



INTRODUCTION TO EFFECTIVE FIELD THEORIES OF QCD

U. van Kolck Institut de Physique Nucléaire d'Orsay and University of Arizona



Supported in part by CNRS, Université Paris Sud, and US DOE

Outline

- Effective Field Theories
- QCD at Low Energies
- Towards Nuclear Structure
 - Contact Nuclear EFT
 - Few-Body Systems
 - No-Core Shell Model
 - Halo/Cluster EFT



Conclusions and Outlook

References:

U. van Kolck, Effective field theory of short-range forces, Nucl.Phys.A645:273-302,1999, nucl-th/9808007

P.F. Bedaque, H.-W. Hammer, and U. van Kolck, **The three-boson system with short-range interactions,** Nucl.Phys.A646:444-466,1999, **nucl-th/9811046**

I. Stetcu, B.R. Barrett, and U. van Kolck, **No-core shell model in an effective-field-theory framework,** Phys.Lett.B653:358-362,2007, **nucl-th/0609023**

P.F. Bedaque, H.-W. Hammer, and U. van Kolck,
Narrow resonances in effective field theory,
Phys.Lett.B569:159-167,2003, nucl-th/0304007

Nuclear physics scales

"His scales are His pride", Book of Job



Lots of interesting nuclear physics at $E \sim 1 \text{ MeV}$ instead of $E \sim 10 \text{ MeV}$

within a few MeV of thresholds:

many energy levels and resonances (cluster structures)

most reactions of astrophysical interest

show universal features,

i.e. to a very good approximation are independent of details of the short-range dynamics

> bonus: same techniques can be used for dilute atomic/molecular systems



contact interactions among local nucleon fields



www.getcliparts.com

Cf. trapped fermions



universality!

Example: square well

$$V(r) = -\frac{\alpha^2}{mR^2} \theta \left(1 - \frac{r}{R} \right)$$

$$\Rightarrow T(k) = -i \left[e^{-2ikR} \frac{\sqrt{\alpha^2 + (kR)^2} \cot \sqrt{\alpha^2 + (kR)^2} + ikr}{\sqrt{\alpha^2 + (kR)^2} \cot \sqrt{\alpha^2 + (kR)^2} - ikr} - 1 \right]$$

zero-energy poles when $\alpha_c \equiv (2n+1)\pi/2$

 $a_0 = \mathbf{R} \left(1 - \frac{\tan \alpha}{\alpha} \right)$

 $r_0 = R \left(1 - \frac{R}{a_0 \alpha^2} - \frac{R^2}{3a_0^2} \right)$

etc.

fine-tuning generic $|1-\alpha/\alpha_c| \ll 1$ $\alpha = \mathcal{O}(1)$ $a_0 = -\frac{R}{\alpha_c^2} \left(1 - \frac{\alpha}{\alpha_c}\right)^{-1} \left\{1 + \ldots\right\} = \frac{1}{\aleph}$ $a_0 \sim \mathbf{R}$ $r_0 = \mathbf{R} \{1 + \ldots\} \sim \mathbf{R}$ $r_0 \sim R$ $\xi \equiv \frac{\left|1 - \alpha/\alpha_{c}\right|}{R} \ll \frac{1}{R}$



$Q \sim \aleph \ll M_{nuc}$

- d.o.f.: nucleons
- symmetries: Lorentz, P, T

$$\mathcal{L}_{EFT} = N^+ \left(i\partial_0 + \frac{\nabla^2}{2m_N} \right) N + C_0 N^+ N N^+ N$$

$$+N^{+} \frac{V^{+}}{8m_{N}^{3}}N + C_{2}N^{+}NN^{+}\nabla^{2}N$$

omitting spin, isospin

pionless EFT

 $+C_2' N^+ \vec{\nabla} N \cdot N^+ \vec{\nabla} N + \dots$

$$\begin{split} \bigvee &\sim iC_0(\Lambda) \\ \swarrow &\sim C_0^2(\Lambda) \int \frac{d^4l}{(2\pi)^4} \frac{1}{l^0 + p^0 - \frac{(\vec{l} + \vec{p})^2}{2m_N} + i\varepsilon} \xrightarrow{1} \frac{1}{-l^0 + p^0 - \frac{(\vec{l} + \vec{p})^2}{2m_N} + i\varepsilon} \xrightarrow{1} \frac{1}{q^2/M} \\ p^0 &= \frac{k^2}{2m_N} = -im_N C_0^2(\Lambda) \int \frac{d^3l}{(2\pi)^3} \frac{1}{\vec{l}^2 - k^2 - i\varepsilon} \xrightarrow{|l|} \\ &= -i\frac{m_N}{2\pi^2} C_0^2(\Lambda) \left\{ \int_0^{\Lambda} dl + k^2 \int_0^{\Lambda} dl \frac{1}{l + k + i\varepsilon} \frac{1}{l - k - i\varepsilon} \right\} \xrightarrow{|l|} \\ &= -im_N C_0^2(\Lambda) \left\{ \frac{1}{2\pi^2} \Lambda + i \frac{k}{4\pi} + O\left(\frac{k^2}{4\pi\Lambda}\right) \right\} = -iC_0^2(\Lambda)I_0(\Lambda) \\ &\swarrow \sim iC_2(\Lambda) k^2 \qquad \text{absorbed in} \qquad \text{non-analytic} \qquad \text{absorbed in} \\ &= tc. \end{split}$$

$$C_0(\Lambda) \to C_0^{(R)} \equiv C_0(\Lambda) \left\{ 1 - \frac{m_N \Lambda}{2\pi^2} C_0(\Lambda) + \dots \right\} = \frac{C_0(\Lambda)}{1 + \frac{m_N \Lambda}{2\pi^2} C_0(\Lambda)}$$

$$C_2(\Lambda) \rightarrow C_2^{(R)} \equiv C_2(\Lambda) - \frac{m_N}{4\pi\Lambda} C_0^2(\Lambda) + \dots$$

Naïve dimensional analysis

$$C_0^{(R)} \equiv \frac{4\pi}{m_N M_0}$$

...

$$C_0^{(R)} \sim \frac{4\pi}{m_N M_{muc}}$$

 $\implies M_0 \sim M_{nuc}$



 $C_2^{(R)} \sim \frac{m_N}{4\pi M_{nuc}} C_0^{(R)2} \implies M_2 \sim M_0$

etc.

But in this case:



⇒ no b.s. at $Q \leq M_{nuc}$, no good: just perturbation theory need one fine-tuning: $M_0 \equiv \aleph \ll M_{nuc}$ assume no other, *e.g.* still $M_2 \sim M_0$, *etc.*

v.K. '97 '99 Kaplan, Savage + Wise '98 Gegelia '98

 $2m_N$

$$= iC_{0}\left\{1 - C_{0}I_{0} + (C_{0}I_{0})^{2} + \ldots\right\} = \frac{i}{\frac{1}{C_{0}} + I_{0}} = \frac{4\pi}{m_{N}} \frac{i}{\frac{4\pi}{m_{N}C_{0}(\Lambda)} + \frac{2\Lambda}{\pi} + ik + \mathcal{O}\left(\frac{k^{2}}{\Lambda}\right)}$$
$$= \frac{4\pi}{m_{N}} \frac{i}{\aleph + ik} \left[1 + \mathcal{O}\left(\frac{k}{\Lambda}, \frac{k^{2}}{\aleph\Lambda}\right)\right] \qquad = \frac{4\pi}{m_{N}C_{0}^{(R)}} \equiv \aleph$$
$$= \frac{4\pi}{m_{N}C_{0}^{(R)}} = \frac{1}{\aleph}$$
$$= \frac{4\pi}{m_{N}C_{0}^{(R)}} = \frac{1}{\aleph}$$
$$= \frac{1}{8}$$
$$= \frac{1}{8}$$

+ ...

 $T^{(0)}$

=

+







Chen, Rupak + Savage '99 fitted $a_1 = 5.42 \text{ fm}(\text{exp})$ $r_1 = 1.75 \text{ fm}(\text{exp})$ predicted $B_d = 1.91 \text{ MeV}(\nu = 0)$ $B_d = 2.22 \text{ MeV}(\text{exp})$





renormalization: quartet s wave







renormalization: doublet s wave -I





renormalization: doublet s wave -I





renormalization: doublet s wave -I



renormalization: doublet s wave -II

Bedaque, Hammer + v.K. '99 '00



Phillips line



+ four-body bound state can be addressed similarly \implies no four-body force at v = -1



Hammer, Meissner + Platter '04

Summary:

Expansion parameter

$$\frac{Q}{M_{nuc}} \sim \frac{\aleph}{M_{nuc}} \sim \frac{r_0}{a_0}$$

LO: two two-nucleon + one three-nucleon interactions

 $C_0^{(0)},\,C_0^{(1)},\,D_0^{(1)}$

- NLO: two more two-nucleon interactions
- etc.
- ~ larger nuclei?

As A grows, given computational power limits number of accessible one-nucleon states IR cutoff $\lambda \ll Q$ in addition to UV cutoff $\Lambda \gg Q$

Finite Volume

. . .



nuclear matter Mueller, Seki, Koonin + v.K. '99 few nucleons Lee et al '05



finite nuclei Stetcu, Barrett + v.K. '07 Rotureau, Stetcu, Barrett + v.K. '12 few atoms Stetcu, Barrett + v.K. '09

W atoms Stetcu, Barrett + v.K. '09 Rotureau, Stetcu, Barrett + v.K. '10 Rotureau et al. '11

Two possible approaches



Both being actively pursued

- many-body systems get complicated rapidly

+ (continuing) focus on simpler <u>halo/cluster nuclei</u>

one or more loosely-bound nucleons around one or more cores

 $\bigotimes \equiv \sqrt{m_N E_p} \ll \sqrt{m_N E_c} \equiv M_c \quad (esp. near driplines)$ particle core
separation energy excitation energy

e.g.

4

He
$$\begin{cases} B_{\alpha^*} \cong 8 \text{ MeV} \\ B_{\alpha} \cong 28 \text{ MeV} \end{cases}$$
 $E_{\alpha} = B_{\alpha} - B_{\alpha^*} \cong 20 \text{ MeV}$

"5 He" $p_{3/2}$ resonance at $E_n \sim 1 \text{ MeV}$

⁶He S_0 bound state at $E_{2n} \sim 1 \text{ MeV}$



$Q \sim \aleph \ll M_c$

- degrees of freedom: nucleons, cores
- symmetries: Lorentz, P, T
- expansion in:

 $\frac{Q}{M_c} = \begin{cases} Q/m_N, Q/m_c \\ Q/m_{\pi}, \cdots \end{cases}$

non-relativistic multipole

halo EFT

simplest formulation: auxiliary fields for core + nucleon states

e.g. ⁴He
$$\mapsto$$
 scalar field φ
⁴He + N
$$\begin{cases} s_{\frac{1}{2}} \equiv 0 + \mapsto \text{spin} - 0 \text{ field } s \\ p_{\frac{1}{2}} \equiv 1 - \mapsto \text{spin} - 1/2 \text{ field } T_1 \\ p_{\frac{3}{2}} \equiv 1 + \mapsto \text{spin} - 3/2 \text{ field } T_3 \\ \vdots \end{cases}$$





More fine-tuning!!!		$a_0 \ (10^3 \ {\rm fm})$	$r_0 \ ({\rm fm})$	$\mathcal{P}_0~(\mathrm{fm}^3)$
	LO	-1.80	1.083	
	NLO	-1.92 ± 0.09	1.098 ± 0.005	-1.46 ± 0.08
		2 /2		

fine-tuning of 1 in 1000! $|a_0| \sim M_c^2 / \aleph^3 \qquad r_0 \sim 1 / M_c$ $|a_0^{E\&M}| = \mathcal{O}(1/2k_c) \simeq 1.8 \text{ fm} \qquad \tilde{r}_0 = -0.13 \text{ fm}$

fine-tuning of 1 in 10

What next

- Coulomb interaction in higher waves: e.g. $p + {}^{4}\text{He} \rightarrow {}^{4}\text{He} + p$ [cf. $p + p \rightarrow p + p$ Kong + Ravndal '99
- three-body bound states: e.g. 1) 6 He = b.s. (4 He + n + n) Cf. 3 H = b.s. (p + n + n) Rotureau + v.K., '12 + in progress Rotureau + v.K., '12 + in progress Bedaque, Hammer + v.K. '99
- reactions: e.g. $p + {}^{7}\text{Be} \rightarrow {}^{8}\text{B} + \gamma$ [cf. $p + n \rightarrow d + \gamma$

Higa + Rupak, in progress

Chen et al. '00

Conclusion

EFT the framework to describe nuclei within the SM
is consistent with symmetries
incorporates hadronic physics
has controlled expansion
many successes so far, but still much to do
grow to larger nuclei!

new, systematic approach to physics near d