



# The ARCADE project

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# ARCADE



## Atmospheric Research for Climate and Astroparticle DEtection

- Approved by the MIUR in 2011
- started in 2012 ended in 2016

EGLI STUD

 involves people from : University / INFN of Naples and Turin, CETEMPS/DPSF L'Aquila + Colorado School of Mines, USA

# Atmosphere and UV light

- UHE cosmic rays entering the atmosphere determine the production of showers of secondary particles (EAS) and of UV light (fluorescence / Cherenkov light)
- UV light can be detected to infer the properties of the primaries

#### The atmosphere is responsible for both the production and attenuation of UV light

Properties of the atmosphere need to be well known : the scattering of light due to aerosols is the most significant and variable phenomenon, influencing the correct determination of the longitudinal development of the EAS in air and of its energy. Clouds distort UV light profiles (enhancing or blocking light).







# The effect of aerosol conditions on air fluorescence measurements

Neglecting the presence of aerosols causes an underestimate in energy on average from 8% (at lower energies) to 25% (at higher energies)

- 20% of showers need a >20% energy correction
- 7% of showers need a >30% energy correction
- 3% of showers need a >40% energy correction



Neglecting the presence of aerosols causes a systematic shift in Xmax from -1 g/cm<sup>2</sup> at lower energies to 8 g/cm<sup>2</sup> at higher energies

from The Pierre Auger Collaboration, Astroparticle Physics 33 (2010) 108-129

## Atmospheric monitoring in the CR community

• The aerosol attenuation in CRs community is usually inferred through elastic LIDARs and / or side-scattering measurements from laser sources. Cloud layers height is measured with the same instruments.

• these measurements are based on assumptions on atmospheric properties (horizontal homogeneity, LIDAR ratio, ...)

#### uncertainty on aerosol VAOD (h) ~ 30 %

 The anelastic Raman LIDAR is the only device performing measurements not based on any assumption, but its operation in the field of view of UV light detectors is limited due to the long acquisition times needed.

ARCADE

# Target of the ARCADE project

Perform measurements of the aerosol attenuation profiles of UV light in atmosphere **simultaneously and on the same air mass** using the typical techniques used in cosmic ray and gamma ray observatories :

1. Side-scattering measurements using a distant laser facility and a UV light telescope

2. Back-scattering measurements using elastic and Raman lidars



# Target of the ARCADE project

Perform measurements of the aerosol attenuation profiles of UV light in atmosphere **simultaneously and on the same air mass** using the typical techniques used in cosmic ray and gamma ray observatories

aiming to : a better understanding of the **systematics** and **limits of applicability** related to each method and **possible enhancements** 

Location : desert – like environment, typical location for cosmic rays / gamma rays experiments : Lamar, Colorado (U.S.A.)



# The experimental setup

LIDAR + telescope for the detection of UV light (AMT)

No interference with any ongoing experiment : free data taking!



designed and realized within this project at the Turin INFN mechanical workshop owned by the Colorado School of Mines, reassembled and improved for this project











## Lidars and side scattering

- LIDARS: remote sensing technology that measures aerosol attenuation profiles by firing a laser beam through the atmosphere and analyzing the back-scattered light.
  - Elastic Lidars measure the backscattered light on molecules and aerosols (needs for some assumption)
  - Raman Lidars measure the backscattered light on specific molecules (N2, H2O, ...)

• Side Scattering Measurements : a telescope for the detection of UV light is positioned far from a UV laser light source. The laser light is scattered and attenuated exactly as the UV light produced during the development of an EAS → the study of the laser light that reaches the telescope is used to infer the aerosol attenuation properties

## Atmospheric Monitoring Telescope Colorado School of Mines

Put in operation after 2 years of incativity. Tested and reassembled in the lab in Golden. Improved DAQ and calibration. Back to operation in the field in June 2014.

Atmospheric Monitoring Telescope

3.5 m<sup>2</sup> mirror (HiRes Optics) 1 Field of View Pixels Power Internet Electronics

UV Pass Filter

# AMT -> Laser Simulation

AugerOffline software

Simulation of laser light emission, transportation and attenuation towards the telescope

Geant4  $\rightarrow$  Telescope simulation



# LED Calibration

Each pixel of the telescope respond in a different way to the same amount of light

Relative calibration needed to make it uniform--> FLAT FIELDING CORRECTION

An uniforme and stable LED source is placed in the center of the 4-petals mirror of the AMT and illuminate uniformly the whole camera.



LED calibrations  $\rightarrow$  before and after each night of data taking



## **Nearby Laser Calibration**

A roving laser placed at 3km far from the AMT was used to obtain an absolute normalization of the simulation to the real data The effects of atmosphere on simulation are negligible at this distance







# The ARCADE Lidar





# Geant4 simulation of the Lidar



#### LIDAR box and optics





#### The laser bench realized



#### The lidar is mounted on a steering mechanism ...



... is housed within an astronomical dome ...



# ... and is hosted on top of a 20 ft shipping container

# **Built for remote operation**





#### Needs only power and internet connection



## **Remote Control**

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ARCADE Pow	ver Control Lidar Control DAQ Log More	♣ Rain: NO ⊀ Wind: 0.5 km/h ≠
Power C	Control	
	Lidar PCs	Turn OFF
	Light	Turn OFF
	Mation control	Turn ON
	Photomultipliers	Turn ON
	Laser	Turn ON

All the devices are remotly controlled :

- → Dome Open/Close
- → High Voltage of PMTs
- → Radiometer
- → Laser
- → Steering
- → Weather Stations
- → Beam Mirror Alignment

## Lidar system



### ELASTIC LIDAR Multi-angle analysis



### RAMAN LIDAR Vertical and inclined shots



Hp: horizontally homogeneous atmosphere

2 equations for two unknowns

## An example of the ARCADE lidar data



## LIDAR -> Raman Analysis

In collaboration with atmospheric group of CETEMPS - L'Aquila

$$P(\lambda_{N_R}, R) = \frac{K(\lambda_R)}{R^2} \cdot \beta(\lambda_R, R) \cdot e^{-\int \alpha(\lambda_0, R) + \alpha(\lambda_R, R) dR}$$

The constant "C" is determined imposing that  $\tau_{aer}(R)$  is a linear function between the ground level and a certain altitude  $R_1$ 

## LIDAR -> Raman Analysis

$$\tau_{aer}(R > R_1) = \alpha(R_1) \cdot R_1 + \tau(R_1, R)$$



The constant "C" is determined imposing that  $\tau_{aer}(R)$  is a linear function between the ground level and a certain altitude R<sub>1</sub>

# Preliminary aerosol profiles with the ARCADE Raman Lidar



Elastic Range Corrected Signal

# Preliminary aerosol profiles with the ARCADE Raman Lidar

#### **CLOUDY DAY**



## Data taking & problems 1 year – June 2014 to June 2015



#### remote shifts from Italy, 1 shifter

- > 20 minutes vertical (0°) Raman acquisition @ 100Hz, full power (6 mJ)
- > 5 minutes @ 100Hz for each of the 5 positions (0°, 30°, 40°, 46°, 51°)
- > 10 minutes @ 90° to test the horizontal homogeneity of the atmosphere
- > a few problems during data taking ...
- 1. Temperatures too low (long periods T<-10°C) laser heating failing when T < -5°C
- 2. hardware failures needed to be fixed
- 3. network very unstable on site
- 4. weather rapidly changing and extremely dusty ...

# July 2015 – umounting the lidar



... very bad weather in June (tornado @ less than 1 mile from the Lidar) broke the electrical panel and the lidar was stuck open under the bad weather for a few hours ...

> the ARCADE Lidar is now back in Italy being upgraded to become part of the CTA



# Conclusions

The target of ARCADE is the measurement of the aerosol attenuation profiles of UV light in atmosphere using different techniques on the same air mass, at the same time, to understand the limits of applicability and the systematics of each technique.

We took data for one year with the ARCADE Raman Lidar and the AMT in Lamar, Colorado.

Simulations of the Lidar and AMT + calibration campaignes of the AMT have been performed. First aerosol profiles using the Raman Lidar have been measured, while the AMT data analysis is in progress.

#### Thanks for your attention!



### An example of the ARCADE lidar data



Fig. 5.7: Top: an example of a Raman signal resulting from the sum of 30000 laser shots. The dashed blue line indicate the altitude from where the analysis is performed. Middle: VAOD profile resulting from the analysis of the Raman signal before the evaluation of the integration constant. Bottom: corrected VAOD profile; continous red line is a smoothing obtained using a central running average, dashed blue lines are obtained shifting the smoothed profile by the associated uncertainty.



Fig. 5.8: A cloud is present above the lidar: it appears as a bump in the elastic signal (top) and as a depression in the Raman one (middle). The value of the VAOD increases at the altitude correspondig to the cloud base (bottom)

# Receiver alignment



# The LIDAR signal



# LIDAR data analysis: VAOD(h)

LIDARs provide a VAOD(h) estimate using the multiangle analysis



Auxiliary function S(r) which is the ratio between the Lidar signal at distance r and  $r_n$ 

# **Multiangle analysis**

#### A.Filipcic et al., AstroparticlePhysics 18 (2003)

based on the assumption of a horizontally uniform atmosphere :  $\mathbf{r} = \mathbf{h}/\mathbf{cos}\theta$ 

#### Multiangle analysis strategy

$$S(r) = \ln \frac{P(r)r^2}{P_n r_n^2} = \ln \left[\frac{\beta(r)}{\beta_n}\right] - 2\tau(r, r_n) = \ln \left[\frac{\beta(r)}{\beta_n}\right] - 2\int_{r_n}^r \alpha(r') dr'$$

Hp: the atmosphere is horizontally homogeneous

$$S(h, \sec \theta) = \ln \left[ \frac{\beta(h)}{\beta_n} \right] - 2\tau(h; h_n) \sec \theta$$



#### 6.2. Multi-angle reconstruction

For the ideal atmosphere, with true horizontal invariance, the  $\xi$  dependence of *S*-function is particularly simple,

$$S(h,\xi) = \ln[\beta(h)/\beta_0] - 2\xi\tau(h;h_0),$$
(23)

with the backscatter coefficient  $\ln[\beta/\beta_0]$  as offset, and OD  $\tau$  as the slope of the resulting linear function in  $\xi$ . Therefore, the optical properties of the atmosphere can be alternatively obtained from the analysis of the *S*-function behavior for scanning lidar measurements.

# **The Elastic Lidar**

Elastic scattering on both molecules and aerosols

 $L^{\lambda 0}(s) = L^{\lambda 0}_{0}O(r) \bullet$   $T^{\lambda 0}_{mol}(s) T^{\lambda 0}_{aer}(s) T^{\lambda 0}_{abs}(s) \bullet$   $[\sigma^{\lambda 0}_{mol}(\pi) n_{mol}(s) + \beta^{\lambda 0}_{aer}] d\Omega/4\pi \bullet$   $T^{\lambda 0}_{mol}(s) T^{\lambda 0}_{aer}(s) T^{\lambda 0}_{abs}(s)$ 

Emitting laser intensity

Attenuation (ongoing path)

Backscattering volume

Attenuation (backscattered light)

Assumption needed on atmospheric properties: Lidar Ratio =  $\alpha^{\lambda 0}_{aer}$  /  $\beta^{\lambda 0}_{aer}$ 

 $T^{\lambda}_{aer}(s) = exp(-\int \alpha^{\lambda}_{aer}(s)(ds))$ 

# The Raman Lidar

Raman Scattering : anelastic collision on N2, O2, H<sub>2</sub>O producing a frequency shift of the backscattered photons

$$L^{\lambda i = \lambda 0 + \Delta \lambda i}(s) = L^{\lambda 0}_{0} O(r) \bullet$$
  

$$T^{\lambda 0}_{mol}(s) T^{\lambda 0}_{aer}(s) T^{\lambda 0}_{abs}(s) \bullet$$
  

$$[\sigma^{\lambda i}_{Raman}(\pi) n_{i}(s)] d\Omega/4\pi \bullet$$
  

$$T^{\lambda i}_{mol}(s) T^{\lambda i}_{aer}(s) T^{\lambda i}_{abs}(s)$$

No more  $\beta_{aer}^{\lambda_0} \rightarrow no$  need to make assumptions on the Lidar Ratio, but  $T_{aer}^{\lambda}$  appears at 2 different wavelengths

 $(\alpha^{\lambda_0}_{aer} / \alpha^{\lambda_i}_{aer}) = (\lambda_0 / \lambda_i)^k$ : 2 Raman channels are used to extrapolate k

Uncertainty on aerosol extinction is lower than with the elastic lidar BUT Raman cross section is 3 orders of magnitude lower than elastic  $\rightarrow$  longer acquisition time !