



Fermi

Gamma-ray Space Telescope

A NEW MEASUREMENT
OF THE COSMIC-RAY
ELECTRON AND
POSITRON SPECTRUM
WITH THE LARGE
AREA TELESCOPE

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on behalf of the *Fermi*-LAT
collaboration

Pisa, October 17, 2016

MORE THAN AN UPDATE

Previous Fermi measurement of the spectrum from 7 GeV to 1 TeV published in 2010 [*Ackermann et al., Phys. Rev. D 82, 092004*]

The analysis is new in many respects

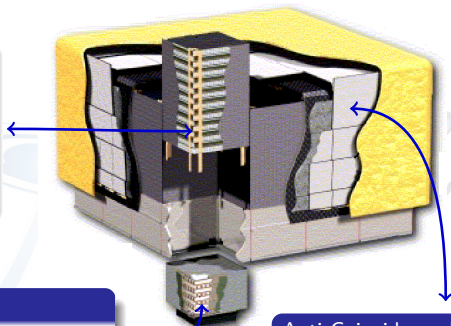
- ▶ Almost seven times the amount of data available
- ▶ Completely revised event reconstruction provided by the *Fermi* collaboration (Pass 8)
- ▶ Upper energy range extended to 2 TeV
- ▶ Improved selection algorithms
- ▶ Better understanding of the environment in which the LAT operates
- ▶ Better modelling of the systematic uncertainties
- ▶ New detailed study of the geomagnetic field effects in the GeV energy range

FERMI LAT

- ▶ Launched by NASA on June 2008 at 565 km altitude, 25.6° inclination
- ▶ Large field of view (20% of the sky), full sky observed every ~ 3 h
- ▶ Designed to detect EM showers, naturally including e^+ and e^-

Tracker/Converter (TKR)

- ▶ 18 x-y planes of silicon strip detectors
- ▶ ~ 10 k sensors, ~ 80 m² of silicon active area, 1M readout channels



Calorimeter (CAL)

- ▶ 1536 CsI(Tl) crystal (96 per tower); 8.6 radiation lengths on-axis
- ▶ Hodoscopic, 3D shower profile reconstruction for leakage correction

Anti-Coincidence Detector (ACD)

- ▶ Segmented (89 tiles) as to minimize self-veto at high energy
- ▶ 0.9997 average MIP detection efficiency

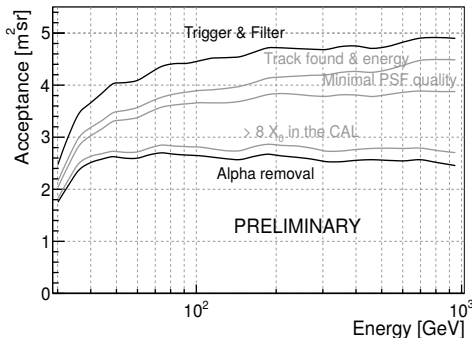
Two slightly different analyses, for two different datasets:

- ▶ High Energy (HE): 42 GeV - 2 TeV using events from the standard on-board filter
Accept all events with more than 20 GeV released in the CAL (fully efficient for electrons above ~ 40 GeV)
- ▶ Low Energy (LE): 7 GeV - 70 GeV using events from the diagnostic (DGN) filter
Unbiased sample of all triggers, pre-scaled by a factor 250

Overall live time is 4.68 years - outside SAA and in standard operation modes

Field of view reduced to $\leq 60^\circ$ from the instrument zenith (avoid Earth limb photons contamination).

CRE SELECTION



- ▶ Event-quality cuts (good direction and energy reconstruction)
- ▶ Removal of particles with $Z > 1$, tagged by ionization in ACD and TKR
- ▶ Boosted Decision Trees (BDTs) selection, for proton rejection
- ▶ LE: additional orbital selection for geomagnetic field effects

PROTON REMOVAL SELECTION

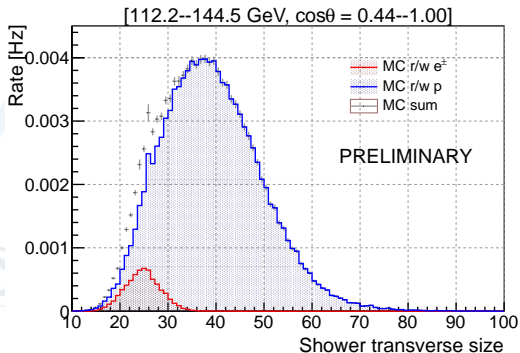
- ▶ Exploits differences between leptonic and hadronic events in the detector:

- ▶ Shower development in the CAL
- ▶ Number of δ -rays produced in the TKR

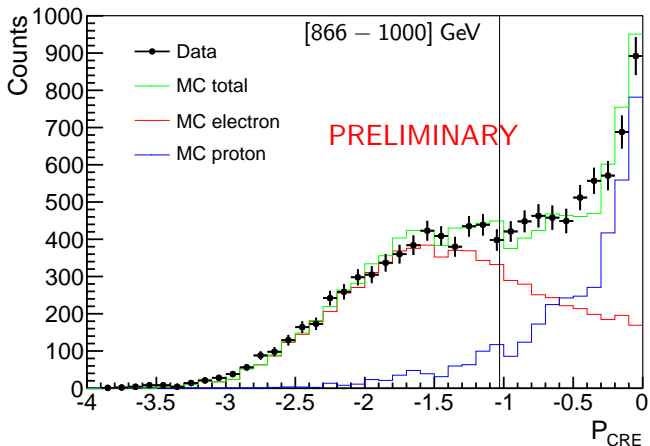
- ▶ A few variables have been individually rescaled, to improve their data/MC agreement

- ▶ The BDT combines all the information into a single estimator of 'leptonicity' (P_{CRE}).
- ▶ 8 BDTs used for HE analysis, optimized for different energy ranges (plus 1 for LE)

Example:

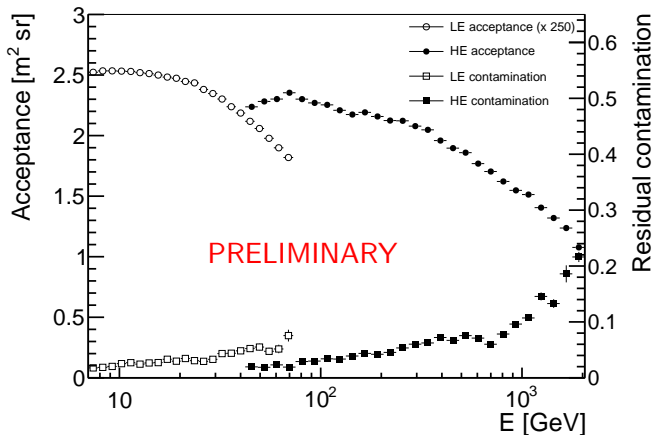


BDT OUTPUT



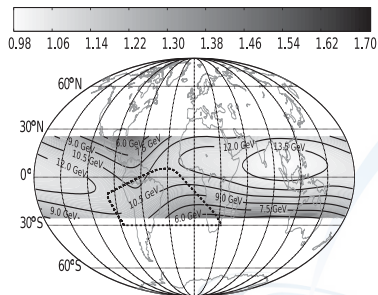
- ▶ Cut on P_{CRE} optimized in each energy bin
- ▶ Normalization of Monte Carlo templates fitted to data for residual background estimation

PERFORMANCE



- ▶ Selection increasingly difficult at higher energy
- ▶ Contamination below 20% in the whole energy range

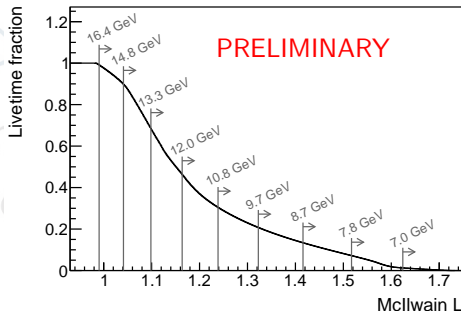
LE ORBITAL SELECTION



- ▶ GeV range: shielding effect of the Earth magnetic field
- ▶ LAT orbit: 0.98 - 1.73 McIlwain L, vertical rigidity cut-off from ~ 6 GeV to ~ 14 GeV.

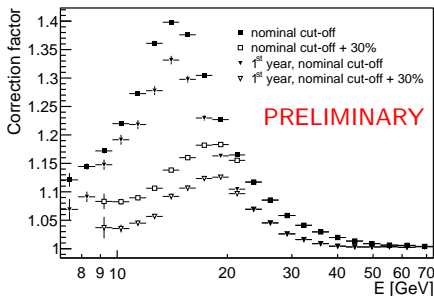
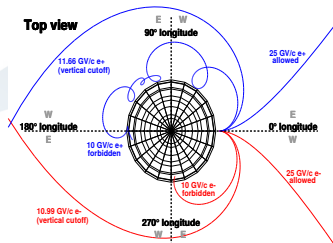
McIlwain L and rigidity cut-off values across the LAT orbit

- ▶ Energy dependent selection: in each bin accept only orbital regions where the whole bin is above the cut-off
- ▶ Optimized directly on data, by fitting the count spectrum in several McIlwain L bins



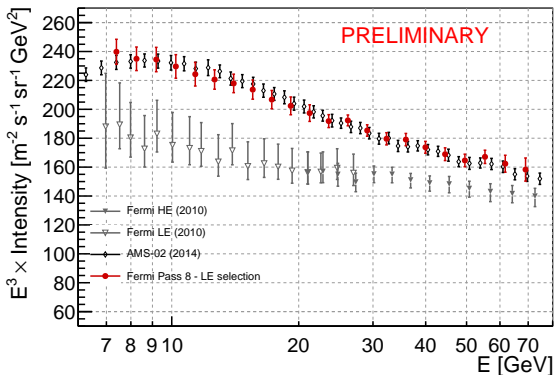
GEOMAGNETIC CORRECTION

- ▶ Cut-off is smooth: residual particle loss even after the selection
- ▶ Recovered with a tracing technique: simulated realistic flux of $e^+ + e^-$ in the LAT, traced back in a model of the Earth magnetic field
- ▶ Estimated correction factors (up to 40%) applied to the observed flux
- ▶ Separate corrections for first year of data (due to rocking angle increase from 35 to 50 degree in 2009)



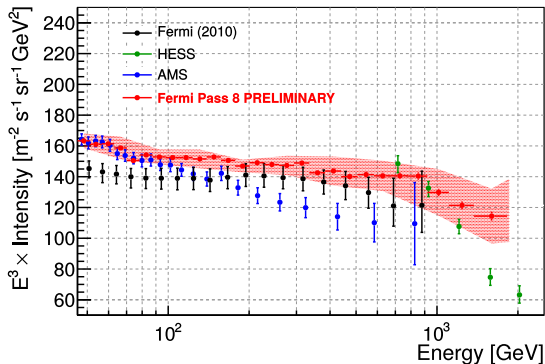
- ▶ **Acceptance uncertainty:**
 - ▶ Studied with a scan on the selection cut
- ▶ **Residual background uncertainty:**
 - ▶ Given by MC simulation uncertainties in reproducing the most electron-like hadronic showers.
- ▶ **Energy scale bias uncertainty:**
 - ▶ At 10 GeV given by comparison of observed CRE cut-off with tracer prediction
 - ▶ At 1 TeV study of data/MC disagreement of key energy reconstruction variables
- ▶ **Individual variable calibration uncertainty (HE analysis):**
 - ▶ Estimated by bracketing with different sets of corrections
- ▶ **Geomagnetic corrections uncertainty (LE analysis):**
 - ▶ Studied by varying the orbital selection and recomputing the corrections

RESULTS: LE SPECTRUM



- ▶ Significant difference wrt the previous Fermi publication
 - ▶ Geomagnetic correction (new)
 - ▶ 'Ghost' signals effect not included in previous simulation (corrected in Pass 8)

RESULTS: HE SPECTRUM



- ▶ Shaded band is sum in quadrature of systematic uncertainties
- ▶ Energy scale uncertainty not included
- ▶ Spectrum is harder than that reported by AMS-02
- ▶ Our systematic errors do not allow to confirm or exclude the cut-off observed by H.E.S.S.

CONCLUSIONS

- ▶ New measurement of the CRE spectrum, pushing the Fermi LAT to its limits in all metrics
- ▶ New event-level analysis (Pass 8)
- ▶ Energy range extended to 2 TeV (first direct measurement above 1 TeV)
- ▶ Better handling of the geomagnetic effects in the evaluation of the low-energy spectrum
- ▶ Better characterization of the systematic uncertainties
- ▶ Paper submitted to the journal (stay tuned)
- ▶ CRE selection will be used for anisotropies study update, currently under publication (again, stay tuned!)