

Investigating nuclear reactions at astrophysical energies with γ -ray beams and an active-target TPC

Chiara Mazzocchi

on behalf of the ELITPC collaboration:

M. Bieda, J.S. Bihałowicz, M. Ćwiok, W. Dominik, Z. Janas, Ł. Janiak,
J. Mańczak, T. Matulewicz, C.M., M. Pfützner, P. Podlaski, S. Sharma,
M. Zaremba (*University of Warsaw*), M. Gai (*University of Connecticut*),
D. Balabanski, A. Bey, D.G. Ghita, O. Tesileanu (*IFIN-HH/ELI-NP, Romania*)

Overview

✓ Physics goal

✓ How?

- Direct capture versus photo-disintegration reactions
- Monochromatic γ -ray beams @ELI-NP
- An active-target Time Projection Chamber (ELITPC)
- Status of the project: R&D and demonstrators

✓ Outlook

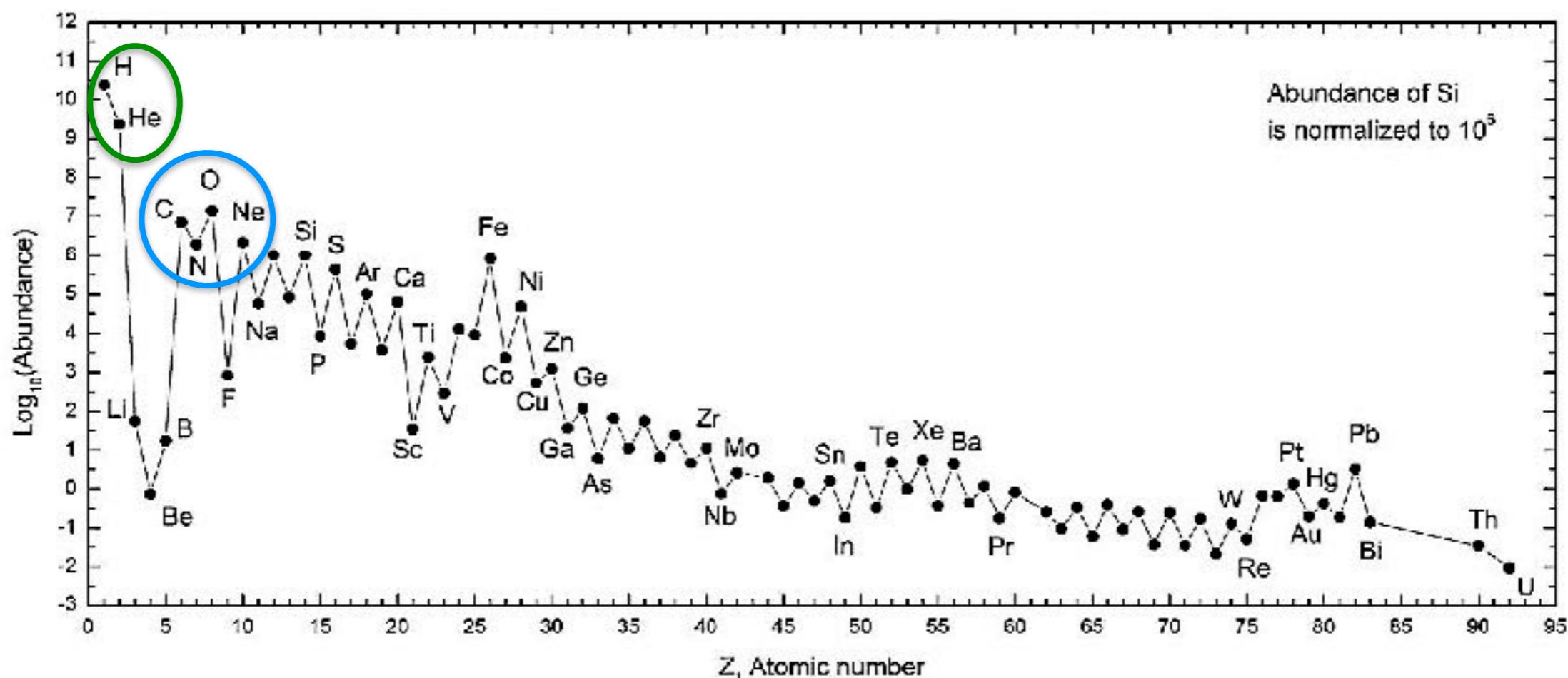
Physics motivations

- ✓ Abundance of the elements in the Universe

- in weight: H - 74%, He - 24%, O - 0.85%, C - 0.39%, ...

- ✓ Abundance of elements in the human body:

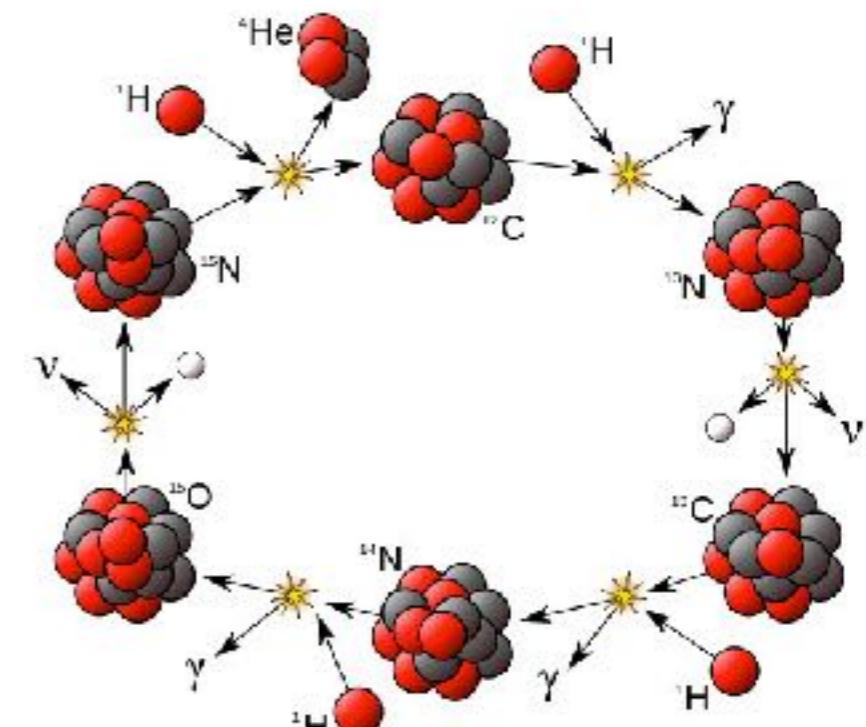
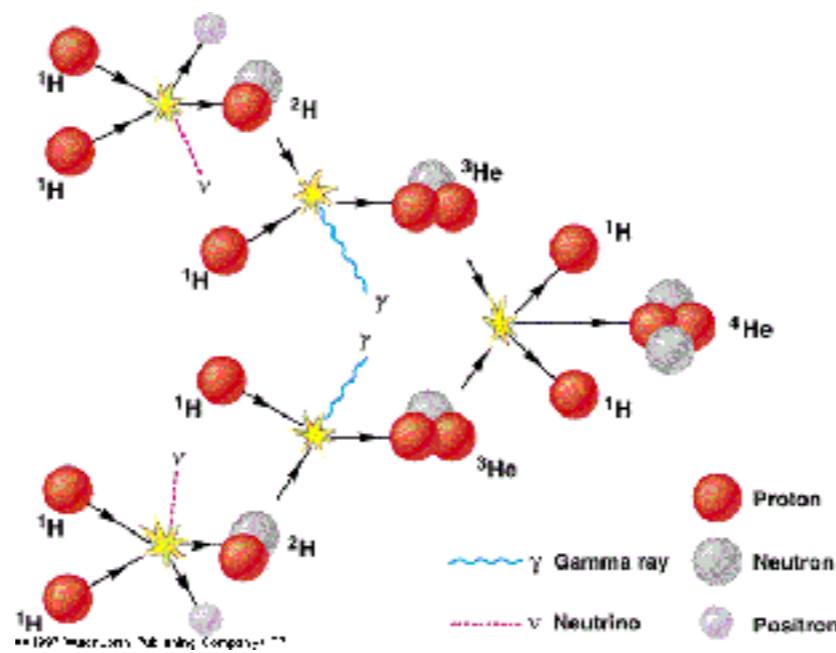
- in weight: O - 65%, C - 18%, H - 10%, N - 3%, other 4%



Physics motivations

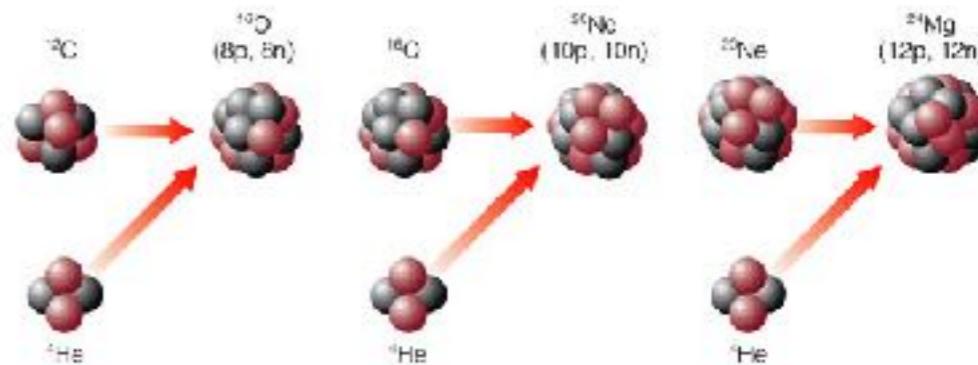
Synthesis of He in *H*-burning reactions

- pp-chain, CNO cycle, hot-CNO, NeNa cycle, MgAl cycle,...
- $4p \rightarrow {}^4He + 2e^+ + 2\nu$



Synthesis of C,O, Ne in He-burning

- $3\alpha \rightarrow {}^{12}C; {}^{12}C(\alpha, \gamma){}^{16}O, {}^{16}O(\alpha, \gamma){}^{20}Ne, {}^{20}Ne(\alpha, \gamma){}^{24}Mg$

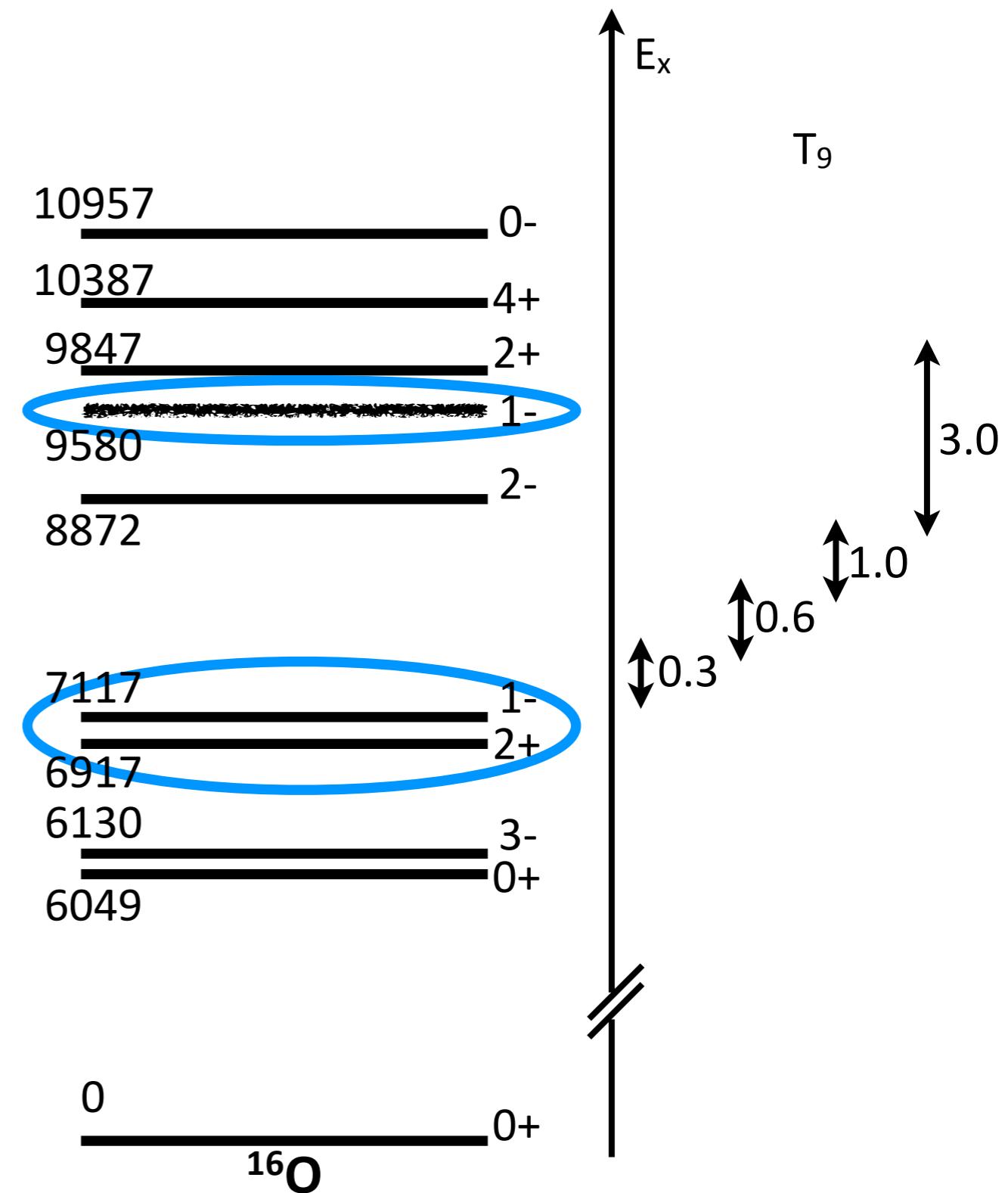
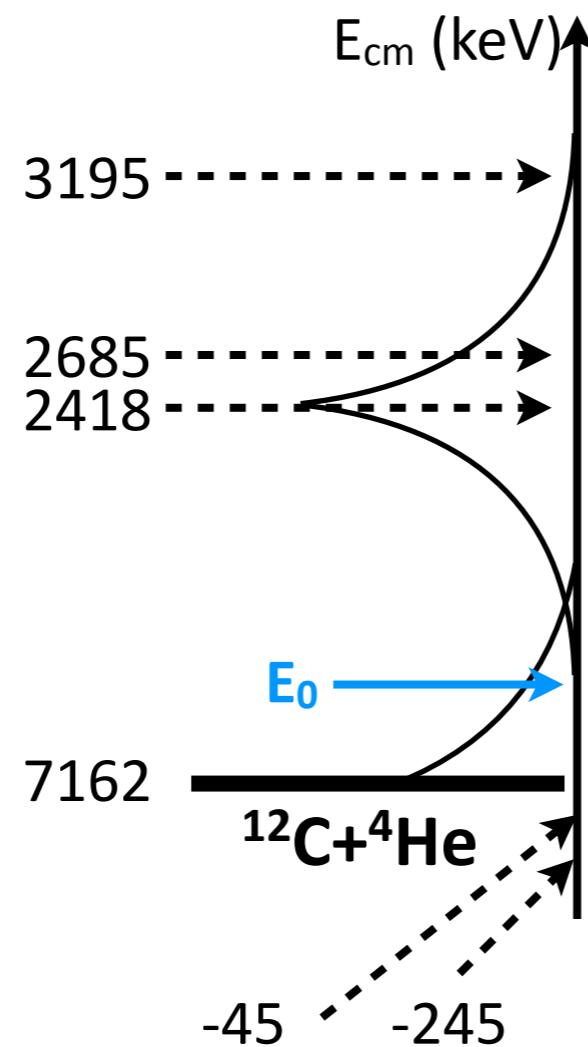


Cross-section measurement of (α, γ) and (p, γ) at astrophysical energies

Physics motivations: the flagship $^{12}\text{C} + \alpha \rightarrow ^{16}\text{O} + \gamma$

Survival of ^{12}C

- ✓ α -bruning in $^{12}\text{C} + \alpha \rightarrow ^{16}\text{O} + \gamma$
- ✓ nuclear structure properties of ^{16}O

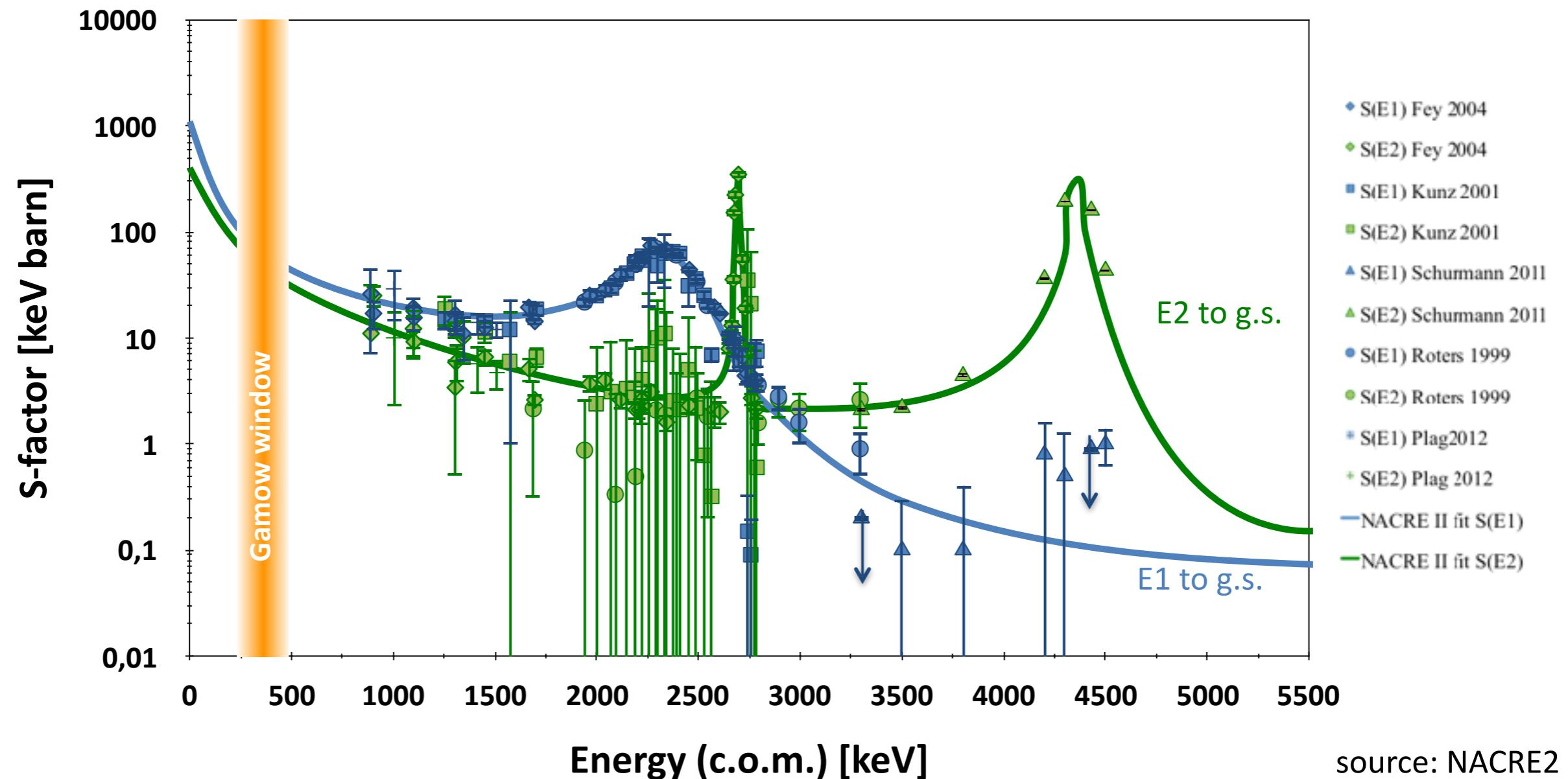


Two reaction mechanisms available

- non-resonant direct-capture process
- non-resonant type of capture into the tails of nearby resonances

The flagship $^{12}\text{C} + \alpha \rightarrow ^{16}\text{O} + \gamma$: status of the experimental knowledge

Extrapolated astrophysical S-factor for p-wave (E1) & d-wave (E2) $\alpha + ^{12}\text{C}$ capture
for the Gamow peak in red giants (300 MK):
40 – 80% uncertainty



Goal of the project

- ✓ Accurate measurements of (very small) cross sections of (α, γ) and (p, γ) nuclear reactions
- ✓ Particular effort is channeled towards the measurement of the $^{12}\text{C}(\alpha, \gamma)^{16}\text{O}$ reaction at low energies
- ✓ Estimated cross section for $^{12}\text{C}+\alpha$ at $E_{\text{cm}}=1.0$ MeV is $\sim 50\text{pbarn}$

Experimental challenges: how to improve the accuracy

✓ Strong and e.m. interactions invariant with respect to time reversal

✓ photo-disintegration vs capture reaction $B(b,\gamma)A \rightleftharpoons A(\gamma,b)B$

✓ principle of detailed balance in nuclear reactions:

$$\sigma_{b\gamma} \cdot g_{b\gamma} \cdot p_{b\gamma}^2 = \sigma_{\gamma b} \cdot g_{\gamma b} \cdot p_{\gamma b}^2$$

$$\sigma_{b\gamma} = \sigma_{\gamma b} \cdot \frac{g_{\gamma b}}{g_{b\gamma}} \cdot \frac{p_{\gamma b}^2}{p_{b\gamma}^2} = \sigma_{\gamma b} \cdot \frac{2J_{CN} + 1}{(2J_b + 1)(2J_B + 1)} \cdot \frac{E_\gamma^2}{E_{CM}} \cdot \frac{1}{\mu_{bB} c^2}$$

$g_{b\gamma}, g_{\gamma b}$ = spin factors

Experimental challenges: how to improve the accuracy

- ✓ direct capture vs photo-disintegration reaction:



at 1.0 MeV

factor 50!!

⇒ measure the cross section for the α -capture reaction
by means of the inverse photo-disintegration reaction

Experimental challenges: how to improve the accuracy

- ✓ Advantages of measuring the cross section for the p- and α -capture reactions by means of the inverse photo-disintegration reaction:
 - inherently low background measurements
 - different systematic uncertainty w.r.t. charged-particle induced reactions at low energies
 - target and its deterioration
 - (effective) beam energy definition
 - etc.
- ✓ **intense monochromatic and focussed γ -ray beams will be used!**

Monochromatic γ -ray beams @ELI-NP

✓ Production of monochromatic γ -ray beams: Gamma Beam System (GBS)

Compton Back Scattering of photons on ultra-relativistic electrons
(the most efficient frequency amplifier)

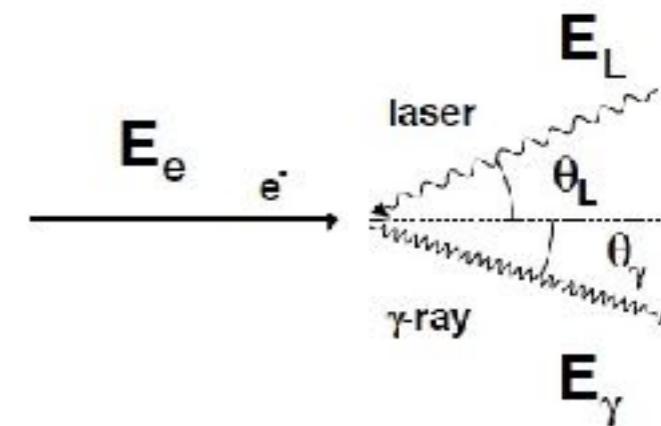
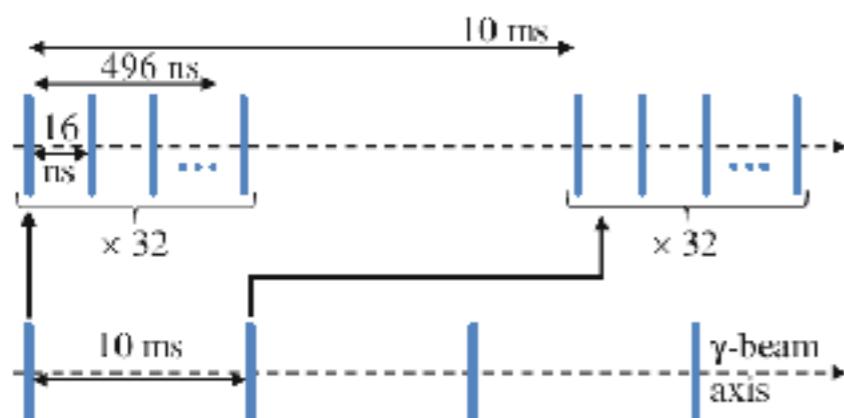
γ beams: high-brilliance, good bandwidth & collimation

intensity: $10^{13} \text{ } \gamma/\text{s}$ [$>10^4 \text{ } \gamma/(\text{s eV})$]

energy bandwidth: $\leq 0.5\%$!

energy range: $0.2 \text{ MeV} < E_\gamma < 19.5 \text{ MeV}$

linear polarization: >95%

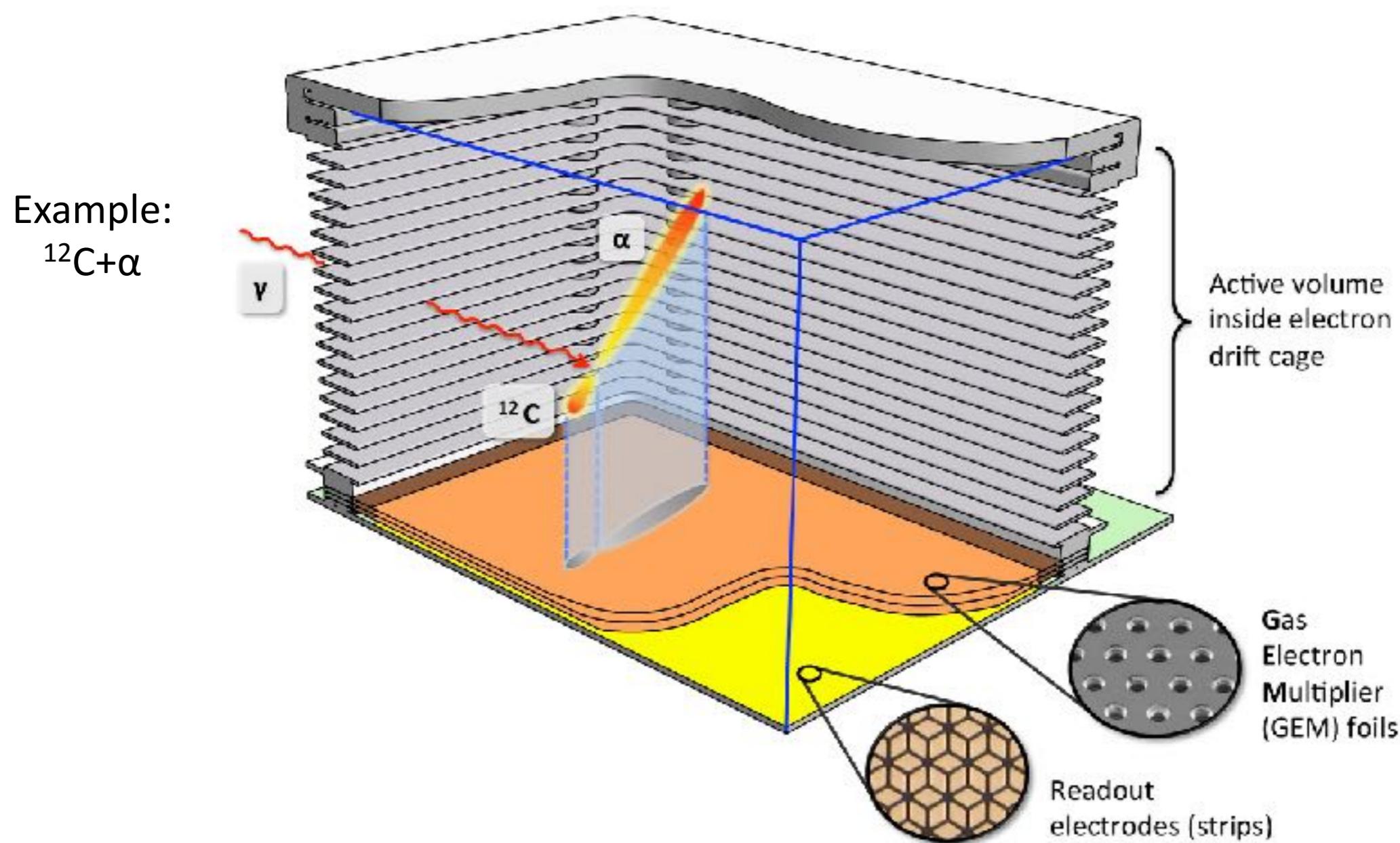


$$E_\gamma = 2\gamma_e^2 \cdot \frac{1 + \cos\theta_L}{1 + (\gamma_e \theta_\gamma)^2 + \frac{4\gamma_e E_L}{mc^2}} \cdot E_L \approx 4\gamma_e^2 E_L$$

$\theta_L \ll 1$

The ELITPC project

- ✓ An **active-target TPC** (ELITPC) to study reaction cross-sections of astrophysical interest where the reaction products are charged particles
 - full unambiguous reconstruction of multiple-particle events is possible



The ELITPC project: detector concept

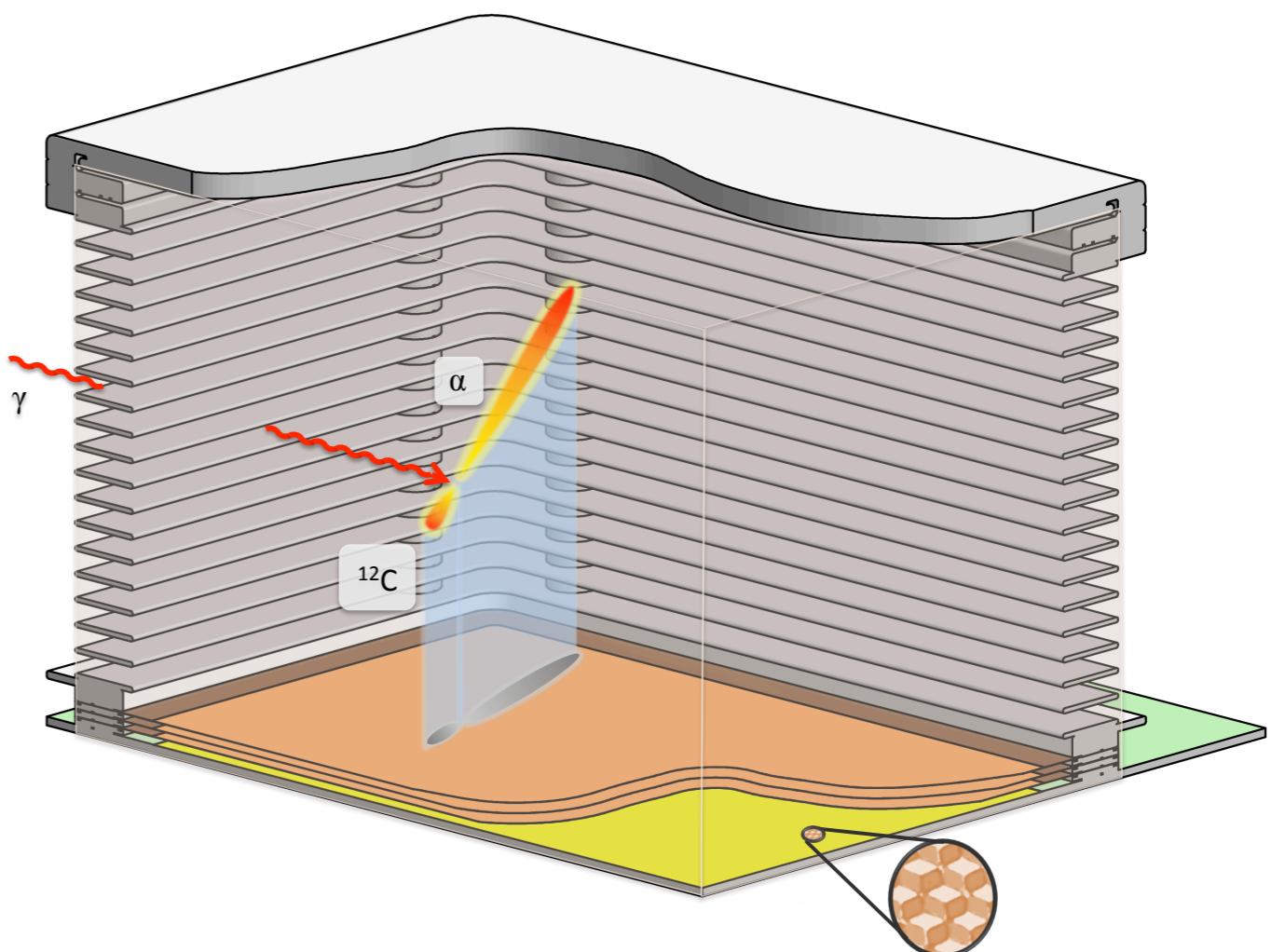
✓ Active target:

- active volume: 35 cm x 20 cm x 20 cm
- under-pressured (~ 100 mbar of CO₂): low-energy particles!
- gas-mixture and pressure tailored for the reaction of interest (containing CO₂ for measuring oxygen photo-disintegration)

✓ Charge-amplification:

- 3 or 4 GEM structures

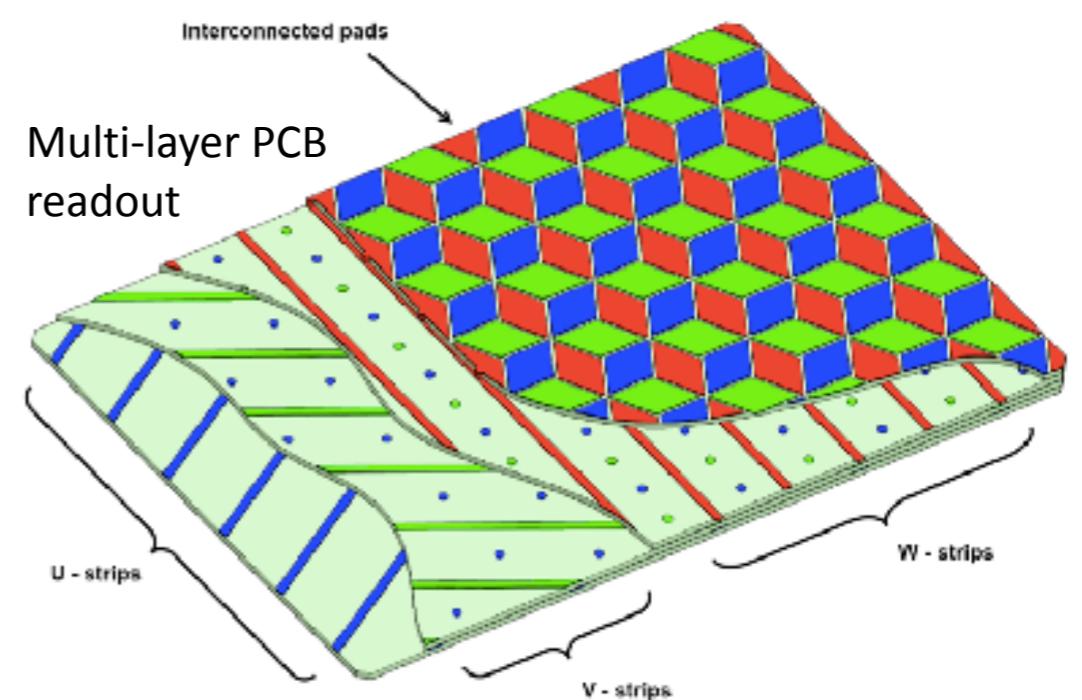
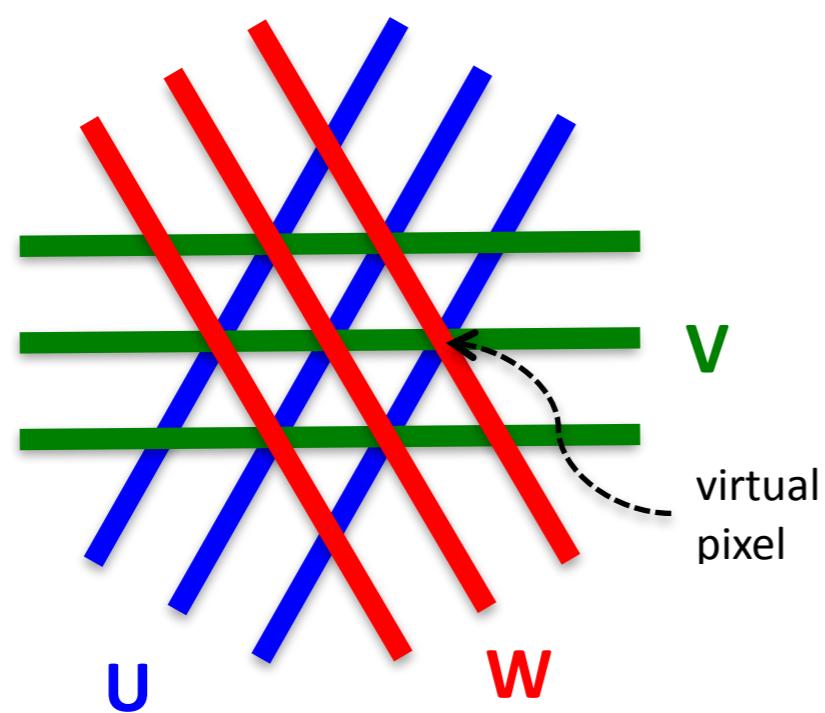
Example:



The ELITPC project: detector concept

✓ Read-out:

- 3-coordinate, planar, redundant electronic readout: 3 independent linear sets of strips (u-v-w): 1.5 mm pitch
- needs only ~1000 channels —> moderate cost of electronics
- u-v-w strip arrays for hit disambiguation in 2D —> virtual pixels
- z-coordinate from timing information
- aimed for relatively simple event topologies —> few tracks per event

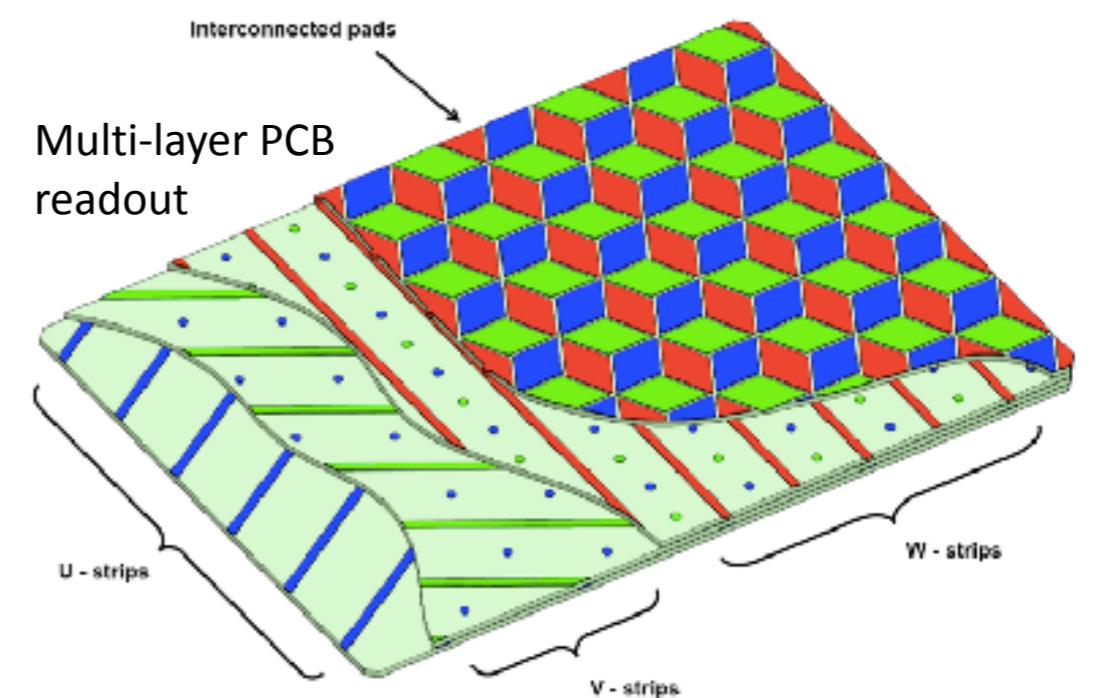
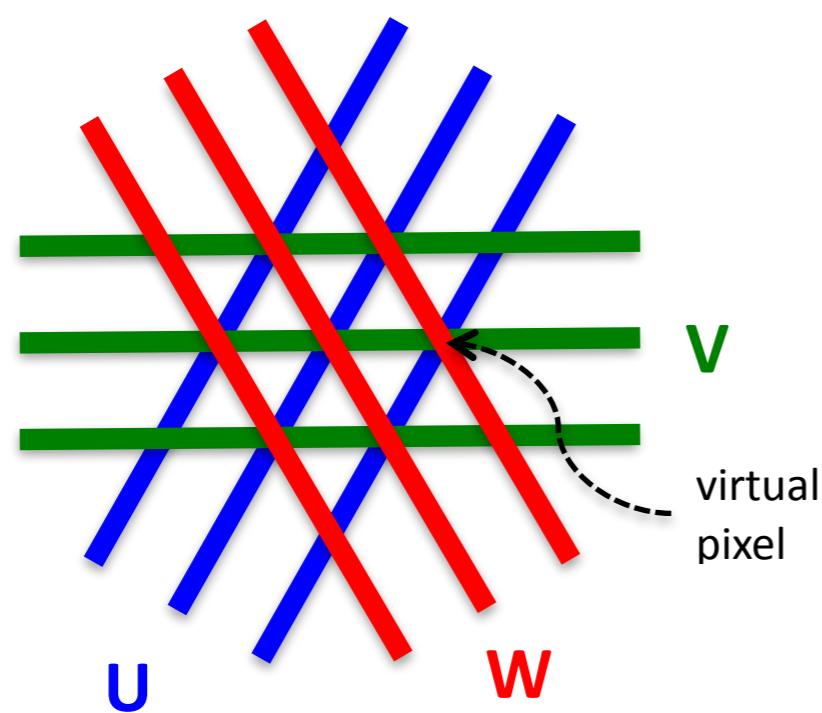


The ELITPC project: detector concept

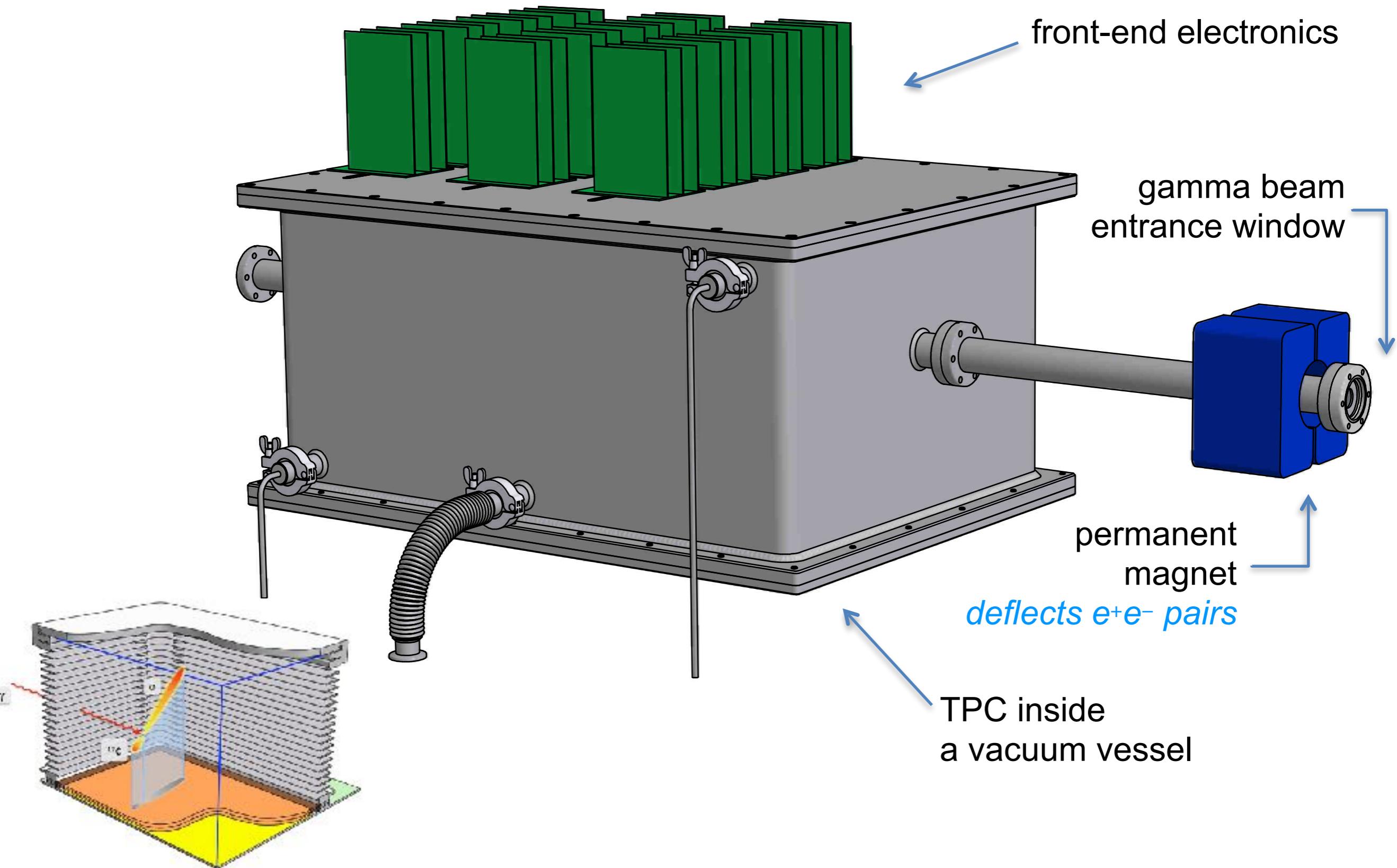
✓ Read-out:

- General Electronics for TPCs (GET) for signal amplification & digitization:
 - flexible sampling frequency: 1-100 MHz
 - adjustable gain & filtering per channel

✓ external trigger from the time-structure of the γ beam (100 Hz)



The ELITPC project: detector concept

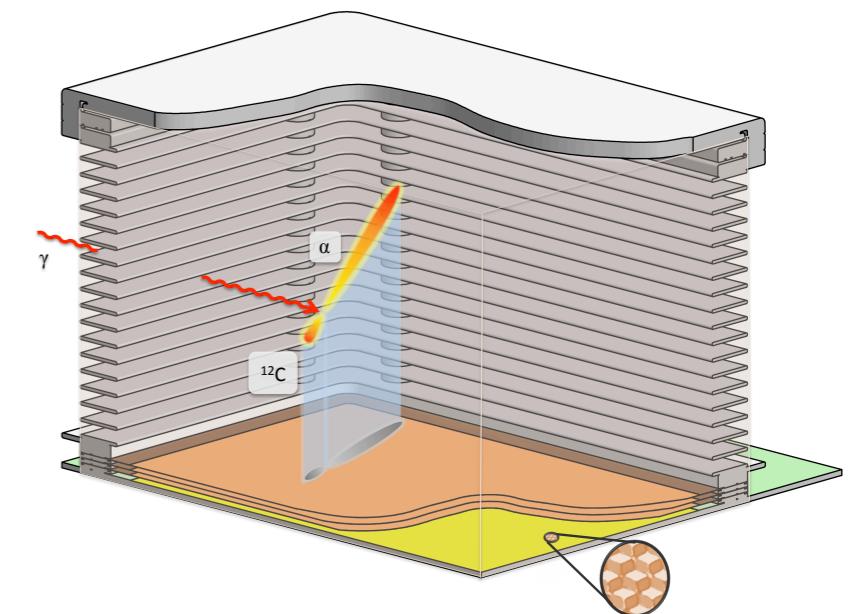


The ELITPC project: expected yields for the flagship $^{16}\text{O}(\gamma,\alpha)^{12}\text{C}$

✓ Time-reverse reaction $\rightarrow ^{16}\text{O}(\gamma,\alpha)^{12}\text{C}$

- measure energy & angular distributions of charged particles
- obtain accurate values of E2/E1 components
- reduce uncertainty from 40-80% to 10%

✓ Target CO_2 @ 100 mbar



E_γ MeV	E_{cm} MeV	# events/week
8.26	1.1	~500
8.36	1.2	~900
8.46	1.3	~1900
8.56	1.4	~3000

beam intensity on target:
 $2.5 \times 10^4 \text{ } \gamma/\text{s/eV}, 0.5\% \text{ bandwidth} \rightarrow 10^9 \text{ } \gamma/\text{s}$
(D. Filipescu et al., Eur. Phys. J. A (2015) 51: 185)

$Q=7.162 \text{ MeV for } ^{12}\text{C}(\alpha,\gamma)^{16}\text{O}$

ELITPC: a versatile detector

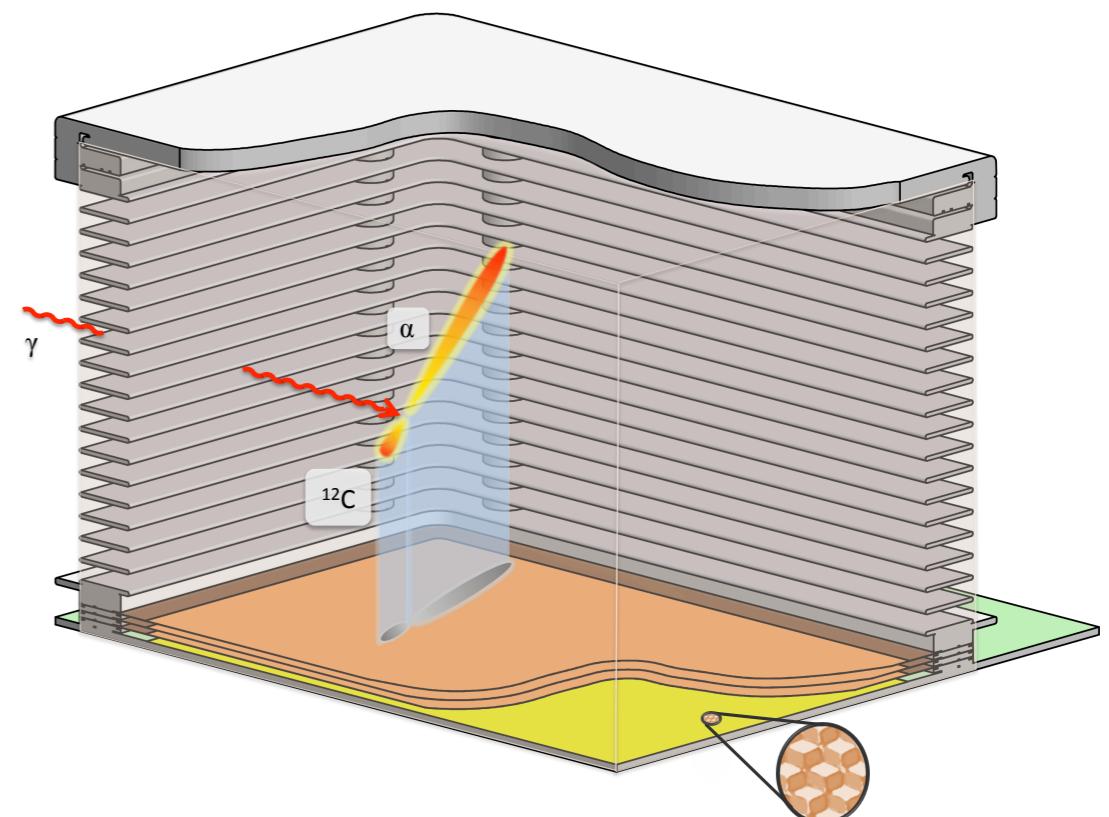
- ✓ **Day1 experiments:** measure decay products of nuclear photo-dissociation reactions

Reaction	Target	Astrophysical relevance
$^{16}\text{O}(\gamma,\alpha)^{12}\text{C}$	CO_2	ratio C/O
$^{21}\text{Ne}(\gamma,\alpha)^{17}\text{O}$	^{21}Ne	role of ^{16}O as neutron poison
$^{22}\text{Ne}(\gamma,\alpha)^{18}\text{O}$	^{22}Ne	ratio $^{16}\text{O}/^{18}\text{O}$, CNO-cycle synthesis of ^{22}Ne

The ELITPC project: background

✓ Background:

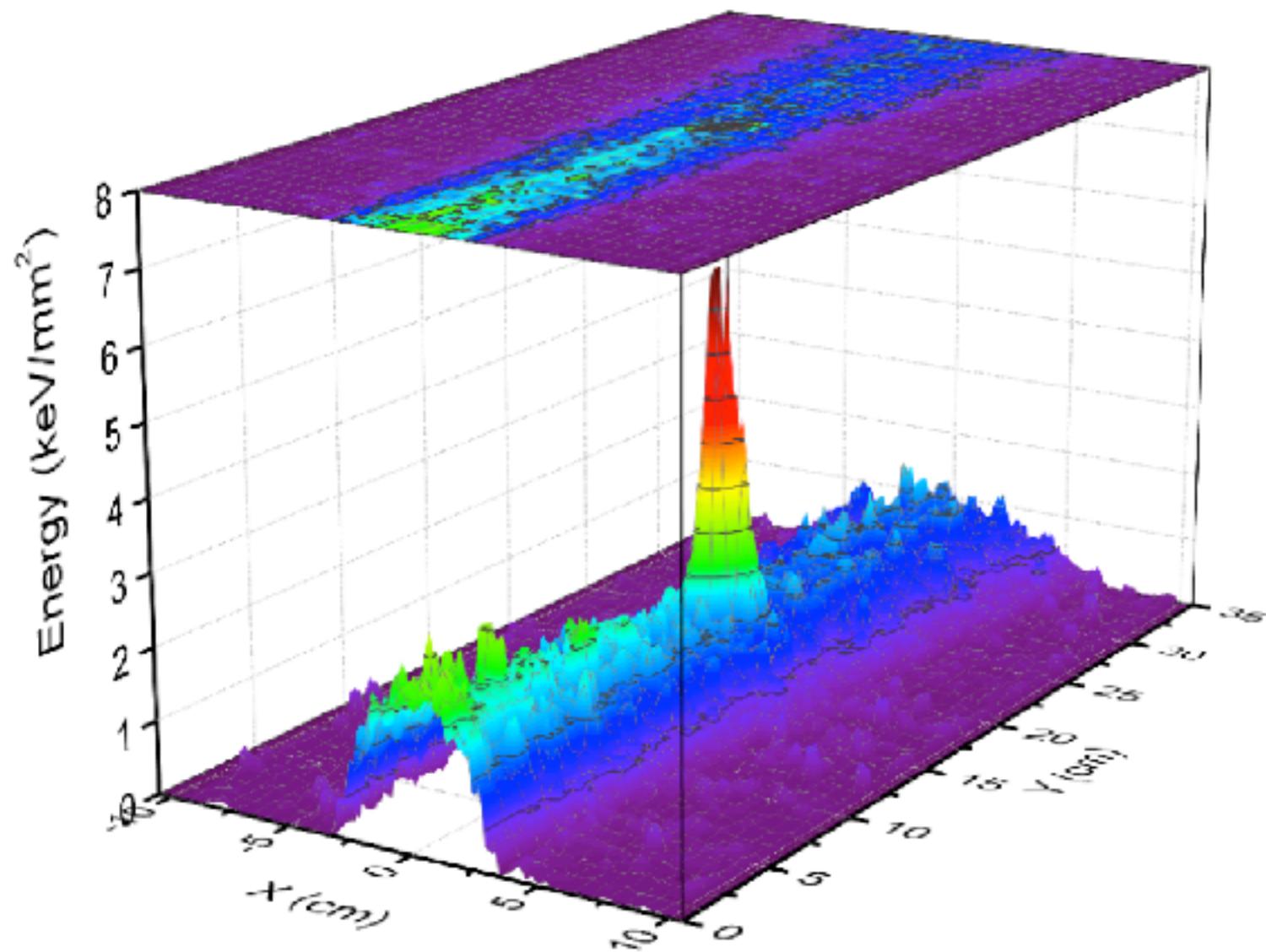
- main sources of background:
 - Compton electrons
 - e^+e^- pairs in gaseous target & in thin mylar/kapton entrance window
 - at higher γ energies isotopic contaminants
- very small w.r.t. direct (α,γ) reaction experiments



The ELITPC project: background

✓ Background: GEANT4 simulation of a single ELI-NP macro-pulse:

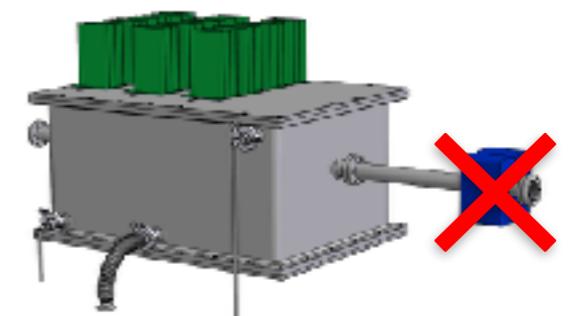
- $10^7 \gamma$ -rays at 8 MeV
- CO₂ @ 100 mbar
- 0.5 MeV α -particle track added artificially to mimic ¹⁶O photo-dissociation



Colour: integrated energy
deposits on the 2D readout plane

The most unfavourable scenario is shown:

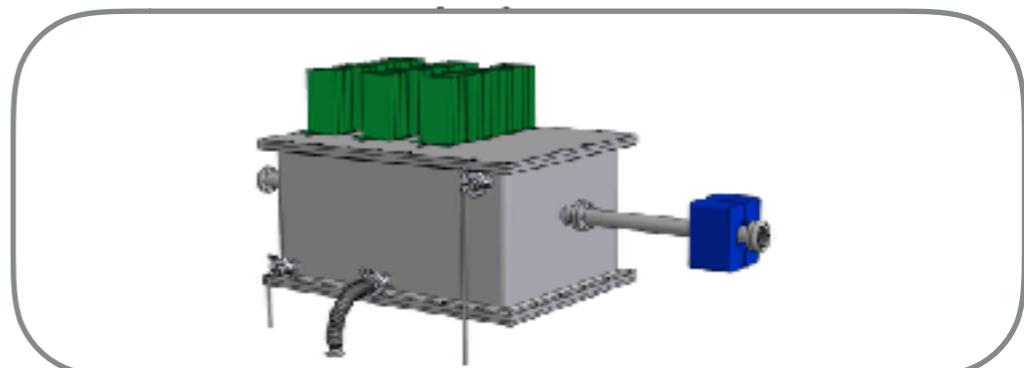
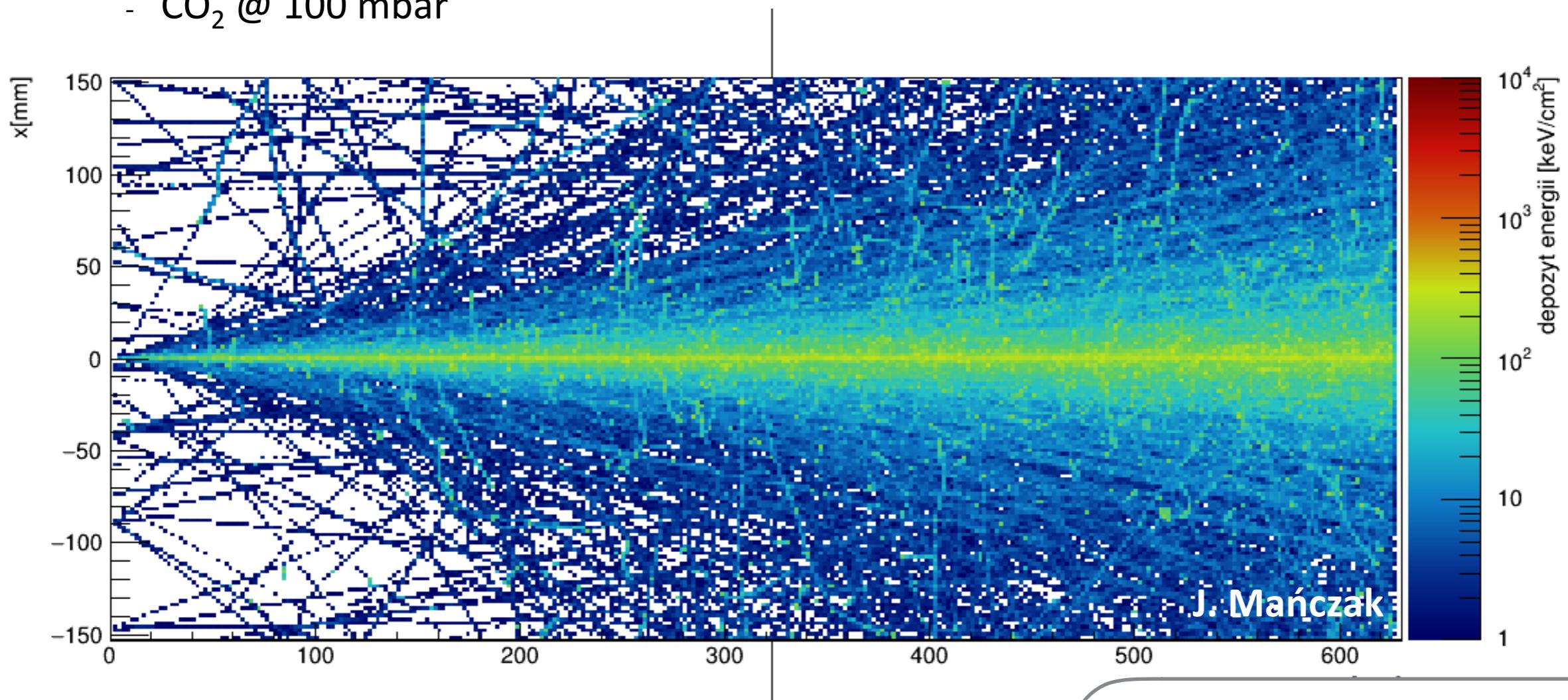
- entrance window very close to the active volume
- no permanent magnet after the entrance window



The ELITPC project: background

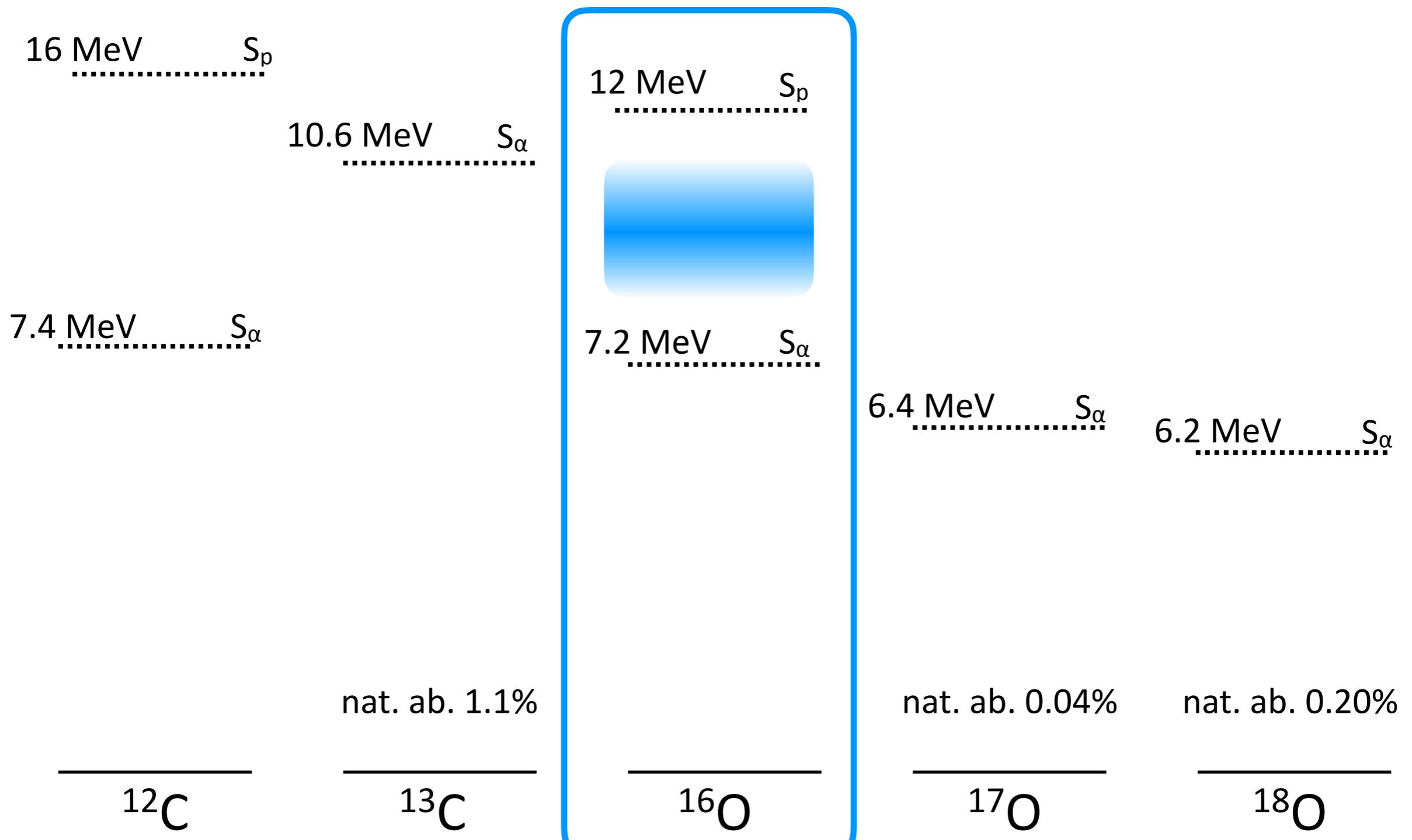
✓ Background: GEANT4 simulation of a single ELI-NP macro-pulse:

- $10^7 \gamma$ -rays at 10 MeV
- CO₂ @ 100 mbar



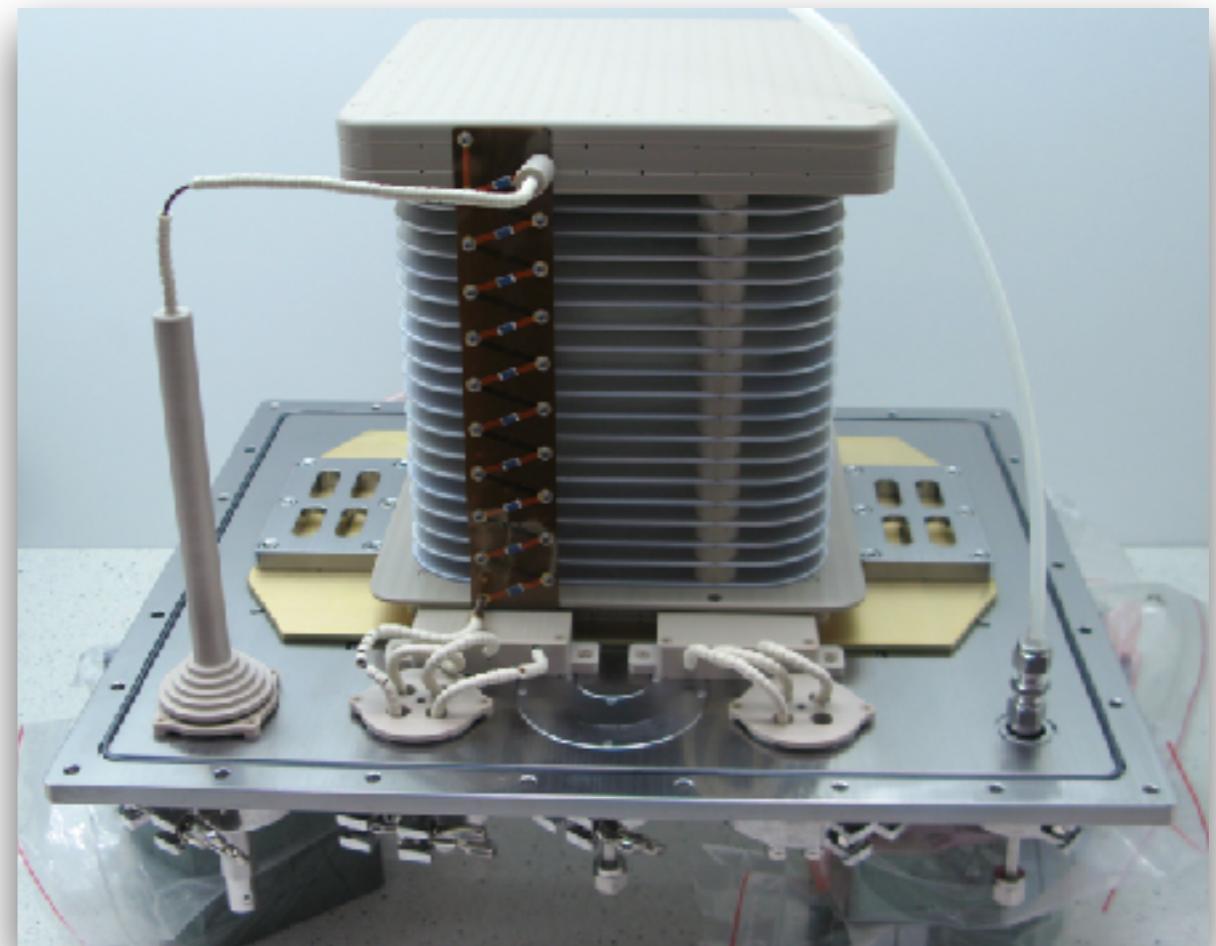
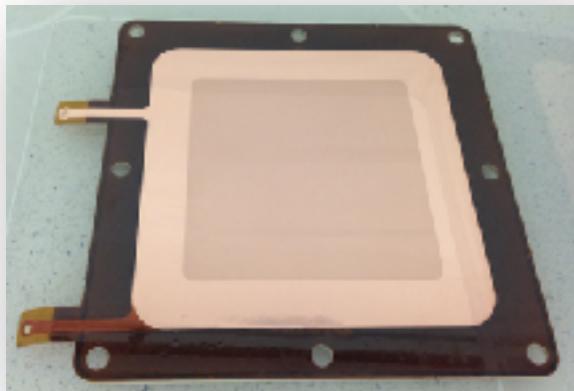
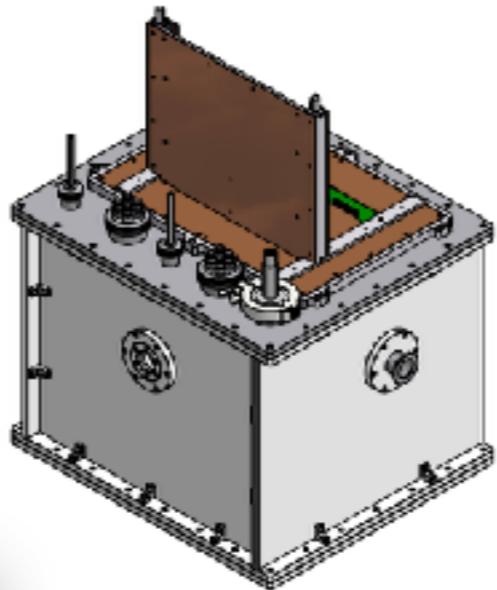
The ELITPC project: isotopic contaminants background for $^{16}\text{O}(\gamma,\alpha)^{12}\text{C}$

- ✓ Background: **isotopic contaminants** in the active gas



Demonstrator(s)

- ✓ Readout area: $10 \times 10 \text{ cm}^2$, drift: 20 cm
- ✓ GET electronics: 256 channels
- ✓ Operating at atmospheric pressure

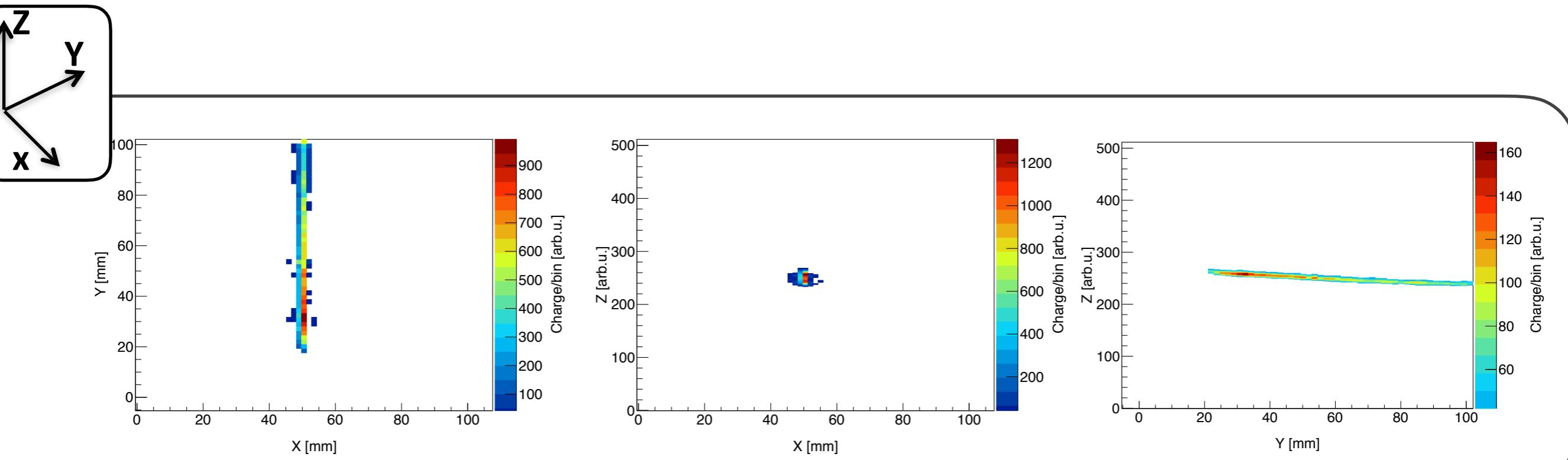
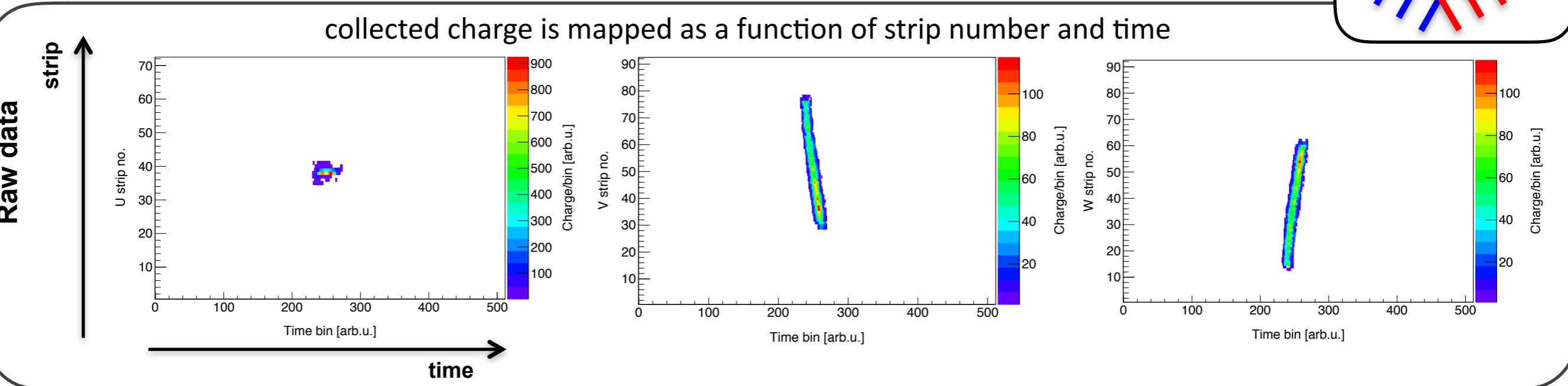
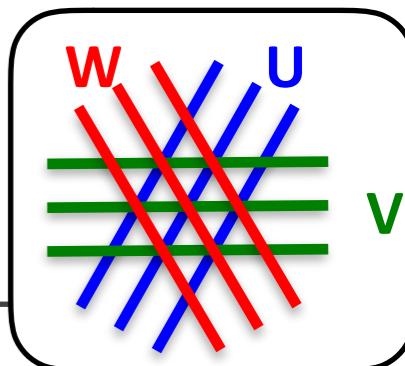


- ✓ Tested @ 9 MV Tandem (IFIN-HH, Romania) with 15 MeV α -particle beam (April 2016)
- ✓ Gas mixture: He+CO₂ (70:30) @1 atm

Demonstrator(s)

✓ Example 1: Single tracks from α -particle source:

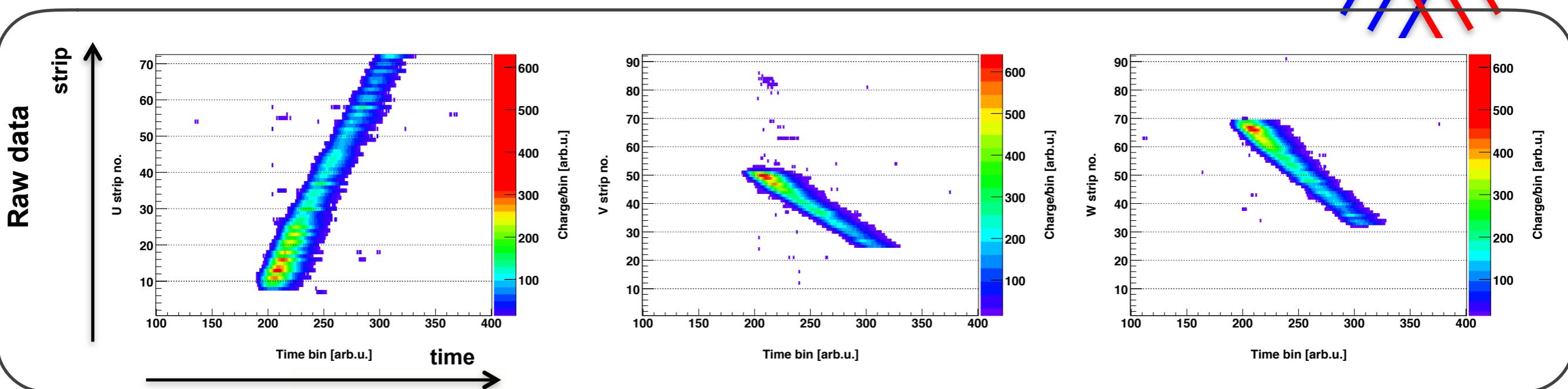
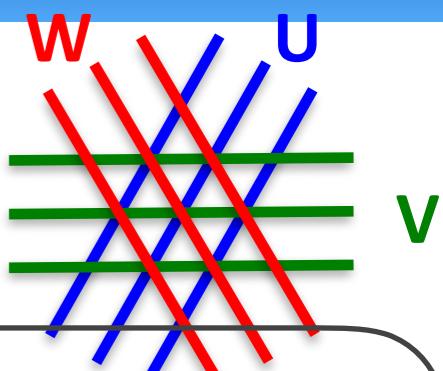
- Gas mixture: He+CO₂ (90:10) @1 atm



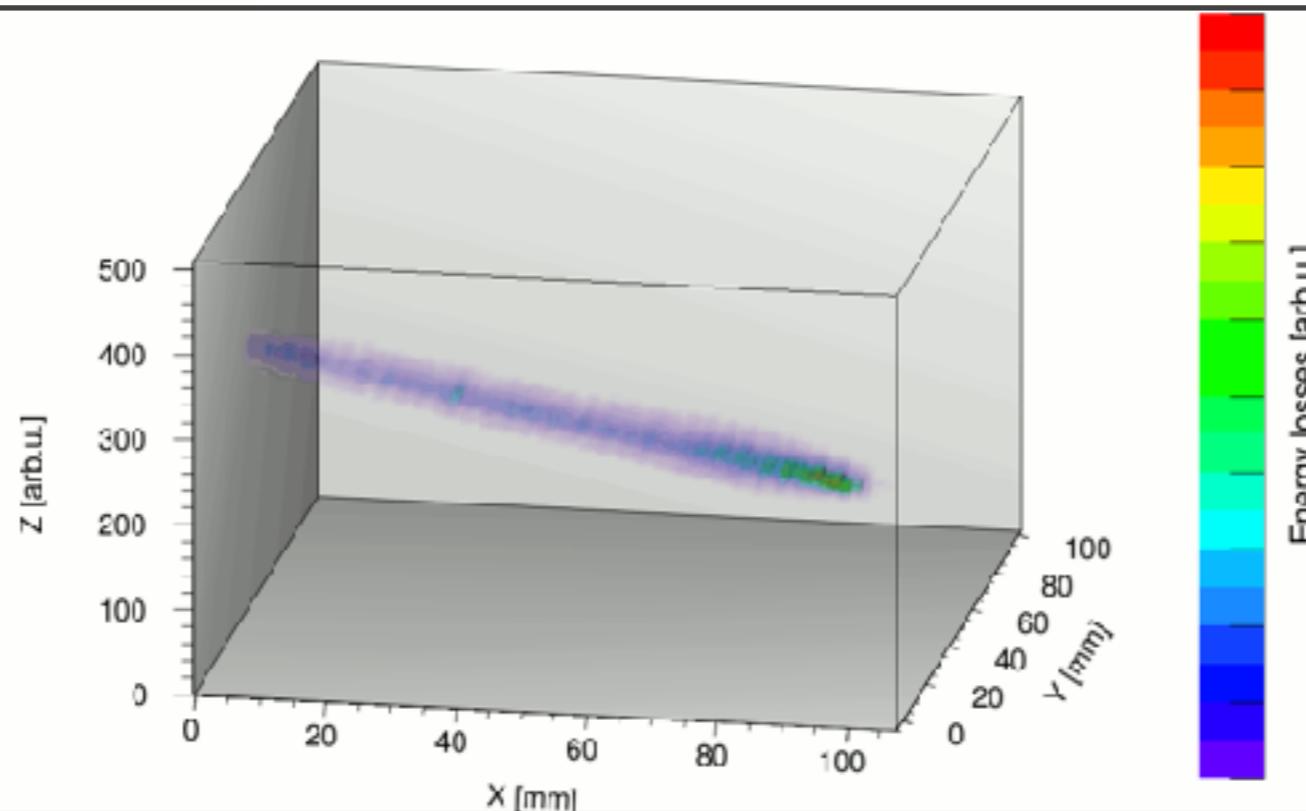
Demonstrator(s)

- ✓ Example 2: Single tracks from α -particle beam:

- Gas mixture: He+CO₂ (70:30) @1 atm



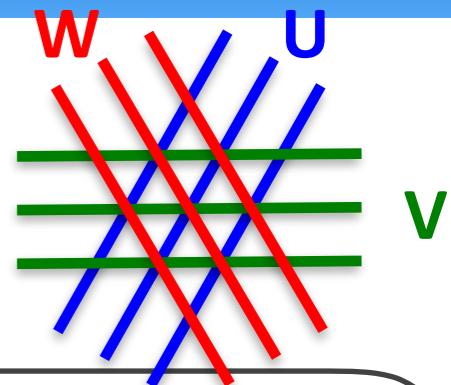
Reconstructed α -particle track in 3D



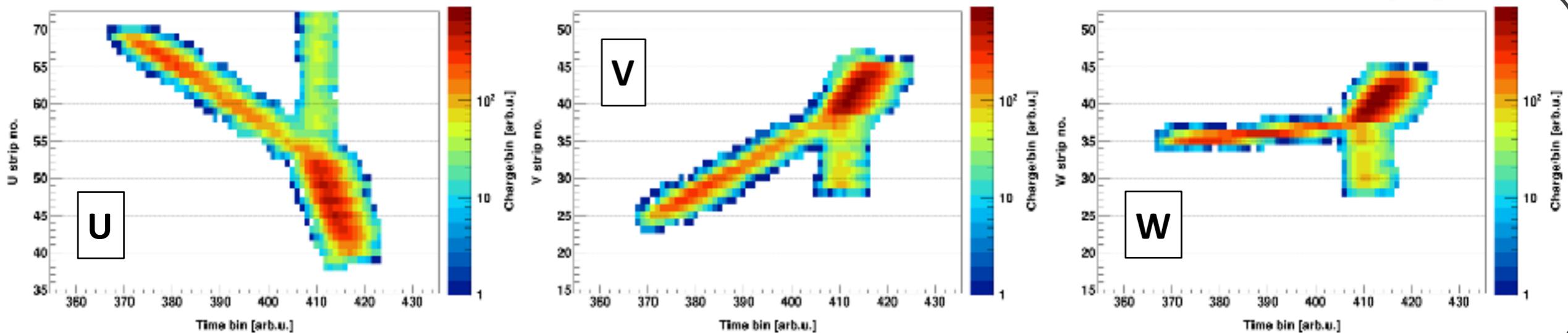
Demonstrator(s)

✓ Example 3: events with 3 tracks → ${}^4\text{He} + {}^{16}\text{O}$ scattering

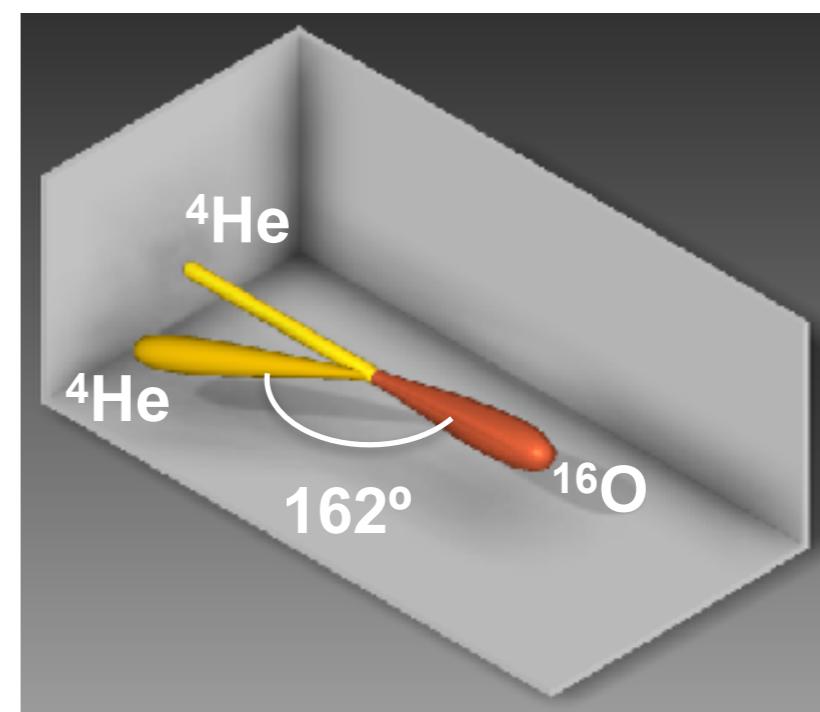
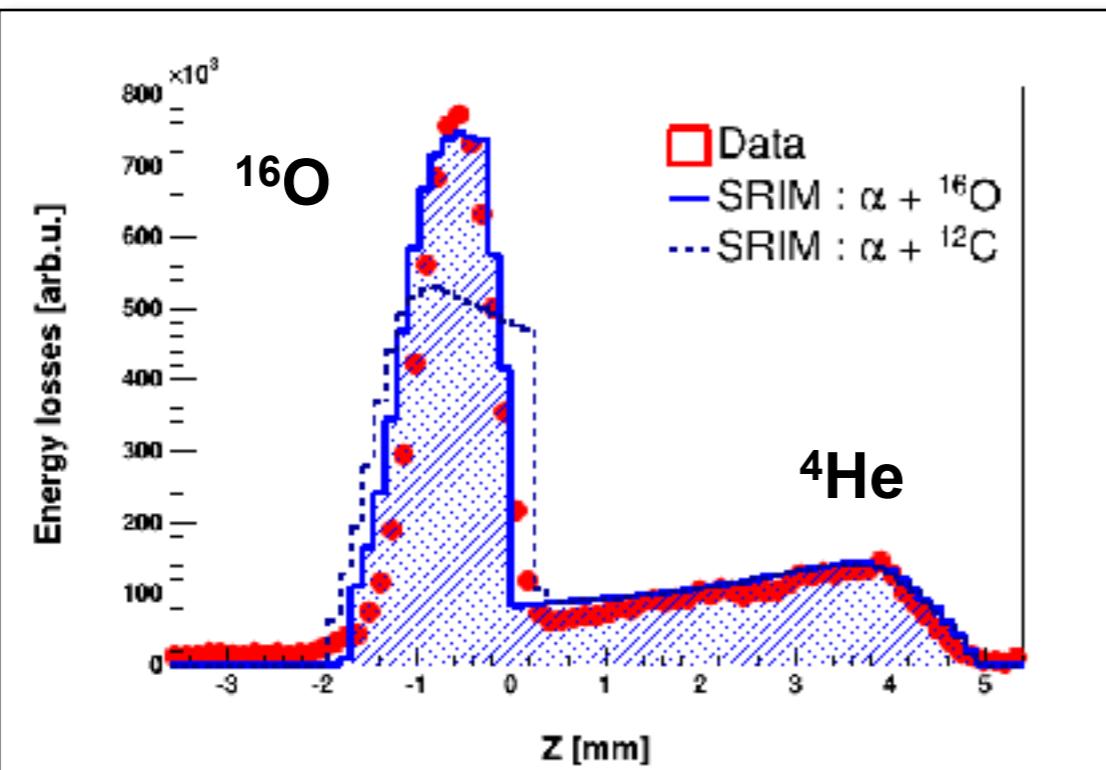
- Gas mixture: He+CO₂ (70:30) @1 atm



Raw data

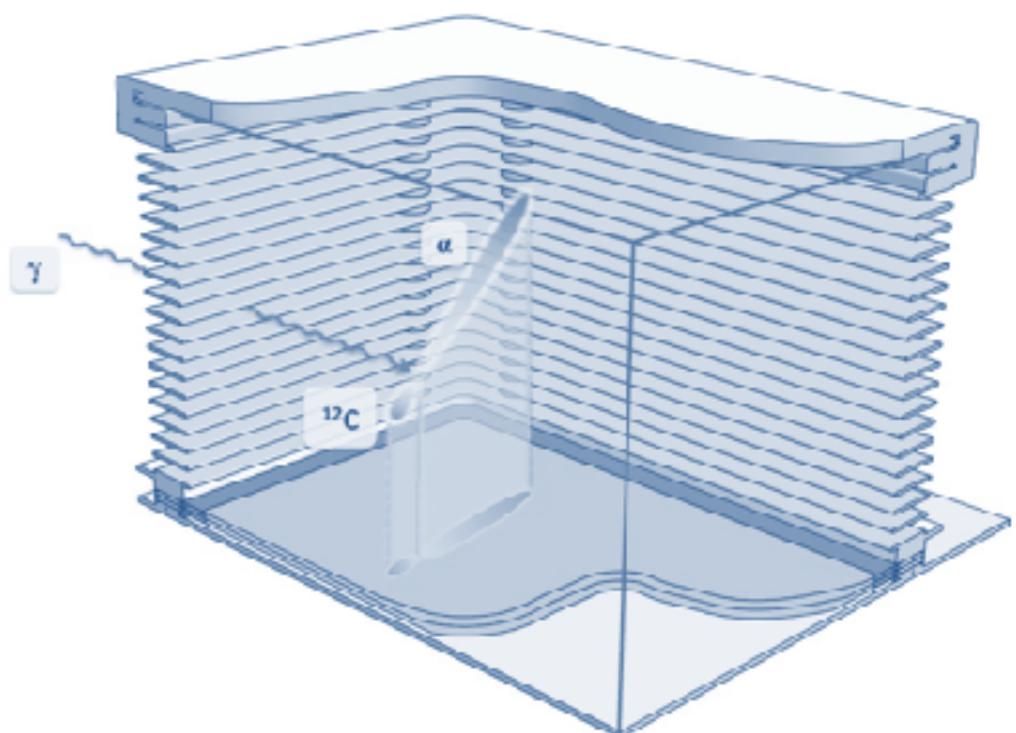


Reconstruction



Summary

- ✓ $^{12}C + \alpha \rightarrow ^{16}O + \gamma$ reaction cross section measurement is **flagship experiment** for the ELITPC collaboration at ELI-NP
- ✓ ELITPC detector among **DAY-1** experiments for ELI-NP
- ✓ 256 chn demonstrator detector is operational:
 - readout development
 - equipped GET electronics
 - reconstruction algorithms tests
- ✓ Other physics cases with ELITPC @ ELI-NP:
 - other astrophysical reactions (different gas targets)
 - nuclear structure physics (clustering phenomena)



Outlook

- ✓ R&D continues

- tests at lower-than-atmospheric pressure
- electronics developments
- simulations

- ✓ First commissioning experiments in 2018

