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## Limits on <sup>60</sup>Fe/<sup>26</sup>Al nucleosynthesis ratios from deep-sea sediment AMS measurements

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% % Nuclear Physics in Astrophysics 8 template for abstract % % Format: LaTeX2e. % % Rename this file to name.tex, where 'name' is the family name % of the first author, and edit it to produce your abstract. % \documentstyle[11pt]{article} % % PAGE LAYOUT: % \textheight=9.9in \textwidth=6.3in \voffset -0.85in \hoffset -0.35in \topmargin 0.305in \oddsidemargin +0.35in \evensidemargin -0.35in %\renewcommand{\rmdefault}{ptm} % to use Times font  $\label{long} \label{long} \la$  $\log\left(\frac{1 \#2}{1 \#2}\right)$ \begin{document} {\small \it Nuclear Physics in Astrophysics 8, NPA8: 18-23 June 2017, Catania, Italy} \vspace{12pt} \thispagestyle{empty} \begin{center} %%% %%% Title goes here. %%% \TITLE{Limits on <sup>60</sup>Fe/<sup>26</sup>Al nucleosynthesis ratios from deep-sea sediment AMS measurements}\\[3mm] %%% %%% Authors and affiliations are next. The presenter should be %%% underlined as shown below. %%%  $\label{eq:authors} $$ AUTHORS{J. Feige^{1,2}, A. Wallner^{2,3}, L.~K. Fifield^3, R. Golser^2, S. Merchel^4, G. Rugel^4, P. Steier^2, \ S.~G. Tims^3, S.~R. Winkler^{2,5} $$$ %%% {\small \it

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The long-lived radionuclide  $^{26}$ Al (t<sub>1/2</sub>\,=\,0.7\,Myr) has been observed throughout our galaxy, reflecting ongoing nucleosynthesis over the past few million years [1]. It is produced and ejected into the interstellar medium by stellar winds and during supernova explosions. A nearby supernova may leave an imprint of  $^{26}$ Al in terrestrial archives, complementing the observation of supernova-produced  $^{60}$ Fe in deep-sea samples. \\ The same set of sediment samples from the Indian Ocean that showed a distinct  $^{60}$ Fe-signature in layers of ages between 1.7 and 3.2\,Myr [2] was also analyzed for  $^{26}$ Al. However, additional terrestrial sources producing  $^{26}$ Al on Earth, such as cosmogenic production in the atmosphere and in-situ production within the sediment, may obscure a supernova imprint. \\

We used our experimental <sup>26</sup>Al data to infer lower limits on <sup>60</sup>Fe/<sup>26</sup>Al nucleosynthesis ratios by comparing the width and the strength of the previously measured <sup>60</sup>Fe-signal to our <sup>26</sup>Al data. We find that our results generally favour the higher theoretical isotopic supernova ratios and deviate from the observed galactic  $^{60}$ Fe/<sup>26</sup>Al flux ratio by 2-3 times of the measurement uncertainty.

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\noindent [1] Diehl et al., New Astron. Rev., 52, 440 (2008);

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[2] Wallner, Feige et al., Nature, 532, 69 (2016).}
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