

# Testing the Standard Model with the lepton $g-2$

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# Outline

- 1.  $\mu$ : The muon g-2: recent theory progress
- 2.  $e$ : Testing the Standard Model with the electron g-2
- 3.  $\tau$ : The tau g-2: opportunities or fantasies? Surprises?

# Lepton magnetic moments: the basics

## The beginning: $g = 2$

- Uhlenbeck and Goudsmit in 1925 proposed:

$$\vec{\mu} = g \frac{e}{2mc} \vec{s}$$
$$g = \underline{2} \quad (\text{not } 1!)$$

- Dirac 1928:

$$(i\partial_\mu - eA_\mu) \gamma^\mu \psi = m\psi$$

- A Pauli term in Dirac's eq would give a deviation...

$$a \frac{e}{2m} \sigma_{\mu\nu} F^{\mu\nu} \psi \quad \rightarrow \quad g = 2(1 + a)$$

...but there was no need for it!  $g=2$  stood for ~20 yrs.

# Theory of the g-2: Quantum Field Theory

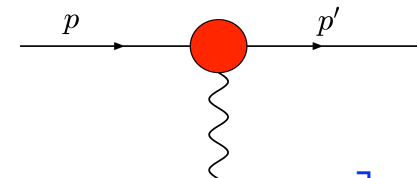
- Kusch and Foley 1948:

$$\mu_e^{\text{exp}} = \frac{e\hbar}{2mc} (1.00119 \pm 0.00005)$$

- Schwinger 1948 (triumph of QED!):

$$\mu_e^{\text{th}} = \frac{e\hbar}{2mc} \left(1 + \frac{\alpha}{2\pi}\right) = \frac{e\hbar}{2mc} \times 1.00116$$

- Keep studying the lepton- $\gamma$  vertex:

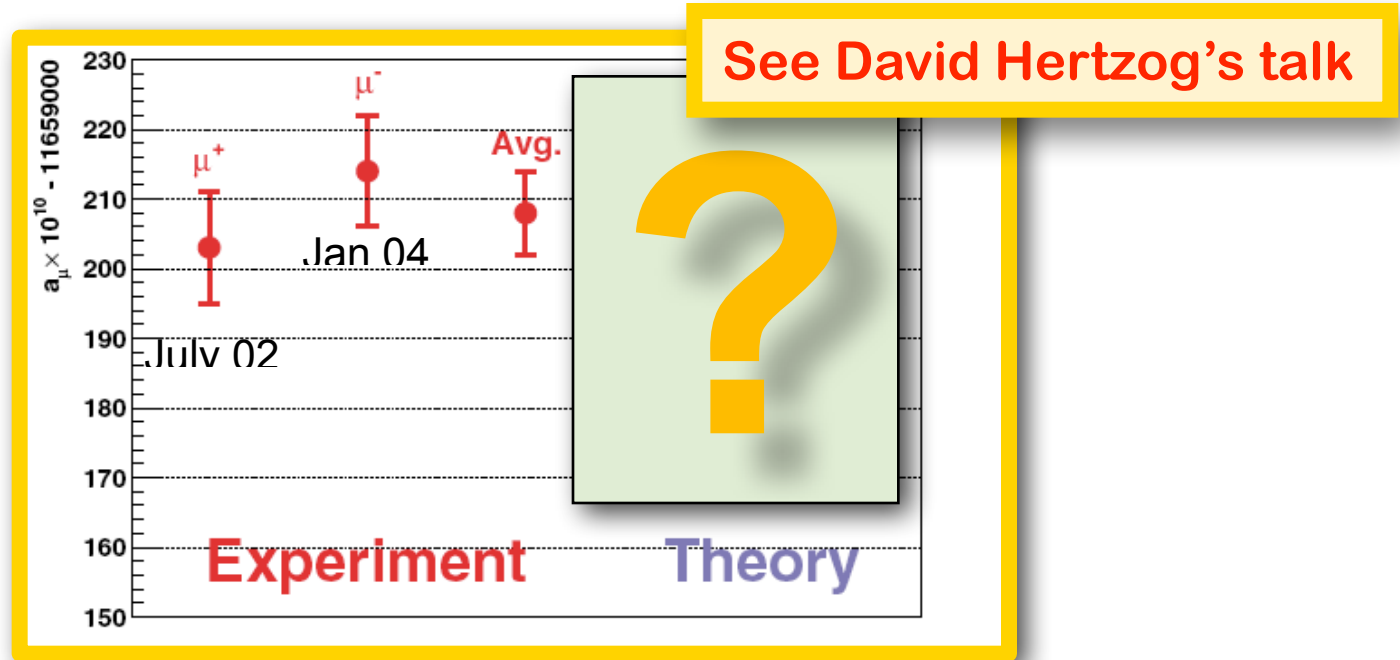


$$\bar{u}(p')\Gamma_\mu u(p) = \bar{u}(p') \left[ \gamma_\mu F_1(q^2) + \frac{i\sigma_{\mu\nu}q^\nu}{2m} F_2(q^2) + \dots \right] u(p)$$

$$F_1(0) = 1 \quad F_2(0) = a_l$$

A pure “quantum correction” effect!

# The muon $g-2$ : recent theory progress



- Today:  $a_\mu^{\text{EXP}} = (116592089 \pm 54_{\text{stat}} \pm 33_{\text{sys}}) \times 10^{-11}$  [0.5ppm].
- Future: new muon g-2 experiments proposed at:
  - Fermilab E989, aiming at  $\pm 16 \times 10^{-11}$ , ie 0.14ppm
  - J-PARC proposal aiming at 0.1 ppm
- Are theorists ready for this (amazing) precision? No(t yet)

# The muon g-2: the QED contribution



$$a_{\mu}^{\text{QED}} = (1/2)(\alpha/\pi) \quad \text{Schwinger 1948}$$

$$+ 0.765857426 (16) (\alpha/\pi)^2$$

Sommerfield; Petermann; Suura&Wichmann '57; Elend '66; MP '04

$$+ 24.05050988 (28) (\alpha/\pi)^3$$

Remiddi, Laporta, Barbieri ... ; Czarnecki, Skrzypek; MP '04;  
Friot, Greynat & de Rafael '05, Mohr, Taylor & Newell 2012

$$+ 130.8773 (61) (\alpha/\pi)^4$$

Kinoshita & Lindquist '81, ... , Kinoshita & Nio '04, '05;  
Aoyama, Hayakawa, Kinoshita & Nio, 2007, Kinoshita et al. 2012 & 2015;  
Lee, Marquard, Smirnov2, Steinhauser 2013 (electron loops, analytic),  
Kurz, Liu, Marquard, Steinhauser 2013 ( $\tau$  loops, analytic);  
Steinhauser et al. 2015 (electron loops, analytic); work in progress

$$+ 752.85 (93) (\alpha/\pi)^5 \text{ COMPLETED!}$$

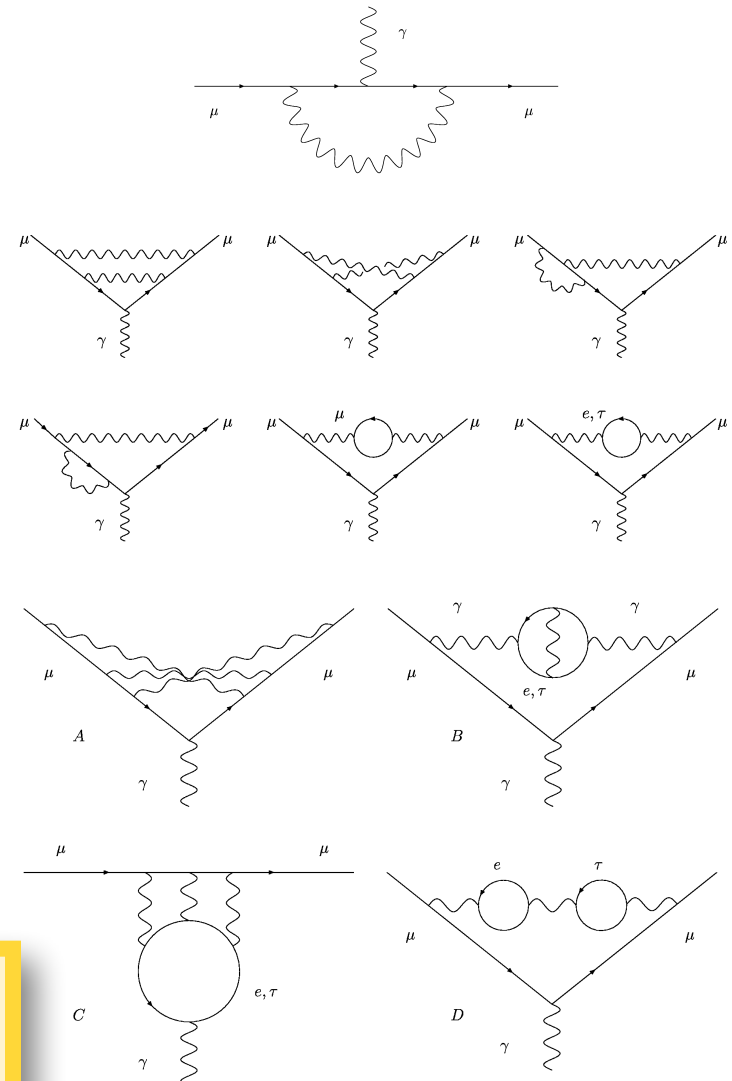
Kinoshita et al. '90, Yelkhovsky, Milstein, Starshenko, Laporta,  
Karshenboim, ..., Kataev, Kinoshita & Nio '06; Kinoshita et al. 2012 & 2015

Adding up, we get:

$$a_{\mu}^{\text{QED}} = 116584718.941 (21)(77) \times 10^{-11}$$

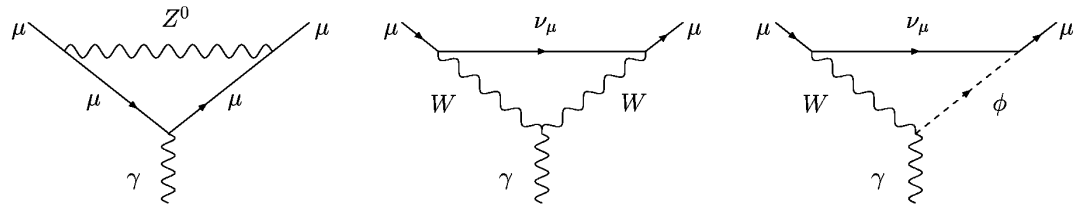
from coeffs, mainly from 4-loop unc  $\leftarrow$   $\rightarrow$  from  $\delta\alpha(\text{Rb})$

with  $\alpha=1/137.035999049(90)$  [0.66 ppb]





## One-loop term:



$$a_{\mu}^{\text{EW}}(1\text{-loop}) = \frac{5G_{\mu}m_{\mu}^2}{24\sqrt{2}\pi^2} \left[ 1 + \frac{1}{5} (1 - 4\sin^2\theta_W)^2 + O\left(\frac{m_{\mu}^2}{M_{Z,W,H}^2}\right) \right] \approx 195 \times 10^{-11}$$

1972: Jackiv, Weinberg; Bars, Yoshimura; Altarelli, Cabibbo, Maiani; Bardeen, Gastmans, Lautrup; Fujikawa, Lee, Sanda; Studenikin et al. '80s

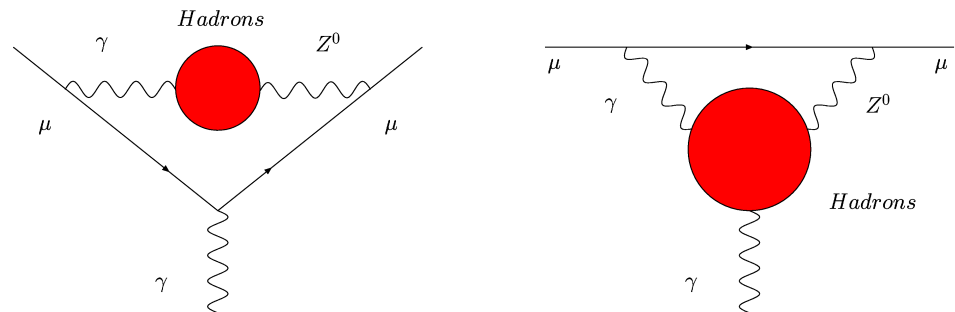
## One-loop plus higher-order terms:

$$a_{\mu}^{\text{EW}} = 153.6 (1) \times 10^{-11}$$

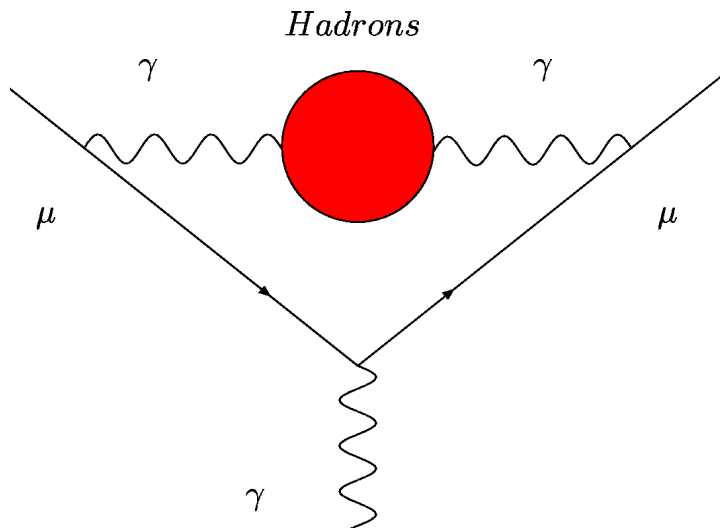
with  $M_{\text{Higgs}} = 125.6 (1.5) \text{ GeV}$

Hadronic loop uncertainties and 3-loop nonleading logs.

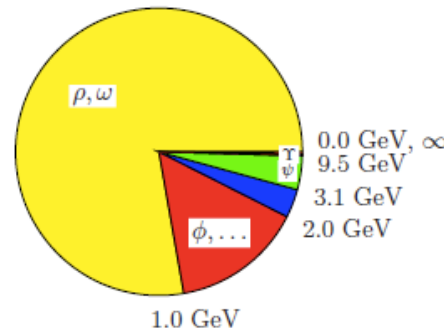
Kukhto et al. '92; Czarnecki, Krause, Marciano '95; Knecht, Peris, Perrottet, de Rafael '02; Czarnecki, Marciano and Vainshtein '02; Degrossi and Giudice '98; Heinemeyer, Stockinger, Weiglein '04; Gribouk and Czarnecki '05; Vainshtein '03; Gnendiger, Stockinger, Stockinger-Kim 2013.



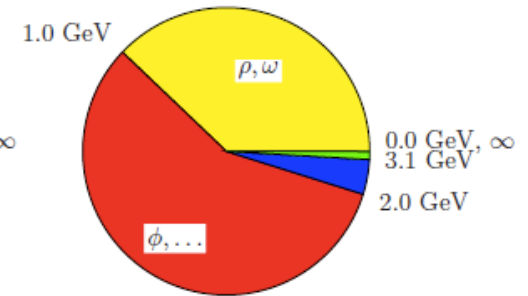
# The muon g-2: the hadronic LO contribution (HLO)



Central values



Errors<sup>2</sup>



F. Jegerlehner and A. Nyffeler, Phys. Rept. 477 (2009) 1

$$K(s) = \int_0^1 dx \frac{x^2(1-x)}{x^2 + (1-x)(s/m^2)}$$

$$a_\mu^{\text{HLO}} = \frac{1}{4\pi^3} \int_{4m_\pi^2}^\infty ds K(s) \sigma^{(0)}(s) = \frac{\alpha^2}{3\pi^2} \int_{4m_\pi^2}^\infty \frac{ds}{s} K(s) R(s)$$

$$a_\mu^{\text{HLO}} = 6903 (53)_{\text{tot}} \times 10^{-11}$$

F. Jegerlehner, A. Nyffeler, Phys. Rept. 477 (2009) 1

$$= 6923 (42)_{\text{tot}} \times 10^{-11}$$

Davier et al, EPJ C71 (2011) 1515 (incl. BaBar & KLOE10 2π)

$$= 6949 (37)_{\text{exp}} (21)_{\text{rad}} \times 10^{-11}$$

Hagiwara et al, JPG 38 (2011) 085003



Alternatively, exchanging the  $x$  and  $s$  integrations in  $a_\mu^{\text{HLO}}$  one gets:

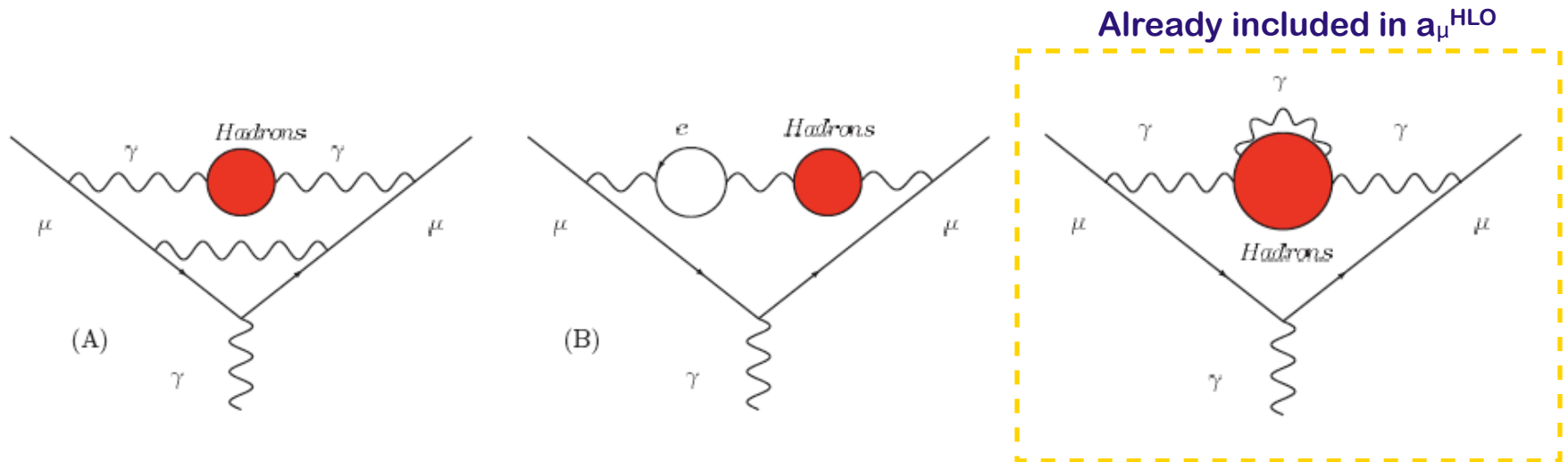
$$a_\mu^{\text{HLO}} = \frac{\alpha}{\pi} \int_0^1 dx (1-x) \Delta\alpha_{\text{had}}[t(x)] \quad t(x) = \frac{x^2 m_\mu^2}{x-1} < 0$$

involving the hadronic contribution to the running of  $\alpha$  in the **space-like region**, which can be extracted from Bhabha scattering data!



Requires measuring the Bhabha cross section at relatively small angles at better than  $10^{-4}$  accuracy. Challenging, but with dedicated work it may become feasible at flavor factories.

## • HNLO: Vacuum Polarization



$O(\alpha^3)$  contributions of diagrams containing hadronic vacuum polarization insertions:

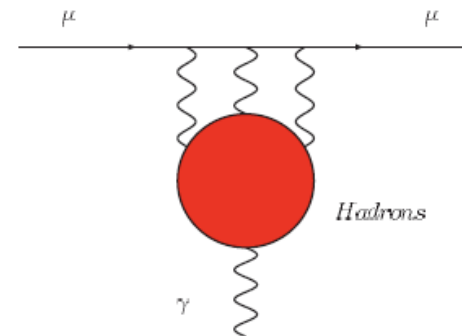
$$a_\mu^{\text{HNLO}}(\text{vp}) = -98 (1) \times 10^{-11}$$

Krause '96, Alemany et al. '98, Hagiwara et al. 2011

## ● HNLO: Light-by-light contribution

🔗 Unlike the HLO term, the hadronic l-b-l term relies at present on theoretical approaches.

🔗 This term had a troubled life! Latest values:



$a_{\mu}^{\text{HNLO}}( b ) = +80 (40) \times 10^{-11}$	Knecht & Nyffeler '02
$a_{\mu}^{\text{HNLO}}( b ) = +136 (25) \times 10^{-11}$	Melnikov & Vainshtein '03
$a_{\mu}^{\text{HNLO}}( b ) = +105 (26) \times 10^{-11}$	Prades, de Rafael, Vainshtein '09
$a_{\mu}^{\text{HHO}}( b ) = +116 (39) \times 10^{-11}$	Jegerlehner & Nyffeler '09

Results based also on Hayakawa, Kinoshita '98 & '02; Bijnens, Pallante, Prades '96 & '02

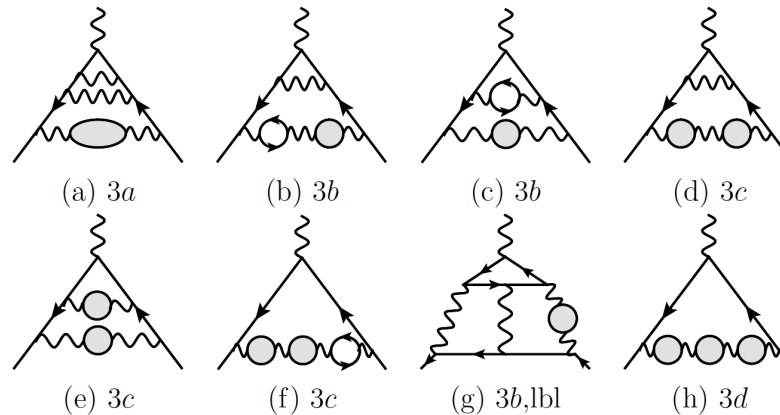
🔗 **“Bound”**  $a_{\mu}^{\text{HNLO}}(|b|) < \sim 160 \times 10^{-11}$  Erler, Sanchez '06, Pivovarov '02; also Boughezal, Melnikov '11

🔗 **Pion exch. in holographic QCD agrees.** D.K.Hong, D.Kim '09; Capiello, Catà, D'Ambrosio '11

🔗 **Lattice? Very hard but promising** Tom Blum et al. 2015

🔗 **Dispersive approach proposed** Colangelo, Hoferichter, Procura, Stoffer 1402.7081, 1506.01386

## HNNLO: Vacuum Polarization



$O(\alpha^4)$  contributions of diagrams containing hadronic vacuum polarization insertions:

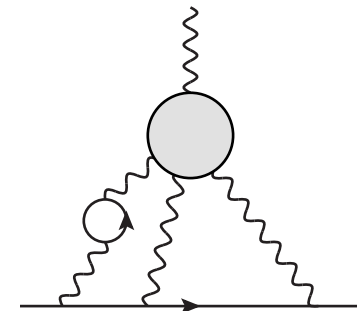
$$a_{\mu}^{\text{HNNLO}}(\text{vp}) = 12.4 (1) \times 10^{-11}$$

Kurz, Liu, Marquard, Steinhauser 2014

## HNNLO: Light-by-light

$$a_{\mu}^{\text{HNNLO}}(|b|) = 3 (2) \times 10^{-11}$$

Colangelo, Hoferichter, Nyffeler, MP, Stoffer 2014



Adding up all contributions, we get the following SM predictions and comparisons with the measured value:

$$a_\mu^{\text{EXP}} = 116592091 (63) \times 10^{-11}$$

E821 – Final Report: PRD73 (2006) 072 with latest value of  $\lambda=\mu_\mu/\mu_p$  from CODATA'10

$a_\mu^{\text{SM}} \times 10^{11}$	$\Delta a_\mu = a_\mu^{\text{EXP}} - a_\mu^{\text{SM}}$	$\sigma$
116 591 809 (66)	$282 (91) \times 10^{-11}$	3.1 [1]
116 591 829 (57)	$262 (85) \times 10^{-11}$	3.1 [2]
116 591 855 (58)	$236 (86) \times 10^{-11}$	2.8 [3]

with the “conservative”  $a_\mu^{\text{HNLO}}(|b|) = 116 (39) \times 10^{-11}$  and the LO hadronic from:

- [1] Jegerlehner & Nyffeler, Phys. Rept. 477 (2009) 1
- [2] Davier et al, EPJ C71 (2011) 1515 (includes BaBar & KLOE10  $2\pi$ )
- [3] Hagiwara et al, JPG38 (2011) 085003 (includes BaBar & KLOE10  $2\pi$ )

**Note that the th. error is now about the same as the exp. one**

- $\Delta a_\mu$  can be explained by errors in QED, EW, HNLO, g-2 EXP, HLO, or, we hope, by **New Physics!**
- Can  $\Delta a_\mu$  be due to **hypothetical mistakes** in the hadronic  $\sigma(s)$ ?
- An upward shift of  $\sigma(s)$  also induces an increase of  $\Delta\alpha_{\text{had}}^{(5)}(M_Z)$ .
- **Consider:**

$$\begin{aligned}
 a_\mu^{\text{HLO}} &\rightarrow a = \int_{4m_\pi^2}^{s_u} ds f(s) \sigma(s), & f(s) &= \frac{K(s)}{4\pi^3}, \quad s_u < M_Z^2, \\
 \Delta\alpha_{\text{had}}^{(5)} &\rightarrow b = \int_{4m_\pi^2}^{s_u} ds g(s) \sigma(s), & g(s) &= \frac{M_Z^2}{(M_Z^2 - s)(4\alpha\pi^2)},
 \end{aligned}$$

and the increase

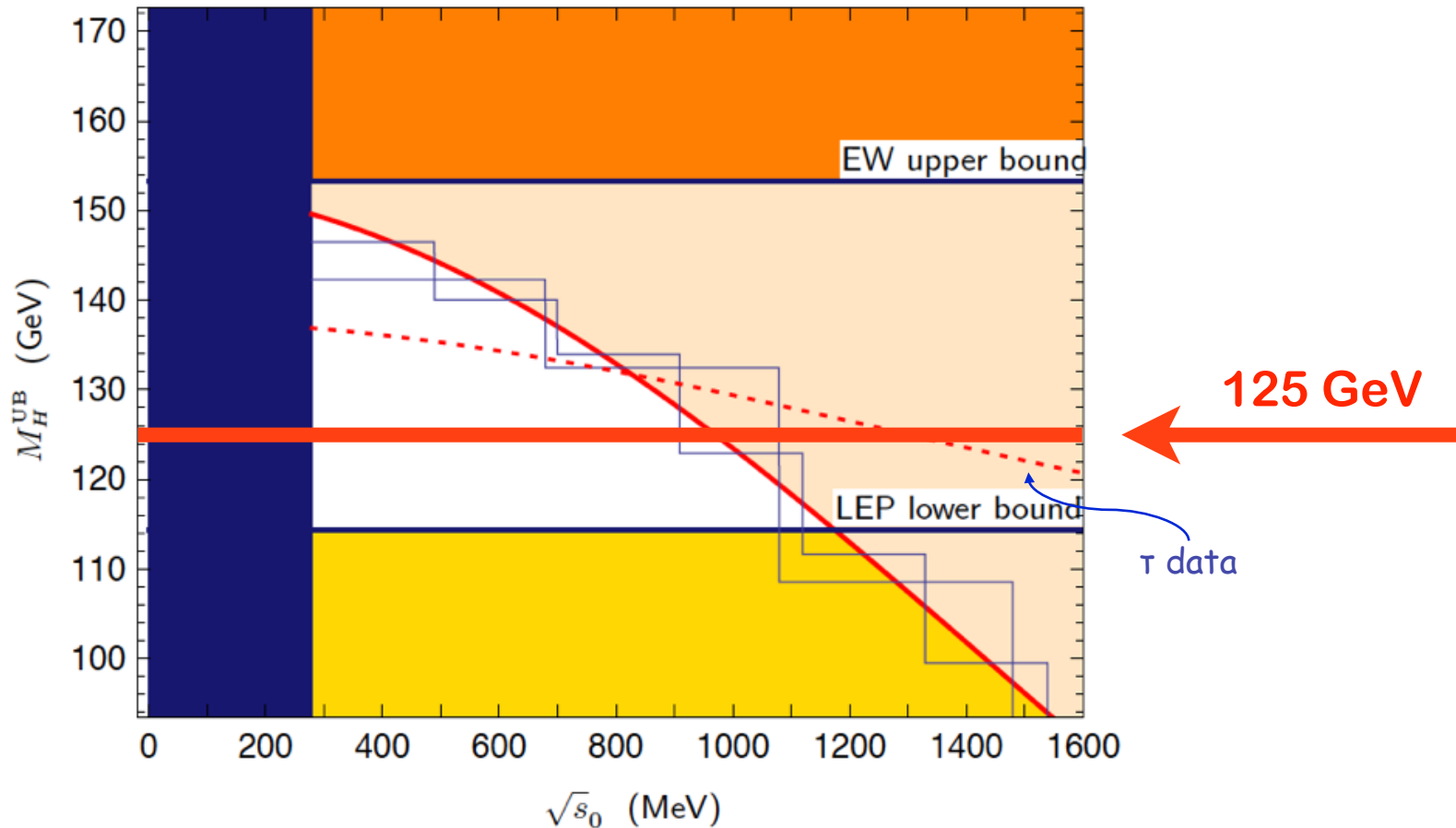
$$\Delta\sigma(s) = \epsilon\sigma(s)$$

( $\epsilon > 0$ ), in the range:




$$\sqrt{s} \in [\sqrt{s_0} - \delta/2, \sqrt{s_0} + \delta/2] \quad \longrightarrow$$



- How much does the  $M_H$  upper bound from the EW fit change when we shift  $\sigma(s)$  by  $\Delta\sigma(s)$  [and thus  $\Delta\alpha_{\text{had}}^{(5)}(M_Z)$ ] to accommodate  $\Delta a_\mu$  ?



W.J. Marciano, A. Sirlin, MP, 2008 & 2010

-  Given the quoted exp. uncertainty of  $\sigma(s)$ , the possibility to explain the muon g-2 with these very large shifts  $\Delta\sigma(s)$  appears to be very unlikely.
-  Also, given a 125 GeV SM Higgs, these hypothetical shifts  $\Delta\sigma(s)$  could only occur at very low energy (below  $\sim 1$  GeV) where  $\sigma(s)$  is precisely measured.
-  Vice versa, assuming we now have a SM Higgs with  $M_H = 125$  GeV, if we bridge the  $M_H$  discrepancy in the EW fit decreasing the low-energy hadronic cross section, **the muon g-2 discrepancy increases.**

# Limiting two-Higgs-doublet models with the muon $g-2$

A. Broggio, E.J. Chun, MP, K. Patel, S. Vempati

arXiv:1409.3199 (JHEP 2014)

## ● One-loop contribution:

$$\delta a_\mu^{2\text{HDM}}(1\text{loop}) = \frac{G_F m_\mu^2}{4\pi^2 \sqrt{2}} \sum_{j=h,H,A,H^\pm} (y_\mu^j)^2 r_\mu^j f_j(r_\mu^j)$$

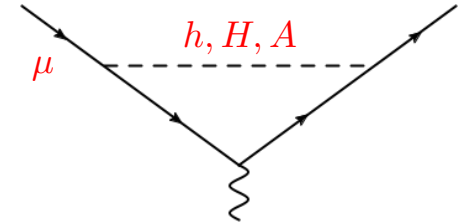
For  $r_\mu^j = m_\mu^2/M_j^2 \ll 1$ :

$$f_{h,H}(r) \sim -\ln r - 7/6 + O(r) > 0$$

$$f_A(r) \sim +\ln r + 11/6 + O(r) < 0$$

$$f_{H^\pm}(r) \sim -1/6 + O(r) < 0$$

roughly scales with  $m_\mu^4$ !



## ● Two-loop Barr-Zee type diagrams:

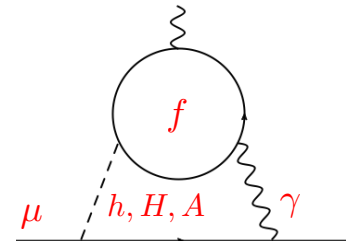
$$\delta a_\mu^{2\text{HDM}}(2\text{loop} - \text{BZ}) = \frac{G_F m_\mu^2}{4\pi^2 \sqrt{2}} \frac{\alpha_{\text{em}}}{\pi} \sum_{f; i=h,H,A} N_f^c Q_f^2 y_\mu^i y_f^i r_f^i g_i(r_f^i)$$

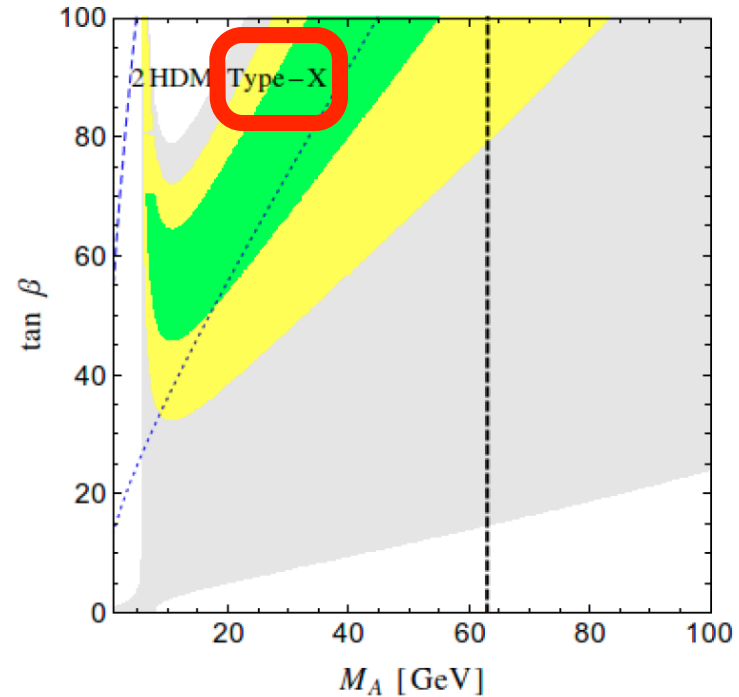
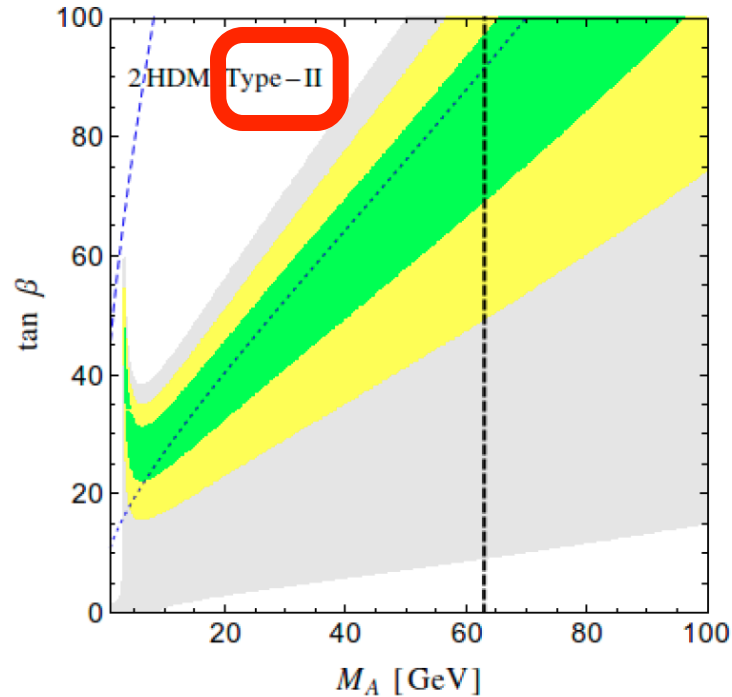
$m_f^2/m_\mu^2$  w.r.t. 1 loop

$$g_{h,H}(r) < 0$$

$$g_A(r) > 0$$

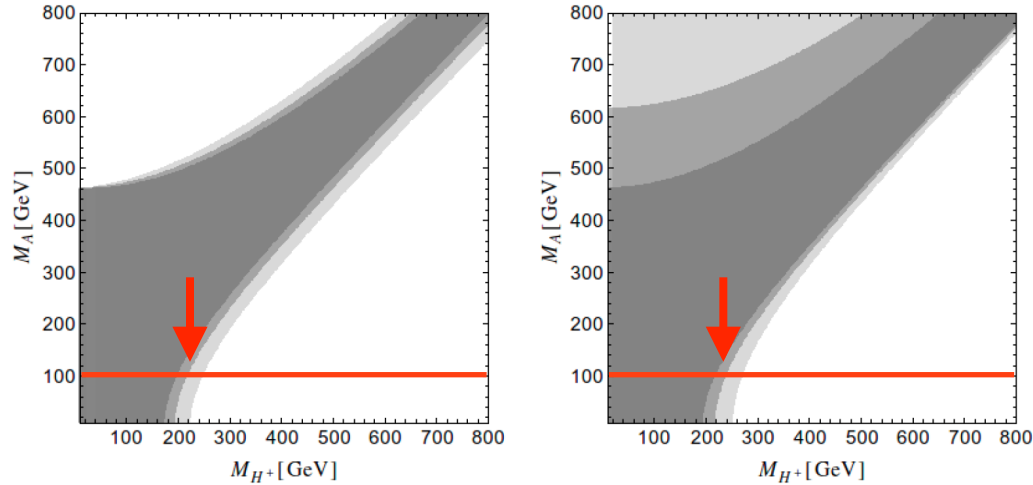
**In type II models, for  $M_A \gtrsim 3\text{GeV}$ :  
(2loop) $_A >$  (1loop) $_A$ . Similar for X.**





The  $1\sigma$ ,  $2\sigma$  and  $3\sigma$  regions allowed by  $\Delta a_\mu$  in the  $M_A$ - $\tan\beta$  plane taking the limit of  $\beta - \alpha = \pi/2$  and  $M_{h(H)} = 126$  (200) GeV in type II (left panel) and type X (right panel) 2HDMs. The regions below the dashed (dotted) lines are allowed at  $3\sigma$  ( $1.4\sigma$ ) by  $\Delta a_e$ . The vertical dashed line corresponds to  $M_A = M_h/2$ .

The contribution of the  $\tau$  loop is enhanced by a factor  $\tan^2\beta$  both in type II and in X models; it is suppressed by  $1/\tan^2\beta$  in models of type I and Y.

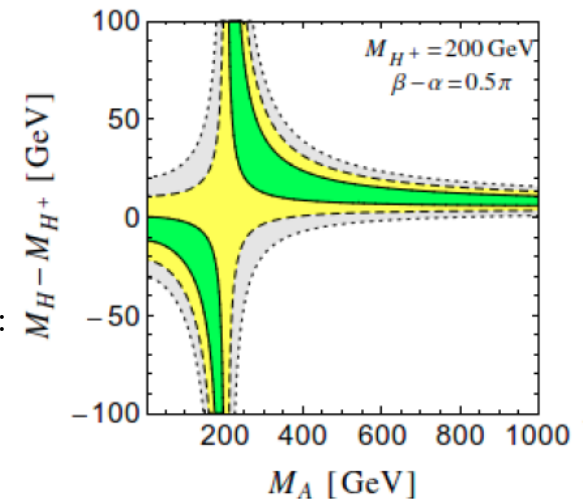


Vacuum stability and perturbativity constraints ( $\beta - \alpha = \pi/2$  and  $M_h = 126$  GeV). Left: allowed regions for  $\Delta M \equiv M_H - M_{H^\pm} = \{20, 0, -30\}$  GeV (darker to lighter),  $\lambda_{\max} = \sqrt{4\pi}$ . Right:  $\lambda_{\max} = \{\sqrt{4\pi}, 2\pi, 4\pi\}$ ,  $\Delta M = 0$ .  $\tan \beta = 50$ , but negligible change for  $\tan \beta \in [5, 100]$ .

**$M_A \lesssim 100$  GeV  $\rightarrow M_{H^\pm} \lesssim 200$  GeV**

**All values of  $M_A$  are allowed by EW precision tests if  $M_H \sim M_{H^\pm}$ .**

Regions allowed by EW precision constraints: green, yellow, gray for  $\Delta\chi_{\text{EW}}^2 < 2.3, 6.2, 11.8$ , i.e. 68.3, 95.4, 99.7% confidence intervals.



- Type I and Y models cannot account for the present value of  $\Delta a_\mu$  due to their lack of  $\tan^2\beta$  enhancements.
- In type II (and Y) models the  $\text{BR}(b \rightarrow s\gamma)$  sets a strong lower bound on  $M_{H^\pm}$  of order 380 GeV [Hermann, Misiak, Steinhauser 2012] → hardly any space left for the light A required by  $\Delta a_\mu$ .
- In type X models, no such strong bounds on  $M_{H^\pm}$  from  $\text{BR}(b \rightarrow s\gamma)$ , only model-indep. LEP bound  $M_{H^\pm} \gtrsim 80$  GeV.
- Therefore, out of type I, II, X, Y models, only **type X** is consistent with all the constraints we considered, provided that  $M_A \lesssim 100$  GeV,  $80 \lesssim M_{H^\pm} \lesssim 200$  GeV,  $M_H \sim M_{H^\pm}$ , large  $\tan\beta$ .  
Work in progress for additional constraints.

# Testing the SM with the electron $g-2$

G.F. Giudice, P. Paradisi & MP

arXiv:1208.6583 (JHEP 2012)



# The QED prediction of the electron g-2

$$a_e^{\text{QED}} = + (1/2)(\alpha/\pi) - 0.328\,478\,444\,002\,55(33)(\alpha/\pi)^2$$

Schwinger 1948 Sommerfeld; Petermann; Suura&Wichmann '57; Elend '66; CODATA Mar '12

$$A_1^{(4)} = -0.328\,478\,965\,579\,193\,78\dots$$

$O(10^{-18})$  in  $a_e$

$$A_2^{(4)}(m_e/m_\mu) = 5.197\,386\,68(26) \times 10^{-7}$$

$$A_2^{(4)}(m_e/m_\tau) = 1.837\,98(33) \times 10^{-9}$$

$$+ 1.181\,234\,016\,816(11)(\alpha/\pi)^3$$

$O(10^{-19})$  in  $a_e$

Kinoshita; Barbieri; Laporta, Remiddi, ... , Li, Samuel; MP '06; Giudice, Paradisi, MP 2012

$$A_1^{(6)} = 1.181\,241\,456\,587\dots$$

$$A_2^{(6)}(m_e/m_\mu) = -7.373\,941\,62(27) \times 10^{-6}$$

$$A_2^{(6)}(m_e/m_\tau) = -6.5830(11) \times 10^{-8}$$

$$A_3^{(6)}(m_e/m_\mu, m_e/m_\tau) = 1.909\,82(34) \times 10^{-13}$$

$$- 1.91206(84)(\alpha/\pi)^4$$

$0.2 \cdot 10^{-13}$  in  $a_e$

Kinoshita & Lindquist '81, ... , Kinoshita & Nio '05; Aoyama, Hayakawa, Kinoshita & Nio 2012 & 2015; Kurz, Liu, Marquard & Steinhauser 2014: analytic heavy virtual lepton part.

$$+ 7.79(34)(\alpha/\pi)^5 \quad \text{Complete Result! (12672 mass indep. diagrams!)}$$

Aoyama, Hayakawa, Kinoshita, Nio, PRL 109 (2012) 111807; PRD 91 (2015) 3, 033006

$0.2 \cdot 10^{-13}$  in  $a_e$  NB:  $(\alpha/\pi)^6 \sim O(10^{-16})$

The SM prediction is:

$$a_e^{\text{SM}}(\alpha) = a_e^{\text{QED}}(\alpha) + a_e^{\text{EW}} + a_e^{\text{HAD}}$$

The EW (1&2 loop) term is: Czarnecki, Krause, Marciano '96 [value from CODATA10]

$$a_e^{\text{EW}} = 0.2973(52) \times 10^{-13}$$

The Hadronic contribution, at LO+NLO+NNLO, is:

Nomura & Teubner '12, Jegerlehner & Nyffeler '09; Krause'97; Kurz, Liu, Marquard & Steinhauser 2014

$$a_e^{\text{HAD}} = 17.10(17) \times 10^{-13}$$

$$a_e^{\text{HLO}} = +18.66(11) \times 10^{-13}$$

$$a_e^{\text{HNLO}} = [-2.234(14)_{\text{vac}} + 0.39(13)_{\text{lbl}}] \times 10^{-13}$$

$$a_e^{\text{HNNLO}} = +0.28(1) \times 10^{-13}$$

Which value of  $\alpha$  should we use to compute  $a_e^{\text{SM}}$ ?

- The 2008 measurement of the electron g-2 is:

$$a_e^{\text{EXP}} = 11596521807.3 (2.8) \times 10^{-13} \quad \text{Hanneke et al, PRL100 (2008) 120801}$$

vs. old (factor of 15 improvement,  $1.8\sigma$  difference):

$$a_e^{\text{EXP}} = 11596521883 (42) \times 10^{-13} \quad \text{Van Dyck et al, PRL59 (1987) 26}$$

- Equate  $a_e^{\text{SM}}(\alpha) = a_e^{\text{EXP}}$  → best determination of alpha (2015):

$$\alpha^{-1} = 137.035\,999\,157 (33) \quad [0.24 \text{ ppb}]$$

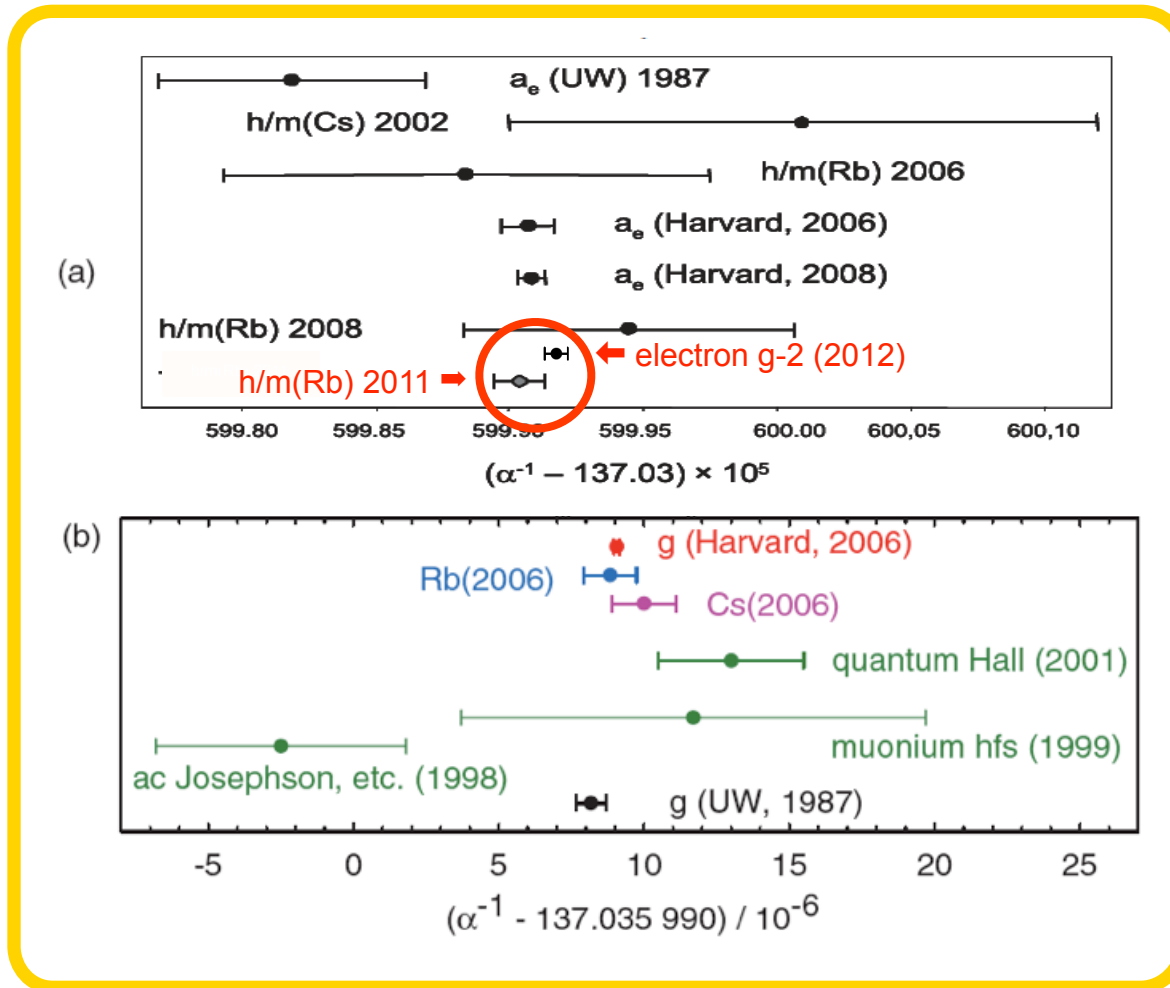
- Compare it with other determinations (independent of  $a_e$ ):

$$\alpha^{-1} = 137.036\,000\,0 (11) \quad [7.7 \text{ ppb}] \quad \text{PRA73 (2006) 032504 (Cs)}$$

$$\alpha^{-1} = 137.035\,999\,049 (90) \quad [0.66 \text{ ppb}] \quad \text{PRL106 (2011) 080801 (Rb)}$$

**Excellent agreement → beautiful test of QED at 4-loop level!**

# Old and new determinations of alpha



Gabrielse, Hanneke, Kinoshita, Nio & Odom, PRL99 (2007) 039902  
 Hanneke, Fogwell & Gabrielse, PRL100 (2008) 120801  
 Bouchendiria et al, PRL106 (2011) 080801

- Using  $\alpha = 1/137.035\,999\,049\,(90)$  [ $^{87}\text{Rb}$ , 2011], the SM prediction for the electron g-2 is

$$a_e^{\text{SM}} = 115\,965\,218\,16.5\,(0.2)\,(0.2)\,(0.2)\,(7.6) \times 10^{-13}$$

$\delta C_4^{\text{qed}}$     $\delta C_5^{\text{qed}}$     $\delta a_e^{\text{had}}$    from  $\delta\alpha$

- The EXP-SM difference is (note the negative sign):

$$\Delta a_e = a_e^{\text{EXP}} - a_e^{\text{SM}} = -9.2\,(8.1) \times 10^{-13}$$

The SM is in very good agreement with experiment ( $1\sigma$ ).  
 NB: The 4-loop contrib. to  $a_e^{\text{QED}}$  is  $-556 \times 10^{-13} \sim 70 \delta\Delta a_e!$   
 (the 5-loop one is  $6.2 \times 10^{-13}$ )

- The present sensitivity is  $\delta\Delta a_e = 8.1 \times 10^{-13}$ , ie ( $10^{-13}$  units):

$$(0.2)_{\text{QED4}}, \quad (0.2)_{\text{QED5}}, \quad (0.2)_{\text{HAD}}, \quad (7.6)_{\delta\alpha}, \quad (2.8)_{\delta a_e^{\text{EXP}}}$$

$$(0.4)_{\text{TH}} \leftarrow \text{may drop to } 0.2$$

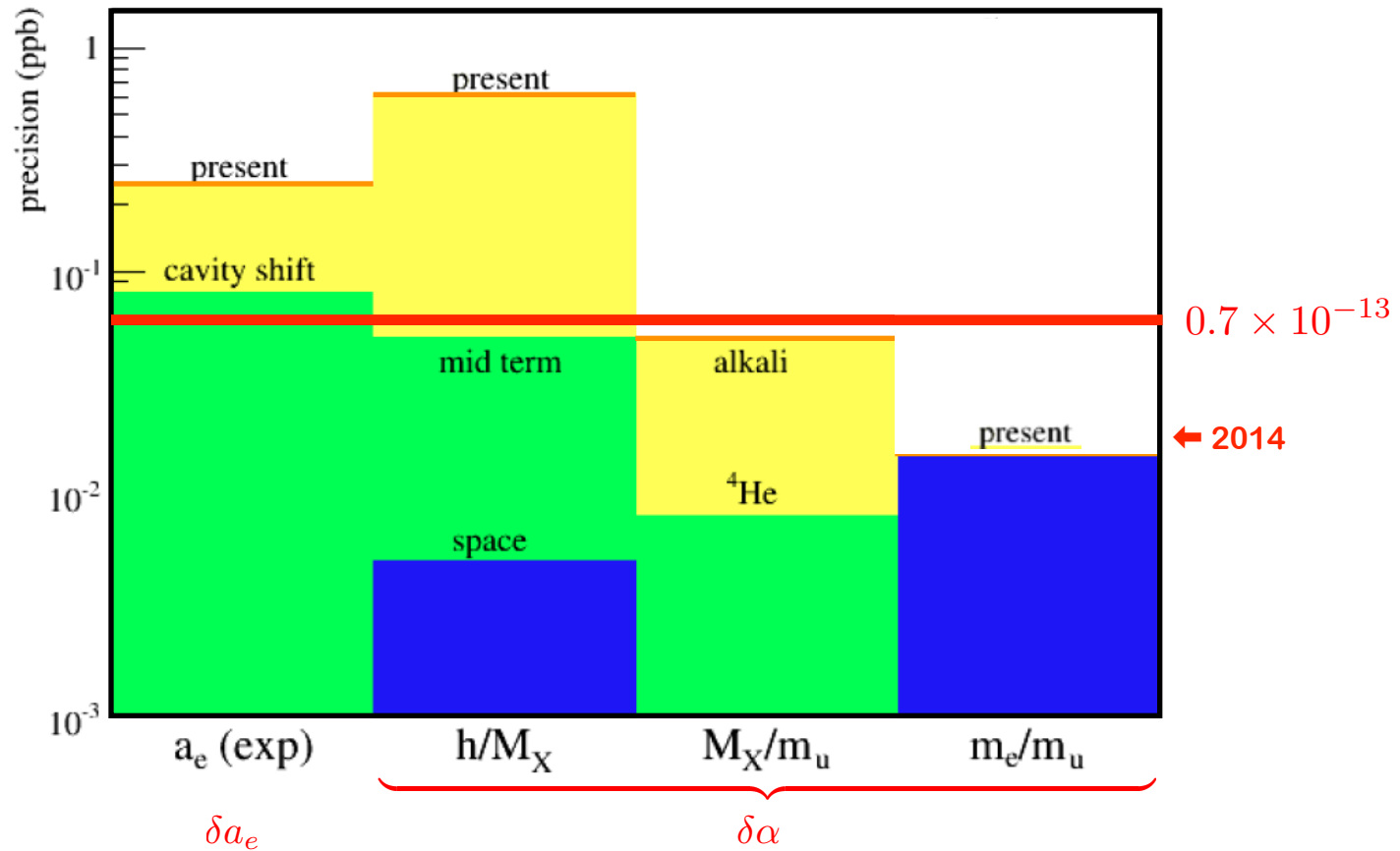
- The  $(g-2)_e$  exp. error may soon drop below  $10^{-13}$  and work is in progress for a significant reduction of that induced by  $\delta\alpha$ .

→ sensitivity of  $10^{-13}$  may be reached with ongoing exp. work

- In a broad class of BSM theories, contributions to  $a_l$  scale as

$$\frac{\Delta a_{\ell_i}}{\Delta a_{\ell_j}} = \left( \frac{m_{\ell_i}}{m_{\ell_j}} \right)^2 \quad \text{This Naive Scaling leads to:}$$

$$\Delta a_e = \left( \frac{\Delta a_\mu}{3 \times 10^{-9}} \right) 0.7 \times 10^{-13}; \quad \Delta a_\tau = \left( \frac{\Delta a_\mu}{3 \times 10^{-9}} \right) 0.8 \times 10^{-6}$$



Summary of the exp. contributions to the relative uncertainty of  $\delta \alpha$  and  $\delta a_e$  (in ppb).

F. Terranova & G.M. Tino, PRA89 (2014) 052118

- The experimental sensitivity in  $\Delta a_e$  is not very far from what is needed to **test if the discrepancy in  $(g-2)_\mu$  also manifests itself in  $(g-2)_e$**  under the naive scaling hypothesis.
- NP scenarios exist which **violate Naive Scaling**. They can lead to larger effects in  $\Delta a_e$  and contributions to EDMs, LFV or lepton universality breaking observables.
- Example: In the MSSM with non-degenerate but aligned sleptons (vanishing flavor mixing angles),  $\Delta a_e$  can reach  $10^{-12}$  (at the limit of the present exp sensitivity). For these values one typically has breaking effects of lepton universality at the few per mil level (within future exp reach).

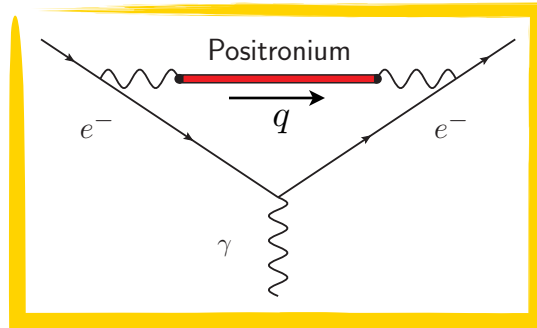


# Positronium contribution to the electron $g-2$

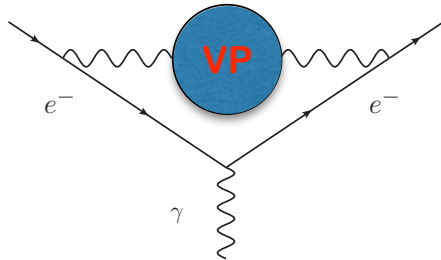
M. Fael & MP, [arXiv:1402.1575](https://arxiv.org/abs/1402.1575) (PRD 2014)

● The leading contribution of positronium to  $a_e$  comes from:

Mishima 1311.7109; Fael & MP 1402.1575; Melnikov et al. 1402.5690; Eides 1402.5860; Hayakawa 1403.0416



● The  $e^+e^-$  bound states appear as poles in the vac. pol.  $\Pi(q^2)$  right below the branch-point  $q^2 = (2m)^2$ . Their contribution is:

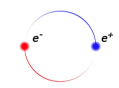


$$\Rightarrow a_e(\text{vp}) = \frac{\alpha}{\pi^2} \int_0^\infty \frac{ds}{s} \text{Im} \Pi(s + i\epsilon) K(s)$$

$$a_e^{\text{P}} = \frac{\alpha^5}{4\pi} \underbrace{\zeta(3) \left( 8 \ln 2 - \frac{11}{2} \right)}_{K(4m^2)} = 0.9 \times 10^{-13} = 1.3 \left( \frac{\alpha}{\pi} \right)^5$$

Mishima 1311.7109

● Of the same magnitude of the exp. unc. of  $a_e$  & the “naively rescaled” muon  $\Delta a_\mu$ . Of the same order of  $\alpha$  as the 5-loop term!



- Melnikov, Vainshtein & Voloshin (MVV) 1402.5690 determined a nonpert. contrib. of the  $e^+e^-$  continuum right above threshold that cancels **one-half** of  $a_e^P$ :

$$a_e(\text{vp})^{\text{cont,np}} = -\frac{|\alpha|^5}{8\pi} \zeta(3) \left( 8 \ln 2 - \frac{11}{2} \right)$$

- Check: the **total positronium poles + continuum** nonperturbative contribution to  $a_e$  arising from the threshold region at LO in  $\alpha$  is:

$$a_e^{\text{thr}}(\text{vp}) = -\frac{\alpha}{\pi} K(4m^2) \text{Re} A(1)$$

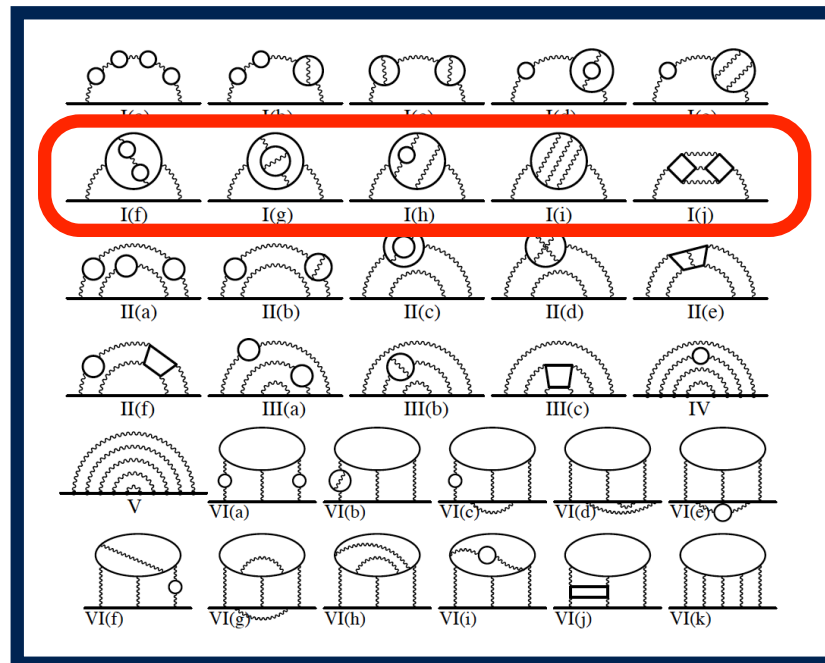
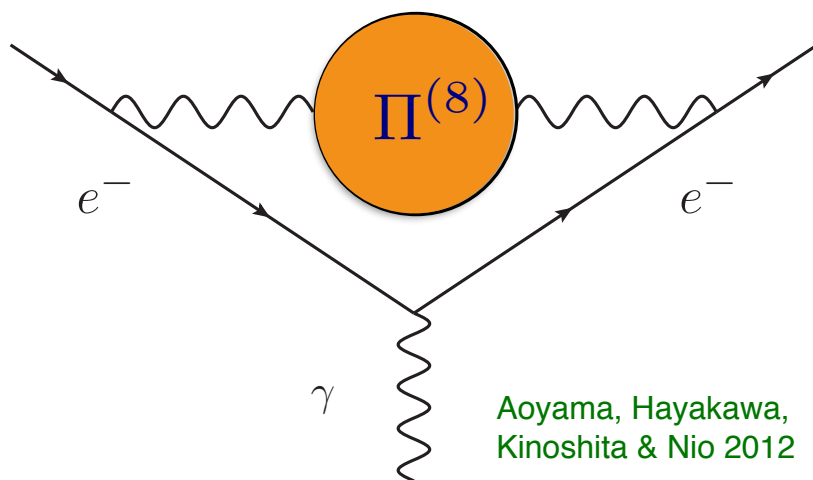
with

$$A(\beta) = -\frac{\alpha^2}{2} \left[ \gamma + \psi \left( 1 - \frac{i\alpha}{2\beta} \right) \right] = \frac{\alpha^2}{2} \sum_{k=1}^{\infty} \zeta(k+1) \left( \frac{i\alpha}{2\beta} \right)^k$$

so that

$$a_e^{\text{thr}}(\text{vp}) = \frac{\alpha^5}{8\pi} \zeta(3) K(4m^2) = \frac{a_e^P}{2}$$

- So, should we add this total threshold contribution  $a_e^P/2$  to the perturbative QED 5-loop result of Kinoshita and collaborators?
- Using the Coulomb Green's function, MVV 1402.5690 argued that it is already contained in the contribution of  $O(\alpha^5)$ .
- Hayakawa 1403.0416 claimed that positronium contributes to  $a_e$  only through a specific class of diagrams of  $O(\alpha^7)$ .
- To address this question we studied the 5-loop QED contribution to  $a_e$  arising from the insertion of the 4-loop VP in the photon line. This has been computed via:

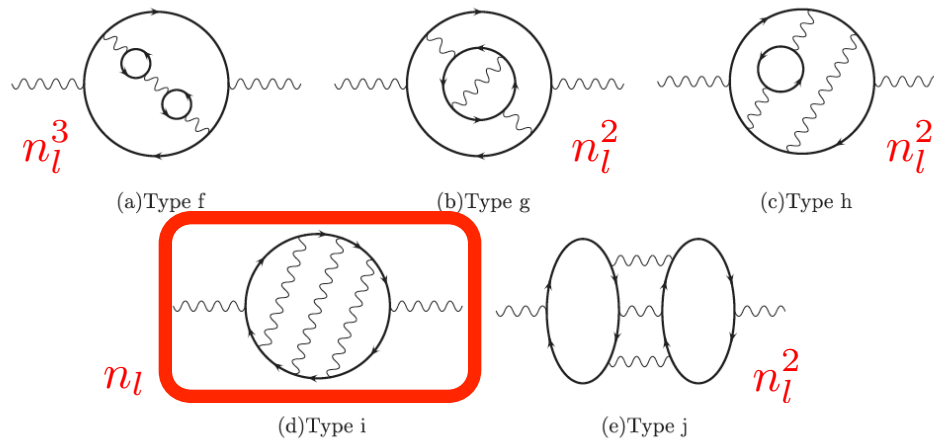


- Using explicit expressions for  $\Pi^{(8)}(q^2)$  (Baikov, Maier, Marquard '13) in

$$a_e^{(10)}(\text{vp}) = -\frac{\alpha}{\pi} \int_0^1 dx (1-x) \Pi^{(8)}\left(-\frac{m^2 x^2}{1-x}\right)$$

we obtain:

$$a_e^{(10)}(\text{vp}) = n_e \frac{\alpha^5}{8\pi} \zeta(3) K(4m^2) + \dots = \frac{a_e^{\text{P}}}{2} + \dots$$



- $a_e^{\text{P}}/2$  is already included in the 5-loop contribution of class I(i).
- There is no additional contrib of QED bound states beyond PT!

M.A. Braun 1968; Barbieri, Christillin, Remiddi 1973

# The tau g-2: opportunities or fantasies?

Work in progress in collaboration with  
S. Eidelman, D. Epifanov, M. Fael, L. Mercolli

arXiv:1301.5302  
arXiv:1310.1081  
arXiv:1506.03416

The Standard Model prediction of the tau g-2 is:

$$\begin{aligned}
 a_{\tau}^{\text{SM}} &= 117324 \quad (2) && \times 10^{-8} && \text{QED} \\
 &+ 47.4 \quad (0.5) && \times 10^{-8} && \text{EW} \\
 &+ 337.5 \quad (3.7) && \times 10^{-8} && \text{HLO} \\
 &+ 7.6 \quad (0.2) && \times 10^{-8} && \text{HHO (vac)} \\
 &+ 5 \quad (3) && \times 10^{-8} && \text{HHO (lbl)}
 \end{aligned}$$

$$a_{\tau}^{\text{SM}} = 117721 (5) \times 10^{-8}$$

Eidelman & MP  
2007

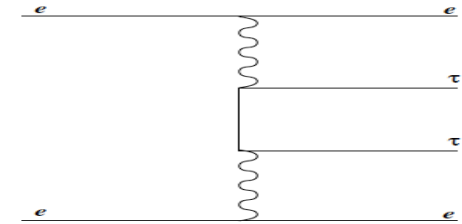
$(m_{\tau}/m_{\mu})^2 \sim 280$ : great opportunity to look for New Physics, and a “clean” NP test too...

	Muon	Tau
$a_{\text{EW}}/a_{\text{H}}$	1/45	1/7
$a_{\text{EW}}/\delta a_{\text{H}}$	3	10

... if only we could measure it!!

- The very short lifetime of the tau makes it very difficult to determine  $a_\tau$  measuring its spin precession in a magnetic field.

- DELPHI's result, from  $e^+e^- \rightarrow e^+e^-\tau^+\tau^-$  total cross-section measurements at LEP 2 (the PDG value):



$$a_\tau = -0.018 (17)$$

PDG 2014

- With an effective Lagrangian approach, using data on tau lepton production at LEP1, SLC, and LEP2:

$$-0.007 < a_\tau^{NP} < 0.005 \quad (95\% \text{ CL})$$

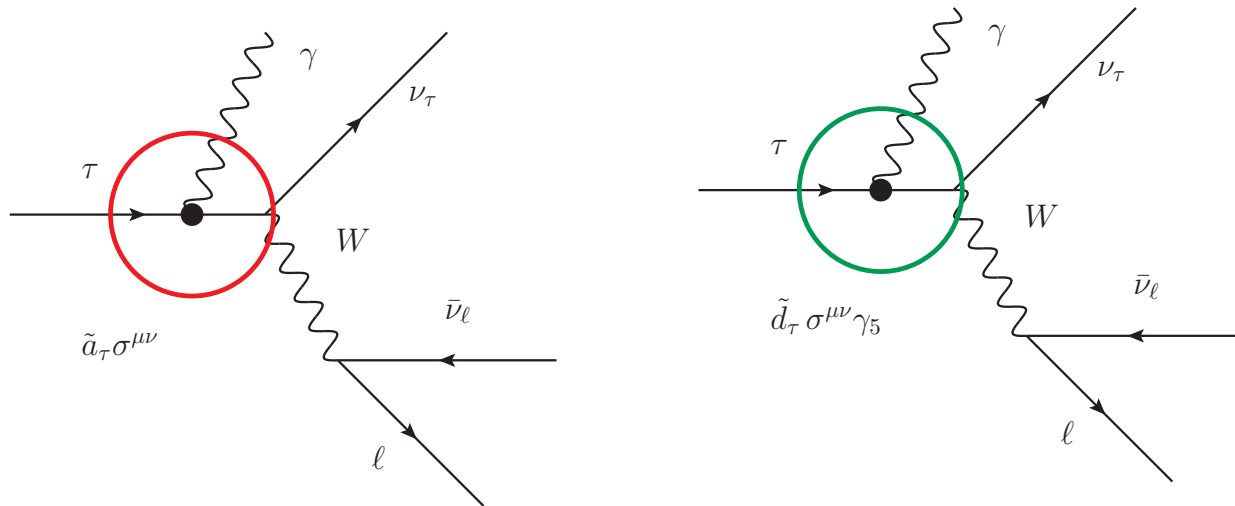
González-Sprinberg et al 2000

- Bernabéu et al, propose the measurement of  $F_2(q^2=M_\gamma^2)$  from  $e^+e^- \rightarrow \tau^+\tau^-$  production at B factories. NPB 790 (2008) 160



- ▶ Must employ the polarized differential decay rate.
- ▶ Study the full phase space.

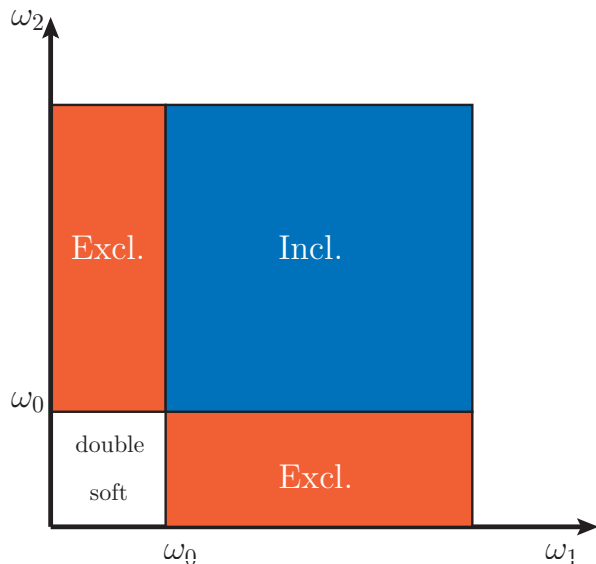
$$d\Gamma = d\Gamma_0 + \left(\frac{m_\tau}{M_W}\right)^2 d\Gamma_W + \frac{\alpha}{\pi} d\Gamma_{\text{NLO}} + \tilde{a}_\tau d\Gamma_a + \tilde{d}_\tau d\Gamma_d$$



- ▶ Under investigation for Belle-II analysis. Should at least improve the Delphi bound. Work in progress.

The branching ratio of radiative  $\mu$  and  $\tau$  leptonic decays for a minimum photon energy  $\omega_0$ :

$$\mathcal{B}(\omega_0) \propto \int d\Phi_4 (d\Gamma_{\text{LO}} + d\Gamma_{\text{virt}}) + \int d\Phi_5 d\Gamma_{\gamma\gamma}$$



- ▶  $\mathcal{B}^{\text{Exc}}(\omega_0)$ : only one  $\gamma$  of energy  $\omega > \omega_0$ , additional second soft photon  $\omega' < \omega_0$ .

$$\mathcal{B}^{\text{Exc}}(\omega_0) = \blacksquare$$

- ▶  $\mathcal{B}^{\text{Inc}}(\omega_0)$ : at least one  $\gamma$  of energy  $\omega > \omega_0$ .

$$\mathcal{B}^{\text{Inc}}(\omega_0) = \blacksquare + \blacksquare$$

B.R. of radiative  $\tau$  leptonic decays ( $\omega_0 = 10$  MeV)

	$\tau \rightarrow e\bar{\nu}\nu\gamma$	$\tau \rightarrow \mu\bar{\nu}\nu\gamma$
$\mathcal{B}_{\text{LO}}$	$1.834 \times 10^{-2}$	$3.663 \times 10^{-3}$
$\mathcal{B}_{\text{NLO}}^{\text{Inc}}$	$-1.06 (1)_n (10)_N \times 10^{-3}$	$-5.8 (1)_n (2)_N \times 10^{-5}$
$\mathcal{B}_{\text{NLO}}^{\text{Exc}}$	$-1.89 (1)_n (19)_N \times 10^{-3}$	$-9.1 (1)_n (3)_N \times 10^{-5}$
$\mathcal{B}^{\text{Inc}}$	$1.728 (10)_{\text{th}} (3)_{\tau} \times 10^{-2}$	$3.605 (2)_{\text{th}} (6)_{\tau} \times 10^{-3}$
$\mathcal{B}^{\text{Exc}}$	$1.645 (19)_{\text{th}} (3)_{\tau} \times 10^{-2}$	$3.572 (3)_{\text{th}} (6)_{\tau} \times 10^{-3}$
$\mathcal{B}_{\text{EXP}}^{\dagger}$	$1.847 (15)_{\text{st}} (52)_{\text{sy}} \times 10^{-2}$	$3.69 (3)_{\text{st}} (10)_{\text{sy}} \times 10^{-3}$

( $n$ ): numerical errors

( $N$ ): uncomputed NNLO corr.

$$\sim (\alpha/\pi) \ln r \ln(\omega_0/M) \times \mathcal{B}_{\text{NLO}}^{\text{Exc/Inc}}$$

$\dagger$  BABAR - PRD 91 (2015) 051103

(th): combined ( $n$ )  $\oplus$  ( $N$ )

( $\tau$ ): experimental error of  $\tau$

lifetime:  $\tau_{\tau} = 2.903(5) \times 10^{-13}$  s

	$\tau \rightarrow e\bar{\nu}\nu\gamma$	$\tau \rightarrow \mu\bar{\nu}\nu\gamma$
$\Delta^{\text{Exc}}$	$2.02 (57) \times 10^{-3} \rightarrow 3.5\sigma$	$1.2 (1.0) \times 10^{-4} \rightarrow 1.1\sigma$

# Conclusions

- The muon discrepancy is  $\Delta a_\mu \sim 3 \div 3.5 \sigma$ . Is it NP? New upcoming g-2 experiment: QED & EW terms ready for the challenge; how about the hadronic one? Future of hadronic LBL??
- Could  $\Delta a_\mu$  be due to **mistakes in the hadronic  $\sigma(s)$** ? Very unlikely. Also, given a 125 GeV SM Higgs, these hypothetical shifts  $\Delta\sigma(s)$  could only occur at very low energies (below  $\sim 1\text{GeV}$ ).
- Can  $\Delta a_\mu$  be solved by **2HDMs**? Not by type I, II, and Y. Type **X** is still allowed by all the constraints we considered if  $M_A \lesssim 100\text{GeV}$ ,  $80 \lesssim M_{H^\pm} \lesssim 200\text{GeV}$ ,  $M_H \sim M_{H^\pm}$  & large  $\tan\beta$ .
- The sensitivity of the electron g-2 has improved. It may soon be possible to **test if  $\Delta a_\mu$  manifests itself also in the electron g-2!** A robust and ambitious experimental program is under way to improve  $\alpha$  &  $a_e$ . **The positronium contribution shouldn't be added!**
- The tau g-2 is essentially unknown: we propose to measure it at Belle II via its radiative leptonic decays. BaBar's recent precise measurement of  $\mathcal{B}(\tau \rightarrow e\bar{\nu}\nu\gamma)$  differs from our prediction by  **$3.5\sigma$** !

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**The End**