

M. Pavan on behalf of

the CUPID-0* Collaboration

LUCIFER

Outline

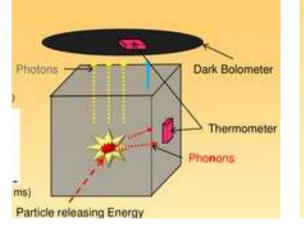
- ♦ transition from LUCIFER → CUPID-0
- status of the ZnSe LUCIFER tower
- CUPID-0 phase II ?

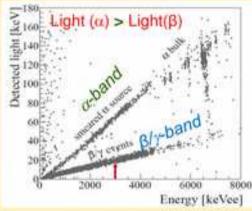
* the name wants to echo the CUORE Upgrade with Particle Identification https://arxiv.org/abs/1504.03612v1

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LUCIFER

ZnSe scintillating bolometers











LUCIFER (ERC Adv. grant) is over

with **30 crystal ingots of Zn⁸²Se** ready to be operated as a scintillating bolometer array in Hall A refrigerator





LUCIFER transition to a new Collaboration and new Funding Institutions

CUPID-0

Italy: Roma1+LNGS+Genova+Milano-Bicocca+LNL France: Orsay+Saclay

- increased number of participants
- main funding today from INFN and CALDER (ERC Starting grant) + LUCIFER heritage
- enlarged scientific program

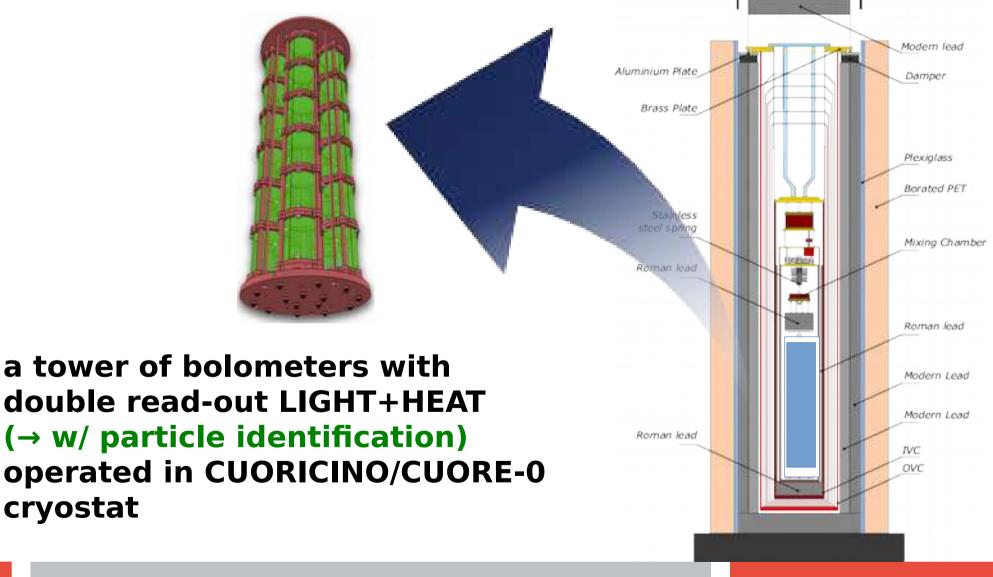
CUPID-0 Collaboration ... today

CUPID-0

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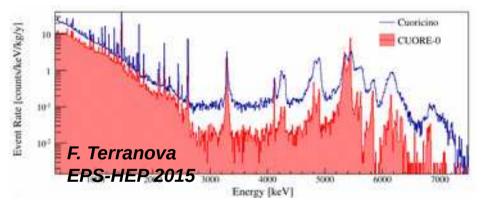
CUPID-0



cryostat

PURPOSE OF THE EXPERIMENT

• prove the scalability to large arrays of the α rejection technique (based on the Heat+Light read-out)



though impressively reduced in CUORE-0 α-induced background is still dominant for E>2.6 MeV

study 0vββ with LUCIFER
 Zn⁸²Se crystals



• & (phase II) add other detectors

ZnSe BOLOMETERS



⁸²Se isotopic enrichment ~ 95.4% (enrichment @ URENCO Netherland)



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Zn⁸²Se crystal mass ~ 0.4 kg

- mass defined by cut and polishing, still ongoing (30 ingots grown in Ucraina ISMA)
- the detector holder allows the use of variable size crystals

25 crystals of Zn⁸²Se 5.5 kg of ⁸²Se

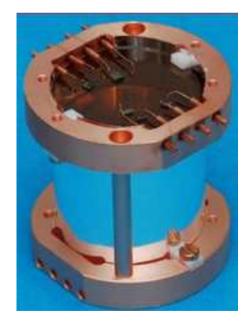
Ge LIGHT DETECTORs

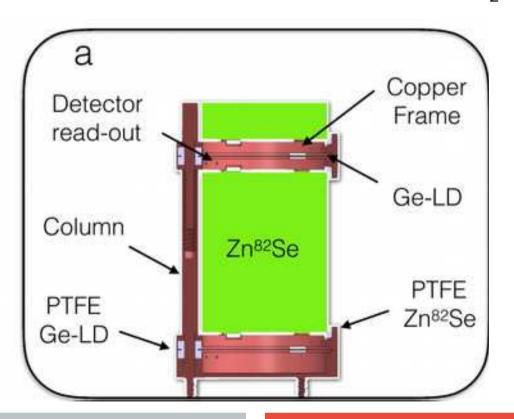


2 Ge light detectors

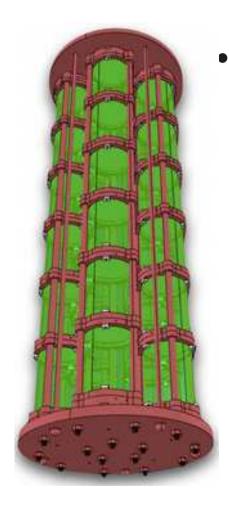
 \emptyset 44.5 mm h=0.15 mm coating with 60 nm SiO₂

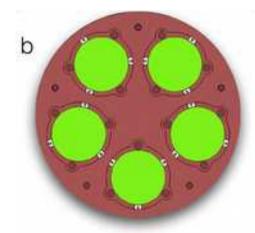
3M VM 2002 reflecting foil





ZnSe ARRAY



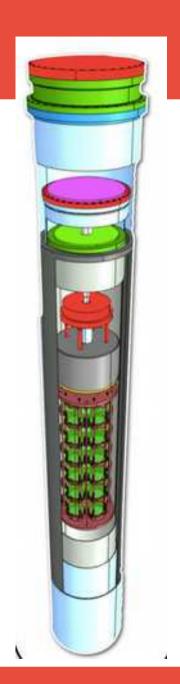


high purity Cu Holder

 $2 < \mu Bq/kg^{232}Th$ 0.6 < $\mu Bq/kg^{238}U$

Roman Lead shield

 $0.3 < \mu Bq/kg^{232}Th$ $0.4 < \mu Bq/kg^{238}U$



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CUPID-0 WORK IN PROGRESS

→ Hall A refurbishing

- + cabling
- + improvement of the anti-Rn system
- + installation of LHe refill

Crystal cut and polishing

Holder Cu parts production and cleaning

(CUORE-0-like surface treatment)

→ time-schedule detector commissioning before end of June operation from 1th July

Zn⁸²Se CRYSTALS : HALL C TEST

3 enriched crystals	crystals masses [g]	440.2 + 442.1 + 439.3
	crystals average dimensions	h = 55 mm, Ø = 43.7 mm
	total detector mass [g]	1321.6
	isotopic abundance	96.3%
	total 82Se emitters	5.11x10 ²⁴

tested in Hall C @ T~ 20 mK

cryogenic problem prevented the cryostat from reaching the usual base T of ~ 5 mK

too high temperature spoiled performances



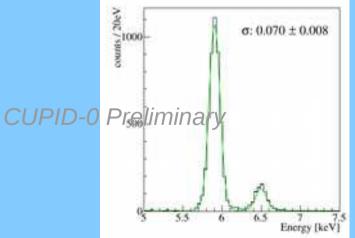
Zn⁸²Se CRYSTALS : HALL C TEST



LIGHT

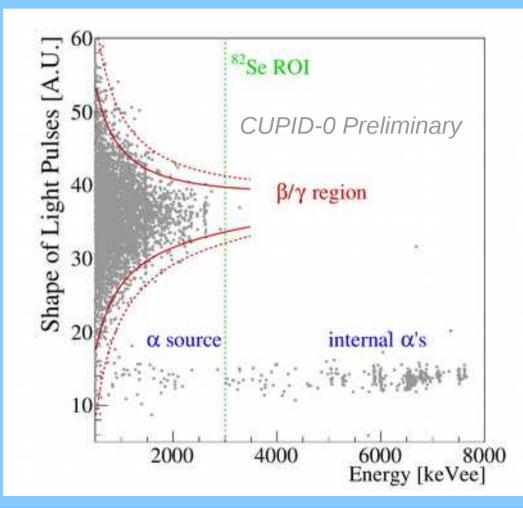


σ_{baseline} ~ 40 eV σ @ 5.5 keV ~ 70 eV



Zn⁸²Se CRYSTALS - HALL C TEST

	O Prelimina CG-1 [µBq/kg]	CG-2 [µBq/kg]	CG-3 [µBq/kg]
²³⁸ U	17 ± 4	20 ± 5	<10
234U	23±5	18 ± 5	<10
²³⁰ Th	18±5	19 ± 5	17 ± 4
²²⁶ Ra	20 ± 5	25 ± 5	21 ± 5
²¹⁰ Po	100 ± 11	250 ± 17	100 ± 12
²³² Th	13±4	13±4	<5
²²⁸ Th	36 ± 7	30 ± 6	26 ± 6



CUPID-0 GOALS

BACKGROUND (CUORE-0 E>2.6 MeV = $2 \ 10^{-2} \ c/keV/kg/y$)

 <u>from detector</u> (Cu+ZnSe+small parts) after rejection of α-background

< 10⁻³ c/keV/kg/y

- from cryostat/environment
 - ~ 10⁻³ c/keV/kg/y

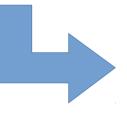
depending on the ²¹⁴Bi content of cryostat and Pb shields

CUPID-0 GOALS



NEMO3 $\tau_{1/2}^{0v} > 0.3 \ 10^{24} \text{ y}$

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~ 25 crystals = 11 kg ZnSe → 5.5 kg <sup>82</sup>Se
FWHM ~ 30 keV
background ~ 10^{-3} c/keV/kg/y
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0.3 counts in 1 y in 30 keV nearly zero bkg approximation $S^{0\nu\beta\beta}_{1\nu} \sim 0.9 \ 10^{25} \text{ y} \ 90\% \text{ C.L.}$

2v\beta\beta ⁸²Se NEMO3 $\tau_{1/2} = (9.2 \pm 0.24_{-0.59}^{+0.67}) 10^{19} y$

use two natural crystal for bkg subtraction

CUPID-0 not only ZnSe

a phase II of Cupid-0 is possible

option (1) add a Molibdate array (scintillating bolometers)

option (2) add a TeO, array (Cerenkov light read-out)



¹⁰⁰Mo option

- ✓ Q_{BB} = 3034 keV (above 2.6 MeV ²⁰⁸TI line)
- \checkmark LiMoO₄/ZnMoO_{4 are} good scintillators w/ better resolution than ZnSe
- Iarge scale scalability (purity, reproducibility and cost) still to be assessed

LiMoO₄ selected for a ~7 kg technology demonstrator (LUCINEU) because of:

Excellent energy resolution (4 – 7 keV FWHM at 2615 keV) **Easier crystallization** procedure

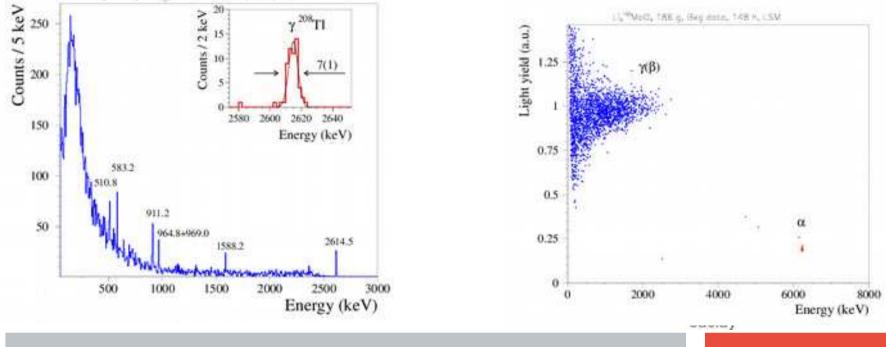
Systematic production of 40 enriched crystals will start in May 2016 (MoU INFN/IN2P3/ITEP) \rightarrow 20 to be installed in LSM and 20 in LNGS



¹⁰⁰Mo option

- \sim Q_{BB} = 3034 keV (above 2.6 MeV ²⁰⁸Tl line)
- \checkmark LiMoO₄/ZnMoO_{4 are} good scintillators w/ better resolution than ZnSe
- large scale scalability (purity, reproducibility and cost) still to be assessed

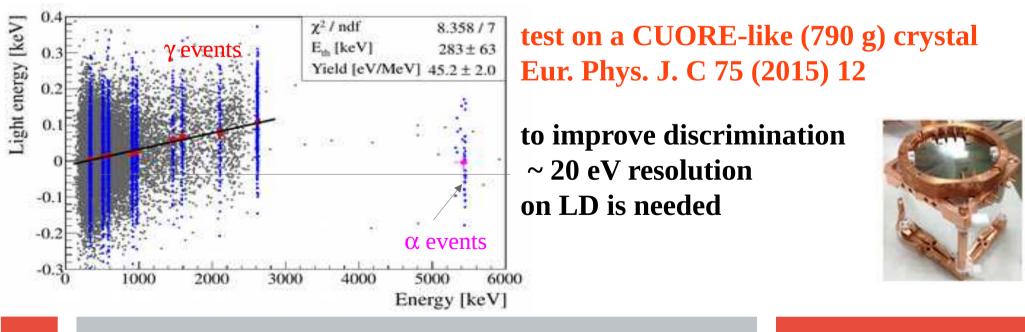
Test is ongoing on a 186 g enriched LMO crystal in LSM (EDELWEISS cryostat) at 20 mK





¹³⁰Te option

- TeO₂ same crystal of CUORE !!
- easy and cheap (already done for MiDBD and Cuoricino)
- low Q_{BB} (2530 keV) environmental γ 's are important (2615 keV)
- no scintillation observed one possibility is **Cerenkov light**



INFN-Cupid activities

LIGHT DETECTOR

requirements for the optimal LD

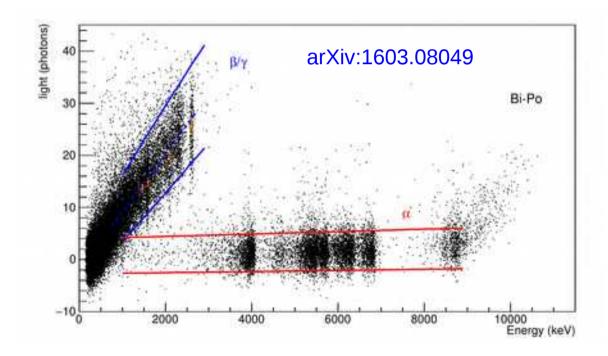
- reproducibility
- easy production and instrumentation as bolometers
- 20 eV energy threshold

Si wafers w/ Luke-Neganov effect &/or MKIDs

phase II demonstrator scintillation/Cerenkov

INFN-Cupid activities

Si wafers w/ Luke-Neganov effect



caution: result obtained on 1 cm³ crystal

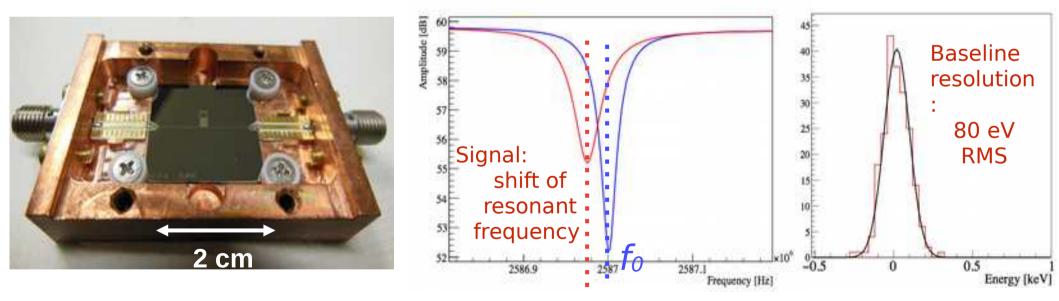


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INFN-Cupid activities

CALDER R&D (ERC StG 2014-2018) www.roma1.infn.it/exp/calder Alternate sensor to NTD and TES: Microwave Kinetic Inductance Detector (MKID). Pros of MKIDs: high scalability and multiplexing, no microphonic noise.

• Phase I - completed: single pixel, high-Q (1.5x10⁵) Aluminum resonator.



- Phase II ongoing: test more sensitive superconductors (TiAl, TiN and Ti+TiN).
 - Goal: 20 eV RMS resolution. *TiAl preliminary: 55 eV RMS.*
- Phase III 2017: test at LNGS with TeO₂ / ZnSe bolometers.

Conclusions: from CUPID-0 to CUPID

Cupid-0 aims at exploring the (at least some) possible options for the CUPID detector

a 1-ton 0-background experiment requires much more than this

- improvement in purity control of materials
- enrichment and crystal production
- improvement in main detector performances

new ideas & collaborators needed and WELCOME !

LUCIFER -> CUPID0



$S_{1y}^{0\nu\beta\beta} \sim 0.9 \ 10^{25} y \ 90\% \ C.L.$

= 2.44 @ 90% C.L.

- T = 1 y n_L = Poisson maximum number of counts compa
- A = 82 $M_{\rm bb} = 5.5 \, \rm kg$

$$\epsilon$$
 = detection efficiency ~ 0.8

$$F_{0\nu}^{ZB} = \ln 2N_{\beta\beta}\epsilon \frac{1}{n_L} = \ln 2 \times \frac{M_{\mu}crr_A}{A} \frac{m_L}{n_L}$$
$$= \ln 2 \times \frac{\epsilon N_A}{A_{\beta\beta}} \frac{M_{\beta\beta}T}{n_L}.$$

$$\begin{aligned} \xi_{0\nu}^{\text{ZB}} &= \ln 2N_{\beta\beta}\epsilon \frac{T}{n_L} = \ln 2 \times \frac{x\eta\epsilon N_A}{A} \frac{MT}{n_L} \\ &= \ln 2 \times \frac{\epsilon N_A}{A} \frac{M_{\beta\beta}T}{m_L}. \end{aligned}$$