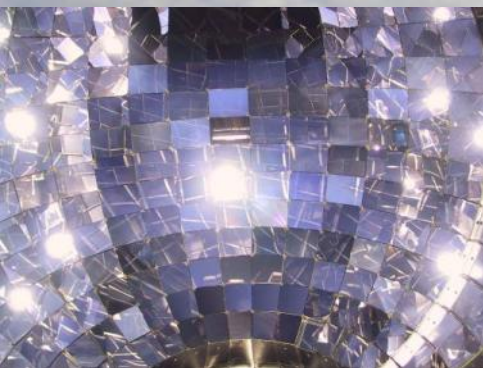


A new method for the determination of very small Γ_γ partial widths

G.Cardella
for the
NEWCHIM collaboration



Istituto Nazionale di Fisica Nucleare

Sezione di Catania



Outline

Status
of the
art

What is known on ^{12}C formation

Measured and evaluated decay widths

Old and New experimental methods
adopted

The
new
method

Description of the method

The CHIMERA detector status and
ongoing upgrading

Kinematic coincidences

Gamma-ray detection

obtained
results &
future

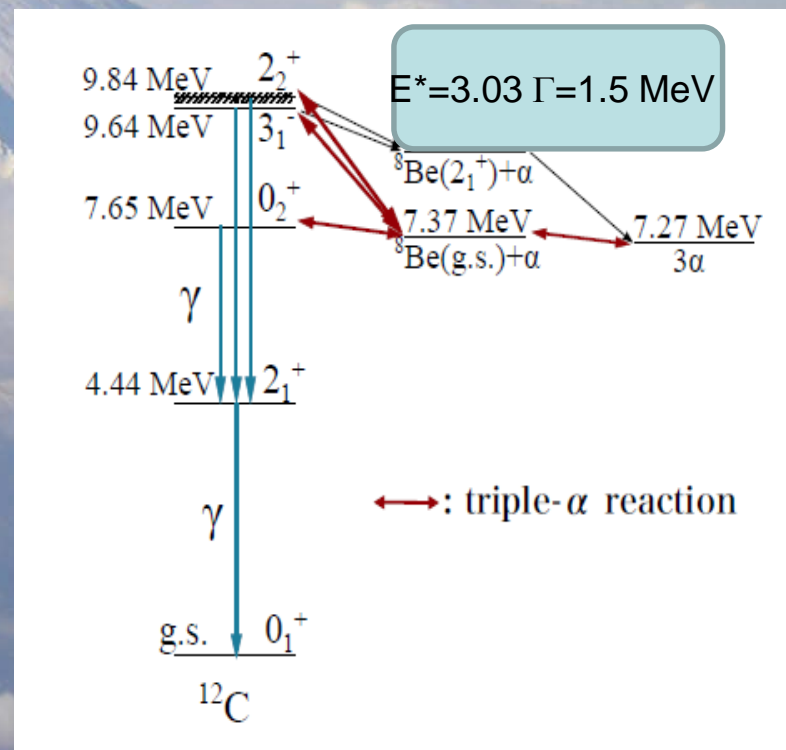
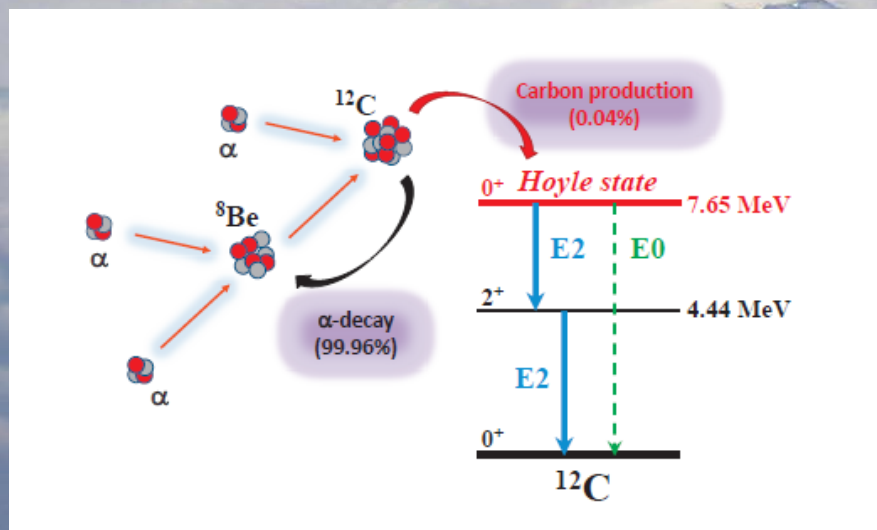
^{68}Ni PYGMY resonance

Future perspectives

Status of the art: ^{12}C production in the Universe

The 3- α process is the main responsible of formation of ^{12}C in stars
 see for instance AGB stars F.Hervig et al PRC
73, 025802 (2006)

At higher temperature, 10^9k or larger, (core collapse supernova) also the ^{12}C 3^- level at 9.64 MeV give some contribution C.Angulo et al NPA 656 (1999) 3 (NACRE compilation)

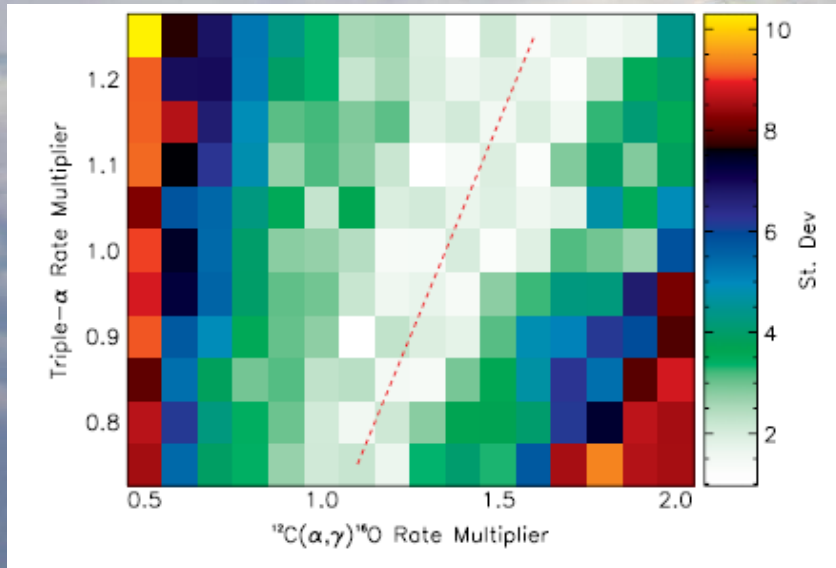


B Alshahrani et al EPJ Web 63,01022(2013)

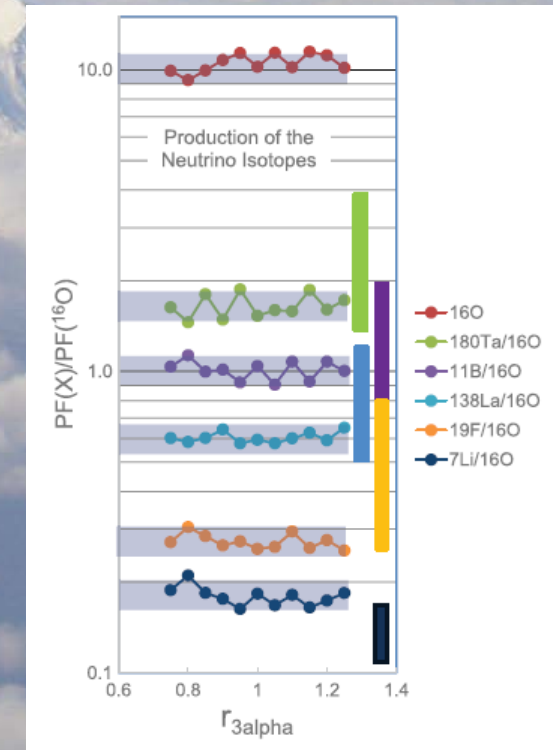
M.Tsumura et al Jou. of phys. Conf.S.
 569(2014)012051

Status of the art: Effective values for the 3α and $^{12}\text{C}(\alpha,\gamma)^{16}\text{O}$ processes

Recent calculations Sam Austin et al PRL 112(2014)111101 show that a very important parameter is the ration between the 3α process and the $^{12}\text{C}(\alpha,\gamma)^{16}\text{O}$ one. This is also due to various unknown processes as in convection not fully governed by simulations



With this Effective values it was possible to describe the production of neutrino emitter in Supernova explosions with rather good precision



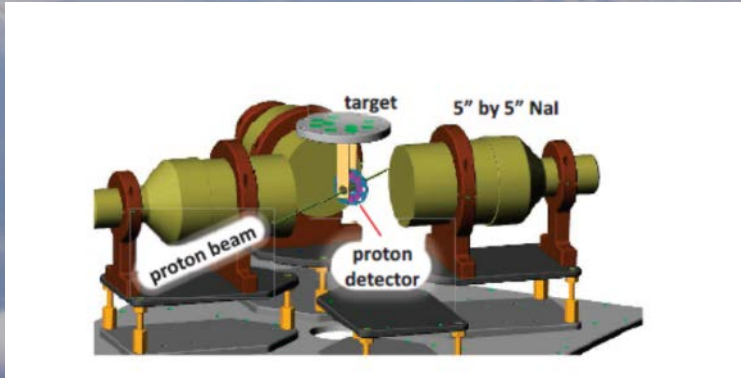
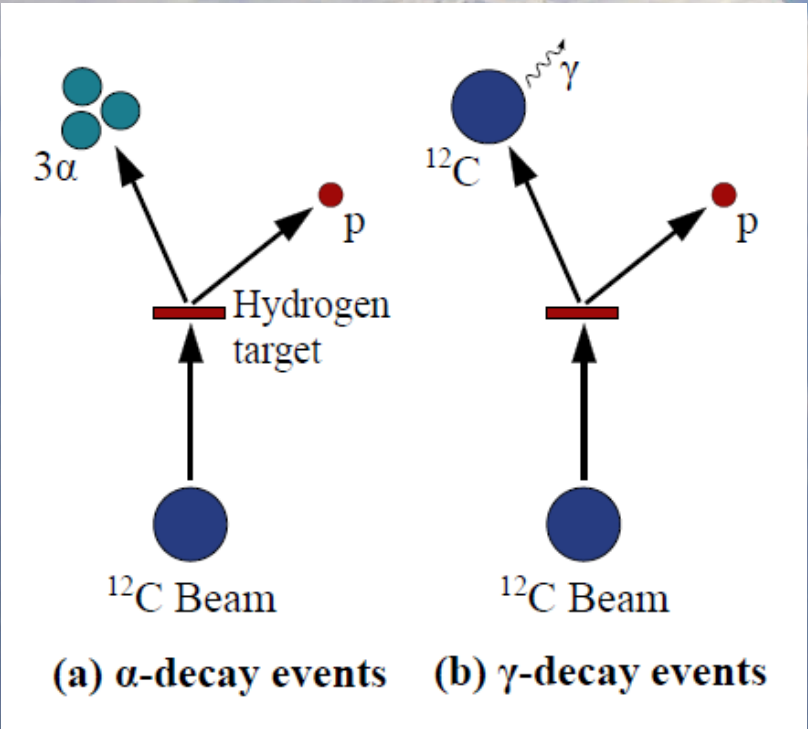
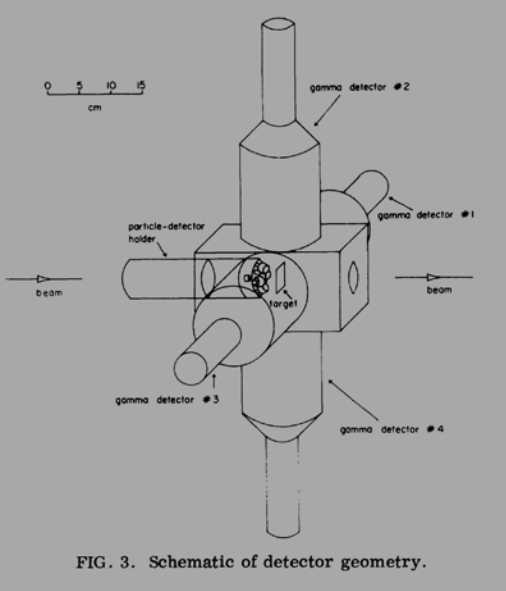
Similar results were obtained this year by the same author on gamma emitting nuclei Al-26 / Ti44 / Fe60 AJL 839(2017)

Status of the art: measurements performed

Hoyle state $\Gamma_{\text{rad}}/\Gamma$ of $4.12 \pm 0.11 \cdot 10^{-4}$

Most of the measurements done looking for coincidence events $p+^{12}\text{C}$ following the methods of D. Chamberlin et al Phys.Rev.C 9(1974)69

But also search of $\gamma-\gamma$ coincidences A. W. Obstt and W. J. Braithwaite phys.Rev.C 13(1976)2033



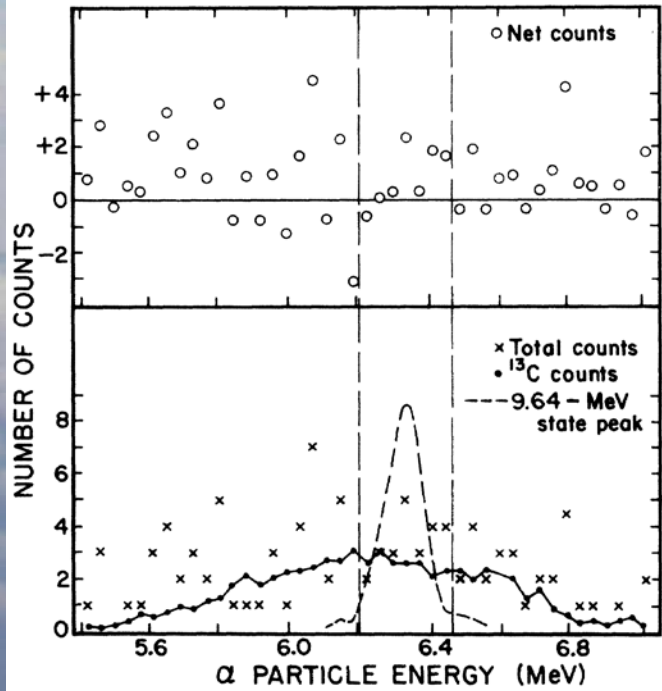
B Alshahrani et al

EPJ Web of Conferences 63, 01022 (2013)

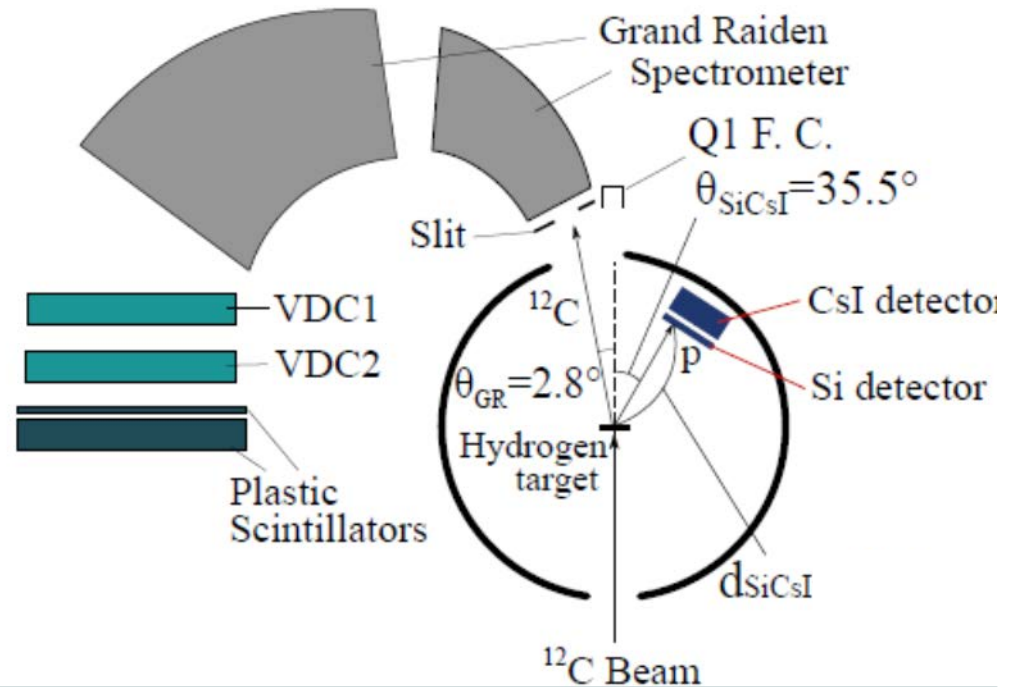
Status of the art: measurements performed

9.64 (3⁻)

Γ	Γ_γ	Γ_γ/Γ
34(5) keV	$> 0.31(4)$ meV	$> 9.1 \times 10^{-9}$
	< 14 meV(1 σ C.L.)	$< 4.1 \times 10^{-7}$



Upper limits given by Chamberlin et al due to ¹³C contaminants in the target in experiment based on the recoil measurement α induced reactions at 24 MeV

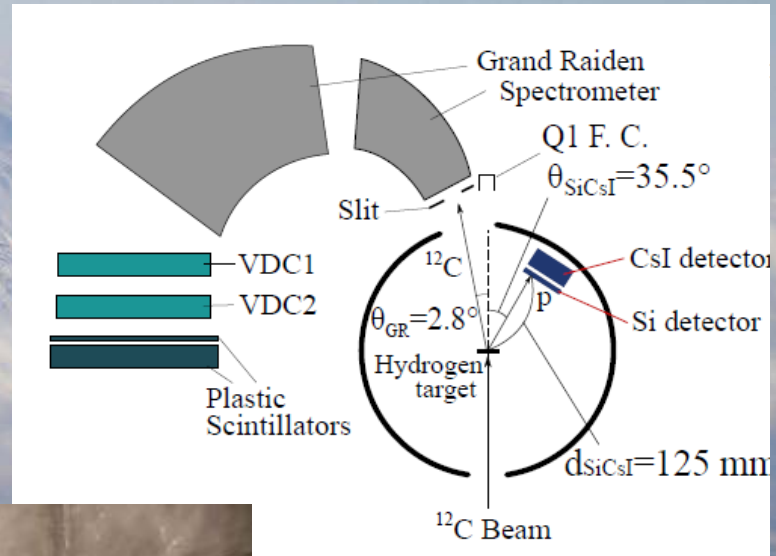
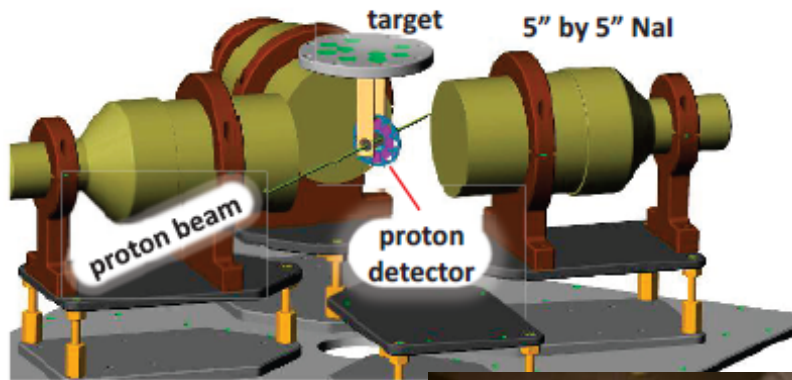


To improve such result a new experiment is in preparation in Japan using ¹²C beam detected by the Grand Raiden Spectrometer

M.Tsumura et al *Journal of Physics: Conference Series* **569** (2014) 012051

The new method: description

The Basic idea is to join both detection methods using the CHIMERA detector in order to improve the sensitivity of the measurement : **4 fold coincidence measurement**

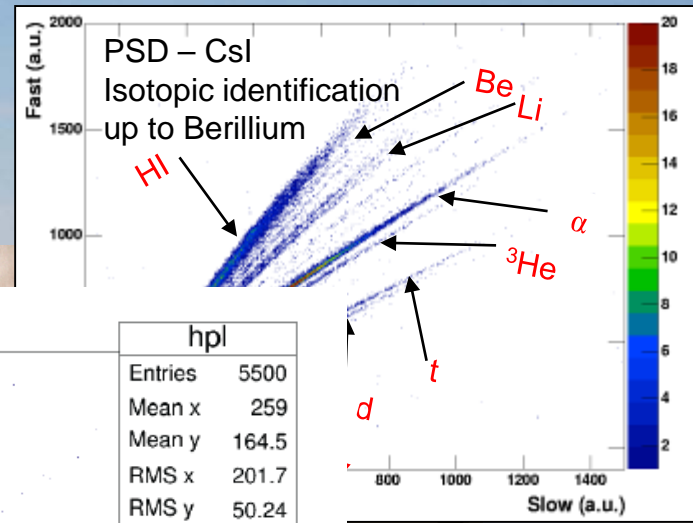
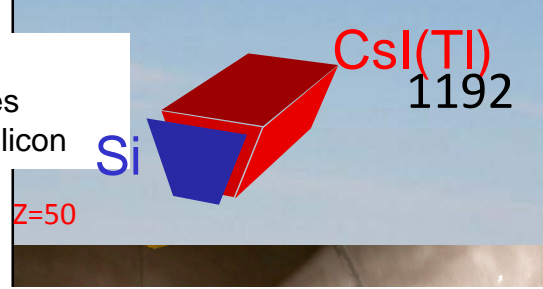
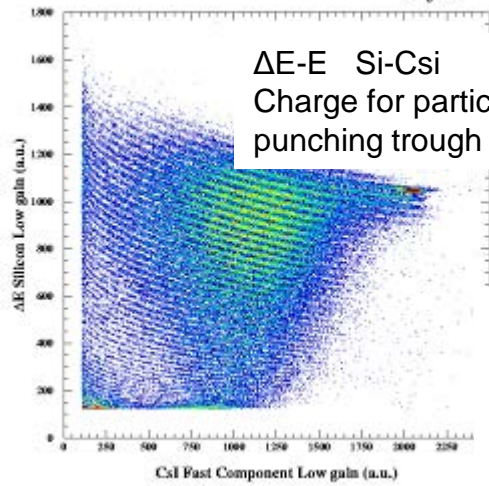


γ -ray detection

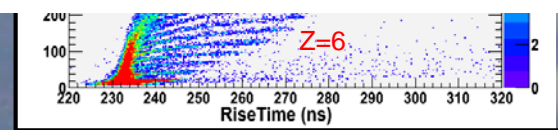
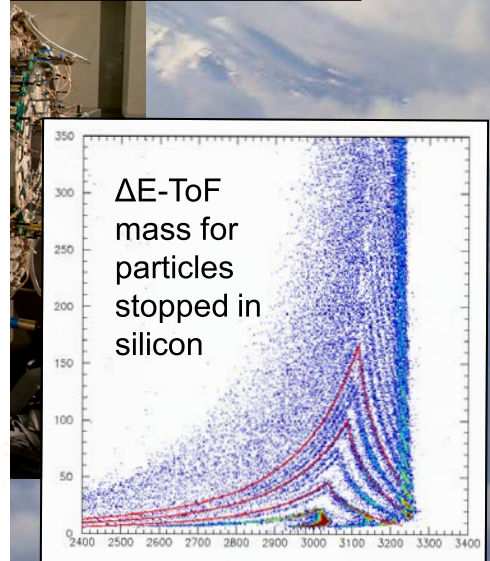
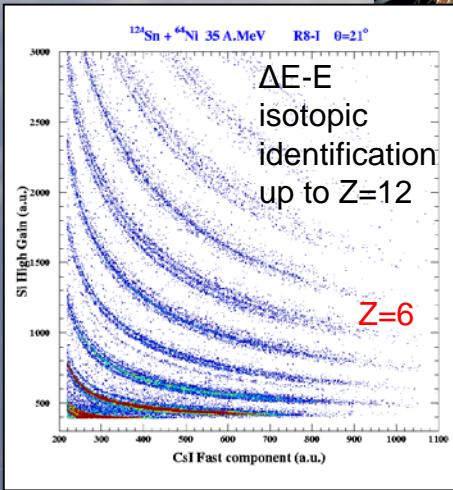
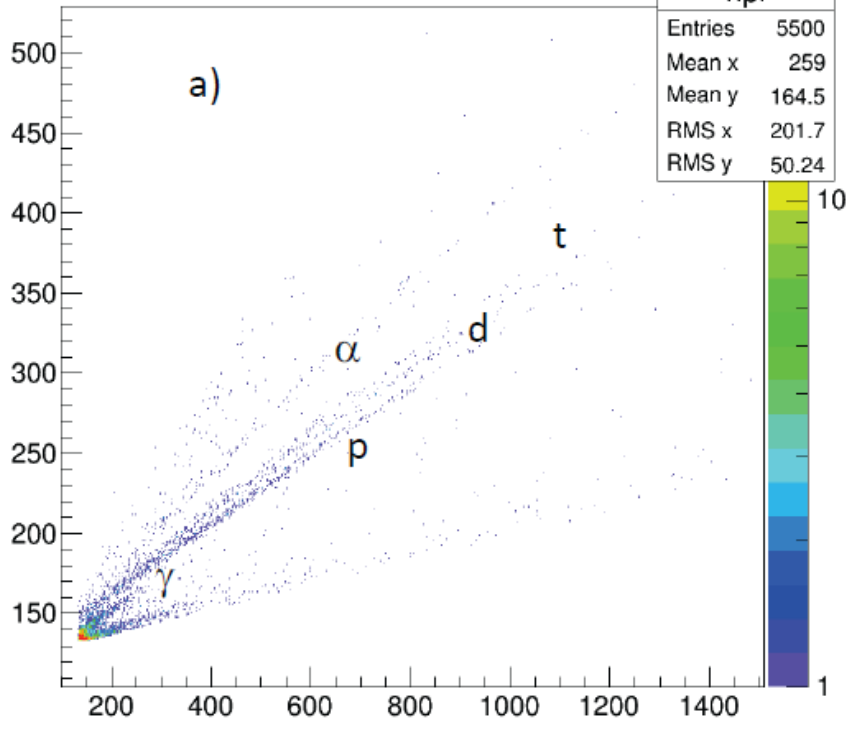


Kinematic
coincidence with
 ^{12}C recoil

The new method: Characteristics of the CHIMERA detector



hpl	
Entries	5500
Mean x	259
Mean y	164.5
RMS x	201.7
RMS y	50.24



The new method I: kinematic coincidences with CHIMERA

Nuclear Instruments and Methods in Physics Research A 715 (2013) 56–61

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Nuclear Instruments and Methods in
Physics Research A

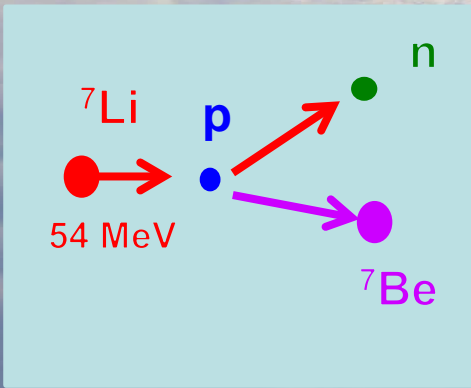
journal homepage: www.elsevier.com/locate/nima



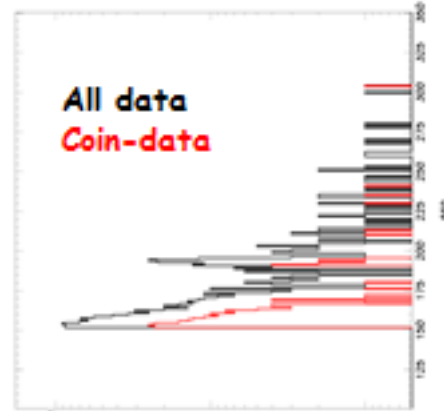
Detection efficiency around 100%

Kinematical coincidence method in transfer reactions

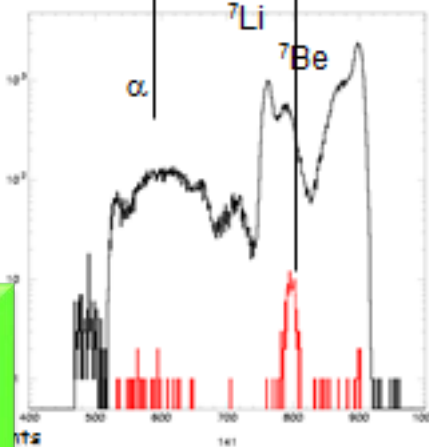
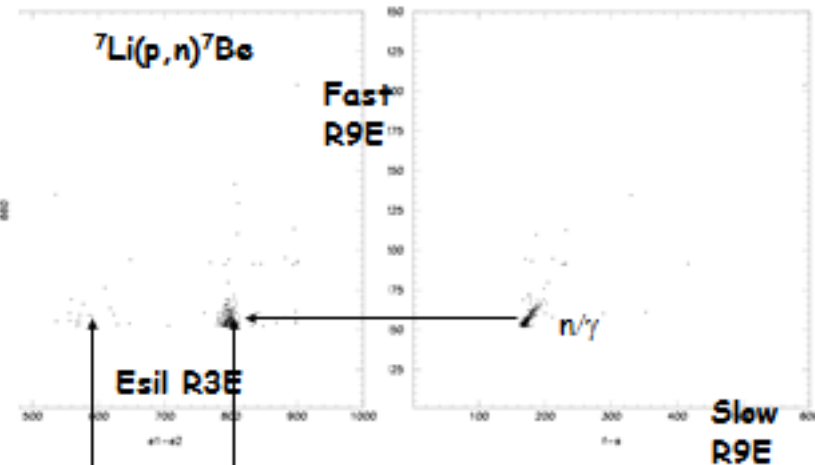
L. Acosta^b, F. Amorini^b, L. Auditore^d, I. Berceanu^h, G. Cardella^{a,*},
L. Francalanza^{b,c}, R. Gianì^{b,c}, L. Grassi^{a,k}, A. Grzeszczuk^j, E. La Gui
D. Loria^d, T. Minniti^d, E.V. Pagano^{b,c}, M. Papa^a, S. Pirrone^a, G. Pi
E. Rosato^f, P. Russotto^{b,c}, S. Santoro^d, A. Trifirò^d, M. Trimarchi^d,



$$\Delta\phi (n-{}^7\text{Be}) = 180^\circ$$



Another advantage:
relatively high
detection efficiency
for neutrons



Coincidences
between
ring3E and 9E
detectors at
 $\Delta\phi = 180^\circ$

More than 3 order of magnitude background suppression only with the request of a kinematical coincidence (look to elastic scattering) (no time window no identification)



The new method II: γ -detection with CsI(Tl)

Nuclear Instruments and Methods in Physics Research A 799 (2015) 64–69


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Physics Research A

journal homepage: www.elsevier.com/locate/nima



Particle gamma correlations in ^{12}C measured with the CsI(Tl) based detector array CHIMERA



Using proton beams on C target we excite the 4.44 MeV level – this level decay emitting a γ -ray that we can see in our CsI(Tl)

The CsI(Tl) of the CHIMERA sphere have a relatively good detection efficiency for γ -rays that we computed with GEANTIV and we have experimentally verified

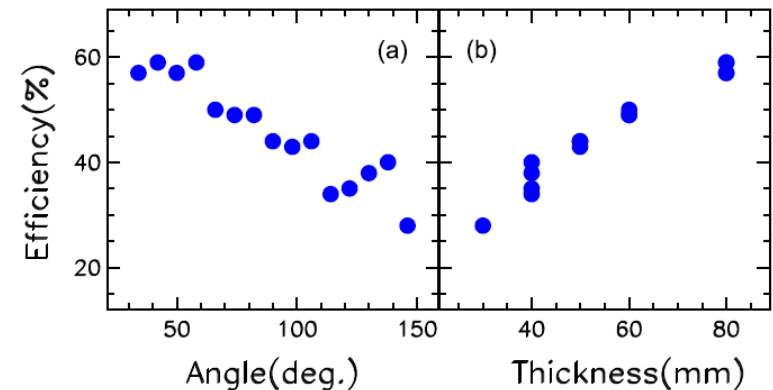
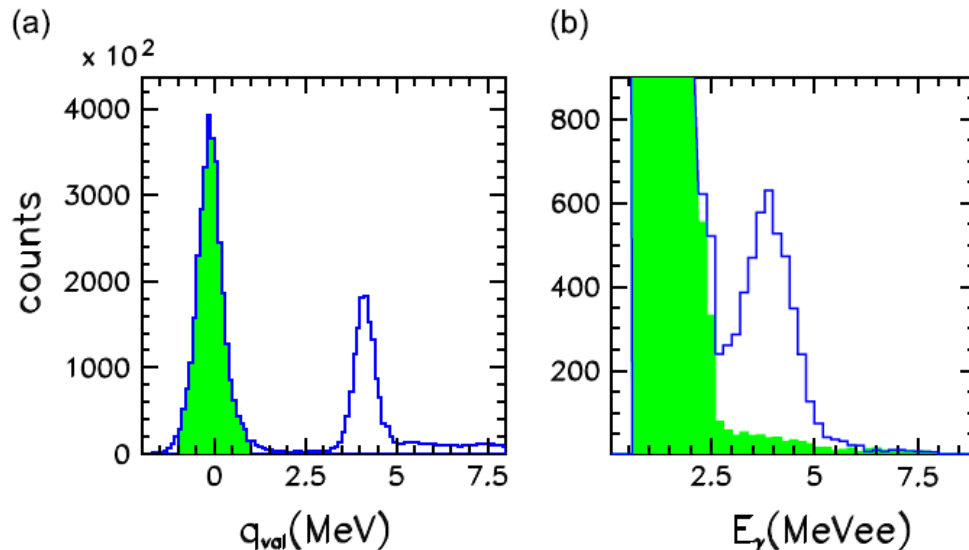
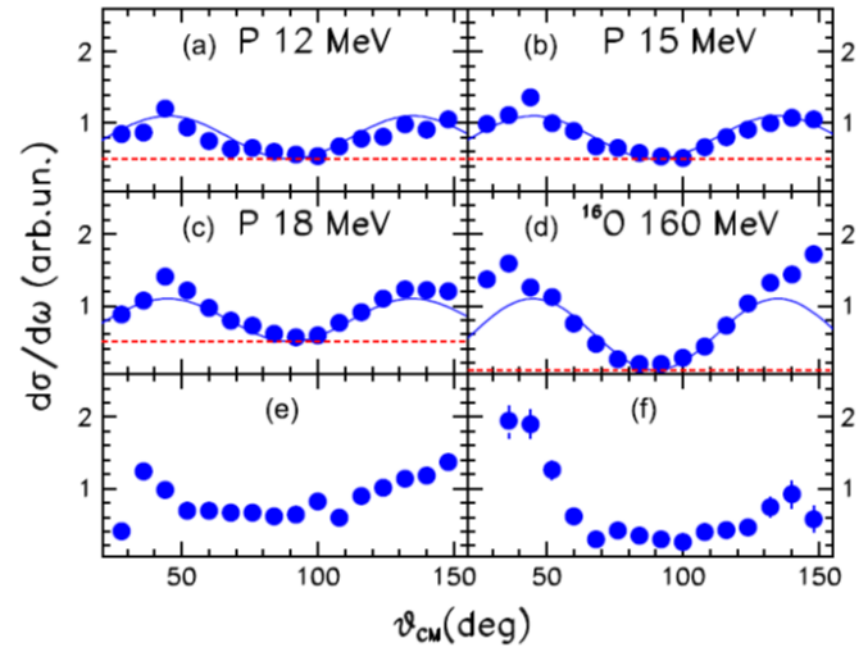
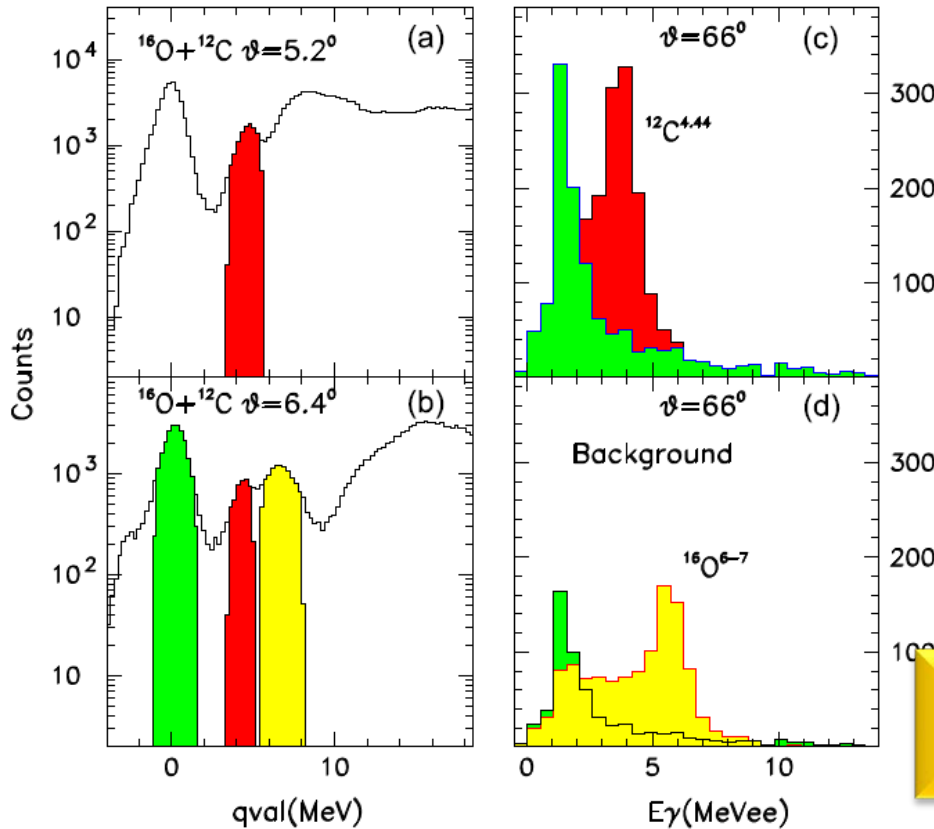


Fig. 1. CsI(Tl) efficiency to γ -rays of 4.44 MeV computed with Geant4 as a function (a) of detection angle, (b) of the thickness of scintillator.

The new method: γ -detection with CsI(Tl)

To verify the E2 character of the transition we can see the angular distribution in the reference frame of emitting recoil nucleus ^{12}C

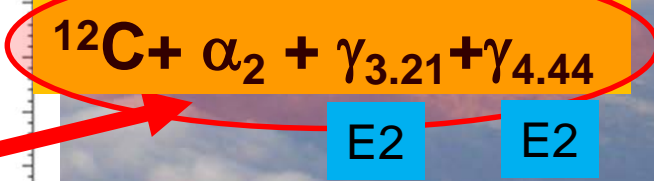
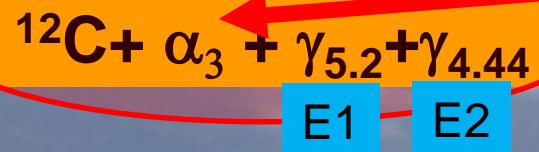
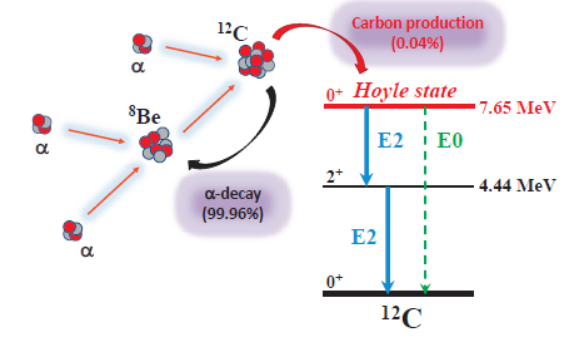
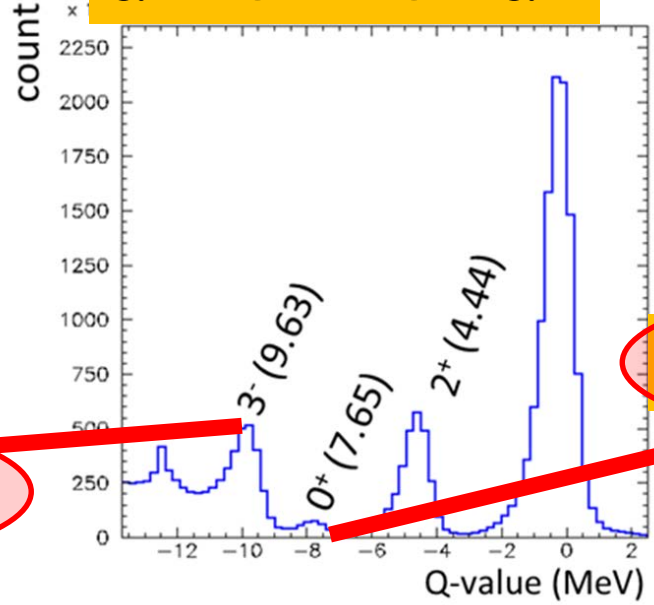
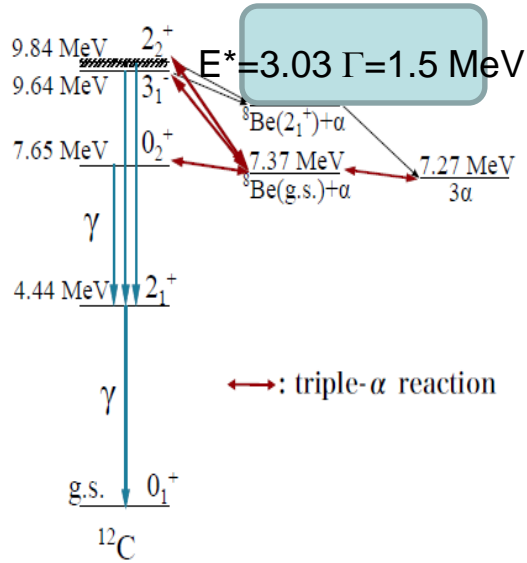


E2 behavior is verified by the minimum at 90° and maxima at 45° and 135° - the minimum is not zero because of the proton spin $\frac{1}{2}$

using zero spin beam (^{16}O) we get more deep minimum

Note that with ^{16}O beam we can also see γ -rays around 6-7 MeV

The method: CHIMERA is able to see the 4 fold coincidence with good efficiency



Around 9-10 MeV it was localized also a 2^+ state interpreted as excited 2^+ Hoyle status. If excited this could decay with a γ -ray of about 2 MeV E2 to the 0^+ Hoyle status we will also look for this channel

The method: rejection factor

condition

Spurious coincidence suppression factor

4-fold coincidence

from time resolution Si and CsI 10^{-13}

particle and γ identification

from count rate at different angles $\approx 3-100$

kinematic coincidence

from $\Delta\theta-\Delta\phi \approx 30$

Qvalue - measurement

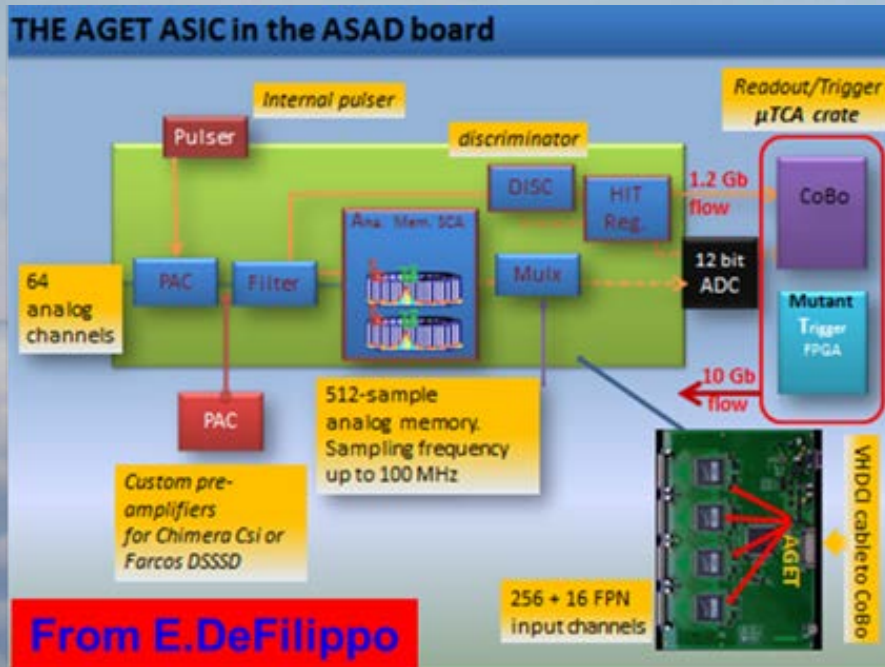
from Energy resolution $\approx 10-20$

total γ -ray energy

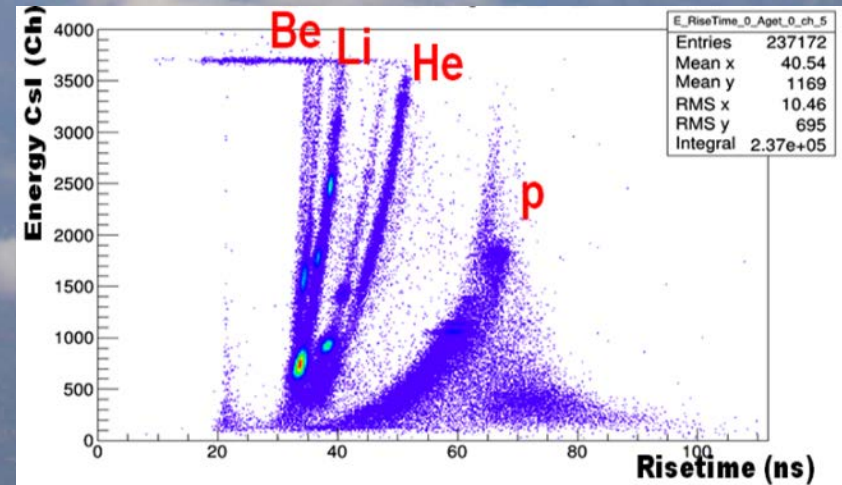
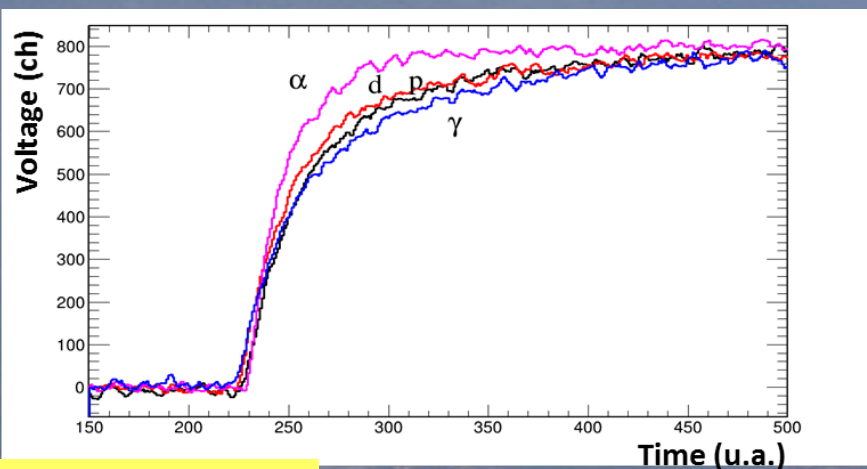
from Energy resolution $\approx 5-10$

10^{-18} should be enough to well detect 10^{-9} decay probability of the 9.64 3^- level

Recent upgradings : γ -detection with CsI(Tl) NEW GET electronics

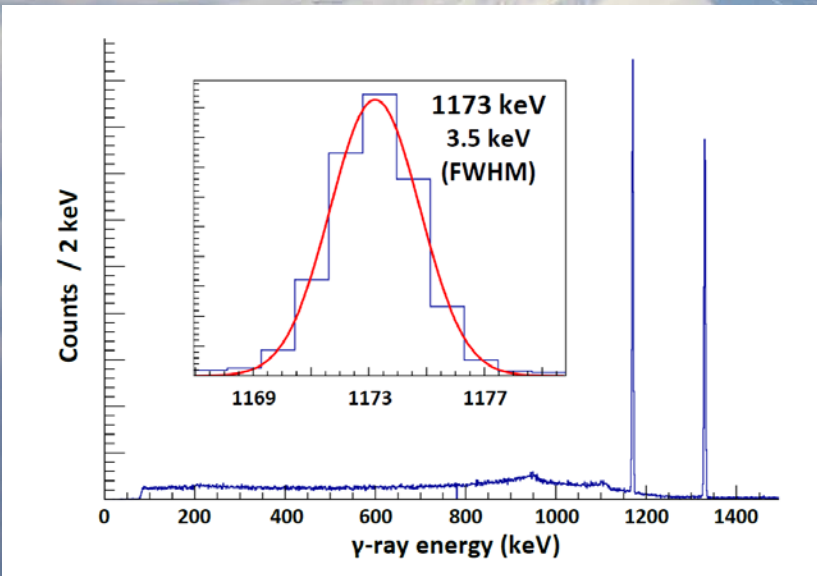


We are upgrading the CHIMERA electronics with GET – Using GET we can have the full digitalization of CsI(Tl) signals therefore we can perform a simple identification of γ -rays. Moreover we have also a trigger information from CsI(Tl) missing in the old CHIMERA electronics

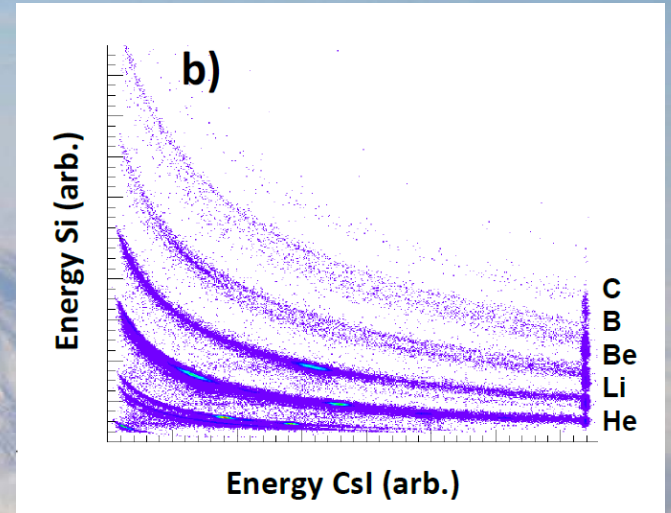


Recent upgradings : γ -detection with CsI(Tl) NEW GET electronics

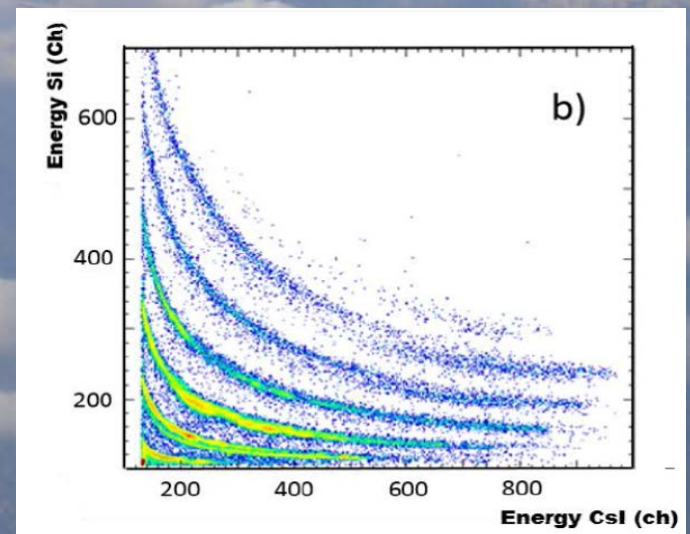
The quality of the new electronics was checked by using signals from a Germanium detector



only 1 keV worse than standard 14 bit Adc with Ortec amplifier

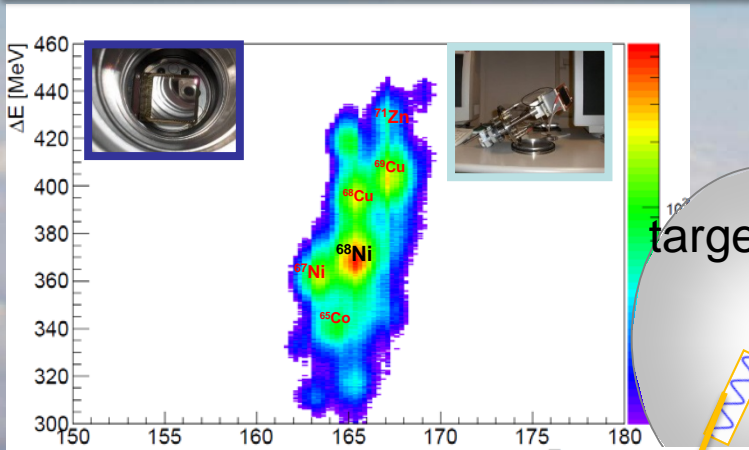


better mass identification in ΔE -E scatter plots (same telescope same reaction)

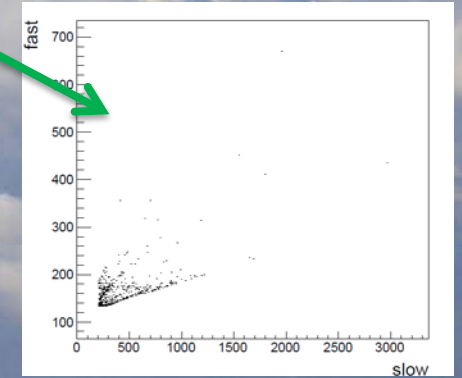
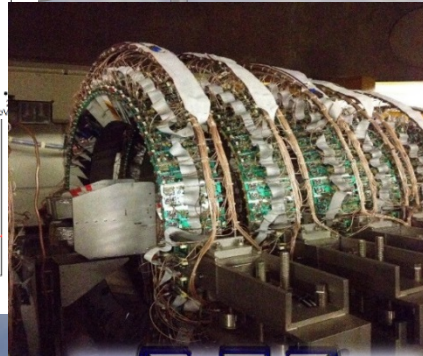
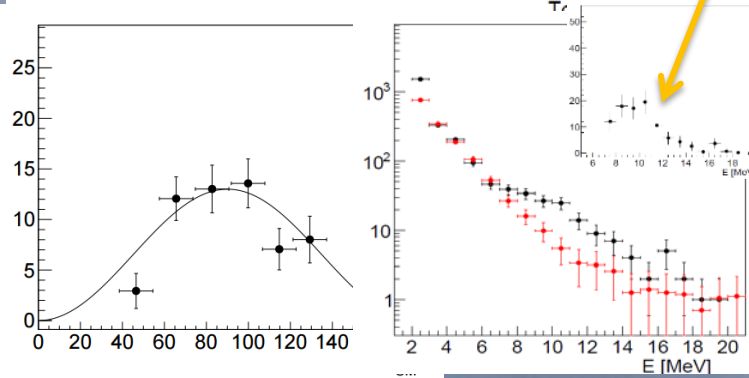
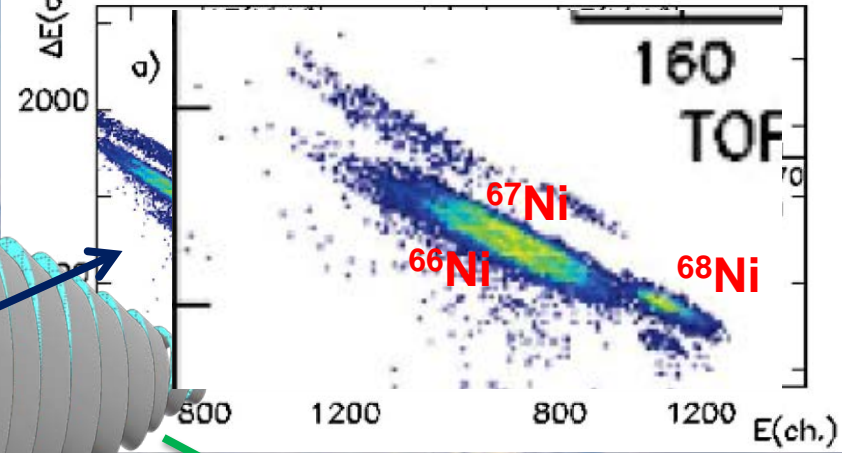


γ -detection with CsI(Tl) : Observation of isoscalar excitation of PYGMY resonance in ^{68}Ni

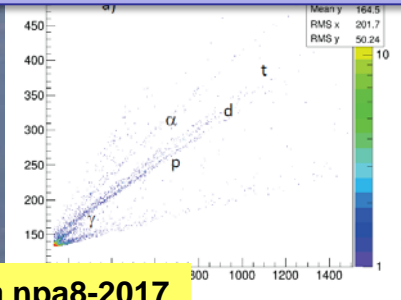
^{68}Ni fragmentation beam produced at LNS and event by event identified



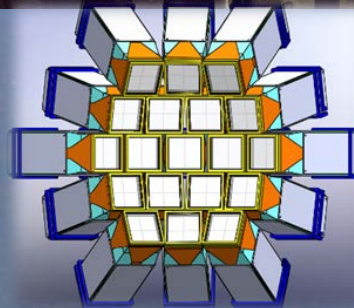
^{68}Ni identification on FARCOS silicon detectors



γ -rays on the CsI(Tl) of the sphere



neutrons on the CsI(Tl) of CHIMERA rings covered by FARCOS

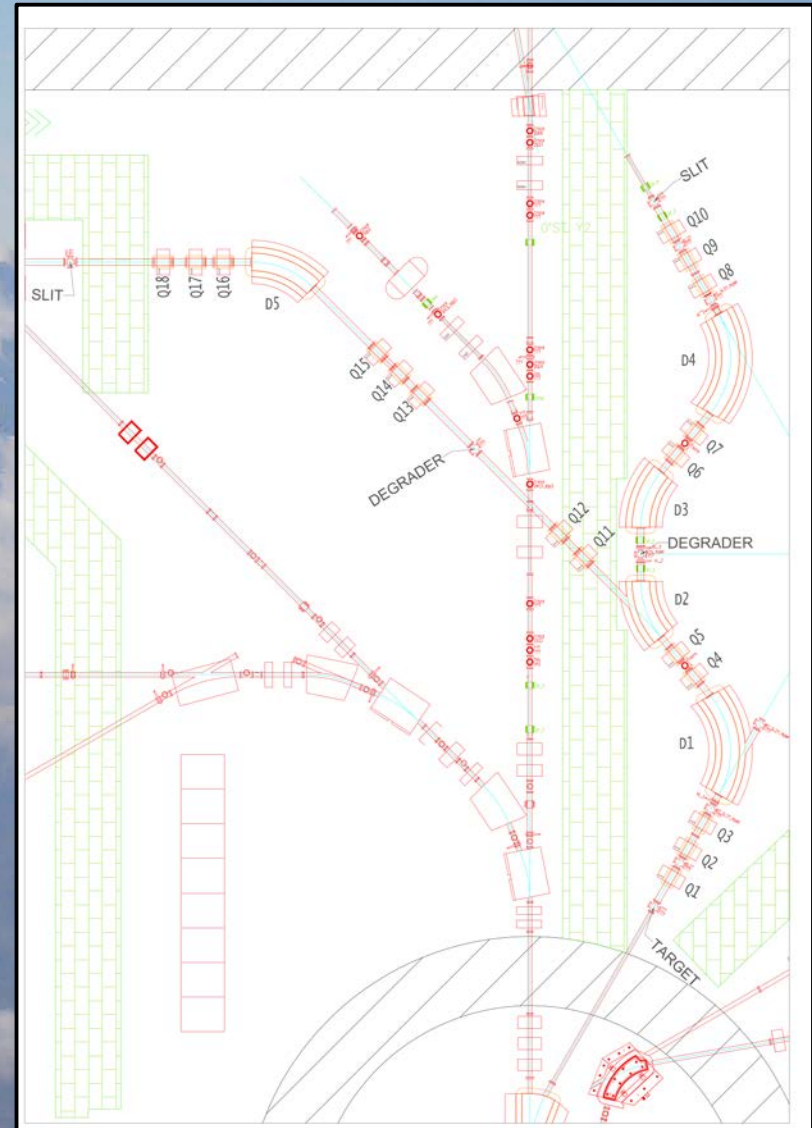


Future perspectives: upgraded Cyclotron at LNS

An upgrading of the cyclotron is going at LNS – high power beam will be available up to 10 MW

A new fragment separator is going to be implemented in order to use this high intensity beam – we will produce in few years very intense exotic radioactive beams in the range of 10^5 - 10^6 particle/s

With such beautiful beam and CHIMERA upgraded we hope to perform exciting experiments you are invited to join us with your smart ideas



Summary and outlook

Use of
CHIMERA
detector

Chimera show its versatility

With the new upgrading improved
resolution

Beautiful perspectives for γ -particle
coincidences with radioactive and stable beams

We hope to have beautiful data at the end
of this year

The Hoyle
measurement

we should be able to improve the lower
limit for the Γ_γ/Γ for the 9.64 3^- level

We will also search for signals from the
decay of the 2^+ excited Hoyle state

Future
opportunities

The new fragment separator at LNS will
provide high intensity radioactive beams to
perform beautiful experiments

The collaboration

L.Acosta^{1,8}, L.Auditore⁴, G.C.¹, E. De Filippo¹, D.Dell'Aquila⁶, S. De Luca⁴, N.Gelli⁹, B.Gnoffo¹, B.Fornal¹⁰, G.Lanzalone^{2,7}, S.Leoni⁵, I.Lombardo⁶, C.Maiolino², N.Martorana², A.Nannini⁹, S.Norella⁴, A.Pagano¹, E.V.Pagano^{2,3}, M.Papa¹, S.Pirrone¹, G.Politi^{1,3}, L.Quattrocchi⁴, F.Rizzo^{2,3}, P.Russotto¹, D.Santonocito², A.Trifirò⁴, M.Trimarchi⁴, M.Vigilante⁶

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