



# Constrained superfields in SUSY and SUGRA

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## Why constrained superfields?

- Non-linear realization of broken supersymmetry.
- Consistency with superspace methods.
- Offer a "hint" to the UV theory.

## *The goldstino sector*

# Nilpotent goldstino superfield

*Rocek '78, Lindstrom, Rocek '79, Casalbuoni, De Curtis, Dominici, Feruglio, Gatto '89, Komargodski, Seiberg '09*

- ▶ We can break SUSY with a chiral superfield

$$X = A + \sqrt{2} \Theta G + \Theta^2 F.$$

- ▶ SUSY broken:  $\delta G_\alpha = -f \epsilon_\alpha + \dots$
- ▶ In the formal limit  $m_A \rightarrow \infty$  the scalar decouples.
- ▶ This is described by **imposing the superspace constraint**

$$X^2 = 0 \rightarrow A = \frac{G^2}{2F}.$$

# Polonyi model and constrained superfields

- ▶ The same setup works in supergravity

$$\mathcal{L} = \int d^2\Theta 2\mathcal{E} \left[ \frac{3}{8}(\overline{\mathcal{D}}^2 - 8\mathcal{R})e^{-K/3} + W \right] + c.c.$$

with  $K = X\overline{X}$  and  $W = fX + W_0$ .

- ▶ In the  $G = 0$  gauge the component form is

$$e^{-1}\mathcal{L} = -\frac{1}{2}R + \frac{1}{2}\epsilon^{klmn}(\overline{\psi}_k\overline{\sigma}_l\mathcal{D}_m\psi_n - \psi_k\sigma_l\mathcal{D}_m\psi_n) \\ - \overline{W}_0\overline{\psi}_a\overline{\sigma}^{ab}\overline{\psi}_b - W_0\psi_a\sigma^{ab}\psi_b - |f|^2 + 3|W_0|^2.$$

*Dudas, Ferrara, Kehagias, Sagnotti '15, Bergshoeff, Freedman, Kallosh, Van Proeyen '15, Hasegawa and Yamada '15, Antoniadis, Markou '15, Dall'Agata, FF '15*

*Removing matter component fields*

# Constrained chiral superfields

## Remove complex scalar

- ▶ For a chiral superfield  $Y$ , remove the scalar component  $y$  by imposing

$$X Y = 0 \rightarrow y = \frac{G \chi^y}{F} - \frac{G^2}{2F^2} F^y.$$

*Brignole, Feruglio, Zwirner '97, Komargodski, Seiberg '09*

- ▶ Naturally describes light fermions. *Cribiori, Dall'Agata, FF '16*

## Keep only real scalar

- ▶ For a chiral superfield  $\mathcal{A}$

$$X \mathcal{A} = X \bar{\mathcal{A}} \rightarrow \mathcal{A}| = \phi + \text{fermions}$$

*Komargodski, Seiberg '09*

- ▶ Naturally describes inflaton.

# Removing any selected component field

*Dall'Agata, Dudas, FF '16*

- ▶ For a generic superfield

$$Q = q + \theta \chi^q + \dots$$

- ▶ We propose the constraint

$$X \bar{X} Q = 0.$$

- ▶ This removes only the lowest component

$$q = \frac{G \chi^q}{\sqrt{2F}} + \dots$$

- ✓ Use to eliminate **more** components ( $|X|^2 DQ = 0$  for  $\chi^q$ ).
- ✓ Reproduces all known constraints.



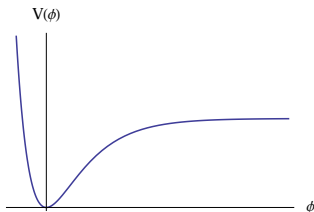
## *Applications to supergravity cosmology*

# Inflation in supergravity

- ▶ Inflation can be implemented by a **single** scalar field

$$e^{-1}\mathcal{L} = -\frac{1}{2}R - \frac{1}{2}\partial\phi\partial\phi - V(\phi)$$

with rather flat potential.



## Usual issues

- ▶ Strongly stabilize other scalars.
- ▶ Build the potential.

# SUGRA models with $X$ and $\mathcal{A}$

*Kahn, Roberts, Thaler '15, Ferrara, Kallosh, Thaler '15, Carrasco, Kallosh, Linde '15, Dall'Agata, FF '15*

- ▶ A simple example is

$$K = X\bar{X} - \frac{1}{4}(\mathcal{A} - \bar{\mathcal{A}})^2, \quad W = g(\mathcal{A}) + X f(\mathcal{A})$$

with  $\overline{f(z)} = f(\bar{z})$  and  $\overline{g(z)} = g(\bar{z})$ .

- ▶ In the  $G = 0$  gauge, the **full Lagrangian** is

$$e^{-1} \mathcal{L} = -\frac{1}{2} R + \epsilon^{klmn} \bar{\psi}_k \bar{\sigma}_l \mathcal{D}_m \psi_n - g(\phi) \left( \bar{\psi}_a \bar{\sigma}^{ab} \bar{\psi}_b + \psi_a \sigma^{ab} \psi_b \right) \\ - \frac{1}{2} \partial\phi\partial\phi - \left( f(\phi)^2 - 3g(\phi)^2 \right).$$

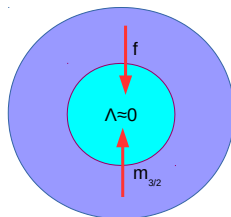
- ▶ Model building becomes extremely easy, but there is almost no predictive power.
- ▶ Supersymmetry must be broken in the vacuum.

# Bousso-Polchinski mechanism in SUGRA

*FF, Kehagias, Racco, Riotto '16*

- ▶ Nilpotent three-form superfields  $\rightarrow$  scanning of SUSY breaking scale.
- ▶ Three-form supergravity  $\rightarrow$  scanning of gravitino mass!
- ▶ Small cosmological constant and large gravitino mass

$$\Lambda = \sum_i (f^i)^2 - 3(m_{3/2})^2.$$



## *Summary - Outlook*

- ▶ We have discussed the constrained superfields formalism both in SUSY and SUGRA.
- ▶ We have presented applications in supergravity cosmology.
- ▶ **More applications?**

Thank you!