

# Shell-model applications to gamma-ray strength function and level density

Yutaka Utsuno

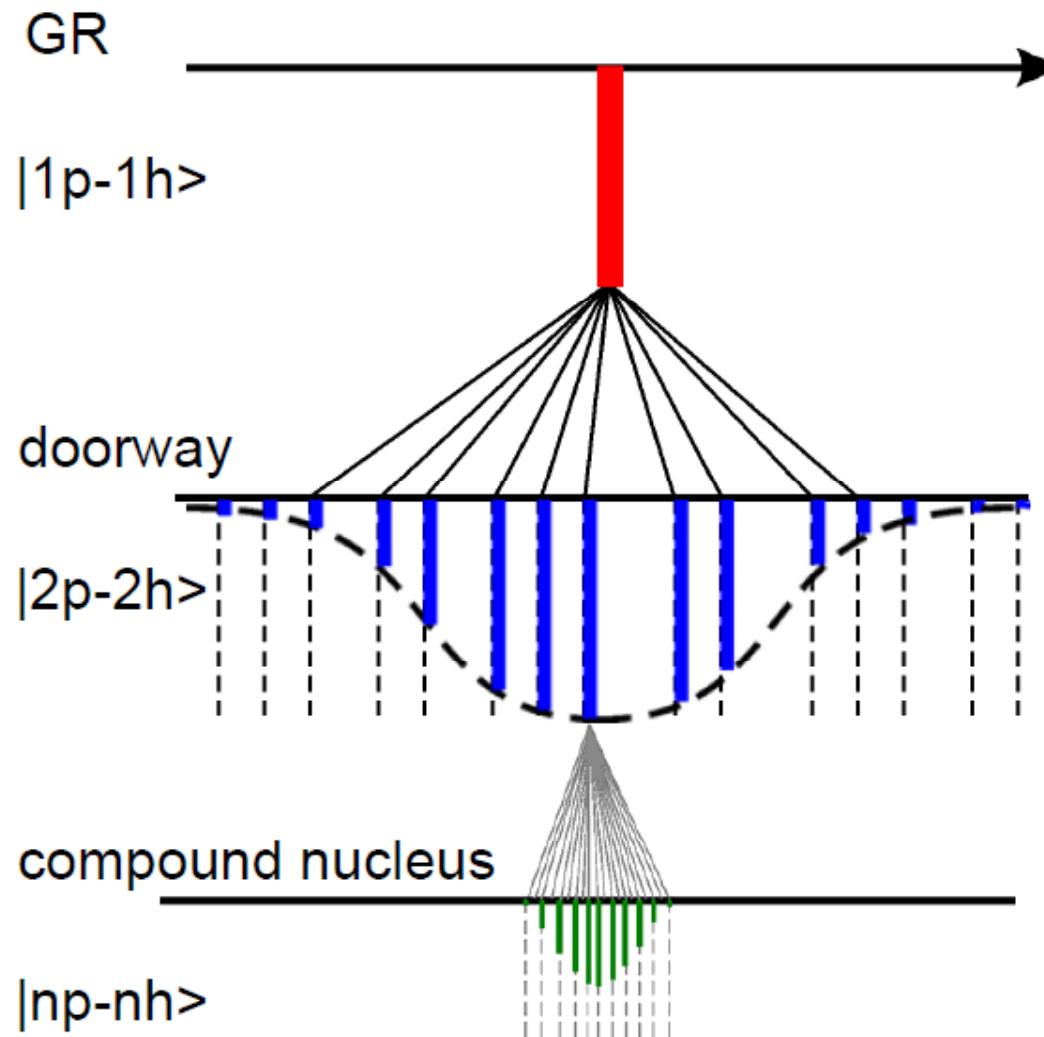
*Advanced Science Research Center, Japan Atomic Energy Agency  
Center for Nuclear Study, University of Tokyo*

## Collaborators

N. Shimizu (CNS), T. Otsuka (Tokyo), M. Honma (Aizu),  
T. Mizusaki (Senshu), Y. Futamura (Tsukuba), T. Sakurai (Tsukuba)

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# Nuclear structure in the GDR region



- Close relation among
  - ❑ Fine structure
  - ❑ Width of GDR
  - ❑ Level density

- Shell model (CI)
  - All the levels in a given model space
  - Transitions between excited states

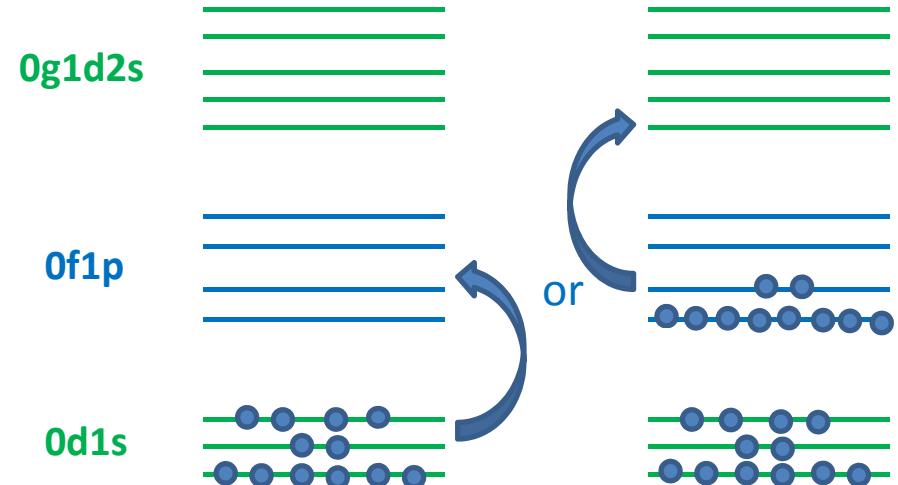
Taken from a figure by A. Richter

# Outline of the present talk

- Studying the gamma-ray strength function and level density on the same footing using large-scale shell-model calculations
  - 1. E1 strength function including transitions among excited levels
    - GDR of Ca isotopes:  $1\hbar\omega$  vs.  $(1+3)\hbar\omega$
    - Two-phonon state probed with decay probabilities
    - Pygmy dipole resonances
    - Examining the Brink hypothesis
  - 2. Level density
    - New method of calculating level density in the shell model
    - Application to  $^{58}\text{Ni}$ : equilibration of positive- and negative-parity levels

# Framework

- Objectives
  - *pf*-shell nuclei (e.g. Ca isotopes)
- Valence shell
  - Full *sd-pf-sdg* shell
  - $0\hbar\omega$  and  $1\hbar\omega$  states for the ground and  $1^-$  levels, respectively
- Effective interaction
  - same as the one used for  $3^-_1$  levels in Ca isotopes (Y. Utsuno et al., PTP Suppl., 2012)
  - USD (*sd*) + GXPF1B (*pf*) +  $V_{MU}$  (remaining)
    - Successful in *sd-pf* shell calculations including exotic nuclei (e.g.  $^{42}\text{Si}$ ,  $^{44}\text{S}$ )
  - $g_{9/2}$  SPE: fitted to  $9/2^+_1$  in  $^{51}\text{Ti}$ 
    - Reasonable agreement with a recent C<sup>2</sup>S data in  $^{49}\text{Ca}$  (A. Gade et al., 2016)
- Removal of spurious center-of-mass motion with Lawson method

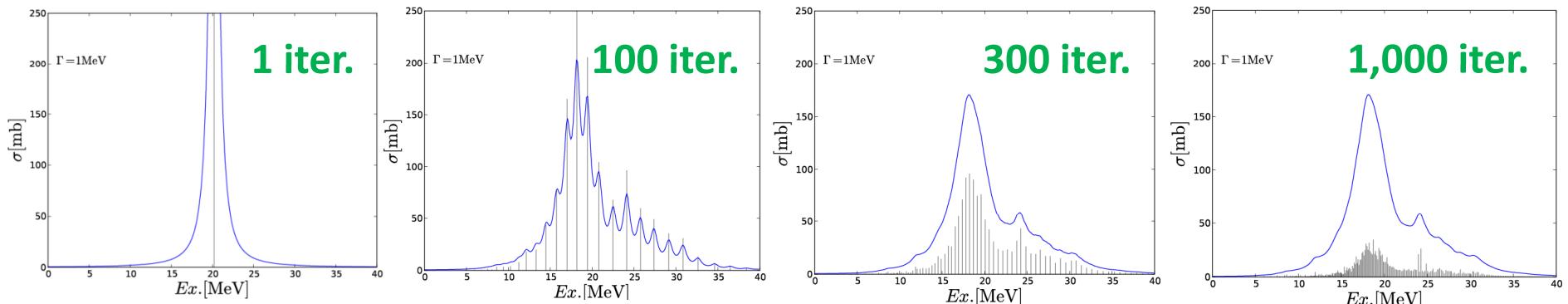


# Lanczos strength function method

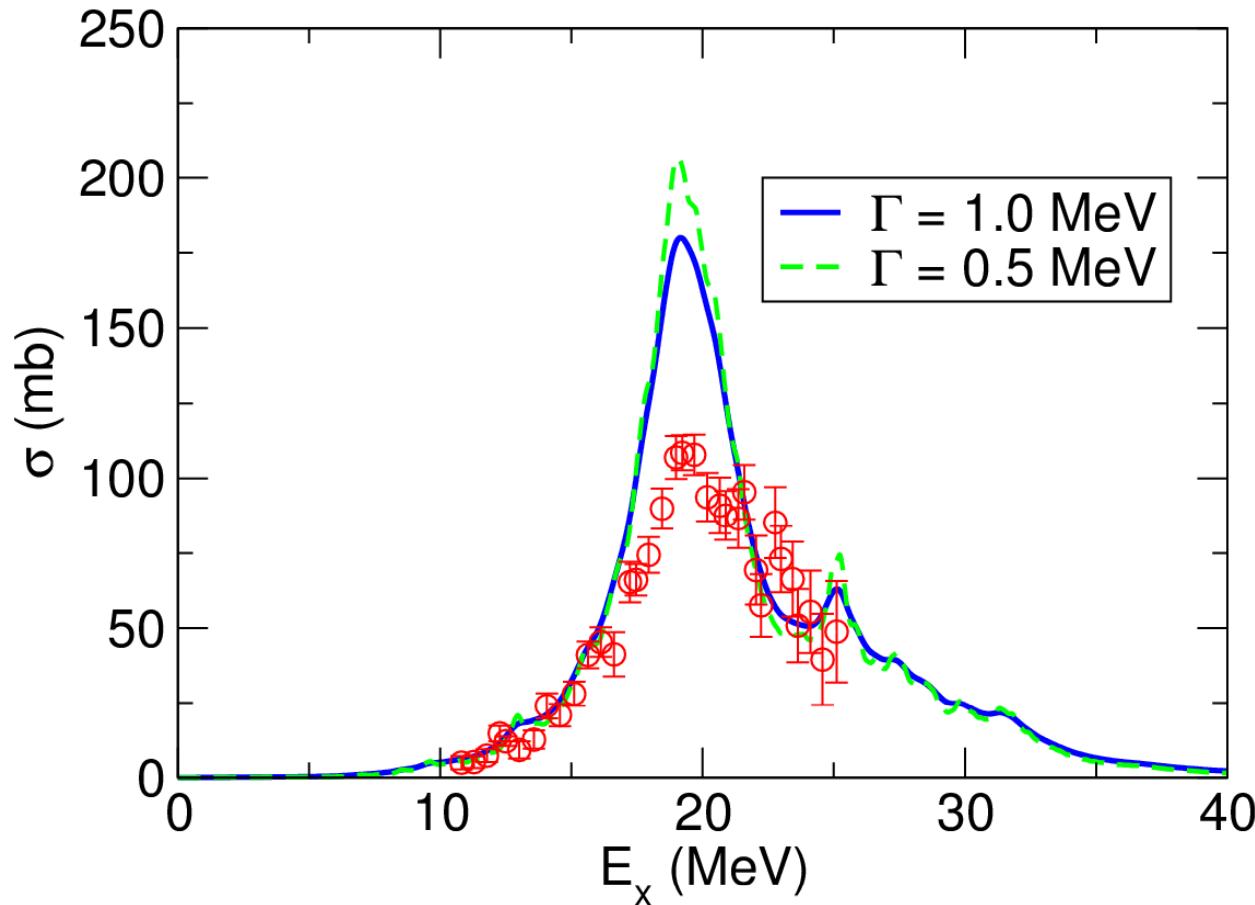
Efficient way to avoid calculating all the eigenstates

1. Take an initial vector:  $\vec{u}_1 = T(E1)|g.s.\rangle$
2. Follow the usual Lanczos iterations
  - $H\vec{u}_k = \beta_{k-1}\vec{u}_{k-1} + \alpha_k\vec{u}_k + \beta_k\vec{u}_{k+1}$  : defining a normalized vector  $\vec{u}_{k+1}$
3. Calculate the strength function  $\sum_\nu B(E1; g.s. \rightarrow \nu) \frac{1}{\pi} \frac{\Gamma/2}{(E-E_\nu+E_0)^2+(\Gamma/2)^2}$  by summing up all the eigenstates  $\nu$  in the Krylov subspace with an appropriate smoothing factor  $\Gamma$  until good convergence is achieved.

Example of convergence with  $\Gamma = 1$  MeV

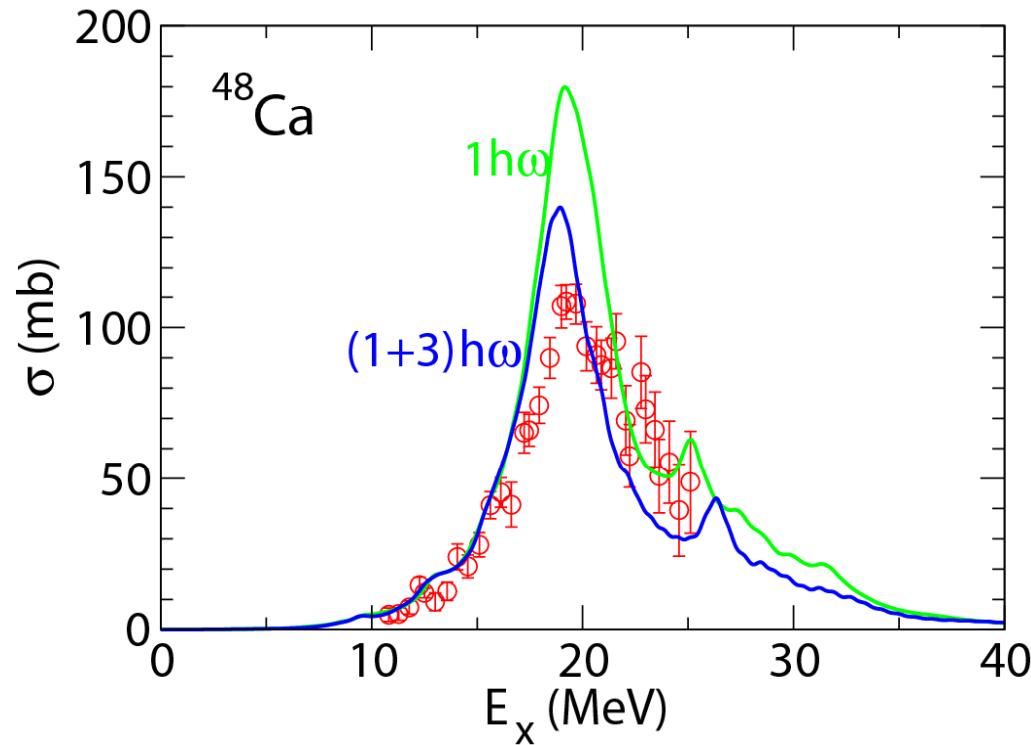


# Photo-absorption cross sections for $^{48}\text{Ca}$



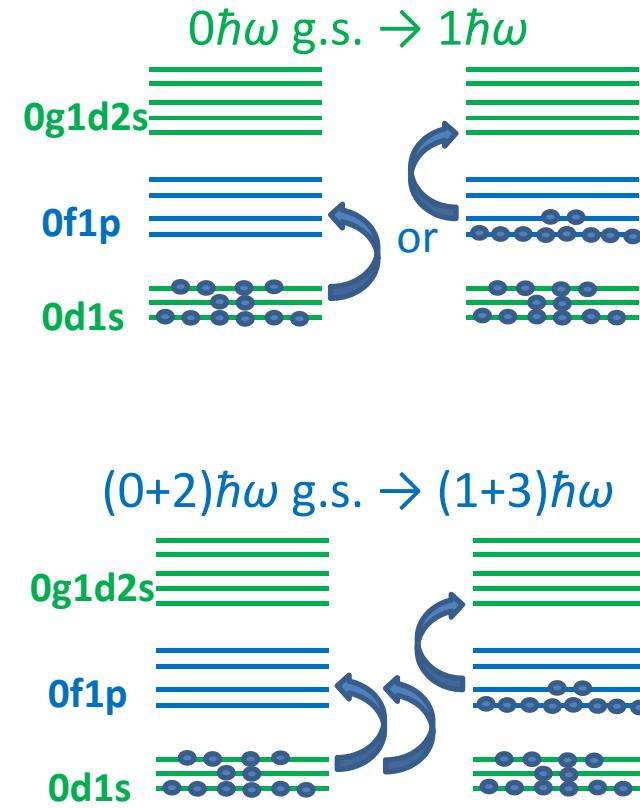
- GDR peak position: good
- GDR peak height: overestimated
- GDR tail: weak dependence on the choice of  $\Gamma$   $\rightarrow$  fine structure

# $E1$ strength function: $1\hbar\omega$ vs. $(1+3)\hbar\omega$



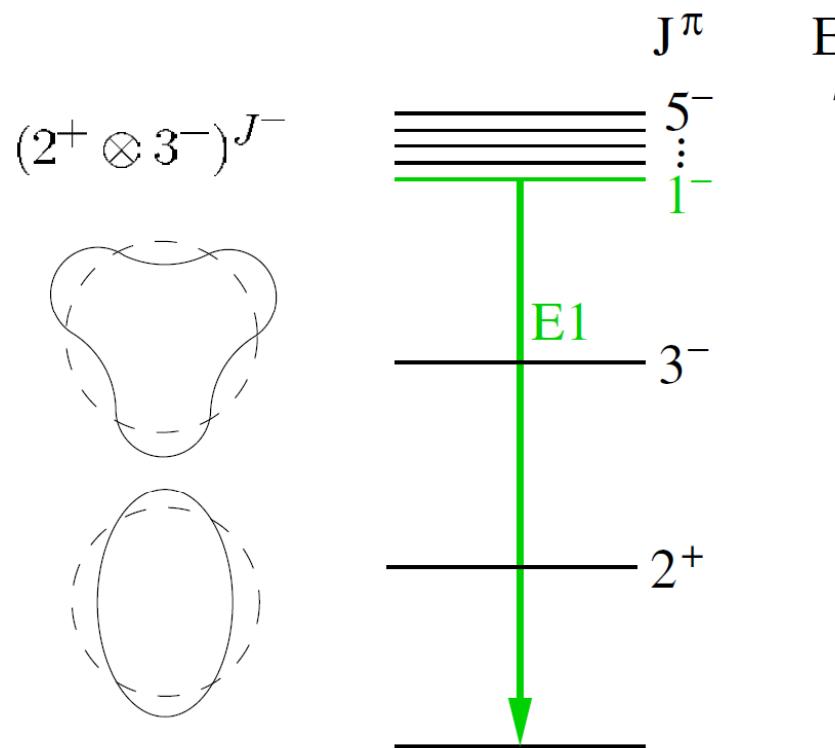
Expt.: G.J. O'Keefe et al., Nucl. Phys. A 469, 239 (1987)

- Improved GDR peak height
- Low-energy tail is almost unchanged.
  - The following results are restricted to  $1\hbar\omega$ , concentrated on the tail region.



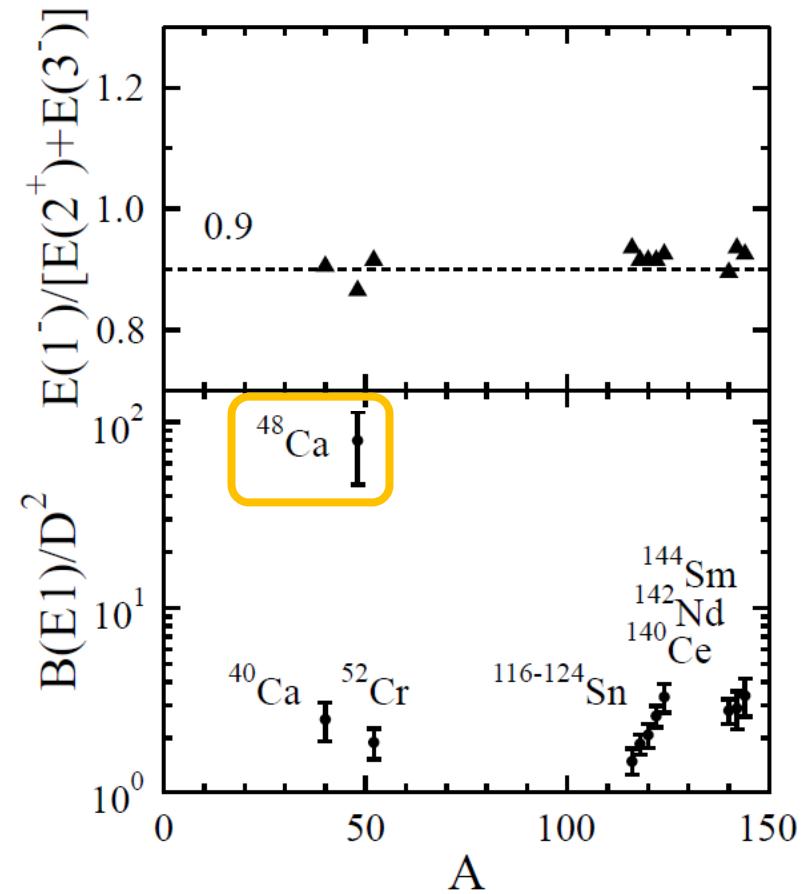
Up to  $10^{10}$   $m$ -scheme dimensions  
in Ca isotopes

# $2^+ \otimes 3^-$ two-phonon states



- Two-phonon states in  $^{48}\text{Ca}$ 
  - The proposed state has much stronger  $B(E1)$  than other isotopes.
  - No E2 or E3 measurement

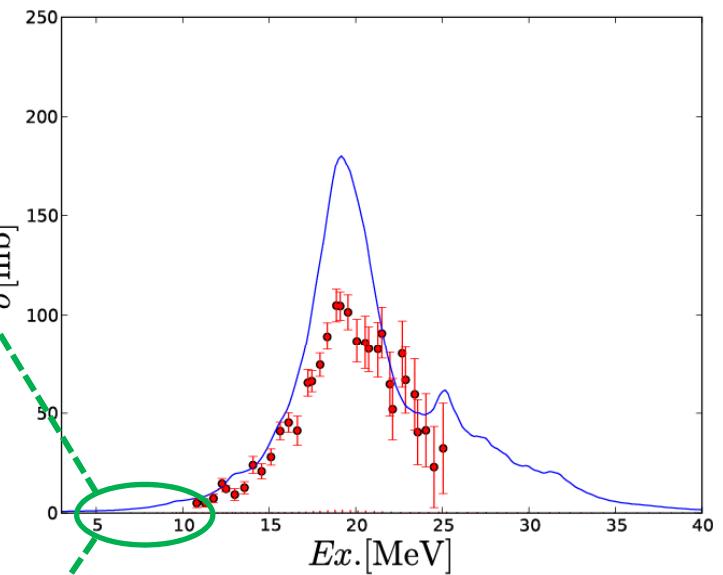
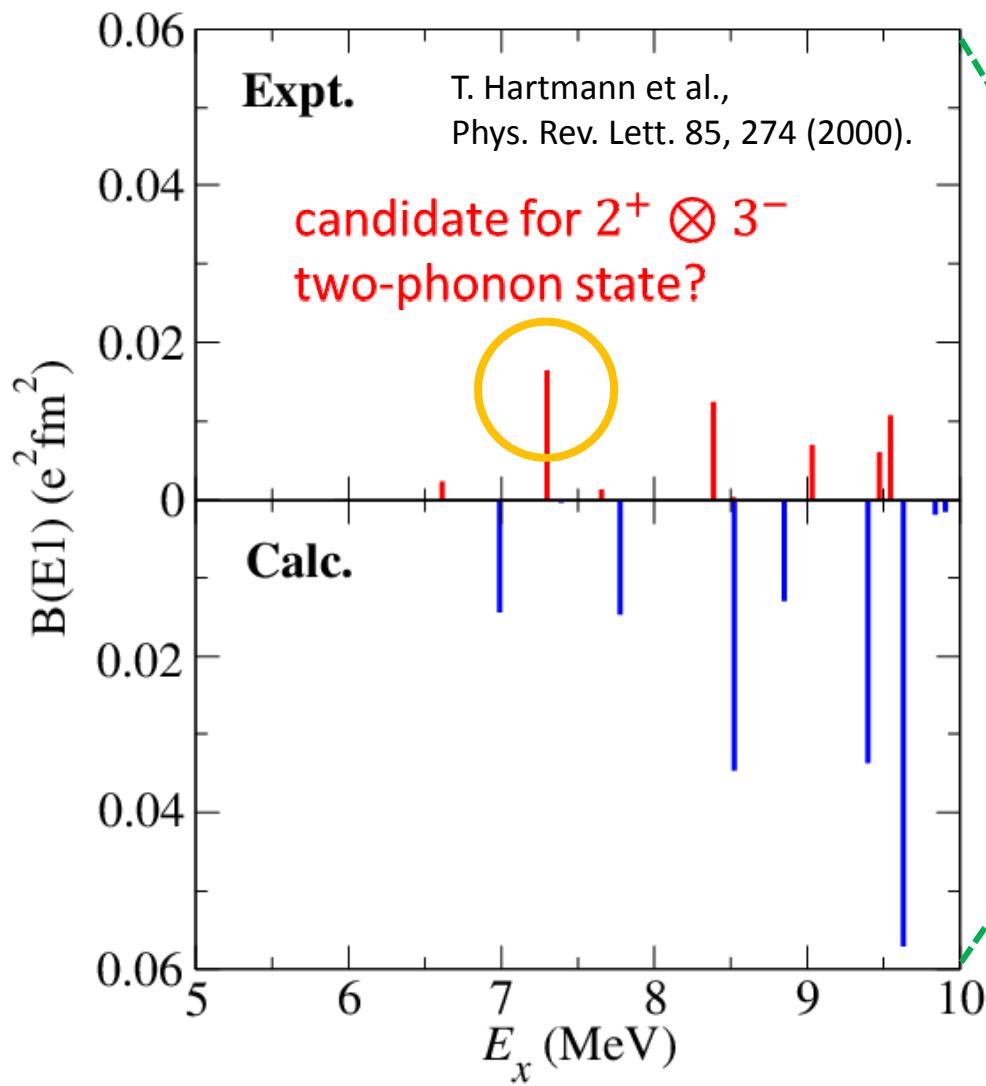
Systematics of two-phonon 1<sup>-</sup> states



$$D = 5.367 \times 10^{-4} (Z + N) Z \beta_2 \beta_3 \text{ (efm)}$$

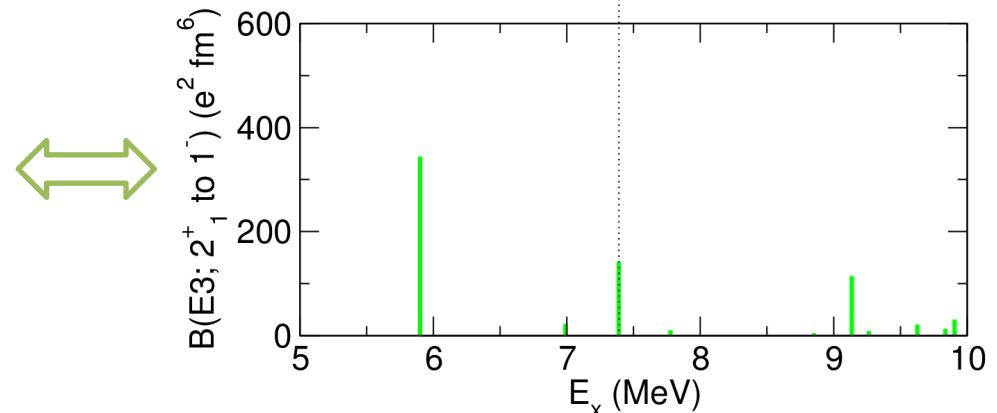
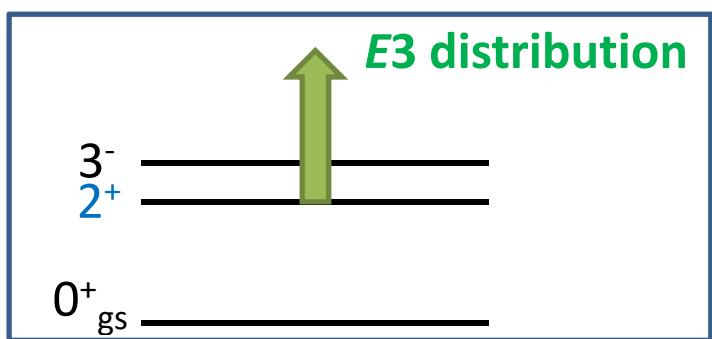
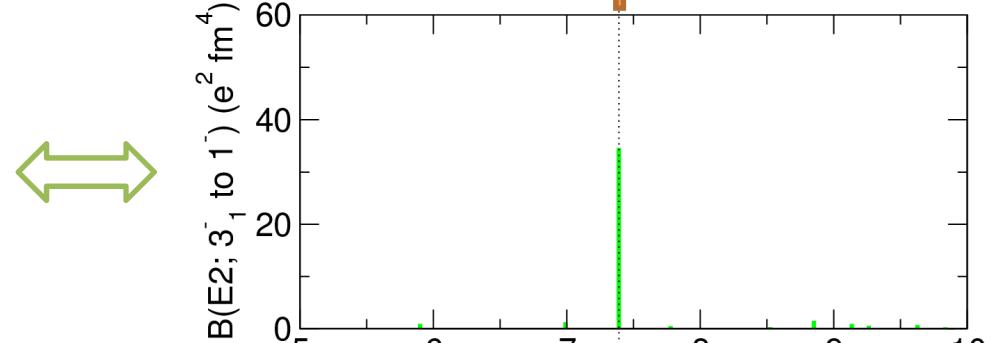
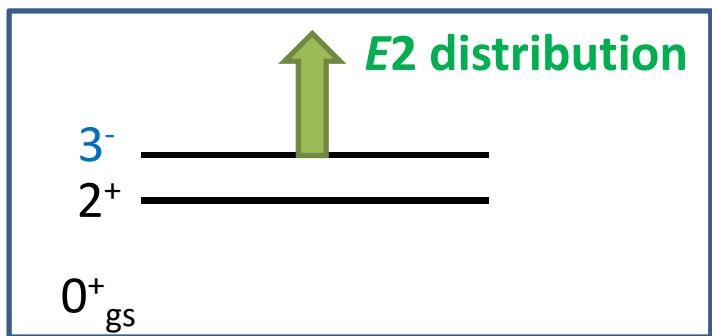
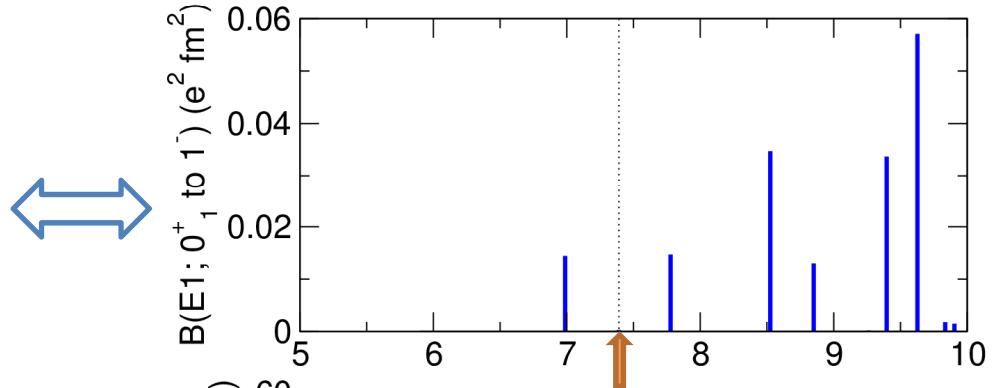
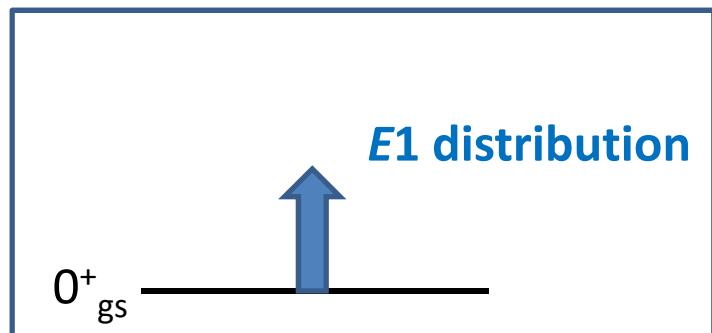
T. Hartmann et al., Phys. Rev. Lett. 85, 274 (2000).

# Which $1^-$ level is the two-phonon state?



- Discrete levels are calculated with the exact diagonalization.
- A few hundred keV upward shift of E1 strength gives excellent agreement.

# Probing $2^+ \otimes 3^-$ two-phonon character

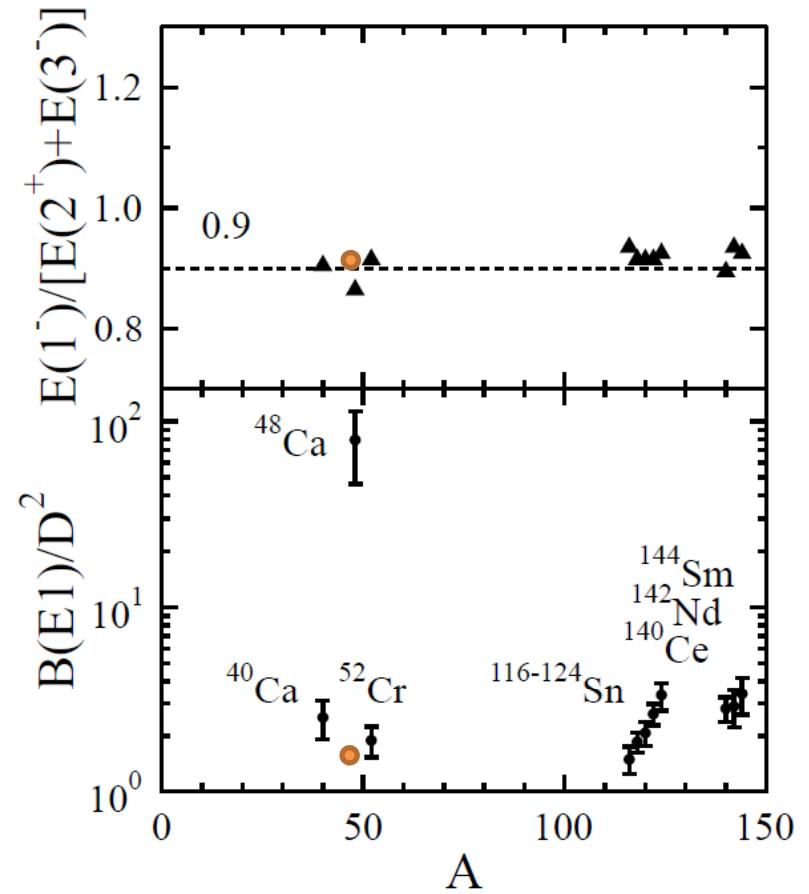
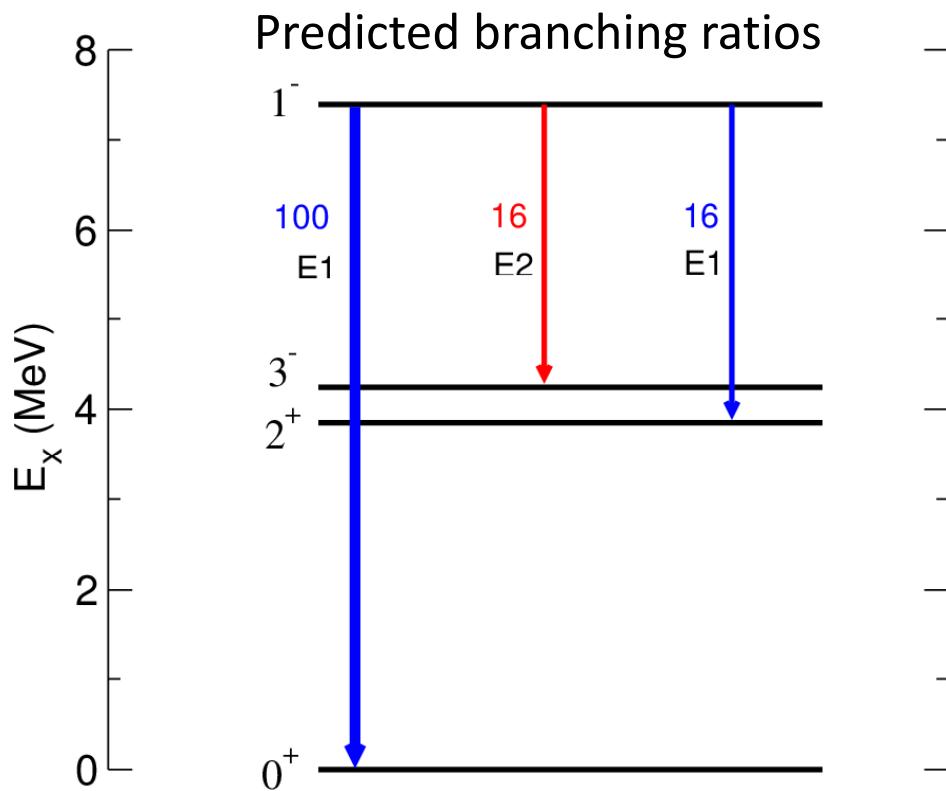


- Two-phonon state: the  $1^-$  level with a very small  $B(E1)$

# Properties of the $1^-_{2\text{ph}}$ state

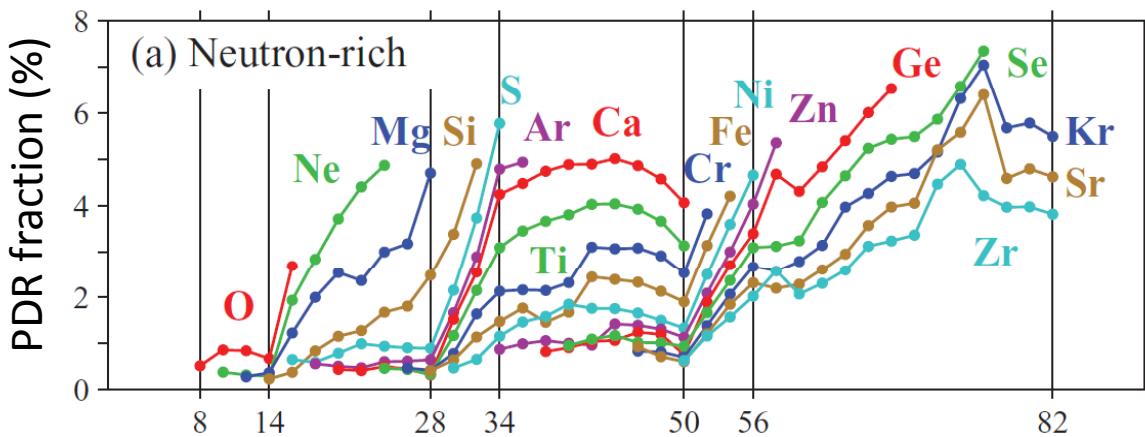
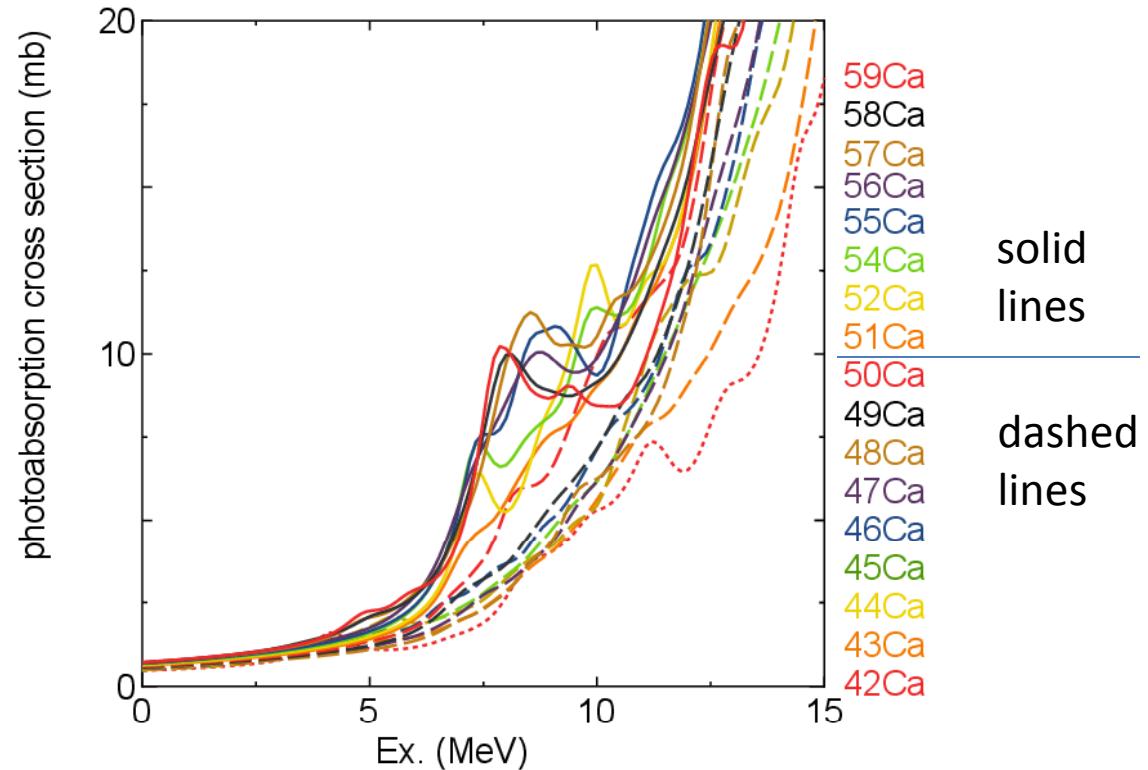
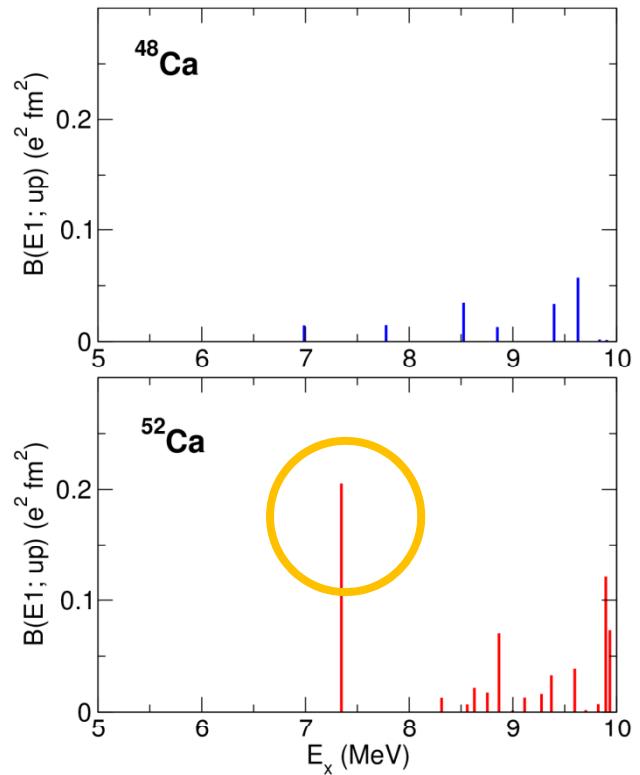
- Agreement with the systematics

- Position:  $\frac{E(1^-_{2\text{ph}})}{[E(2^+) + E(3^-)]} = 0.91$
- $B(E1)$ :  $B(E1)/D^2 = 1.6$



E2 decay competes with E1.

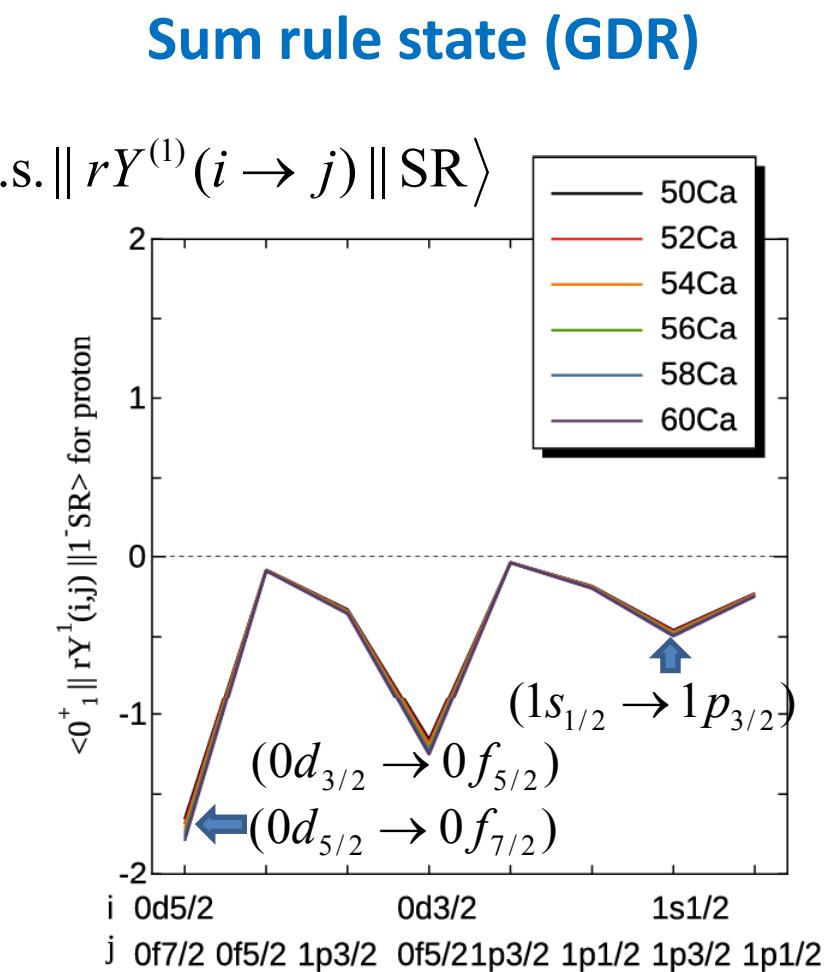
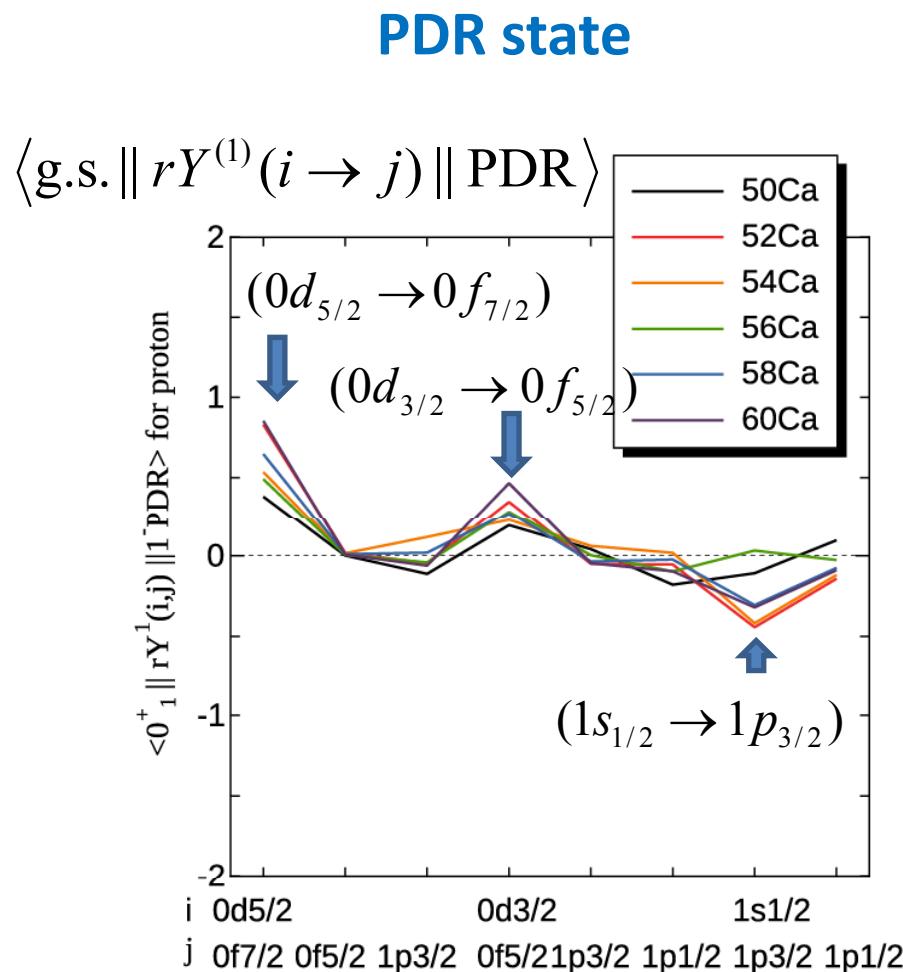
# Development of pygmy dipole resonance



T. Inakura et al., Phys. Rev. C 84, 021302(R) (2011)  
pointed out strong correlation with the occupation of the  $p$  orbitals

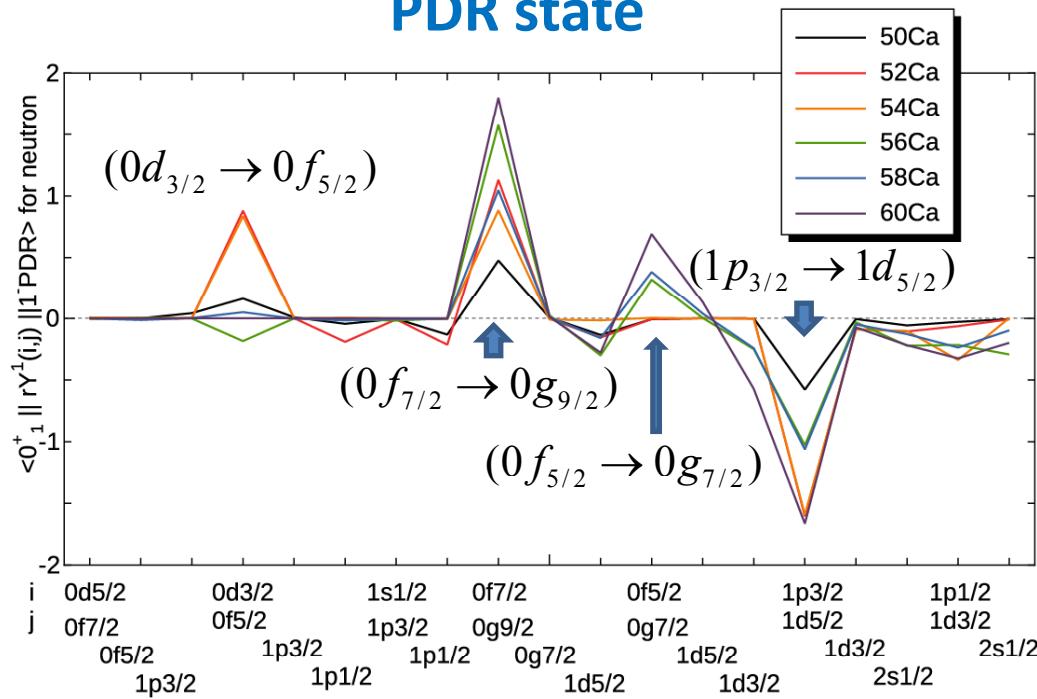
# Proton $rY^{(1)}$ matrix elements

- Decomposing the E1 mat. ele. into  $\sum_{i,j} \langle g.s. | rY^{(1)}(i \rightarrow j) | 1^- \rangle$



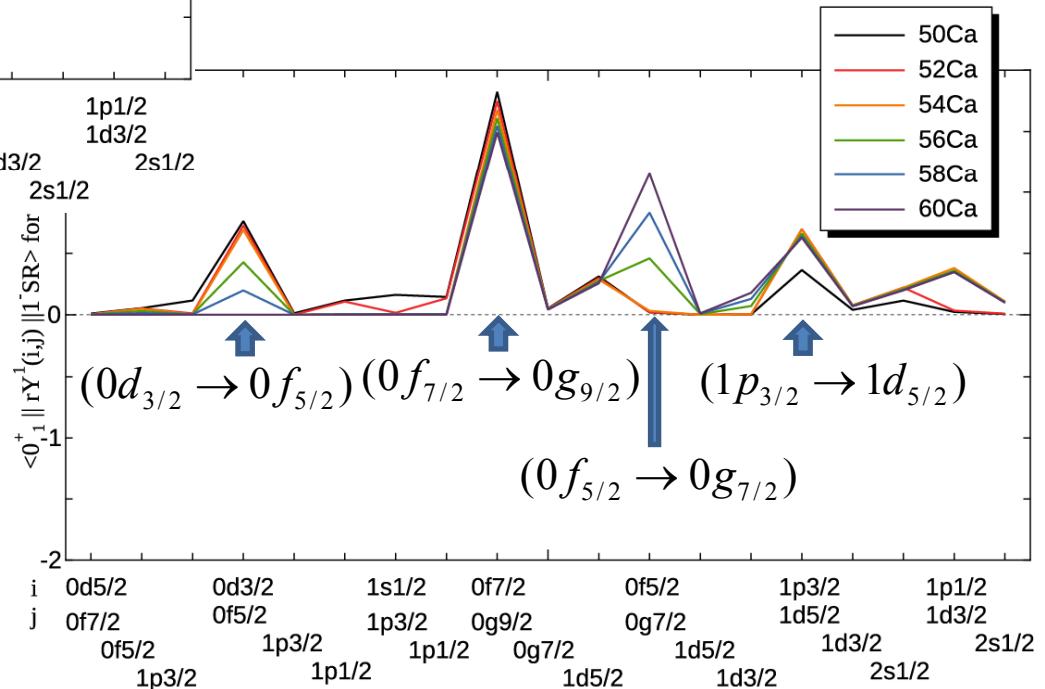
# Neutron $rY^{(1)}$ matrix elements

PDR state



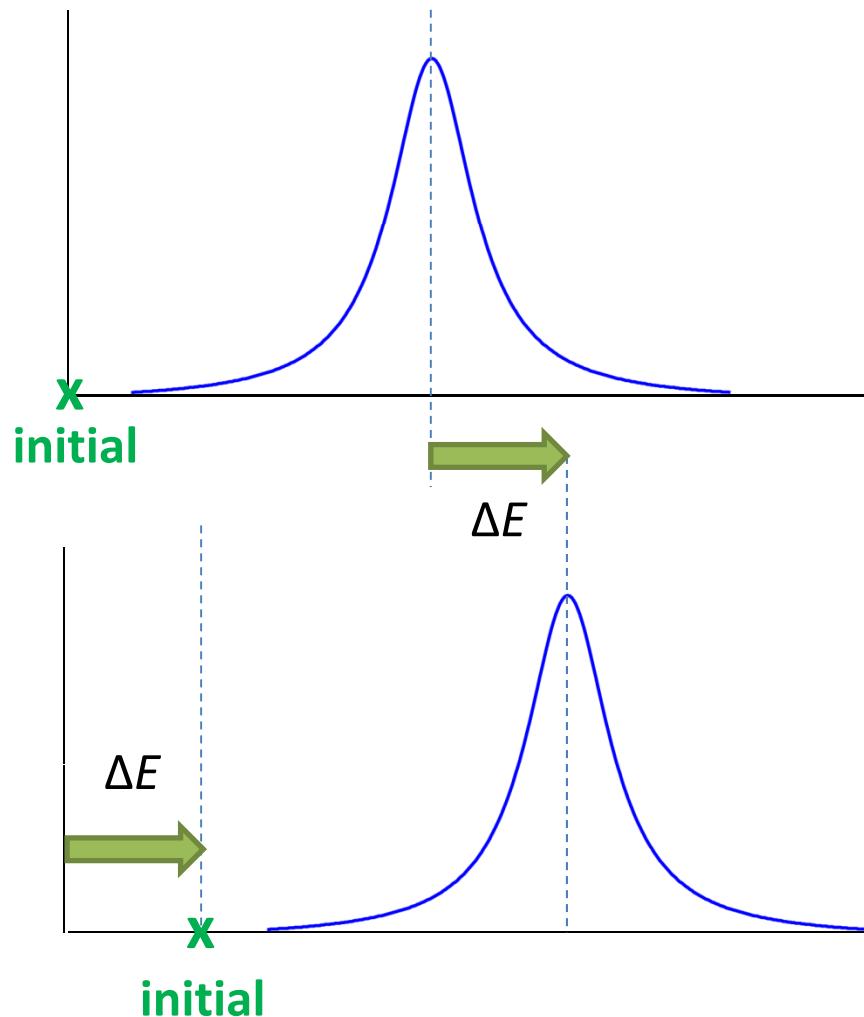
Incoherence of higher-node transition

Sum rule state (GDR)

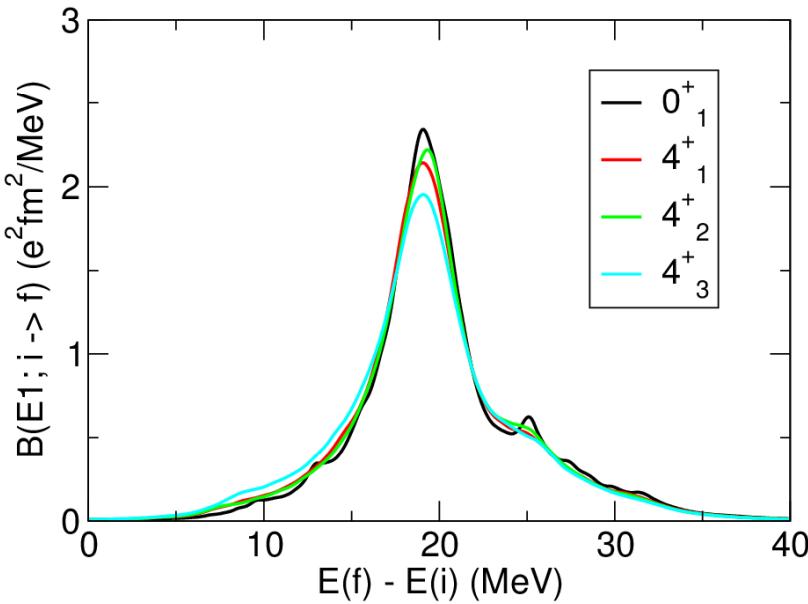
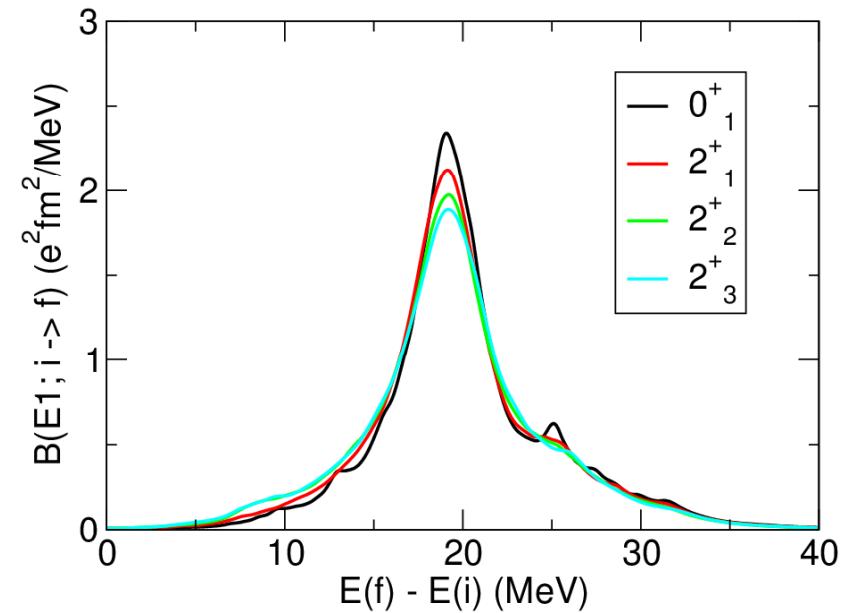
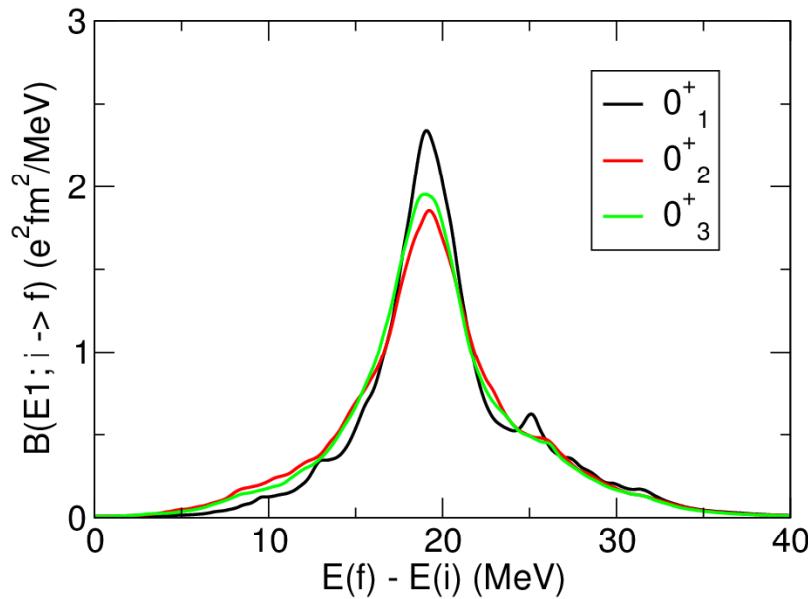


# Brink hypothesis

- GDR built on excited states
  - Assumed to be identical with that of the ground state by shifting Ex. of the initial state
    - geometric nature of GDR
  - Practically important to evaluate  $(n, \gamma)$  cross sections
- Recent experiments
  - Generalized Brink hypothesis valid in  $^{238}\text{Np}$  (M. Guttormsen et al., 2016)
  - Possible deviation from the Brink hypothesis based on  $^{89}\text{Y}(p, \gamma)^{90}\text{Zr}$  (L. Netterdon et al., 2015)

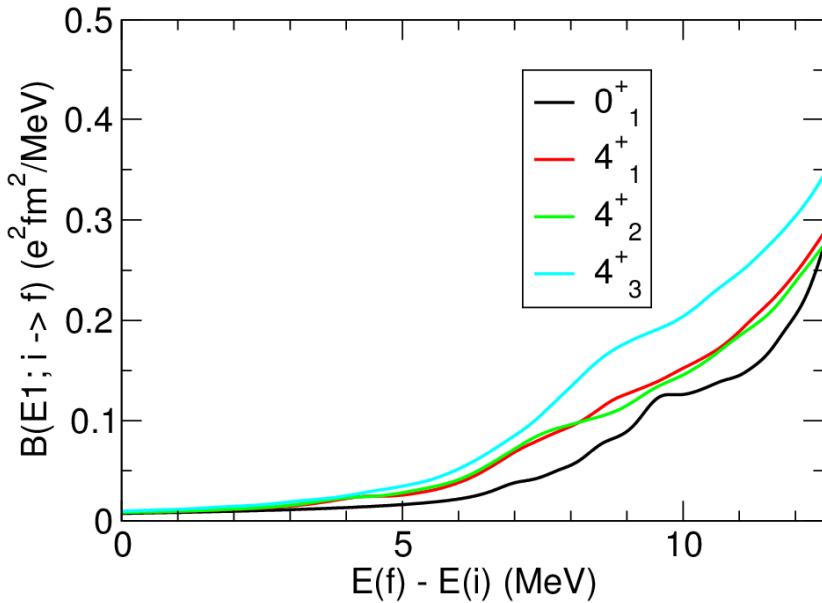
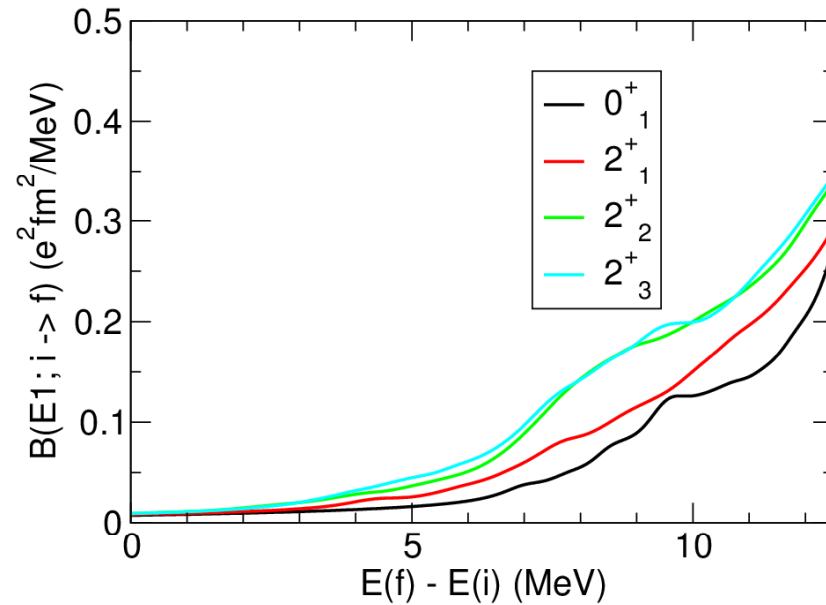
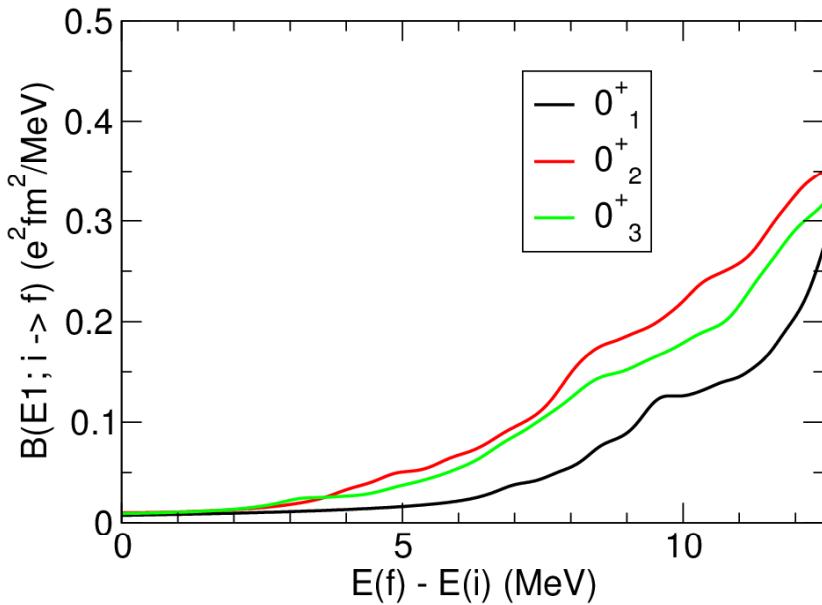


# Examining the Brink hypothesis in $^{48}\text{Ca}$



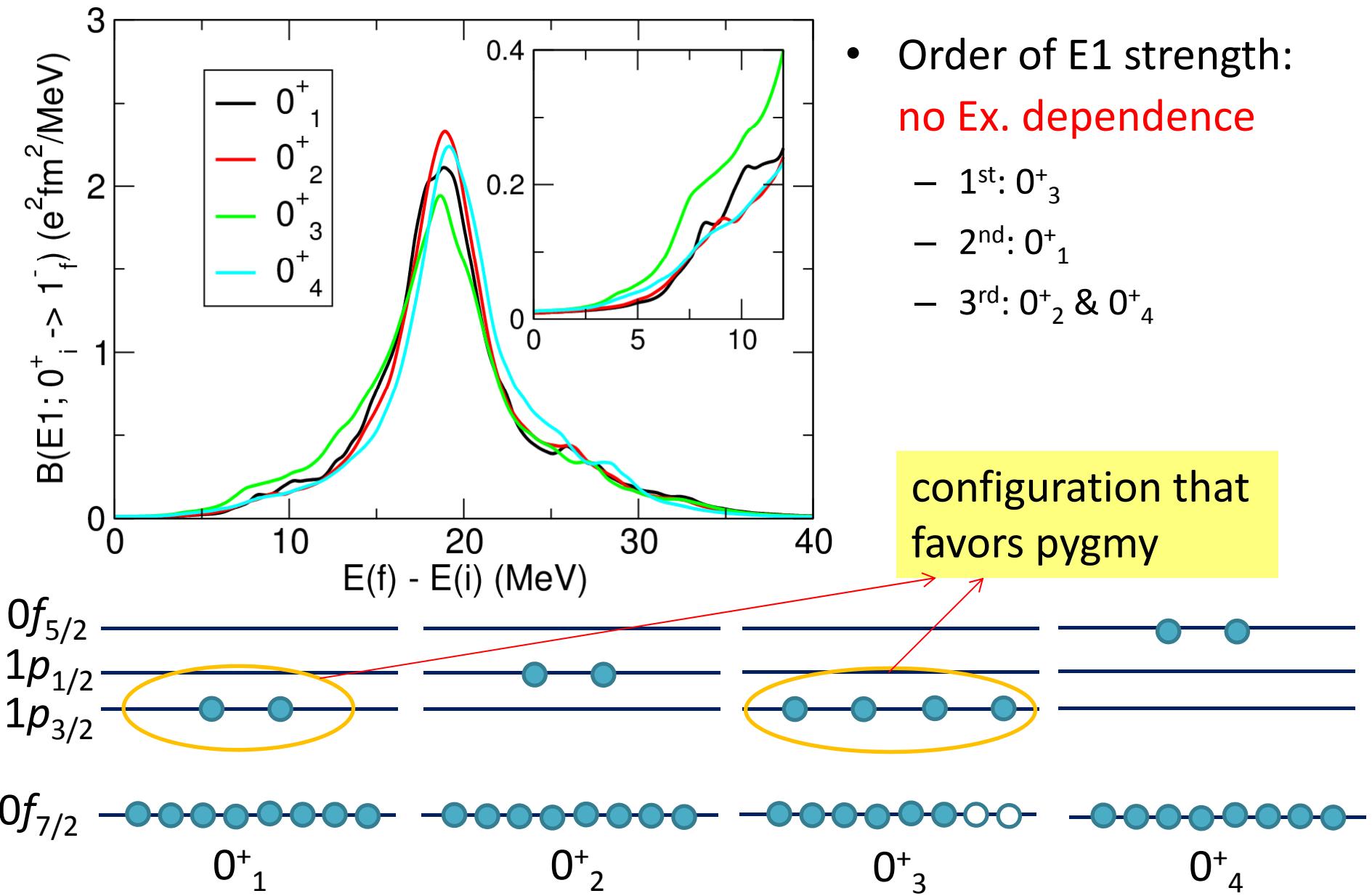
- The “generalized” Brink hypothesis (independent of excitation energy and spin) looks valid as far as the GDR region is concerned.

# Brink hypothesis in the tail region: $^{48}\text{Ca}$



- Excited states have larger E1 strengths in the tail region.
- What dominates the difference? Excitation energy?

# Brink hypothesis in the tail region: $^{50}\text{Ca}$



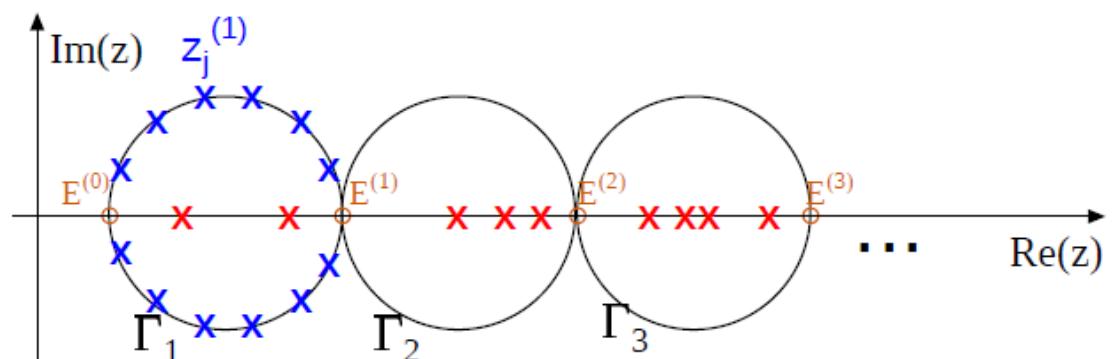
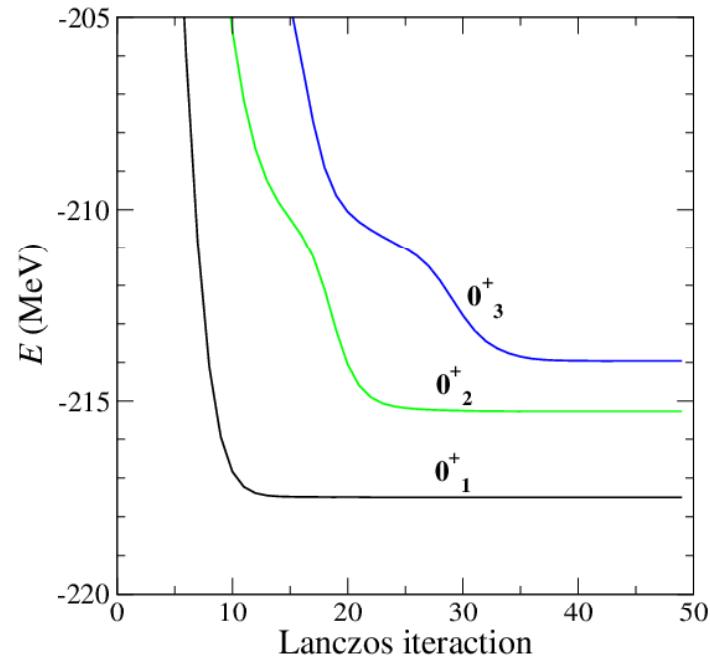
# Shell-model calculation for level density

- Direct counting with Lanczos diag.
  - Practically impossible because high-lying levels are very slow to converge
- New method (Shimizu, Futamura, Sakurai)
  - Utilizing contour integral

$$\mu_k = \frac{1}{2\pi i} \oint_{\Gamma_k} dz \sum_i^D (z - \lambda_i)^{-1}$$

$$= \frac{1}{2\pi i} \oint_{\Gamma_k} dz \operatorname{tr}(z - H)^{-1}$$

Typical convergence pattern



# Stochastic estimate of the trace

- Remaining task: estimating

$$\text{tr}(z - H)^{-1}$$

$$= \sum_i \mathbf{e}_i^T (z - H)^{-1} \mathbf{e}_i$$

dimension  $D$

- Stochastic sampling

$$\simeq \frac{1}{N_s} \sum_s \mathbf{v}_i^T (z - H)^{-1} \mathbf{v}_i$$

# sampling  $N_s$

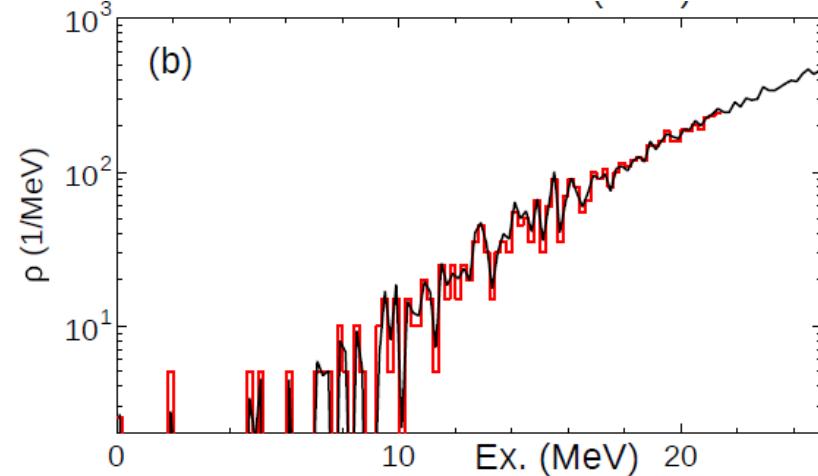


Solve  $\mathbf{v}_i = (z - H)\mathbf{x}_i$  using the conjugate gradient method

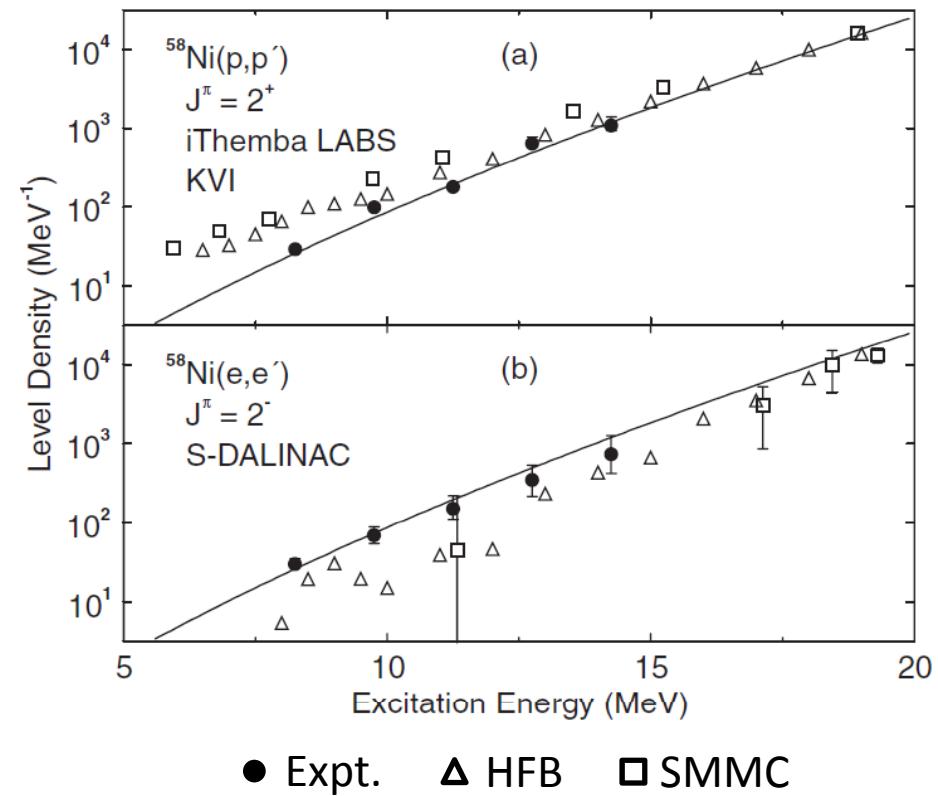
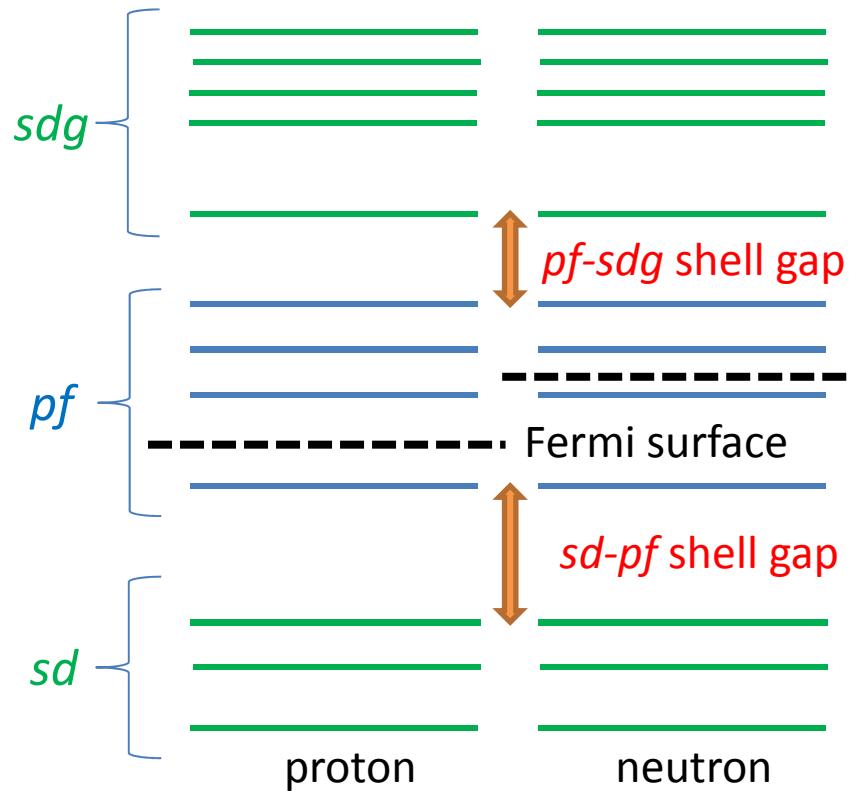
When  $\mathbf{v}_i$ 's are chosen to have good quantum numbers ( $J, \pi$ ), spin-parity dependent level densities are obtained.

The trace can be excellently estimated with a small number of sampling (known in computational mathematics).

Exact vs. stochastic estimate ( $N_s=32$ ) for the level density in  $^{28}\text{Si}$



# Application: $2^+$ and $2^-$ level density in $^{58}\text{Ni}$



Y. Kalmykov et al., Phys. Rev. Lett. 99, 202502 (2007).

- **Middle of the *pf* shell:** Large energy is needed to excite a nucleon across the major shells. It is thus not easy to reproduce the equilibration of  $2^+$  and  $2^-$  levels observed recently. **Is it possible to simultaneously describe low-lying spectroscopic strengths and the parity equilibration?**

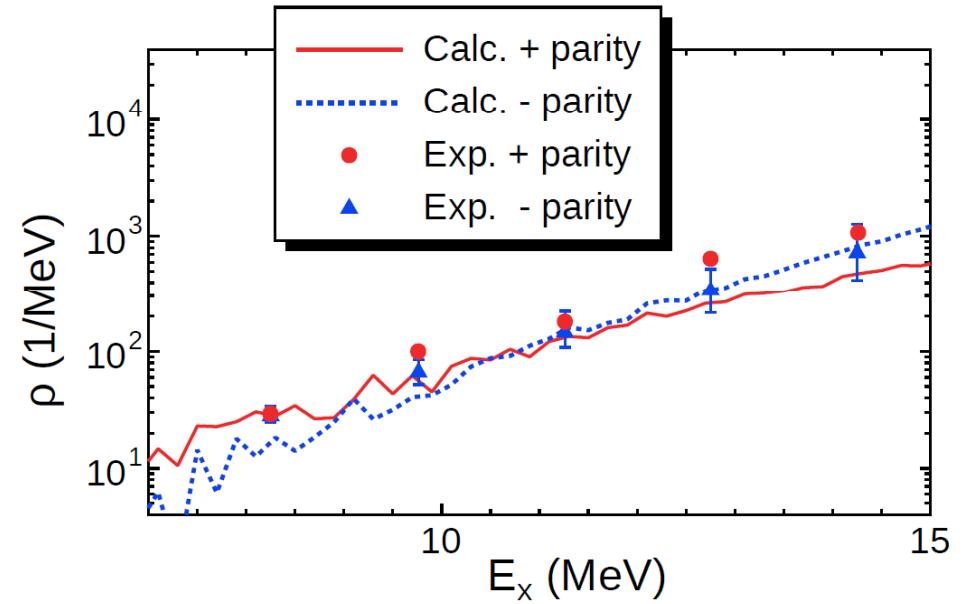
# Results of the shell-model calculation

- Very large-scale shell-model calculation
  - $0\hbar\omega$  and  $1\hbar\omega$  calc.:  $M$ -scheme dimension =  $1.5 \times 10^{10}$  for  $1\hbar\omega$
  - SPEs of the  $sdg$  orbitals are fine-tuned to obtain good spectroscopic strengths.

## Spectroscopic strengths

Nucl.	$J^\pi$	$E_x$ (MeV)		$C^2 S$		
		Cal.	Exp.	$j$	Cal.	
$^{57}\text{Co}$	$7/2^-_1$	0	0	$\pi 0f_{7/2}^{-1}$	5.28	4.27, 5.53
	<b>-1p</b>	1.037	2.981	$\pi 1s_{1/2}^{-1}$	0.98	1.05, 1.31
	$3/2^+_1$	3.565	3.560	$\pi 0d_{3/2}^{-1}$	1.70	1.50, 2.33
$^{57}\text{Ni}$	$3/2^-_1$	0	0	$\nu 1p_{3/2}^{-1}$	1.14	1.04, 1.25, 0.96
	<b>-1n</b>	5.581	5.580	$\nu 1s_{1/2}^{-1}$	0.51	0.62, 1.08
	$3/2^+_1$	5.579	4.372	$\nu 0d_{3/2}^{-1}$	0.29	0.01
	$3/2^+_2$	6.093	6.027	$\nu 0d_{3/2}^{-1}$	0.22	0.66, 0.54
$^{59}\text{Cu}$	$3/2^-_1$	0	0	$\pi 1p_{3/2}^{+1}$	0.53	0.46, 0.49, 0.25
	<b>+1p</b>	3.139	3.023	$\pi 0g_{9/2}^{+1}$	0.26	0.24, 0.32, 0.27
$^{59}\text{Ni}$	$3/2^-_1$	0	0	$\nu 1p_{3/2}^{+1}$	0.51	0.82, 0.33
	<b>+1n</b>	3.053	3.054	$\nu 0g_{9/2}^{+1}$	0.63	0.84, 0.39
	$5/2^+_1$	4.088	3.544	$\nu 1d_{5/2}^{+1}$	0.04	0.03
	$5/2^+_2$	4.595	4.506	$\nu 1d_{5/2}^{+1}$	0.30	0.23, 0.14
	$1/2^+_1$	4.399	5.149	$\nu 2s_{1/2}^{+1}$	0.00	0.09
	$1/2^+_2$	5.492	5.569	$\nu 2s_{1/2}^{+1}$	0.18	0.02
	$1/2^+_3$	5.589	5.692	$\nu 2s_{1/2}^{+1}$	0.02	0.13

## $2^+$ and $2^-$ level densities in $^{58}\text{Ni}$



N. Shimizu et al., Phys. Lett. B 753, 13 (2016).

# Summary

- Large-scale shell-model calculations are carried out to study gamma-ray strength function and level density on the same footing.
  - **Coupling to non-collective levels** is automatically included in the shell-model framework, thus leading to good GDR tail properties.
  - **Transitions between excited levels** can be obtained.
  - Some topics in Ca isotopes
    - Two-phonon state
    - Pygmy dipole resonances
    - Examining the Brink hypothesis
      - good for the GDR region, not necessarily perfect for the tail region due to configuration dependence in PDR
  - Level density
    - New method which enables large-scale calculations
    - $^{58}\text{Ni}$ : equilibration of positive- and negative-parity levels