

# Cosmic-ray studies using the ALICE detector at LHC

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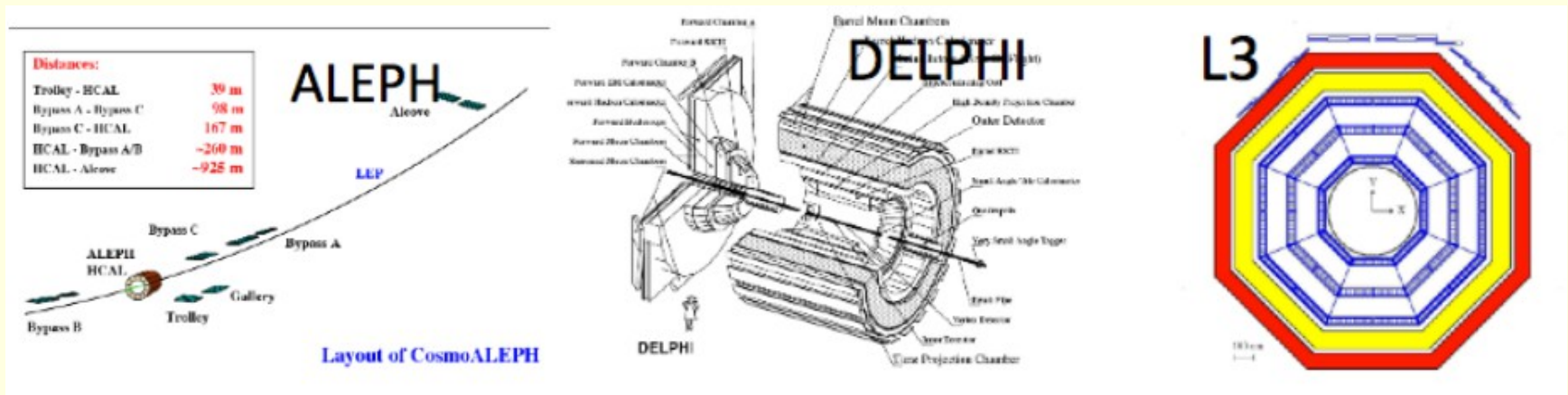
# Outline

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- Motivations
- The ALICE Experiment and the atmospheric muons
  - Triggering and tracking detectors
- The Muon Multiplicity Distribution
- High Muon Multiplicity events
- Monte Carlo and data comparison
- Conclusions

# Motivations

- Use of collider detectors for cosmic-ray studies was pioneered by LEP experiments ALEPH, DELPHI and L3

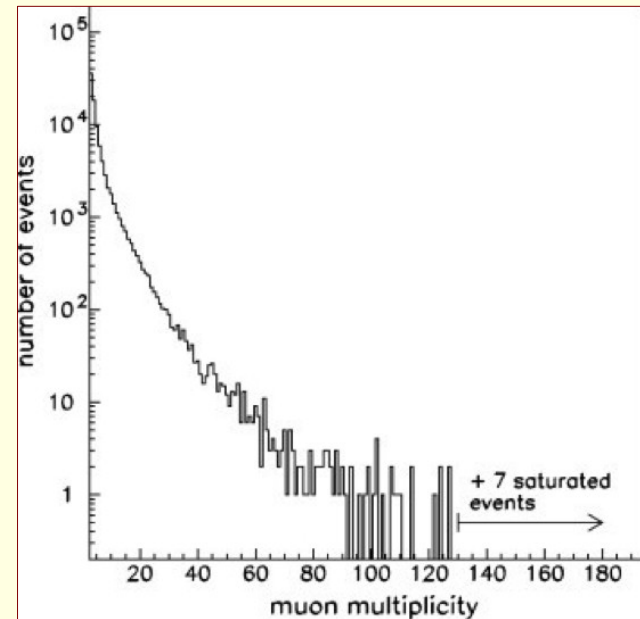
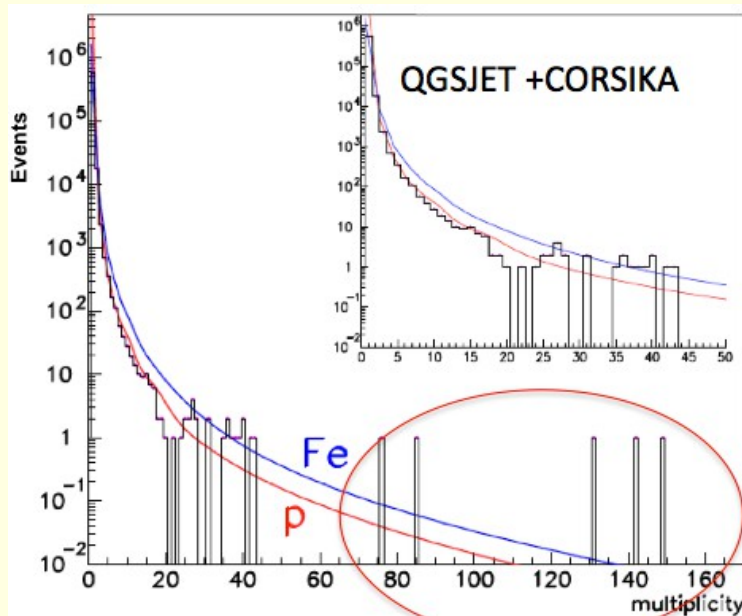


- ✗ Small apparatus
- ✗ Muons crossing the rock

- ✓ High performance detectors
  - tracking
  - magnetic field

# Motivations

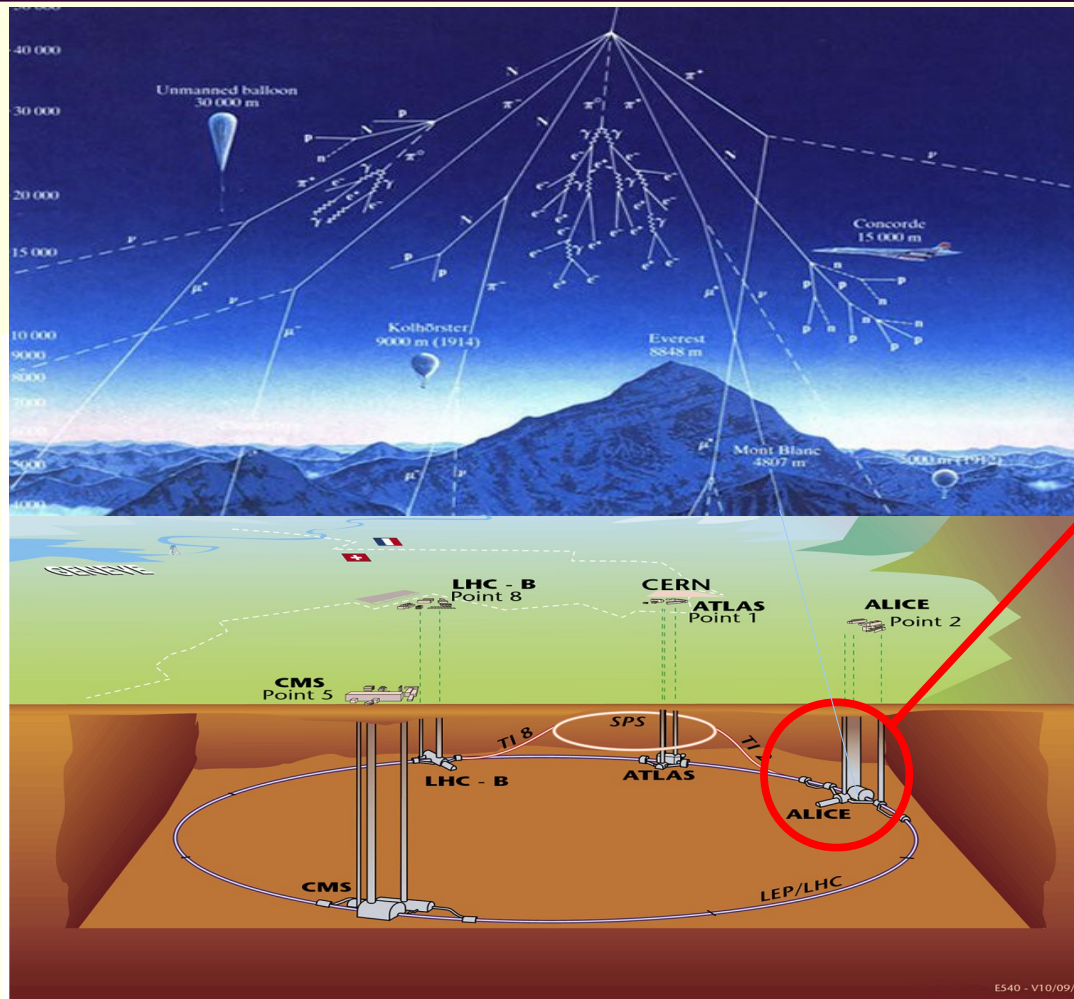
- All LEP results were consistent with standard hadronic interaction models except the observation of high multiplicity muon bundles
  - even under the assumption of highest measured flux and pure iron spectrum



ALEPH Coll., *Astrop. Phys.* **19** (2003) 513  
May 26<sup>th</sup>, 2016

DELPHI Coll., *Astrop. Phys.* **28** (2007) 273  
Vulcano Workshop 2016

# Detection of cosmic muons at LHC

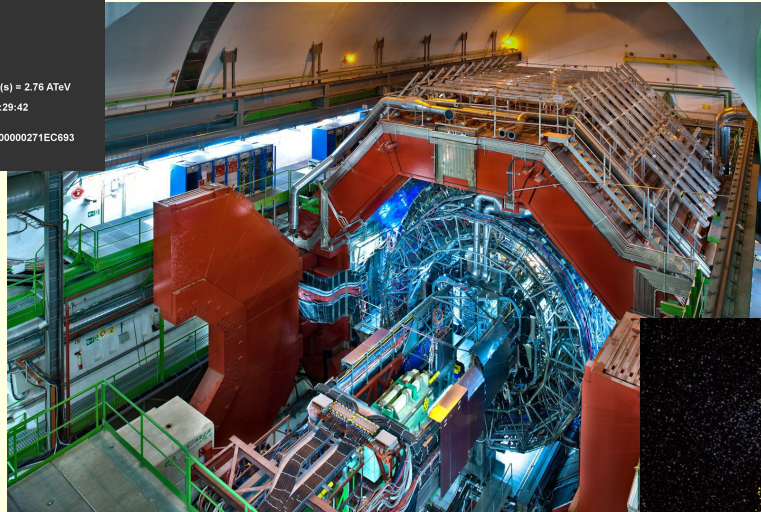
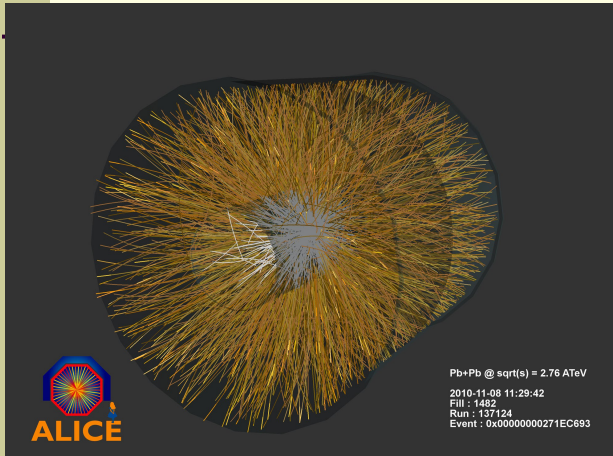


- ▶ ALICE is located at LHC Point 2 52m underground (28m rock above)

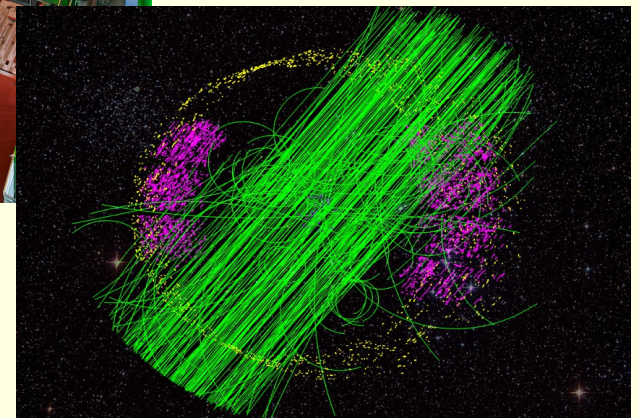
- ▶ Muon energy threshold  $\sim 16$  GeV

# ALICE Experiment

ALICE is mainly devoted to the study of strongly interacting matter in  $pp$ ,  $pA$  and  $AA$  collisions at ultra-relativistic energies



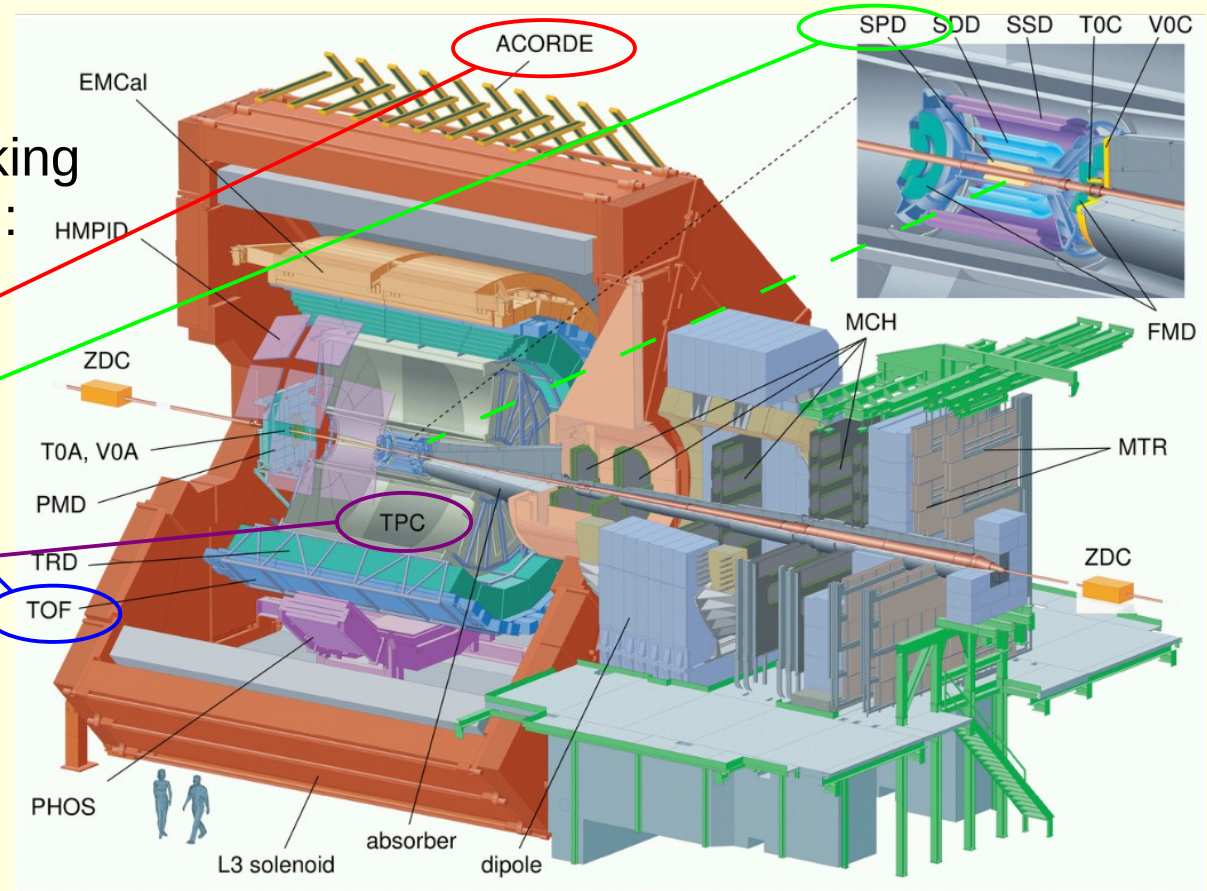
Besides the Heavy-Ion Physics program, ALICE has a dedicated physics group devoted to cosmic-ray studies



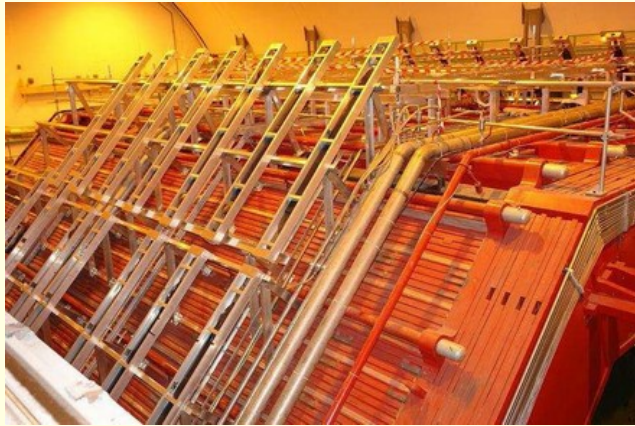
# The ALICE detectors

Detectors used for cosmic-ray data taking in the central barrel:

- ▶ Trigger
  - ACORDE
  - TOF
  - SPD
- ▶ Tracking
  - TPC



# Trigger detectors for cosmic muons



## ACORDE

Array of 60 scintillator modules located in the three top octants of the magnet

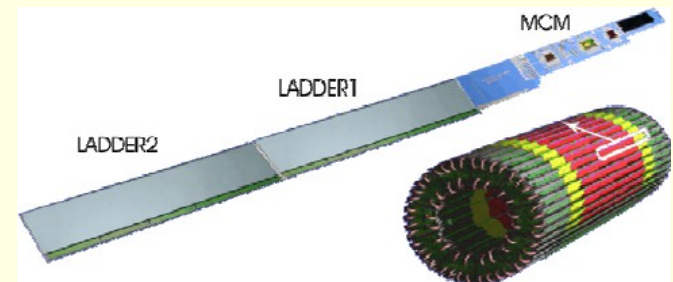
Each module is made of two plastic scintillators with effective area of  $0.38 \text{ m}^2$ .

Configurable from 2-fold coincidence (1Hz) onward.

## SPD

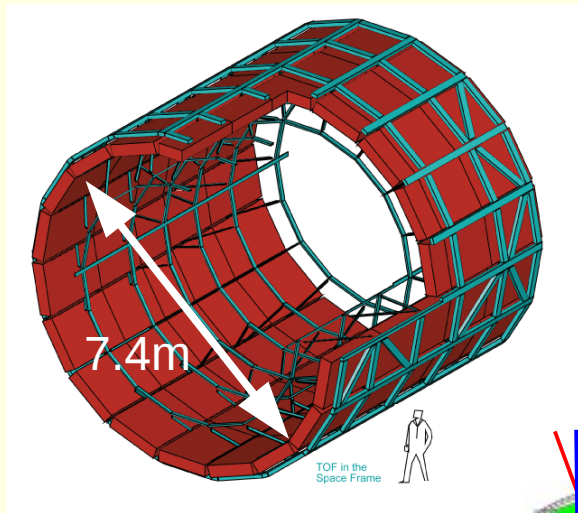
Two innermost coaxial cylinders of ITS, around the beam pipe.

10 M pixels segmented in 120 modules provide trigger and particle position.





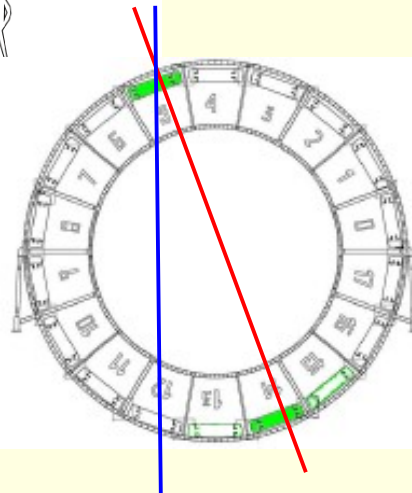
# Trigger detectors for cosmic muons



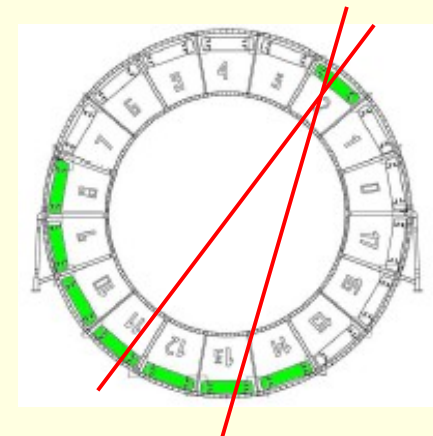
## TOF

Array of 1638 MRPC pads (18  $\phi$  sectors with 5 modules each) around TPC.

Full  $\phi$  coverage,  $45^\circ < \theta < 135^\circ$ , time resolution 100ps,  $\sim 95\%$  efficiency



Back to back

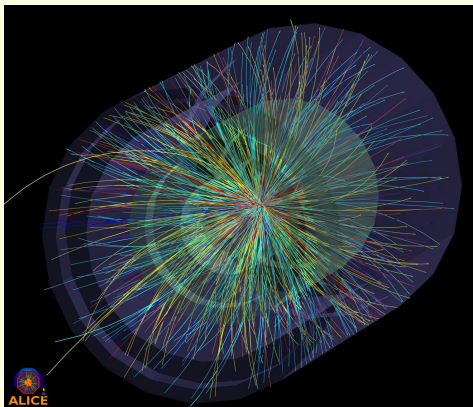
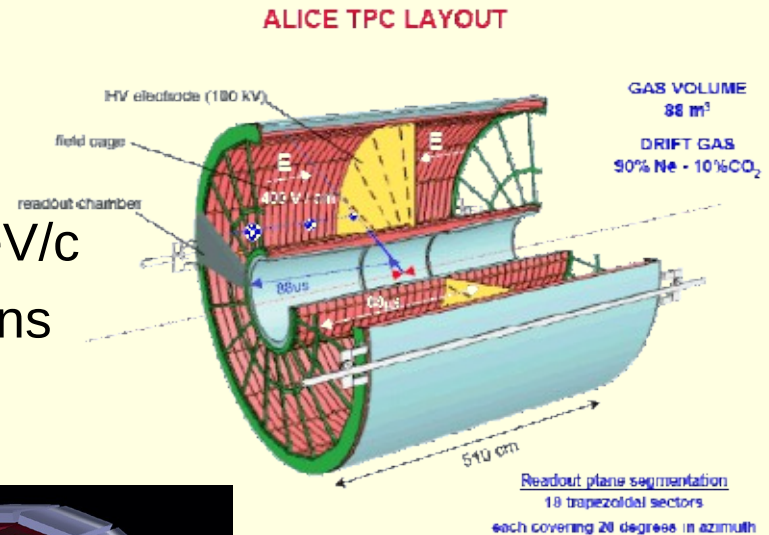


Back to back  $\pm 3$  pads

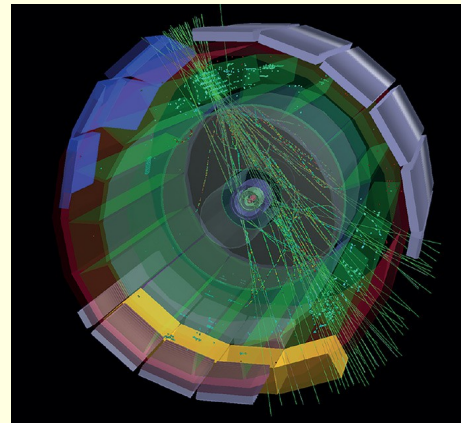
# Tracking detector for cosmic muons

## TPC

- Main tracking device with excellent capabilities for high-track density
- 557k readout channels
- Moment resolution  $\sim 1\%$  for  $p_t < 2 \text{ GeV}/c$   
 $\sim 20\%$  for  $p_t = 100 \text{ GeV}/c$  in HI collisions
- Tracking efficiency 90%
- $dE/dx$  resolution  $< 10\%$

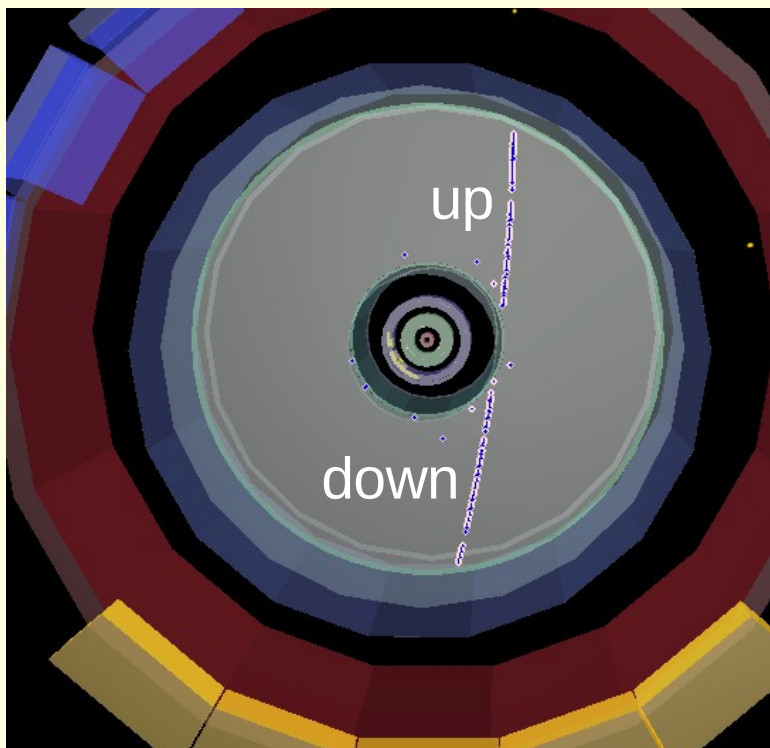


May 26<sup>th</sup>, 2016



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# Atmospheric muon reconstruction



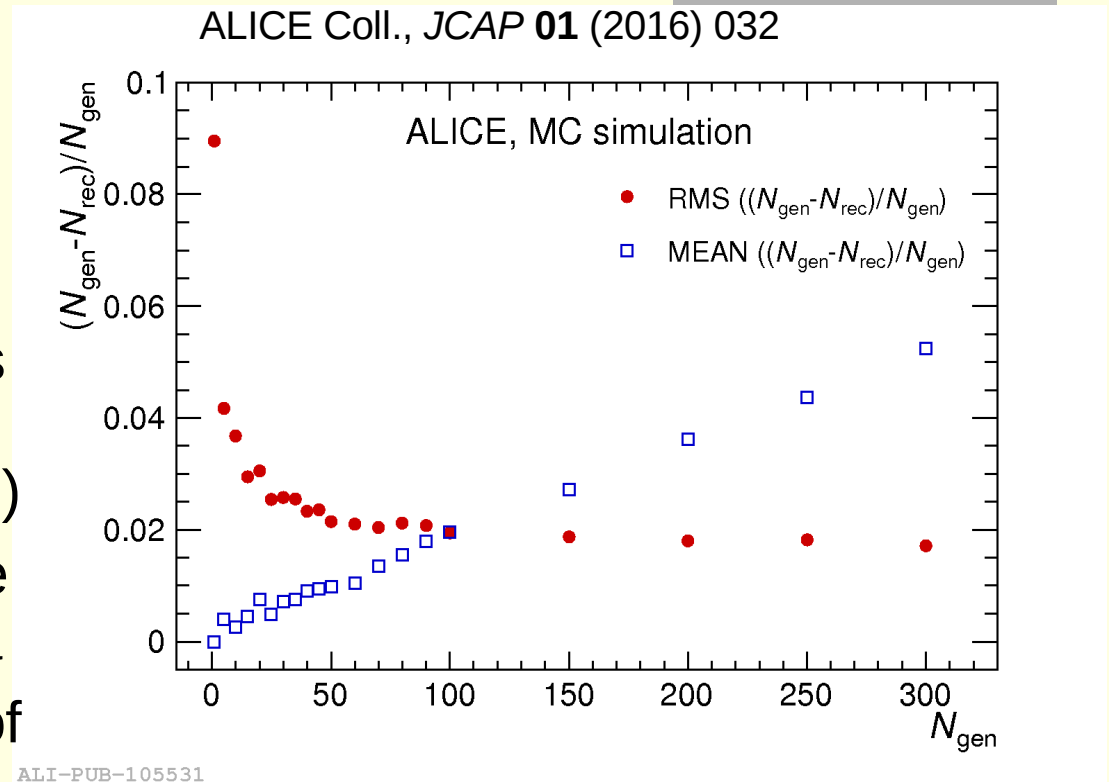
- The TPC reconstructs a single muon as two tracks (up and down)
- A specific algorithm was worked out to match the two tracks as a single one
- Monte Carlo events and data of high multiplicity have been used to optimize the parameters of the matching algorithm

# Analysis cuts

- To accept a track
  - $> 50$  space points in the TPC (out of a maximum of 159)
  - $p > 0.5 \text{ GeV}/c$
  - if multi-muon, parallelism cut  $\cos(\Delta\Psi) > 0.990$
- To match an up track with a down track
  - $d_{ca} < 3\text{cm}$  in the mid horizontal TPC plane
- Matched muon: up and down tracks matched
- Single-track muon: a track satisfying all cuts but distance  $d_{ca}$

# Efficiency of analysis cuts

- generate 1000 events for 20 different muon multiplicities (1 to 300)
- reconstruct with same algorithm as real data
- plot mean and RMS of

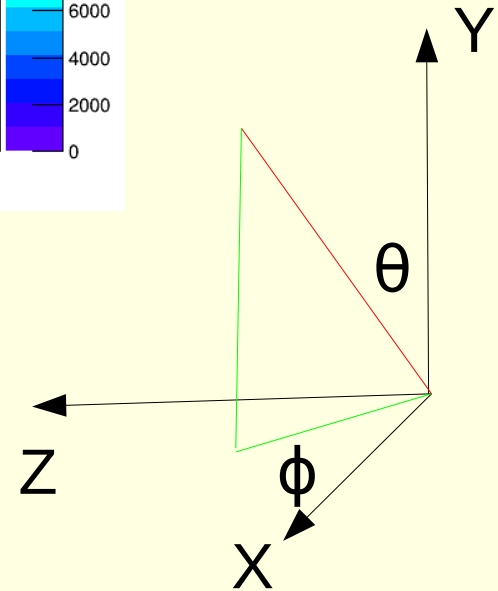
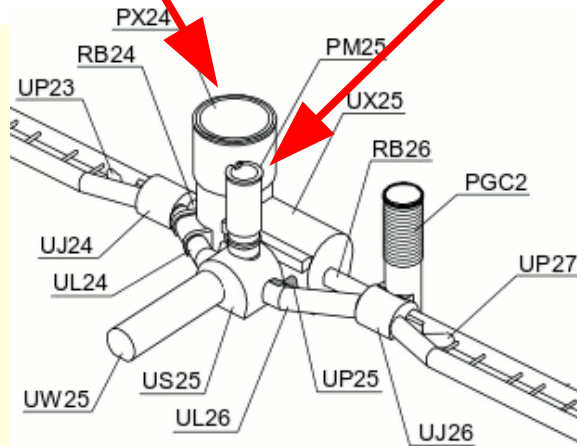
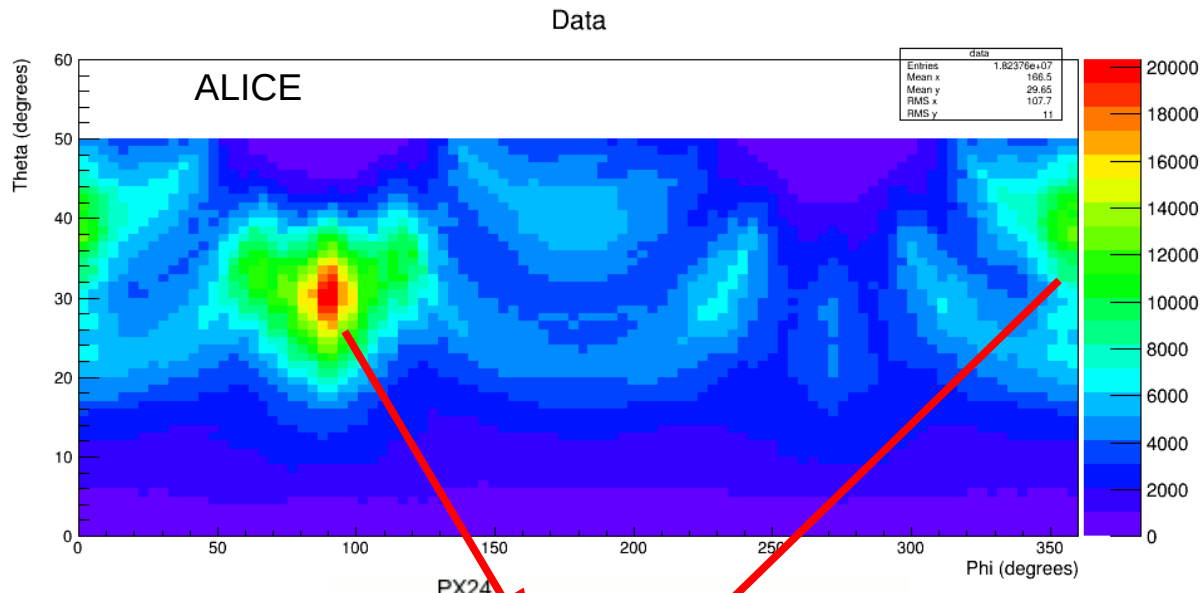


$$\left( \# \text{ generated } \mu - \# \text{ reconstructed } \mu \right) / \left( \# \text{ generated } \mu \right)$$

# Data sample for cosmic-ray studies

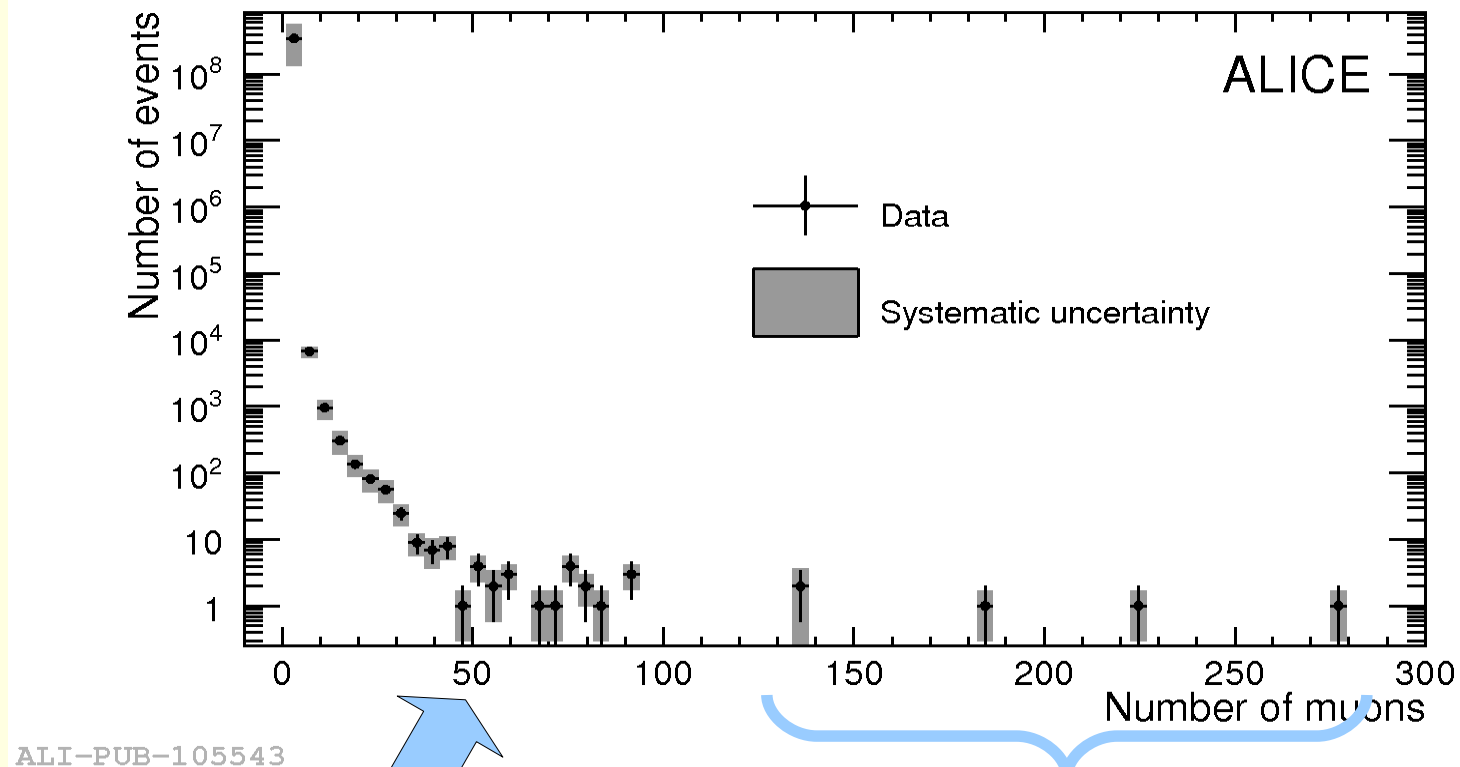
- Data taken between 2010 and 2013 during no-beam periods
  - OR of ACORDE, TOF and SPD triggers
  - with and without magnetic field (0.5 T)
  - integrated live time 30.8 days
  - ~ 22.6M events with at least 1 reconstructed muon in TPC
- Multi-muon event:  $N_{\mu} > 4$  in TPC
  - 7487 multi-muon events

# Muon angular distribution



# Muon multiplicity distribution (MMD)

ALICE Coll., JCAP **01** (2016) 032



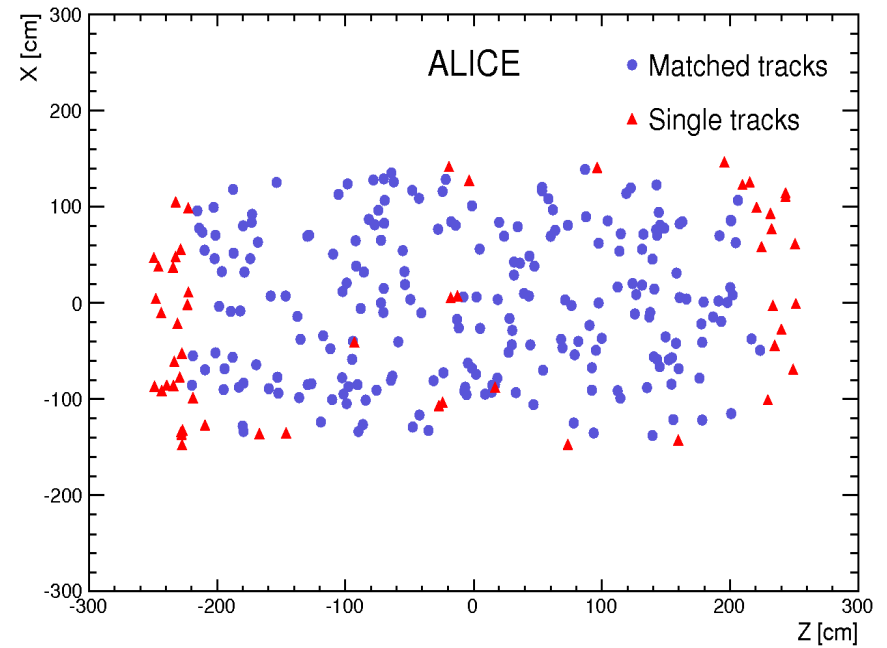
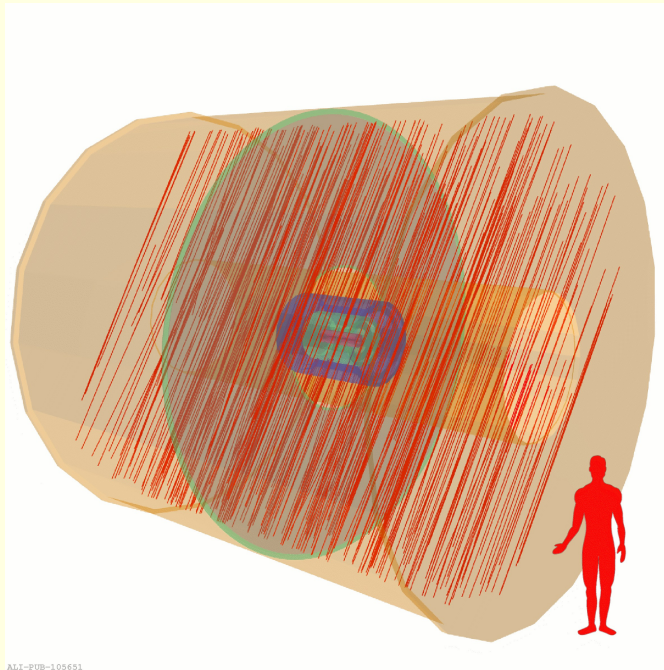
Smooth distribution up to  $N_\mu \sim 70$

5 High Muon Multiplicity events  
(defined when  $N_\mu > 100$ )



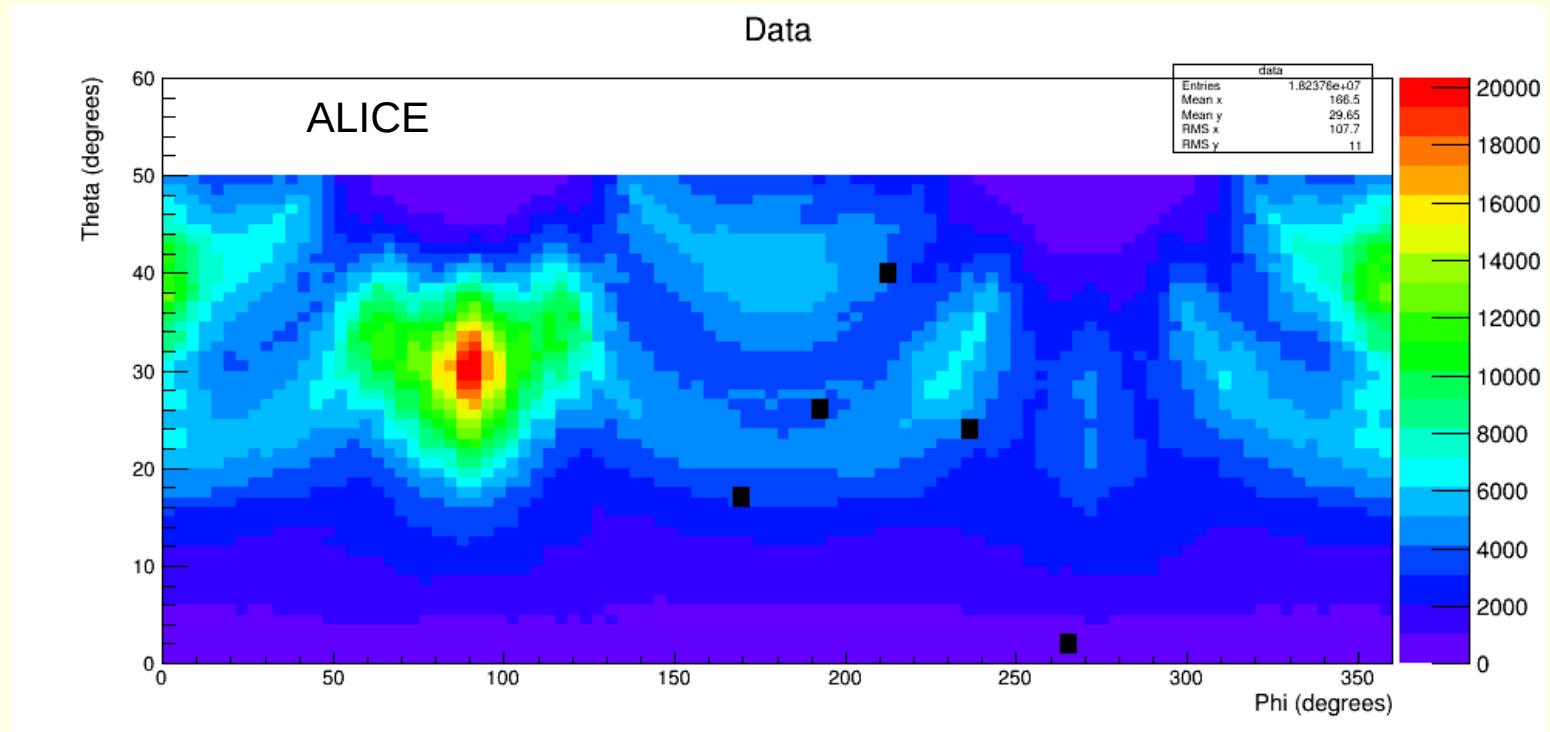
# High Muon Multiplicity event display

ALICE Coll., *JCAP* **01** (2016) 032



High Muon Multiplicity (HMM) event with 276 muons

# Location of HMM events

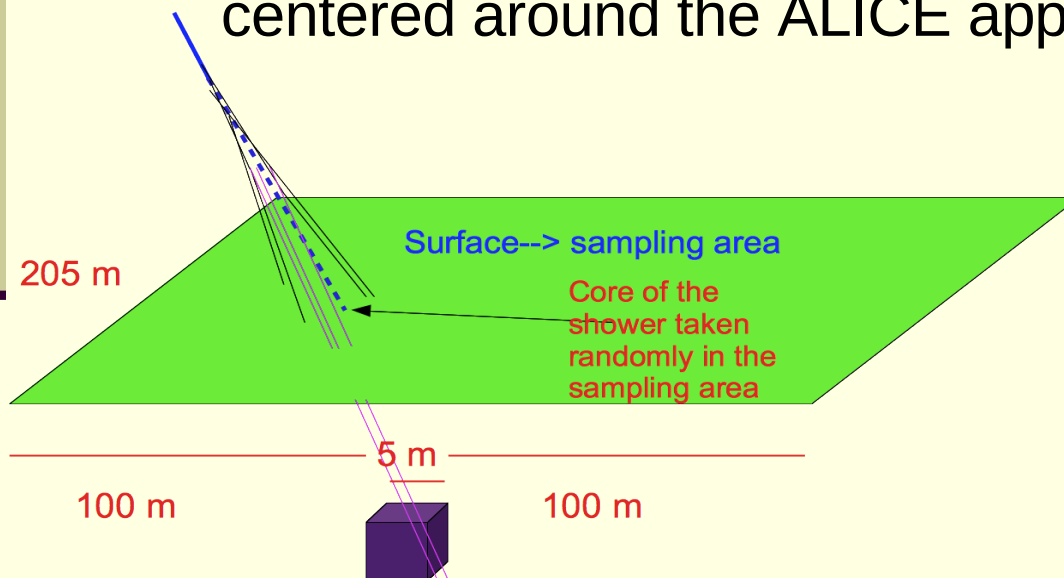


# Monte Carlo simulation

- Simulated events equivalent to 30.8 days live time were generated
  - CORSIKA 6990 and QGSJET II-03 were used to model low-intermediate MMD and study HMM events
  - CORSIKA 7350 and QGSJET II-04 were used to check and confirm results for HMM events
  - two samples: pure  $p$  (light composition) and pure Fe (extremely heavy composition)
  - primary cosmic-ray energy  $10^{14} < E < 10^{18}$  eV
  - usual power law energy spectrum  $E^{-\gamma}$  with  $\gamma = 2.7$  below the knee ( $3 \times 10^{15}$  eV) and  $\gamma = 3.0$  above

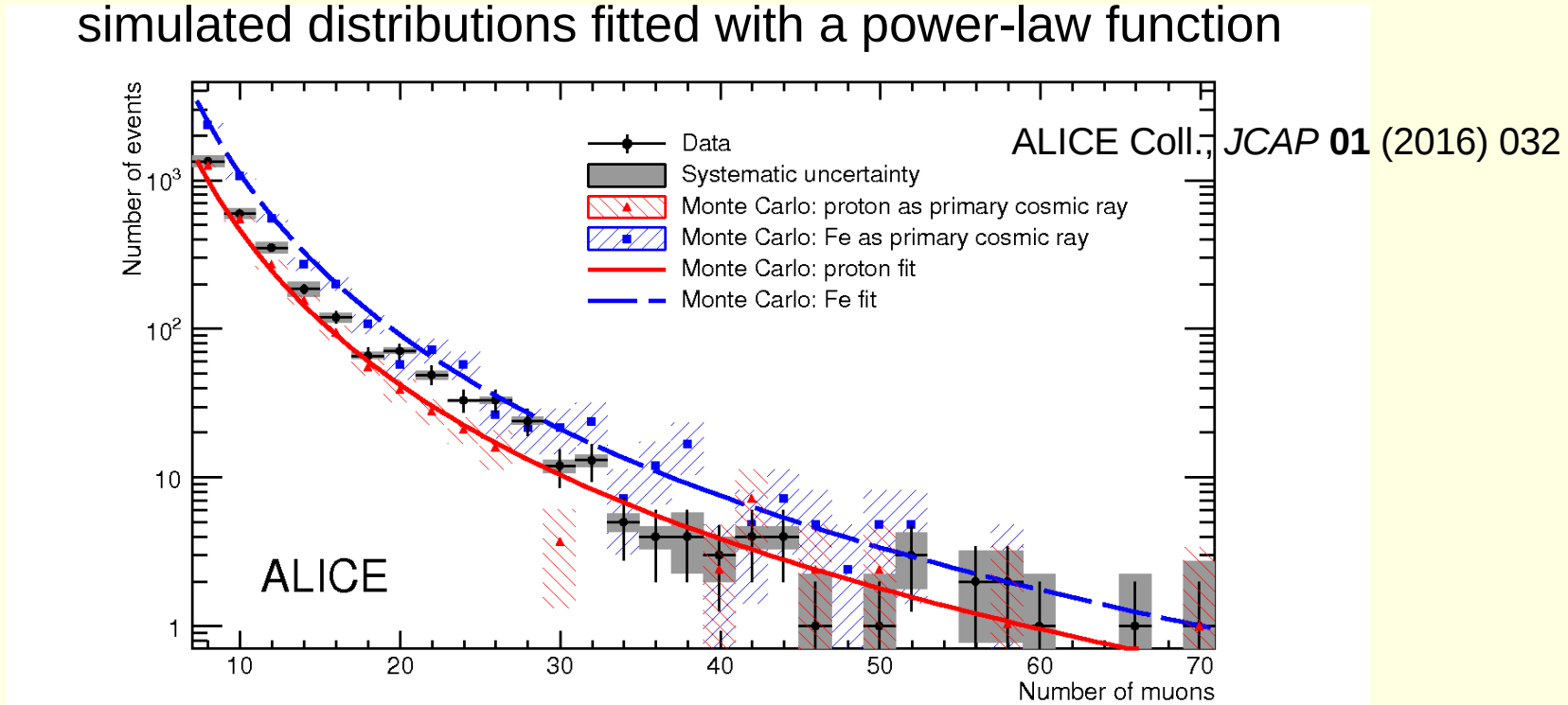
# Monte Carlo simulation

- the total (all-particle) absolute flux of cosmic rays was extracted from J. R. Hörandel, *Astrop. Phys.* **19** (2003) 193-220
- the core of each shower was scattered at surface level with a flat random distribution in an area of  $205 \times 205 \text{ m}^2$  centered around the ALICE apparatus



# Comparison Data-MonteCarlo

Compare the MMD in the range  $7 < N_{\mu} < 70$  with the simulated distributions fitted with a power-law function



As expected the data are between the pure  $p$  composition (approaching it at low multiplicity) and the pure Fe composition (at higher multiplicity)

# Monte Carlo study of HMM events

- In 30.8 days, 5 HMM events were recorded, corresponding to a rate of  $1.9 \times 10^{-6}$  Hz
- To estimate the rate of these events
  - a simulation of 1 year of live time was performed
  - simplified Monte Carlo simulations show that only primaries with  $E > 10^{16}$  eV contribute to HMM events reconstructed in the ALICE TPC
  - so only primaries with  $10^{16} < E < 10^{18}$  eV are generated, and  $\theta < 50^\circ$
  - samples for both  $p$  and Fe primaries

# Monte Carlo results

Number of HMM events in 365 days of data taking

	Simplified MC		Full MC	
	proton	iron	proton	iron
CORSIKA 6990 QGSJET II-03	40	61	27	51
CORSIKA 7350 QGSJET II-04	41	72	30	52

- To reduce the statistical fluctuations, four additional simulations were performed, reusing the EAS sample and randomly assigning the core of each shower over the 205x205 m<sup>2</sup> area

# Monte Carlo results

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- Averaging the 5 samples
  - estimate the number of HMM in 1 year
  - reduce the statistical fluctuations
- Uncertainties are dominated by
  - statistical errors on real data
  - systematic errors on MC data
- Two sources of systematic errors in MC
  - the uncertainties in the generation parameters
  - the muon reconstruction algorithm

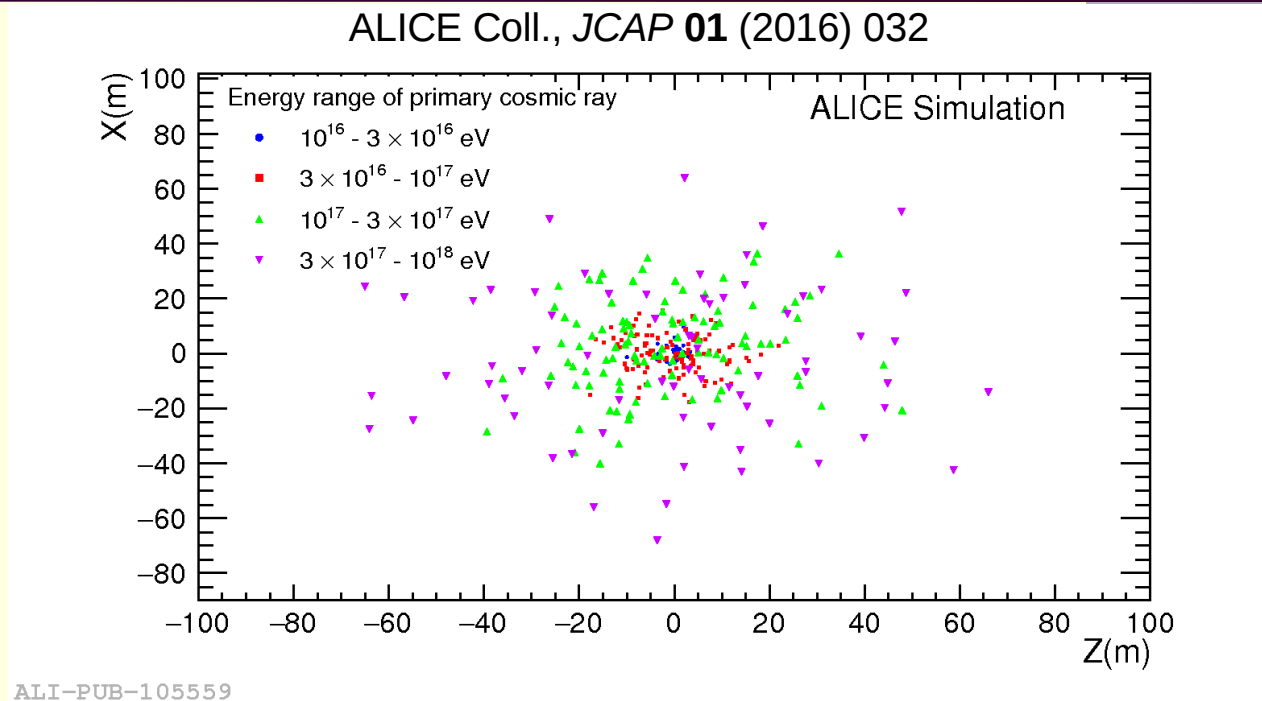


# HMM: comparison data-MC

HMM events	CORSIKA 6990		CORSIKA 7350		Data
	QGSJET II-03		QGSJET II-04		
	$p$	Fe	$p$	Fe	
Period [days per event]	15.5	8.6	11.6	6.0	6.2
Rate [ $\times 10^{-6}$ Hz]	0.8	1.3	1.0	1.9	1.9
Uncertainty (syst+stat) (%)	25	25	22	28	49

- The rate of HMM events can be reproduced using the latest hadronic interaction models and a reasonable CR primary flux
- Pure Fe primary composition (i.e. heavy nuclei composition) seems to better reproduce the HMM events
  - though the large uncertainty in the measured rate prevents a firm conclusion about the origin of these events
- Consistent with the fact that HMM events are generated by primaries with  $E > 10^{16}$  where the composition is dominated by heavier elements

# Core location of MC HMM events



Core location of simulated EAS giving  $> 100$  muons in the ALICE TPC, for Fe primaries with  $10^{16} < E < 10^{18}$  eV corresponding to 5 years of data taking

on average shower core falls farther from ALICE location for  $E > 3 \times 10^{17}$  eV

# Conclusions – I

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- In 2010-2013 ALICE took 30.8 days of cosmic-ray data
- The MMD at low and intermediate multiplicity is well reproduced by Monte Carlo simulations using CORSIKA 6990 with QGSJET II-30 model
- ALICE results suggest a mixed ion primary CR composition with an average mass increasing with energy

# Conclusions – II

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- 5 HMM events ( $N_{\mu} > 100$ ) were recorded in the same 30.8 days data taking period
- The observed rate is consistent with the predictions of CORSIKA 7350 with QGSJET II-04 model using a pure Fe primary composition and energy  $> 10^{16}$  eV
- For the first time the rate of HMM events has been well reproduced using conventional hadronic interaction models

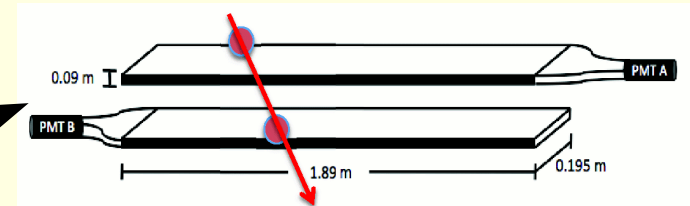
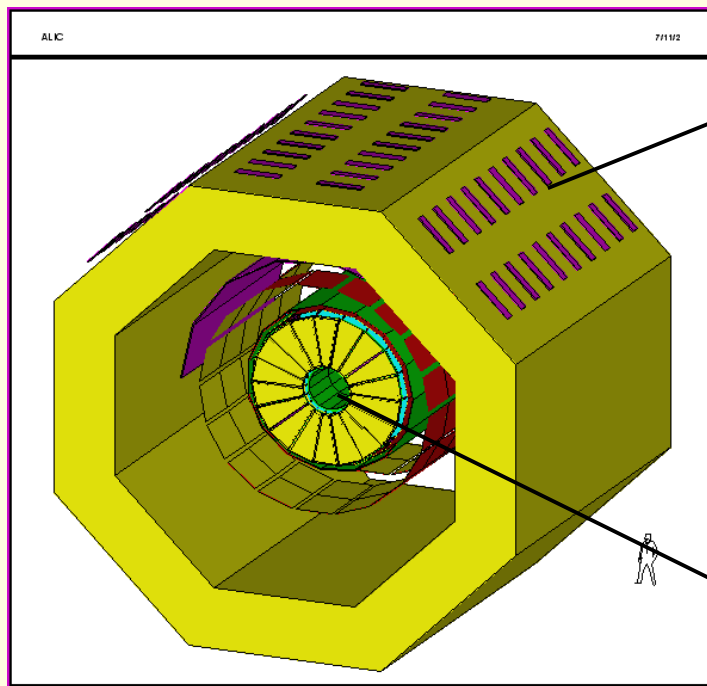
# Outlook

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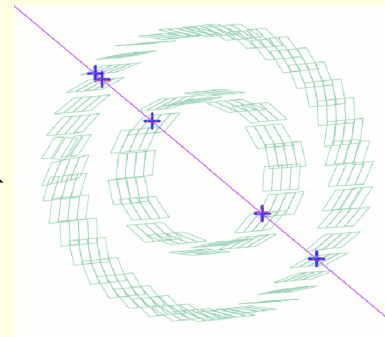
- ALICE will continue to take cosmic data during LHC Run2 (2015 onward)
  - during no-beam periods
  - a dedicated trigger to select HMM events during  $pp$  collision runs has been implemented (and tested during Run1)
    - in 2015 about 40 days of live time were collected
- The aim is to study HMM events in greater detail
- Other cosmic-ray topics (e.g. cosmic muon charge ratio) could also be performed with larger statistics

# BACKUP

# ACORDE and SPD triggers

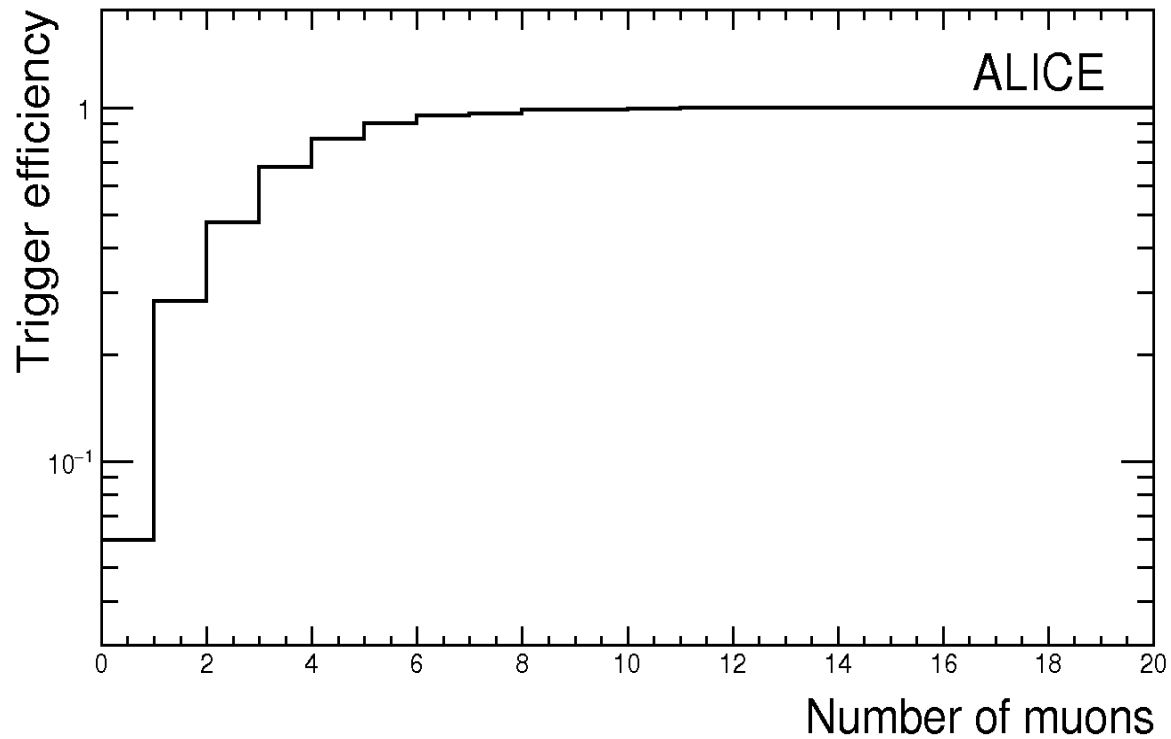


The ACORDE cosmic trigger requires a 4-fold coincidence of modules



The SPD cosmic trigger requires a coincidence between top and bottom halves of outer layer

# TOF trigger efficiency



ALI-PUB-105539



# Monte Carlo results

Number of HMM events in 365 days of data taking

Run	CORSIKA 6990 QGSJET II-03				CORSIKA 7350 QGSJET II-04			
	Simple MC		Full MC		Simple MC		Full MC	
	$\rho$	Fe	$\rho$	Fe	$\rho$	Fe	$\rho$	Fe
1	40	51	27	51	41	72	30	52
2	40	64	24	42	42	88	32	71
3	31	43	25	31	48	78	29	62
4	26	52	20	34	46	84	35	61
5	33	64	22	53	36	83	31	58

Two sources of systematic errors:

- the muon reconstruction algorithm
- the uncertainties in the generation parameters