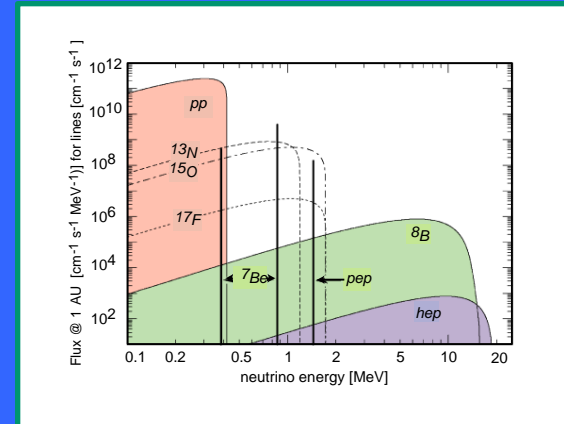
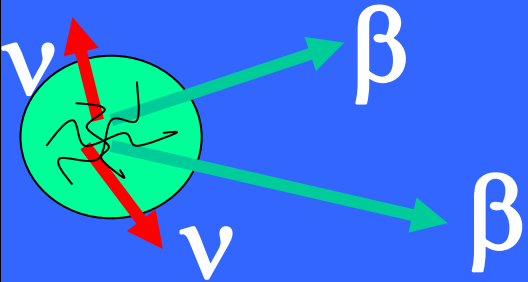


Charge-exchange reactions

GT-transitions, $\beta\beta$ -decay

and

things beyond



NUMEN2015

Outline

➤ Chargex-reactions ($^3\text{He},t$) & ($d,^2\text{He}$)

- highlights & features of $2\nu\beta\beta$ nuclear matrix elements (NME)

^{76}Ge , ^{82}Se , ^{96}Zr , ^{100}Mo , ^{136}Xe

fragmentation - smallest/largest NME



➤ the $0\nu\beta\beta$ decay nuclear matrix elements

1st forbidden NME's and 2⁻ states

➤ solar ν SNU rates and ($^3\text{He},t$) reaction

$^{71}\text{Ga}(^3\text{He},t)$, $^{82}\text{Se}(^3\text{He},t)$

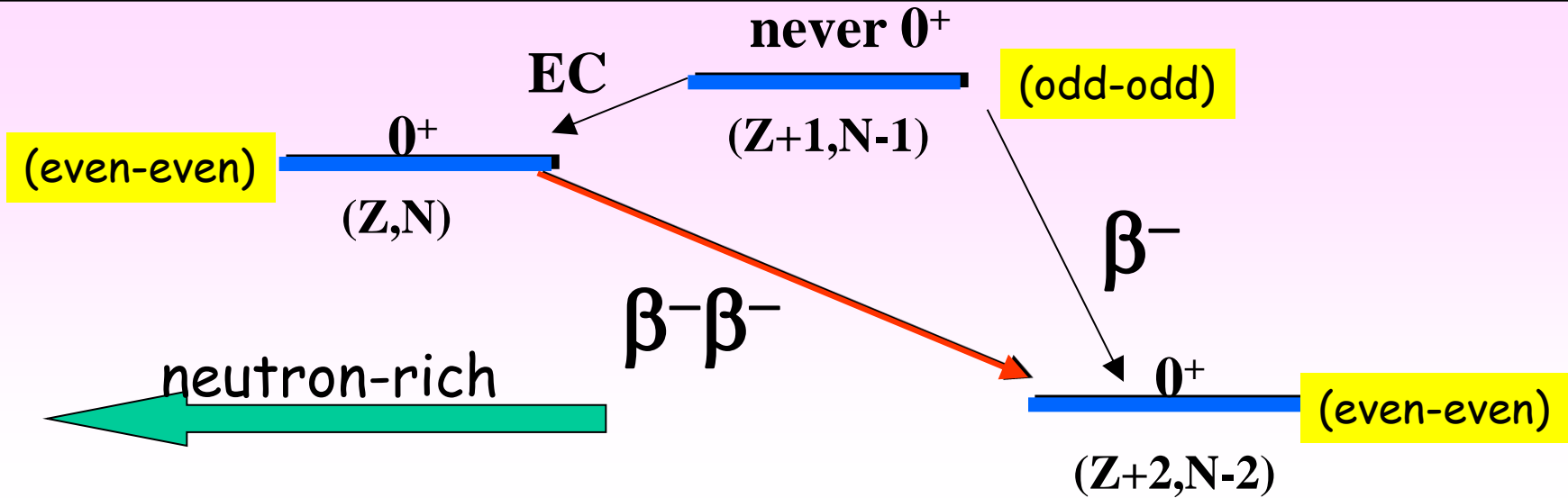
➤ the A=96 system

the $^{96}\text{Zr} (\beta^-) \rightarrow ^{96}\text{Nb}$ Q-value
and a direct test of $0\nu\beta\beta$ NME



JYVÄSKYLÄN YLIOPISTO
UNIVERSITY OF JYVÄSKYLÄ

$\beta^-\beta^-$ decay



$2\nu\beta^-\beta^-$ decay:

$$T_{1/2} \approx 10^{19-21} \text{ y}$$

$$\Gamma = (\text{ph-spc}) \times \left| \begin{array}{c} NME \\ \text{5-body} \\ \text{allowed} \end{array} \right|^2$$

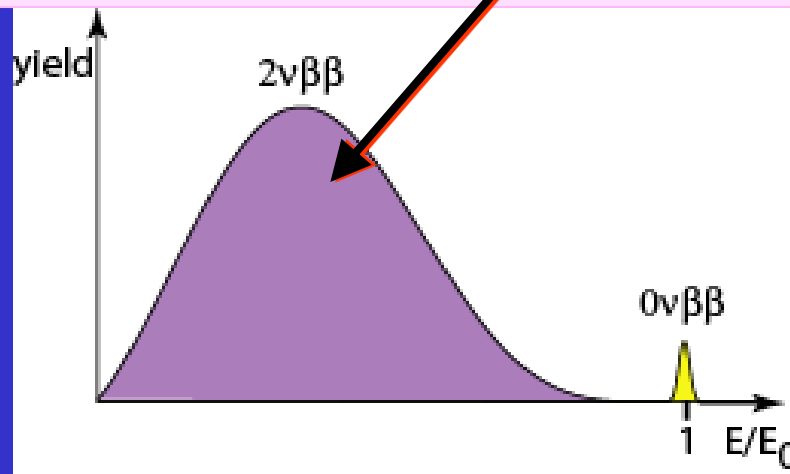
$0\nu\beta^-\beta^-$ decay:

$$T_{1/2} > 10^{24} \text{ y}$$

$$\Gamma = (\text{ph-spc}) \times \left| \begin{array}{c} NME \\ \text{3-body} \\ \text{any degree} \end{array} \right|^2 \times \left| \sum_{i=1}^3 U_{ei}^2 m_i \right|^2$$

N_{ucl.} **M**_{atrix} **E**_{lements}

2νβ-β- decay



**q-transfer like in ordinary
β-decay**

($q \sim 0.01 \text{ fm}^{-1} \sim 2 \text{ MeV}/c$)

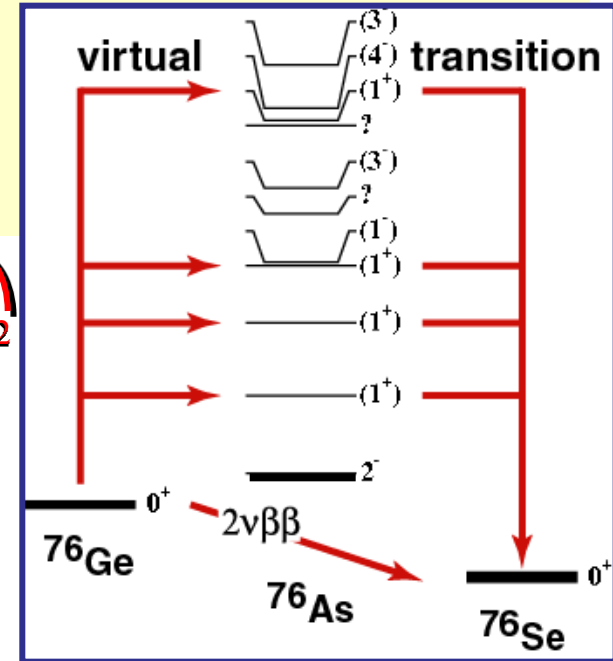
i.e. only allowed transitions possible

$$\Gamma_{(\beta^-\beta^-)}^{2\nu} = \frac{C}{8\pi^7} \left(\frac{G_F g_A}{\sqrt{2}} \cos(\Theta_C) \right)^4 \left| M_{\text{DGT}}^{(2\nu)} \right|^2 \mathcal{F}_{(-)}^2 f(\mathbf{Q})$$

$$= G^{2\nu}(\mathbf{Q}, Z) \left| M_{\text{DGT}}^{(2\nu)} \right|^2$$

$$\propto Q^{11} \cdot Z^2$$

$\exp \approx 10^{-3} \text{ MeV}^{-2}$
extracted from
half-life

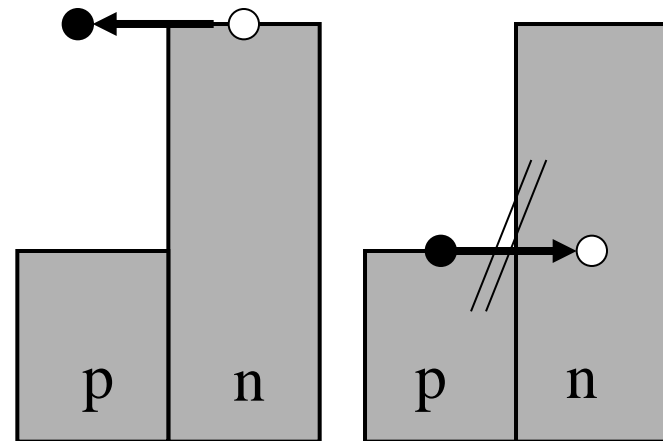


favorable:

1. high Q-value
2. large Z

unfavorable (but cannot be changed):

1. large neutron excess
(Pauli-blocking)



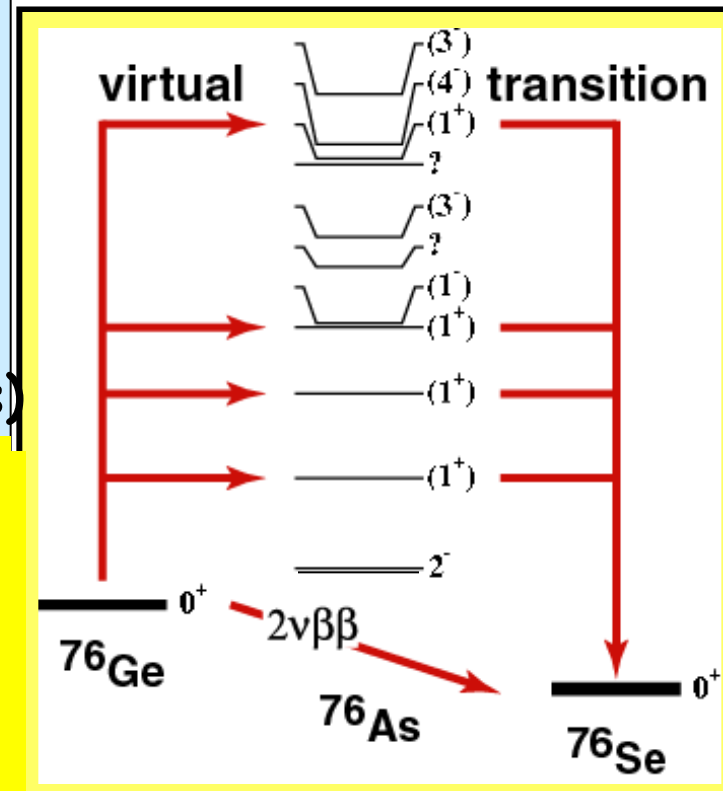
$$M_{\text{DGT}}^{(2\nu)} = \sum_m \frac{\langle \mathbf{0}_{g.s.}^{(f)} | \sum_k \sigma_k \tau_k^- | \mathbf{1}_m^+ \rangle \langle \mathbf{1}_m^+ | \sum_k \sigma_k \tau_k^- | \mathbf{0}_{g.s.}^{(i)} \rangle}{\frac{1}{2} Q_{\beta\beta}(\mathbf{0}_{g.s.}^{(f)}) + E(\mathbf{1}_m^+) - E_0}$$

$$= \sum_m \frac{M_m \quad GT^+ \quad M_m \quad GT^-}{E_m}$$

to remember:

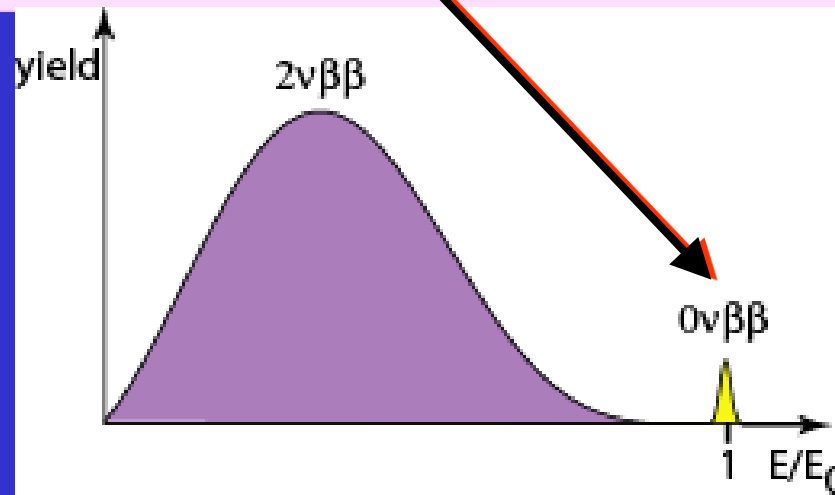
1. 2 sequential & „allowed“ β^- -decays of „Gamow-Teller“ type
2. „1, 2, 3, ... forbidden“ decays negligible
3. Fermi-transitions do not contribute (because of different isospin-multiplets)

Can be determined via charge-exchange reactions in the (n,p) and (p,n) direction (e.g. (d, ^2He) or (^3He , t))



N_{ucl.} **M**_{atrix} **E**_{lements}

$0\nu\beta\beta$ decay



neutrino is a virtual particle

$$q \sim 0.5 \text{ fm}^{-1} (\sim 100 \text{ MeV}/c)$$

(due to Heisenberg $\Delta q \cdot \Delta x \sim 1$)

degree of forbiddenness is lifted

$$\Gamma_{(\beta^-\beta^-)}^{0\nu} = G^{0\nu}(Q,Z) g_A^4 \left| M_{\text{DGT}}^{(0\nu)} - \left(\frac{g_V}{g_A} \right)^2 M_{\text{DF}}^{(0\nu)} \right|^2 |m_{\nu_e}|^2$$

$$\propto Q^5 \cdot Z^4$$

theory $\approx 10 !!$
 largely independent of (A,Z)
 (except near magic nuclei)

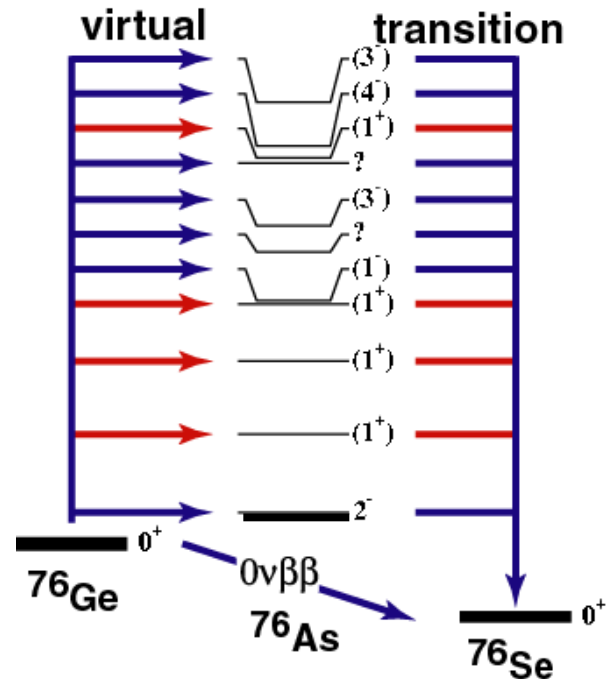
mass of Majorana- ν !

to remember:

1. „higher-fold forbidden“ transitions possible
2. Fermi-transitions important
3. „Pauli-blocking“ important
4. large Q important

NUMEN

Not accessible via charge exchange reactions

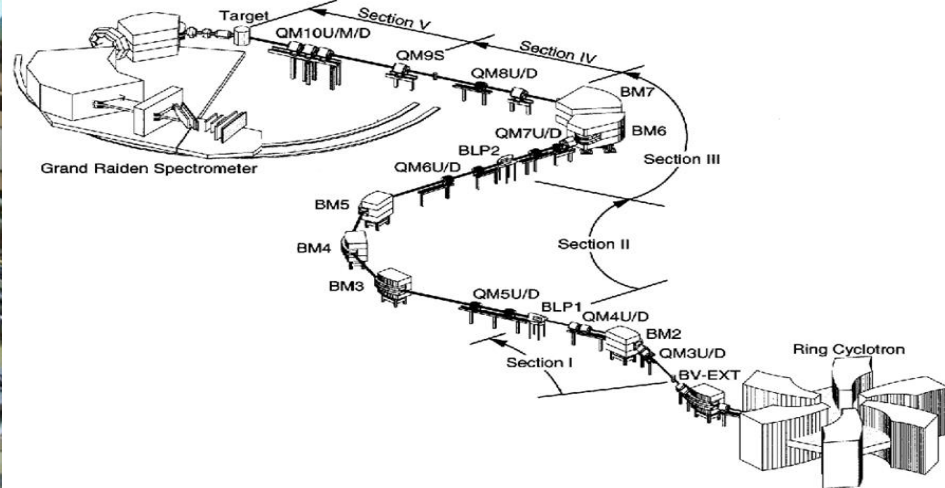


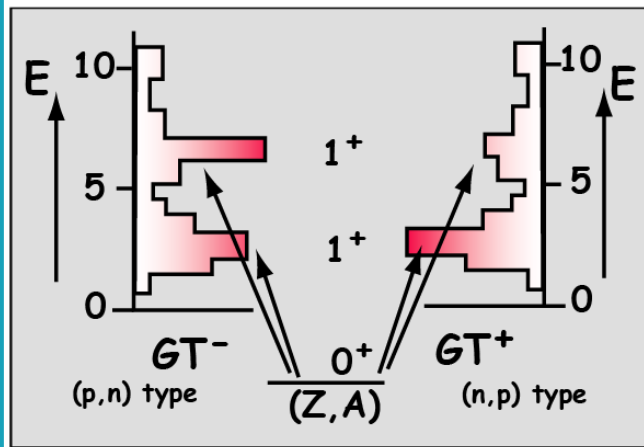
Charge-exchange reactions

Grand Raiden Magnetic Spectrometer



$\Delta E/E \sim 5 \times 10^{-5}$ ~ 25 keV
at 420 MeV (^3He)





Q: what is the connection between „weak $\sigma\tau$ operator“ and the hadronic reaction

A: dominance of the $V_{\sigma\tau}$ effective interaction at medium energies

$$M(GT) = \langle 1^+ || \sigma\tau^{\pm} || 0_{g.s.}^+ \rangle$$

$$B(GT) = \frac{1}{2J_i+1} | M(GT) |^2$$

hadronic probes: (n, p), (d, ^2He), (t, ^3He)

or (p, n), (^3He , t)

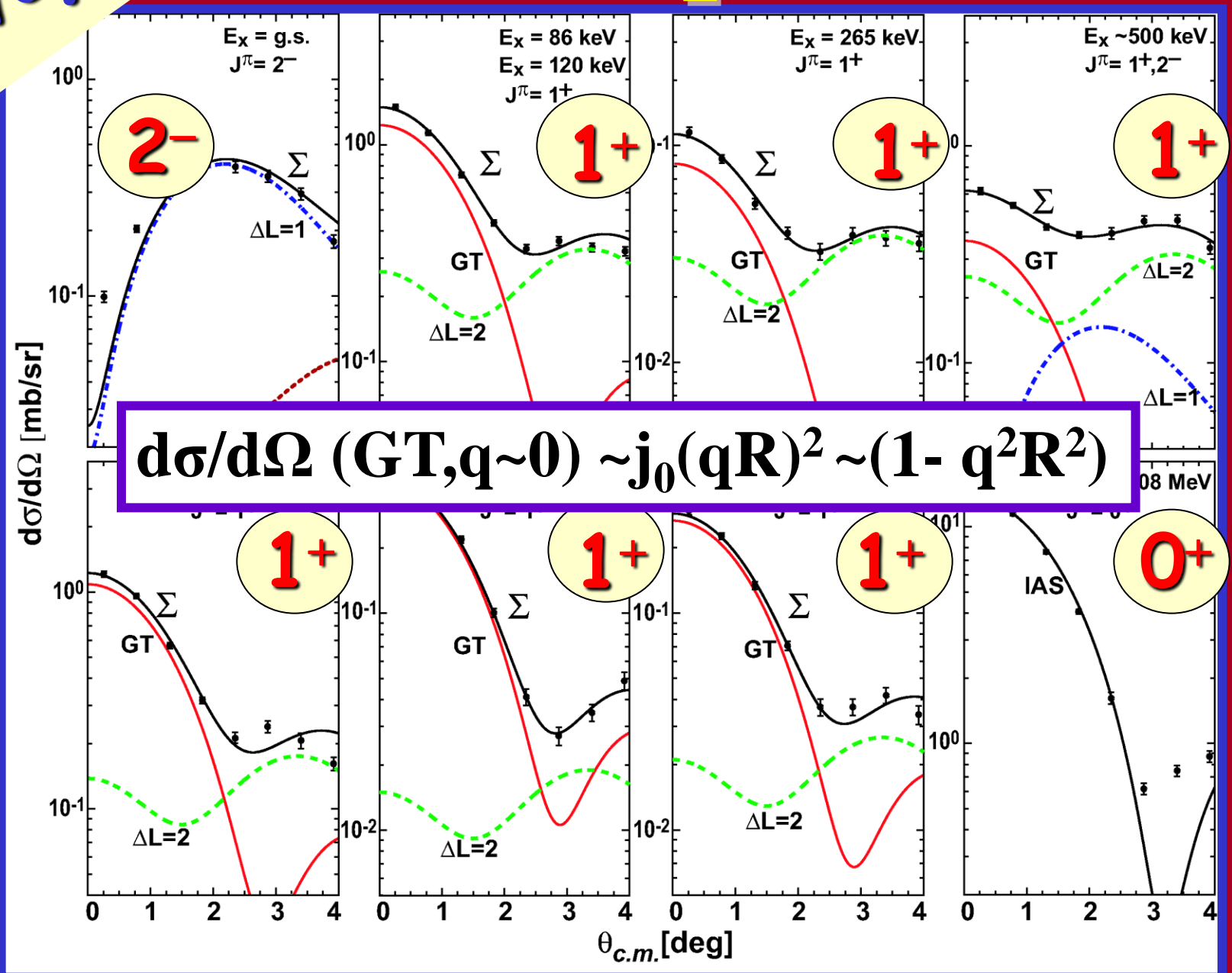
$$\left[\frac{d\sigma}{d\Omega} \right] = \left[\frac{\mu}{\pi\hbar} \right]^2 \frac{k_f}{k_i} N_d |V_{\sigma\tau}|^2 |\langle f | \sigma\tau | i \rangle|^2$$

$q = 0!!$

largest at 100 - 200 MeV/A

$^{76}\text{Ge}-^{76}\text{As}$

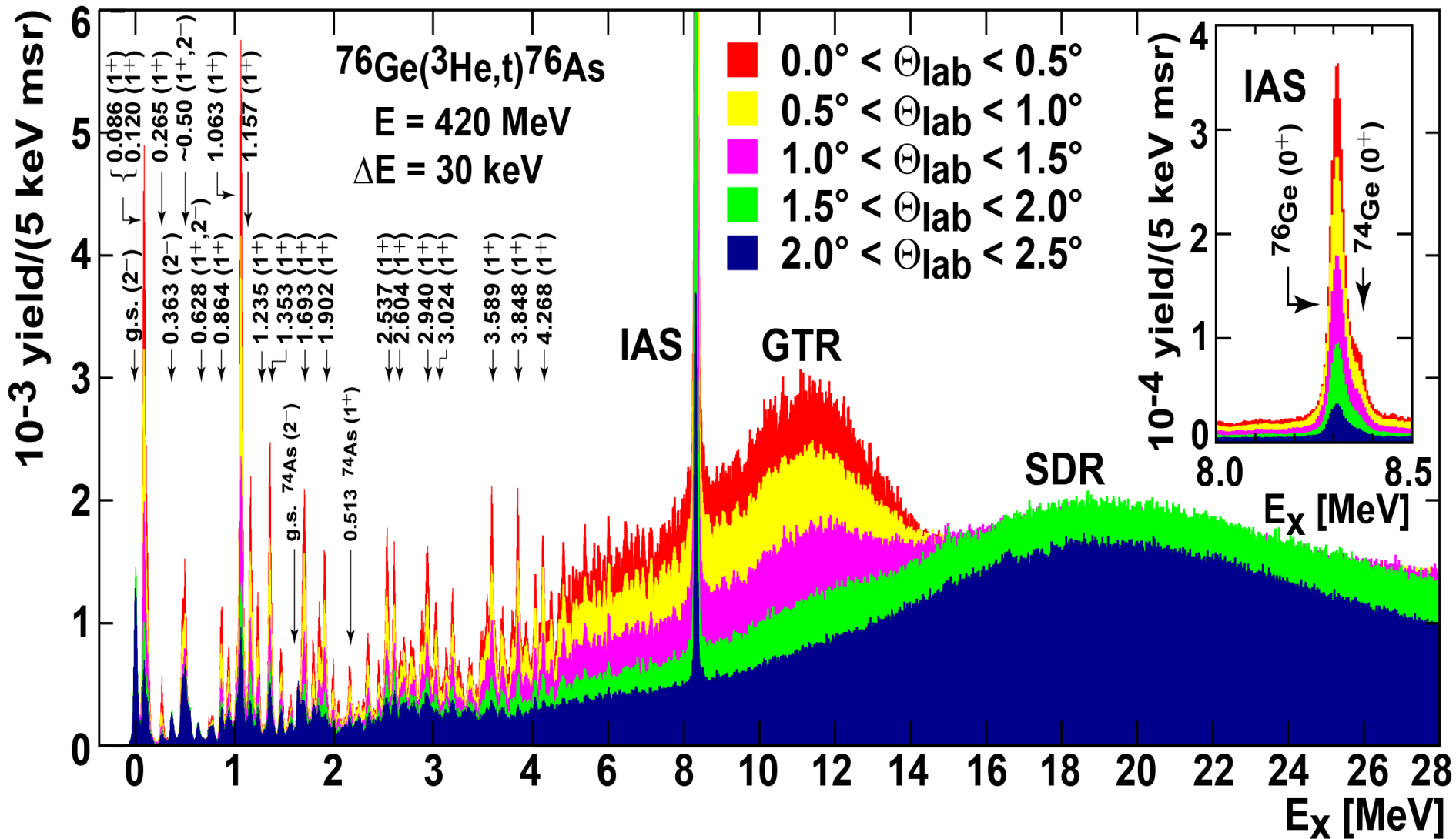
examples



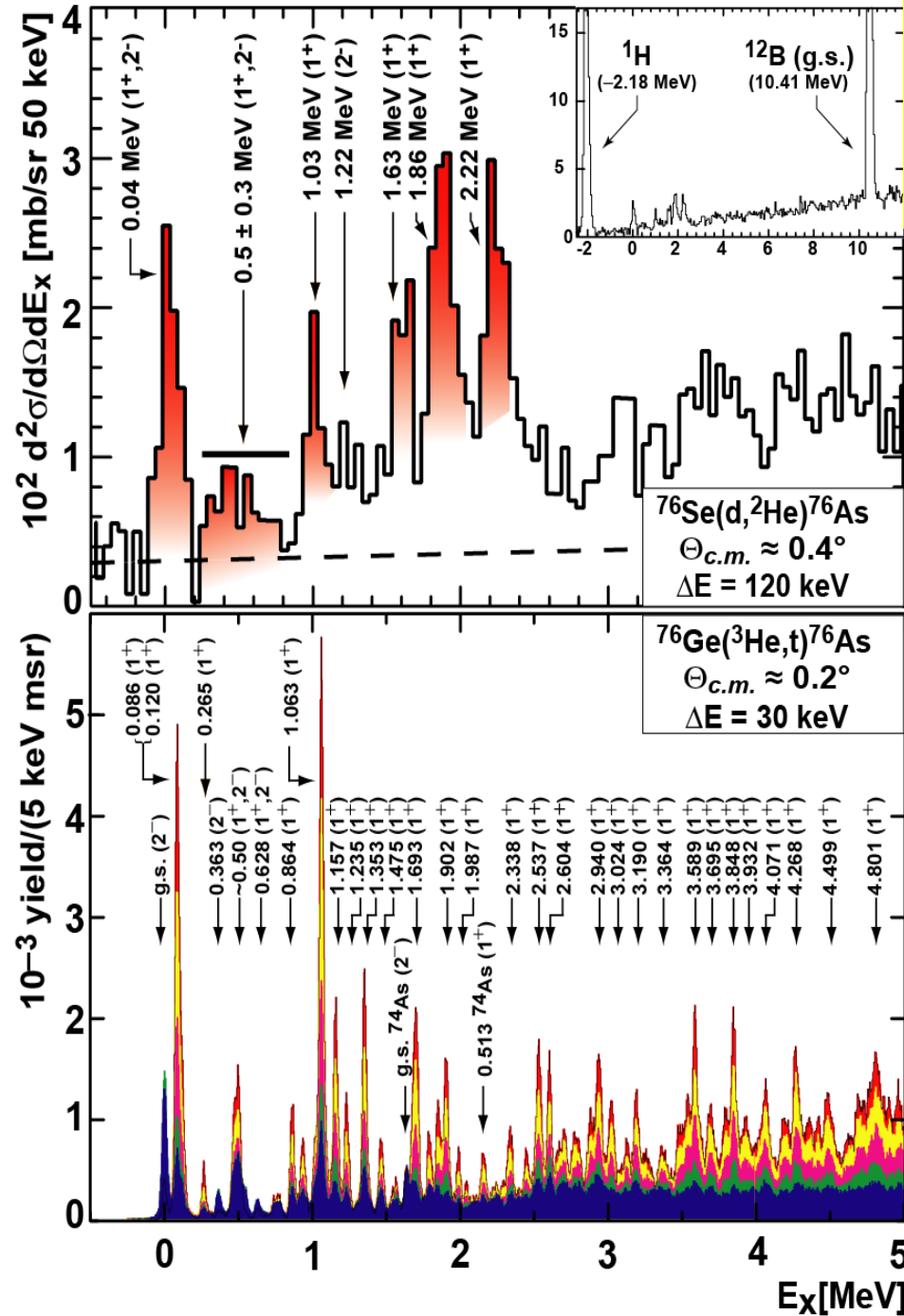
^{76}Ge

$N-Z=10$

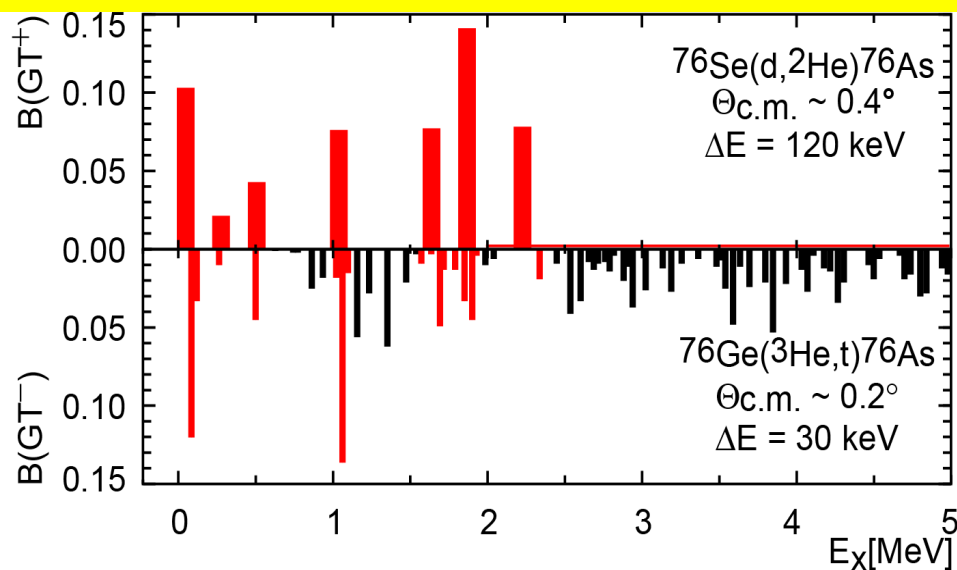
Resolution is the key !!!



**almost 70 !! resolved single states up to 5 MeV
 identified as GT 1^+ transitions !!!**



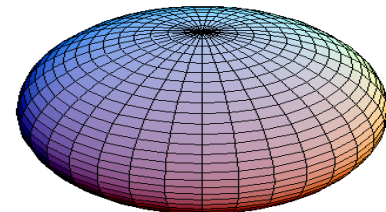
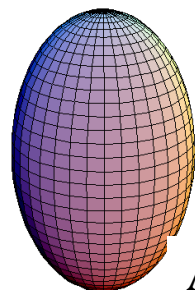
**~ 70 !! single states
 up to 5 MeV !!!
 ??? anti-correlation ???**



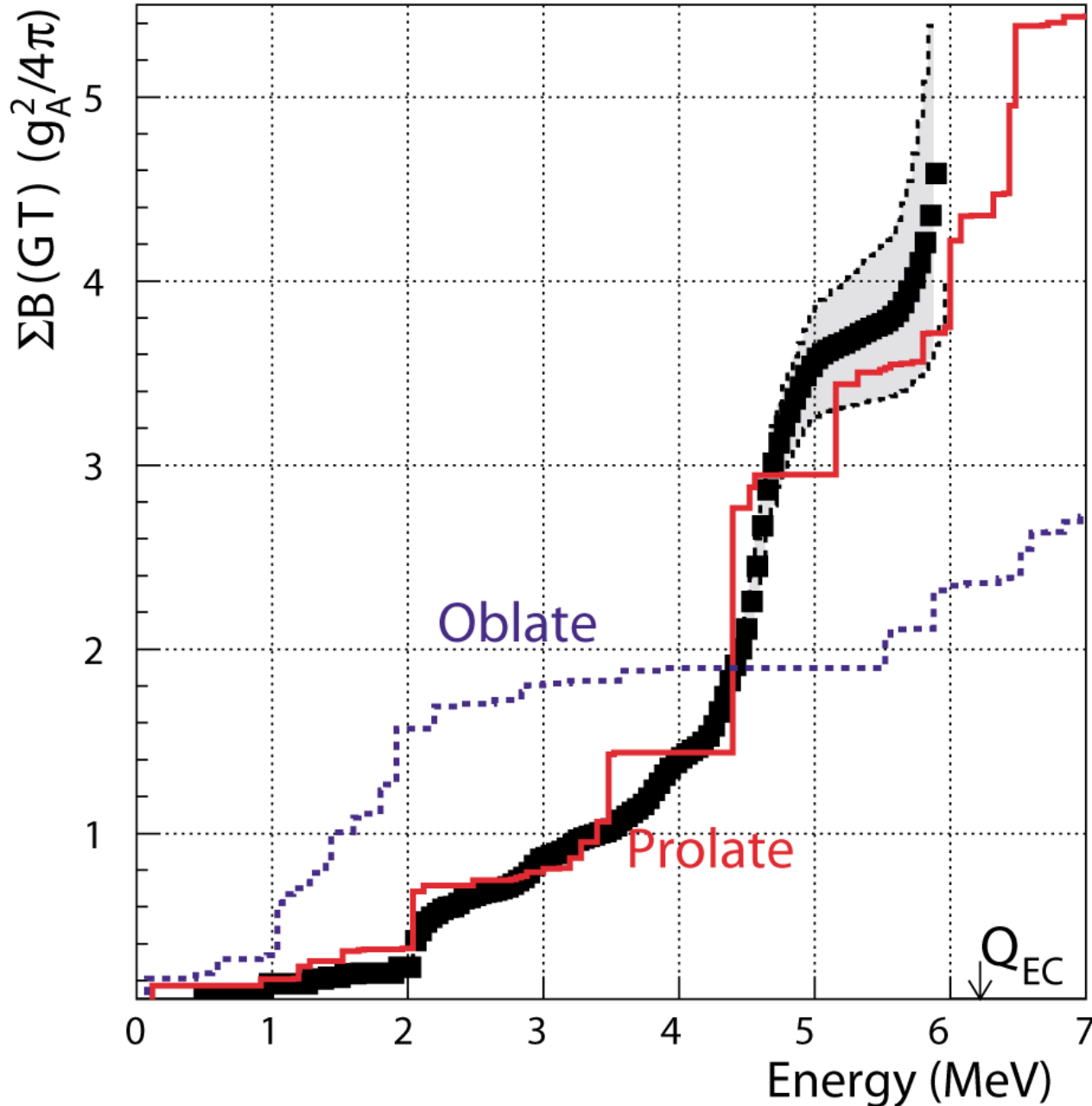
**is the anti-correlation a
 property of deformation ??**

^{76}Ge

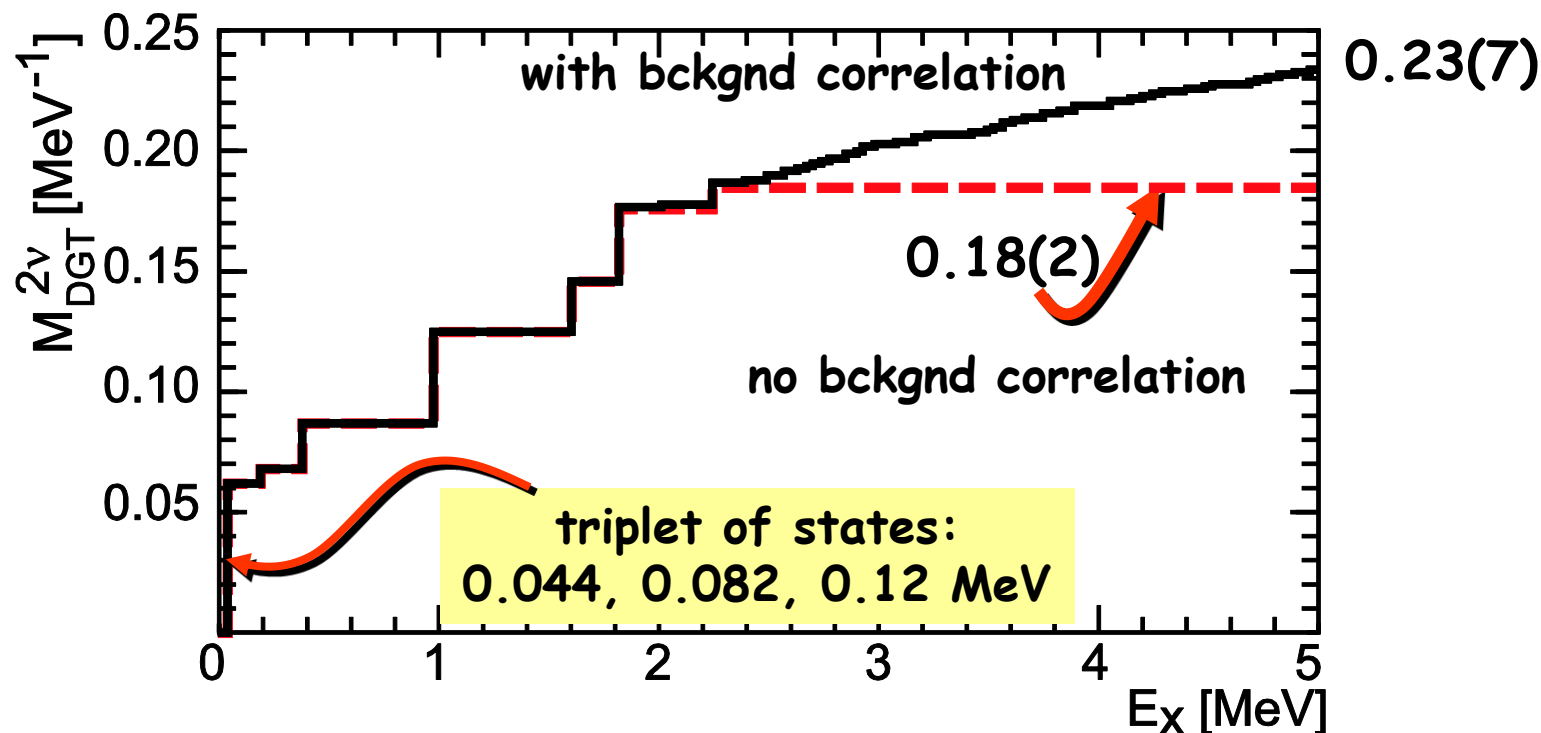
^{76}Se



A. Faessler et al. PRC70 (2004)



^{76}Sr
prolate
or
oblate



Low-E part of M_{DGT} makes up
~100(+)% of $2\nu\beta\beta$ -ME

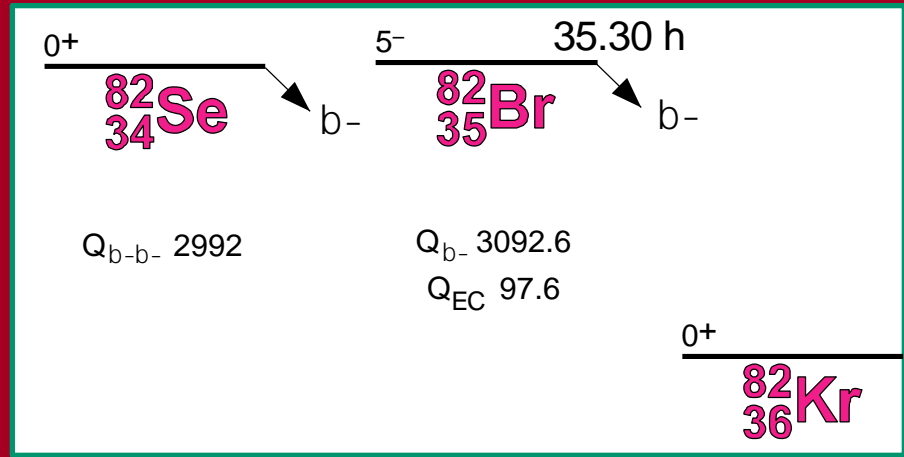
$$M_{DGT} = 0.12 + 0.01 \text{ MeV}^{-1}$$

$$T_{1/2} = (2.1 + 0.1) \times 10^{21} \text{ yr}$$

No need for GT giant resonance contribution
(note: 0.06 MeV⁻¹ are due to low-E triplet, which may not be correlated!!)

^{82}Se

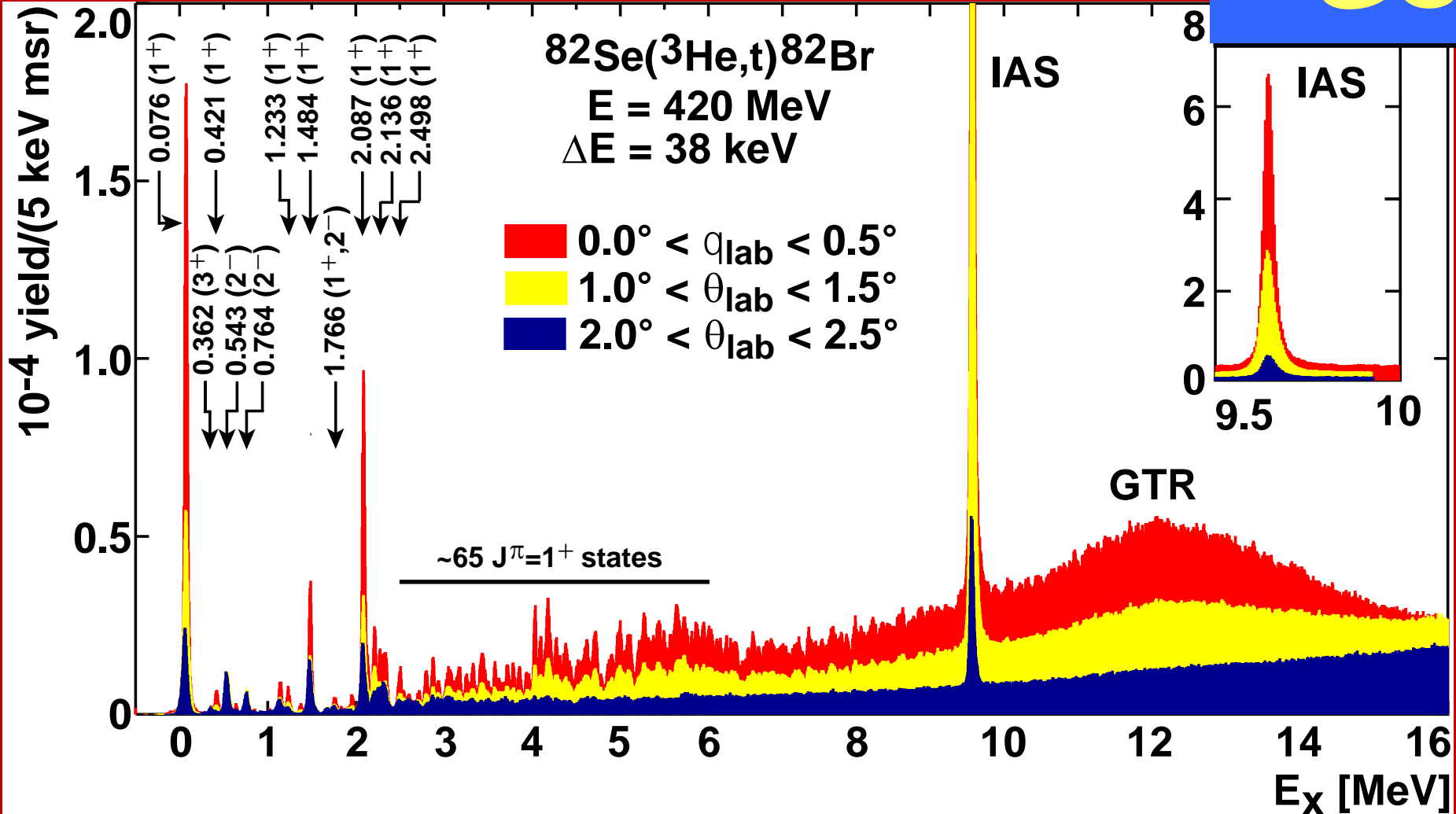
$N-Z=14$



Resolution is the key !!!

**possibly useful for solar
neutrino detection**

^{82}Se



3 isolated GT transition below 2 MeV -
fragmentation recedes to GT resonance

^{96}Zr

$N-Z=16$

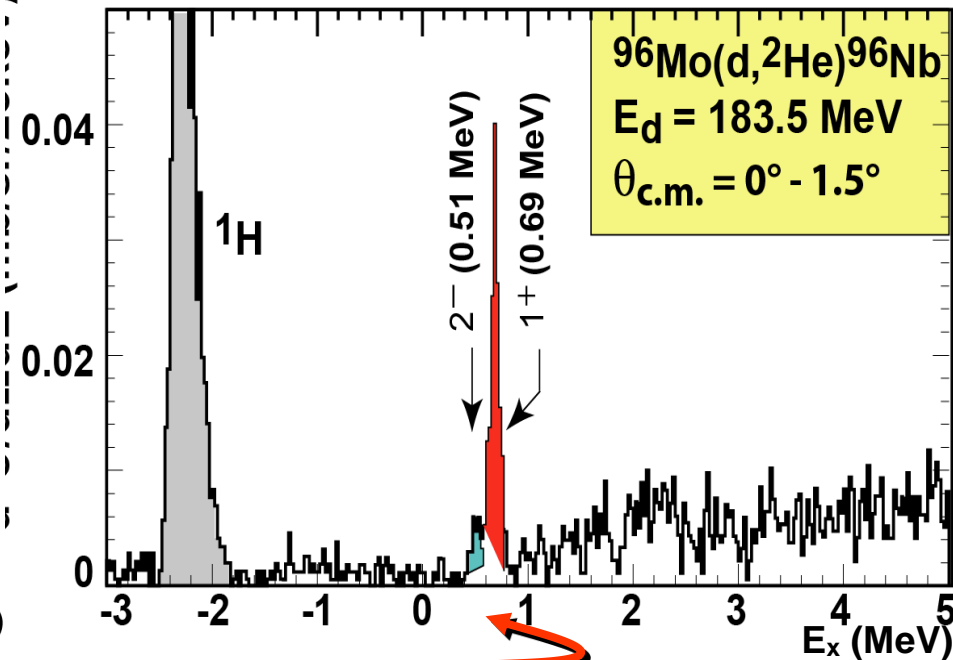
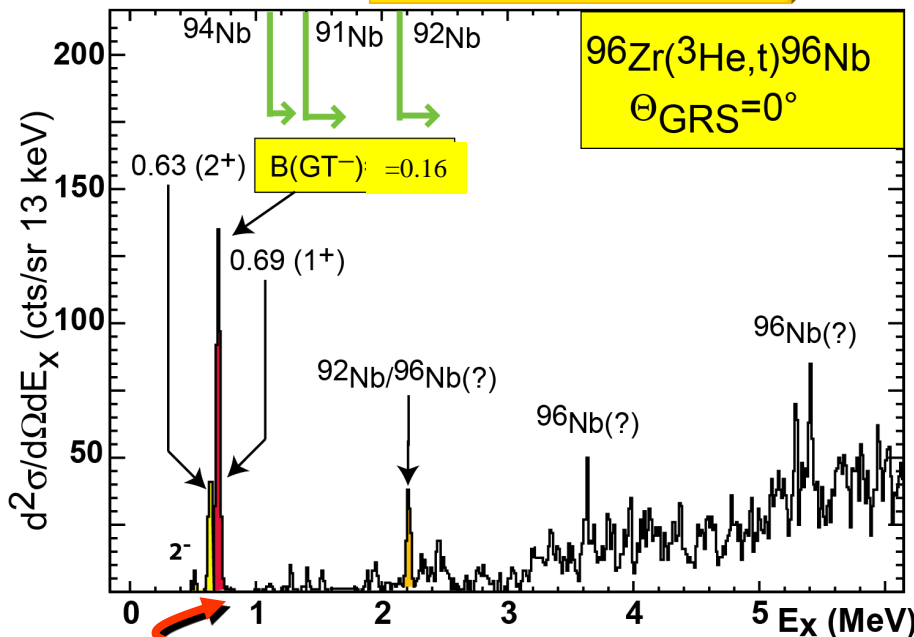
Remember: $B(\text{GT})_{\text{tot}} = 3(N-Z) \sim 50!$

$B(\text{F}) = (N-Z)$

$(^3\text{He}, t)$

$(d, ^2\text{He})$

RCNP 2007/08



$$B(\text{GT}^-) = 0.16$$

$$B(\text{GT}^+) = 0.3$$

Fascination: With only 1 state:

$$T_{1/2}^{\text{calc.}}(2\nu\beta\beta) = (2.1 \pm 0.4) \cdot 10^{19} \text{ years}$$

$$T_{1/2}^{\text{exp.}}(2\nu\beta\beta) = (2.3 \pm 0.2) \cdot 10^{19} \text{ years (NEMO3-result)}$$

100 Mo

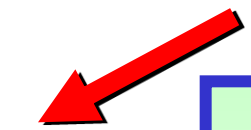
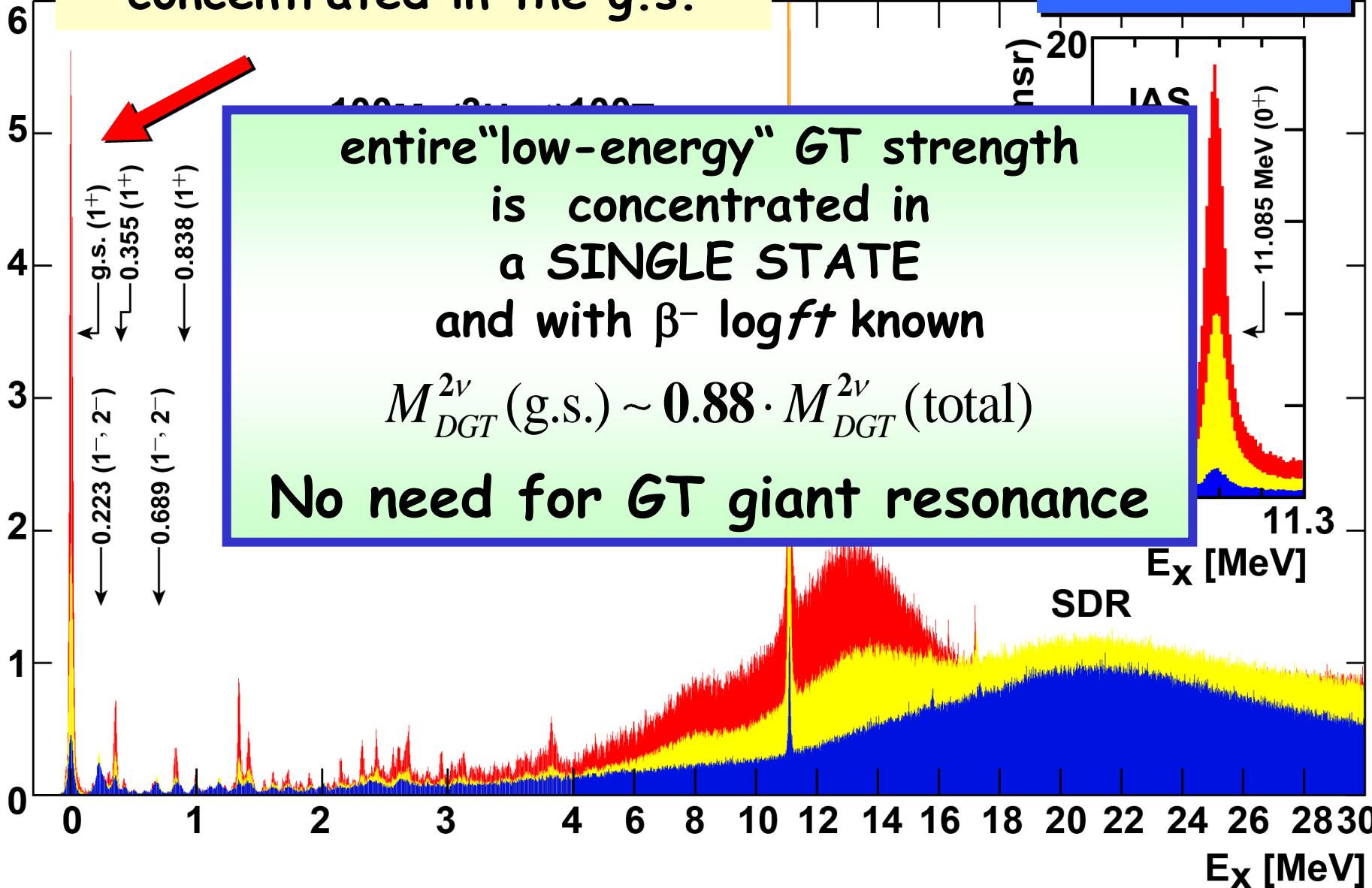
N-Z=16

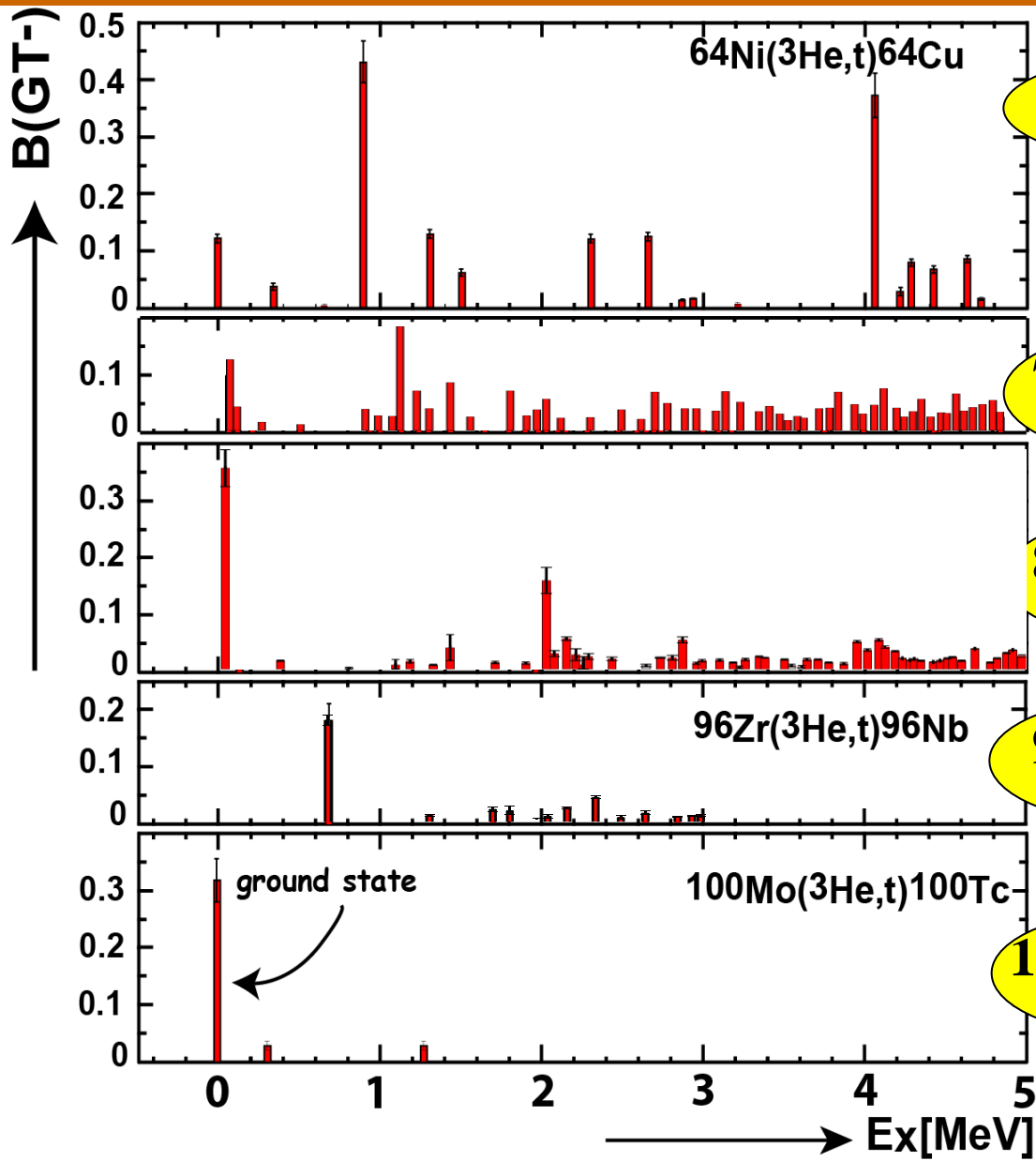
**useful as SN neutrino detector
(sensitive to ν temperature in SN)**

HERE: almost the entire low-E GT strength is concentrated in the g.s.

100Mo

10⁻³ yield/(5 keV msr)





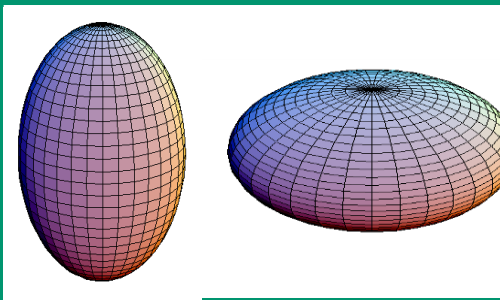
$^{64}\text{Zn}(\epsilon\epsilon, \epsilon\beta^+)$

$^{76}\text{Ge}(\beta\beta^-)$

$^{82}\text{Se}(\beta\beta^-)$

$^{96}\text{Zr}(\beta\beta^-)$

$^{100}\text{Mo}(\beta\beta^-)$



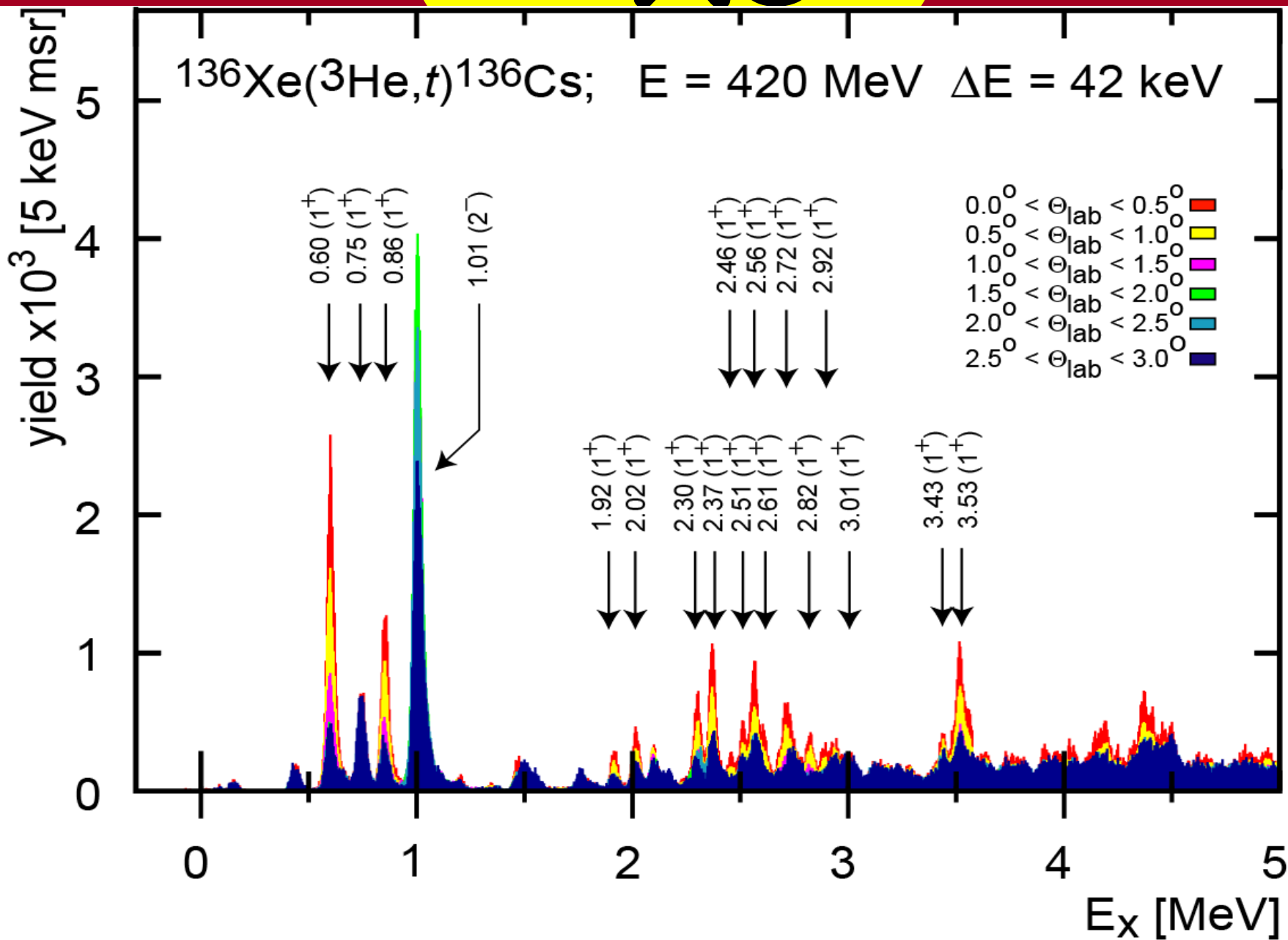
reduced fragmentation of GT strength

^{136}Xe

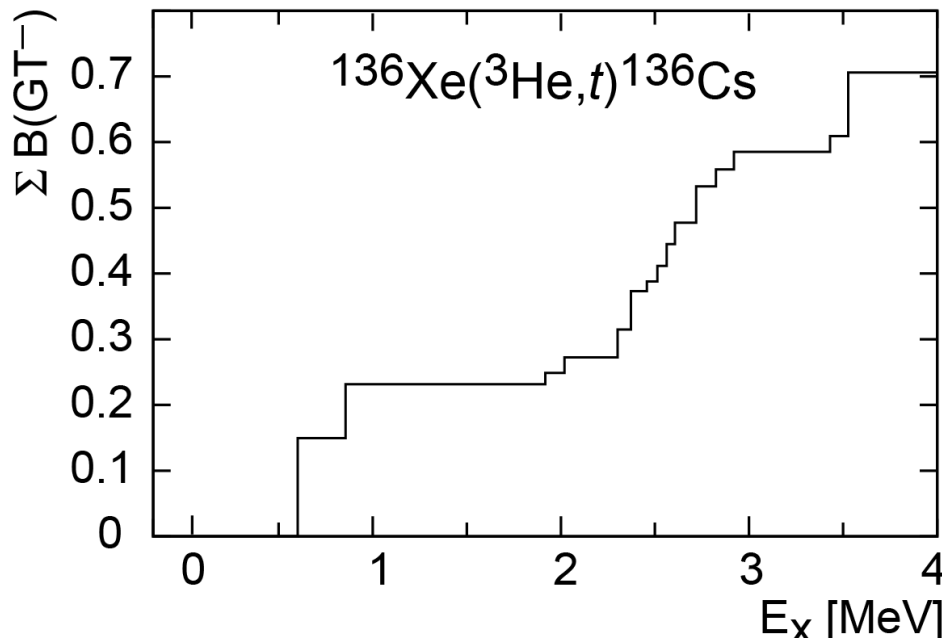
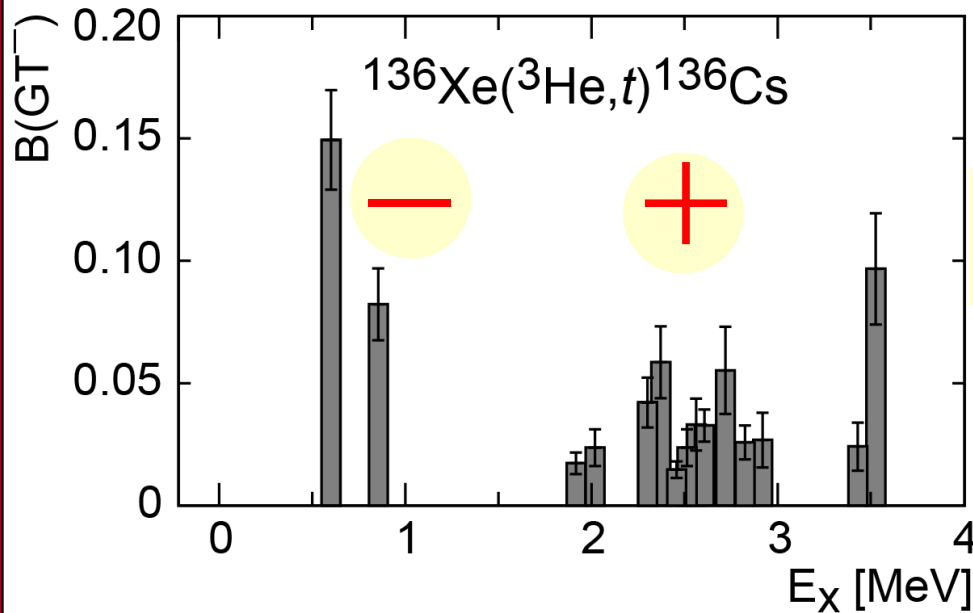
$N-Z=28$

question: why so stable !!!

^{136}Xe



What's the size of the NME?



$$T_{1/2}^{2\nu} = 2.2 \cdot 10^{21} \text{ yr}$$

$$M_{\text{DGT}}^{(2\nu)} \sim 0.019 \text{ MeV}^{-1}$$

all signs positive \rightarrow

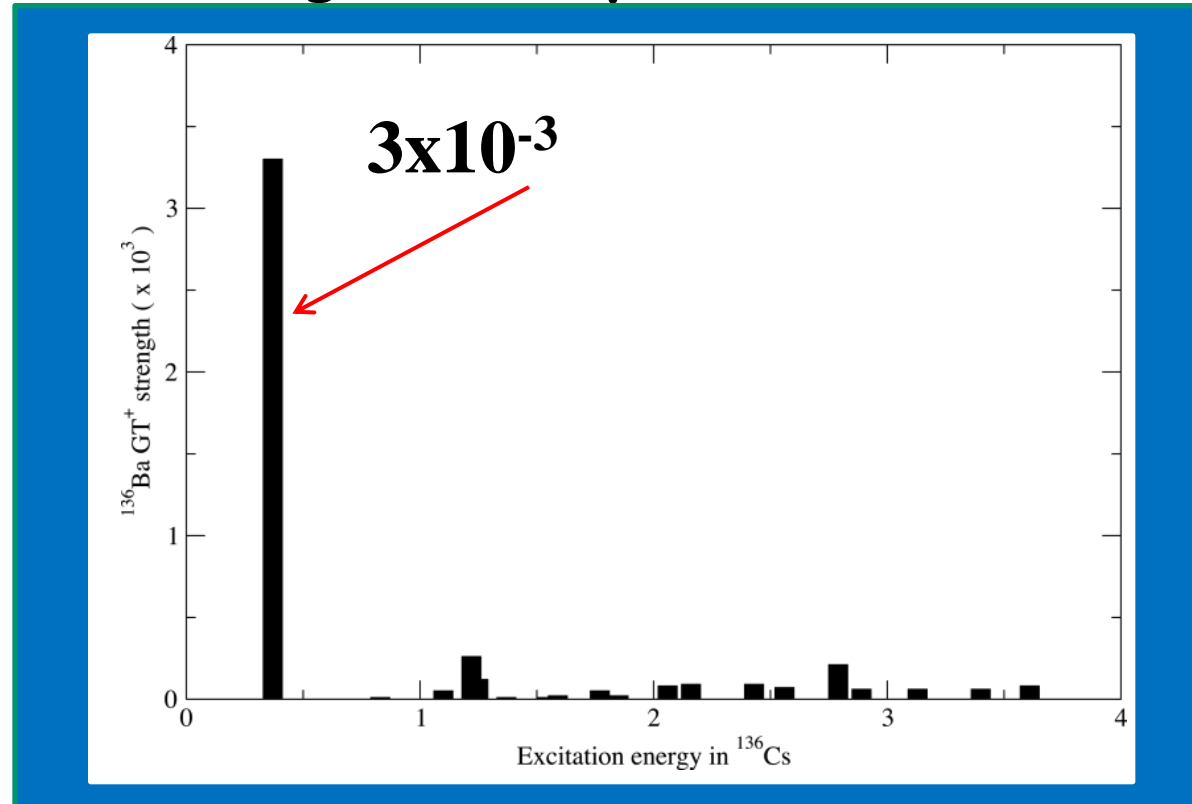
$$B_m \text{ } GT^+ \approx 10^{-2} \cdot B_m \text{ } GT^-$$

$$B_m \text{ } GT^+ \approx 10^{-3} \text{ !!!!}$$

A. Poves (simultaneous to our publication):

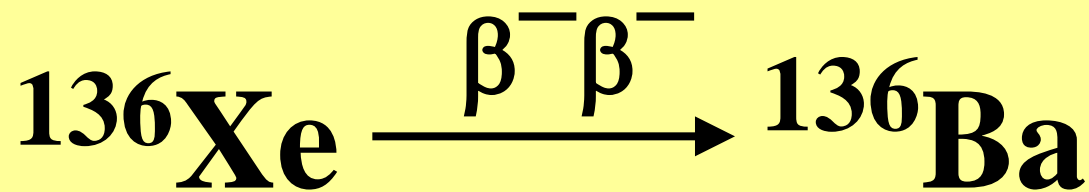
there is no $B(GT^+)$ strength, except for lowest 1^+ state

Recall:
 ^{136}Xe is almost
doubly magic!!



Shell model provides explanation for the deemed „pathologically“ long half-life of ^{136}Xe .

Expt'l test: $^{136}\text{Ba}(d, ^2\text{He})^{136}\text{Cs}$

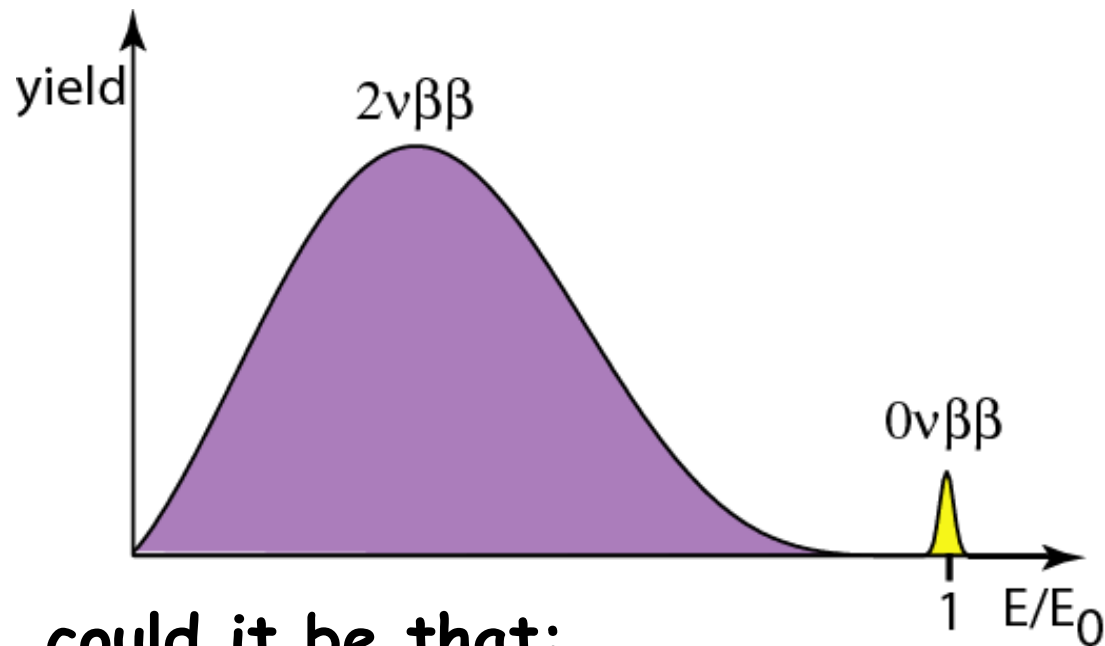


expmt:

$2\nu\beta\beta$ NME is exceptionally small

question:

how does the ME scale in the case of $0\nu\beta\beta$ decay?



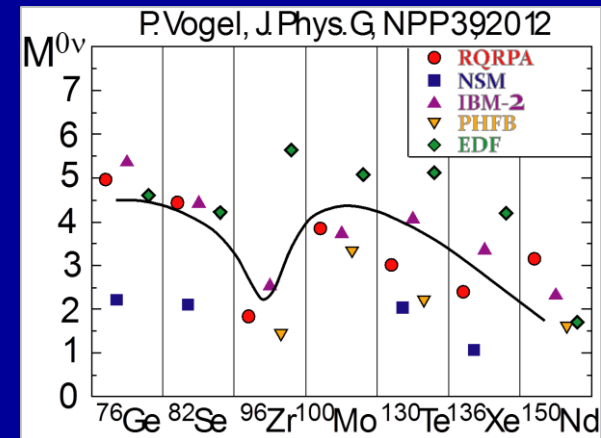
could it be that:

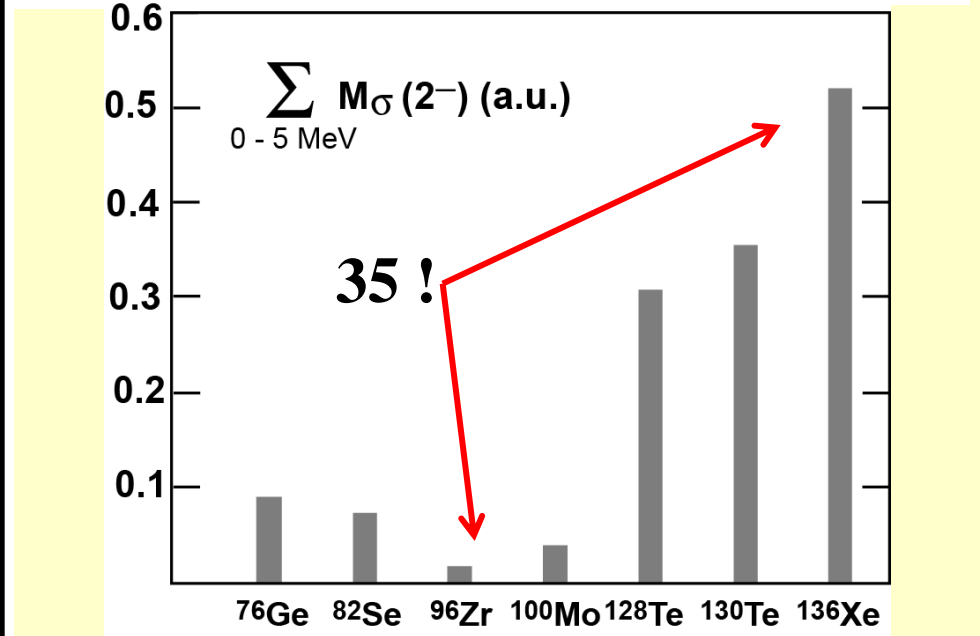
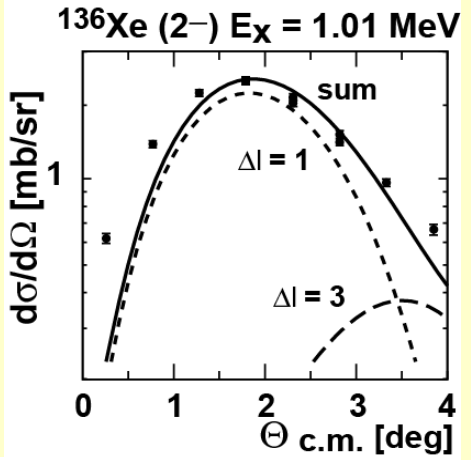
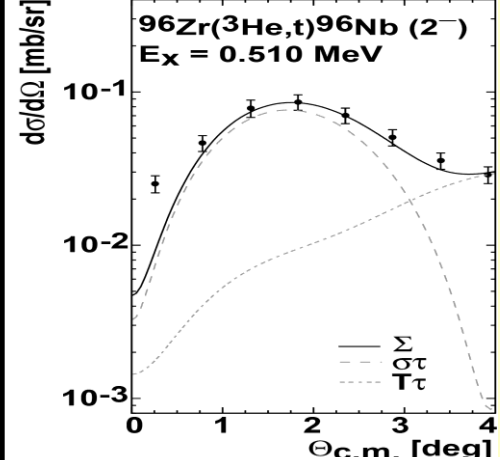
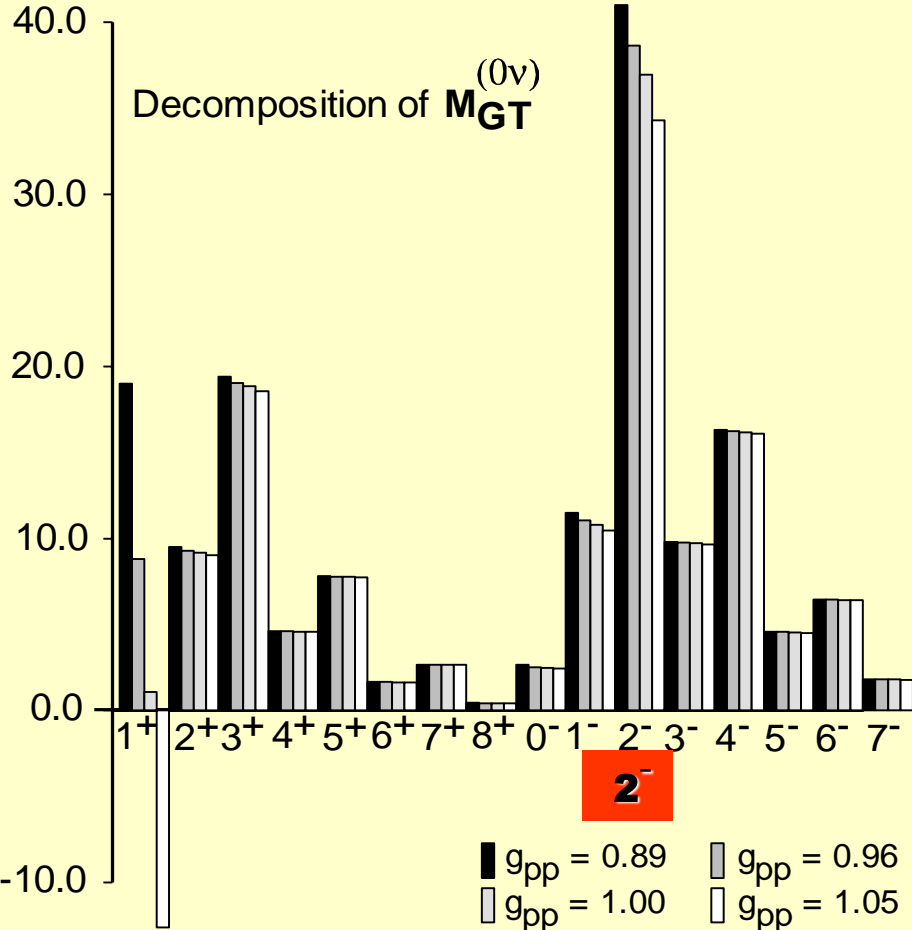
$2\nu\beta\beta$ ME is suppressed **AND**

$0\nu\beta\beta$ ME is enhanced ???

Experiments towards the $0\nu\beta\beta$ NMEs

Here:
2- states and occupation
vacancy numbers
via chargex reactions





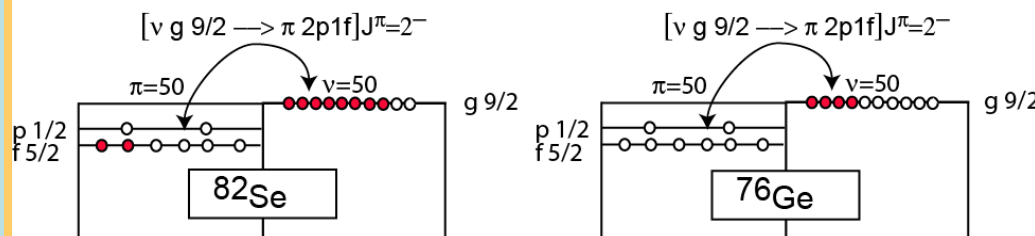
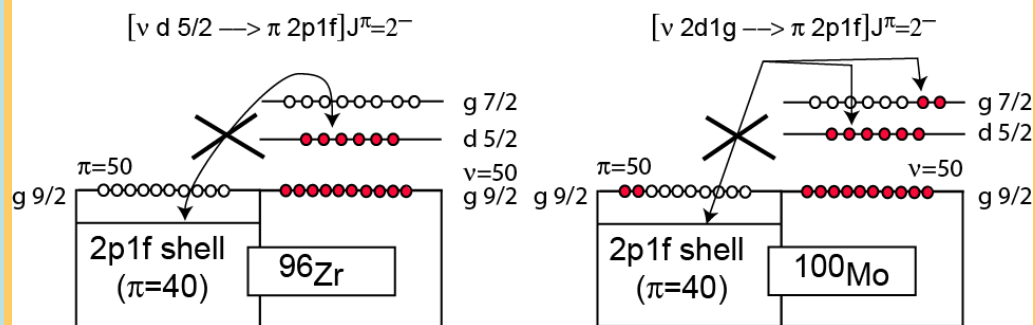
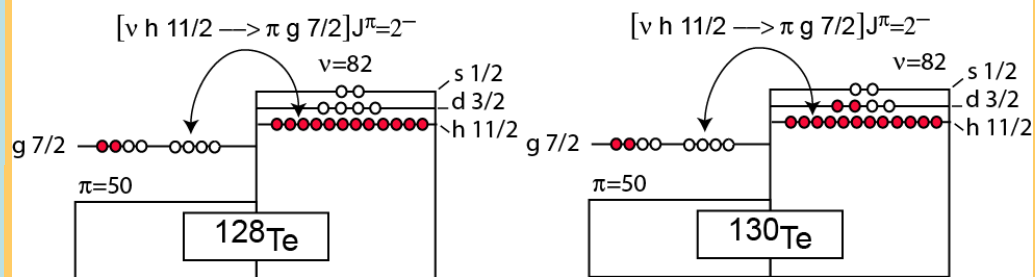
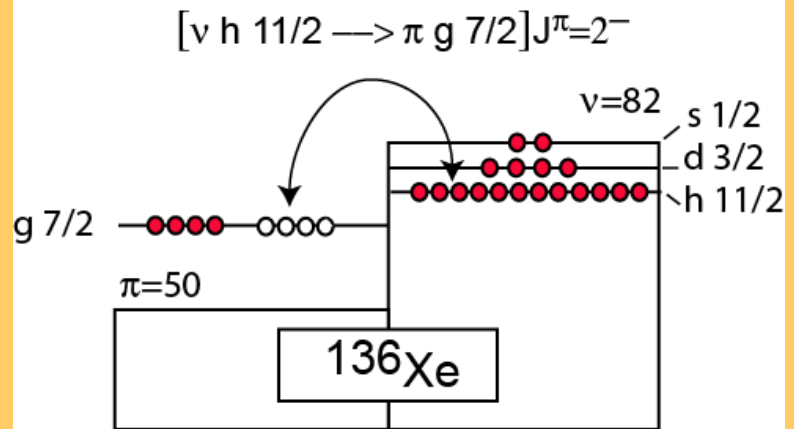
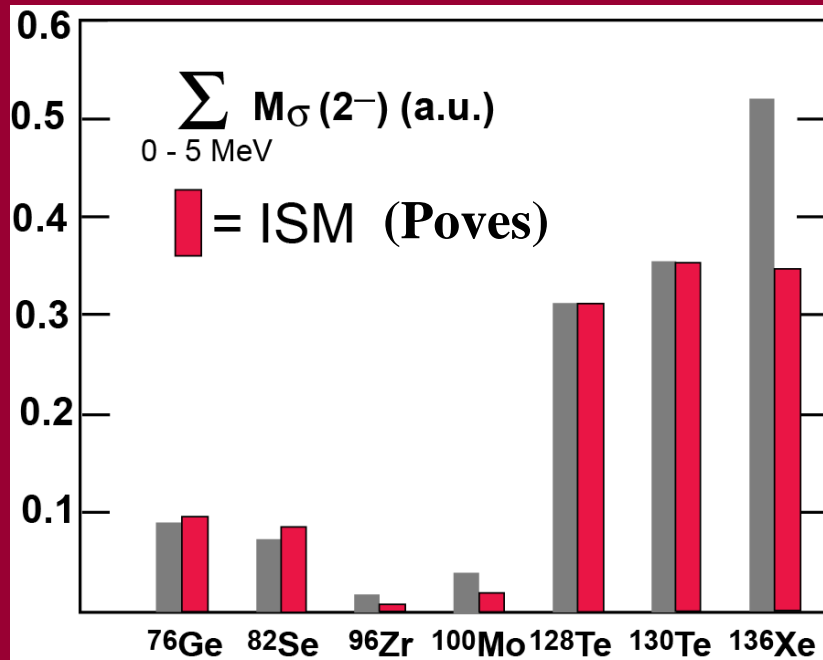
Theory:

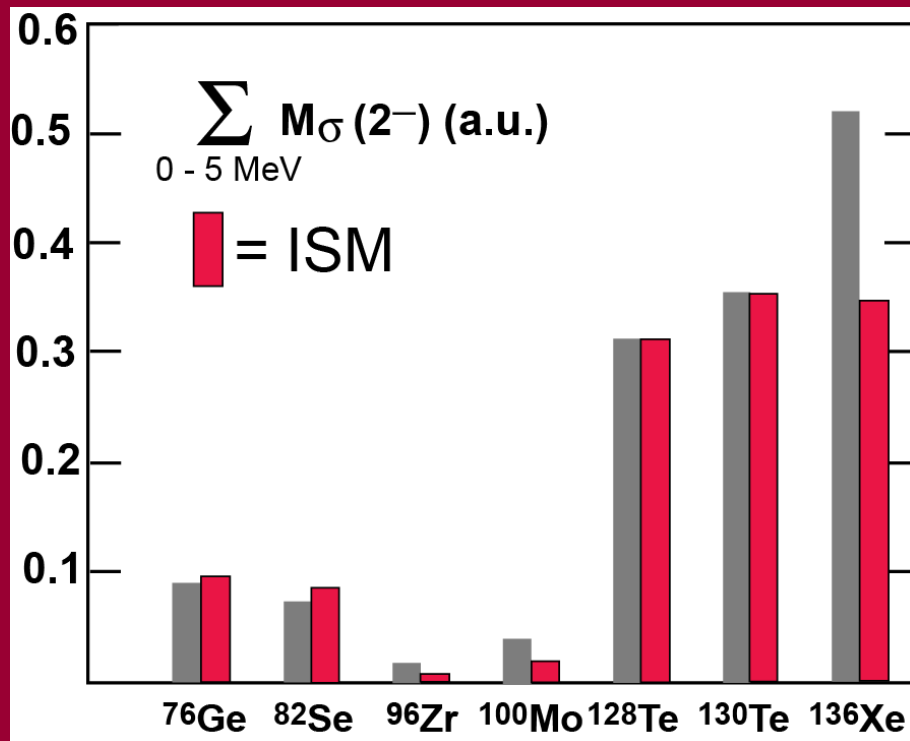
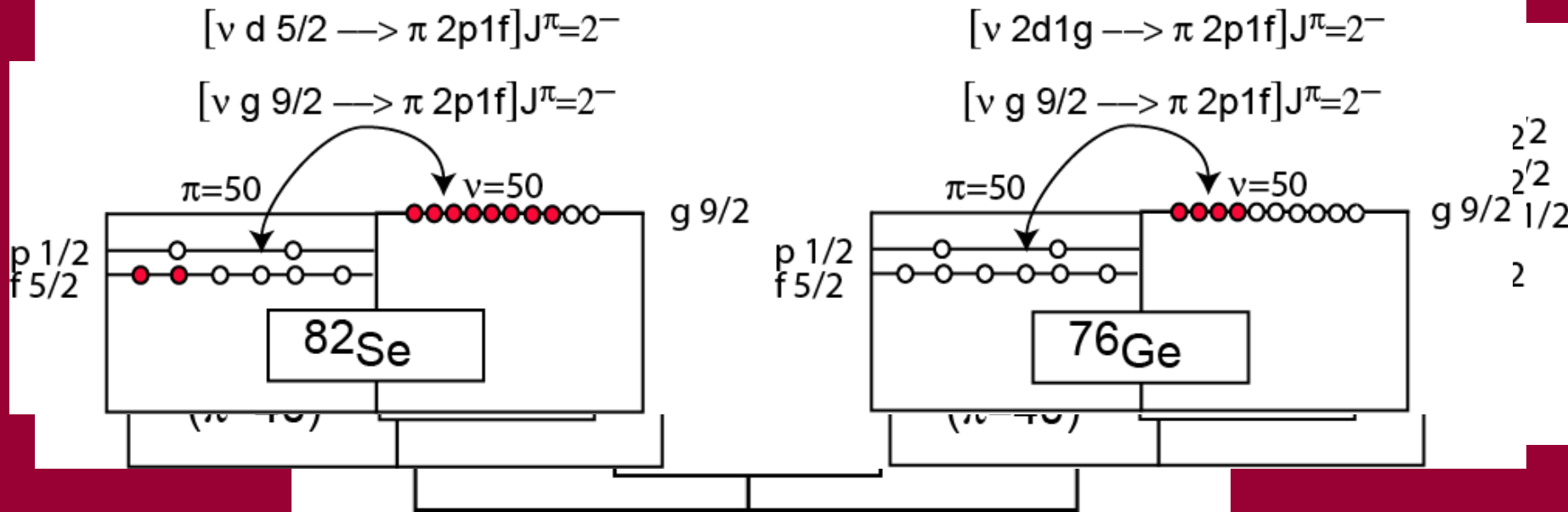
The 2^- strength makes up
 ~ 20-30% of the $0\nu\beta\beta$ ME!!

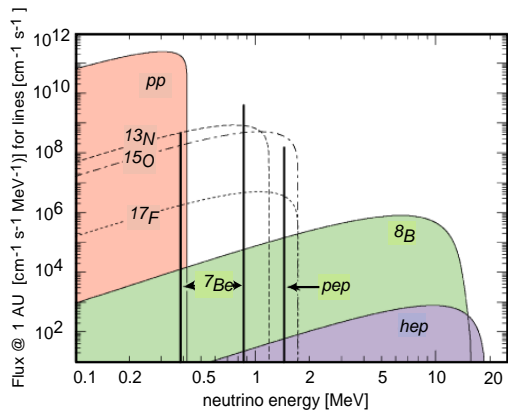
J. Suhonen, Phys. Lett B607, 87 (2005)

Expmt:

^{136}Xe exhibits largest 2^- strength
 $0\nu\beta\beta$ ME enhanced???



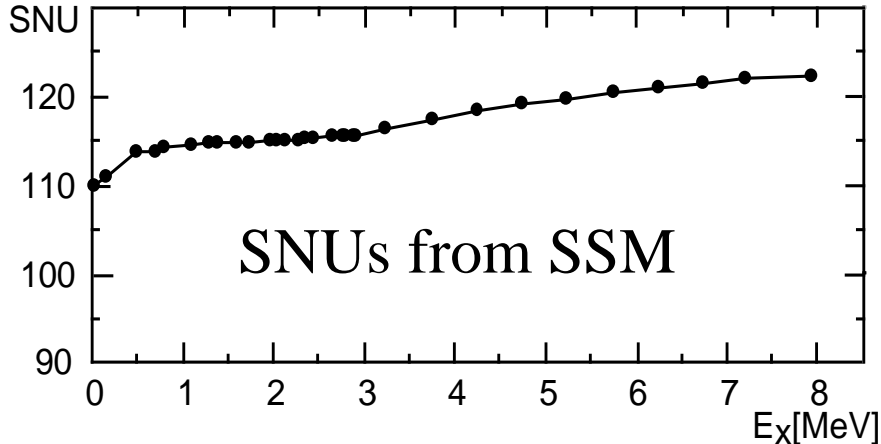
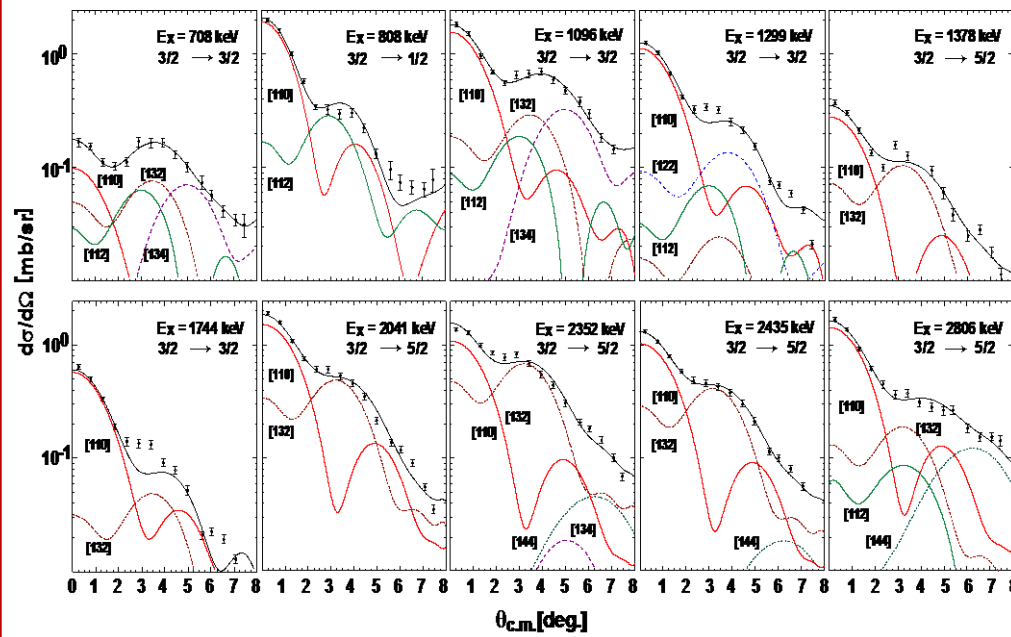
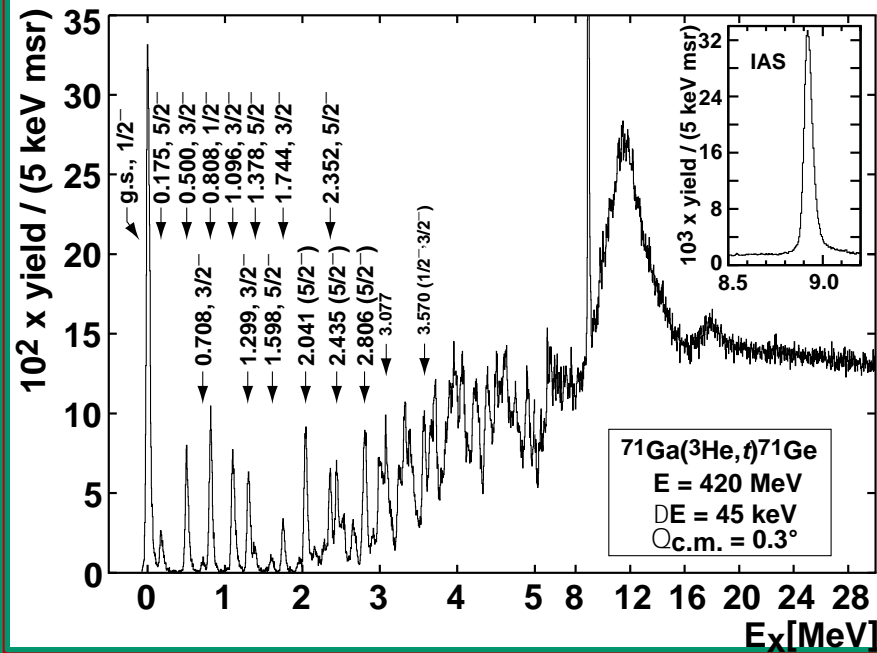




solar neutrino
rates via (${}^3\text{He}, t$)

${}^{71}\text{Ga}(\nu_{\odot}, e^{-})$ SNU from
 ${}^{71}\text{Ga}({}^3\text{He}, t){}^{71}\text{Ge}$ charge-ex reaction

$^{71}\text{Ga}(\nu_{\odot}, e^-)$ SNU from $(^3\text{He}, t)$ charge-exchange reaction

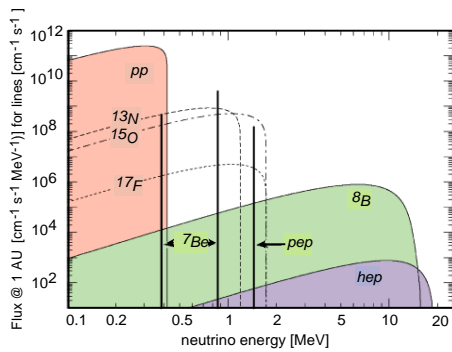


$^{71}\text{Ga}(\nu_{\odot}, e^-)$
 $R = 122.4 \pm 3.4(\text{stat}) \pm 1.1(\text{sys})$
 stat. err. mostly due to CNO ν 's

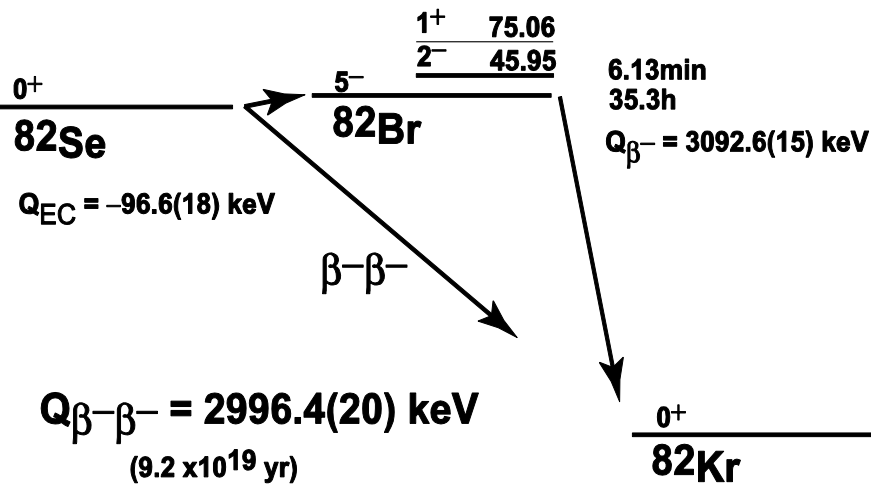
prev'ly: 132 ± 18

DF et al, PRC91,2015

solar neutrino rates via ($^3\text{He}, t$)



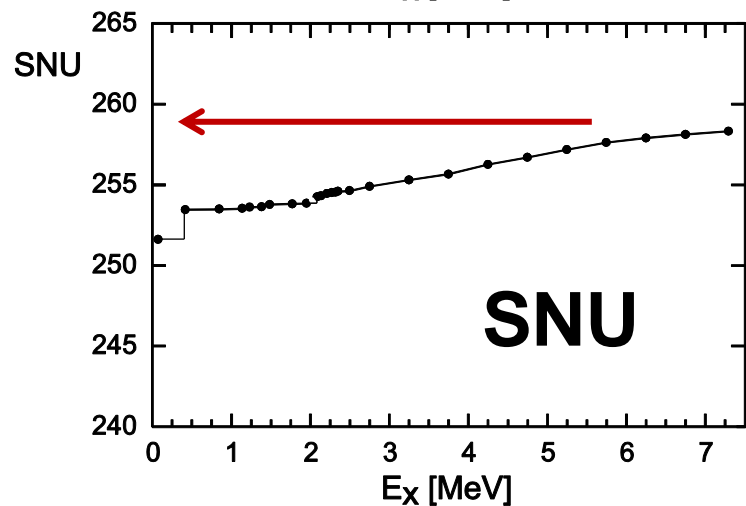
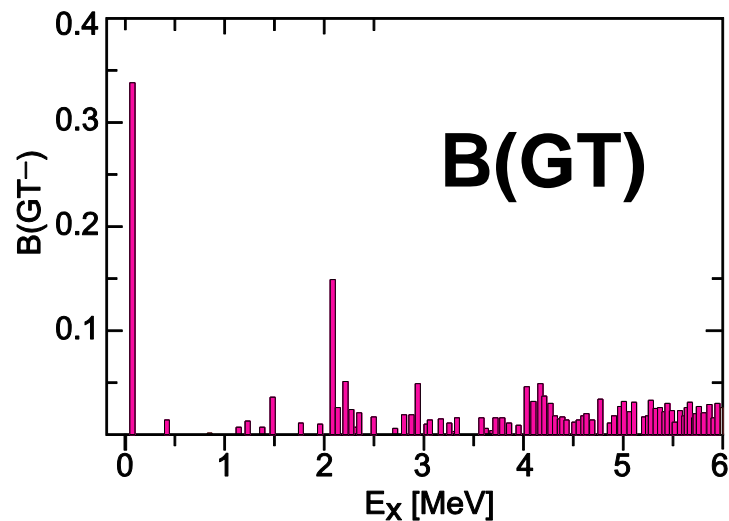
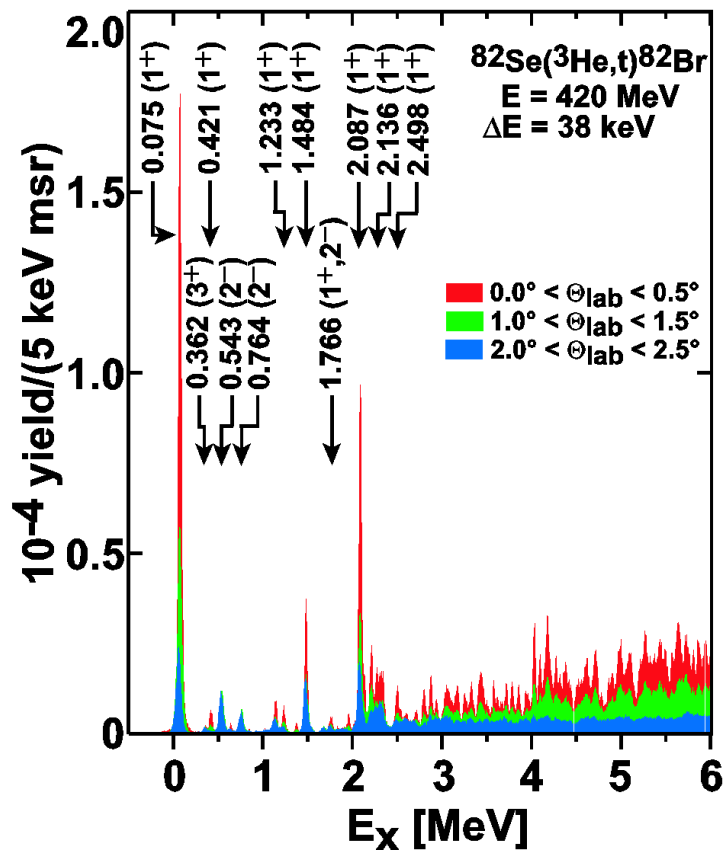
$^{82}\text{Se}(\nu_{\odot}, e^{-})$ SNU's from $^{82}\text{Se}({}^3\text{He}, t){}^{82}\text{Br}$ charge-ex reaction



Advantages:

- low threshold
- enhanced sensitivity to pp-neutrinos
- short life-time against β -decay (35h)
- pp- ν 's in „real time“
- γ -emission, easy to detect

$^{82}\text{Se}(^3\text{He}, t)$ spectrum



Total rate:

258 SNU

Population of 1st 1⁺ state:

97%

pp ν fraction:

76%

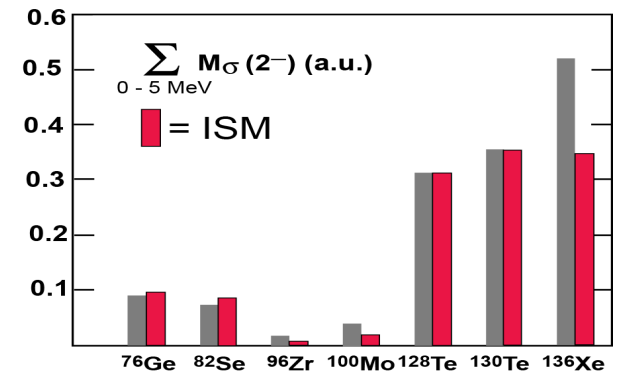
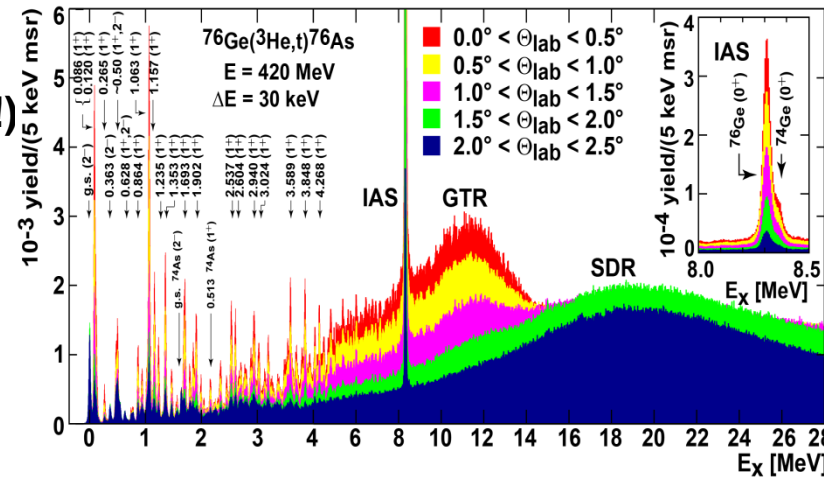
preliminary

Future perspectives of chargex-reactions

- **$\beta\beta$ -decay and nuclear matrix elements**
 - Resolution is key issue (RCNP gives the lead!)
 - need 20 - 30 keV for $(^3\text{He,t})$ & $(\text{d},^2\text{He})$
 - Need to explore proportionality between chargex x-section and $\Delta L \neq 0$ transitions (e.g. 2^- states) in weak interaction (resol'n is key)

- **ν -physics and chargex-reactions**
 - Hadronic chargex and weak-interaction x-sections are fortuitously connected -- exploit this!!
 - solar neutrinos, SN-neutrinos, element synthesis

- **Need to address $0\nu\beta\beta$ decay and the quenching issue urgently!!**
 - New ideas to tackle the problem needed
 - DCX could play a pivotal role for $\beta\beta$ decay
 - DCX reaction could also shed light on the quenching of g_A



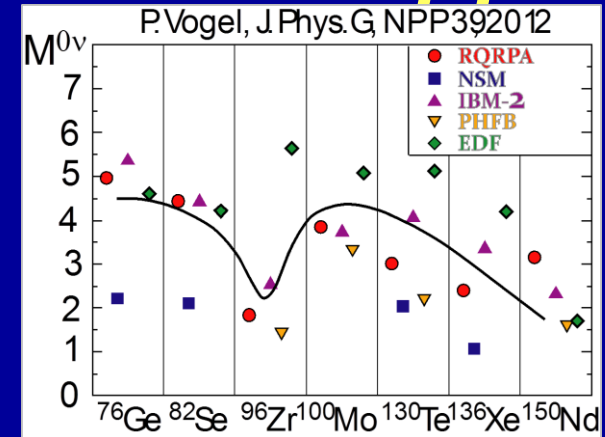
$$^{71}\text{Ga}(\nu_{\odot}, e^-)$$

$$R = 122.4 \pm 3.4 \pm 1.1 \text{ SNU}$$

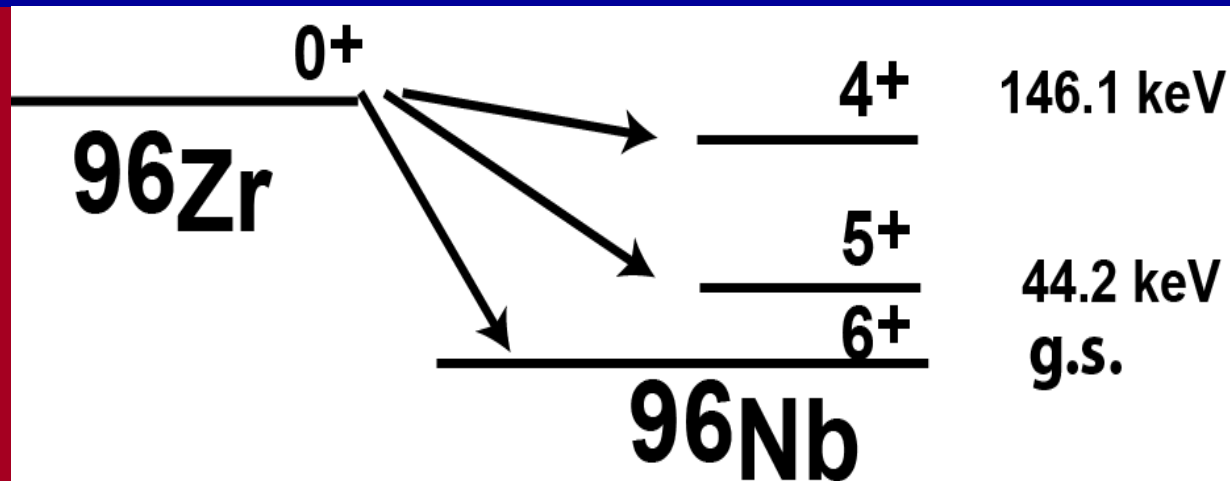
$$^{82}\text{Se}(\nu_{\odot}, e^-)$$

$$R = 258.4 \text{ SNU}$$

experiments towards the $0\nu\beta\beta$ NMEs



the $A=96$ system in $\beta\beta$ decay



Competition between β & $\beta\beta$ decay in ^{96}Zr

two conflicting half-lives:

NEMO-3: $T_{1/2}^{2\nu\beta\beta} = (2.3 \pm 0.2) \times 10^{19} \text{ y}$
 geo-chem: $T_{1/2}^{\beta} = (0.94 \pm 0.32) \times 10^{19} \text{ y}$ ①

can this difference be reconciled ?
 yes, if single β competes with $\beta\beta$ decay

$$(T_{1/2})^{-1} = (T_{1/2}^{2\nu\beta\beta})^{-1} + (T_{1/2}^{\beta})^{-1}$$

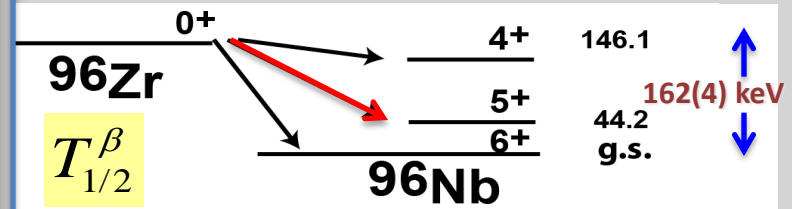
expected $T_{1/2}^{\beta} = (1.6 \pm 0.9) \times 10^{19} \text{ y}$

experiment $T_{1/2}^{\beta} > 2.6 \times 10^{19} \text{ y}$ ②

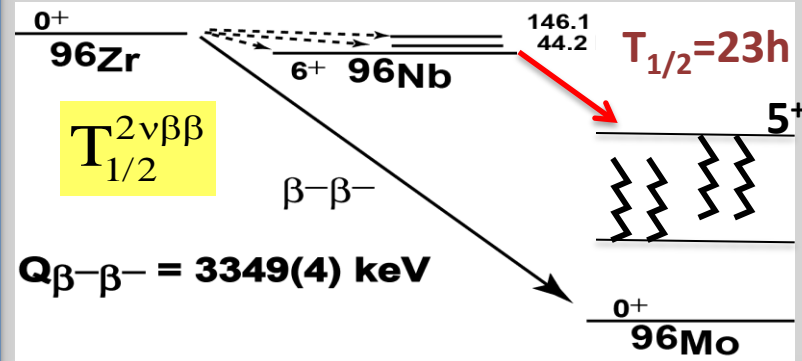
pred. (QRPA) $T_{1/2}^{\beta} = 24 \times 10^{19} \text{ y}$ ③

BUT

$$(T_{1/2}^{\beta})^{-1} \propto o(Q^{13}) \langle M_{\beta}^{4u} \rangle^2$$



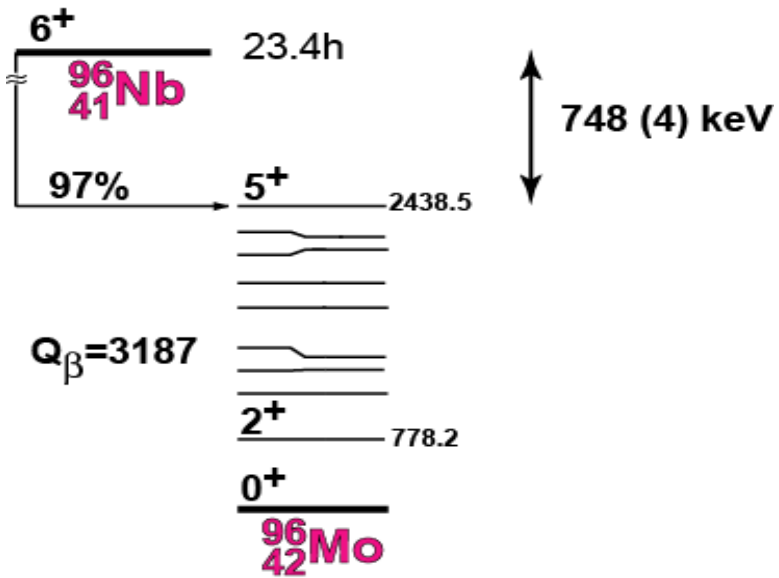
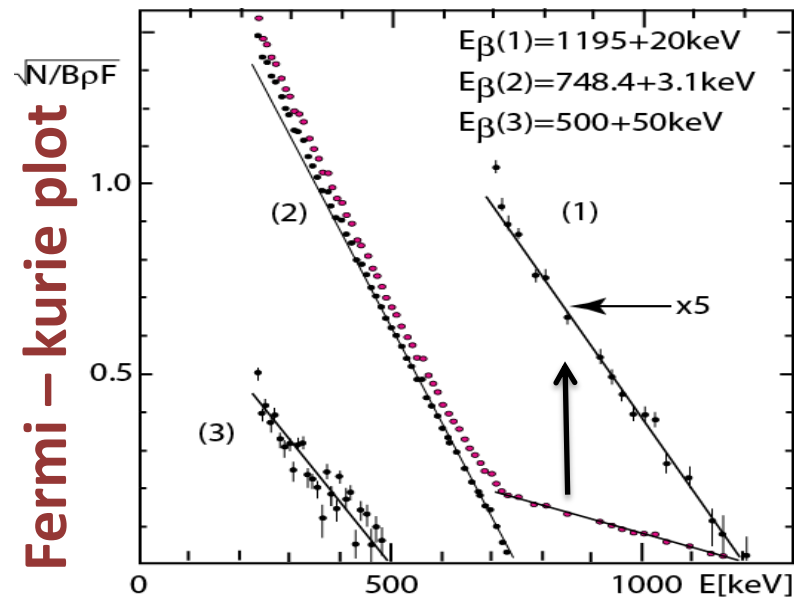
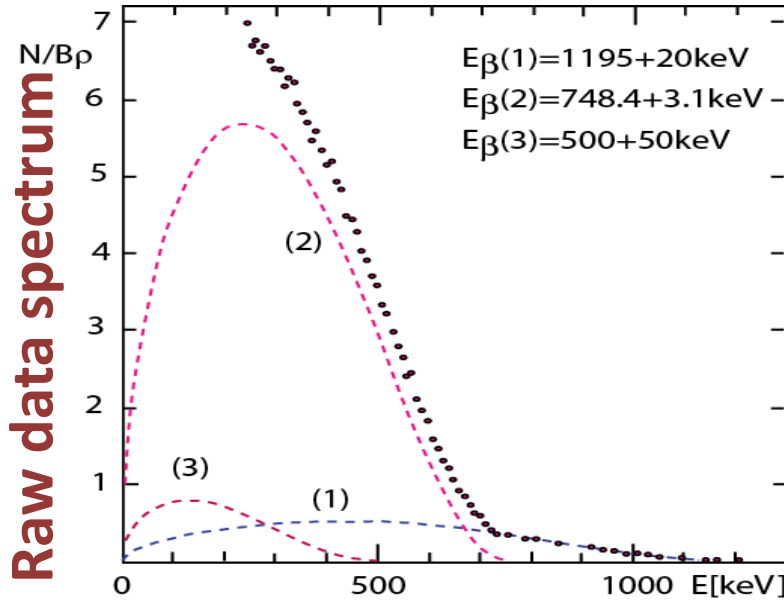
- $0^+ \rightarrow 6^+$ 6-fold non-unique (unobservably long)
- $0^+ \rightarrow 5^+$ 4-fold unique (possible)
- $0^+ \rightarrow 4^+$ 4-fold non-unique (no phase space)



Q-value

$$\longrightarrow M_{\beta}^{4u} \longrightarrow (T_{1/2}^{0\nu\beta\beta})^{-1} \propto Q^5 |M_{\beta\beta}^{0\nu}|^2 \langle m_{\beta\beta} \rangle^2$$

The only ^{96}Nb mass: deduced from β^- decay end point energy {Antman et al. NPA 110 1968}



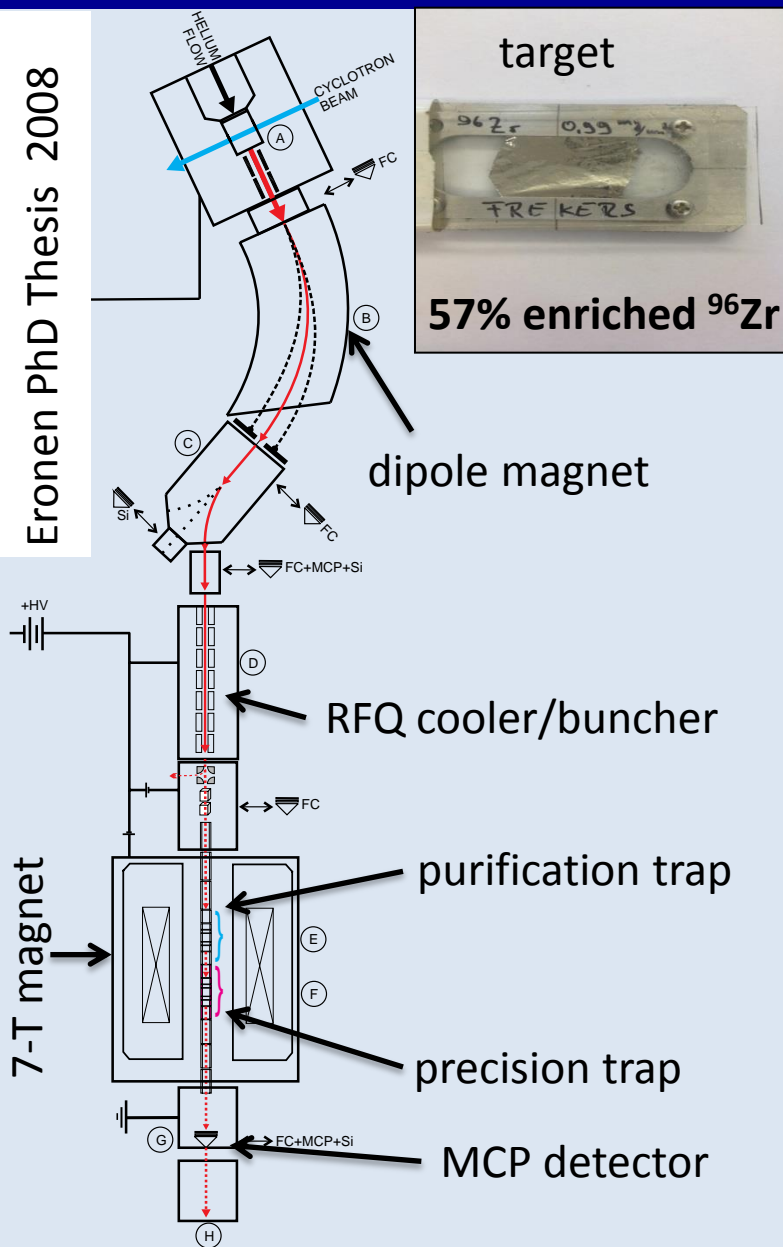
the 1195 keV is NOT a ^{96}Nb long-lived isomer!!

the 500 keV transition does NOT exist.
 → fit depends on 2 transitions, which don't exist!!

→ Q-value is in doubt !!

IGISOL/JYFLTRAP mass measm'nts

Eronen PhD Thesis 2008



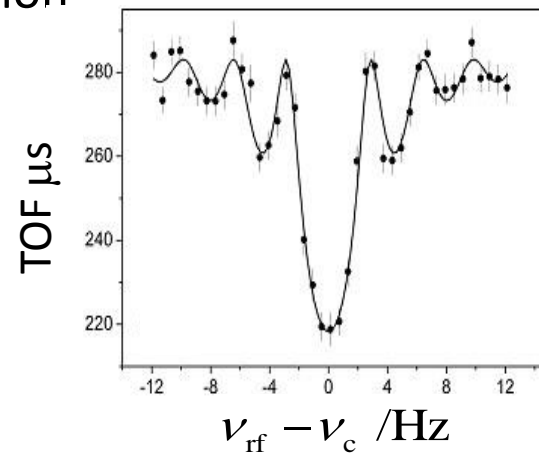
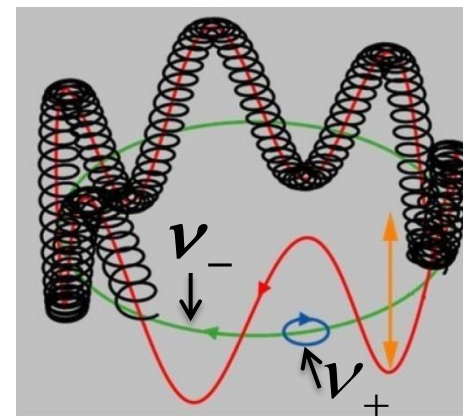
$^{96}\text{Zr} (p, n)^{96}\text{Nb}$ reaction
for production of ^{96}Nb

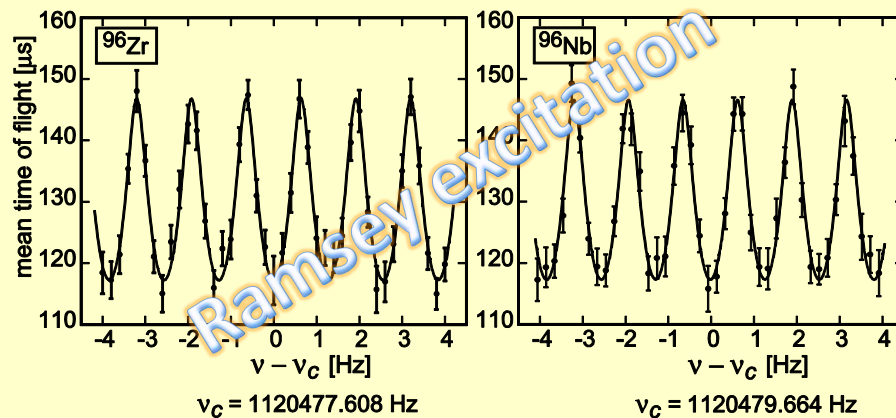
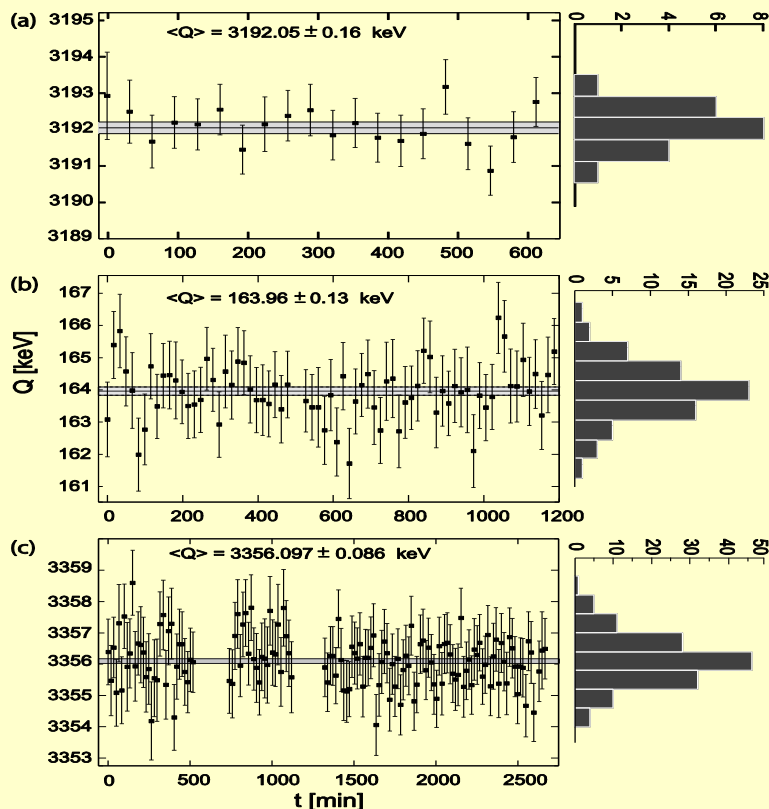
- performing accurate mass measurements via cyclotron frequency

$$v_c = \frac{1}{2\pi} \frac{q}{m} \cdot B$$

$$v_c = v_- + v_+$$

- frequency determination done by **TOF-ICR** technique



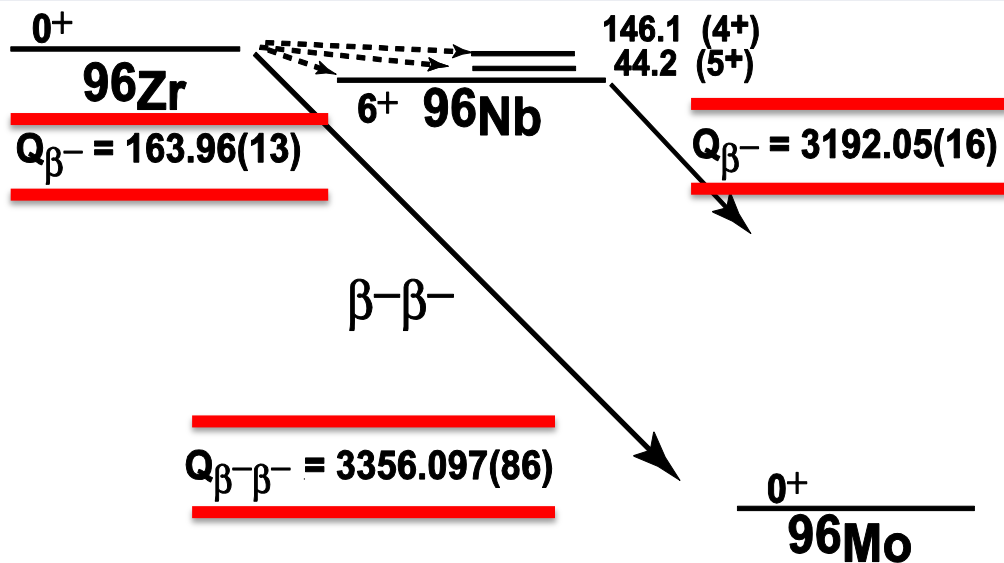


96Zr

$$Q_{\beta\beta} = 3356.097 \pm 0.086 \text{ keV}$$

7.1 keV higher than AME2012

$$Q_{\beta} = 163.96 \pm 0.13 \text{ keV}$$



$$T_{1/2}(\text{QRPA}) = 24 \times 10^{19} \text{ yr}$$

$$T_{1/2}(\text{SM}) = \frac{11}{g_A^2} \times 10^{19} \text{ yr}$$

$$T_{1/2}(\text{exp}) > 2.3 \times 10^{19} \text{ yr}$$

