

THE SHiP EXPERIMENT AT CERN



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SHiP

On behalf of the SHiP Collaboration

CERN-SPSC-2015-017
SPSC-P-350-ADD-1
9 April 2015



CERN-SPSC-2015-016
SPSC-P-350
8 April 2015

arXiv:1504.04956v1.

SHiP 240 physicists, 15 Countries

Search for Hidden Particles

Steered west-southwest, and encountered a heavier sea than they had met with before in the whole voyage. Saw particles and a green ruck near the vessel. The crew of the Pinta was a cane and a log, they also picked up a stick which appeared to have been carved with an iron tool, a piece of cane, a plant which grows on land, and a board. The crew of the Niña saw other signs of land, and a stork loaded with rose berries. These signs encouraged them, and they all grew cheerful. Sailed this day till sunset, twenty-seven leagues.

After sunset steered their original course west and called twelve miles an hour till two hours after midnight, going ninety miles, which are twenty-two leagues and a half and as the Pinta was the wisest caller, and kept ahead of the Admiral,

she discovered land



Technical Proposal



SHiP

arXiv:1504.04855v1.

85 theorists

Search for Hidden Particles

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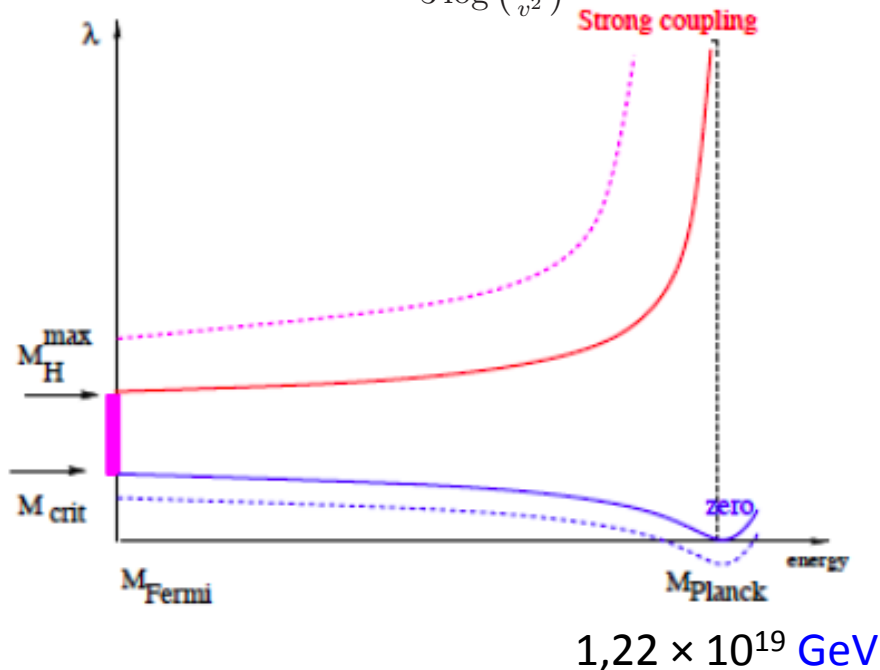
Physics Proposal

SM may well be a consistent effective theory all the way up to the Plank scale

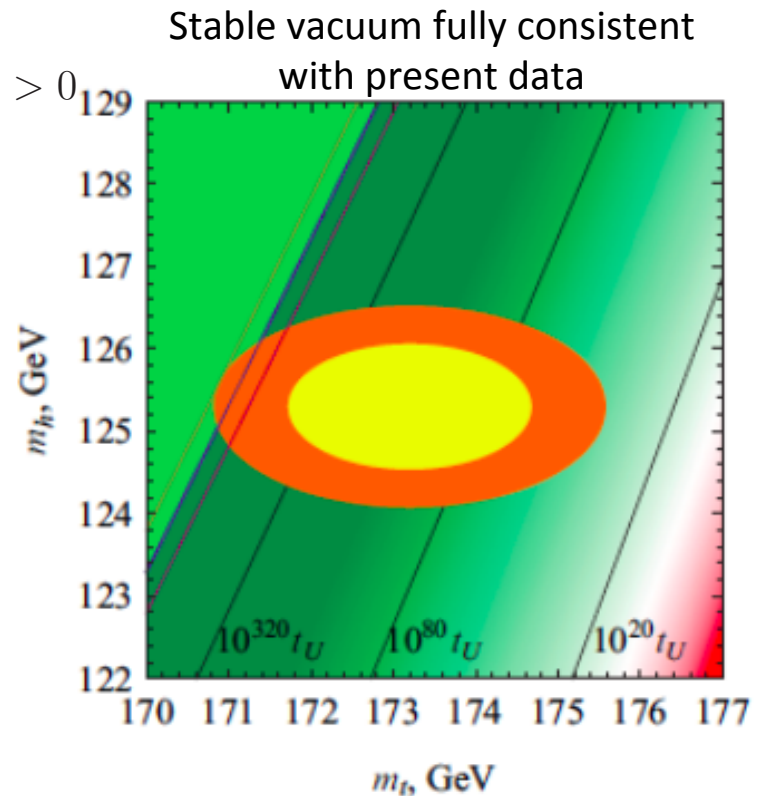
- ✓ $M_H < 175 \text{ GeV} \rightarrow$ SM is a weakly coupled theory up to the Plank energies!
- ✓ $M_H > 111 \text{ GeV} \rightarrow$ EW vacuum is stable or metastable with a lifetime greatly exceeding the age of our Universe (Espinosa et al)

S. Heinemeyer, Higgs Physics, arXiv:1405.3781

$$\lambda(\Lambda) < \infty \Rightarrow M_H^2 \leq \frac{8\pi^2 v^2}{3 \log\left(\frac{\Lambda^2}{v^2}\right)}$$



$$\lambda(\Lambda) > 0$$



G. Degraasi et al., Higgs mass and vacuum stability

- ✓ No sign of New Physics seen FPCapri2016 in the SM at NNLO, JHEP 1208 (2012) 098

Nevertheless, many open questions in particle physics!

Among the most relevant ones:

Why is the Higgs boson so light (so-called “naturalness” or “hierarchy” problem) ?

What is the origin of the matter-antimatter asymmetry in the Universe ?

Why 3 fermion families ? Why do neutral leptons, charged leptons and quarks behave differently ?

What is the origin of neutrino masses and oscillations ?

What is the composition of dark matter (~25% of the Universe) ?

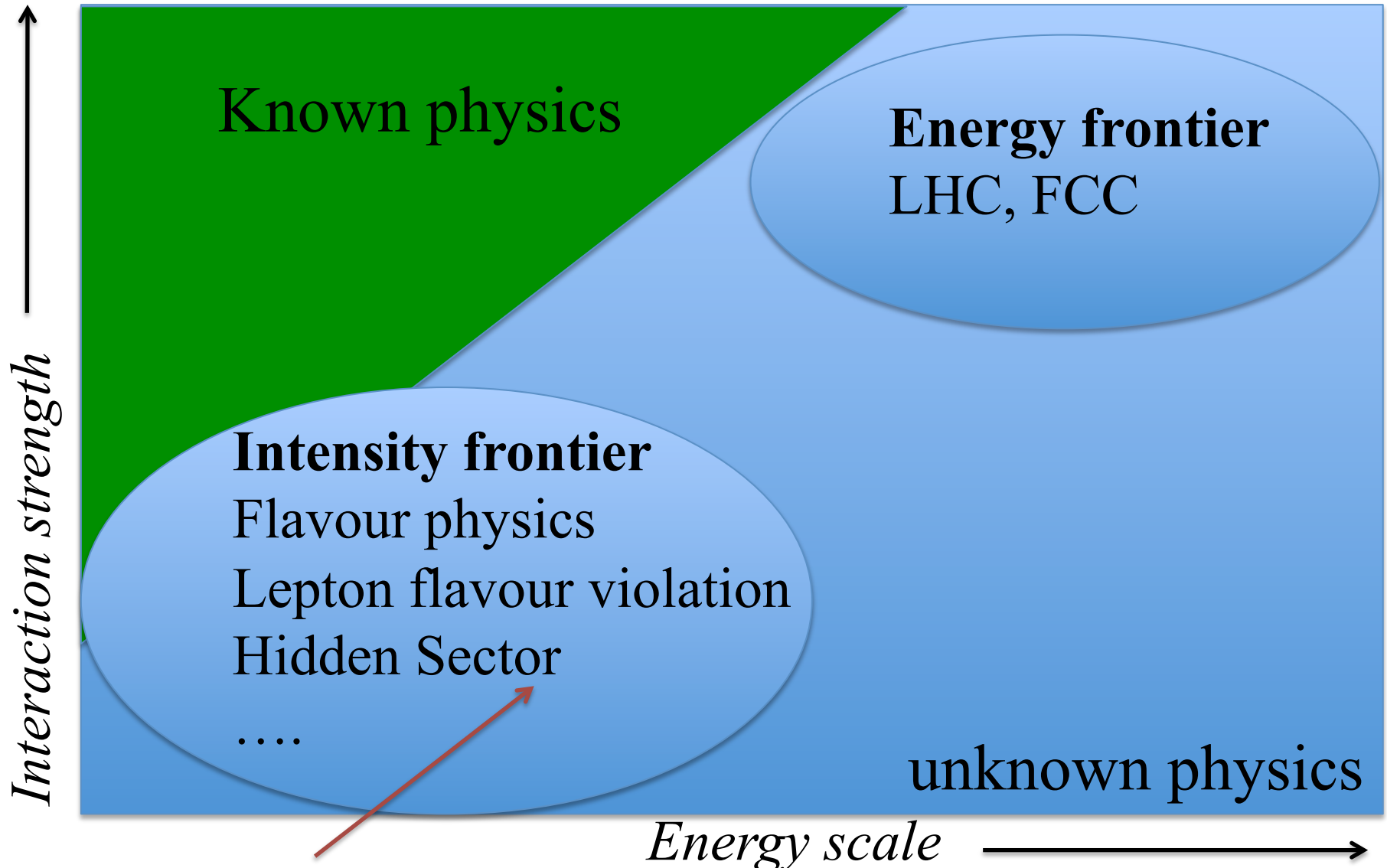


However: there is NO direct evidence for new particles (yet...)
from the LHC or other facilities

Where is the New Physics ?

i.e. at what E scale(s) will we find the answers to these questions ?

High Intensity Frontier



This talk

Energy scale

FPCapri2016

Search for Hidden Sector (HS)

or very weakly interacting NP

$$L = L_{SM} + L_{mediator} + L_{HS}$$

Visible Sector



Mediators or portals to the HS:
vector, scalar, axial, neutrino

Hidden Sector

Naturally accommodates Dark Matter
(may have very complicated structure)

- ✓ HS production and decay rates are strongly suppressed relative to SM
 - Production branching ratios $O(10^{-10})$
 - Long-lived objects
 - Travel unperturbed through ordinary matter

Models	Final states
HNL, SUSY neutralino	$l^+\pi^-, l^+K^-, l^+\rho^- \rho^+ \rightarrow \pi^+\pi^0$
Vector, scalar, axion portals, SUSY sgoldstino	l^+l^-
HNL, SUSY neutralino, axino	$l^+l^-\nu$
Axion portal, SUSY sgoldstino	$\gamma\gamma$
SUSY sgoldstino	$\pi^0\pi^0$

Full reconstruction and PID are essential to minimize model dependence

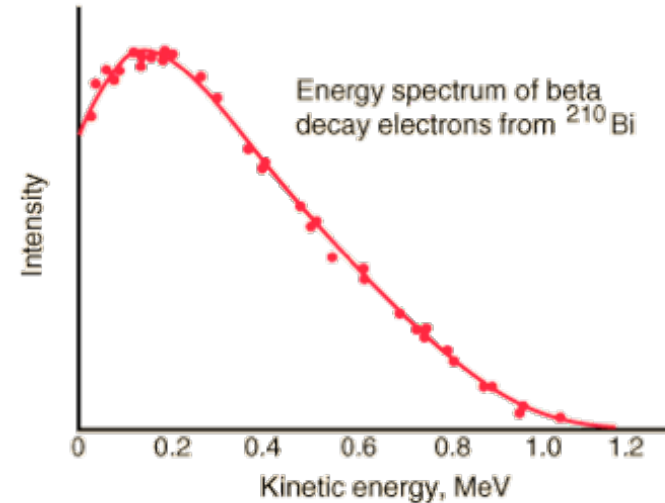
Experimental challenge is background suppression
→ requires $O(0.01)$ carefully estimated

History lesson - 1930s:

- Back then, the “Standard Model” was photon, electron, nucleons

- Beta decay: $n \rightarrow p + e^{-}$

Continuous spectrum!



- Pauli proposes a radical solution - the neutrino!



- Great example of a hidden sector!

- neutrino is electrically neutral (QED gauge singlet)
- very weakly interacting and light
- interacts with “Standard Model” through “portal” -

$$(\bar{p}\gamma^{\mu}n)(\bar{e}\gamma_{\mu}\nu)$$

Search for dark photons

- Assuming no lighter hidden particles, γ' decay into SM particles through a virtual photon:

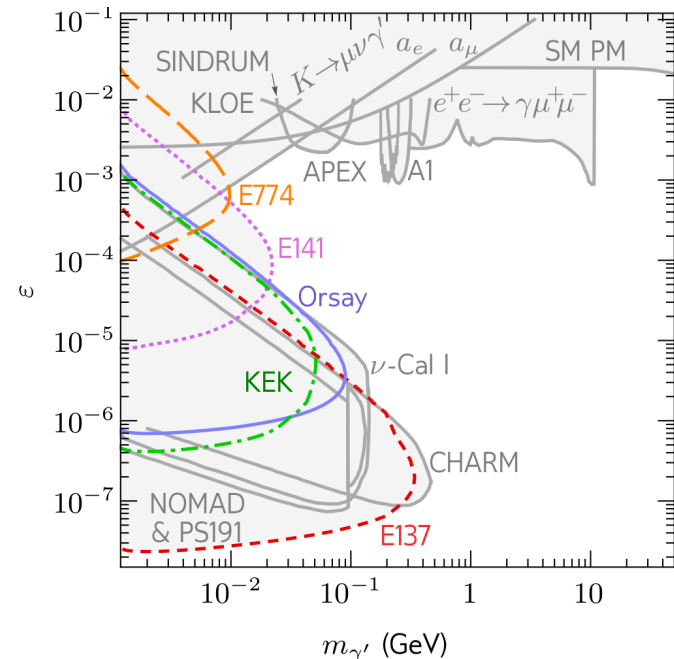
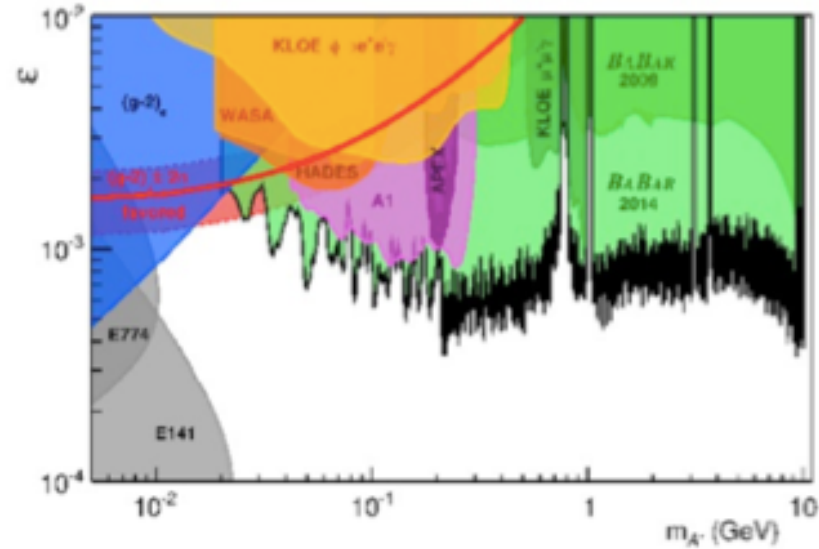
$$\gamma' \rightarrow e^+e^-, \quad \mu^+\mu^-, \quad q\bar{q}, \dots$$

- decay length $c\tau \sim \varepsilon^{-2}m_{\gamma'}^{-1}$
- cosmological constraints (nucleo-synthesis):
 $\tau < 0.1 \text{ s} \Rightarrow \varepsilon^2 m_{\gamma'} > 10^{-21} \text{ GeV}$

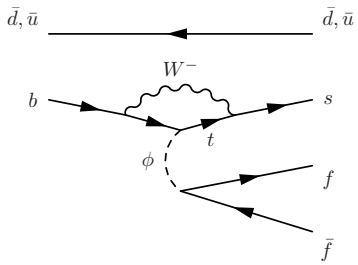
γ' production

- proton bremsstrahlung:
 - initial-state radiation from the incoming proton, followed by a hard proton-nucleus interaction
- secondary particles decay:

Mass interval (GeV)	Process	$n_{\gamma'}/p.o.t$
$m_{\gamma'} < 0.135$	$\pi^0 \rightarrow \gamma\gamma'$	$\varepsilon^2 \times 5.41$
$0.135 < m_{\gamma'} < 0.548$	$\eta \rightarrow \gamma\gamma'$	$\varepsilon^2 \times 0.23$
$0.548 < m_{\gamma'} < 0.648$	$\omega \rightarrow \pi^0\gamma'$	$\varepsilon^2 \times 0.07$
$0.648 < m_{\gamma'} < 0.958$	$\eta' \rightarrow \gamma\gamma'$	$\varepsilon^2 \times 10^{-3}$



Higgs (scalar) portal: production and decay modes

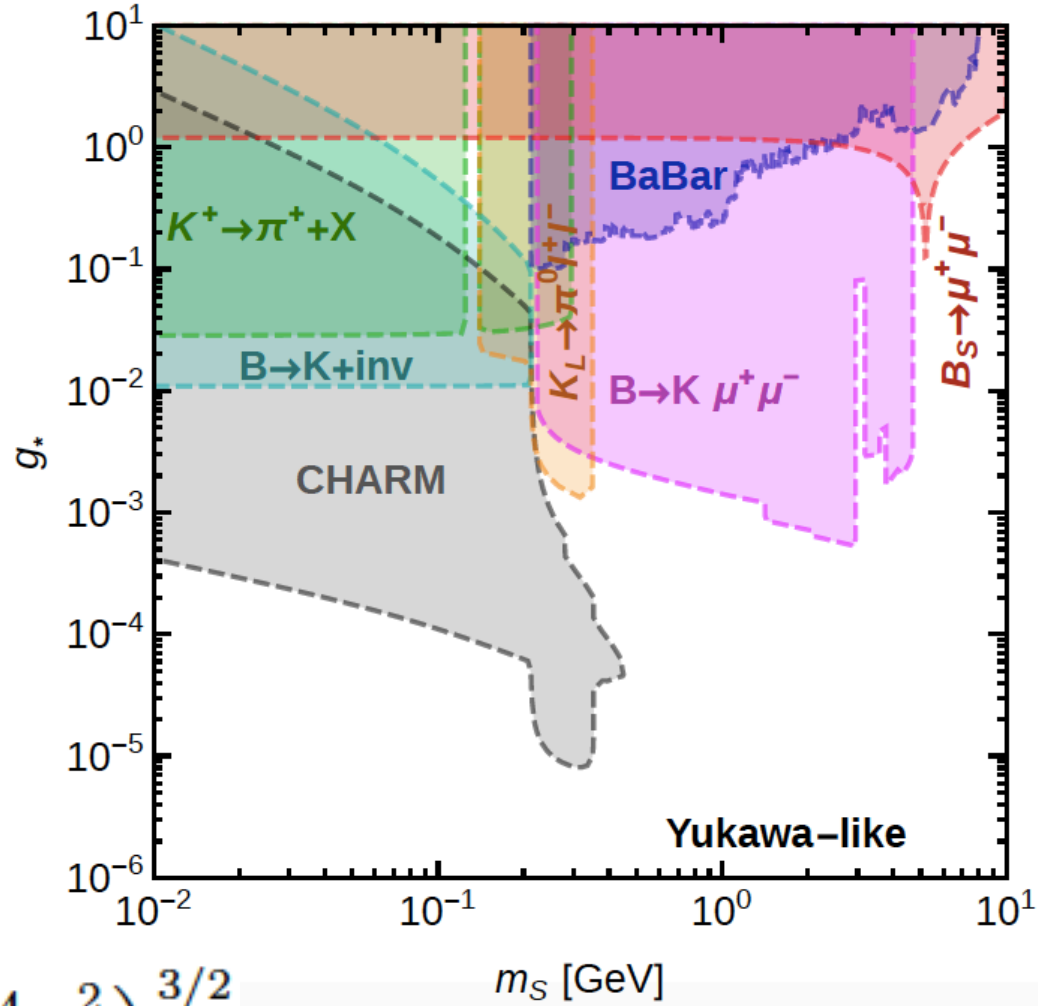


Rare B meson decays mediated by a light scalar ϕ

$$\Gamma(D \rightarrow \pi\phi) \sim (m_b^2 |V_{cb}^* V_{ub}|)^2 \propto m_b^4 \lambda^5$$

$$\Gamma(B \rightarrow K\phi) \sim (m_t^2 |V_{ts}^* V_{tb}|)^2 \propto m_t^4 \lambda^2$$

B decays favoured compared to D



$$\Gamma(S \rightarrow \ell\bar{\ell}) = \frac{g_*^2 m_\ell^2 m_S}{8\pi v^2} \left(1 - \frac{4m_\ell^2}{m_S^2}\right)^{3/2}$$

Motivation for Heavy Neutral Leptons

See-saw mechanism for neutrino masses

Most general renormalisable Lagrangian of SM particles (+3 singlets wrt SM gauge group):

$$L_{\text{singlet}} = i\bar{N}_I \partial_\mu \gamma^\mu N_I - Y_{I\alpha} \bar{N}_I^c \tilde{H} L_\alpha - M_I \bar{N}_I^c N_I + h.c.$$

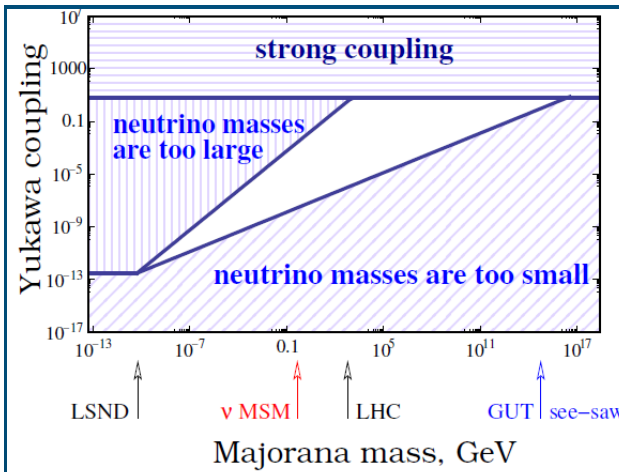
Yukawa term: mixing of N_I with active neutrinos to explain oscillations

Majorana term which carries no gauge charge

$$v \sim 246 \text{ GeV}$$

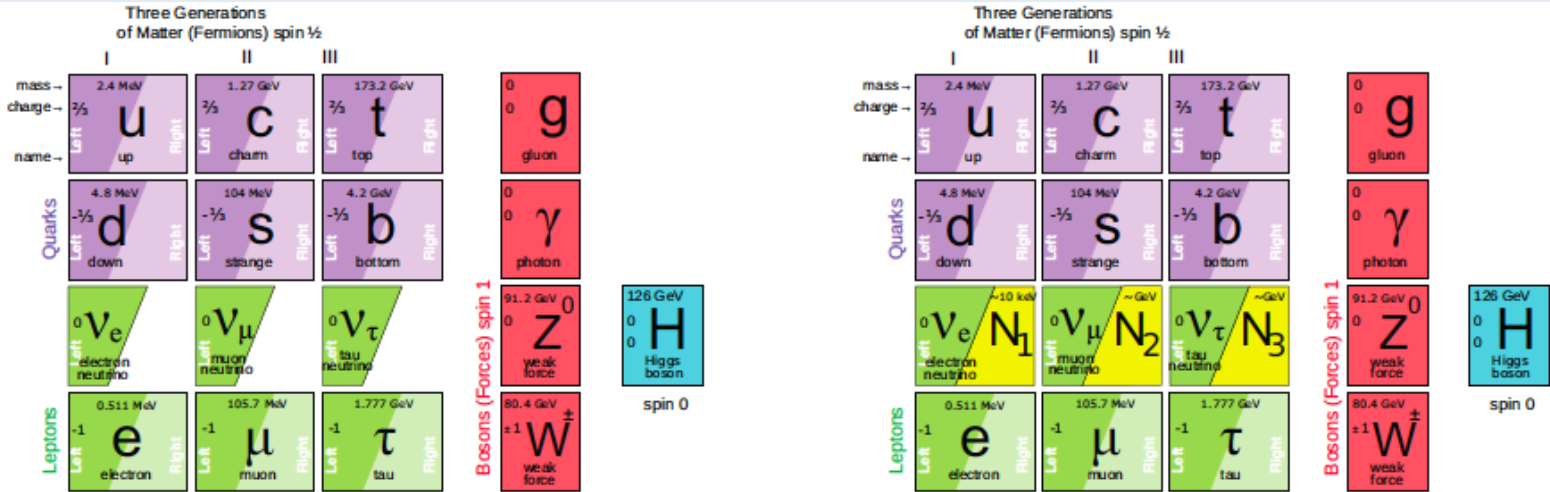
The scale of the active neutrino mass is given by the see-saw formula: $m_\nu \sim \frac{m_D^2}{M}$ where $m_D \sim Y_{I\alpha} v$ - typical value of the Dirac mass term

Four “popular” N mass ranges



	N mass	ν masses	eV ν anomalies	BAU	DM	M_H stability	direct search	experiment
GUT see-saw	10^{-16} - 10 GeV	YES	NO	YES	NO	NO	NO	-
EWSB	10^{-2} - 10 GeV	YES	NO	YES	NO	YES	YES	LHC
ν MSM	keV - GeV	YES	NO	YES	YES	YES	YES	a'la CHARM
ν scale	eV	YES	YES	NO	NO	YES	YES	a'la LSND

The ν MSM model: leptogenesis and dark matter



N = Heavy Neutral Lepton - HNL

Role of N_1 with mass in keV region: dark matter

Role of N_2 , N_3 with mass in 100 MeV – GeV region: “give” masses to neutrinos and produce baryon asymmetry of the Universe

Role of the Higgs: give masses to quarks, leptons, Z and W and inflate the Universe.

**ν MSM: T.Asaka, M.Shaposhnikov PL B620 (2005) 17
M.Shaposhnikov Nucl. Phys. B763 (2007) 49**

global lepton-number symmetry broken at the level of $O(10^{-4})$ leads to the required pattern of sterile neutrino masses consistent with neutrino oscillations data

Masses and couplings of HNLs

- $M(N_2) \approx M(N_3) \sim$ a few GeV \rightarrow CPV can be increased dramatically to explain **Baryon Asymmetry of the Universe (BAU)**

Very weak $N_{2,3}$ -to- ν mixing ($\sim U^2$) $\rightarrow N_{2,3}$ are much longer-lived than SM particles

- Produced in semi-leptonic decays,

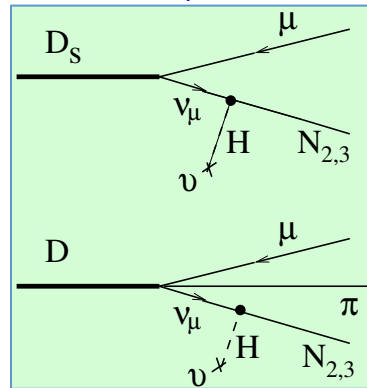
$$K \rightarrow \mu\nu, D \rightarrow \mu\pi\nu, B \rightarrow D\mu\nu$$

- $\propto \sigma_D \times U^2$

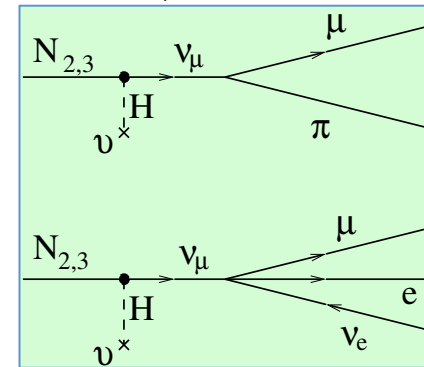
- $U_2^2 = U_{2,\nu_e}^2 + U_{2,\nu_\mu}^2 + U_{2,\nu_\tau}^2$

Example:

$N_{2,3}$ production in charm



and subsequent decays



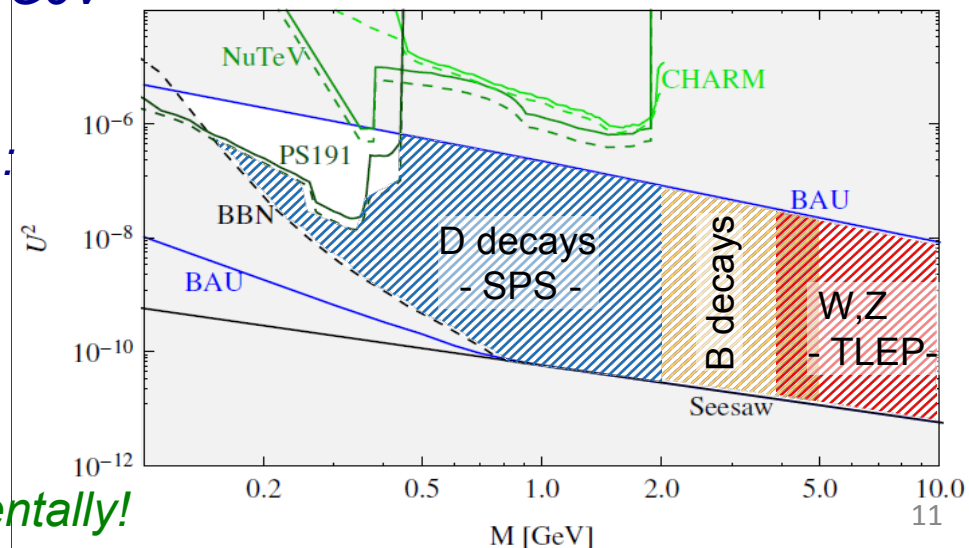
- Typical lifetimes $> 10 \mu\text{s}$ for $M(N_{2,3}) \sim 1 \text{ GeV}$
Decay distance O(km)

- Typical BRs (depend on flavour mixing):

$$Br(N \rightarrow \mu/e \pi) \sim 0.1 - 50\%$$

$$Br(N \rightarrow \mu/e^- \rho^+) \sim 0.5 - 20\%$$

$$Br(N \rightarrow \nu\mu e) \sim 1 - 10\%$$



Domain only marginally explored, experimentally!



Common experimental features of Hidden Sector (HS)

✓ Production through hadron decays (π , K , D , B , proton bremsstrahlung, ...)

✓ Decays:

Models	Final states
Neutrino portal, SUSY neutralino	$l^\pm \pi^\mp, l^\pm K^\mp, l^\pm \rho^\mp, \rho^\pm \rightarrow \pi^\pm \pi^0$
Vector, scalar, axion portals, SUSY sgoldstino	$l^+ l^-$
Vector, scalar, axion portals, SUSY sgoldstino	$\pi^+ \pi^-, K^+ K^-$
Neutrino portal, SUSY neutralino, axino	$l^+ l^- \nu$
Axion portal, SUSY sgoldstino	$\gamma\gamma$
SUSY sgoldstino	$\pi^0 \pi^0$

✓ Full reconstruction and PID are essential to minimize model dependence

✓ Production and decay rates are strongly suppressed when compared to SM

- Production branching ratios $O(10^{-10})$

- Long-lived objects

- Travel unperturbed through ordinary matter

✓ **Challenge is background suppression \rightarrow requires $O(0.01)$ carefully estimated**

✓ **Physics with ν_τ produced in D_s decays share many of these features**

ν_τ STUDIES

- Less known particle in the Standard Model
- **First observation** by DONUT at Fermilab in 2001 with 4 detected candidates, *Phys. Lett. B504 (2001) 218-224*
- 9 events (with an estimated background of 1.5) reported in 2008 with looser cuts

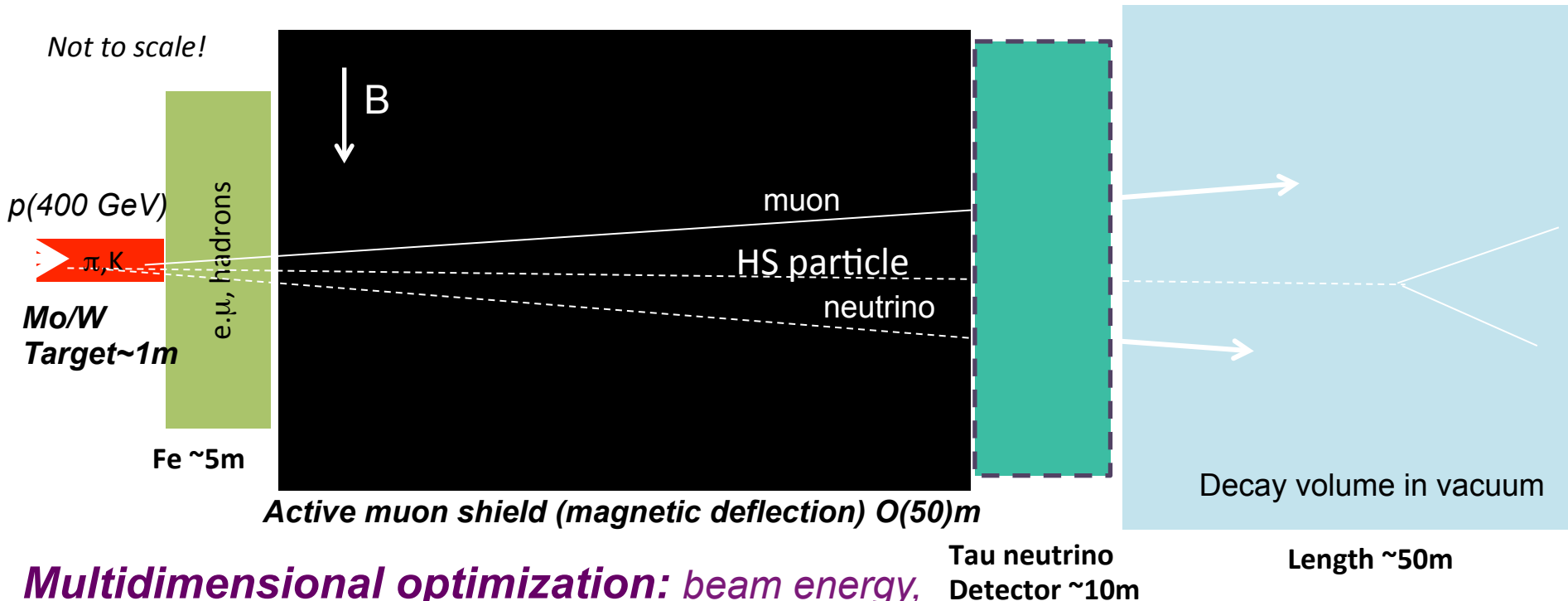
$$\sigma^{\text{const}}(\nu_\tau) = (0.39 \pm 0.13 \pm 0.13) \times 10^{-38} \text{ cm}^2 \text{ GeV}^{-1}$$

- 5 ν_τ candidates reported by OPERA for the discovery (5.1 σ result) of **ν_τ appearance** in the CNGS neutrino beam PRL 115 (2015) 121802
- Tau anti-neutrino never observed

General experimental requirements

Initial reduction of beam induced backgrounds

- Heavy target to maximize Heavy Flavour production (large A) and minimize production of neutrinos in $\pi/K \rightarrow \mu\nu$ decays (short λ_{int})
- Hadron absorber
- Effective muon shield (without shield: muon rate $\sim 10^{10}$ per spill of 4×10^{13} pot)
- Slow (and uniform) beam extraction $\sim 1s$ to reduce occupancy in the detector

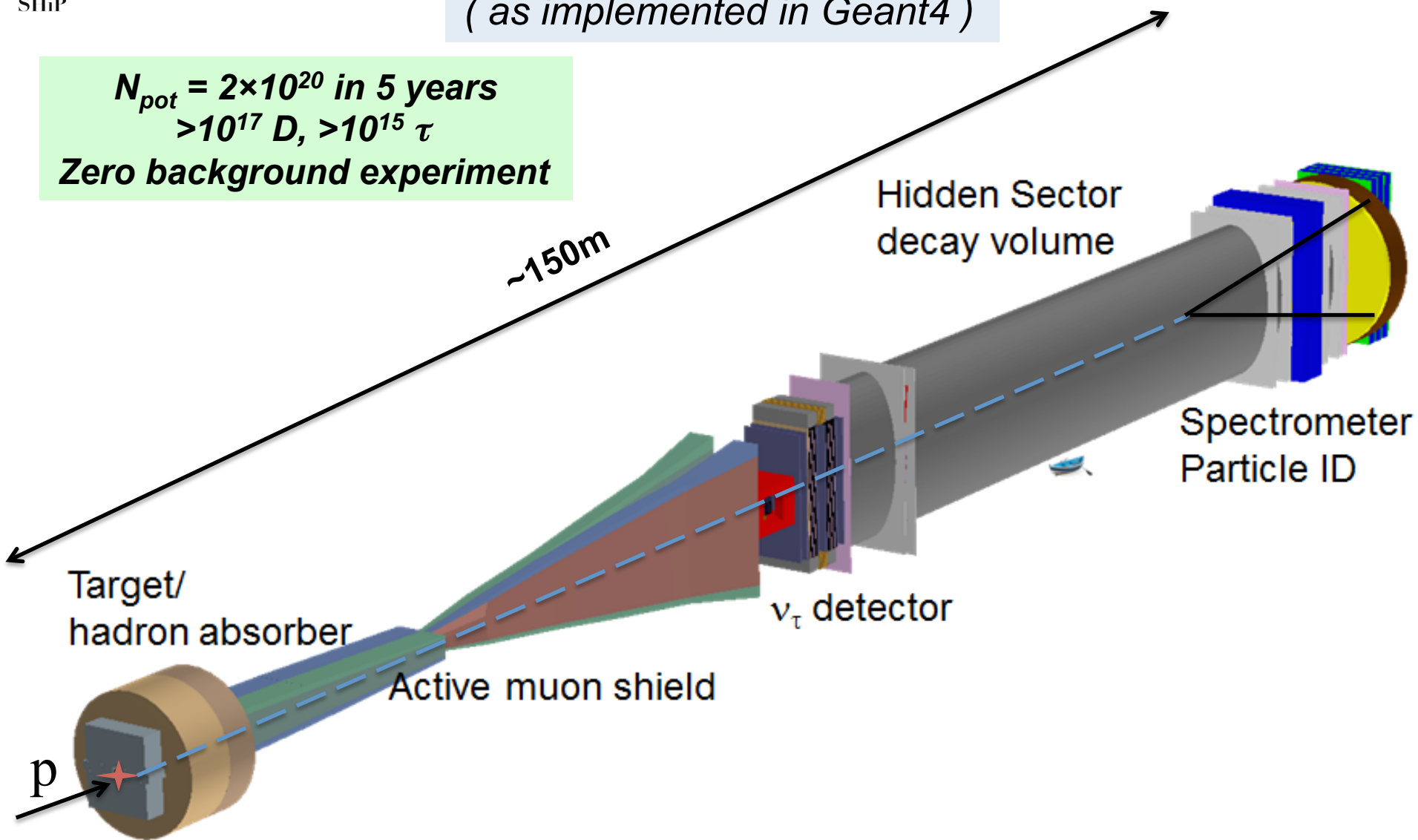


Multidimensional optimization: beam energy, beam intensity, background conditions and detector acceptance



The SHiP experiment (as implemented in Geant4)

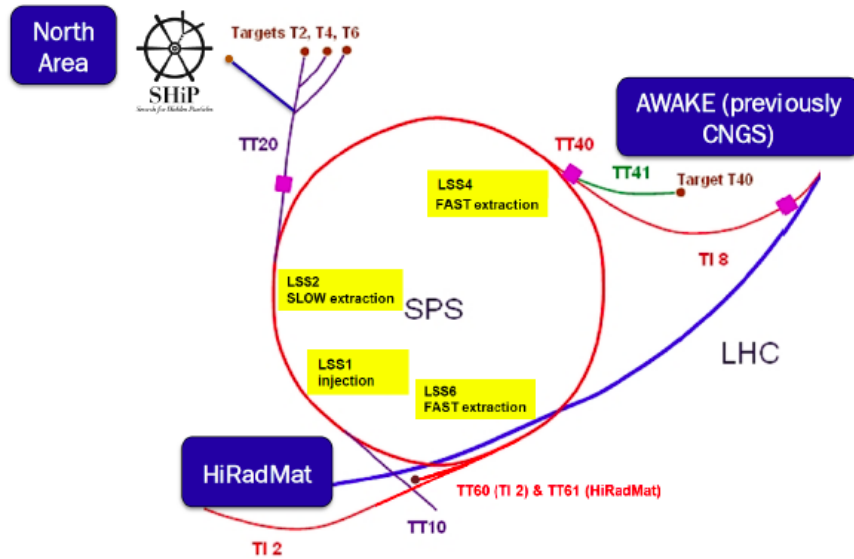
$N_{pot} = 2 \times 10^{20}$ in 5 years
 $> 10^{17} D, > 10^{15} \tau$
Zero background experiment





The Fixed-target facility at the SPS: Preveessin North Area site

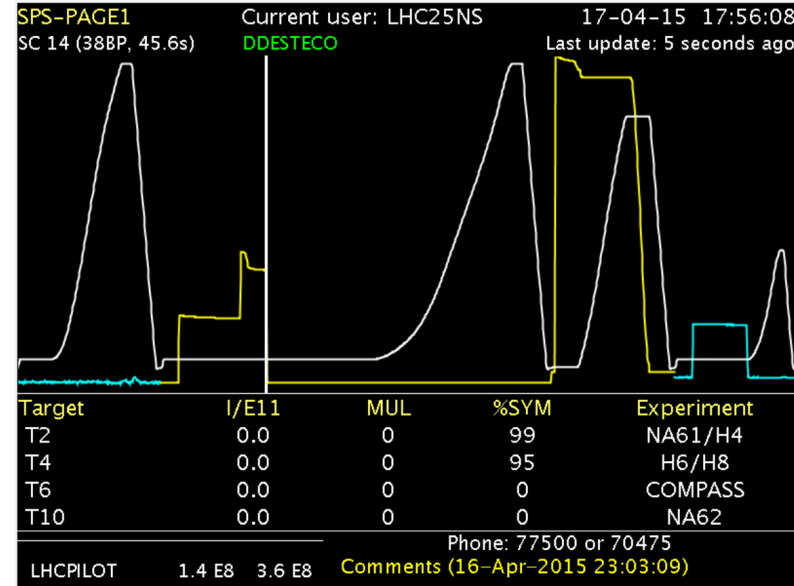
Proposed implementation based on minimal modification of the SPS complex
High-intensity proton beam: $4 \cdot 10^{13}$ ppp, $4 \cdot 10^{19}$ pot/yr, 5 years run (as for CNGS)



The SHiP facility is located on the North Area, and shares the TT20 transfer line and slow extraction mode with the fixed target programmes

R&D at CERN for extraction and beam lines

- Deployment of the new SHiP cycle
- Extraction loss characterisation and optimisation
 - Reduce p density on septum wires
 - Probe SPS aperture limits during slow extraction
- Development of new TT20 optics
 - Change beam at splitter on cycle-to cycle basis
- Characterisation of spill structure
- R&D and development of laminated splitter and dilution (sweep) magnets

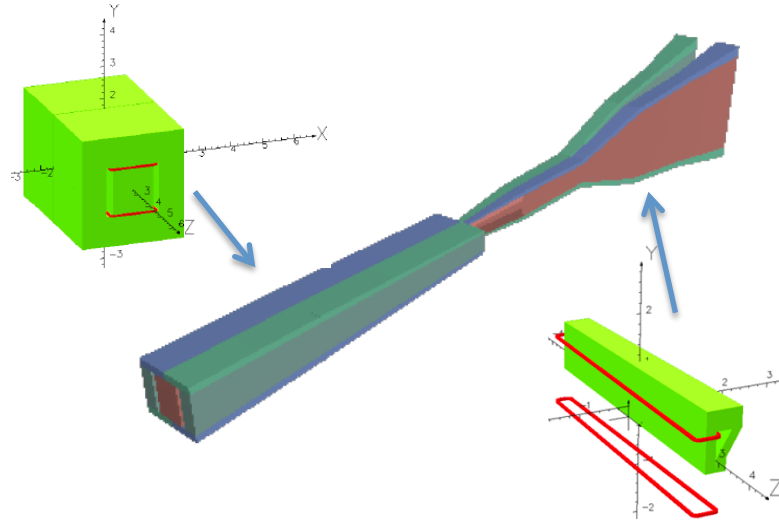


Successful test in April 2015

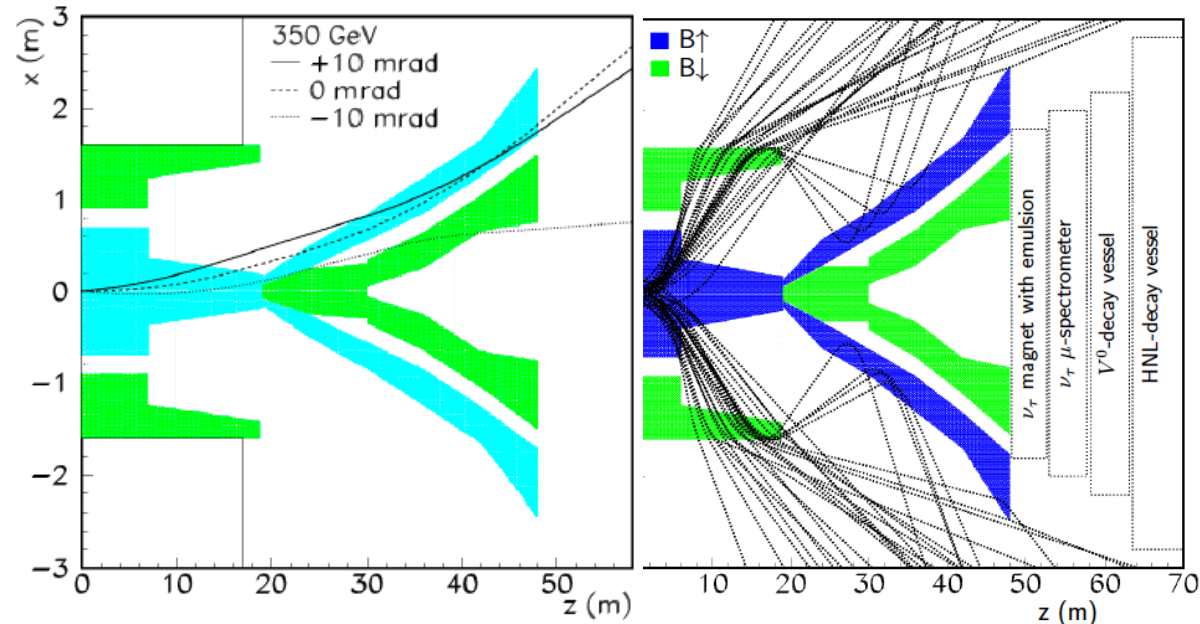


SHiP muon shield

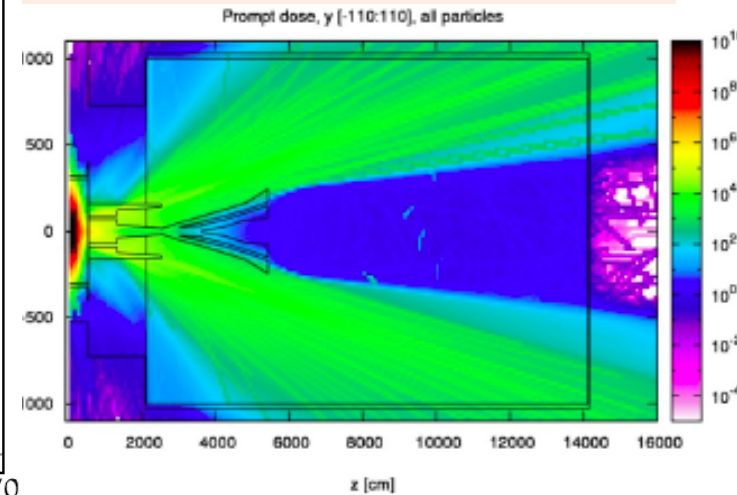
- ✓ Muon flux limit driven by HS background and emulsion-based neutrino detector
- ✓ Active muon shield based entirely on magnet sweeper with a total field integral $B_y = 86.4 \text{ Tm}$
- Realistic design of sweeper magnets in progress
- Challenges: flux leakage, constant field profile, modeling magnet shape
- ✓ $< 7\text{k muons / spill } (E_\mu > 3 \text{ GeV}), \text{ from } 10^{10}$
- ✓ Negligible flux in terms of detector occupancy



Magnetic sweeper field



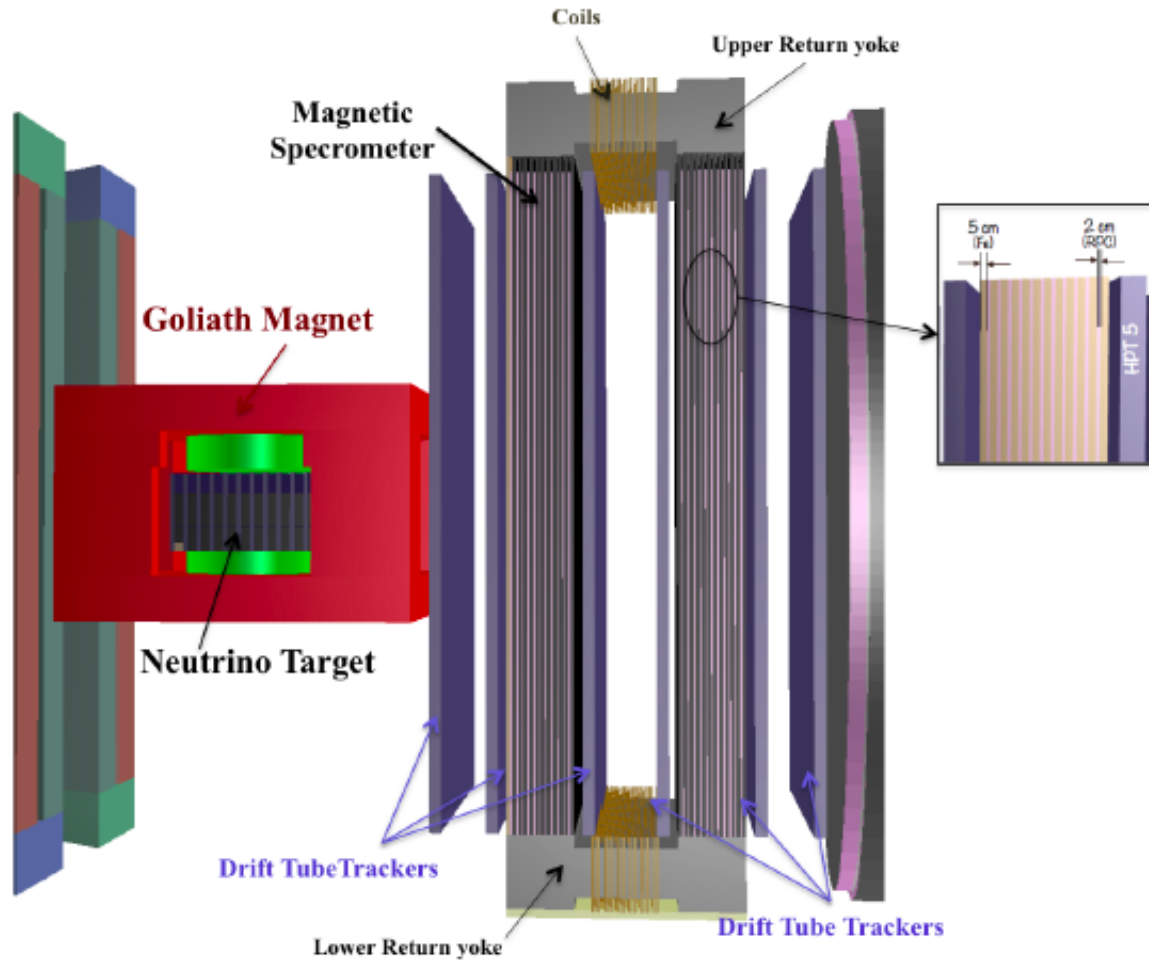
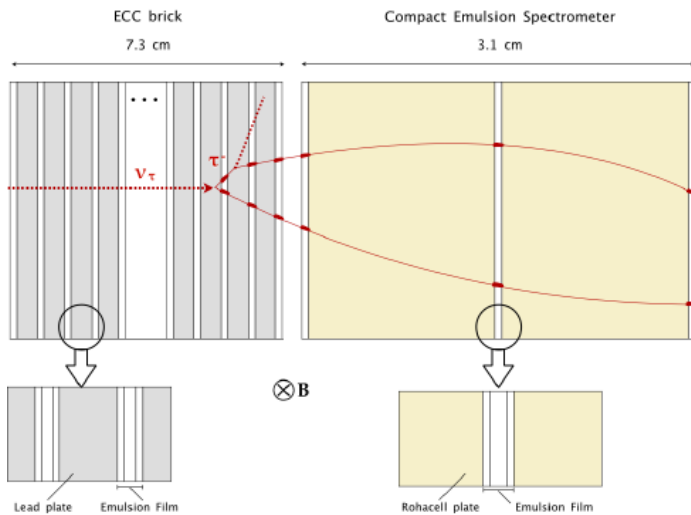
Dose rate ($\mu\text{Sv/h}$) in the SHiP hall



ν_τ detector follows the OPERA concept

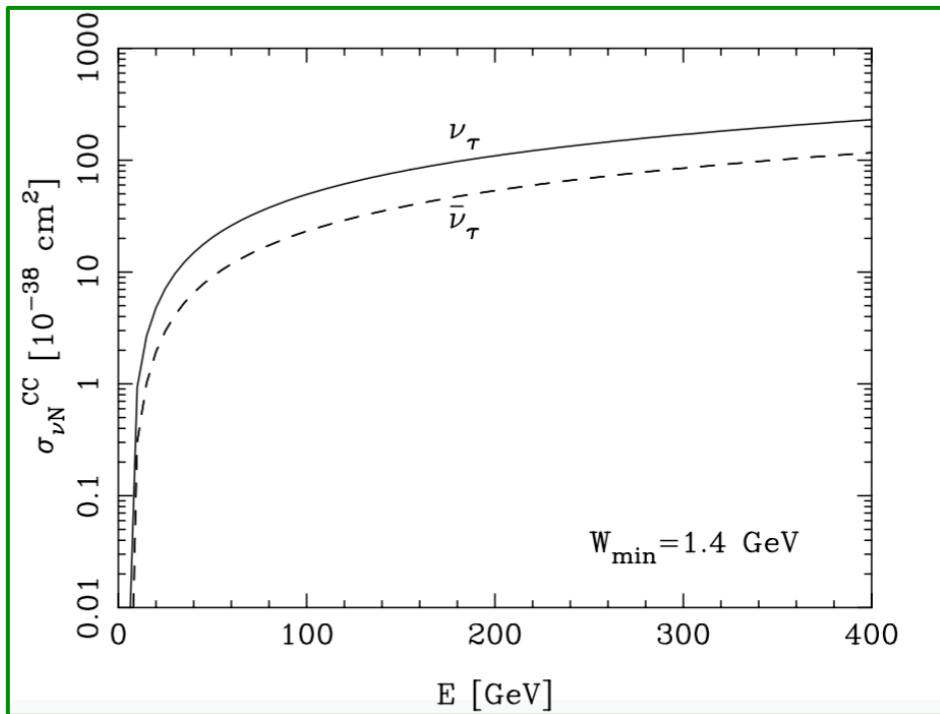


Emulsion Cloud Chamber is the key element of ν_τ detection



ν_τ INTERACTIONS IN THE TARGET

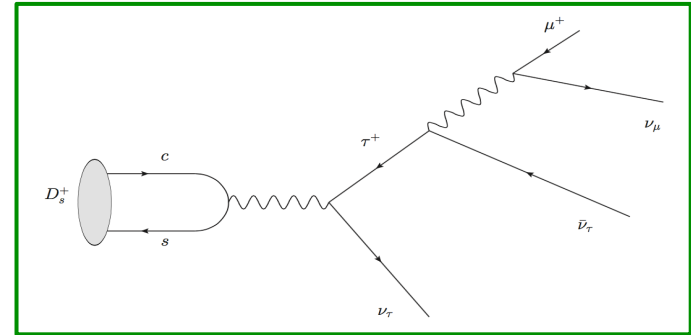
$$N_{\nu_\tau + \bar{\nu}_\tau} = 4N_p \frac{\sigma_{c\bar{c}}}{\sigma_{pN}} f_{D_s} Br(D_s \rightarrow \tau) = 2.85 \times 10^{-5} N_p = 6.6 \times 10^{15}$$



M. H. Reno, Phys. Rev. D74 (2006) 033001

Uncertainty ($\lesssim 10\%$) from:

- Scale choices
- Pdf
- Target mass correction



Expected number of interactions*

*in 5 years run (2×10^{20} pot)
target mass ~ 9.6 ton (Pb)

$$N_{\nu_\tau} \simeq 7.6 \times 10^3$$

$$N_{\bar{\nu}_\tau} \simeq 3.9 \times 10^3$$

20% uncertainty mainly
from scale variations in
c-cbar differential cross-section

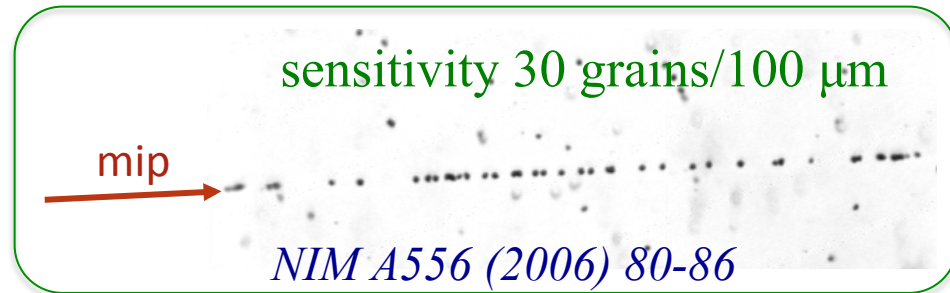
ν_τ DETECTOR

THE UNITARY CELL

Emulsion Cloud Chamber (ECC)

BRICK

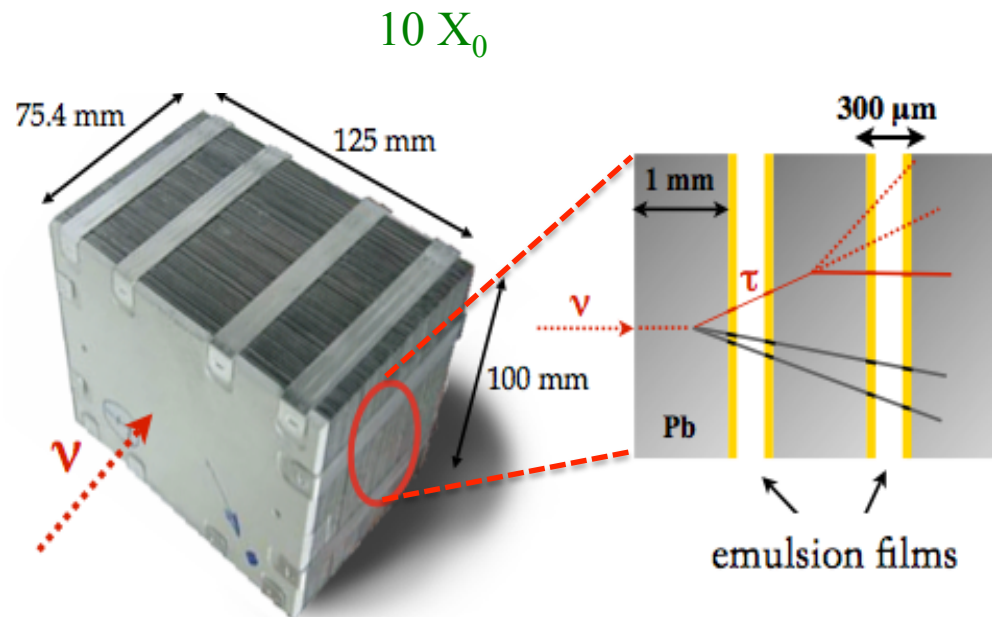
- passive material \rightarrow lead
(*massive target*)
- tracking device \rightarrow nuclear emulsions
(*high resolution*)



PERFORMANCES

- Primary and secondary **vertex definition** with μm resolution
- **Momentum measurement** by Multiple Coulomb Scattering
 - largely exploited in the OPERA experiment
- **Electron identification**: shower ID through calorimetric technique

OPERA: 1 event in 1 brick
SHIP: ~ 230 events/brick



ν_τ /ANTI- ν_τ SEPARATION

THE COMPACT EMULSION SPECTROMETER

Magnetised target \rightarrow charge and momentum measurement for hadrons

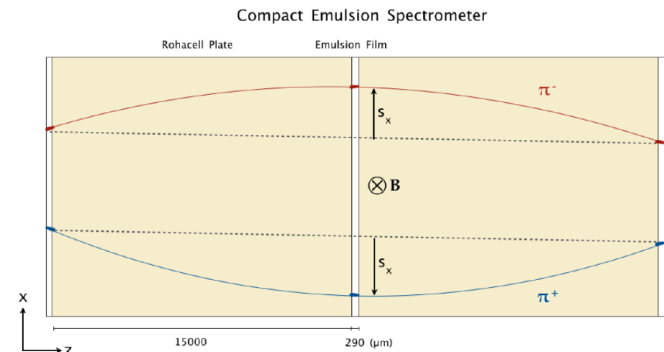
BR($\tau \rightarrow$ hadrons) \sim 65%

Use Compact Emulsion Spectrometer (CES) \rightarrow R&D going on

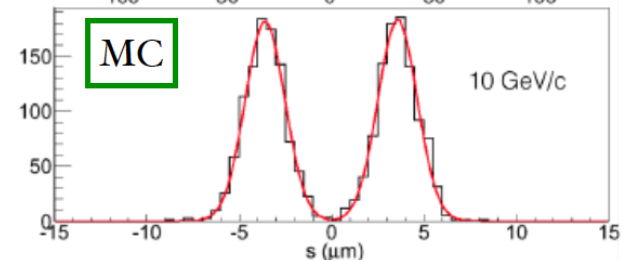
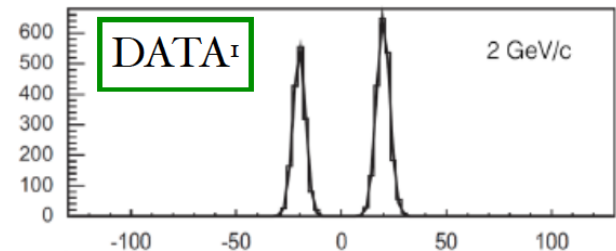
- 1T field
- 3 films interleaved with 2 Rohacell layers (15 mm)
- Thin chamber: 3cm in total
- 90% efficiency for hadronic τ daughters reaching the CES
- Sagitta to discriminate between positive and negative charge

Performances to be achieved

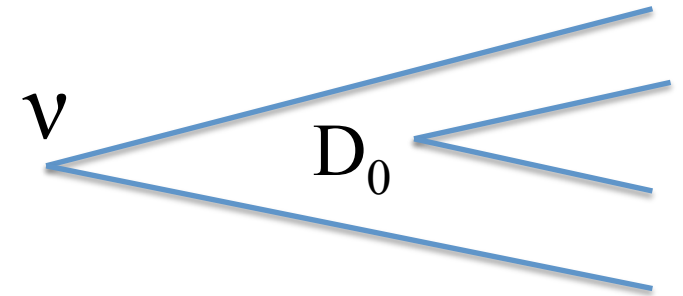
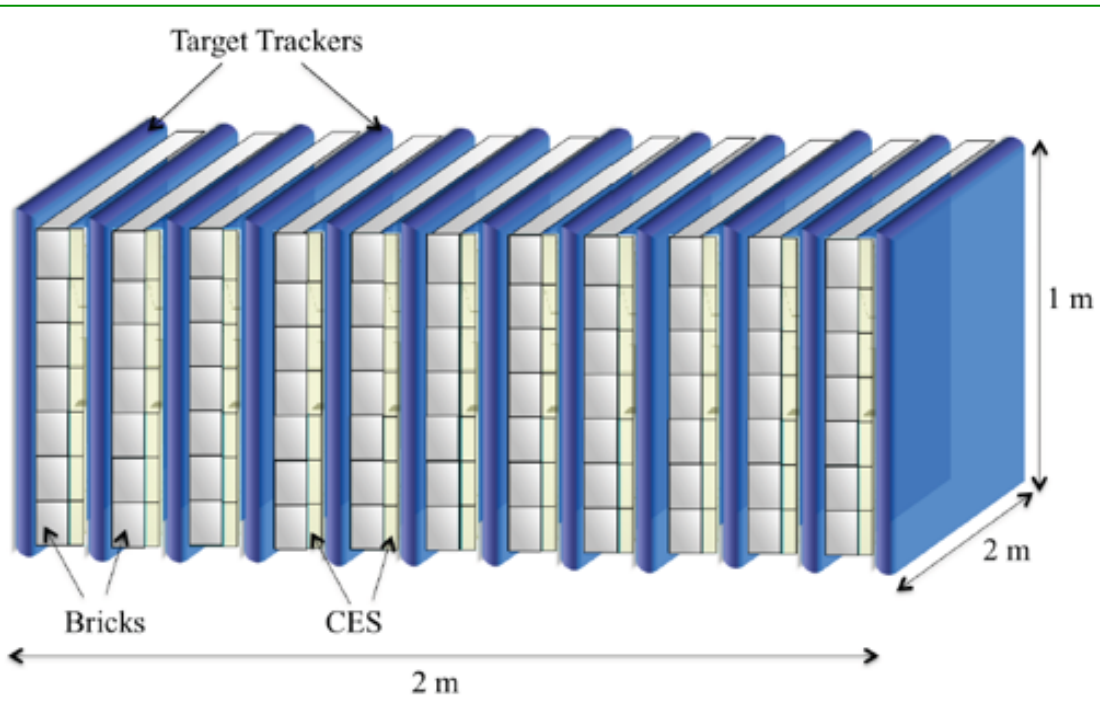
- charge measured up to 10 GeV/c (3 sigma level)
- $\Delta p/p < 20\%$ up to 12 GeV/c



NIM A 592 (2008) 56–62



THE TARGET TRACKER



- 12 target tracker (TT) planes interleaving the 11 brick walls
- first TT plane used as veto
- Transverse size $\sim 2 \times 1 \text{ m}^2$

FEATURES

- Provide time stamp
- Link muon track information from the target to the magnetic spectrometer

REQUIREMENTS

- Operate in 1T field
- X-Y position resolution $< 100 \mu\text{m}$
- high efficiency ($>99\%$) for angles up to 1 rad

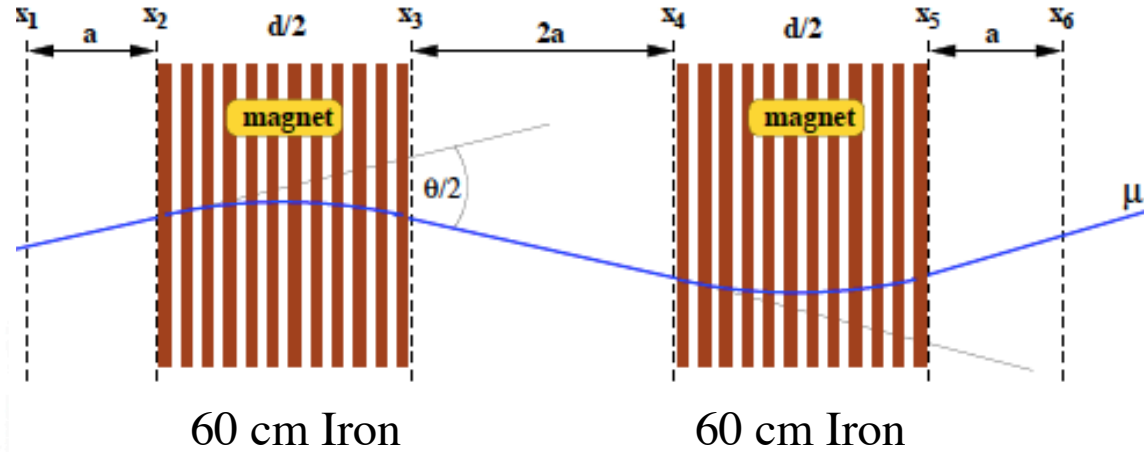
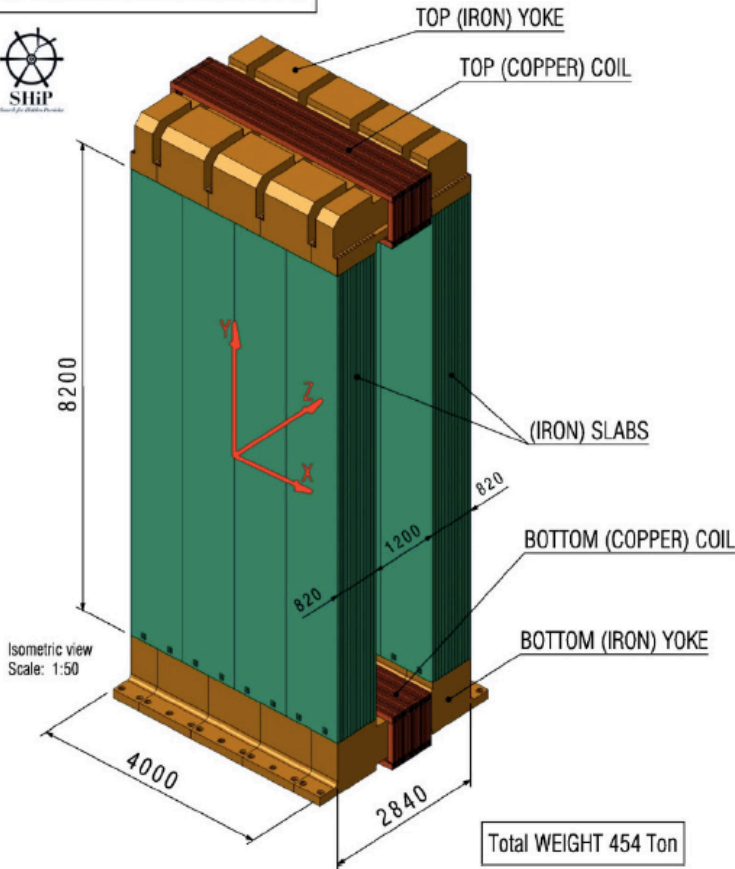
TARGET TRACKER PLANES

POSSIBLE OPTIONS

- Scintillating fibre trackers
- Micro-pattern gas detectors (GEM, Micromegas)

Tracking stations inside the magnet

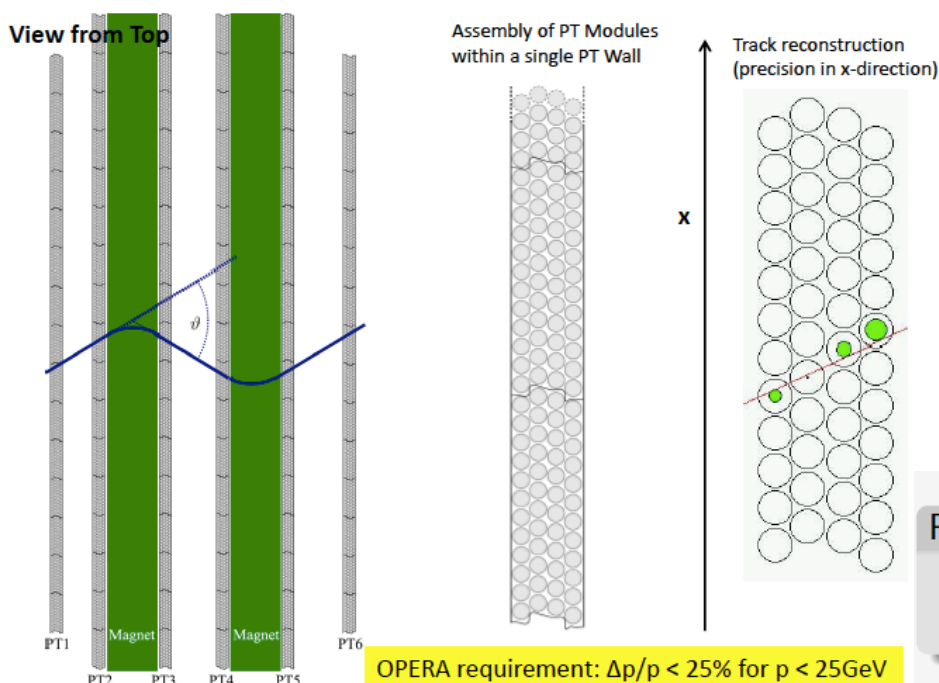
SPECTROMETER MAGNET



- Requirements:
 - coarse (1 cm) tracking inside the magnetised volume
 - 1ns time resolution
 - Muon rate $\sim 5\text{kHz}/\text{m}^2$ rate
 - Electron rate ~ 1 order higher
- RPC technology is one option
- Streamer versus avalanche to be studied

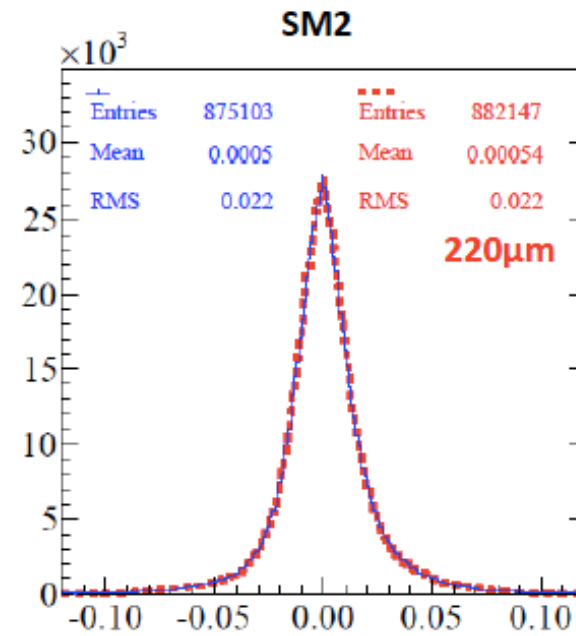
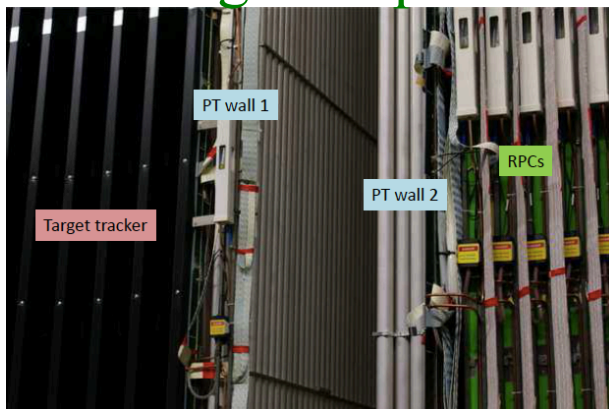


Muon momentum measurement and identification



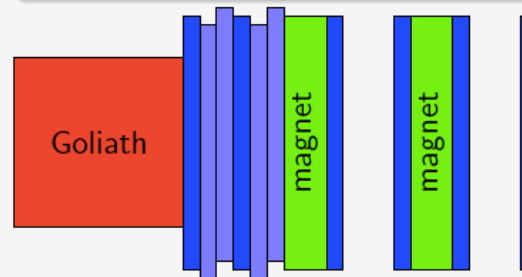
OPERA requirement: $\Delta p/p < 25\%$ for $p < 25\text{GeV}$

OPERA-like drift tubes are a good option

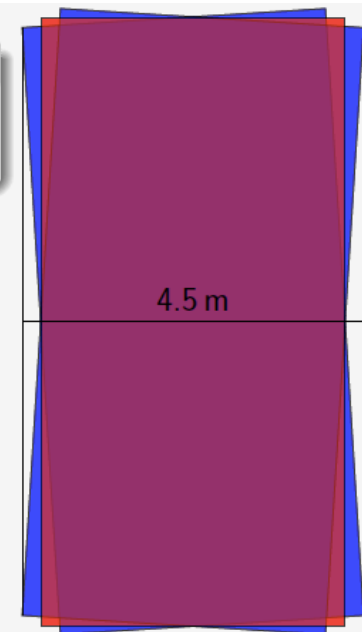


Requirements for 3d reconstruction

- Stereo angle between planes
- 3 projections to avoid ambiguities



- 2×3 projections
- 2 projections rotated by $\pm 3.6^\circ$
- maximal width: 4.5 m

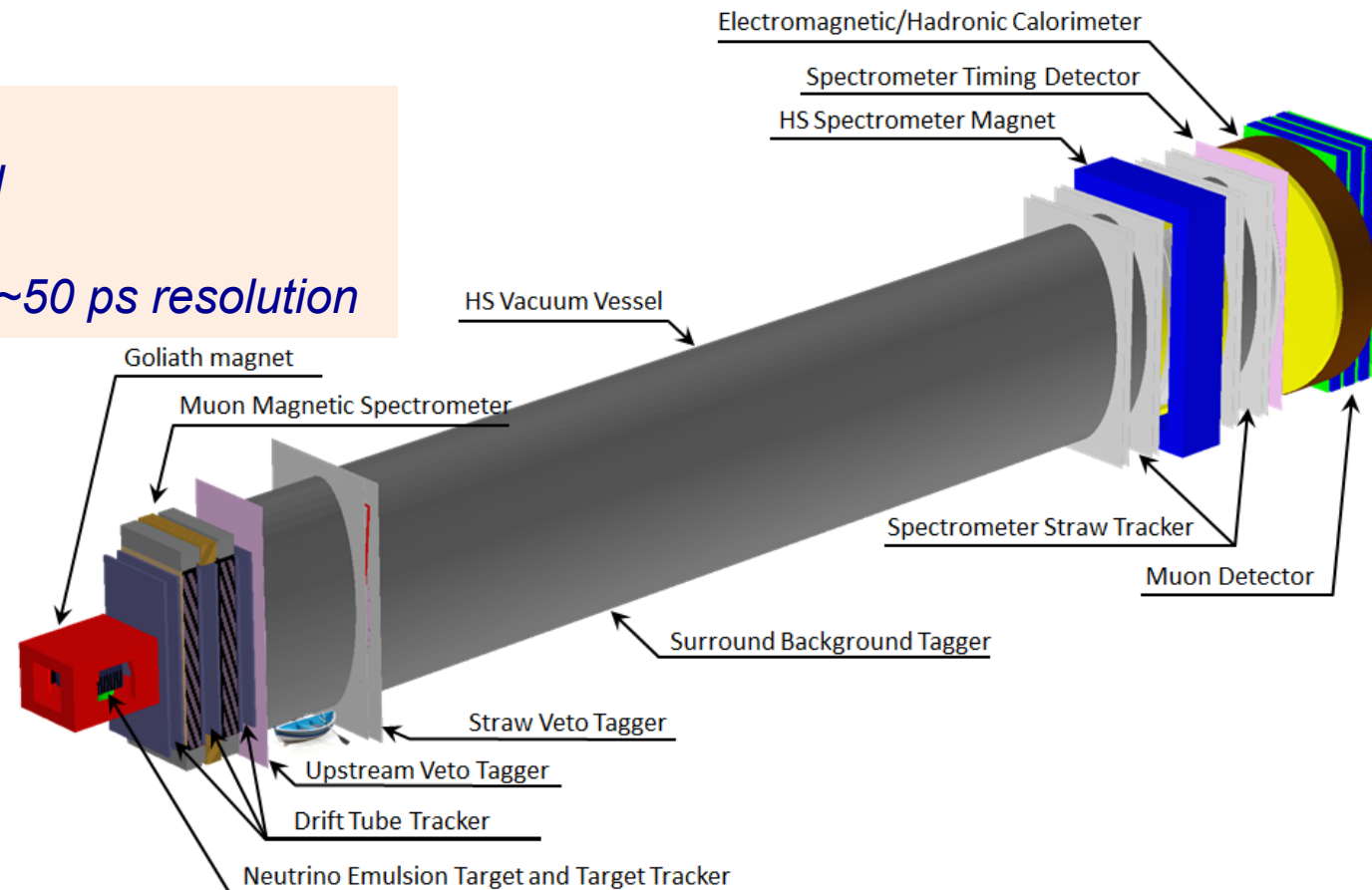


Hidden Sector detector concept

- ✓ *Reconstruction of HS decays in all possible final states*
Long decay volume protected by various Veto Taggers, Magnetic Spectrometer followed by the Timing Detector, and Calorimeters and Muon systems.
All heavy infrastructure is at distance to reduce neutrino / muon interactions in proximity of the detector

Challenges:

- Large vacuum vessel
- 5 m long straw tubes
- Timing detector with ~ 50 ps resolution

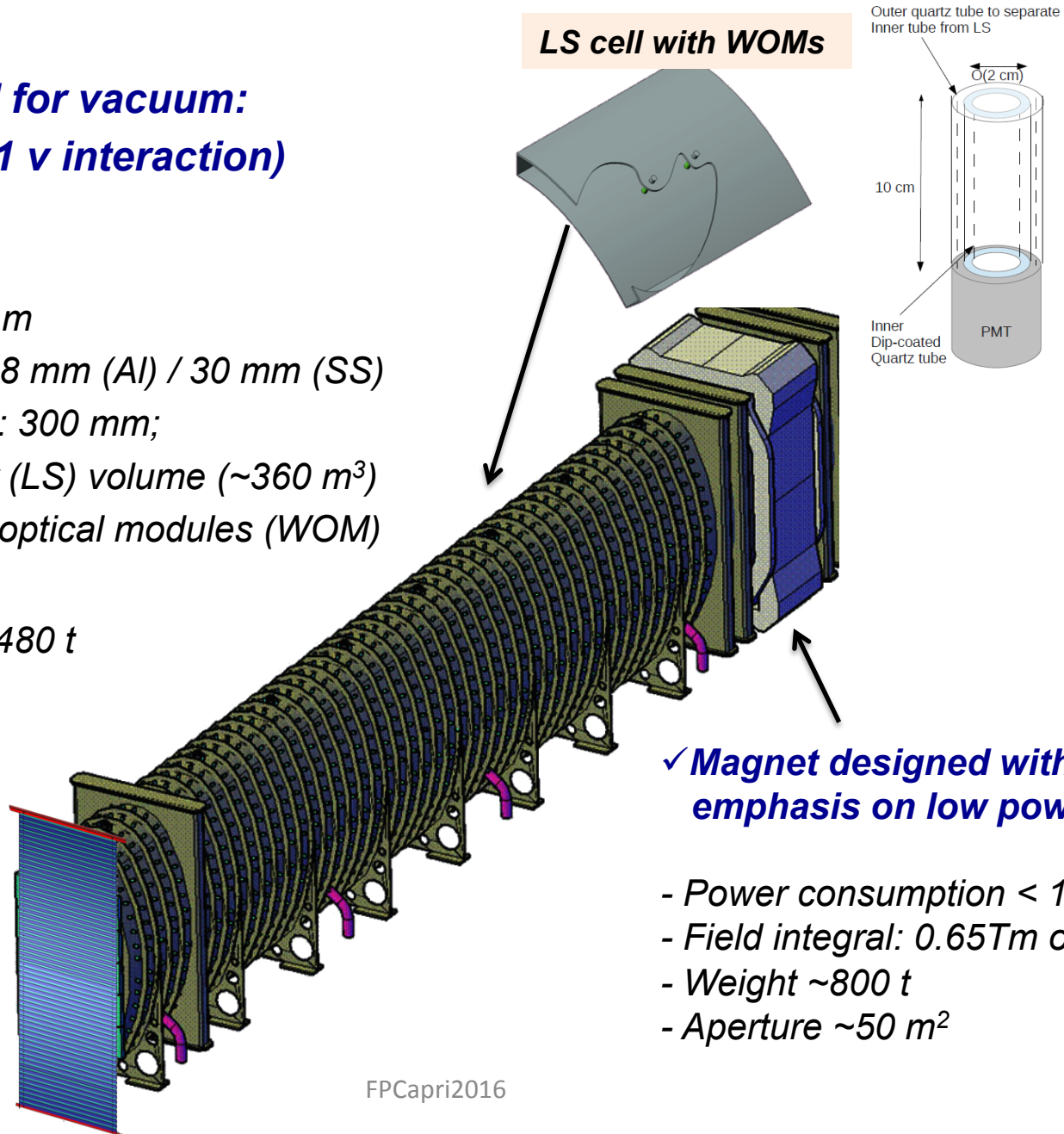


HS decay volume and spectrometer magnet

✓ **Estimated need for vacuum:**
 $\sim 10^{-3}$ mbar (<1 v interaction)

✓ **Vacuum vessel**

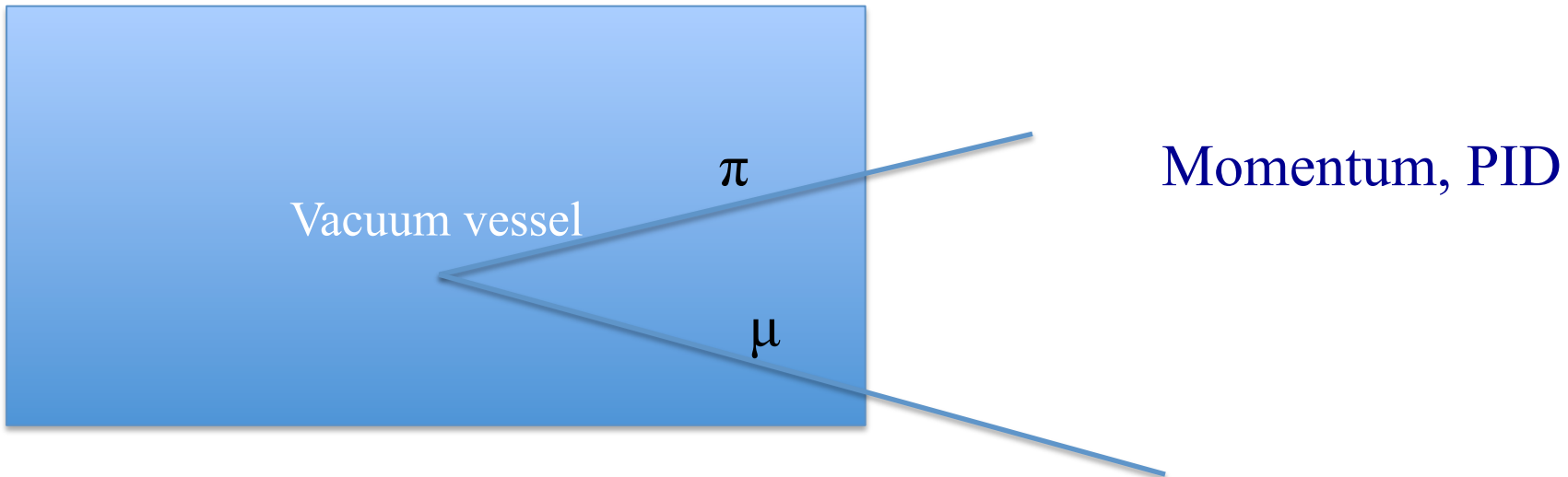
- 10 m x 5 m x 60 m
- Walls thickness: 8 mm (Al) / 30 mm (SS)
- Walls separation: 300 mm;
- Liquid scintillator (LS) volume (~ 360 m³) readout by WLS optical modules (WOM) and PMTs
- Vessel weight ~ 480 t



✓ **Magnet designed with an emphasis on low power**

- Power consumption < 1 MW
- Field integral: 0.65Tm over 5m
- Weight ~ 800 t
- Aperture ~ 50 m²

Signal features



- Main background: neutrino interactions, (muon) combinatorial
- Reduce this background by:
 - IP cut
 - Invariant mass
- Important to
 - Measure precisely the momentum
 - Identify the particle
- Reduce combinatorial background by precise timing



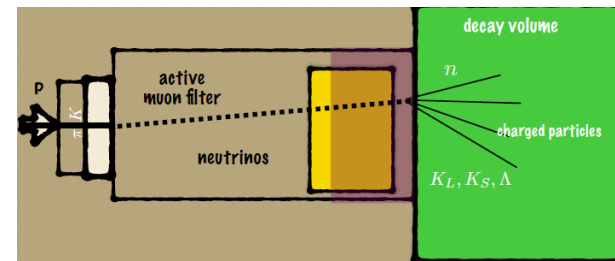
HS Backgrounds (1)

Main sources of background

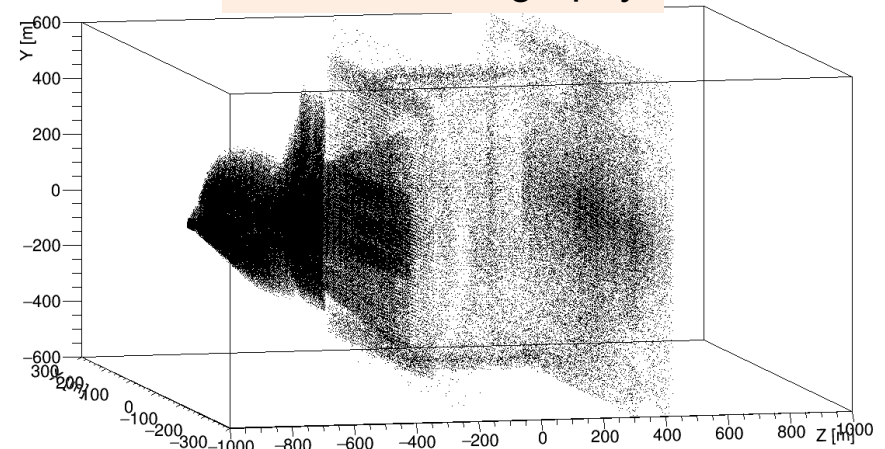
✓ **Neutrino DIS interactions with material in the vicinity of the HS decay volume** (interactions of ν with air in the decay volume are negligible at 10^{-3} mbar)

Origin of neutrino interactions

- Walls of the decay volume (>80%)
- Tau neutrino detector
- HS tracking system

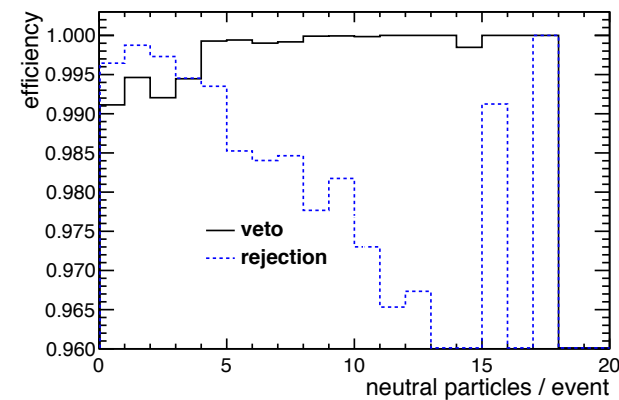
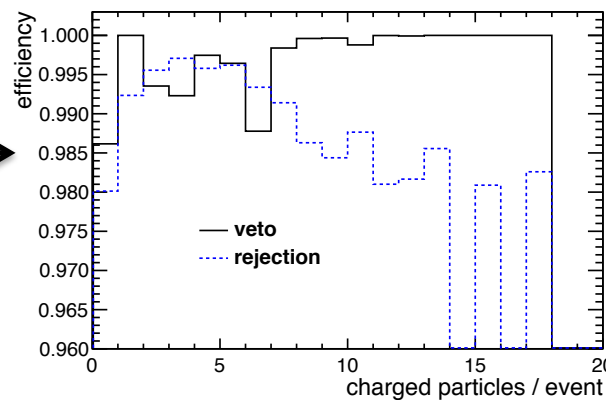


Neutrino tomography



Combination of veto and selection cuts reduces the ν -induced background to zero

Veto efficiency increases with event multiplicity →



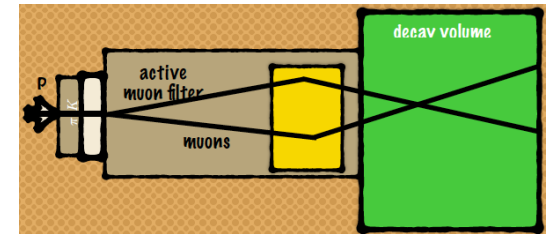
HS Backgrounds (2)

✓ Muon combinatorial background

Simulation predicts $O(10^{12})$ muon pairs in the decay volume in 5 years of data taking

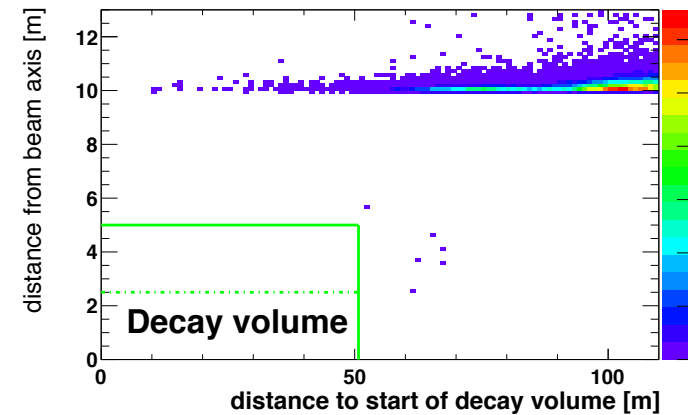
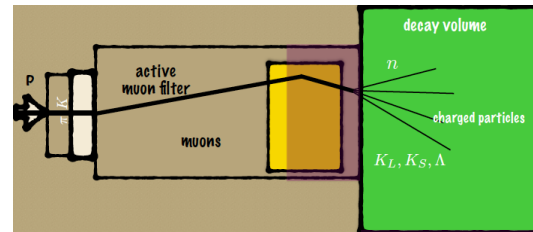
Suppressed by:

- Basic kinematic and topological cuts $\sim 10^4$
- Timing veto detectors $\sim 10^7$
- Upstream veto and surrounding veto taggers $\sim 10^4$



✓ Muon DIS interactions

- V^0 s produced in the walls of the cavern
- DIS close to the entry of the decay volume
→ smaller than neutrino induced background



✓ Cosmics

Background summary: no evidence for any irreducible background



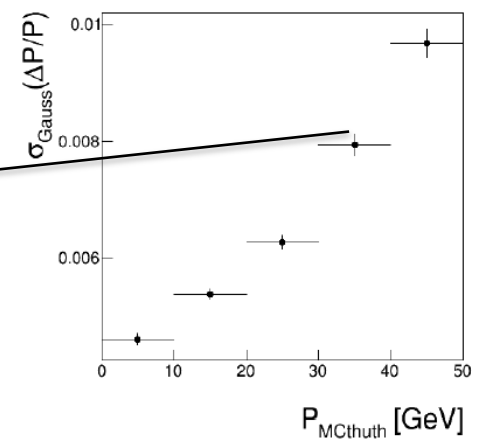
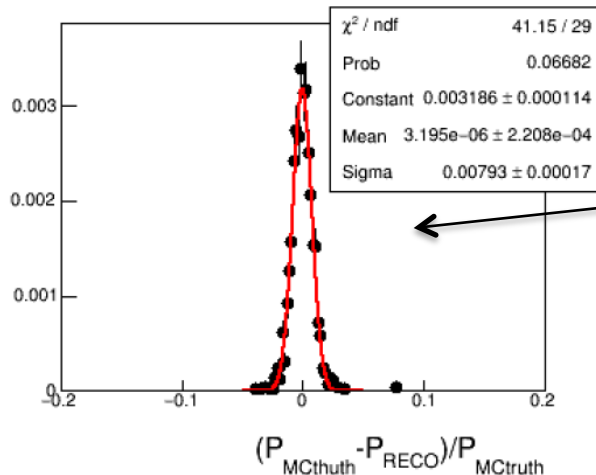
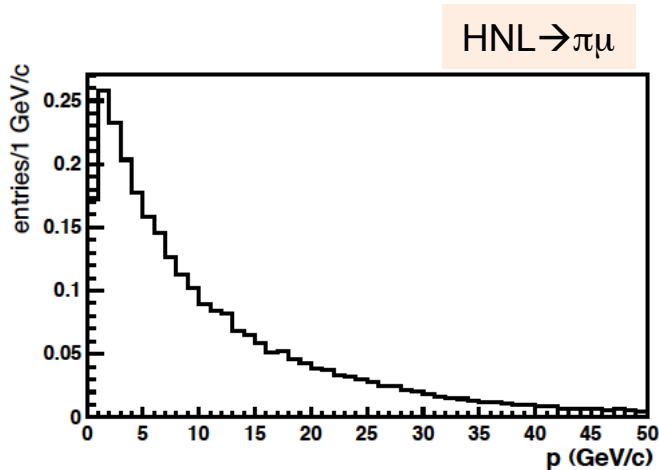
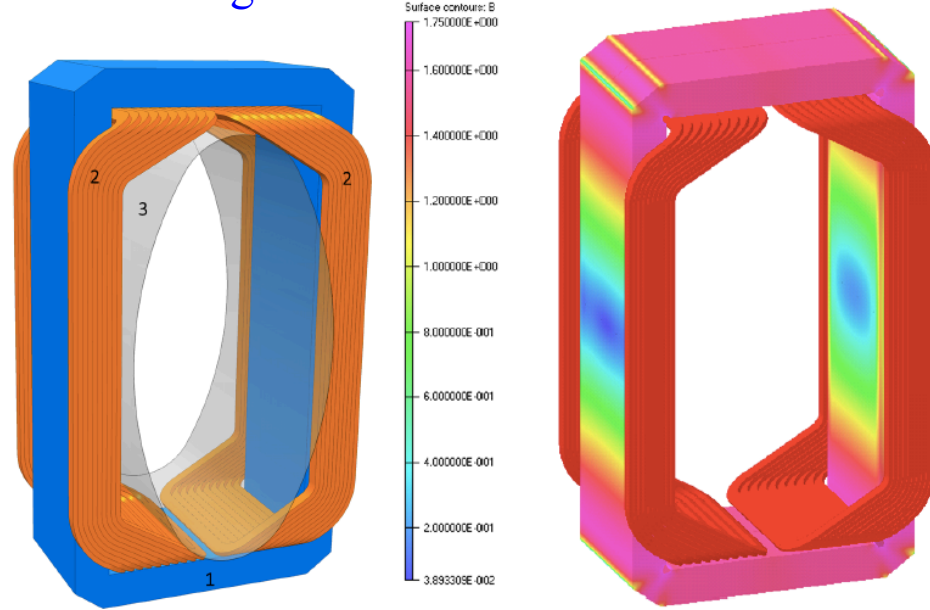
Momentum resolution of the HS (straw tubes) tracker

Magnet with vacuum vessel

- material budget per station $0.5\% X_0$
- position resolution $120 \mu\text{m}$ per straw, 8 hits per station on average

$$\left(\frac{\sigma_p}{p}\right)^2 \approx [0.49\%]^2 + [0.022\%/(\text{GeV}/c)]^2 \cdot p^2$$

Momentum resolution is dominated by multiple scattering below 22 GeV/c
 (For $\text{HNL} \rightarrow \pi\mu$, 75% of both decay products have $P < 20 \text{ GeV}/c$)



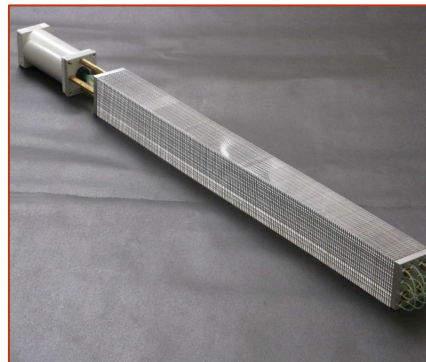
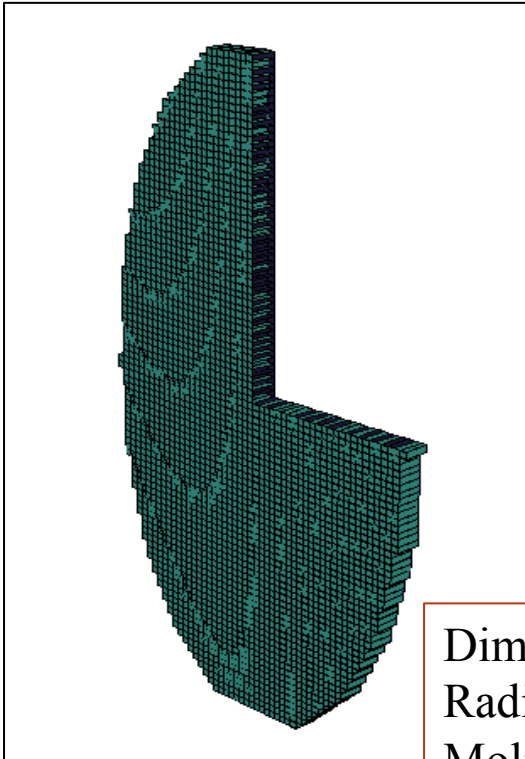
Vertex resolution (also driven by multiple scattering and $\Delta P/P$):

$$\sigma_{xy} \sim \mathcal{O}(\text{mm}), \sigma_z \sim \mathcal{O}(\text{cm})$$

Calorimeters

ECAL

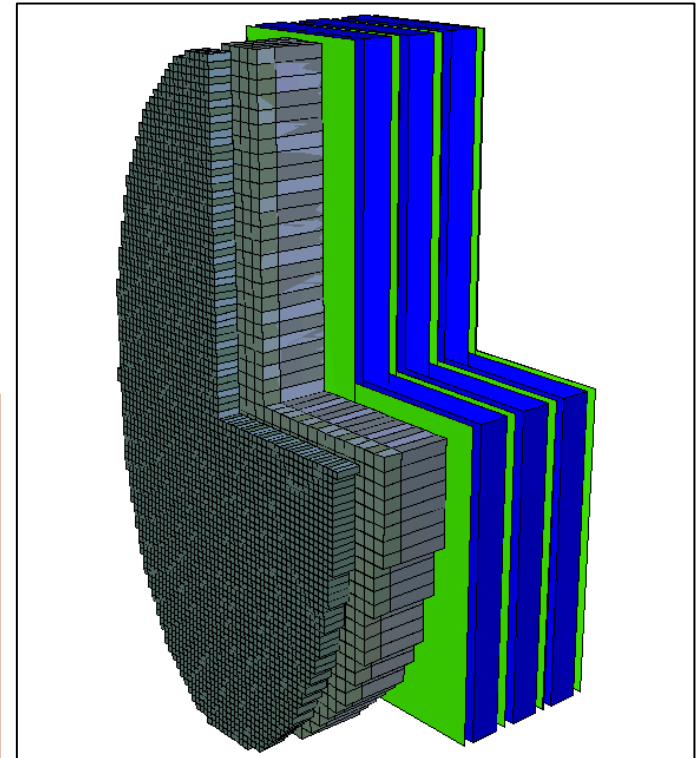
- Almost elliptical shape (5 m x 10 m)
- 2876 Shashlik modules
- 2x2 cells/modules, width=6 cm
- 11504 independent readout channels



Dimensions	60x60 mm ²
Radiation length	17 mm
Moliere radius	36 mm
Radiation thickness	25 X ₀
Scintillator thickness	1.5 mm
Lead thickness	0.8 mm
Energy resolution	1%

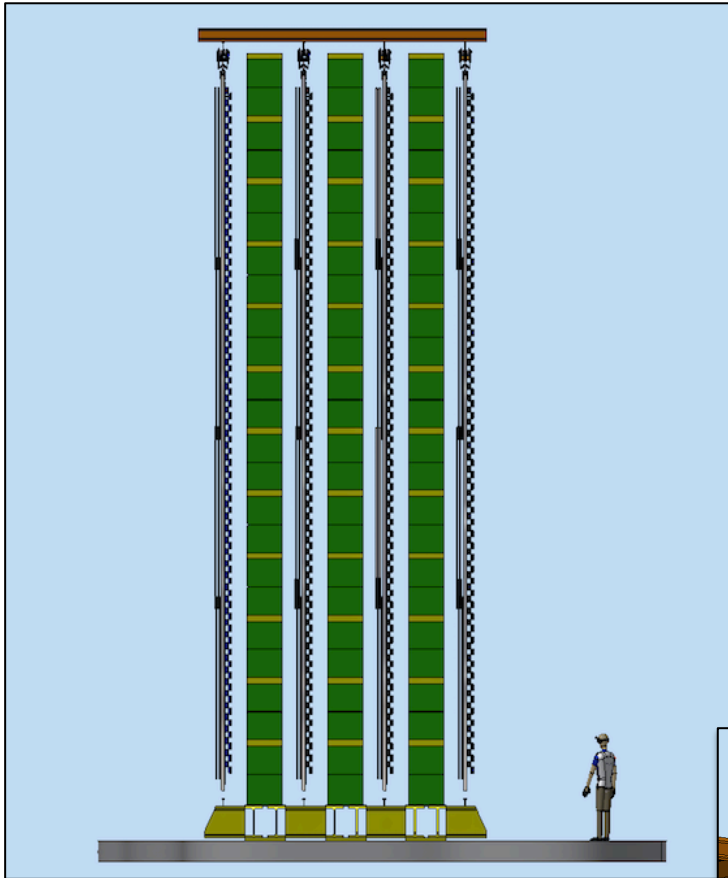
HCAL

- Matched with ECAL acceptance
- 2 stations
- 5 m x 10 m
- 1512 modules
- 24x24 cm² dimensions
- Stratigraphy: N x (1.5 cm steel+0.5 cm scint)
- 1512 independent readout channels



Muon System

Based on scintillating bars, with WLS fibers and SiPM readout

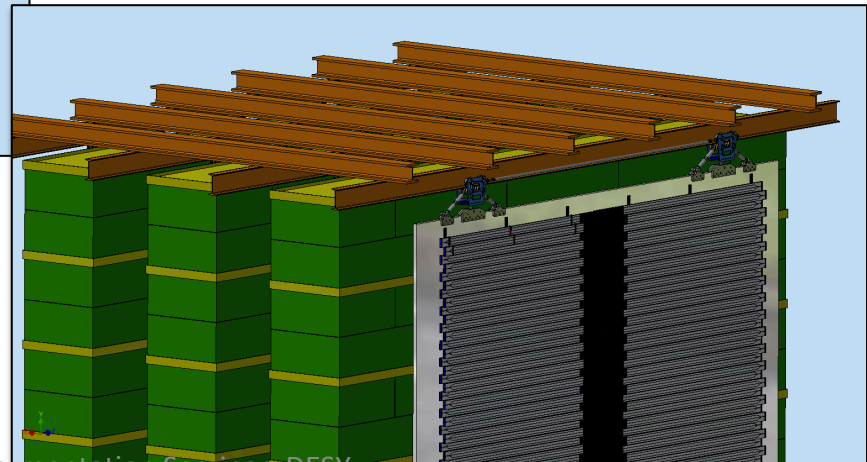


Requirements:

- High-efficiency identification of muons in the final state
- Separation between muons and hadrons/electrons
- Complement timing detector to reject combinatorial muon background

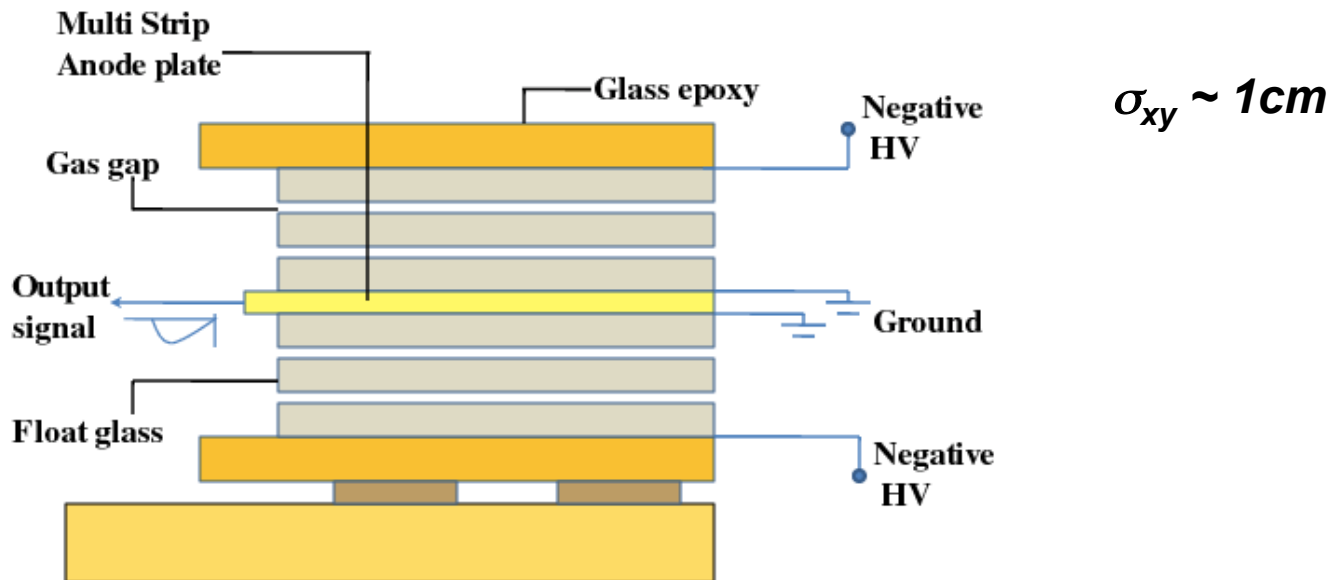
Technical Proposal (preliminary design)

- 4 active stations
- transverse dimensions: $1200 \times 600 \text{ cm}^2$
- x,y view
- 3380 bars, $5 \times 300 \times 2 \text{ cm}^3$ /each
- 7760 FEE channels
- 1000 tons of iron filters



Timing detector (< 100ps)

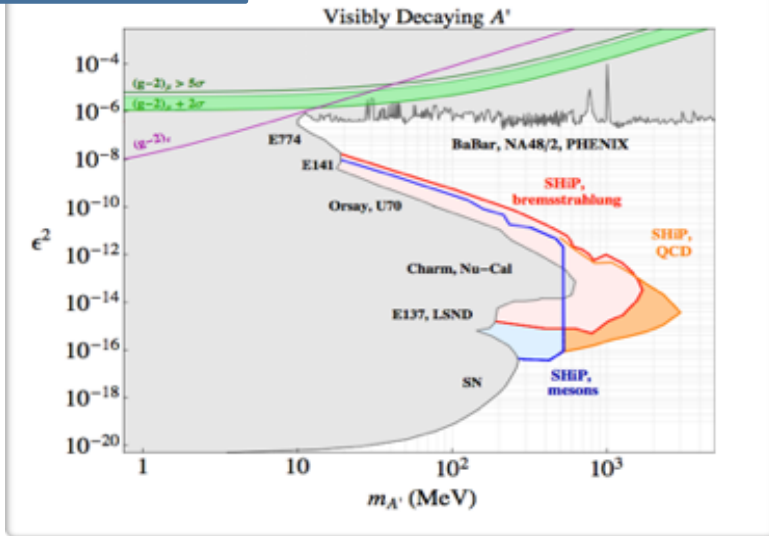
Multi-gap RPC is one option



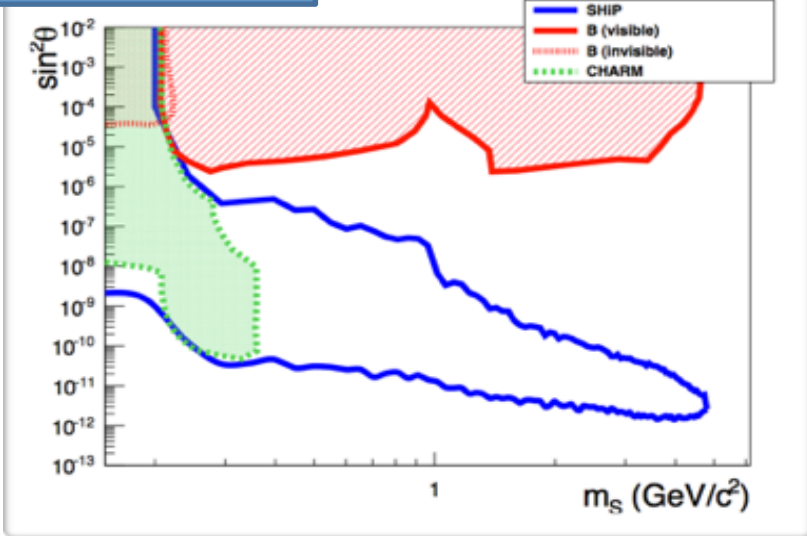
SHiP sensitivity to Hidden Sector

Based on 2×10^{20} pot
@400 GeV in 5 years

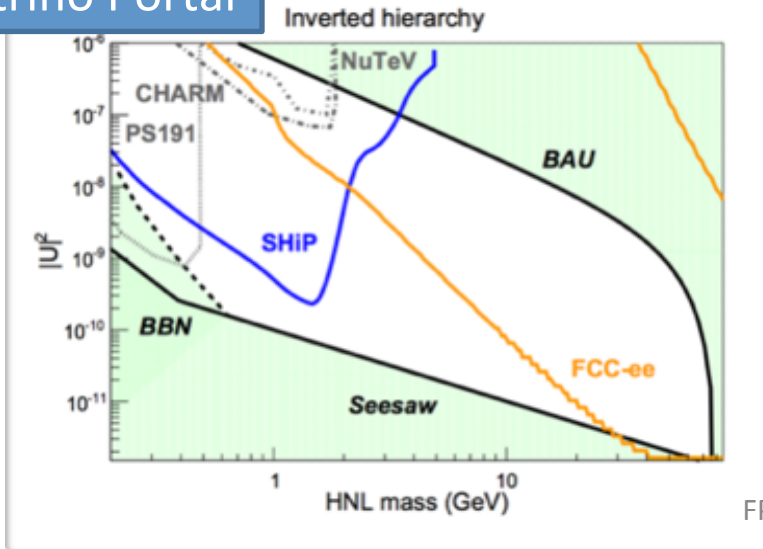
Vector Portal



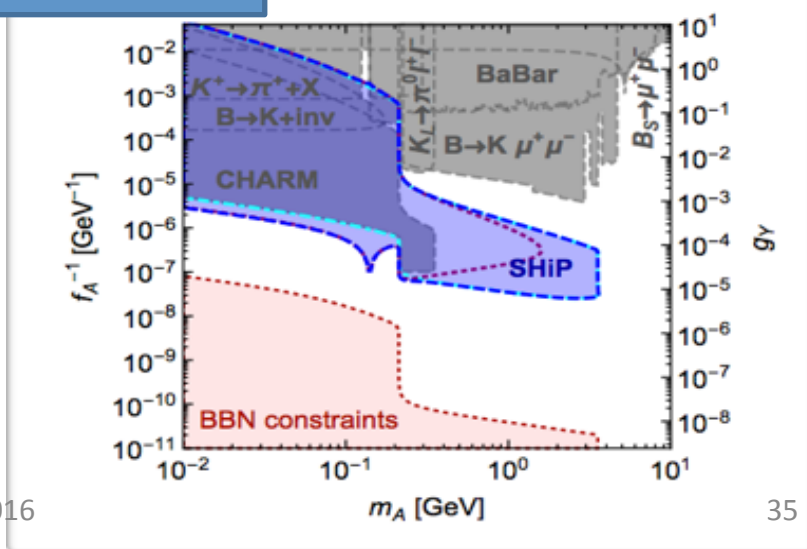
Scalar Portal



Neutrino Portal



Axion Portal

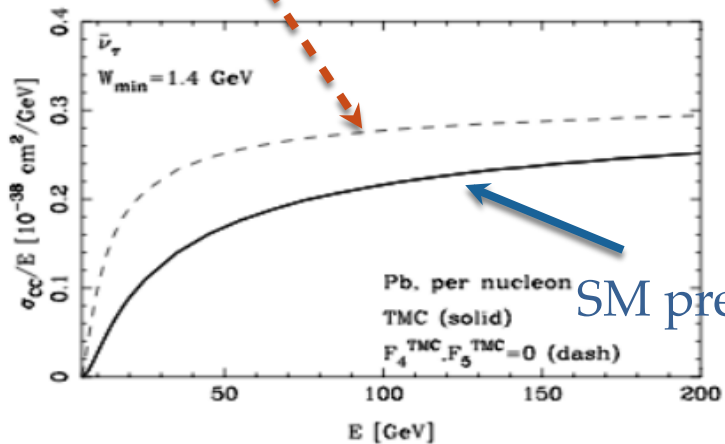


F₄ AND F₅ STRUCTURE FUNCTIONS

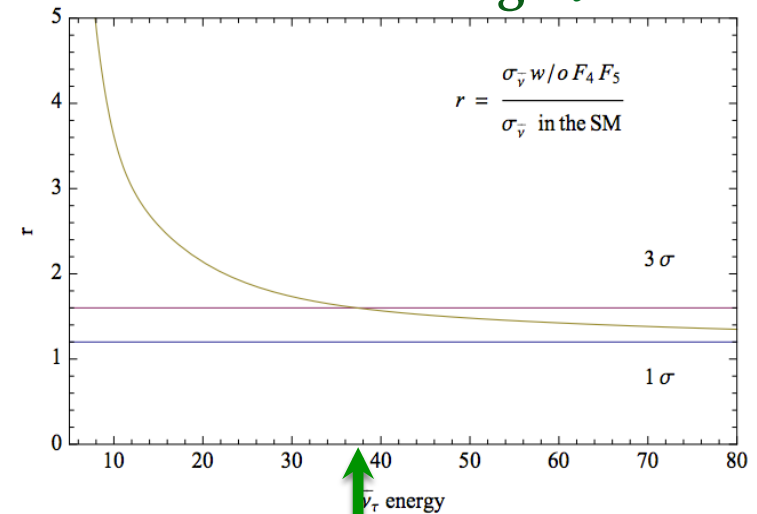
First evaluation of F₄ and F₅, not accessible with other neutrinos

$$\frac{d^2\sigma^{\nu(\bar{\nu})}}{dxdy} = \frac{G_F^2 M E_\nu}{\pi(1 + Q^2/M_W^2)^2} \left((y^2x + \frac{m_\tau^2 y}{2E_\nu M}) F_1 + \left[(1 - \frac{m_\tau^2}{4E_\nu^2}) - (1 + \frac{Mx}{2E_\nu}) \right] F_2 \right. \\ \left. \pm \left[xy(1 - \frac{y}{2}) - \frac{m_\tau^2 y}{4E_\nu M} \right] F_3 + \frac{m_\tau^2(m_\tau^2 + Q^2)}{4E_\nu^2 M^2 x} F_4 - \frac{m_\tau^2}{E_\nu M} F_5 \right),$$

F₄ = F₅ = 0



CC interacting $\bar{\nu}_\tau$

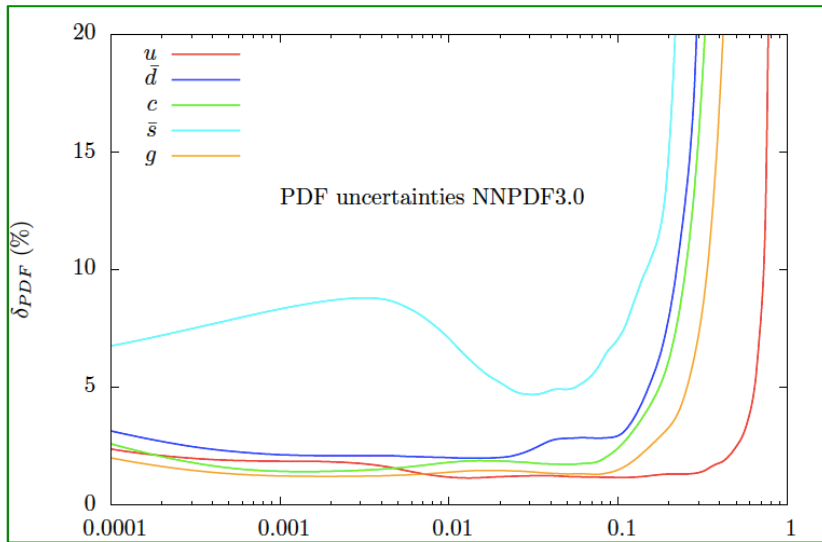
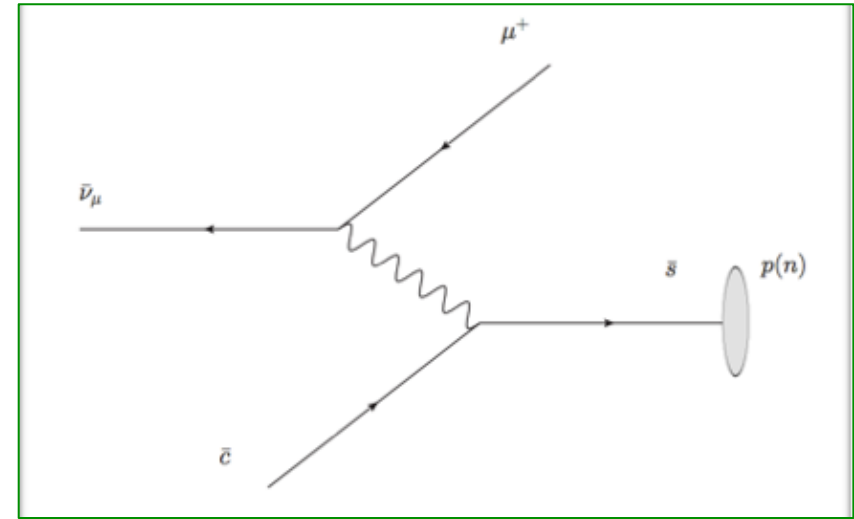


$E(\bar{\nu}_\tau) < 38 \text{ GeV}$

- At LO $F_4 = 0$, $2xF_5 = F_2$
- At NLO $F_4 \sim 1\%$ at 10 GeV

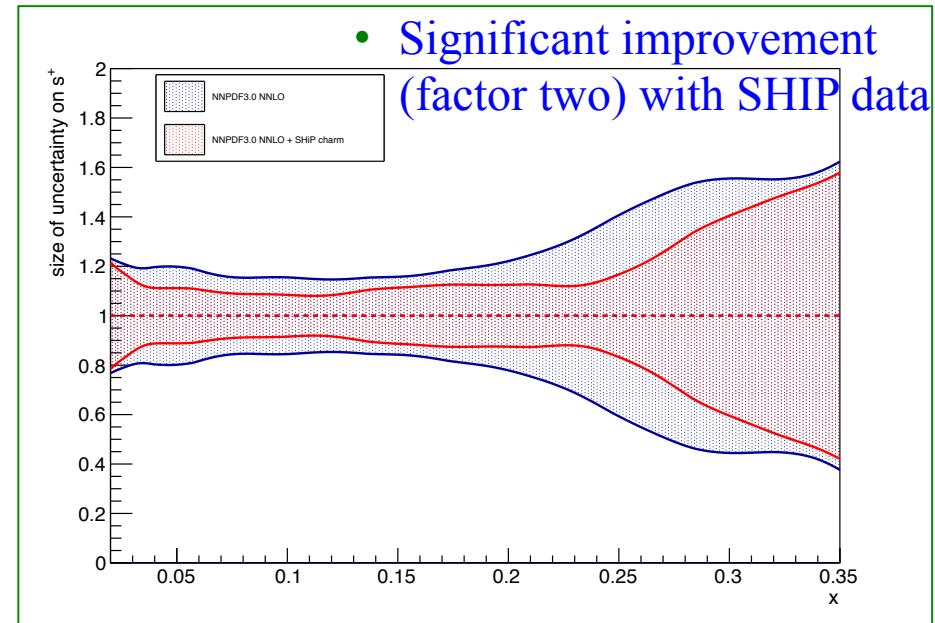
STRANGE QUARK NUCLEON CONTENT

- Charmed hadron production in anti-neutrino interactions selects anti-strange quark in the nucleon
- Strangeness important for precision SM tests and for BSM searches
- W boson production at 14 TeV: 80% via $u\bar{d}$ and 20% via $c\bar{s}$



Phys. Rev. D91 (2015) 113005

Fractional uncertainty of the individual parton densities $f(x; m^2_W)$ of NNPDF3.0



• Significant improvement (factor two) with SHIP data

$$s^+ = s(x) + \bar{s}(x)$$

DARK MATTER SEARCH

WITH THE NEUTRINO DETECTOR

χ produced by a dark photon decay

$$\chi e^- \rightarrow \chi e^-$$

P. deNiverville, D. McKeen, and A. Ritz,

Phys.Rev. D86 (2012) 035022

α' = dark photon coupling with χ

SIGNAL SELECTION

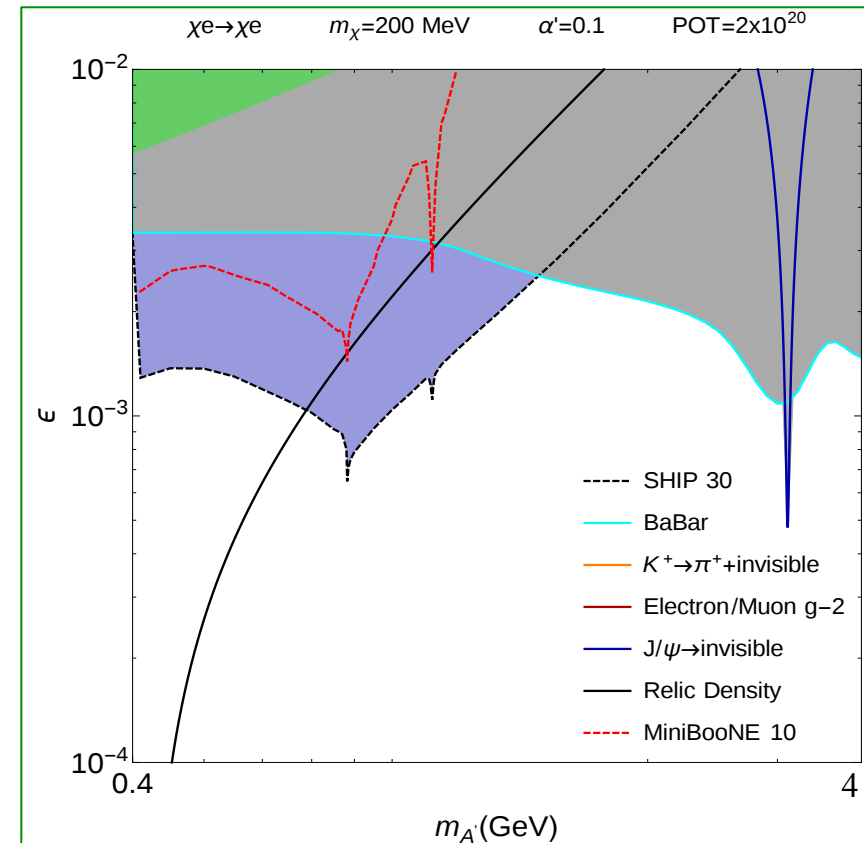
$$\left\{ \begin{array}{l} 0.01 < \theta < 0.02 \\ E < 20 \text{ GeV} \end{array} \right.$$

BACKGROUND PROCESSES

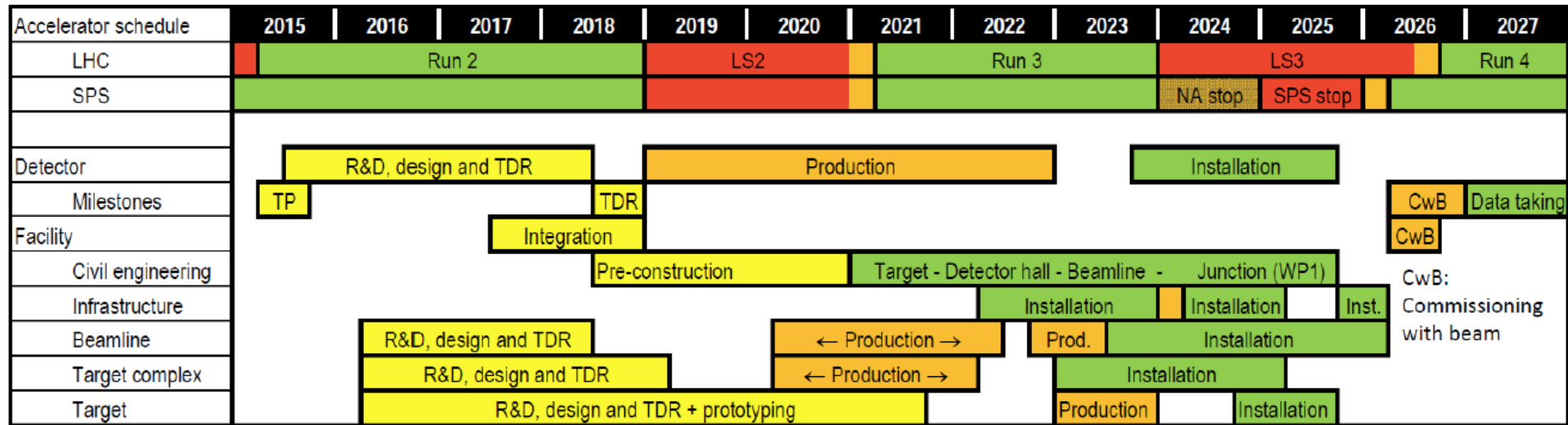
	ν_e	$\bar{\nu}_e$	ν_μ	$\bar{\nu}_\mu$	all
Elastic scattering on e^-	16	2	20	18	56
Quasi - elastic scattering	105	73			178
Resonant scattering	13	27			40
Deep inelastic scattering	3	7			10
Total	137	109	20	18	284

ϵ = dark photon coupling with e.m. current

m_A = dark photon mass



Project schedule



- *10 years from TP to data taking*
- ✓ *Schedule optimized to avoid interference with operation of North Area*
 - ➔ *Preparation of facility in four clear and separate work packages (junction cavern, beam line, target complex and detector hall)*
 - ➔ *Use LS3 for junction cavern and first short section of SHiP beam line*
- ✓ *Comprehensive Design Study 2016-2018: Starting now! ➔ Update of European HEP strategy 2018*
- ✓ *Construction/production 2021-*
 - ➔ ***Data taking 2026***



Summary

SHiP to complement searches for New Physics at CERN in the largely unexplored domain of new, very weakly interacting particles with masses $O(10)$ GeV

- ✓ *Unique opportunity for ν_τ physics*
- ✓ *Sensitivity improves past experiments by $O(10000)$ for Hidden Sector and by $O(\sim 1000)$ for ν_τ physics*
- ✓ *Recommendations of the SPS Committee at CERN in January 2016: The SPSC supports the motivation for the search for hidden particles, which will explore a domain of interest for many open questions in particle physics and cosmology, and acknowledges the interest of the measurements foreseen in the neutrino sector. SHiP could therefore constitute a key part of the CERN Fixed Target programme in the HL-LHC era. SPSC **recommends** that the SHiP proponents proceed with the preparation of a Comprehensive Design Report (CDR), and that this preparation be made in close contact with the planned Fixed Target working group.*
- ✓ *CERN DG has launched the “Beyond Colliders Physics” Working Group. The beam dump is the first item. Kick-off meeting on September 6th-7th*
- ✓ *Design optimisation going on: your contribution is very welcome!*