Dark photon searches with PADME

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The long quest for dark matter

Zwicky, Coma cluster (1933)

 $M/L \approx 660 \ M_{\odot}/L_{\odot}$

Hubble Space Telescope (2007) Cluster Cl 0024+17

Lensing of background galaxies









 $\frac{1}{2}mv^2 = GMm/R$

Virial theorem

For a system of *N* particles, the moment of inertia and its derivatives are: $I = \frac{1}{2} \sum_{i} m_{i} \mathbf{r}_{i} \cdot \mathbf{r}_{i}$

 $dI/dt = \Sigma_i m_i d\mathbf{r}_i/dt \cdot \mathbf{r}_i$

 $d^{2}I/dt^{2} = \sum_{i} m_{i} \left(\frac{d\mathbf{r}_{i}}{dt} \cdot \frac{d\mathbf{r}_{i}}{dt} + \mathbf{r}_{i} \cdot \frac{d^{2}\mathbf{r}_{i}}{dt^{2}} \right)$

Equation of motion: $m_i d^2 r_i / dt^2 = -\sum_{j \neq i} G m_i m_j / |r_i - r_j|^3 (r_i - r_j)$

Kinetic energy; $2T = \sum_{i} m_{i} (dr_{i}/dt \cdot dr_{i}/dt)$

 $d^{2}I/dt^{2} - 2T = -\sum_{i} \sum_{j \neq i} Gm_{i}m_{j} / |\mathbf{r}_{i} - \mathbf{r}_{j}|^{3} \mathbf{r}_{i} \cdot (\mathbf{r}_{i} - \mathbf{r}_{j}) = \dots$ $= \frac{1}{2} \sum_{i} \sum_{j} Gm_{i}m_{j} / |\mathbf{r}_{i} - \mathbf{r}_{j}| = U$

 $d^2I/dt^2 = 2T + U$

Virial equilibrium: 2<*T*> + <*U*> = 0







Estimate masses using velocities

 $\Sigma_{i} m_{i} < v_{i}^{2} > = \Sigma_{i} \Sigma_{j < i} G m_{i} m_{j} 1 / < |r_{i} - r_{j}| >$

 $< v_{r,i}^{2} >_{\Omega} = 1/3 v_{i}^{2}$

projected along radial direction, averaged over solid angle Ω

- We see only radial component of motion, $\langle v_i \rangle \approx \sqrt{3} v_r$
- We see projected radii: $r = \theta d$

 $1/\langle |\mathbf{r}_i - \mathbf{r}_j| \rangle = 1/|\mathbf{r}_i - \mathbf{r}_j| \langle 1/\sin\theta_{ij} \rangle_{\Omega}$

Assuming N equal masses $\Sigma_i m_i = N m$

 $M_{VT} = 3/2\pi \ G^{-1} \ N \sum_{i} v_{i}^{2} / \sum_{j < i} 1/r_{ij}$

Coma cluster (Zwicky): $\sigma \approx 1000 \text{ km/s}, R \approx 3Mpc, M_{VT} = 3 \ 10^{15} M_{\odot}$ $L = 5 \ 10^{12} L_{\odot}$







Many pieces of evidence for dark matter

Rotation curves



Large scale structures



Cosmic Microwave Background



$\Lambda\text{-}\mathsf{CDM}$ model



Lensing



Colliding clusters (Chandra)





Dark photon searches with PADME



The dark matter problem

Original drawing by Stacy McGaugh (1995)



Several hypothesized solutions



Roots are the empirical observations





A new kind of matter?

Dark matter dominating in the early Universe





- Standard Model only includes <20% of the matter in the Universe
 - We only know dark matter interacts gravitationally
- Many open questions
 - What is dark matter made of?
 - How dark matter interact, if it does, with SM particles?
 - Does one or more new dark force exist?
 - How complex is the dark sector spectrum?









Where to search for dark matter

- Without modifying the SM structure: U(1)_Y+SU(2)_L+SU(3)_C
 - Dark matter can't be strong interacting (scattering cross section too high)
 - Cannot be electrically charged, otherwise it would not be dark!
 - It can be weakly interacting and massive!
- The WIMP has all the characteristics needed to solve the dark matter problem...
- But so far* more than 20 years of unsuccessful attempt to detect WIMPs
 - Strong constraints from the LHC and direct searches at masses up to 1TeV
 - *Some hints however...



What about introducing a **new force**?









The WIMP "miracle": m_{χ} =100 GeV, g_{χ} =0.6, Ω_{χ} = 0.1



Portals to secluded sector







Dark photon

- The simplest hidden sector model just introduces one extra U(1) gauge symmetry and a corresponding gauge boson: the "dark photon" or U boson or heavy photon (γ' or A')
- An extra U(1) symmetry implied in many Standard Model extensions, some classes of string theory, etc.







Dark photon

- <u>Two types</u> of interactions with SM particles should be considered
 - As in QED, generates interactions of the type:

 ${\cal L}~\sim~g'q_far{\psi}_f\gamma^\mu\psi_f U'_\mu$

- Not all the SM particles need to be charged under this new symmetry
- In the most general case q_f is different in between leptons and quarks and can even be 0 for quarks. (P. Fayet, Phys. Lett. B 675, 267 (2009).)
- Couples to SM hypercharge through kinetic mixing operator:
 - $\epsilon/2 F_{\mu\nu}^{\gamma}F^{\mu\nu}$, where $F^{\mu\nu}=\partial_{\mu}A'_{\nu}$
 - $A_{\mu} \rightarrow A_{\mu} + \epsilon a_{\mu}$; $\alpha' = \epsilon^2 \alpha$
 - The dark photon acquires a (small) SM charge



e





Dark photon production



A' can be produced in electron collision on target by:

- Bremsstrahlung: $e N \rightarrow e N A'$
- Annihilation: $e^+ e^- \rightarrow \gamma A'$
- Meson decays





Dark photon visible decays

- Assume that no additional lighter states exists in the dark sector with $m_{\gamma} < m_{A'}/2$
- Dark photon couples to SM particles through kinetic mixing only (with same coupling εq)
- For $m_{A'} < 2 m_{\mu}$ only decays to e^+e^-









Dark photon invisible decays

- If χ state with U(1) charge q_U and coupling constant g_U exists in the dark sector with $m_{\chi} < m_{A'}/2$, the coupling to the A' will be: $q_U g_U$
- $A' \rightarrow \overline{\chi}\chi$ will be dominant wrt to visible decays for $\alpha_D > \alpha$, i.e. $|q_U g_U| > \varepsilon e$









A dark matter "messenger"



Dark Matter scattering on nuclei





Dark Matter annihilation...



Dark photon searches with PADME





Dark sector with dark Higgs

- Model assumes the existence of an elementary dark Higgs boson h', which spontaneously breaks the U(1) symmetry.
 PRD 79, 115008 (2009)
- A' boson produced together with a dark Higgs h' through a Higgs-strahlung $e^+e^- \rightarrow A' h'$
 - Cross section =20fb×(α/α_{D})($\epsilon^{2}/10^{-4}$)(10GeV)²/s
 - For light h' and A' ($M_{U,h'}$ <2M μ) final state with 3(e⁺e⁻ pair) are predicted
 - Background events with 6 leptons are very rare at this low energies
 - Due to A',h'being very narrow resonances strong kinematical constraints are available on lepton pair masses
- Experimental search by BaBar and KLOE-2 for A' masses above 200 MeV









Dark photon + dark Higgs searches



BaBar Phys. Rev. Lett. 108, 211801 (2012)



KLOE-2 arXiv:1501.06795

- No data available below 200 MeV in M_{A'}
- Production mechanism being Bremsstrahlung, PADME can reach $M_{A'}$ >100MeV
- PADME can provide sensitivity in unexplored parameter region







and the ATLAS excess of course...





Photons, Photon Jets and Dark Photons at 750 GeV and Beyond, arXiv:1602.04692

Dark sector shining through 750 GeV dark Higgs boson at the LHC, arXiv:1601.02490





The DAMA-Libra effect



4000

4500

5000





- Nuclear recoil by the exchange of a dark photon
 - Independent of χ mass value



3500

Time (day)





Particle astrophysics: PAMELA, AMS



- Positron eccess: PAMELA, FERMI, AMS-02
- No significant excess in antiprotons
 - Consistent with pure secondary production
- Leptofilic dark matter annihilation?
- If DM is the explanation, the mediator should be light, < 2m_{proton}



...naturally leptophilic







Muon g-2 SM discrepancy



g-2 in the Standard Model







About 3σ discrepancy between theory and experiment (3.6 σ , if taking into account only e+ e- \rightarrow hadrons) Additional diagram with dark photon exchange can fix the discrepancy (with sub GeV A' masses)

Contribution to g-2 from dark photon

$$\Delta a_{\mu} = \frac{\varepsilon^2 \alpha}{2\pi} \times \begin{cases} 1 & \text{for} \quad m_{\mu} \ll m_{A'} \\ \frac{2m_{\mu}^2}{3m_{A'}^2} & \text{for} \quad m_{\mu} \ll m_{A'} \end{cases}$$







g-2 electron

Caution with $(g-2)_e$ constraint

- The two most precise determinations of fine structure constant disagree at 1.5σ level
- One can reasonably argue for a more conservative constraint

$$\Delta a_e = (-1.05 \pm 0.82) \times 10^{-12}$$
Aoyama et al. 1205.5368

Or just using error

$$\Delta a_e = \pm 0.82 \times 10^{-12}$$



Important to also have a direct probe of this region of parameter space!

Brian Batell





Where to look for dark photons?







Where to look for dark photons?

- Coupling expected in the range $\varepsilon \sim 10^{-2} 10^{-3}$ but can be further • suppressed by an enhanced symmetry
- Depending on the model, mass scales like:
 - $m_{A'}/m_{W} \sim \varepsilon \varepsilon^{1/2}$

leading to a MeV-GeV mass scale









Dark photon experiments

Thick target (beam-dump)

- Absorb all SM backgrounds
- Look for visible decays (e+ e⁻, μ + μ ⁻, ...)

• Thin target + decay of dark photon:

- Decay to visible particles (e+ e⁻, μ + μ ⁻, ...)
 - "Bump hunting", looking for a peak in the invariant mass
 - Displaced vertices, looking for long-lived particles
- Decay to invisible particles
 - Look for missing mass
 - DM particles recoil
- Meson decays
- Dark particles scattering





Why fixed target?







Fixed target experiments



- Main backgrounds: SM Bremsstrahlung + Bethe-Heitler
- Kinematics:
 - A' takes nearly all the beam energy E_0 (sharply peaked at x \approx 1)
 - Electron takes a small energy $\approx m_{A'}$
 - A' emission almost collinear to the beam: $\theta_{A'} = (m_{A'}/E_0)^{3/2}$
 - Electron going at "wide" angle: $\theta_e = (m_{A'}/E_0)^{1/2}$
 - A' decay products open by $\theta \approx m_{A'}/E_0$

F'



 E_0

Target

θ



Electron beam-dump experiments



In addition to cross section advantage





Electron beam-dump experiments



$$\begin{split} N_{\gamma'} &= \sigma_{\gamma'} N_e \frac{N_0}{A} \rho_{\rm sh} L_{\rm sh} & \begin{array}{l} & \begin{array}{l} \mbox{Electron energy} \\ \mbox{distribution due to} \\ \mbox{the interaction in} \\ \mbox{target+shield} \end{array} & \begin{array}{l} \mbox{Decay probability of } \gamma' \\ \mbox{after shield} \end{array} \\ \\ \hline \\ \frac{dN_{\gamma'}}{dx_0 \ dz} &= N_e \frac{N_0 X_0}{A} \int_{E_{\gamma'}+m_e}^{E_0} dE_e \int_0^{T_{\rm sh}} dt_{\rm sh} \Biggl[I_e(E_0, E_e, t_{\rm sh}) \left. \frac{E_0}{E_e} \left. \frac{d\sigma}{dx_e} \right|_{x_e = \frac{E_{\gamma'}}{E_e}} \frac{dP(z - \frac{X_0}{\rho_{\rm sh}} t_{\rm sh})}{dz} \Biggr] \\ \\ T_{\rm sh} \equiv \rho_{\rm sh} L_{\rm sh}/X_0 & \begin{array}{l} \mbox{dofdx for } \gamma' \text{ production} \\ \mbox{by Bremsstrahlung} \end{array} \end{split}$$





Limits from electron beam-dump experiments







Limits from electron beam-dump experiments

- Beam-dump experiments: looking for decay products of "rare penetrating particles" behind a **stopped electron beam**
- SLAC E141 (1987) and SLAC E137 (1988), Fermilab E774 (1991)







Proton beam dump experiments

Use data of the search of $v_H \rightarrow ve+e^-$ for looking for $P \rightarrow \gamma A'$ Pseudoscalar decaying to spin 0 or ½ particles negligibly small







Limits from past experiments: proton beam dump







Limits from past experiments: proton beam dump

NOMAD and **PS191** looked for decay of and heavy neutrino $v_H \rightarrow ve+e^-$ Look for $\pi^0 \rightarrow \gamma A'$

NOMAD: 4.1·10¹⁹ POT E>4 GeV, m_{ee}<95 MeV PS191: 0.89·10¹⁹ POT

$$Br(\pi^0 o \gamma A') = 2\epsilon^2 Br(\pi^0 o \gamma \gamma) \left(1 - \frac{M_{A'}^2}{M_{\pi^0}^2}\right)$$







Dark photon experiments map







Thin target experiments

Running:

- APEX at JLAB Hall-A, test run done, full run coming
- A1 at MAMI
- HPS at JLAB Hall-B, first run done in 2015

Coming soon:

PADME at Frascati (approved)

Proposed:

- DarkLight at JLAB FEL (electron on gas jet target)
- VEPP3 (electron on gas jet target)
- Cornell (positron extracted from CESR on H₂ target)




APEX







APEX test

Background rejection and final dataset

Reducible backgrounds

- Electron singles from inelastic or electron-nucleon scattering
- · Pions from virtual photon decays
- Proton singles
- Accidental e⁺e⁻ coincidences
- $ightarrow e^+e^-$ pairs from real photon conversions

Pion rejection:

- + Production ratio in right HRS: $e^+/\pi^+ > 1/100$
- Online pion rejection: factor of 30
- Offline rejection > 1/100 using both gas Cherenkov and calorimeters

Final event sample trigger:

Double coincidence gas Cherenkov signal within 12.5 ns window in each arm

Final data sample consisted of 770500 true e^+e^- coincident events with 0.9% (7.4%) meson (accidental e^+e^- coincidence) contamination







MAMI A1





JLAB Hall-A APEX

n×1.1 GeV, continuous, 200 μA beam

MAMI A1

855 MeV, continuous, 90 μA beam







Summary of limits from visible decays

Practically, all the $(g-2)_{\mu}$ favored band already excluded But still large interest for excluding the uncovered parameter space









HPS

- Increase acceptance wrt double arm spectrometer
- Look for displaced vertex
- θ of the decay is small:
 - put detectors as close as possible, good forward coverage
 - Aim at minimum $\theta_{A'} \approx 15$ mrad



- Bump hunting needs good momentum/mass resolution
- Good tracking and analyzing magnet
- Aim at $\Delta m/m \approx 1\%$ and $\Delta z \approx 1$ mm
- Trigger with a high rate calorimter
- Magnet+calorimeter to select e⁺ and e⁻
- Magnet+muon detector to select $\mu^{\scriptscriptstyle +}$ and $\mu^{\scriptscriptstyle -}$







HPS





Detectors split in two halves to let the beam pass through







HPS



The challenge – operating Layer-1 Si-tracker at 500 μm from the high intensity beam!

Small size (σ_y ~50 µm), very stable (~50 µm) beam is needed

The operational goals have been achieved during the first engineering run







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DarkLight

FEL electron beam, 100 MeV, continuous, 10 mA, onto $10^{19} H_2/cm^2$ gas jet target

- Proton recoil detector. Full reconstruction of event for background rejection
- Vertexing and low momentum lepton tracker: TPC
- Outer trackers



Test performed on prototype vacuum chamber to assess beam transport feasibility





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Summary of limits from visible decays

Practically, all the $(g-2)_{\mu}$ favored band already excluded Still large interest for excluding the uncovered parameter space







Invisible decays: a dark matter beam





Scattering on nuclei





Elastic scattering on electrons







Back to past experiments



SLAC E-137 20 GeV electron beam 30 C dumped on Aluminum target Shower calorimeter, 400 m distance Re-analysis (Batell, Essig, Surujon) constrains $m_{A'}$ vs ε , dependent on α_{D} and m_{χ}



arXiv:1406.2698v1

 χe elastic scattering

LSND

 π^0 decays to $\gamma A'$ from LAMPF 800 MeV protons $10^{23}~\text{POT}$

30 m off-axis detector, 170 ton mineral oil





Possible future proton beam dumps: SHIP at SPS









A new scattering experiment: BDX at JLAB



Backgrounds:

- Neutrino production
- Cosmogenic muons and neutrons

LOI presented to JLAB PAC

Scintillator 1 m³ 1 MeV/10 MeV e^+e^- detection threshold





Scattering experiment: BDX at JLAB

High energy beam advantages:

- Higher cross sections
- χ beam boosted, larger acceptance







Cross section dependence from A' mass, χ mass, coupling constant







BDX experiment (Hall-A)





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Invisible decays, model dependence of limits









Combine visible and invisible decays









Model independent limits from invisible decays







P-348 at CERN SPS

H4 high purity electron beam, <1% contamination required (tertiary, from γ conversions) Use synchrotron radiation tagging





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P-348 at CERN SPS

- N_e=10¹² requested (3 months run)
- Main backgrounds:
 - punch-through of primary energy into ECAL1
 - Beam-related background (mis-identified electrons): muon and hadronic events









P-348 at CERN SPS

• Also proposal for $A' \rightarrow$ invisible search

 $S_{A'} = \text{ECAL1} \times \overline{\text{V1} \times \text{S1} \times \text{S2} \times \text{ECAL2} \times \text{V2} \times \text{HCAL}}$

- Main backgrounds:
 - punch-through of e^- or γ
 - Non-hermeticity of HCAL
 - Low energy tail of e⁻ beam
 - e⁻ induced photo-nuclear reactions
 - Muon events





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BaBar limit

Invisible Dark Photon: $e^+e^- \rightarrow \gamma + invisible$





Peaking background from $e^+e^-\rightarrow\gamma\gamma$, with one of the photons missing the EM calorimeter. Veto such events by detecting activity in the muon detector (IFR).

- $\Upsilon(3S) \rightarrow \gamma + \text{invisible}$ (arXiv:0808.0017)
- Require a single photon with $E_{\gamma}^*>2.2 \text{ GeV}$
- No charged tracks
- No additional energy in EMC above 100 MeV
- Missing momentum points to EMC
- No activity in IFR aligning with missing momentum
- No signal found: limits on ε of order O(10⁻³-10⁻²)
- Updated analysis in progress

LDMA2015

BABAR New Physics Searches

Alberto Lusiani



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The PADME approach

- At present all experimental results rely on at least one of the following model dependent assumptions:
 - A' decays to e^+e^- (visible decay assumption BR(A' $\rightarrow e^+e^- = 1$)
 - A' couples with the same strength to all fermions ($\varepsilon_q = \varepsilon_l$) (kinetic mixing)
- In the **most general** scenario
 - A' can decay to dark sector particles lighter than the A' BR($A' \rightarrow e^+e^- \ll 1$)
 - Dump and meson decay experiment only limit $\epsilon^2 BR(A' \rightarrow e^+e^- << 1)$
 - A' can couple to quark with a coupling constant smaller ϵ_{l} or even 0
 - Suppressed or no production at hadronic machines and in mesons decays
- PADME aims at detecting A' produced in e^+e^- annihilation and decaying into invisibles by searching for missing mass in $e^+e^- \rightarrow \gamma A'$, $A' \rightarrow XX$
 - No assumption on the A' decays products and coupling to quarks
 - Only minimal assumption: A' bosons couples to leptons
 - PADME will limits the coupling of **any new light particle** produced in e^+e^- collision (scalars (H_d), vectors (A' and Z_d)







What we need: signal



In order to compute $M_{\text{miss}}^2 = (\underline{P}_{\gamma} - \underline{P}_{e+})^2$ we need:

- a) A **positron** beam with a well defined four-momentum
 - 1. Small energy and angular spread
 - 2. Small transverse spot
 - 1+2 = small emittance
 - 3. Tunable intensity (in order to optimize annihilation vs. Bremsstrahlung)
- b) Measure precisely the photon (tri-)momentum (angle and energy)





What do we need: background



We need to fight the backgrounds i.e. **one photon** + something else, eventually going undetected:









$\mathsf{DA}\Phi\mathsf{NE}$ complex in Frascati

- DAΦNE, replacing ADONE (operational until 1993), has been running as e⁺e⁻ collider at 1,02 GeV since 1999, for KLOE, DEAR, FINUDA, Siddharta, and now KLOE/2 ...
- Synchrotron light source operational with 3 lines (X, UV, IR)
- High current electron/positron linac + damping ring + test facility







LINAC parameters

The "shotgun" of the system is of course the high-current linac



	Design	Operational
Electron beam final energy	800 MeV	510 MeV
Positron beam final energy	550 MeV	510 MeV
RF frequency	2856 MHz	
Positron conversion energy	250 MeV	220 MeV
Beam pulse rep. rate	1 to 50 Hz	1 to 50 Hz
Beam macropulse length	10 nsec	1 to 40 nsec
Gun current	8 A	8 A
Beam spot on positron converter	1 mm	1 mm
norm. Emittance (mm. mrad)	1 (electron) 10 (positron)	< 1.5
RMS energy spread	0.5% (electron) 1.0% (positron)	0.5% (electron) 1.0% (positron)
electron current on positron converter	5 A	5.2 A
Max output electron current	>150 mA	350 mA
Max output positron current	36 mA	100 mA max
Trasport efficiency from capture section to linac end	90%	90%
Accelerating structure	SLAC-type, CG, 2π/3	
RF source	4 x 45 MWp SLED-ed klystrons TH2128C	









The beam test facility







DA Φ NE Beam Test Facility (BTF)

- Longer Duty Cycle
 - Standard BTF duty cycle = 50*10 ns = $5x10^{-7}$ s
 - Already obtained upgrade 50*40ns= $20x10^{-7}$ s
 - Work in progress to reach **150 ns** (new pulser) ...
 - ... Up to 250 ns (double phase inversion at the SLED) ...
 - ... and beyond (no SLED, or SLED detuning), in principle up to 4 μs
- Energy upgrade planned in 2017.
 - Region from 0-22 MeV can be explored with a positron beam of 550 MeV
 - The accessible $M_{A'}$ region is limited by beam energy
 - e.g. $M_{A'}$ up to **28 MeV** with 750 positron beam









BTF beam summary

- Energy spread Δp/p ~1%
- Beam spot: <1 mm RMS
- Divergence: 1 1.5 mrad
 - Effect of multiple scattering and Bremsstrahlung on the Beryllium exit window and in air has to be considered
 - Both size and divergence depend on the **optics** ullet
- Beam position: 0.25 mm RMS
- Pulse duration: 1.5 40 ns
 - 10 ns during collider operations











Beam spot size

BTF experimental hall





Approximately **<5.5 m total** length (<3 m lateral width)





BTF users

Luckily enough, BTF is already extensively used by many experimental groups in HEP and astro-particles...



Non Italian institutions (2011-2014)



Lab. De Phys. Grenoble, 2

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Laboratoire de l'Accelerateur Lineaire, 4



BTF beam-line upgrade







The PADME experiment



- 10³-10⁴ e⁺ on target per bunch, at 50 bunches/s (10¹³-10¹⁴ e⁺/year), limited by pile-up, mainly due to Bremsstrahlung events
- Active target, thin: e.g. 50-100μm diamond with strips
 - Optimize by looking at annihilation vs. Bremsstrahlung cross section
- Magnetic spectrometer/veto ~ 1m length × 0.5 T for sweeping away 550 MeV beam
 - Conventional magnet with large gap for gaining acceptance
 - Possibility to increase field for energy upgrade to ~ 1 GeV
 - Available from CERN, spare of MBP dipoles of SPS transfer line
- Cylindrical crystal calorimeter
 - Optimize radius vs. distance by looking at background rejection vs. acceptance
 - In order to have an acceptable rate, central hole and
- Small angle detector for Bremsstrahlung veto
- Vacuum pipe





Starting from the magnet, build the layout around it





Adjustable gap by adding/removing iron insets





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The PADME magnet



- MBP-S series, on loan from CERN
 - Many thanks to TE-MSC-MNC, R. Lopez, D. Tommasini
 - Shipped to Frascati in Dec. 2015
- Poles: 100 cm length, 52 cm width
- Variable gap 11 to 20 cm, we further extended to 23 cm gap
- Preliminary field mapping:
 - Good B field quality
 - Fringe field not negligible, even outside the coils, relevant for beam control upstream of the active target








The PADME magnet



Available power supply: 80 V/400 A







The PADME magnet



Magnet gap (– vacuum pipe depth) Fixes the maximum (vertical) acceptance together with the target position



Given the gap, the minimum magnetic field is given by **beam momentum** and **sweeping angle**





First task, measure the recoil photon

Our main detector is of course a calorimeter, with two basic requirements:

1. Measure E_{γ} , θ_{γ}

- Good energy resolution: 1-2%/VE[GeV])
 - High Photo-statistics
 - Containment
- Good angular resolution: ≈1 mrad

2. Fight pile-up

Sub-ns timing resolution





- The material choice fixes light yield & time resolution, Moliére radius & X₀
- The Moliére radius determines granularity
- The granularity + required angular resolution, set the distance from the target
- Given the distance, the lateral size fixes the angular coverage (i.e. acceptance)

But...

...we have to take into account two important constraints:

- The overall size of the experiment is the hall length (<5 m)
- Another important bound is the cost, which is driven by the material, size and granularity (i.e. the number of channels)

So the message is "keep it compact!"







The calorimeter

Parameter Units:	$r: \rho$ g/cm ³	MP °C	X_0^* cm	R_M^* cm	dE^*/dx MeV/cm	λ_I^* cm	$ au_{ m decay}$ ns	$\lambda_{ m max}$ nm	n^{\natural}	$\operatorname{Relative}_{\operatorname{output}^{\dagger}}$	Hygro- scopic?	d(LY)/dT $\%/^{\circ}C^{\ddagger}$
NaI(Tl)	3.67	651	2.59	4.13	4.8	42.9	245	410	1.85	100	yes	-0.2
BGO	7.13	1050	1.12	2.23	9.0	22.8	300	480	2.15	21	no	-0.9
BaF_2	4.89	1280	2.03	3.10	6.5	30.7	650^{s}	300^{s}	1.50	36^s	no	-1.9^{s}
							0.9^{f}	220^{f}		4.1^{f}		0.1^{f}
CsI(Tl)	4.51	621	1.86	3.57	5.6	39.3	1220	550	1.79	165	slight	0.4
CsI(pure)	4.51	621	1.86	3.57	5.6	39.3	30^s	420^{s}	1.95	3.6^{s}	slight	-1.4
							6^{f}	310^{f}		1.1^{f}		
$PbWO_4$	8.3	1123	0.89	2.00	10.1	20.7	30^s	425^{s}	2.20	0.3^{s}	no	-2.5
							10^{f}	420^{f}		0.077^{f}		
LSO(Ce)	7.40	2050	1.14	2.07	9.6	20.9	40	402	1.82	85	no	-0.2
LaBr ₃ (Ce)) 5.29	788	1.88	2.85	6.9	30.4	20	356	1.9	130	yes	0.2

Small Moliére radius and high light yield: LYSO and BGO

- Granularity $\approx R_{M} \rightarrow 2 \text{ cm}$
- a=2 cm → **point resolution**: 2 cm/V12=6 mm
- $\sigma_{\text{point}}=6 \text{ mm} \rightarrow 1 \text{ mrad at } 6 \text{ m distance} \rightarrow \text{too much}!$

But...

...we have clusters!

- Center of gravity should have a better resolution
- Most of the energy will be in a single crystal, pulling the cog towards the center of the most energetic one)

LYSO(Ce): high LY, high ρ , small X_0 and small R_M , short τ_{decay}

- Performance:
 - $\sigma(E)/E$ =1.1%/VE \oplus 0.4%/E \oplus 1.2%

BGO: high LY, high ρ , small X_0 and small R_M , long τ_{decay}

■ Resolution also in 1-2%/VE range

	- HILLING		
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		Mit.	

2 cm crystals							
d	$\langle d_{exp} - d \rangle$	RMS					
0.0	0.00	0.18					
-0.2	0.13	0.18					
-0.4	0.24	0.20					
-0.6	0.33	0.24					
-0.8	0.33	0.29					
-1.0	0.10	0.40					

Results with a Geant4 "photon gun", E=500 MeV $\sigma_{cluster} \approx 4.5$ mm (including the systematic shift), better than 6 mm





Longitudinal containment





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Signal acceptance in calorimeter



Calorimeter hole: 20 mrad

 θ_{max} =65, 75, 83 mrad from 58% to 65% acceptance







Calorimeter 2y and 3y backgrounds



NiewAl Constant NiewAl 1+2+3 NiewAl 1+2+3 NiewAl 1+2+3 NiewAl 1+2+3 NiewAl 1+2+3 NiewAl 1+2+3

NSeeLAV

NSeeSAC

No 2y events



Recoil γ definition: 10 MeV < E < 400 MeV 30 mrad < θ < 65 mrad

Per-mil 3γ background





-

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Residual background

Tighter signal definition: in fiducial region **and** 150 MeV < E < 450 MeV







Calorimeter layout



- Length >20 cm
- Cell side 2 cm
- Make simpler (square) hole, radius≈6 cm
- Fix outer radius at 28.5 cm, fiducial radius=25 cm
- Given 300 cm distance:
 - Fiducial acceptance 25/300=83 mrad
 - Angular resolution 4.5 mm/3 m≈1.5 mrad
- Total of 616 crystals
- 616×20×2×2 = 50000 cm³

Now...

... at 20€/cc such a calorimeter will cost 1 M€ only for buying the crystals

- + photosensors
- + readout





L3 BGO crystals



- 140 BGO crystals from former L3 electromagnetic calorimeter
- Asked **500 more** to the L3 collaboration
- Cut from trapezoidal prism shape to square section 20×20 mm², 220 mm long







Missing mass resolution



The minimum energy is also very important







BGO performance (electron beam)



PADME vacuum vessel



Vacuum mandatory for three purposes:

- 1. Not to spoil **beam quality** before hitting the target
- 2. To minimize **photon interactions** before reaching the calorimeter
- 3. To minimize **positron interactions** before hitting the veto detector (in particular showers!)

Different possibilities under study to minimize the material thickness, i.e. increase acceptance (given the magnet gap) for the vessel, with the following requirements:

- Hold the vacuum
- Host the scintillating bars for positron veto detectors
- Interface to target box (upstream) and straight section before calorimeter (downstream)





Small angle calorimeter

- BGO calorimeter cannot tolerate the Bremsstrahlung rate in the very central crystals
 - Inner hole 4-8 cm radius
- Small angle calorimeter aim to tolerate a rate of the order of 10 clusters (40 ns bunch length)
- The only fast enough inorganic crystal is BaF₂ with a fast PMT readout
- A possible alternative: Cherenkov detector











The diamond target



- Diamond is the rigid material with the best ee(γγ)/Bremsstrahlung ratio (Z=6)
- Measure charge and position of 5000-10000 positrons/bunch
 - Below mm precision in x-y coordinates
 - Better than 10% charge measurement
- Polycrystalline diamonds 50-100 mm thickness:
 - 16x1mm² strip and x-y readout in a single detector
 - Readout strips are graphitized by using a laser to avoid metallization
 - PADME prototype 50 μm thick, 20×20mm² sample produced and tested on beam







50 µm, silver painter Estimated CCD=10-

The diamond target



- Bonding to the readout board
- Connect to amplifier
- Digitize signal

- Step motor to move target in and out of the beam
- Possibly, add a (thick) Silicon pixel detector in order to have a more accurate transverse image of the incoming positron beam







The diamond target



Beam-test, end of 2015



Digitized strips signals



Center of gravity

Profile X Event - Charge Integral

Average position





Mean FMS y²/ ndf Constan Mean Scota 9.277 2.299 15.8.19 34.95 + 2.96 5.20 + 0.28 2.697 + 0.174

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Positron veto

- Time resolution better than 500ps
- Momentum resolution of few % based on impact position
- Efficiency better than 99.5% for MIPs
- Low energy part inside the magnet gap
- High energy part close to not interacting beam









Positron veto



Low momentum losses are reduced for E_{γ} <400 MeV Interesting positron energy starting at ~150 MeV

Which granularity?

- 1 cm scintillator bars, readout by SiPM
- Few % momentum resolution in a large part of the spectrum





PADME TDAQ

- Readout based on digitizers CAEN V1742
- ~1000 channels
- ~33 FADC boards involved



- Trigger and clock distribution to the 33 boards
- Online FADC zero suppression (L0)
- FADC boards synchronization to few 100ps needed



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Monte Carlo simulation







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Search in annihilation production









Decay to invisible signal selection

- Only one cluster in calorimeter...
 - Rejects $e+e^- \rightarrow \gamma \gamma$, $e+e^- \rightarrow \gamma \gamma(\gamma)$ final states
- ... in the fiducial region 30 mrad < θ_{CI} < 65 mrad
 - Ensure shower containment, σ(E)/E
- Positron veto: no tracks in ±2 ns
 - Reject background from Bremsstrahlung identifying primary positrons
- Photon veto: no γ with E_γ>50 MeV in time in ±1ns in the small angle veto (SAV), covering the hole acceptance
- Cluster energy within: E_{min}(M_{A'}) < E_{Cluster} < E_{max}(M_{A'})
 - Removes low energy bremsstrahlung photons and piled up clusters
- Missing mass in the region: M²_{miss} ± σ(M²_{miss})







Signal vs. background



Resolution is the result of combination of angular resolution, energy resolution and angle-energy correlation due to production







Background estimates



- Main background sources are: $e^+e^- \rightarrow \gamma\gamma$, $e^+e^- \rightarrow \gamma\gamma(\gamma)$, $e^+N \rightarrow e^+N\gamma$, **Pile-up**
- Pile-up contribution is important but rejected by the maximum cluster energy cut and M²_{miss}
- Veto inefficiency at high missing mass (E_{e+} ≃ E_{beam})
 - Additional Veto detector introduced to reject residual background





Missing mass resolution



- Improvement mainly due to better angular resolution when the calorimeter distance increase
- Depending on dark photon mass angular resolution is no more the dominant contribution to the M²_{miss} resolution and the improvement is reduced (smaller for higher dark photon masses, i.e. lower energy γ)
- Impact of beam angular divergence to be taken into account!





Improving BG rejection



- Move the front veto detector inside the beam or add a new detector
- Thanks to the dispersion included by the dipole magnetic field, a high resolution front detector can enhance the background rejection for events with a positron emitting a soft photon







ala 666.944

PADME-invisible decay sensitivity



- Based on 2.5x10¹⁰ fully GEANT4 simulated 550MeV e+ on target events
 - Number of BG events is extrapolated to 1x10¹³ electrons on target
- Using N(A'γ)=s(N_{BG})
- δ enhancement factor $\delta(M_{A'}) = \sigma(A' \gamma)/\sigma(\gamma\gamma)$ with $\epsilon=1$

$$\frac{\Gamma(e^+e^- \to U\gamma)}{\Gamma(e^+e^- \to \gamma\gamma)} = \frac{N(U\gamma)}{N(\gamma\gamma)} * \frac{Acc(\gamma\gamma)}{Acc(U\gamma)} = \epsilon^2 * \delta$$

PADME 2 years of data taking at 50% efficiency with bunch length of 40 ns 10^{13} EOT = 6000 e⁺/bunch × 3.1·10⁷s · 49 Hz

PADME can explore in a *model-independent way the* favourite by $(g-2)_{\mu}$ band up to $M^{2}_{A'} = 2m_{e}E_{e+}$

 E_{e+} =550 MeV: $M_{A'}$ < 23.7 MeV/ c^2

 E_{e+} =750 MeV: $M_{A'}$ < 27.7 MeV/ c^2

 E_{e+} =1 GeV: $M_{A'}$ < 32 MeV/ c^2





The $\gamma\gamma$ normalization selection



$$N_{\gamma\gamma}^{tot} = \frac{N_{\gamma\gamma}}{Acc_{\gamma\gamma}} = Flux(e^+) \cdot \sigma_{\gamma\gamma}$$

- Number of calorimeter clusters = 2
- Cluster energy: 100MeV<E_{cl}<400 MeV
- Cluster radial position cut
- $\gamma\gamma$ invariant mass 20 MeV < M_{$\gamma\gamma$} < 26 MeV

$$M_{\gamma\gamma} = \frac{\sqrt{[(X_{\gamma 1} - X_{\gamma 2}) + (Y_{\gamma 1} - Y_{\gamma 2})]E_{\gamma 2}E_{\gamma 2}}}{Z_{EMcal} - Z_{Target}}$$

- Acceptance_{$\gamma\gamma$} = 7%
- Contamination from Bremsstrahlung < 1‰





Search in bremsstrahlung production









Visible search experiment



- Search for the process: $e^-N \rightarrow Ne^-A' \rightarrow Ne^-e^-e^+$
- 750 MeV electron beam on a ~0.5 mm tungsten target
- Measure in the spectrometer only the $P_{e^-}^4 P_{e^+}^4$
- Compute the $M_{A'}^2 = (P_{e^-}^4 + P_{e^+}^4)^2$ and decay vertex position
 - Search for peaks in the e^+e^- invariant mass







Indication on visible decay sensitivity



- Production cross section calculated with MADGraph code
- Final state is more constrained by invariant mass of the e⁺e⁻ pair
- Indication of a limit down to $\varepsilon^2 \sim 10^{-7}$ is expected at low masses
 - Density of tracks in the spectrometer is the crucial point to be clarified
 - Design of the spectrometer not yet finalized







Electron dumps experiments

Experiment	target	E_0	N_{e}	$N_{ m el}$		$L_{\rm dec}$	Ν.	N
		$[\mathrm{GeV}]$	electrons	Coulomb	[m]	[m]	IVobs	1V95%up
E141 47	W	9	2×10^{15}	$0.32 \ \mathrm{mC}$	0.12	35	1126^{+1312}_{-1126}	3419
E137 48	Al	20	$1.87{ imes}10^{20}$	30 C	179	204	0	3
E774 49	W	275	5.2×10^{9}	0.83 nC	0.3	2	0^{+9}_{-0}	18
KEK [39]	W	2.5	1.69×10^{17}	$27 \mathrm{mC}$	2.4	2.2	0	3
Orsay 40	W	1.6	2×10^{16}	$3.2 \mathrm{mC}$	1	2	0	3
PADME dump	W	1.2	2 • 10²⁰	~ 30 C	~0.1	1		





High intensity







Radioprotection limit: <n> = 3.125×10¹⁰ particles/s

But...

- $\diamond\,$ Much higher charge on positron converter
- \diamond 8 A (12 A) from gun cathode

A few measurements on the maximum LINAC charge, driven by beam-dump experiments requirements







Bunch charge vs. length



Trying to put all together: WCM readout saturated at 16 nC...







How many electrons on target?

- Let's compute how many eot/y^{*} for 10 nC/pulse so we can scale easily with the charge available from the LINAC
 - **10 nC** = $10^{-8}/1.6 \times 10^{-19}$ = **6.25**×10¹⁰
 - At 49 Hz (1 pulse to spectrometer line) = 3×10^{12} e/s
 - 2 orders of magnitude more than present BTF authorization
 - Standard year = 1 y^{*} = 120 days at 100% efficiency (10⁷ s)
 - 3.175×10¹⁹ eot/y*
- 25 nC translates in 0.8×10²⁰ eot/y*
 - Considering <u>measurements</u> at 725 MeV, 40 ns, in the **present LINAC** configuration and quite conservative assumptions
 - Further extension of the pulse to 150 ns seems feasible with the present RF configuration, and should bring us to ≈100 nC, i.e. 3×10²⁰ eot/y^{*}

Where can we dump 3×10^{12} to 3×10^{13} e/s ?






LINAC beam dump



50 cm





Dark photon searches with PADME

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30 cm

LINAC beam dump



DHPTT02



DR pumps hall



PADME dump toy Monte Carlo

- Try to evaluate driving design parameters for the PADME dump
- Toy MC includes:
 - Production cross section calculated by MADgraph

$$\begin{aligned} \frac{d\sigma_{\gamma'}}{dx_e \ d\cos \theta_{\gamma'}} \ &= \ 8\alpha^3 \chi^2 E_e^2 x_e \ \xi(E_e, m_{\gamma'}, Z, A) \ \sqrt{1 - \frac{m_{\gamma'}^2}{E_e^2}} \\ & \left[\frac{1 - x_e + \frac{x_e^2}{2}}{U^2} + \frac{(1 - x_e)^2 m_{\gamma'}^4}{U^4} - \frac{(1 - x_e) x_e m_{\gamma'}^2}{U^3} \right], \end{aligned}$$

Evaluate the produced number of dark photons

$$N_{\gamma'} \ = \ \sigma_{\gamma'} \ N_e \ n_{\rm sh} \ L_{\rm sh} \ = \ \sigma_{\gamma'} \ N_e \ \frac{N_0}{A} \ \rho_{\rm sh} \ L_{\rm sh} \ , \qquad \qquad \frac{dP(l)}{dl} = \frac{1}{l_{\gamma'}} \ e^{-l/l_{\gamma'}}$$

- Scale by decay length acceptance
- Scale by electron acceptance in the detector using kinematical distribution from a toy MC
 - Distribution have been compared with MADGraph for several M_{A^\prime}
- Not yet implemented in depth production of the A'







PADME dump main parameters <u>±10 cm at 1m</u> ±10 cm at 1m



Decay length acceptance applied



Acceptance as function of MU



Electron angular acceptance



Dump comparison

1•10²⁰, 1.2 GeV electrons; 20 cm aperture at 50 cm from 8 cm W dump Zero background hypothesis, in depth production to be refined, not yet a sensitivity plot



BDX @ LNF



Beam energy **1.2 GeV** (e⁻) CsI detector 60×60×225 cm³ built with crystals from dismounted BaBar ECal?





ALP physics at PADME

Primakoff



PADME can search for invisible decaying or long living ALP by searching for $1 \gamma + M_{miss}^2$ final states

Bremsstrahlung



Annihilation



INFN Physrev D 38 11 1998

In the visible final state a->γγ all production mechanisms can be explored extending the mass range in the region of ~100MeV

The observables at PADME will be: eγγ or γγγ



nnecy Seminar

Limits on ALPs coupling to photons



Energy upgrade



BTF upgrade

 10^{-4} a and is excluded OE KI 10-5 10⁻⁵ BaBar a_{μ} A' is 'welcome' APEX MAMI 10-6 PADME Energy 10⁻⁶ 1013 eot 10-7 ε2 ε^{2} Energy E774 10⁻⁷ Intensity **Duty cycle** 10^{-8} IIII E141 **10**⁻⁸ 10-9 ŧ PADME dump E137 10¹⁶ eqt Energy 10-10 **Intensity** 10⁻⁹ 10⁻³ Ē $\textit{M}_{\textit{A}'}(\text{GeV/c}^2)^{10^{-1}}$ 10^{-2}

Decays to lepton pairs

Decays to invisible















Add 4 sections + 2 SLED-ed klystrons







Add 4 sections + 4 SLED-ed klystrons





"Low cost" energy upgrade



The reduction factor depends on the energy selection: E and ΔE

Can we reach 10⁴ e⁺/40 ns with acceptable energy spread and spot size?





Test of positron beam

 E_0 =510 MeV electrons, Q=1 nC

E=447 MeV=88% *E*₀, N≈2000 positrons

10 ns bunch width

BTF target at 1.7 x_0

 σ_{v}

- ≈0.8 mm (2 mm FWHM)
- Dominated by multiple scattering on 0.5 mm Be window + 20 cm of air
- Can be pushed down to 0.6-0.7 mm (with Be window)
- Can be further improved operating in vacuum ۲
- σ,
- Dominated by momentum spread, TB2 slits (before selecting dipole) + TB4 (after)
- Can be improved by using an optimized (thinner) target and by closing the slits
- A thinner target also allows to run closer to the primary energy



FITPIX 15×15 mm





PADME status

- The PADME construction phase is just starting
 - Magnet delivered at LNF, being mapped
 - First diamond target tests (20x20mm² e 50um di spessore)
 - First calorimeter test beam with just 3x3 crystal matrix
- Approval status
 - Project has been endorsed by both LNF scientific committee and INFN CSN1
 - Crystals from L3 collaborations obtained
 - 2016 budget secured in INFN CSN1 by "What Next" program
- New physics cannels for PADME identified (ALP)
- Refinement of Monte Carlo and sensitivity estimates ongoing
- Additional dark sector searches identified waiting to be explored





A new hope

- Growing interest in the low mass dark matter/hidden sectors/portals
- Careful with assumptions and model-dependent limits!
- Lots of limits from old experiments re-analysis
- JLAB strong investment
- Opportunities with Frascati electron/positron beam
- PADME-invisible could start at the beginning of 2018







