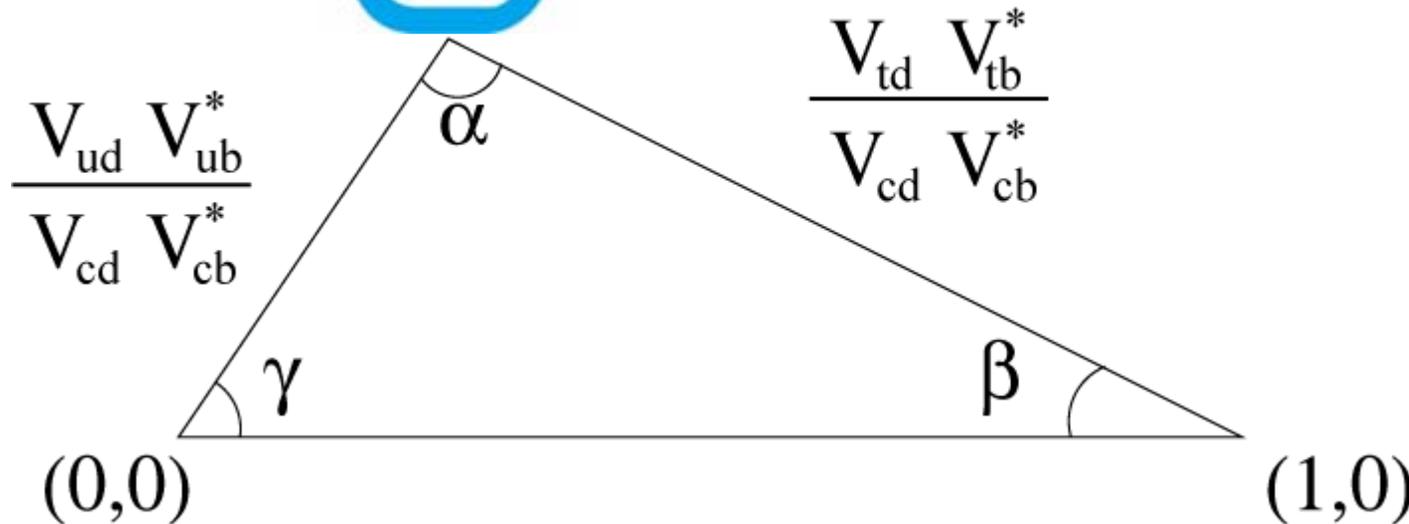


Measuring α at SuperB



Adrian Bevan

a.j.bevan@qmul.ac.uk



Overview

- The experimental Landscape Now
 - SU(2) $\pi\pi$, $\rho\rho$, $\rho\pi$
 - SU(3) $\rho\rho$, + $a_1\pi$ (soon)
- Projections to 75ab^{-1} and 250ab^{-1} ($1/4 \text{zb}^{-1}$)
- Conclusions
- Relevant Past talks by AB:
 - Paris '07, Elba '08

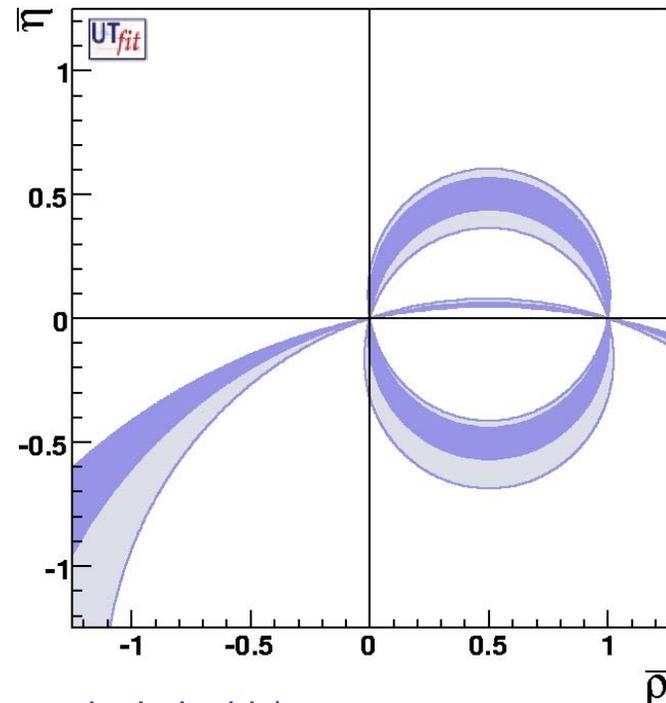
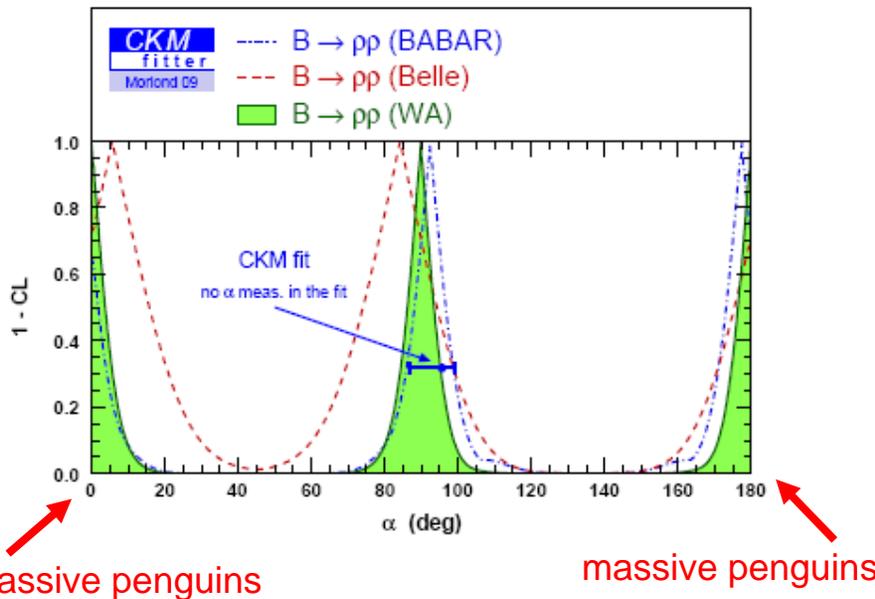




Experimental Landscape

Averaging Group	direct	indirect
UTfit	$(92 \pm 7)^\circ$ U $(166 \pm 2)^\circ$	92.4 ± 4.1
CKM fitter	$89.0^{+4.4}_{-4.2}$	$95.6^{+3.3}_{-8.8}$
UTfit $\rho\rho$ only	$(90 \pm 8)^\circ$ U $[167, 197]^\circ$...

- Better than 4° now
- $\rho\rho$ dominates precision
- **BECAUSE** $\rho\rho$ triangle is flat!

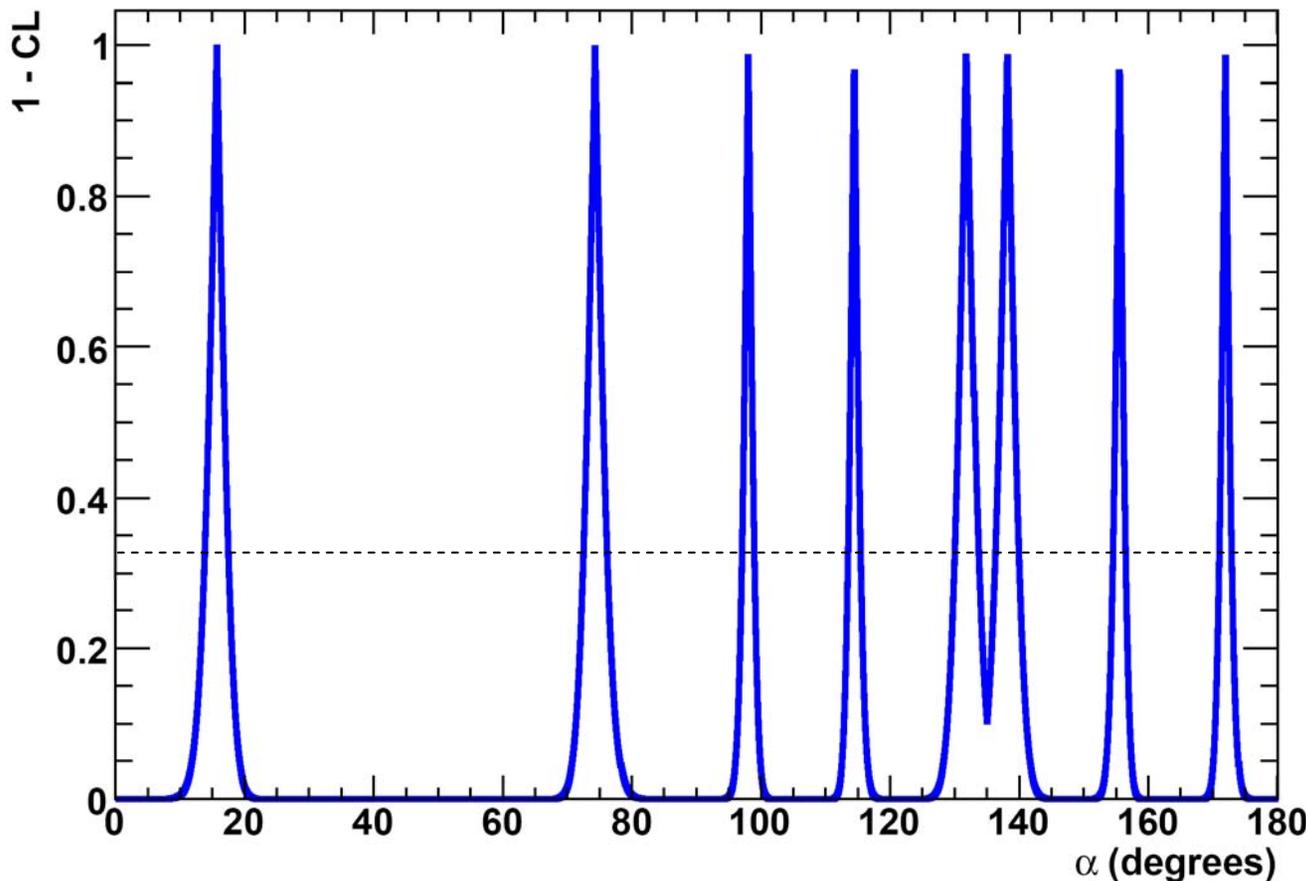


- Take care with SU(2) analysis – allows for unphysical $|P/T|$ at π and 0.
- Experiment should correct for EWP directly (currently we don't do this).
- At high statistics different statistical interpretations of data should give the same results, so don't prefer Bayesian over Frequentist or visa versa.





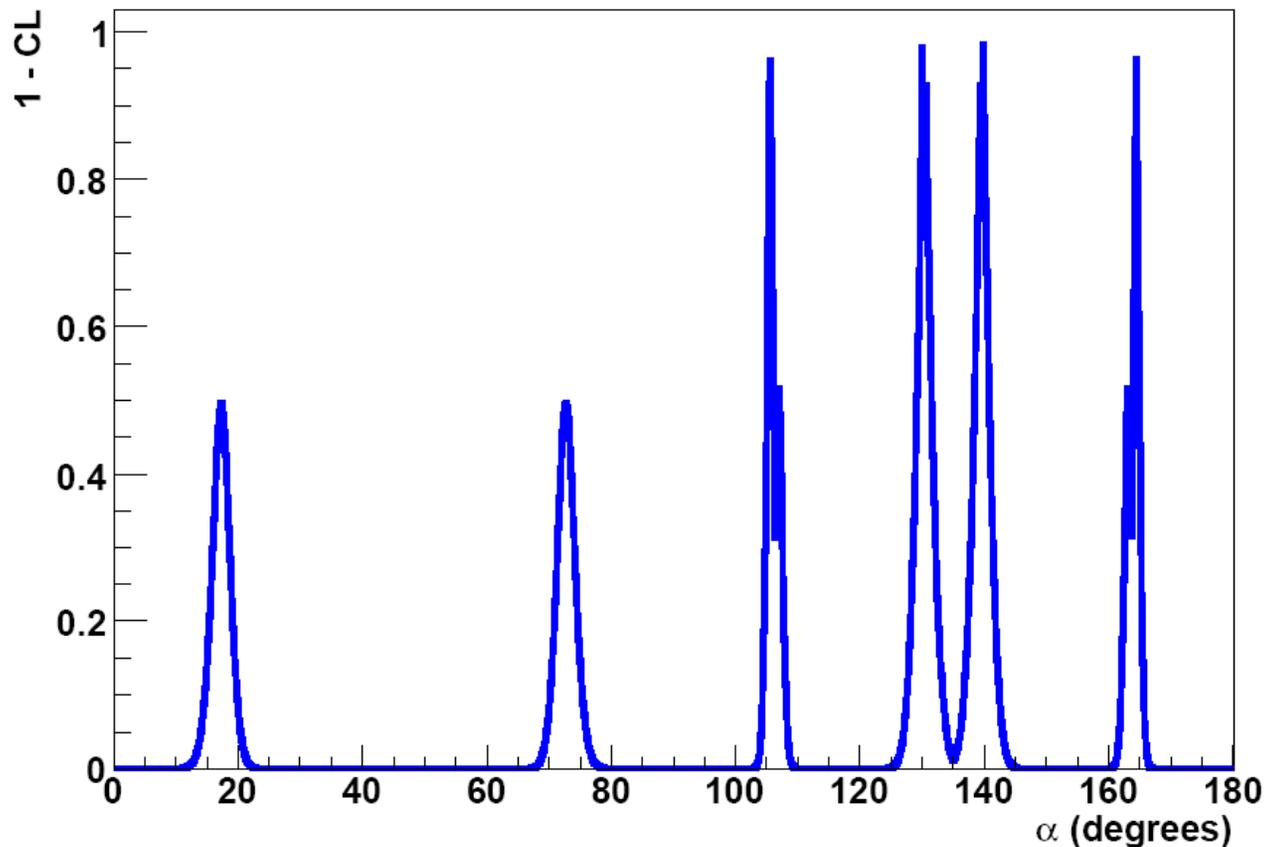
75ab⁻¹



Predictions using the GL method.

- All ambiguities are distinct solutions at 75ab⁻¹ data samples.
- Precision of solutions vary from O(0.9) – O(1.9) degrees.
- Can measure S^{00} to ~ 0.23 with 50ab⁻¹ (not included in this projection)
- Including S^{00} will help resolve the ambiguities.

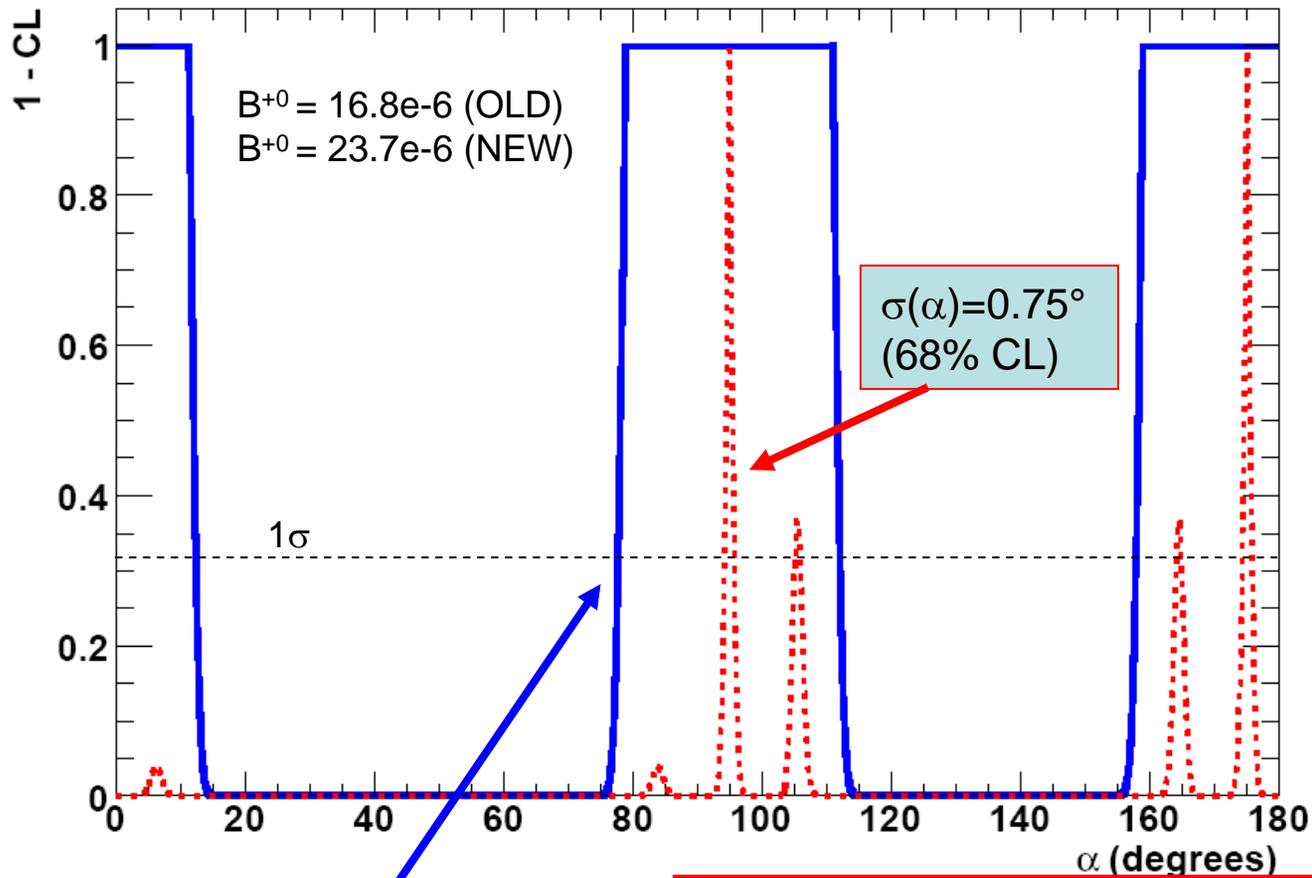


**250ab⁻¹**

- Prediction using $C^{00}=0$ to illustrate the effect of this parameter on ambiguities.
- Precision of central solutions is better than 1° .
- S^{00} (not included in this projection)



75ab⁻¹



Ignoring S^{00} and C^{00}
 Don't resolve ambiguities.

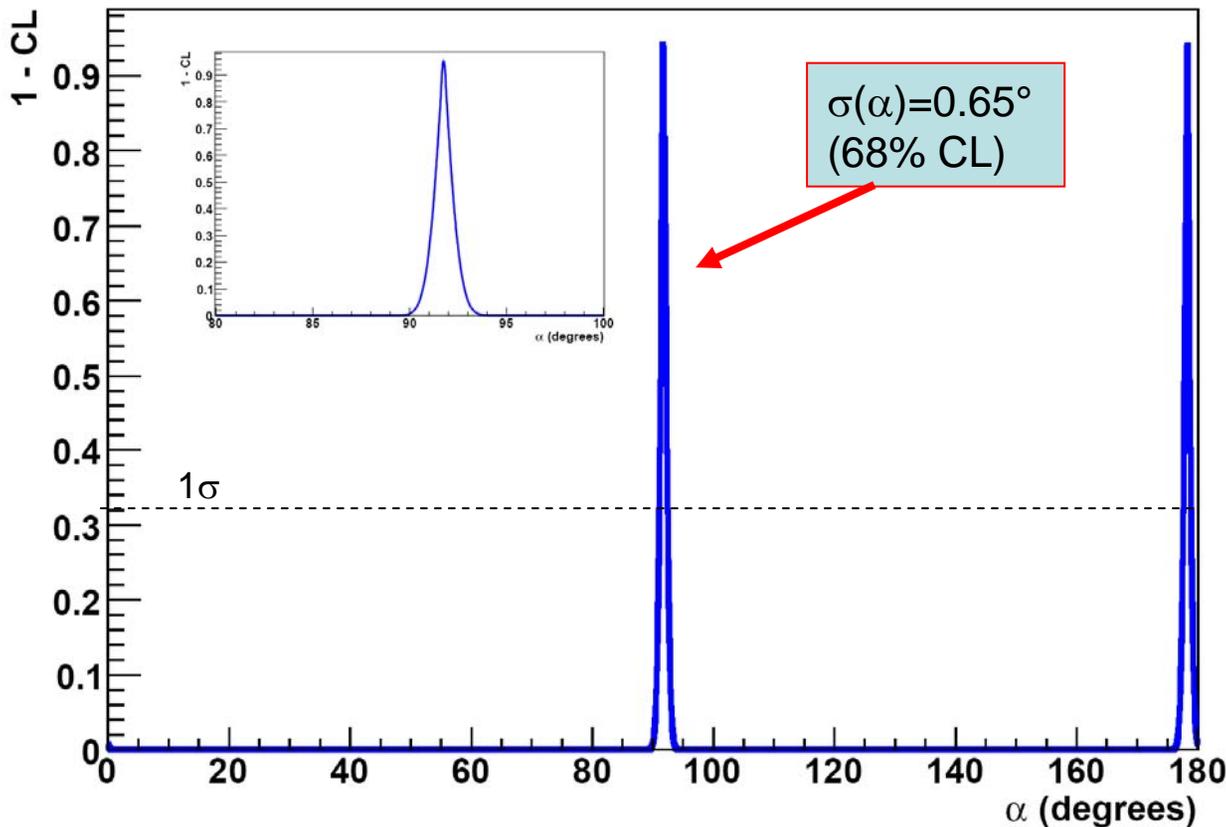
Using S^{00} and C^{00} we obtain the red curve with the OLD $\rho^+\rho^0$ branching fraction.

With the new $\rho^+\rho^0$ branching fraction we obtain only the central solution from each grouping





250ab⁻¹



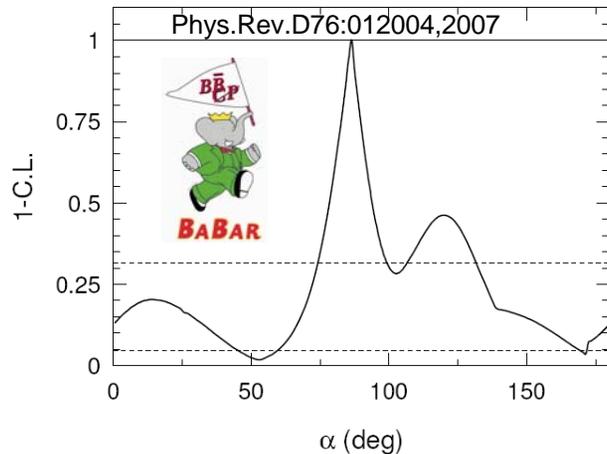
- Using the new $\rho^+\rho^0$ branching fraction WA removes the ambiguities seen in the previous slide.
- Only slight improvement in error on α when going to 250ab⁻¹.
- This analysis will be systematics limited at this luminosity.

- SuperB will need to work on the analysis technique to reduce $\sigma(\alpha)$ beyond 0.65°:
 - SCF, DCSD interference, Fit Bias are the main culprits.

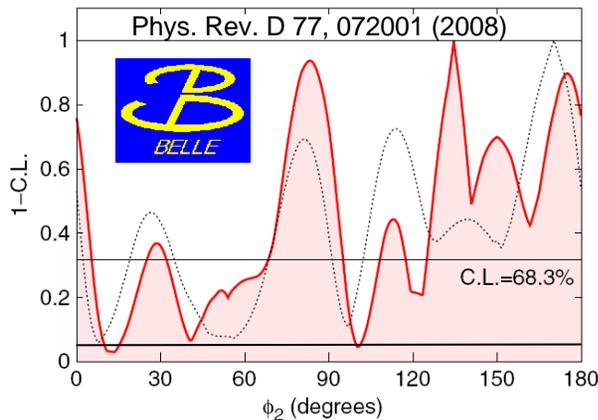




- Many parameters fitted to a complex Dalitz structure: Hard to make definite predictions now.
 - The dominant role of $\rho\pi$ is to reduce ambiguity.



- Currently:
 - Useful 1σ constraint.
 - Almost no constraint at 2σ .
 - No constraint at 3σ .
- Will select a solution (based on current data, most probably the SM solution).
- Will improve overall precision slightly.



- *Conservatively* Ignore the effects of $\rho\pi$ on the precision of α and just assume we are able to select with certainty the SM solution.
- **This mode plays a very important role!**



Limiting uncertainties at Super-B

SU(2)

▪ $\pi\pi$

$$\sigma_{\pi\pi}(\alpha) = 0.9 \sim 1.9 \text{ (expt)} \pm \text{(th.)}$$

- $\Delta I=5/2$ amplitudes to test for.
- SU(2) breaking in π^0 - η - η' mixing.
- Should be dominated by experimental uncertainty.

See the following Refs. (not an exhaustive list) for more details on sources of theoretical uncertainty: Gronau and Zupan PRD71 074017 (2005); Gardner PRD72 034015 (2005); hep-ph/9906269; Botella et al. PRD73 071501 (2006) ...

▪ $\rho\rho$

$$\sigma_{\rho\rho}(\alpha) = 0.75 \text{ (expt)} \pm 0.4 \text{ (EWP)}$$

- $I=1$ amplitudes to test for.
- $\Delta I=5/2$ amplitudes to test for.
- ρ - ω mixing to include in $\rho^+\rho^0$ and $\rho^0\rho^0$ measurements.
- SU(3) method has a precision of $O(2)^\circ$: Limited by theory uncertainties.

75ab⁻¹

▪ $\rho\pi$

$$\sigma_{\rho\pi}(\alpha) = {}^{+45}_{-13} \text{ (Current BaBar Error)}$$

- Too complicated to make sensible projections for at the moment.

▪ **Experimental** precision of $\sim 1^\circ$ is achievable.

Essential to have control of all theoretical uncertainties so that the worst case scenario is comparable to or better than experiment.





- $\rho\pi$ TDDP not included \rightarrow this will select one of the solutions (assume $\alpha \sim \pi/2$ is selected).
- The precision of $\pi\pi$ and $\rho\rho$ IAs will be comparable at the % level.
 - Differences in central value will help resolve ambiguities.
 - This is a function of C^{00} , S^{00} simultaneously for both $\pi\pi$ and $\rho\rho$.
 - Neglected the fact that we can measure S^{00} to a precision of ~ 0.23 with $50ab^{-1}$ using photon conversions (see Ishino et al. hep-ex/0703039). This will help a little in resolving ambiguities.
- There is room for improvement:
 - The measurement of S^{+-} and C^{+-} for $\rho\rho$ will be systematically limited using the current methods.
 - These limitations can probably be overcome with some thought. Given the data it is reasonable to assume a sub 0.5° measurement is possible.
 - Fitting lepton tagged events only should overcome the systematic error limitation and result in a similar precision on α to that quoted here.
 - These extrapolations assume the published tagging performance $Q \sim 30\%$. The final performance has $Q \sim 33\%$.



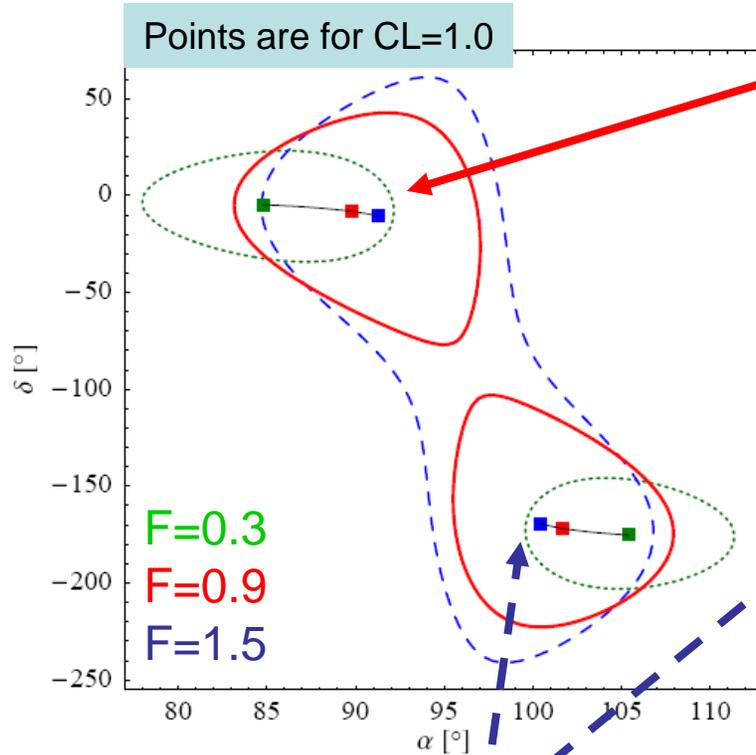
SU(3) based methods: $\rho\rho$

SU(3)

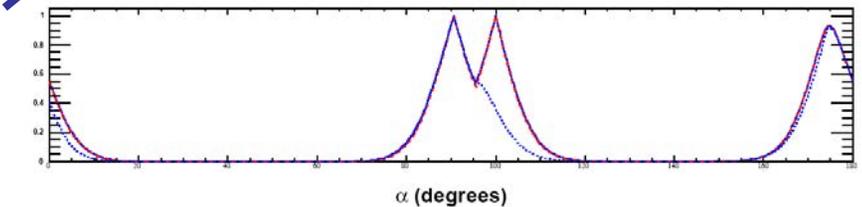
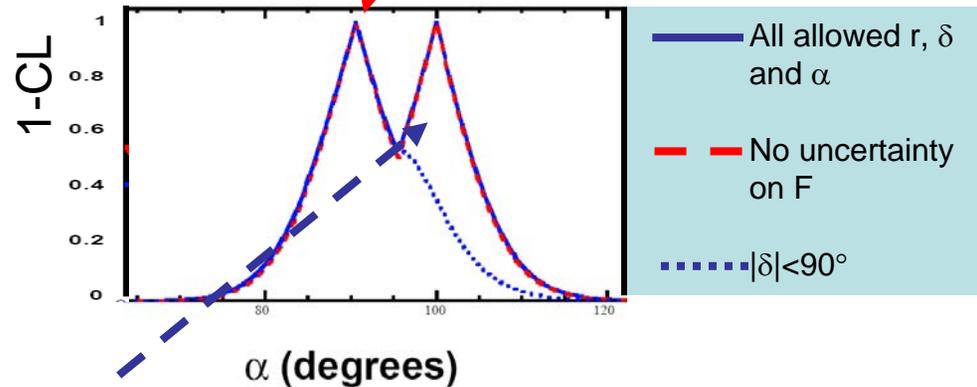
- Relate $K^*0\rho^+$ to the penguin contribution in $\rho^+\rho^-$ (BGRS):

- Gives a precise determination of α : independent of B^{+0} .

0.35ab⁻¹



$\alpha \sim 90^\circ$ is preferred for $|\delta| < 90^\circ$

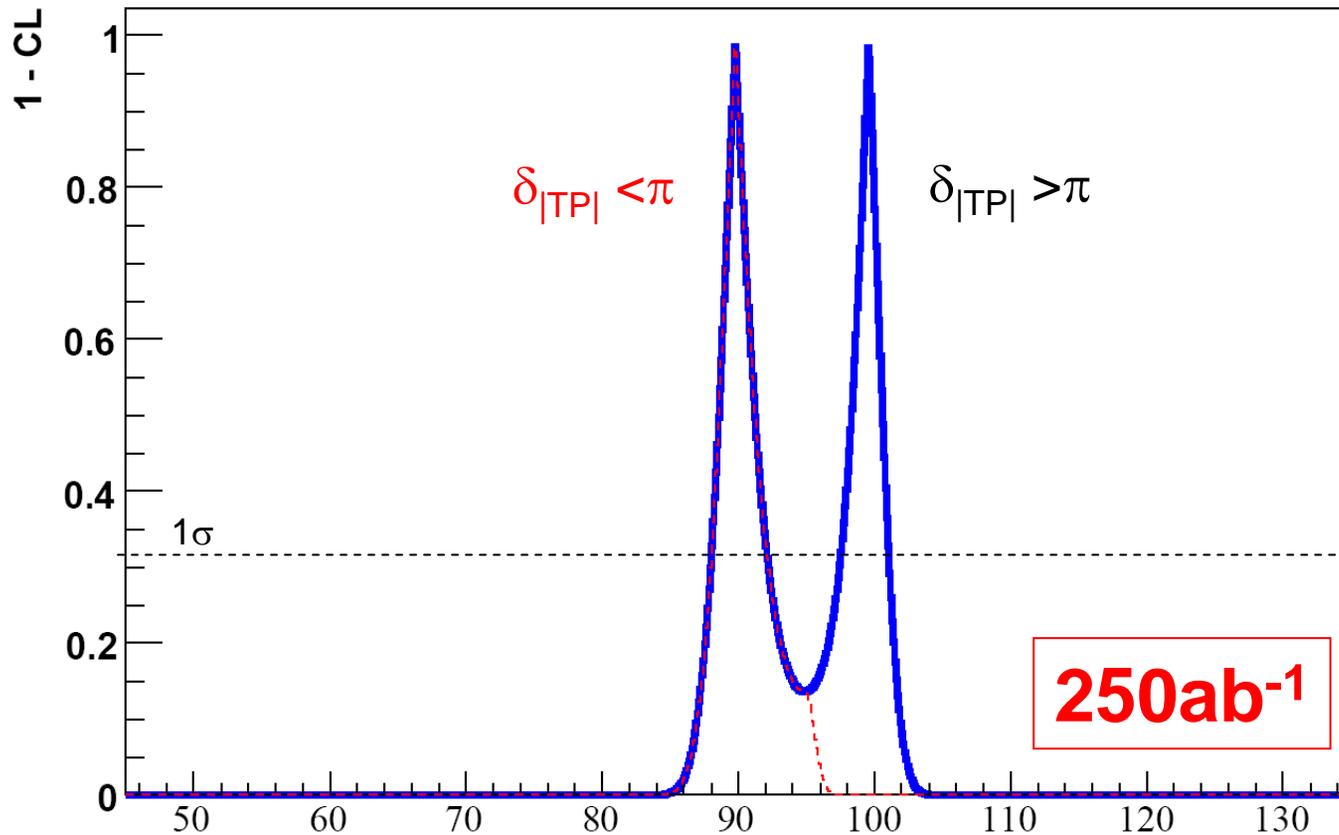


$\alpha \sim 100^\circ$ is preferred for $|\delta| > 90^\circ$

$\sigma(\alpha) \sim 7^\circ (\text{expt.}) \pm 1.5^\circ (\text{th.})$

Beneke et al., Phys.Lett. B638 (2006) 68-73





- Precise determination of α with a few degrees uncertainty.
- Ultimate precision depends on systematic limitations in S^{+-} and C^{+-} .



Other SU(3) based methods

- $a_1\pi$:
 - $\sigma(\alpha_{\text{eff}}) \sim 7^\circ$
 - Need $K_1\pi$ and $a_1\pi$ branching fractions to constrain penguins using SU(3).
 - It doesn't matter what value $\delta\alpha$ is in this mode – it WILL be measurable at SuperB, so we will be able to use this in our α determination.
 - Will have all of the ingredients from current B-factories to extrapolate this method to higher luminosities during the TDR period...
- $a_1\rho$, $\omega\rho$, $b_1\rho$, ... Looked for ... nothing seen so far!
- What about a_1a_1 ?
 - If found what can we do with it? This modes is also related to polarisation puzzle!





Conclusions

- Note: These estimates are extrapolated from the current B-factories.
- % level measurement is achievable, as sub-% level measurement is quite possible from the experimental side.
 - ... can theory do the same?
- There is clearly room for improvement.
 - Also need to look at how SuperB Detector Geometry affects precision. Can we make further optimisations to help this measurement.
- What are the prospects of improving the theoretical control of SU(2) and SU(3) uncertainties? [See Jure]
- Are there higher order effects that would hamper an α version of the ΔS measurement?
- α @ SuperB will be a game of $\pi\pi$, $\rho\rho$, $\rho\pi$, $a_1\pi$, + ...

