

# Sensitivity to the Higgs Sector of the SUSY-Seesaw Models in the LFV $\tau \rightarrow \mu f_0$ decay

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- M. J. Herrero, J. Portoles, A.M. Rodriguez-Sanchez, [[arXiv:0903.5151](#) [hep-ph]]
- E. Arganda, M. J. Herrero, J. Portoles, A. Rodriguez-Sanchez and A. M. Teixeira, [[arXiv:0810.0163](#) [hep-ph]].
- E. Arganda, M. J. Herrero, J. Portoles, A. Rodriguez-Sanchez and A. M. Teixeira, [[arXiv:0812.2692](#) [hep-ph]]

# Why Lepton Flavour Violation (LFV)?

- **LFV occurs in Nature:**  
 $\nu_i - \nu_j$  oscillations DO NOT conserve Lepton Flavour Number
- In SM: no LFV if  $m_\nu = 0$  and extremely suppressed if  $m_\nu \neq 0$
- **LFV is very sensitive to SUSY:** if Seesaw Mechanism for  $m_\nu$  generation with Majorana  $N_R \Rightarrow Y_\nu$  can be  $O(1)$ . Large  $Y_\nu$  induce, via SUSY loops, large LFV rates.
- **Challenging exp. bounds** : present/future sensitivities (?):  
MEGA, SINDRUM, BaBar, Belle / MEG, SuperB fact.,  
**PRISM/PRIME**

$$\text{BR}(\tau \rightarrow \mu\gamma) < 4.5 \times 10^{-8}/10^{-9} \quad \text{BR}(\tau \rightarrow \mu f_0) < 3.4 \times 10^{-8}/10^{-9}$$
$$\text{BR}(\tau \rightarrow 3\mu) < 3.2 \times 10^{-8}/10^{-9} \quad \text{BR}(\tau \rightarrow \mu\eta) < 5.1 \times 10^{-8}/10^{-9}$$

In the  $\tau - \mu$  sector the semileptonic channels are already competitive with the leptonic ones.

- **LFV bounds  $\Rightarrow$  Bounds on SUSY and  $\nu$  parameter space**

# Constrained SUSY-Seesaw models

- MSSM-Seesaw introduces too **many new** parameters due to the SOFT SUSY breaking terms.
- **SOFT- SUSY breaking universality** at the gauge coupling unification scale  $M_x = 2 \times 10^{16}$  GeV
- We work in **CMSSM** ( $M_0, M_{1/2}, A_0, \tan \beta, \text{sign}(\mu)$ ) and **NUHM** (previous and  $M_{H_1}^2 = M_0^2(1 + \delta_1), M_{H_2}^2 = M_0^2(1 + \delta_2)$ )  
The main difference is that in NUHM a light Higgs sector can be obtained even for heavy SUSY.
- The low energy parameters are obtained by solving the RGE in two steps:
  - The full set of equations is run **from  $M_x$  to  $m_M$** . At  $m_M$  the  $\nu_R$  as well as  $\tilde{\nu}_R$  decouple.
  - The RGE without the equations for the  $\nu_R$  and  $\tilde{\nu}_R$  are run **from  $m_M$  to  $M_{EW}$** . The masses and couplings are computed.

# How to generate LFV via SUSY loops?

- Need non vanishing **off diagonal slepton mass entries**.
- The flavor off diagonal mass entries  $M_1^{ij}$  and  $M_{\tilde{\nu}}^{ij}$  ( $i \neq j$ ) at  $M_{EW}$  are generated via **RGE**-running of  $Y_{\nu}$ .

The LL off-diagonal entry of the slepton mass matrix in the Leading-Logarithmic (LLog) approximation:

$$M_{LL}^{ij2} = -\frac{1}{8\pi^2} (3 M_0^2 + A_0^2) (Y_{\nu}^{\dagger} L Y_{\nu})_{ij}; L_{kl} \equiv \log\left(\frac{M_X}{m_{M_k}}\right) \delta_{kl}$$

- Flavor changing sleptons propagators into loops then generate LFV
- $\delta_{LL}^{ij}$  useful phenomenological parameter that encodes the LFV in the i-j sector:

$$\delta_{LL}^{ij} = \frac{M_{LL}^{ij2}}{M_{SUSY}^2}$$

- Prediction of the LFV  $\tau \rightarrow \mu f_0$  branching ratio in Constrained MSSM-Seesaw Models: CMSSM-NUHM.
- Full one loop computation of LFV rates. SPHENO 2.2.2. We do not use LLog nor MI approx.
- Require compatibility with  $\nu$  data.
- Compare our prediction with present LFV bound.
- Explore sensitivity to SUSY, Higgs and heavy  $\nu_R$ .
- Provide an approximate formulae useful for future analysis.
- Comparison with other Higgs's sensitive channels:  
 $\tau \rightarrow 3\mu, \tau \rightarrow \mu\eta$ .

# Seesaw mechanism with 3 $\nu_R$ versus neutrino data

**SeeSaw Mechanism** with 3  $\nu_R$ : ( $m_{\nu_1}, m_{\nu_2}, m_{\nu_3}, m_{N_1}, m_{N_2}, m_{N_3}$ )

$$m_\nu = -m_D^T m_N^{-1} m_D; m_N = m_M; m_D = Y_\nu < H_2 >$$

**Solution:**

$$m_D = i \sqrt{m_N^{diag}} R \sqrt{m_\nu^{diag}} U_{PMNS}^\dagger \quad [\text{Casas, Ibarra ('01)}]$$

$R$  is a  $3 \times 3$  complex matrix and orthogonal

$$R = \begin{pmatrix} c_2 c_3 & -c_1 s_3 - s_1 s_2 c_3 & s_1 s_3 - c_1 s_2 c_3 \\ c_2 s_3 & c_1 c_3 - s_1 s_2 s_3 & -s_1 c_3 - c_1 s_2 s_3 \\ s_2 & s_1 c_2 & c_1 c_2 \end{pmatrix}$$

$c_i = \cos \theta_i$ ,  $s_i = \sin \theta_i$ ,  $\theta_{1,2,3}$  complex

**Parameters:**  $\theta_{ij}, \delta, \alpha, \beta, m_{\nu_i}, m_{N_i}, \theta_i$  (18);  $m_{N_i}, \theta_i$  drive the size of  $Y_\nu$ .

Hierarchical  $\nu$ 's :

$$m_{\nu_1}^2 \ll m_{\nu_2}^2 = \Delta m_{sol}^2 + m_{\nu_1}^2 \ll m_{\nu_3}^2 = \Delta m_{atm}^2 + m_{\nu_1}^2$$

2 scenarios :

- **Degenerate N's**  $\rightarrow m_{N_1} = m_{N_2} = m_{N_3} = m_N$
- **Hierarchical N's**  $\rightarrow m_{N_1} \ll m_{N_2} \ll m_{N_3}$

# Our choice of input parameters

## Constrained MSSM + $3\nu_R$ (Majorana) + $3\tilde{\nu}_R$

- **CMSSM:**

- SUSY parameters:  $M_0, M_{1/2}, A_0$ .
- $\tan \beta < H_2 > / < H_1 >$  (at EW scale)
- $\text{sign}(\mu)$  ( $\mu$  derived from EW breaking)

- **NUHM**

- CMSSM parameters:  $M_0, M_{1/2}, A_0, \tan \beta$  and  $\text{sign}(\mu)$ .
- Non Universal Higgs masses  
 $M_{H_1}^2 = M_0^2(1 + \delta_1), M_{H_2}^2 = M_0^2(1 + \delta_2)$

- **Seesaw parameters**

- $m_{\nu_{1,2,3}}$  and  $U_{MNS}$  (set by data)
- $m_{N_{1,2,3}}$  and  $R(\theta_1, \theta_2, \theta_3)$  (input)

- For numerical estimates:  $(\Delta m^2)_{12} = \Delta m_{\text{sol}}^2 = 8 \times 10^{-5} \text{ eV}^2$

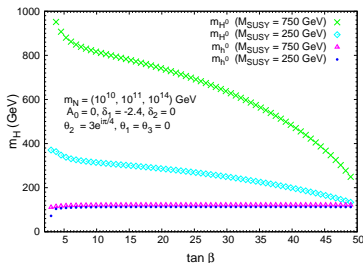
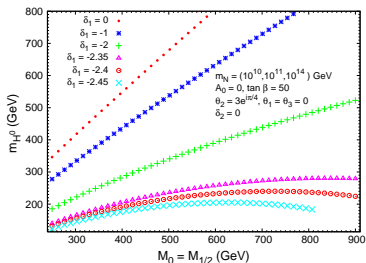
$$(\Delta m^2)_{23} = \Delta m_{\text{atm}}^2 = 2.5 \times 10^{-3} \text{ eV}^2$$

$$\theta_{12} = 30^\circ; \theta_{23} = 45^\circ; \delta = \alpha = \beta = 0; 0 \leq \theta_{13} \leq 10^\circ$$

$$250 \text{ GeV} < M_0, M_{1/2} < 1000 \text{ GeV},$$

$$-500 \text{ GeV} < A_0 < 500 \text{ GeV} \quad 5 < \tan \beta < 50, \quad -2 < \delta_{1,2} < 2$$

# Potential Higgs sensitivity in NUHM versus CMSSM



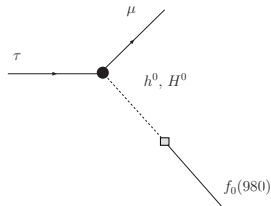
- In CMSSM a heavy soft SUSY spectrum  $\Rightarrow$  heavy  $H_0$ .
- In NUHM, a proper choice of these non-universal parameters,  $\delta_1$  and  $\delta_2$ , can lead us to light Higgs particles even for very large soft SUSY masses of  $\mathcal{O}(1 \text{ TeV})$  if  $\tan\beta$  is large.
- $m_{h^0}$  is independent of  $\tan\beta$  or  $M_{SUSY}$ .
- $m_{H^0}$  becomes lighter with the increase of  $\tan\beta$ .



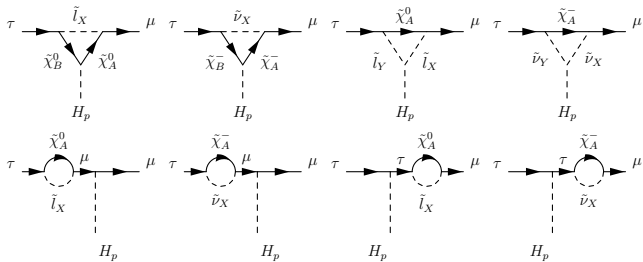
## Results for $BR(\tau \rightarrow \mu f_0)$

- Analytical
  - Full
  - Approximate at large  $\tan\beta$  and large  $M_{SUSY}$
- Numerical
- Comparison with other channels

# Analytical computations



Higgs mediated  $\tau \rightarrow \mu f_0$  decay



One loop SUSY LFV diagrams

# Full Analytical Results

$$\text{BR}(\tau \rightarrow \mu f_0) = \frac{1}{4\pi} \frac{(m_\tau^2 + m_\mu^2 - m_{f_0}^2)^2 - 4m_\tau^2 m_\mu^2}{m_\tau^2 \Gamma_\tau} \frac{1}{2} \sum_{i,f} |T_H|^2,$$

$$\frac{1}{2} \sum_{i,f} |T_H|^2 = \frac{(m_\mu + m_\tau)^2 - m_{f_0}^2}{4 m_\tau} |c_{h^0} + c_{H^0}|^2.$$

$$c_p = \frac{g}{2m_W} \frac{1}{2M_{H_p}^2} \left( J_L^{(p)} + J_R^{(p)} \right) \left( H_R^{(p)} + H_L^{(p)} \right),$$

$H_{L,R}^{(h^0, H^0)} \rightarrow \tau \mu H^{(h^0, H^0)}$  SUSY one-loop LFV vertex functions

$J_{L,R}^{(h^0, H^0)} \rightarrow$  hadron form factors

Hadronisation  $\rightarrow$  substitution of quarks bilinears by scalar currents

# $f_0(980)$ state and hadron form factors

$f_0(980)$  isosinglet state  $\rightarrow$  rotation of the octet  $R_8$  and singlet  $R_0$  components of the  $R(0^+)$  nonet of resonances in the  $N_C \rightarrow \infty$  limit:

$$\begin{pmatrix} R_8 \\ R_0 \end{pmatrix} = \begin{pmatrix} \cos \theta_S & \sin \theta_S \\ -\sin \theta_S & \cos \theta_S \end{pmatrix} \begin{pmatrix} f_0(1500) \\ f_0(980) \end{pmatrix}$$

$\theta_S$  mixing angle uncertain  $\rightarrow \theta_S = 7^\circ$  and  $\theta_S = 30^\circ$

$$J_L^{(H^0)} = \sqrt{2} c_m \left\{ \frac{\sin \alpha}{\sin \beta} \left[ \frac{1}{2\sqrt{3}} \sin \theta_S + \frac{2}{3} \cos \theta_S \right] m_\pi^2 + \frac{-\cos \alpha}{\cos \beta} \left[ \frac{\sqrt{3}}{2} \sin \theta_S m_\pi^2 - \left( \frac{1}{\sqrt{3}} \sin \theta_S - \frac{2}{3} \cos \theta_S \right) 2 m_K^2 \right] \right\}$$

$J_R^{(H^0)} = J_L^{(H^0)*}$  ;  $c_m = F/2$  where  $F \sim F_\pi =$  pion decay constant

In the isospin limit:

$$B_0 m_s = m_K^2 - \frac{1}{2} m_\pi^2 ; H^0 - f_0 \text{coupling} \sim m_K^2 (H^0_{ss} \propto m_s)$$

# Approximate formulae of $\tau \rightarrow \mu f_0$ branching ratio

Approximate formulae valid at **large  $\tan \beta$**  and **heavy  $m_{SUSY}$**

- In this limit  $H_L \gg H_R$  and  $H_L^{H^0} \gg H_L^{h^0}$ . We neglect  $H_R$  and  $H_L^{h^0}$ .
- This limit + MI approx  $\Rightarrow$  chargino/neutralino contribution:

$$H_{L,c}^{(H^0)} = \frac{g^3}{16\pi^2} \frac{m_\tau}{12m_W} \delta_{32} \tan^2 \beta ; \quad H_{L,n}^{(H^0)} = \frac{1}{2}(1 - 3 \tan^2 \theta_W) H_{L,c}^{(H^0)}$$

**Non-decoupling of SUSY in Higgs mediated LFV processes**

$$\begin{aligned} BR(\tau \rightarrow \mu f_0(980))_{\text{approx}} &= \frac{1}{16\pi m_\tau^3} (m_\tau^2 - m_{f_0}^2)^2 \left| \frac{g}{2m_W} \frac{1}{m_{H^0}^2} J_L^{(H^0)} H_{L,c}^{(H^0)} \right|^2 \frac{1}{\Gamma_\tau} \\ &= \left( \begin{array}{l} 7.3 \times 10^{-8} \ (\theta_S = 7^\circ) \\ 4.2 \times 10^{-9} \ (\theta_S = 30^\circ) \end{array} \right) |\delta_{32}|^2 \left( \frac{100}{m_{H^0}(\text{GeV})} \right)^4 \left( \frac{\tan \beta}{60} \right)^6. \end{aligned}$$

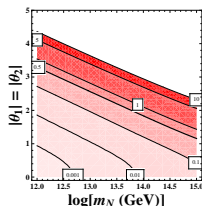
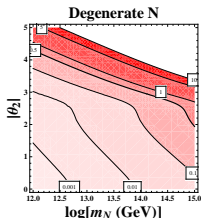
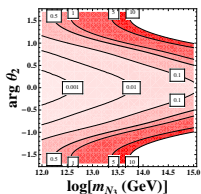
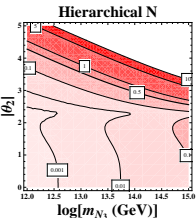
In contrast to  $BR(\tau \rightarrow \mu \gamma) \propto \left( \frac{m_W}{M_{SUSY}} \right)^4$

## Numerical results

# Size of $\delta_{32}$ in Constrained SUSY-Seesaw models

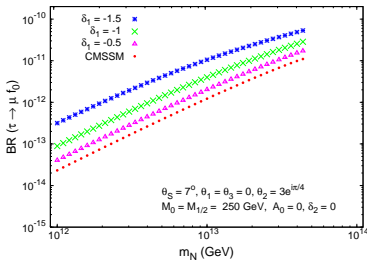
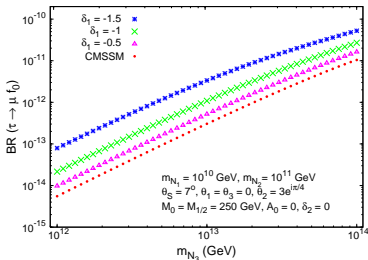
Hierarchical N

Degenerate N



- Large size of  $|\delta_{32}|$  for large  $\theta_{1,2}$  and/or large  $m_N$  in the degenerate and large  $m_{N_3}$  in the hierarchical neutrino case.
- Complex  $\theta_{1,2}$ , with large modulus ( $2 < |\theta_{1,2}| < 3$ ) and argument ( $\pi/4 < \arg\theta_{1,2} < 3\pi/4$ ), and  $m_{N_3}$  between  $10^{14} - 10^{15}$  GeV  $\Rightarrow |\delta_{32}| \sim 1 - 10$ .
- perturbativity in all the gauge and Yukawa couplings  $\Rightarrow |Y_\nu|^2/(4\pi) < 1.5$  and  $|\delta_{32}| < 0.5$

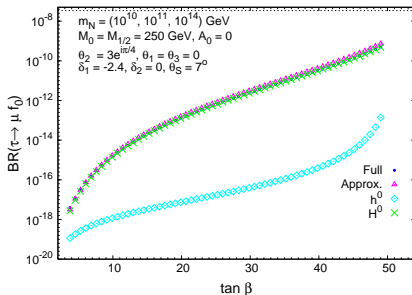
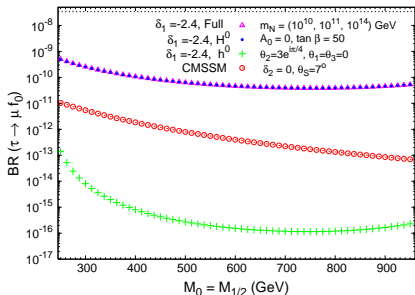
# Results for $BR(\tau \rightarrow \mu f_0)$ : CMSSM/NUHM, hierarchical/degenerate



- $BR(\tau \rightarrow \mu f_0)_{NUHM} > BR(\tau \rightarrow \mu f_0)_{CMSSM}$  due to  $m_{H^0}|_{NUHM} < m_{H^0}|_{CMSSM}$
- $BR(\tau \rightarrow \mu f_0)$  grows with  $m_{N_3}/m_N$
- Independence of  $BR(\tau \rightarrow \mu f_0)$  with  $m_{N_1}$  and  $m_{N_2}$  if  $m_{N_1} < m_{N_2} < m_{N_3}$
- $BR(\tau \rightarrow \mu f_0)_{deg} \geq BR(\tau \rightarrow \mu f_0)_{hierch}$  but hierarchical neutrino scenario more appealing for BAU

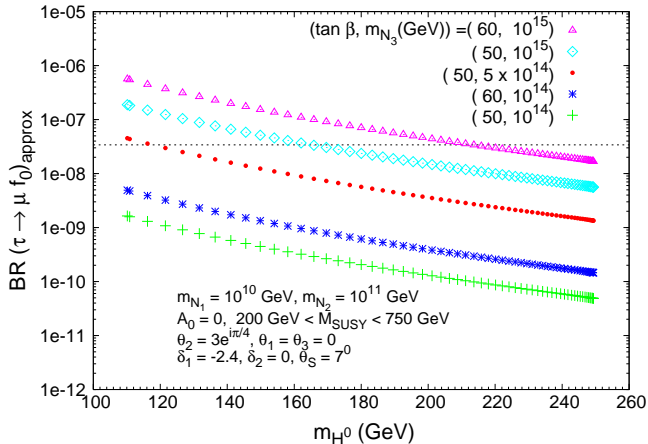


# Getting larger $BR(\tau \rightarrow \mu f_0)$



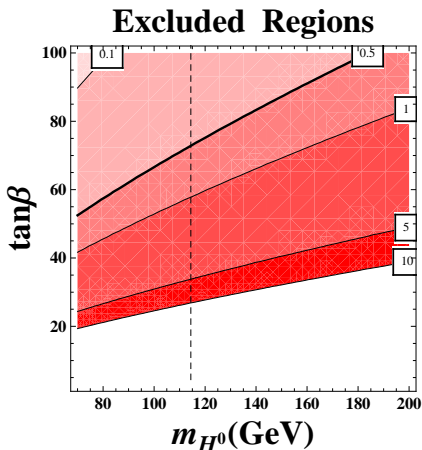
- Large BR for large  $\tan \beta \sim 50$
- The total rates do **not decrease with  $M_{SUSY}$**  in the NUHM  $\Rightarrow$  **SUSY particles do not decouple at large  $M_{SUSY}$**  in this observable
- $BR(\tau \rightarrow \mu f_0)_{H^0} \gg BR(\tau \rightarrow \mu f_0)_{h^0}$ . At large  $\tan \beta$  the  $H^0$  contributions are enhanced by a  $\tan^6 \beta$  factor whereas the  $h^0$  ones are suppressed
- $BR(\tau \rightarrow \mu f_0)_{Approx} \sim BR(\tau \rightarrow \mu f_0)_{Full}$

# Sensitivity to $H^0$ in $\text{BR}(\tau \rightarrow \mu f_0)$ in the NUHM



- There is Higgs sensitivity in this channel. For large  $m_{N_3} \sim 5 \times 10^{14} - 10^{15}$  GeV and large  $\tan \beta \sim 50 - 60$  the rates are at the present experimental reach

# Constraining the model parameters



- Sensitivity to Higgs sector  $\Rightarrow$  constraining mainly  $\tan\beta$  and  $m_{H^0}$
- For fixed  $\delta_{32}$ , comparison with present exp bound  $\Rightarrow$  limits on large  $\tan\beta$  and light  $m_{H^0}$

# Comparison with other LFV $\tau$ decays

$$\begin{aligned} \text{BR}(\tau \rightarrow \mu f_0(980))_{\text{approx}} &= \frac{1}{16\pi m_\tau^3} (m_\tau^2 - m_{f_0}^2)^2 \left| \frac{g}{2m_W} \frac{1}{m_{H^0}^2} J_L^{(H^0)} H_{L,c}^{(H^0)} \right|^2 \frac{1}{\Gamma_\tau} \\ &= \left( \begin{array}{l} 7.3 \times 10^{-8} \quad (\theta_S = 7^\circ) \\ 4.2 \times 10^{-9} \quad (\theta_S = 30^\circ) \end{array} \right) |\delta_{32}|^2 \left( \frac{100}{m_{H^0}(\text{GeV})} \right)^4 \left( \frac{\tan \beta}{60} \right)^6. \end{aligned}$$

$$\begin{aligned} \text{BR}(\tau \rightarrow \mu \eta)_{H_{\text{approx}}} &= \frac{1}{8\pi m_\tau^3} (m_\tau^2 - m_\eta^2)^2 \left| \frac{g}{2m_W} \frac{F}{m_{A^0}^2} B_L^{(A^0)}(\eta) H_{L,c}^{(A^0)} \right|^2 \frac{1}{\Gamma_\tau} \\ &= 1.2 \times 10^{-7} (\theta = -18^\circ) |\delta_{32}|^2 \left( \frac{100}{m_{A^0}(\text{GeV})} \right)^4 \left( \frac{\tan \beta}{60} \right)^6 \end{aligned}$$

$$\begin{aligned} \text{BR}(\tau \rightarrow 3\mu)_{H_{\text{approx}}} &= \frac{G_F^2}{2048\pi^3} \frac{m_\tau^7 m_\mu^2}{\Gamma_\tau} \left( \frac{1}{m_{H^0}^4} + \frac{1}{m_{A^0}^4} + \frac{2}{3m_{H^0}^2 m_{A^0}^2} \right) \left| \frac{g^2 \delta_{32}}{96\pi^2} \right|^2 (\tan \beta)^6 \\ &= 1.2 \times 10^{-7} |\delta_{32}|^2 \left( \frac{100}{m_{A^0}(\text{GeV})} \right)^4 \left( \frac{\tan \beta}{60} \right)^6 \end{aligned}$$

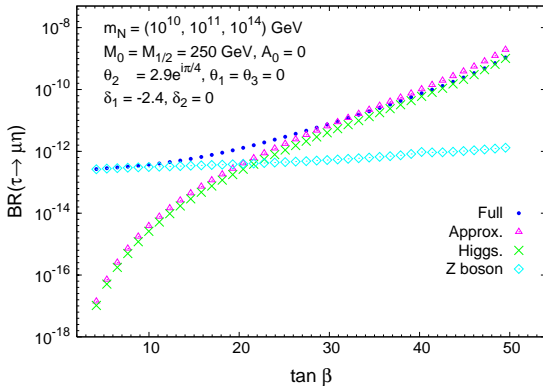
- $\tau \rightarrow \mu f_0$  is dominated by  $H^0 \forall \tan \beta \Rightarrow$  more sensitive to  $H^0$
- $\tau \rightarrow \mu \eta$  is dominated by  $A^0$  (versus  $Z$ ) only for  $\tan \beta > 20$
- $\tau \rightarrow 3\mu$  is dominated by  $\gamma$ .  $H^0$  and  $A^0$  compete with  $\gamma$  only at large  $\tan \beta > 60$

# Conclusions

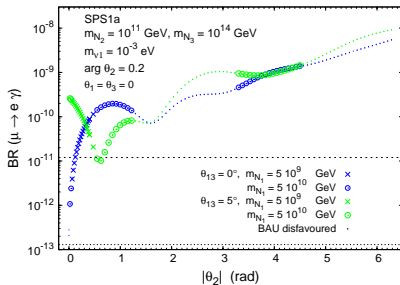
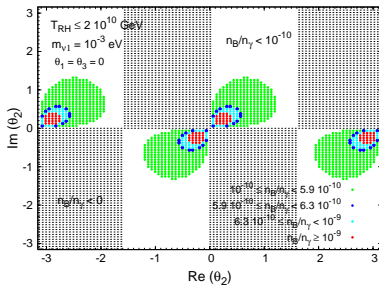
- $BR(\tau \rightarrow \mu f_0)$  grows with  $\tan \beta$  as  $\tan^6 \beta$ , with  $1/m_{H^0}$  as  $(1/m_{H^0})^4$ , and it is approximately constant with  $M_{\text{SUSY}}$ . The dependence with  $m_{N_3}$  and  $\theta_{1,2}$  goes via the  $\delta_{32}$  parameter as  $BR \sim |\delta_{32}|^2 \sim |m_{N_3} \log m_{N_3}|^2$
- Much larger rates in the NUHM-seesaw than in the CMSSM-seesaw, due mainly to the lighter Higgs mass  $m_{H^0}$  in the NUHM-seesaw
- Challenging future sensitivities with SuperB factories

## Additional transparencies

# $BR(\tau \rightarrow \mu\eta)$ versus $\tan\beta$



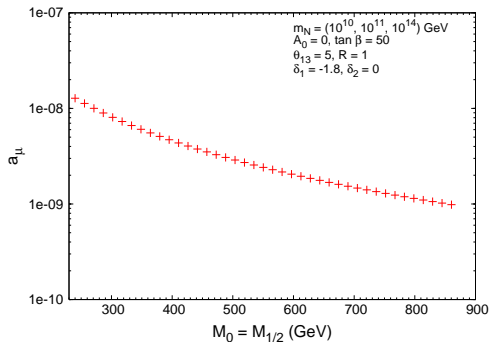
# Constraints from 'viable' BAU



- BAU requires complex  $R \neq 1 \Rightarrow$  complex  $\theta_i \neq 0$ . Most relevantly  $\theta_2$
- $n_B/n_\gamma \in [10^{-10}, 10^{-9}] \Rightarrow (Re(\theta_2), Im(\theta_2)) \in$  area ('ring') (WMAP in darkest ring)
- 'Optimal'  $m_{N_1}$  not far from  $10^{10}$  GeV
- The BAU [fav] windows occur at small ( $\neq 0$ )  $|\theta_2| \lesssim 1.5$
- **smaller**  $|\theta_2| \Rightarrow$  **smaller LFV rates**



# Contributions to $\Delta a_\mu^{SUSY}$



- $\Delta a_\mu^{SUSY} \in [10^{-8}, 10^{-9}]$ : compatible with  $a_\mu^{EXP} - a_\mu^{SM} = 3.32 \times 10^{-9} (3.8\sigma)$

The results of this work have been summarised in the following references:

- E. Arganda, M. J. Herrero, J. Portoles, A. Rodriguez-Sanchez and A. M. Teixeira, “LFV in semileptonic  $\tau$  decays and  $\mu - e$  conversion in nuclei in SUSY-seesaw,” Preprint IFT-UAM/CSIC-08-59, FTUAM/08-19. [arXiv:0810.0163 [hep-ph]].
- E. Arganda, M. J. Herrero, J. Portoles, A. Rodriguez-Sanchez and A. M. Teixeira, “Lepton Flavour Violation in charged leptons within SUSY-seesaw” Preprint IFT-UAM/CSIC-08-85, FTUAM/08-23. [arXiv:0812.2692 [hep-ph]]
- M. J. Herrero, J. Portoles, A.M. Rodriguez-Sanchez “Sensitivity to the Higgs sector of the SUSY-Seesaw Models in the Lepton Flavour Violating  $\tau \rightarrow \mu f_0$  decay,” Preprint IFT-UAM/CSIC-08-84, FTUAM/08-24.

# Analytical results of BR( $\tau \rightarrow \mu f_0$ )

$$\begin{aligned} T_H &= T_{h^0} + T_{H^0} \\ &= \sum_{h^0, H^0} \frac{1}{M_{H_p}^2} \left\{ H_L^{(p)} S_{L,q}^{(p)} [\bar{u}_\mu P_L u_\tau] [\bar{u}_q P_L v_q] + H_R^{(p)} S_{R,q}^{(p)} [\bar{u}_\mu P_R u_\tau] [\bar{u}_q P_R v_q] \right. \\ &\quad \left. + H_L^{(p)} S_{R,q}^{(p)} [\bar{u}_\mu P_L u_\tau] [\bar{u}_q P_R v_q] + H_R^{(p)} S_{L,q}^{(p)} [\bar{u}_\mu P_R u_\tau] [\bar{u}_q P_L v_q] \right\} \end{aligned}$$

The Higgs boson couplings to quarks:

$$i (S_{L,q}^{(p)} P_L + S_{R,q}^{(p)} P_R)$$

$$S_{L,q}^{(p)} = \frac{g}{2m_W} \left( \frac{-\sigma_2^{(p)*}}{\sin \beta} \right) m_q, \quad q = u, S_{R,q}^{(p)} = S_{L,q}^{(p)*}$$

$$S_{L,q}^{(p)} = \frac{g}{2m_W} \left( \frac{\sigma_1^{(p)*}}{\cos \beta} \right) m_q, \quad q = d, s, S_{R,q}^{(p)} = S_{L,q}^{(p)*}.$$

$$\sigma_1^{(p)} = \begin{pmatrix} \sin \alpha \\ -\cos \alpha \\ i \sin \beta \end{pmatrix}, \quad \sigma_2^{(p)} = \begin{pmatrix} \cos \alpha \\ \sin \alpha \\ -i \cos \beta \end{pmatrix}, \quad H_p = h^0, H^0, A^0$$

# Why seesaw mechanism for $m_\nu$ generation

- The seesaw is the simplest mechanism to **explain small  $m_\nu$**
- If Majorana  $\nu$ , the seesaw allows for **large  $Y_\nu$  couplings**
- If Majorana  $\nu$ , **L not preserved, viable BAU via Leptogenesis**

$$-\mathcal{L}_{Y+M} = Y^e \bar{l}_L e_R H_1 + Y^\nu \bar{l}_L \nu_R H_2 + \frac{1}{2} m_M \nu_R^T C \nu_R + h.c.$$

where  $m_D = Y_\nu \langle H_2 \rangle$ ,  $\langle H_2 \rangle = v_2 \sin \beta$

Both Dirac mass  $m_D$  Majorana mass  $m_M$  involved  $\leftrightarrow M_\nu$

$$\mathbf{M}_\nu = \begin{pmatrix} 0 & m_D^T \\ m_D & m_M \end{pmatrix}$$

$m_D \ll m_M \Rightarrow$  seesaw:  $m_\nu = -m_D^T m_M^{-1} m_D$  (light),  $m_N = m_M$  (heavy)

For  $Y_\nu \sim \mathcal{O}(1)$ ,  $m_M \sim 10^{14}$  GeV  $\Rightarrow m_\nu \sim 0.1$  eV (OK)  $m_N \sim 10^{14}$  GeV

Generalization to three generations also OK with data

# MSSM spectrum and experimental constraints

Extended Standard Model spectrum	SUSY particles			
	$SU(3)_C \times SU(2)_L \times U(1)_Y$ interaction eigenstates		Mass eigenstates	
	Notation	Name	Notation	Name
$q = u, d, s, c, b, t$ $l = e, \mu, \tau$ $\nu = \nu_e, \nu_\mu, \nu_\tau$	$\tilde{q}_L, \tilde{q}_R$ $\tilde{l}_L, \tilde{l}_R$ $\tilde{\nu}$	squarks sleptons sneutrino	$\tilde{q}_1, \tilde{q}_2$ $\tilde{l}_1, \tilde{l}_2$ $\tilde{\nu}$	squarks sleptons sneutrino
g	$\tilde{g}$	gluino	$\tilde{g}$	gluino
$W^\pm$ $H_1^+ \supset H^+$ $H_2^- \supset H^-$	$\tilde{W}^\pm$ $\tilde{H}_1^+$ $\tilde{H}_2^-$	wino higgsino higgsino	$\tilde{\chi}_i^\pm (i=1,2)$	charginos
$\gamma$ $Z$ $H_1^0 \supset h^0, H^0, A^0$ $H_2^0 \supset h^0, H^0, A^0$ $W^3$ $B$	$\tilde{\gamma}$ $\tilde{Z}$ $\tilde{H}_1^0$ $\tilde{H}_2^0$ $\tilde{W}^3$ $\tilde{B}$	photino zino higgsino higgsino wino bino	$\tilde{\chi}_j^0 (j=1, \dots, 4)$	neutralinos

- Mass bounds (95% C.L.) from direct searches (PDG 2008) in GeV

$$m_{h^0} > 92.8, m_{A^0} > 93.4, m_{\tilde{e}} > 73, m_{\tilde{\mu}} > 94,$$

$$m_{\tilde{\tau}} > 81.9, m_{\tilde{\nu}} > 94, m_{\tilde{\chi}_1^0} > 46, m_{\tilde{\chi}_1^\pm} > 94$$

- Motivation/Introduction
  - ★ Constrained SUSY Models
  - ★ Neutrino masses and Seesaw Mechanism
  - ★ Theoretical framework for LFV in semileptonic tau decays
  - ★ Hadronization of quark bilinears
  - ★ Experimental data (bound)
- Results:
  - ★ Analytical results
  - ★ Numerical results
  - ★ Full versus Approximate results
- Conclusions

# Why SUSY?

- **Experimental evidence of Physics Beyond the SM**

- ★ Neutrino oscillations  $\Rightarrow$  Neutrinos are massive
- ★ The SM can not explain the Baryon Asimmetry of the Universe
- ★ The SM does not incorporate gravitation
- ★ No understanding of dark matter and dark energy

- **SUSY solves the hierarchy problem of the SM and SM-Seesaw**

- ★ SUSY introduces a new symmetry between bosons and fermions

fermionic dof = bosonic dof

SUSY  $\Rightarrow$  Cancellation of quadratic divergences  
of the Higgs mass

# No experimental evidence of SUSY yet

- **Direct searches of SUSY particles at colliders**

- best {
- ★ But, not seen at LEP, Tevatron..  $\Rightarrow$  Experimental constraints
  - ★ Some sparticles may not be produced at LHC if too heavy

- **Indirect searches of SUSY via radiative corrections**

- ★ If SUSY particles not seen: Complementary to direct searches
- ★ Similar to past LEP hints on top quark via  $\Delta \rho$  etc

## Hints on SUSY: virtual SUSY particles propagate into the loops



### Look for observables enhanced in SUSY respect to SM prediction

- ★ Involving **Higgs sector**

- ★ **FC and LFV** suppressed in SM



Non-decoupling of heavy SUSY



Window to new physics



# Seesaw mechanism with $3 \nu_R$

For 3 generations  $\Rightarrow$  6 physical neutrinos: 3  $\nu$  light, 3  $N$  heavy

$$U^{\nu T} M^{\nu} U^{\nu} = \hat{M}^{\nu} = \text{diag}(m_{\nu_1}, m_{\nu_2}, m_{\nu_3}, m_{N_1}, m_{N_2}, m_{N_3}).$$

$$m_D \ll m_M, m_D = Y_{\nu} \langle H_2 \rangle \Rightarrow$$

$$m_{\nu}^{\text{diag}} = U_{\text{PMNS}}^T m_{\nu} U_{\text{PMNS}} = \text{diag}(m_{\nu_1}, m_{\nu_2}, m_{\nu_3}), \quad (1)$$

$$m_N^{\text{diag}} = m_N = \text{diag}(m_{N_1}, m_{N_2}, m_{N_3}),$$

All,  $Y_{\nu}$ ,  $m_D$ ,  $m_M$ ,  $U_{\text{PMNS}}$ , are  $3 \times 3$  matrices;  $c_{ij} \equiv \cos(\theta_{ij})$ ,  $s_{ij} \equiv \sin(\theta_{ij})$

$$U_{\text{PMNS}} = \begin{pmatrix} c_{12} c_{13} & s_{12} c_{13} & s_{13} e^{-i\delta} \\ -s_{12} c_{23} - c_{12} s_{23} s_{13} e^{i\delta} & c_{12} c_{23} - s_{12} s_{23} s_{13} e^{i\delta} & s_{23} c_{13} \\ s_{12} s_{23} - c_{12} c_{23} s_{13} e^{i\delta} & -c_{12} s_{23} - s_{12} c_{23} s_{13} e^{i\delta} & c_{23} c_{13} \end{pmatrix} \times \text{diag}(1, e^{i\alpha}, e^{i\beta}). \quad (2)$$

**Pontecorvo-Maki-Nakagawa-Sakata matrix:**  $\theta_{12}, \theta_{13}, \theta_{23}, \delta, \alpha, \beta$