Nuclear Physics in Astrophysics VIII



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Study of nuclear physics input-parameters via high-resolution γ -ray spectroscopy

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\textwidth=6.3in
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 \oddsidemargin +0.35in
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{\small \it Nuclear Physics in Astrophysics 8, NPA8: 18-23 June 2017, Catania, Italy}
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\begin{center}
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 %%% Title goes here.
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 \TITLE{Study of nuclear physics input-parameters via high-resolution \gamma-ray spectroscopy}\\[3mm]
 %%% Authors and affiliations are next. The presenter should be
 %%% underlined as shown below.
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\AUTHORS{P. Scholz<sup>1</sup>, F. Heim<sup>1</sup>, J. Mayer<sup>1</sup>, M. Spieker, and A. Zilges<sup>1</sup>}
{\boldsymbol \lambda \in \mathbb{N} \ it}
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% Enter contact e-mail address here.
\centerline{Contact email: {\it pscholz@ikp.uni-koeln.de}}
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Nuclear reaction cross sections are one of the main ingredients for the understanding of nucleosynthesis processes in stellar environments. For isotopes heavier than those in the iron-peak region, reaction rates are often calculated using the Hauser-Feshbach statistical model. The accuracy of the predicted cross sections strongly depend on the uncertainties of the nuclear-physics input-parameters. These are nuclear-level densities, γ -strength functions, and particle+nucleus optical-model potentials.

The precise measurement of total and partial reaction cross sections at sub-Coulomb energies and their comparison to statistical model calculations are used to constrain or exclude different nuclear-physics models.

This talk is going to introduce experimental methods and present recent experiments performed at the Cologne 10 MV FN-Tandem accelerator and the high-efficiency HORUS γ -ray spectrometer. Results for cross-section measurements of α induced reactions on the p nucleus 108 Cd [1] and the 85 Rb(p, γ) reaction will be presented. In addition, preliminary results of γ -strength function studies applying the method of two-step cascades [2] for the reactions 92 Mo(p, $\gamma\gamma$) and 63 Cu(p, $\gamma\gamma$) will be shown.\newline

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[1] P. Scholz \textit{et al.}, Phys. Lett. B \textbf{761} (2016) 247.
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\noindent
[2] F. Be\u{c}v\'{a}\u{r} \textit{et al.}, Phys. Rev. C \textbf{46} (1992) 1276.
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