Nuclear Physics in Astrophysics VIII



Contribution ID: 56 Type: Oral

Direct study of the 22 Ne({\it p, γ }) 23 Na reaction in inverse kinematics at DRAGON

Monday, 19 June 2017 17:30 (20 minutes)

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%%% Authors and affiliations are next. The presenter should be
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for the DRAGON collaboration}
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\AFFILIATION{1}{TRIUMF, Vancouver, Canada}
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\vspace{12pt} % Do not modify
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The 22 Ne({\it p, γ }) 23 Na reaction largely impacts the abundance of the only stable sodium isotope, 23 Na, in various stellar environments, such as AGB stars, massive enough to undergo hot-bottom burning, type Ia supernovae and novae. However, the 22 Ne({\it p, γ }) 23 Na reaction rate still carries one of the highest uncertainties among the astrophysical reactions involved in the NeNa cycle, thereby also affecting the abundance predictions of elements between 20 Ne and 27 Al.

Reducing the uncertainties of abundance predictions for NeNa cycle elements by constraining the relevant reaction rates experimentally has received increased attention with the discovery of the anticorrelation between sodium and oxygen abundances in globular cluster stars.

The thermonuclear reaction rate for the 22 Ne({\it p, γ }) 23 Na proton capture reaction is dominated by a number of narrow resonances within the Gamow window.

Recently, a study with the objective to directly measure the strengths of the most relevant resonances in the $^{22}\mathrm{Ne}(\{\mathrm{hit}\ \mathrm{p},\gamma\})^{23}\mathrm{Na}$ reaction in inverse kinematics was carried out using the DRAGON (Detector of Recoils and Gammas Of Nuclear Reactions) recoil separator at TRIUMF. Resonances within an energy range from $\mathrm{E}_{c.m.}$ =178~keV to $\mathrm{E}_{c.m.}$ =1.222~keV were investigated.

In this contribution the astrophysical motivation behind this measurement, as well as preliminary results of the first inverse kinematics study of the 22 Ne({\it p, γ }) 23 Na reaction will be presented.

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Session Classification: Direct measurements 1

Track Classification: Stellar evolution and nucleosynthesis