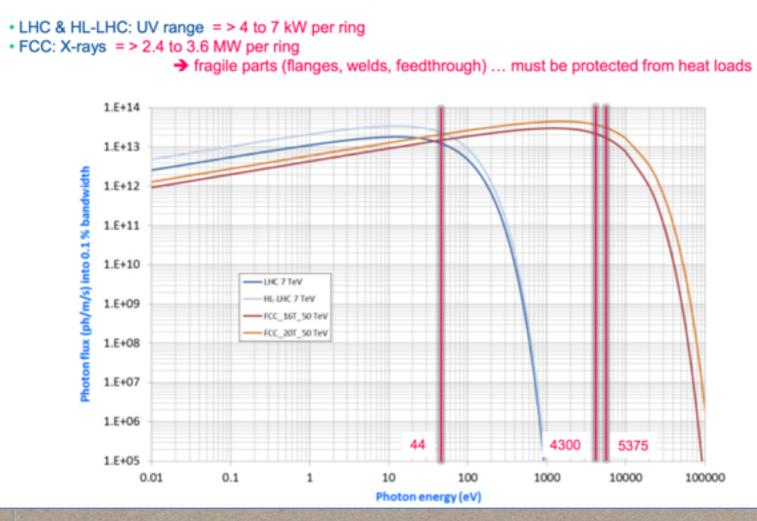
SR Spectrum



$$\begin{split} E_{FCC_{hh}} &\propto 7 \times E_{LHC} \\ P_{SR_{FCC_{hh}}} &\propto 170 \times P_{SR_{LHC}} \\ E_{Crit_{FCC_{hh}}} &\propto 100 \times E_{Crit_{LHC}} \\ & \checkmark \\ \end{split}$$

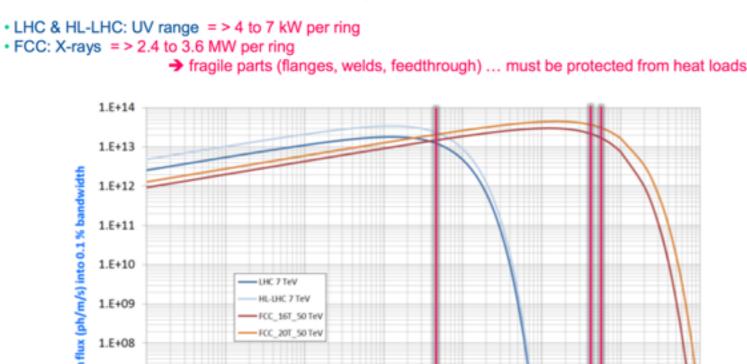
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The total radiated power will probably be still low, but the much greater photon energy demands for a careful evaluation

5375

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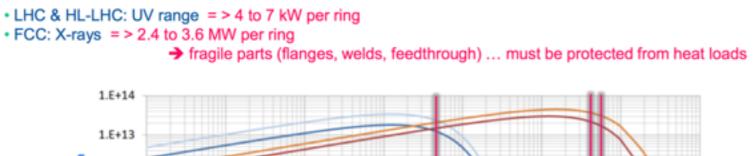
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 Energy distribution of synchrotron radiation photons for LHC and FCC-hh:

SR Spectrum



1.E+12 1.E+11 1.E+10 LHC 7 TeV 1.E+09 FCC 16T 50 TeV FCC 20T 50 TeV 1.E+08 1.E+07 1.E+06 44 4300 5375 1.E+05 0.01 0.1 100 1000 10000 100000 1 10 Photon energy (eV)

$$\begin{split} E_{FCC_{hh}} &\propto 7 \times E_{LHC} \\ P_{SR_{FCC_{hh}}} &\propto 170 \times P_{SR_{LHC}} \\ E_{Crit_{FCC_{hh}}} &\propto 100 \times E_{Crit_{LHC}} \\ & & & & & & \\ & & & & & & & \\ & & & & & & & \\ & & & & & & & \\ & & & & & & & \\ & & & & & & & \\ & & & & & & & \\ & & & & & & & \\ & & & & & & & \\ & & & & & & & \\ & & & & & & & \\ & & & & & & & \\ & & & & & & & \\ & & & & & \\ & & & & & & \\ & & & & & \\ & & & & & & & \\ & & & & & & \\ & & & & & & & \\ & & & & & & \\ & & & & &$$

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More, and more energetic photons!

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 - Edge effect: rise in the critical frequency at the borders of the magnets due to magnetic gradient

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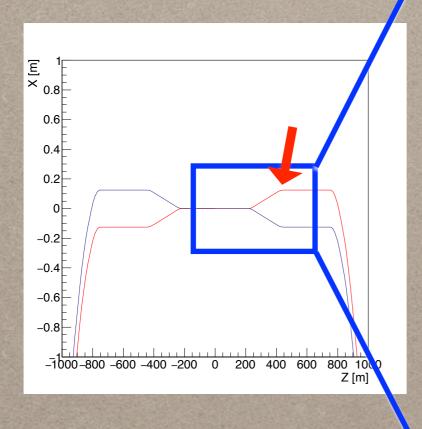
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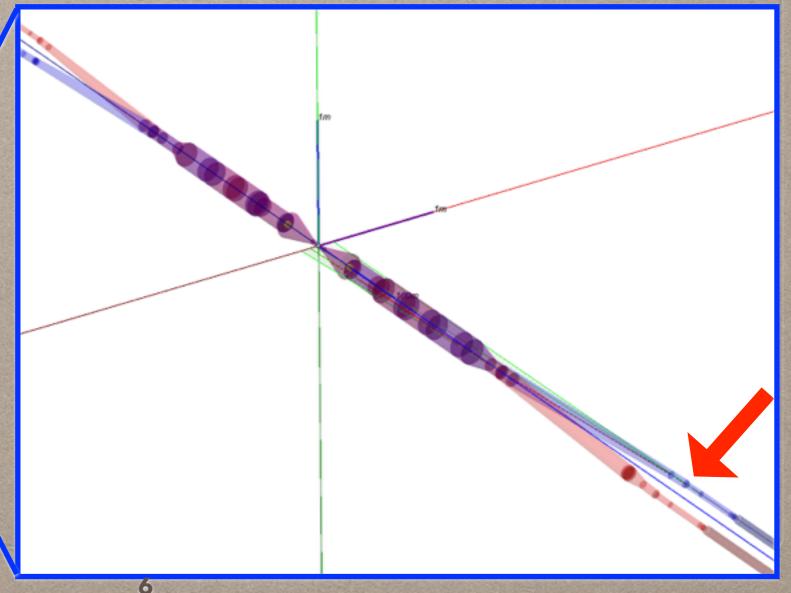
→ The radiation cone is very narrow:

$$\gamma_p = \frac{E_p}{m_p} = \frac{50TeV}{938MeV} \sim 5 \times 10^4$$
$$\frac{1}{\gamma_p} \sim 1.9 \times 10^{-5} rad \sim 10^{-3} grd$$

MY TASK

- My (current) task is to evaluate the contribution of synchrotron radiation photons emitted in the last bending magnets into the interaction region
- Tool used: MDISim

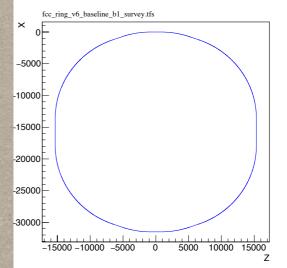




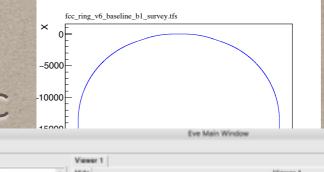
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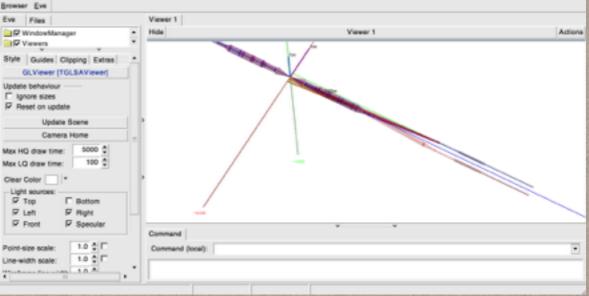
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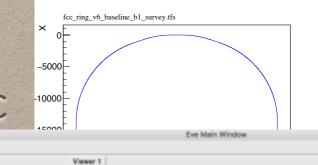
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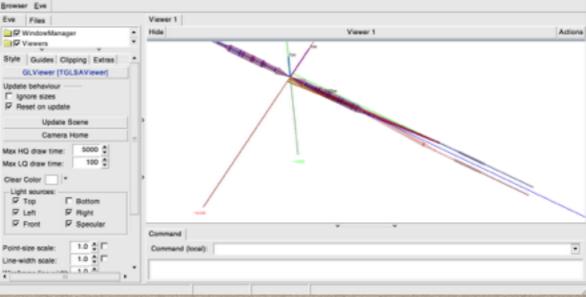




iele	NAME KEYWORD	S	L	Angle	Ecrit ngamBend	rho	B BETX	SIGX	divx	Power	frac>10MeV	ngam*npart Egamt	ot Emean
		m	m		keV	m	T m	mm	mrad	kW		GeV	keV
23	MBXA.A4RA.H SBEND	231.3	12.5	-0.0003199	1.146 0.1795	39079.0	-4.267862523.1459	1.6066	0.0037	0.03221	Θ	1.8e+10 6.34e+0	3 0.353
25	MBXA.B4RA.H SBEND	245.3	12.5	-0.0003199	1.146 0.1795	39079.0	-4.267858576.9378	1.5551	0.0037	0.03221	Θ	1.8e+10 6.34e+0	3 0.353
27	MBRD.A4RA.H1 SBEND	426.9	15	0.0003199	0.9552 0.1795	46894.8	3.556519052.8783	0.8869	0.0037	0.02684	Θ	1.8e+10 5.28e+0	3 0.294
29	MBRD.B4RA.H1 SBEND	443.4	15	0.0003199	0.9552 0.1795	46894.8	3.556516533.0927	0.8262	0.0037	0.02684	Θ	1.8e+10 5.28e+0	3 0.294
51	MBS.A8RA.H1 SBEND	767.1	13.4	0.00128	4.279 0.7183	10468.8	15.9313 133.1204	0.0741	0.0013	0.481	θ	7.18e+10 9.46e+04	1.32
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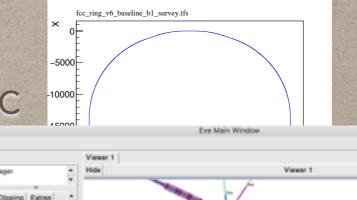
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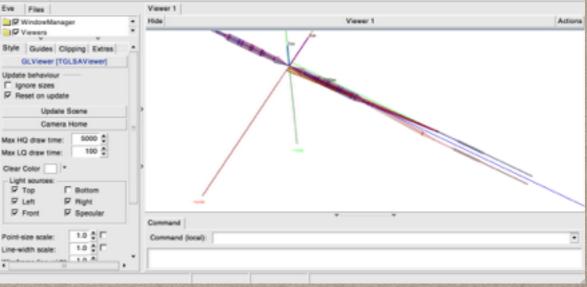




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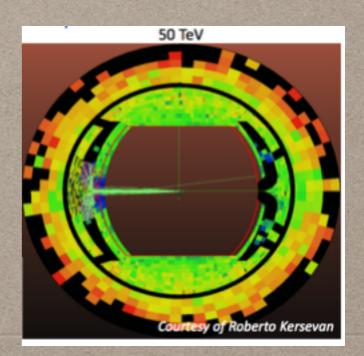


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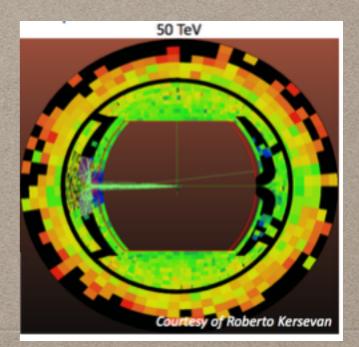
Browser Eve

Import geometry and SR in Geant to perform full simulation

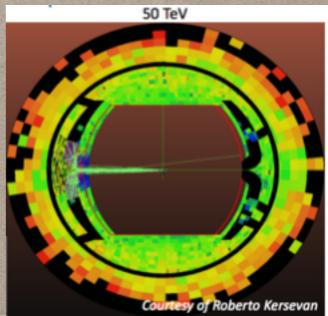
- Evaluate the solid angle of photons reaching the interaction region, assuming "worst case scenario" to see if the contribution is already negligible
 - Is "total absorption of the pipe" a "too good scenario"!?



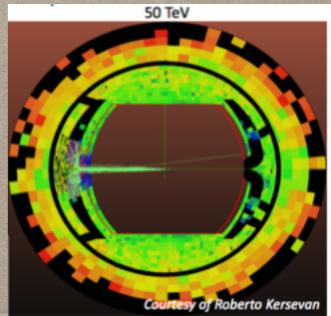
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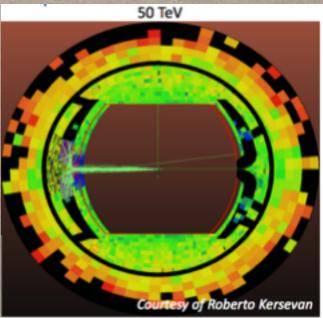
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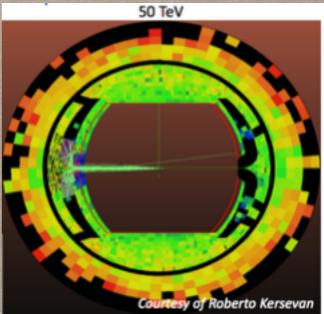
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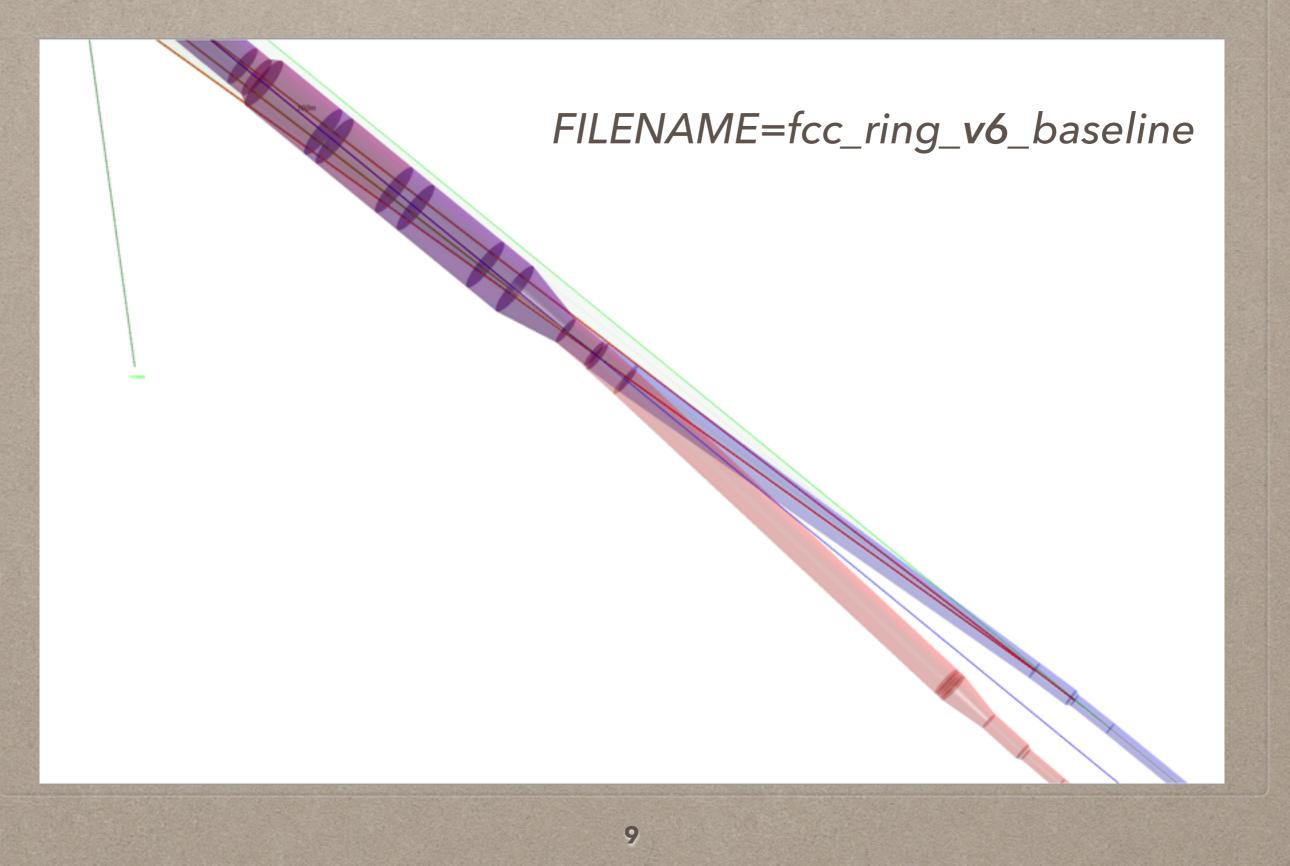
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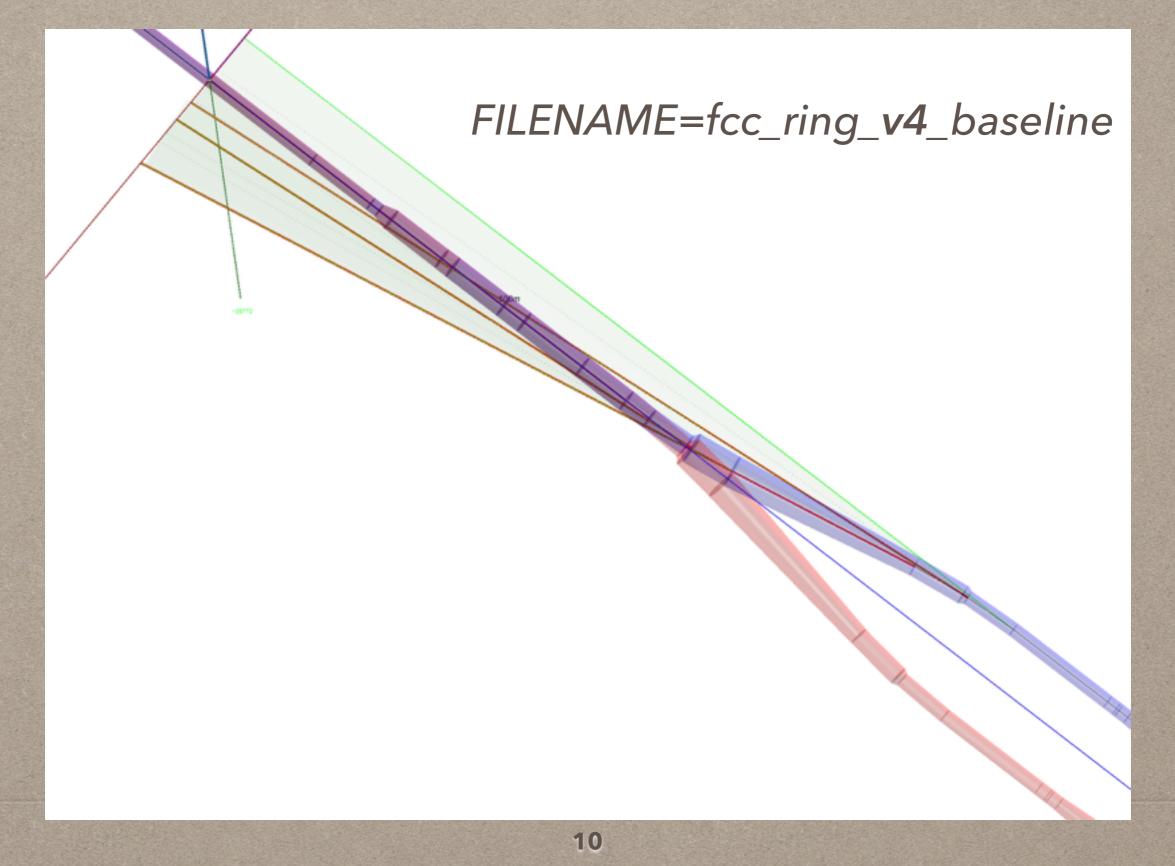
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 - Polarization of photons



"CONE" DIRECTION



"CONE" DIRECTION





LHC, HL-LHC and FCC-hh Parameters

	LHC	Design	HL-LHC	F	cc		
	Nominal	Ultimate	Nominal	16 T	20 T		
Energy [TeV]		7		50			
Luminosity [x10 ³⁴ cm ⁻² .s ⁻¹]	1.0	2.3	5*	5 to	30		
Current [mA]	584	860	1090	509	609		
Proton per bunch [x10 ¹¹]	1.15	1.7	2.2	1.0			
Number of bunches	2	808	2736	10600	8900		
Bunch spacing [ns]		25	25 (then 5 ?)				
Critical energy [eV]		44.1		4300	5375		
Photon flux [ph/m/s]	1 10 ¹⁷	1.5 10 ¹⁷	1.9 10 ¹⁷	1.7 10 ¹⁷	2.6 10 ¹⁷		
SR power [W/m]**	0.22	0.33	0.42	36.3	68.0		
Photon dose [ph/m/year]	1 10 ²⁴	1.5 10 ²⁴	1.9 10 ²⁴	1.7 1024	2.6 10 ²⁴		

* Levelled luminosity

** to be multiplied by 0.8 to get the average power in the arc taking into account the quadrupoles and interconnects lenghts



* During MD periods OLAV-IV, Hsinchu, Taiwan, April 1-4, 2014

