



TFI 2015

Theories of the Fundamental Interactions

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Born-Infeld/Gravity Duality

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Based on:

G. Grignani, T. Harmark, AM and M. Orselli – arXiv:1512.xxxx [hep-th]

Outline

- 1 Overview
- 2 Born-Infeld side
- 3 Gravity side
- 4 Conclusions

- New duality between 4d Born-Infeld (BI) theory and 5d classical gravity
- manifestation of the open/closed string duality
- Born-Infeld side \rightarrow low energy effective theory for open strings ending on a D3-brane in a slowly varying background Kalb-Ramond field
- gravity side \rightarrow gravitational (closed string) description of D3-branes in the same background Kalb-Ramond field
- the duality is a correspondence between **effective theories** in a similar sense as in the fluid/gravity correspondence
- it can be used to derive higher derivative correction to the BI action

- N coincident D3-branes in 10d Minkowski background

Mink ¹⁰	x_0	x_1	x_2	x_3	x_4	x_5	x_6	x_7	x_8	x_9
D3	×	×	×	×	—	—	—	—	—	—

- flat embedding $\rightarrow x_a(\sigma) = \sigma_a$, for $a = 0, 1, 2, 3$, and $x_i(\sigma) = 0$ for $i = 4, 5, \dots, 9$
- IIB SUGRA background has $\phi = 0$ and $H_{abc} = 0$
- turn on B_{ab} || D3 world-volume (while $B_{ij} = 0$)
- $\mathcal{F}_{ab} = B_{ab} \leftrightarrow$ gauge invariant field strength on the D3
- IIB SUGRA EOMs $\rightarrow dB_{(2)} = 0 \rightarrow \partial_{[a}\mathcal{F}_{bc]} = 0$

Setup

- Assume \mathcal{F}_{ab} slowly varying over the D3

$$\frac{R}{\ell_s} \gg (g_s N)^{1/4} \quad \text{and} \quad \frac{R}{\ell_s} \gg 1$$

R = minimal length scale of variation of \mathcal{F}_{ab}

- when $g_s N \ll 1 \rightarrow$ open string description \rightarrow Born-Infeld action
- when $g_s N \gg 1 \rightarrow$ closed string description \rightarrow type IIB SUGRA

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Born-Infeld theory

BI action: low energy effective action for open strings ending on a single D3 valid when $g_s \ll 1$ (perturbative string theory)

$$S_{\text{BI}} = -T_{\text{D3}} \int d^4\sigma \sqrt{-\det(\eta_{ab} + \mathcal{F}_{ab})}$$

$$\mathcal{F}_{ab} = B_{ab} + 2\pi\ell_s^2 F_{ab} \quad \text{and} \quad T_{\text{D3}} = (g_s \ell_s^4 (2\pi))^{-1}$$

$$\partial_{[a} \mathcal{F}_{bc]} = 0 \quad \leftrightarrow \quad d\mathcal{F} = 0$$

■ useful notation:

$$\mathcal{F}_a{}^b = \mathcal{F}_{ac} \eta^{cb} \quad G_a{}^b = \delta_a^b - \mathcal{F}_a{}^c \mathcal{F}_c{}^b \quad m^4 = \det(G_a{}^b)$$

$$\text{BI action} \quad \rightarrow \quad S_{\text{BI}} = -T_{\text{D3}} \int d^4\sigma m^2$$

- EOMs by varying the action

$$\partial^b(mG^{-1}\mathcal{F})_{ba} = 0$$

- Born-Infeld energy-momentum tensor

$$\tau_{ab} = -T_{D3}m(G^{-1})_{ab}$$

- energy-momentum tensor conservation

$$\partial^b\tau_{ba} = 0 \quad \longrightarrow \quad \partial^b(m(G^{-1})_{ba}) = 0$$

→ these eq.s along with $d\mathcal{F} = 0$ → constraints on how \mathcal{F}_{ab} can vary along the D3

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- SUGRA description of N D3-branes in the background of a slowly varying Kalb-Ramond potential

Type IIB Supergravity Lagrangian (bosonic part)

$$\mathcal{L} = \sqrt{-g} \left(\mathcal{R}^{(10)} - \frac{1}{2} \partial_\mu \phi \partial^\mu \phi - \frac{1}{12} e^{-\phi} H_{\mu\nu\rho} H^{\mu\nu\rho} - \frac{1}{2} e^{2\phi} \partial_\mu \chi \partial^\mu \chi - \frac{1}{12} e^\phi F_{\mu\nu\rho} F^{\mu\nu\rho} - \frac{1}{4 \cdot 5!} F_{\mu\nu\rho\lambda\sigma} F^{\mu\nu\rho\lambda\sigma} \right) + \frac{1}{8 \cdot 4!} \epsilon^{\mu_1 \dots \mu_{10}} A_{\mu_1 \dots \mu_4} \partial_{\mu_5} B_{\mu_6 \mu_7} \partial_{\mu_8} A_{\mu_9 \mu_{10}}$$

boundary conditions

$$\lim_{r \rightarrow 0} B_{ab} = 0 \quad \lim_{r \rightarrow \infty} B_{ab} = \mathcal{F}_{ab} \quad \text{with } d\mathcal{F} = 0$$

$$\lim_{r \rightarrow \infty} g_{\mu\nu} = \eta_{\mu\nu} \text{ (Mink}^{10}\text{)} \quad \lim_{r \rightarrow \infty} \phi, \chi, A^{(2)} = 0$$

r = radial coordinate in the transverse space of the D3 (D3 are at $r = 0$)

Gravity setup

- using symmetries and gauge freedom \rightarrow ten-dimensional metric can be written as

$$g_{\mu\nu} dx^\mu dx^\nu = e^{\frac{1}{2}\eta} (h_{mn} dx^m dx^n + e^{2\rho} d\Omega_5^2) \quad m, n = 0, \dots, 4$$

with $\rho = \log r$ and h_{mn} and η only depends on x^m

- impose that we have N D3-branes \rightarrow set the RR 5-form field strength to

$$F^{(5)} = 4r_c^4 (\omega_5 + *\omega_5) \quad r_c^4 = \frac{N}{2\pi^2 T_{D3}}$$

r_c sets the length scale of the thickness of the D3

- only the metric and $F^{(5)}$ involve the S^5 directions \rightarrow 5d reduction

$$\frac{\partial_m \left(\sqrt{-h} e^{2\phi+2\eta+5\rho} \partial^m \chi \right)}{e^{5\rho} \sqrt{-h}} = -\frac{1}{6} e^{\phi+\eta} F_{mnp} H^{mnp}$$

$$\frac{\partial_m \left(\sqrt{-h} e^{2\eta+5\rho} \partial^m \phi \right)}{e^{5\rho} \sqrt{-h}} = -\frac{1}{12} e^{\eta-\phi} H_{mnp} H^{mnp} + \frac{1}{12} e^{\eta+\phi} F_{mnp} F^{mnp} + e^{2\eta+2\phi} \partial_m \chi \partial^m \chi$$

$$\partial_m \left(\sqrt{-h} e^{\eta+5\rho} (e^{-\phi} H^{mnp} - \chi e^{\phi} F^{mnp}) \right) = -\frac{4r_c^4}{6} \epsilon^{npq} (F_{lpq} + \chi H_{lpq})$$

$$\partial_m \left(\sqrt{-h} e^{\eta+5\rho} e^{\phi} F^{mnp} \right) = \frac{4r_c^4}{6} \epsilon^{npq} H_{lpq}$$

$$4e^{-2\rho} - 5\partial^m \rho \partial_m \rho - \frac{1}{2} \partial^m \eta \partial_m \eta - \frac{13}{4} \partial^m \eta \partial_m \rho - D^m D_m \rho - \frac{1}{4} D^m D_m \eta$$

$$= -\frac{1}{48} e^{-\phi-\eta} H_{mnp} H^{mnp} - \frac{1}{48} e^{\phi-\eta} F_{mnp} F^{mnp} + 4r_c^8 e^{-2\eta-10\rho}$$

$$(\mathcal{R}^{(5)})_{mn} = 5\partial_m \rho \partial_n \rho + 5D_m D_n \rho - \frac{1}{2} \partial_m \eta \partial_n \eta + 2D_m D_n \eta + \frac{1}{2} h_{mn} \partial^k \eta \partial_k \eta + \frac{1}{4} h_{mn} D^k D_k \eta$$

$$+ \frac{5}{4} h_{mn} \partial^k \eta \partial_k \rho + \frac{1}{2} \partial_m \phi \partial_n \phi + \frac{1}{2} e^{2\phi} \partial_m \chi \partial_n \chi + \frac{1}{12} e^{-\phi-\eta} (3H_m{}^{kl} H_{nkl}$$

$$- \frac{1}{4} h_{mn} H^{klp} H_{klp}) + \frac{1}{12} e^{\phi-\eta} (3F_m{}^{kl} F_{nkl} - \frac{1}{4} h_{mn} F^{klp} F_{klp}) - 4r_c^8 e^{-2\eta-10\rho} h_{mn}$$

D3 with constant \mathcal{F}_{ab}

Known solutions

■ D3 solution $\rightarrow \mathcal{F}_{ab} = 0$

■ F1-D3 solution $\rightarrow \vec{E} \neq 0, \vec{B} = 0$

■ D1-D3 solution $\rightarrow \vec{B} \neq 0, \vec{E} = 0$

■ (F1||D1)-D3 solution $\rightarrow \vec{E} \times \vec{B} = 0$

■ (F1⊥D1)-D3 solution $\rightarrow \vec{E} \cdot \vec{B} = 0$

$$\mathcal{F}_{ab} = \begin{pmatrix} 0 & E_1 & E_2 & E_3 \\ -E_1 & 0 & B_3 & -B_2 \\ -E_2 & -B_3 & 0 & B_1 \\ -E_3 & B_2 & -B_1 & 0 \end{pmatrix}$$

$$\vec{E} = (E_1, E_2, E_3)$$

$$\vec{B} = (B_1, B_2, B_3)$$

Grignani, Harmark, AM, Orselli arXiv:1311.3834

D3 with constant \mathcal{F}_{ab}

New (general) solution

→ generalization of the previous solutions

$$h_{mn}dx^m dx^n = h_{ab}dx^a dx^b + dr^2 \quad (a, b = 0, \dots, 3) \quad \rho = \log r$$

$$h_{ab} = \{(I + mr_c^4 r^{-4} G^{-1})^{-1}\}_a^c \eta_{cb} \quad \sqrt{-h} = e^{-2\eta}$$

$$e^{2\eta} = 1 + \frac{\text{Tr}(G)}{2m} \frac{r_c^4}{r^4} + \frac{r_c^8}{r^8} \quad e^{2\phi} = e^{-2\eta} \left(1 + m \frac{r_c^4}{r^4}\right)^2 \quad \chi = \frac{\text{Tr}(\mathcal{F}^* \mathcal{F})}{4} \frac{\frac{r_c^4}{r^4}}{1 + m \frac{r_c^4}{r^4}}$$

$$B_{ab} = \mathcal{F}_a^c h_{cb} \quad A_{ab} = \frac{r_c^4}{r^4} {}^* \mathcal{F}_a^c h_{cb}$$

- low energy description of D3 dynamics → $r_c \gg \ell_s \rightarrow g_s N \gg 1$
- low energy effective description of the **strongly coupled** Born-Infeld theory

Derivative expansion

- consider the D3 solution with general \mathcal{F}_{ab}
- let \mathcal{F}_{ab} be a **slowly varying** function of the world-volume coordinates x_a
↔ analogy with fluid/gravity correspondence
- set up a perturbative procedure to build a corrected solution →
world-volume derivative expansion
- zeroth order → D3 with constant \mathcal{F}_{ab}

- neglect in the EOMs all the terms with more than one world-volume derivative
- equations with one world-volume derivative correction are

$$\partial_b (e^{-\eta-\phi} H^{bar} - e^{-\eta+\phi} \chi F^{bar}) = \frac{4r_c^4}{6r^5} \epsilon^{abcd} (F_{bcd} + \chi H_{bcd}) \quad (1)$$

$$\partial_b (e^{-\eta+\phi} F^{bar}) = -\frac{4r_c^4}{6r^5} \epsilon^{abcd} H_{bcd} \quad (2)$$

$$\begin{aligned} \frac{1}{2} \partial_b (h^{bc} \partial_r h_{ca}) - \frac{1}{2} \Gamma_{ab}^c h^{bd} \partial_r h_{dc} &= -\frac{1}{2} \partial_a \eta \partial_r \eta + \frac{1}{2} \partial_a \phi \partial_r \phi + \frac{1}{2} e^{2\phi} \partial_a \chi \partial_r \chi \\ &+ \frac{1}{4} e^{-\phi-\eta} H_a{}^{bc} H_{bcr} + \frac{1}{4} e^{\phi-\eta} F_a{}^{bc} F_{bcr} \end{aligned} \quad (3)$$

First order

Consider the above equations to leading order at large r

■ Eq. (1) $\rightarrow \partial_b H^{bar} = 0 \rightarrow \partial^b (mG^{-1}\mathcal{F})_{ba} = 0$

■ Eq. (2) $\rightarrow \partial_b F^{bar} = 0 \rightarrow \partial^b (*\mathcal{F})_{ba} = 0 \leftrightarrow d\mathcal{F} = 0$

■ Eq. (3) $\rightarrow \partial^b \partial_r h_{ba} = 0 \rightarrow \partial^b (mG^{-1})_{ba} = 0$

these are the constraints on the variation of \mathcal{F}_{ab}

\rightarrow Exactly the same eqs we got from the BI!

Higher-derivative correction

- starting from second order \rightarrow corrections to the zeroth order solution
- general corrected solution

$$Y_I(x^a, r) = Y_I^{(0)}(x^a, r) + Y_I^{(2)}(x^a, r) + Y_I^{(4)}(x^a, r) + \mathcal{O}(\partial^6)$$

$$Y_I = (\chi, \phi, A_{ab}, B_{ab}, \eta, h_{ab})$$

- no corrections with an odd number of world-volume derivatives \leftrightarrow no covariant expression in terms of \mathcal{F}_{ab}

Second order

- Eq.s that determine $Y_I^{(2)}$ involve two world-volume derivatives

$$H_{(0)}^{m,I} Y_I^{(2)} + H_{(1)}^{m,I} \partial_r Y_I^{(2)} + H_{(2)}^{m,I} \partial_r^2 Y_I^{(2)} = K^{m,ab,I} \partial_a \partial_b Y_I^{(0)} + L^{m,ab,IJ} \partial_a Y_I^{(0)} \partial_b Y_J^{(0)}$$

- solution in a large r expansion has the form

$$Y_I^{(2)} = \sum_{n=0}^{\infty} \frac{C_{I,n}^{(2)}}{r^{2+4n}}$$

- to leading order at large r

$$\begin{aligned} \chi^{(2)} &= \frac{r_c^4}{16r^2} \partial_a \partial^a (\mathcal{F}^* \mathcal{F}) + \mathcal{O}(r^{-6}) & \phi^{(2)} &= \frac{r_c^4}{4r^2} \partial_a \partial^a \left(m - \frac{\text{Tr}G}{4m} \right) + \mathcal{O}(r^{-6}) \\ B_{ab}^{(2)} &= -\frac{r_c^4}{4r^2} 3\partial^c \partial_{[c} (mG^{-1} \mathcal{F})_{ab]} + \mathcal{O}(r^{-6}) & A_{ab}^{(2)} &= \frac{r_c^4}{4r^2} 3\partial^c \partial_{[c} (\mathcal{F}^* \mathcal{F})_{ab]} + \mathcal{O}(r^{-6}) \\ \eta^{(2)} &= \frac{r_c^4}{16r^2} \partial^c \partial_c \left(\frac{\text{Tr}G}{m} \right) + \mathcal{O}(r^{-6}) & h_{ab}^{(2)} &= -\frac{r_c^4}{4r^2} \partial^c \partial_c (mG^{-1})_{ab} + \mathcal{O}(r^{-6}) \end{aligned}$$

Third order

Third order eq.s to leading order at large r

■ Eq. (1) $\rightarrow \partial_b H^{bar} = 0 \rightarrow \partial_b \partial^c \partial_{[c} (mG^{-1} \mathcal{F})_{ab]} = 0$

■ Eq. (2) $\rightarrow \partial_b F^{bar} = 0 \rightarrow \partial_b \partial^c \partial_{[c} {}^* \mathcal{F}_{ab]} = 0$

■ Eq. (3) $\rightarrow \partial^b \partial_r h_{ba} = 0 \rightarrow \partial^b \partial^c \partial_c (mG^{-1})_{ba} = 0$

\rightarrow these are trivially satisfied assuming the first order constraints

- fourth order corrections at large r

$$Y_I^{(4)} = C_{I,0}^{(4)} \log r + \frac{C_{I,4}^{(4)}}{r^4} + \mathcal{O}\left(\frac{1}{r^8}\right)$$

- one can easily compute $C_{I,0}^{(4)}$

$$\begin{aligned} C_{\chi,0}^{(4)} &= -\frac{r_c^4}{64} \partial^2 \partial^2 (\mathcal{F}^* \mathcal{F}) & C_{\phi,0}^{(4)} &= -\frac{r_c^4}{16} \partial^2 \partial^2 \left(m - \frac{\text{Tr}G}{4m} \right) \\ C_{B_{ab},0}^{(4)} &= \frac{r_c^4}{16} 3 \partial^2 \partial^c \partial_{[c} (m G^{-1} \mathcal{F})_{ab]} & C_{A_{ab},0}^{(4)} &= -\frac{r_c^4}{16} 3 \partial^2 \partial^c \partial_{[c} {}^* \mathcal{F}_{ab]} \\ C_{\eta,0}^{(4)} &= -\frac{r_c^4}{64} \partial^2 \partial^2 \left(\frac{\text{Tr}G}{m} \right) & C_{h_{ab},0}^{(4)} &= \frac{r_c^4}{16} \partial^2 \partial^2 (m G^{-1})_{ab} \end{aligned}$$

where $\partial^2 = \partial_a \partial^a$

- the next-to-leading coefficient $C_{I,4}^{(4)}$ cannot be fixed

Comparison with the BI

- the constraints we got at first and third order are consistent with the BI ones
- what about the corrections to the fields at second and fourth order?
- corrections to the metric \rightarrow corrections to EM tensor
- compare this corrected EM tensor with BI EM tensor \rightarrow BI higher derivative corrections

Reading off the energy-momentum tensor

- EM tensor can be read off from the metric in the asymptotic region
- in our approximation the asymptotic region is $r_c \ll r \ll R$
- linearize the metric

$$h_{ab} = \eta_{ab} + \bar{h}_{ab} \quad \eta = \bar{\eta} \quad \text{with } |\bar{h}_{ab}| \ll 1, |\bar{\eta}| \ll 1$$

- make the gauge choice $\rightarrow \eta^{ab} \bar{h}_{ab} = -4\bar{\eta}$ and $\partial^b \bar{h}_{ba} = 0$
- imagine an infinitely thin 3-brane sitting at $r = 0$ with EM tensor $\tau_{ab}(x^c)$ on the brane
- linearized Einstein equations in the region $r_c \ll r \ll R$

$$\left[\partial_r^2 + \frac{5}{r} \partial_r + \partial^c \partial_c \right] \bar{h}_{ab} = -16\pi G \delta^6(r) \tau_{ab}$$

Linearized Einstein eq.s

- assume τ_{ab} varies slowly along the world-volume
- expand τ_{ab} in higher-derivative corrections

$$\tau_{ab} = \tau_{ab}^{(0)} + \tau_{ab}^{(2)} + \tau_{ab}^{(4)} + \mathcal{O}(\partial^6)$$

- expand \bar{h}_{ab} in higher-derivative contributions

$$\bar{h}_{ab} = \bar{h}_{ab}^{(0)} + \bar{h}_{ab}^{(2)} + \bar{h}_{ab}^{(4)} + \mathcal{O}(\partial^6)$$

- linearized Einstein eq.s become

$$\begin{aligned} \left[\partial_r^2 + \frac{5}{r} \partial_r \right] h_{ab}^{(0)} &= -16\pi G \delta^6(r) \tau_{ab}^{(0)} \\ \left[\partial_r^2 + \frac{5}{r} \partial_r \right] h_{ab}^{(2)} + \partial^c \partial_c h_{ab}^{(0)} &= -16\pi G \delta^6(r) \tau_{ab}^{(2)} \\ \left[\partial_r^2 + \frac{5}{r} \partial_r \right] h_{ab}^{(4)} + \partial^c \partial_c h_{ab}^{(2)} &= -16\pi G \delta^6(r) \tau_{ab}^{(4)} \end{aligned}$$

- Solution of the linearized Einstein eq.s

$$\bar{h}_{ab}^{(0)} = \frac{r_c^4}{r^4} \frac{\tau_{ab}^{(0)}}{NT_{D3}}$$

$$\bar{h}_{ab}^{(2)} = \frac{r_c^4}{r^4} \frac{\tau_{ab}^{(2)}}{NT_{D3}} + \frac{r_c^4}{4r^2} \frac{\partial^c \partial_c \tau_{ab}^{(0)}}{NT_{D3}}$$

$$\bar{h}_{ab}^{(4)} = \frac{r_c^4}{r^4} \frac{\tau_{ab}^{(4)}}{NT_{D3}} + \frac{r_c^4}{4r^2} \frac{\partial^c \partial_c \tau_{ab}^{(2)}}{NT_{D3}} - \frac{r_c^4}{16} \log r \frac{\partial^c \partial_c \partial^d \partial_d \tau_{ab}^{(0)}}{NT_{D3}}$$

Zeroth order EM tensor

- solution of linearized Einstein $\rightarrow \bar{h}_{ab}^{(0)} = \frac{r_c^4}{r^4} \frac{\tau_{ab}^{(0)}}{NT_{D3}}$

- leading correction to flat space metric at zeroth order

$$\bar{h}_{ab}^{(0)} = -\frac{r_c^4}{r^4} m (G^{-1})_{ab}$$

- plug $\bar{h}_{ab}^{(0)}$ into above solution of linearized Einstein eq.

$$\rightarrow \tau_{ab}^{(0)} = -NT_{D3} m (G^{-1})_{ab}$$

- this is the leading order EM-tensor for N D3-branes = $N \times$ Born-Infeld EM-tensor

Second order EM tensor

- solution of linearized Einstein $\rightarrow \bar{h}_{ab}^{(2)} = \frac{r_c^4}{r^4} \frac{\tau_{ab}^{(2)}}{NT_{D3}} + \frac{r_c^4}{4r^2} \frac{\partial^c \partial_c \tau_{ab}^{(0)}}{NT_{D3}}$
- zeroth order EM tensor $\rightarrow \tau_{ab}^{(0)} = -NT_{D3} m(G^{-1})_{ab}$
- second order large r correction to the metric

$$\bar{h}_{ab}^{(2)} = -\frac{r_c^4}{4r^2} \partial^c \partial_c (mG^{-1})_{ab} + \mathcal{O}\left(\frac{1}{r^6}\right)$$

- plug $\tau_{ab}^{(0)}$ and $\bar{h}_{ab}^{(2)}$ into above sol. of linearized Einstein eq.

$$-\frac{r_c^4}{4r^2} \partial^c \partial_c (mG^{-1})_{ab} + \mathcal{O}\left(\frac{1}{r^6}\right) = \frac{r_c^4}{r^4} \frac{\tau_{ab}^{(2)}}{NT_{D3}} - \frac{r_c^4}{4r^2} \partial^c \partial_c (mG^{-1})_{ab}$$

- second order EM tensor $\rightarrow \tau_{ab}^{(2)} = 0$
- Born-Infeld is not corrected at second order

Andreev, Tseytlin (1988)

Fourth order EM tensor?

- solution of linearized Einstein eq.

$$\bar{h}_{ab}^{(4)} = \frac{r_c^4}{r^4} \frac{\tau_{ab}^{(4)}}{NT_{D3}} + \frac{r_c^4}{4r^2} \frac{\partial^c \partial_c \tau_{ab}^{(2)}}{NT_{D3}} - \frac{r_c^4}{16} \log r \frac{\partial^c \partial_c \partial^d \partial_d \tau_{ab}^{(0)}}{NT_{D3}}$$

- zeroth and second order EM tensors

$$\tau_{ab}^{(0)} = -NT_{D3} m (G^{-1})_{ab} \quad \tau_{ab}^{(2)} = 0$$

- fourth order large r correction to the metric

$$\bar{h}_{ab}^{(4)} = \frac{r_c^4}{16} \log r \partial^c \partial_c \partial^d \partial_d (mG^{-1})_{ab} + \mathcal{O}\left(\frac{1}{r^4}\right)$$

- plug $\tau_{ab}^{(0)}$, $\tau_{ab}^{(2)}$ and $\bar{h}_{ab}^{(2)}$ into above sol. of linearized Einstein eq.

~~$$\frac{r_c^4}{16} \log r \partial^c \partial_c \partial^d \partial_d (mG^{-1})_{ab} + \mathcal{O}\left(\frac{1}{r^4}\right) = \frac{r_c^4}{r^4} \frac{\tau_{ab}^{(4)}}{NT_{D3}} + \frac{r_c^4}{16} \log r \partial^c \partial_c \partial^d \partial_d (mG^{-1})_{ab}$$~~

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Conclusions

- explored a new duality between BI theory and 5d gravity
- open/closed string duality for D3-branes in a slowly varying background Kalb-Ramond field \mathcal{F}_{ab}
- world-volume derivative expansion to build the gravity solution for D3 with a slowly varying $\mathcal{F}_{ab} \rightarrow$ dual to BI theory (including derivative corrections)
- proved that the BI action is not corrected up to the second order in the derivative expansion
- what about the fourth order?

Extra slides

Duality chain generating constant \mathcal{F}_{ab} D3 solution

- start with the D3 solution

$$h_{ab} = \frac{\eta_{ab}}{1 + \frac{r_c^4}{r^4}} \quad e^\eta = 1 + \frac{r_c^4}{r^4} \quad \phi = \chi = 0 \quad A_{ab} = B_{ab} = 0$$

↓ T-duality along x_1

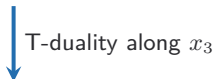
- D2-branes smeared along x_1

↓ rotation in the 12-plane + T-duality along x_1

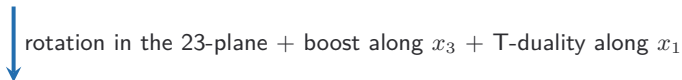
- D1-D3 bound state with D-strings along x_3

Duality chain... cont'd

- D1-D3 bound state with D-strings along x_3



- D0-D2 bound state smeared along x_3



- F1-D1-D3 bound state with F-strings along x_3 and D-strings in the 13-plane

Second order solution

- Eqs that determine $Y_I^{(2)}$

$$H_{(0)}^{m,I} Y_I^{(2)} + H_{(1)}^{m,I} \partial_r Y_I^{(2)} + H_{(2)}^{m,I} \partial_r^2 Y_I^{(2)} = K^{m,ab,I} \partial_a \partial_b Y_I^{(0)} + L^{m,ab,IJ} \partial_a Y_I^{(0)} \partial_b Y_I^{(0)}$$

- particular solution $\rightarrow Y_I^{(2)} = \sum_{n=0}^{\infty} \frac{C_{I,n}^{(2)}}{r^{2+4n}}$

- solution of the homogeneous part $\rightarrow \delta Y_I = \sum_{n=1}^{\infty} \frac{C_{I,n}^{(2)}}{r^{4n}}$

- general second order solution $\rightarrow Y_I^{(2)} + \delta Y_I$

- demand δY_I to be a small perturbation of $Y_I^{(0)}$ and whole solution to be asymptotically flat $\rightarrow \delta Y_I = 0$