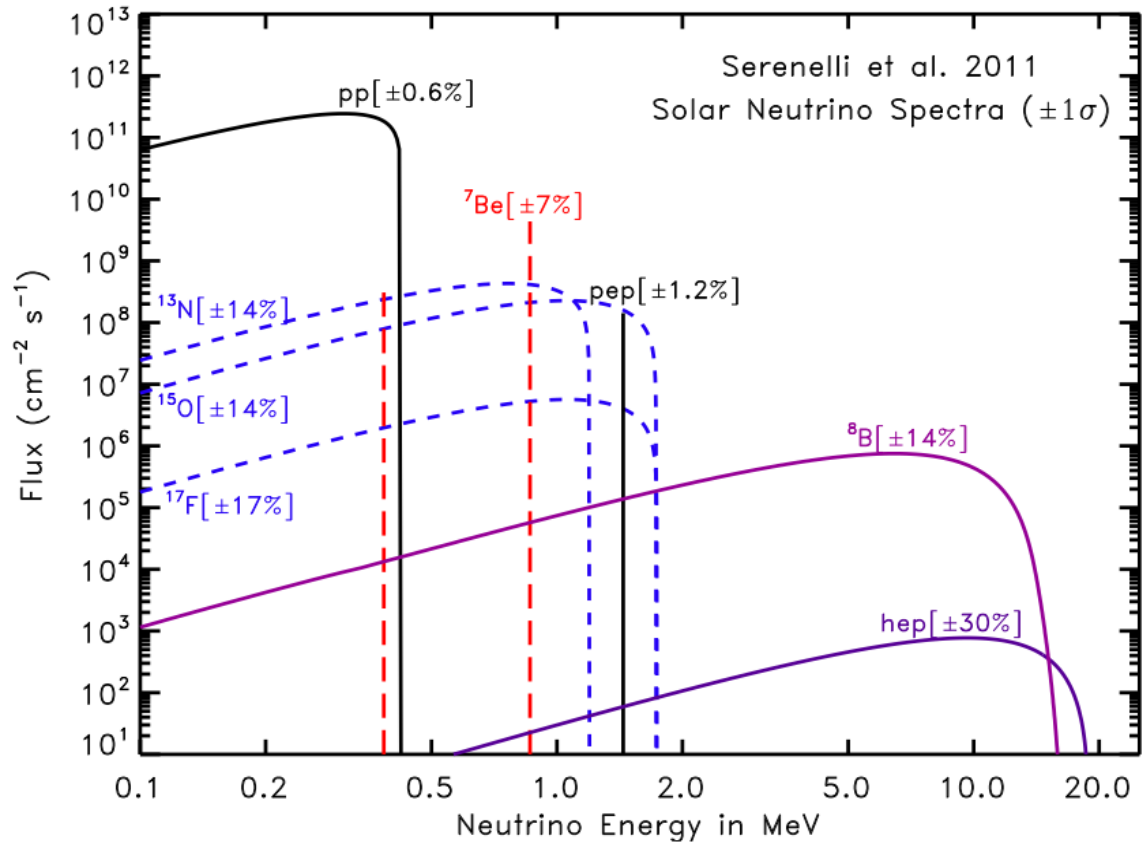


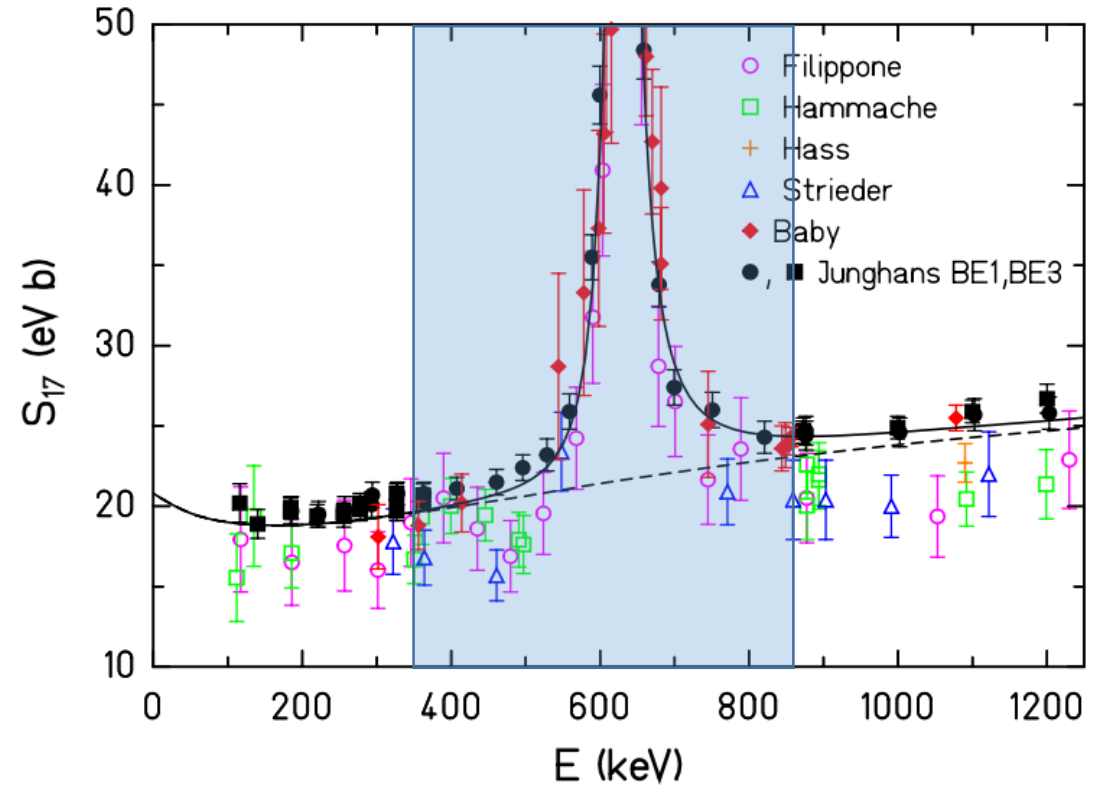
# Absolute measurement of the ${}^7\text{Be}(p,\gamma){}^8\text{B}$ cross section with the recoil separator ERNA

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Università degli Studi della Campania Luigi Vanvitelli  
INFN Naples

# Why ${}^7\text{Be}(p,\gamma){}^8\text{B}$ ?



*Large uncertainty on the predicted solar neutrino flux.*

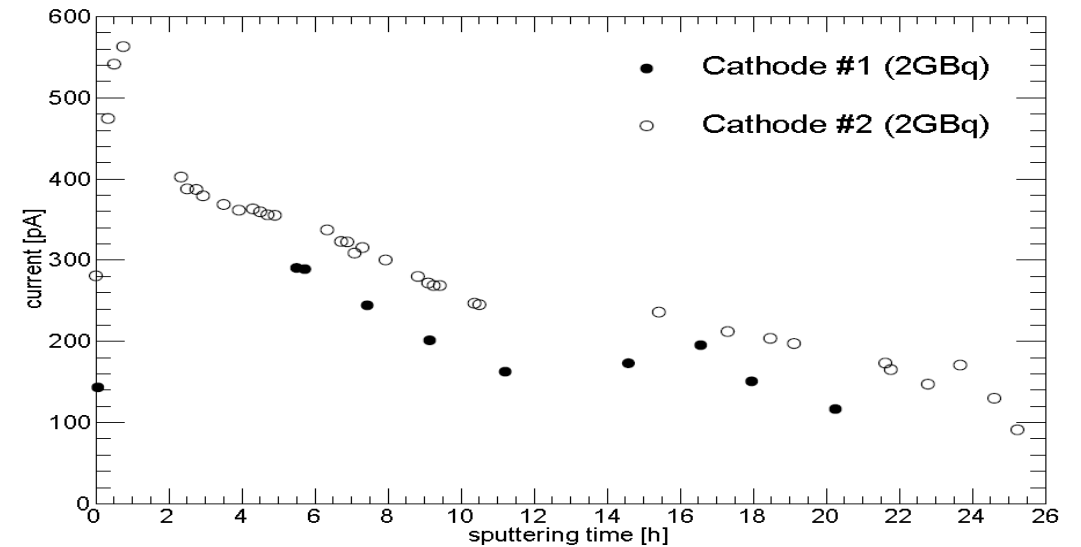
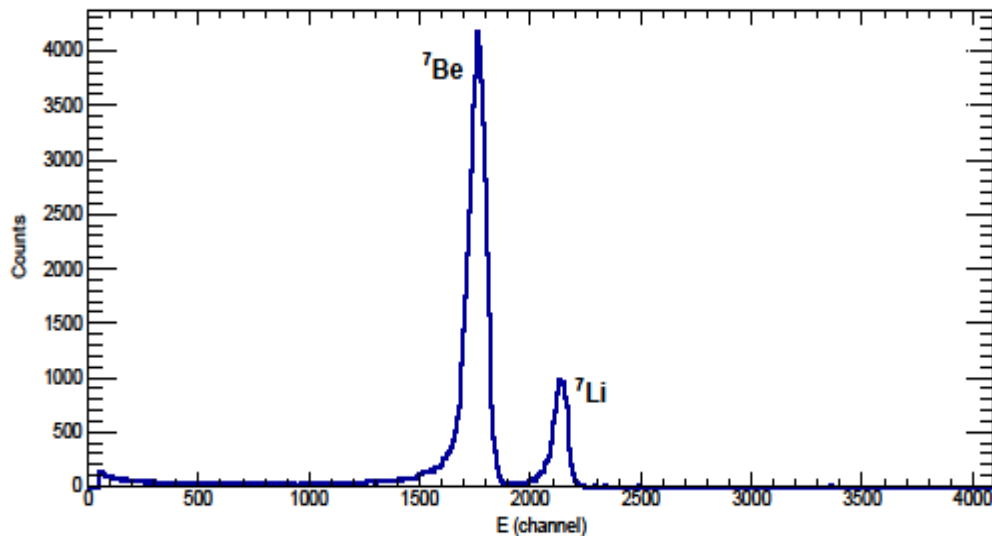
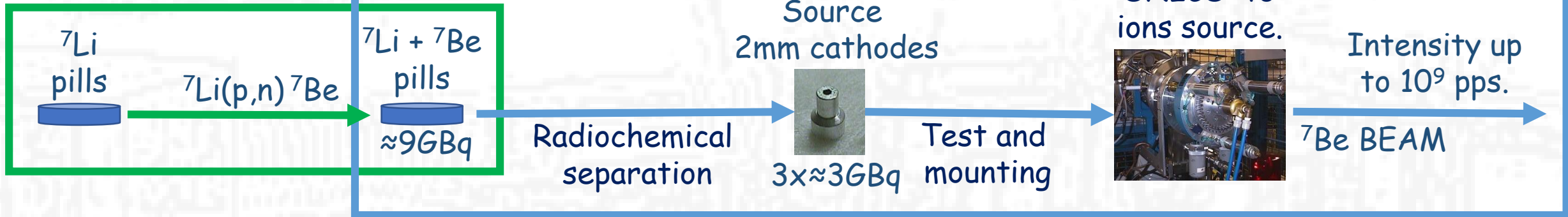


Adelberger et al. 2011

*Discrepancies between existing data sets limits the precision of the extrapolation to solar energy.*

# $^7\text{Be}$ beam

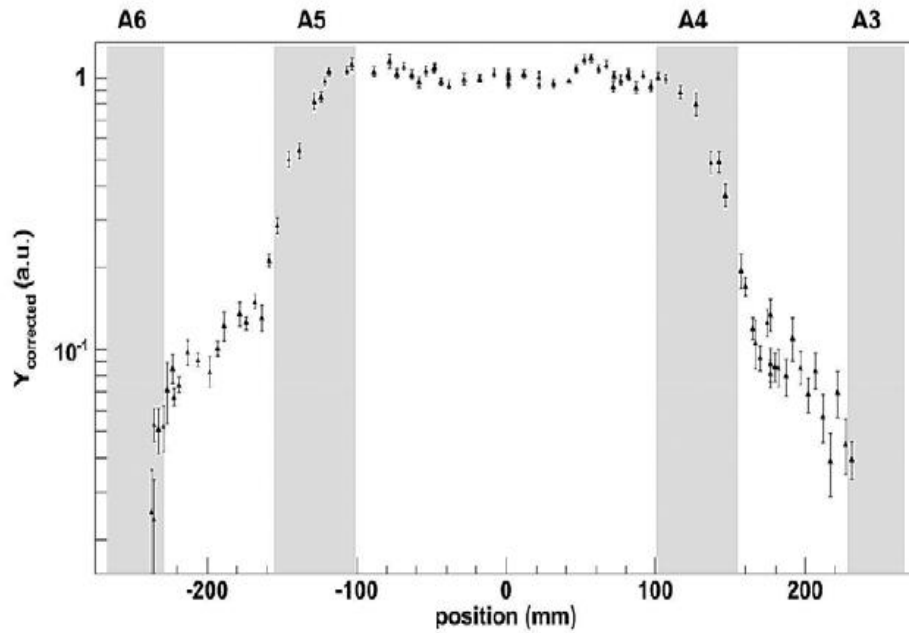
ATOMKI  
Cyclotron Laboratory.



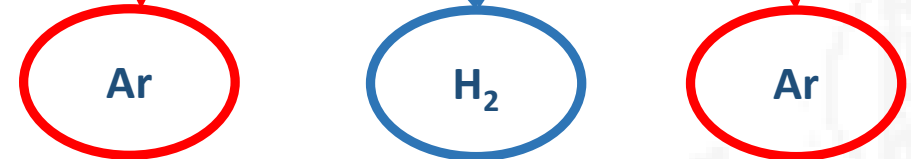
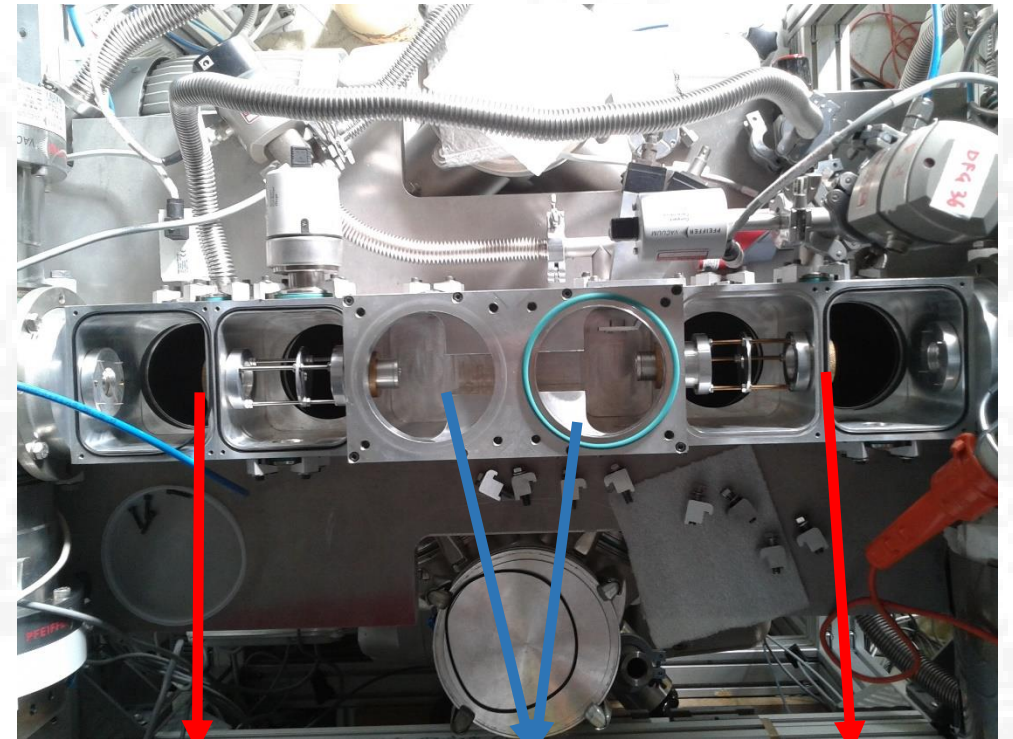
The number of incident projectiles, including lithium contamination, is monitored on line through elastic scattering.

# Windowless gas target

Eur. Phys. J. A (2013) 49: 80



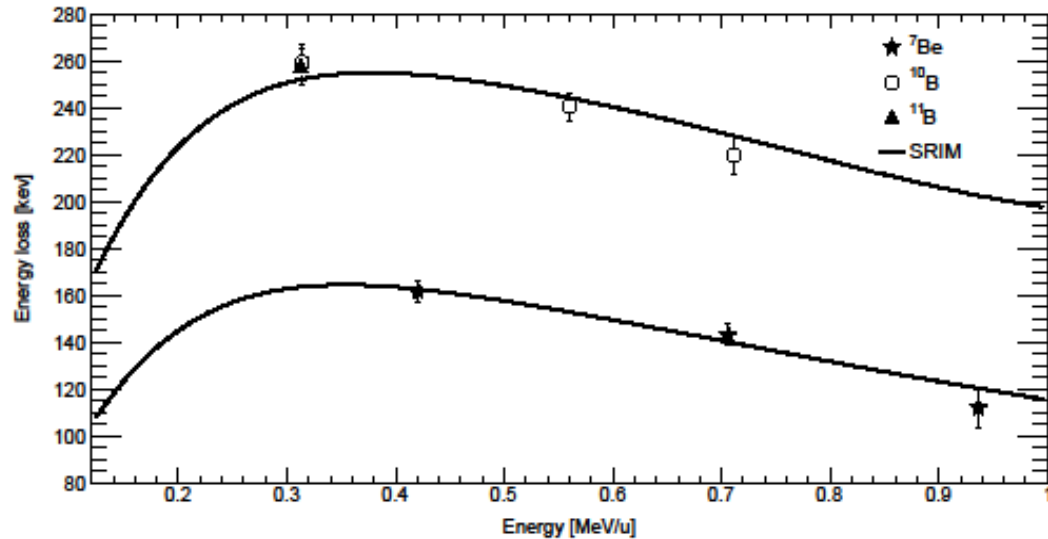
Density profile of the gas target as seen in the yield of the 478 keV  $\gamma$ -ray line from the  ${}^7\text{Li}(p, p){}^7\text{Li}$



Target density  $n = 7.22 \pm 0.15 \cdot 10^{18} \text{ at/cm}^2$

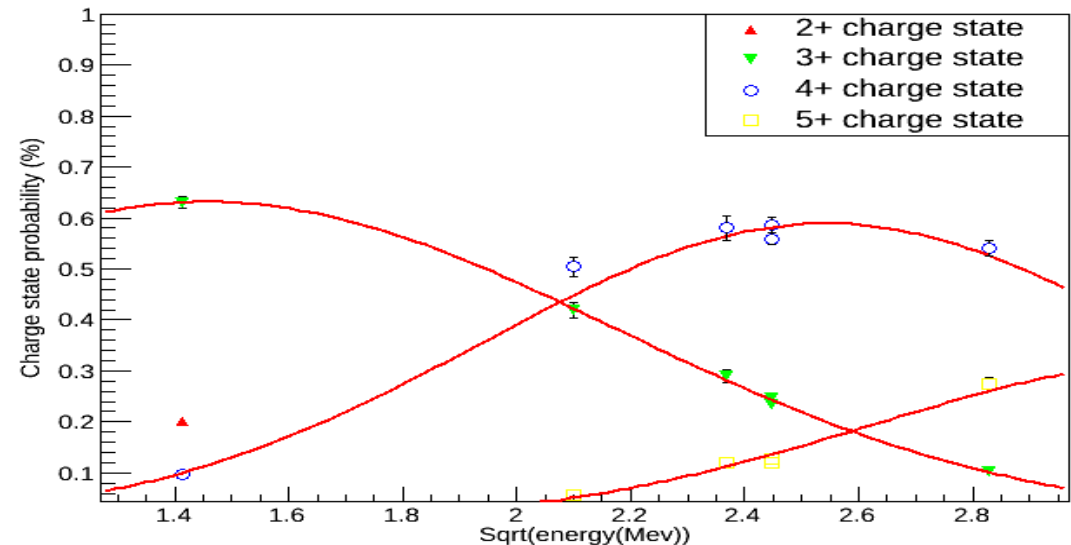
D. Schürmann et al., Eur. Phys. J. A (2013) 49: 80

# Energy loss measurements



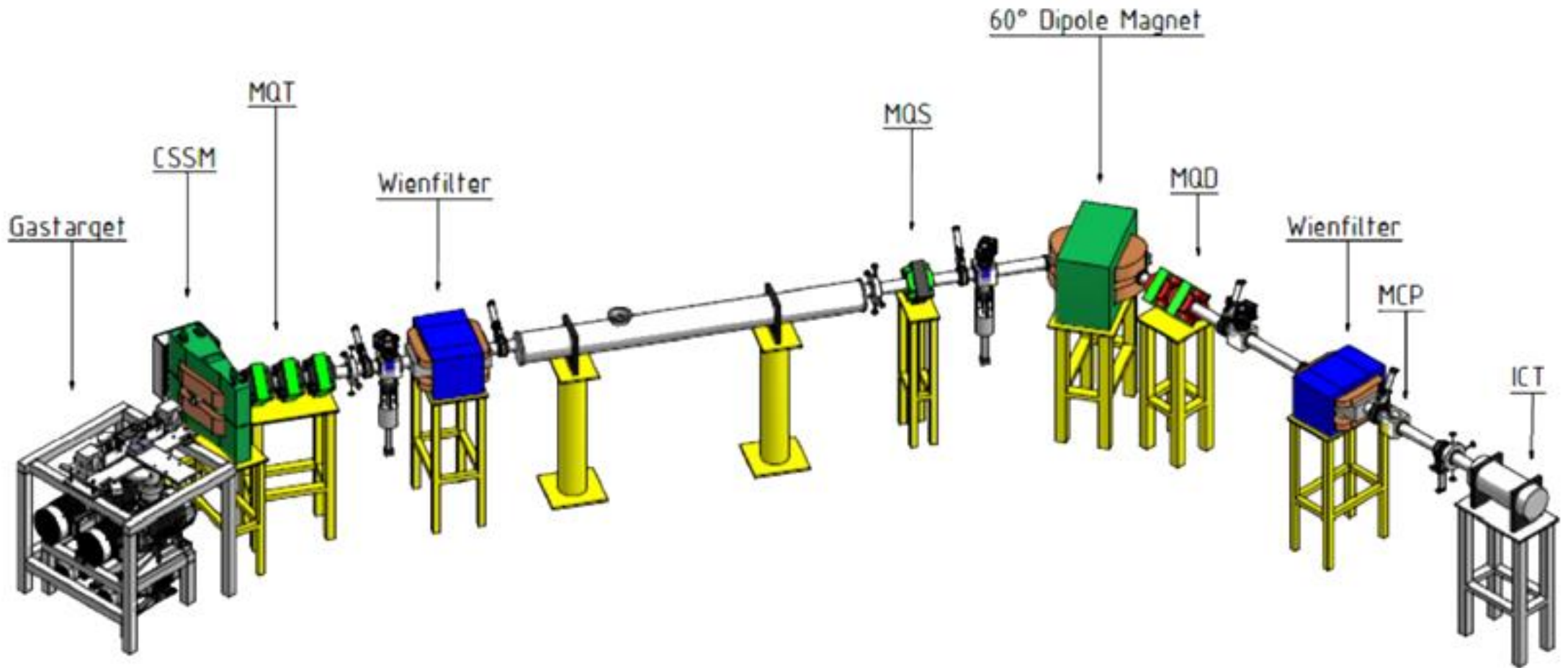
Energy loss of  ${}^7\text{Be}$  projectiles and  ${}^8\text{B}$  recoils in the target. Significant differences in absolute values are found with respect to SRIM (shown renormalized to data).

${}^8\text{B}$  charge state probability. The reaction yield measurements are performed selecting recoils in the 3+ charge state.

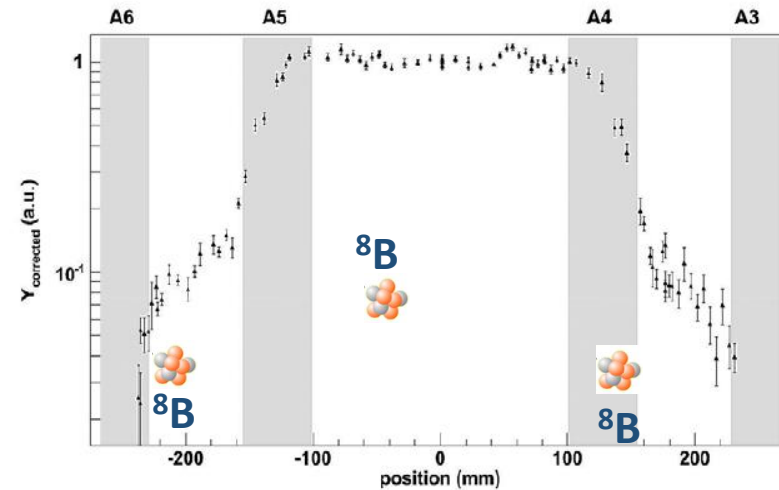
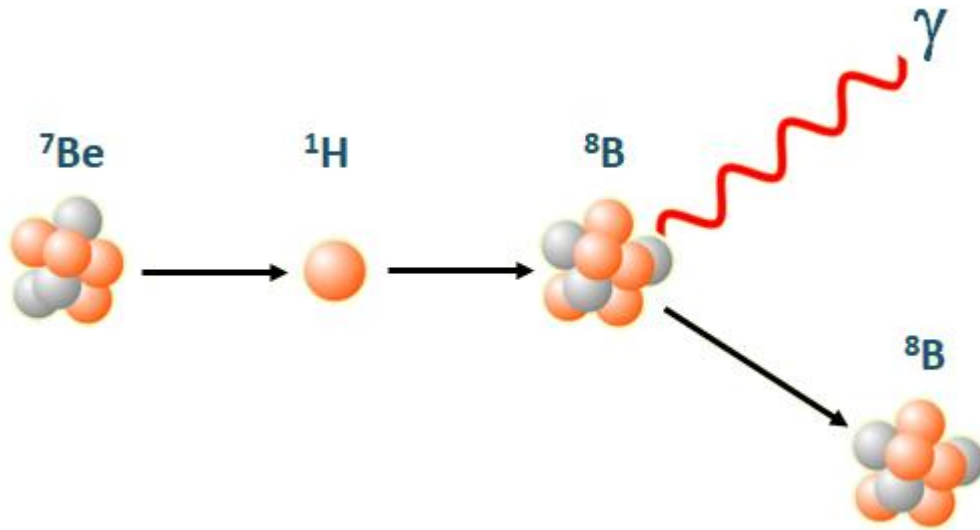




# European Recoil mass separator for Nuclear Astrophysics (ERNA)



# Recoils emittance



Recoils emittance is determined by reaction kinematics and straggling due to interaction with target gas.

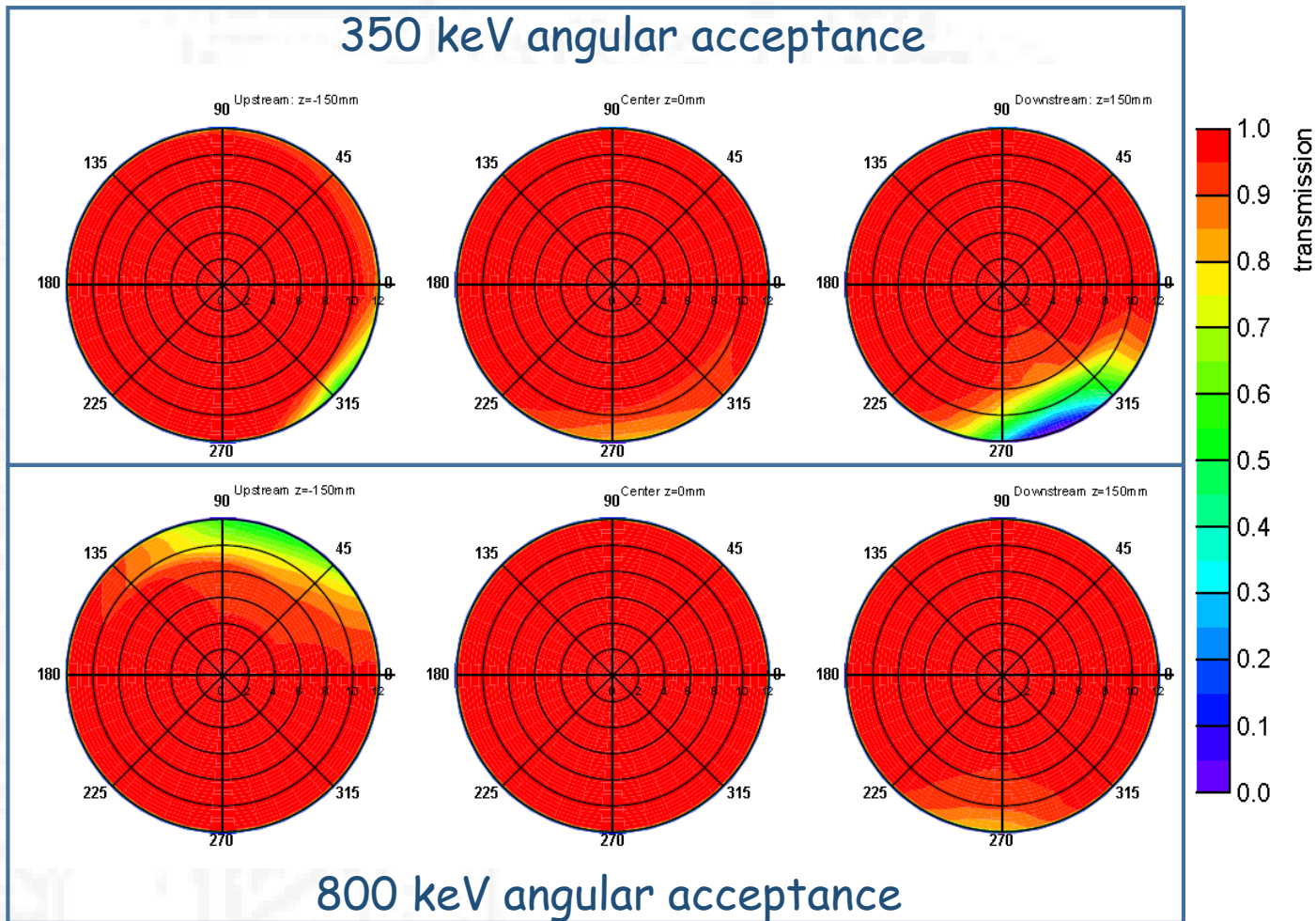
E.g. at  $E_{\text{c.m.}}=350\text{keV}$  maximum angle is 12 mrad and energy broadening is 90 keV

The  ${}^8\text{B}$  can be produced in different part along the gas target.

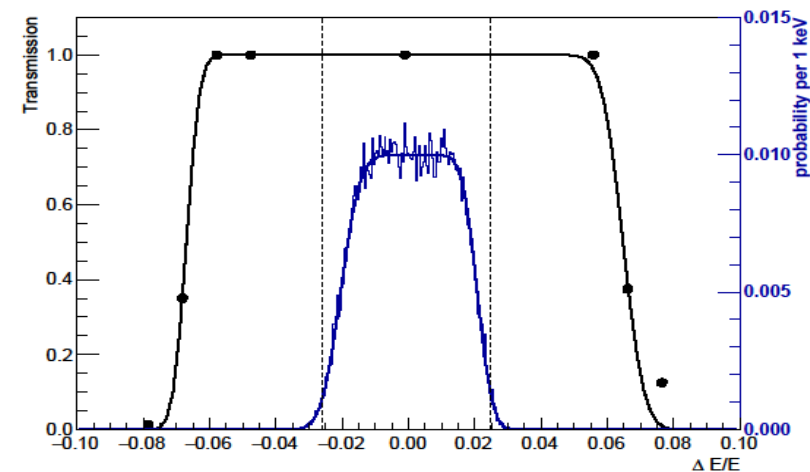
In all conditions 100% transmission to the end detector of the recoils in the selected charge state is mandatory.

# Recoils Acceptance

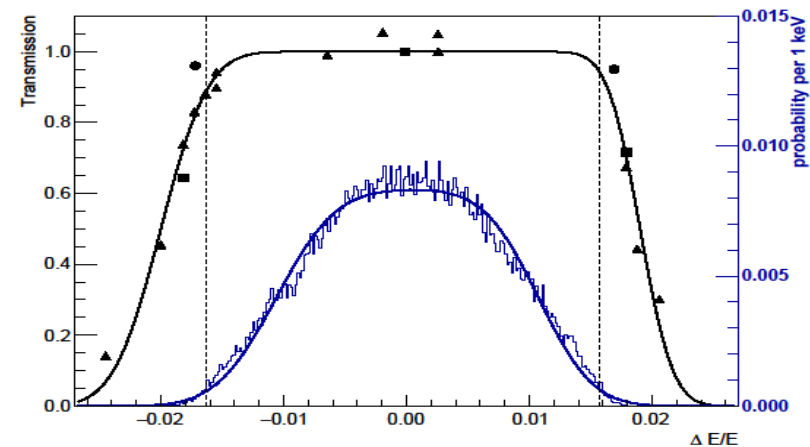
## 350 keV angular acceptance



## 350 keV energy acceptance

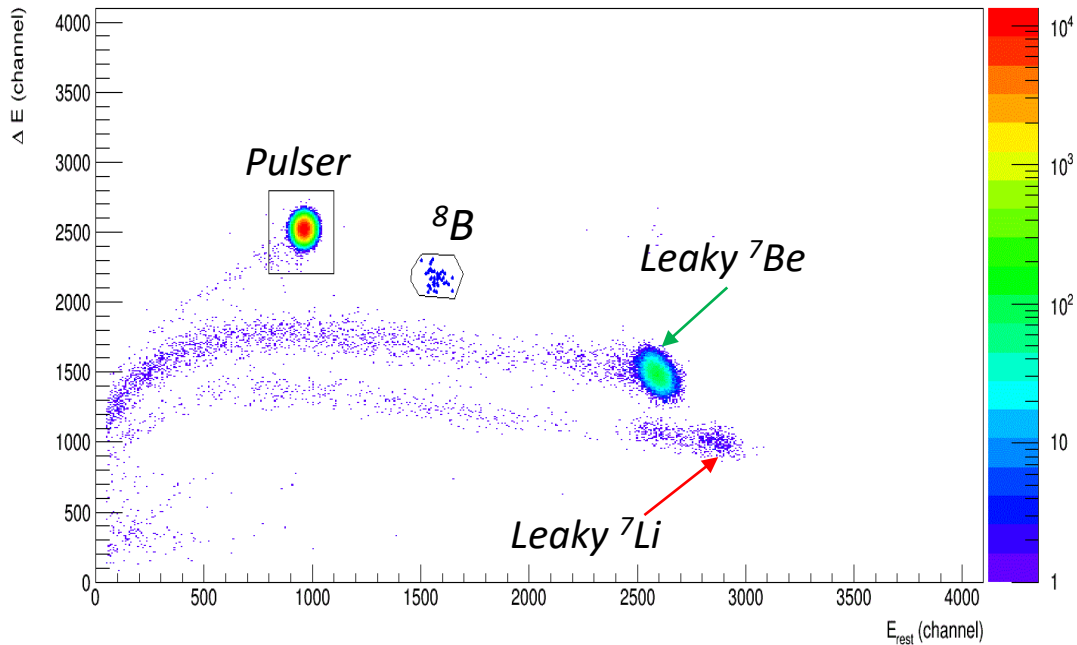


## 800 keV energy acceptance



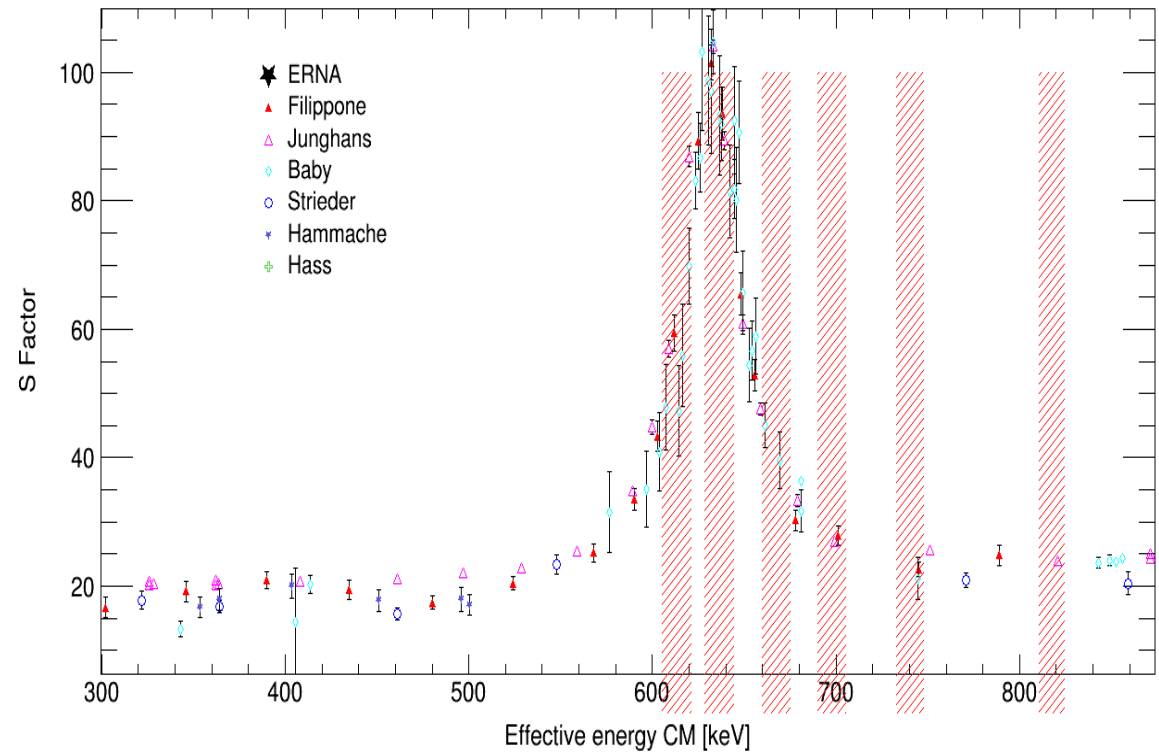


# Preliminary results



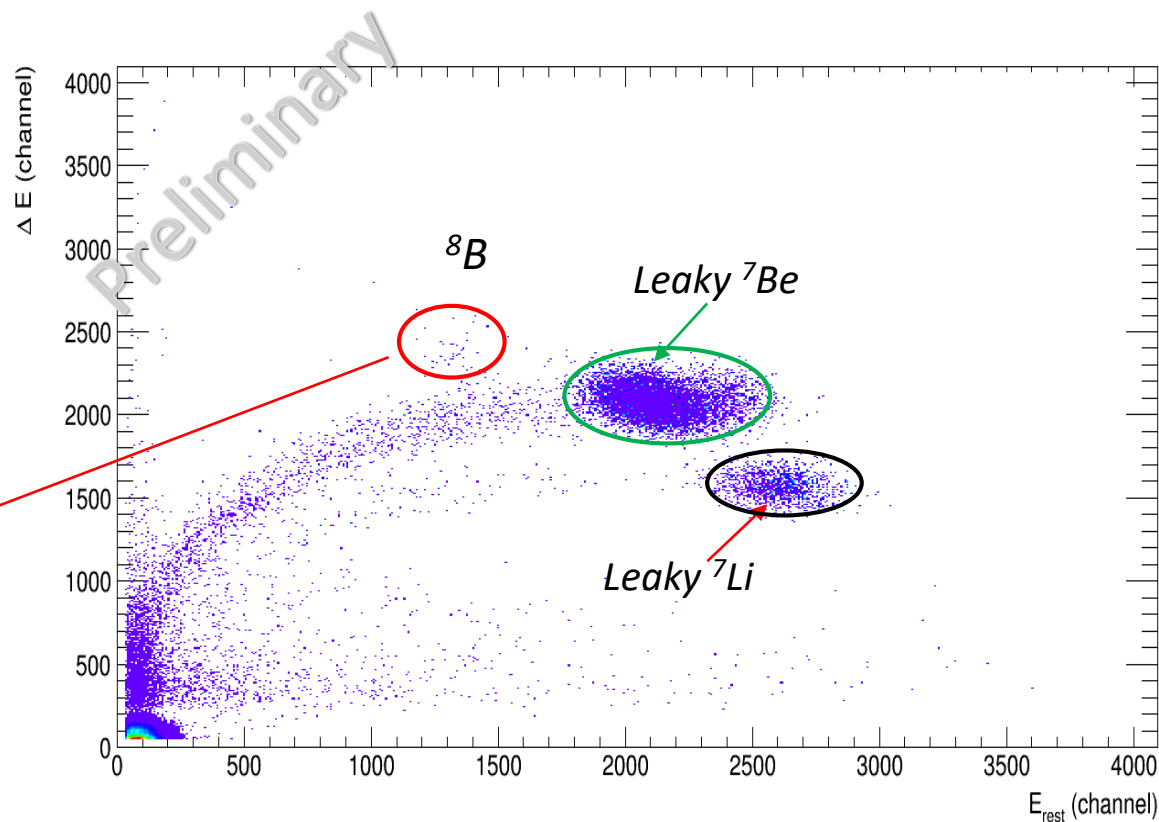
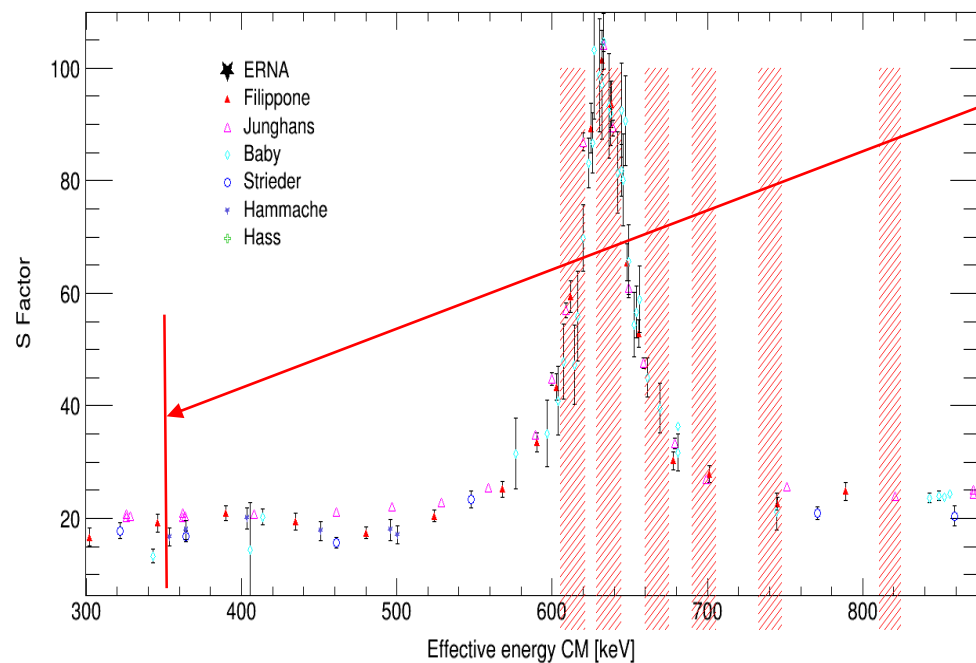
Comparison from the literature data and the energy range (corresponding to the energy loss in the H target) of the measurements already performed at  $E > 600\text{keV}$  center of mass effective energy.

Typical ionization chamber telescope E-DE spectrum.  
The  $8\text{B}$  recoils are well separated from the leaky.



# Ec.m.= 350 keV measurement

Reaction yield measurement performed over 4 days of beam time, analysis ongoing.  
The result will allow to evaluate the slope of the astrophysical S-factor.



E- $\Delta E$  spectrum with ionization chamber telescope. The  $^8\text{B}$  recoil, in the red ellipse.

# Conclusion

- A very intense  ${}^7\text{Be}$  beam, up to  $10^9$  pps, is routinely produced and characterized at CIRCE laboratory.
- The characterization of the extended gas target and the tunings of the separator has been completed at 800 keV, 600 keV and at 350 keV ;
- The measurements of absolute cross section, of the  ${}^7\text{Be}(p,\gamma){}^8\text{B}$  reaction trough the 629 keV resonance and up to the 800 keV has been performed.
- A new measurements at 350 keV has been performed, and the analysis data is ongoing;
- The preliminary results of the data after the 629 keV resonance up to 800 keV shown a better accord with the data sets with high value;
- The analysis of the data can be completed in the next months, including the 350 keV point, and the impact in the extrapolation of the astrophysical factor will be evaluated.



# Center for Isotopic Research on Cultural and Environmental heritage (CIRCE)



**Thanks**

Nuclear Physics in Astrophysics VIII 18-23 June 2017 Laboratori  
Nazionali del Sud