

Perspectives at the Superflavor Factories



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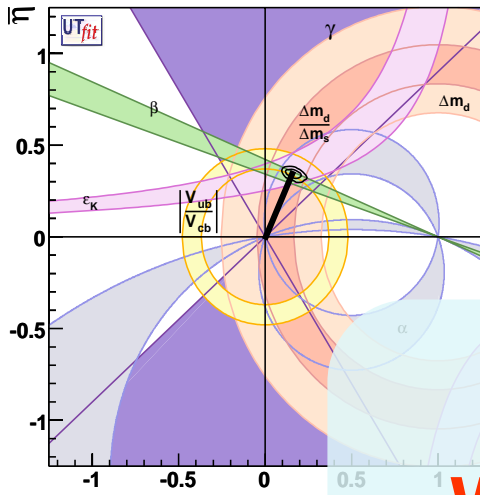
Flavour Physics was recently in the front line of the stage thanks to the e⁺ e⁻ factories.

1. 2001 CP violation in B sector from Time dependent analysis
2. 2003 Discovery of new charm particles start a new spectroscopy season with D_{sJ}
3. 2004 Direct CP violation in B sector
4. 2006 Limits on MSSM parameters from B → τν
5. 2007 Charm mixing from D⁰ D⁰bar decay
6. 2008 η_b discovery



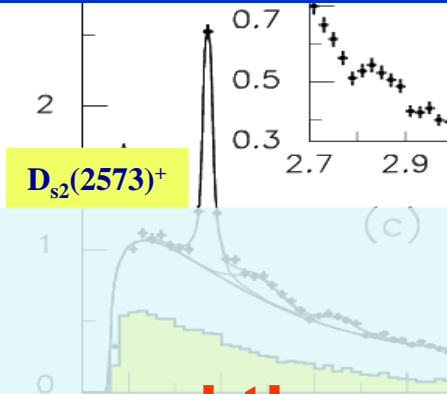
What in the next decade will e^+e^- factories say about new physics and contribute as complementary to LHC and LHCb ?

Unitarity Triangle precision measurements

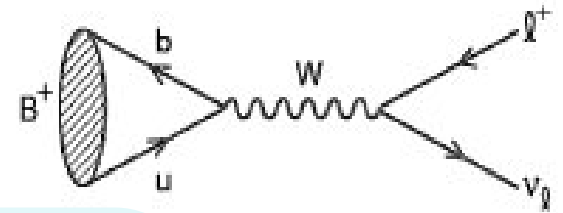


Spectroscopy of new, unexpected states

New DK state(s) at $2.86\text{GeV}/c^2$

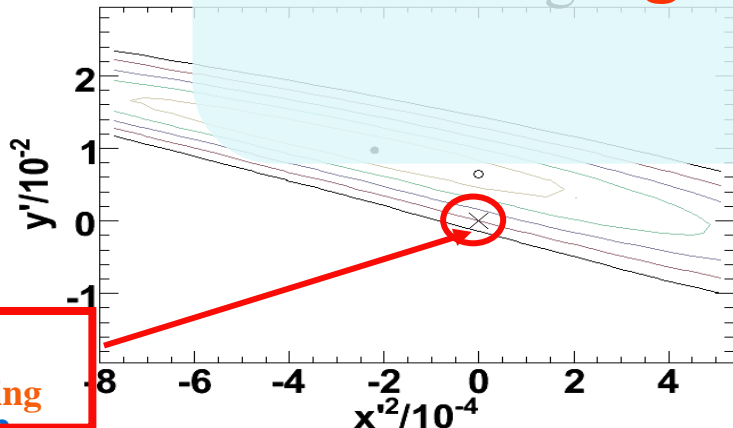


$B \rightarrow \tau \nu$ setting limits on MSSM parameters

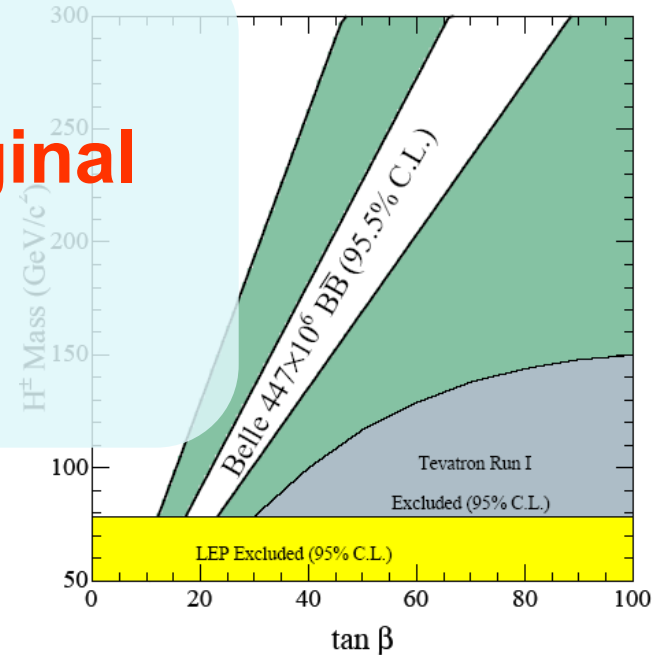


Well beyond the original goals

$D^0 - \bar{D}^0$ mixing



No Mixing



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Constraints on CKM

Precise measurements of CKM parameters in quark sector, challenging new physics can come from a wide spectrum of future projects (some approved, some close to approval, some....):

In Kaons there are projects for a next generation of experiments : $K^+ \rightarrow \pi^+ \nu \bar{\nu}$ at Cern(P326), Kloe at upgraded DaΦne, (a $K^0 \pi^0$ like experiment $K_L^0 \rightarrow \pi^0 \nu \bar{\nu}$????).

In charm, τ and b:

- BES upgrade, KEKB upgrade,
- Super c- τ at Novosibirsk and
- SuperB

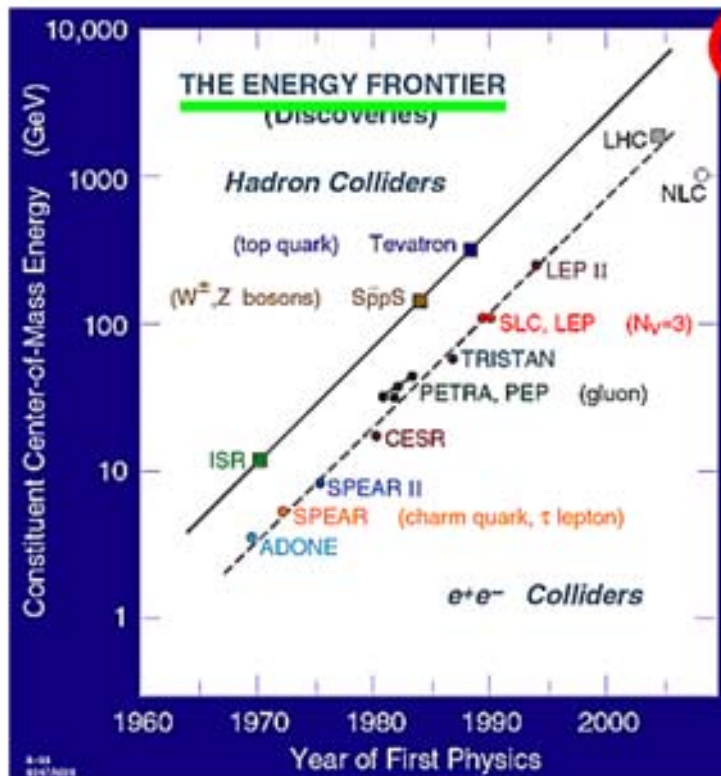
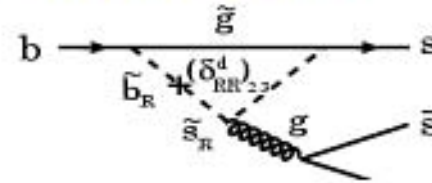


Two frontiers for the exploration of New Physics

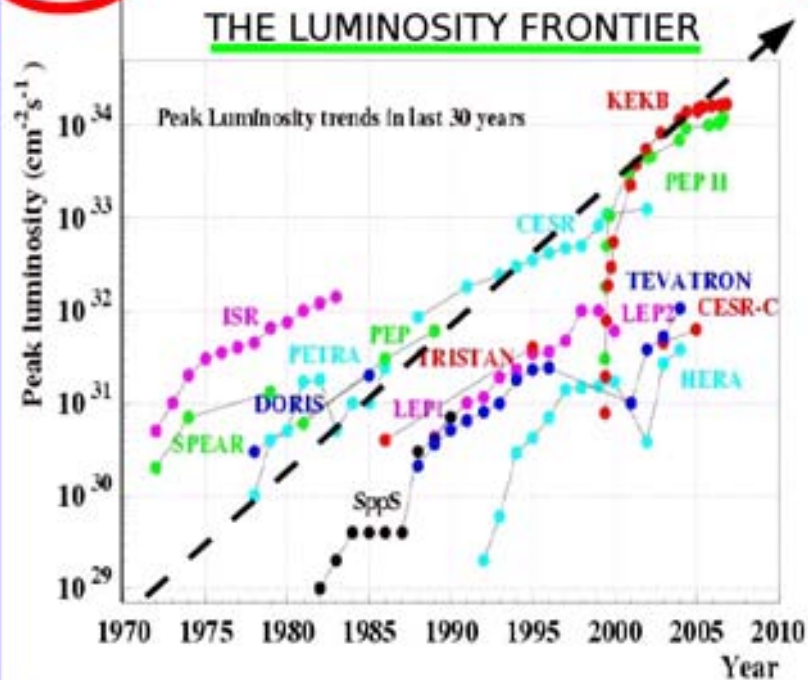
Relativistic way



Quantum way



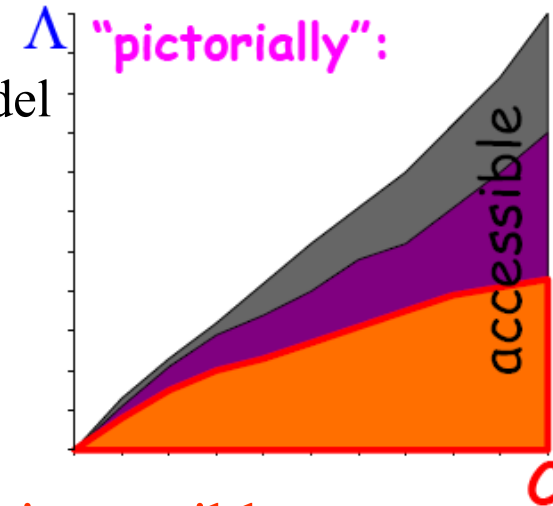
10^{36}



10^{36}



- Flavour precision measurements sensitive to New Physics (NP)
 - Measure interference effect in known processes
 - Measure decays: rare or forbidden in Standard Model
- NP effects governed by
 - New Physics Scale NP(Λ)
 - Effective coupling C
 - Different Intensities (from interactions)
 - Different Patterns (for instance from symmetries)



With 7-10x10¹⁰ pair bb, cc, $\tau\tau$ (75-100 ab⁻¹) it is possible

NP(Λ) found at LHC

- Determine couplings FV e CPV of NP
- Look for heavier states
- Study the flavour structure of NP

NP(Λ) not found at LHC

- Look for indirect signals of NP
- Link them to explaining NP models
- Constrain regions in parameter space with NP(Λ) sensitivity up several tens of TeV.

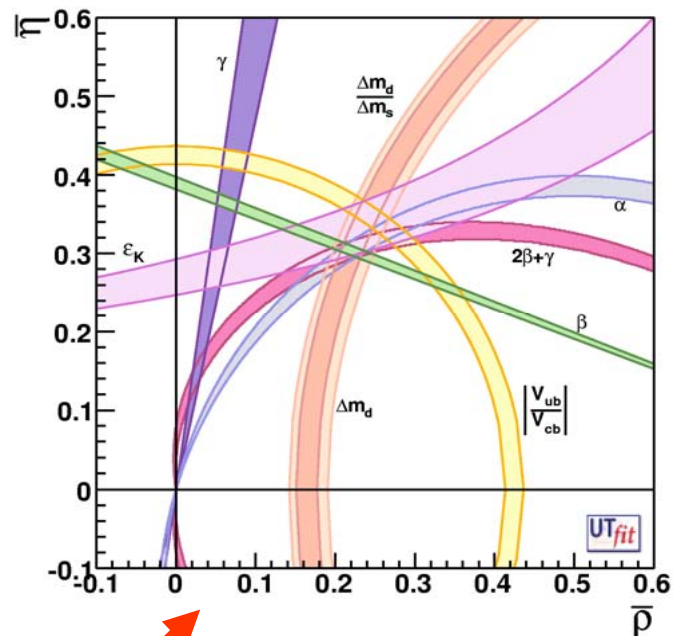
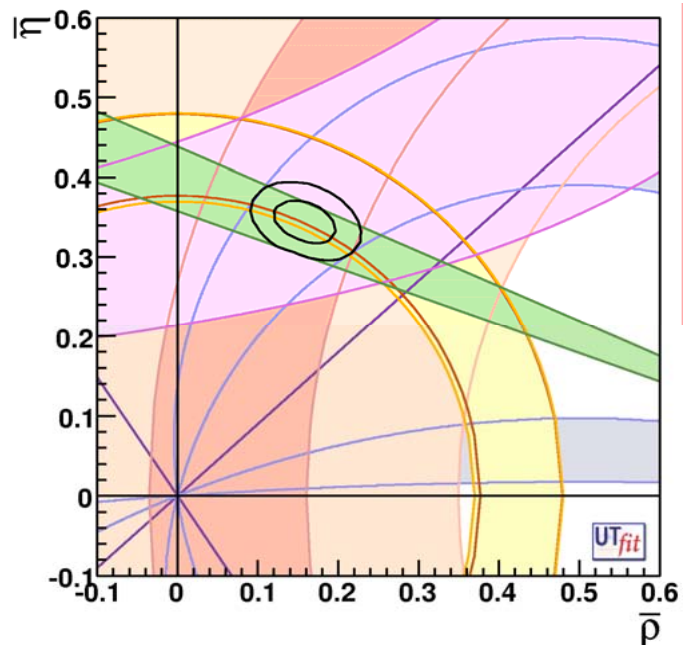
Some channels as τ LFV clear signals of NP



NP sensitivity with 75 ab^{-1}

Today

SuperB+Lattice improvements



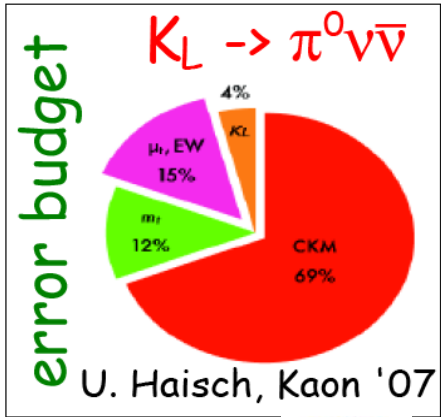
$$\rho = 0.163 \pm 0.028$$

$$\eta = 0.344 \pm 0.016$$

$$\rho = \pm 0.0028$$

$$\eta = \pm 0.0024$$

Improving CKM is crucial to look for NP



B Physics @ Y(4S)

	B Factories (2 ab ⁻¹)	SuperB (75 ab ⁻¹)	Observable	B Factories (2 ab ⁻¹)	SuperB (75 ab ⁻¹)
$\sin(2\beta) (J/\psi K^0)$	0.018	0.005 (†)	$ V_{cb} $ (exclusive)	4% (*)	1.0% (*)
$\cos(2\beta) (J/\psi K^{*0})$	0.30	0.05	$ V_{cb} $ (inclusive)	1% (*)	0.5% (*)
$\sin(2\beta) (Dh^0)$	0.10	0.02	$ V_{ub} $ (exclusive)	8% (*)	3.0% (*)
$\cos(2\beta) (Dh^0)$	0.20	0.04	$ V_{ub} $ (inclusive)	8% (*)	2.0% (*)
$S(J/\psi \pi^0)$	0.10	0.02	$\mathcal{B}(B \rightarrow \tau\nu)$	20%	4% (†)
$S(D^+D^-)$	0.20	0.03	$\mathcal{B}(B \rightarrow \mu\nu)$	visible	5%
$S(\phi K^0)$	0.13	0.02 (*)	$\mathcal{B}(B \rightarrow D\tau\nu)$	10%	2%
$S(\eta' K^0)$	0.05	0.01 (*)	$\mathcal{B}(B \rightarrow \rho\gamma)$	15%	3% (†)
$S(K_S^0 K_S^0 K_S^0)$	0.15	0.02 (*)	$\mathcal{B}(B \rightarrow \omega\gamma)$	30%	5%
$S(K_S^0 \pi^0)$	0.15	0.02 (*)	$A_{CP}(B \rightarrow K^*\gamma)$	0.007 (†)	0.004 († *)
$S(\omega K_S^0)$	0.17	0.03 (*)	$A_{CP}(B \rightarrow \rho\gamma)$	~ 0.20	0.05
$S(f_0 K_S^0)$	0.12	0.02 (*)	$A_{CP}(b \rightarrow s\gamma)$	0.012 (†)	0.004 (†)
$\gamma (B \rightarrow DK, D \rightarrow CP \text{ eigenstates})$	~ 15°	2.5°	$A_{CP}(b \rightarrow (s+d)\gamma)$	0.03	0.006 (†)
$\gamma (B \rightarrow DK, D \rightarrow \text{suppressed states})$	~ 12°	2.0°	$S(K_S^0 \pi^0 \gamma)$	0.15	0.02 (*)
$\gamma (B \rightarrow DK, D \rightarrow \text{multibody states})$	~ 9°	1.5°	$S(\rho^0 \gamma)$	possible	0.10
$\gamma (B \rightarrow DK, \text{combined})$	~ 6°	1-2°	$A_{CP}(B \rightarrow K^* \ell \ell)$	7%	1%
$\alpha (B \rightarrow \pi\pi)$	~ 16°	3°	$A^{FB}(B \rightarrow K^* \ell \ell)_{s_0}$	25%	9%
$\alpha (B \rightarrow \rho\rho)$	~ 7°	1-2° (*)	$A^{FB}(B \rightarrow X_s \ell \ell)_{s_0}$	35%	5%
$\alpha (B \rightarrow \rho\pi)$	~ 12°	2°	$\mathcal{B}(B \rightarrow K \nu \bar{\nu})$	visible	20%
$\alpha (\text{combined})$	~ 6°	1-2° (*)	$\mathcal{B}(B \rightarrow \pi \nu \bar{\nu})$	-	possible
$2\beta + \gamma (D^{(*)\pm} \pi^\mp, D^\pm K_S^0 \pi^\mp)$	20°	5°			

Charm mixing and CP

Mode	Observable	$\Upsilon(4S)$ (75 ab ⁻¹)	$\psi(3770)$ (300 fb ⁻¹)
$D^0 \rightarrow K^+ \pi^-$	x'^2	3×10^{-5}	
	y'	7×10^{-4}	
$D^0 \rightarrow K^+ K^-$	y_{CP}	5×10^{-4}	
$D^0 \rightarrow K_S^0 \pi^+ \pi^-$	x	4.9×10^{-4}	
	y	3.5×10^{-4}	
	$ q/p $	3×10^{-2}	
	ϕ	2°	
$\psi(3770) \rightarrow D^0 \bar{D}^0$	x^2		$(1-2) \times 10^{-5}$
	y		$(1-2) \times 10^{-3}$
	$\cos \delta$		(0.01-0.02)

Charm FCNC

	Sensitivity
$D^0 \rightarrow e^+ e^-, D^0 \rightarrow \mu^+ \mu^-$	1×10^{-8}
$D^0 \rightarrow \pi^0 e^+ e^-, D^0 \rightarrow \pi^0 \mu^+ \mu^-$	2×10^{-8}
$D^0 \rightarrow \eta e^+ e^-, D^0 \rightarrow \eta \mu^+ \mu^-$	3×10^{-8}
$D^0 \rightarrow K_S^0 e^+ e^-, D^0 \rightarrow K_S^0 \mu^+ \mu^-$	3×10^{-8}
$D^+ \rightarrow \pi^+ e^+ e^-, D^+ \rightarrow \pi^+ \mu^+ \mu^-$	1×10^{-8}
$D^0 \rightarrow e^\pm \mu^\mp$	1×10^{-8}
$D^+ \rightarrow \pi^+ e^\pm \mu^\mp$	1×10^{-8}
$D^0 \rightarrow \pi^0 e^\pm \mu^\mp$	2×10^{-8}
$D^0 \rightarrow \eta e^\pm \mu^\mp$	3×10^{-8}
$D^0 \rightarrow K_S^0 e^\pm \mu^\mp$	3×10^{-8}
$D^+ \rightarrow \pi^- e^+ e^+, D^+ \rightarrow K^- e^+ e^+$	1×10^{-8}
$D^+ \rightarrow \pi^- \mu^+ \mu^+, D^+ \rightarrow K^- \mu^+ \mu^+$	1×10^{-8}
$D^+ \rightarrow \pi^- e^\pm \mu^\mp, D^+ \rightarrow K^- e^\pm \mu^\mp$	1×10^{-8}

τ Physics

Observable	Sensitivity
$\mathcal{B}(\tau \rightarrow \mu \gamma)$	2×10^{-9}
$\mathcal{B}(\tau \rightarrow e \gamma)$	2×10^{-9}
$\mathcal{B}(\tau \rightarrow \mu \mu \mu)$	2×10^{-10}
$\mathcal{B}(\tau \rightarrow e e e)$	2×10^{-10}
$\mathcal{B}(\tau \rightarrow \mu \eta)$	4×10^{-10}
$\mathcal{B}(\tau \rightarrow e \eta)$	6×10^{-10}
$\mathcal{B}(\tau \rightarrow \ell K_S^0)$	2×10^{-10}

B_s Physics @ Y(5S)

Observable	B Factories (2 ab ⁻¹)	SuperB (75 ab ⁻¹)
$\Delta\Gamma$	0.16 ps ⁻¹	0.03 ps ⁻¹
Γ	0.07 ps ⁻¹	0.01 ps ⁻¹
β_s from angular analysis	20°	8°
A_{SL}^s	0.006	0.004
A_{CH}	0.004	0.004
$\mathcal{B}(B_s \rightarrow \mu^+ \mu^-)$	-	$< 8 \times 10^{-9}$
$ V_{td}/V_{ts} $	0.08	0.017
$\mathcal{B}(B_s \rightarrow \gamma\gamma)$	38%	7%
β_s from $J/\psi\phi$	10°	3°
β_s from $B_s \rightarrow K^0 \bar{K}^0$	24°	11°

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SOME EXAMPLES

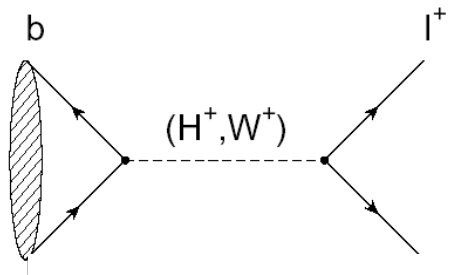
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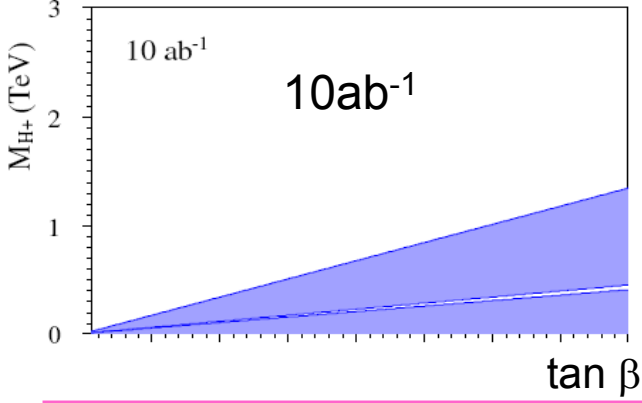
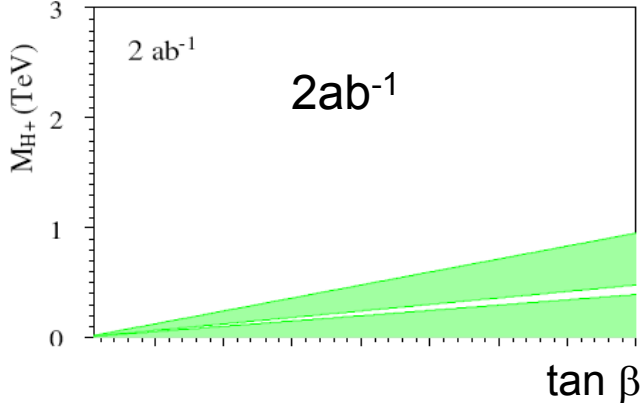


Higgs-mediated NP in MFV at large $\tan\beta$

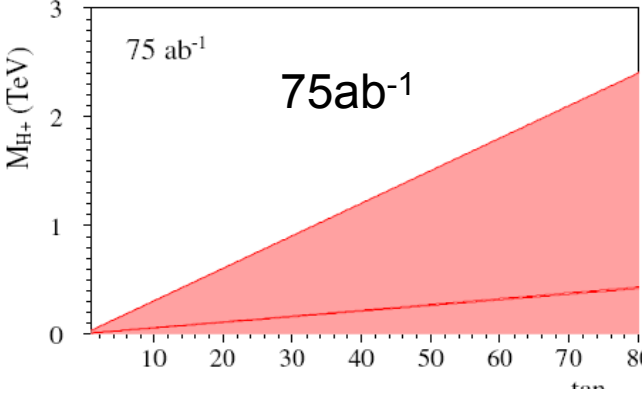


$$BR(B \rightarrow \tau \nu) = BR_{SM}(B \rightarrow \tau \nu) \left(1 - \frac{m_B^2}{M_H^2} \tan^2 \beta \right)^2$$

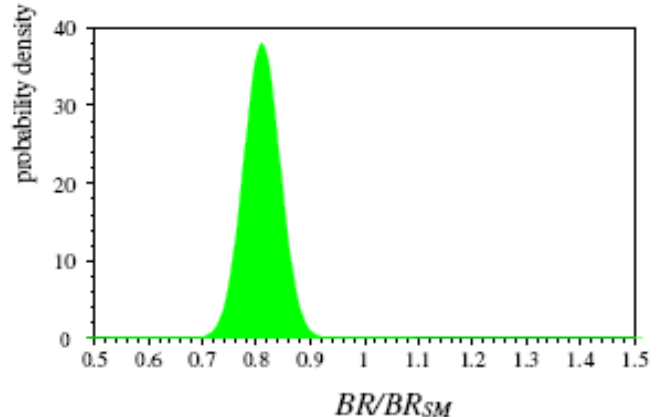
2ab⁻¹
 $M_H \sim 0.4 - 0.8$ TeV
 for $\tan\beta \sim 30 - 60$



SuperB - 75ab⁻¹
 $M_H \sim 1.2 - 2.5$ TeV
 for $\tan\beta \sim 30 - 60$



**How signal would like
 with $M_H = 350$ GeV**

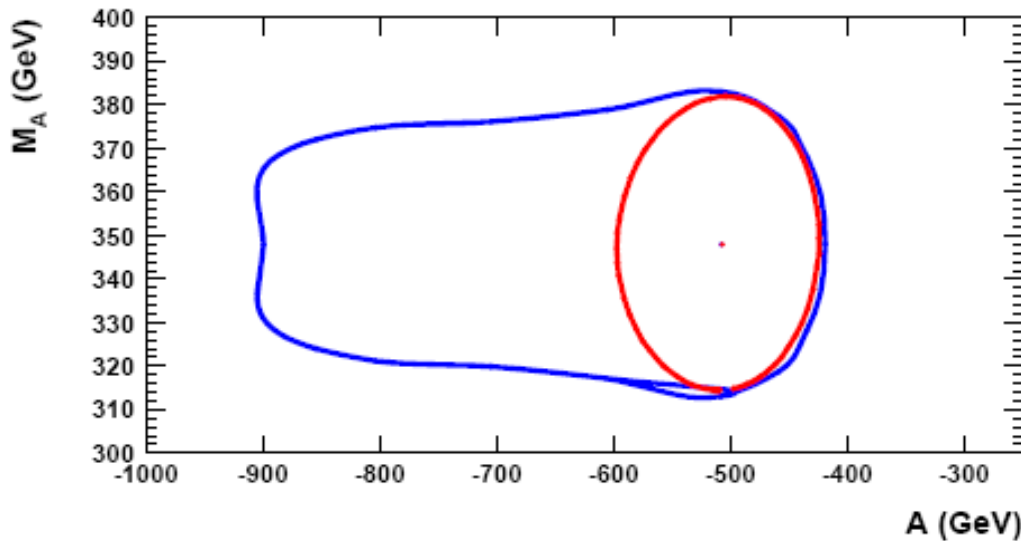
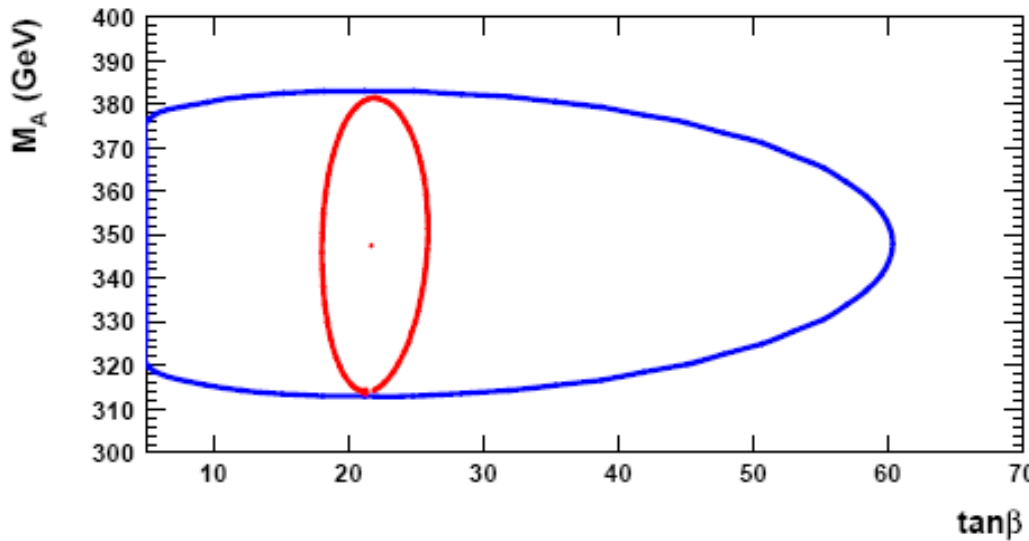


Importance of having very large sample $\geq 75ab^{-1}$



COMPLEMENTARY: LHC and Flavour with 75 ab⁻¹

IF LHC DISCOVERS
SUPERSYMMETRY



Red are LHC+EW constraints+**SuperB**

Blue is LHC alone



MFV : SNOWMASS points

SPS	$M_{1/2}$ (GeV)	M_0 (GeV)	A_0 (GeV)	$\tan\beta$	μ
1 a	250	100	-100	10	> 0
1 b	400	200	0	30	> 0
2	300	1450	0	10	> 0
3	400	90	0	10	> 0
4	300	400	0	50	> 0
5	300	150	-1000	5	> 0

	SPS1a	SPS4	SPS5
$\mathcal{R}(B \rightarrow s\gamma)$	0.919 \pm 0.038	0.248	0.848 \pm 0.081
$\mathcal{R}(B \rightarrow \tau\nu)$	0.968 \pm 0.007	0.436	0.997 \pm 0.003
$\mathcal{R}(B \rightarrow X_s l^+ l^-)$	0.916 \pm 0.004	0.917	0.995 \pm 0.002
$\mathcal{R}(B \rightarrow K\nu\bar{\nu})$	0.967 \pm 0.001	0.972	0.994 \pm 0.001
$\mathcal{B}(B_d \rightarrow \mu^+ \mu^-)/10^{-10}$	1.631 \pm 0.038	16.9	1.979 \pm 0.012
$\mathcal{R}(\Delta m_s)$	1.050 \pm 0.001	1.029	1.029 \pm 0.001
$\mathcal{B}(B_s \rightarrow \mu^+ \mu^-)/10^{-9}$	2.824 \pm 0.063	29.3	3.427 \pm 0.018
$\mathcal{R}(K \rightarrow \pi^0 \nu\bar{\nu})$	0.973 \pm 0.001	0.977	0.994 \pm 0.001

SPS4 ruled out by present values of $B \rightarrow s\gamma$.

SPS1a is the least favorable for flavour, but SuperB and only SuperB can observe 2 σ deviations in several observables



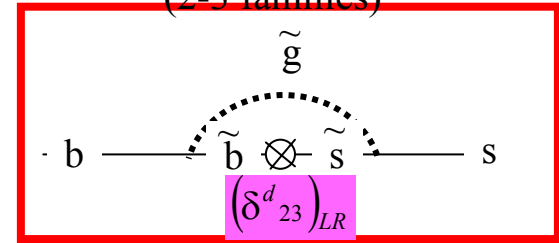
MSSM+generic soft SUSY breaking terms

Flavour-changing NP effects in the squark propagator

→ NP scale SUSY mass $\tilde{m} \sim m_{\tilde{g}}$

→ flavour-violating coupling $(\delta_{ij}^q)_{AB} \equiv \frac{(M_{ij}^2)^q}{\tilde{m}^2}$

New Physics contribution
(2-3 families)



$$|\delta_{23}|_{LR}$$

1

In the red regions the δ are measured with a significance $>3\sigma$ away from zero

$$|\delta_{23}|_{LR} = (0.026 \pm 0.005)$$

$$\text{Arg}(\delta_{23})_{LR} = (44.5 \pm 2.6)^\circ$$

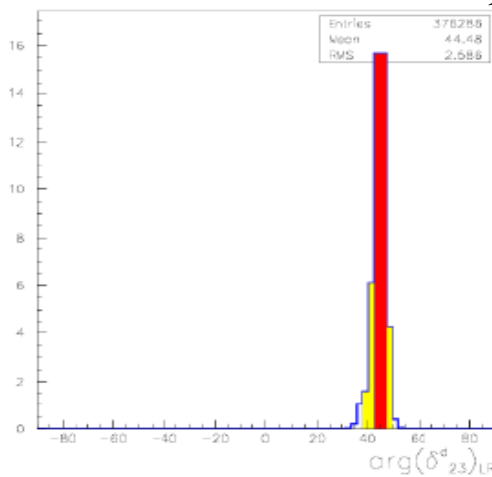
10^{-1}

10^{-2}

1

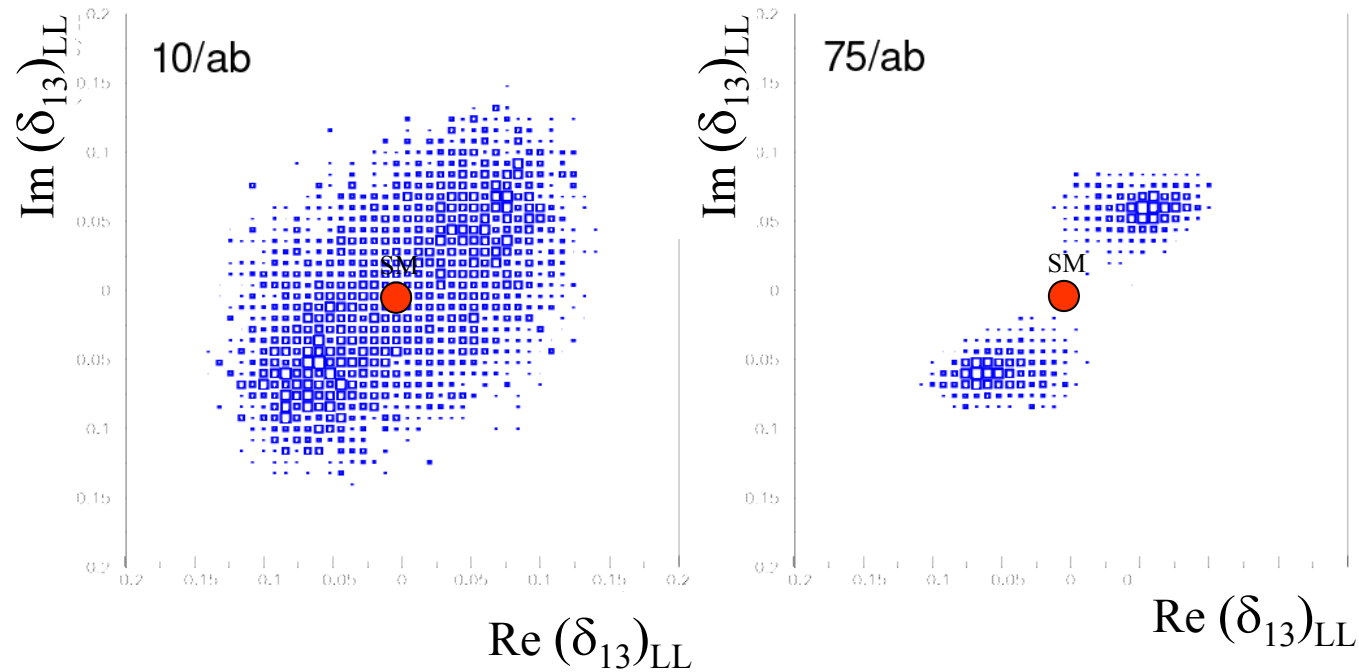
1 TeV

$10 m_{\tilde{g}} \text{ (TeV)}$



Determination of coupling [in this case : $(\delta_{13})_{LL}$]

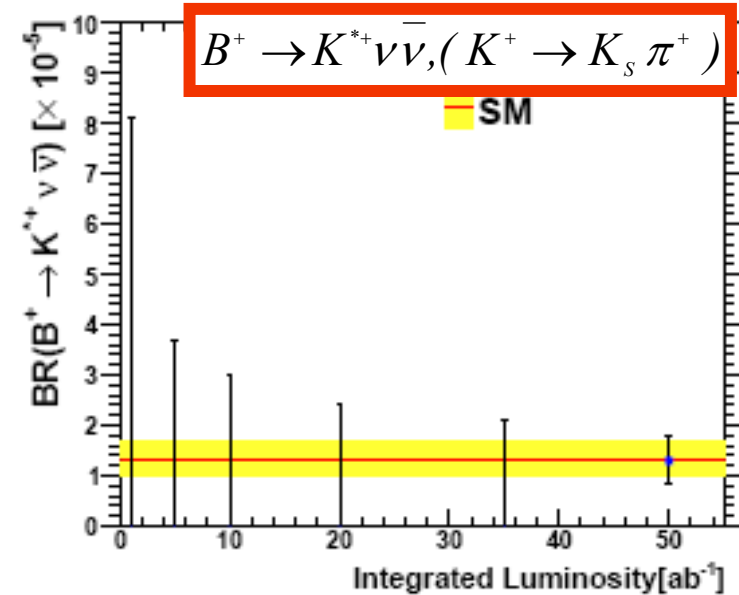
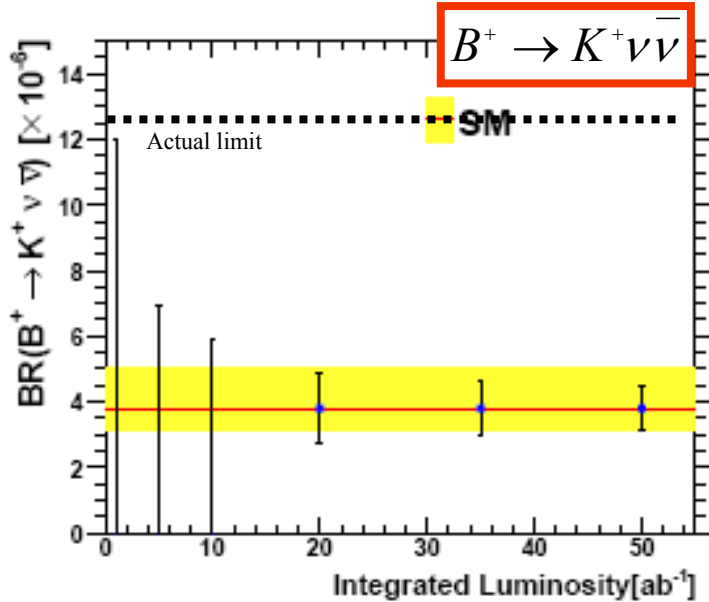
with 10 ab^{-1} and 75 ab^{-1}



Importance of having very large sample $>75\text{ab}^{-1}$



Some comparison: Current $\rightarrow 10\text{ab}^{-1} \rightarrow 75\text{ab}^{-1}$

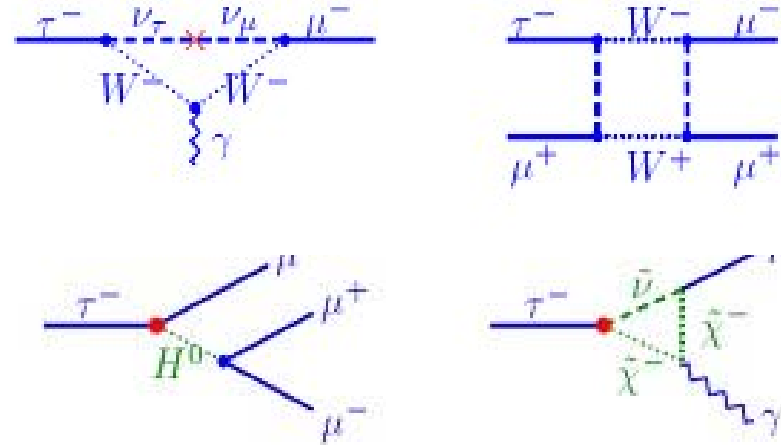


Mode	Sensitivity		
	Current	Expected (10 ab^{-1})	Expected (75 ab^{-1})
$\mathcal{B}(B \rightarrow X_s \gamma)$	7%	5%	3%
$A_{CP}(B \rightarrow X_s \gamma)$	0.037	0.01	0.004–0.005
$\mathcal{B}(B^+ \rightarrow \tau^+ \nu)$	30%	10%	3–4%
$\mathcal{B}(B^+ \rightarrow \mu^+ \nu)$	not measured	20%	5–6%
$\mathcal{B}(B \rightarrow X_s l^+ l^-)$	23%	15%	4–6%
$A_{FB}(B \rightarrow X_s l^+ l^-)_{s_0}$	not measured	30%	4–6%
$\mathcal{B}(B \rightarrow K \nu \bar{\nu})$	not measured	not measured	16–20%
$S(K_S^0 \pi^0 \gamma)$	0.24	0.08	0.02–0.03




SuperB is also a τ factory \rightarrow golden measurement LFV (Complementarity with $\mu \rightarrow e \gamma$)

Process	Sensitivity
$\mathcal{B}(\tau \rightarrow \mu \gamma)$	2×10^{-9}
$\mathcal{B}(\tau \rightarrow e \gamma)$	2×10^{-9}
$\mathcal{B}(\tau \rightarrow \mu \mu \mu)$	2×10^{-10}
$\mathcal{B}(\tau \rightarrow e e e)$	2×10^{-10}
$\mathcal{B}(\tau \rightarrow \mu \eta)$	4×10^{-10}
$\mathcal{B}(\tau \rightarrow e \eta)$	6×10^{-10}
$\mathcal{B}(\tau \rightarrow \ell K_s^0)$	2×10^{-10}




Further improvements if polarized beams.

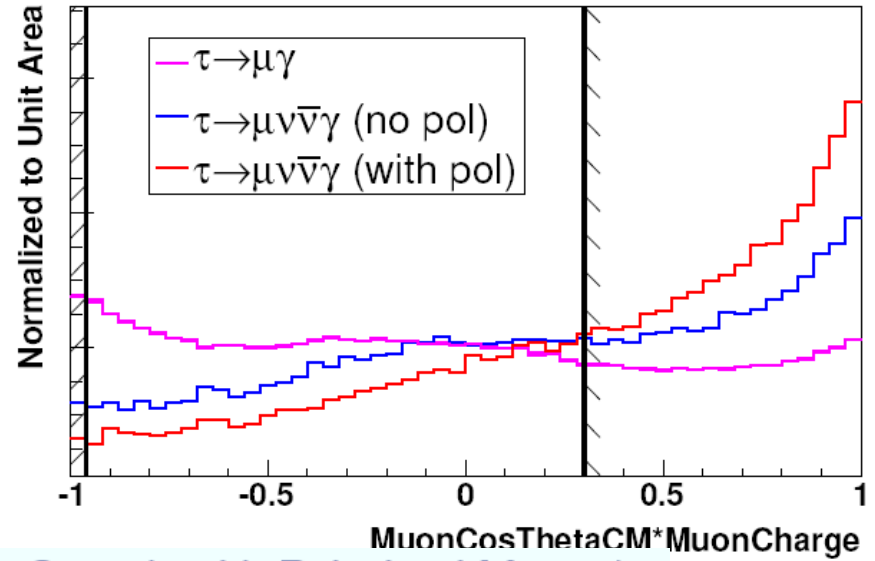
90% CL limits

$\text{Br}(\tau^- \rightarrow e^- \gamma) < 12 \times 10^{-8}$ 

$\text{Br}(\tau^- \rightarrow \mu^- \gamma) < 4.1 \times 10^{-8}$

$\text{Br}(\tau^- \rightarrow e^- \gamma) < 11 \times 10^{-8}$ 

$\text{Br}(\tau^- \rightarrow \mu^- \gamma) < 6.7 \times 10^{-8}$



Optimization of BKG rejection is in progress. Pol. Helps also to discriminate models. In some model there is a strong effect on the angular distribution of μ from signal:

(see hep-ph/9604296, Y.Kuno, Y.Okada, $\mu \rightarrow e \gamma$ Search with Polarized Muons)

Comparison with Snowmass points on Tau using also Polarization

SuperB with 75 ab-1, evaluation assuming the most conservative scenario about syst. errors

SPS	$M_{1/2}$ (GeV)	M_0 (GeV)	A_0 (GeV)	$\tan\beta$	μ
1 a	250	100	-100	10	> 0
1 b	400	200	0	30	> 0
2	300	1450	0	10	> 0
3	400	90	0	10	> 0
4	300	400	0	50	> 0
5	300	150	-1000	5	> 0

◆ NP predictions for experimentally constrained SUSY in a number of standard scenarios
B.C.Allanach *et al.*, hep-ph/0202233

<i>LFV</i>	Snowmass points predictions						SuperB	
	1 a	1 b	2	3	4	5	90% UL	5 σ disc
$\text{BF}(\tau \rightarrow \mu\gamma) \times 10^{-9}$	4.2	7.9	0.18	0.26	97	0.019	1÷2	5
$\text{BF}(\tau \rightarrow 3\mu) \times 10^{-12}$	9.4	18	0.41	0.59	220	0.043	200	880

SuperKEKB worse by a factor 2.5 and 4.5 in $\tau \rightarrow \mu\gamma$ and >5 in $\tau \rightarrow 3\mu$



$\mu \rightarrow e\gamma$ search at PSI: sensitive to 10^{-13} BRs

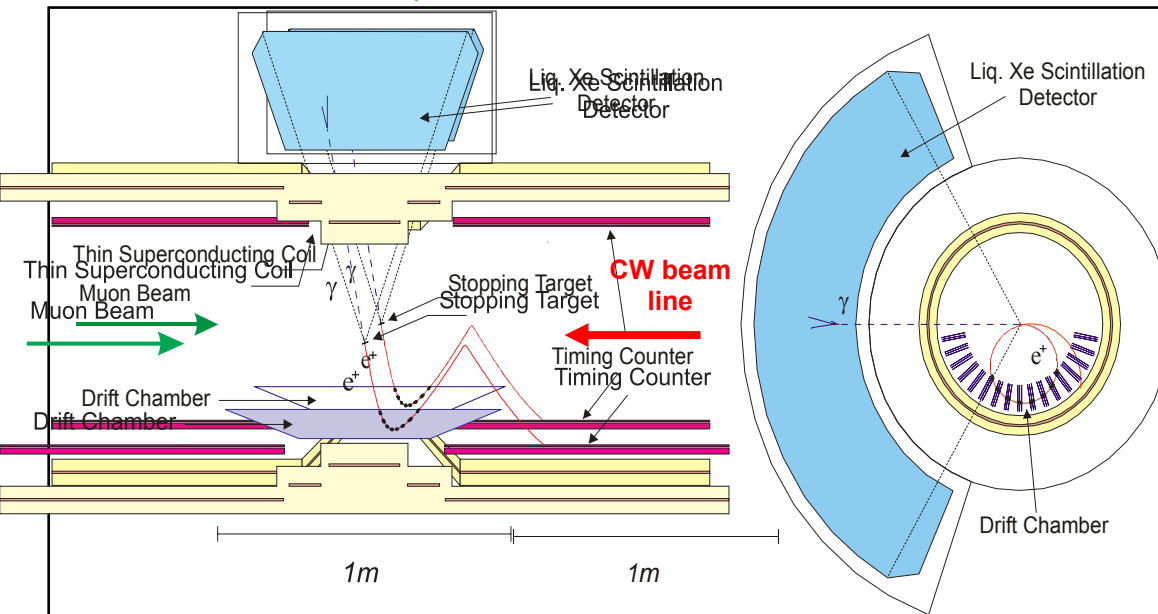
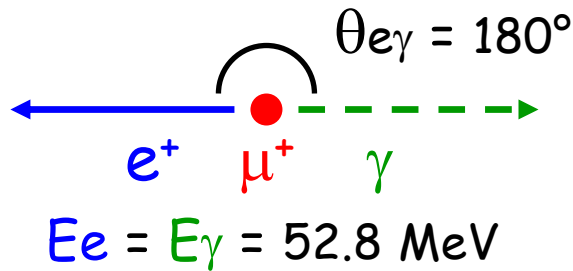
SUSY SU(5) predictions

$BR(\mu \rightarrow e\gamma) \approx 10^{-14} \div 10^{-13}$

SUSY SO(10) predictions

$BR_{SO(10)} \approx 100 BR_{SU(5)}$

Easy signal selection with μ^+ at rest



Detector outline

- Stopped beam of $3 \cdot 10^7 \mu / \text{sec}$ in a $150 \mu\text{m}$ target

Photon detector

- Liquid Xenon calorimeter for γ detection (scintillation)
 - fast: 4 / 22 / 45 ns
 - high LY: $\sim 0.8 \cdot \text{NaI}$
 - short X_0 : 2.77 cm

- 1 MeV Cockroft-Walton for continuous calorimeter calibration

Positron spectrometer

- Solenoid spectrometer & drift chambers for e^+ momentum
- Scintillation counters for e^+ timing



Start with the expt. with μ

$$\Delta a_\mu = a_\mu^{\text{exp}} - a_\mu^{\text{SM}} \approx (3 \pm 1) \times 10^{-9}$$

assume SuperB at 75 fb^{-1} , **80% e^- beam polarization**

extend to all tau decay channels

combine 2 measurement methods for $\text{Re}\{F_2\}$

studies on simulated events show no limiting syst. effects

	Snowmass points predictions						SuperB
	1 a	1 b	2	3	4	5	exp. resolution
$\Delta a_\mu \times 10^{-9}$	3.1	3.2	1.6	1.4	4.8	1.1	
$\Delta a_\tau \times 10^{-6}$	0.9	0.9	0.5	0.4	1.4	0.3	<1

SuperKEKB, without beam polarization, expected worse by factor ≈ 10 , and worse systematics

Make use of all the informations (total x-section, angular distribution, f-b asymmetry.
Measure Re and Im parts



Charm physics using the charm produced at Y(4S)

Charm physics at threshold **0.3 ab⁻¹**

Consider that running 4 month at threshold we will collect 1000 times the stat. of CLEO-C
~ 10 times of future BESIII

Strong dynamics and CKM measurements

@threshold(4GeV)

D decay form factor and decay constant @ 1%
Dalitz structure useful for γ measurement

$\xi \sim 1\%$,
exclusive $V_{ub} \sim \text{few } \%$
syst. error on γ from Dalitz Model $< 1^\circ$

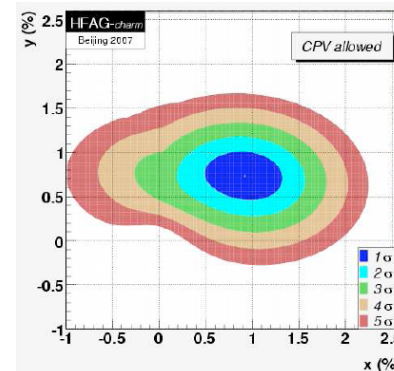
Rare decays FCNC down to 10^{-8}

@threshold(4GeV)

Channel	Sensitivity
$D^0 \rightarrow e^+e^-, D^0 \rightarrow \mu^+\mu^-$	1×10^{-8}
$D^0 \rightarrow \pi^0 e^+e^-, D^0 \rightarrow \pi^0 \mu^+\mu^-$	2×10^{-8}
$D^0 \rightarrow \eta e^+e^-, D^0 \rightarrow \eta \mu^+\mu^-$	3×10^{-8}
$D^0 \rightarrow K_s^0 e^+e^-, D^0 \rightarrow K_s^0 \mu^+\mu^-$	3×10^{-8}
$D^+ \rightarrow \pi^+ e^+e^-, D^+ \rightarrow \pi^+ \mu^+\mu^-$	1×10^{-8}
$D^0 \rightarrow e^\pm \mu^\mp$	1×10^{-8}
$D^+ \rightarrow \pi^+ e^\pm \mu^\mp$	1×10^{-8}
$D^0 \rightarrow \pi^0 e^\pm \mu^\mp$	2×10^{-8}
$D^0 \rightarrow \eta e^\pm \mu^\mp$	3×10^{-8}
$D^0 \rightarrow K_s^0 e^\pm \mu^\mp$	3×10^{-8}
$D^+ \rightarrow \pi^- e^+ e^+, D^+ \rightarrow K^- e^+ e^+$	1×10^{-8}
$D^+ \rightarrow \pi^- \mu^+ \mu^+, D^+ \rightarrow K^- \mu^+ \mu^+$	1×10^{-8}
$D^+ \rightarrow \pi^- e^\pm \mu^\mp, D^+ \rightarrow K^- e^\pm \mu^\mp$	1×10^{-8}

D mixing

Better studied using the high statistics collected at Y(4S)

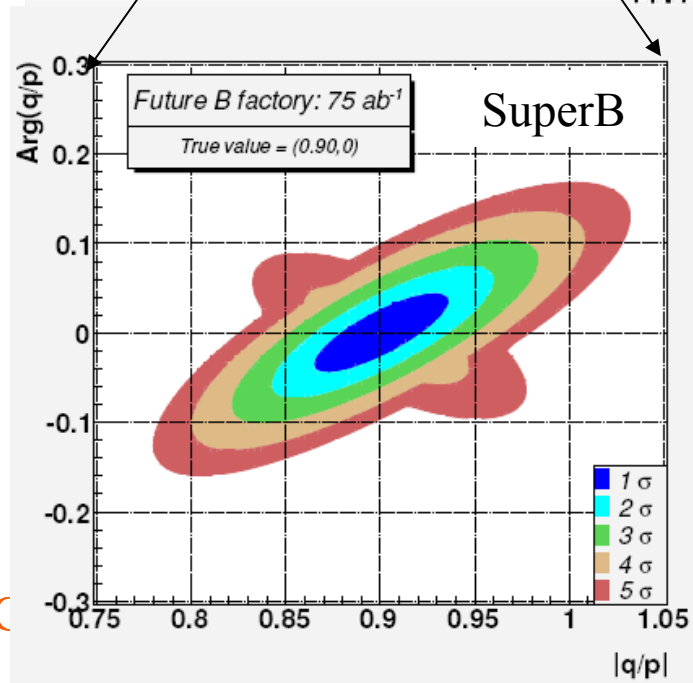
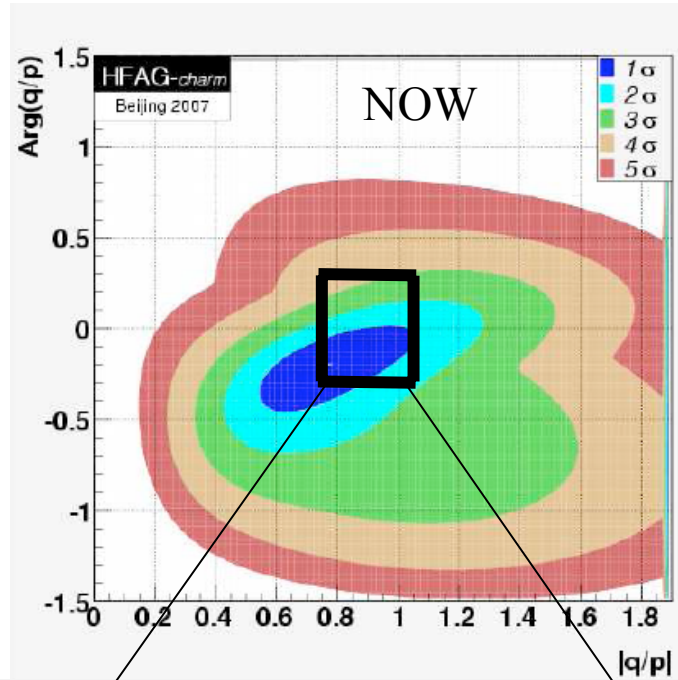


Mode	Observable	B Factories (2 ab ⁻¹)	SuperB (75 ab ⁻¹)
$D^0 \rightarrow K^+ K^-$	y_{CP}	$2-3 \times 10^{-3}$	5×10^{-4}
$D^0 \rightarrow K^+ \pi^-$	y'_D	$2-3 \times 10^{-3}$	7×10^{-4}
	x_D^2	$1-2 \times 10^{-4}$	3×10^{-5}
$D^0 \rightarrow K_s^0 \pi^+ \pi^-$	y_D	$2-3 \times 10^{-3}$	5×10^{-4}
	x_D	$2-3 \times 10^{-3}$	5×10^{-4}
Average	y_D	$1-2 \times 10^{-3}$	3×10^{-4}
	x_D	$2-3 \times 10^{-3}$	5×10^{-4}



CP Violation in charm

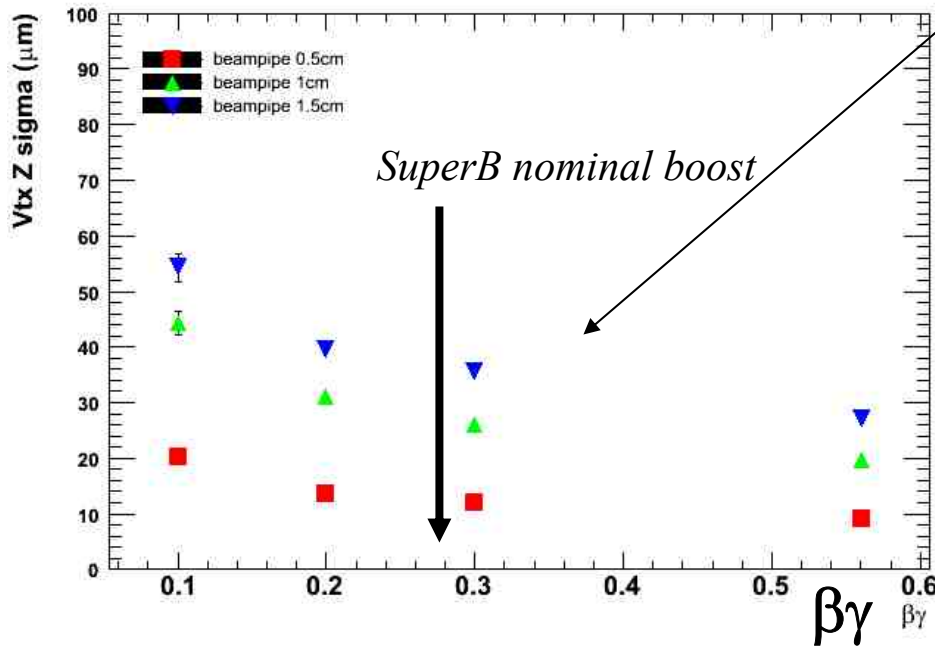
Mode	Observable	$\Upsilon(4S)$ (75 ab^{-1})	$\psi(3770)$ (300 fb^{-1})
$D^0 \rightarrow K^+ \pi^-$	x'^2	3×10^{-5}	
	y'	7×10^{-4}	
$D^0 \rightarrow K^+ K^-$	y_{CP}	5×10^{-4}	
$D^0 \rightarrow K_S^0 \pi^+ \pi^-$	x	4.9×10^{-4}	
	y	3.5×10^{-4}	
	$ q/p $	3×10^{-2}	
	ϕ	2°	
$\psi(3770) \rightarrow D^0 \bar{D}^0$	x^2		$(1-2) \times 10^{-5}$
	y		$(1-2) \times 10^{-3}$
	$\cos \delta$		$(0.01-0.02)$



- Charm events at threshold are very clean: pure DD, no additional fragmentation
- High signal/bkg ratio: optimal for decays with neutrinos.
- Quantum Coherence: new and alternative CP violation measurement wrt to $\Upsilon(4S)$. Unique opportunity to measure D^0 - D^0 relative phase.
- Increased statistics is not an advantage running at threshold: cross-section 3x wrt 10GeV but luminosity 10x smaller.
- SuperB lumi at 4 GeV = 10^{35} cm⁻²s⁻¹ produces $\sim 10^9$ DD pairs per month of running. (using Cleo-c cross-section measurement $\sigma(e^+e^- \rightarrow D^0D^0) \sim 3.6$ nb + $\sigma(e^+e^- \rightarrow D^+D^-) \sim 2.8$ nb ~ 6.4 nb)
- Time-dependent measurements at 4 GeV only possible at SuperB.



- Proper time resolution dominated by decay vertex resolution.
 - Production vertex precisely determined thanks to nm beamspot dimensions





SuperB lumi at 4 GeV = $10^{35} \text{ cm}^{-2}\text{s}^{-1}$
 produces $\sim 10^9$ $D\bar{D}$ per month

$\beta\gamma_{ct} = 0.28 \times 120 \text{ } \mu\text{m} \sim 30 \mu\text{m}$
 Average flight distance similar to
 vertex resolution $\rightarrow \sigma_{\tau} \sim \tau$

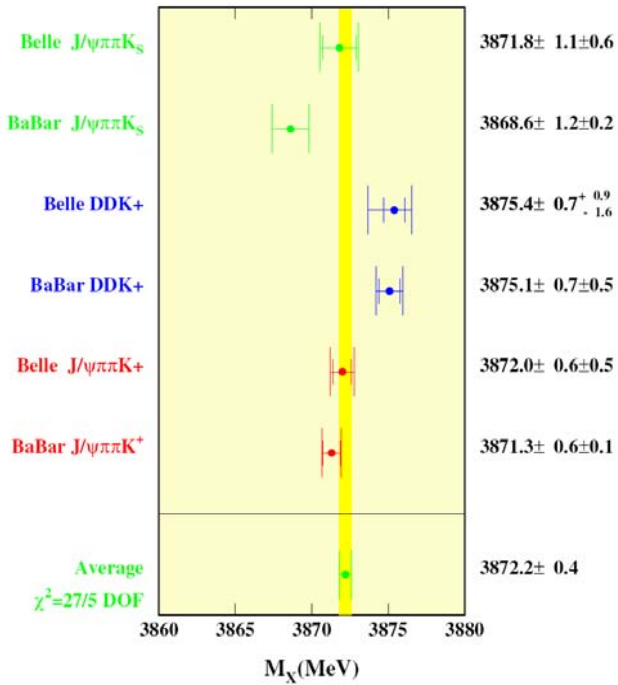
Resolution is still adequate for time dependent measurements

Error on lifetime $\approx \sqrt{\frac{\tau^2 + \sigma_t^2}{N}} = \sqrt{2} \frac{\tau}{\sqrt{N}}$ (wrt perfect resolution)

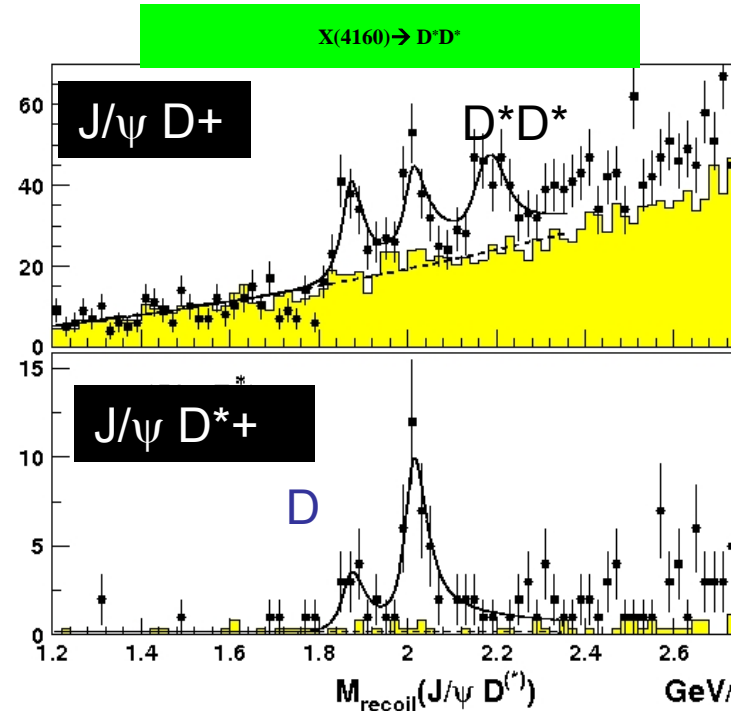
Sept 9, 2008   23

TWO STATES? X(3872) & X(3876) ?

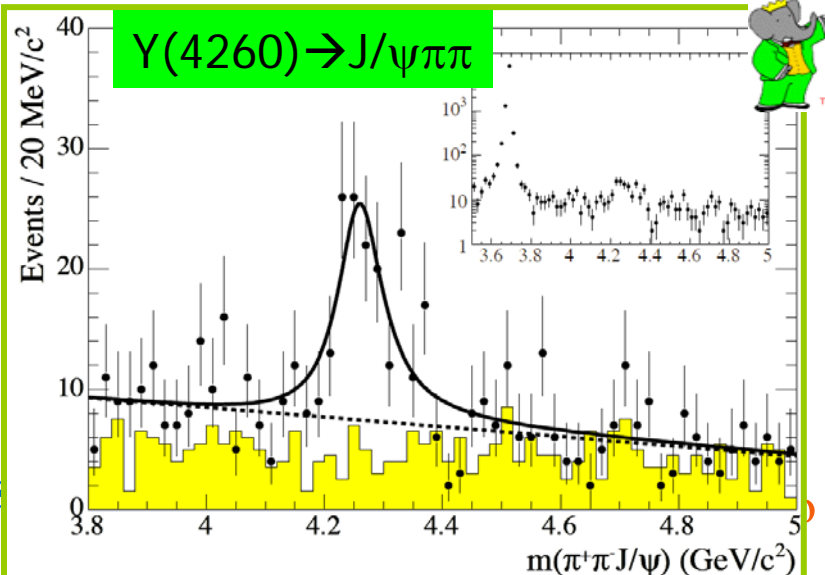
Predicted by tetraquark



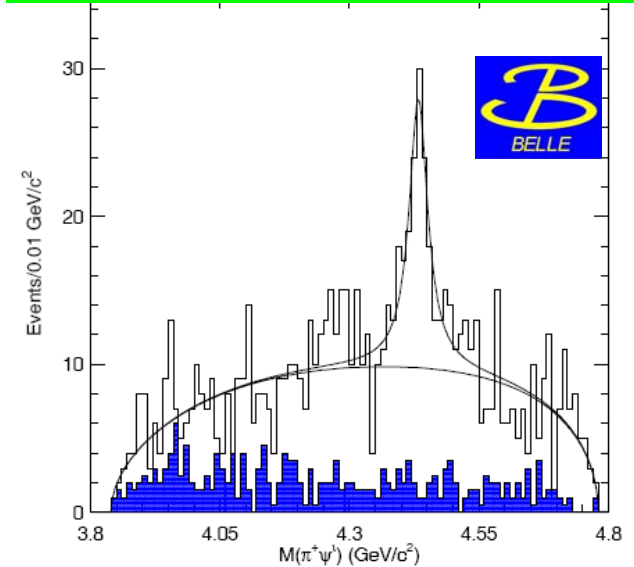
$e^+ e^- \rightarrow J/\psi D^{(*)} D^{(*)}$



The first charged state: Z(4430)!



A. Gic



Super flavor factory projects

Machine project	Cms Energy (GeV)	Mode	Polarization of e^- beam $>80\%$ for τ	Luminosity ($\text{cm}^{-2} \text{s}^{-1}$)
Super c - τ BINP (Russia)	3.0÷4.5	Symmetric	Yes	1÷2 10^{35}
SuperKEKB (Japan)	10.58	Asymmetric	No	2÷8 10^{35}
SuperB- Roma	10.58 4.0	Asymmetric	Yes	1÷4 10^{36} 1 10^{35}



Luminosity

- For gaussian bunches:

$$\mathcal{L} = f_{\text{coll.}} \times \frac{N_{e^+} N_{e^-}}{4\pi \sigma_x \sigma_y} \times R_l$$

geometrical
Reduction
factor

N_{e^+} (N_{e^-}) is the number of positrons (electrons) in a bunch

f_{coll} is the collision frequency

σ_x (σ_y) is the horizontal (vertical) r.m.s. size at the I.P.

R_l is the Luminosity Reduction factor by incomplete overlap: crossing angle and “hour glass” effect.

- **TRADITIONAL** (brute force): increase the numerator Currents increase: from 1A on 2 A up to 4.1 A on 9.4 A- **Wall Plug Power**, HOM, CSR: hard to surpass $5 \cdot 10^{35} \text{ cm}^2\text{s}^{-1}$

Crab Crossing to increase R_l and to optimize beam dynamic

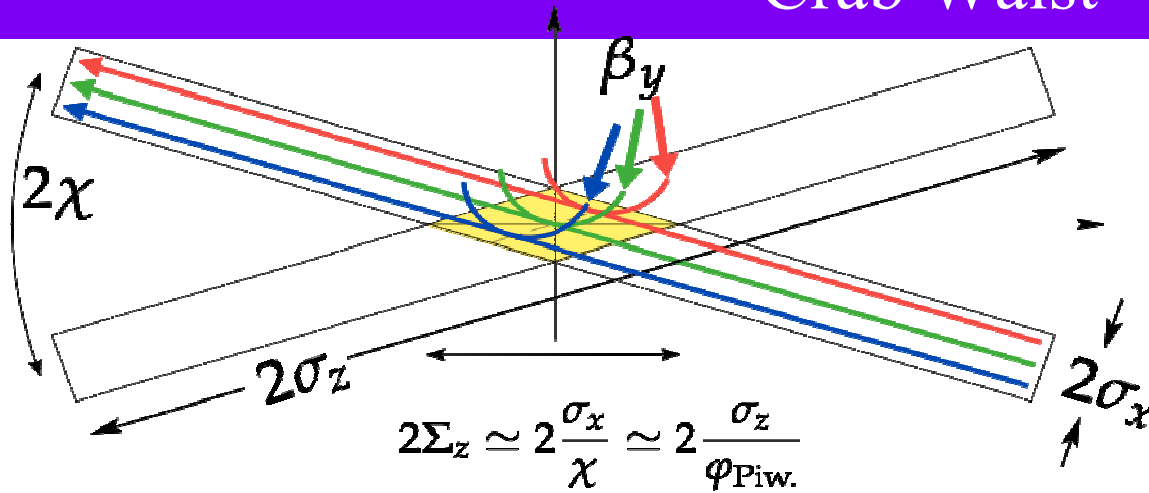
- **SuperB**: decrease the denominator (same currents as PEP-II)

Bunch sizes: from $\sigma_y = 3 \mu\text{m}$ down to $\sigma_y = 40 \text{ nm}$ Luminosity: $10^{36} \text{ cm}^2\text{s}^{-1}$ (baseline) .

Crab Waist and large **Piwinisky** angle to optimize beam dynamic



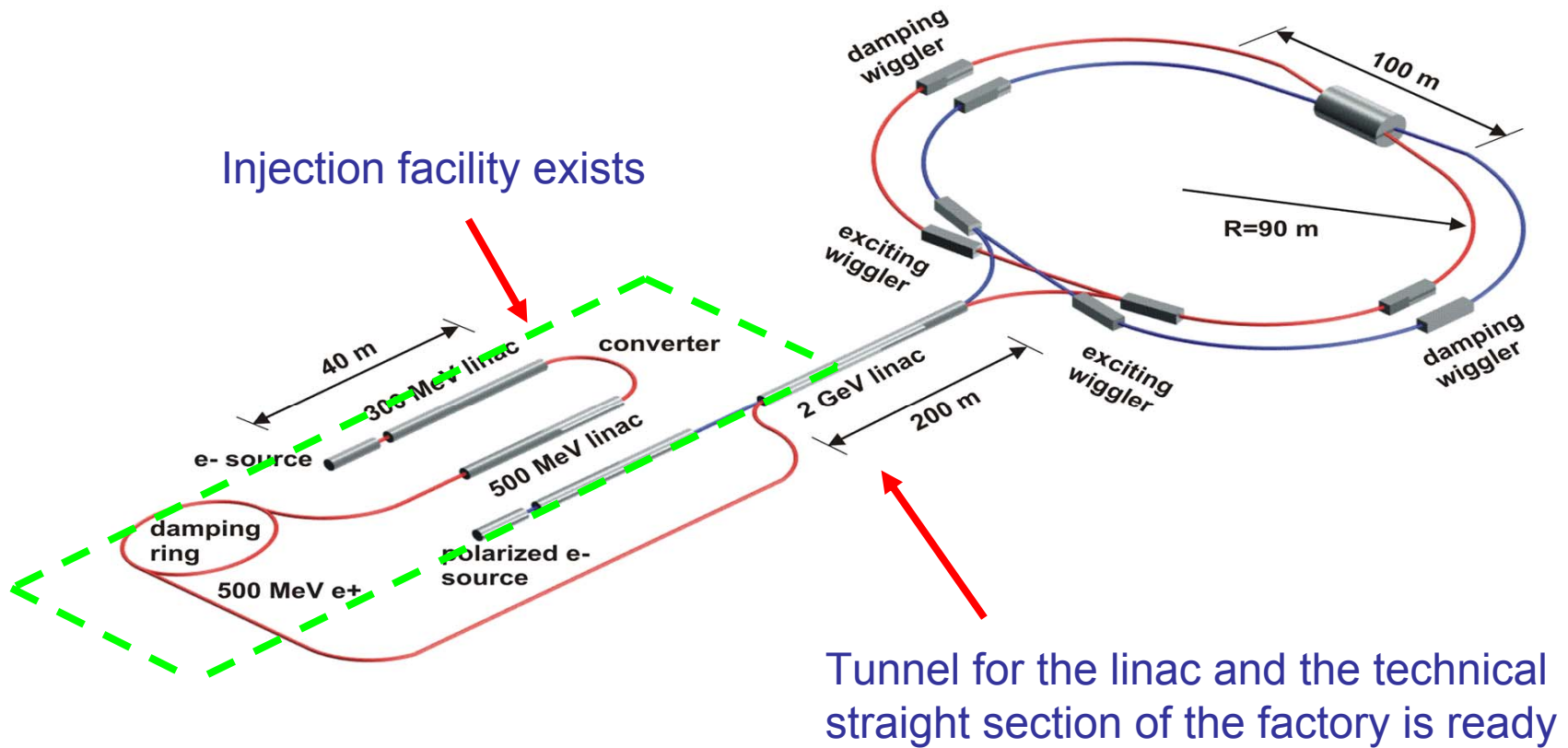
Crab Waist



- Crab waist: modulation of the y-waist position, particles collide at same β_y realized with a sextupole upstream the IP.
- Minimization of nonlinear terms in the beam-beam interaction: reduced emittance growth, suppression of betatron and synchro-betatron coupling
- Maximization of the bunch-bunch overlap: luminosity gain
- Low wall power

SuperB and Super c- τ are based on the crabwaist concept invented in 2006 by P. Raimondi in 2006. Background not exceeding too much the present Babar, thanks to low current, crossing angle and a careful design of the Interaction Region.





Specification for the Super c- τ design

Energy, GeV	2
Beam current, A	1.36
Number of bunches	295
β_x , mm	20
β_y , mm	0.76
ε_x , nm rad	10
Coupling $\varepsilon_y/\varepsilon_x$, %	1
Beam length σ_z , cm	1
Crossing angle, mrad	34

Tune shift ξ_y	0.13
Particles per bunch	$7 \cdot 10^{10}$
Luminosity, $\text{cm}^{-2}\text{sec}^{-1}$	$1 \cdot 10^{35}$
Hour glass $\frac{\sigma_x}{\theta\beta_y}$	1.095
Piwinski angle $\varphi = \frac{\sigma_z\theta}{\sigma_x}$	12.021

- ◆ No bend for incoming beam.
- ◆ No longitudinal field integral over each final focus lens.
- ◆ Longitudinal field is compensated before each final focus lens.
- ◆ Interaction region length less than 100 m.
- ◆ Place for CRAB sextupole.

From the talk by Anton Bogomyagkov

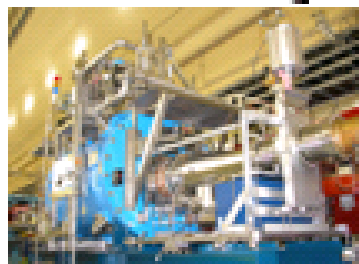


SuperKEKB project

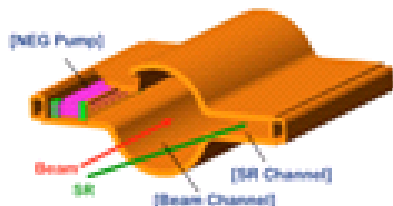


Crab cavities will be installed and tested with beam in 2026.

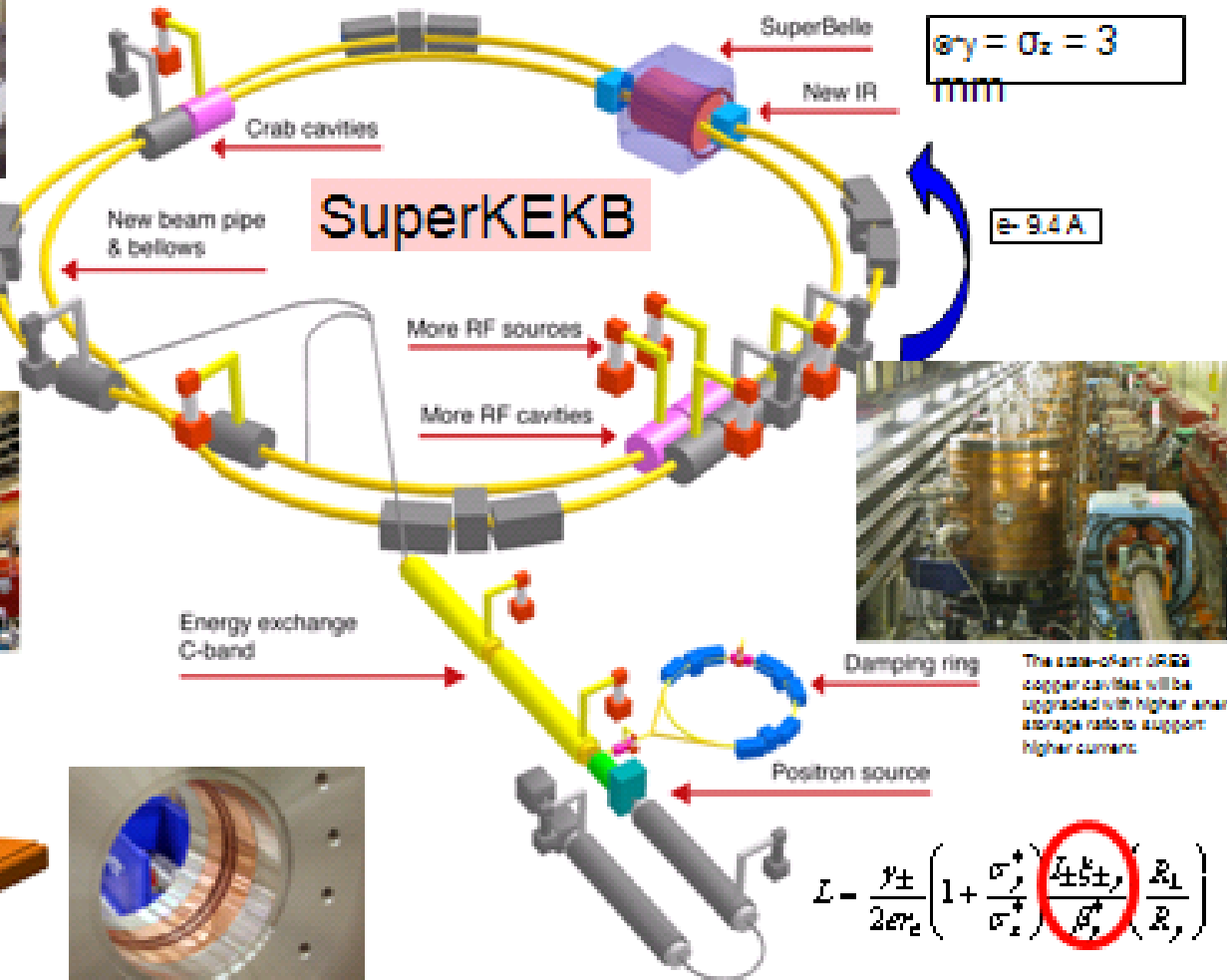
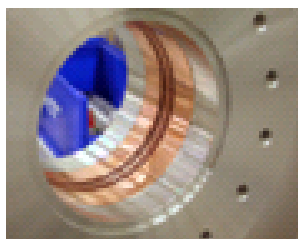
$e^+ 4.1 \text{ A}$



The superconducting cavities will be upgraded to absorb more higher-order mode power up to 50 kW.



The beam pipe and all vacuum components will be replaced with higher-current-proof design.



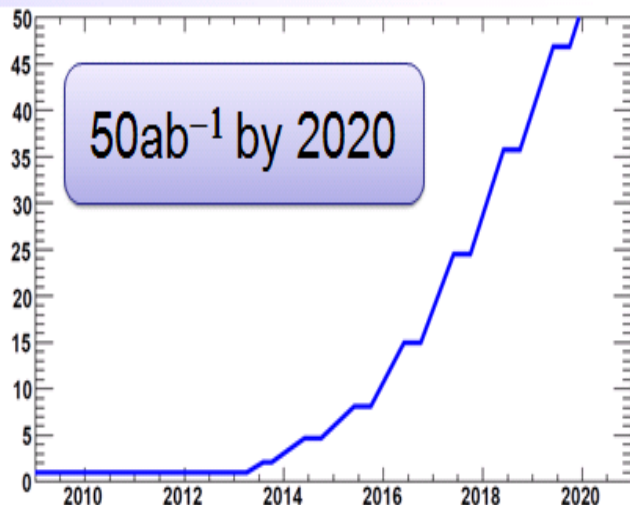
$$I = \frac{y_{\pm}}{2e\sigma_z} \left(1 + \frac{\sigma_y^+}{\sigma_z^+} \frac{I_{\pm 5\pm}}{\beta_s^*} \frac{R_1}{R} \right)$$

will reach $8 \times 10^{35} \text{ cm}^{-2} \text{ s}^{-1}$.



Machine Parameters of SuperKEKB

Luminosity prospect



3

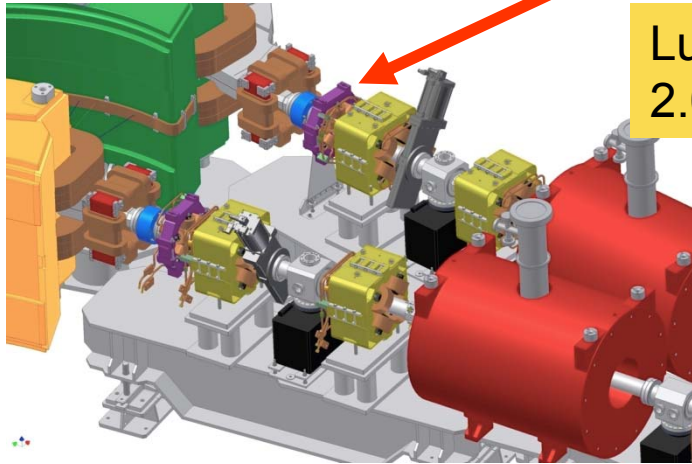
	symbol	LER	HER	unit
Beam Energy	E	3.5	8.0	GeV
Beam current	I	9.4	4.1	A
Circumference	C	3016		m
Number of bunches	n _b	5018		
Number of particles	N/bunch	11.8	5.1	x10 ¹⁰
Emittance	ε _x	9		nm
Emittance ratio	ε _y /ε _x	0.5		%
Beta (hor.) at IP	β _x [*]	200		mm
Beta (ver.) at IP	β _y [*]	3		mm
Bunch length	σ _z	3		mm
Crossing angle	θ _x [*]	30 to 0		mrاد
Beam-Beam (hor.)	ξ _x	0.36		
Beam-Beam (ver.)	ξ _y	0.43		
RF AC plug power	P _{ac}	73		MW
Luminosity	L	8.0		x10 ³³ cm ⁻² s ⁻¹

2

After 7th year integrated Luminosity can grow at rate of ~10÷12 ab⁻¹/year

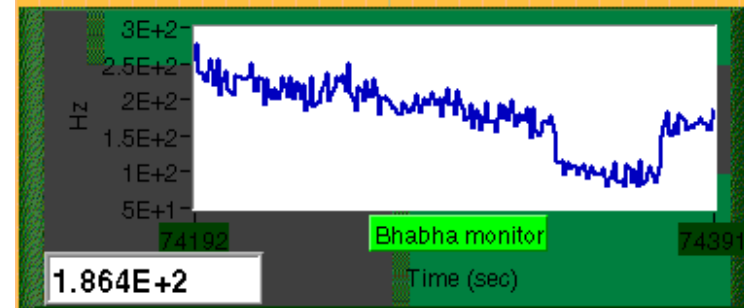
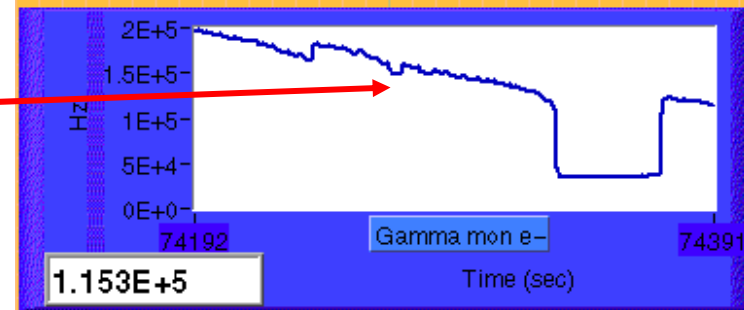
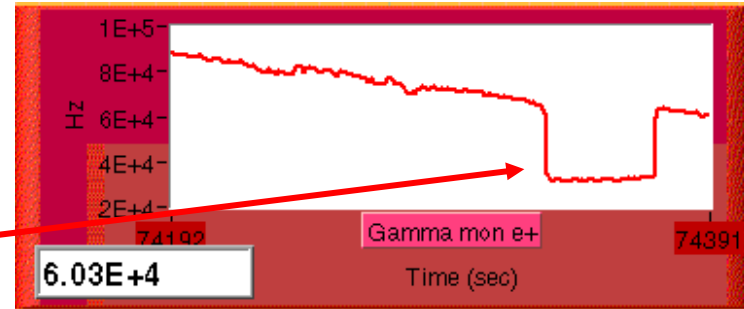


SuperB Test of crab waist at DaΦne (LNF)



Luminosity stable above 2.0×10^{32} reached

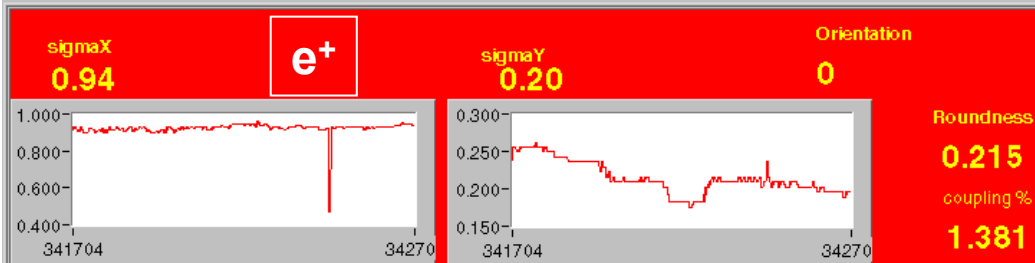
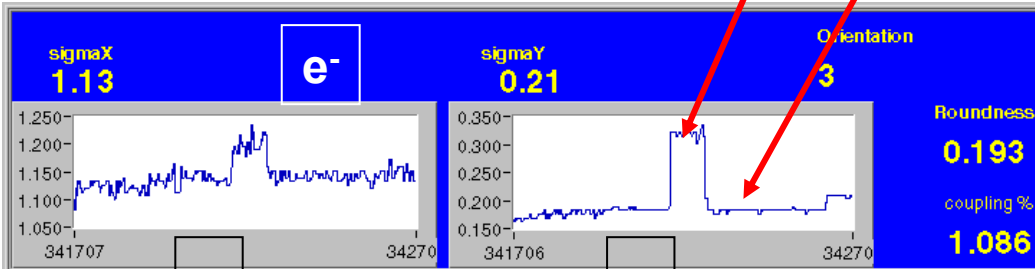
LUMINOMETERS



Transverse beam dimensions at the Synchrotron Light Monitor

Crab OFF

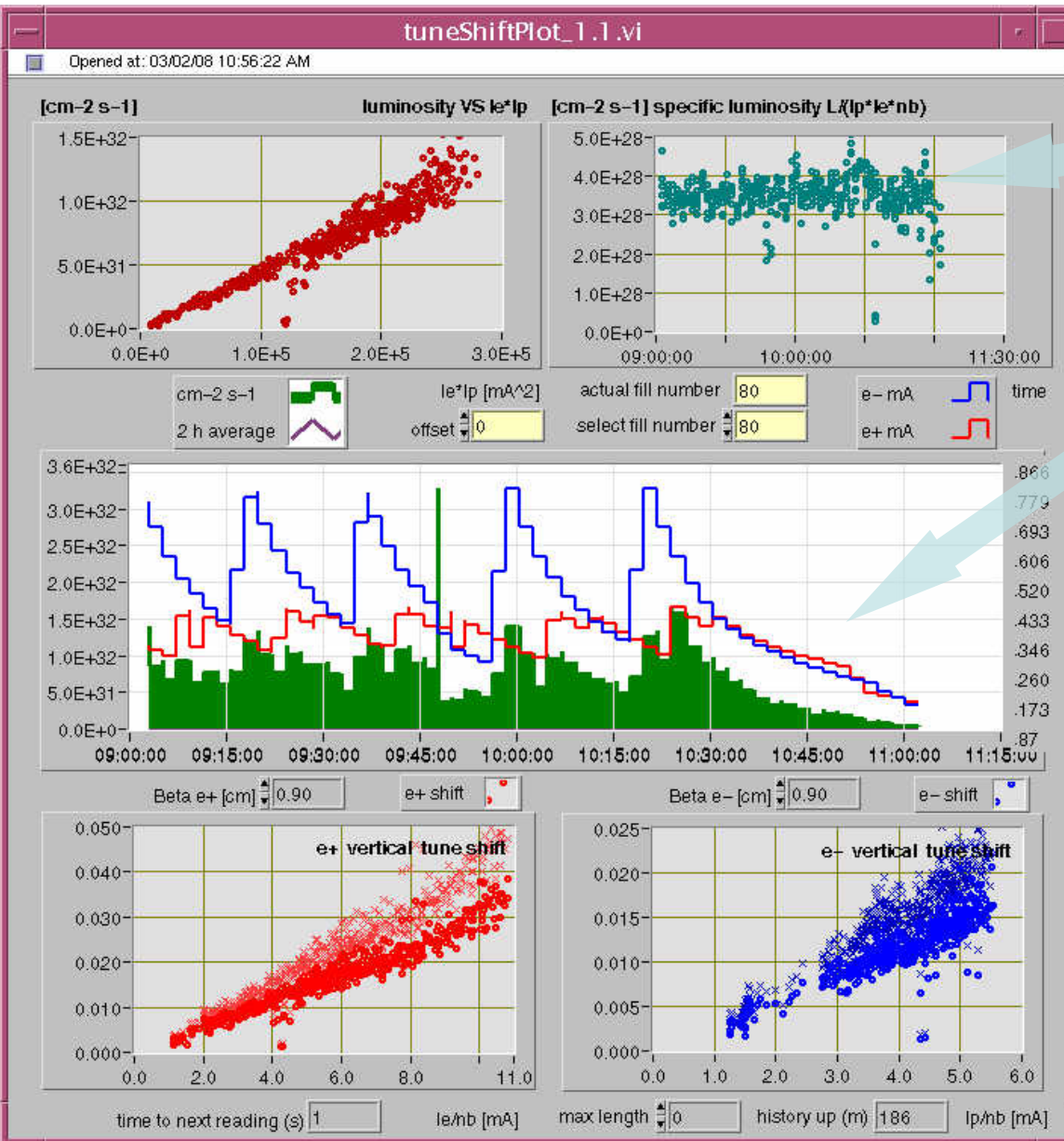
Crab ON



Blow-up in beam sizes and decrease in Bhabha rates observed when crab sexts for one ring OFF (other ring ON)



Specific luminosity



- Specific Luminosity(vs time) does not depends on beam currents
- headrom for luminosity improvements

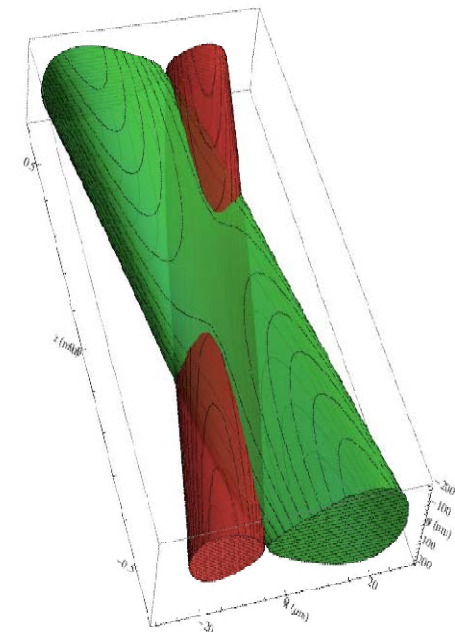
SuperB parameters with higher Wall Power

PARAMETER	Nominal		Upgrade	
	LER (e+)	HER (e-)	LER (e+)	HER (e-)
Energy (GeV)	4	7	4	7
Luminosity $\times 10^{36}$	1.0		2.0	
Circumference (m)	1800	1800		
Revolution frequency (MHz)	0.167			
Eff. long. polarization (%)	0	80		
RF frequency (MHz)	476			
Momentum spread ($\times 10^{-4}$)	7.9	5.6	9.0	8.0
Momentum compaction ($\times 10^{-4}$)	3.2	3.8	3.2	3.8
Rf Voltage (MV)	5	8.3	8	11.8
Energy loss/turn (MeV)	1.16	1.94	1.78	2.81
Number of bunches	1251			
Particles per bunch ($\times 10^{10}$)	5.52			
Beam current (A)	1.85			
Beta y^* (mm)	0.22	0.39	0.16	0.27
Beta x^* (mm)	35	20		
Emit y (pm-rad)	7	4	3.5	2
Emit x (nm-rad)	2.8	1.6	1.4	0.8
Sigma y^* (microns)	0.039	0.039	0.0233	0.0233
Sigma x^* (microns)	9.9	5.66	7	4
Bunch length (mm)	5		4.3	
Full Crossing angle (mrad)	48			
Wigglers (#) 20 meters each	0	0	2	2
Damping time (trans/long)(ms)	40/20	40/20	28/14	28/14
Luminosity lifetime (min)	6.7		3.35	
Touschek lifetime (min)	13	20	6.9	10.3
Effective beam lifetime (min)	4.5	5.1	2.3	2.5
Injection rate pps ($\times 10^{11}$) (100%)	2.6	2.3	5.1	4.6
Tune shift y (from formula)	0.15		0.20	
Tune shift x (from formula)	0.0043	0.0025	0.0059	0.0034
RF Power (MW)	17		25	

150m needed for Polarization

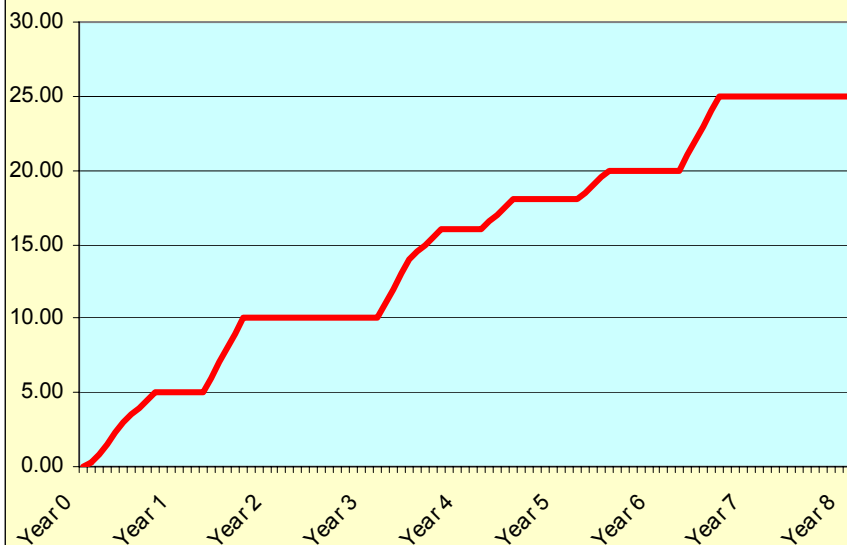
Doubling currents with a factor 2 in Wall power we can double the luminosity

LEB HEB

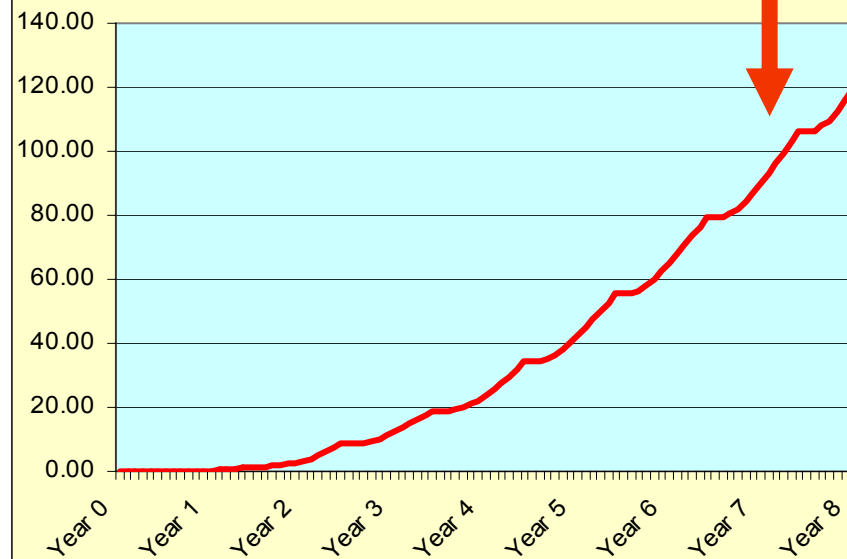


SuperB expected LUMI

Peak Luminosity (10^{35})



Integrated Luminosity ($1/\text{ab}$)

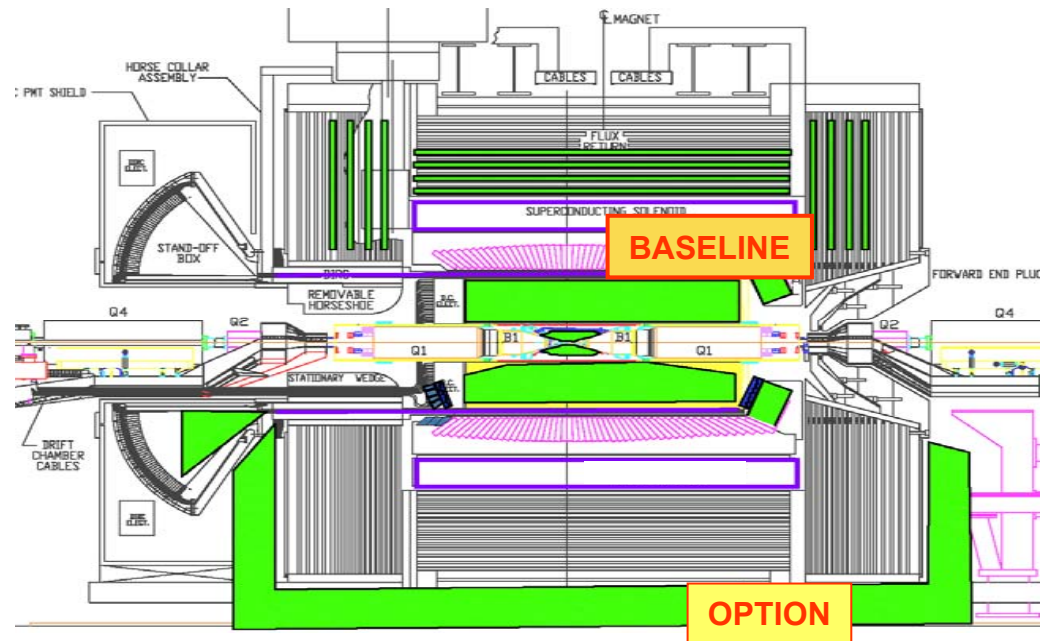
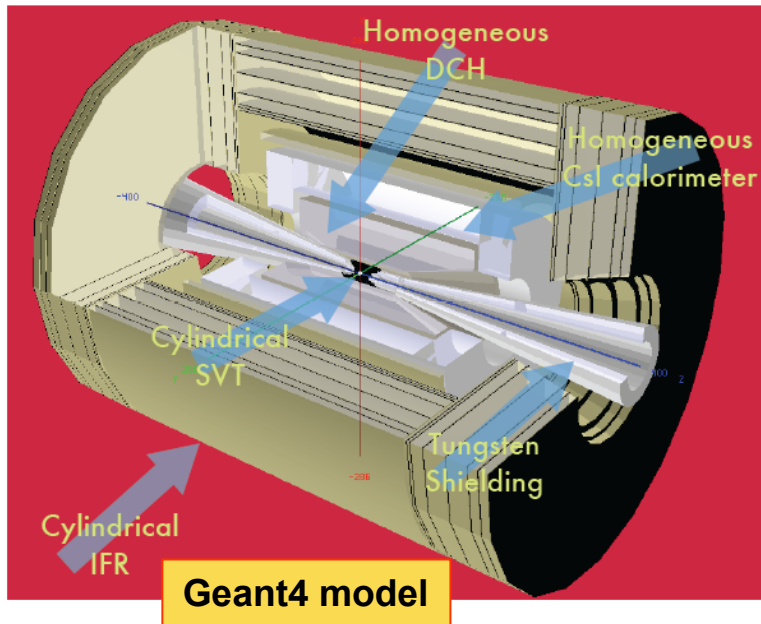


>80ab-1 in 7th year

After 7th year integrated Luminosity can grow at rate of $\sim 40 \div 60 \text{ ab}^{-1}/\text{year}$



Detector Layout – Reuse parts of Babar



Test beam goals for 2008-2010

Silicon Vertex Tracker

MAPS pixel devices: resolution, efficiency, readout speed

(now under test at CERN !)

Advanced trigger systems (Associative Memories)

Drift Chamber

Cell size, shape, and gas mixture

Particle ID system (forward system)

Radiators (Aerogel, NaF)

Photon detector (MCP, MAPMTs, SiPM)

Timing for TOF system

Electromagnetic Calorimeter

Forw: LYSO Crystals leakage, resolution, mechanical structure

Back: Lead-scintillator calorimeter resolution

Instrumented Flux Return

Scintillator, fibers, photon detector, readout electronics



END

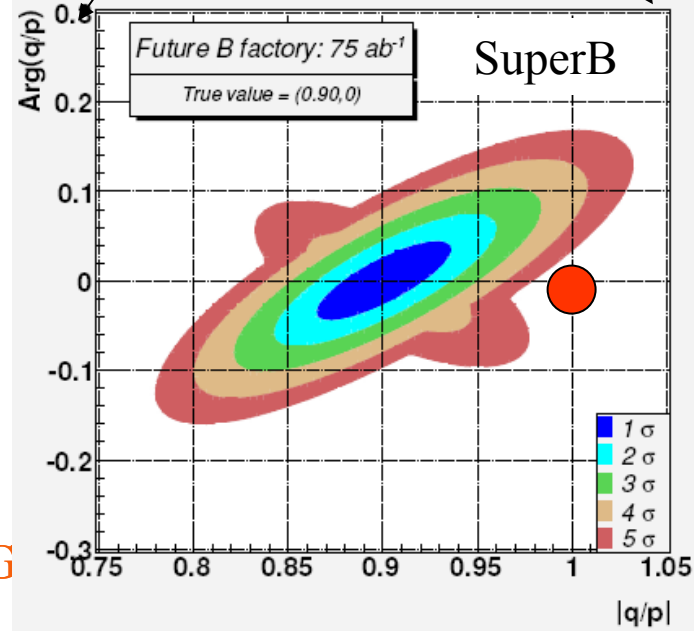
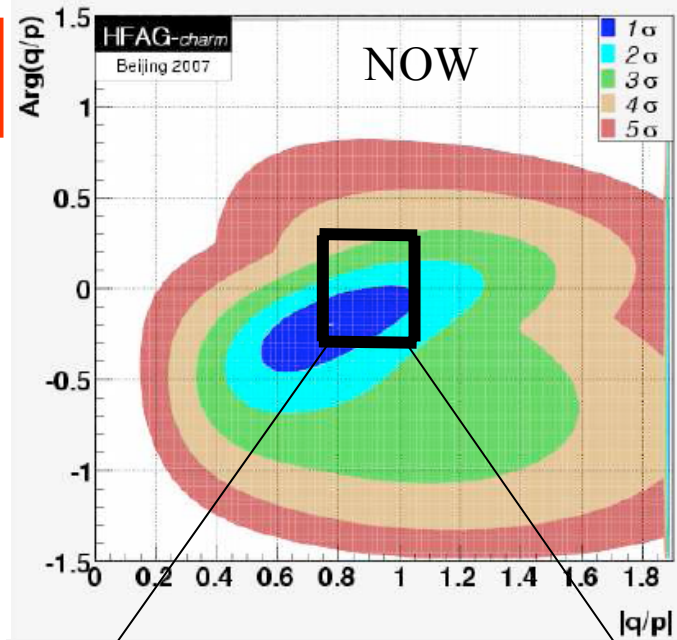


- Charm events at threshold are very clean: pure DD, no additional fragmentation
- High signal/bkg ratio: optimal for decays with neutrinos.
- Quantum Coherence: new and alternative CP violation measurement wrt to $\Upsilon(4S)$. Unique opportunity to measure D^0 - D^0 relative phase.
- Increased statistics is not an advantage running at threshold: cross-section 3x wrt 10GeV but luminosity 10x smaller.
- SuperB lumi at 4 GeV = 10^{35} cm⁻²s⁻¹ produces $\sim 10^9$ DD pairs per month of running. (using Cleo-c cross-section measurement [$\sigma(e^+e^- \rightarrow D^0D^0) \sim 3.6$ nb] + [$\sigma(e^+e^- \rightarrow D^+D^-) \sim 2.8$ nb] ~ 6.4 nb)
- Time-dependent measurements at 4 GeV **only** possible at SuperB to extract Phase .



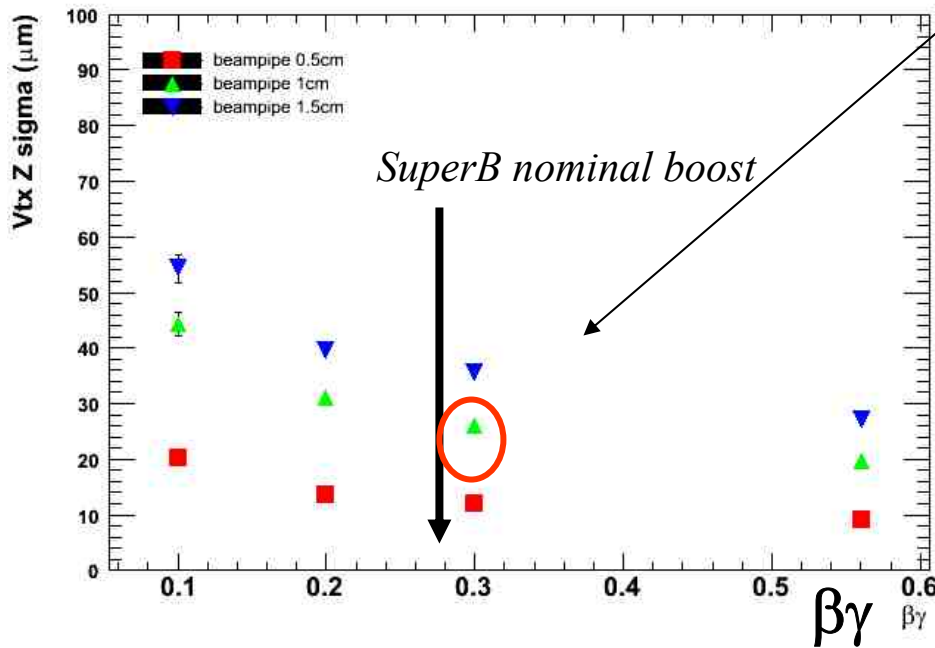
CP Violation in charm from mixing

Mode	Observable	$\Upsilon(4S)$ (75 ab^{-1})	$\psi(3770)$ (300 fb^{-1})
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$\psi(3770) \rightarrow D^0 \bar{D}^0$	x^2		$(1-2) \times 10^{-5}$
	y		$(1-2) \times 10^{-3}$
	$\cos \delta$		$(0.01-0.02)$



Time dependent measurements at DD threshold: only possible at SuperB

- Proper time resolution dominated by decay vertex resolution.
 - Production vertex precisely determined thanks to nm beamspot dimensions



With SuperB lumi at 4 GeV = $10^{35} \text{ cm}^{-2}\text{s}^{-1}$
expected $\sim 10^9 \text{ DD}$ per month

$\beta\gamma ct = 0.28 \times 120 \mu\text{m} \sim 30 \mu\text{m}$
Average flight distance similar to vertex resolution $\rightarrow \sigma_t \sim \tau$

