# Neutrino oscillations: status and perspectives with accelerator beams

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# Outline

- Which Physics and how it is done.
- Fruitful days with present experiments

– OPERA: 
$$\nu_{\mu} \rightarrow \nu_{\tau}$$
, T2K, NOvA:  $\nu_{\mu} \rightarrow \nu_{e} / \nu_{\mu}$ 

- Starting to aim at leptonic CP violation, mass hierarchy
- The challenges for the future:
  - "
     statistics ↓ systematics"
- New ideas and initiatives in EU, JP and US
  - Hyper-K, DUNE, SB program, CERN  $\nu$  plat

#### A vast topic ... not covered:

exotic searches, sterile neutrinos, cross section, R&D experiments, ...

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### v mixing and oscillations

Mass eigenstates  $(v_1, v_2, v_3) \leftrightarrow$  weak eigenstates  $(v_1, v_2, v_3) \leftrightarrow$ "atmospheric"  $|
u_{lpha}(t)
angle = \sum_{i=1}^{3} U_{lpha i}^{*} |
u_{i}(t)
angle$ "solar"  $\Delta m_{21}^2$  $\Delta m^2_{31}$ U: PMNS matrix s = sin, c = cos $\mathsf{U} = \left(\begin{array}{cccc} 1 & 0 & 0 \\ 0 & c_{23} & s_{23} \\ 0 & -s_{23} & c_{23} \end{array}\right) \left(\begin{array}{cccc} c_{13} & 0 & e^{-i\delta}s_{13} \\ 0 & 1 & 0 \\ -e^{i\delta}s_{13} & 0 & c_{13} \end{array}\right) \left(\begin{array}{cccc} c_{12} & s_{12} & 0 \\ -s_{12} & c_{12} & 0 \\ 0 & 0 & 1 \end{array}\right)$ SuperK, K2K, MINOS, (D)CHOOZ, Daya Bay, RENO SuperK, SNO, GNO, **OPERA, T2K, NOvA T2K, MINOS, NOvA** Gallex, Borexino, KamLAND  $\theta_{23} = (45.8 \pm 3.2)^{\circ}$  $\Delta m_{21}^2 = (7.53 \pm 0.18) 10^{-5} eV^2$  $\theta_{12} = (33.4 \pm 0.85)^{\circ}$  $|\Delta m^2_{22}| = (2.44 \pm 0.06) 10^{-3} \text{ eV}^2$  $\theta_{13} = (8.88 \pm 0.39)^{\circ}$ PDG2014 Long baseline experiments

CP violation? mass hierarchy  $(m_{1,2} \le m_3)$ ?  $\theta_{23} = 45^{\circ}$ ? Majorana/Dirac ? Symmetries? Relation with CKM ? Leptogenesis and BAU ?

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# **2012:** new scenarios from large $\theta_{13}$



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#### The future: learning a lot from (precisely!) measuring $v_{\parallel} \rightarrow v_{\perp}$



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### Learning a lot from (precisely!) measuring $v_{\mu} \rightarrow v_{e}$

- $\delta_{CP} \rightarrow$  a modulation in the spectrum of the appeared v
- The direction of the variation is opposite for  $\nu$  and anti- $\nu$  beams  $\rightarrow$  use both
- Mainly a change in normalization
  - accessing the  $2^{nd}$  maximum (at higher L and E)  $\rightarrow$  more spectral info.
- Sub-leading: crucial role of systematics and statistics



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#### v-beams (recent past and present)



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### The long way to appearance

• **Disappearance** a "leading" effect: deficit of atmospheric  $v_{\parallel} \rightarrow$ 

- 1998 discovery of v-oscillations by Super-K, MACRO, K2K  $\ldots$ 

• **Appearance** on the other hand considered difficult:

• At the **solar scale**. Reactors and solar v.

 $\nu_e \! \rightarrow \! \nu_{\mu} \,$  "IMPOSSIBLE":  $\mu$  is below threshold

• At the **atmospheric scale**. Atmospheric-v, artificial beams.

 $v_{\mu} \rightarrow v_{\tau}$  "DIFFICULT" ! $v_{\mu} \rightarrow v_{e}$  "RARE"(...)mass suppression, small  $c\tau$  $\theta_{13}$  suppression ?Today's perspectiveConfirmed difficult, but event<br/>by-event detection achieved by<br/>OPERA & sk (with a much lower S/B)• Reactors: no...  $\theta_{13}$  is BIG !<br/>• Appearance seen by T2K<br/>and NOvA (with few POT)

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### The OPERA road map

An experimental and technological challenge. 732 km baseline. Beam O(10) more energetic (17 GeV) than any other LBL ( $m_{\tau}$ ). A "fine-grained" detector O(100) more massive (1.25 kt) than the precursors SBL (i.e. CHORUS).



# The CNGS beam for $v_{\mu} \rightarrow v_{\tau}$

< E <sub>v</sub> >	<b>17 GeV</b>		
L / < E <sub>v</sub> >	43 km/GeV		

The oscillation peak for L= 732 km at ~ 1.5 GeV (similar to NuMI) but here the goal is to produce  $\tau$  leptons  $\rightarrow$  unbalance at higher energies

$$N(\tau) \sim Pr(\nu_{\mu} \rightarrow \nu_{\tau}) \ge \sigma_{\nu(\tau)CC}(E) \ge flux$$

Fluxes:

$(v_e + \overline{v_e}) / v_{\mu}$	0.9 %
$\overline{\mathbf{v}}_{\mu} \ \mathbf{I} \mathbf{v}_{\mu}$	2.1 %
$v_{\tau}$ prompt (from D <sub>s</sub> )	negligible

Interaction rates (1.8 x  $10^{20}$  pot ):

```
~ 20k \nu_{\mu} CC+NC
66.4 \nu_{\tau} CC (not efficiency corrected)
```

DESIGN: 4.5.10<sup>19</sup> pot/year, 200 days/y per 5 y



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## The $v_{\tau}$ detection challenge





#### Modular detector of "Emulsion Cloud Chambers" (or bricks)

**Reconciles the needs for:** 

- Large mass
  - $N_{\tau} \propto (\Delta m^2)^2 M_{target}$
- Extreme granularity
  - ~ µm

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Super Module 1

Super Module 2



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# $\textbf{OPERA} \, \boldsymbol{v}_{_{\boldsymbol{\tau}}} \, \textbf{appearance}$

Channel	Total background	Expected signal	Observed
$\tau \to 1 h$	$0.04\pm0.01$	$0.52\pm0.10$	3
$\tau \to 3h$	$0.17\pm0.03$	$0.73\pm0.14$	1
$\tau \to \mu$	$0.004 \pm 0.001$	$0.61\pm0.12$	1
$\tau \to e$	$0.03\pm0.01$	$0.78\pm0.16$	0
Total	$0.25\pm0.05$	$2.64\pm0.53$	5

5 candidates fulfilling the kinematic selection defined in the experiment proposal

5.1  $\sigma$  exclusion of the background-only hypothesis

 $\rightarrow$ 



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### T2K

500 members 59 istitutes 11 countries





#### First "off-axis" beam

- $2.5^{\circ} \rightarrow \text{peak at} \sim 0.6 \text{ GeV}$
- Enriched in Quasi-elastic interactions (good measurement of E<sub>v</sub>)
- Reduced instrinsic  $\boldsymbol{v}_{a}$  background
- Reduced NC  $\pi^0$  ~backg. from D.I.S.
- Double detector: 280 m and 295 km

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3000

2500

2000

1500

1000

500

0.5

(Flux × x-section)

OA0°

OA2°

OA3°

0A2.5°



### T2K runs



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#### v disappearance μ

Approximate value of  $\Delta m_{23}^{2}$  known at design phase.

Maximal suppression exactly at peak – not the case f.e. in MINOS.

#### 446 ± 23 exp. (no osc.) 120 obs.



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2

 $E_{\nu}^{rec}$ 

18

16

12

10

8

6

4

2

0

events/0.2GeV



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# **T2K:** anti- $v_e$ appearance



3 events observed in the current sample ( $4 \times 10^{20}$  POT)

Short term (1 year) goal:  $9.5 \times 10^{20}$  POT

- $2\sigma$  rejection for no anti- $v_a$  appearance
- 60% chance of 99 CL observation

Next step: joint v+anti-v fit

First results on anti-numu disappearance shown later on.

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### The NOvA experiment



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### The NOvA detectors



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NOvA: v μ μ

 $2.74 \times 10^{20}$  POT (8% of planned)

- 201 expected without oscillation
- 33 observed

 $\Delta m_{32}^2 = 2.37 + 0.16_{-0.15} \times 10^{-3} \text{ eV}^2$  $\sin^2 2\theta_{23} = (0.51 \pm 0.10)$ 



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### MINOS+

MINOS, MINOS+ Far Detector Data

Prediction, No Oscillations

MINOS, MINOS+ Combined Fit

The MINOS detector continues being illuminated by the on-axis beam in "NOvA configuration" (yielding a higher-E beam than before  $\sim$  7 GeV)  $\rightarrow$  strengthen the measurement of disappearance (especially far from of the oscillation maximum)

 $2.99 \times 10^{20}$  POT v<sub>u</sub>-mode MINOS+

10.71  $\times 10^{20}$  POT  $\nu_{\mu}\text{-mode MINOS}$  3.36  $\times 10^{20}$  POT  $\overline{\nu}_{\mu}\text{-mode MINOS}$ 

**MINOS+** Preliminary

15

20

10

Reconstructed  $v_{\mu}$  Energy (GeV)





1200

1000

800

600

400

200

Events / GeV

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Ratio to No Oscillations

50

30

### Summary on disappearance

#### **Neutrinos**

#### **Anti-neutrinos**



T2K anti-v T2K v

**NEW! T2K result today** 

on the arXiv 1512.02495

- Both for neutrinos and anti-neutrinos
  - MINOS: still leading for  $\Delta m^2$  (~ E of the "dip")
  - T2K: leading for  $\theta_{23}$  (depth of the "dip")
- No significant hints for deviation from maximal mixing or CPT effects

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favored

Calorimetric energy (GeV)

# NOvA: $\delta_{CP}$ with reactor constraint

For all  $\sin^2 2\theta_{23}$  in [0.4,0.6]

#### LEM analysis

- IH disfavored at >  $2.2\sigma$
- NH mildly disfavored (>1 $\sigma$ ) for  $\delta \in [0, \pi]$

#### LID analysis

• IH mildly disfavored (>1 $\sigma$ ) for  $\delta \in [0, 0.8\pi]$ 

Both LEM and LID prefer

- normal hierarchy
- $\delta \sim 3/2\pi \ (= -\pi/2)$

#### In the same direction of T2K

T2K excludes (90% CL):

0.15  $\pi < \delta_{_{\rm CP}} < 0.83 \pi$  (NH) -0.08  $\pi < \delta_{_{\rm CP}} < 1.09 \pi$  (IH)



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### Looking forward

### Near and far detectors

Ideally ... in a near-far double detector oscillation experiment

 $N_{events}(E_{\nu}) = \sigma_{\nu}(E_{\nu})\Phi(E_{\nu})$ 

 $N_{events}^{far}(E_{\nu}) = \sigma_{\nu}(E_{\nu})\Phi(E_{\nu})P_{osc}(E_{\nu})$ 

$$\frac{N_{events}^{far}(E_{\nu})}{N_{events}(E_{\nu})} = P_{osc}(E_{\nu})$$

Neutrino cross sections factorize out But beams are not monochromatic  $\rightarrow$  we need to determine E<sub>1</sub> event by event



Oscillations introduce differences in the flux spectrum: cross-sections do not cancel out



#### We need: **φ(E**, **)**, **σ(E**, **)**, **P(E**, **|E'**, **)**

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# Uncertainty on events at SK after the T2K near detector constraint



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### T2K systematics today

Events prediction at Super-K with the near detector constraints:

- ~ 7 % for v
- $\sim 10\%$  for anti-v

Largest contributions:

- difference in nuclear targets between far and near
- effect of the poorly known multi-nucleon processes (MEC) cross section

		$\nu_{\mu}\text{sample}$	$v_{e}$ sample		$\overline{ u}_{\mu}$ sample	$\overline{\nu}_e$ sample	
$\nu$ flux (with hadroproduction constraints NA61)		16%	11%		7.1%	8%	
$\boldsymbol{v}$ flux and							
cross section	w/ ND measurement	2.7%	3.1%		3.4%	3.0%	
v cross section due to difference of nuclear target btw. near and far		5.0%	4.7%		10%	9.8%	
Final or Second Hadronic Inter	dary action	3.0%	2.4%		2.1%	2.2%	
Super-K detect	or ~7 %	4.0%	2.7%		3.8%	3.0%	
* 2015 uncertainties include additionally MEC							

 $2014 \rightarrow 2015$ 

~10 %

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### **Desiderata for systematics reduction**

#### $\phi(E_v) \times \sigma(E_v)$ at near and far

- Same target nuclei:
  - cancellation of nuclear effects, final state interactions (FSI)
- Same acceptance:
  - $\boldsymbol{\checkmark}$  avoid model dependence in the not common p-0 phase space
- Same flux (shape):
  - a tunable beam (combine different off-axis angles vPRISM) (to mimic oscill. distorsion)
  - a "not too near" NEAR (to reduce finite-distance effects)
- $\sigma(v)$  measured independently (v at near is subdominant):
  - tagged  $\nu_{p}$  beams,  $\mu$  facilities (nuSTORM).

#### **Energy reconstruction P(E**, |E', ) at near and far

- multi-nucleon interactions, π absorption in nuclei (FSI) → non genuine CCQE (CCQE-like) → "QE formula" is applied giving bias, broadening of P(E, |E',)
  - high granularity, low-thresholds (demanding at far!)
  - neutron tagging

### **T2K perspectives, upgrades**

- JPARC Main Ring upgrade approved:  $7.8 \times 10^{21}$  POT (=T2K design), 0.9 MW by 2020
- "T2K×3" (2020-25) phase (2×10<sup>22</sup> POT). Before Hyper-Kamiokande (~2025).
- If sys. < 2-3 % "T2K×3" could give >  $3\sigma$  CPV for any value of  $\theta_{23}$
- Discussions on upgrading the near/intermediate detectors already in 2020



### TITUS

Tokai Intermediate Tank with Unoscillated Spectrum

- 2kt Gd doped (0.1%) water Cherenkov
- $\sim$  2 km from J-PARC, 2.5° off-axis
- Magnetized downstream Muon Range Detector (MRD)
- Small side MRD



0.1% Gd doping:

- 49000 b vs 0.3 b (H)
- 8 MeV  $\gamma$  (4-5 MeV visible)
- 90% capture efficiency

#### NB. > 2018 also SuperKamiokande planned to become Gd-doped (EGADS demonstrator)

Same target, similar acceptance, same flux, sensitivity to multi-nucleon with n-tagging

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### Hyper-Kamiokande



Ring-imaging **water Cherenkov detector** Tochibora mine: 648 m overburden (1.750 mwe) 2.5° at 295 km (= Super-K)

#### 1 Mton mass

99.000 20" PMTs 20% photo-coverage 25.000 8" PMTs Light attenuation > 100 m @ 400 nm

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# Hyper-K: $v_e$ samples & $\delta_{CP}$

Neutrino mode: Appearance

Antineutrino mode: Appearance



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 $\delta_{CP}$  [degree]

### **LBNF and DUNE**



- (staged) 40 kt LAr detector, at the SURF site, 1300 km from FNAL
- high granularity/high precision near detector
- 1.2 MW, tunable v beam produced by the PIP-II upgrade at FNAL by 2024, evolving to a power of 2.3 MW by ~ 2030.





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### **DUNE far detectors**



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### DUNE: $v_{e}$ samples, $\delta_{CP}$ , MH





#### Challenges

- Large cryostats ~ 13000 m<sup>3</sup>
- Deep underground activities
- Long drifts > 3.5 m
- High liquid purity at the ppt level
- High T stability ~0.3° 0.5°
- Cold FE electronics or in gas amplif.
- Low threshold signals
- Large data handling capabilities
- Automatic pattern recognition, tracking



#### To succeed we need to proceed in steps (for cryostats, cryogenics and detectors)

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Baseline

(12.9 cm)

om chower

### **FNAL short baseline program**

#### SBND, $\mu$ BooNE, ICARUS-T600

- LAr TPC R&D
- (together with the reactor's program) likely clarify the scenario of sterile neutrino disproving or confirming previous hints with beams:
  - LSND anomaly?
  - MiniBooNE low-E excess ?
  - Differences v / anti-v?





### **FNAL short baseline layout**



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### **CERN v platform: EHN1**

S. Bertolucci at INFN CSN2

- PLAFOND : an generic R&D framework
- ✓ WA104 : ICARUS as far detector for SBN
- ✓ WA105 : demonstrator + engineering prototype for a double ph. TPC
- ✓ ProtoDUNE : engineering prototype for a single phase TPC
- Baby MIND : a muon spectrometer for the WAGASCI experiment
- ArgonCube : a modular TPC R&D

✓ For the moment CERN is not committing to any neutrino beam at CERN, in view of an agreed road map between all partners

 The CERN Neutrino Platform represents a gateway for the European Neutrino Community towards a global, organized accelerator neutrino program

 In the short- and medium-term, Europe is helping in getting a Short Baseline operational at FNAL with an agreed physics program ... and later a Long Baseline

- HKK detector components R&D
- ✓ Darkside 20K
- ✓ ARIADNE\_
  - LBNF cryostat and LAr cryogenics
  - ✓ SBND cryostat and LAr cryogenics
  - <u>CERN</u> member of DUNE and SBN

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### Conclusions

- Neutrino oscillation physics with accelerator beams is living an exciting phase. Beautiful measurements in done/progress & good future prospects for fundamental physics.
- CPV out of reach if  $\theta_{13}~$  or  $\theta_{12}$  would have been small! Will  $\delta$  "continue the tradition" and finally establish the (somewhat deserved) good luck of neutrino physicists ?
- Or: Will indications for (N.I.,  $\delta = -\pi/2$ ) of T2K/NOvA survive further data?
- Even with the help of Nature ... need for smart experimental designs and inputs from several sides (cross-sections, hadro-production measurements...). New ideas to reduce systematics being proposed for present and future infrastructures.
- The international scenario is getting clearer. Opportunity of a new class of large experiments: will strengthen the reach for CPV and mass hierarchy. Precision neutrino measurements. Boost the potential for unexpected discoveries (Super-K docet).

### **Backup slides**

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### v<sub>e</sub>, $\overline{v}_e$ Cross Section Sensitivity Impact

- Perform sensitivity study where the  $v_e$  and  $\overline{v}_e$  cross sections are assigned two uncorrelated normalization systematic parameters
- The uncertainties on the normalization parameters are varied and the impact on the CPV sensitivity is studied.



- The systematic uncertainty should be controlled to <1-2% to minimize the impact on the CPV discovery sensitivity
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### **DUNE schedule**

#### Indicative schedule



### Hyper-K: CPV reach and $\delta_{_{CP}}$ precision

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Normal mass hierarchy **CPV** discovery  $\sigma = \sqrt{\chi^2}$ Well known detector technology + analysis. 5σ Robust/realistic estimation of 3σ systematic uncertainties 2 CPV:  $\delta_{CP} = 0 \text{ or } \pi$ 50 -150 -100-50 0 100 150  $\delta_{CP}$  [degree] δ<sub>CP</sub> [degree 100 5(  $\delta_{CP}[\%]$ Fraction of values of  $\delta_{cr}$  for which CPV can be discovered 90 45  $\delta = 0$ 80 40  $\delta_{_{CP}}$  precision 70 δ **= 90** 35 60 Fraction of 30 68% CL error of 50 25 5σ 40 20 30 15 3σ 201010 10 8 2 8 10 6 Integrated beam power [MW 10<sup>7</sup> sec] Integrated beam power [MW 10<sup>7</sup> sec]

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#### Two large TCPs prototypes : protoDUNE



#### Single phase LAr TPC

#### Operational in 2017, SPS calibration beams in 2018-19

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#### Two large TCPs prototypes : WA105



#### Double phase LAr TPC

Operational in 2017, SPS calibration beams in 2018-19

Active volume 6x6x6 m<sup>3</sup>

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#### Hyper-K atmospheric data

295 km → small matter effects → limited contribution from CPV induced by matter effects → clean measurement of genuine CPV





Would mass hierarchy be still unknown by the time of Hyper-K: use large samples of atmospheric neutrinos for which matter effects are definitely large.



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# Hyper-K: $\theta_{23}$ octant



# **T2K potential:** $\theta_{23}$ and $\Delta m^2_{32}$

7.8e21 POT+2012 syst. err. + 50-50% v-anti-v





### WAGASCI

Addresses the issues of: acceptance, target definition, external backgrounds.

Goal: 3 % error on cross section ratio (water/CH)

Plastic scintillators + WLS fibers in arrays (water/plastic) filled. Hamamatsu MPPC (SiPM) readout.

Being constructed close to ND280, INGRID



Grooves to mechanically connect orthogonal scintillator bars. Shallow enough to allow fiber housing.

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#### **Off-axis near detector analysis**

Fit of  $v_{\mu}$  spectrum to constrain flux X cross-section ( $v_{\mu}$  also constrain  $v_{e}$  via correlation in the production mechanism). 3 subsamples with final state  $\pi$  "CC  $0\pi$ ", "CC  $1\pi$ " and "CC other"



# **High pressure TPC**

- No passive material (interactions in the gas)
- Low thresholds (5-10 bar pressure) disentangle multi-nucleon processes from CCQE
- Realistic gases: He, Ne, Ar, CF<sub>4</sub>
- H and D would "by-pass" nuclear physics ... not realistic
- In principle more appealing for the US program (Argon). Difficult to use CO<sub>2</sub>, H<sub>2</sub>O (for water)



#### Taken from F. Sanchez



# TITUS + HyperK: impact on $\delta_{CP}$



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### vPRISM

Extract the energy dependence by measuring the rates and final state kinematics over a range of off-axis angles



Detector moved up and down a shaft ~ 1 km baseline: span: 1-4 degrees

WC detector: 6 m diameter x 10 m height 40 % photo-coverage: 3120 8" PMT or 7385 5" PMT



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- An intermediate phase: "T2K x 3"
  - 3x T2K statistics ( $20 \times 10^{21}$  POT)



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#### The tool: accelerator-based v-beams (prehistory)



#### The tool: accelerator-based v-beams (history)



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