

Measuring a_{μ}^{HLO} in the spacelike region

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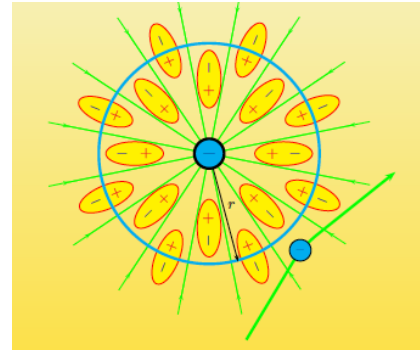
[based on Phys.Lett. B746 (2015) 325-32]

**LFC15: Physics prospects for Linear and other
Future Colliders after the discovery of the Higgs**

Trento, September 7-11, 2015

α_{em} running and the Vacuum Polarization

- Due to Vacuum Polarization effects $\alpha_{em}(q^2)$ is a running parameter from its value at vanishing momentum transfer to the effective q^2 .

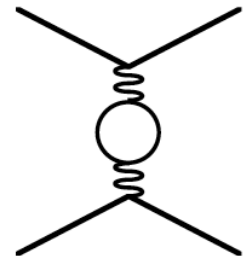
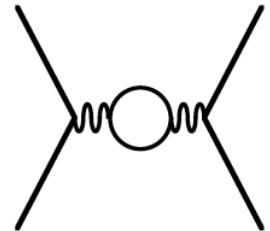


- The “Vacuum Polarization” function $\Pi(q^2)$ can be “absorbed” in a redefinition of an effective charge:

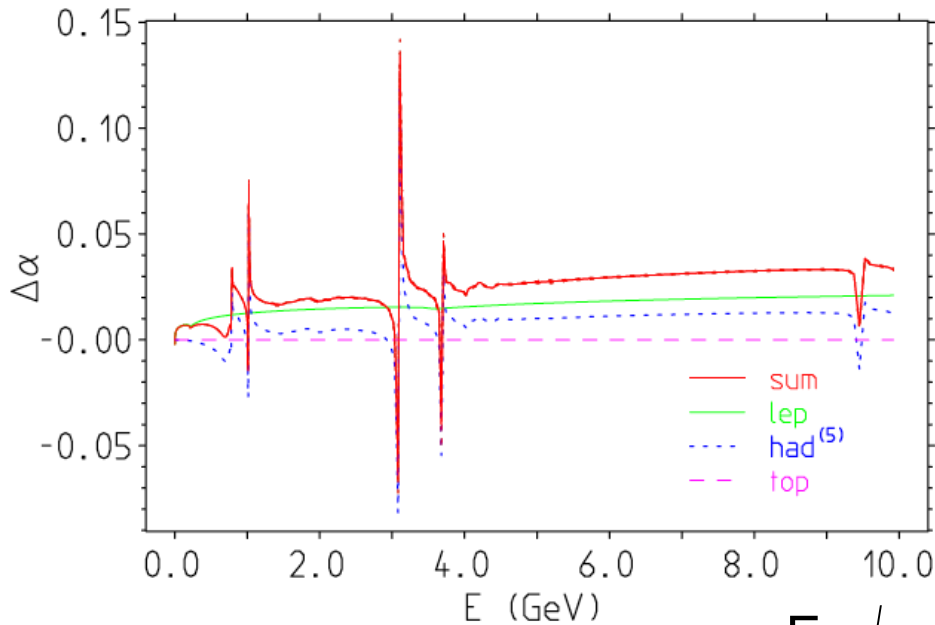
$$e^2 \rightarrow e^2(q^2) = \frac{e^2}{1 + (\Pi(q^2) - \Pi(0))} \quad \alpha(q^2) = \frac{\alpha(0)}{1 - \Delta\alpha}; \quad \Delta\alpha = -\Re e(\Pi(q^2) - \Pi(0))$$

$$\Delta\alpha = \Delta\alpha_l + \Delta\alpha_{had}^{(5)} + \Delta\alpha_{top}$$

- $\Delta\alpha$ takes a contribution by non perturbative hadronic effects ($\Delta\alpha_{had}^{(5)}$) which exhibits a different behaviour in time-like and spacelike region



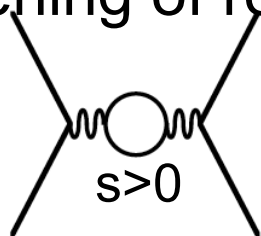
Running of α_{em}



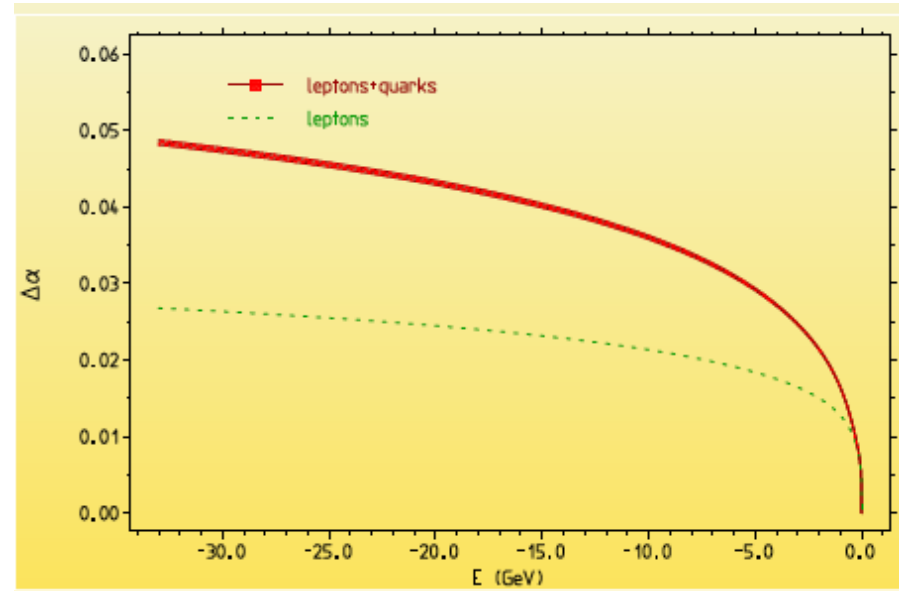
Time-like

$E = \sqrt{s}$

Behaviour characterized by the opening of resonances



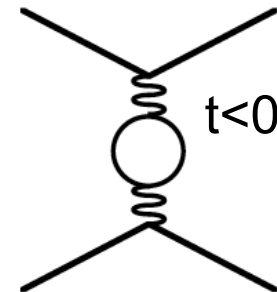
$$\Delta\alpha_{had}^{(5)}(M_Z^2) = -\frac{\alpha M_Z^2}{3\pi} \text{Re} \int_{4m_\pi^2}^{\infty} ds \frac{R(s)}{s(s - M_Z^2 - i\epsilon)}$$



Space-like

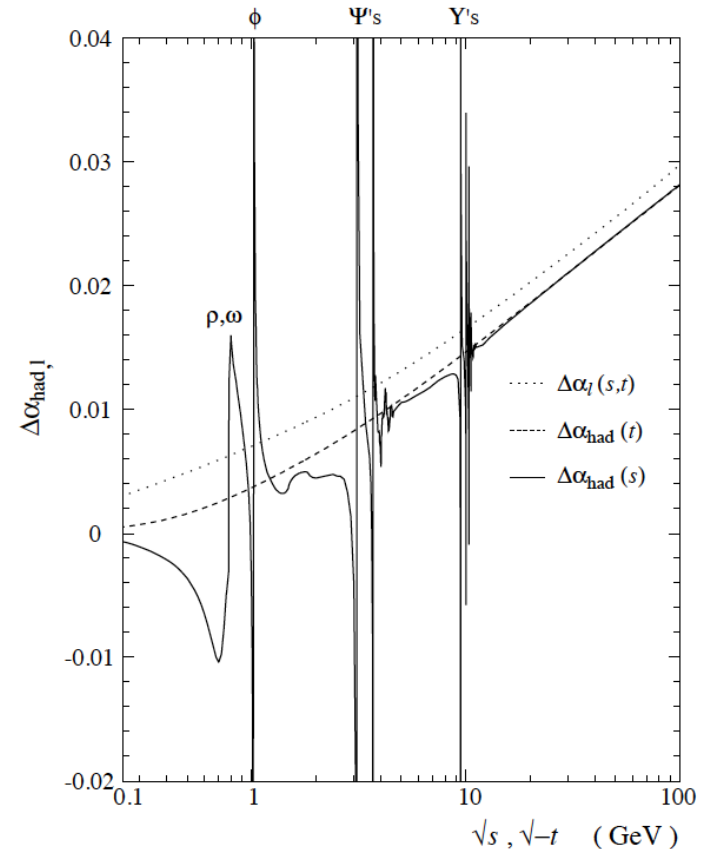
$E = -\sqrt{-t}$

Very smooth behaviour



Measurement of α_{em} running

- A direct measurement of $\alpha_{em}(q^2)$ in space/time like region can prove the running of α_{em}
- It can provide a test of “duality” (fare way from resonances)
- It has been done in past by few experiments at e^+e^- colliders by comparing a “well-known” QED process with some reference (obtained from data or MC)

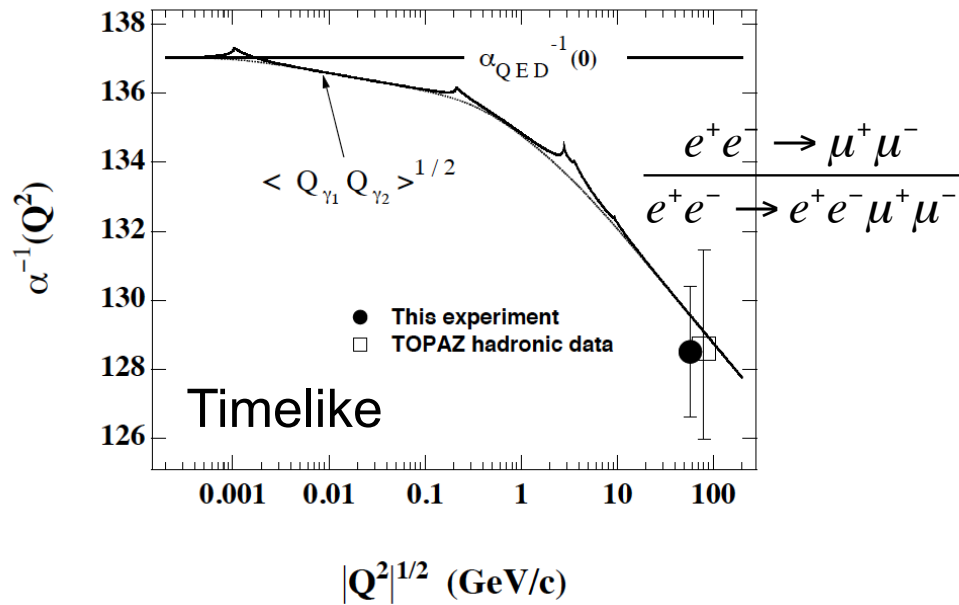


$$\left(\frac{\alpha(q^2)}{\alpha(q_0^2)} \right)^2 \sim \frac{N_{signal}(q^2)}{N_{norm}(q_0^2)}$$

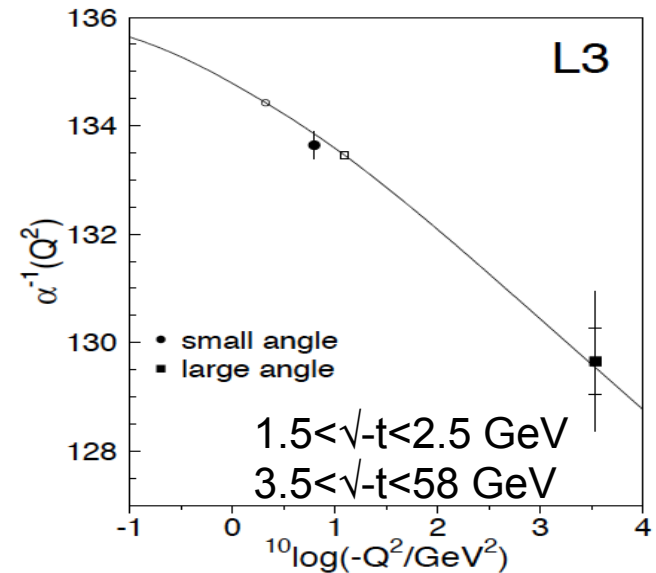
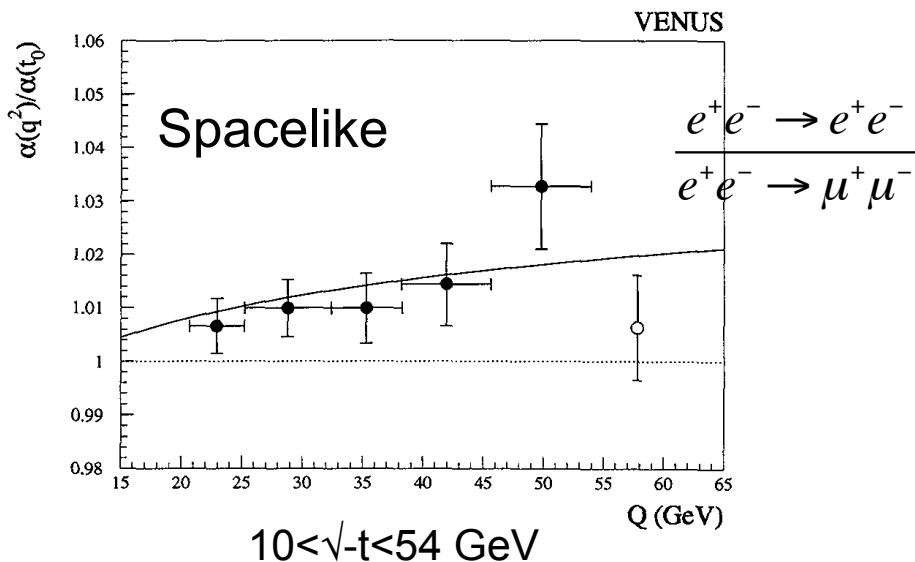
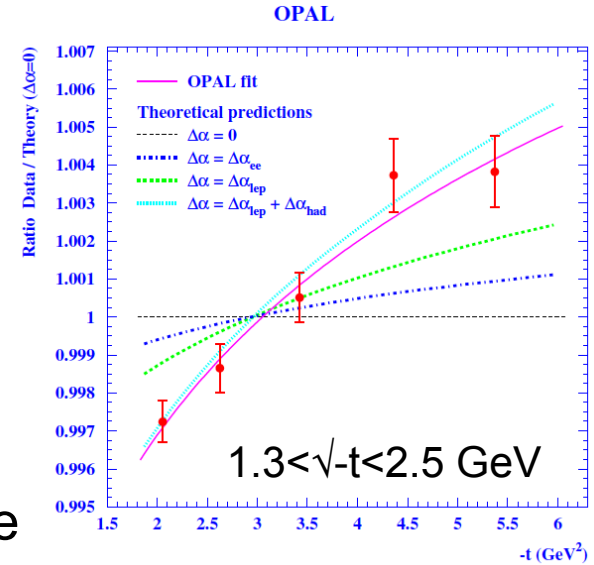
N_{signal} can be Bhabha process, muon pairs, etc...
 N_{signal} can be Bhabha process, $\gamma\gamma$ pairs, Theory, etc...

Measurement of α_{em} running

e+e- collider TRISTAN at $\sqrt{s}=57.8$ GeV,



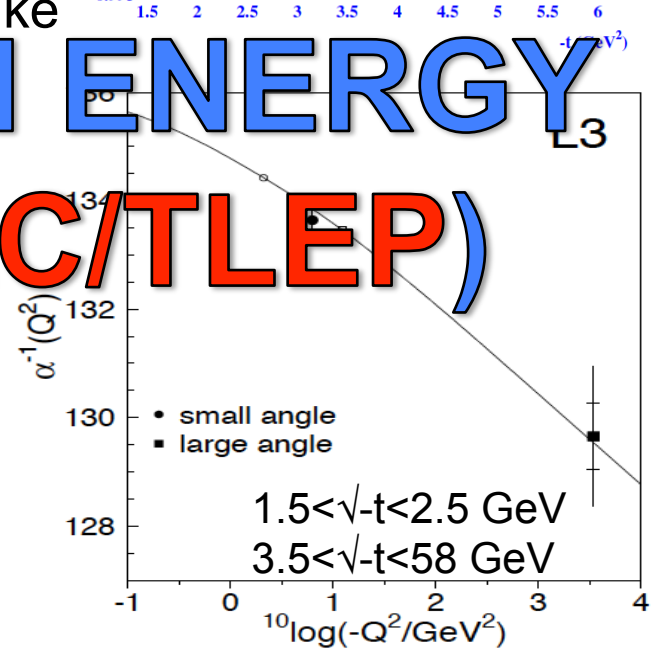
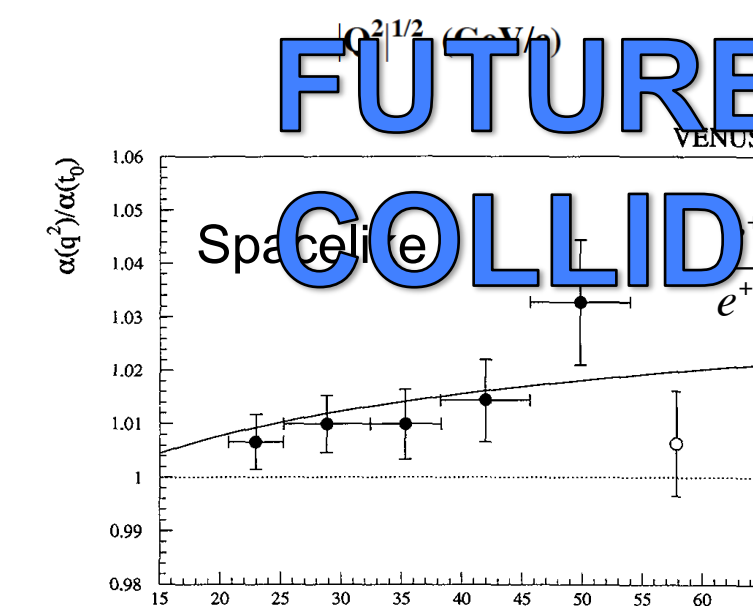
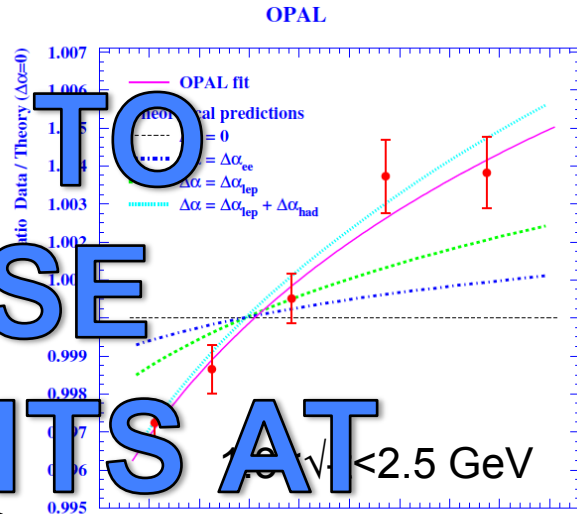
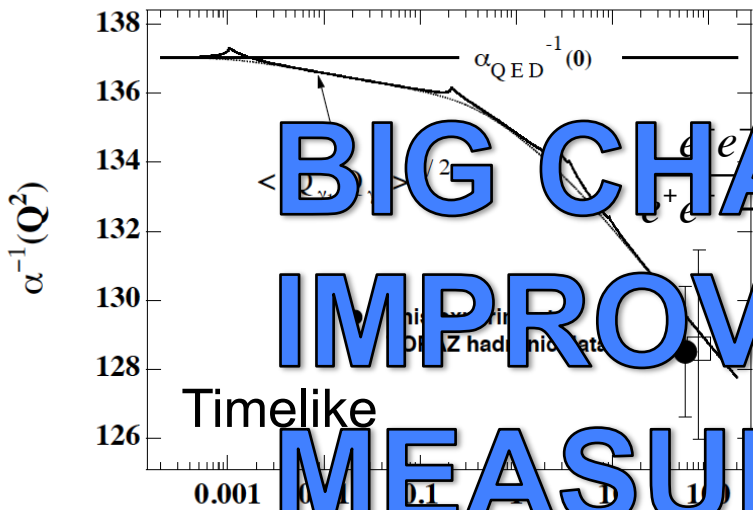
e+e- collider LEP at $\sqrt{s}=189$ GeV, using Bhabha events



Measurement of α_{em} running

e+e- collider TRISTAN at $\sqrt{s}=57.8$ GeV,

e+e- collider LEP at $\sqrt{s}=189$ GeV, using Bhabha events



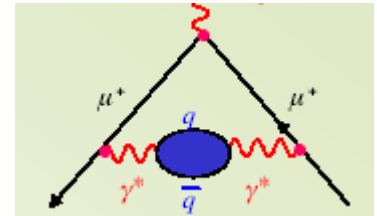
BIG CHANCES TO IMPROVE THESE MEASUREMENTS AT FUTURE HIGH ENERGY COLLIDER (ILC/TLEP)

10 < sqrt(-t) < 54 GeV

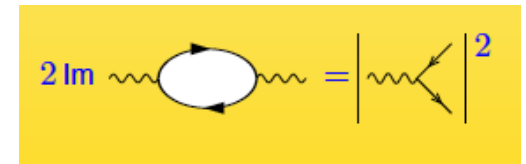
a_μ^{HLO} calculation, traditional way: time-like data

$$a_\mu^{\text{HLO}} = \frac{1}{4\pi^3} \int_{4m_\pi^2}^{\infty} \sigma_{e^+e^- \rightarrow \text{had}}(s) K(s) ds$$

$$a_\mu = (g-2)/2$$



$$a_\mu^{\text{HLO}} = \frac{\alpha}{\pi^2} \int_0^{\infty} \frac{ds}{s} K(s) \text{Im} \Pi_{\text{had}}(s) \sigma_{e^+e^- \rightarrow \text{had}}(s) = \frac{4\pi}{s} \text{Im} \Pi_{\text{had}}(s)$$



$$K(s) = \int_0^1 dx \frac{x^2(1-x)}{x^2 + (1-x)(s/m^2)} \sim \frac{1}{s}$$

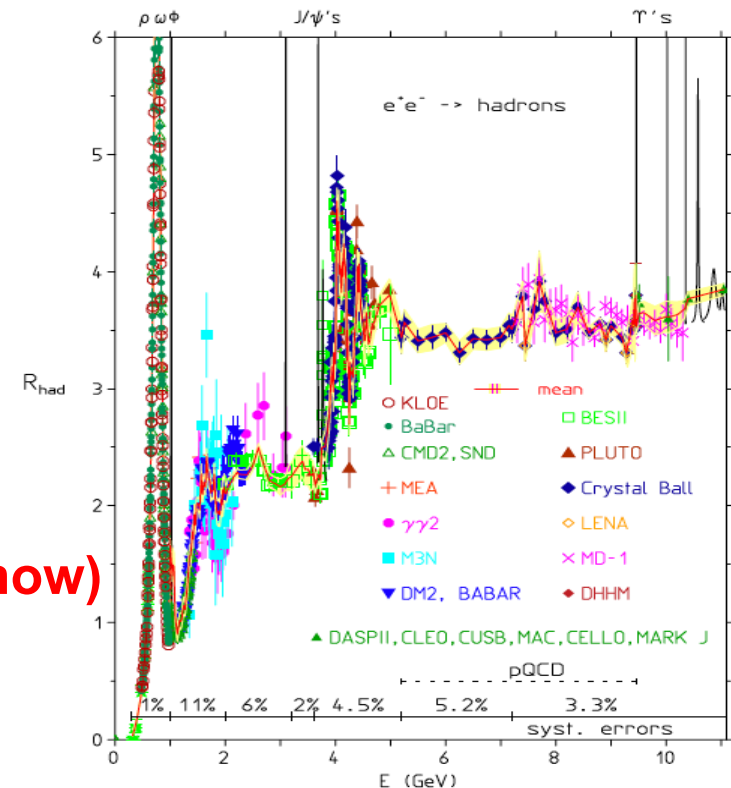
Traditional way: based on precise experimental (time-like) data:

$$a_\mu^{\text{had}} = (689.7 \pm 4.4) \cdot 10^{-10}$$

Main contribution in the low energy region

$\delta a_\mu^{\text{exp}} \rightarrow 1.5 \cdot 10^{-10} = 0.2\%$ on a_μ^{HLO} (from 0.7% now)

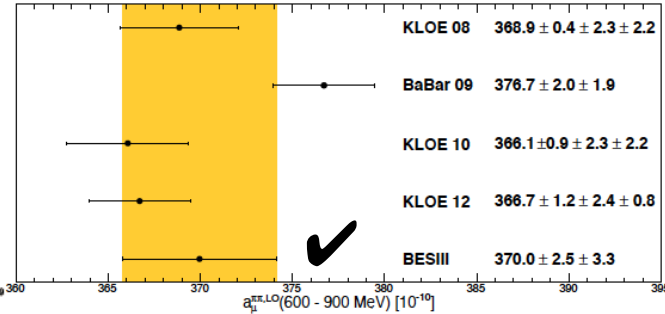
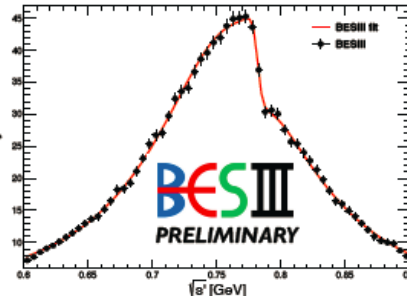
NEW G-2 at FNAL and JPARC



New results on a_μ^{HLO}

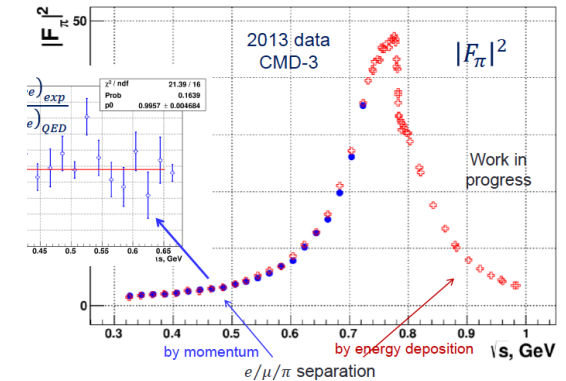
– BES

- New; 1/3rd data; $\pi\pi$ results consistent with others (arxiv:1507.08188)



– VEPP-2000

- New results this year at $\sim 0.6\%$ on $\pi\pi$
- Aim at $\sim 0.3\%$ by 2017 (the ultimate goal)

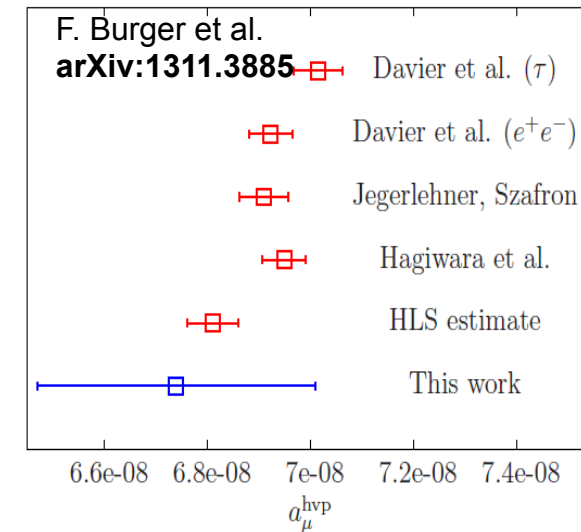


– Lattice

- Results on a_μ^{HLO} with 4% error; 4x worse than the dispersive approach ($\sim 1\%$ uncertainty)

$$a_\mu^{\text{hvp}} = 6.74(21)(18) \cdot 10^{-8} \quad (N_f = 2 + 1 + 1)$$

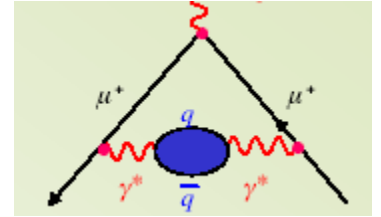
$$a_\mu^{\text{hvp}} = 6.91(01)(05) \cdot 10^{-8} \quad (\text{dispersive analysis}) .$$



a_μ^{HLO} evaluation in spacelike region: alternative approach

$$a_\mu = (g-2)/2$$

$$a_\mu^{HLO} = -\frac{\alpha}{\pi} \int_0^1 (1-x) \Pi_{had} \left(-\frac{x^2}{1-x} m_\mu^2 \right) dx$$

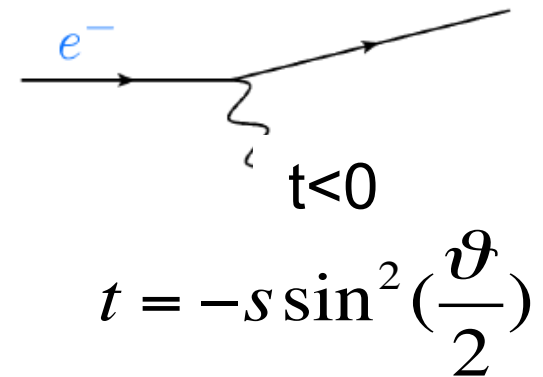


x = Feynman parameter

See also G.Fedotov, proceedings of PHIPSI08

$$t = \frac{x^2 m_\mu^2}{x-1} \quad 0 \leq -t < +\infty$$

$$x = \frac{t}{2m_\mu^2} \left(1 - \sqrt{1 - \frac{4m_\mu^2}{t}} \right); \quad 0 \leq x < 1;$$



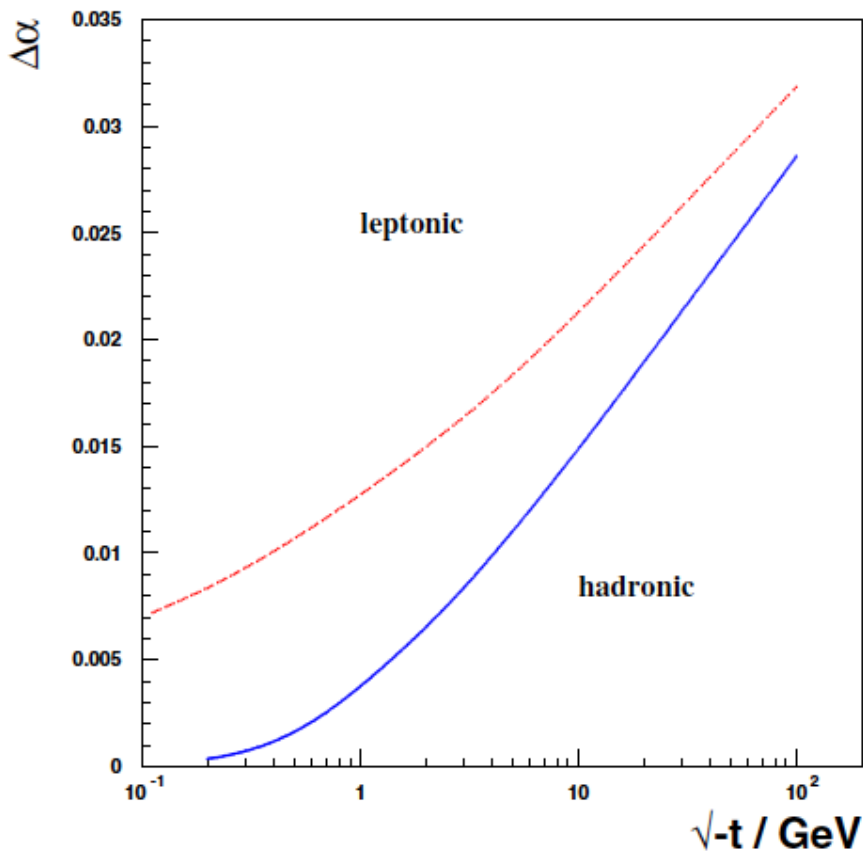
$$t = -s \sin^2 \left(\frac{\vartheta}{2} \right)$$

$$\Delta\alpha_{had}(t) = -\Pi_{had}(t) \quad \text{for } t < 0$$

$$a_\mu^{HLO} = -\frac{\alpha}{\pi} \int_0^1 (1-x) \Delta\alpha_{had} \left(-\frac{x^2}{1-x} m_\mu^2 \right) dx$$

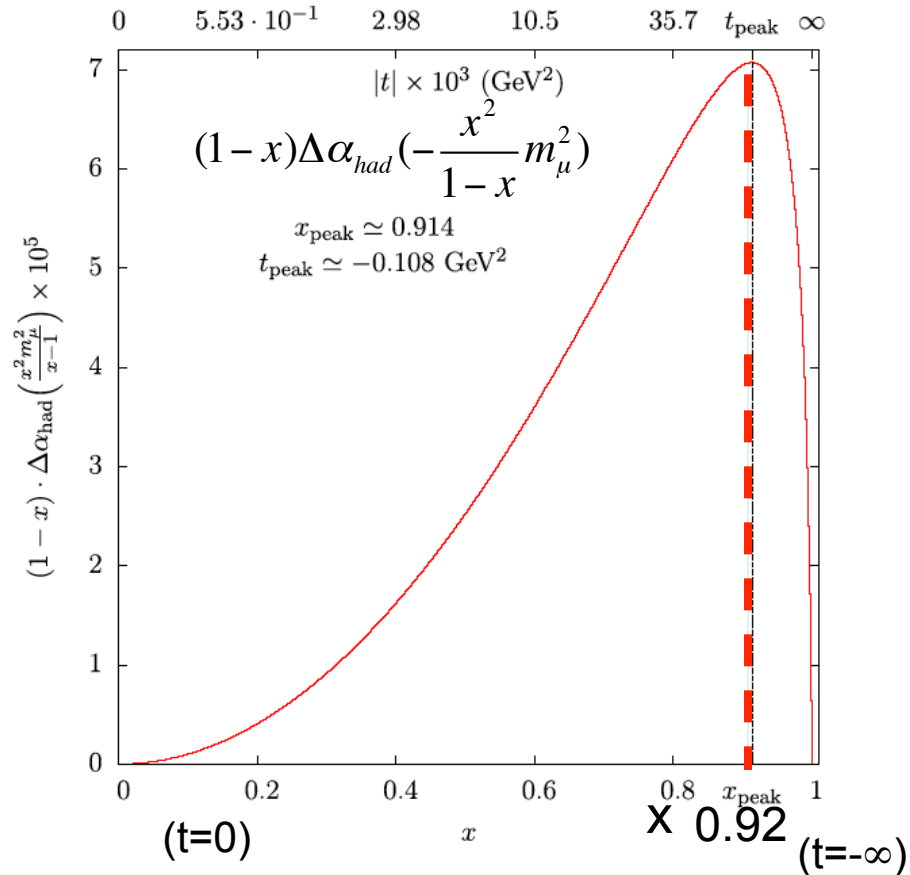
For $t < 0$

Behaviors



$\Delta\alpha \sim \log(-t)$

Dominated at low $|t|$ by leptonic contribution



High $|t|$ -values are depressed by $1-x$
 (a kind of analogy with time-like region)
 The integrand is peaked at $\sim x=0.92$
 $\rightarrow t=-0.11 \text{ GeV}^2$ ($\sim 330 \text{ MeV}$) for which
 $\Delta\alpha_{\text{had}}(\mathbf{0.92}) \sim \mathbf{10^{-3}}$

Experimental considerations

Using Bhabha at small angle (to emphasize t-channel contribution) to extract $\Delta\alpha$:

$$\left(\frac{\alpha(t)}{\alpha(0)}\right)^2 \sim \frac{d\sigma_{ee\rightarrow ee}(t)}{d\sigma_{MC}^0(t)}$$

Where $d\sigma_{MC}^0$ is the MC prediction for Bhabha process with $\alpha(t)=\alpha(0)$, and there are corrections due to RC...

$$\Delta\alpha_{had}(t) = 1 - \left(\frac{\alpha(t)}{\alpha(0)}\right)^{-1} - \Delta\alpha_{lept}(t) \quad \Delta\alpha_{lep}(t) \text{ theoretically well known!}$$

Which experimental accuracy we are aiming at?

$\delta\Delta\alpha_{had} \sim 1/2$ fractional accuracy on $d\sigma(t)/d\sigma_{MC}^0(t)$.

If we assume to measure $\delta\Delta\alpha_{had}$ at 5% at the peak of the integrand ($\Delta\alpha_{had} \sim 10^{-3}$ at $x=0.92$) \rightarrow fractional accuracy on $d\sigma(t)/d\sigma_{MC}^0(t) \sim 10^{-4}$!

Very challenging measurement (one order of magnitude improvement respect to date) for systematic error

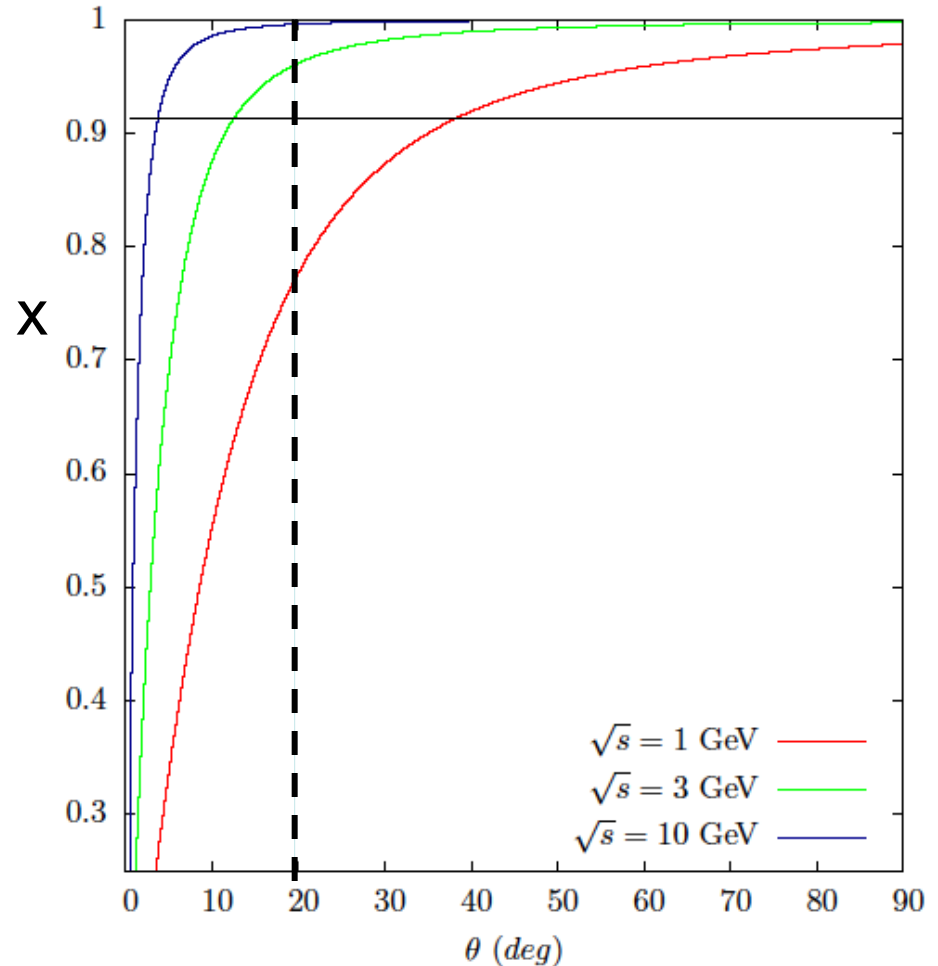
Experimental considerations - II

Most of the region (up to $x \sim 0.98$) can be covered with a low energy machine (like Dafne/VEPP-2000 or tau/charm-B-factories)

Example:

Covering up to 60° at $\sqrt{s}=1$ GeV can arrive at $x=0.95(!)$

A different situation can be obtained at tau/charm/B-factories (and at future ILC/TLEP machines) where smaller angles (below 20°) are needed



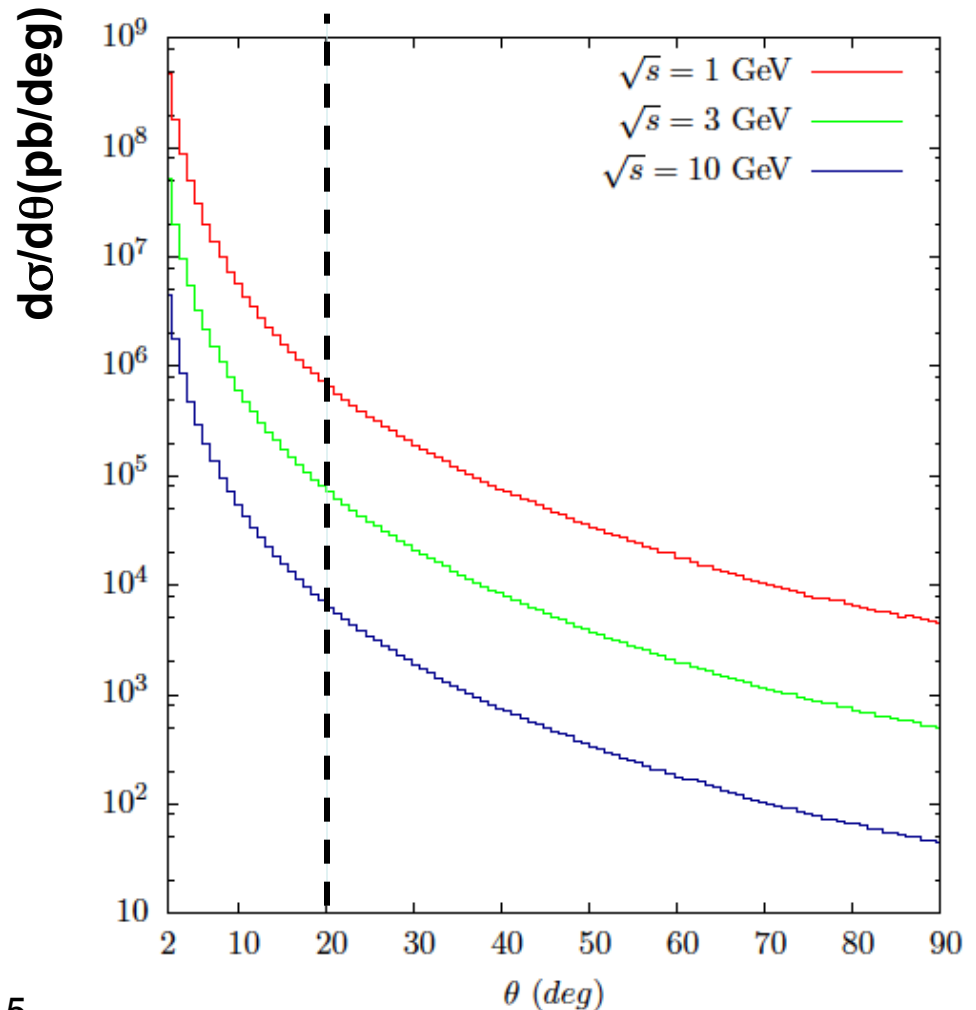
$$t = -s \sin^2\left(\frac{\vartheta}{2}\right)$$

Statistical consideration

10^{-4} accuracy on Bhabha cross section requires at least 10^8 events which at 20° mean at least:

- $O(1) \text{ fb}^{-1}$ @ 1 GeV
- $O(10) \text{ fb}^{-1}$ @ 3 GeV
- $O(100) \text{ fb}^{-1}$ @ 10 GeV

These luminosities are within reach at flavour factories!



Additional considerations: s-channel

At low energy (<10 GeV) above 10^0 there is still a sizeable contribution from s-channel.

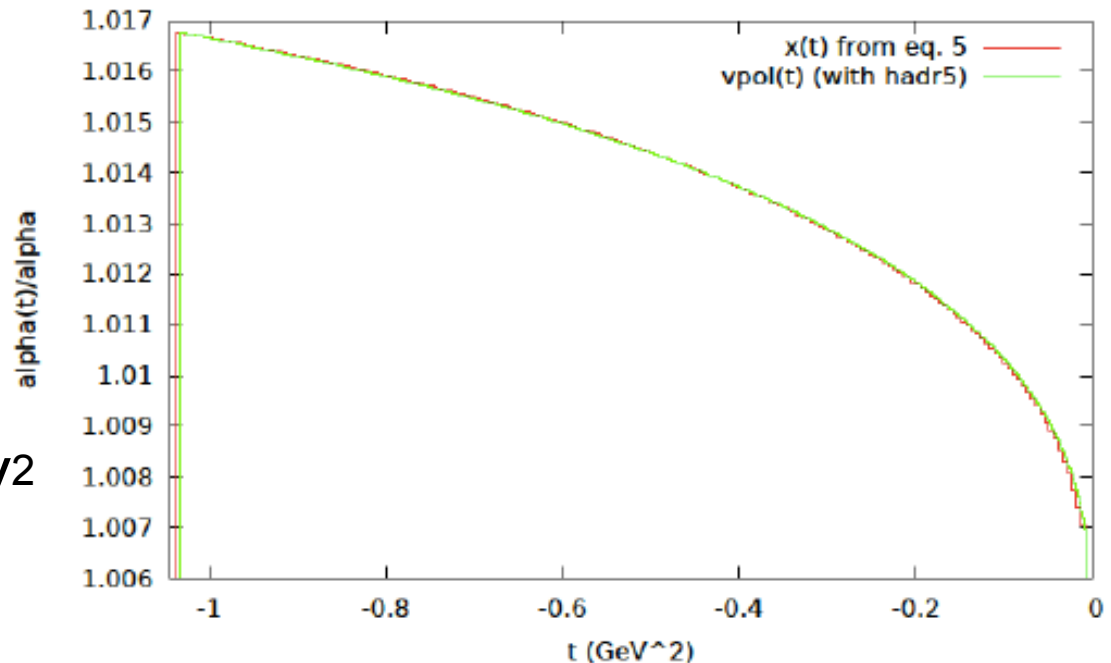
At LO no difficulty to deconvolute the cross section for the s-channel

Test with Babayaga:

$$s=1 \text{ GeV}$$

$$10^\circ < \theta < 170^\circ$$

$$d\sigma_{\text{born}}/dt = 1.52 \text{ mb/GeV}^2$$



However this picture changes with Rad. Corr.

Additional considerations: Rad. Corr.

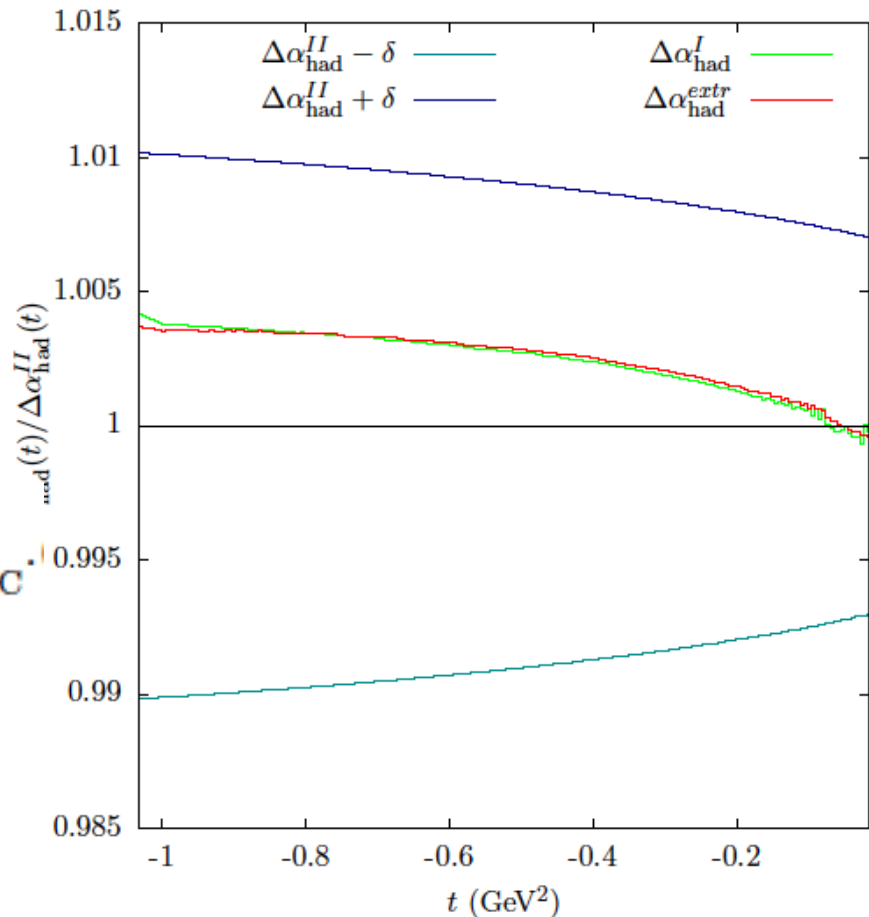
A Monte Carlo procedure has been developed to check if $\Delta\alpha_{\text{had}}(t)$ can be obtained by a minimization procedure with a different $\Delta\alpha_{\text{had}}(t)'$ inside

$$\left. \frac{d\sigma}{dt} \right|_{\text{data}} = \left. \frac{d\sigma}{dt} \left(\alpha(t), \alpha(s) \right) \right|_{\text{MC}},$$

→

$$\left. \frac{d\sigma}{dt} \right|_{j,\text{data}} = \left. \frac{d\sigma}{dt} \left(\bar{\alpha}(t) + \frac{i_j}{N} \delta(t), \alpha(s) \right) \right|_{j,\text{MC}}$$

$\Delta\alpha_{\text{had}}(t)$ is obtained
with $< 10^{-4}$ error !

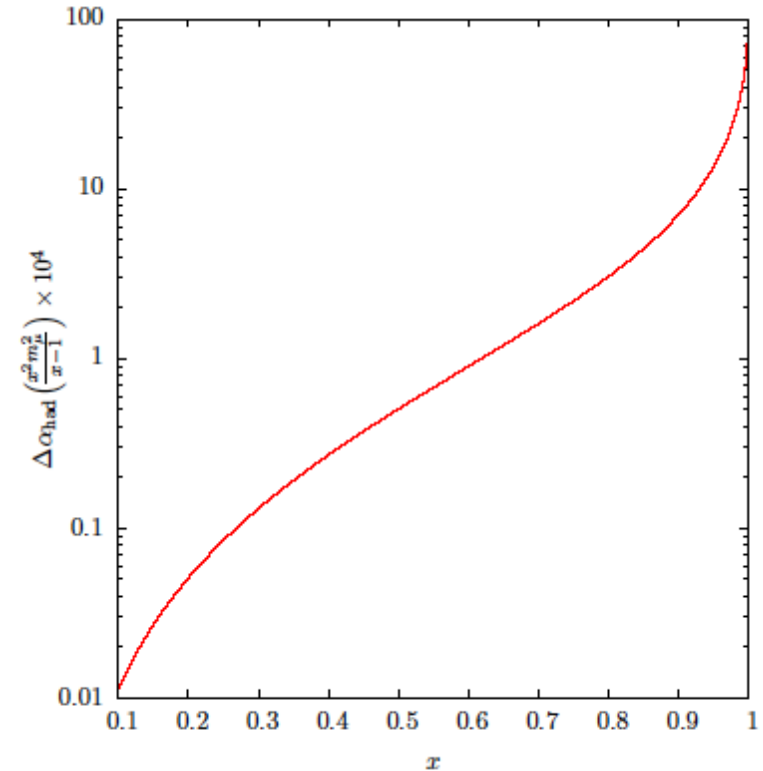


Additional consideration: Normalization

To compare Bhabha absolute cross section from data with MC we need Luminosity of the machine.

Two possibilities:

- 1) Use Bhabha at very small angle where the uncertainty on $\Delta\alpha_{\text{had}}$ can be neglected (for example at $E_{\text{beam}}=1$ GeV and $\theta=5^\circ$, $\Delta\alpha_{\text{had}} \sim 10^{-5}$).
- 2) Use a process with $\Delta\alpha_{\text{had}}=0$, like $e^+e^- \rightarrow \gamma\gamma$. However very difficult to determine it at 10^{-4} accuracy.



Option 1) looks better to us as some of the common systematics cancel in the measurement !

Measurement of DAFNE Luminosity with KLOE/KLOE-2 at 10^{-4} ?

F. Ambrosino et al [KLOE] Eur. Phys. J. C 47, 589–596 (2006)

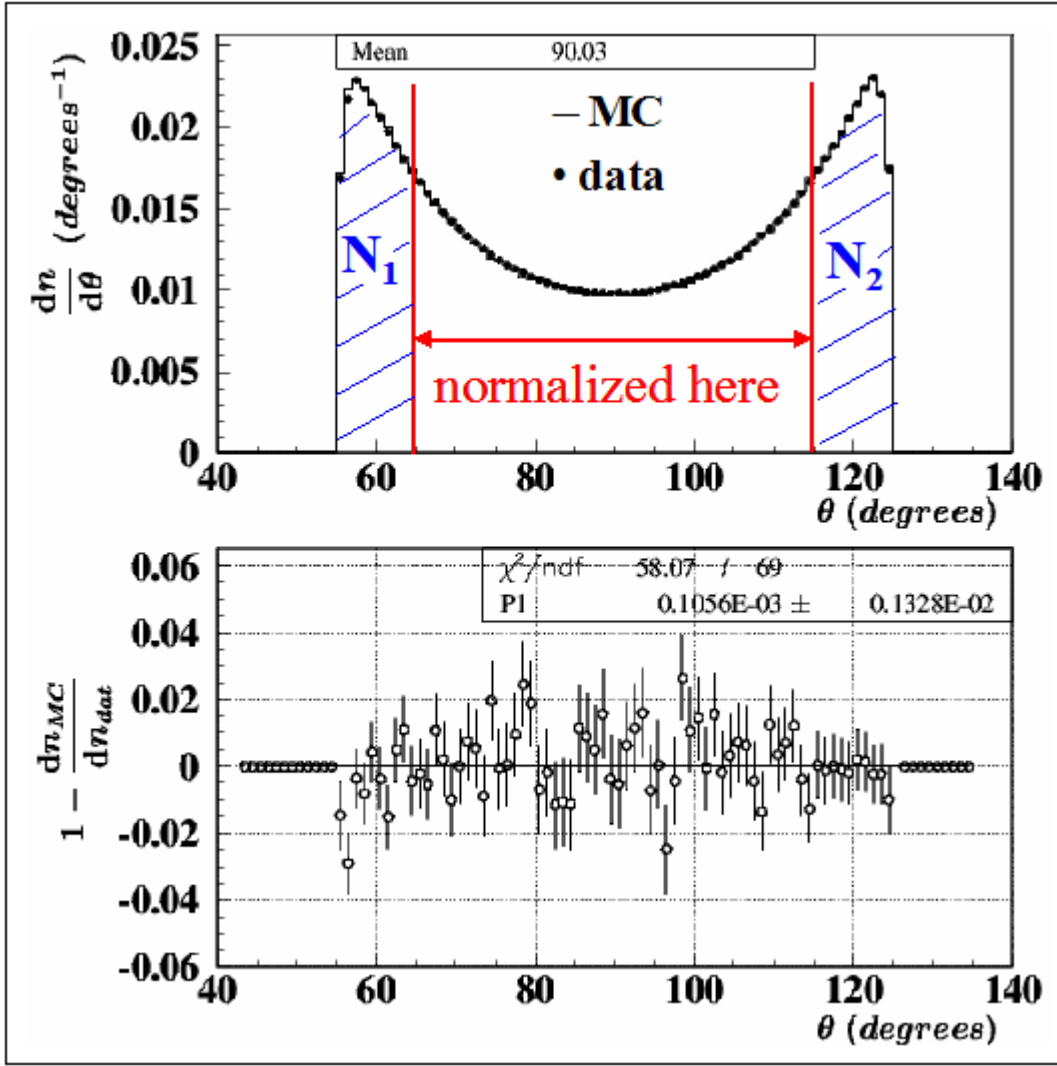
Table 2. Summary of the corrections and systematic errors in the measurement of the luminosity

	correction (%)	systematic error (%)
angular acceptance	+0.25	0.25
tracking	–	0.06
clustering	+0.14	0.11
background	–0.62	0.13
cosmic veto	+0.40	–
energy calibration	–	0.10
center of mass energy	+0.10	0.10
	+0.34	0.32

Adding in quadrature: 0.3 %

(can be improved by a factor 10?)

From F. Nguyen 2006 Polar angle systematics



✓ global agreement is very good

but the cut occurs in a steep region of the distributions
 ⇒ estimate of border mismatches

✓ after normalizing MC to make it coincide with data in the region $65^\circ < \theta < 115^\circ$, we estimate as a systematic error:

$$\frac{N_{[55:65]+[115:125]}^{dat} - N_{[55:65]+[115:125]}^{MC}}{N_{TOT}^{dat}} \sim 0.25\%$$

Can be improved at 10^{-4} ?

A measurement of the Luminosity at 10^{-4} at LEP

Giovanni Abbiendi

INFN - Bologna

Eur. Phys. J. C 45, 1–21 (2006)
Digital Object Identifier (DOI) 10.1140/epjc/s2005-02389-3

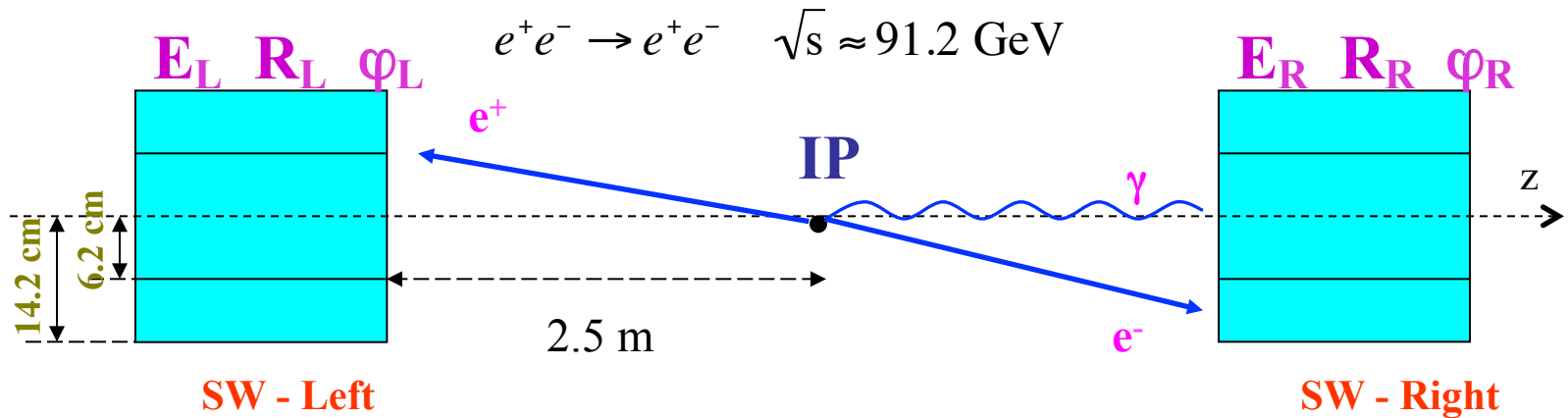
THE EUROPEAN
PHYSICAL JOURNAL C

Measurement of the running of the QED coupling in small-angle Bhabha scattering at LEP

The OPAL Collaboration

G. Abbiendi², C. Ainsley⁵, P.F. Åkesson^{3,y}, G. Alexander²², G. Anagnostou¹, K.J. Anderson⁹, S. Asai²³, D. Axen²⁷, I. Bailey²⁶, E. Barberio^{8,p}, T. Barillari³², R.J. Barlow¹⁶, R.J. Batley⁵, P. Bechtle²⁵, T. Behnke²⁵, K.W. Bell²⁰, P.J. Bell¹, G. Bella²², A. Bellerive⁶, G. Benelli⁴, S. Bethke³², O. Biebel³¹, O. Boeriu¹⁰, P. Bock¹¹, M. Boutemur³¹, S. Braibant², R.M. Brown²⁰, H.J. Burckhart⁸, S. Campana⁴, P. Capiluppi², R.K. Carnegie⁶, A.A. Carter¹³, J.R. Carter⁵, C.Y. Chang¹⁷, D.G. Charlton¹, C. Ciocca², A. Csilling²⁹, M. Cuffiani², S. Dado²¹, G.M. Dallavalle², A. De Roeck⁸, E.A. De Wolf^{8,s}, K. Desch²⁵, B. Dienes³⁰, J. Dubbert³¹, E. Duchovni²⁴, G. Duckeck³¹, I.P. Duerdoth¹⁶, E. Etzion²², F. Fabbri², P. Ferrari⁸, F. Fiedler³¹, I. Fleck¹⁰, M. Ford¹⁶, A. Frey⁸, P. Gagnon¹², J.W. Gary⁴, C. Geich-Gimbel³, G. Giacomelli², P. Giacomelli², R. Giacomelli², M. Giunta⁴, J. Goldberg²¹, E. Gross²⁴, J. Grunhaus²², M. Gruwe⁸, P.O. Günther³, A. Gupta⁹, C. Hajdu²⁹, M. Hamann²⁵, G.G. Hanson⁴, A. Harel²¹, M. Hauschild⁸, C.M. Hawkes¹, R. Hawkings⁸, R.J. Hemingway⁶, G. Herten¹⁰, R.D. Heuer²⁵, J.C. Hill⁵, D. Horváth^{29,c}, P. Igo-Kemenes¹¹, K. Ishii²³, H. Jeremie¹⁸, P. Jovanovic¹, T.R. Junk^{6,i}, J. Kanzaki^{23,u}, D. Karlen²⁶, K. Kawagoe²³, T. Kawamoto²³, R.K. Keeler²⁶, R.G. Kellogg¹⁷, B.W. Kennedy²⁰, S. Kluth³², T. Kobayashi²³, M. Kobel³, S. Komamiya²³, T. Krämer²⁵, P. Krieger^{6,1}, J. von Krogh¹¹, T. Kuhl²⁵, M. Kupper²⁴, G.D. Lafferty¹⁶, H. Landsman²¹, D. Lanske¹⁴, D. Lellouch²⁴, J. Letts^o, L. Levinson²⁴, J. Lillich¹⁰, S.L. Lloyd¹³, F.K. Loebinger¹⁶, J. Lu^{27,w}, A. Ludwig³, J. Ludwig¹⁰, W. Mader^{3,b}, S. Marcellini², A.J. Martin¹³, T. Mashimo²³, P. Mättig^m, J. McKenna²⁷, R.A. McPherson²⁶, F. Meijers⁸, W. Menges²⁵, F.S. Merritt⁹, H. Mes^{6,a}, N. Meyer²⁵, A. Michelini², S. Mihara²³, G. Mikenberg²⁴, D.J. Miller¹⁵, W. Mohr¹⁰, T. Mori²³, A. Mutter¹⁰, K. Nagai¹³, I. Nakamura^{23,v}, H. Nanjo²³, H.A. Neal³³, R. Nisius³², S.W. O’Neale^{1,*}, A. Oh⁸, M.J. Oreglia⁹, S. Orito^{23,*}, C. Pahl³², G. Pásztor^{4,g}, J.R. Pater¹⁶, J.E. Pilcher⁹, J. Pinfold²⁸, D.E. Plane⁸, O. Pooth¹⁴, M. Przybycien^{8,n}, A. Quadt³, K. Rabbertz^{8,r}, C. Rember⁸, P. Renkel²⁴, J.M. Roney²⁶, A.M. Rossi², Y. Rozen²¹, K. Runge¹⁰, K. Sachs⁶, T. Saeki²³, E.K.G. Sarkisyan^{8,j}, A.D. Schaile³¹, O. Schaile³¹, P. Scharff-Hansen⁸, J. Schieck³², T. Schörner-Sadenius^{8,z}, M. Schröder⁸, M. Schumacher³, R. Seuster^{14,f}, T.G. Shears^{8,h}, B.C. Shen⁴, P. Sherwood¹⁵, A. Skuja¹⁷, A.M. Smith⁸, R. Sobie²⁶, S. Söldner-Rembold¹⁶, F. Spano⁹, A. Stahl^{3,x}, D. Strom¹⁹, R. Ströhmer³¹, S. Tarem²¹, M. Tasevsky^{8,s}, R. Teuscher⁹, M.A. Thomson⁵, E. Torrence¹⁹, D. Toya²³, P. Tran⁴, I. Trigger⁸, Z. Trócsányi^{30,e}, E. Tsur²², M.F. Turner-Watson¹, I. Ueda²³, B. Ujvári^{30,e}, C.F. Vollmer³¹, P. Vannerem¹⁰, R. Vértési^{30,e}, M. Verzocchi¹⁷, H. Voss^{8,q}, J. Vossebeld^{8,h}, C.P. Ward⁵, D.R. Ward⁵, P.M. Watkins¹, A.T. Watson¹, N.K. Watson¹, P.S. Wells⁸, T. Wengler⁸, N. Wermes³, G.W. Wilson^{16,k}, J.A. Wilson¹, G. Wolf²⁴, T.R. Wyatt¹⁶, S. Yamashita²³, D. Zer-Zion⁴, L. Zivkovic²⁴

Small-angle Bhabha scattering in OPAL



2 cylindrical calorimeters encircling the beam pipe at $\pm 2.5 \text{ m}$ from the Interaction Point

19 Silicon layers

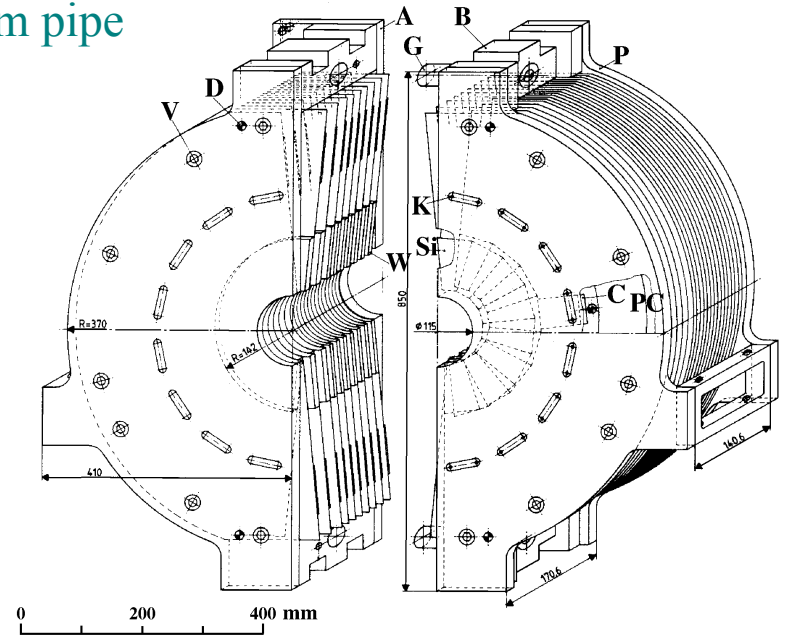
Total Depth $22 X_0$

18 Tungsten layers

(14 cm)

Each detector layer divided into 16 overlapping wedges

Sensitive radius: 6.2 – 14.2 cm,
corresponding to scattering angle
of **25 – 58 mrad** from the beam line



Final Error on Luminosity

After all the effort on Radial reconstruction the dominant systematic error is related to Energy (mostly tail in the E response and nonlinearity)

Quantitatively:

(OPAL Collaboration, Eur.Phys.J. C14 (2000) 373)

	Systematic Error ($\times 10^{-4}$)
Energy	1.8
Inner Anchor	1.4
Radial Metrology	1.4

Total Experimental Systematic Error : 3.4×10^{-4}

Theoretical Error on Bhabha cross section: 5.4×10^{-4}

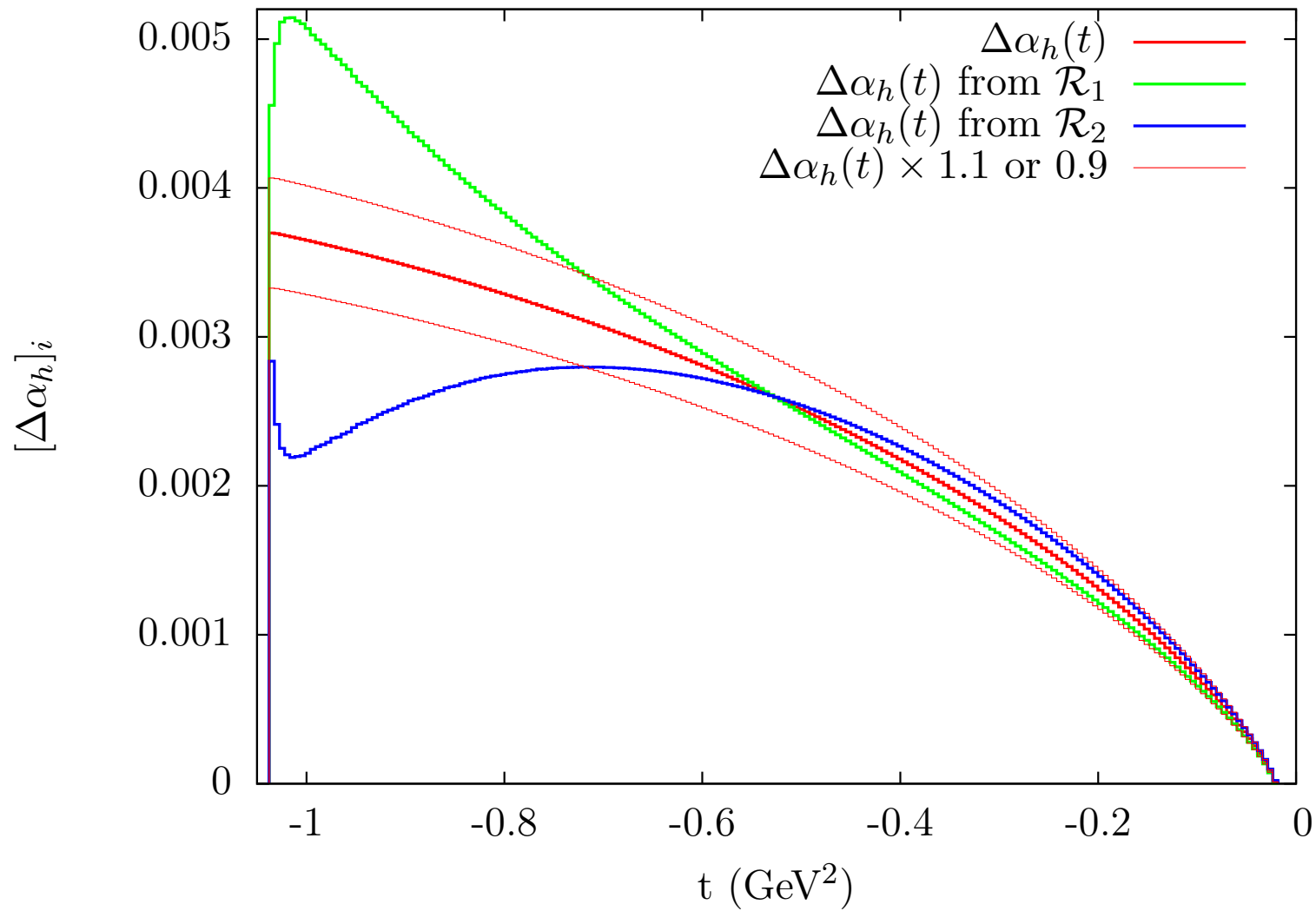
Conclusions

- Measuring α_{em} running in time-like and space like region appears to be very interesting. (Relatively) high q^2 -values can be explored at ILC/TLEP
- An alternative formula for a_{μ}^{HLO} in spacelike region has been studied in details. It emphasizes low values of t ($<1 \text{ GeV}^2$) and can be explored at low energy e^+e^- machines (VEPP2000/DAFNE, τ /charm, B-factories)
- It requires to measure the Bhabha cross section at relatively small angles at (better than) 10^{-4} accuracy!
- Reaching such an accuracy demands a dedicated experimental and theoretical work for the next few years
- Can this method apply also at other (e^-e^- ; fixed target) machines?

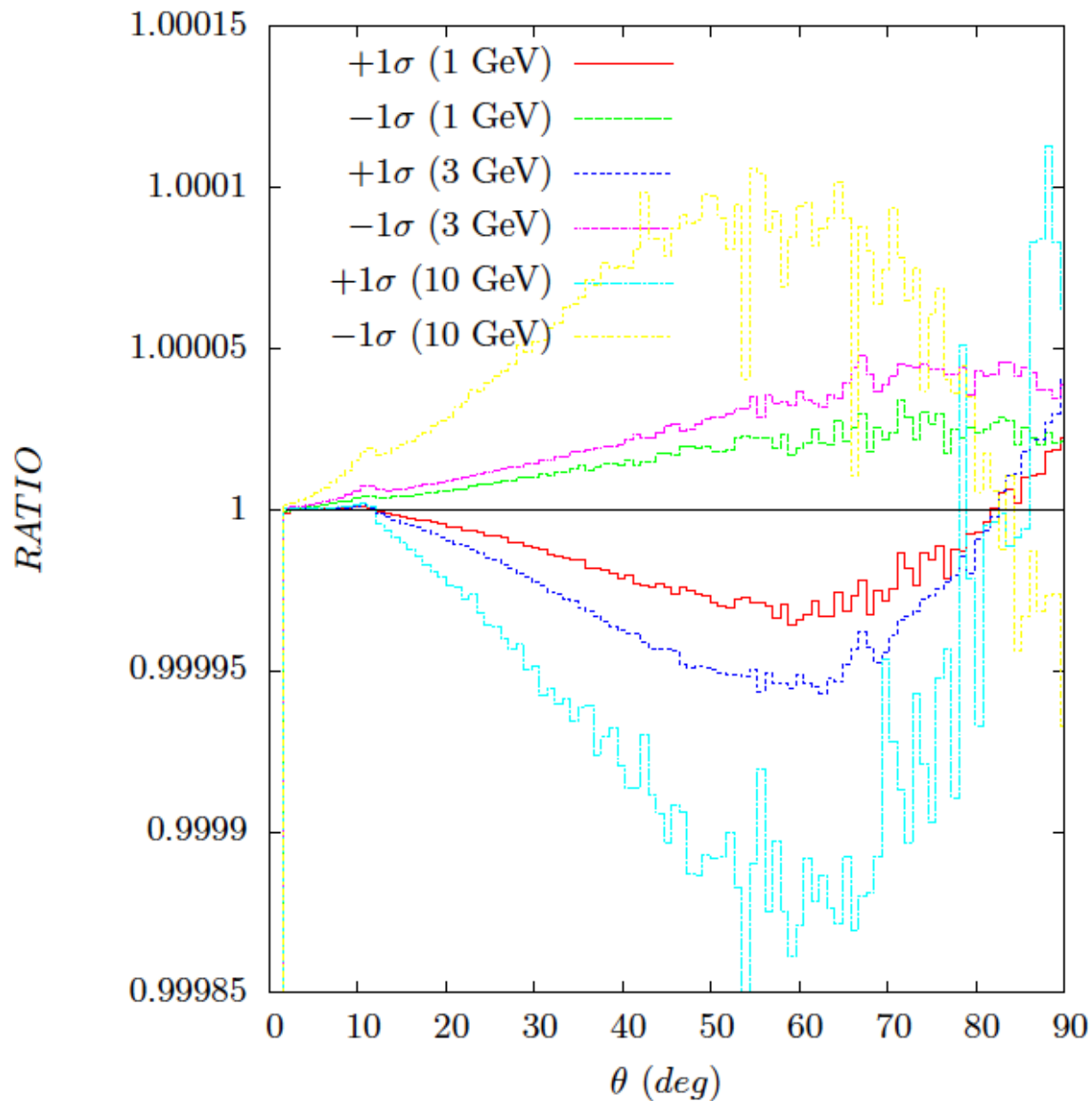
Thanks!

END

test



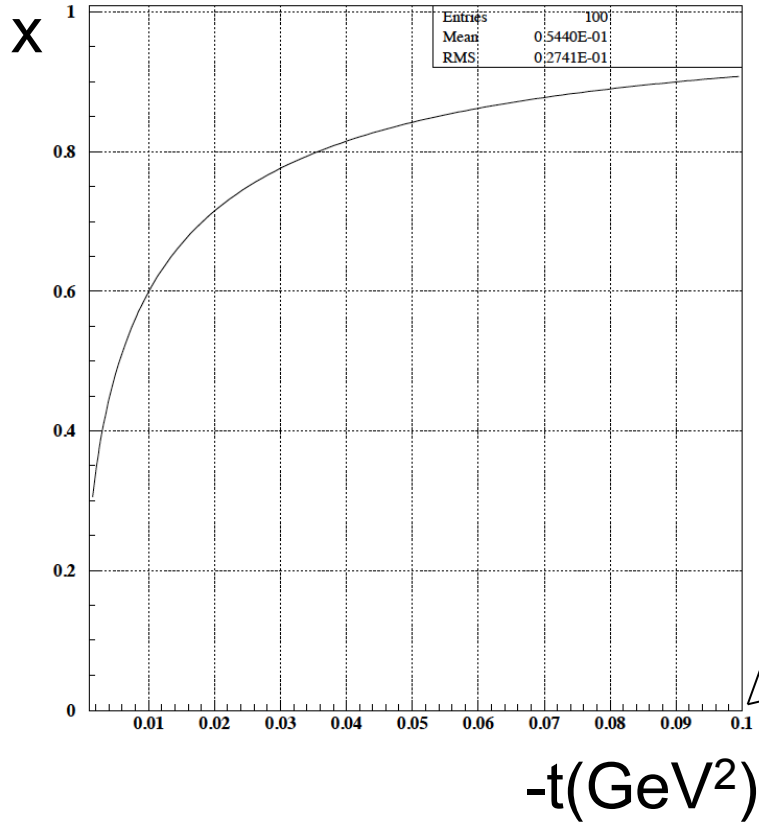
$\Delta\alpha_{em}^{\text{HAD}}(s)$ dependence



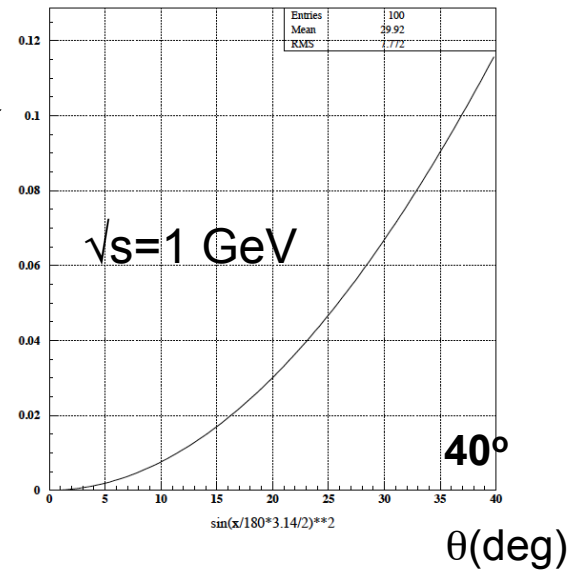
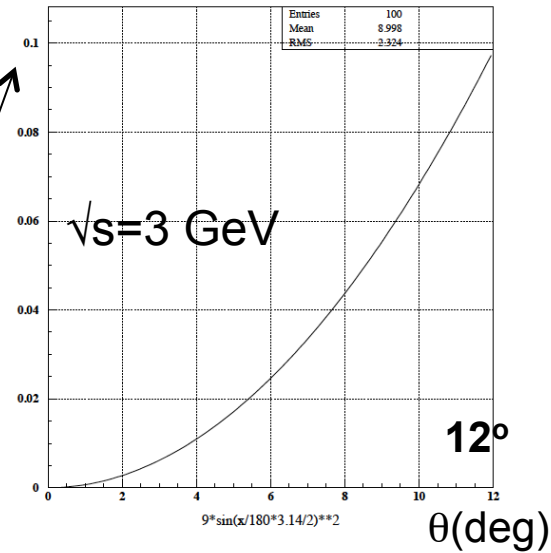
Which is the best energy/angle configuration?

$$x = \frac{t}{2m_\mu^2} \left(1 - \sqrt{1 - \frac{4m^2}{t}}\right)$$

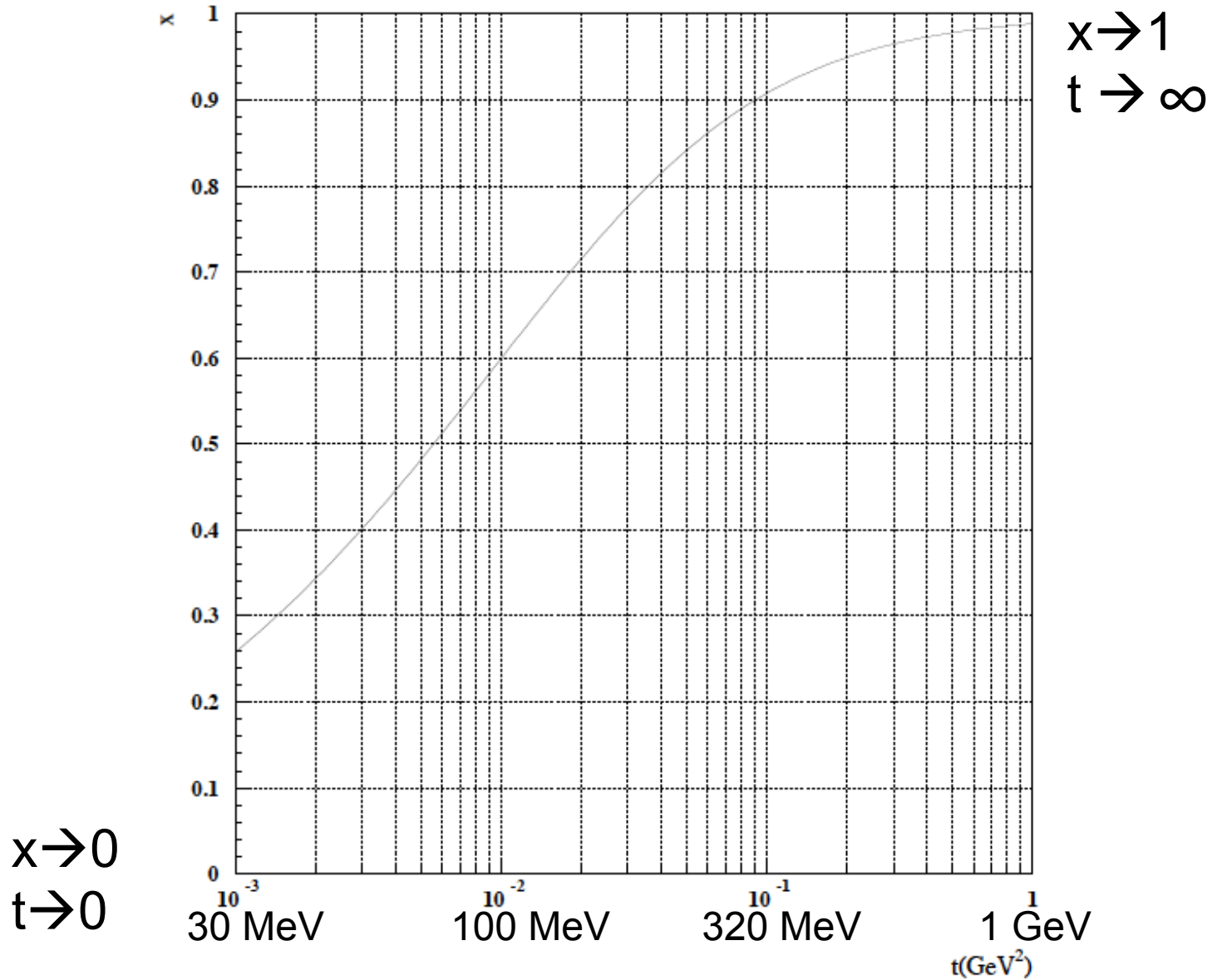
2013/06/24 00:37



$$-t = 9(1 - \cos\theta)/2$$



x vs t behaviour



Measuring a_{μ}^{HLO} in the spacelike region

C.M.C. Calame¹, M. Passera², L. Trentadue³, G. Venanzoni⁴

[\[arXiv:1504.02228\]](https://arxiv.org/abs/1504.02228), Phys.Lett. B746 (2015) 325-32]

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**LFC15: Physics prospects for Linear and other
Future Colliders after the discovery of the Higgs**

Trento, September 7-11, 2015



ECT*



EUROPEAN CENTRE FOR THEORETICAL STUDIES
IN NUCLEAR PHYSICS AND RELATED AREAS

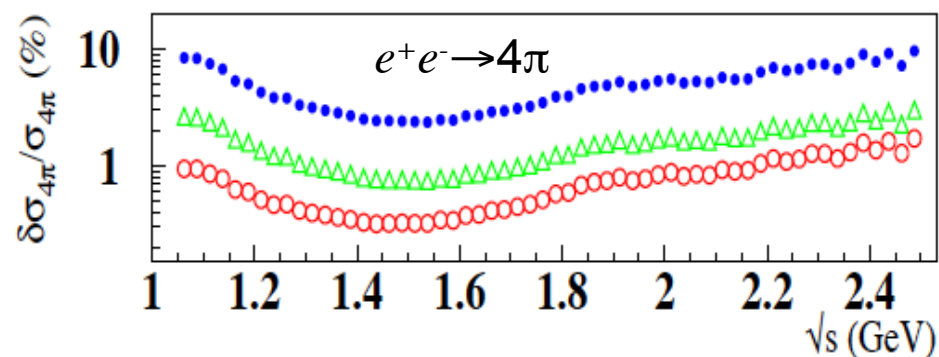
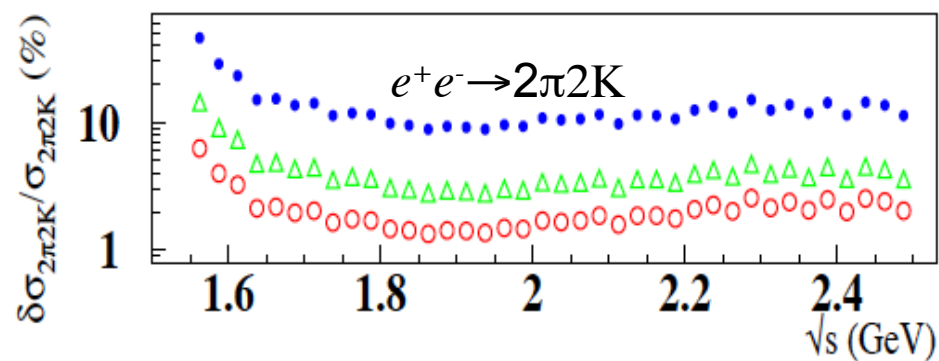
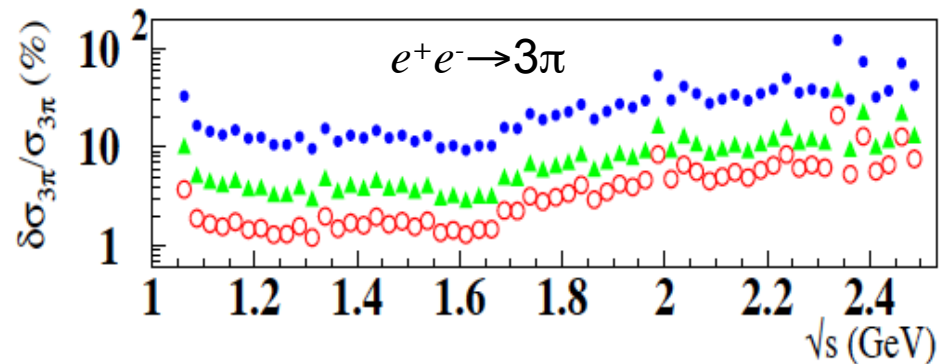
TRENTO, ITALY

Institutional Member of the ESF Expert Committee NuPECC

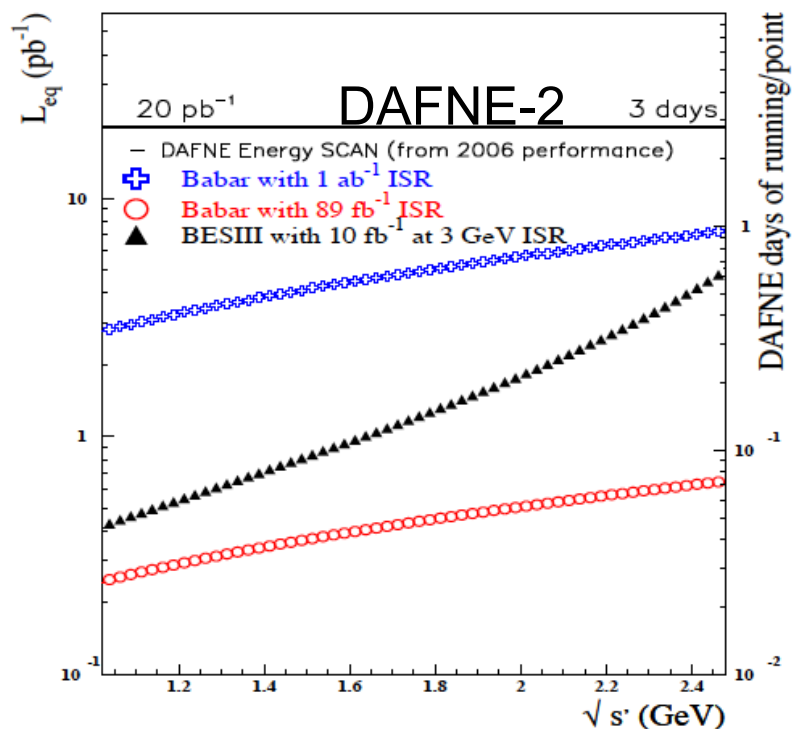
Muon g-2 Physics Workshop, Seattle, 14 July 2015

Impact of DAFNE-2 on exclusive channels in the range [1-2.5] GeV with a scan (Statistics only)

arXiv:1007.521



- Published BaBar results: 89 fb^{-1} (ISR)
- ▲ “BaBar” $\times 10$ (890 fb^{-1})
- KLOE-2 energy scan: $20 \text{ pb}^{-1}/\text{point}$
@ $L = 10^{32} \text{ cm}^{-2} \text{ s}^{-1}$, 25 MeV bin
 \Rightarrow 1 year data-taking



DAFNE-2 is statistically equivalent to $5 \div 10 \text{ ab}^{-1}$ (Super)B-factory