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The 750 GeV resonance and Dark Matter as heavy pions

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Antipin, Redi, Strumia, Vigiani (arXiv:1503.08749) Redi, Strumia, Tesi, Vigiani (arXiv:1602.07297)

Outline

Vectorlike Confinement

Phenomenology of techni-pions as dark matter and digamma candidates

Conclusions

Vectorlike Confinement

Kilic, Okui, Sundrum (arXiv:0906.0577)

$$Q = \sum_i (N_{\mathrm{TC}}, r_i) \oplus (\bar{N}_{\mathrm{TC}}, \bar{r}_i)$$

new fermions in a vectorial rep of the SM charged under a new strong interaction that confines without breaking the SM:

$$\langle \mathcal{Q} \bar{\mathcal{Q}} \rangle \sim \,$$
 SM singlet

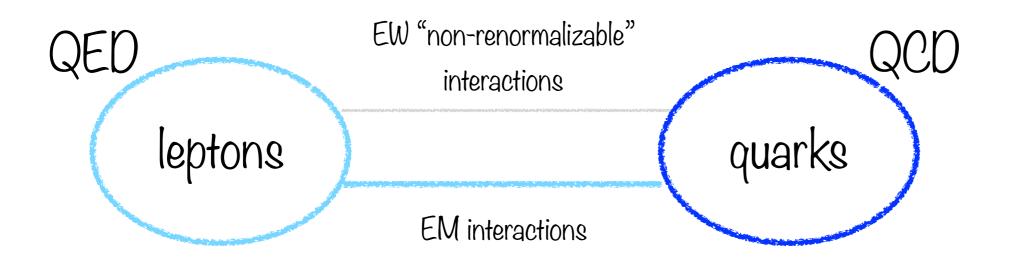
let's call it "techni-color" (\neq old Technicolor theories !!!)

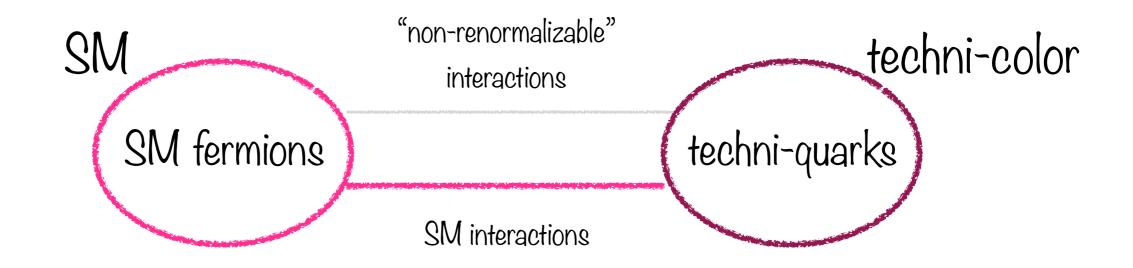
EWPT MFV

DM candidates:

stable thanks to accidental symmetries

VC is like seeing QCD from the point of view of QED:





 $\mathrm{SU}(N_{\mathrm{TC}})$ with N_{TF} techni-flavors



techni-baryons

techni-color singlets made of N_{TC} techni-quarks

fermions if N_{TC} odd bosons if N_{TC} even

SSB of the chiral symmetry

$$SU(N_{TF})_L \otimes SU(N_{TF})_R \to SU(N_{TF})$$

 $N_{
m TF}^2-1$ techni-pions $({\cal Q}{ar {\cal Q}})$ pseudo Goldstone Bosons

$$m_{\mathrm{TC}\pi}^2 \sim m_{\mathcal{Q}} \Lambda_{\mathrm{TC}} + \underbrace{\Delta_{\mathrm{gauge}}}_{\sim 1/100\Lambda_{\mathrm{TC}}^2}$$

mass of techni-pions $\sim 1/10$ of the mass of techni-baryons

Accidental Symmetries

techni-baryons as DM

Antipin, Redi, Strumia, Vigiani (arXiv:1503.08749)

* techni-baryon number : $\mathcal{Q} \to e^{i\alpha}\mathcal{Q}$ the lightest techni-baryon is stable

see A. Mitridate's talk tomorrow

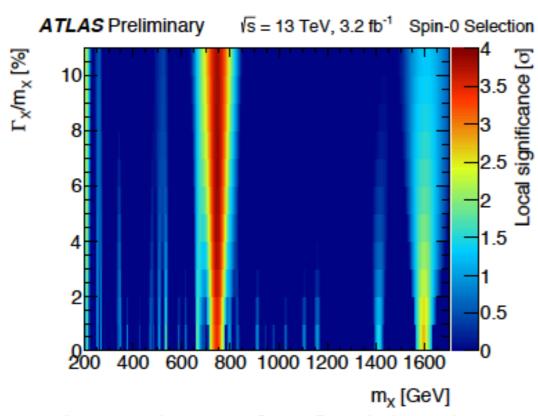
techni-pions as DM

Redi, Strumia, Tesi, Vigiani (arXiv:1602.07297)

* species number : the lightest $\mathcal{Q}_i\mathcal{Q}_j$, $i \neq j$ is stable in analogy with $\pi^+(u\bar{d})$

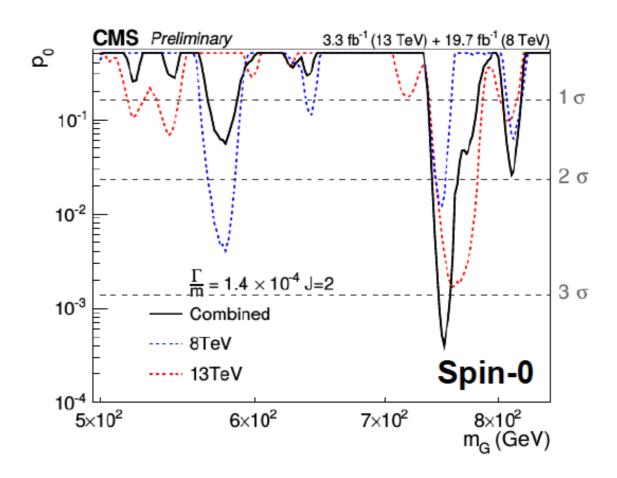
imes G-parity : $\mathcal{Q} \to e^{i\pi T_2} \mathcal{Q}^c$ techni-pions with even (odd) isospin are G-even (odd)

The \digamma particle



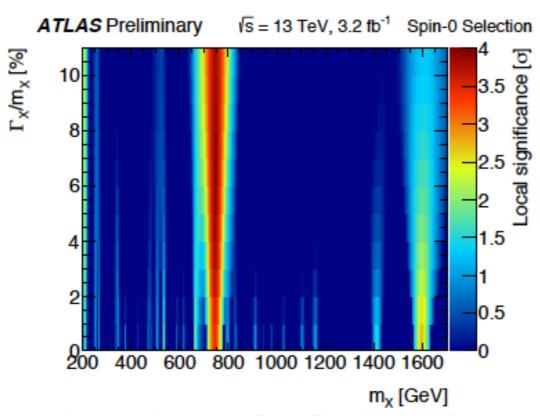
- Largest deviation from B-only hypothesis
 - \checkmark m_X ~ 750 GeV, Γ_X ~ 45 GeV (6%)
 - \checkmark Local Z = 3.9 σ
 - ✓ Global $Z = 2.0 \sigma$

from Moriond 2016 Conference



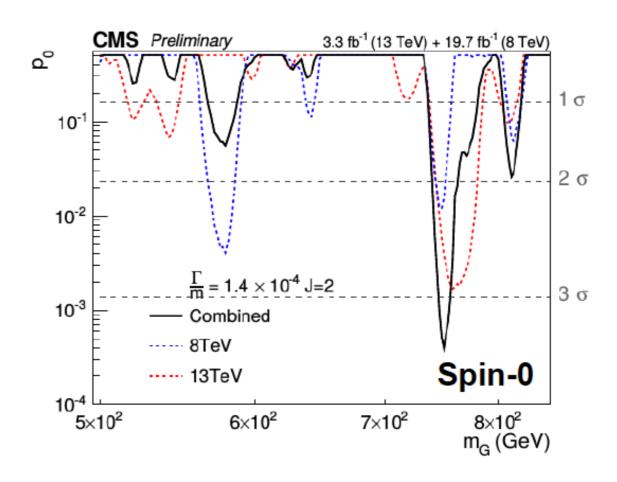
- Largest excess observed at $m_x = 750$ GeV and for narrow width.
 - Local significance: 3.4σ
 - Taking into account mass range 500-3500GeV (and all signal hypotheses), "global" significance becomes 1.6σ

The F particle



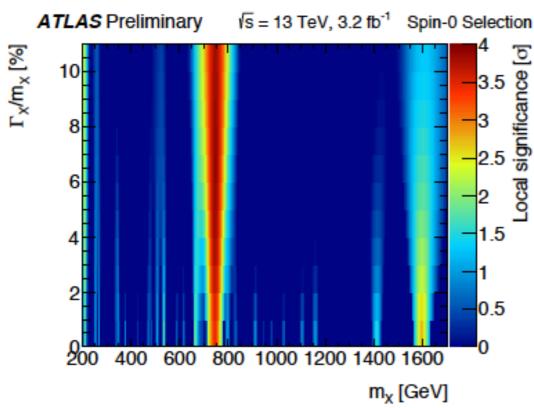
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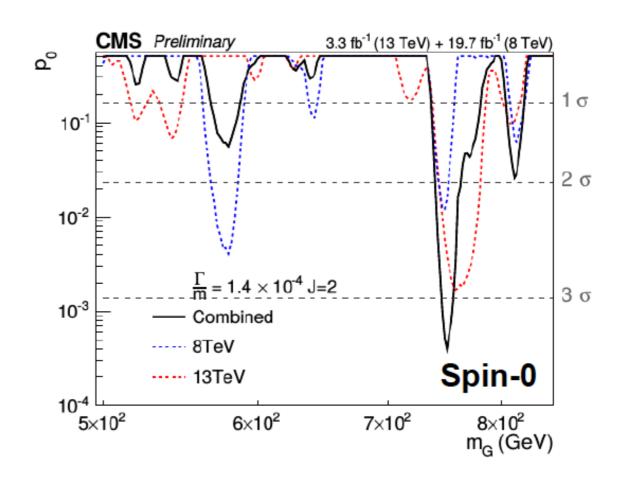
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The F particle



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 - \checkmark m_X ~ 750 GeV, Γ _X ~ 45 GeV (6%)
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 - ✓ Global Z = 2.0 σ

from Moriond 2016 Conference



- not enough to claim a discovery, but enough to be excited!
- \blacktriangleright Largest excess observed at $m_x = 750 \text{GeV}$ and for narrow width.
 - Local significance: 3.4σ
 - Taking into account mass range 500-3500GeV (and all signal hypotheses), "global" significance becomes 1.6σ

Some properties...

spin 0 or spin 2? different analysis by CMS and ATLAS

production mechanism the background grows by 2.3 at $750\,\mathrm{GeV}$, the signal by

 $gg,b\overline{b},c\overline{c},s\overline{s}$ favored by the compatibility with Run-I

CP odd, CP even (or CP is violated)? Franceschini et al. (arXiv:1604.06446)

$$F \rightarrow \gamma^* \gamma^* \rightarrow 4l$$

 $F \to \gamma^* \gamma^* \to 4l$ the rate is only 10^{-3} of the digamma rate

$$F \to \gamma \gamma \to e^+ e^- e^+ e^-$$
 in matter, small $e^+ e^-$ angle

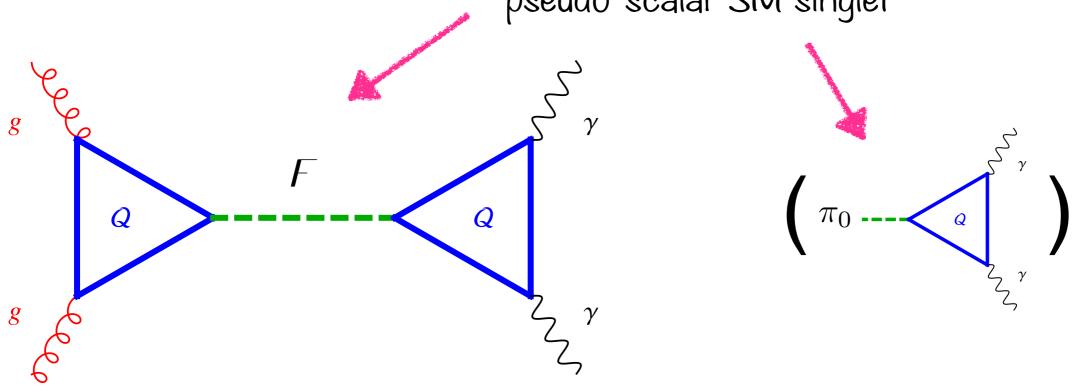
$$pp \rightarrow Fjj$$



pp o Fjj from the di-jet angular distribution

Simple and predictive





$$\mathcal{L}_{\text{anomalies}} \supset -\frac{1}{16\pi^2} \frac{F}{f} \left[g_1^2 c_B B_{\mu\nu} \tilde{B}^{\mu\nu} + g_2^2 c_W W_{\mu\nu}^a \tilde{W}_a^{\mu\nu} + g_3^2 c_G G_{\mu\nu}^a \tilde{G}_a^{\mu\nu} \right]$$

$$c_B = 2N_{\mathrm{TC}}\mathrm{Tr}(T_{\digamma}Y^2) \ c_W\delta^{ab} = 2N_{\mathrm{TC}}\mathrm{Tr}(T_{\digamma}T^aT^b) \ c_G\delta^{AB} = 2N_{\mathrm{TC}}\mathrm{Tr}(T_{\digamma}T^AT^B)$$

$$\frac{\Gamma(F \to VV)}{M_F} = c_V^2 \frac{\kappa_V \alpha_V^2}{64\pi^3} \frac{M_F^2}{f^2}$$

Q	$N_{ m TF}$	$rac{c_B^{\eta}}{N_{ m TC}}$	$rac{c_W^{\eta}}{N_{ m TC}}$	$rac{c_G^{\eta}}{N_{ m TC}}$	$\frac{\Gamma^{\eta}_{\gamma Z}}{\Gamma^{\eta}_{\gamma \gamma}}$	$\frac{\Gamma^{\eta}_{ZZ}}{\Gamma^{\eta}_{\gamma\gamma}}$	$\frac{\Gamma^{\eta}_{GG}}{\Gamma^{\eta}_{\gamma\gamma}}$	$\frac{f({ m GeV})}{N_{ m TC}}$	$\frac{c_B^{\eta'}}{N_{ m TC}}$	$\frac{{c_W^{\eta'}}}{N_{ m TC}}$	$rac{c_G^{\eta'}}{N_{ m TC}}$	$\frac{\Gamma^{\eta'}_{\gamma Z}}{\Gamma^{\eta'}_{\gamma \gamma}}$	$\frac{{\Gamma_{ZZ}^{\eta'}}}{{\Gamma_{\gamma\gamma}^{\eta'}}}$	$\frac{\Gamma_{GG}^{\eta'}}{\Gamma_{\gamma\gamma}^{\eta'}}$	$\frac{f({ m GeV})}{N_{ m TC}}$
$D \oplus L$	5	$\frac{1}{6}\sqrt{\frac{5}{3}}$	$\frac{1}{2}\sqrt{\frac{3}{5}}$	$-\frac{1}{\sqrt{15}}$	1.8	4.7	240	96	$\frac{1}{3}\sqrt{\frac{5}{2}}$	$\frac{1}{\sqrt{10}}$	$\frac{1}{\sqrt{10}}$	0.23	1.9	180	_
$D \oplus U$	6	$\frac{1}{\sqrt{3}}$	ò	0	0.57	0.082	0	-	$\frac{5}{3\sqrt{3}}$	0	$\frac{1}{\sqrt{3}}$	0.57	0.082	470	150
$D \oplus E$	4	$\frac{4}{3}\sqrt{\frac{2}{3}}$	0	$-\frac{1}{2\sqrt{6}}$	0.57	0.082	46	170	$\frac{2\sqrt{2}}{3}$	0	$\frac{1}{2\sqrt{2}}$	0.57	0.082	180	_
$D\oplus Q$	9	$-\frac{1}{6}$	$\frac{1}{2}$	0	17	22	0	_	$\frac{1}{3\sqrt{2}}$	$\frac{1}{\sqrt{2}}$	$\frac{1}{\sqrt{2}}$	2.9	6.1	740	150
$D\oplus T$	6	$\frac{8}{3\sqrt{3}}$	$\frac{2}{\sqrt{3}}$	$-\frac{1}{2\sqrt{3}}$	0.43	2.4	15	430	$\begin{array}{r} \frac{1}{3\sqrt{2}} \\ \frac{10}{3\sqrt{3}} \end{array}$	$\frac{\frac{1}{\sqrt{2}}}{\frac{2}{\sqrt{3}}}$	$\frac{\frac{1}{\sqrt{2}}}{\frac{1}{2\sqrt{3}}}$	0.23	1.9	12	_
$L \oplus U$	5	$\frac{7}{6\sqrt{15}}$	$-\frac{1}{2}\sqrt{\frac{3}{5}}$	$\frac{1}{\sqrt{15}}$	200	180	12000	_	$\frac{11}{3\sqrt{10}}$	$\frac{1}{\sqrt{10}}$	$\frac{1}{\sqrt{10}}$	0.0027	0.83	60	230
$L\oplus Q$	8	$ \begin{array}{r} 6\sqrt{15} \\ -\frac{2}{3\sqrt{3}} \end{array} $	0	$\frac{1}{2\sqrt{3}}$	0.57	0.082	740	61	$\frac{1}{3}$	1	$\frac{1}{2}$	2.9	6.1	180	210
$L \oplus S$	8	$\frac{7}{12\sqrt{3}}$	$-\frac{\sqrt{3}}{4}$	$\frac{5}{4\sqrt{3}}$	200	180	74000	_	$\frac{19}{12}$	$\frac{1}{4}$	$\frac{5}{4}$	0.095	0.47	610	290
$U \oplus E$	4	$\frac{5}{3\sqrt{6}}$	0	$-\frac{1}{2\sqrt{6}}$	0.57	0.082	120	110	$\frac{7}{3\sqrt{2}}$	0	$\frac{1}{2\sqrt{2}}$	0.57	0.082	60	260
$U\oplus Q$	9		$\frac{1}{2}$	0	32	17	0	_	$\frac{1}{\sqrt{2}}$	$\frac{1}{\sqrt{2}}$	$\frac{1}{\sqrt{2}}$	0.79	3.0	330	220
$U \oplus V$	6	$-\frac{4}{3\sqrt{3}}$	$\frac{\frac{1}{2}}{\frac{2}{\sqrt{3}}}$	$-\frac{1}{2\sqrt{3}}$	83	82	740	-	$ \begin{array}{r} \frac{1}{\sqrt{2}} \\ \frac{4}{3\sqrt{3}} \\ \frac{2\sqrt{2}}{3} \end{array} $	$\frac{\frac{1}{\sqrt{2}}}{\frac{2}{\sqrt{3}}}$	$ \frac{\frac{1}{2\sqrt{2}}}{\frac{\sqrt{2}}{\sqrt{2}}} $ $ \frac{1}{2\sqrt{3}} $	1.5	4.1	29	310
$U\oplus N$	4	$-\frac{2}{3}\sqrt{\frac{2}{3}}$	0	$-\frac{1}{2\sqrt{6}}$	0.57	0.082	180	87	$\frac{2\sqrt{2}}{3}$	0	$\frac{1}{2\sqrt{2}}$	0.57	0.082	180	_
$E \oplus Q$	7	$-\frac{2}{3}\sqrt{\frac{2}{3}}$ $-\frac{5}{6}\sqrt{\frac{7}{3}}$	$\frac{1}{2}\sqrt{\frac{3}{7}}$	$\frac{1}{\sqrt{21}}$	3.6	0.52	70	150	$\frac{1}{3}\sqrt{\frac{7}{2}}$	$\frac{3}{\sqrt{14}}$	$\sqrt{\frac{2}{7}}$	1.2	3.7	180	_
$E \oplus S$	7	$-\frac{10}{3\sqrt{21}}$	0	$\frac{5}{2\sqrt{21}}$	0.57	0.082	740	120	$\frac{11}{3}\sqrt{\frac{2}{7}}$	0	$5\sqrt{14}$	0.57	0.082	610	310
$S \oplus V$	9	$-\frac{8}{9}$	$\frac{4}{3}$	$-\frac{5}{6}$	83	82	4600	_	$\frac{8\sqrt{2}}{9}$	$\frac{2\sqrt{2}}{3}$	$\frac{5}{3\sqrt{2}}$	0.43	2.4	380	350
$S\oplus N$	7	$-\frac{8}{3\sqrt{21}}$	0_	$-\frac{5}{2\sqrt{21}}$	0.57	0.082	1200	93	$\frac{8}{3}\sqrt{\frac{2}{7}}$	0	$\frac{5}{\sqrt{14}}$	0.57	0.082	1200	230
$N\oplus Q$	7	$\frac{1}{6\sqrt{21}}$	$\frac{1}{2}\sqrt{\frac{3}{7}}$	$\frac{1}{\sqrt{21}}$	4.9	8.5	470	58	$\frac{1}{3\sqrt{14}}$	$\frac{3}{\sqrt{14}}$	$\sqrt{\frac{2}{7}}$	4.9	8.5	470	140

Digamma candidates in VC models

Q	$N_{ m TF}$	$rac{c_B^{\eta}}{N_{ m TC}}$	$rac{c_W^{\eta}}{N_{ m TC}}$	$rac{c_G^{\eta}}{N_{ m TC}}$	$\frac{\Gamma^{\eta}_{\gamma Z}}{\Gamma^{\eta}_{\gamma \gamma}}$	$\frac{\Gamma_{ZZ}^{\eta}}{\Gamma_{\gamma\gamma}^{\eta}}$	$\frac{\Gamma^{\eta}_{GG}}{\Gamma^{\eta}_{\gamma\gamma}}$	$rac{f({ m GeV})}{N_{ m TC}}$
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$D \oplus E$	4	$\frac{4}{3}\sqrt{\frac{2}{3}}$	0	$-\frac{1}{2\sqrt{6}}$	0.57	0.082	46	170
$D \oplus T$	6	$\frac{\frac{8}{3\sqrt{3}}}{\frac{1}{3\sqrt{3}}}$	$\frac{2}{\sqrt{3}}$	$-\frac{1}{2\sqrt{3}}$	0.43	2.4	15	430
$U\oplus N$	4	$-\frac{2}{3}\sqrt{\frac{2}{3}}$	0	$-\frac{1}{2\sqrt{6}}$	0.57	0.082	180	87
$E \oplus Q$	7	$-\frac{5}{6}\sqrt{\frac{7}{3}}$	$\frac{1}{2}\sqrt{\frac{3}{7}}$	$\frac{1}{\sqrt{21}}$	3.6	0.52	70	150

the digamma candidate is η , expected an heavier singlet $\sim {
m TeV}$

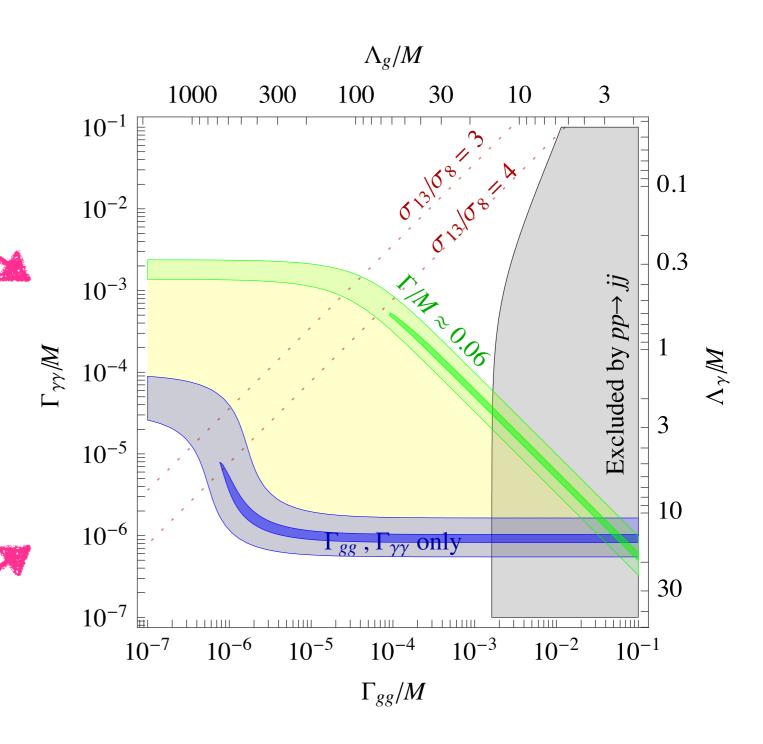
- >> bounds on extra singlets from digamma searches
- \varkappa bounds on colored objects from di-jet searches ($\lesssim {\rm TeV}$)

- + other models where the digamma candidate is η' and a lighter singlet η is expected
- more difficult to threat

Reproducing the signal

another decay channel maybe DM?

$$\Gamma_{\mathcal{F}} \approx \Gamma_{gg} < 1 \text{GeV}$$

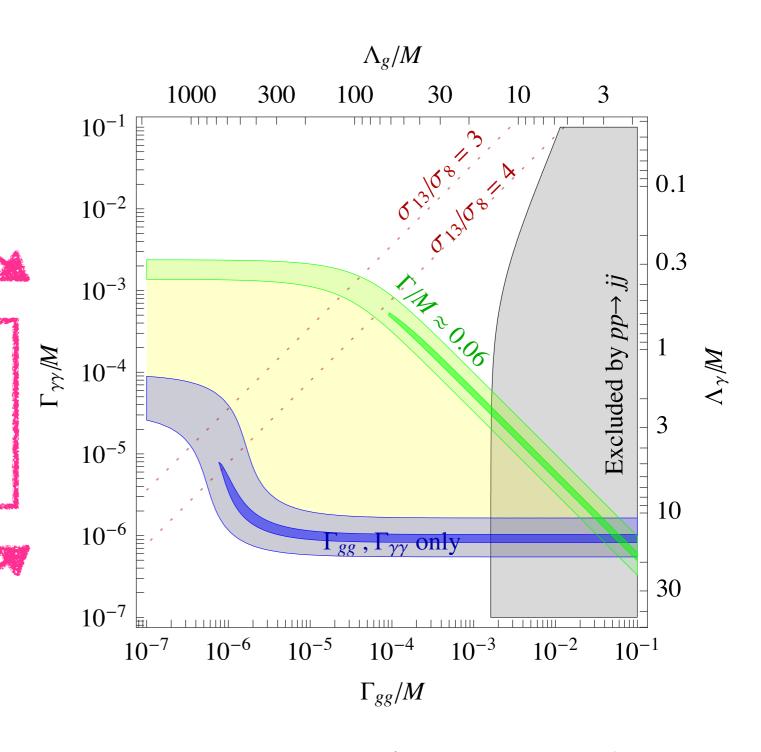


Franceschini et al. (arXiv:1512.04933)

Reproducing the signal

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$$\Gamma_F \approx \Gamma_{gg} < 1 \text{GeV}$$



Franceschini et al. (arXiv:1512.04933)

An explicit model:

DM from species number

$$Q = N_1 \oplus N_2 \oplus U = (1,1)_0 \oplus (1,1)_0 \oplus (\bar{3},1)_{-2/3}$$

pseudo-scalars with given anomalous couplings to SM gauge bosons, up to mixing effects

$$\eta_1 \sim U\bar{U} - 3/2(N_1\bar{N}_1 + N_2\bar{N}_2)$$

$$\eta_2 \sim N_1\bar{N}_1 - N_2\bar{N}_2$$

$$\eta' \sim U\bar{U} + N_1\bar{N}_1 + N_2\bar{N}_2$$

also heavier colored objects $(m_U > m_{N_1, N_2})$

$$m_{\eta'} \sim \Lambda_{\rm TC}$$

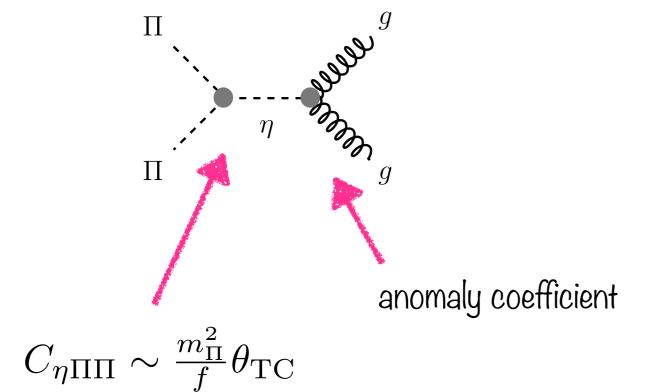
digamma resonance
$$m_{\eta_1}^2 \approx 4/5\,B_0 m_U\,\,(B_0 \sim \Lambda_{
m TC})$$
 $m_{\eta_2}^2 \lesssim m_\Pi^2 = B_0 (m_{N_1} + m_{N_2})$

$$\Gamma_{\eta_1 \to \Pi\Pi^*} \sim \theta_{\rm TC}^2 {\rm GeV} < 45 {\rm GeV}$$

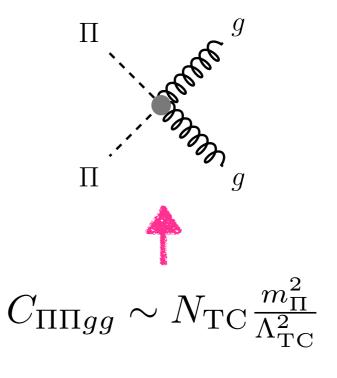
DM interactions

$$\mathcal{L}_{\text{DM}} = C_{\eta \Pi \Pi} \frac{\eta \Pi^2}{2} + \frac{g_3^2}{16\pi^2} C_{\Pi \Pi gg} \frac{\Pi^2}{f^2} G_{\mu\nu}^a G^{a,\mu\nu}$$

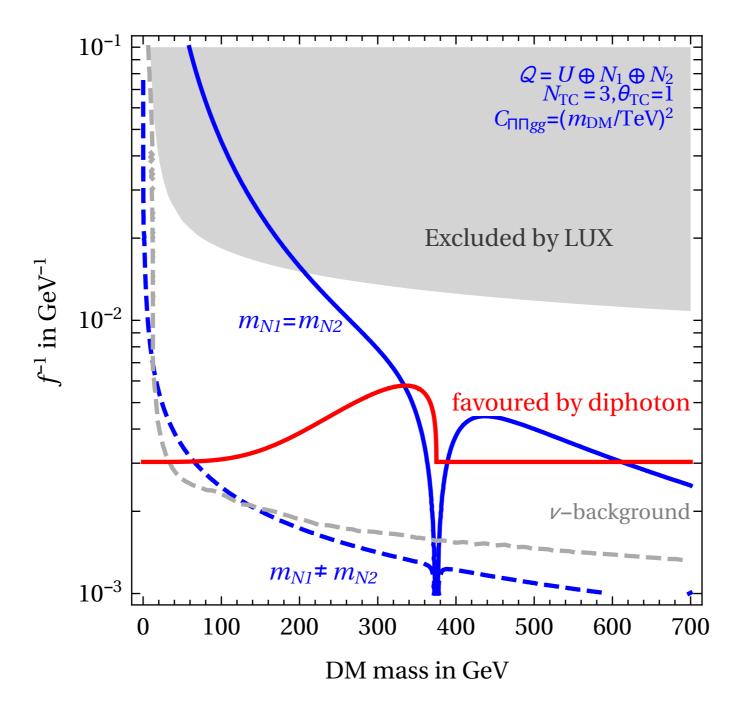
resonant scattering

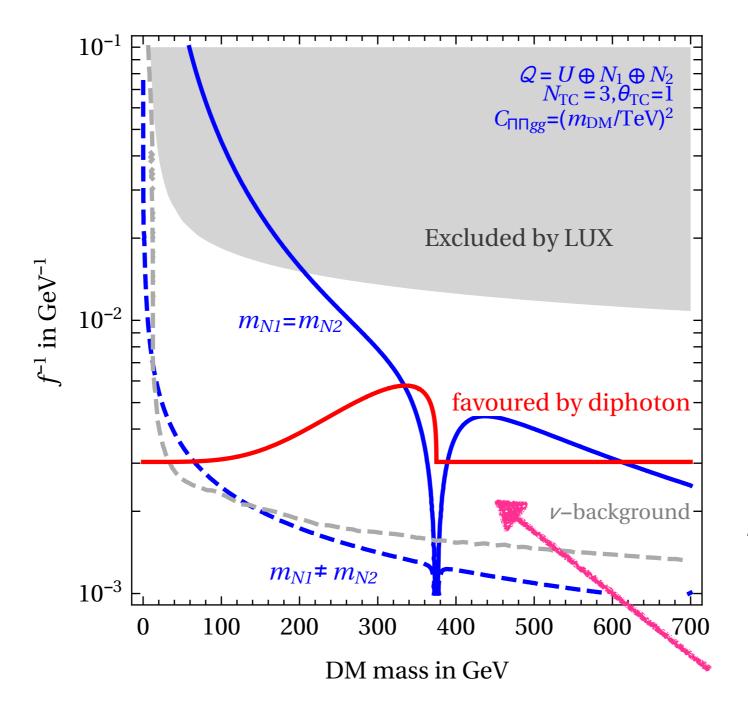


Rayleigh interactions



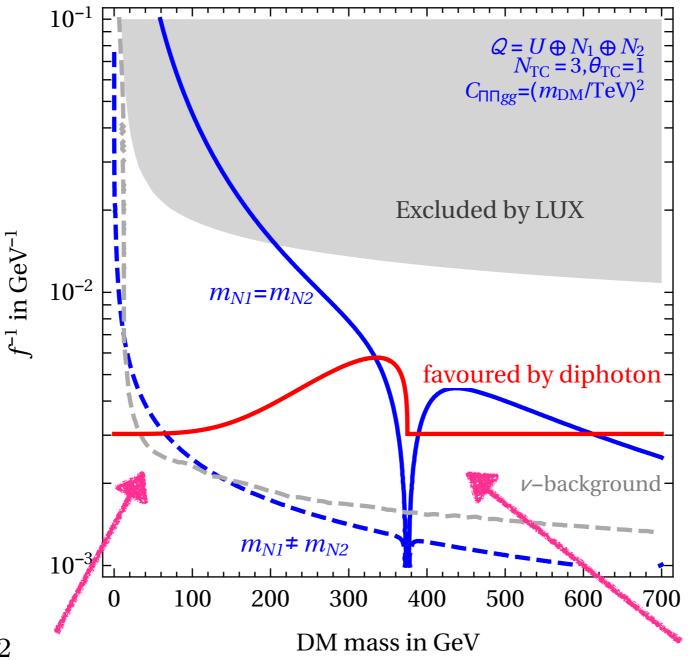
if allowed, also techni-pion/techni-pion scattering can be important





 $m_{\eta_2}=m_\Pi$ and $\eta_2,~\Pi,~\Pi^*$ becomes a triplet of DM candidates

 $\Pi\Pi^* o \eta_1 o gg$ dominates, relic abundance reproduced for $m_\Pi \approx 750\,{
m GeV}/2$

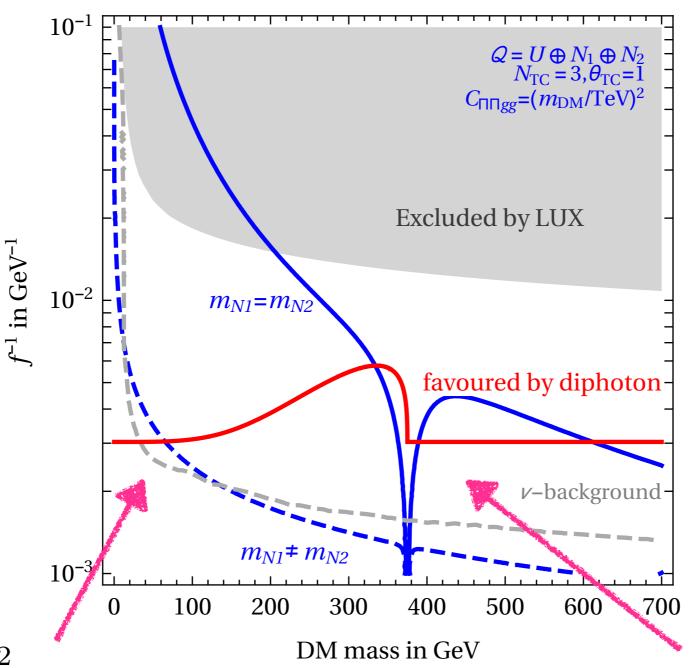


 $m_{\eta_2}=m_\Pi$ and $\eta_2,~\Pi,~\Pi^*$ becomes a triplet of DM candidates

 $\Pi\Pi^* \to \eta_2 \eta_2$

dominates, relic abundance reproduced for $m_\Pi < 100\,{
m GeV}$

$$\Pi\Pi^* o \eta_1 o gg$$
 dominates, relic abundance reproduced for $m_\Pi \approx 750\,{
m GeV}/2$



 $m_{\eta_2}=m_\Pi$ and $\eta_2,~\Pi,~\Pi^*$ becomes a triplet of DM candidates

ethoresonances,

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dominates, relic abundance reproduced for $m_\Pi < 100\,{
m GeV}$

 $\Pi\Pi^* o \eta_1 o gg$ dominates, relic abundance reproduced for $m_\Pi \approx 750\,{
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Conclusions

- Vectorlike Confinement as interesting scenario for BSM physics
- Dark Matter can be a composite techni-baryon
- Techni-pions as Dark Matter and digamma candidates: both digamma signal and relic abundance are reproduced, small width

If the digamma excess will survive, it would be the first BSM discovery at LHC!

VC would be one of the simplest and most predictive scenarios, with more data we would be able to get the right model!

Conclusions

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VC would be one of the simplest and most predictive be patient! with more data we would be able to get the right model.

Backnik

Techni-baryons as DM candidates

SU(5)	$SU(3)_c$	$\mathrm{SU}(2)_L$	$\mathrm{U}(1)_Y$	name	$\Delta b_3/N_{ m TC}$	$\Delta b_2/N_{\mathrm{TC}}$	$\Delta b_Y/N_{ m TC}$
1	1	1	0	N	0	0	0
$\bar{5}$	$\bar{3}$	1	1/3	D	1/3	0	2/9
	1	2	-1/2	L	0	1/3	1/3
10	$\bar{3}$	1	-2/3	U	1/3	0	8/9
	1	1	1	E	0	0	2/3
	3	2	1/6	Q	2/3	1	1/9
15	3	2	1/6	Q	2/3	1	1/9
	1	3	1	T	0	4/3	2
	6	1	-2/3	S	5/3	0	8/9
24	1	3	0	V	0	4/3	0
	8	1	0	G	2	0	0
	$\bar{3}$	2	5/6	X	2/3	1	25/9
	1	1	0	N	0	0	0

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Techni-baryons as DM candidates

 $\Delta b_2/N_{\mathrm{TC}}$

0

0

1/3

0

0

SU(5)	$SU(3)_c$	$\mathrm{SU}(2)_L$	$\mathrm{U}(1)_Y$	name	$\Delta b_3/N_{ m T}$
1	1	1	0	N	0
$\bar{5}$	$\bar{3}$	1	1/3	D	1/3
	1	2	-1/2	L	0
10	$\bar{3}$	1	-2/3	U	1/3
	1	1	1	E	0
	3		$\mathrm{SU}(N_{\mathrm{TC}})$) techni-	color Y
15	3			ni-quarks	
	1		N_{r}	$_{\Gamma \mathrm{F}}=3$	
	6			Q = V	
24	1			$= N \oplus L$	
	8			$_{\Gamma \mathrm{F}}=4$	
	$\bar{3}$		Q =	$V \oplus N$	~

$SU(N_{TC})$ techni-color Techni-quarks	Yukawa couplings	Allowed	Techni- pions	Techni-	under
$N_{\mathrm{TF}} = 3$	coupings	$N_{ m TC}$	8	baryons $8, \bar{6}, \dots$ for $N_{\text{TC}} = 3, 4, \dots$	$SU(3)_{TF}$
Q = V	0	3	3	VVV = 3	$SU(2)_L$
$\mathcal{Q} = N \oplus L$	1	3 - 14	unstable	$(N^{N_{\rm TC}})^* = 1$	=
$N_{\mathrm{TF}} = 4$	-	0 11	15	$\frac{\overline{20},20',\ldots}{20}$	$\mathrm{SU}(4)_{\mathrm{TF}}$
$Q = V \oplus N$	0	3	3×3	VVV, VNN = 3, VVN = 1	$SU(2)_L$
$Q = N \oplus L \oplus \tilde{E}$	2	3, 4, 5	unstable	$(N^{N_{\rm TC}})^* = 1$	=
$N_{\mathrm{TF}} = 5$, ,	24	$\frac{40,50}{}$	$\mathrm{SU}(5)_{\mathrm{TF}}$
$Q = V \oplus L$	1	3	unstable	VVV = 3	$\mathrm{SU}(2)_L$
$\mathcal{Q}=N\oplus L\oplus ilde{L}$	2	3	unstable	$NL ilde{L}=1$	=
=	2	4	unstable	$NNL\tilde{L}, L\tilde{L}L\tilde{L}=1$	=
$N_{\mathrm{TF}}=6$			35	$70, \overline{105'}$	$SU(6)_{TF}$
$Q = V \oplus L \oplus N$	2	3	unstable	VVV, VNN = 3, VVN = 1	$\mathrm{SU}(2)_L$
$\mathcal{Q} = V \oplus L \oplus \tilde{E}$	2	3	unstable	VVV = 3	=
$\mathcal{Q} = N \oplus L \oplus \tilde{L} \oplus \tilde{E}$	3	3	unstable	$NL ilde{L}, ilde{L} ilde{L} ilde{E}=1$	=
=	3	4	unstable	$NNL\tilde{L}, L\tilde{L}L\tilde{L}, N\tilde{E}\tilde{L}\tilde{L} = 1$	=
$N_{ m TF}=7$			48	112	$SU(7)_{TF}$
$Q = L \oplus \tilde{L} \oplus E \oplus \tilde{E} \oplus N$	4	3	unstable	$LLE, \tilde{L}\tilde{L}\tilde{E}, L\tilde{L}N, E\tilde{E}N = 1$	$\mathrm{SU}(2)_L$
$\mathcal{Q} = N \oplus L \oplus \tilde{E} \oplus V$	3	3	unstable	$VVV, VNN = 3, \ VVN = 1$	=
$N_{\mathrm{TF}} = 9$			80	240	$SU(9)_{TF}$
$\mathcal{Q} = Q \oplus ilde{D}$	1	3	unstable	$QQ\tilde{D}=1$	$\mathrm{SU}(2)_L$
$N_{\mathrm{TF}} = 12$			143	572	$SU(12)_{TF}$
$Q = Q \oplus \tilde{D} \oplus \tilde{U}$	2	3	unstable	$QQ\tilde{D}, \tilde{D}\tilde{D}\tilde{U} = 1$	$\mathrm{SU}(2)_L$

 $\Delta b_Y/N_{\mathrm{TC}}$

0

2/9

1/3

8/9

2/3

Backrib

Techni-baryons as DM candidates

 $\Delta b_3/N_{\mathrm{TC}}$

0

1/3

0

1/3

0

 $\mathrm{U}(1)_Y$

0

1/3

-1/2

-2/3

name

N

D

L

U

 \boldsymbol{E}

 $\Delta b_2/N_{\mathrm{TC}}$

0

0

1/3

0

0

 $\Delta b_Y/N_{\mathrm{TC}}$

0

2/9

1/3

8/9

2/3

SU(5)	$SU(3)_c$	SU(2)
1	1	1
$\bar{5}$	$\bar{3}$	1
	1	2
10	$\bar{3}$	1
	1	1
	3	
15	3	
	1	
	6	
24	1	
	$\frac{8}{3}$	
	$\bar{3}$	
	1	

 $M_{
m DM} \sim 100\,{
m TeV}$ if DM as a thermal relic

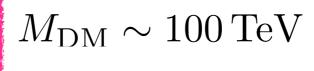
1 2			-/ -	1	
$SU(N_{TC})$ techni-color	Yukawa	Allowed	Techni-	Techni-	
Techni-quarks	couplings	$N_{ m TC}$	pions	baryons	under
$N_{\mathrm{TF}}=3$			8	$8, \bar{6}, \dots \text{ for } N_{\text{TC}} = 3, 4, \dots$	$\mathrm{SU}(3)_{\mathrm{TF}}$
$\mathcal{Q} = V$	0	3	3	VVV = 3	$\mathrm{SU}(2)_L$
$\mathcal{Q} = N \oplus L$	1	3 - 14	unstable	$(N^{N_{\rm TC}})^* = 1$	=
$N_{ m TF}=4$			15	$\overline{20}, 20', \dots$	$SU(4)_{TF}$
$\mathcal{Q} = V \oplus N$	0	3	3×3	$VVV, VNN = 3, \ VVN = 1$	$\mathrm{SU}(2)_L$
$\mathcal{Q}=N\oplus L\oplus ilde{E}$	2	3, 4, 5	unstable	$(N^{N_{ m TC}})^* = 1$	=
$N_{ m TF}=5$			24	$\overline{40}, \overline{50}$	$\mathrm{SU}(5)_{\mathrm{TF}}$
$\mathcal{Q} = V \oplus L$	1	3	unstable	VVV = 3	$\mathrm{SU}(2)_L$
$\mathcal{Q}=N\oplus L\oplus ilde{L}$	2	3	unstable	$NL ilde{L}=1$	=
=	2	4	unstable	$NNL\tilde{L}, L\tilde{L}L\tilde{L}=1$	=
$N_{\mathrm{TF}}=6$			35	$70, \overline{105'}$	$SU(6)_{TF}$
$Q = V \oplus L \oplus N$	2	3	unstable	VVV, VNN = 3, VVN = 1	$\mathrm{SU}(2)_L$
$\mathcal{Q} = V \oplus L \oplus \tilde{E}$	2	3	unstable	VVV = 3	=
$\mathcal{Q} = N \oplus L \oplus \tilde{L} \oplus \tilde{E}$	3	3	unstable	$NL ilde{L}, ilde{L} ilde{L} ilde{E}=1$	=
=	3	4	unstable	$NNL\tilde{L}, L\tilde{L}L\tilde{L}, N\tilde{E}\tilde{L}\tilde{L} = 1$	=
$N_{ m TF}=7$			48	112	$SU(7)_{TF}$
$\mathcal{Q} = L \oplus \tilde{L} \oplus E \oplus \tilde{E} \oplus N$	4	3	unstable	$LLE, \tilde{L}\tilde{L}\tilde{E}, L\tilde{L}N, E\tilde{E}N = 1$	$SU(2)_L$
$Q = N \oplus L \oplus \tilde{E} \oplus V$	3	3	unstable	$VVV, VNN = 3, \ VVN = 1$	=
$N_{\mathrm{TF}} = 9$			80	240	$SU(9)_{TF}$
$\mathcal{Q} = Q \oplus ilde{D}$	1	3	unstable	$QQ\tilde{D}=1$	$\mathrm{SU}(2)_L$
$N_{\mathrm{TF}}=12$			143	572	$SU(12)_{TF}$
$Q = Q \oplus \tilde{D} \oplus \tilde{U}$	2	3	unstable	$QQ\tilde{D}, \tilde{D}\tilde{D}\tilde{U} = 1$	$\mathrm{SU}(2)_L$

CKUP

DM phenomenology

DM as a thermal relic: non relativistic annihilation of techni-baryons to techni-pions via techni-strong interactions

$$\sigma_{p\bar{p}}^{\rm ann}\,v\sim 100/m_p^2$$



too high for minimal direct detection interactions! but fermionic DM, even if a SM singlet, can have typical interactions with photons



Magnetic and Electric Dipole Moments

$$\frac{1}{2}\bar{\Psi}\gamma_{\mu\nu}(\mu_M + id_E\gamma_5)\Psi F^{\mu\nu}$$

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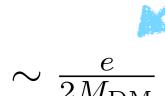
Magnetic and Electric Dipole Moments
$$\label{eq:polesym} \frac{1}{2}\bar{\Psi}\gamma_{\mu\nu}(\mu_M + id_E\gamma_5)\Psi F^{\mu\nu}$$

$$\mathcal{L} = \mathcal{L}_{SM} - \frac{1}{4g_{TC}^2} \mathcal{G}_{\mu\nu}^A \mathcal{G}_A^{\mu\nu} + \frac{\theta_{TC}}{32\pi^2} \mathcal{G}_{\mu\nu}^A \tilde{\mathcal{G}}_A^{\mu\nu} + \bar{\mathcal{Q}}_i (i\not\!\!D - m_i) \mathcal{Q}_i + \left[H\bar{\mathcal{Q}}_i (y_{ij}^L P_L + y_{ij}^R P_R) \mathcal{Q}_j + \text{h.c.} \right]$$

CKAL

DM phenomenology

$$\frac{1}{2}\bar{\Psi}\gamma_{\mu\nu}(\mu_M + id_E\gamma_5)\Psi F^{\mu\nu}$$

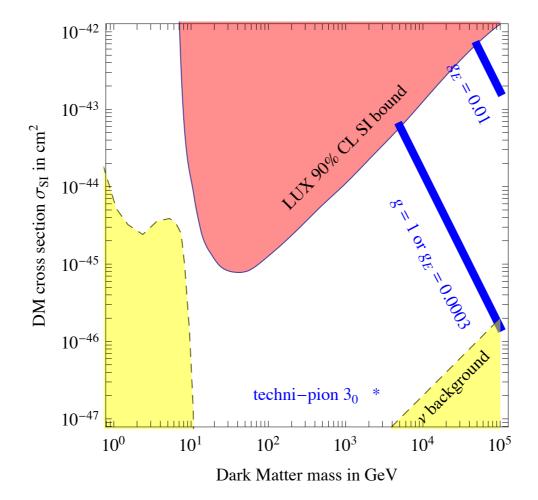


$$\sim heta_{\mathrm{TC}} \, rac{e}{2M_{\mathrm{DM}}} \, rac{\min[m_{\mathcal{Q}}]}{M_{\mathrm{DM}}}$$

direct detection:

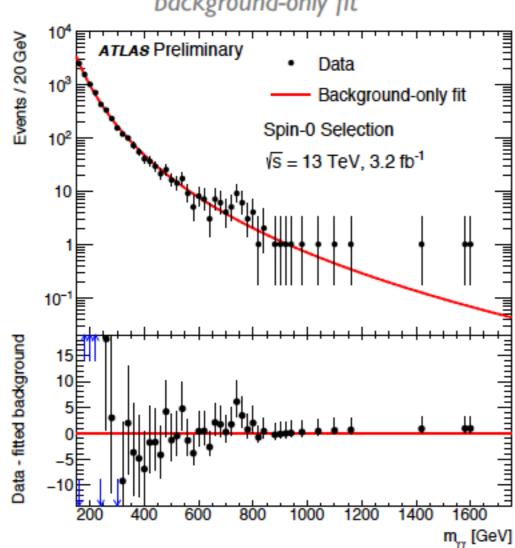
$$\frac{d\sigma}{dE_R} \approx \frac{e^2 Z^2}{4\pi E_R v^2} \left(\mu_M^2 v^2 + d_E^2 \right)$$

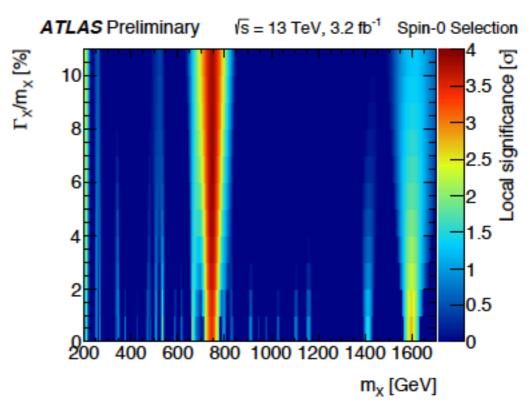
Dirac techni-baryon DM



SPIN-0 ANALYSIS

background-only fit

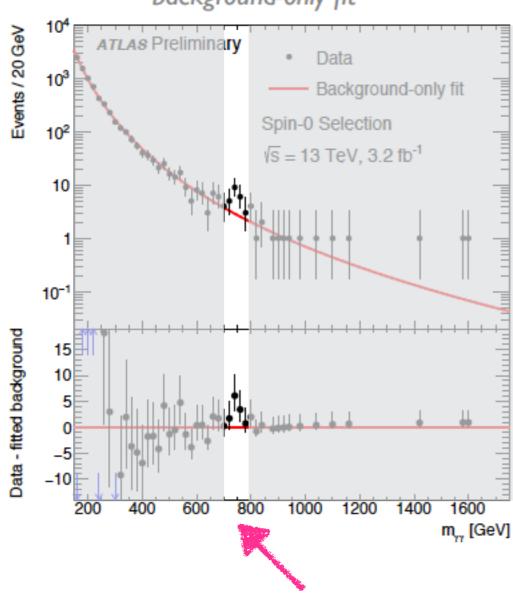




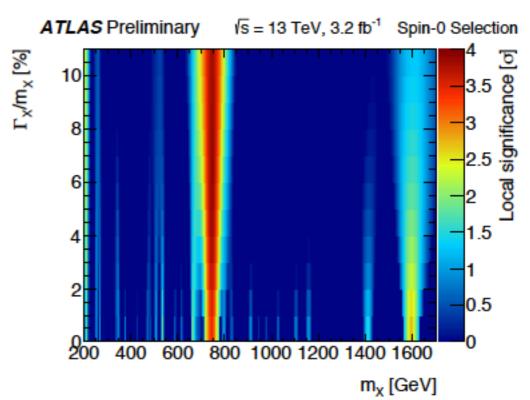
- Largest deviation from B-only hypothesis
 - \checkmark m_X ~ 750 GeV, Γ_X ~ 45 GeV (6%)
 - ✓ Local Z = 3.9 σ
 - $\sqrt{\text{Global Z}} = 2.0 \sigma$

SPIN-0 ANALYSIS

background-only fit



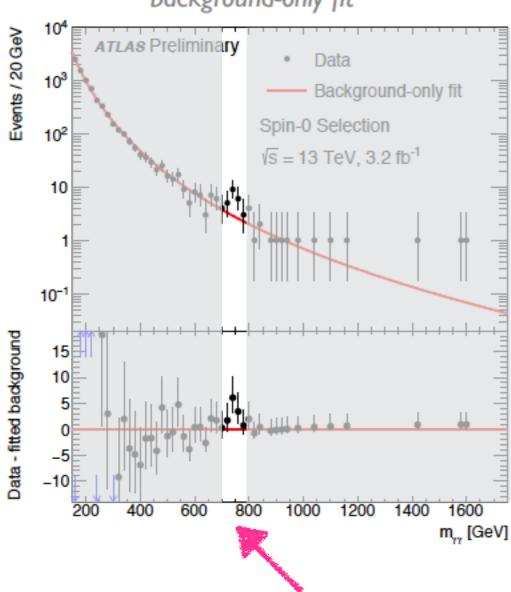
 $750\,\mathrm{GeV}$



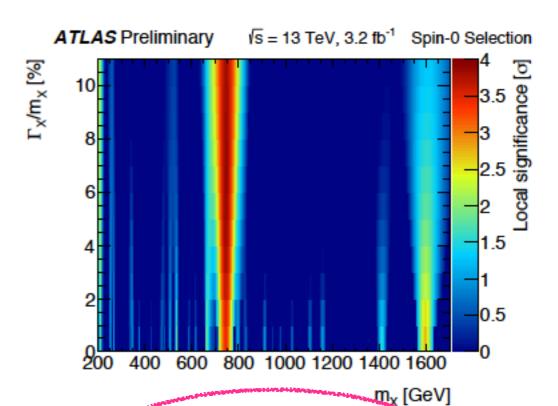
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background-only fit



 $750\,\mathrm{GeV}$



Largest deviation from B-only hypothesis

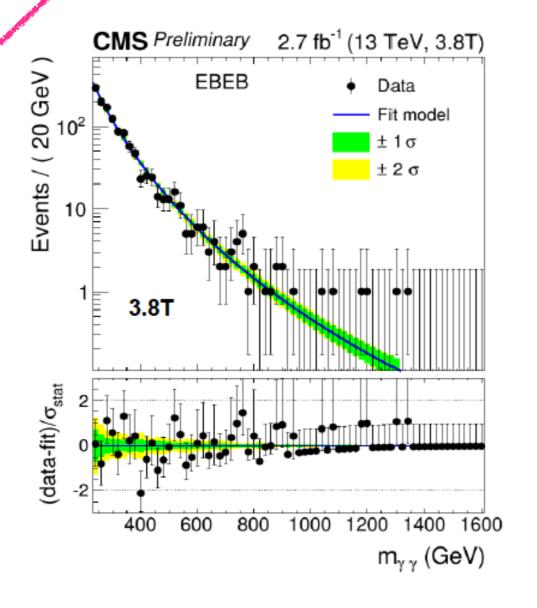
$$\checkmark$$
 m_X ~ 750 GeV, Γ_X ~ 45 GeV (6%)

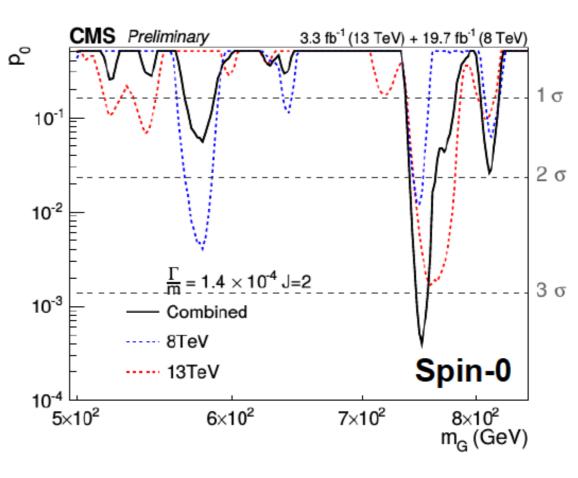
✓ Local Z = 3.9
$$\sigma$$

✓ Global
$$Z = 2.0 \sigma$$

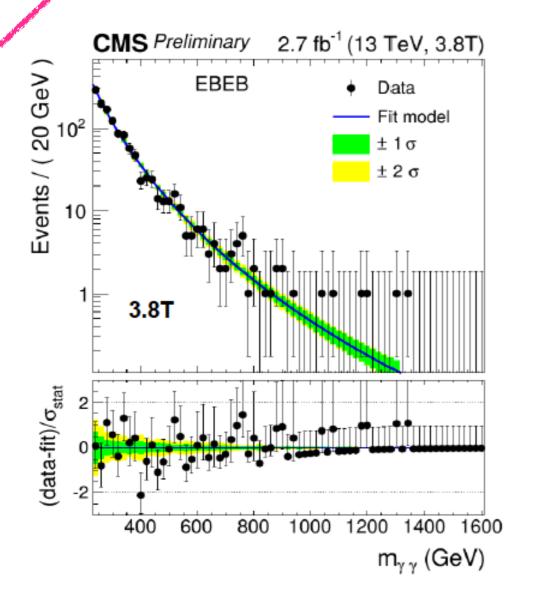
Backup

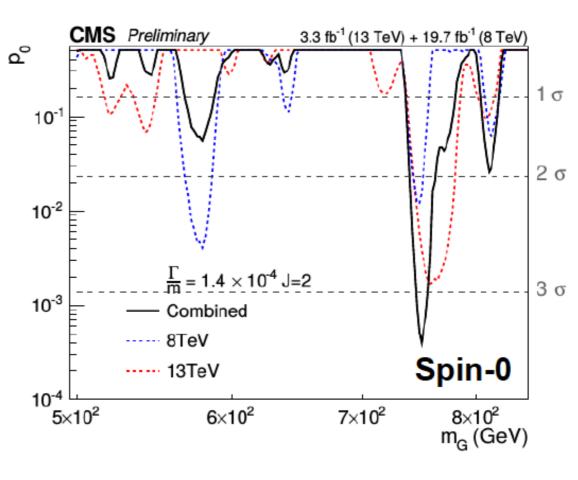
Di-photon excess





- ightharpoonup Largest excess observed at $m_x = 750 \text{GeV}$ and for narrow width.
 - Local significance: 3.4σ
 - Taking into account mass range 500-3500GeV (and all signal hypotheses), "global" significance becomes 1.6σ



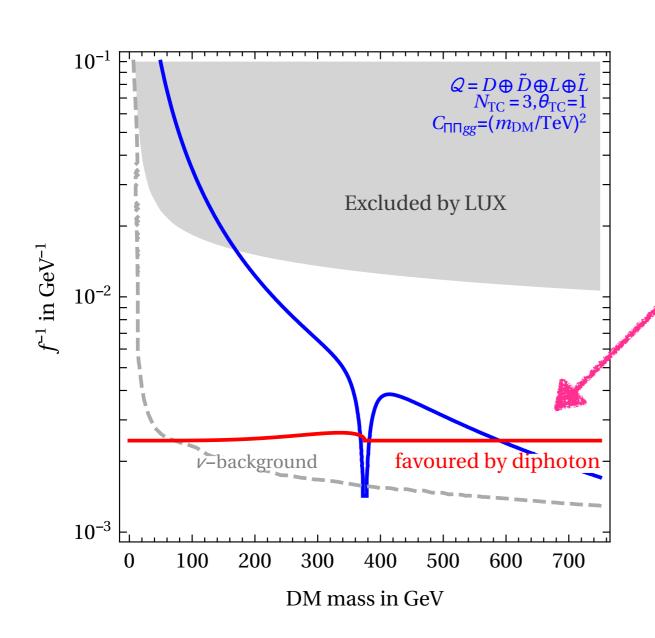


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 - Local significance: 3.4σ
 - Taking into account mass range 500-3500 GeV (and all signal hypotheses), "global" significance becomes 1.6σ

$$Q = D \oplus \tilde{D} \oplus L \oplus \tilde{L} = (\bar{3}, 1)_{1/3} \oplus (3, 1)_{-1/3} \oplus (1, 2)_{-1/2} \oplus (1, 2)_{1/2}$$

$$m_D = m_{\tilde{D}}, \ m_L = m_{\tilde{L}}$$

DM from a generalized G-parity $D\leftrightarrow \tilde{D} \quad L\leftrightarrow \tilde{L}$



 $1_A^-1_A^- \to \eta \to gg$ dominates, relic abundance reproduced near the resonant region

narrow width, look for extra resonances

Backnik