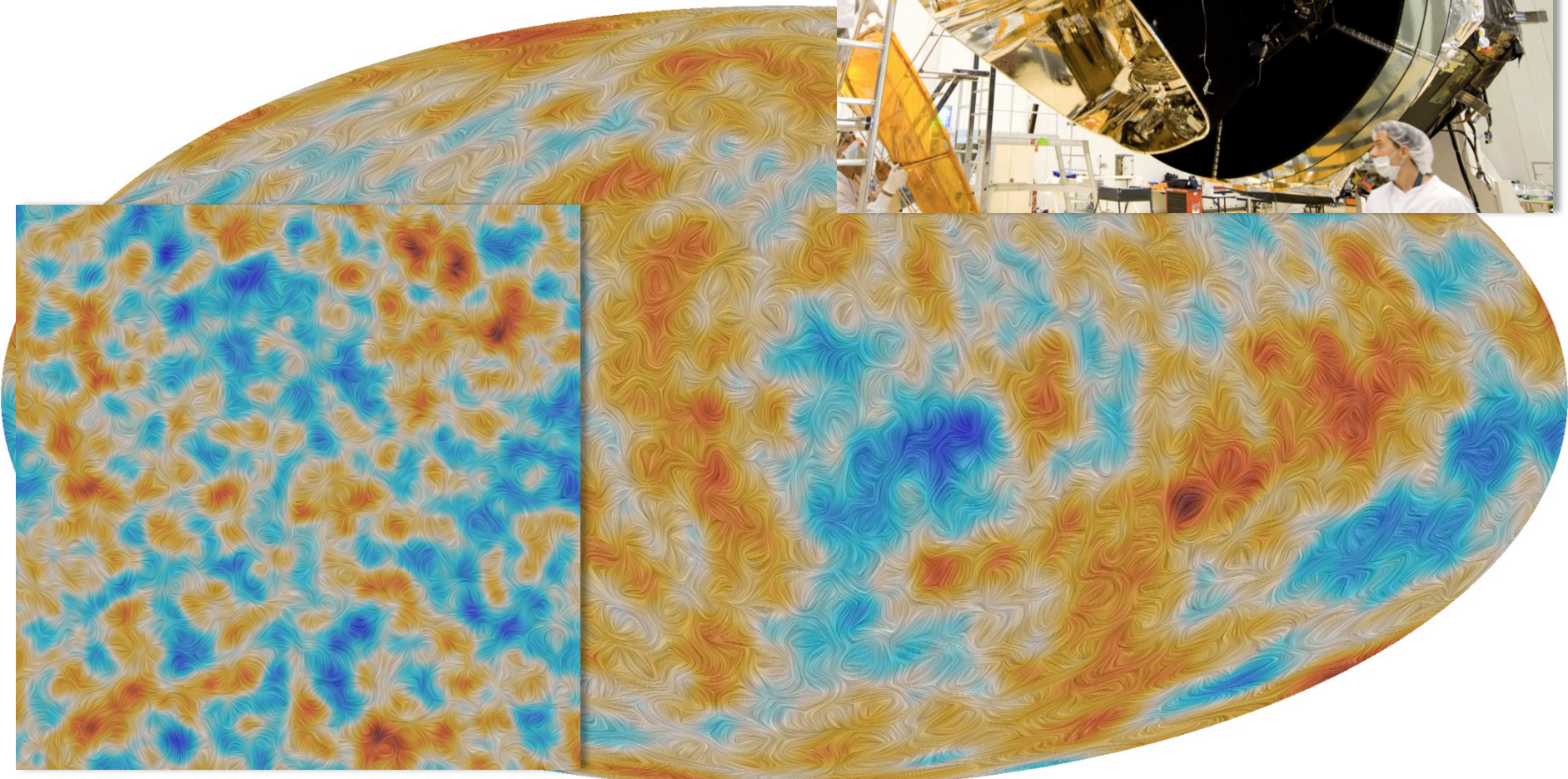


Investigating the *Dark* Universe with the Euclid satellite

Luca Valenziano
INAF/IASF-Bologna
& INFN – Bologna

On behalf of the EC and Euclid Italia
INFN Bologna - 2015

Planck & CMB



Beyond Planck

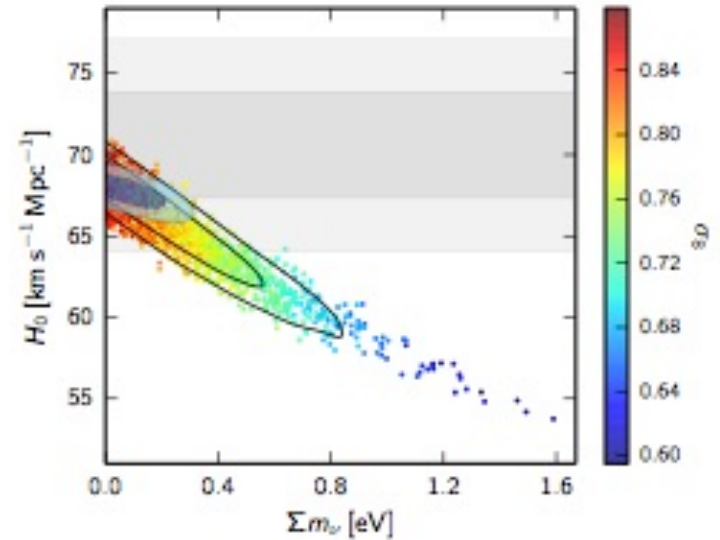
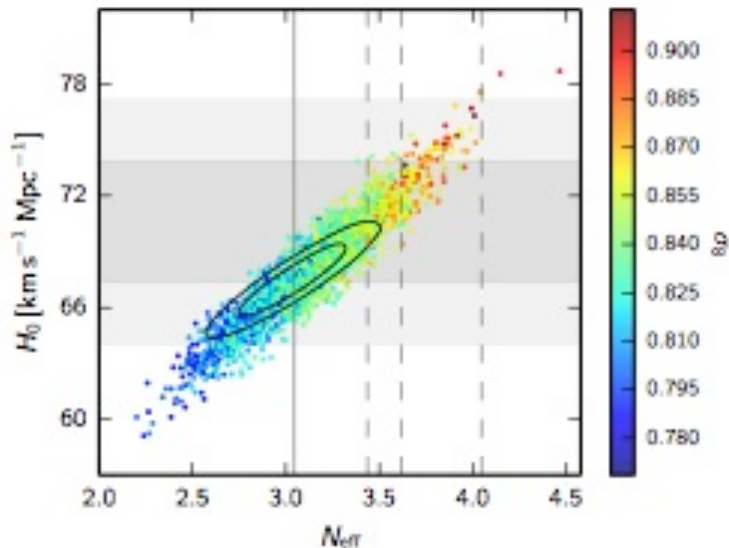
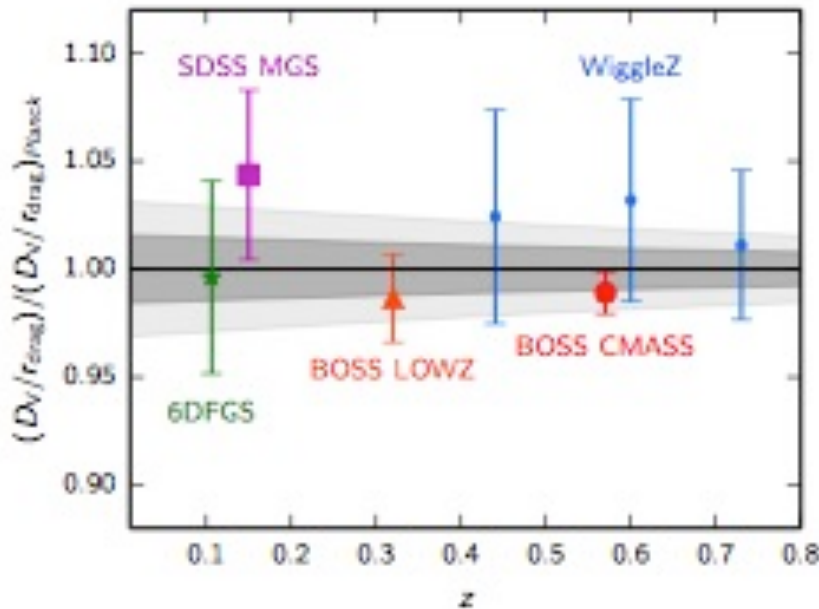
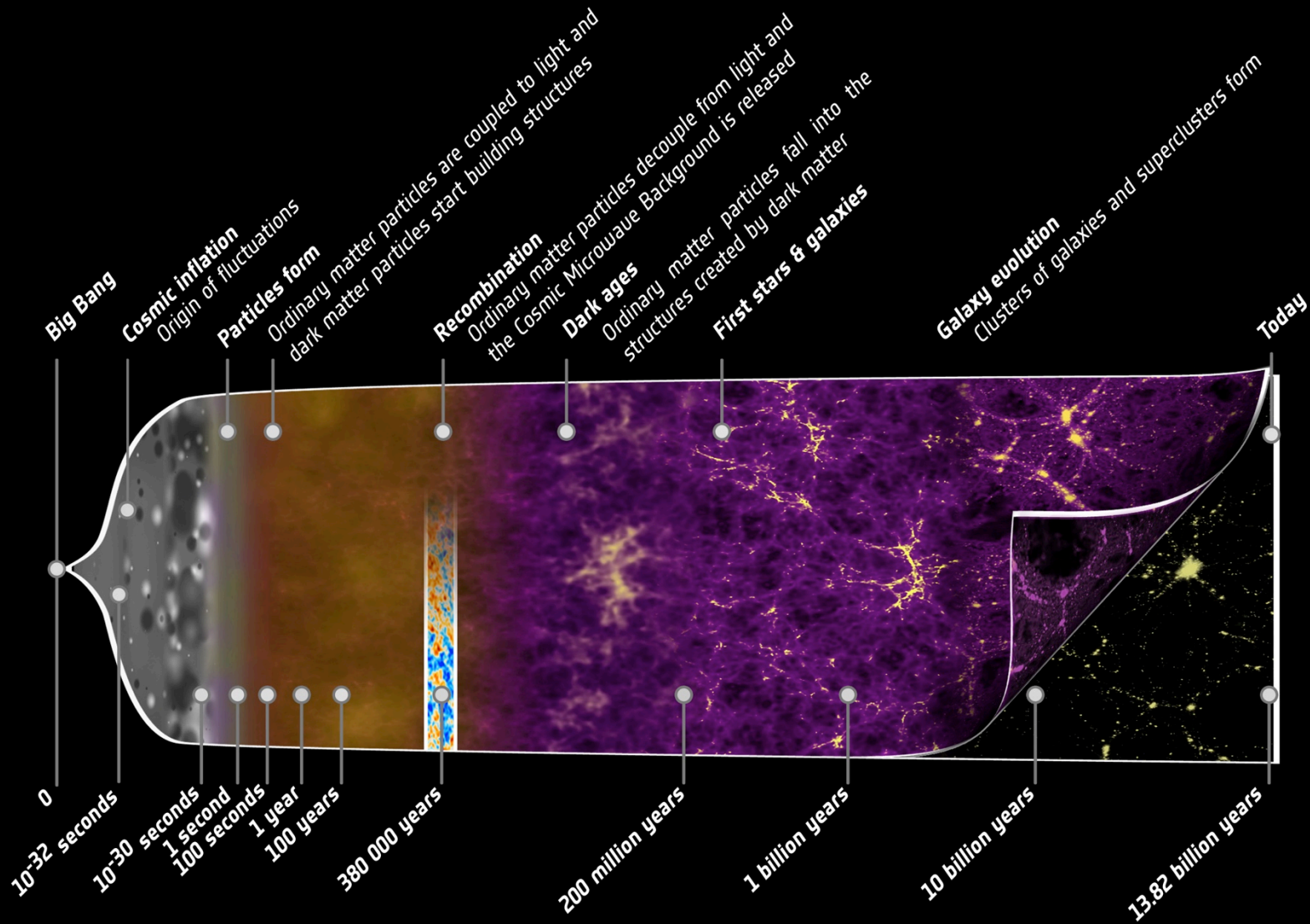


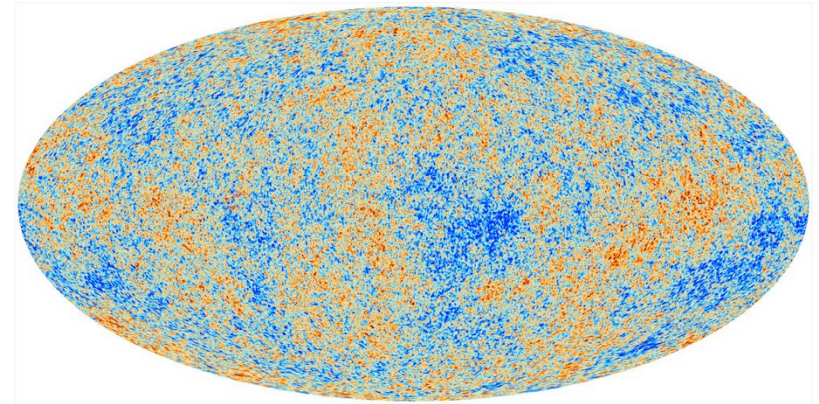
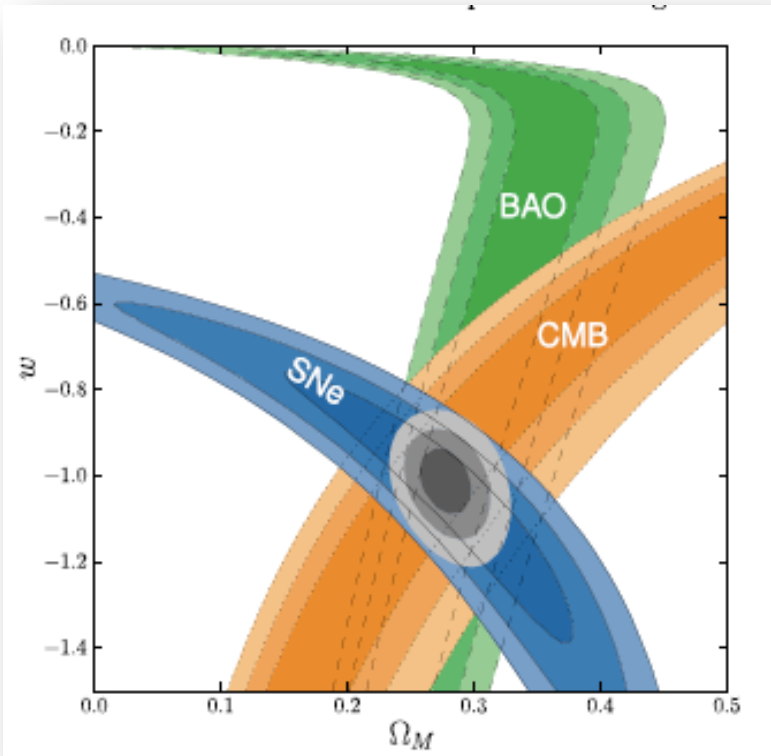
Fig. 29. Samples from the *Planck* TT+lowP posterior in the Σm_ν - H_0 plane, colour-coded by σ_8 . Higher Σm_ν damps the matter fluctuation amplitude σ_8 , but also decreases H_0 (grey bands show the direct measurement $H_0 = (70.6 \pm 3.3)$ km s⁻¹ Mpc⁻¹, Eq. 30). Solid black contours show the constraint from *Planck* TT+lowP+lensing (which mildly prefers larger masses), and filled contours show the constraints from *Planck* TT+lowP+lensing+BAO.



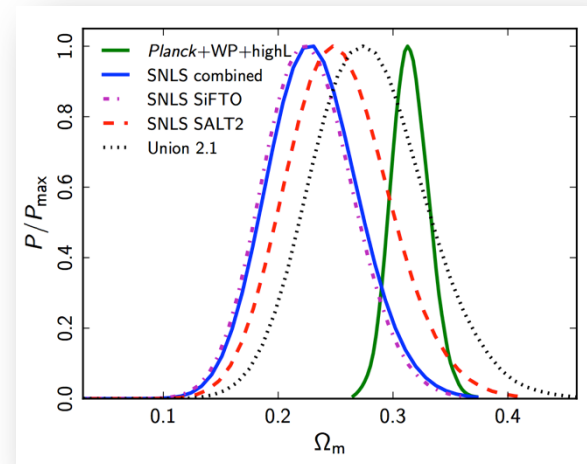
Euclid cosmological motivation



An accelerating Λ -dominated Universe: "concordance", but with a few key open questions...



Planck Collaboration 2013, paper XVI



Amanullah et al. 2010 (Union supernovae)

Q1: Is cosmic acceleration produced by a cosmological constant or by an evolving scalar field?

Evolving equation of state of DE:

e.g. $w(a) = w_0 + w_a(1 - a)$

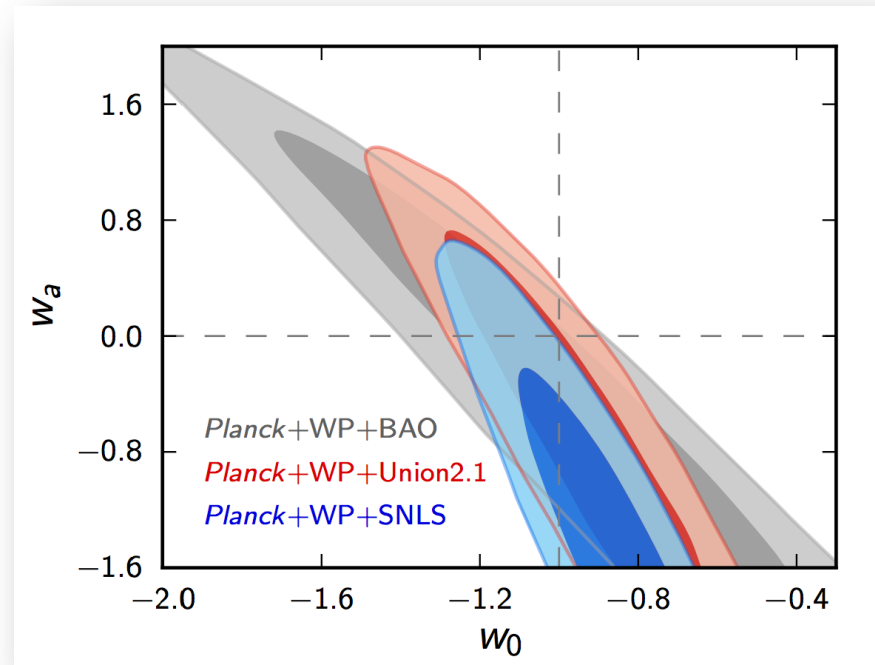
DETF (Albrecht et al. 2006): characterize experiments through a Figure of Merit in (w_0, w_a) plane (or similar):

$$\text{FoM} = 1/(\Delta w_0 \times \Delta w_a)$$

But this reflects chosen parameterization

→ FoMs should be taken with a big grain of salt (e.g. NASA/DOE/ESA FoMSWG report, Albrecht et al. 2009): there is much more science in a galaxy survey

Parameterizing our ignorance



Planck Collaboration 2013, XVI

Q2: Does General Relativity still work on cosmological scales?

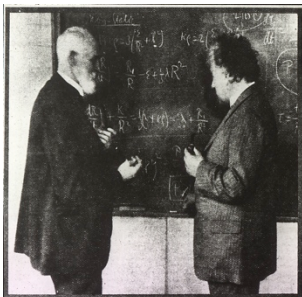
“Dark Gravity” instead of Dark Energy...?

A story with two sides...

$$R_{\mu\nu} - \frac{1}{2} g_{\mu\nu} R = -\frac{8\pi G}{c^2} T_{\mu\nu} + \Lambda g_{\mu\nu}$$

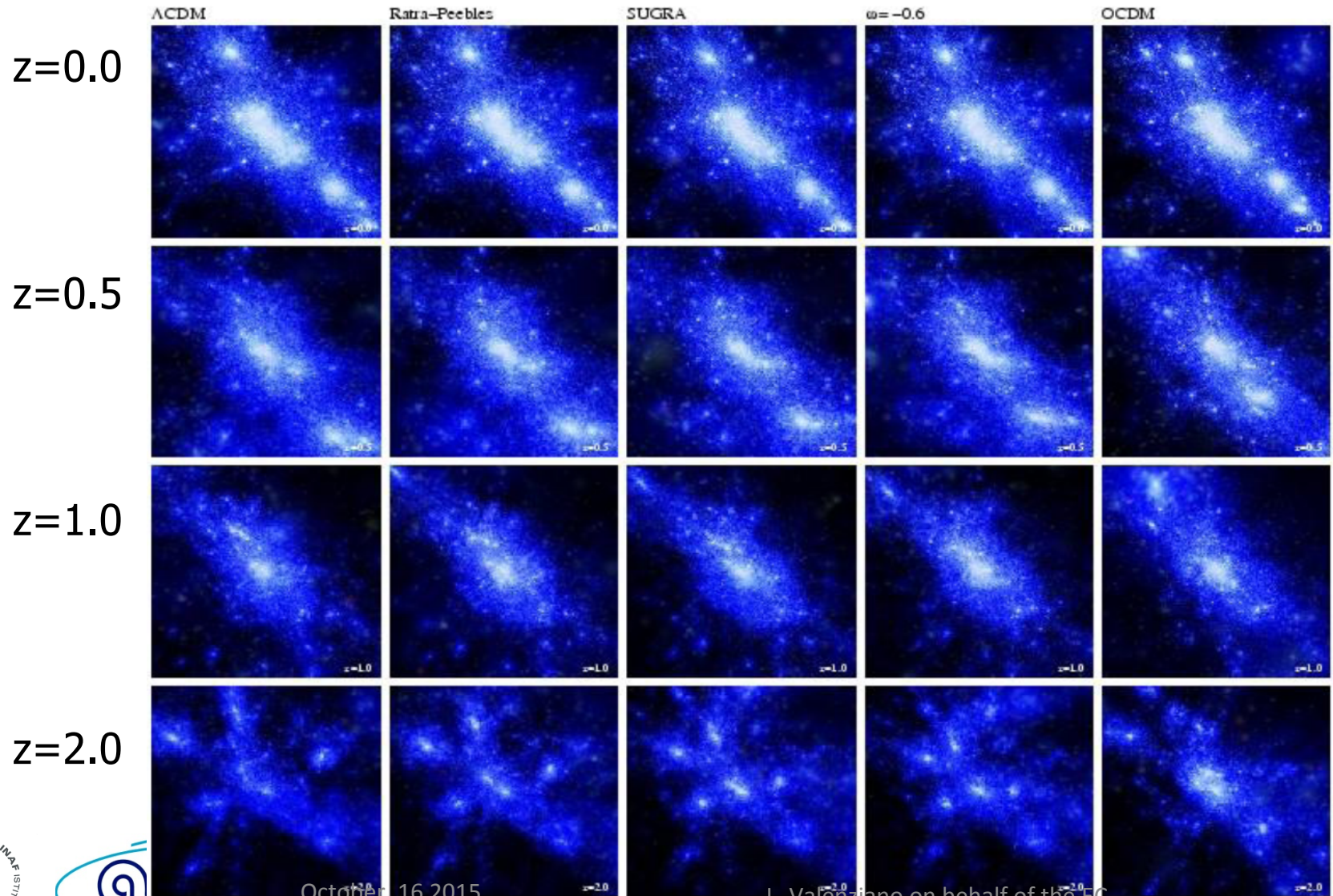
Modify gravity theory [e.g. $R \rightarrow f(R)$]

Add dark energy



→ Distinguish by measuring both background expansion $H(z)$ and growth rate of structure $f(z)$

Tiny differences between DE/MG models

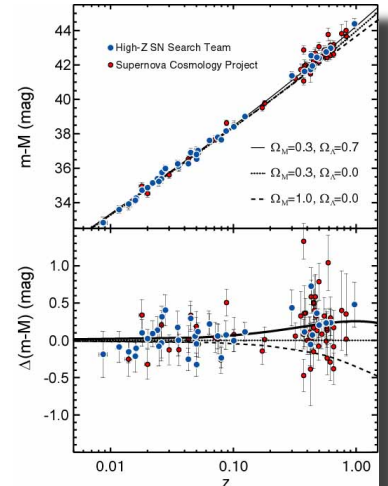


Euclid: answer both questions

1. Measure *the expansion history $H(z)$* to high accuracy, as to detect percent variations of DE *equation of state $w(z)$* with robust control of systematics:

Achieve this through **two probes**:

- A. Using the scale of Baryonic Acoustic Oscillations (BAO) in the clustering pattern of galaxies as a standard rod
- B. Using shape distortions induced by Weak Gravitational Lensing



Euclid: answer both questions

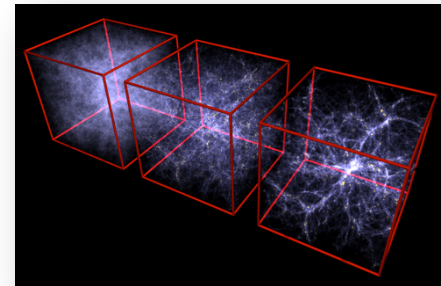
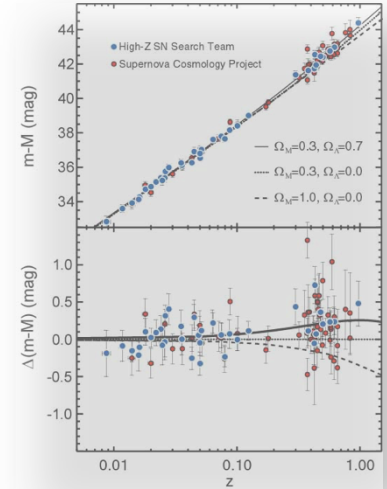


1. Measure the *growth rate of structure* from the same probes, to detect modifications of gravity:
2. Measure at the same time *the growth rate of structure* from the same probes, to detect modifications of gravity:

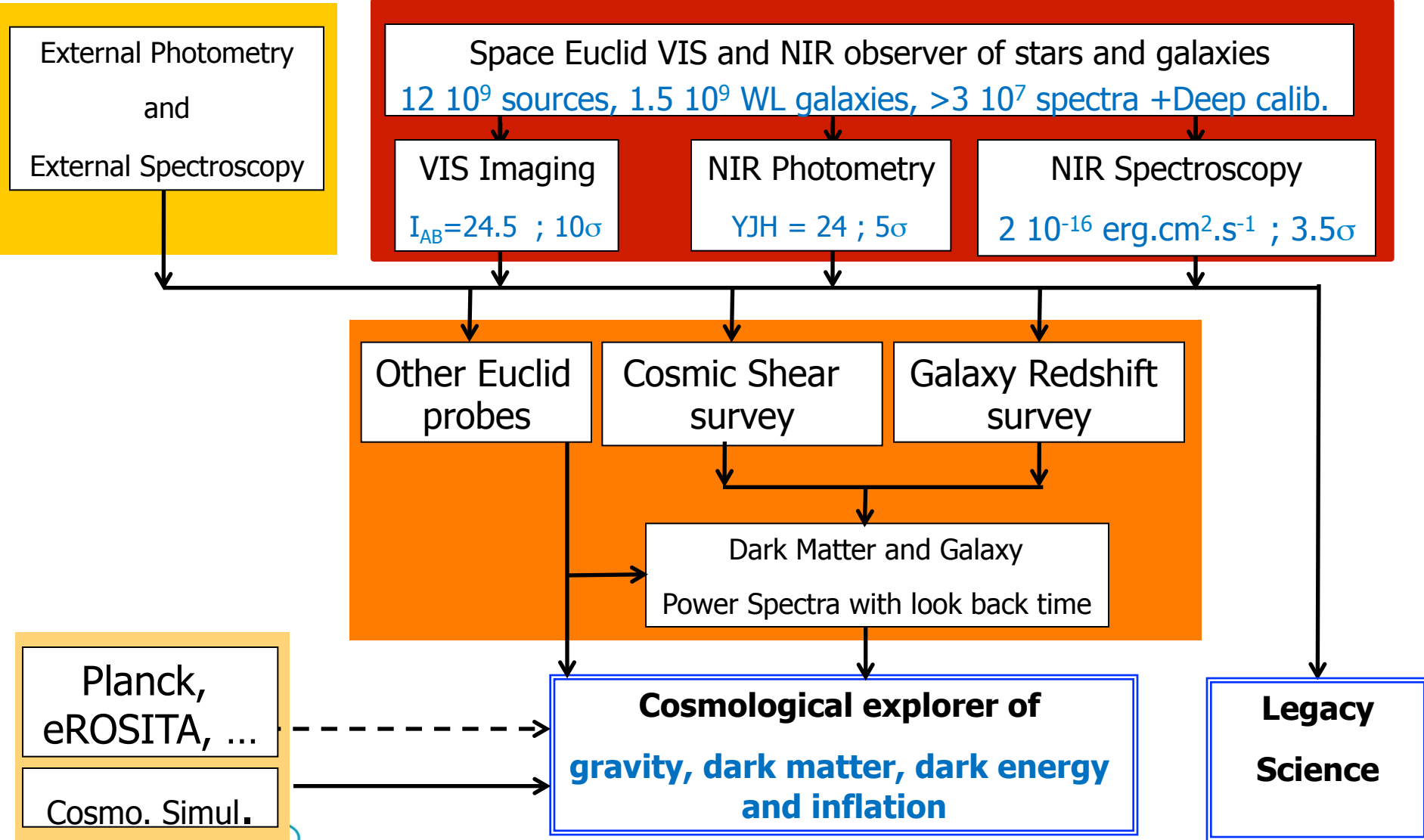
A. Clustering redshift-space distortions (RSD)

B. Weak Lensing (WL) Tomography

•→ These two probes are differently sensitive to the Ψ and Φ potentials of the perturbed metric, i.e. to deformations of time and space



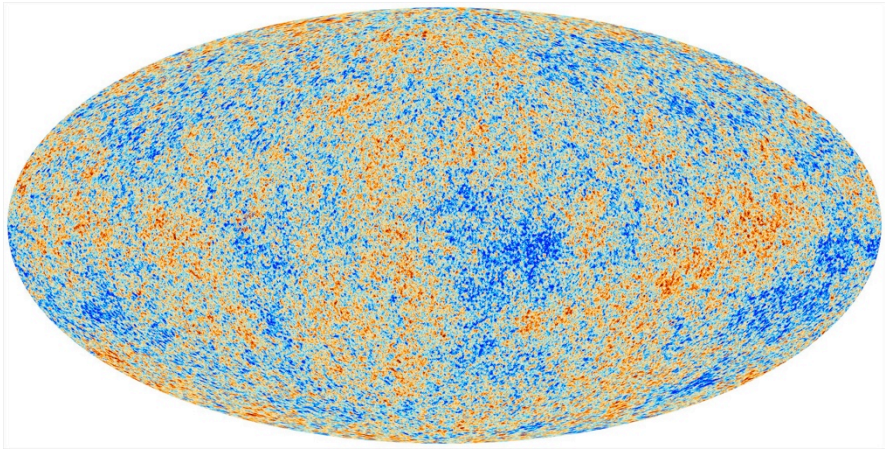
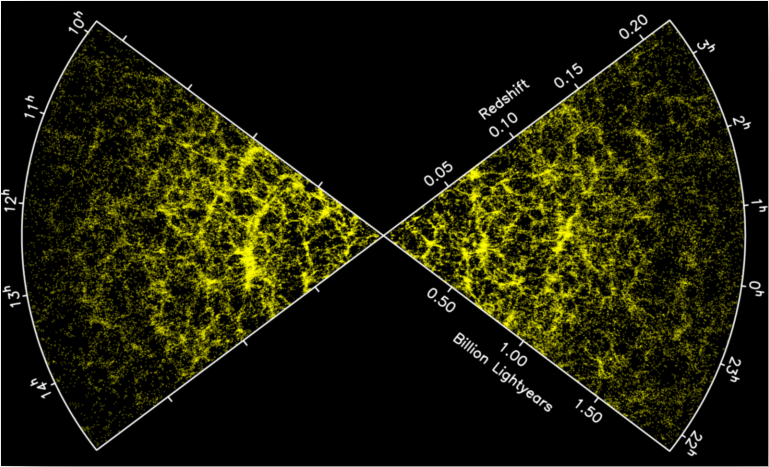
The Euclid Wide Survey Machine



Baryonic Acoustic Oscillations

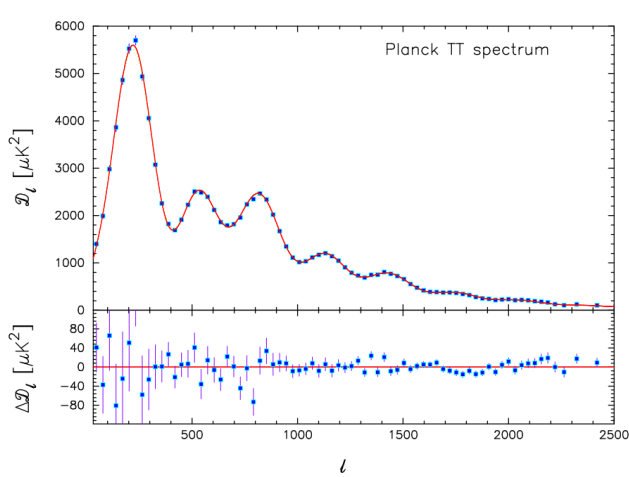
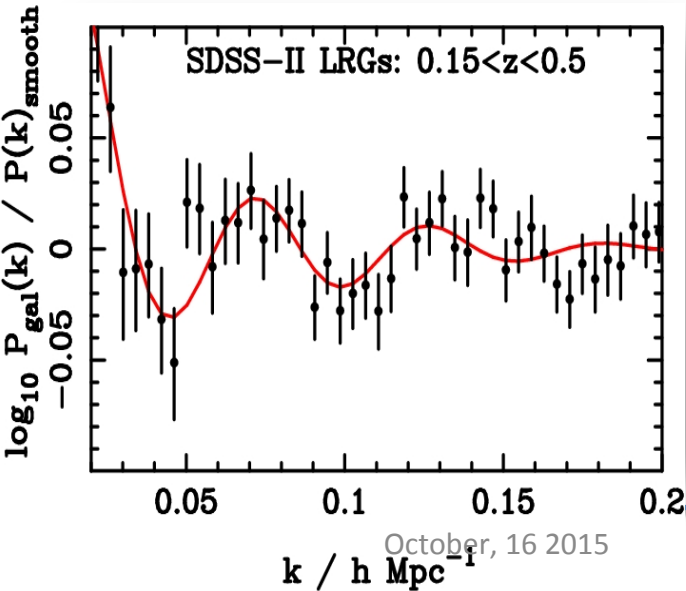
Galaxies

Microwave background

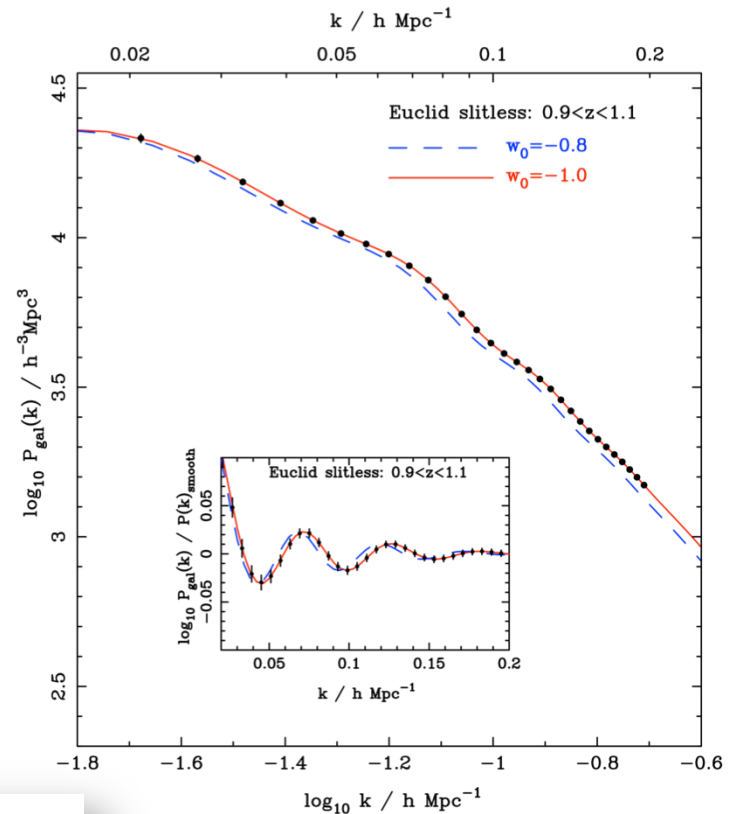
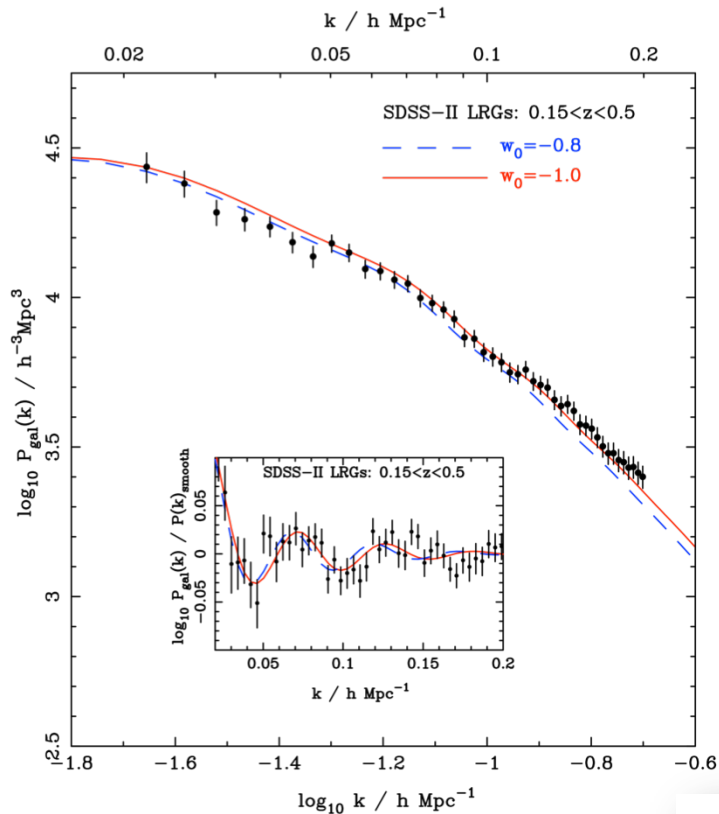


Percival et al. (2007, 2009, 2010); Anderson et al 2012

Planck 2013, XVI: Cosmological results

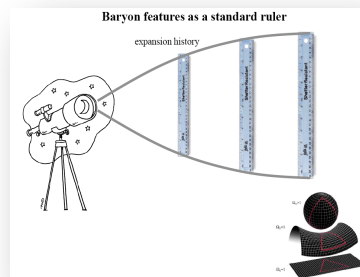


w(z) from Baryonic Acoustic Oscillations



SDSS LRGs at $z \sim 0.35$

Total effective volume
 $V_{eff} = 0.26 \text{ Gpc}^3 h^3$

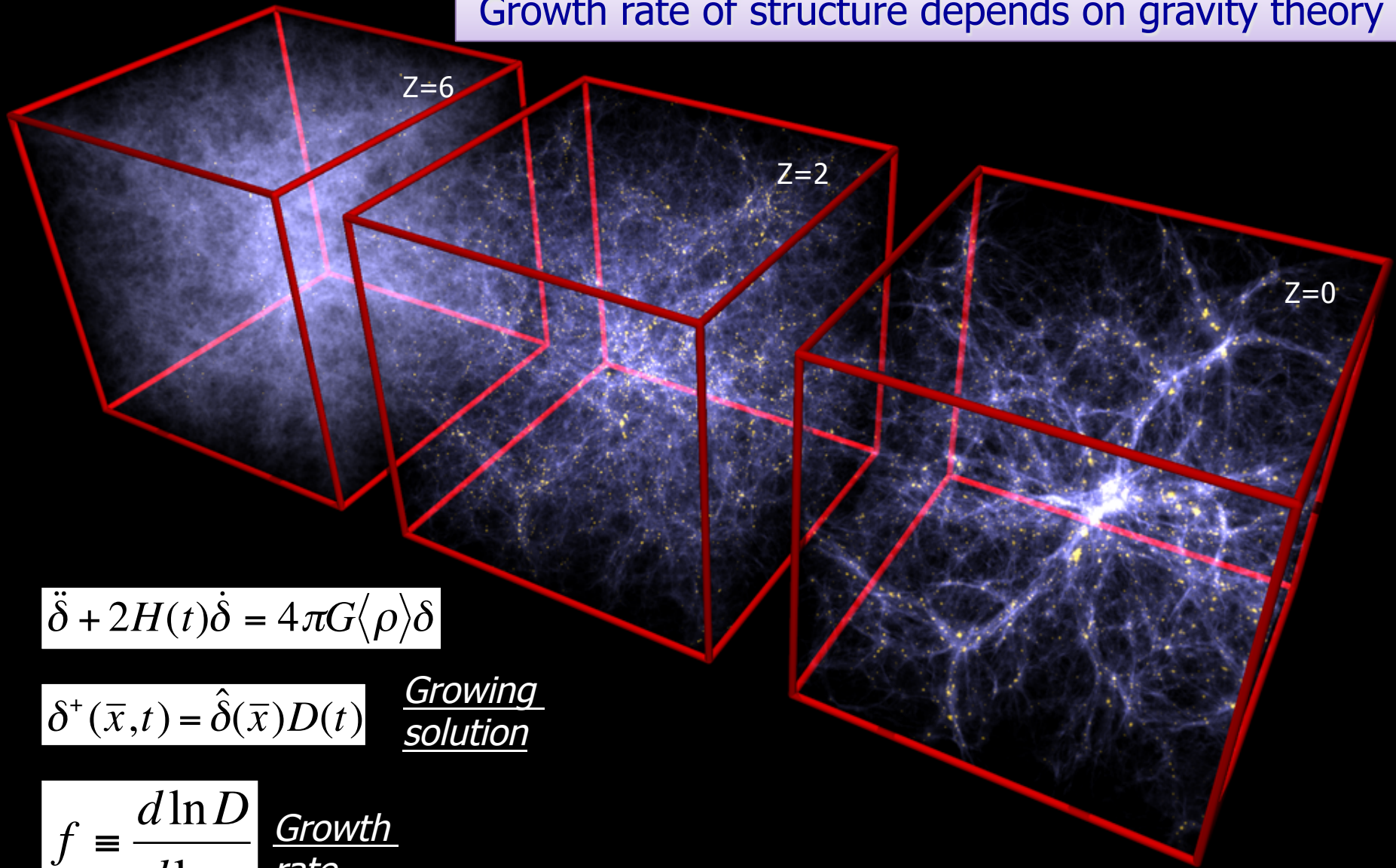


20% of the Euclid slitless data at $z \sim 1$

Total effective volume (of Euclid)
 $V_{eff} = 19.7 \text{ Gpc}^3 h^3$



Growth rate of structure depends on gravity theory



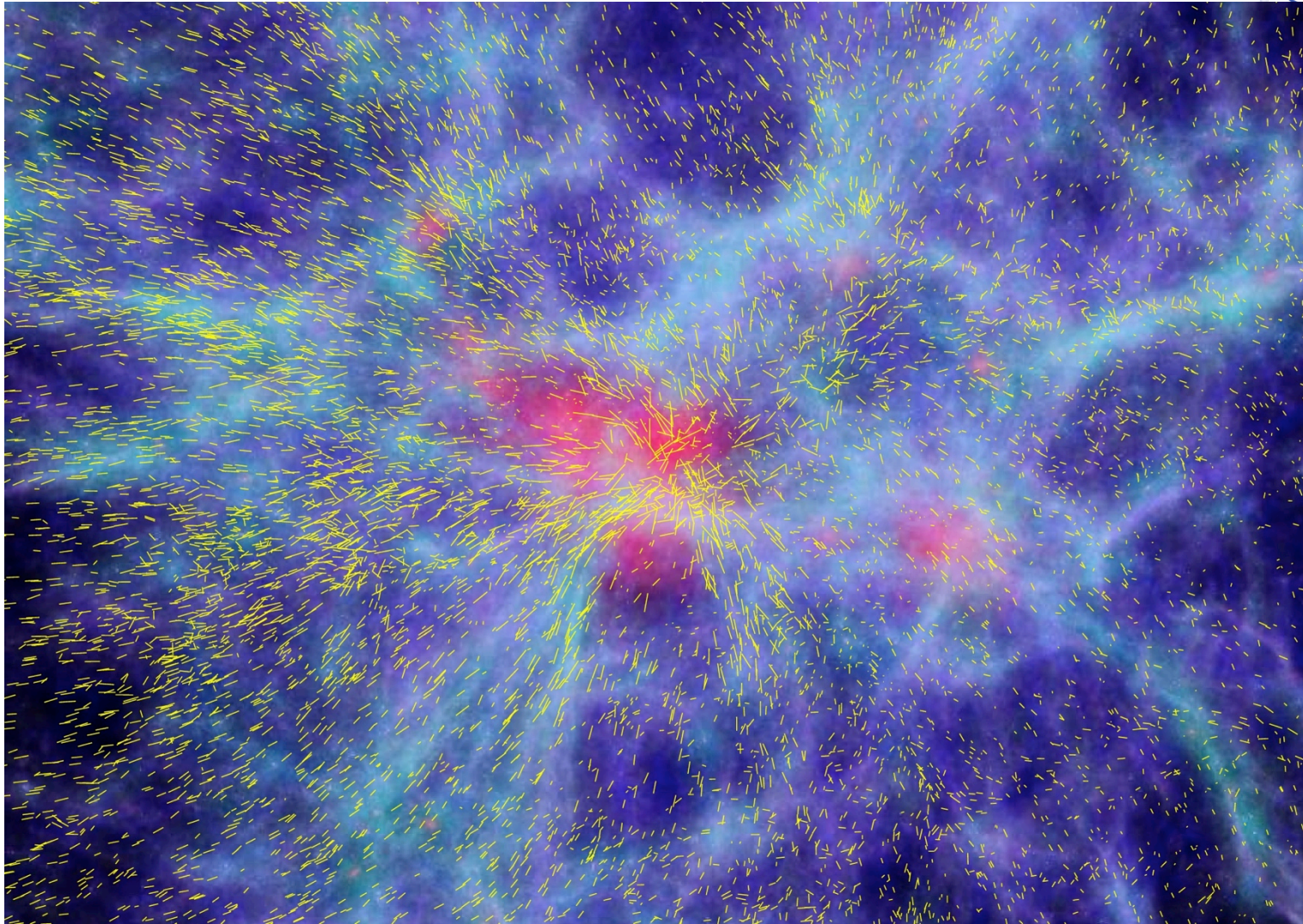
$$\ddot{\delta} + 2H(t)\dot{\delta} = 4\pi G\langle\rho\rangle\delta$$

$$\delta^+(\bar{x}, t) = \hat{\delta}(\bar{x})D(t) \quad \text{Growing solution}$$

$$f \equiv \frac{d \ln D}{d \ln a} \quad \text{Growth rate}$$

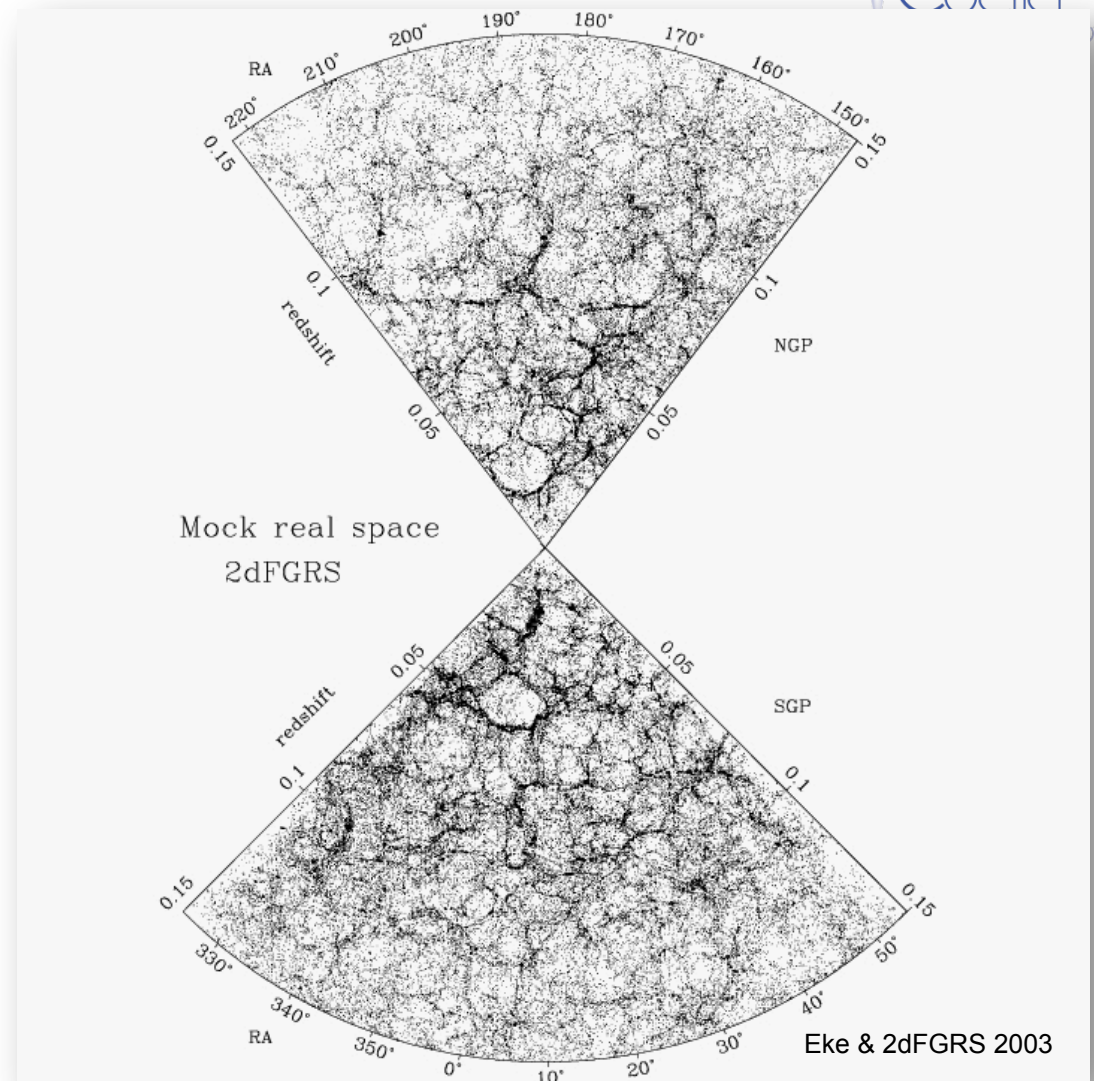
(Image credit: V. Springel)

Growth produces motions: galaxy peculiar velocities

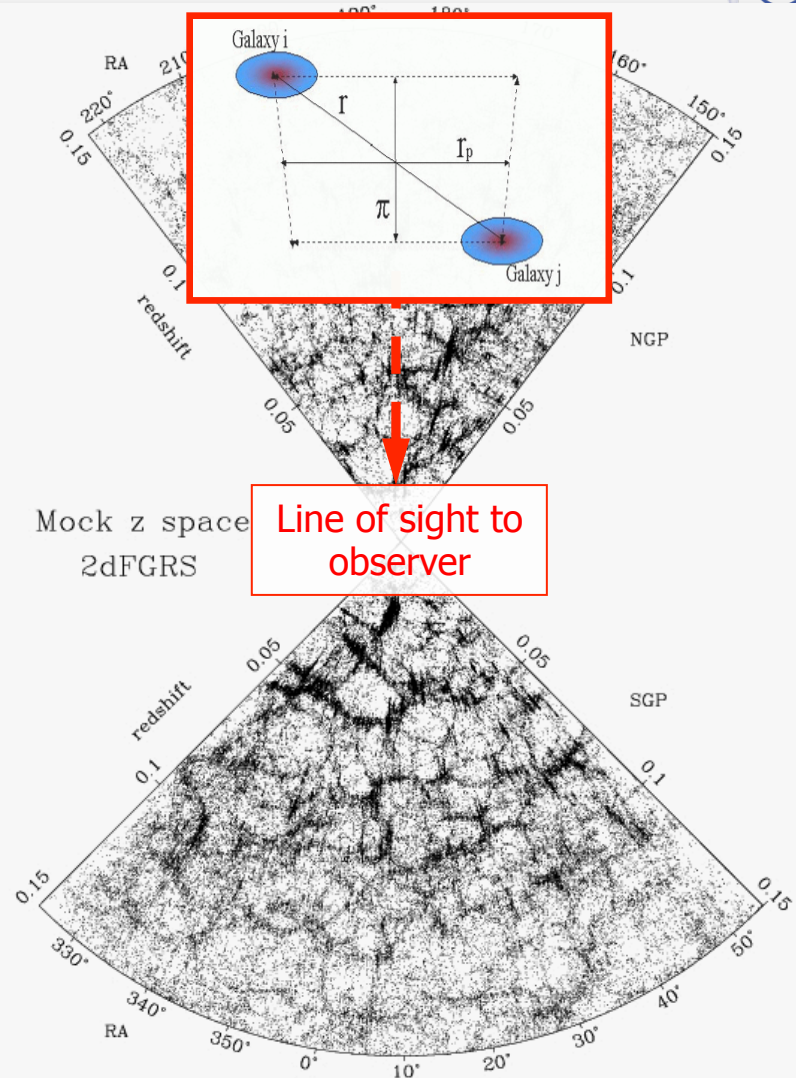
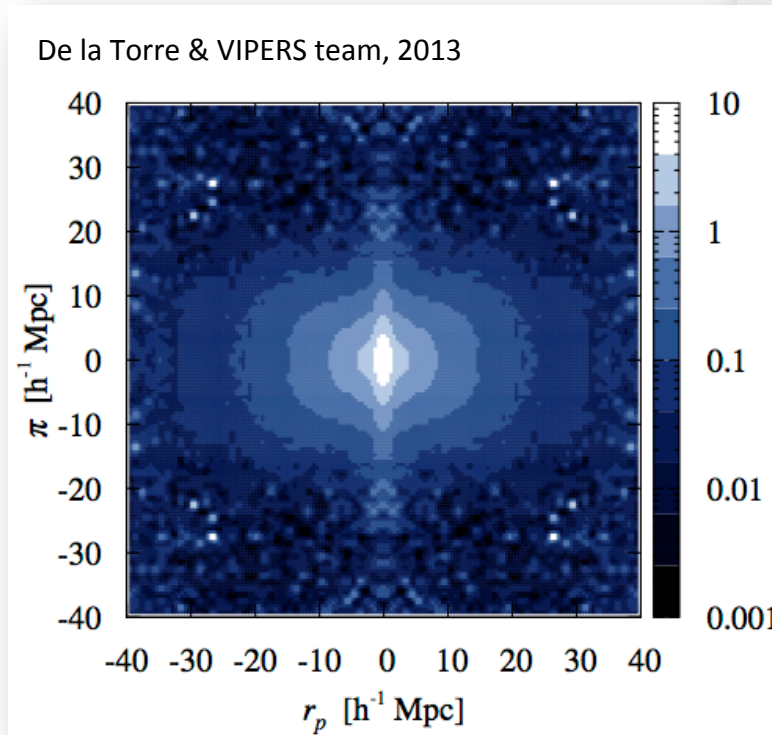


In galaxy redshift surveys peculiar velocities manifest themselves as redshift-space distortions (Kaiser 1987)

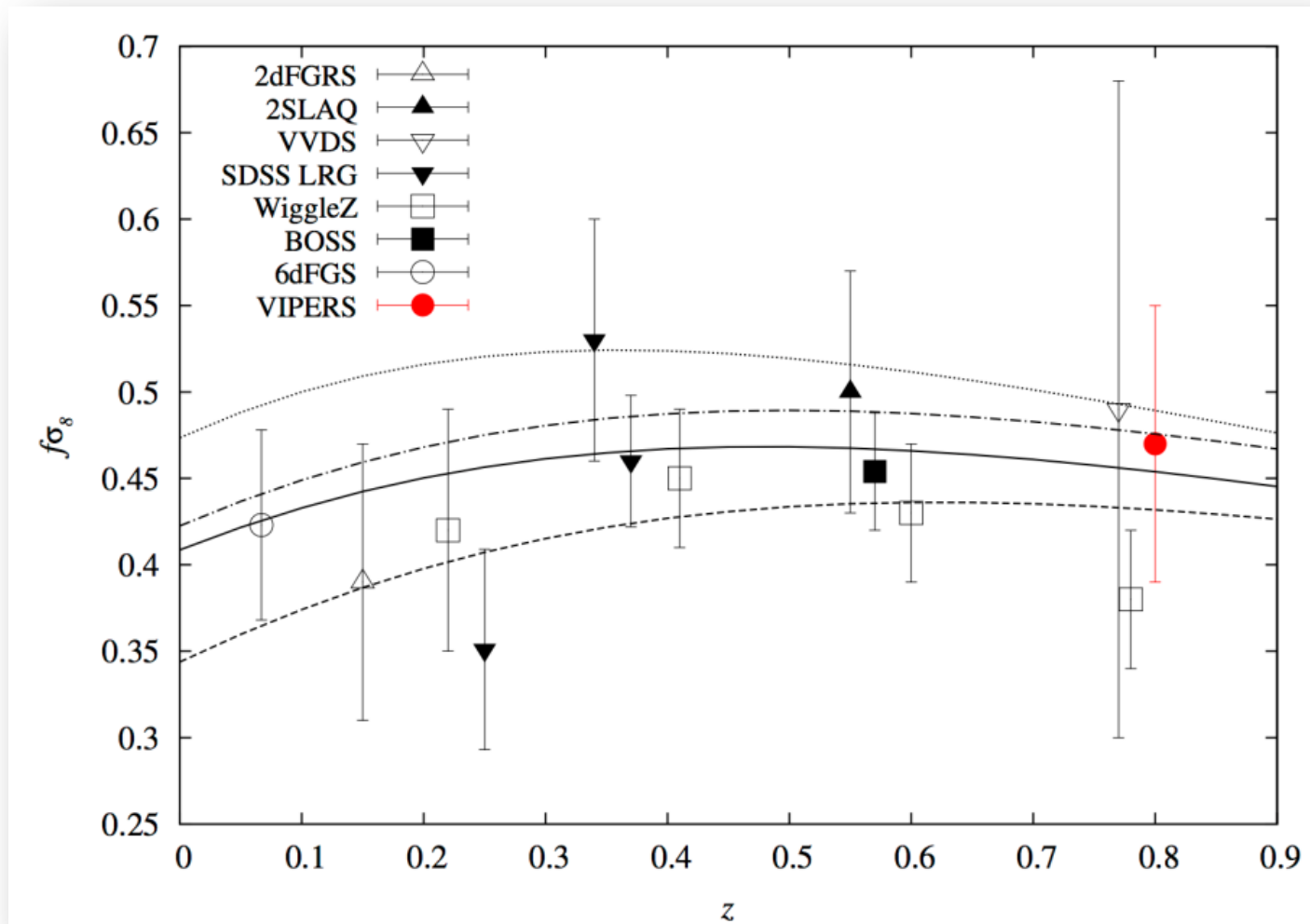
real space



In galaxy redshift surveys peculiar velocities manifest themselves as redshift-space distortions (Kaiser 1987)

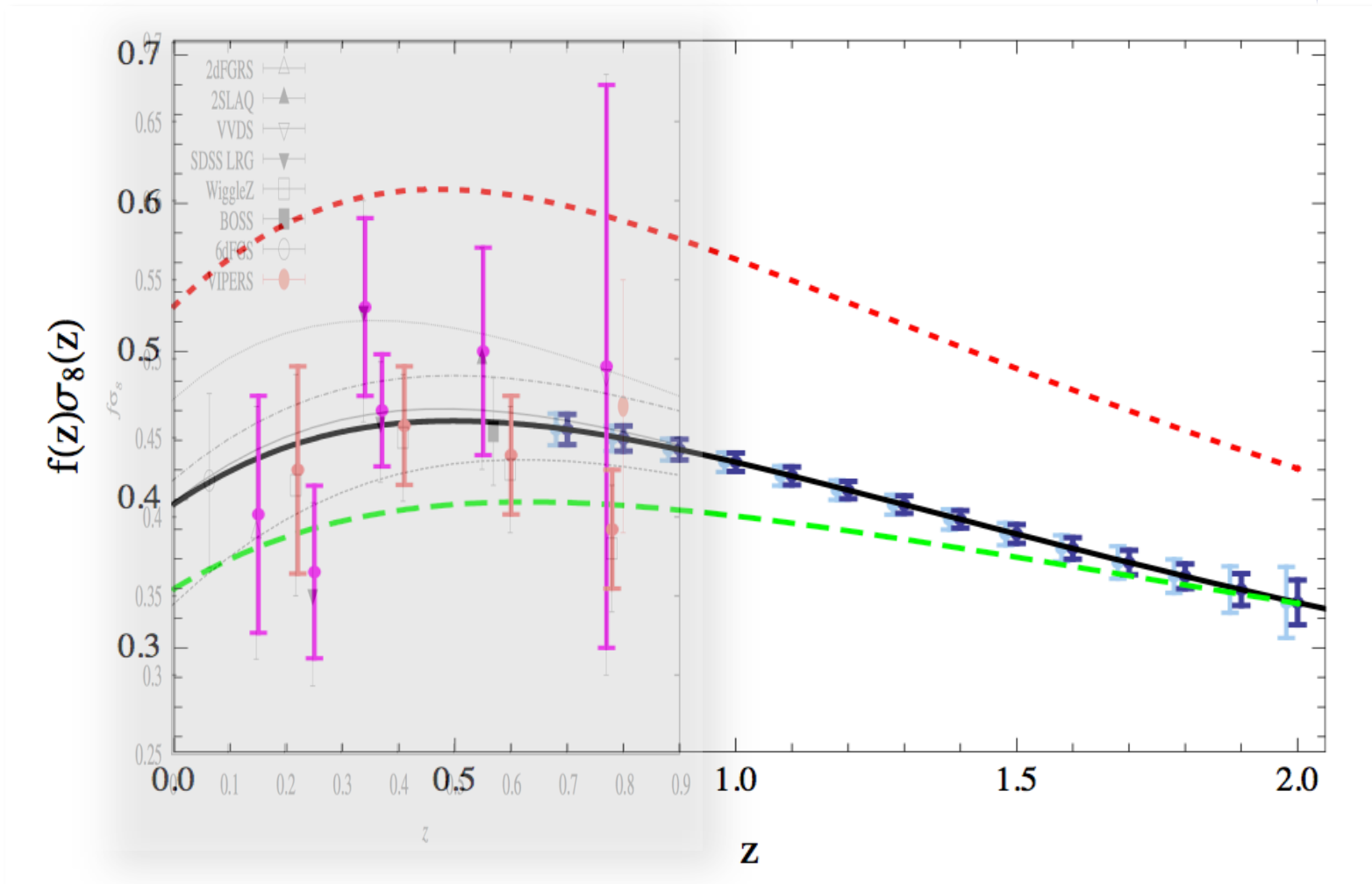


State-of-the art

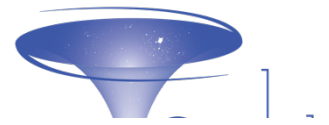


De la Torre, Guzzo & VIPERS team, 2013

Euclid potential



Majerotto, Guzzo, Samushia, Percival, Wang, et al. 2012

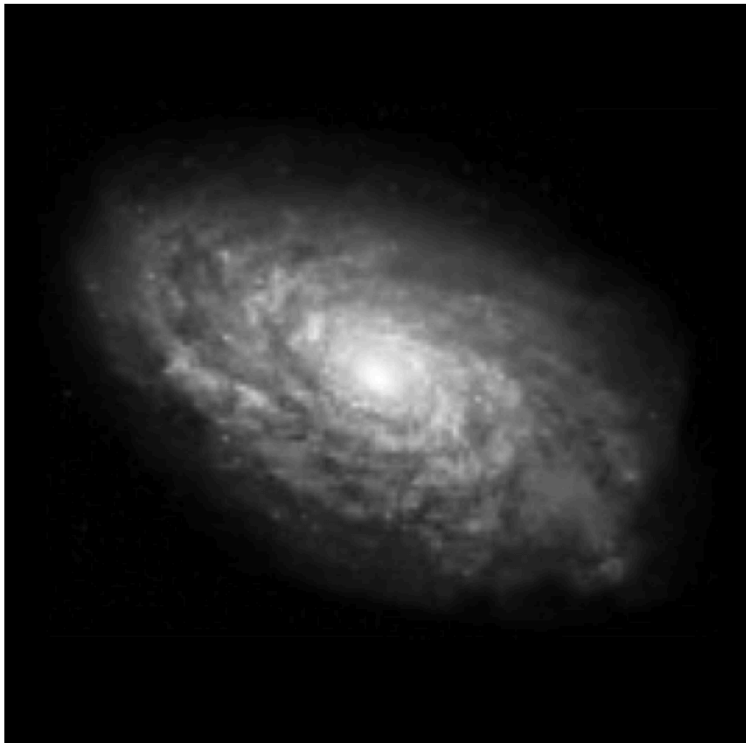


Weak Lensing increases galaxy ellipticity

$$\langle \epsilon \rangle = \langle \epsilon^S \rangle + \langle g \rangle$$

$$\langle \epsilon \rangle \approx \langle g \rangle.$$

Assuming galaxy orientations are randomly distributed

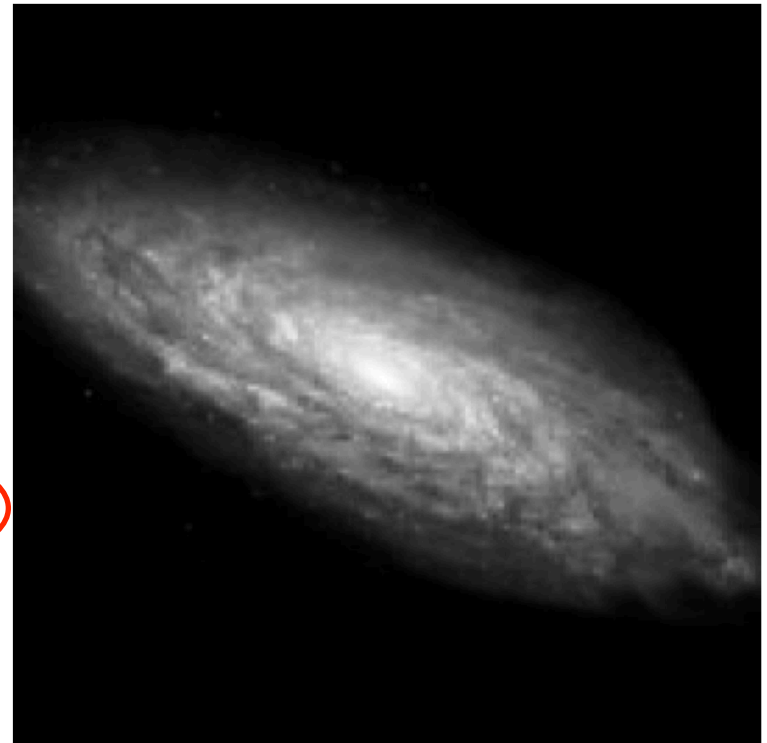


Dark matter



$g_i \sim 0.2$

Real data:
 $g_i \sim < 0.03$

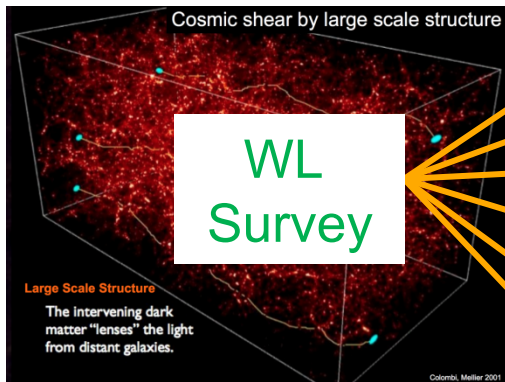


Kitching 2010

$$\begin{pmatrix} x_u \\ y_u \end{pmatrix} = \begin{pmatrix} 1 - g_1 & -g_2 \\ -g_2 & 1 + g_1 \end{pmatrix} \begin{pmatrix} x_l \\ y_l \end{pmatrix}$$

Cosmology with the Euclid

Weak Lensing survey



What is the expansion rate of the Universe?

What is the expansion rate of the Universe?

How does structure form within this background?

What are the neutrino masses, matter density?

What is f_{nl} , which quantifies non-Gaussianity?
GR-horizon effects

Does the potential change along line-of-sight to CMB

Understanding Dark Energy

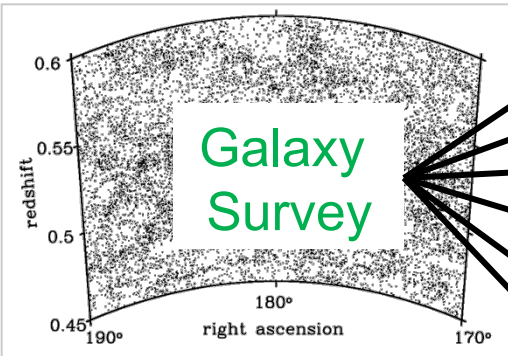
Understanding energy-density, gravity

Understanding energy-density

Understanding Inflation, GR

Understanding DE, GR

Cosmology with the Euclid Redshift Survey



Baryon Acoustic Oscillations

Alcock-Paczynski effect

Redshift-Space Distortions

Comoving clustering

Large-scale shape

ISW effect

What is the expansion rate of the Universe?

What is the expansion rate of the Universe?

How does structure form within this background?

What are the neutrino masses, matter density?

What is f_{nl} , which quantifies non-Gaussianity?
GR-horizon effects

Does the potential change along line-of-sight to CMB

Understanding Dark Energy

Understanding energy-density, gravity

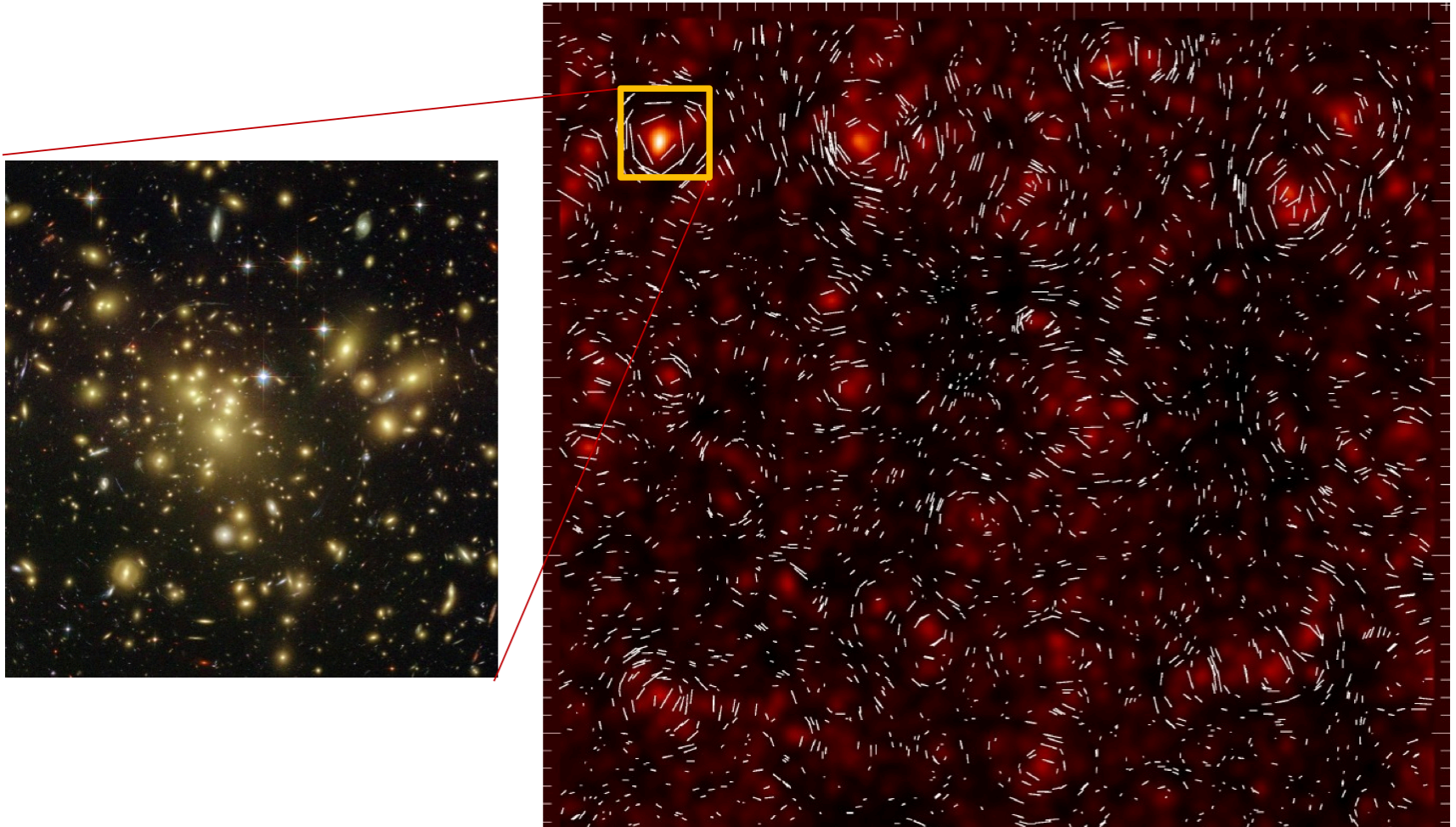
Understanding energy-density

Understanding Inflation, GR

Understanding DE, GR



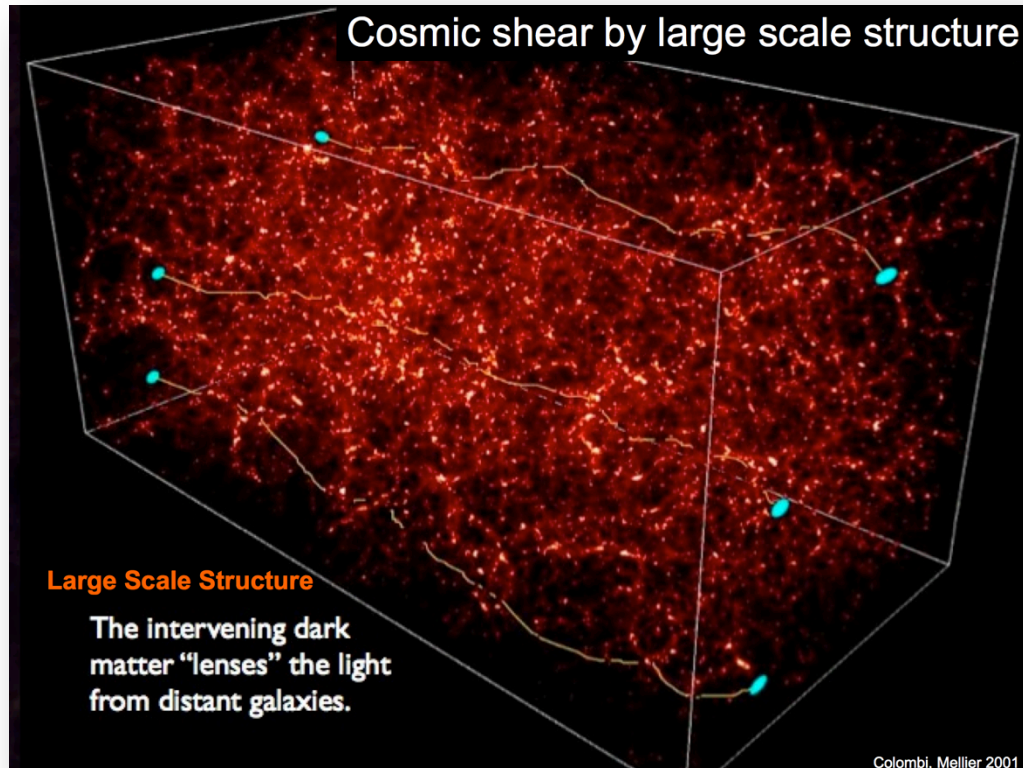
Cosmological distortion field projected on the sky



Weak Lensing tomography: get matter $P(k,z)$



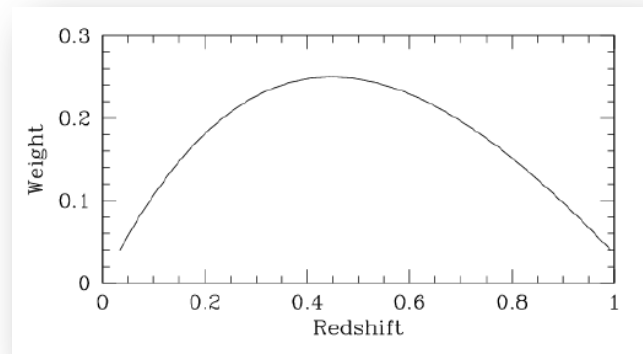
- The lensing kernel is most sensitive to structure halfway between the observer and the source. But the kernel is broad: we do not need precise redshifts for the sources: **photometric redshifts are fine**
- Also, since the kernel is broad the tomographic bins are very correlated. The gain saturates quickly with the number of bins: **not many z bins**



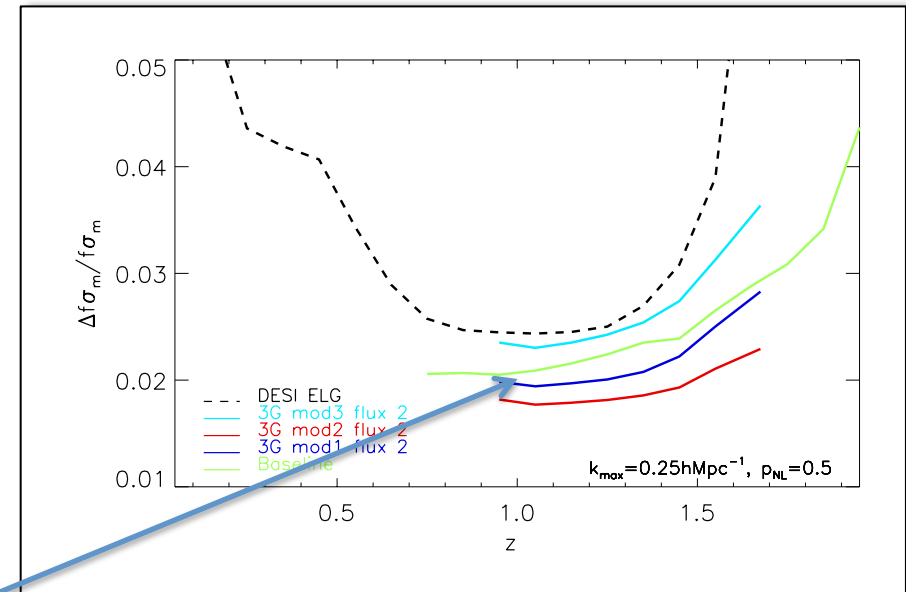
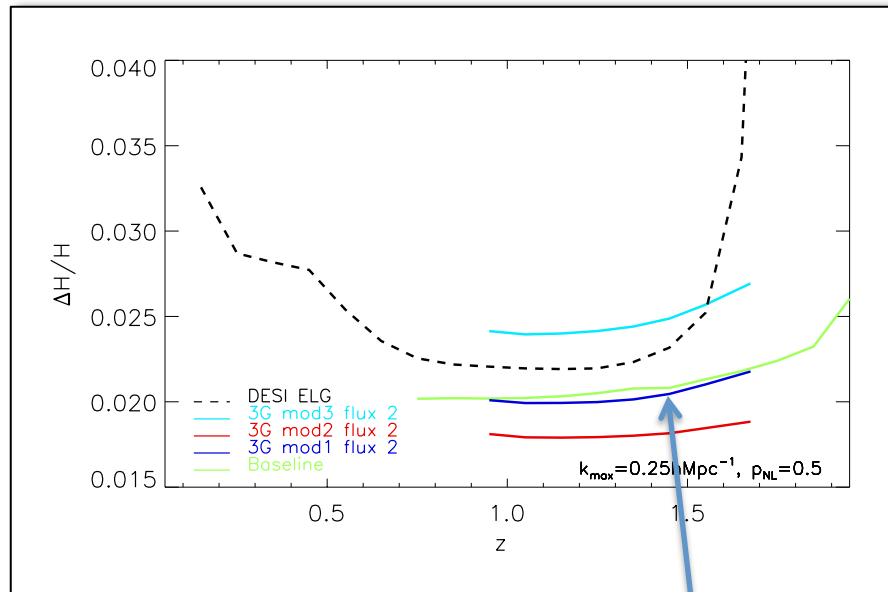
$$C_{ij}(\ell) = \int_0^{r_H} dr W_{ij}^{GG}(r) P_{\delta\delta}\left(\frac{\ell}{S_k(r)}; r\right)$$

Geometry Large Scale Structure

Lensing Power Spectrum



Forecast on $H(z)$ and growth errors



Flux $> 2 \times 10^{-16} \text{ erg cm}^{-2} \text{ s}^{-1}$

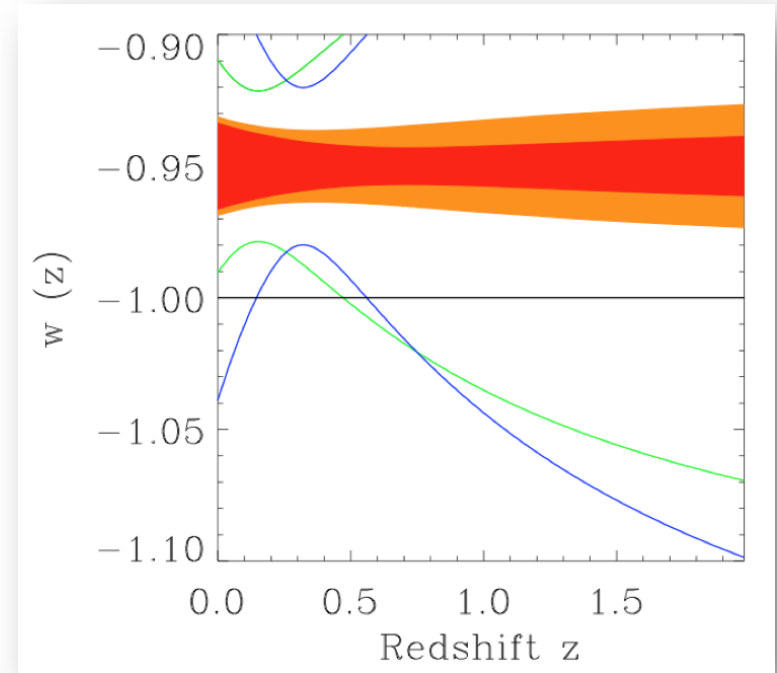
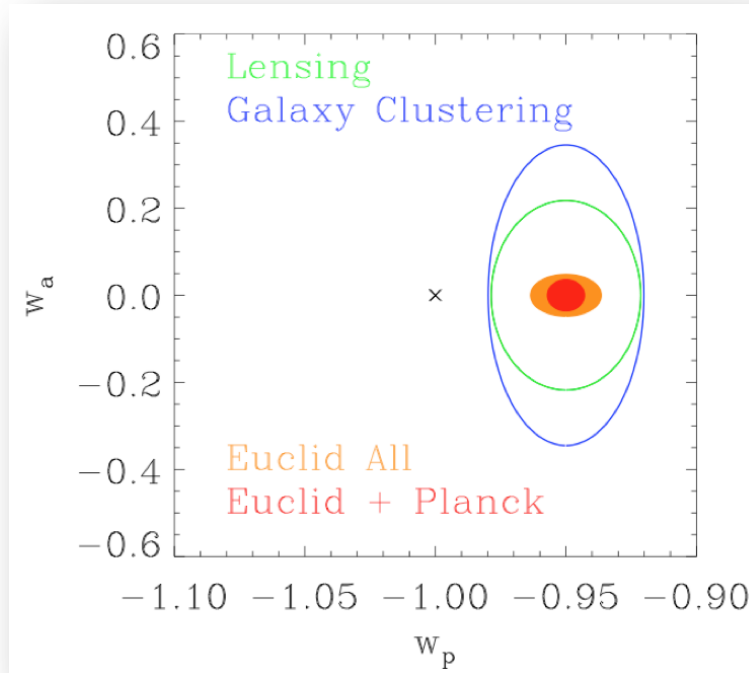
$z=0.9$ start, Mod. 1, No. of galaxies = 26.2M FoM = 433 (γ fixed)

- Deep flux limit: $2 \times 10^{-16} \text{ erg cm}^{-2} \text{ s}^{-1}$
- 3 up-to-date models for dN/dz by Pozzetti, Geach & Hirata
- Forecast code by R. Bean

The full power of EUCLID (RedBook reference)



1. Dark Energy equation of state from combined Weak Lensing and Galaxy Clustering (BAO)



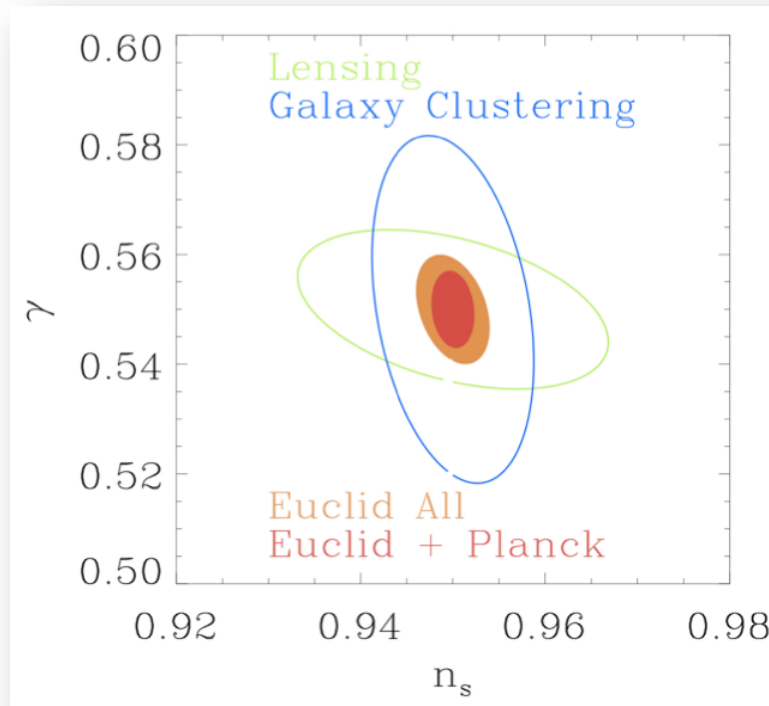
Answering Euclid key science question 1: Is dark energy simply a cosmological constant, or is it a field that evolves dynamically with the expansion of the Universe?

The full power of EUCLID



2. Growth rate of structure from combined Weak Lensing (tomography) and Galaxy Clustering (redshift-space distortions)

$$f(z)=[\Omega_m(z)]^\gamma$$



$$P_o(k)=Ak^n$$

Answering Euclid key science question 2: Is is the apparent acceleration instead a manifestation of a breakdown of General Relativity on the largest scales?

Euclid performance

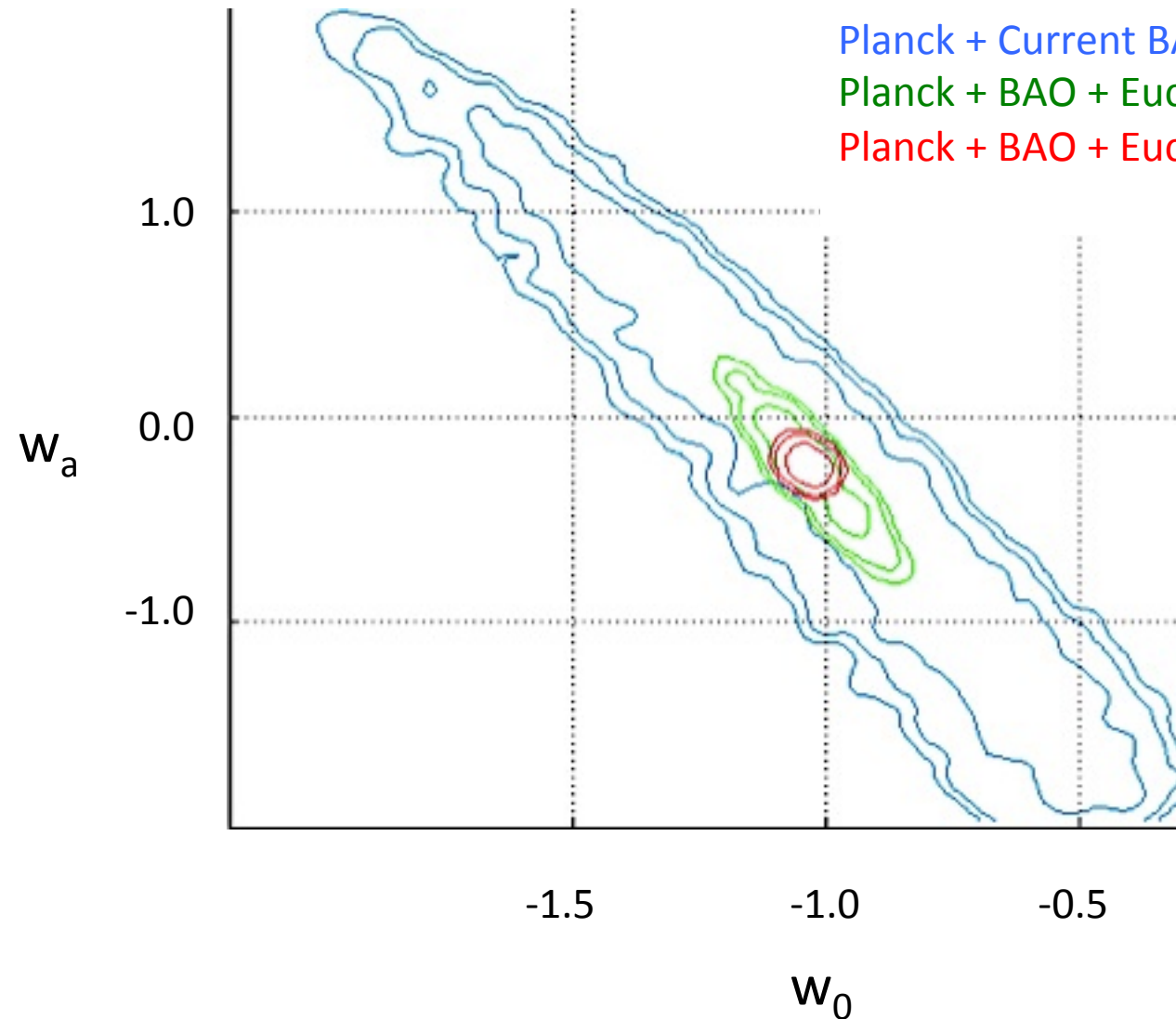
	Modified Gravity	Dark Matter	Initial Conditions	Dark Energy		
Parameter	γ	m_ν/eV	f_{NL}	w_p	w_a	FoM
Euclid Primary	0.010	0.027	5.5	0.015	0.150	430
Euclid All	0.009	0.020	2.0	0.013	0.048	1540
Euclid+Planck	0.007	0.019	2.0	0.007	0.035	4020
Current (09/2011)	0.200	0.580	100	0.100	1.500	~10
Improvement Factor	30	30	50	>10	>50	>300

Euclid addresses most aspects of the current cosmological paradigm

Planck + Current BAO MCMC Chains

Planck + BAO + Euclid Weak Lensing

Planck + BAO + Euclid Weak Lensing + Euclid BAO

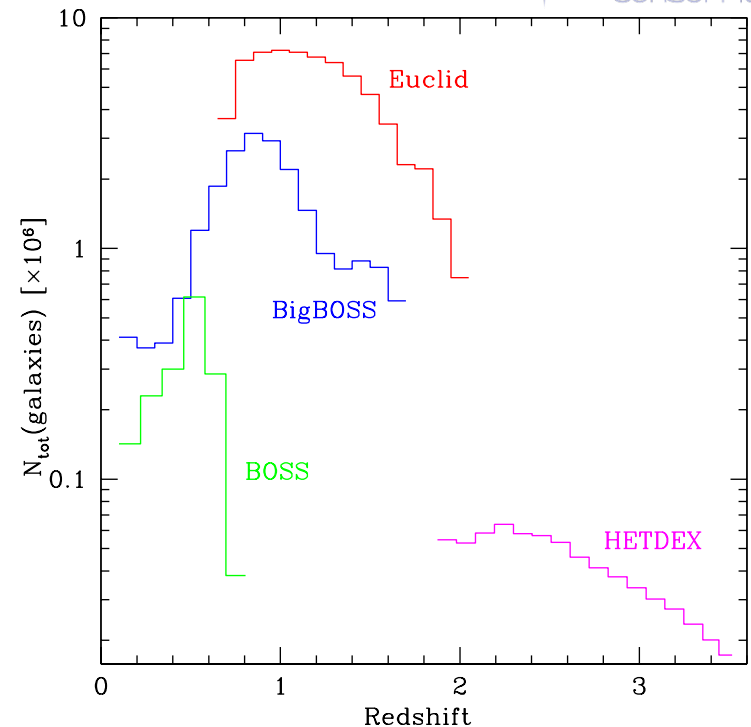


- Fisher Matrix (Gaussian Likelihood) predictions
 - Part of “*Inter-SWG Taskforce*” (IST) on forecasts (provided by T. Kitching)
- For WL
 - including physical IA Model
- For GC
 - Updates from Wang, Chuang, Hirata (2013)
- Combined with *Real* Planck MCMC chains

(T. Kitching & Euclid Forecast IST)

- Wide survey
 - 15,000deg²
 - 4 dithers
 - NIR Photometry
 - Y, J, H
 - 24mag, 5 σ point source
 - NIR slitless spectroscopy
 - 1100-2000nm
 - 3×10^{-16} erg cm⁻² s⁻¹ 3.5 σ line flux
 - 2 dispersion directions, 2 wavebands
 - 52M galaxies

- Deep survey
 - 40deg²
 - 48 dithers
 - 12 passes, as for wide survey
 - dispersion directions for 12 passes >10deg apart



Euclid Mission Requirements



	Wide survey	Deep survey
Survey: 6 years		
size	15, 000 deg ²	40 deg ² N/S
VIS imaging		
Depth	$n_{gal} > 30/\text{arcmin}^2$ $M_{AB} = 24.5, 10\sigma$ for gal size $0.3 \gg$ $\rightarrow \langle z \rangle \sim 0.9$	$M_{AB} = 26.5$
PSF size knowledge	$\sigma[R^2]/R^2 < 10^{-3}$	
Multiplicative bias in shape	$\sigma[m] < 2 \cdot 10^{-3}$	
Additive bias in shape	$\sigma[c] < 2 \cdot 10^{-4}$	
Ellipticity RMS	$\sigma[e] < 2 \cdot 10^{-4}$	
NIP photometry: YJH		
Depth	24 M_{AB}	26 M_{AB}
NIS spectroscopy: 4 R exp., 3 R orientations		
Flux limit (erg/cm ² /s)	$2 \cdot 10^{-16}$	$5 \cdot 10^{-17}$
Completeness	> 45 %	>99%
Purity	>80%	>99%
Confusion	3 rotations	>12 rotations

WL and systematics

$$\gamma^{obs} = (1 + m) \times \gamma^{true} + c$$

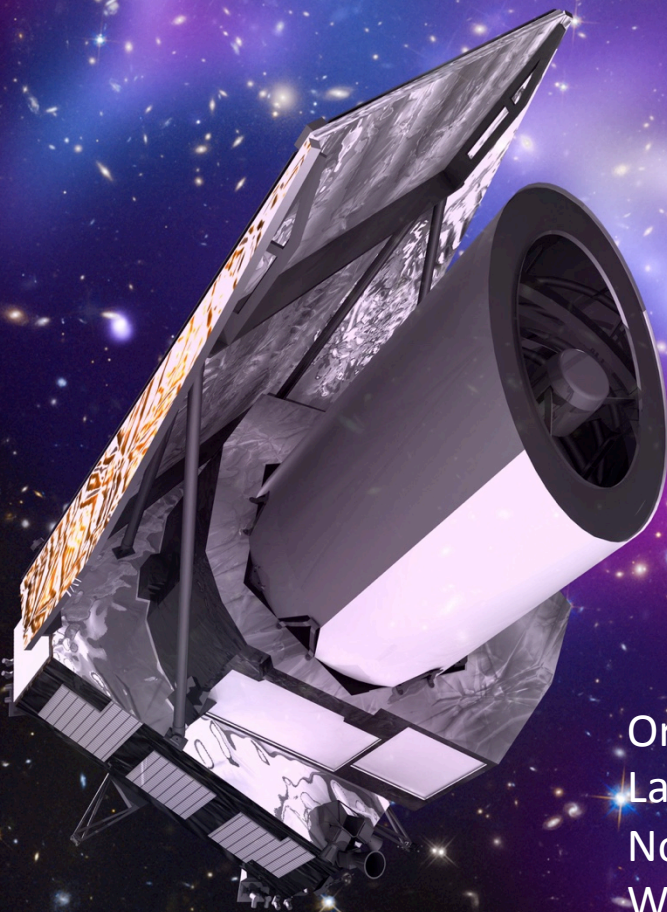
$$C_l^{true} \approx [1 + 2\langle m \rangle] \times C_l^{obs} + \langle c^2 \rangle$$

- Small PSF, **Knowledge** of the PSF size
- Knowledge of distortion
- Method to correct distortion
- Method to correct Non-convolutive PSF
- Stability in time \rightarrow space telescope
- Visible photom photo-z accuracy: $0.05x(1+z)$
- Catastrophic $z < 10\%$

GC and systematics

- Understand selection \rightarrow Deep field (photo+spectro)
 - Completeness
 - Purity

Euclid



- Payload:
Airbus
Defence & Space
- Spacecraft:
Thales
AleniaSpace
Torino
- Instruments:
Euclid Consortium
 - 13 EU countries + NASA+ US labs
 - >120 institutes/labs
 - >1100 members

Orbit: L2

Launch: Q4 2020

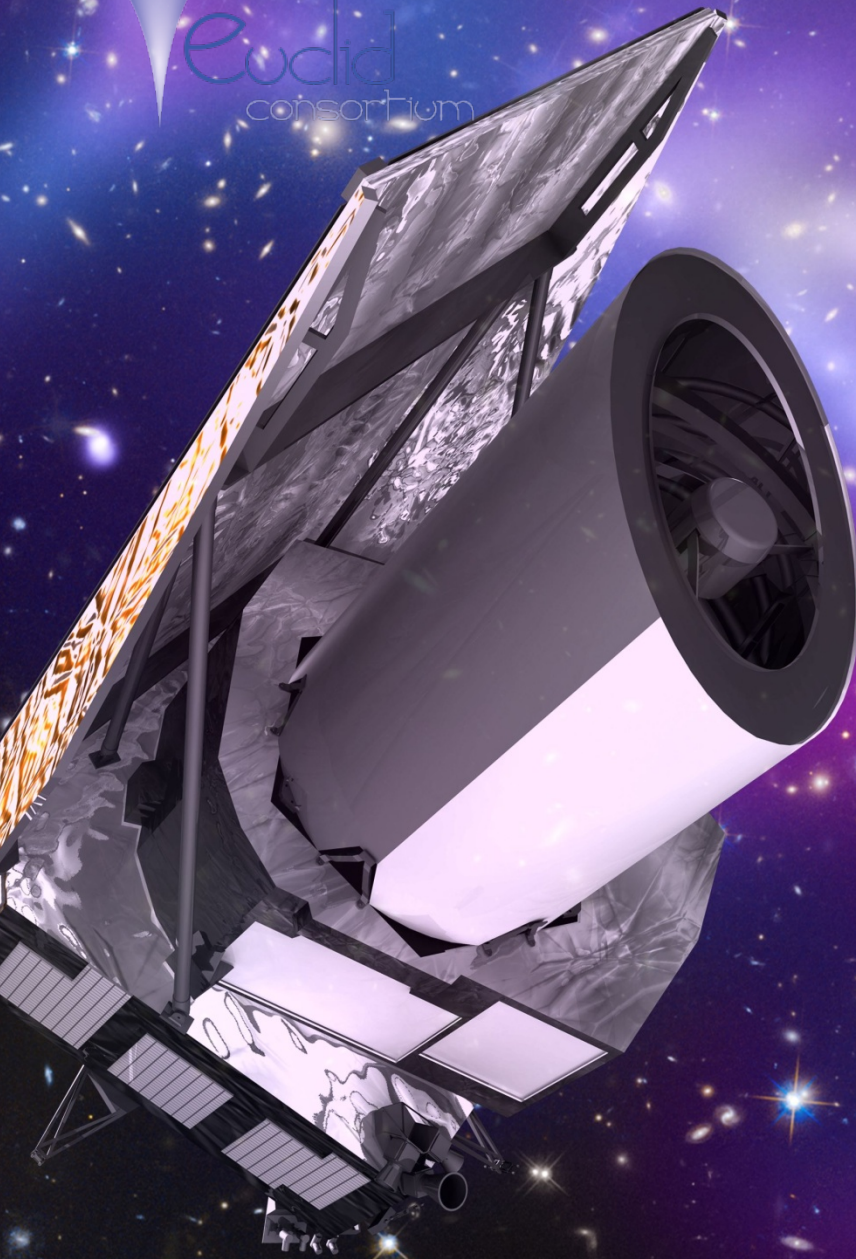
Nominal Mission: 6.5 y

Wide survey: 15000 sq deg



Euclid

Euclid
consortium



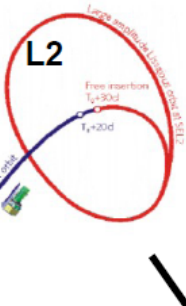
- **ESA-led mission**
- **Selected in Oct. 2011 - Fully funded**
- **Partners: ESA, TAS, Airbus DS, Euclid Consortium (EC)**
- **Cost: ~ 850 Meuros (~ 2/3 ESA; ~1/3 EC)**
- **Overall Mass: ~2020 kg**
- **Power : 1710 W**
- **Telescope :**
 - 1.2m aperture primary ,3 mirror anastigmat
- **2 Instruments (VIS, NISP); 2 channels**
 - Wide field instrument, VIS: 36 e2v 4kx4k CCDs
 - 576 M pixels, 0.11 arcsec/pix, 0.53 deg² FoV
 - Photom.+spectrom.: 16 H2GR HgCdTe detectors;
 - 64 Mpixels, 0.30 arcsec/pix, 0.53 deg² FoV (=VIS)
 - Grism slitless spectro (1B + 3R grisms)
- **Downlink Rate: X/X + K-band to Ground Station 55 Mbits/s. 850 Gbit/day to transfer 4hr/day.**
- **Ground Segment: ESA (50%,) EC (50%, EC leads science and external data): 1.5 billion galaxies for WL, 30 million redshifts, 12 billion sources (3sigma)**
- **L2 orbit**
- **Launch Vehicle – Soyuz-Fregat**
- **Launch date 2020, from Kourou space port**
- **6.5 years mission+additional surveys (exopl, SN)**
- **Main surveys: 15,000 deg²+40 deg² 2 mag. deeper**
- **Science drivers: DE**
- **Science leads: Euclid Consortium**

The Euclid Mission in one slide

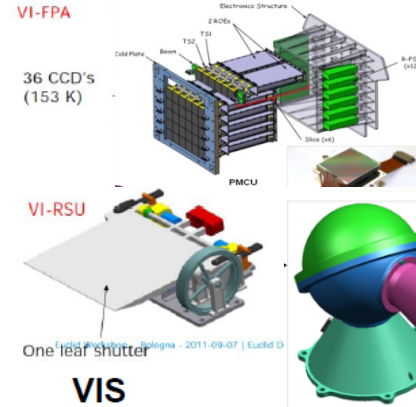
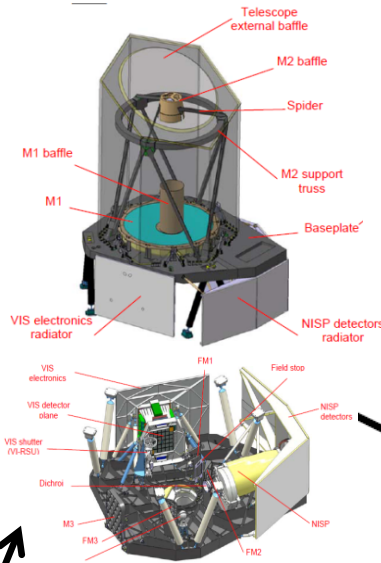


Soyuz@Kourou

Q1 2020



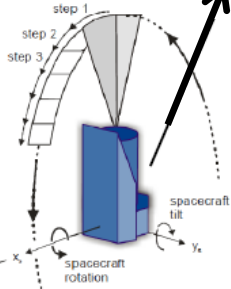
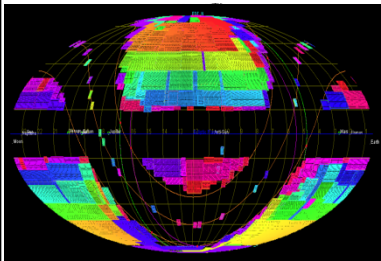
PLM+SVM: 2010-2019



VIS imaging: 2010-2020

(VIS team)

Surveys: 2010-2028 (Survey WG)



6 yrs - 15,000 deg²

Commissioning – SV

Euclid opération:

5.5 yrs:Euclid Wide+Deep

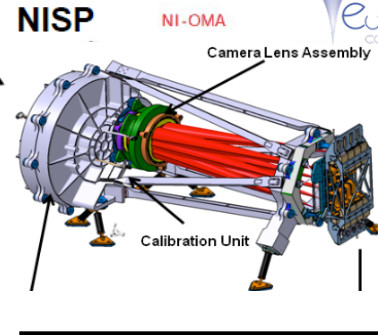
+ : SNIa, mu-lens, MW?

Ground data



October, 16-2015

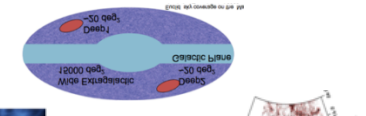
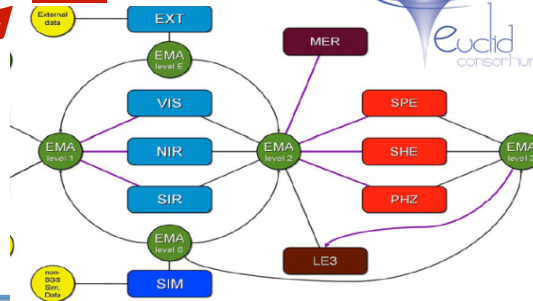
NISP



NIR spectro-imaging

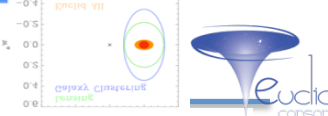
2010-2020 (NISP team)

SGS: 2010-2028



SWG:

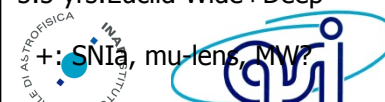
2019-2028



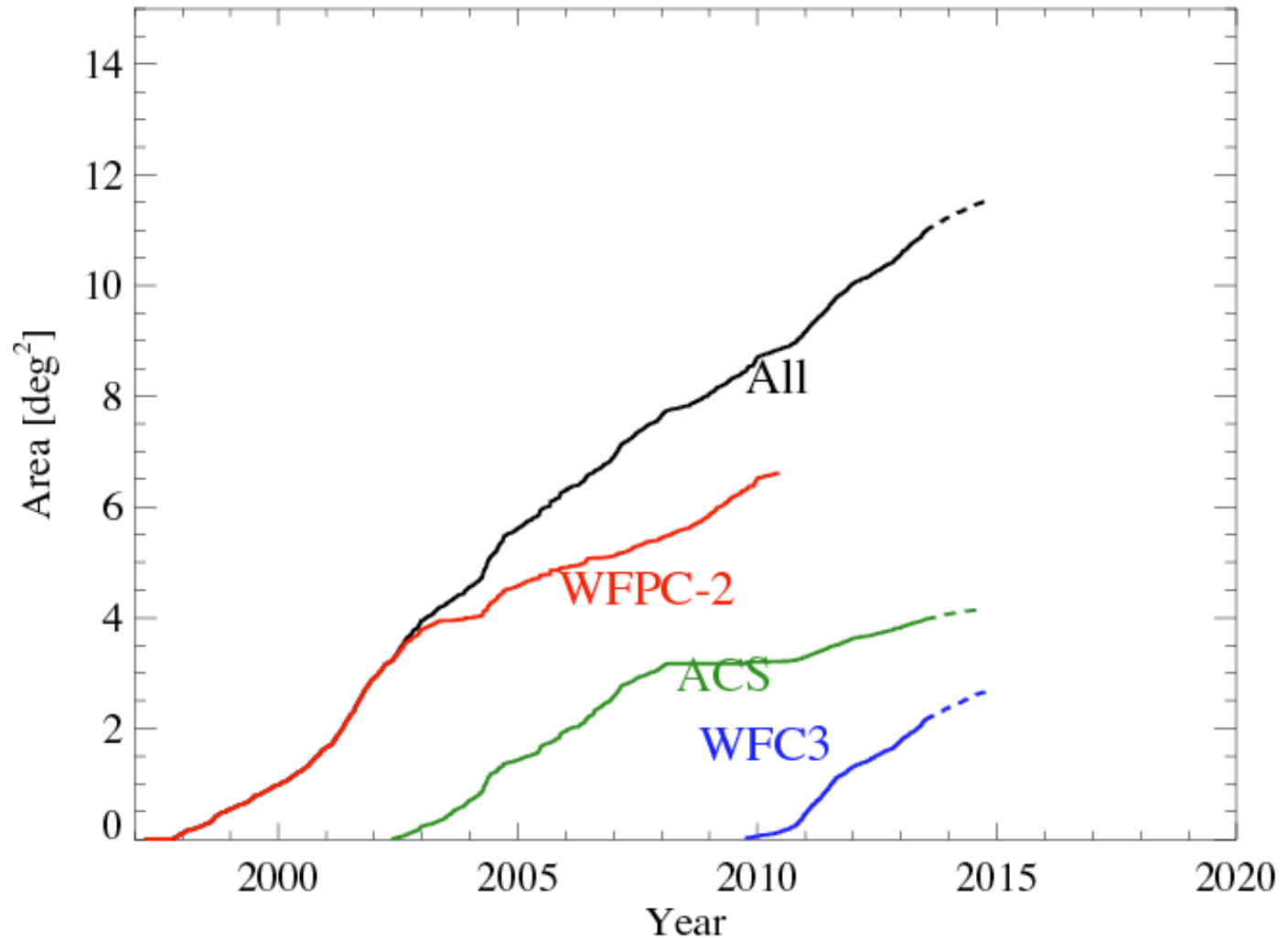
20-30 PB data processing (EC-SGS team)

Science analyses

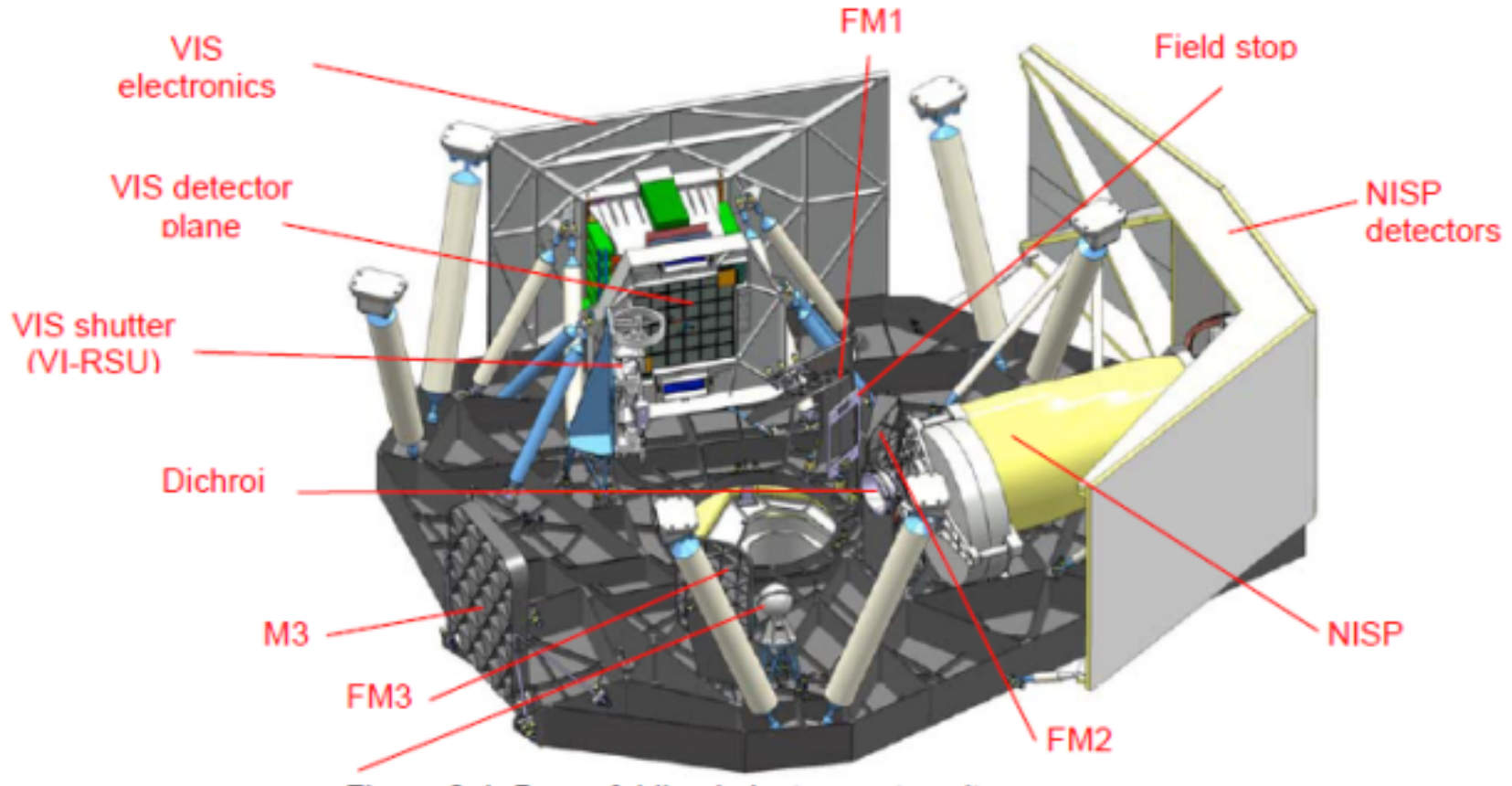
L. Valenziano on behalf of the EC



HST survey amplitude

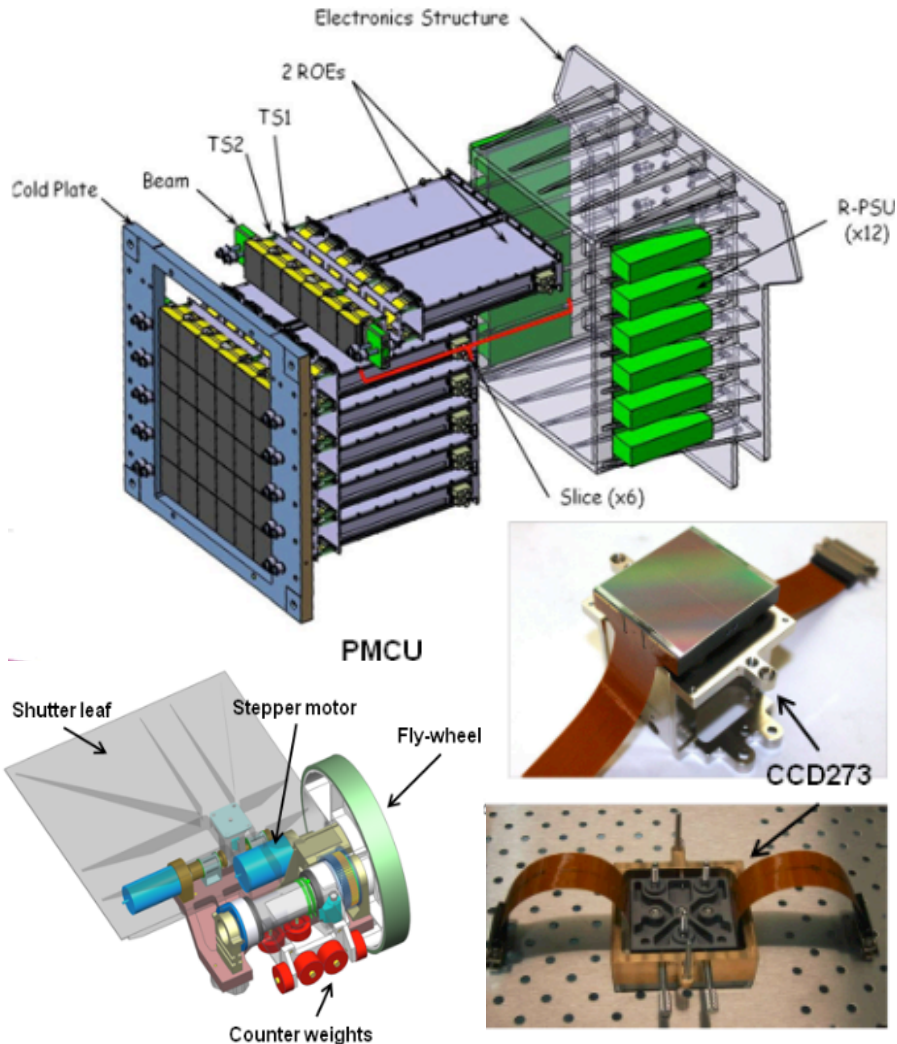


Euclid Instruments

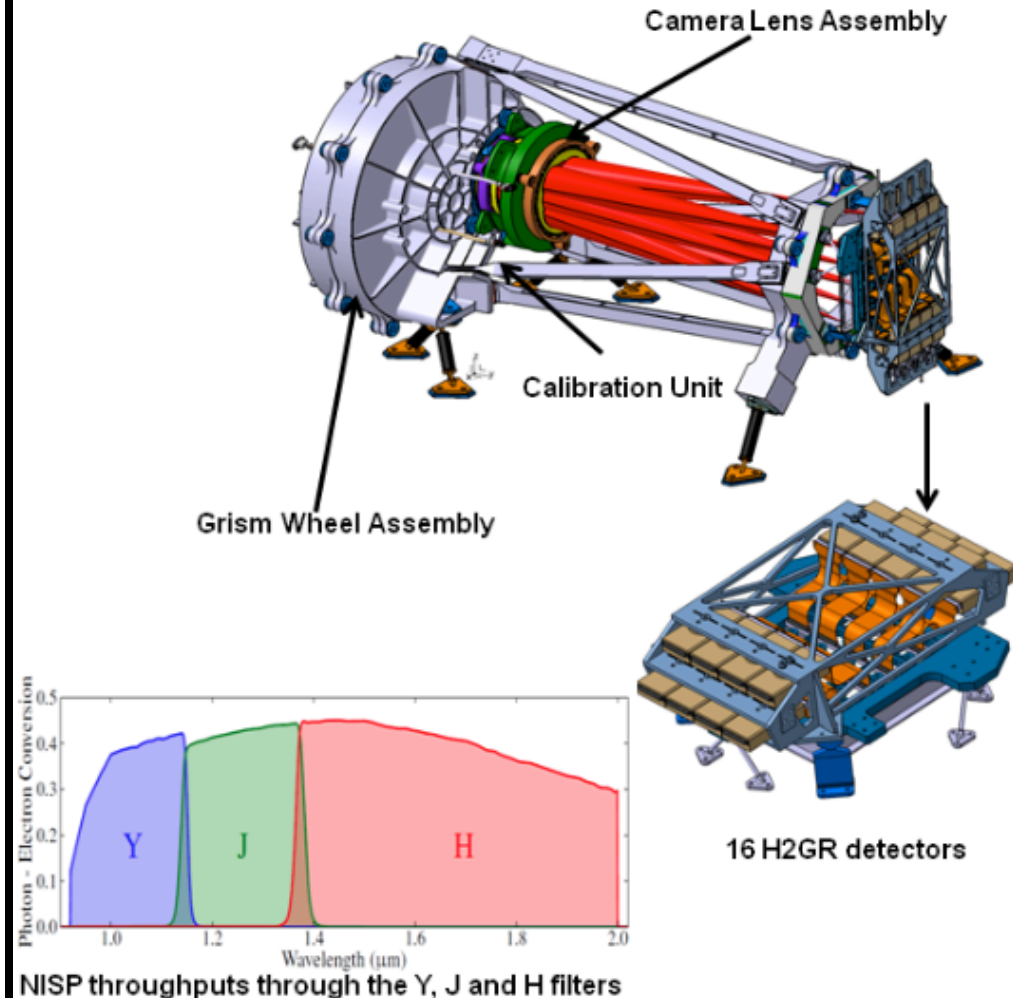


VIS and NISP Scientific Instrument

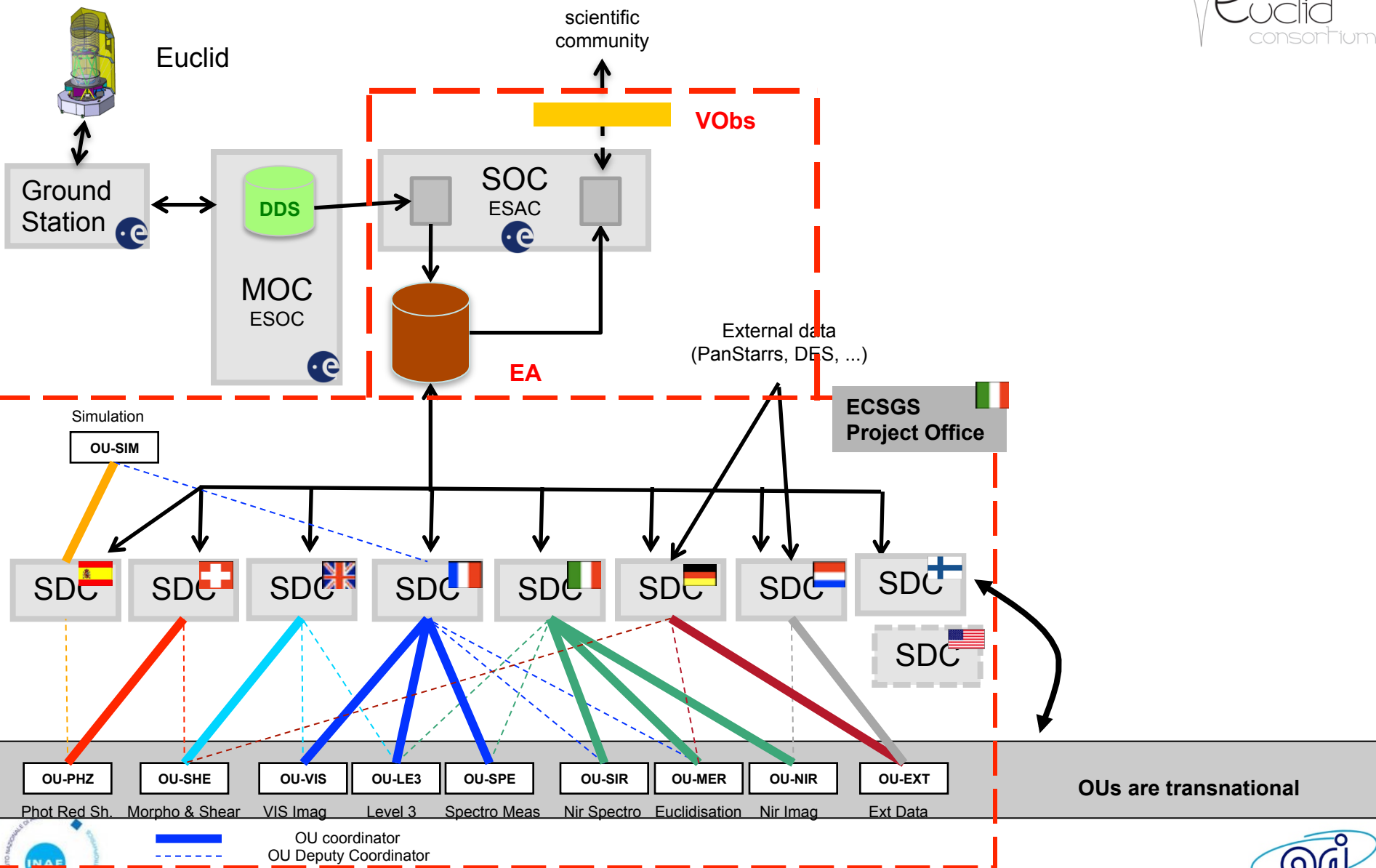
Courtesy: S. Pottinger, M. Cropper and the VIS team



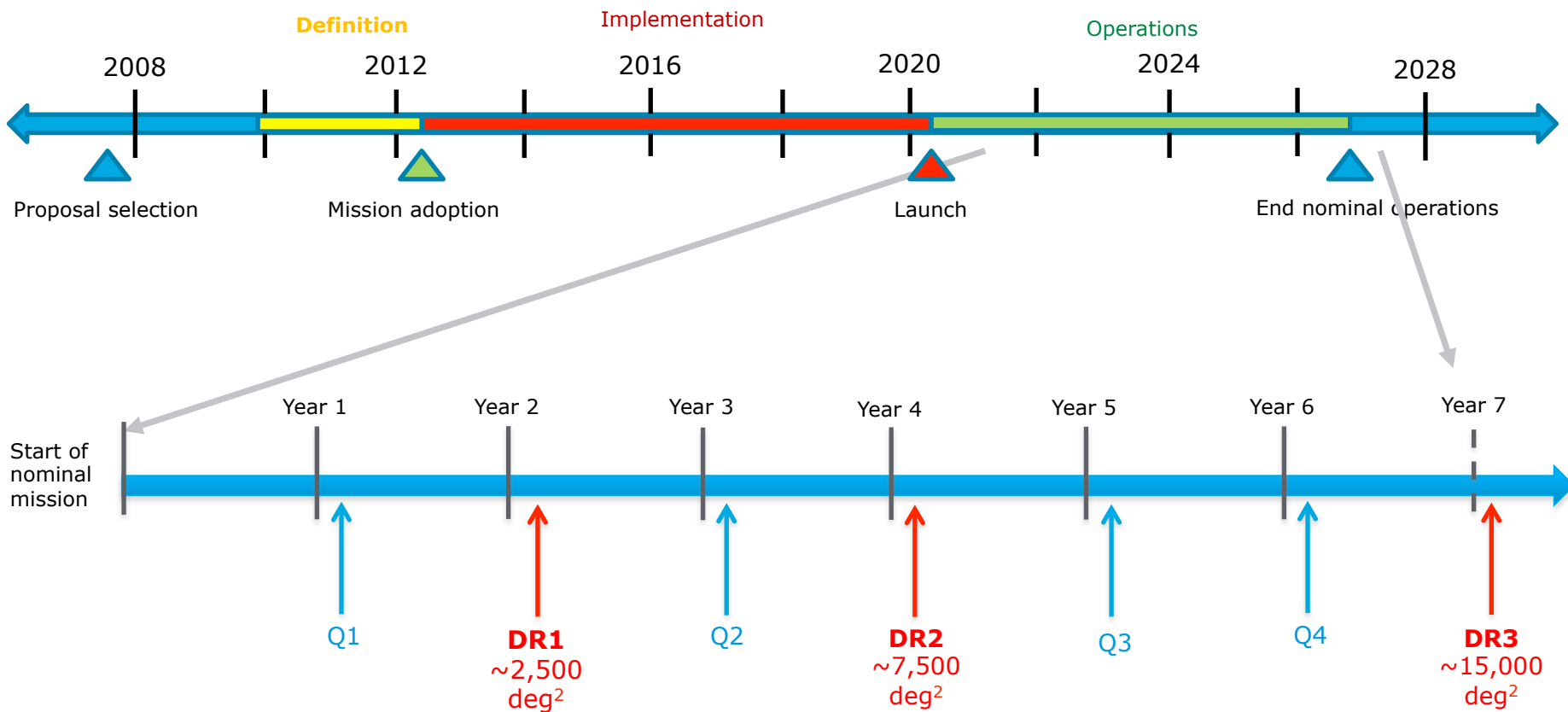
Courtesy: T. Maciaszek and the NISP team



The Ground Segment at a glance



Euclid Schedule

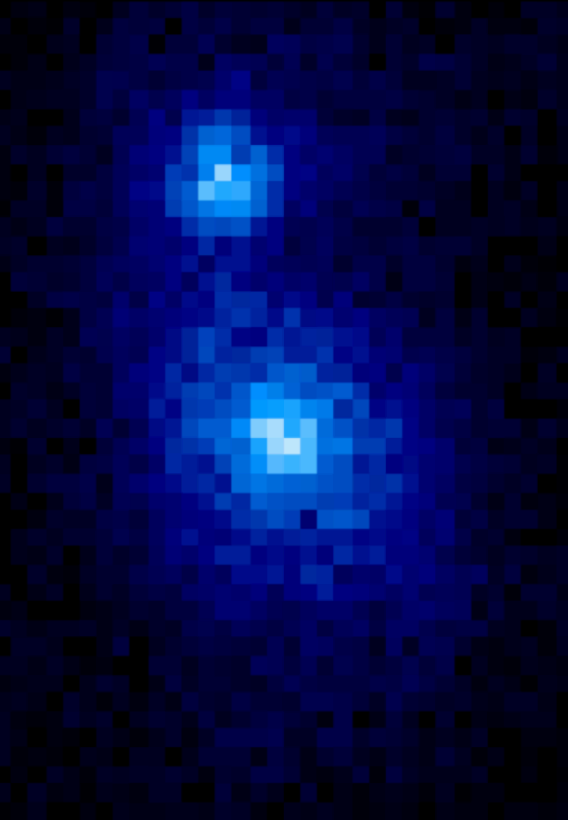


Euclid Imaging

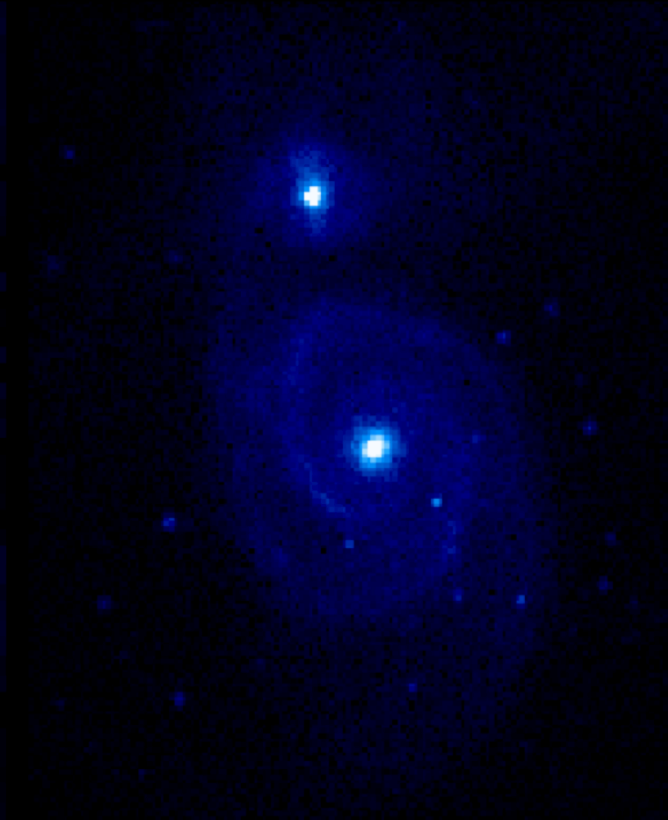


Courtesy J. Brinchmann,
Steve Warren

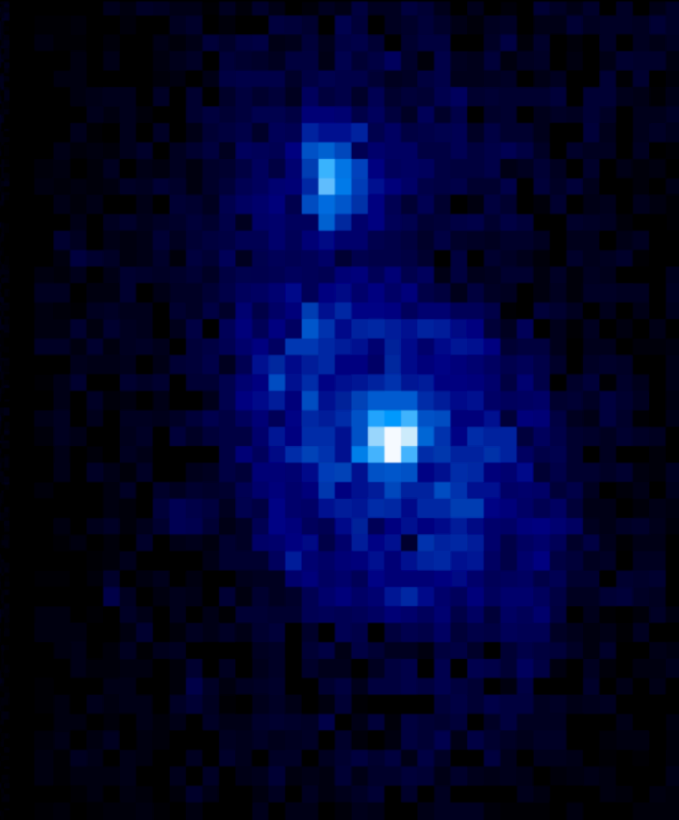
M51



SDSS @ $z=0.1$



Euclid @ $z=0.1$



Euclid @ $z=0.7$

Télescope Canada France Hawaii



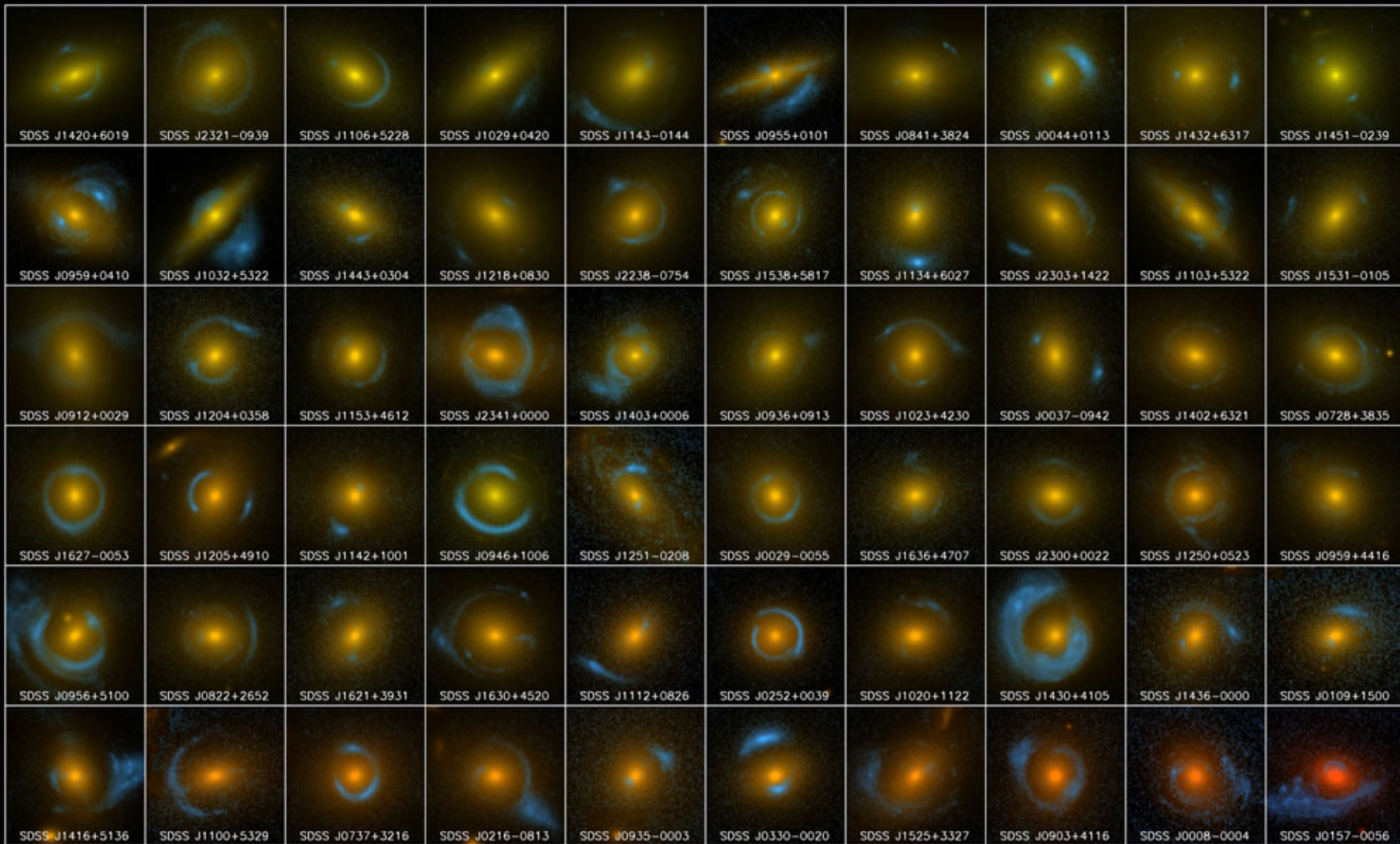
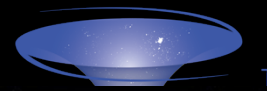
October, 16 2015

Euclid



L. Valenziano on behalf of the EC

SLACS (~2010 - HST)



SLACS: The Sloan Lens ACS Survey

www.SLACS.org

A. Bolton (U. Hawai'i IfA), L. Koopmans (Kapteyn), T. Treu (UCSB), R. Gavazzi (IAP Paris), L. Moustakas (JPL/Caltech), S. Burles (MIT)

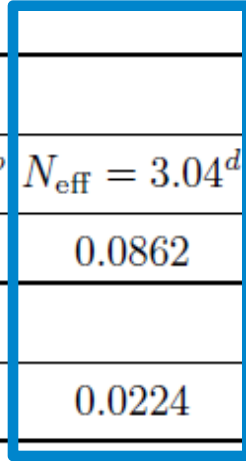
October, 16 2015

L. Valenziano on behalf of the SLACS team and NASA/ESA

SLACS

Euclid VIS Legacy : after 2 months
(66 months planned)

Neutrinos and Relativistic Species



Amendola et al 2013	General cosmology					$N_{\text{eff}} = 3.04^d$
	$\Sigma = 0.3 \text{ eV}^a$	$\Sigma = 0.2 \text{ eV}^a$	$\Sigma = 0.125 \text{ eV}^b$	$\Sigma = 0.125 \text{ eV}^c$	$\Sigma = 0.05 \text{ eV}^b$	
fiducial \rightarrow						
EUCLID+Planck	0.0361	0.0458	0.0322	0.0466	0.0563	0.0862
Λ CDM cosmology						
EUCLID+Planck	0.0176	0.0198	0.0173	0.0218	0.0217	0.0224

^a for degenerate spectrum: $m_1 \approx m_2 \approx m_3$; ^b for normal hierarchy: $m_3 \neq 0, m_1 \approx m_2 \approx 0$

^c for inverted hierarchy: $m_1 \approx m_2, m_3 \approx 0$; ^d fiducial cosmology with massless neutrinos

Changes in neutrino mass and species make tiny changes in $P(k)$ and in the Alcock-Paczynski test.

- If $\Sigma > 0.1 \text{ eV}$ then Euclid spectroscopic survey will be able to determine the neutrino mass scale independently of the model cosmology assumed.
- If $\Sigma < 0.1 \text{ eV}$, the sum of neutrino masses, and in particular the minimum neutrino mass required by neutrino oscillations, can be measured in the context of the Λ -CDM

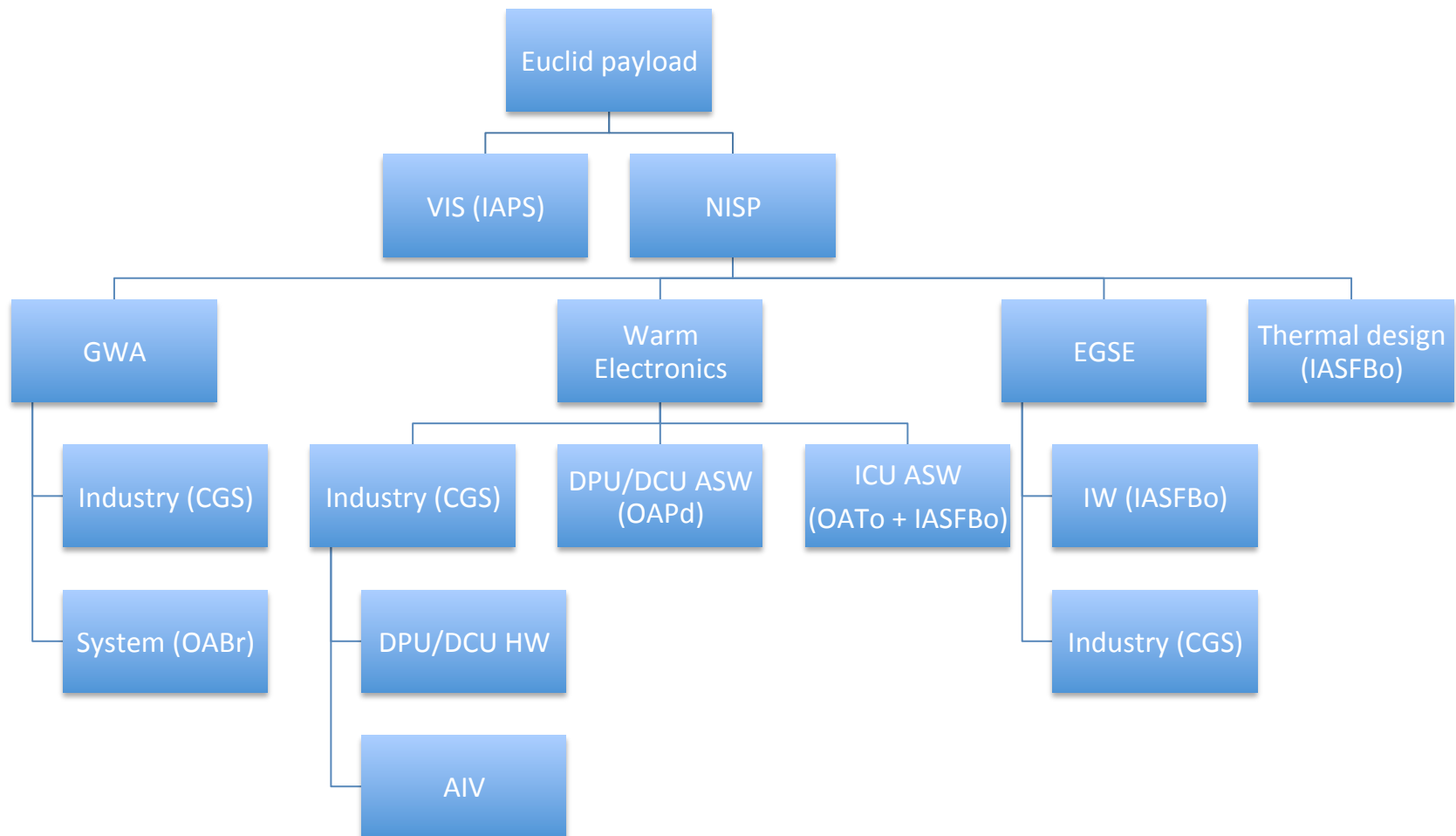
Forecast for the Primary Program



	Modified Gravity	Dark Matter	Initial Conditions	Dark Energy		
Parameter	γ	m_ν / eV	f_{NL}	w_p	w_a	FoM <small>= $1/(\Delta w_0 \times \Delta w_a)$</small>
Euclid primary (WL+GC)	0.010	0.027	5.5	0.015	0.150	430
EuclidAll (clusters, ISW)	0.009	0.020	2.0	0.013	0.048	1540
Euclid+Planck	0.007	0.019	2.0	0.007	0.035	4020 → 6000
Reference value (2009)	0.200	0.580	100	0.100	1.500	~10
Improvement Factor	30	30	50	>10	>40	>400

Assume systematic errors are under control

Italian Contribution to Euclid instruments



INFN@Euclid in 2016



BOLOGNA 5.4 FTE, 13 persone		PADOVA 5.9 FTE, 16 persone	
L. Patrizzii (INFN) Resp. Nazionale		S.Dusini (INFN) Resp. Locale Coord. AIV WP Coord. attività DPU	
G. Sirri (INFN) Resp.Locale Coord. attività ICU		F. Dal Corso (INFN)	
T. Chiarusi (INFN)		M. Roda (Dottorando)	
F. Giacomini (CNAF)		C.Sirignano (UniPd)	
F. Fornari (dottorando)		L.Stanco (INFN)	
A. Margiotta (PA UniBO)		S.Ventura (INFN)	
N. Mauri (AR INFN)		N.Bartolo (UniPd teorico)	
L. Pasqualini (Dottoranda)		M.Liguori (UniPd teorico)	
M. Spurio (UniBo)		S.Matarrese (UniPd teorico)	
M. Tenti AR CNAF		D.Karagiannis (UniPd teorico)	
F. Tronconi (INFN, grIV)		G. Naletto (CISAS)	
F. Finelli (INAF)		C. Bonoli (INAF)	
L. Valenziano (INAF)		F.Bortoletto (INAF)	
		A. Balestra (INAF)	
		Dottorando da Ottobre	
		RTD su fondi ASI	



Contributo all'application software della ICU:

sviluppo del modulo SW di comunicazione con le DPU attraverso il bus di comunicazione MIL1553

- inoltro di telecomandi
- richiesta telemetrie
- gestione bus 1553 (BC = bus controller)

Processore: LEON2

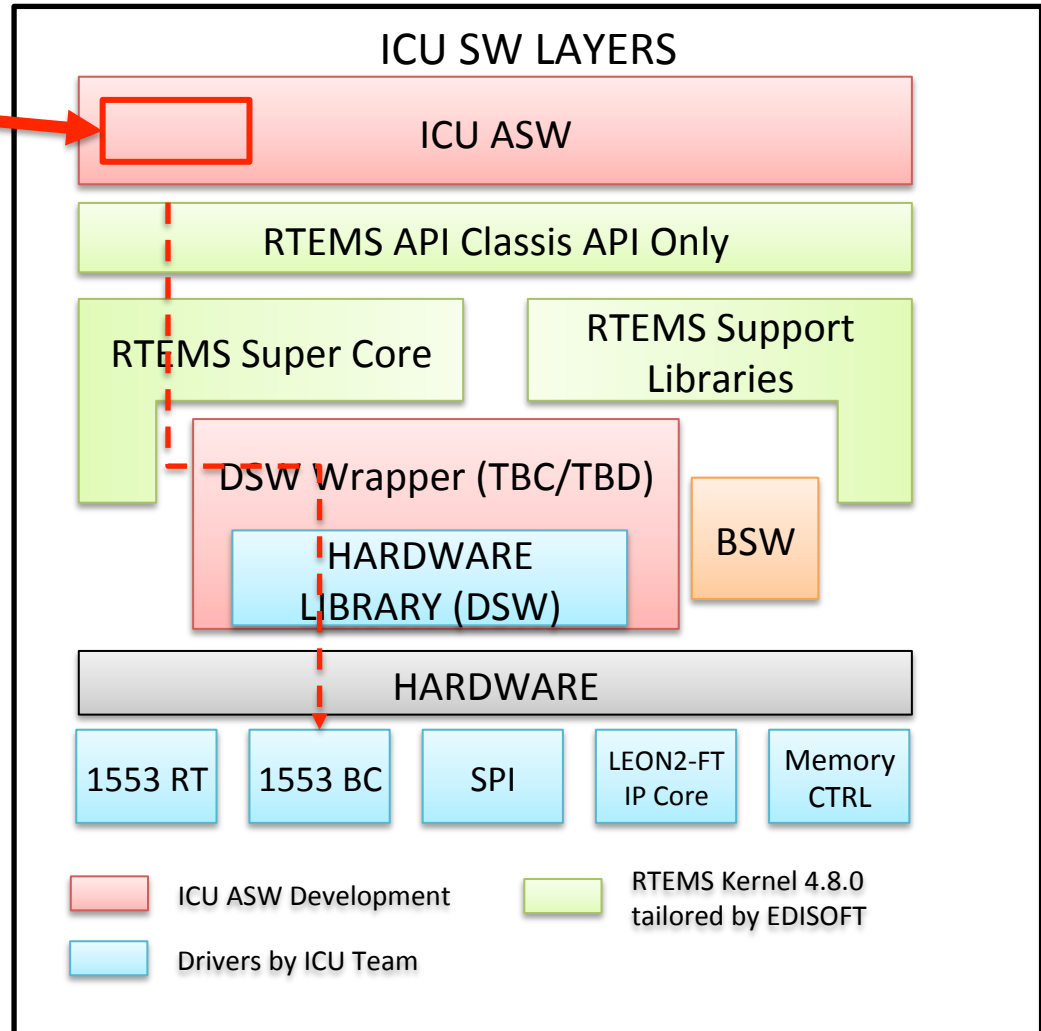
OS : RTEMS (versione spazio)

Linguaggio : C

Supporto allo sviluppo:

- repository
- Unit test automatico
- Sviluppo procedure di test

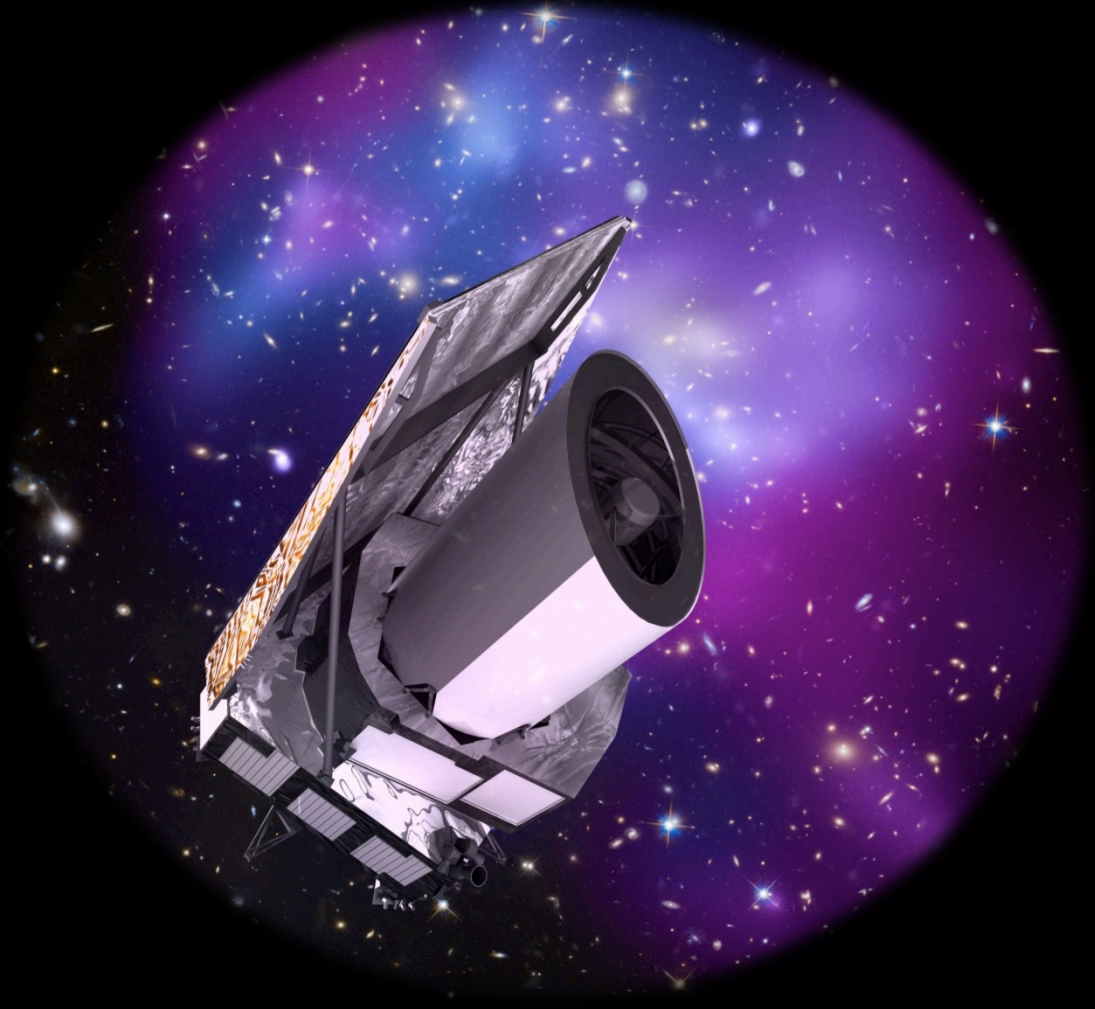
[con l'aiuto del CNAF]



Summary



1. Euclid is an experiment combining GC and WL: an unprecedented match of an imaging and redshift survey from space, building a sample of $>10^9$ galaxy shapes and $\sim 5 \cdot 10^7$ galaxy distances (and much more).
2. Euclid results may well revolutionize our understanding of physics: for sure it will provide a huge database for unexpected discoveries (legacy).
3. Ideal complementarity to CMB observations.
4. Euclid is one of the most sophisticated scientific instruments ever launched: large cryo optics, large focal planes, the most powerful on-board data processing.
5. Italy: one of the major contributors, leading the SGS and providing critical elements to the payload.
The Italian participation is funded by ASI.



October, 16 2015

L. Valenziano on behalf of the EC

Acknowledgements



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