



The **DarkSide** physics program and its recent results

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(on behalf of the DarkSide collaboration)

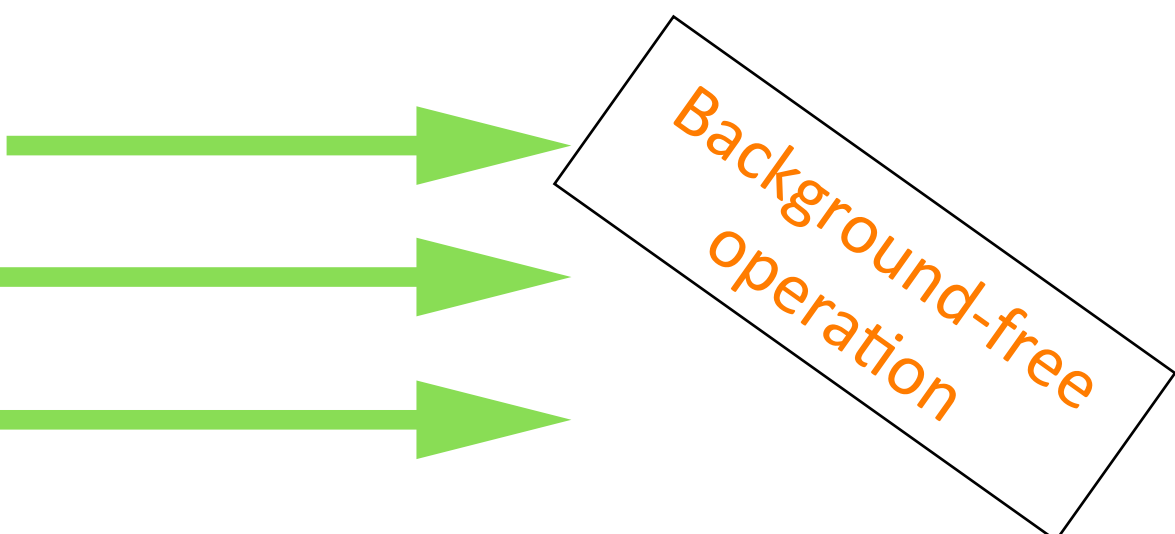
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07 Mar 2016 - Les Rencontres de Physique de la Vallée d'Aoste – La Thuile (Italy)

DarkSide Keywords

- ➔ Direct detection of dark matter
- ➔ Wimp-nucleus scattering in **liquid Argon**
- ➔ Dual-phase Time Projection Chambers (**TPC**)
- ➔ **Multi-stage** approach
- ➔ At Laboratori Nazionali del Gran Sasso (**LNGS**) in central Italy
rock coverage: ~3800m w.e.
- ➔ **Very low** intrinsic **background** levels
- ➔ Electron recoil **discrimination**
- ➔ Neutron **active** suppression



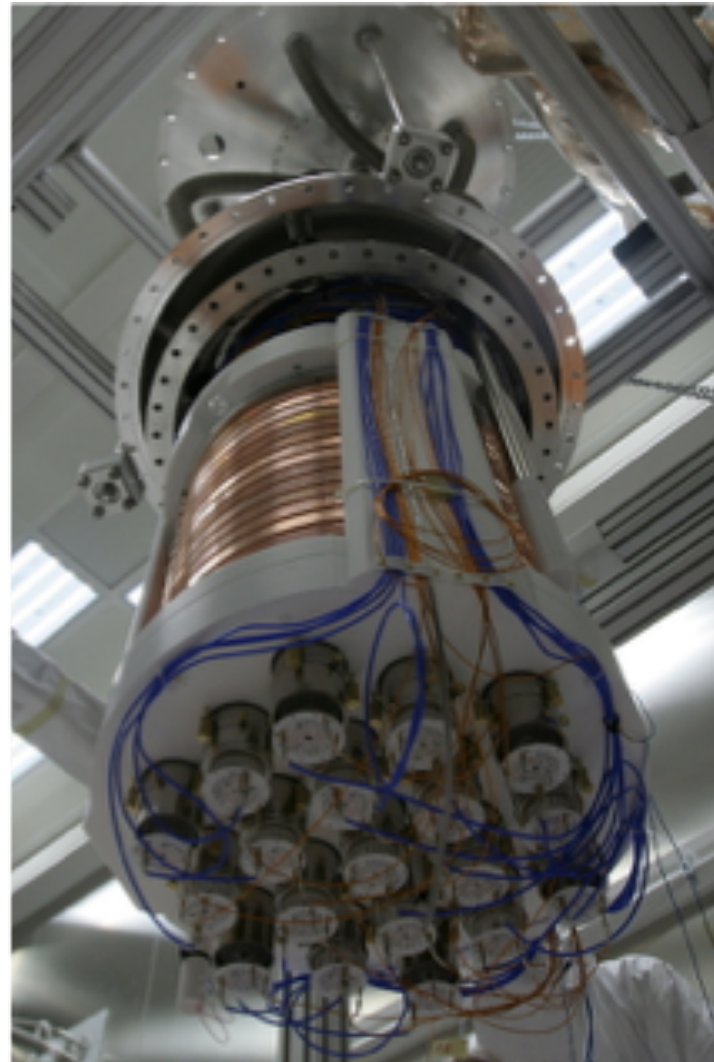
Background-free
operation

DarkSide Multi-stage Program



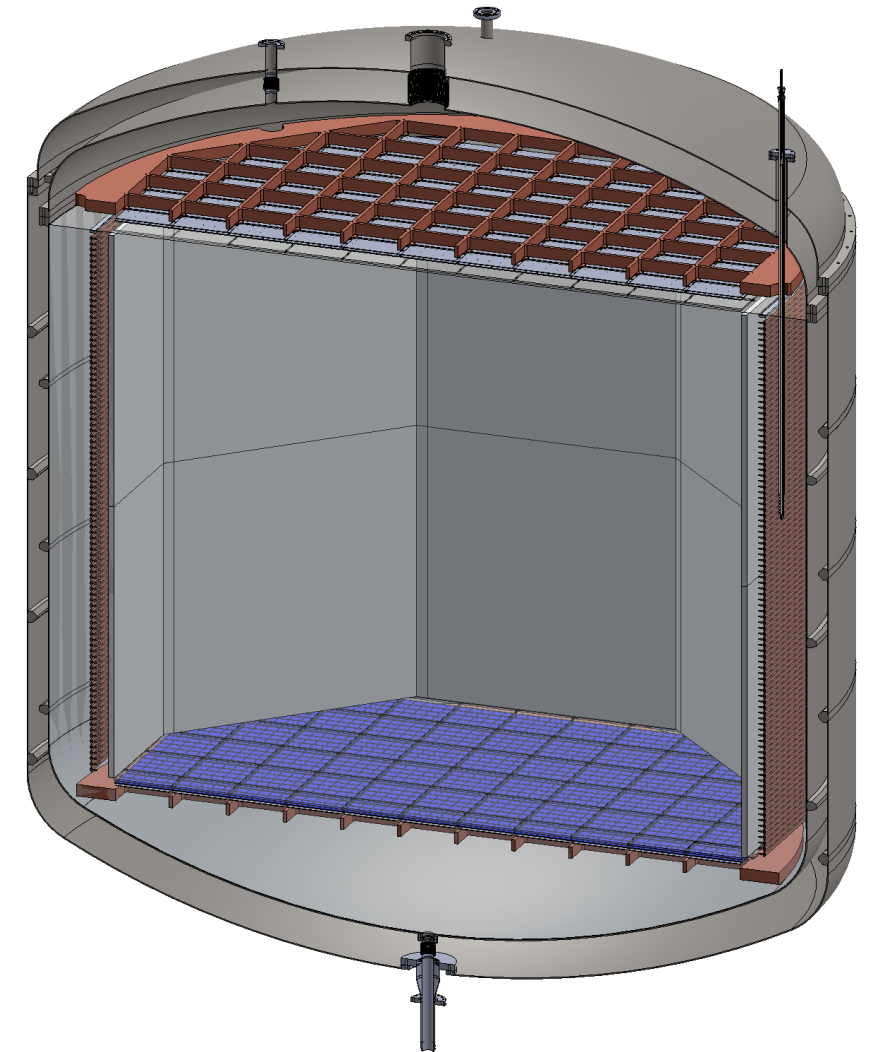
DarkSide-10

Prototype detector



DarkSide-50

First physics detector
 $\sim 10^{-45} \text{cm}^2$ @100GeV

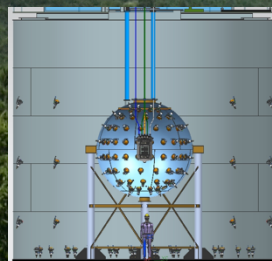


DarkSide-20k

Future **multi-ton** detector
 $\sim 10^{-47} \text{cm}^2$ @100GeV

Gran Sasso

3800 m w. e.



Deep underground location at LNGS, Italy.

DarkSide-50

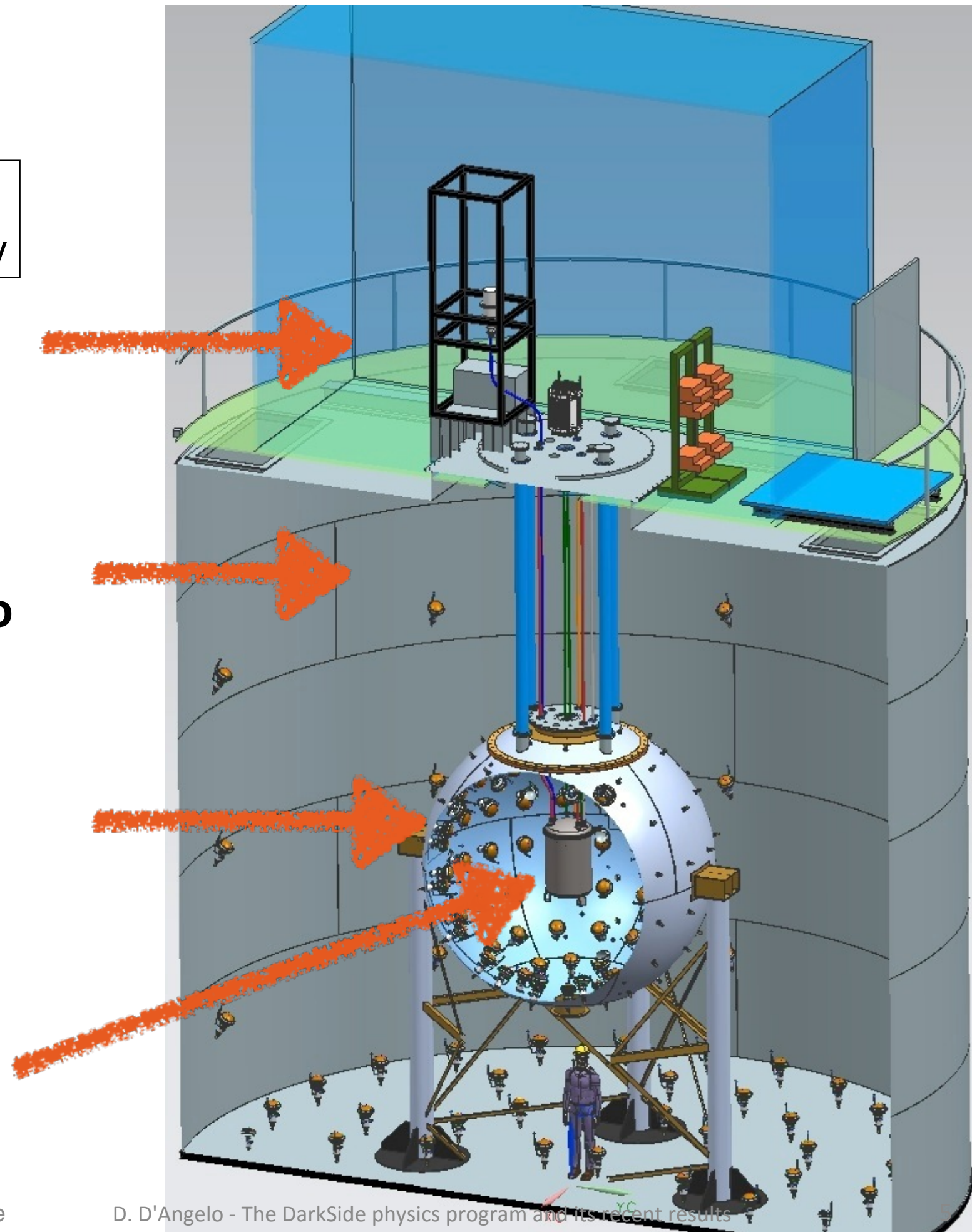
Housed in CTF
Borexino Counting Test Facility

Radon-free (Rn levels < 5 mBq/m³)
Clean Room

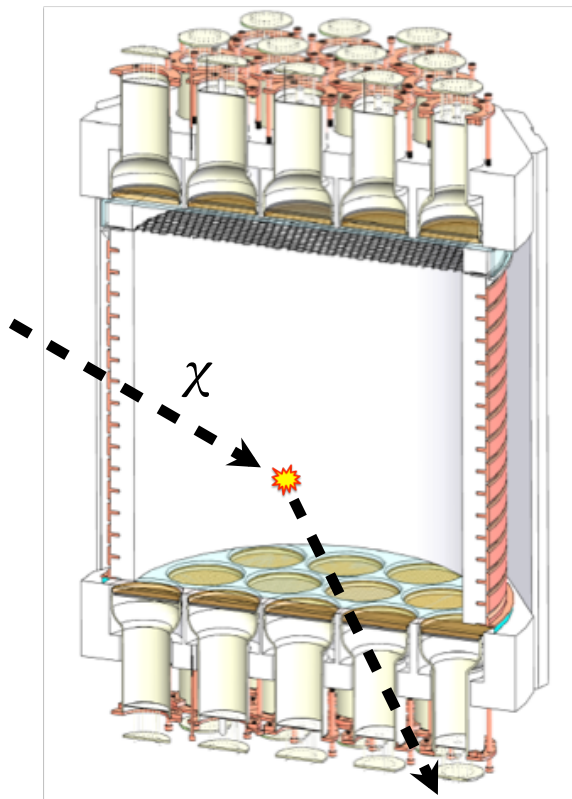
1,000-tonne Water-based
Cherenkov **Cosmic Ray Veto**

30-tonne Liquid Scintillator
Neutron and γ 's Veto

Inner detector **TPC**

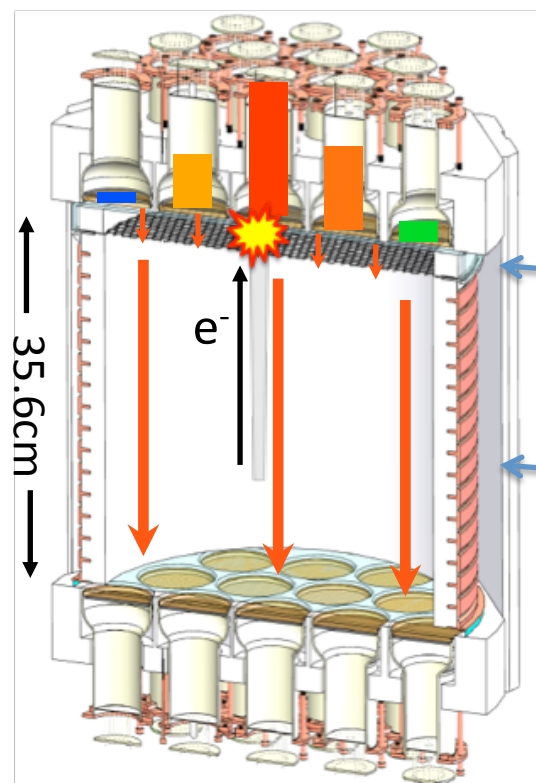
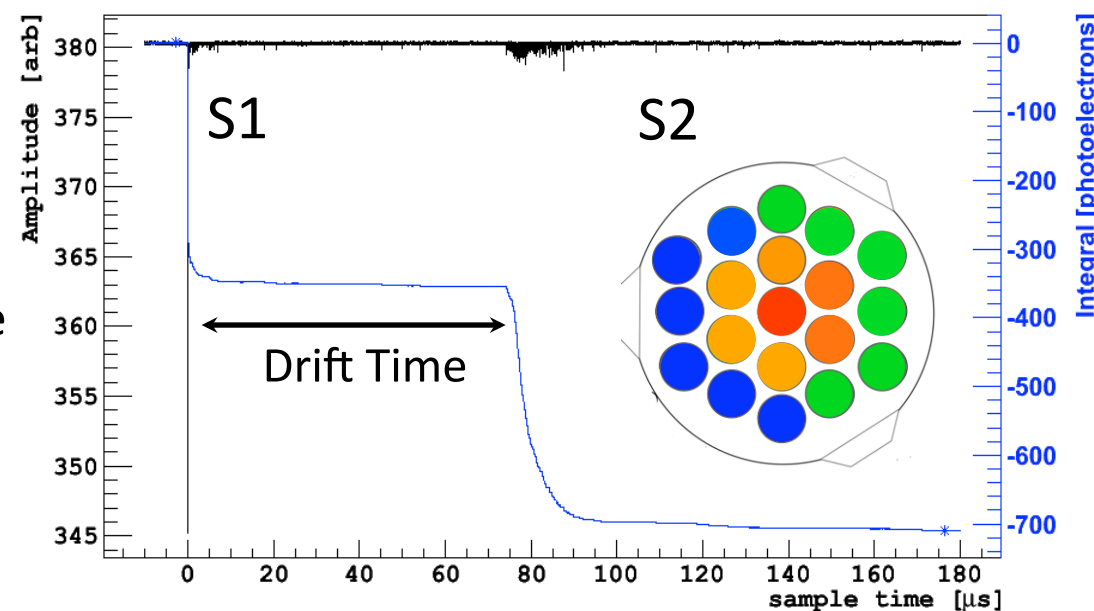


Two Phase Argon TPC



Total Ar mass: 145kg
Active mass: 49.4kg
 Fiducial mass: 44.9kg

19x2 3" PMTs
 ~20% photocathode coverage
 ~60% of end plate surface



Gas Ar ($E_{lum} \sim 4200$ V/cm)

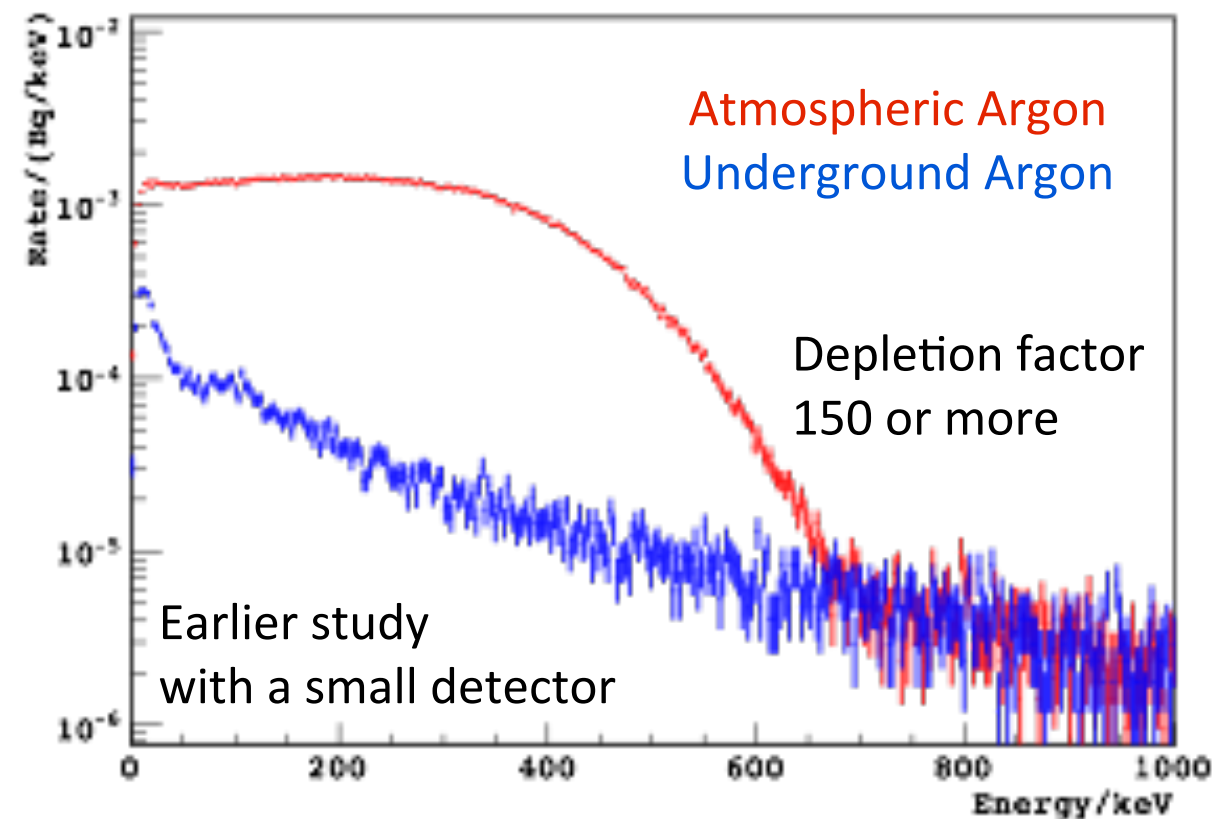
Liquid Ar ($E_{drift} \sim 200$ V/cm)

- Nuclear Recoil **excites** and **ionizes** the liquid argon, producing **scintillation** light (S1) that is detected by the photomultipliers
- The electrons are extracted into the gas region, where they induce **electroluminescence** (S2)
- The time between the S1 and S2 signals gives the vertical position.
- x-y position of events are reconstructed from fraction of S2 in each PMT.

Electron drift lifetime > 5 ms, compared to max. drift time of ~ 375 μ s.
 Electron drift speed = 0.93 mm/ μ s

Underground Ar

- Intrinsic ^{39}Ar radioactivity in **atmospheric argon** (~ 1 Bq/kg) is the primary background for argon-based detectors
- ^{39}Ar activity sets the dark matter detection threshold at low energies (where pulse shape discrimination is ineffective)
- ^{39}Ar is a cosmogenic isotope, produced via $^{40}\text{Ar}(n,2n)^{39}\text{Ar}$. It is a β emitter with $Q=565$ keV and $T_{1/2}=269$ years
- The activity in argon from **underground sources** can be significantly lower
- Identified source in Colorado, measured to have < 6.5 mBq/kg, i.e. > 150 times lower rate compared to atmospheric argon.
- Plant (including cryogenic distillation at FNAL) produces ~ 0.5 kg/d
- First results with UAr in this talk



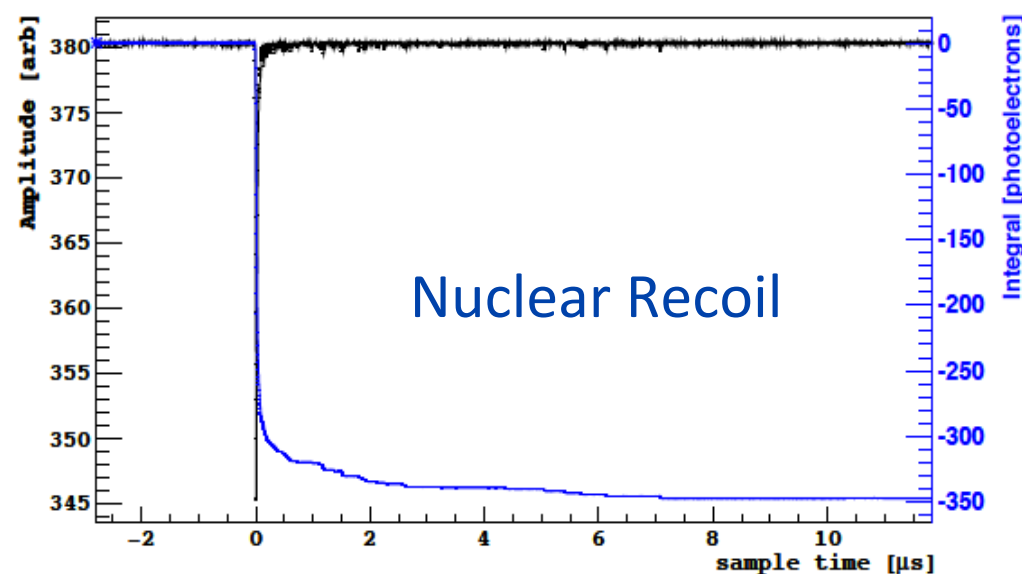
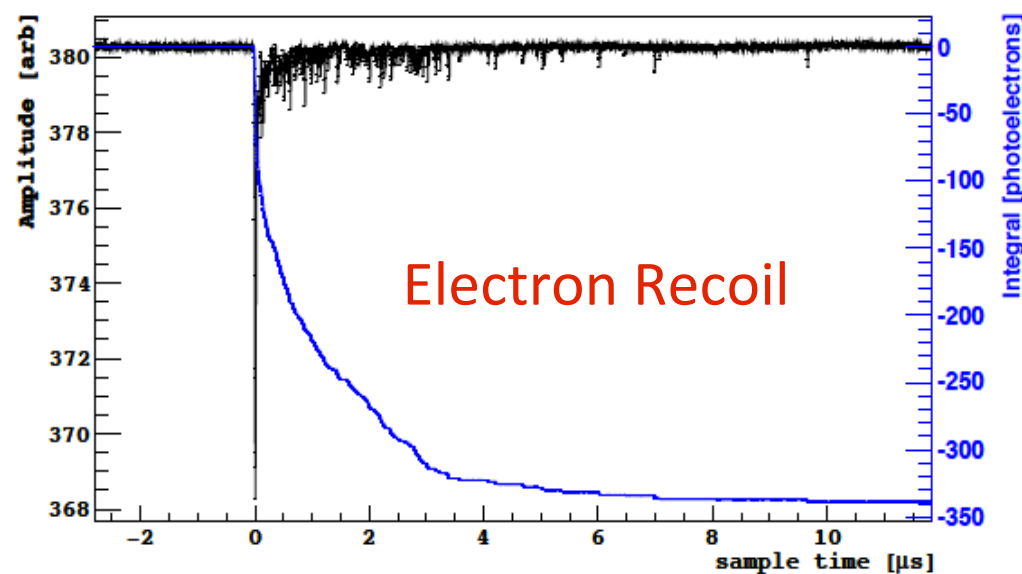
[arXiv:1204.60111 \[physics.ins-det\]](https://arxiv.org/abs/1204.60111)

Pulse Shape Discrimination

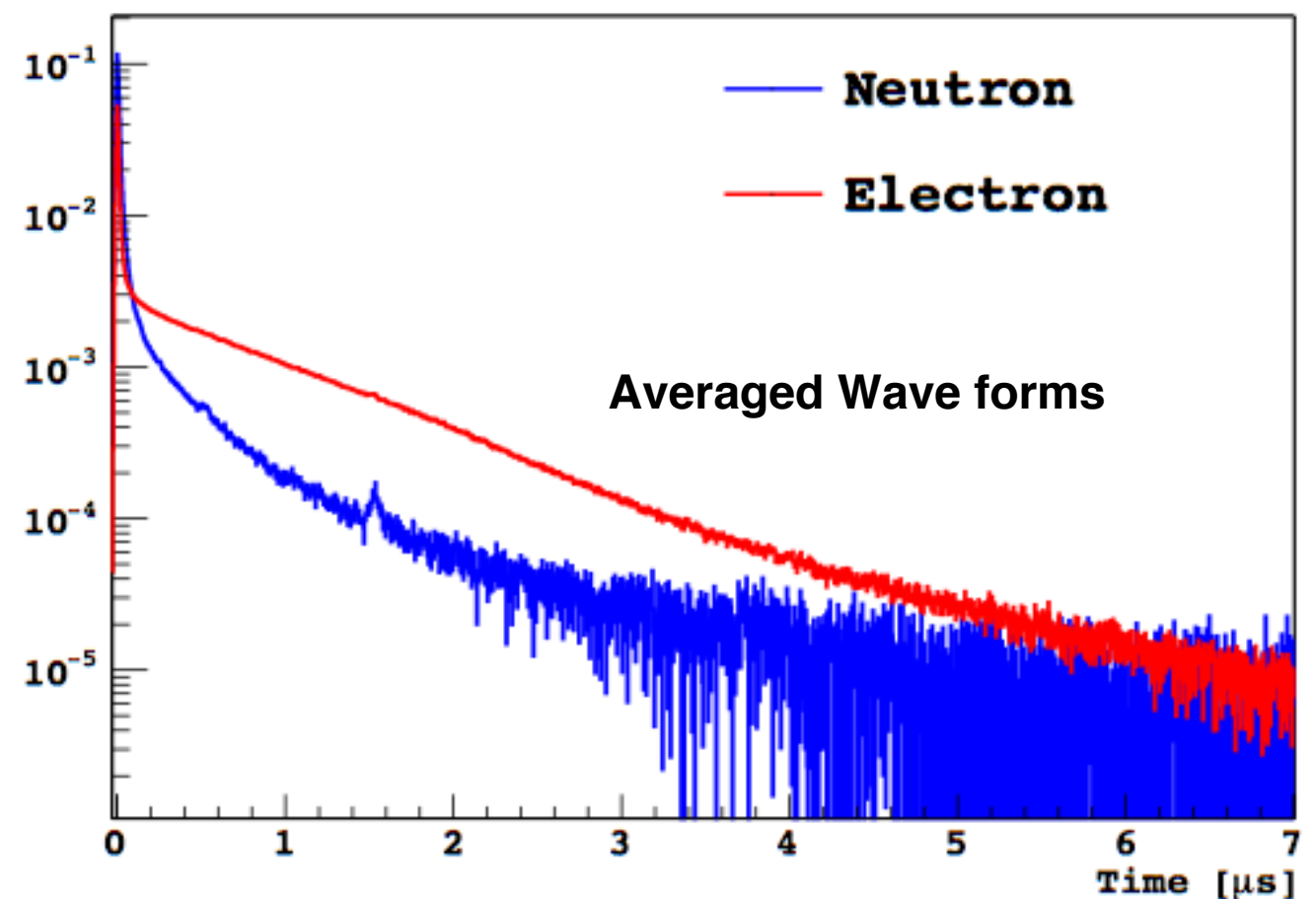
Electron and nuclear recoils produce different excitation densities in the argon, leading to different **ratios of singlet and triplet excitation states**

$T_{\text{singlet}} \sim 7 \text{ ns}$

$T_{\text{triplet}} \sim 1500 \text{ ns}$



$$F_{90} = \frac{\int_0^{90 \text{ ns}} dt f(t)}{\int_0^{\infty} dt f(t)} = \begin{cases} 0.3 \text{ ER} \\ 0.7 \text{ NR} \end{cases}$$

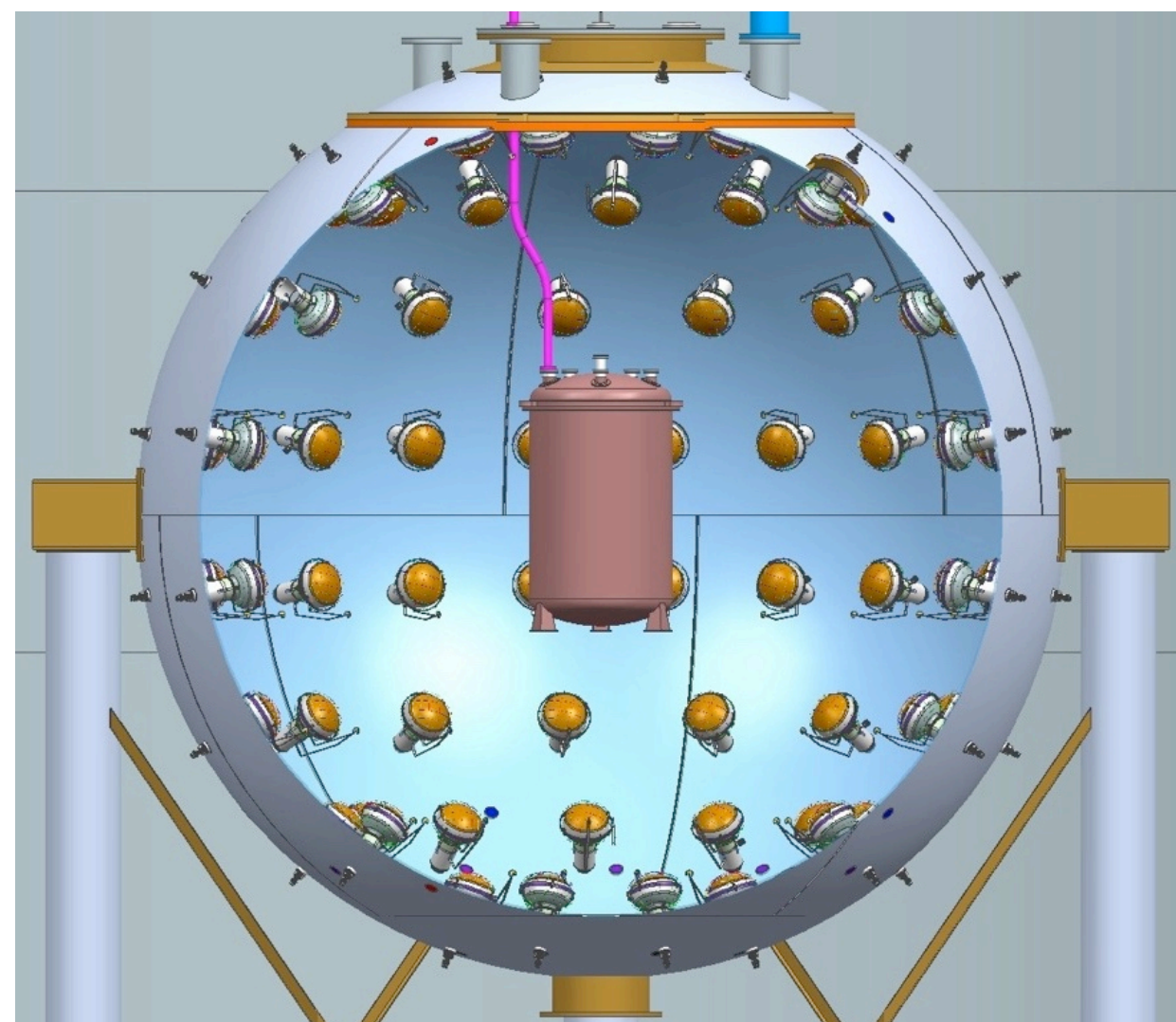


Liquid Scintillator Veto

Neutron
Rejection

Liquid scintillator allows coincident veto of **neutrons (and γ 's)** in the TPC and provides *in situ* measurement of the neutron background rate

- 4 m diameter sphere containing PC + TMB scintillator
- Instrumented with 110 8" PMTs



^{14}C content was too high ($\sim 98\%$ efficiency) to achieve design efficiency ($\sim 99.5\%$) after the first fill.

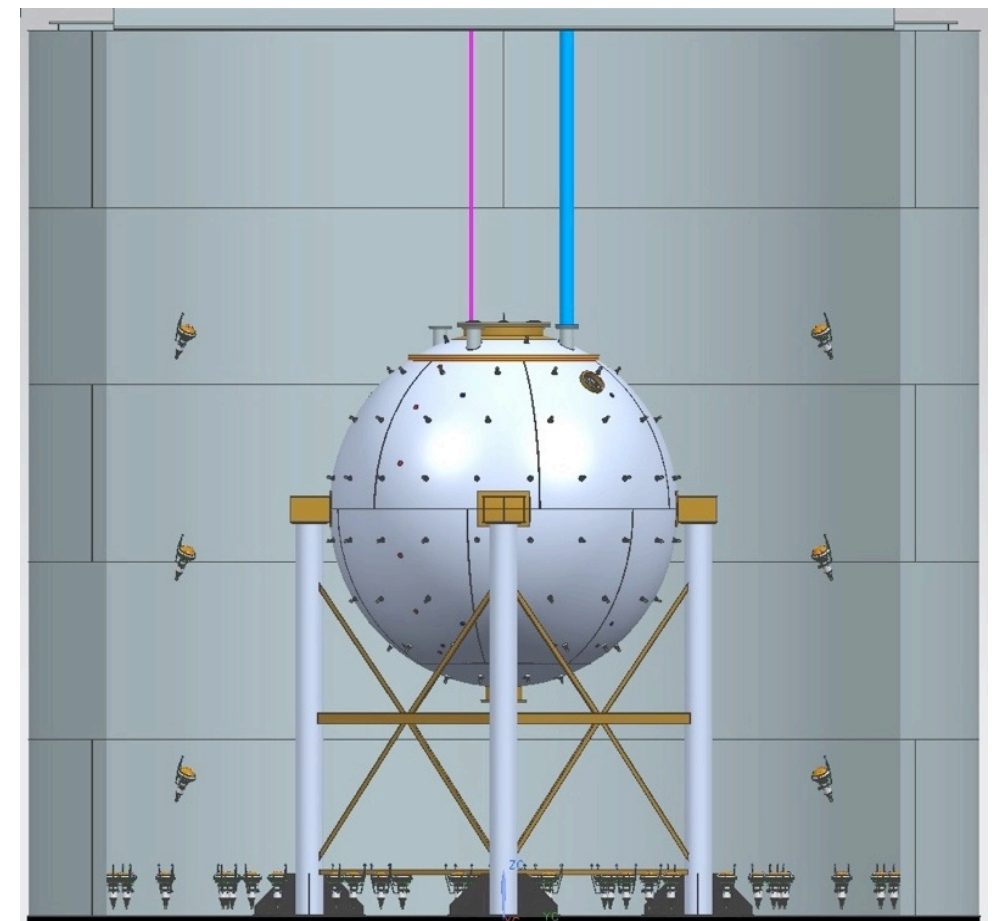
The TMB was replaced with new low ^{14}C TMB (Jan. 2015).

^{14}C activity decreased from **150 kBq** to **0.3 kBq**.

External Water tank

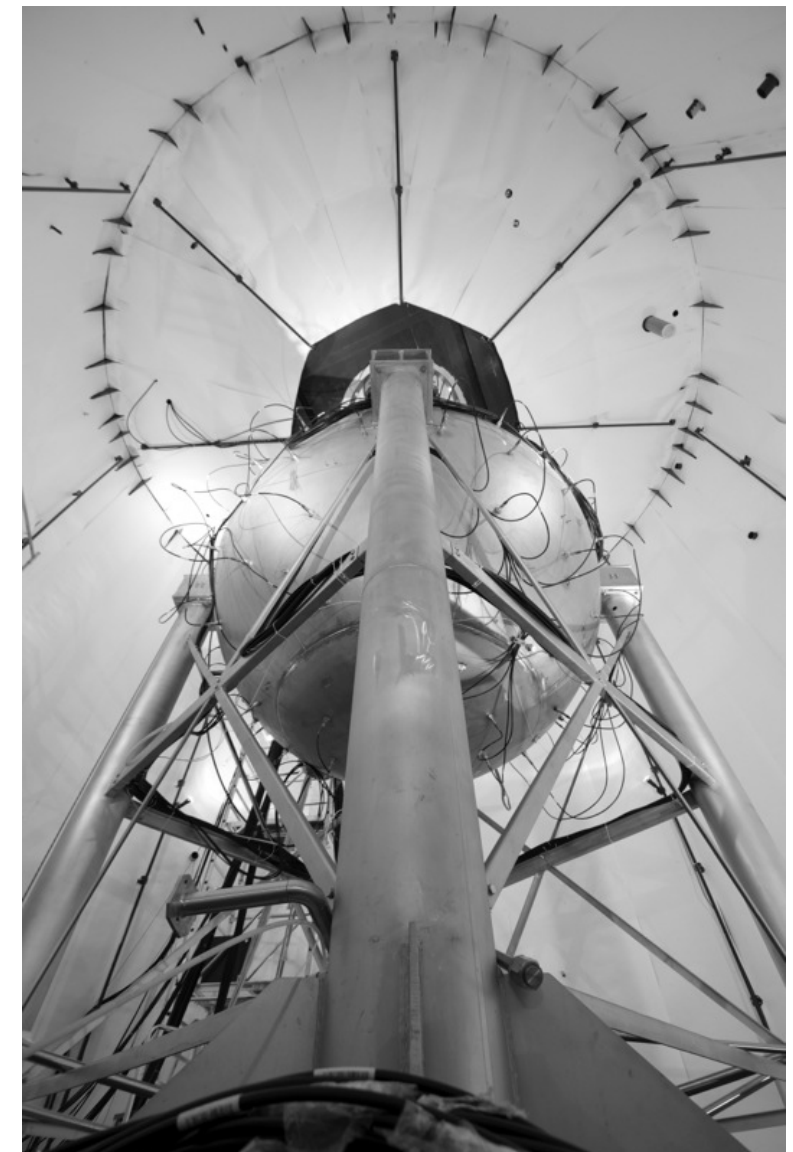
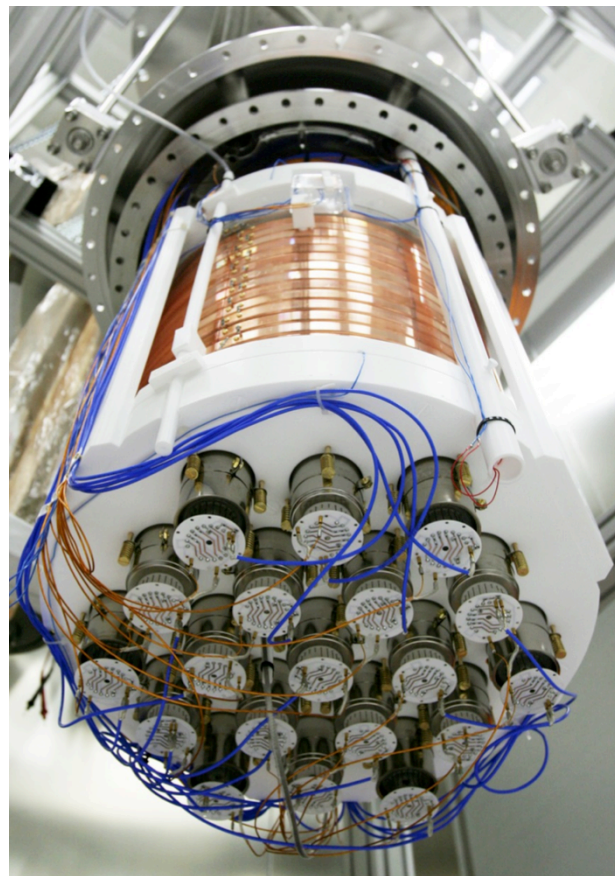
Cosmogenics
Rejection

- 80 PMTs within water tank (11 m diameter x 10 m height)
- Acts as a **muon and cosmogenic veto** (~ 99% efficiency)
- Provides **passive γ 's and neutron shielding**



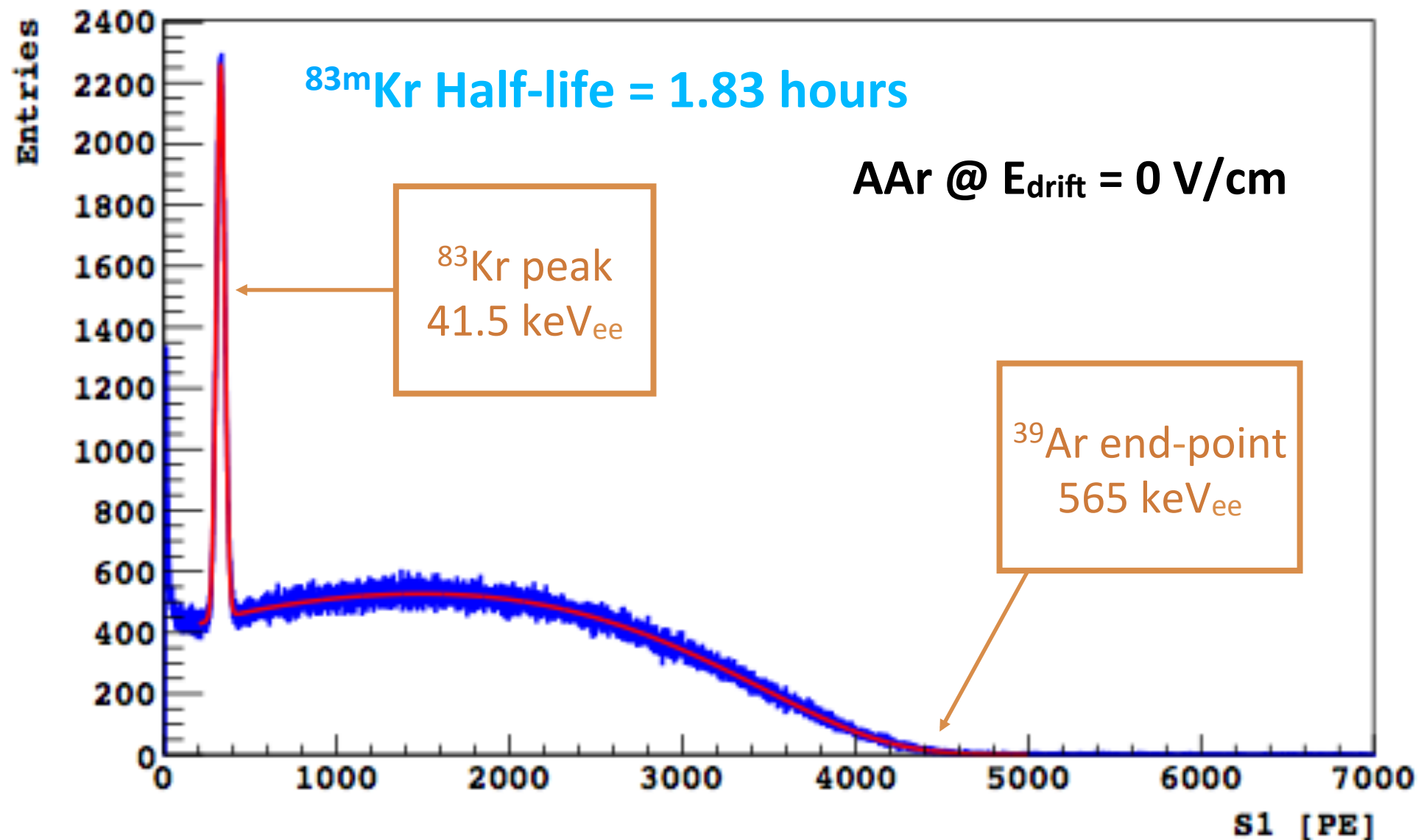
DS-50 Assembly

Sept - Oct 2013



TPC: Electron Recoils calibration

- ^{39}Ar (565 keV_{ee} end-point) present in AAr
- $^{83\text{m}}\text{Kr}$ gas deployed into detector (41.5 keV_{ee})

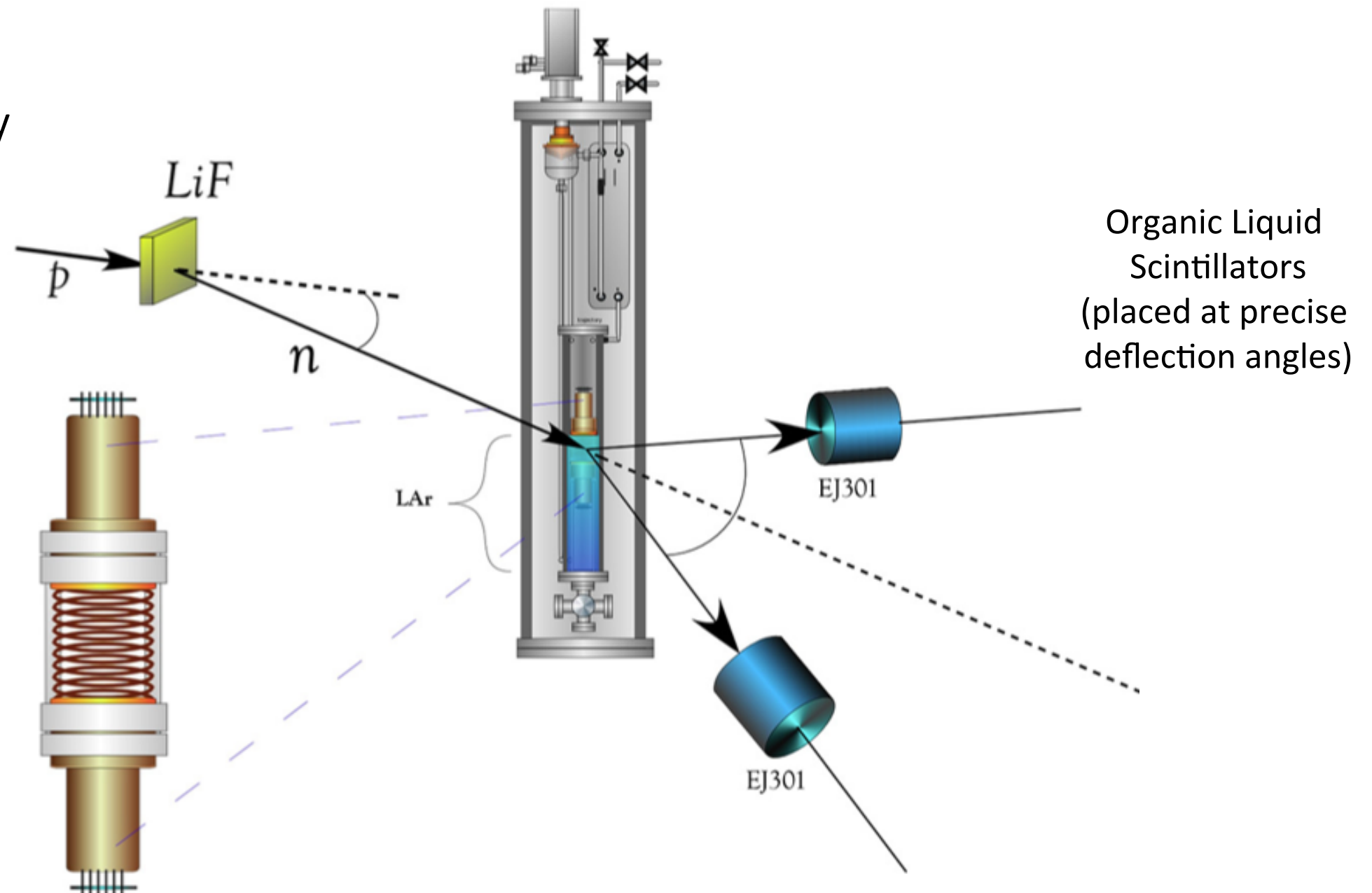


LIGHT YIELD: $8.1 \pm 0.2 \text{ PE/keV}_{\text{ee}}$ at zero field and $7.0 \pm 0.3 \text{ PE/keV}_{\text{ee}}$ at 200 V/cm

SCENE

(Scintillation Efficiency of Nuclear Recoils in Noble Elements)

Proton Beam at University
of Notre Dame



Liquid Argon TPC
(based on DarkSide design)

${}^7\text{Li}(p, n){}^7\text{Be}$ reaction produces low energy mono-energetic neutrons TOF measurement between target, LAr and organic scintillators allows clean identification of elastic neutron interactions of known energy

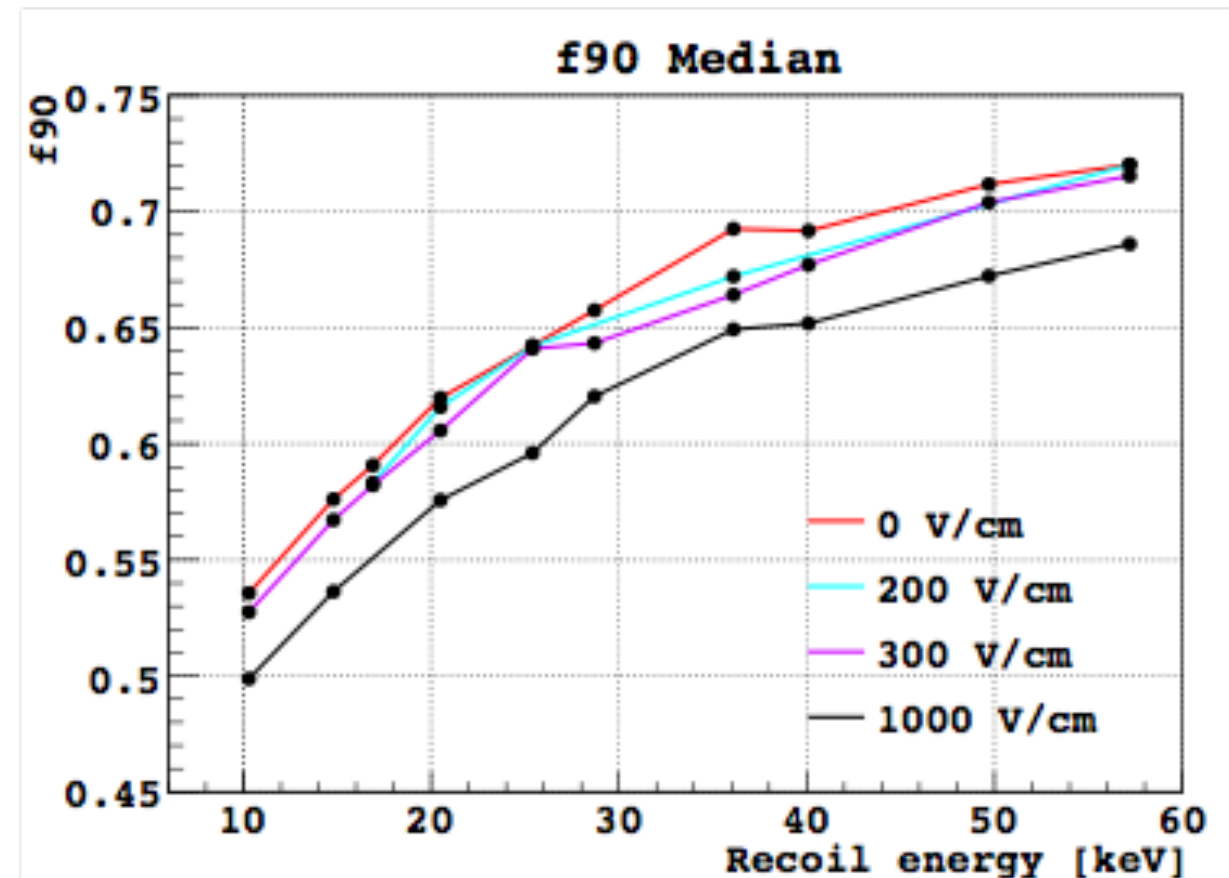
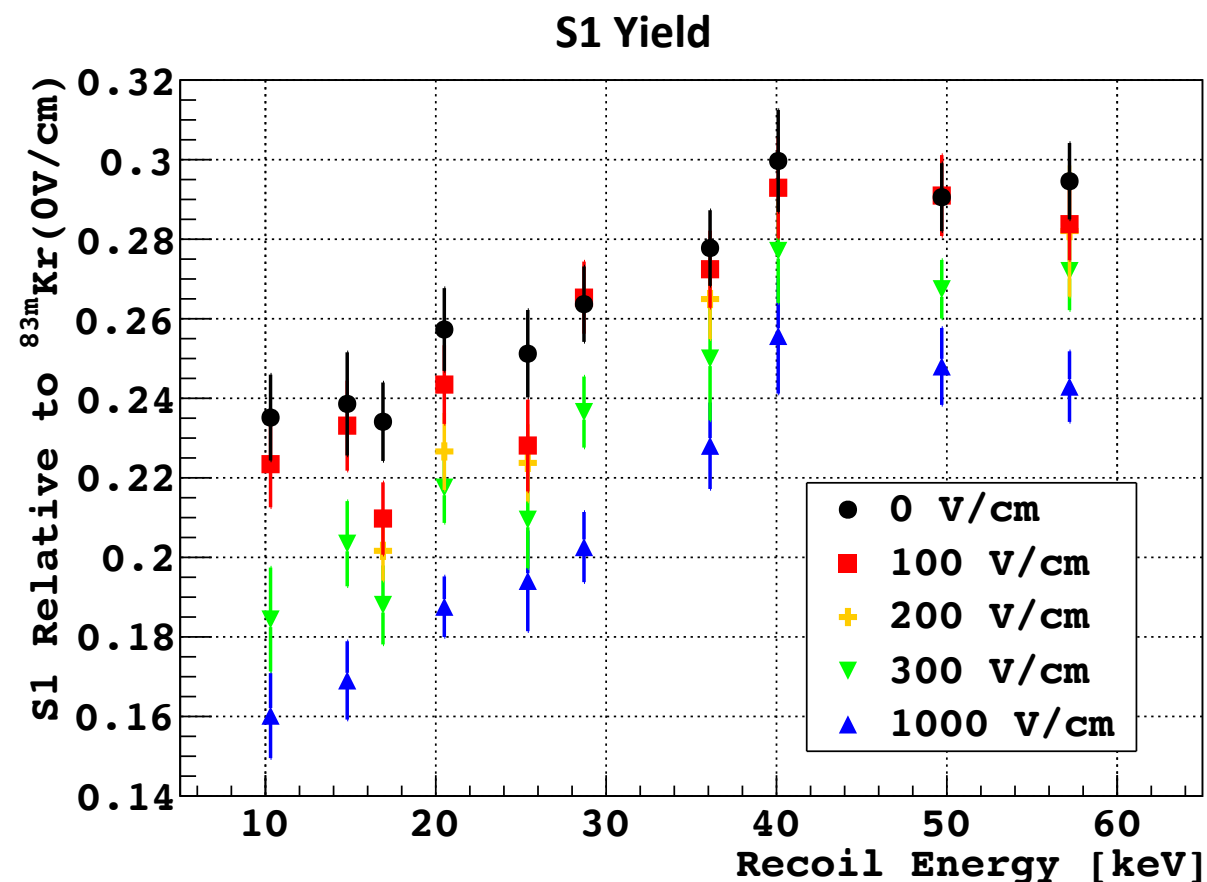
Nuclear Recoil

From **SCENE**:

1. nuclear recoil quenching
2. the F90 distribution

by processing SCENE data with DS-50 code and extrapolated to DS-50 detector along with the systematics.

Phys. Rev. D 91, 092007

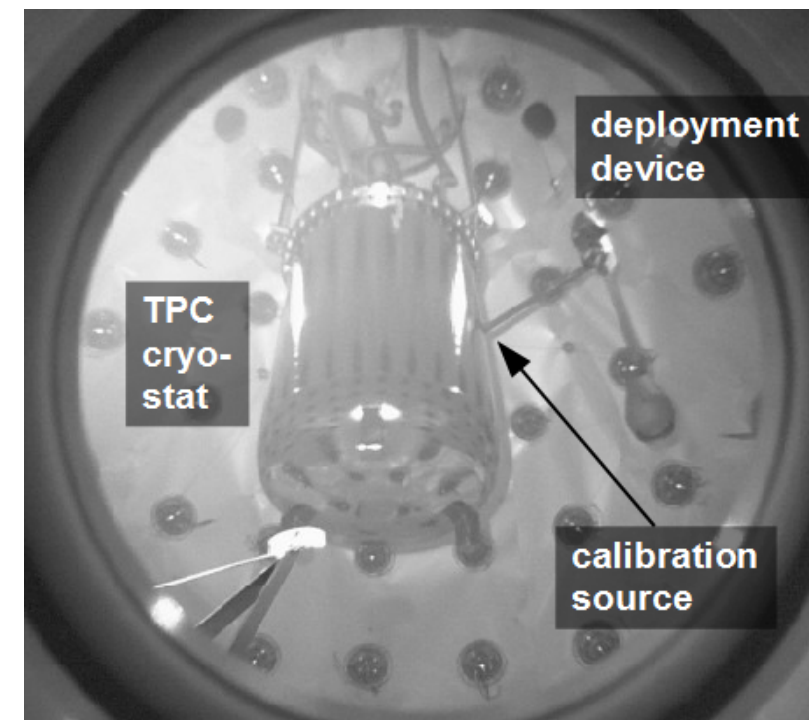


CALIS

CALibration Insertion System

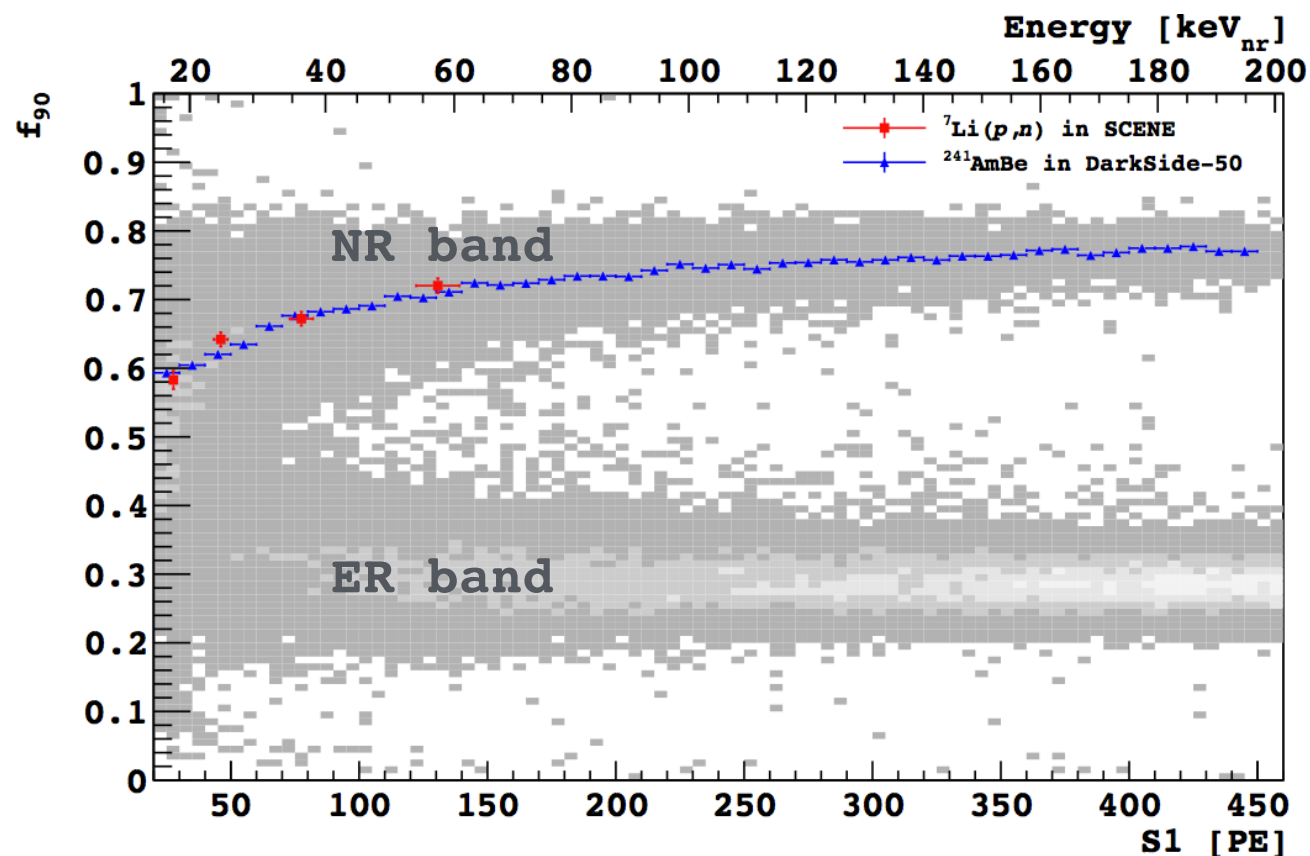
Calibrate both **TPC** and **Neutron veto**

- **Gamma sources:** ^{57}Co (122 keV), ^{133}Ba (356 keV), ^{137}Cs (663 keV)
- **Neutron source:** AmBe



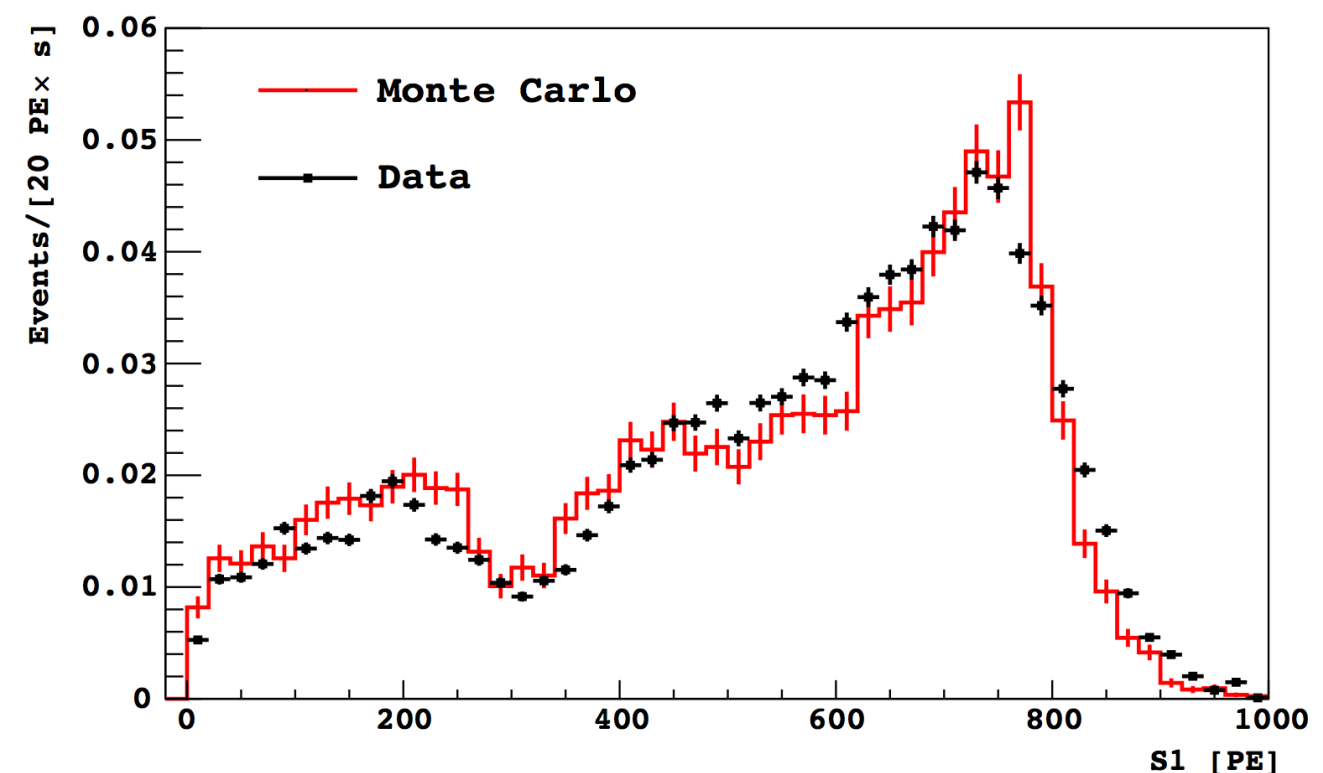
NR band matches with the points extrapolated from SCENE.

NR study (crosscheck of SCENE data)



Test of the MC code

DATA-MC comparison: ^{57}Co source next to the cryostat



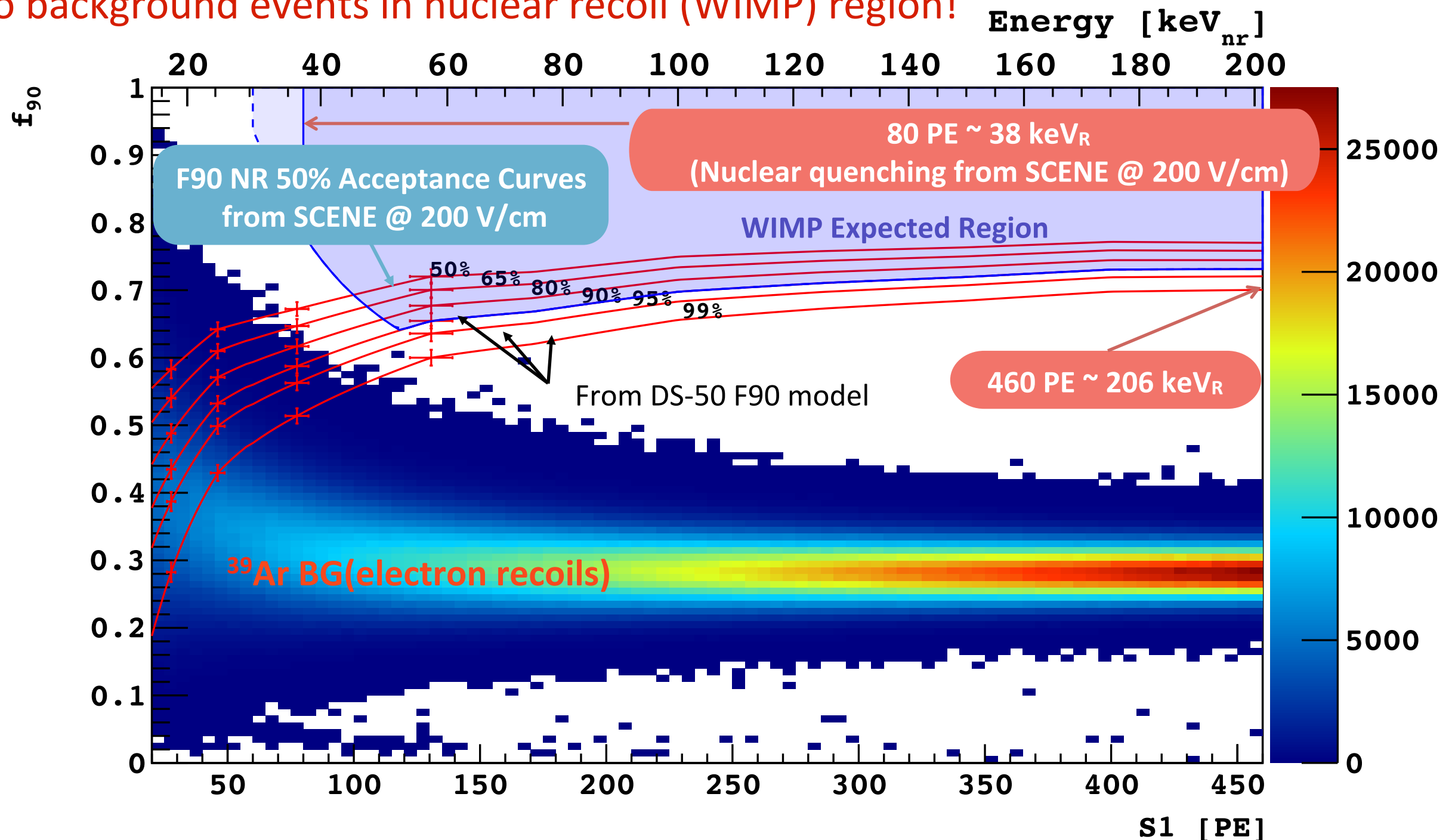
DarkSide-50 Timeline

- **Oct. 2013**: LArTPC, Neutron Veto and Muon Veto commissioned.
 - TPC filled with **atmospheric argon** (AAr).
- **Up to Jun. 2014**: data taken with high ^{14}C content in LSV.
 - **47.1 live days** (1422 kg·day fiducial) for the first physics result.
 - TMB (^{14}C) was removed to reduce the ^{14}C rate.
- **Oct. to Dec. 2014**: Calibration of TPC with radioactive sources.
- **Jan. 2015**: Add radiopure TMB at 5% concentration.
- **Mar. to Apr. 2015**: Fill with **UAr** and re-commissioning the detector.
- **Apr. to Aug. 2015**: Accumulate first data with **UAr** for **dark matter search**.
- Data taking is foreseen to last until approximately the end of 2017

Atmospheric Argon run (2014)

Exposure of $1422 \pm 67 \text{ kg} \cdot \text{day}$

No background events in nuclear recoil (WIMP) region!



Phys. Lett. B 743 (2015) 456

three cuts only: single-hit (1) interactions in the Fiducial Volume (2) and no energy deposition in the veto (3).

Underground Ar



Plant at Colorado

1. Extraction at Colorado (CO₂ Well)

Extract a crude argon gas mixture (Ar, N₂, and He)

2. Purification at Fermilab

Separate Ar from He and N₂



Distillation Column at Fermilab



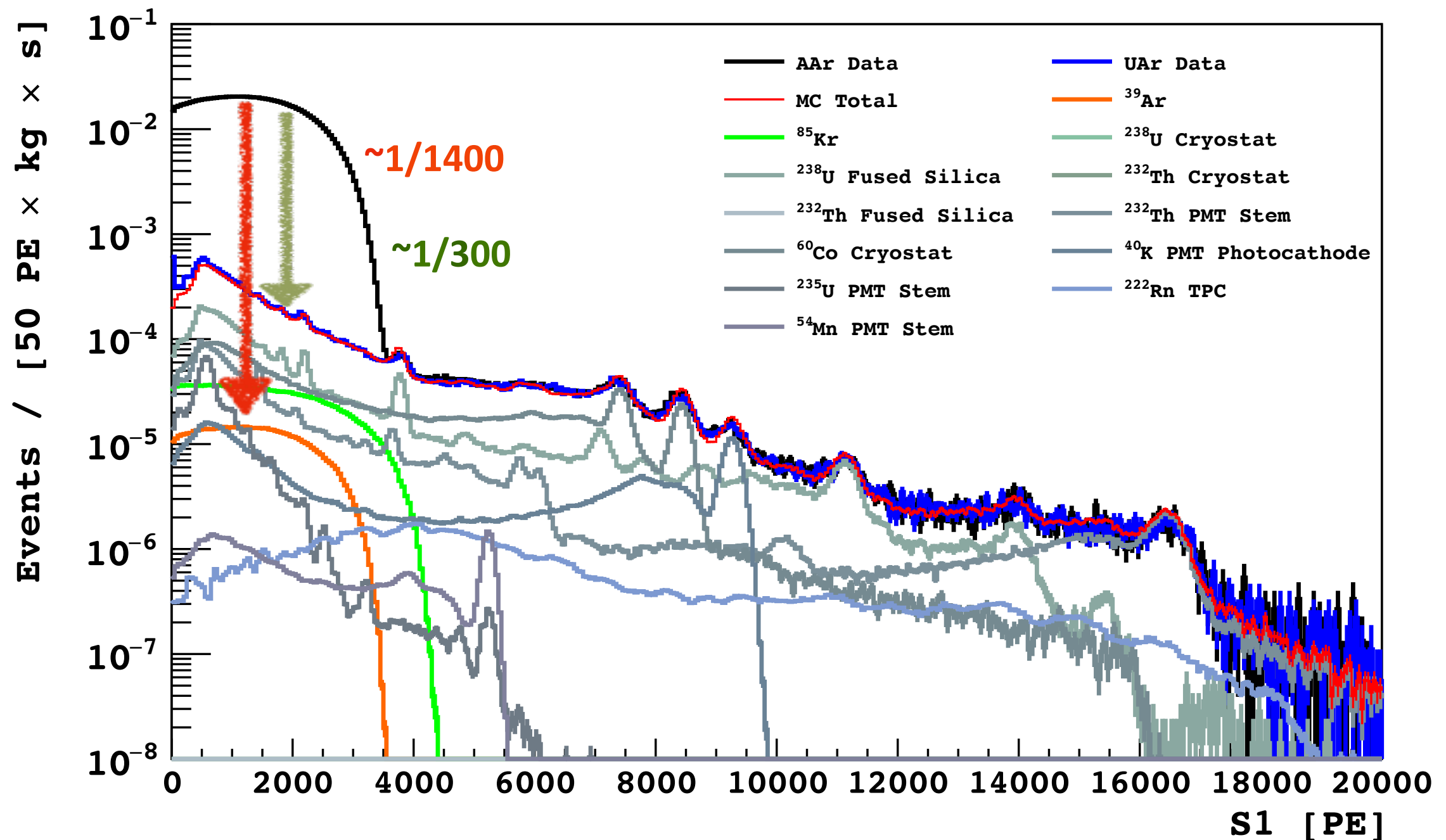
UAr bottles at LNGS

3. Arrived at LNGS

Ready to fill into DS-50

UAr First Results

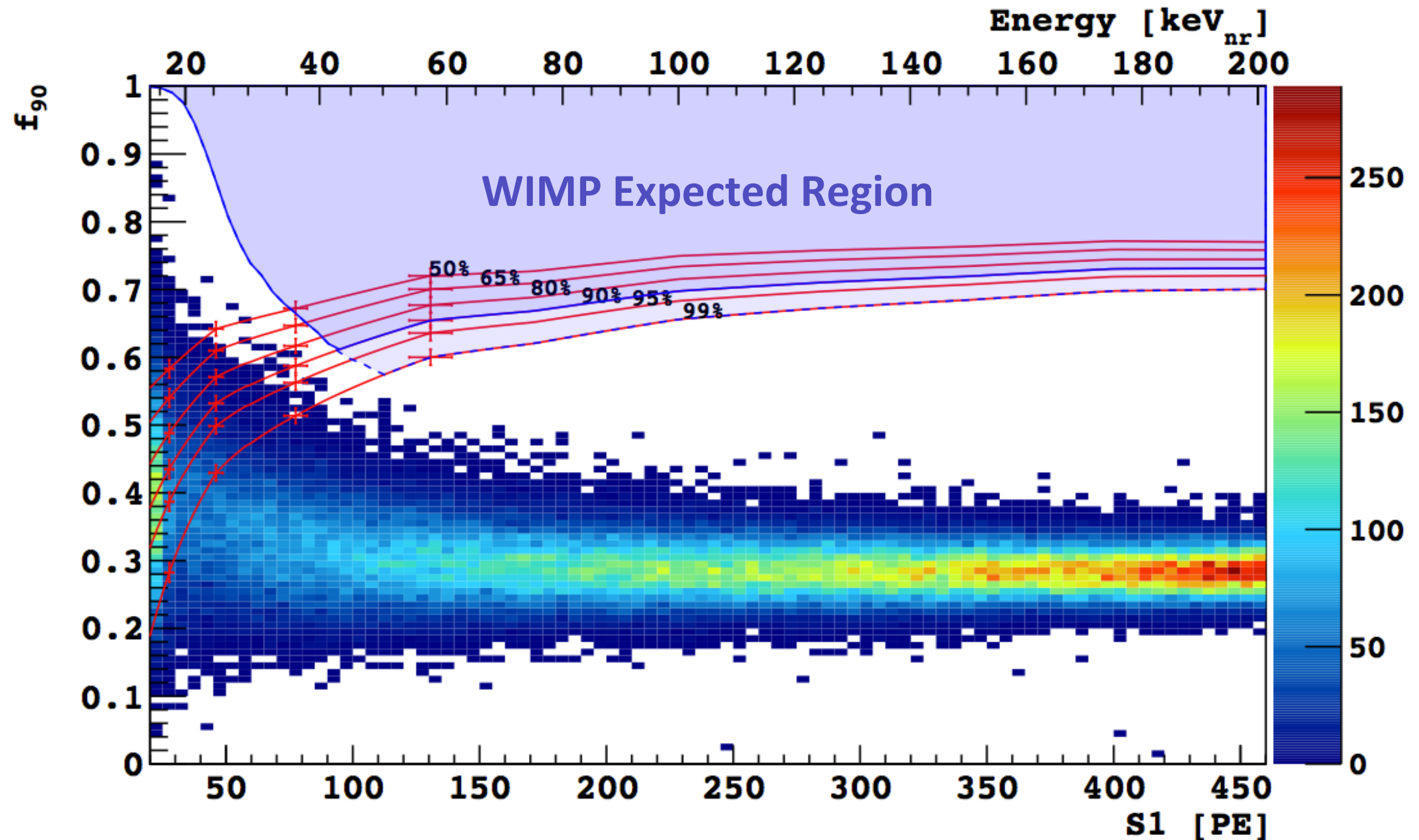
AAr vs UAr. Live-time-normalized S1 pulse integral spectra at Zero field.



Low level of ^{39}Ar allows extension of DarkSide program to ton-scale detector.

UAr First Results

No background events in nuclear recoil (WIMP) region!

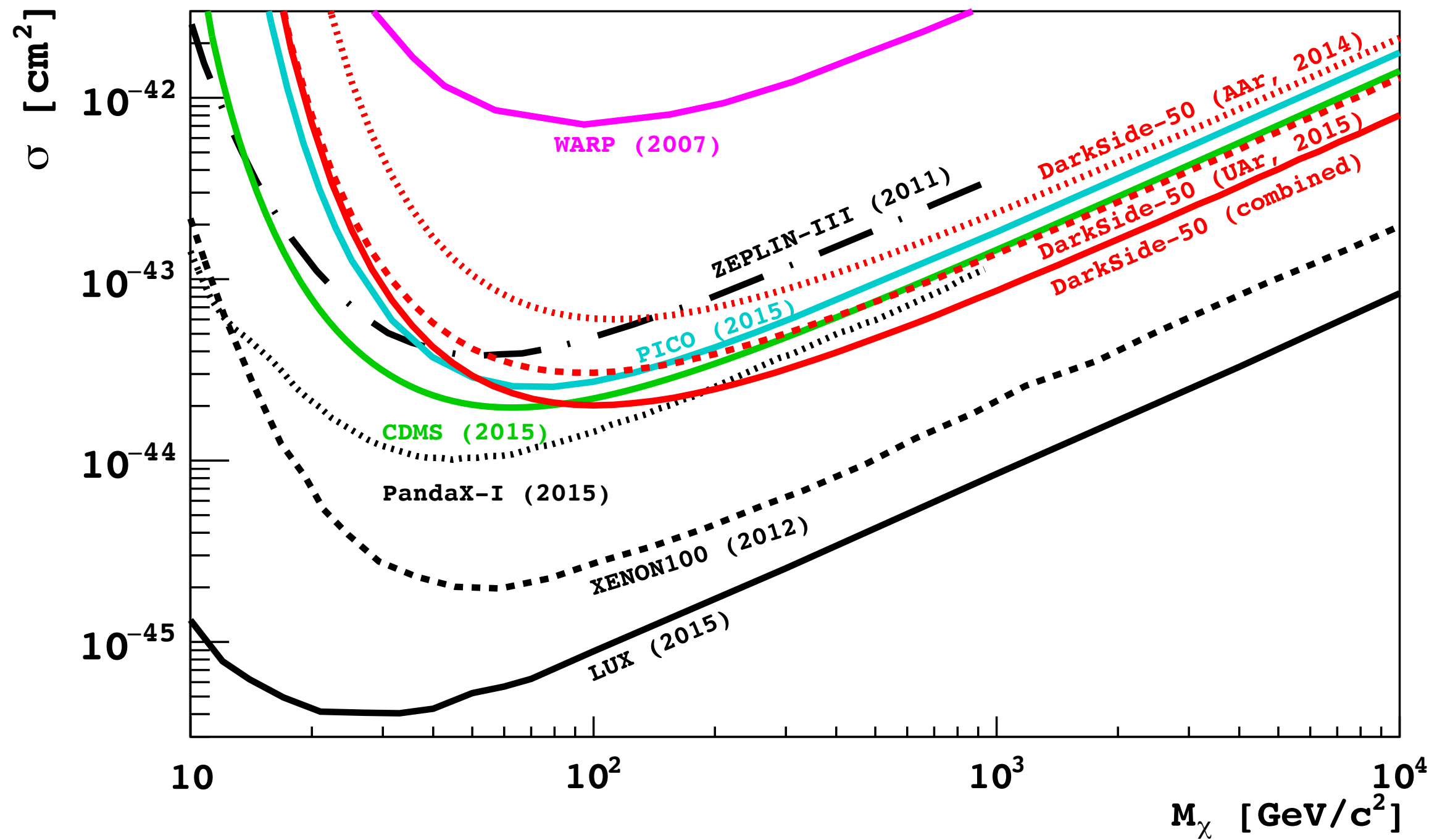


71 live-days after all cuts. (2616 ± 43) kg day exposure.

Single-hit interactions in the TPC, no energy deposition in the veto

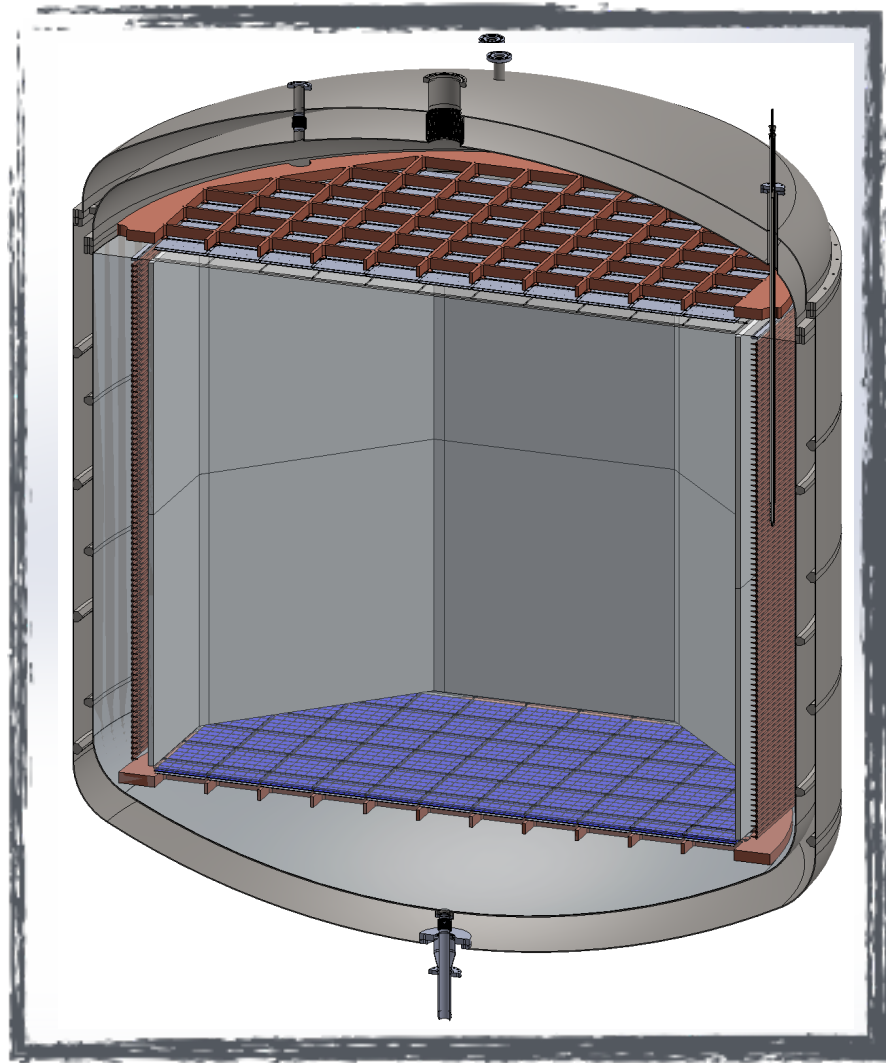
[arXiv:1510.00702](https://arxiv.org/abs/1510.00702)

UAr First Results



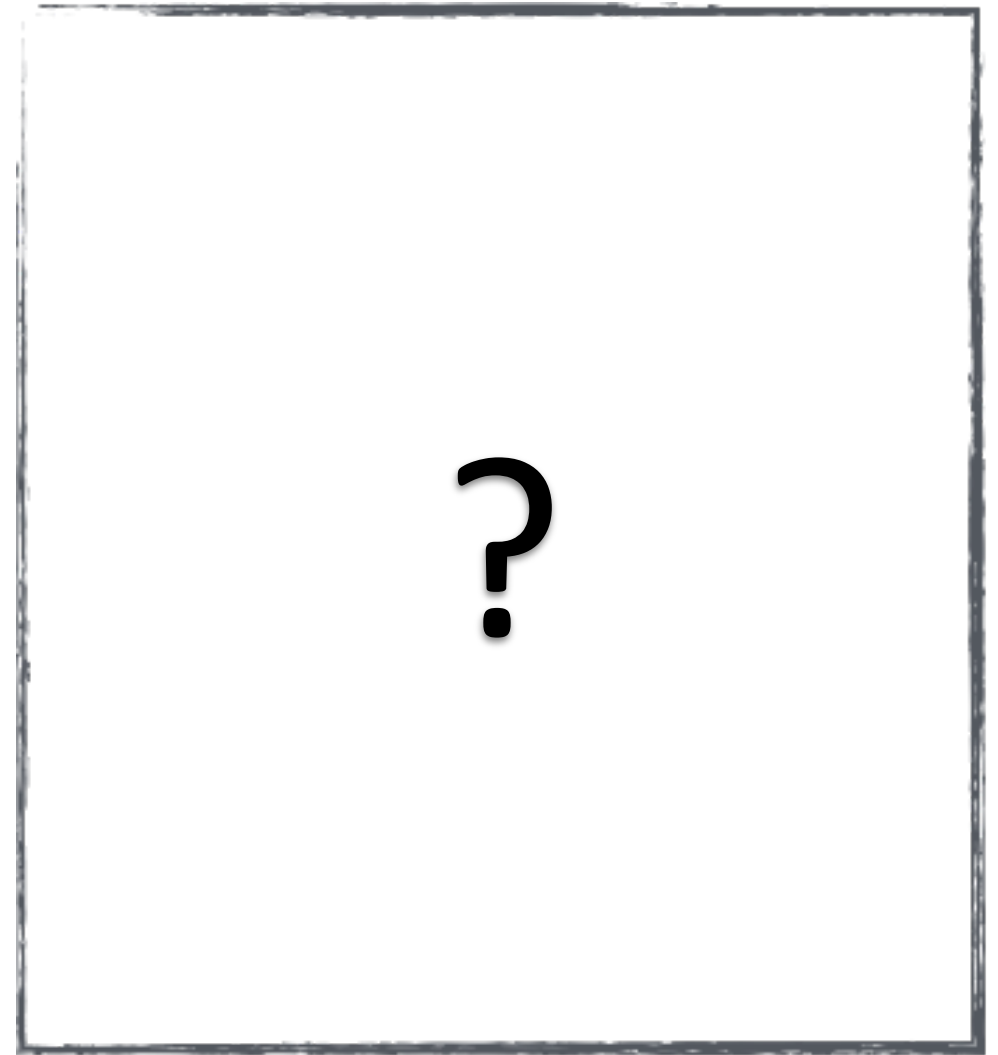
Best limit to date, with argon target, third best limit behind LUX & Xenon100.

Future Detectors



DS-20k

30 tonne (20 tonne fiducial) detector



ARGO

300 tonne (200 tonne fiducial) detector

Further Depletion of Ar

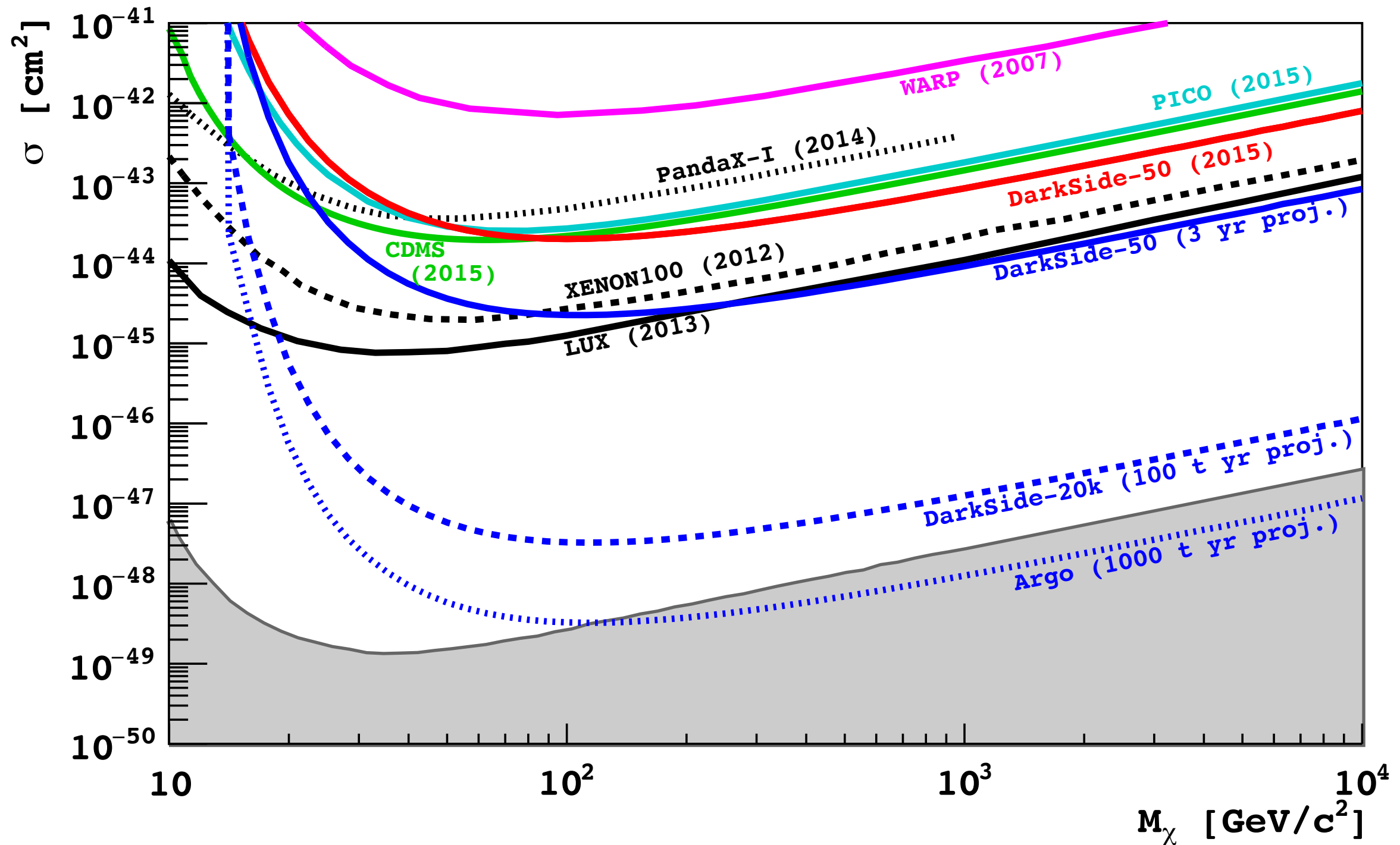
Urania (Underground Argon):

- Expansion of the argon extraction plant in Cortez, CO, to reach capacity of **100 kg/day** of Underground Argon

Aria (UAr Purification):

- Very tall column in the Seruci mine in Sardinia, Italy, for high-volume chemical and isotopic purification of Underground Argon

Expected sensitivity



Summary

- **Background free:**
 - ^{39}Ar BG from **47.1 live days** (1422 kg·day fiducial) of AAr corresponds to that expected in **38.7 years of UAr** DS-50 run ($\gg$ planning physics run time, 3 years).
 - Concentration of ^{39}Ar in **UAr** is **1400** times lower than in **AAr**.
 - DarkSide demonstrates ^{39}Ar **BG** rejection at level of **5.5 tonne·year** with UAr.
- Elastic spin-independent WIMP-nucleon interaction: the non observation of events in the ROI in DarkSide-50 excludes cross-sections larger than $2 \cdot 10^{-44} \text{ cm}^2$ for 100 GeV WIMPs.
 - Best limit obtained with Argon target
 - Third best limit after Xenon detectors
- Future detectors are planned with sensitivities of 10^{-47} and 10^{-48} cm^2 (@100GeV)
 - Letter of Intent was submitted to LNGS April 27 2015.

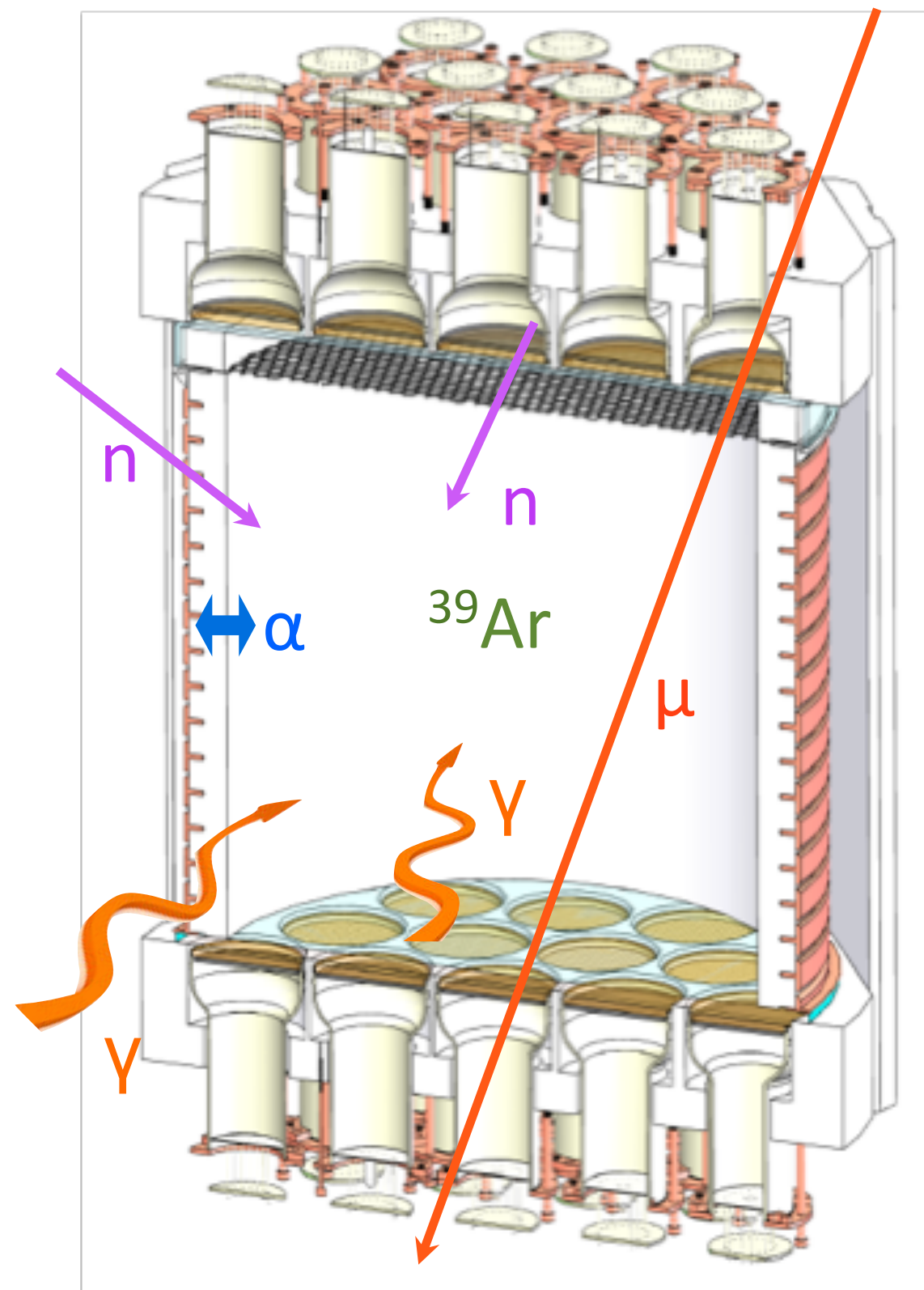
THE END

Backgrounds

ELECTRON RECOILS

^{39}Ar
 $\sim 9 \times 10^4 \text{ evt/kg/day}$

γ
 $\sim 1 \times 10^2 \text{ evt/kg/day}$



NUCLEAR RECOILS

μ
 $\sim 30 \text{ evt/m}^2/\text{day}$

Radiogenic n
 $\sim 6 \times 10^{-4} \text{ evt/kg/day}$

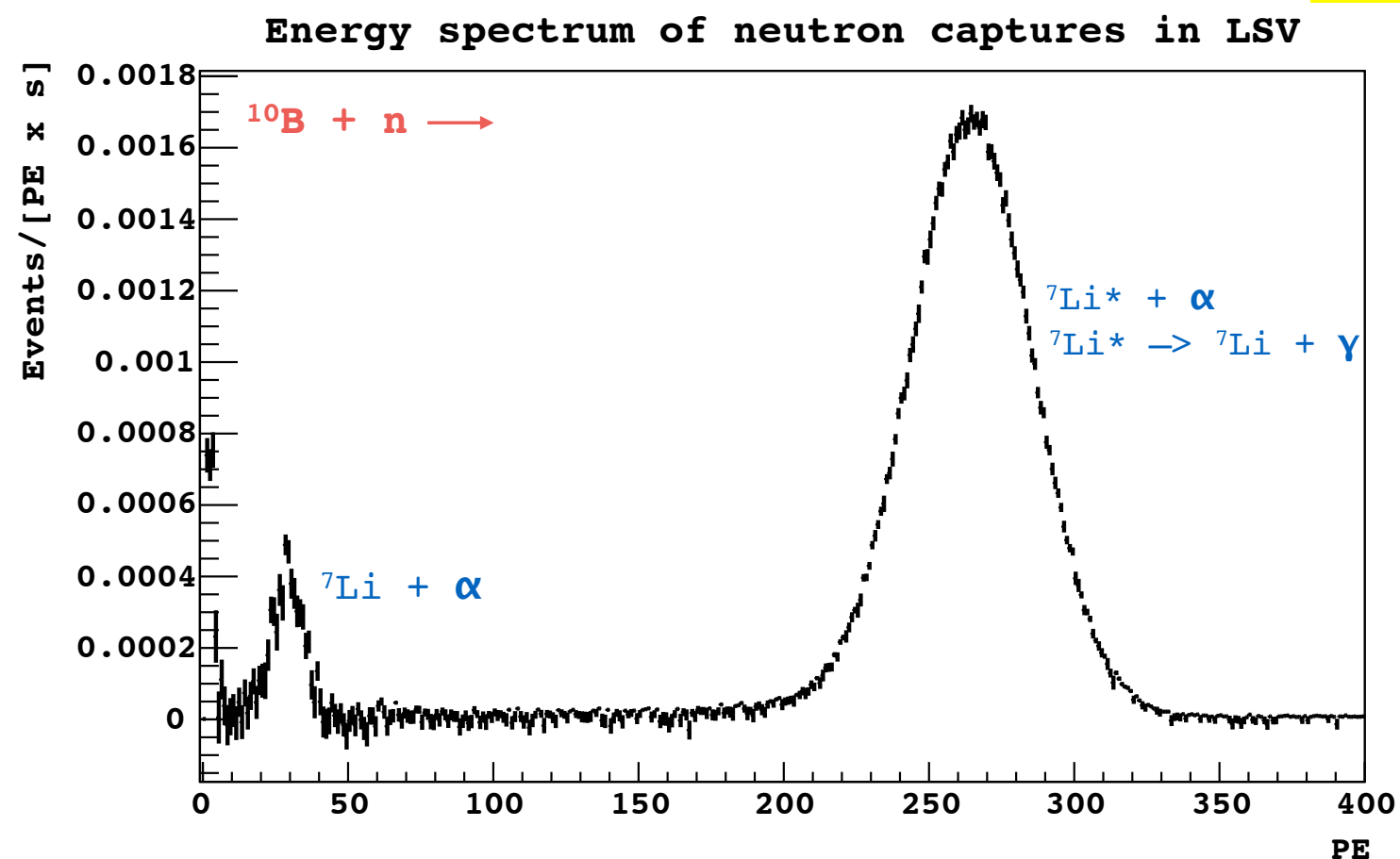
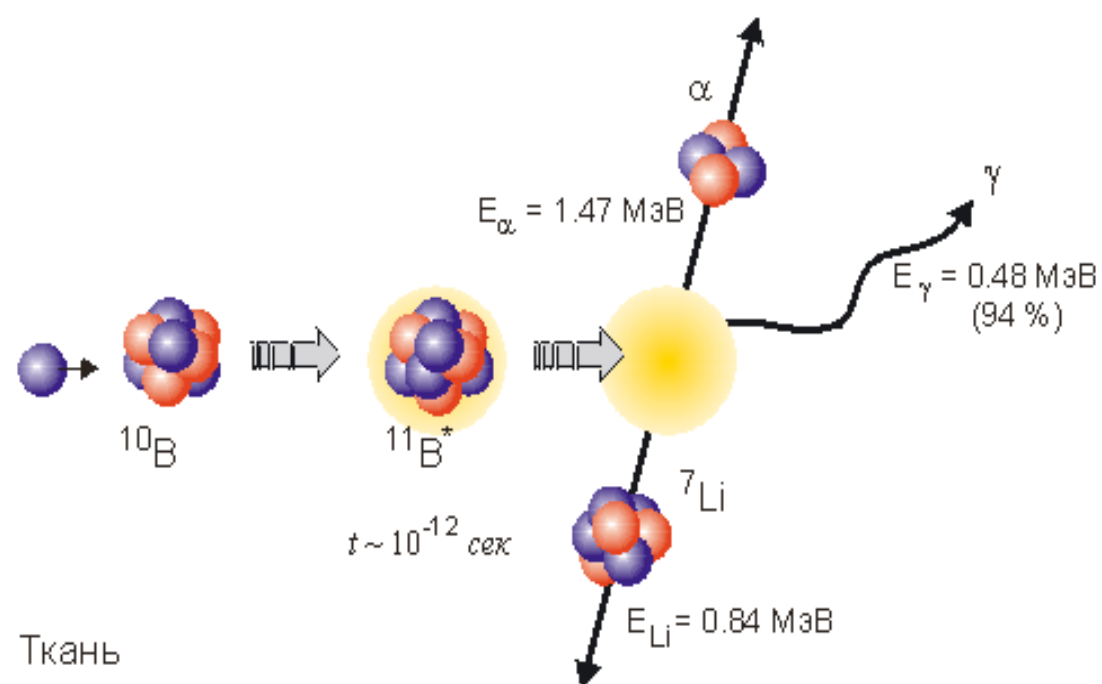
α
 $\sim 10 \text{ evt/m}^2/\text{day}$

100 GeV, 10^{-45} cm^2 WIMP Rate $\sim 10^{-4} \text{ evt/kg/day}$

Borated Liquid Scintillator

- High neutron capture cross section on Boron allows for compact veto size
- Short capture time **2.3 μ s** reduces dead time loss. Cmp. $\sim 260\mu$ s in pure scintillator
- Capture results in 1.47 MeV α particle, quenched to **~ 50 keV**: it must be detected with high efficiency!

	Veto Efficiency (MC)
Radiogenic Neutrons	> 99%
Cosmogenic Neutrons	> 95%



PMTs with cold-amplifiers

- 3" PMTs
- Hamamatsu R11065 series
- The “-20” have **good background levels** but show problems at nominal gain at LAr temperature
- Require **low PMT Gain**
 $\sim 4 \times 10^5$
- Custom cold amplifiers:
Noise $\sim 3 \mu\text{V}$ on 200 MHz



Requirements for DS-20k

Neutron Background:

- **Cosmogenic**: Veto system
- **Radiogenic**: radiopure SiPM & ultra-clean Titanium (TPC cryostat)

β/γ background:

- ^{39}Ar : Underground Argon (Urania Project) & Depleted Argon (Aria Project)
- γ : SiPM & ultra-clean Titanium

Sensitivity Comparison

Experiment	σ [cm ²] @1 TeV/c ²	σ [cm ²] @10 TeV/c ²
LUX [10k kg×day Xe]	1.1×10^{-44}	1.2×10^{-43}
XENON [7.6k kg×day Xe]	1.9×10^{-44}	1.9×10^{-43}
DS-50 [1.4k kg×day Ar]	2.3×10^{-43}	2.1×10^{-42}
ArDM [1.5 tonnexyr Ar]	8×10^{-45}	7×10^{-44}
DEAP-3600 [3.0 tonnexyr Ar]	5×10^{-46}	5×10^{-45}
XENON-1ton [2.7 tonnexyr Xe]	3×10^{-46}	3×10^{-45}
LZ [15 tonnexyr Xe]	5×10^{-47}	5×10^{-46}
DS-20k [100 tonnexyr]	9×10^{-48}	9×10^{-47}
1 Neutrino Event [400 tonnexyr Ar or 300 tonnexyr Xe]	2×10^{-48}	2×10^{-47}
ARGO [1,000 tonnexyr]	9×10^{-49}	9×10^{-48}

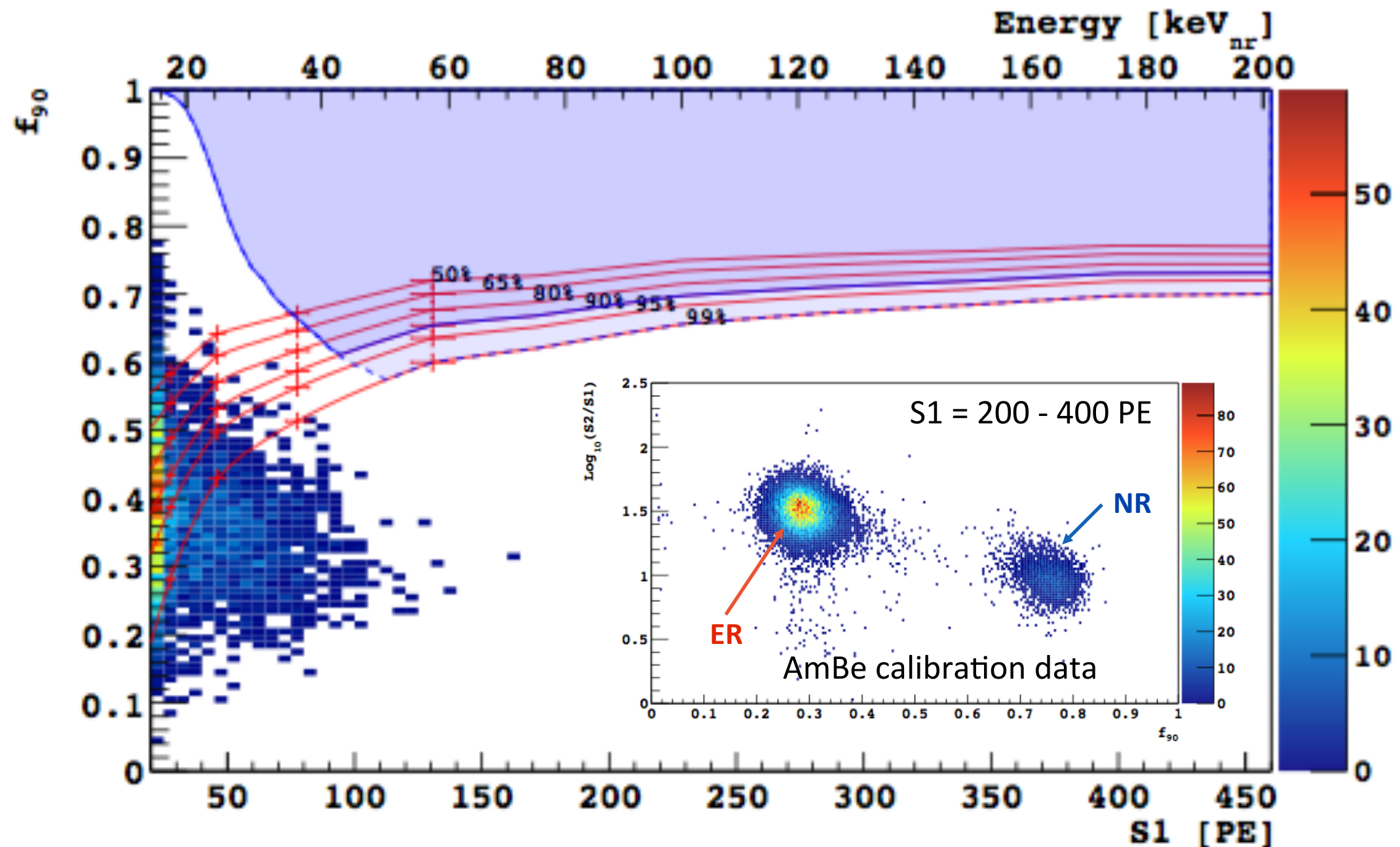
DarkSide-20k and Argo Lol Signatories

D. Franco, A Tonazzo - [APC Paris](#)
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UAr First Results with S2/S1 cut

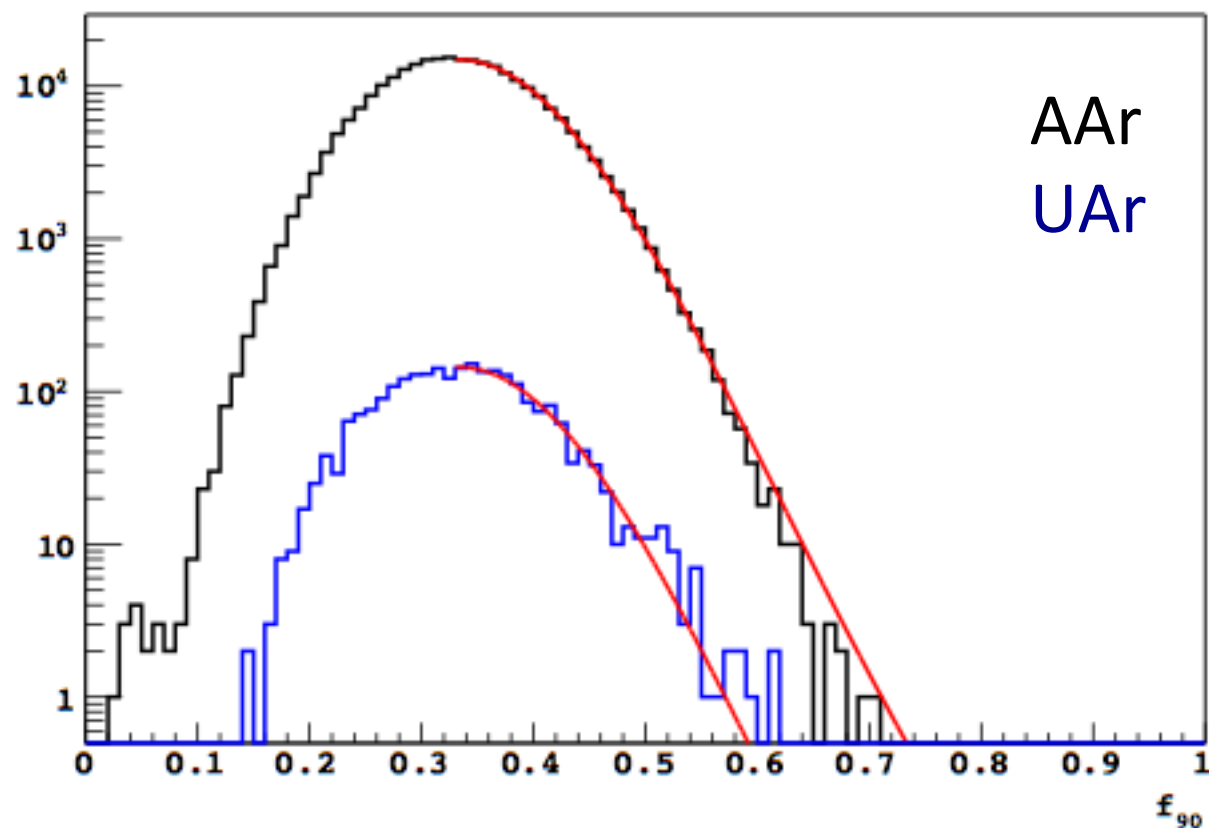
We have another discrimination power to suppress ERs (S2/S1 cut w/ 50% acceptance of NRs).



Sample F90 models

- Use analytic model for F90 distributions
- Fit to high statistics AAr data
- Scale to UAr data
- Derive 0.01 ER leakage events / S1 bin

S1: [60.00, 70.00] PE



S1: [250.00, 260.00] PE

