



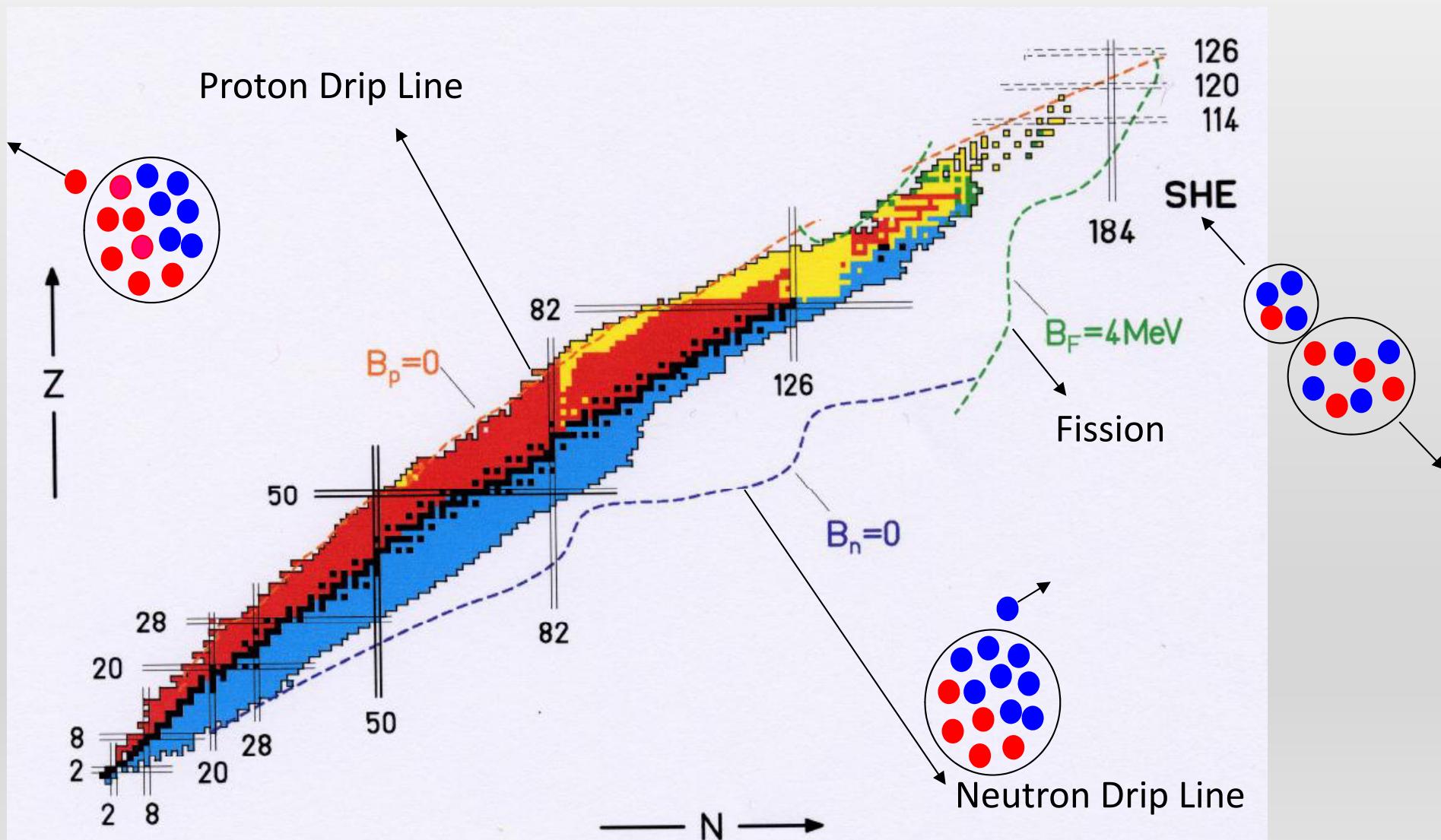
Beta decay studies using the gamma Total Absorption Technique

Berta Rubio

IFIC-Valencia (Spain)

Many of the slides by A. Algora (planned speaker), also J.L. Taín and V. Guadilla

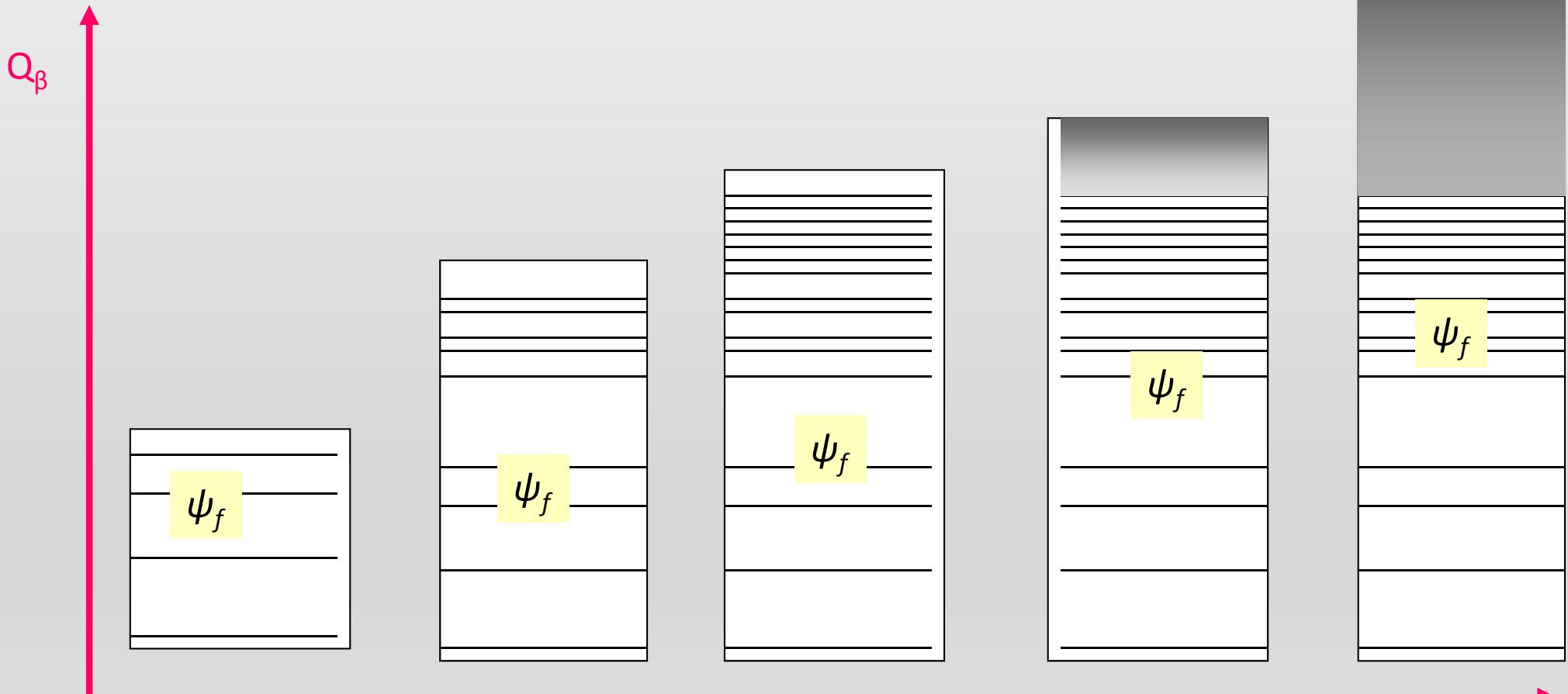
Most unstable nuclei decay
by either β^+ or β^- decay



Most of the beta decay are dominated by Gamow Teller transitions

Definition of Gamow teller strength:

$$B(GT) = \left| \frac{1}{\sqrt{2}} \langle \psi_f | \sum_{\mu} \sum_k \sigma_k^{\mu} \tau_k^{\pm} | \psi_i \rangle \right|^2$$



The beta Gamow Teller strength or transition probability Theory versus experiment

Theoretically

$$B(GT) = \left| \frac{1}{\sqrt{2}} \langle \psi_f | \sum_{\mu} \sum_k \sigma_k^{\mu} \tau_k^{\pm} | \psi_i \rangle \right|^2$$

Experimentally

$$B(GT) = k \frac{I_{\beta}(E)}{f(Q_{\beta} - E, Z) T_{1/2}} = k \frac{1}{ft}$$

From the experiment



$$I_{\beta}(E)$$

Beta feeding to states
in the daughter nucleus



$$Q_{\beta}$$

Q-value

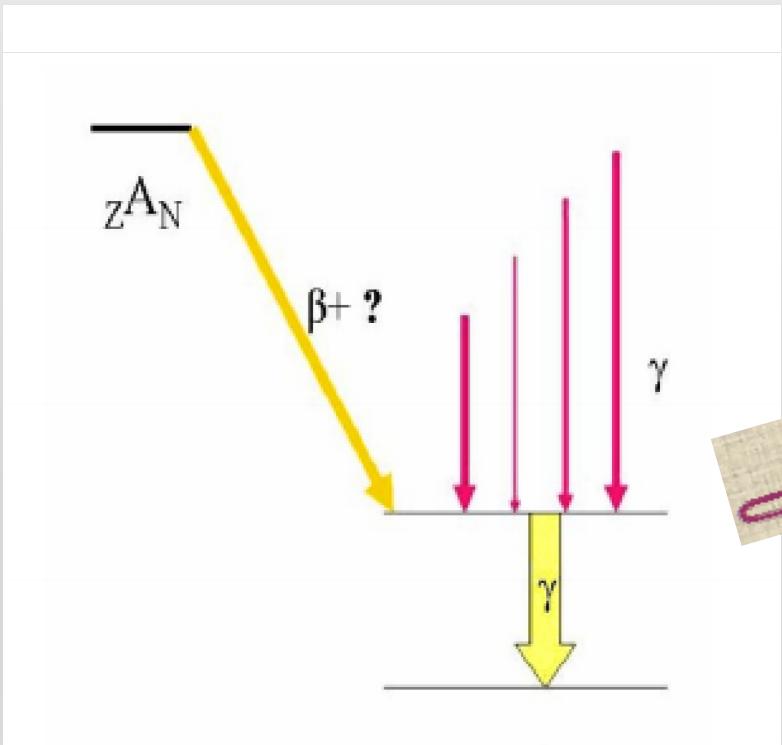


$$T_{1/2}$$

Parent half life

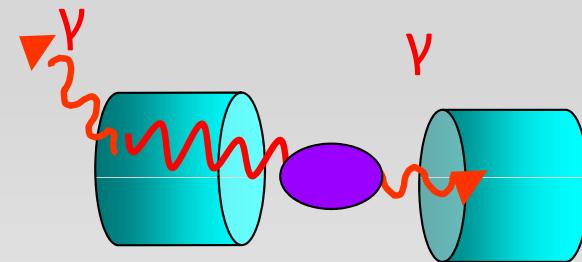
This talk

How to measure beta feeding intensity To each of the levels in the daughter nucleus

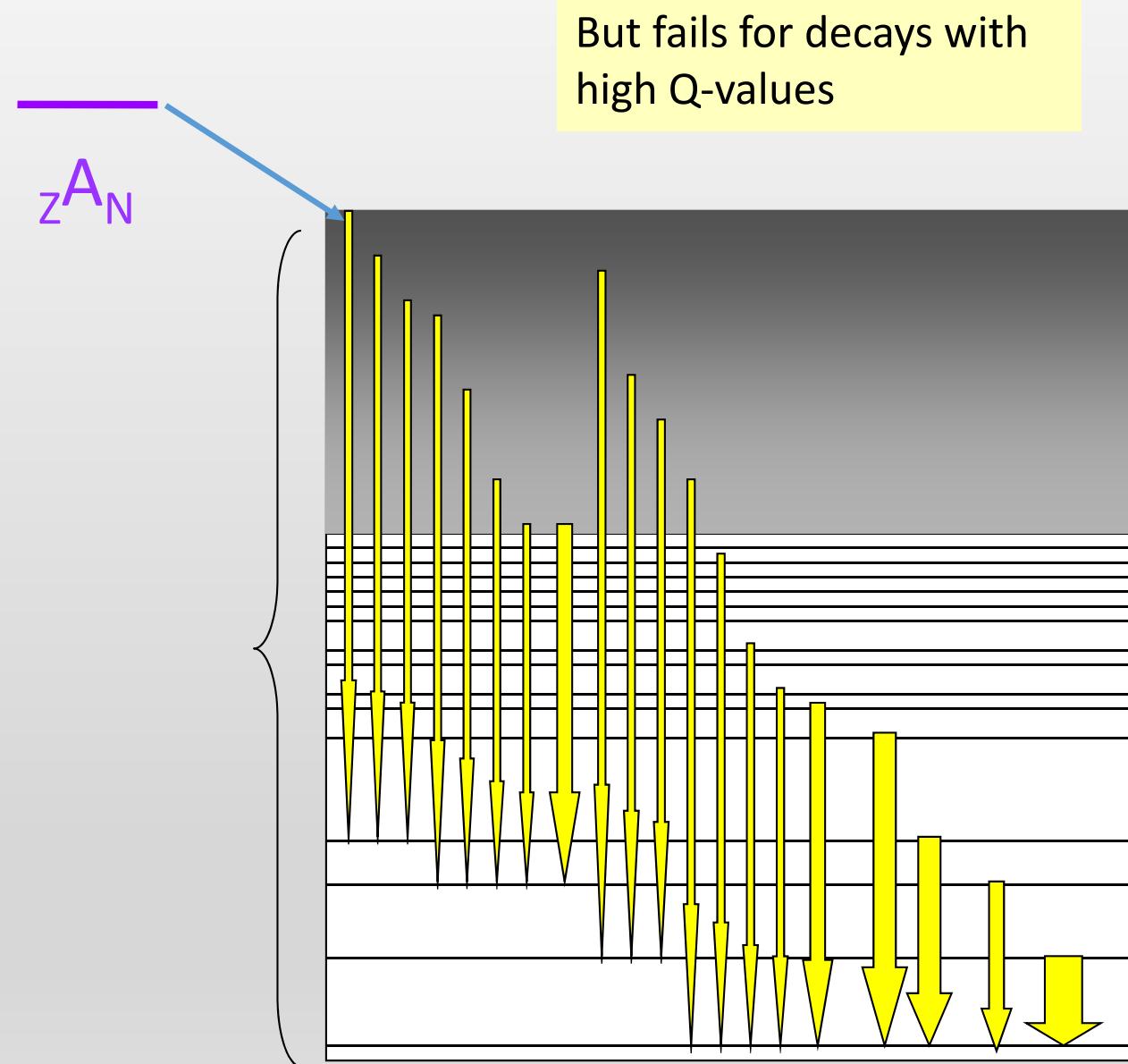
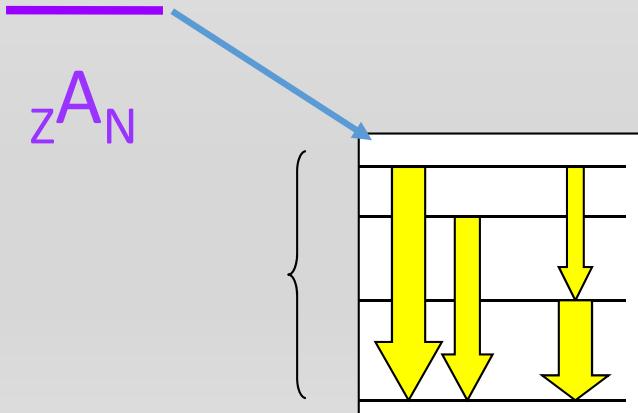


$$B(GT) = k \frac{I_\beta(E)}{f(Q_\beta - E, Z) T_{1/2}}$$

$I_\beta(E)$



This works very well
close to the stability





Volume 71B, number 2

PHYSICS LETTERS

21 November 1977

**THE ESSENTIAL DECAY OF PANDEMONIUM:
A DEMONSTRATION OF ERRORS IN COMPLEX BETA-DECAY SCHEMES**

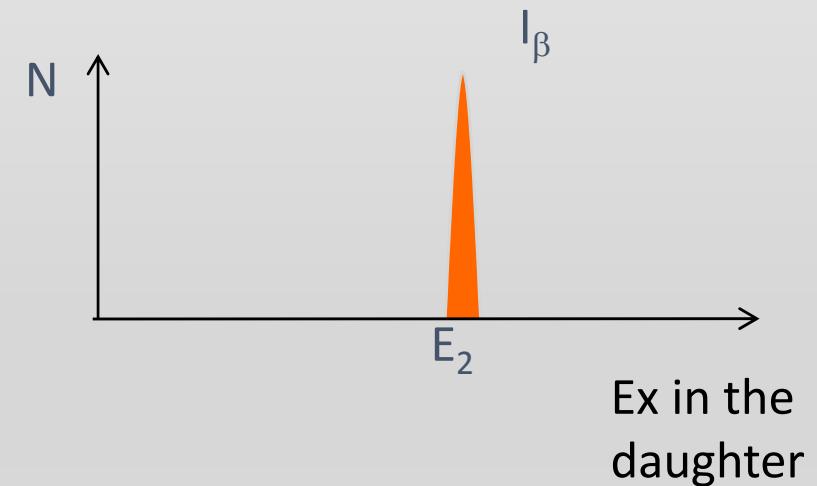
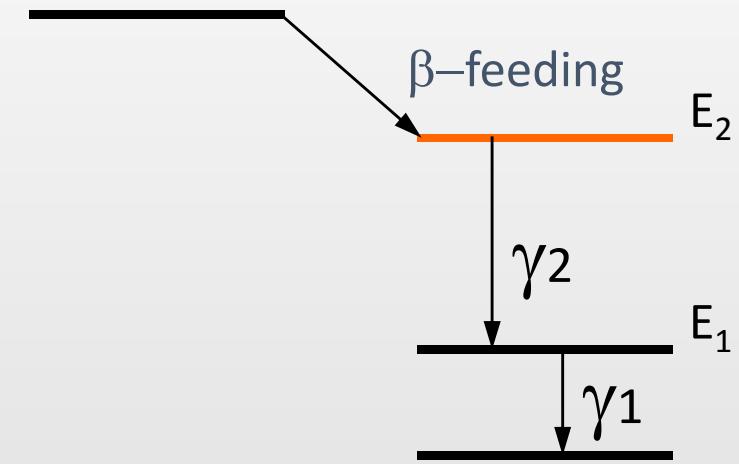
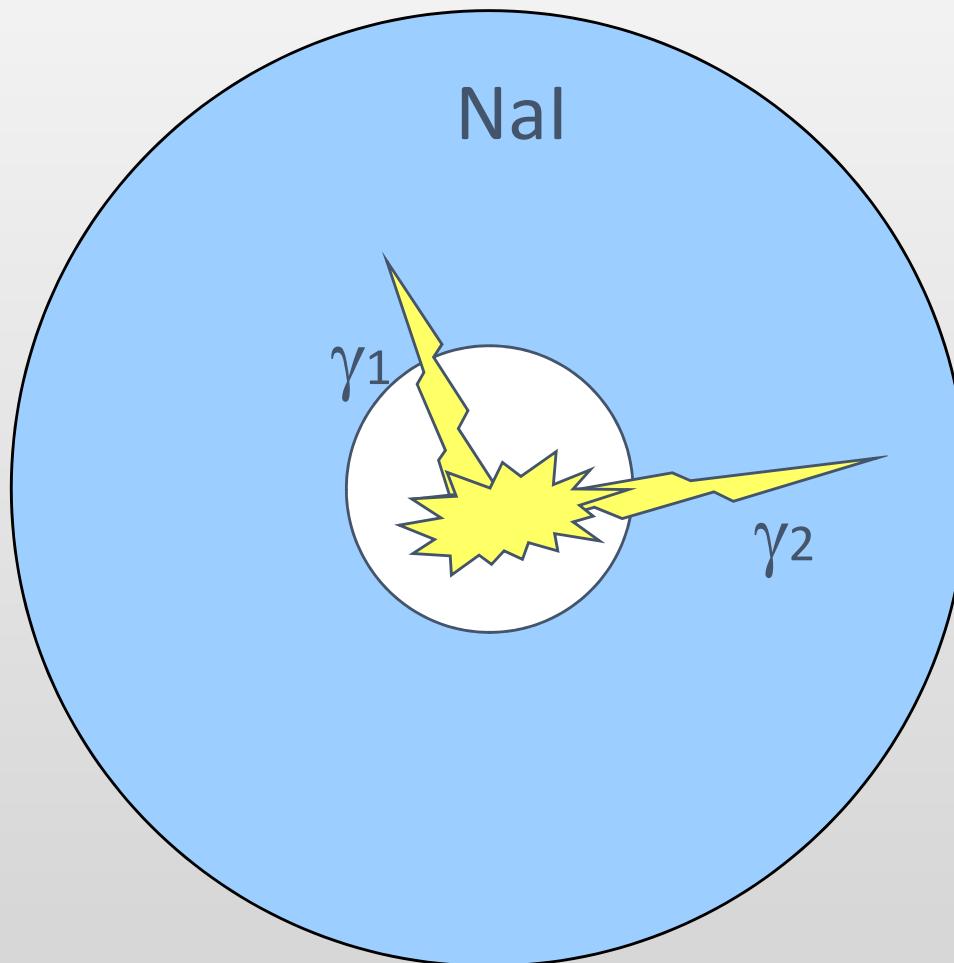
J.C. HARDY *, L.C. CARRAZ, B. JONSON ‡ and P.G. HANSEN §

CERN, Geneva, Switzerland



J. Milton, *Paradise Lost, Book I* (1667)

Total Absorption spectroscopy



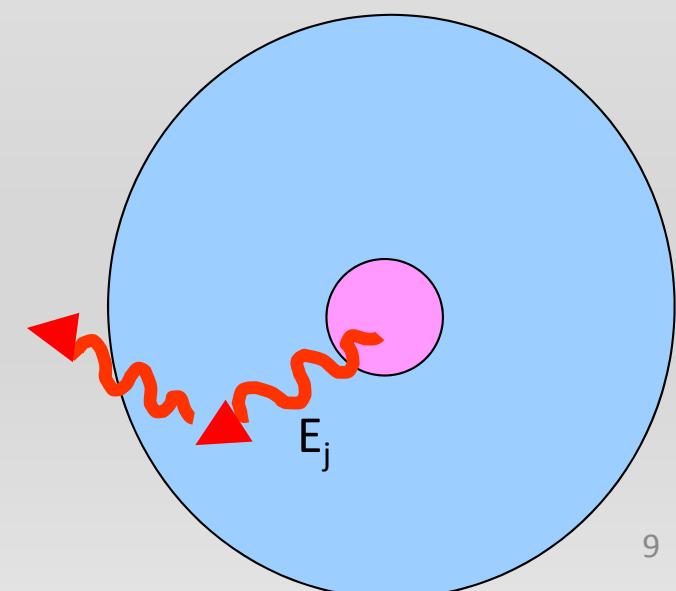
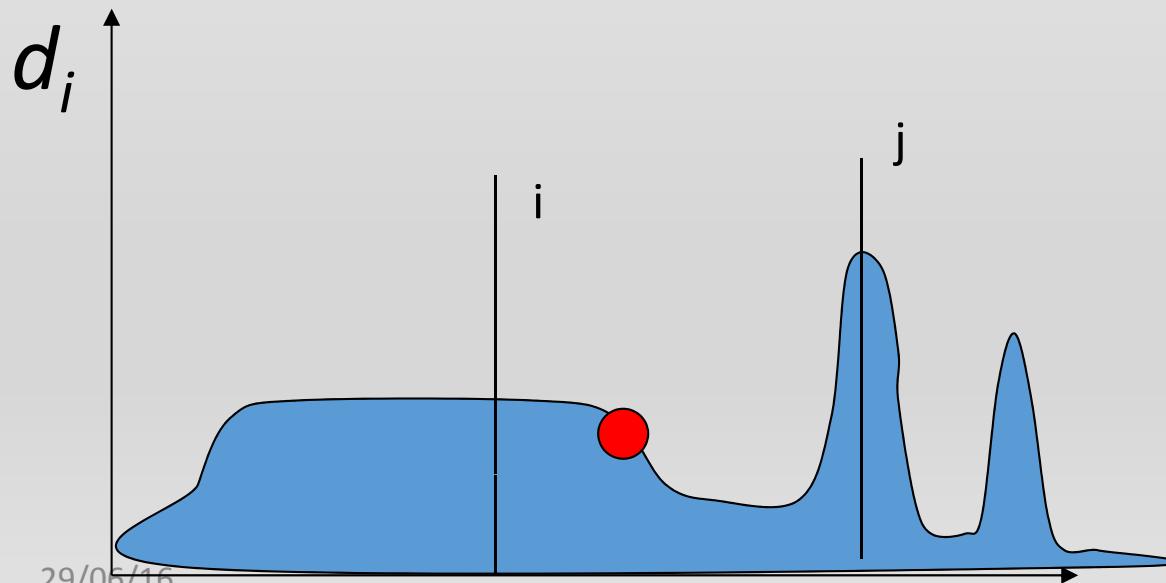
Ideal case

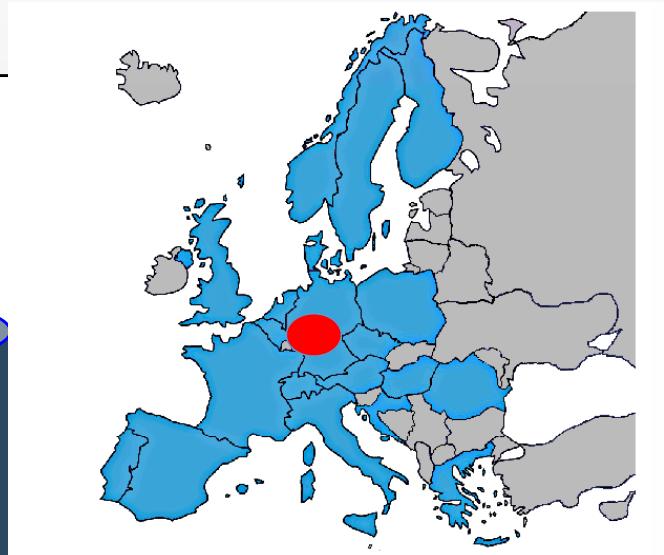
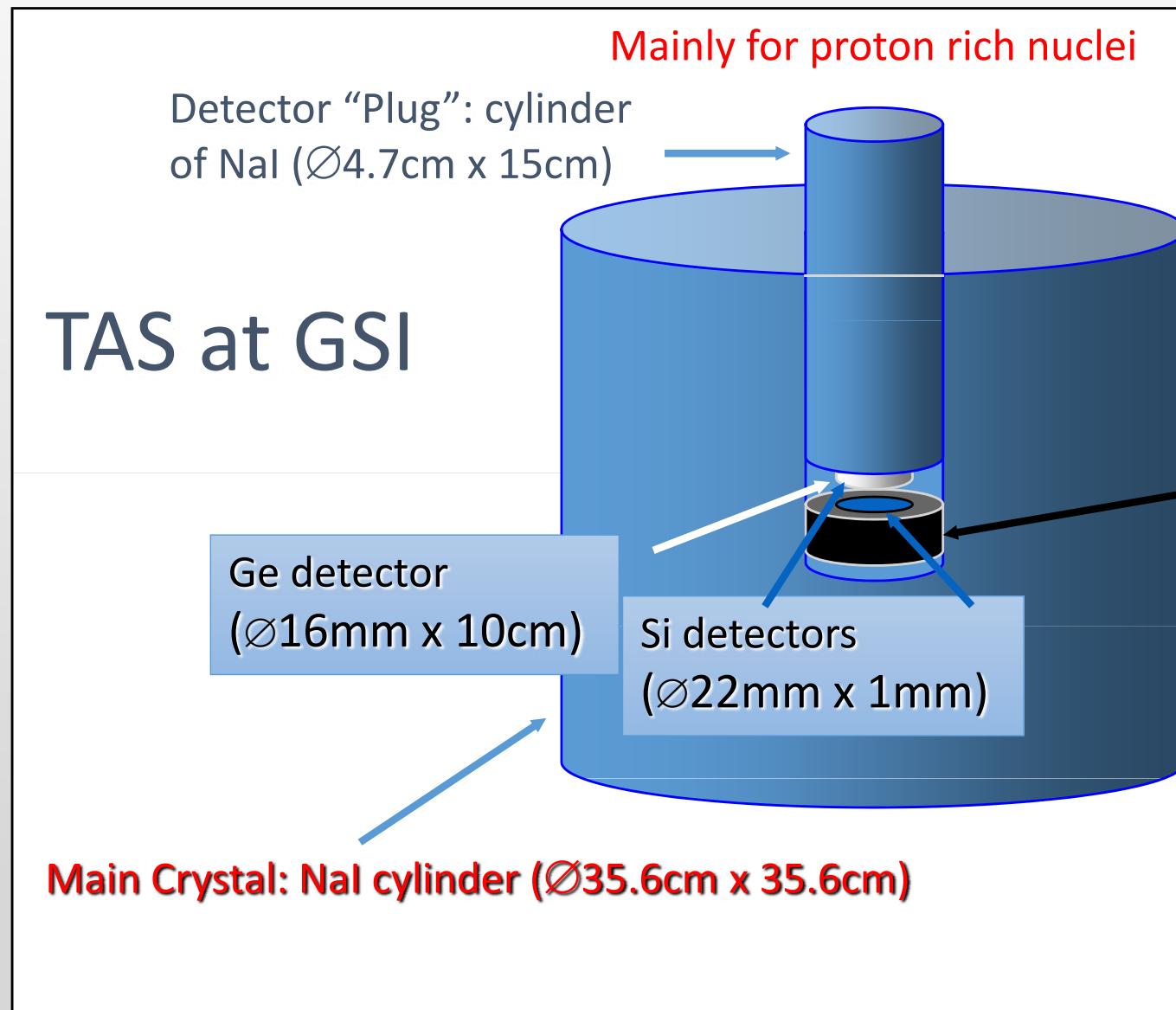
The main problem is not the intrinsic efficiency
but the material placed inside the detector or the holes
Needed to put the activity inside

One has to solve the following equation:

$$d_i = \sum_{j=1}^{j_{\max}} R_{ij} f_j$$

J.L.Tain, D. Cano–Ott, Nucl. Instr. and Meth. in Phys. Res. A **571**, 719 (2007)
J.L.Tain, D. Cano–Ott, Nucl. Instr. and Meth. in Phys. Res. A **571**, 728 (2007)





Positron absorber:
polyethylene
($\varnothing 51\text{mm} \times 21\text{mm}$)



Spokesperson B. Rubio

Observations of the Gamow-Teller resonance in the rare-earth nuclei above ^{146}Gd populated in β decay

P.h.D E. Nacher

E. Nácher,^{1,*} B. Rubio,^{1,†} A. Algora,^{1,‡} D. Cano-Ott,^{1,§} J. L. Taín,¹ A. Gadea,¹ J. Agramunt,¹ M. Gierlik,² M. Karny,² Z. Janas,² E. Roeckl,³ A. Blazhev,^{3,||} R. Collatz,³ J. Döring,^{3,¶} M. Hellström,³ Z. Hu,³ R. Kirchner,^{3,#} I. Mukha,³ C. Plettner,³ M. Shibata,³ K. Rykaczewski,⁴ L. Batist,⁵ F. Moroz,⁵ V. Wittmann,^{5,**} and J. J. Valiente-Dobón^{6,††}

¹IFIC, CSIC-Universidad Valencia, Apartado Postal E-22085, 46071 Valencia, Spain

²Faculty of Physics, University of Warsaw, PL-02-093 Warsaw, Poland

³Gesellschaft für Schwerionenforschung, D-64291 Darmstadt, Germany

⁴Physics Division, Oak Ridge National Laboratory, Oak Ridge, Tennessee 37831-6371, USA

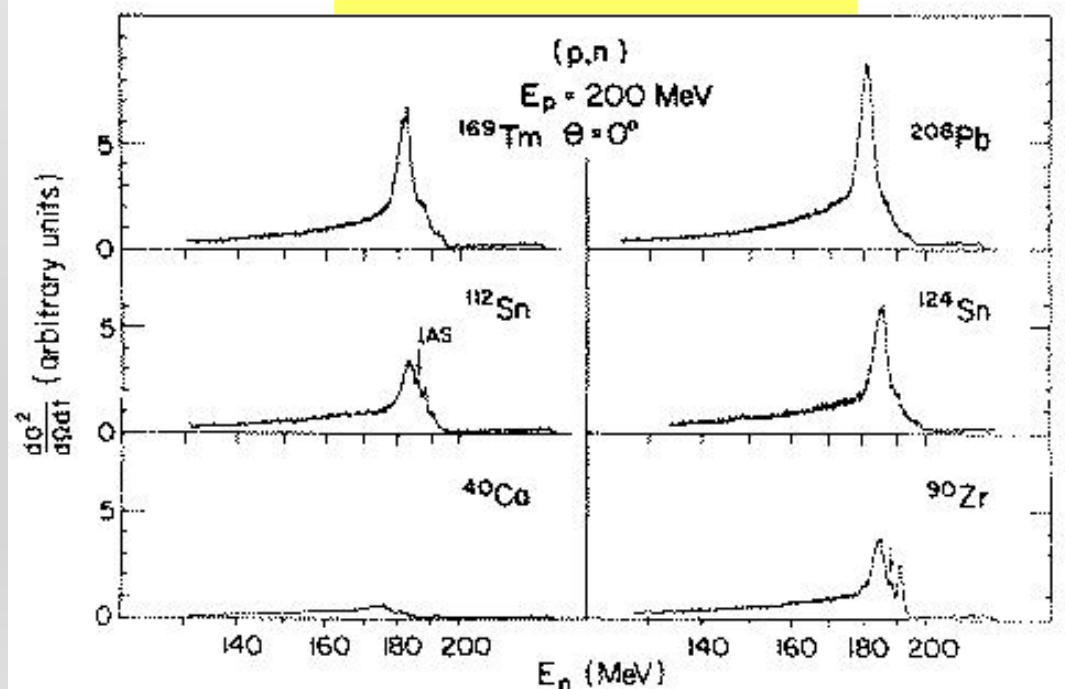
⁵St. Petersburg Nuclear Physics Institute, 188-350 Gatchina, Russia

⁶Department of Physics, University of Surrey, Guildford GU2 5XH, United Kingdom

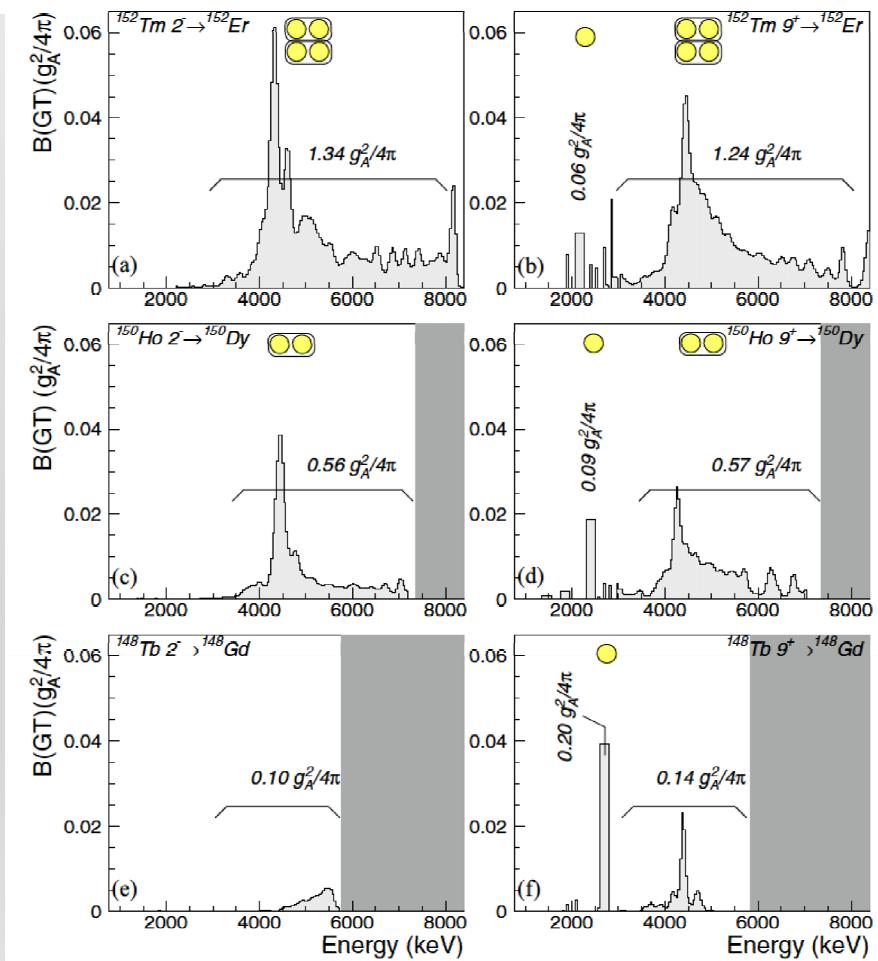
(Received 26 October 2015; published 13 January 2016)

Beta-decay!!!

(p,n) reaction



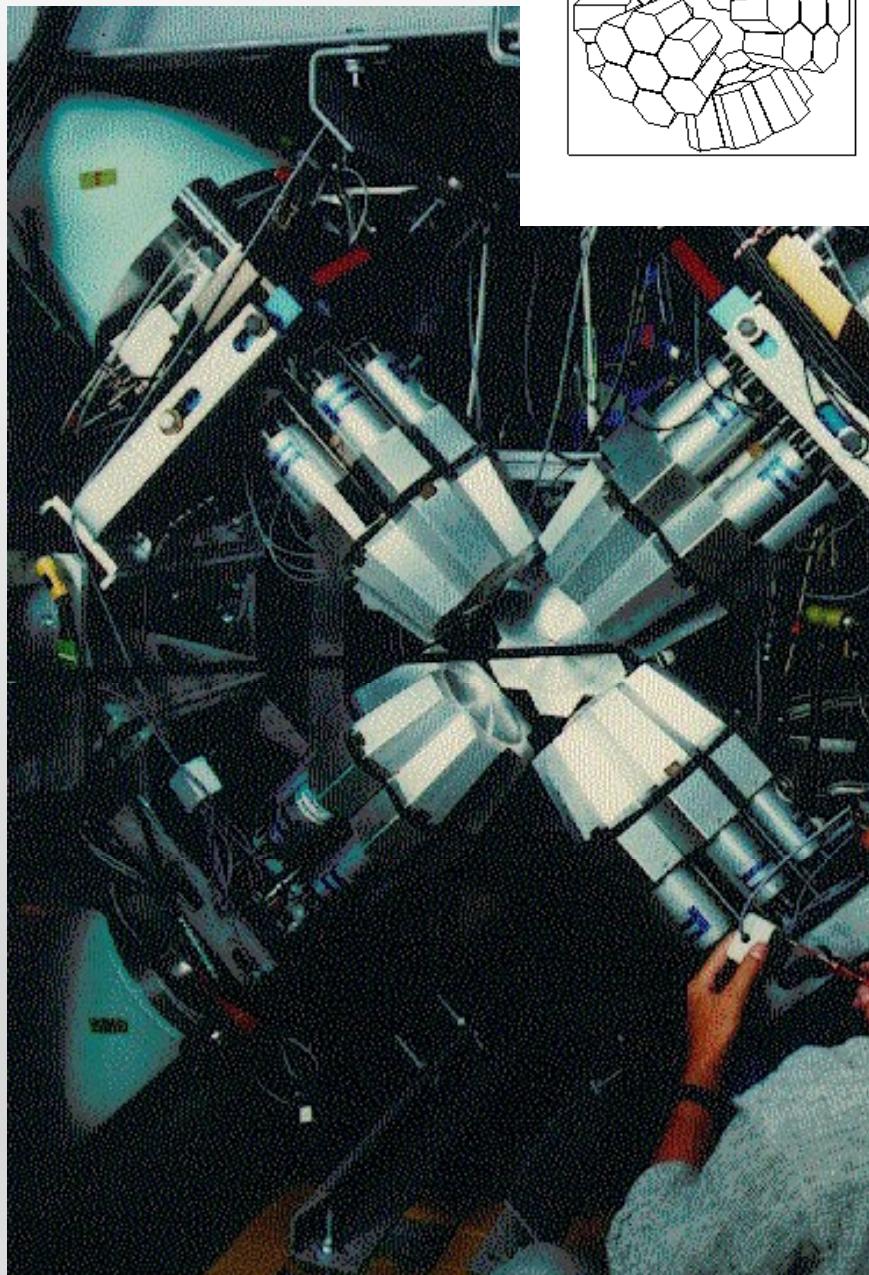
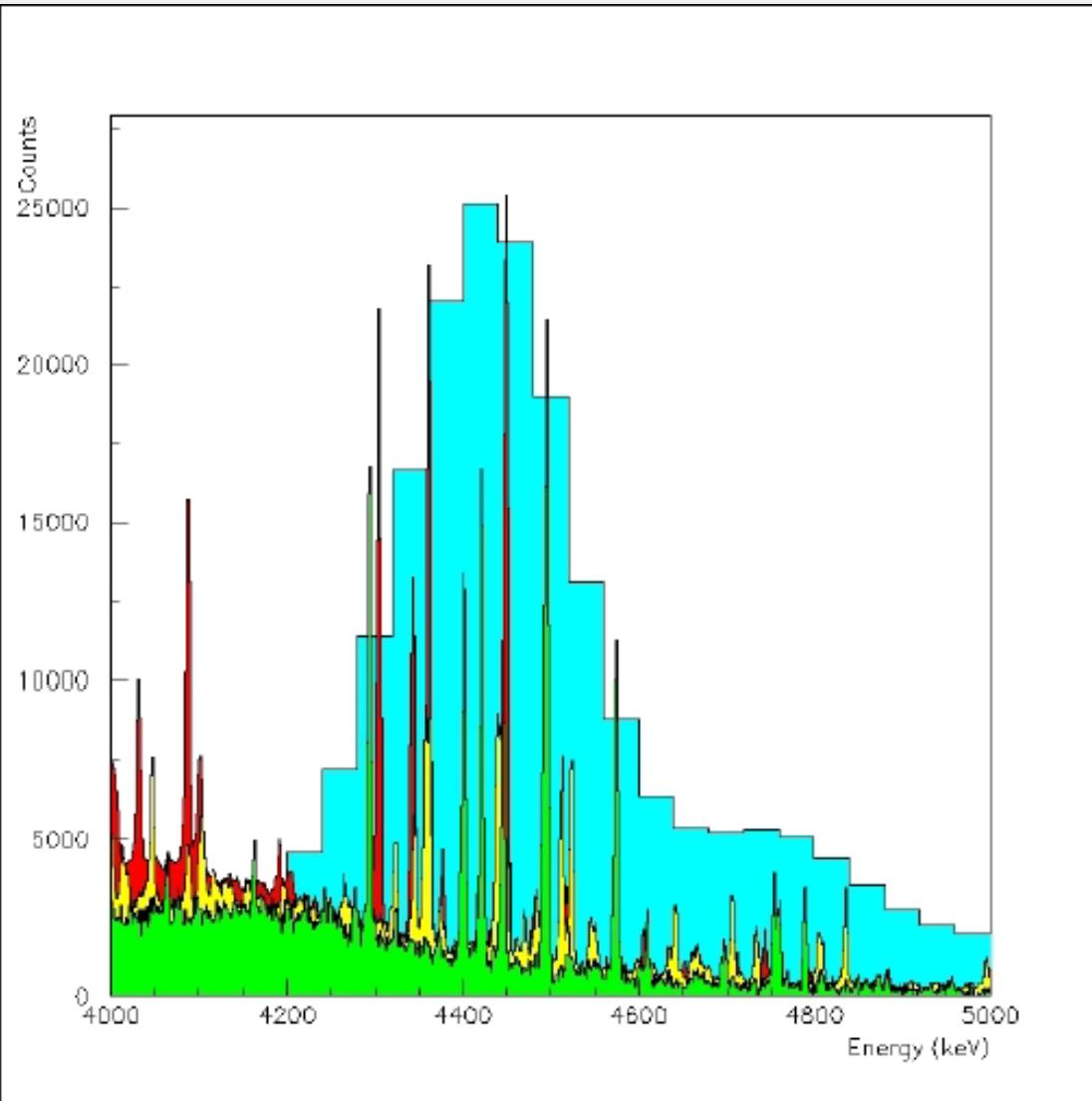
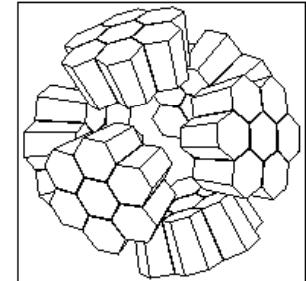
Gaarde NPA369(1981)258



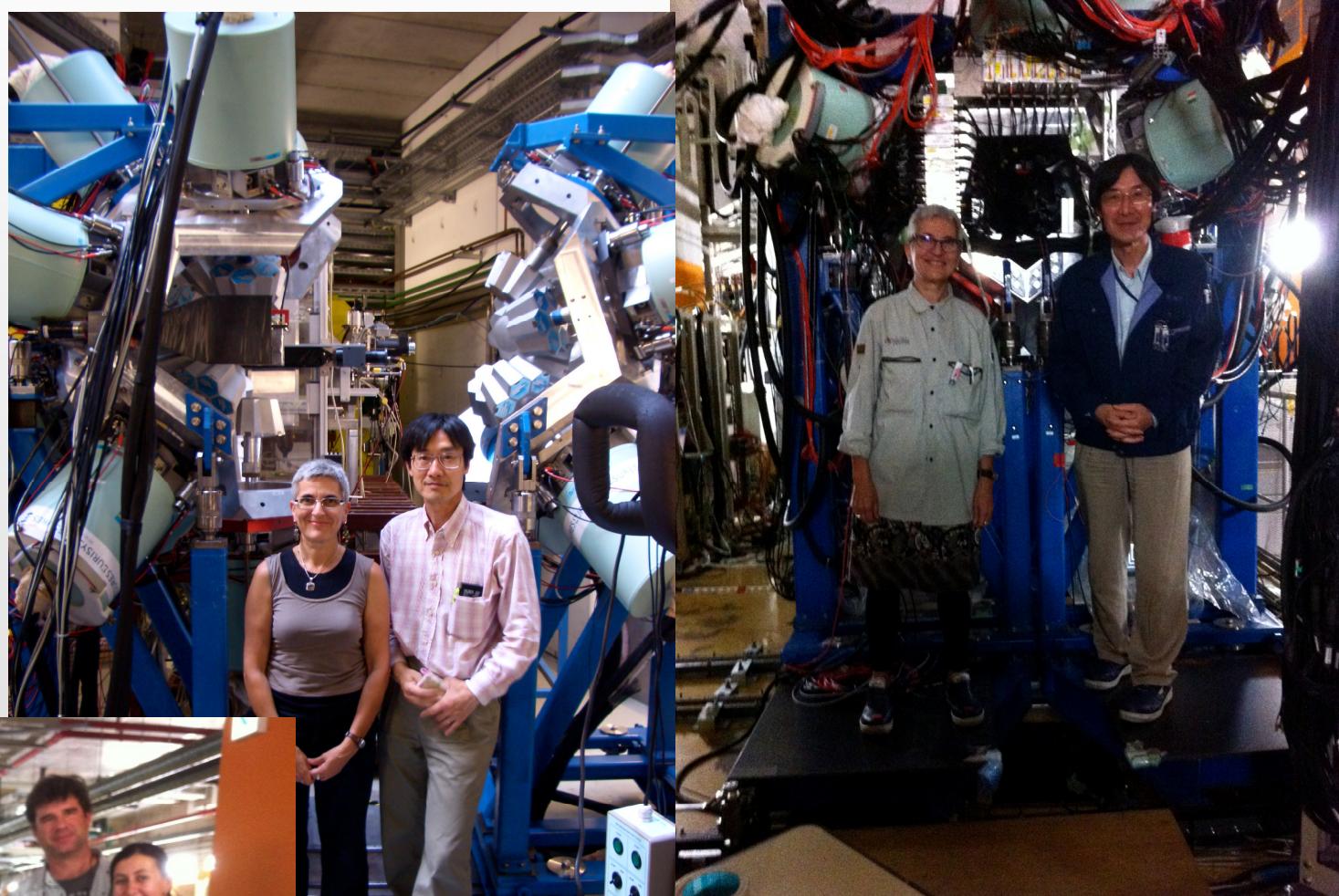
Open Parenthesis

COMPARISON WITH THE HIGH RESOLUTION METHOD

Six EUROBALL CLUSTER Ge detectors
in close geometry



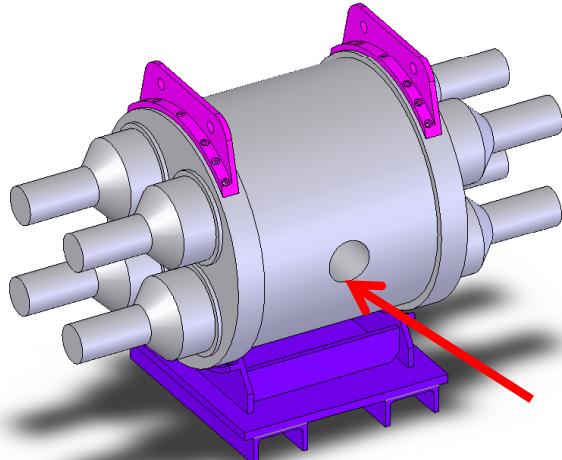
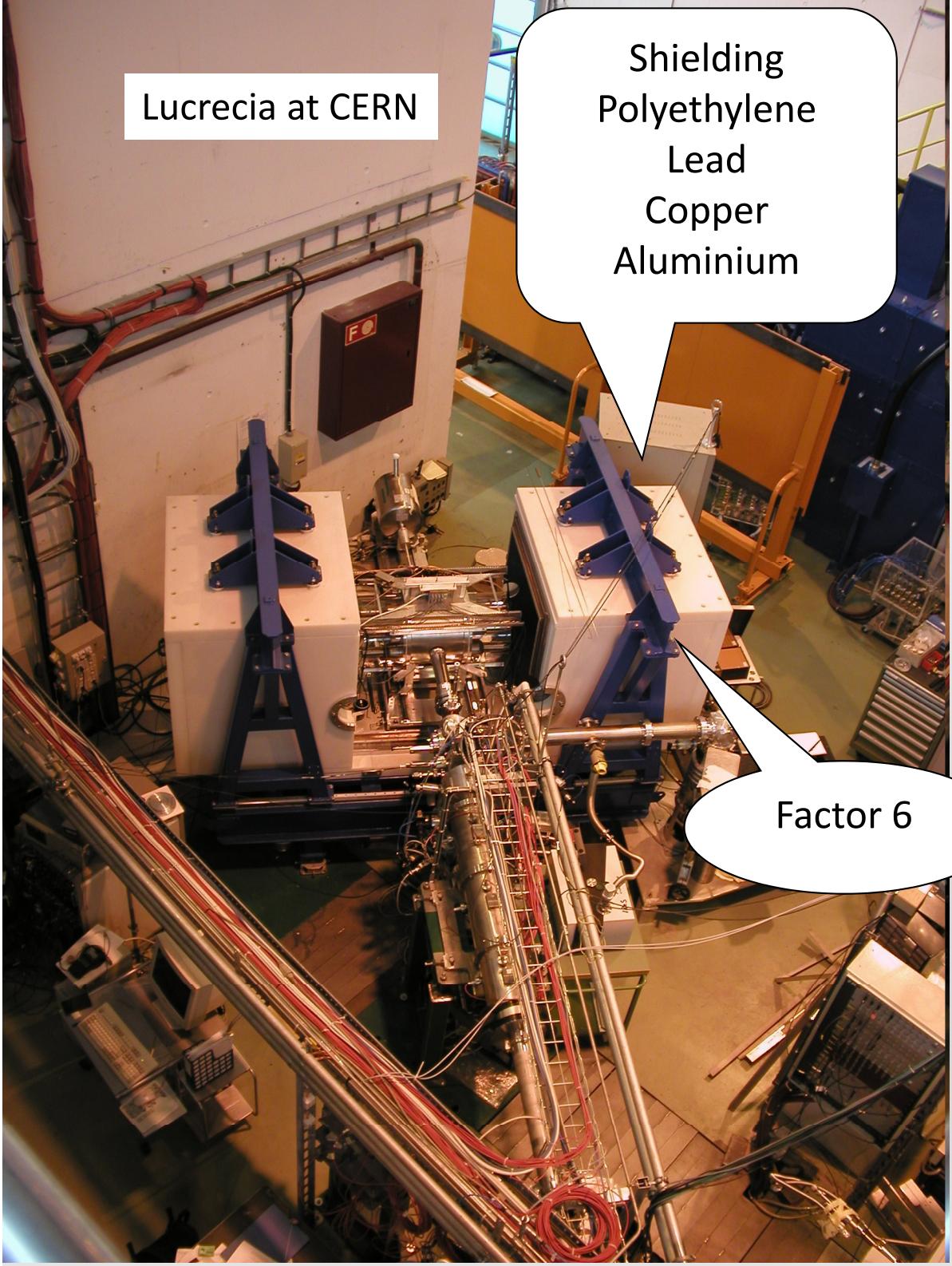
**Tribute to the Euroball
CLUSTER Detectors and to
Our chairman
RISING (GSI) and EURICA (RIKEN)
Campaigns for DECAY Studies**



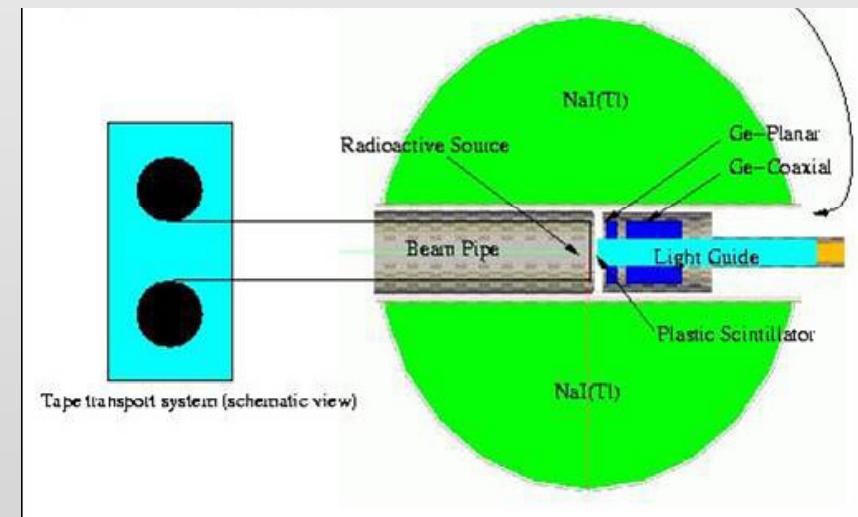
Close Parenthesis

Lucrecia at CERN

Shielding
Polyethylene
Lead
Copper
Aluminium

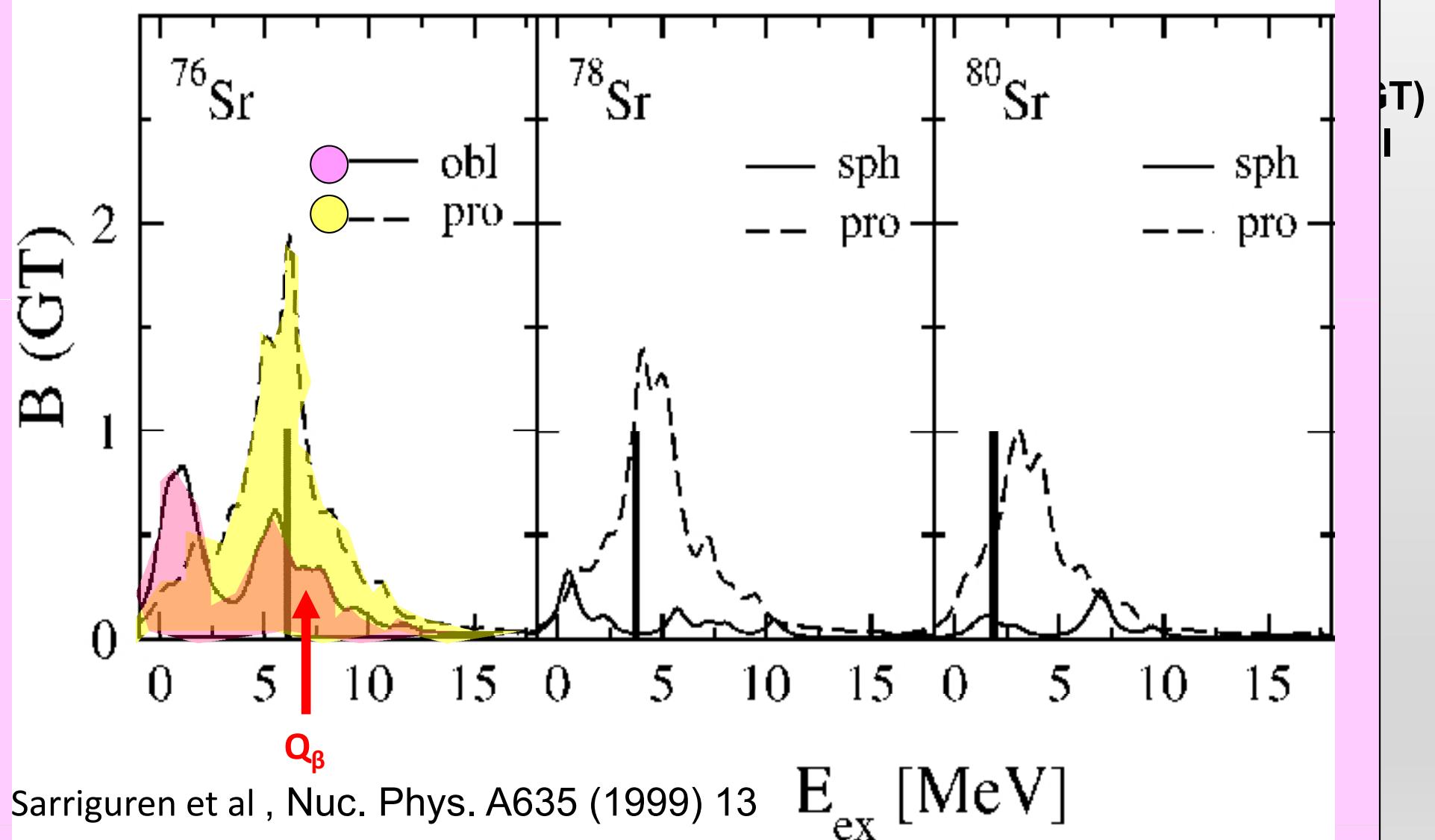


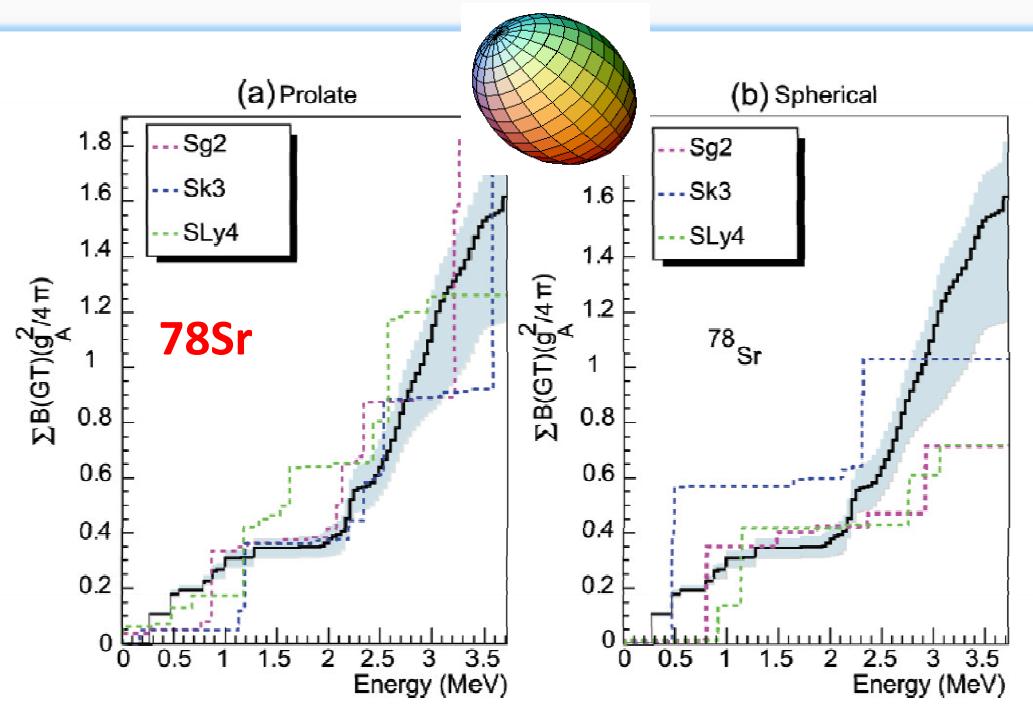
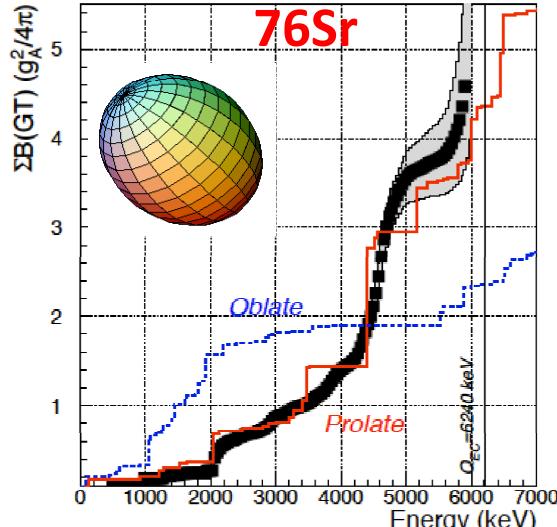
Rad. beam



The Lucrecia TAS at Isolde
is a NaI single-crystal $\Phi 38\text{cm} \times 38\text{cm}$
(+ ancillary detectors)Ge for X-rays
and plastic for β particles

Deducing the deformation of the ground state of ^{76}Sr from its B(GT) distribution



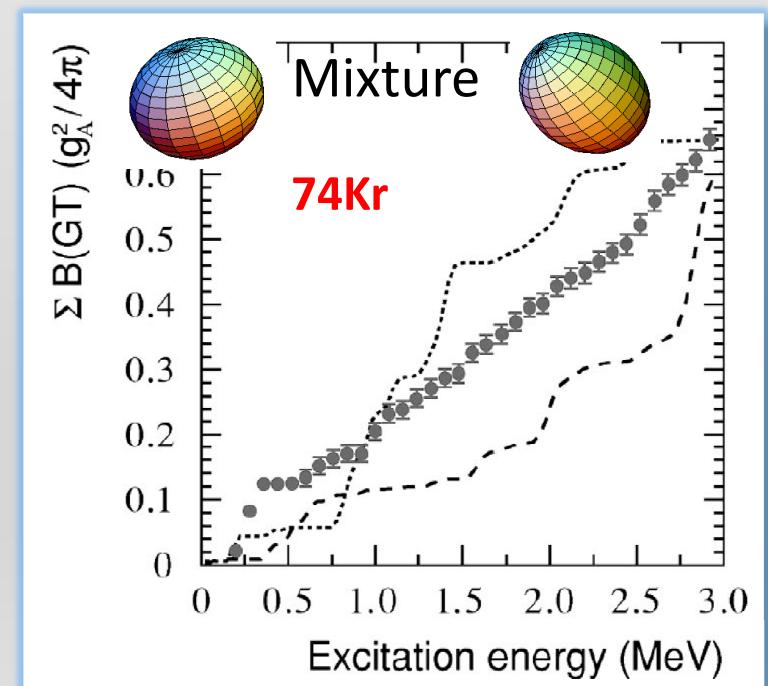
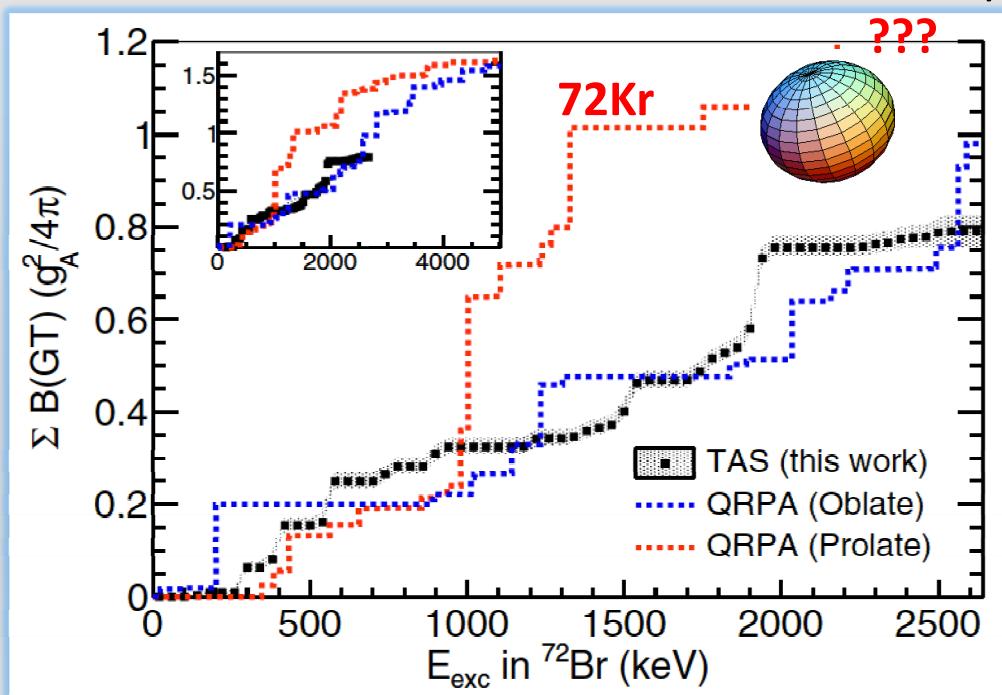


Ph.D's
E.Poirier
E. Nacher
A. Pérez-Cerdán
B. J. Briz

Calculations by
Sarriguren

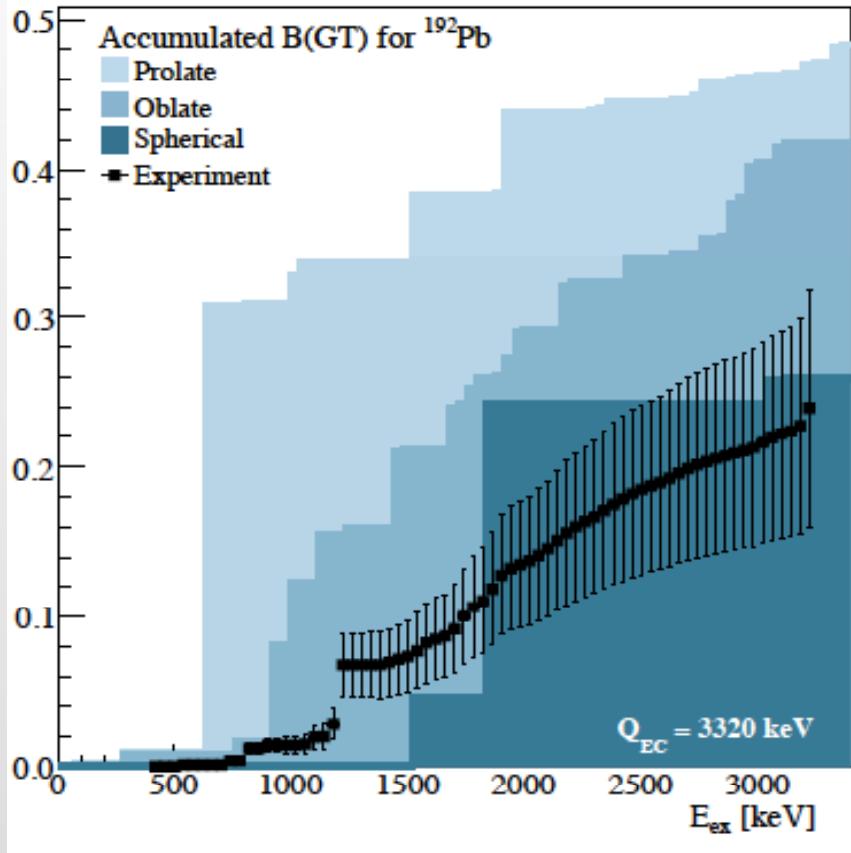
E. Nacher *et al.*, PRL.92, 232501 (2004)

A. Pérez-Cerdán *et al.* PRC88 (2013) 014324

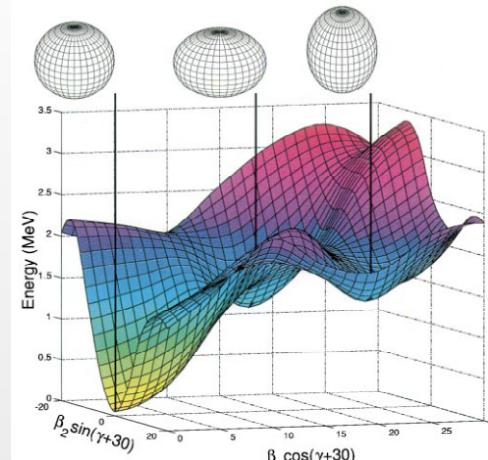


J. Briz *et al*, PRC 92(2016) 054326

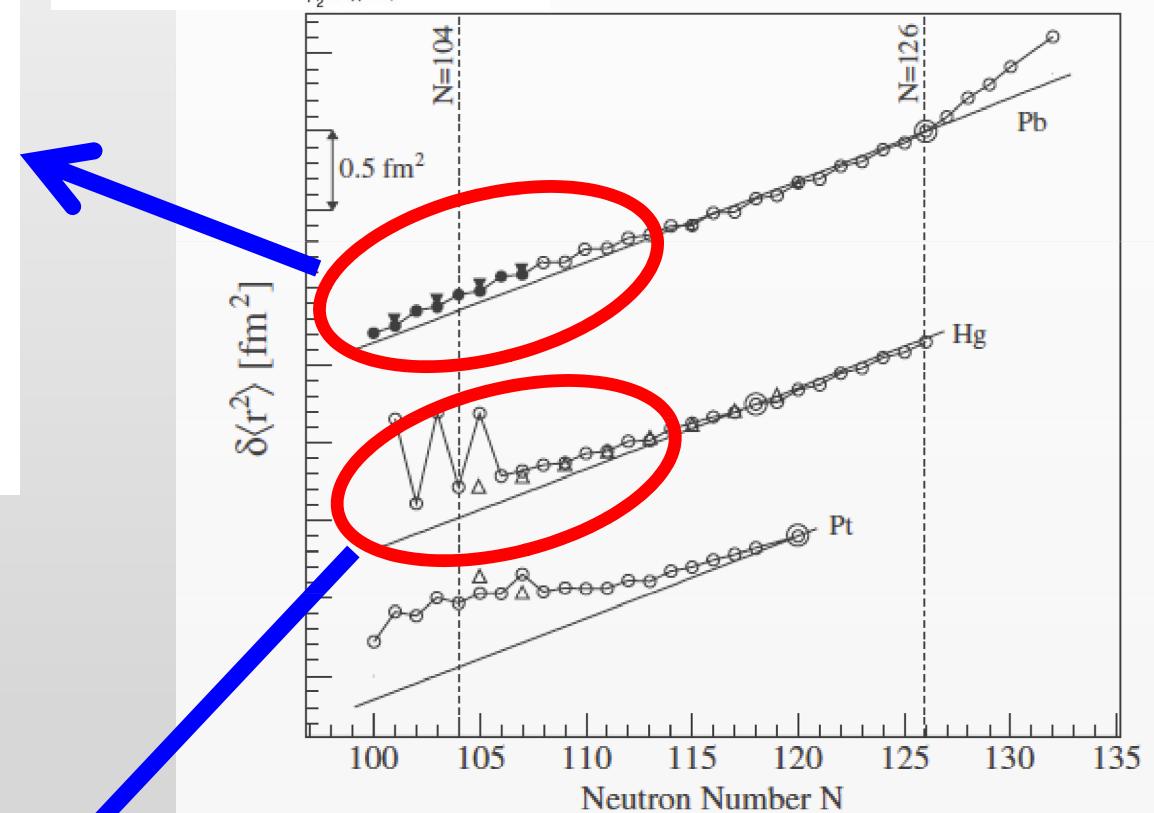
E. Poirier *et al.*, PRC 69, 034307 (2004)



E. Estévez et al. PRC 92, 044321
(2015)



A. Andreyev
Nature 405 (2000) 430



Also T. Cocolios et al. PRL 106,
052503
H. De Witte et al.
PRL 98, 0112502

Spokespersons

A. Algora, E. Nácher, L. M. Fraile

- X-ray Bursts: luminosity curves highly dependent on the EC/ β^+ ratios
- EC component neglected so far in network calculations (naked atoms at this T...)
- Calculations on the cEC process prove that this is a mistake → need for a measurement

Experiment performed at Isolde
May 2016

EUROPEAN ORGANIZATION FOR NUCLEAR RESEARCH

Proposal to the ISOLDE and Neutron Time-of-Flight Committee

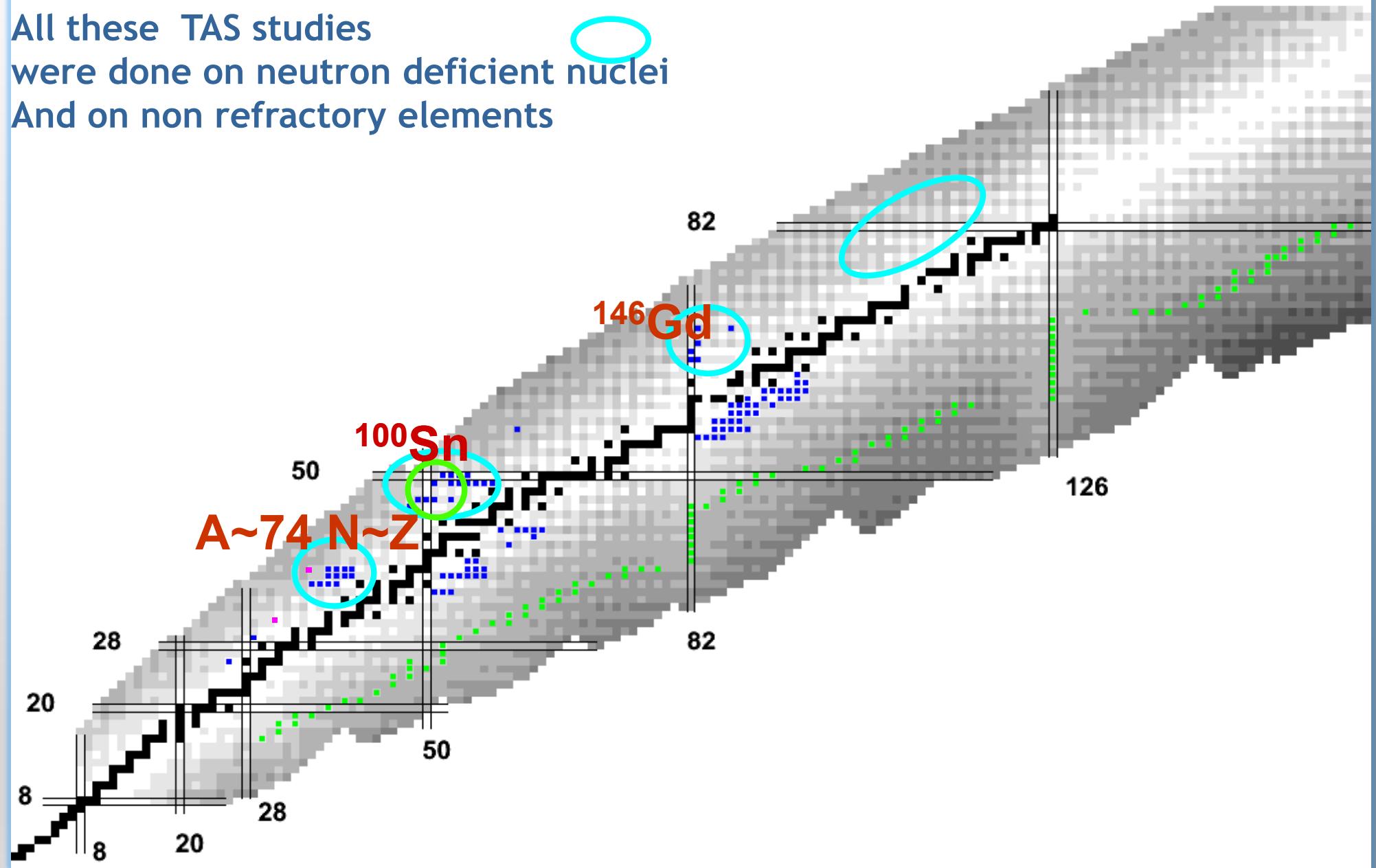
Beta decay of the $N=Z$, rp-process waiting points: ^{64}Ge , ^{68}Se and the $N=Z+2$: ^{66}Ge , ^{70}Se for accurate stellar weak-decay rates

[May 29th - 2013]

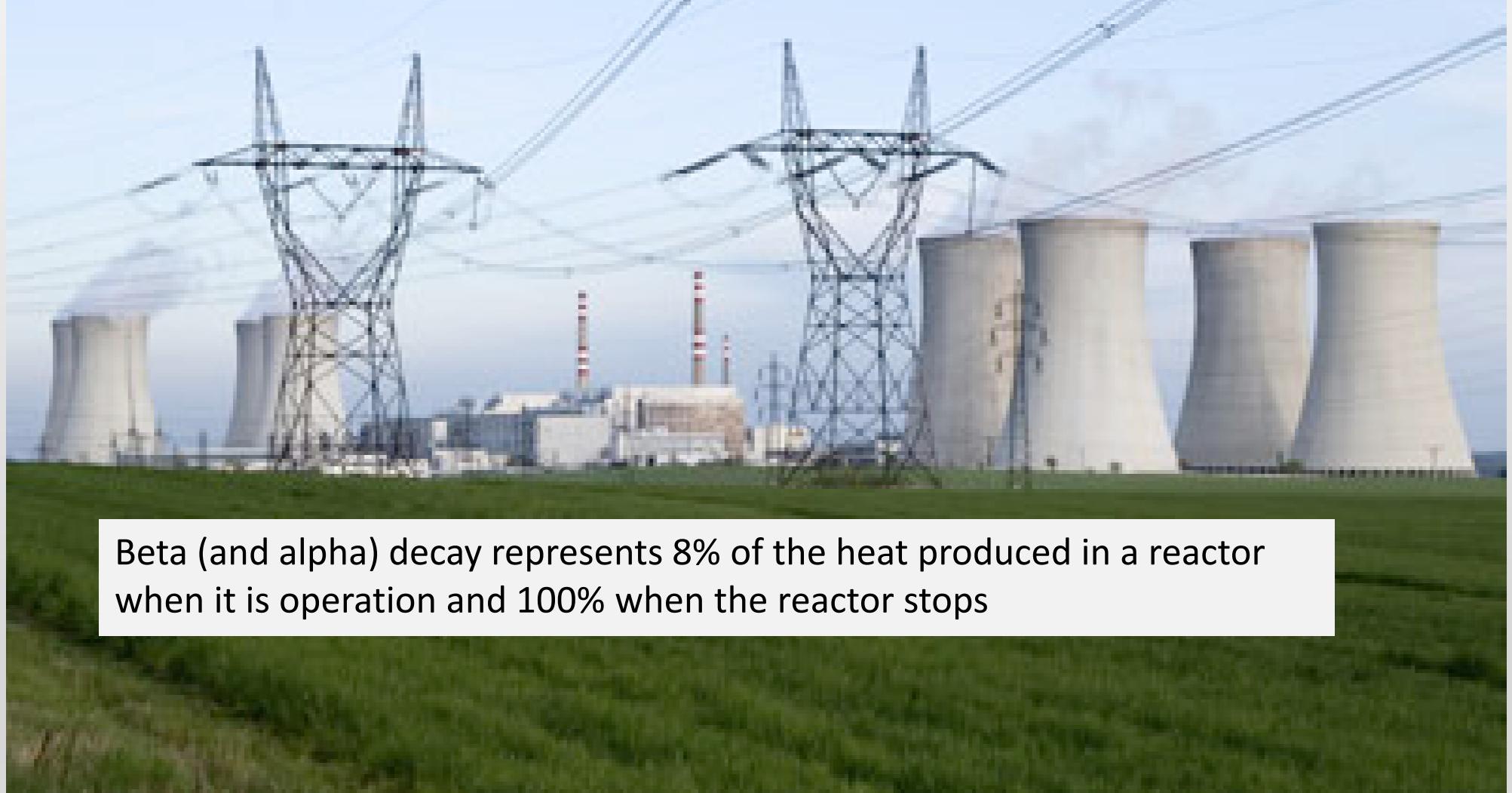
E. Nácher, J.A. Briz, M. Carmona, A. Illana, A. Jungclaus, A. Perea, V. Pesudo, G. Ribeiro, J. Sánchez-del-Río, P. Sarriuguren, J. Taprogge, O. Tengblad
Instituto de Estructura de la Materia – CSIC, Madrid (Spain)

C. Domingo, A. Algora, J. Agramunt, G. Giubrone, V. Guadilla, A. Montaner, S.E.A. Orrigo, B. Rubio, J. L. Tain, E. Valencia
Instituto de Física Corpuscular, CSIC – Universidad de Valencia (Spain)

All these TAS studies
were done on neutron deficient nuclei
And on non refractory elements



TAS measurements for Reactor decay Heat calculations
Motivated by the work by Yoshida.

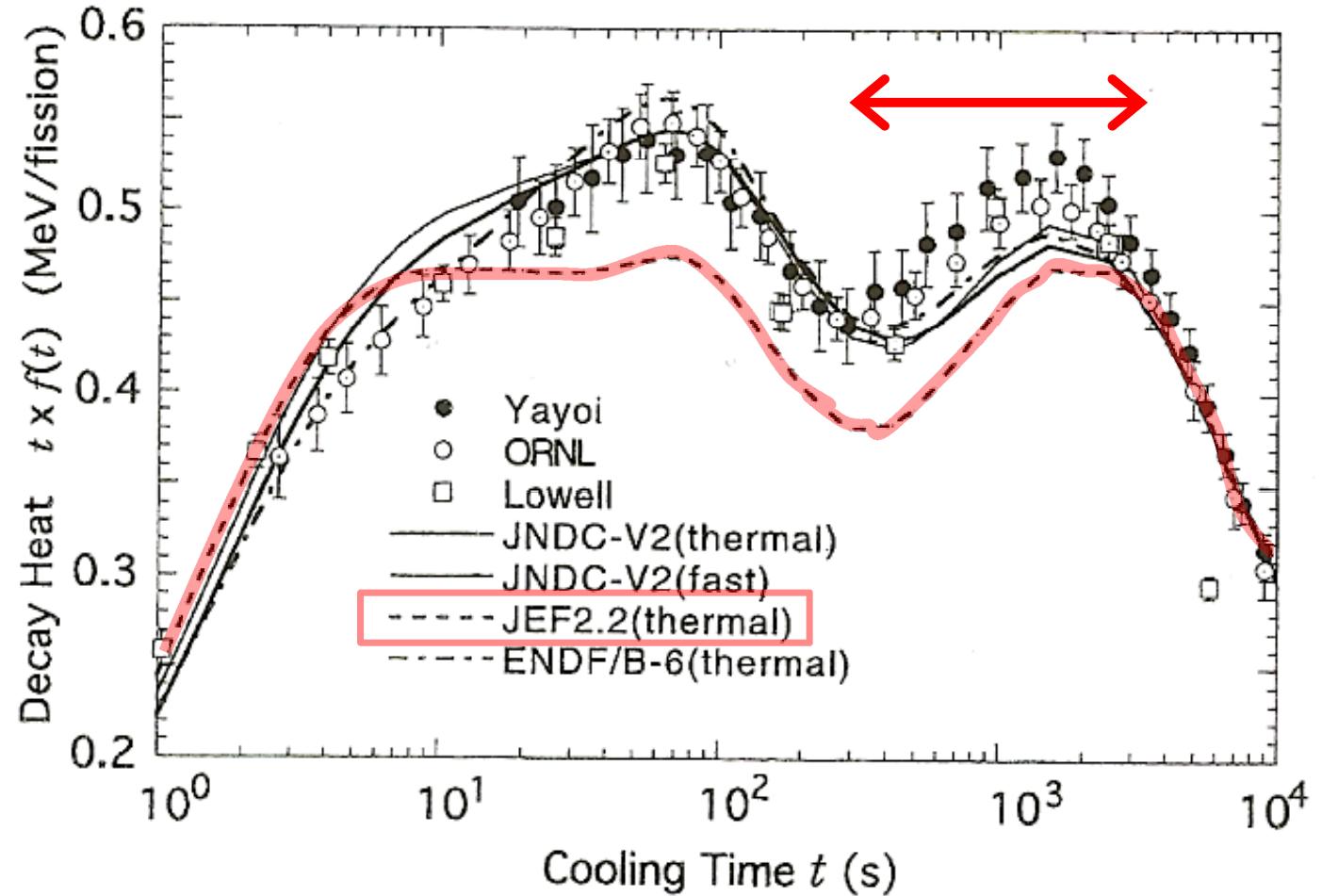


Beta (and alpha) decay represents 8% of the heat produced in a reactor when it is operation and 100% when the reactor stops

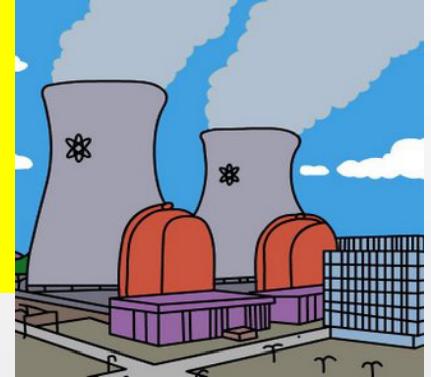
A long standing problem: “The γ -ray discrepancy between decay heat summation calculations and experimental data for reactors decay heat calculations”

^{239}Pu example (γ component)

Yoshida and
co-workers (Journ. of
Nucl. Sc. and Tech.
36 (1999) 135)



Decay heat: summation calculations



$$f(t) = \sum_i E_i \lambda_i N_i(t)$$

E_i Decay energy of the nucleus i (gamma, beta or both)

λ_i Decay constant of the nucleus i $\lambda = \frac{\ln(2)}{T_{1/2}}$

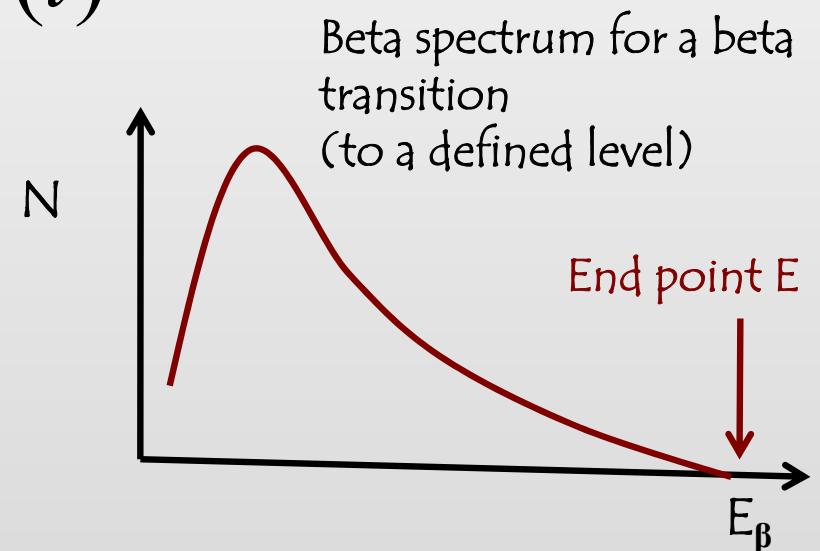
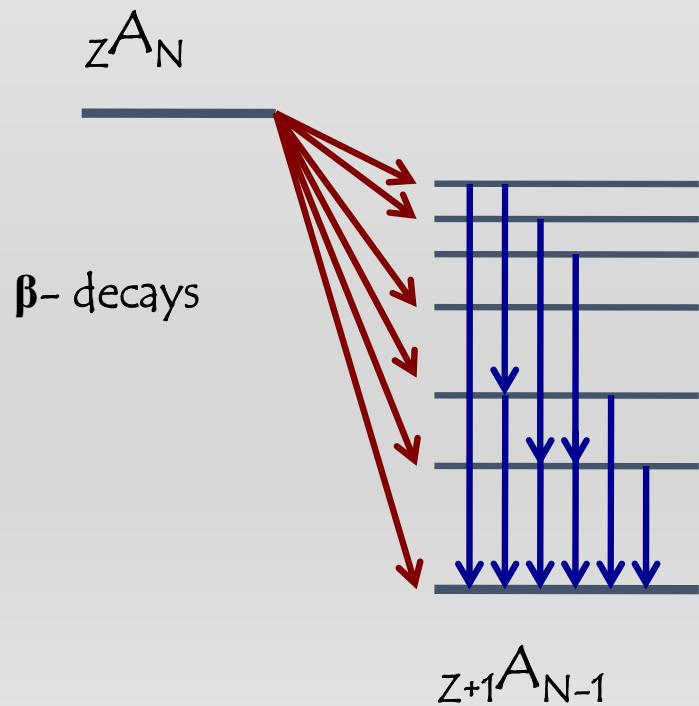
N_i Number of nuclei i at the cooling time t

Requirements for the calculations: large databases that contain all the required information (**half-lives**, mean α - and β -energies released in the decay, n-capture cross sections, fission yields, this last information is needed to calculate the inventory of nuclides)

How the mean energies are determined ?

$$f(t) = \sum_i E_i \lambda_i N_i(t)$$

DATA bases:
feeding or beta
decay prob.
distributions



$$\bar{E}_\beta = \sum_i I_\beta(E_i) \langle E_{\beta,i} \rangle$$

$$\bar{E}_\gamma = \sum_i I_\beta(E_i) E_i$$

Why JYFL: IGISOL + JYFLTRAP

space for post-trap decay spectroscopy setups

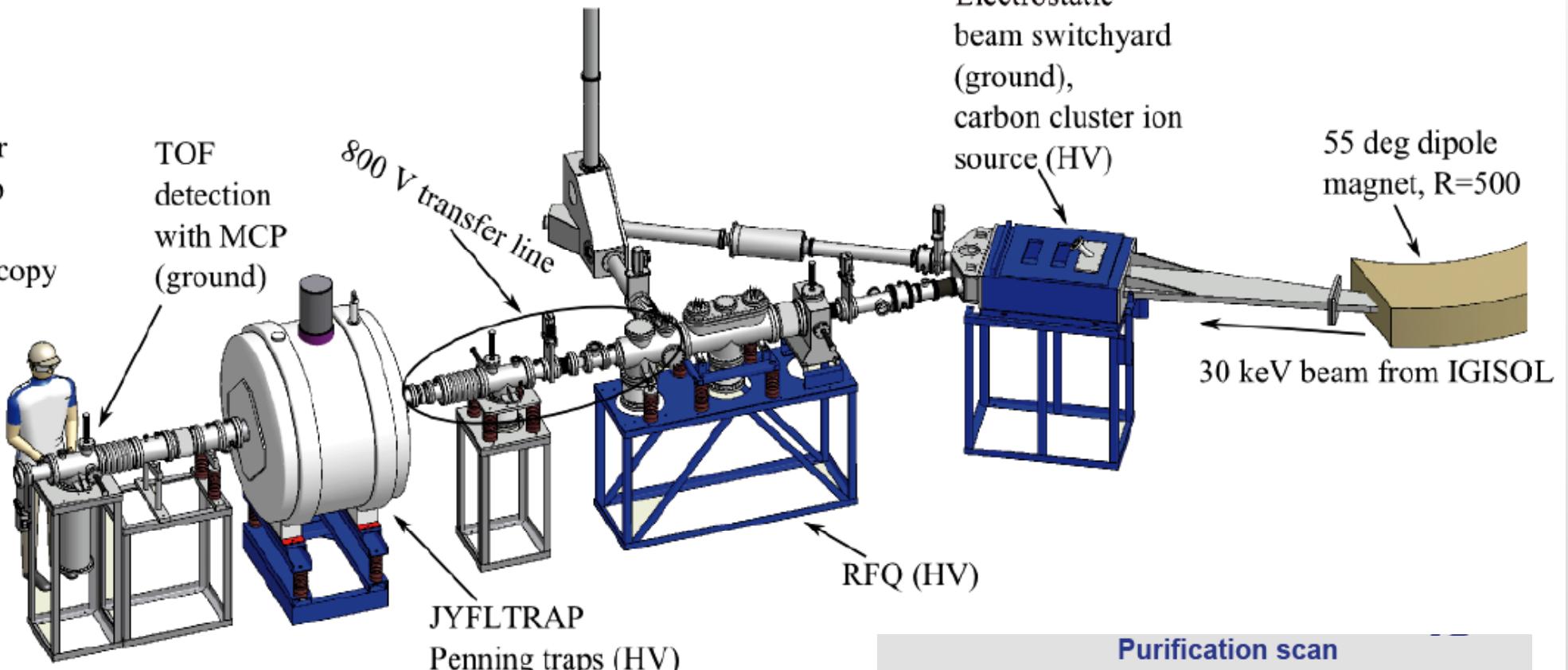
TOF detection with MCP (ground)

800 V transfer line

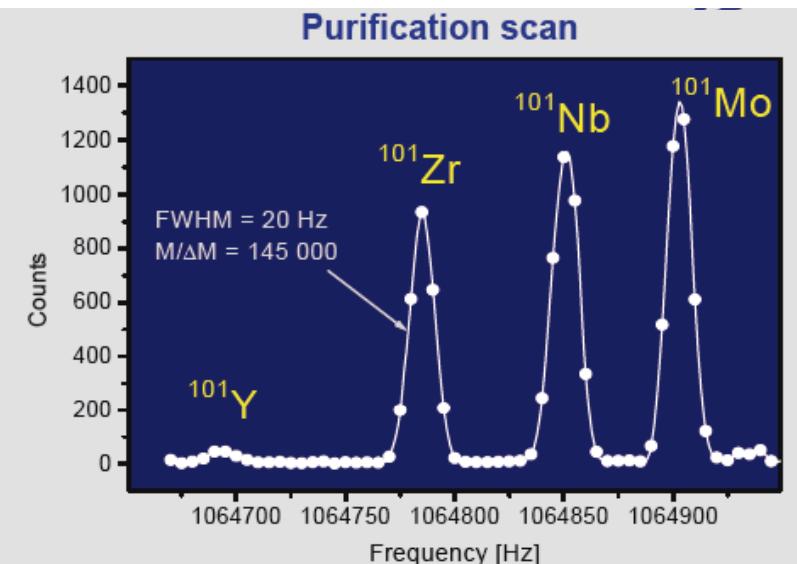
Electrostatic beam switchyard (ground), carbon cluster ion source (HV)

55 deg dipole magnet, $R=500$

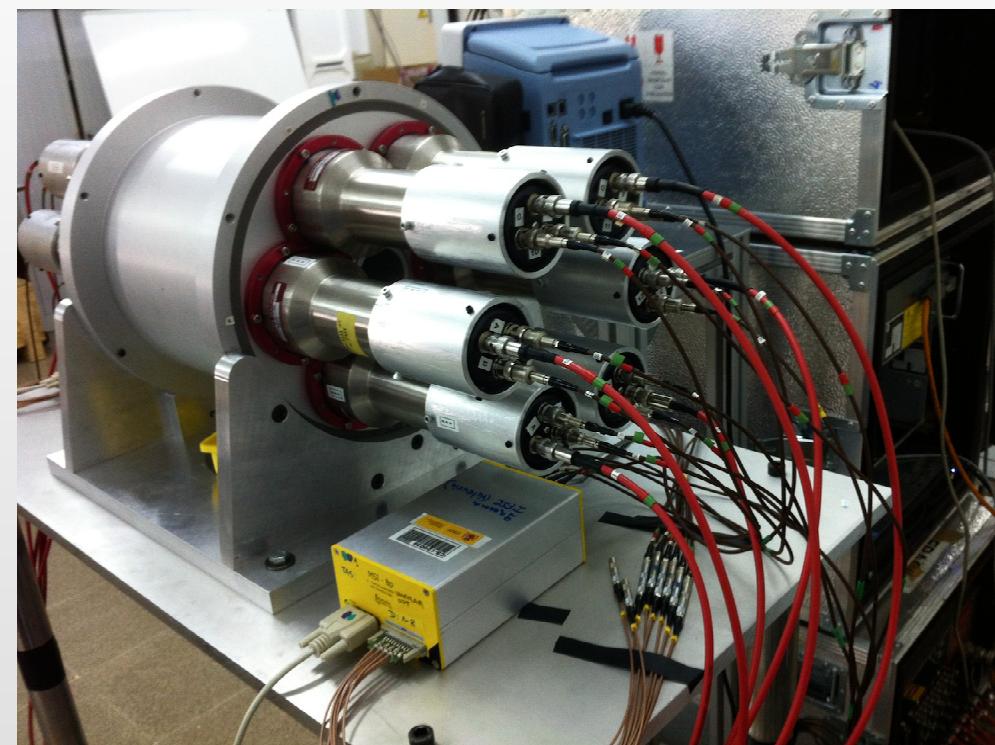
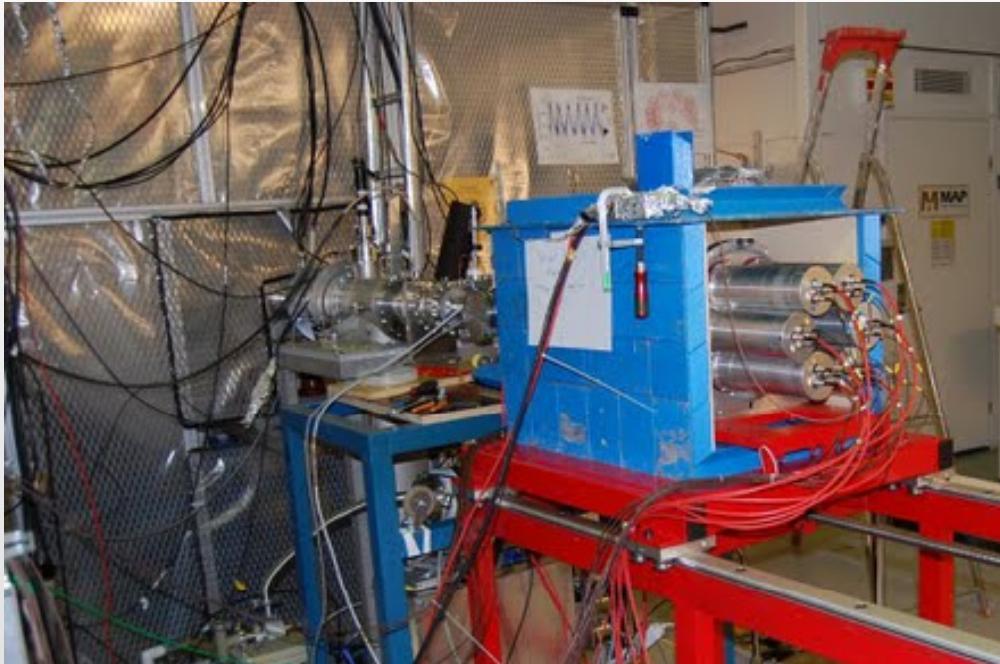
30 keV beam from IGISOL



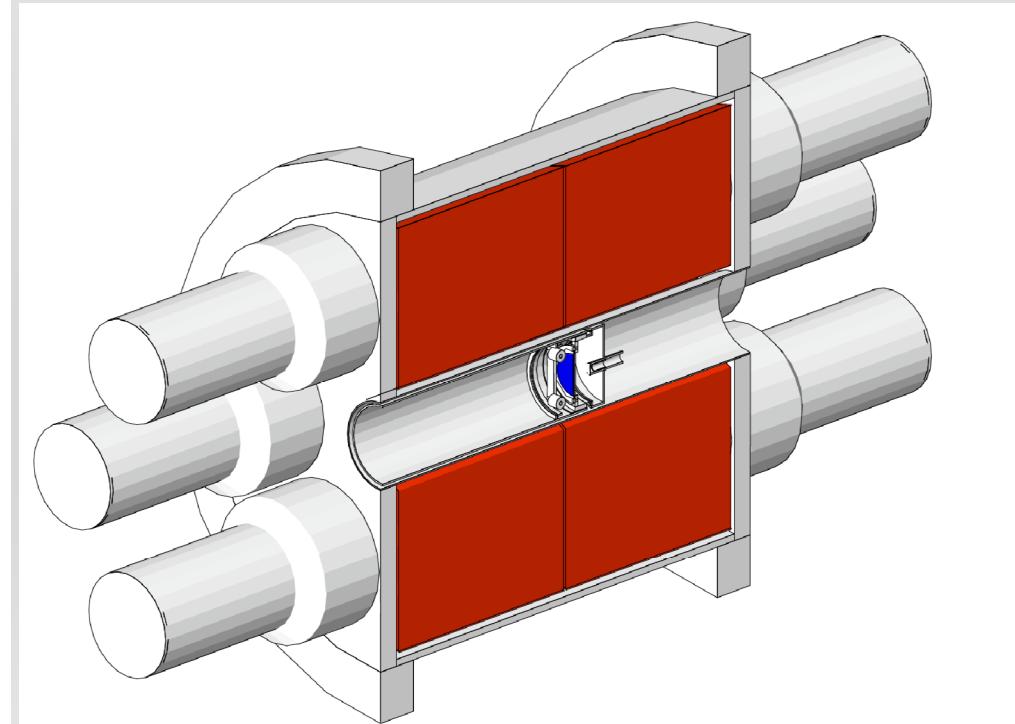
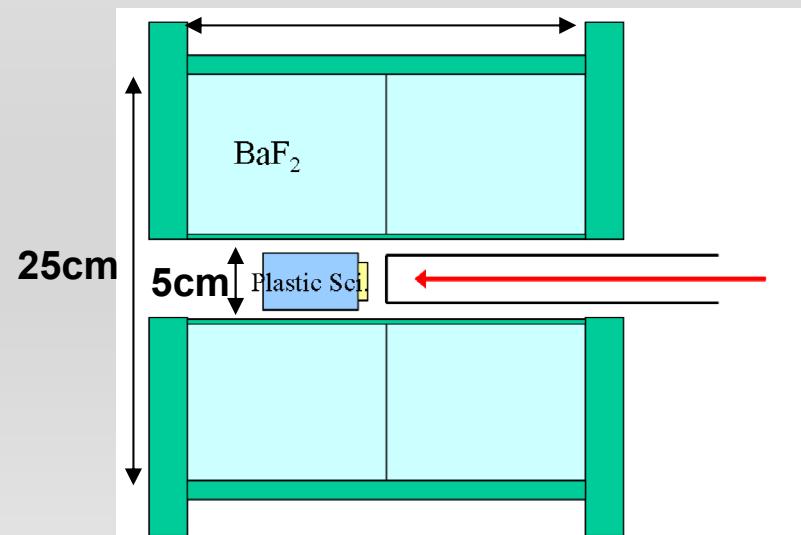
The main reasons are the chemical insensitivity (ion guide technique), high purity by means of purification of the beam using the JYFLTRAP and acceptable yields!



Segmented BaF₂ Total Absorption Spectrometer: Rocinante (Valencia-Surrey)



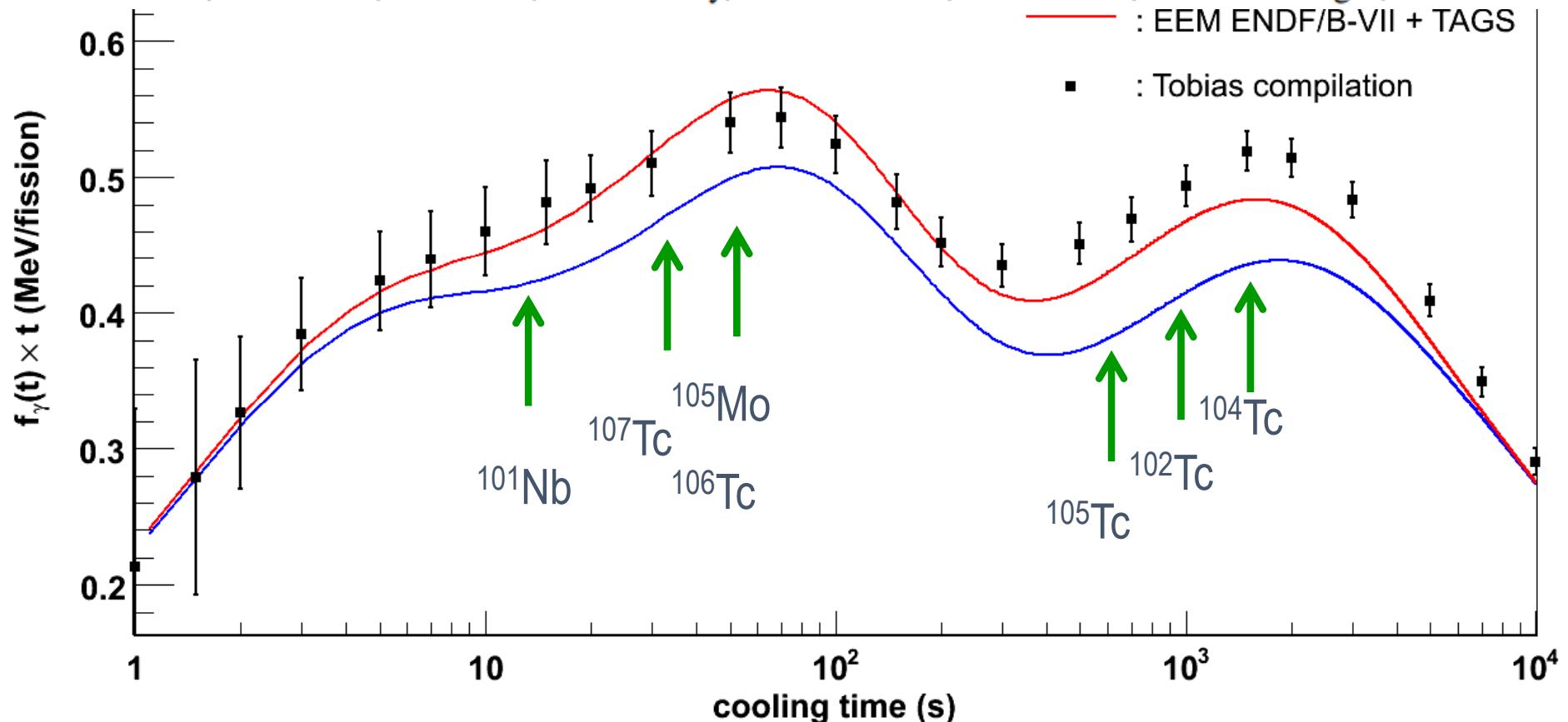
: 12× BaF₂ crystals / 3" Q-PMT
 cascade multiplicity information
25cm





Reactor Decay Heat in ^{239}Pu : Solving the γ Discrepancy in the 4–3000-s Cooling Period

A. Algora,^{1,2,*} D. Jordan,¹ J. L. Taín,¹ B. Rubio,¹ J. Agramunt,¹ A. B. Perez-Cerdan,¹ F. Molina,¹ L. Caballero,¹ E. Nácher,¹ A. Krasznahorkay,² M. D. Hunyadi,² J. Gulyás,² A. Vitéz,² M. Csatlós,² L. Csige,² J. Äystö,³ H. Penttilä,³ I. D. Moore,³ T. Eronen,³ A. Jokinen,³ A. Nieminen,³ J. Hakala,³ P. Karvonen,³ A. Kankainen,³ A. Saastamoinen,³ J. Rissanen,³ T. Kessler,³ C. Weber,³ J. Ronkainen,³ S. Rahaman,³ V. Elomaa,³ S. Rinta-Antila,³ U. Hager,³ T. Sonoda,³ K. Burkard,⁴ W. Hüller,⁴ L. Batist,⁵ W. Gelletly,⁶ A. L. Nichols,⁶ T. Yoshida,⁷ A. A. Sonzogni,⁸ and K. Peräjärvi⁹

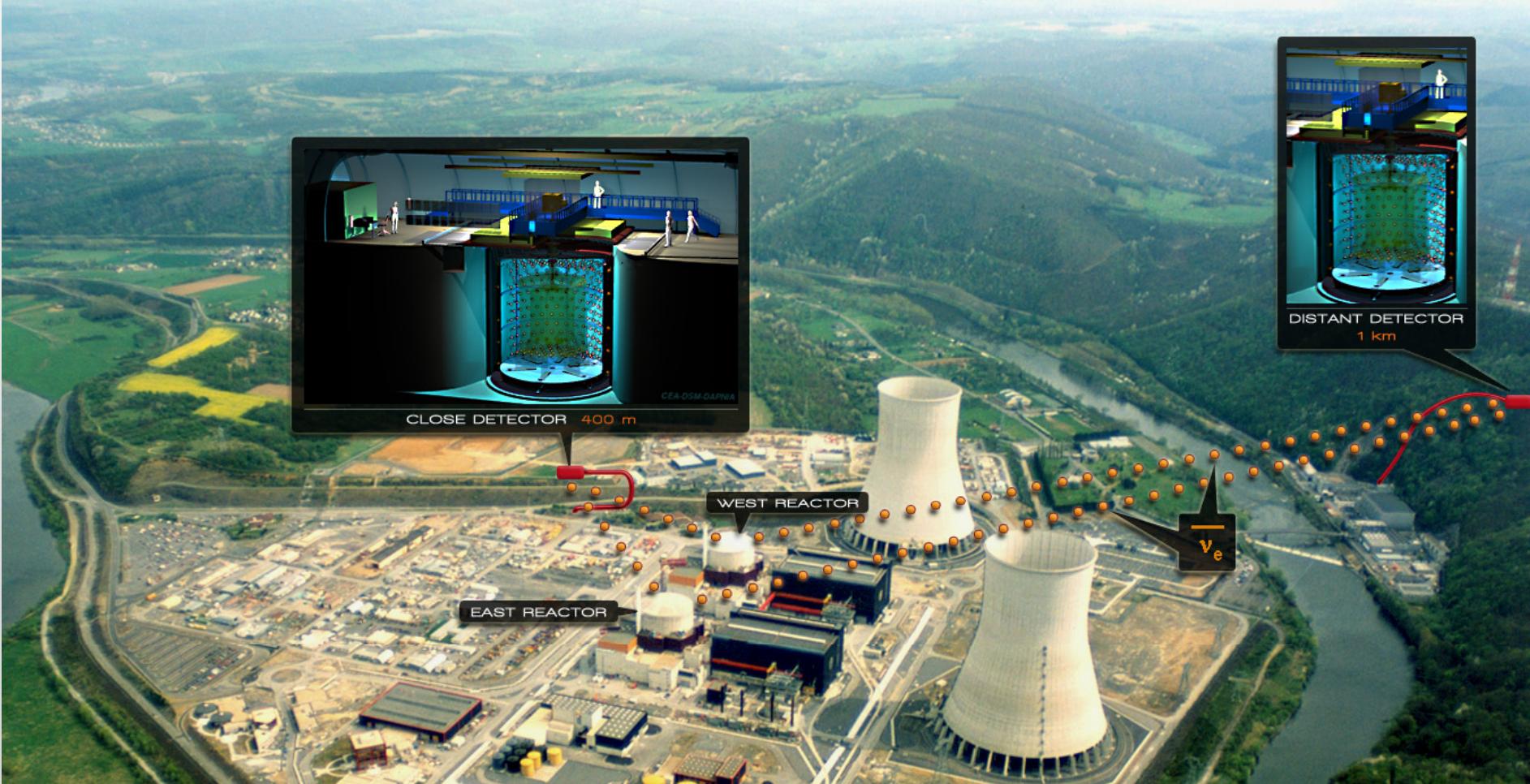


Spokespersons

A. Algora, J.L. Taín

Courtesy A. Sonzogni

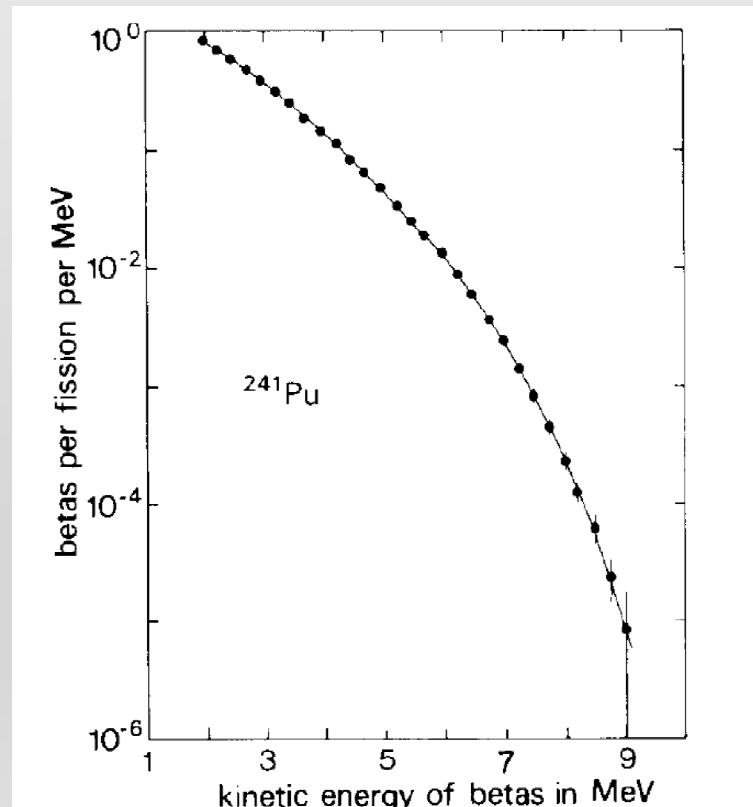
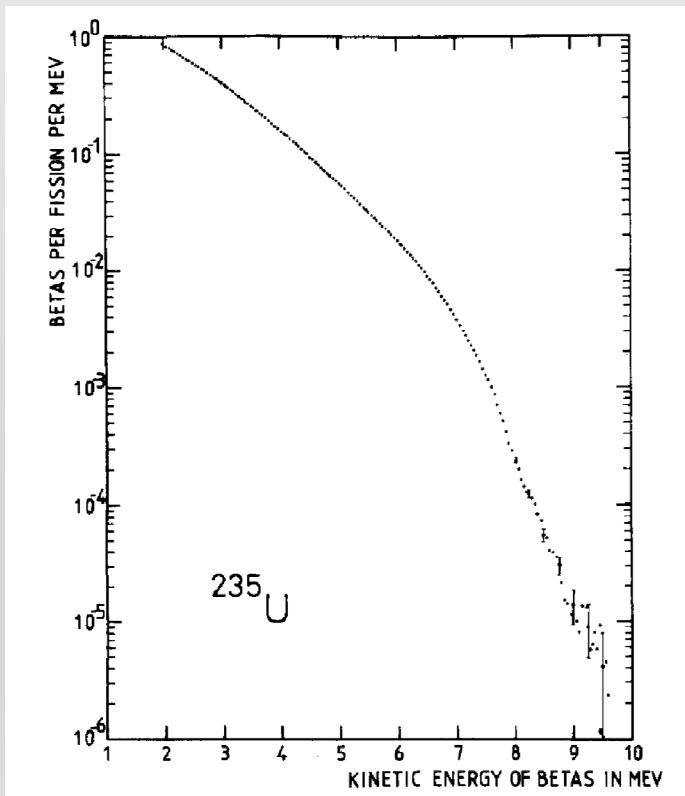
Another impact of our data



TAS measurements for the prediction of the neutrino spectrum from Reactors. Important for experiments of neutrino oscillations such as DOUBLE CHOOZ. Impact in the energy range where the antineutrino anomaly has been observed
Spokespersons: M. Fallot, A. Algara, J.L. Taín...

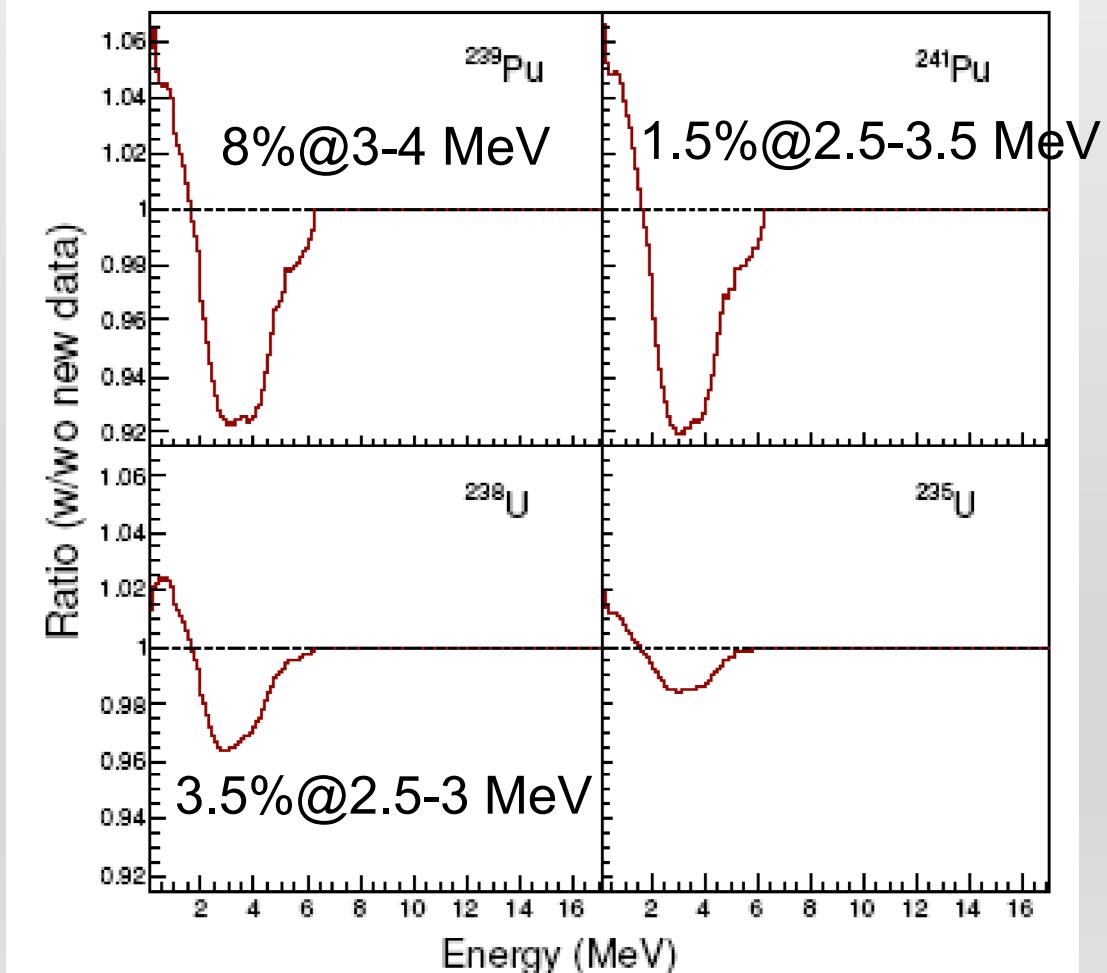
Determination of the primary neutrino spectrum: neutrino physics, non-proliferation applications

- Using the beta spectrum measured by Schreckenbach et al. from different fissile nuclides (^{235}U , $^{239,241}\text{Pu}$) and more recently ^{238}U (Haag et al.), which requires complex conversion procedures



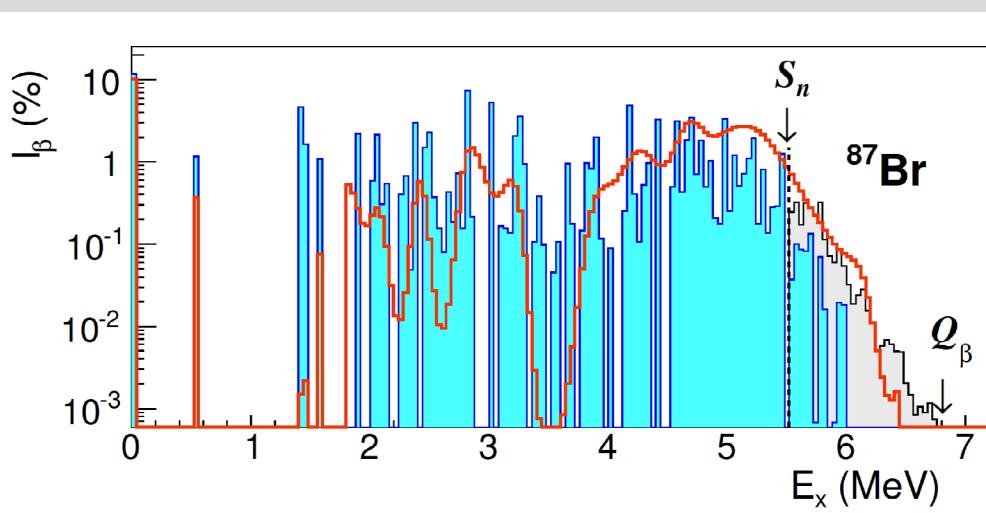
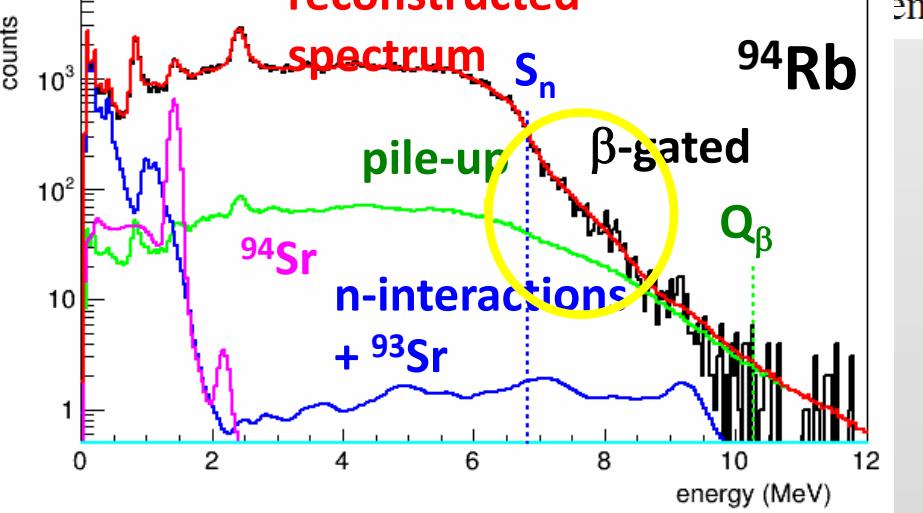
New Antineutrino Energy Spectra Predictions from the Summation of Beta Decay Branches of the Fission Products

M. Fallot,¹ S. Cormon,¹ M. Estienne,¹ A. Algora,^{2,3} V. M. Bui,¹ A. Cuocoanes,¹ M. Elnimr,¹ L. Giot,¹ D. Jordan,² J. Martino,¹ A. Onillon,¹ A. Porta,¹ G. Pronost,¹ A. Remoto,¹ J. L. Taín,² F. Yermia,¹ and A.-A. Zakari-Issoufou¹



Enhanced γ -Ray Emission from Neutron Unbound States Populated in β Decay

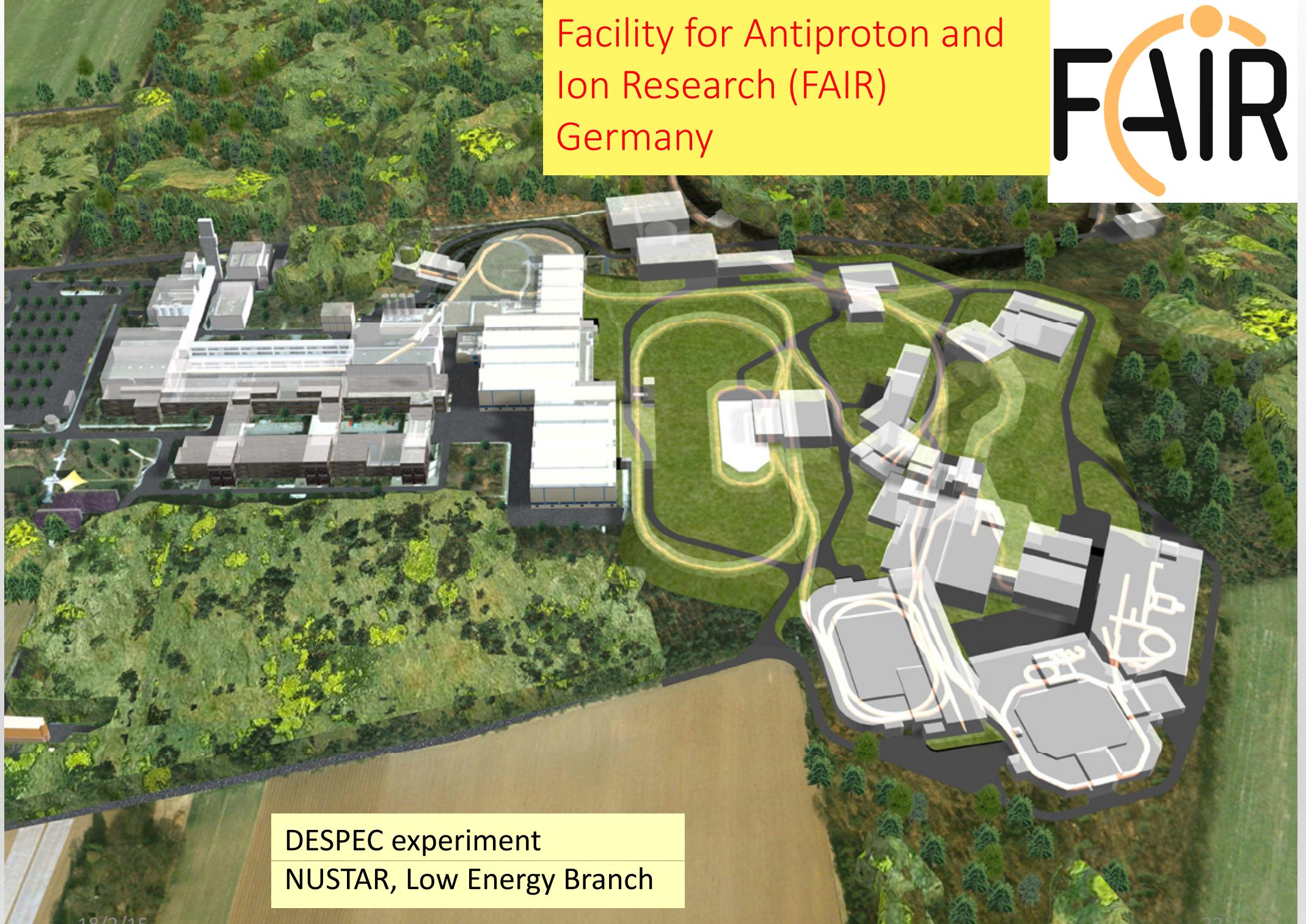
J. L. Tain,^{1,*} E. Valencia,¹ A. Algara,¹ J. Agramunt,¹ B. Rubio,¹ S. Rice,² W. Gelletly,² P. Regan,² A.-A. Zakari-Issoufou,³ M. Fallot,³ A. Porta,³ J. Rissanen,⁴ T. Eronen,⁴ J. Äystö,⁵ L. Batist,⁶ M. Bowry,² V. M. Bui,³ R. Caballero-Folch,⁷ D. Cano-Ott,⁸ V.-V. Elomaa,⁴ E. Estevez,¹ G. F. Farrelly,² A. R. Garcia,⁸ B. Gomez-Hornillos,⁷ V. Gorlychev,⁷ J. Hakala,⁴ M. D. Jordan,¹ A. Jokinen,⁴ V. S. Kolhinen,⁴ F. G. Kondev,⁹ T. Martinez,⁸ E. Mendoza,⁸ I. Moore,⁴ H. Penttilä,⁴ en,⁴ V. Sonnenschein,⁴ and A. A. Sonzogni¹⁰



β -delayed neutron emitters:
 $^{87}\text{Br}, ^{88}\text{Br}, ^{93}\text{Rb}, ^{94}\text{Rb}$ (1st exp.)
+ $^{95}\text{Rb}, ^{137}\text{I}, ^{138}\text{I}$ (2nd exp.)

Zr 90	Zr 91	Zr 92	Zr 93	Zr 94	Zr 95	Zr 96	Zr 97	Zr 98
51.45	11.22	17.15	$1.5 \cdot 10^6$ a	17.38	64.0 d	$2.8 \cdot 10^6$ a	16.8 h	30.7 s
(β^- , n)								
$\tau = 0.014$	$\tau = 1.2$	$\tau = 0.2$	$\tau = 0.06$	$\tau = 0.049$	$\tau = 0.11$	$\tau = 1.9$	$\tau = 8.8$	$\tau = 2.3$
$\nu = 0.0058$								
$Y\ 89$	$Y\ 90$	$Y\ 91$	$Y\ 92$	$Y\ 93$	$Y\ 94$	$Y\ 95$	$Y\ 96$	$Y\ 97$
16.0 y	16.1 h	48.7 m	3.54 h	10.1 h	18.7 m	10.3 m	5.34 s	3.78 s
$\nu = 0.0058$								
$Sr\ 88$	$Sr\ 89$	$Sr\ 90$	$Sr\ 91$	$Sr\ 92$	$Sr\ 93$	$Sr\ 94$	$Sr\ 95$	$Sr\ 96$
82.58	50.5 d	28.64 a	9.5 h	2.7 h	7.45 m	7.4 s	24.4 s	1.0 s
$\nu = 0.0058$								
$Rb\ 87$	$Rb\ 88$	$Rb\ 89$	$Rb\ 90$	$Rb\ 91$	$Rb\ 92$	$Rb\ 93$	$Rb\ 94$	$Rb\ 95$
27.83	17. m	15.2 m	8.3 m	3.8 s	4.5 s	5.8 s	7.5 s	37. ms
$\nu = 0.0058$								
$Kr\ 86$	$Kr\ 87$	$Kr\ 88$	$Kr\ 89$	$Kr\ 90$	$Kr\ 91$	$Kr\ 92$	$Kr\ 93$	$Kr\ 94$
17.279	7.63 m	2.84 m	3.18 m	32.3 s	8.6 s	1.84 s	1.29 s	0.20 s
$\nu = 0.0058$								
$Br\ 85$	$Br\ 86$	$Br\ 87$	$Br\ 88$	$Br\ 89$	$Br\ 90$	$Br\ 91$	$Br\ 92$	$Br\ 93$
2.87 m	1 s	5.7 s	16.3 s	4.40 s	4.40 s	0.64 s	343 ms	102 ms
$\nu = 0.0058$								
$Se\ 84$	$Se\ 85$	$Se\ 86$	$Se\ 87$	$Se\ 88$	$Se\ 89$	$Se\ 90$	$Se\ 91$	$Se\ 92$
3.1 m	33 s	14.1 s	5.8 s	1.5 s	0.4 s	>300 ns	0.27 s	>300 ns
$\nu = 0.0058$								
$As\ 83$	$As\ 84$	$As\ 85$	$As\ 86$	$As\ 87$	$As\ 88$	$As\ 89$	$As\ 90$	$As\ 91$
$\nu = 0.0058$								

$^{86}\text{Kr}(n,\gamma)$: Raman+, PRC28, 1983



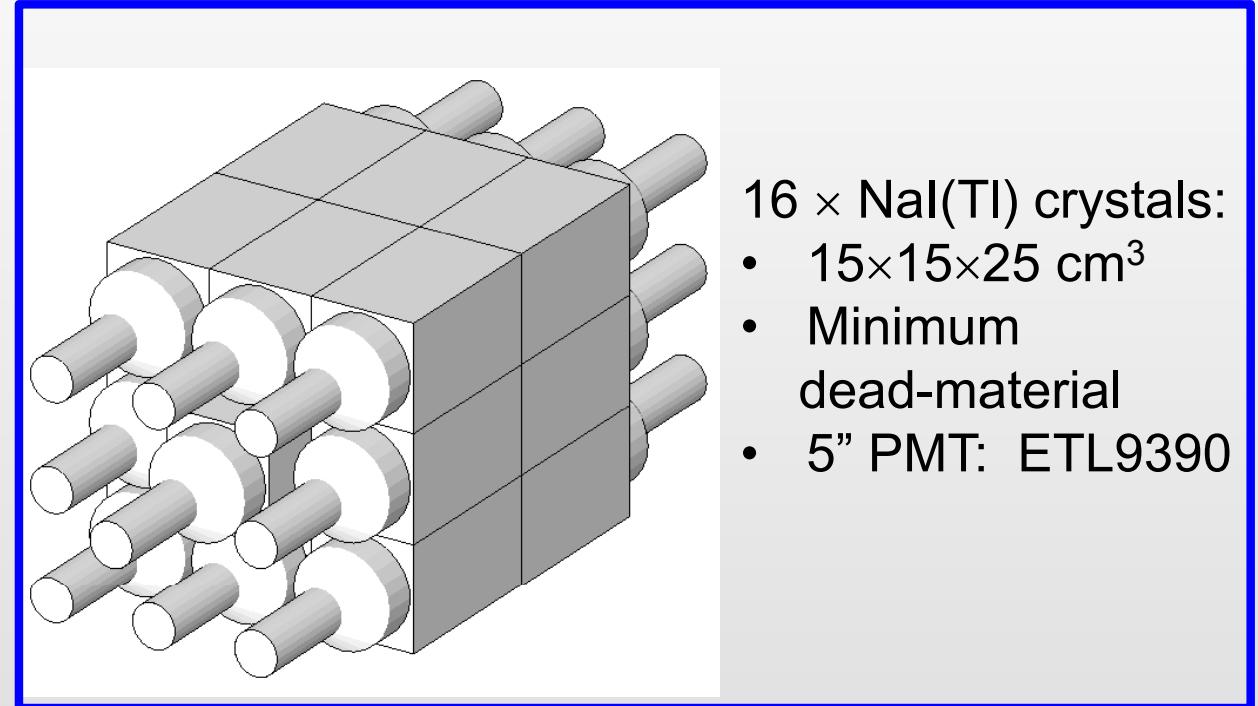
Facility for Antiproton and
Ion Research (FAIR)
Germany



DESPEC experiment
NUSTAR, Low Energy Branch

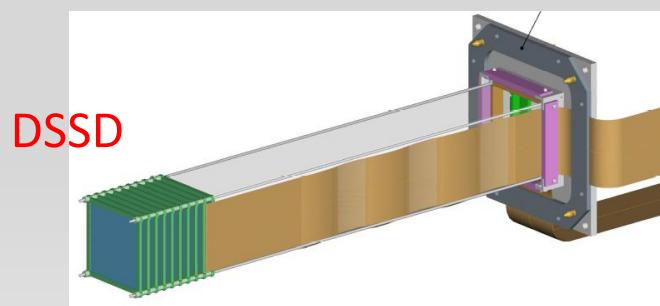
DTAS 17 independent NaI crystals (Valencia-Surrey-Madrid)

A new DTAS was designed and constructed to work at FAIR (DESPEC experiment)

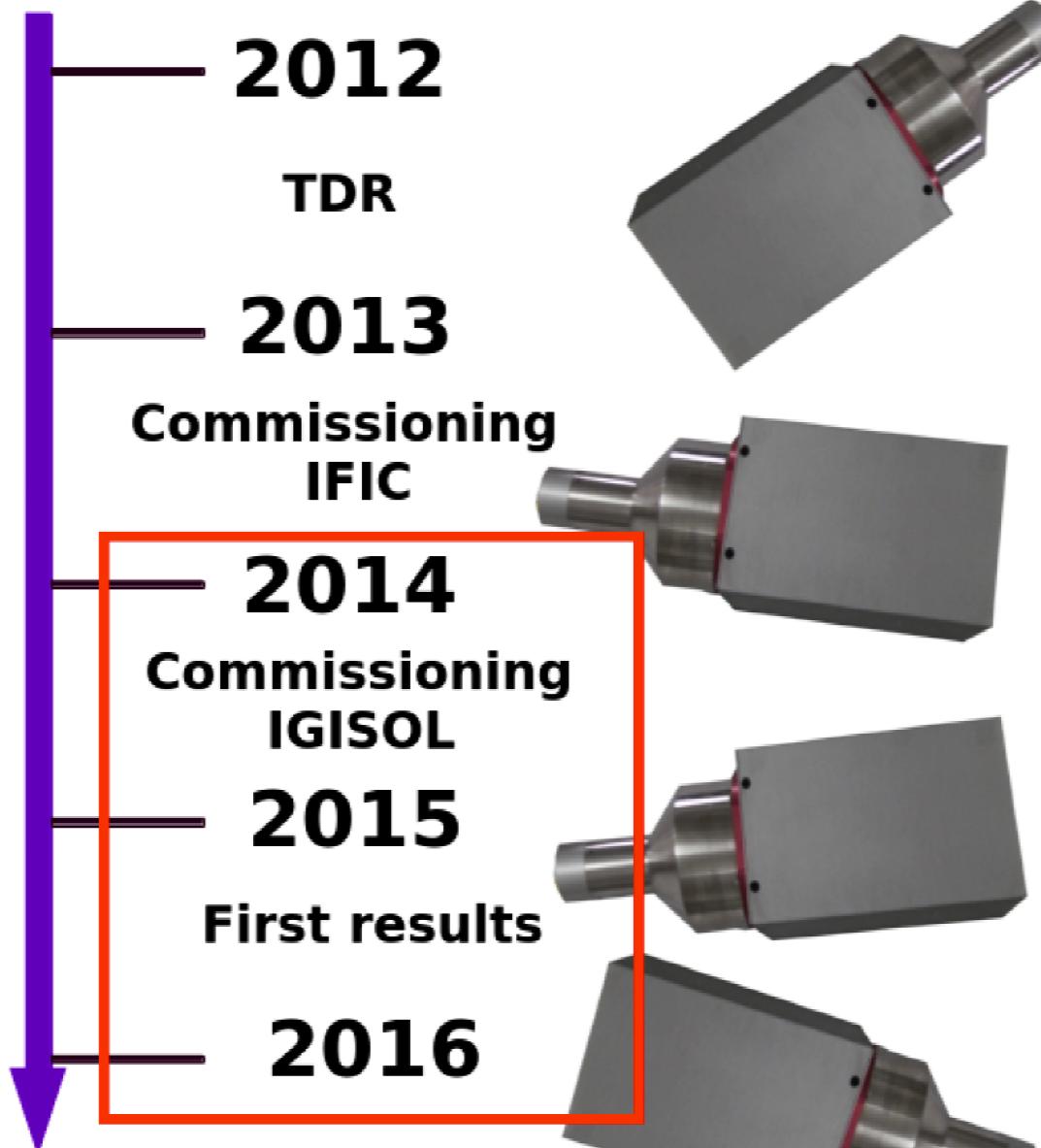
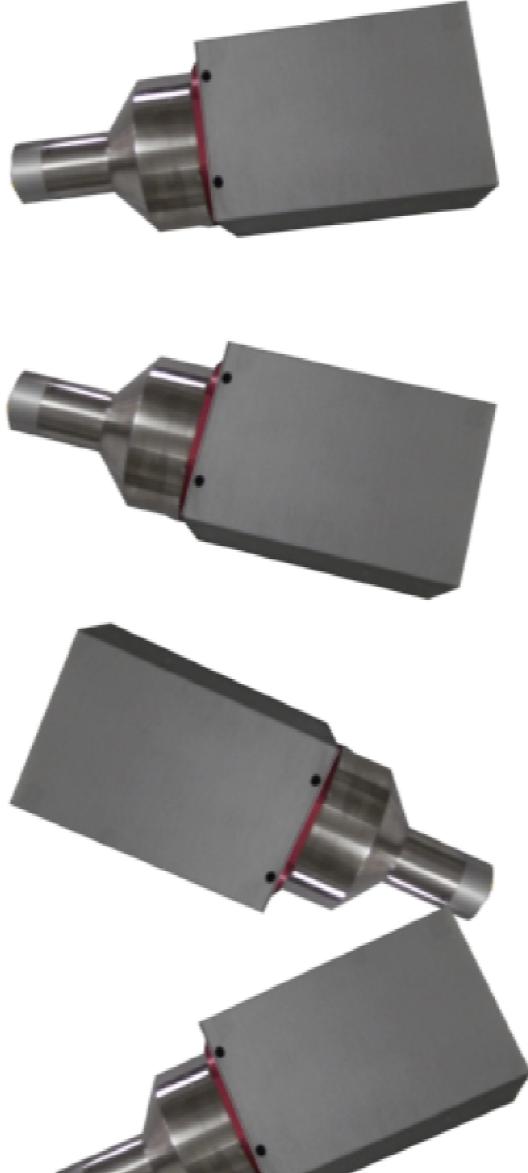


- 16 × NaI(Tl) crystals:
- $15 \times 15 \times 25 \text{ cm}^3$
 - Minimum dead-material
 - 5" PMT: ETL9390

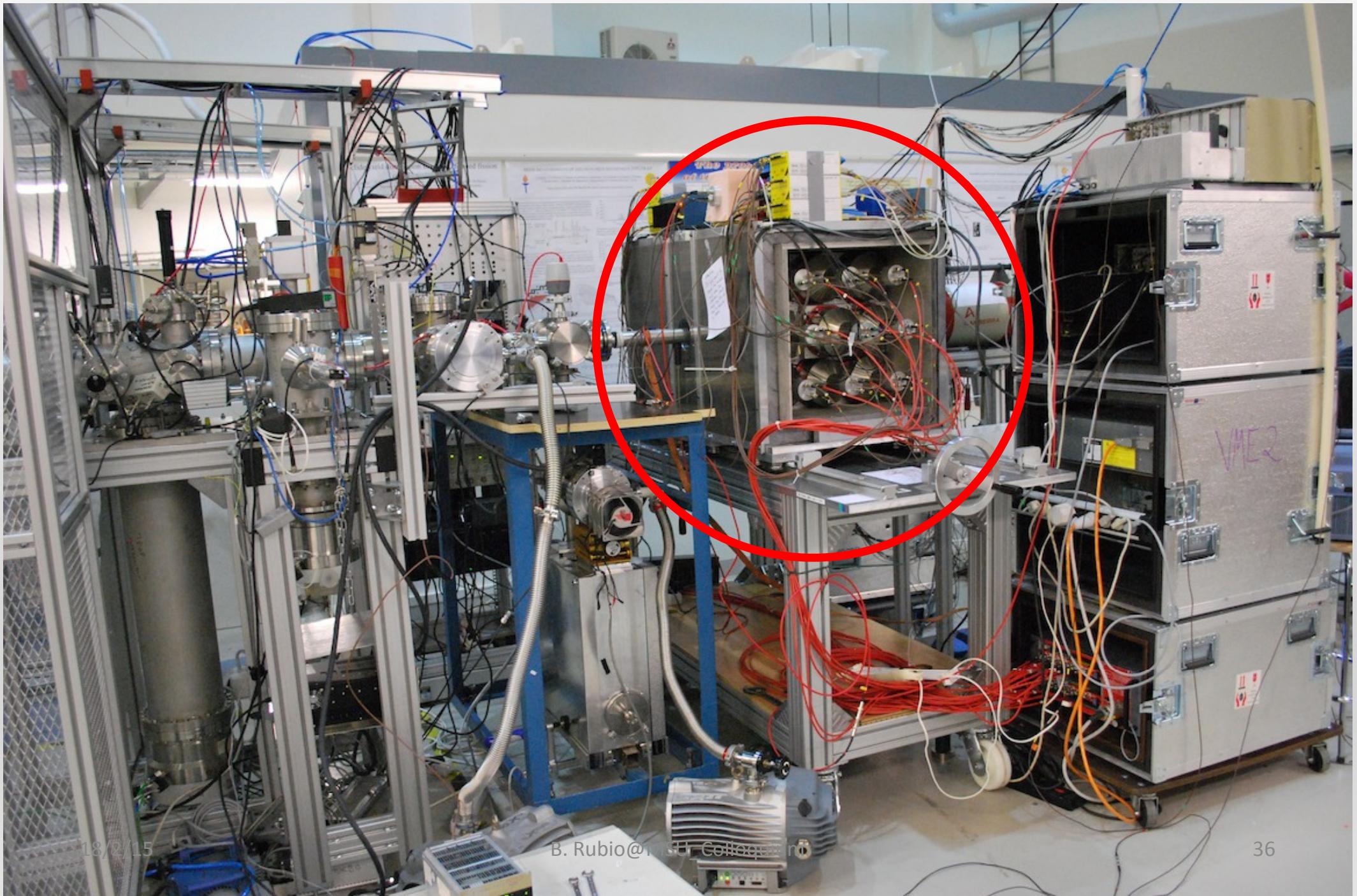
Designed to be coupled to AIDA implantation detector



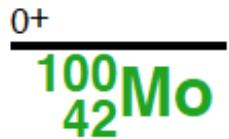
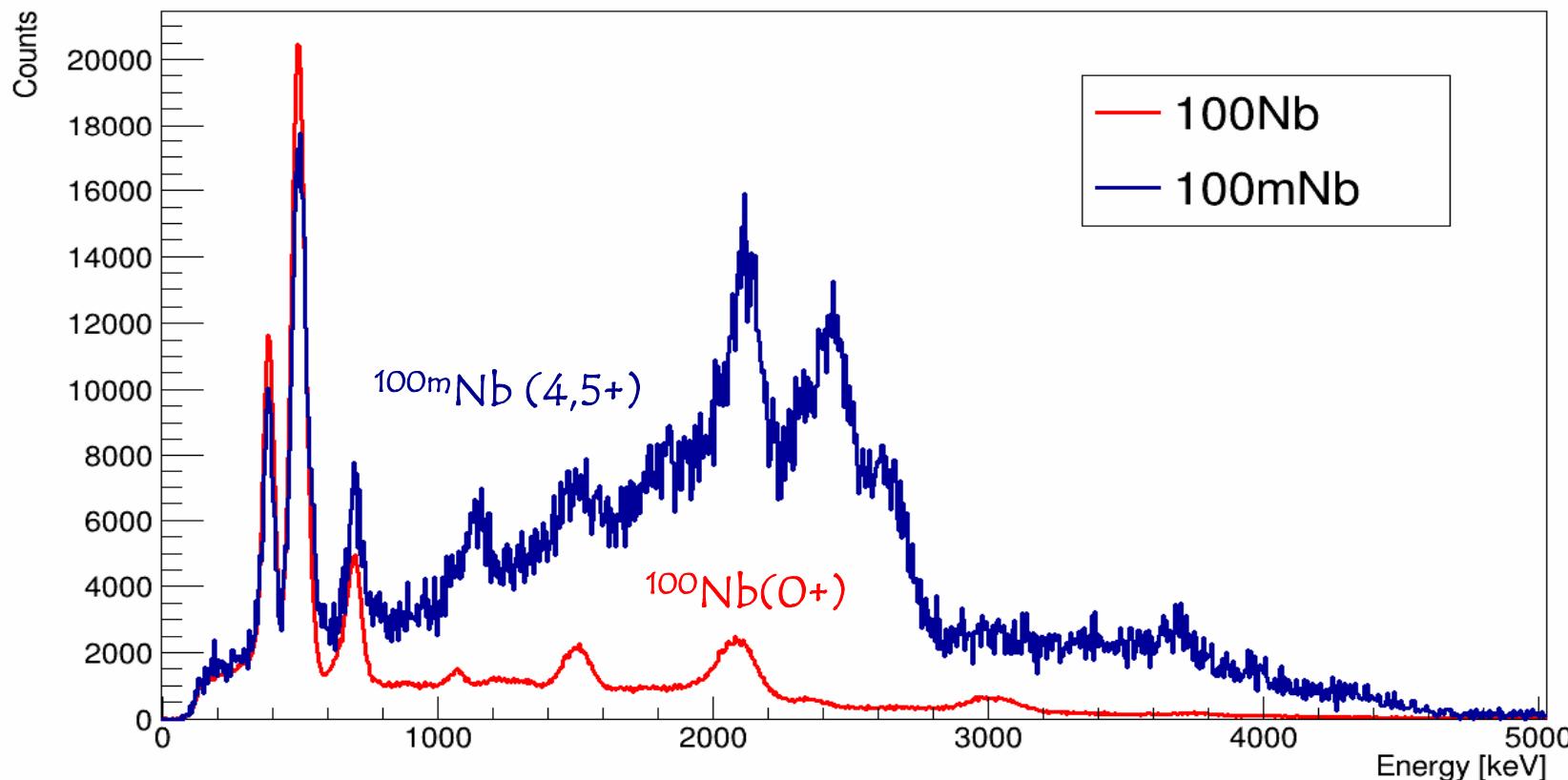
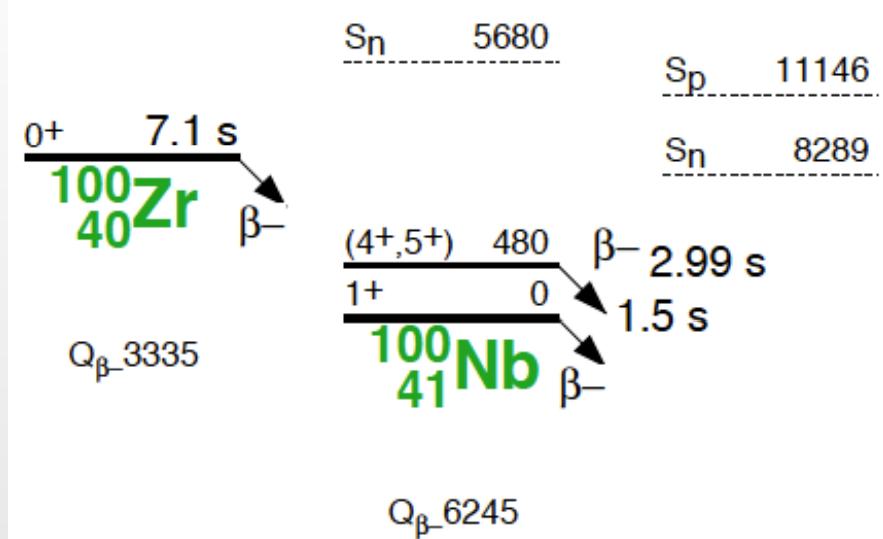
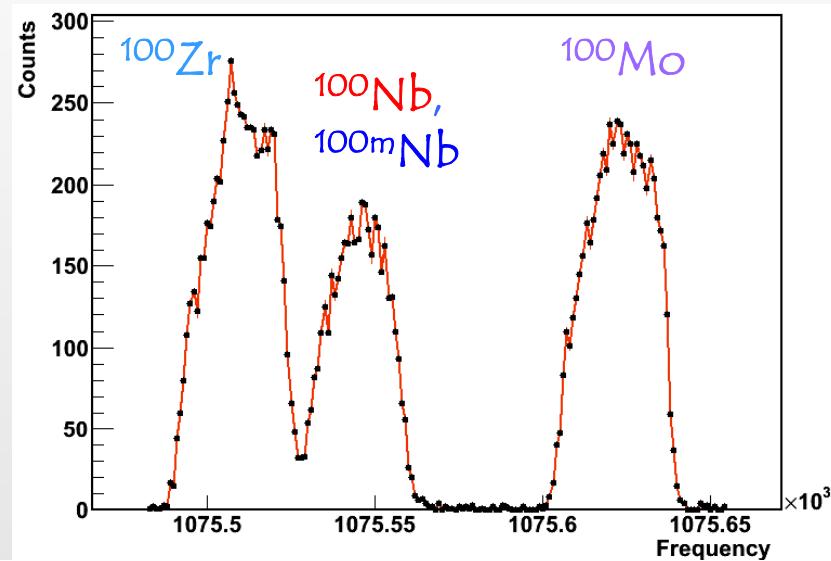
Chronology



March 2014



JYFLTRAP, A=100 Mass Scan



Beta gated
TAS
spectra of
the two
isomers

V. Guadilla
courtesy

New questions: reactor anomaly ?

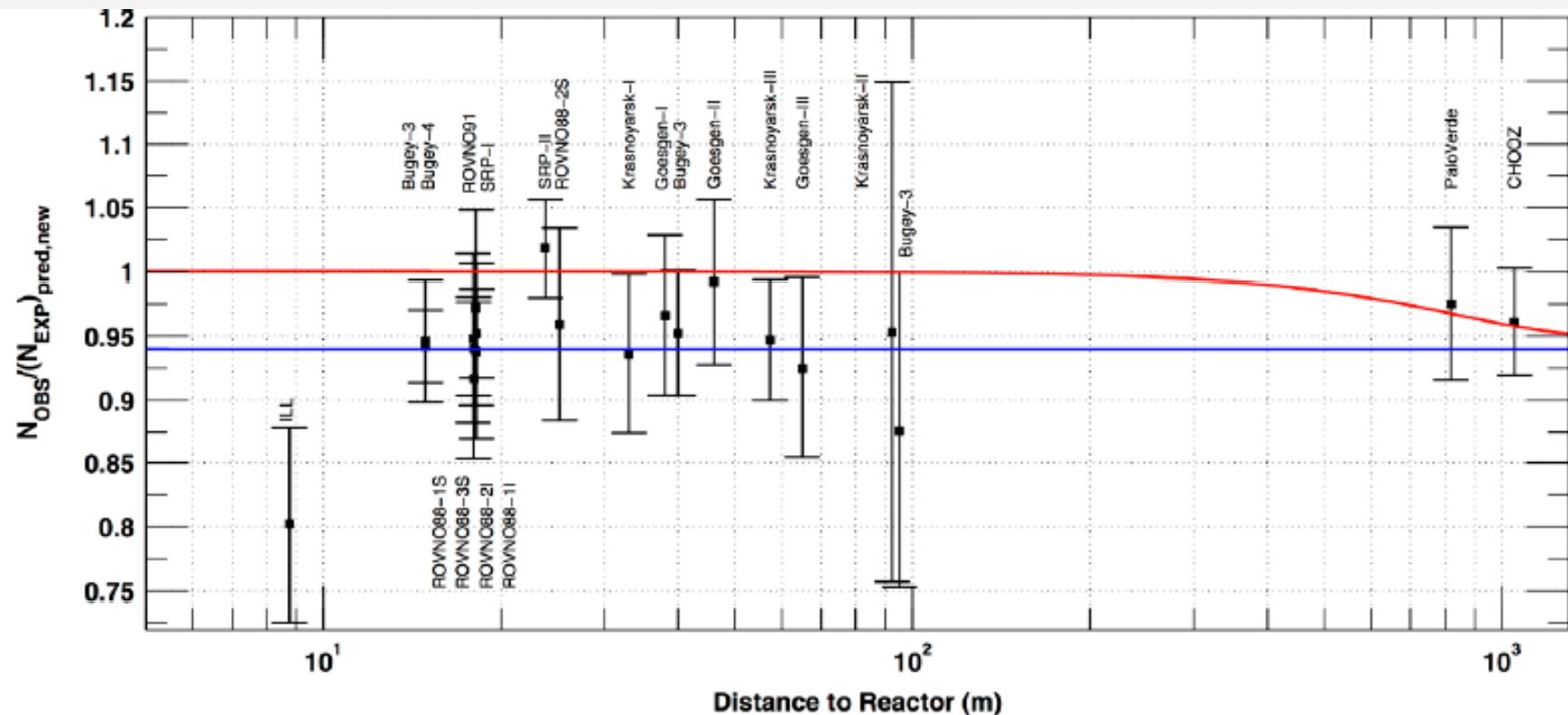


Illustration of the reactor anomaly. Rates in various experiments are compared with the expectations based on the Mueller et al. (2011) spectrum. The mean is 0.943 ± 0.023 .

Possible explanation:

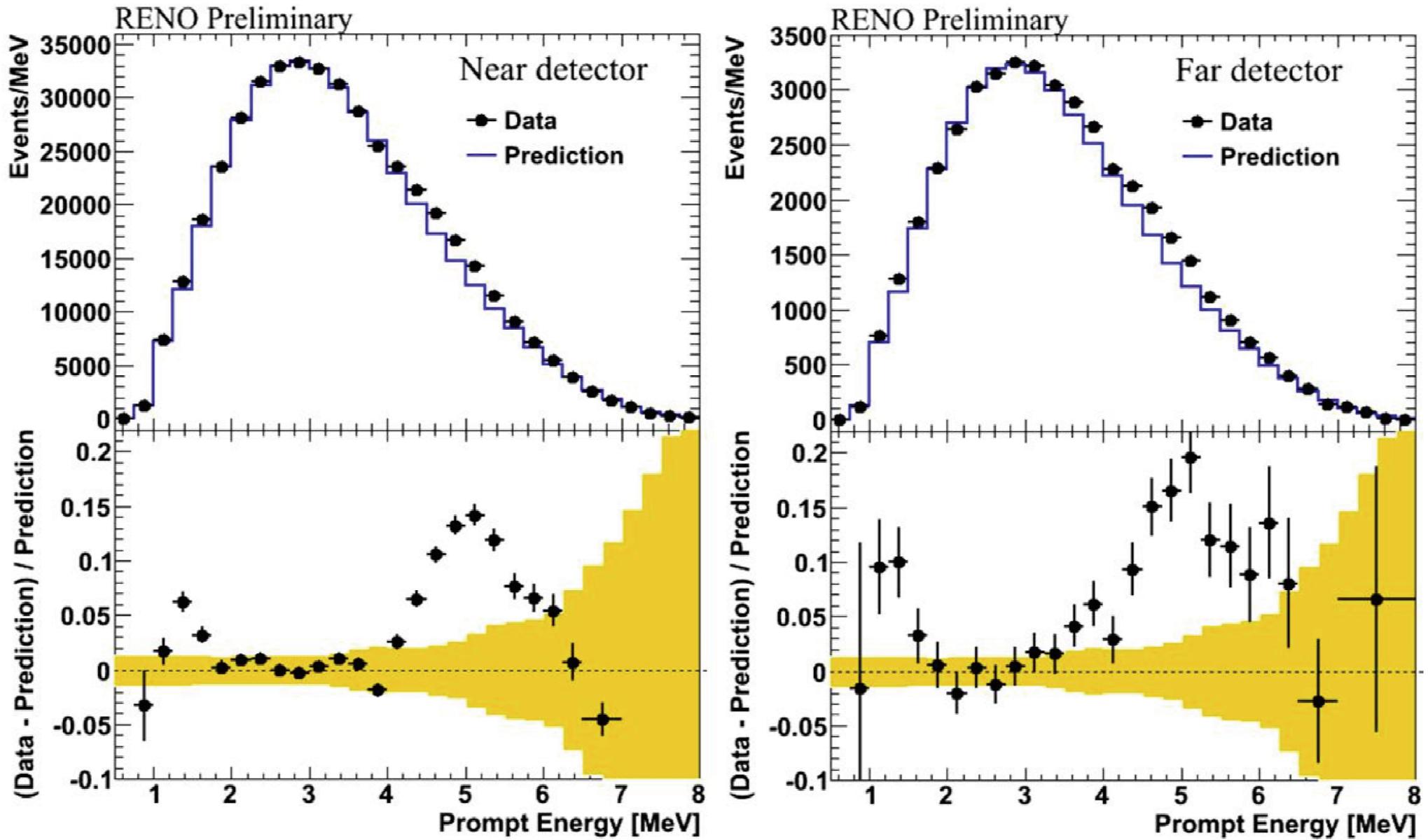
- 1) Wrong reactor flux or its error
- 2) Bias in all experiments
- 3) New physics at short baseline involving a sterile 4th neutrino

ν_{new} with $\Delta m^2 \sim 1 \text{ eV}^2$ and mixing with ν_e with $\theta_{\text{new}} \sim 10^\circ$

From P. Vogel

The explanation 3) could be supported by several other, so far unconfirmed anomalies. It would involve unexpected but significant "New Physics"

A new famous bump !!!

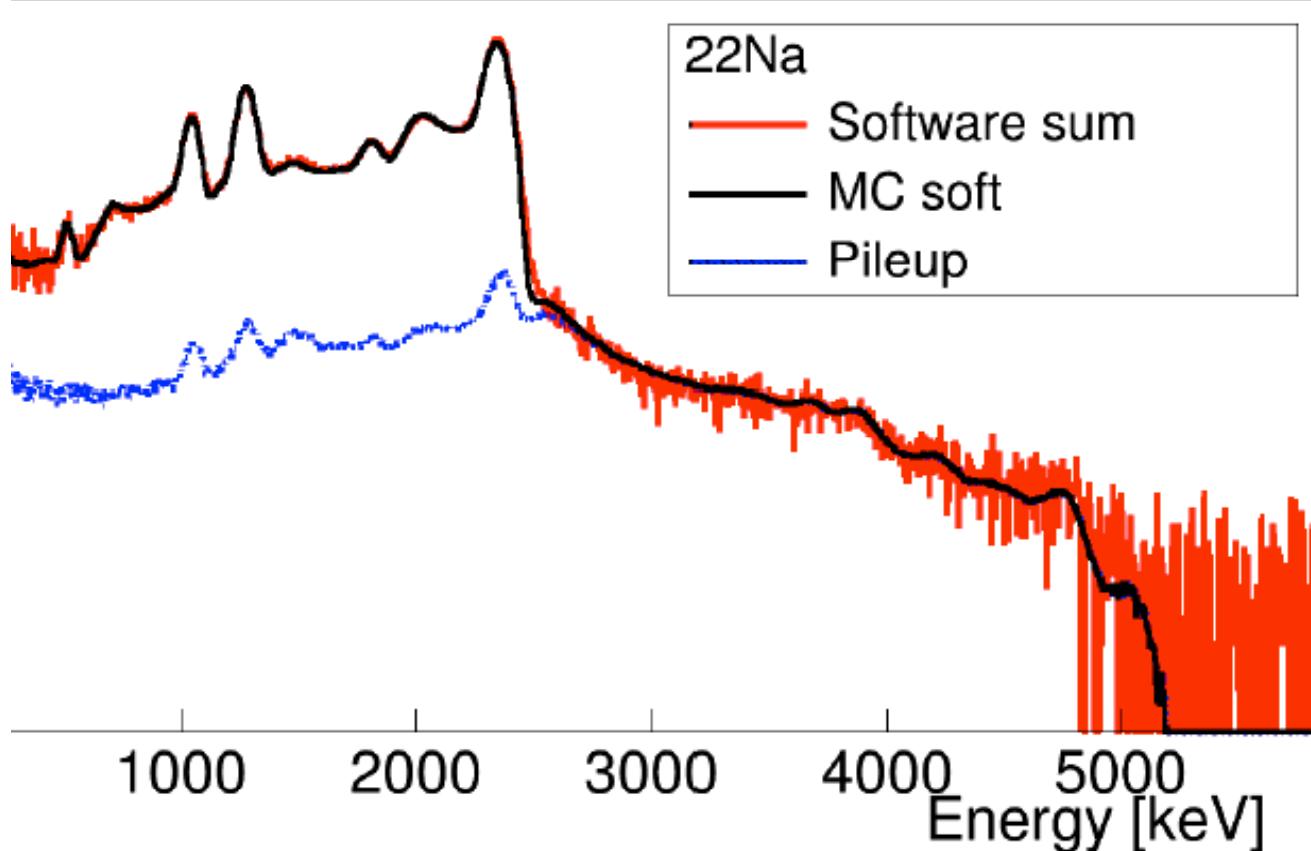
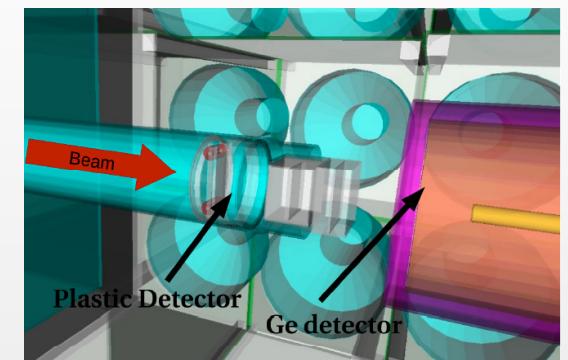


RENO results, arXiv:1504.08268

Also seen in DAYA BAY

Monte Carlo simulation of full detector response

- Geant 4.9.6.p03
- Inclusion of light yield non-proportionality function
- Detailed geometry description

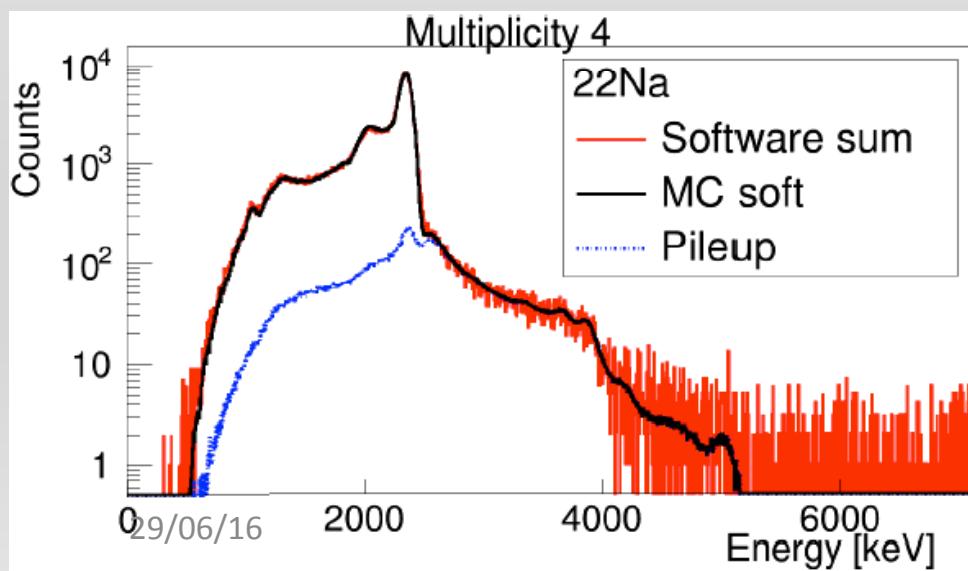
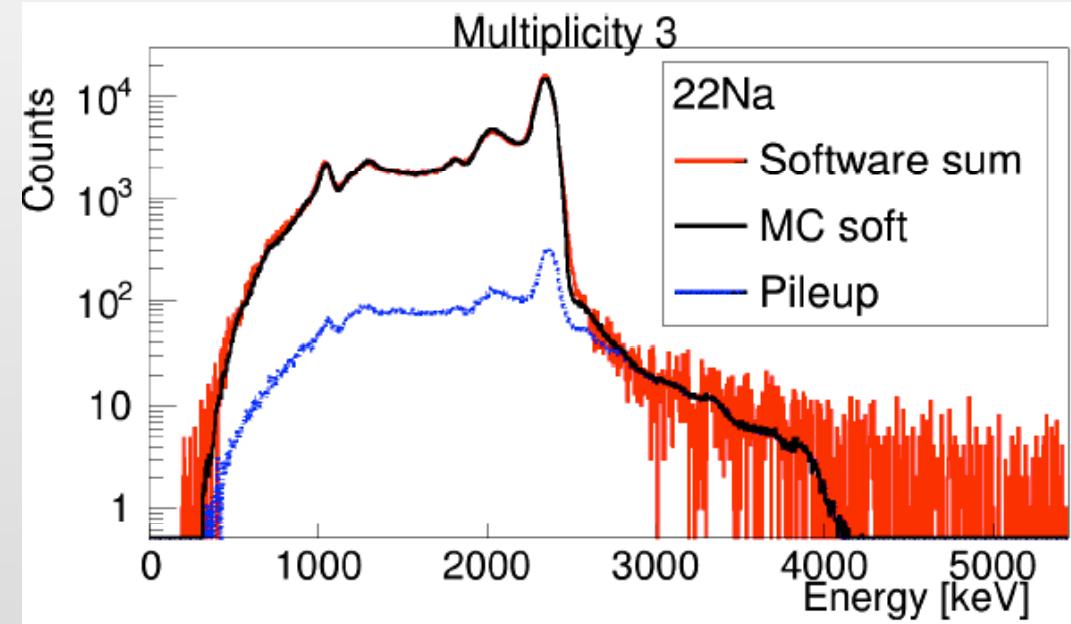
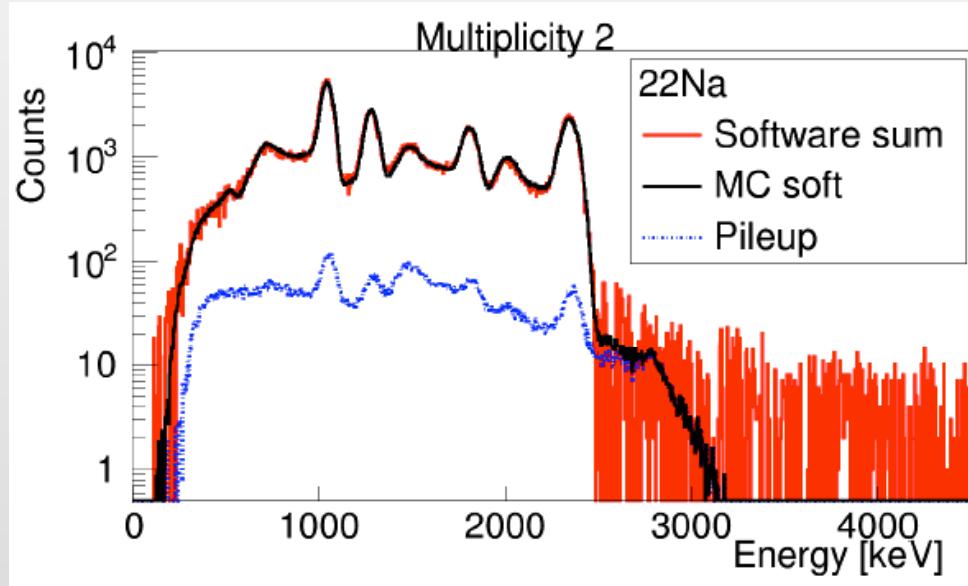


Summing-pileup calculation

- Monte Carlo method base recorded events or Geant4 generated events
- Uses the measured pulse shape

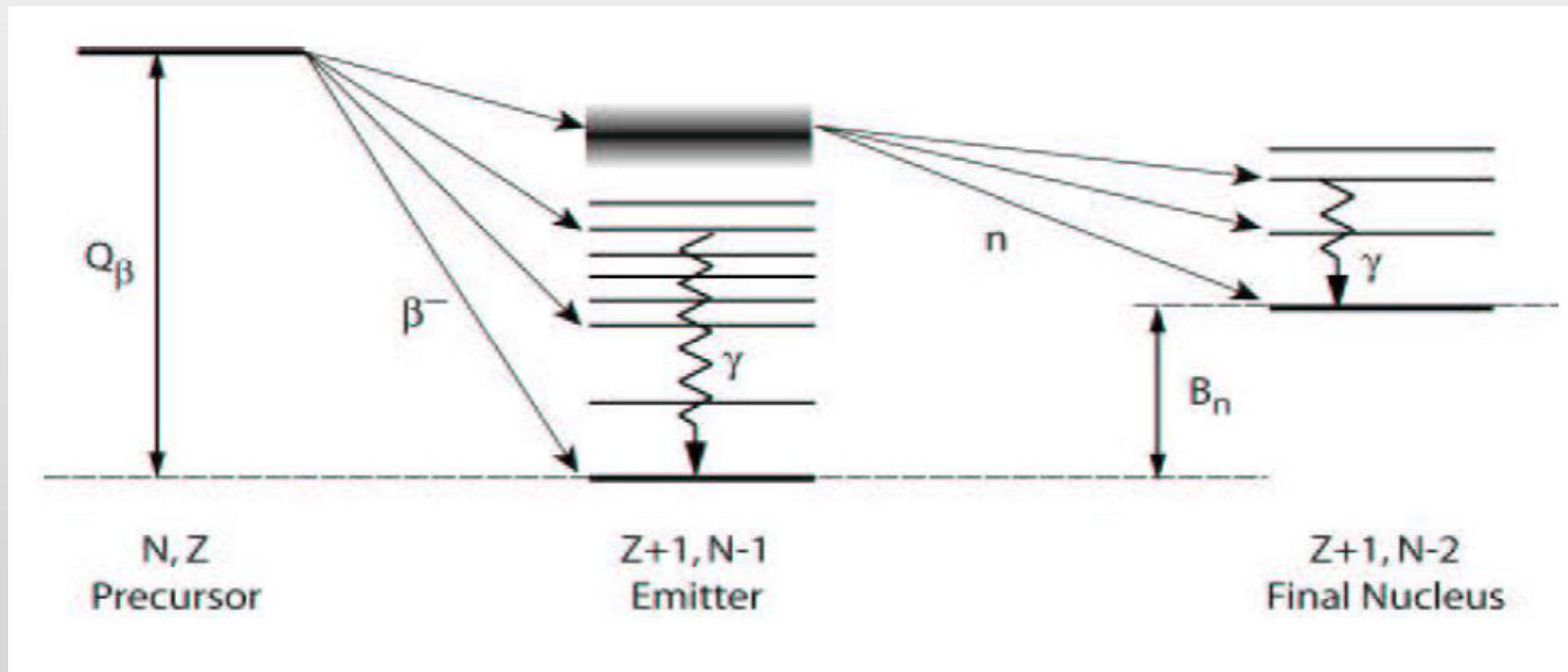
Monte Carlo simulation of crystal multiplicity response

- More stringent test of Monte Carlo simulation



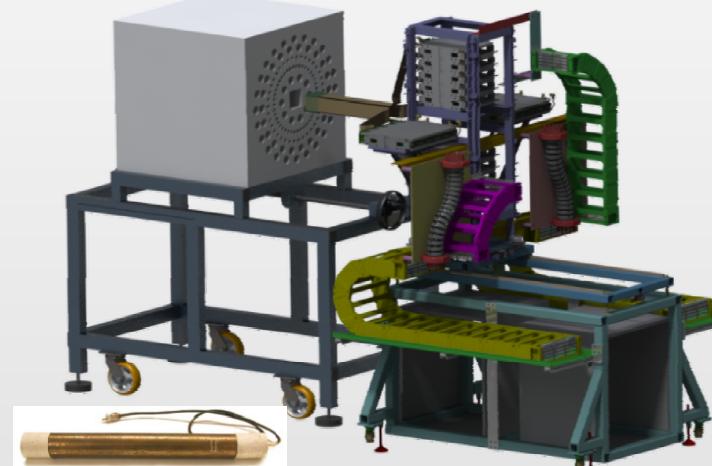
- The sensitivity to cascade gamma-ray multiplicity will add accuracy to the response calculation for deconvolution process, allowing adjustments to the unknown part of b.r.

One can also measure beta-delayed neutrons directly



BRIKEN Project: Beta delayed neutron measurements at RIKEN

→ Largest ${}^3\text{He}$ array ever built (182 tubos de ${}^3\text{He}$)



${}^3\text{He}$ -Proportional Counter Mega-Array

➤ Astrophysics, nuclear structure, reactor technology

✓ 1st BRIKEN Workshop, Valencia 17-18/12/2012

<http://indico.ific.uv.es/indico/event/briken>

✓ 2nd BRIKEN Workshop, Tokyo 30-31/07/2013

<http://indico.ific.uv.es/indico/event/briken2>

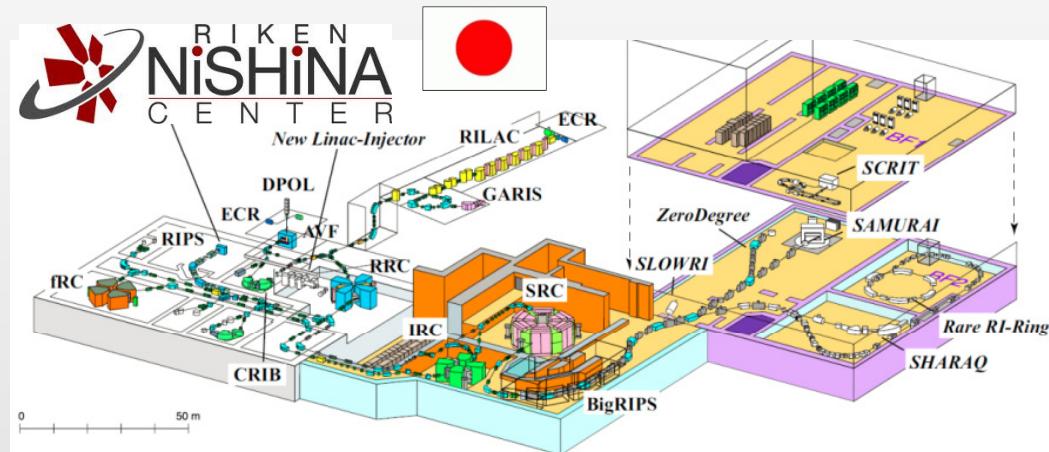
✓ 3rd BRIKEN Workshop, Valencia 22-24/07/2015

<http://indico.ific.uv.es/indico/event/briken3>

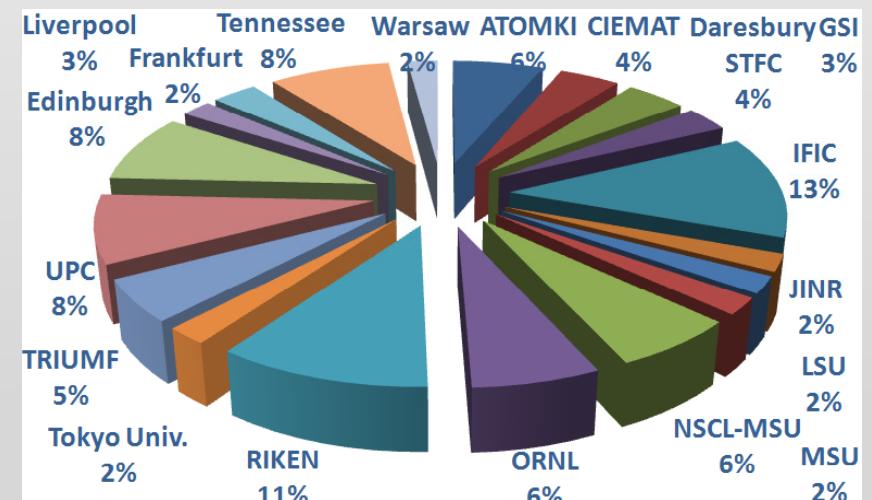
✓ Construction Proposal

✓ 3 Exp. Proposals approved 2014-2015

✓ 23 days of beam-time approved at RIKEN / BigRIPS!



60 scientists from 24 institutions



BRIKEN Project: Beta delayed neutron measurements at RIKEN

arXiv.org > physics > arXiv:1606.05544

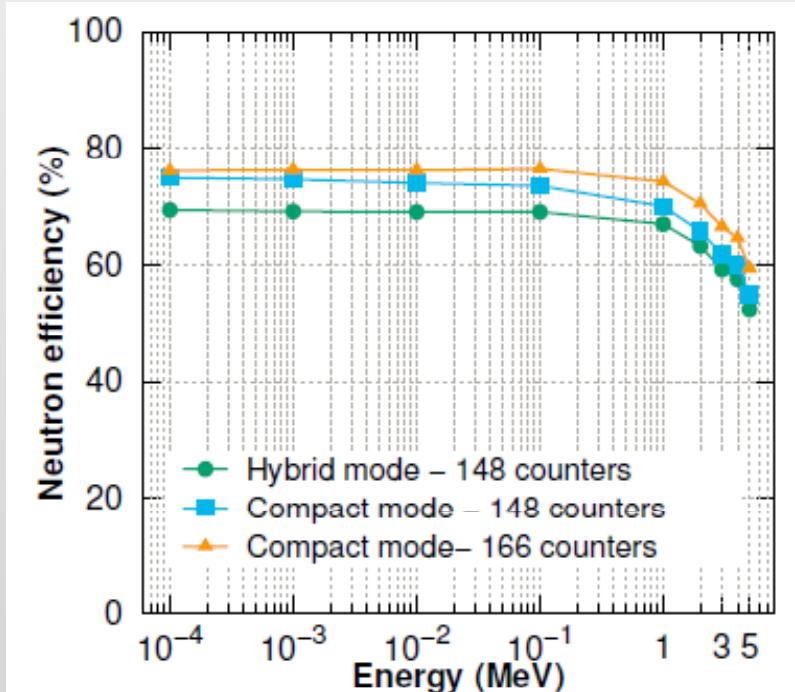
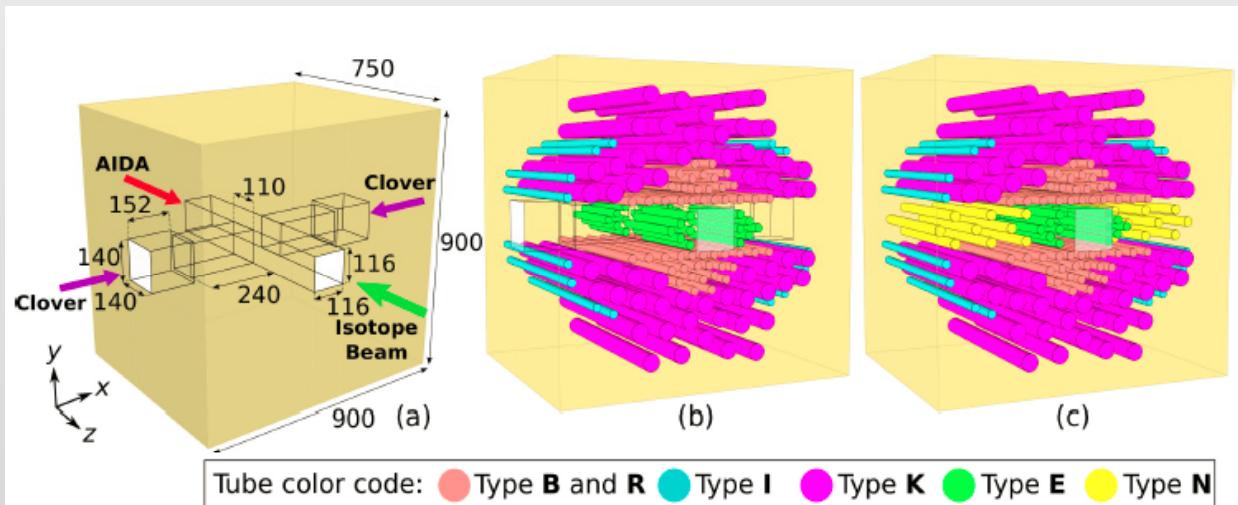
Search or Article

Physics > Instrumentation and Detectors

Conceptual design of the BRIKEN detector: A hybrid neutron-gamma detection system for nuclear physics at the RIB facility of RIKEN

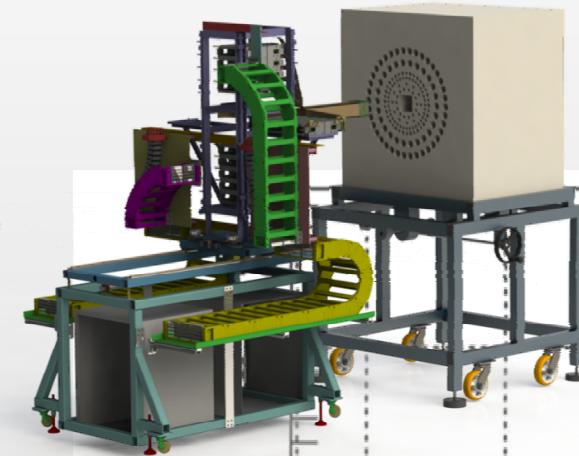
A. Tarifeño-Saldivia, J. L. Tain, C. Domingo-Pardo, F. Calviño, G. Cortes, V. H. Phong, A. Riego, The BRIKEN collaboration

<http://arxiv.org/abs/1606.05544>



The Spanish 3He tubes belong to the UPC

BRIKEN Project: Beta delayed neutron measurements at RIKEN



GSI:S410 "Measurement of β -delayed neutrons around the 3rd r-process peak" C. Domingo-Pardo et al., Sept. 2011

Measurement of β -delayed neutron emission probabilities around the doubly magic ^{78}Ni , K. Ryckaczewski, J.L. Tain, R. Griwacz, I. Dillmann (**BRIKEN** Col.)

Proton number Z

70

60

50

40

30

20

10

GSI: S323 "Beta-decay of very neutron-rich Rh, Pd, Ag nuclei including the r-process waiting point ^{128}Pd ". F. Montes et al. Sept. 2011

Decay properties of r-process nuclei in deformed region around A=100-125, S. Nishimura, A. Algora (**BRIKEN** Col.)

β -delayed neutron emission on the precursors of the REP, C. Domingo-Pardo (**BRIKEN** Col.)

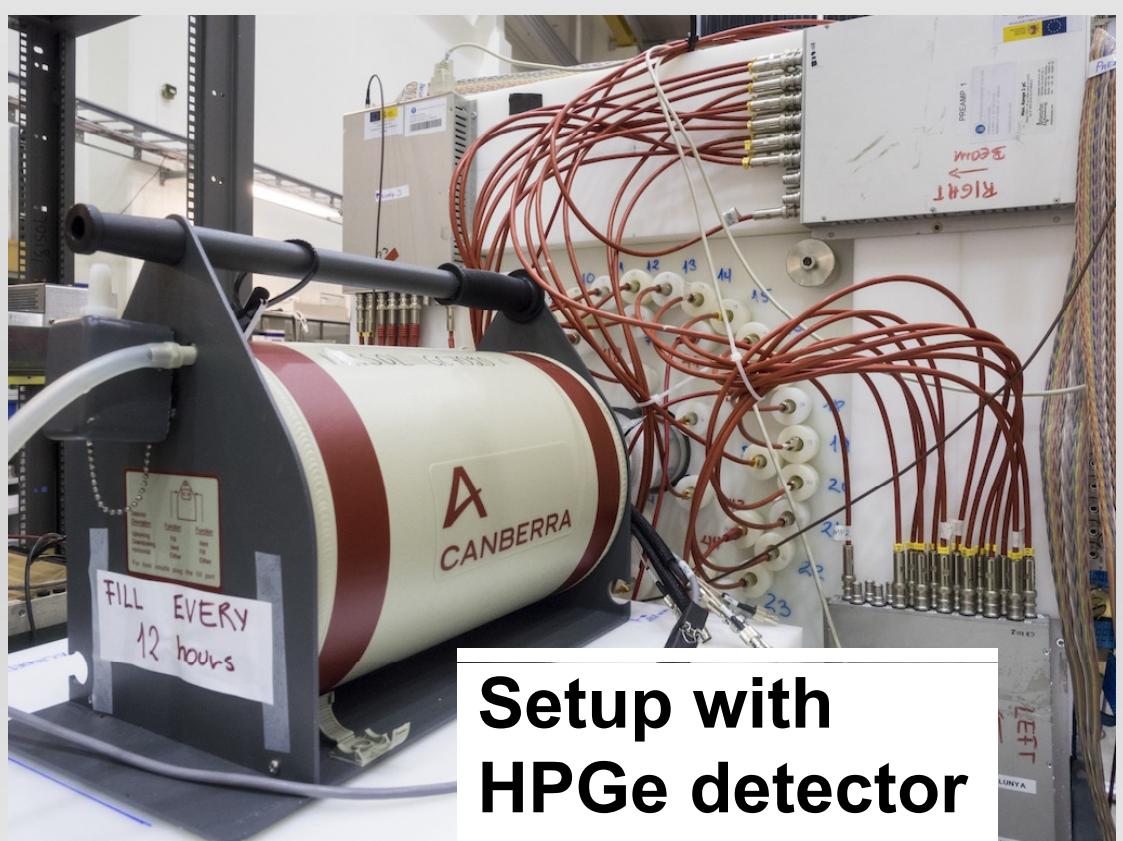
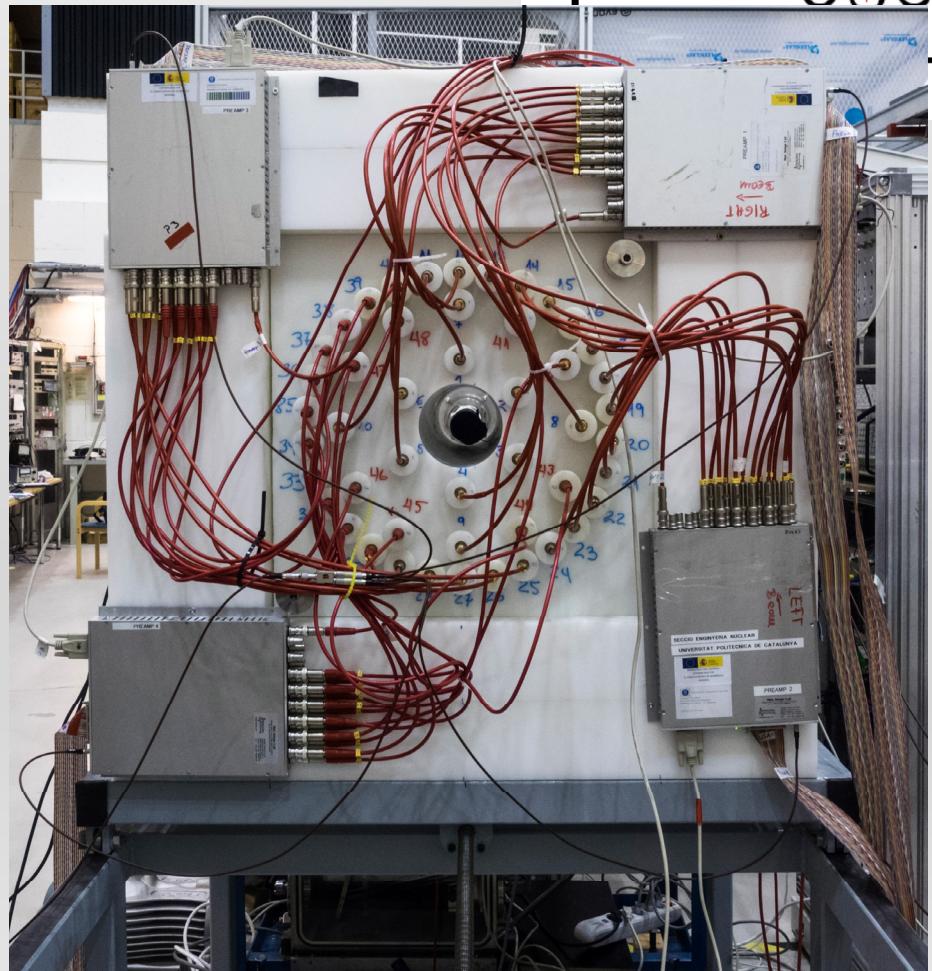
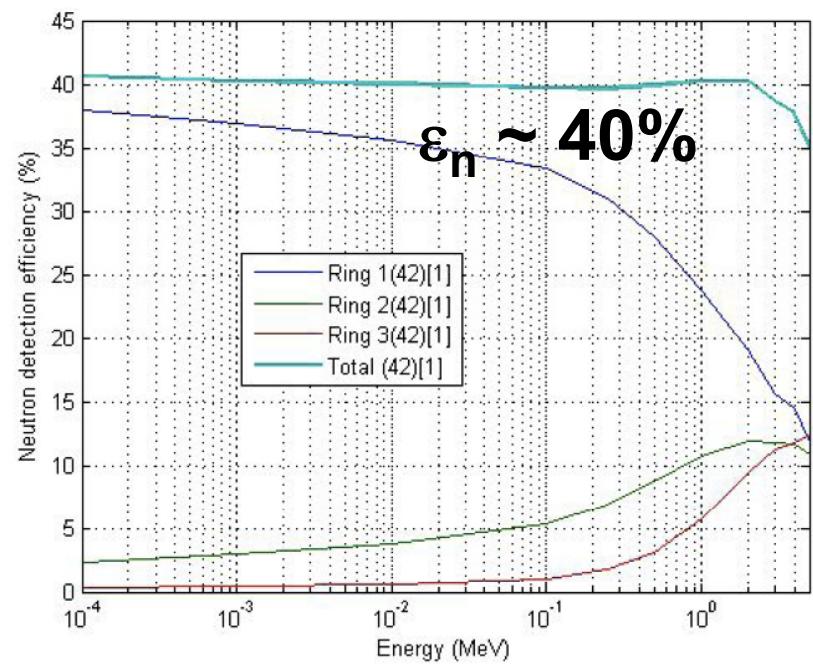
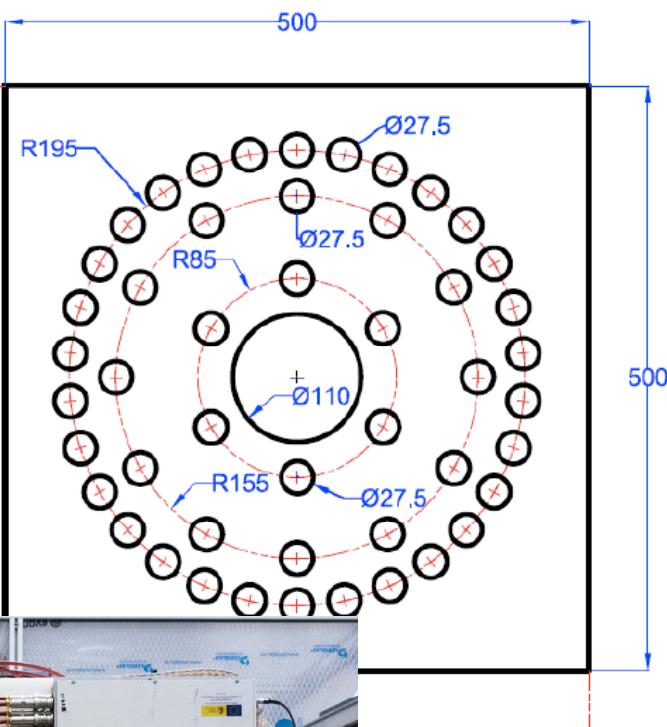
Measurement of β -delayed neutron emission probabilities relevant for the A=130 r-process abundance peak, A. Estrade, G. Lorusso, F. Montes (**BRIKEN** Col.)

JYFL: I162 "Delayed neutron measurements for advanced reactor technologies and astrophysics" JL Tain

Neutron number N

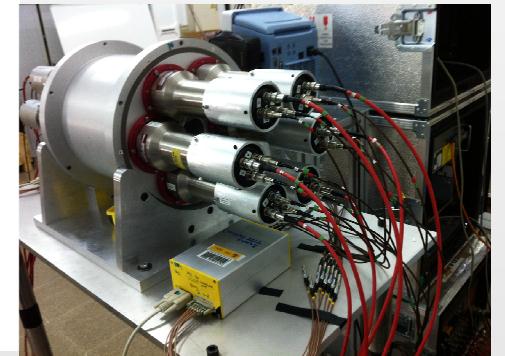
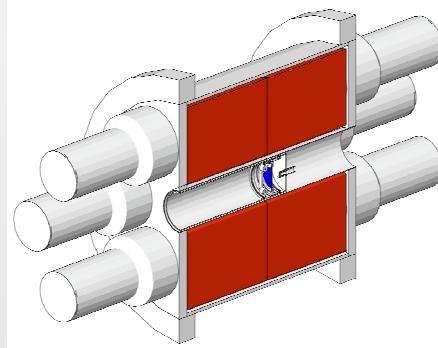
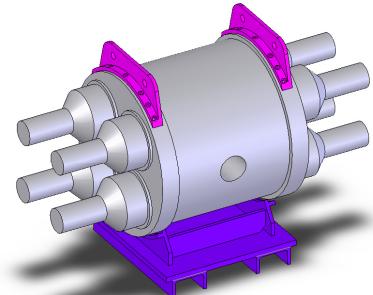
I162: BELEN-48a

48x 3He tubes
@ 8 & 10 atm



**Setup with
HPGe detector**

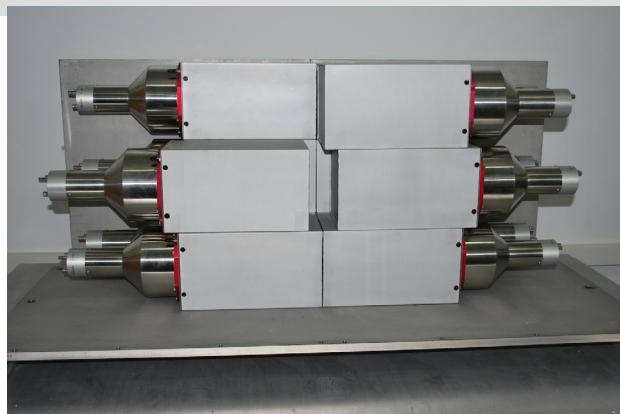
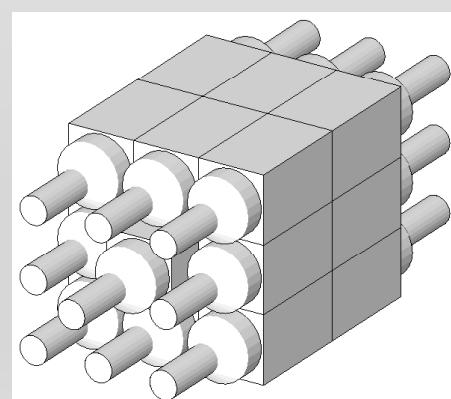
Future



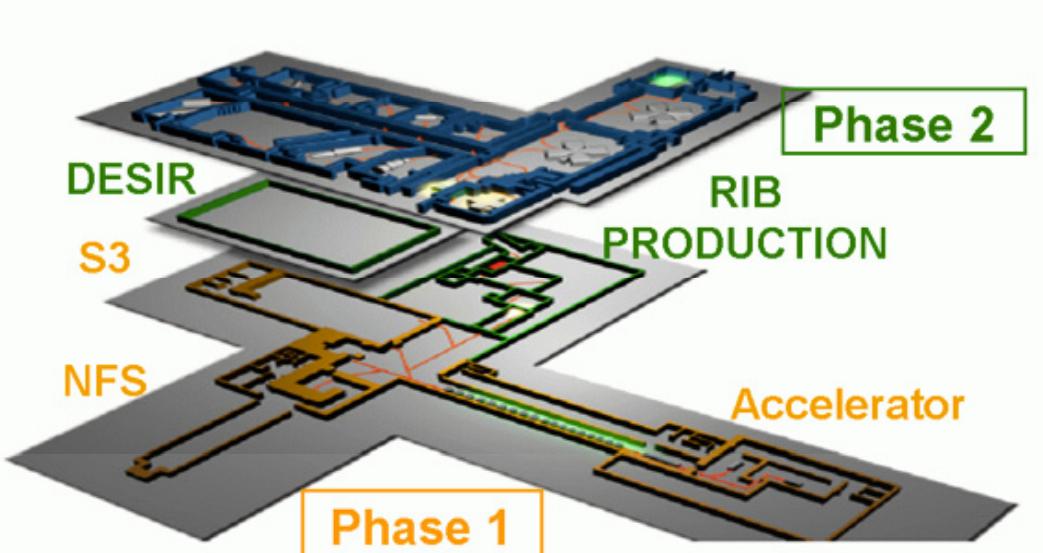
Rocinante, 12 BaF crystals

In a common canning, moving

Lucrecia, fix setup
At Isolde, on-going
programme



DTAS, 18 NaI crystals, modular
Moving detector



TOTAL ABSORPTION SPECTROSCOPY FOR NUCLEAR STRUCTURE, NUCLEAR ASTROPHYSICS, NEUTRINO AND REACTOR PHYSICS

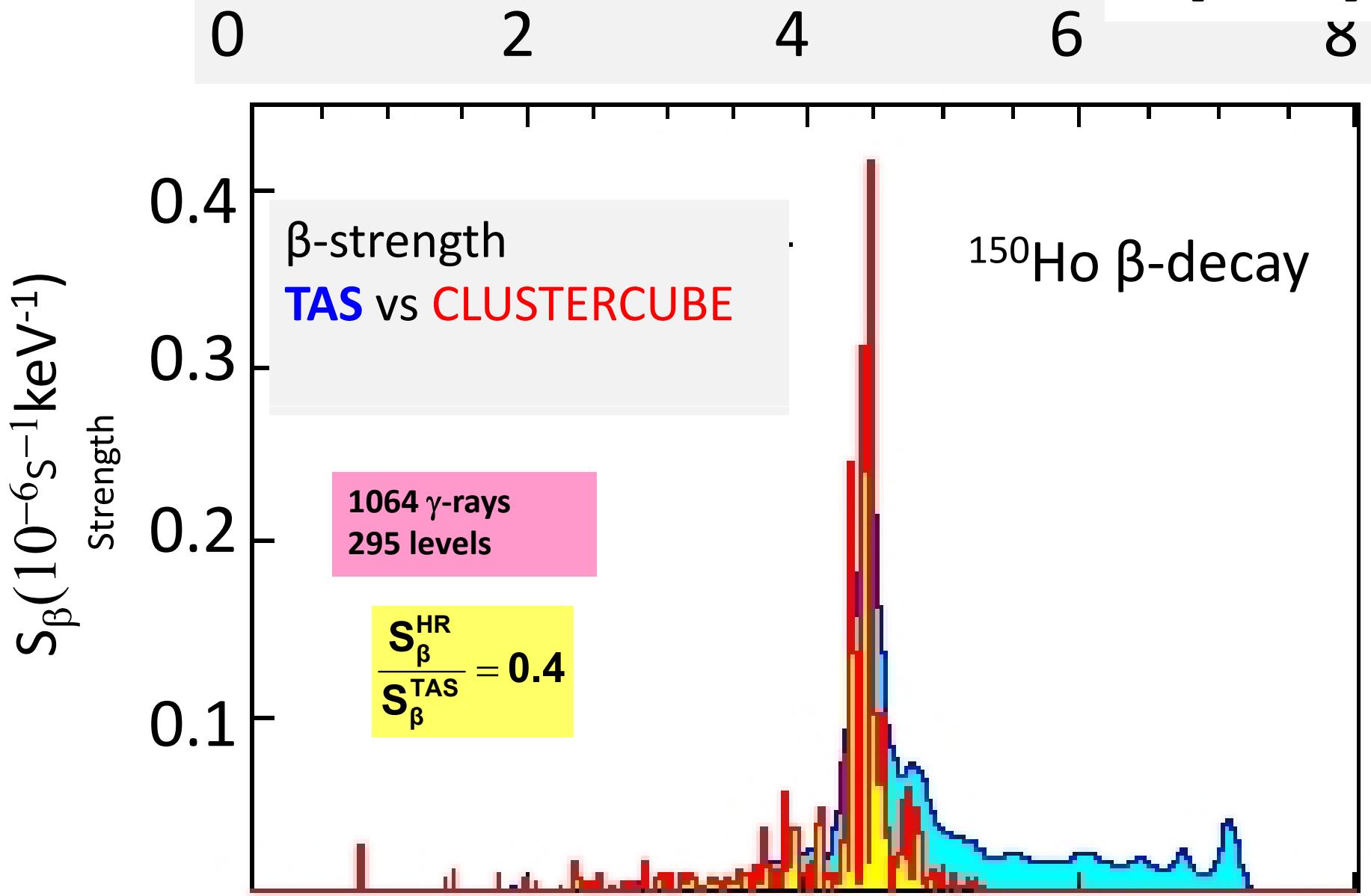
Spokespersons: A. Algora, M. Fallot, A. Porta, B. Rubio, J.-L. Tain



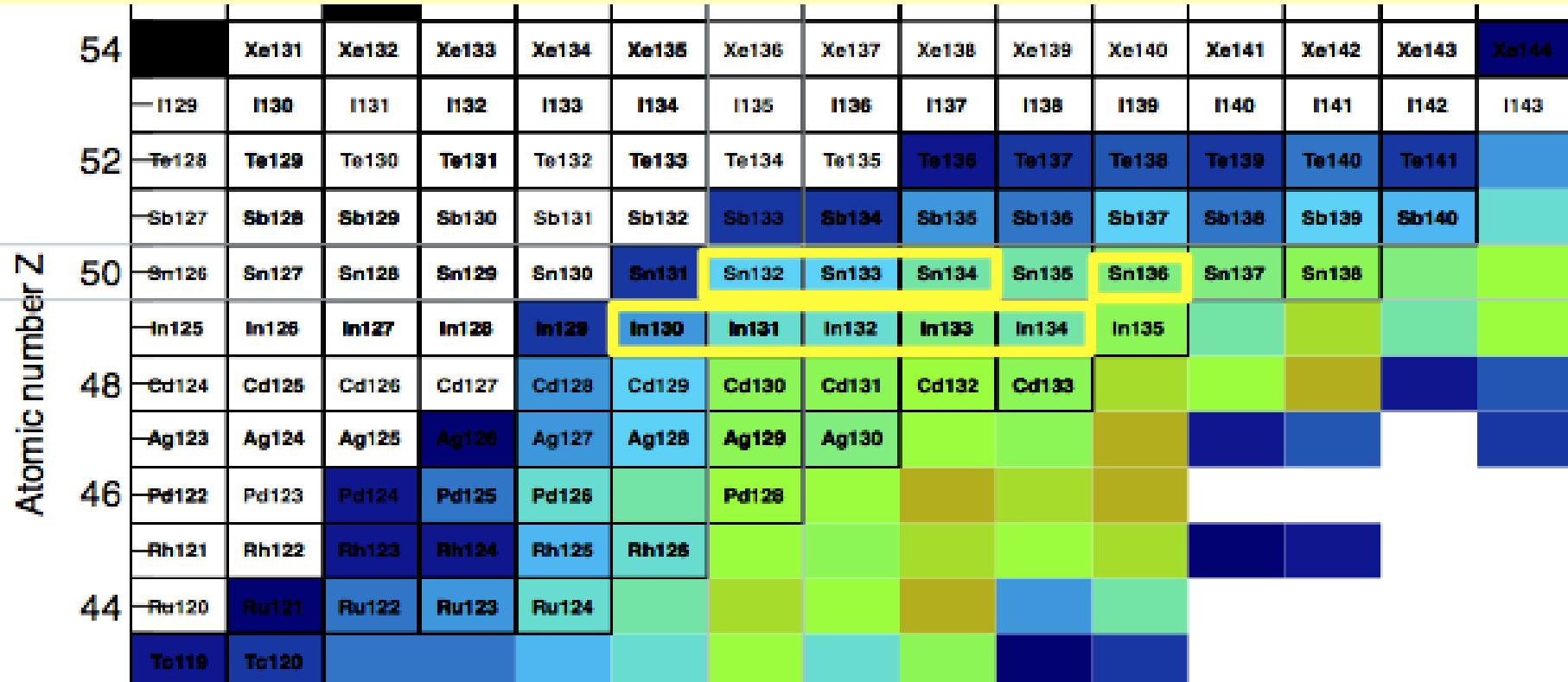


FIN

Mr. Miller estimated that demand for helium 3 was about 65,000 liters per year through 2013 and that total production by the only two countries that produce it in usable form, the United States and Russia, was only about 20,000 liters. In a letter to [President Obama](#), he called the shortage “a national crisis” and said the price had jumped to \$2,000 a liter from \$100 in the last few years, which threatens scientific research.



Proposal at ALTO (2014), its feasibility is based on ionization schemes for the desired species



TOTAL ABSORPTION SPECTROSCOPY FOR NUCLEAR STRUCTURE, NUCLEAR ASTROPHYSICS, NEUTRINO AND REACTOR PHYSICS

Spokespersons: A. Algora, M. Fallot, A. Porta, B. Rubio, J.-L. Tain

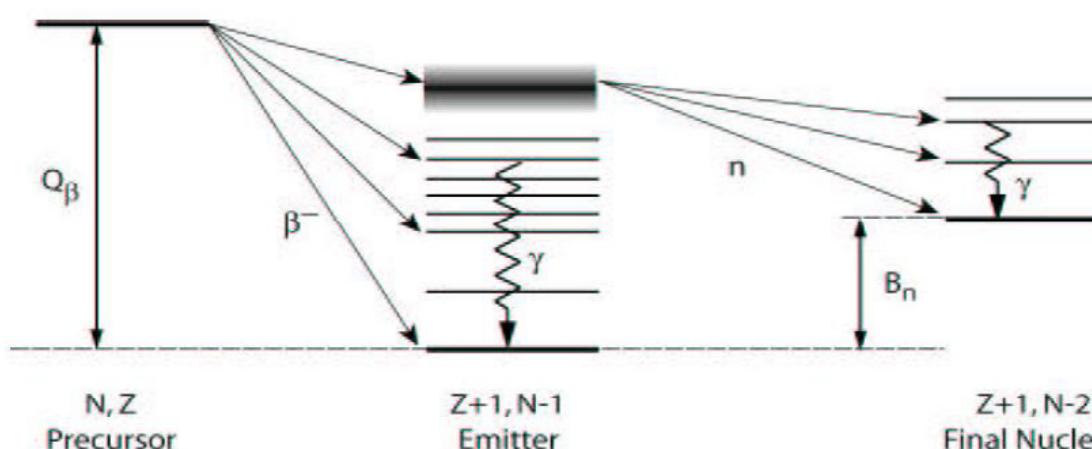
Motivation of recently analyzed cases: ^{87}Br , ^{88}Br



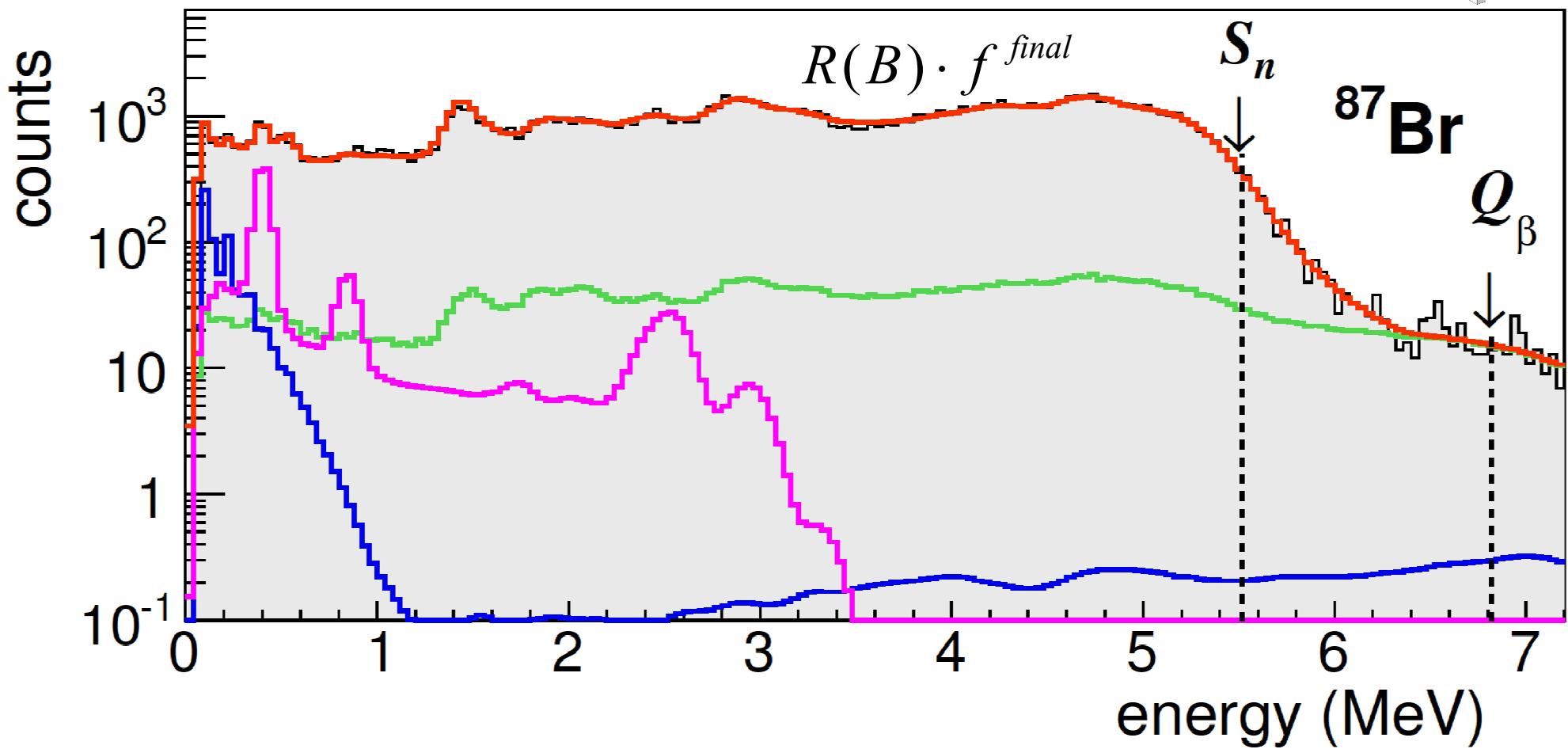
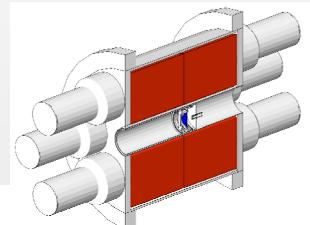
- Priority one in the IAEA list (decay heat)
- Moderate fission yields
- Pandemonium cases ?
- Interest from the structure point of view: vicinity of n closed shell
- Measurement of competition between gamma and neutron emission above the Sn value, which is conventionally neglected in astrophysical calculations

$$\frac{1}{T_{1/2}} = \int_0^{Q_\beta} S_\beta(E_x) \cdot f(Q_\beta - E_x) dE_x$$

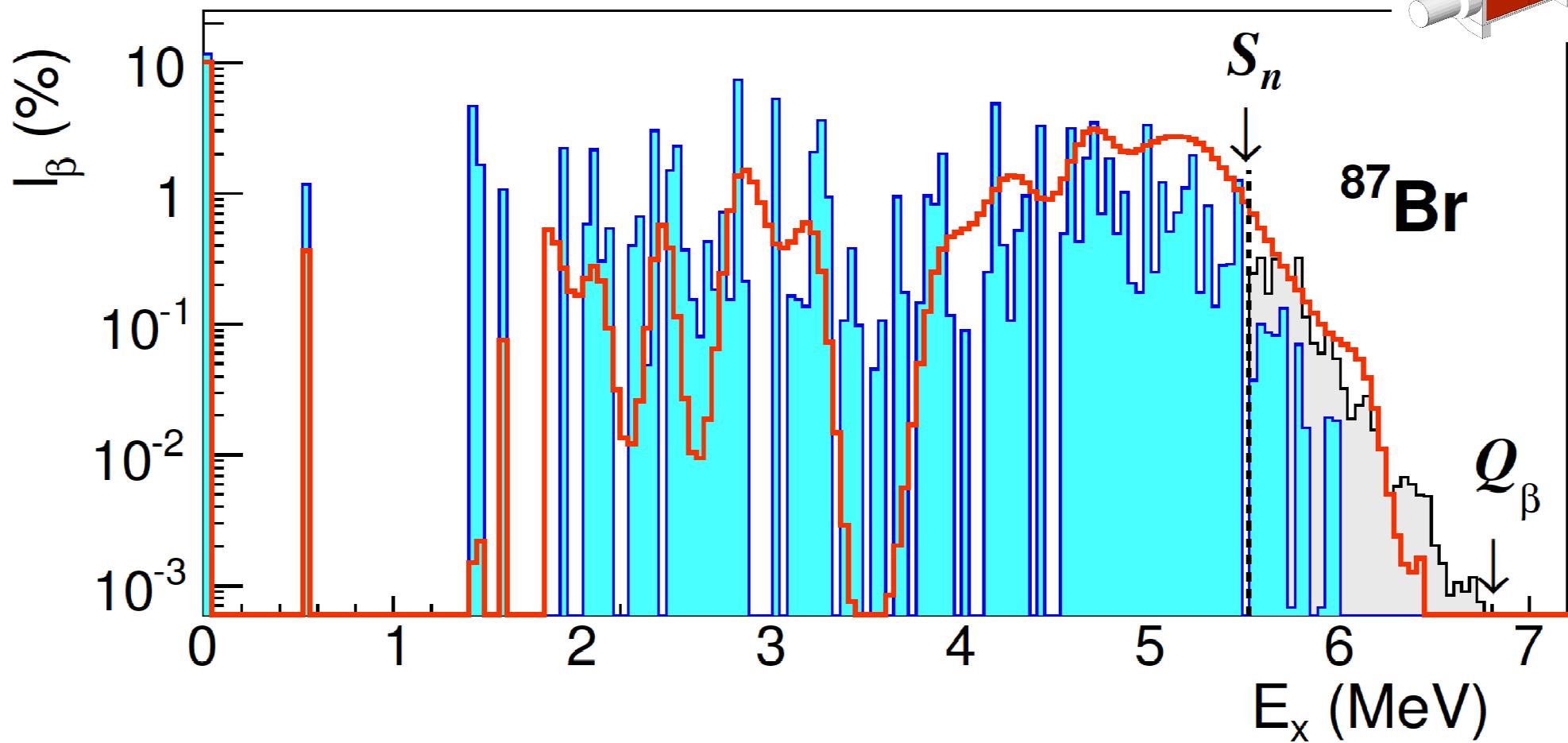
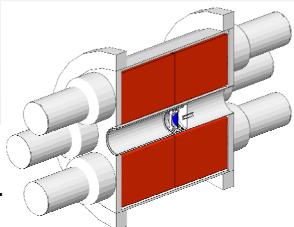
$$P_n = \frac{\int_{S_n}^{Q_\beta} S_\beta(E_x) \cdot f(Q_\beta - E_x) \cdot \frac{\Gamma^n}{\Gamma^n + \Gamma^\gamma} dE_x}{\int_0^{Q_\beta} S_\beta(E_x) \cdot f(Q_\beta - E_x) dE_x}$$



Beta delayed neutron emitters, example: ^{87}Br

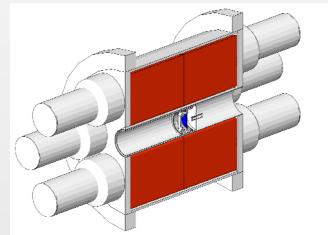


Beta delayed neutron emitters, example: ^{88}Br



Impact of the studied (bdn) cases with strong astrophysical implications

E. Valencia, JL Tain, A. Algora et al, in preparation



Isotope	\bar{E}_γ (keV)		\bar{E}_β (keV)	
	ENSDF	TAGS	ENSDF	TAGS
^{87}Br	3009	3938^{+40}_{-67}	1599	1159^{+32}_{-19}
^{88}Br	2892	4609^{+78}_{-67}	2491	1665^{+32}_{-38}
^{94}Rb	1729	4063^{+62}_{-66}	2019	2329^{+32}_{-30}

Isotope	\bar{E}_β (keV)	
	This work	Ref. [58]
^{87}Br	1170^{+32}_{-19}	1410 ± 10
^{88}Br	1706^{+32}_{-38}	1680 ± 10
^{94}Rb	2450^{+32}_{-30}	2830 ± 70

Isotope	P_γ (%)	P_n (%)
^{87}Br	3.50^{+49}_{-40}	$2.60(4)$
^{88}Br	1.59^{+27}_{-22}	$6.4(6)$
^{94}Rb	0.53^{+33}_{-22}	$10.18(24)$

Astrophysics implications discussed in
Tain et al. PRL 115, 062502
Can provide constraints for (n,γ) cross
section far from stability (r-process)

Shielding
Polyethylene
Lead
Copper
Aluminium

Factor 6

