



Quantum Gravity, from the atoms of space to cosmology

The Universe as a Quantum Condensate

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Quantum spacetime: the difficult path from microstructure to macrophysics

Quantum Gravity problem:

identify microscopic d.o.f. of quantum spacetime and their fundamental dynamics

derive effective (QG-inspired) models for macroscopic continuum physics: explain features of early Universe, obtain testable QG predictions

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identify microscopic d.o.f. of quantum spacetime and their fundamental dynamics

various approaches: group field theory, loop quantum gravity

derive effective (QG-inspired) models for (quantum) cosmology: explain features of early Universe, obtain testable QG predictions

various models: loop quantum cosmology,

task is daunting (compare with analogue problem in condensed matter theory)

Matrix models



Tensor models











Quantum field theories over group manifold G (or corresponding Lie algebra) $\varphi:G^{ imes d} o\mathbb{C}$

relevant classical phase space for "GFT quanta":

 $\left(\mathcal{T}^*G\right)^{\times d} \simeq \left(\mathfrak{g} \times G\right)^{\times d}$

can reduce to subspaces in specific models depending on conditions on the field

d is dimension of "spacetime-to-be"

example: d=4 $\varphi(g_1, g_2, g_3, g_4) \leftrightarrow \varphi(B_1, B_2, B_3, B_4) \rightarrow \mathbb{C}$

can be defined for any (Lie) group and dimension d, any signature,

very general framework; interest rests on specific models/use

Fock vacuum: "no-space" ("emptiest") state | 0 >

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single field "quantum": spin network vertex or tetrahedron ("building block of space")





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generic quantum state: arbitrary collection of spin network vertices (including glued ones) or tetrahedra (including glued ones)

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classical action: kinetic (quadratic) term + (higher order) interaction (convolution of GFT fields)

$$S(\varphi,\overline{\varphi}) = \frac{1}{2} \int [dg_i] \overline{\varphi(g_i)} \mathcal{K}(g_i) \varphi(g_i) + \frac{\lambda}{D!} \int [dg_{ia}] \varphi(g_{i1}) \dots \varphi(\overline{g}_{iD}) \mathcal{V}(g_{ia},\overline{g}_{iD}) + c.c.$$

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combinatorics of field arguments in interaction: gluing of 5 tetrahedra across common triangles, to form 4-simplex ("building block of spacetime")

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= stranded diagrams dual to cellular complexes of arbitrary topology

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$$S_{int}[\varphi_{\ell}] = \lambda \int [dg_i]^6 \varphi_1(g_1, g_2, g_3) \varphi_2(g_3, g_4, g_5) \varphi_3(g_5, g_2, g_6) \varphi_4(g_6, g_4, g_1) + \lambda \int [dg_i]^6 \overline{\varphi_4}(g_1, g_4, g_6) \overline{\varphi_3}(g_6, g_2, g_5) \overline{\varphi_2}(g_5, g_4, g_3) \overline{\varphi_1}(g_3, g_2, g_1)$$

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discretization of: $S(e, \omega) = \int Tr(e \wedge F(\omega))$



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spin foam formulation of 3d gravity/BF theory

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$$\underbrace{\text{discrete 1st order path integral for 3d gravity/BF theory on simplicial complex dual to GFT Feynman diagram}$$

spin foam formulation of 3d gravity/BF theory

GFT models of 4d gravity:

based on classical (Plebanski) formulation of GR as BF theory + (simplicity) constraints

start from GFT formulation of 4d BF theory

+ impose simplicity constraints (geometricity of simplicial structures)

(Barbieri, Baez, Barrett, Crane, Reisenberger, Perez, De Pietri, Engle, Pereira, Freidel, Krasnov, Rovelli, Livine, Speziale, Baratin, DO,)

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second quantized version of Loop Quantum Gravity but dynamics not derived from canonical quantization of GR

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QFT methods (i.e. GFT reformulation of LQG and spin foam models) useful to address physics of large numbers of LQG d.o.f.s, i.e. many and refined graphs (continuum limit)

(superpositions of "many-vertices" states, refinement as creation of new vertices, etc)

1. making sense of quantum dynamics and LQG partition function (correlations)

- 2. understanding LQG phase structure
- 3. extracting effective continuum dynamics

1st message

we have a solid candidate formalism for a theory of quantum gravity (a QFT for the "atoms of quantum space")

grounded in LQG (and discrete gravity, tensor models)

rigorous mathematics, clear pre-geometric meaning promising fundamental dynamical models

lots of results

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various models: loop quantum cosmology,

task is daunting (imagine analogue problem in condensed matter theory)

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GFTs are a formulation of LQG/spin foams that is most suited to tackle this problem, thanks to QFT tools









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for a non-spatio-temporal QG system (e.g. LQG in GFT formulation),

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Renormalization Group is crucial tool (mathematical, conceptual, physical)

renormalization is not about "curing or hiding divergences", but taking into account the physics of more and more d.o.f.s

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in specific GFT case:

• fundamental formulation of QG = QFT, defined perturbatively around "no-space" (degenerate) vacuum

need to prove consistency of the theory: perturbative GFT renormalizability

ned to understand effective dynamics at different "GFT scales": RG flow of effective actions & phase structure & phase transitions

Geometrogenesis in LQG/GFT

idea of "geometrogenesis" in LQG/GFT : starting from degenerate phase, continuum geometric physics in new phase new phase can be "condensate" phase of QG "atoms of space"

in canonical LQG context: T. Koslowski, 0709.3465 [gr-qc] in covariant SF/GFT context: DO, 0710.3276 [gr-qc] also in tensor models V. Rivasseau, '13
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some experience and results in tensor models and GFTs

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other phases and phase transition -not- physical, just formal: theory makes sense only in geometric phase this is point of view in CDT (J. Ambjorn, R. Loll, ...), but see J. Mielczarek, '14

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natural hypothesis: very early Universe - big bang as QG phase transition

GFT is QG analogue of QFT for atoms in condensed matter system

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 - continuum spacetime as a peculiar quantum fluid
 - more specific hypothesis: continuum spacetime is GFT condensate GR-like dynamics from GFT condensate hydrodynamics
 - simple candidates for physical (geometric) vacuum: GFT condensates

what is their definition? do they have a continuum geometric interpretation? what is their effective quantum dynamics? does it relate to GR?

> DO, L. Sindoni, '10; S. Gielen, DO, L. Sindoni, 1303.3576 [gr-qc], 1311.1238 [gr-qc]; S. Gielen, '14; G. Calcagni, '14; L. Sindoni, '14; S. Gielen, DO, '14

canonical LQG:

purely kinematical, inequivalent representations (phases) of quantum algebra of observables



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spin foam models (without GFT framework)

- rewrite them as lattice gauge theory path integrals
- define (background independent) coarse graining procedure
- look for flow of effective actions and fixed points

technically (numerically) very challenging

many results, mainly in simplified models (simpler algebraic data, dimensionally reduced models)

work by:

B. Bahr, B. Dittrich, F. Eckert, F. Hellmann, W. Kaminski, M. Martin-Benito, S. Steinhaus, - '09-'15

power counting and radiative corrections in GFT models (cut-off of fields in representation space)

topological simplicial GFT models (BF theory):

partial power counting and scaling theorems - large-N scaling

L. Freidel, R. Gurau, DO, '09; J. Magnen et al., '09; J. Ben Geloun, J. Magnen, V. Rivasseau, '10; R. Gurau, '11; S. Carrozza, DO, '11,'12

radiative corrections of 2-point function: need for Laplacian kinetic term

J. Ben Geloun, V. Bonzom, '11

super-renormalizability in abelian case (3d, with Laplacian)

J. Ben Geloun, '13

4d gravity models

super-renormalizability of some versions of BC model

A. Perez, C. Rovelli, '00, '01

radiative corrections of 2-point function in EPRL-FK model

T. Krajewski, J. Magnen, V. Rivasseau, A. Tanasa, P. Vitale, '10; A. Riello, '13

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- Functional Renormalization Group for TGFTs

D. Benedetti, J. Ben Geloun, DO, '14; J. Ben Geloun, R.Martini, DO, '15

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 $\begin{array}{rcl} & \text{D. Benedetti, J. Ben Geloun, DO, '14} \\ \hline \textbf{an example:} & \varphi(g_1, g_2, g_3) = \sum_{p_1, p_2, p_3} \varphi_{p_1 p_2 p_3} e^{ip_1 \theta_1} e^{ip_2 \theta_2} e^{ip_3 \theta_3} \in \mathbb{R} & g_i = e^{i\theta_i} \in U(1) \quad \theta_i \in [-\pi, \pi) \quad p_i \in \mathbb{Z} \\ \hline \Gamma_N(\varphi) & = & \frac{Z_N}{2} \operatorname{Tr}_2(\varphi \cdot K \cdot \varphi) + \frac{m_N}{2} \operatorname{Tr}_2(\varphi^2) + S^{\operatorname{int}} & \Delta S_N(\phi) & = & \frac{1}{2} \operatorname{Tr}(\phi \cdot R_N \cdot \phi) \\ S'^{\operatorname{int}} & = & \frac{\lambda_N}{4} \Big(\operatorname{Tr}_{4;1}(\varphi^4) + \operatorname{Sym}(1 \to 2 \to 3) \Big) & R_N(\{p_i\}; \{p'_i\}) = \delta_{p_i, p'_i} Z_N \left(N - \frac{1}{3} \sum_{i=1}^3 p_i \right) \Theta(N - \frac{1}{3} \sum_{i=1}^3 p_i) \\ \operatorname{Tr}_2(\varphi \cdot K \cdot \varphi) & = & \sum_{p_i \in \mathbb{N}} \varphi_{123}(\frac{1}{3} \sum_i p_i) \varphi_{123} \\ \operatorname{Tr}_2(\varphi^2) & = & \sum_{p_i \in \mathbb{N}} \varphi_{123}^2 \varphi_{1'2'3'} \varphi_{1'2'3'} \\ \operatorname{Tr}_{4;1}(\varphi^4) & = & \sum_{p_i, p'_i \in \mathbb{N}} \varphi_{123} \varphi_{1'2'3'} \varphi_{1'2'3'} \\ \end{array}$

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flow equations:

$$\partial_t \mu_N = -\mu_N \eta - \frac{\lambda_N N}{(N+\mu_N)^2} \left\{ \frac{9}{2} (3N+2)(N+1) + \frac{\eta}{2} \left(11 + 18N + 9N^2 \right) \right\}$$

$$\partial_t \lambda_N = -\frac{1}{3} \lambda_N^2 \frac{\eta}{(N+\mu_N)^3} N(3N+1)(6N+13) + 9\lambda_N^2 \frac{(\eta+1)}{(N+\mu_N)^3} N(N+1)(N+2) - \frac{1}{3} \lambda_N^2 \frac{\eta}{(N+\mu_N)^3} N(3N+1)(6N+13) + 9\lambda_N^2 \frac{(\eta+1)}{(N+\mu_N)^3} N(N+1)(N+2) - \frac{1}{3} \lambda_N^2 \frac{\eta}{(N+\mu_N)^3} N(3N+1)(6N+13) + 9\lambda_N^2 \frac{(\eta+1)}{(N+\mu_N)^3} N(N+1)(N+2) - \frac{1}{3} \lambda_N^2 \frac{\eta}{(N+\mu_N)^3} N(3N+1)(6N+13) + 9\lambda_N^2 \frac{(\eta+1)}{(N+\mu_N)^3} N(N+1)(N+2) - \frac{1}{3} \lambda_N^2 \frac{\eta}{(N+\mu_N)^3} N(3N+1)(6N+13) + 9\lambda_N^2 \frac{(\eta+1)}{(N+\mu_N)^3} N(N+1)(N+2) - \frac{1}{3} \lambda_N^2 \frac{\eta}{(N+\mu_N)^3} N(3N+1)(6N+13) + 9\lambda_N^2 \frac{(\eta+1)}{(N+\mu_N)^3} N(N+1)(N+2) - \frac{1}{3} \lambda_N^2 \frac{\eta}{(N+\mu_N)^3} N(3N+1)(6N+13) + 9\lambda_N^2 \frac{(\eta+1)}{(N+\mu_N)^3} N(N+1)(N+2) - \frac{1}{3} \lambda_N^2 \frac{\eta}{(N+\mu_N)^3} N(3N+1)(6N+13) + 9\lambda_N^2 \frac{(\eta+1)}{(N+\mu_N)^3} N(N+1)(N+2) - \frac{1}{3} \lambda_N^2 \frac{\eta}{(N+\mu_N)^3} N(3N+1)(6N+13) + 9\lambda_N^2 \frac{(\eta+1)}{(N+\mu_N)^3} N(N+1)(N+2) - \frac{1}{3} \lambda_N^2 \frac{\eta}{(N+\mu_N)^3} N(3N+1)(6N+13) + 9\lambda_N^2 \frac{(\eta+1)}{(N+\mu_N)^3} N(N+1)(N+2) - \frac{1}{3} \lambda_N^2 \frac{\eta}{(N+\mu_N)^3} N(3N+1)(6N+13) + 9\lambda_N^2 \frac{(\eta+1)}{(N+\mu_N)^3} N(N+1)(N+2) - \frac{1}{3} \lambda_N^2 \frac{\eta}{(N+\mu_N)^3} N(3N+1)(6N+13) + 9\lambda_N^2 \frac{(\eta+1)}{(N+\mu_N)^3} N(N+1)(N+2) - \frac{1}{3} \lambda_N^2 \frac{\eta}{(N+\mu_N)^3} N(3N+1)(6N+13) + \frac{1}{3} \lambda_N^2 \frac{(\eta+1)}{(N+\mu_N)^3} N(N+1)(N+2) - \frac{1}{3} \lambda_N^2 \frac{(\eta+1)}{(N+\mu_N)^3} N(N+1) + \frac{1}{3} \lambda_N^2 \frac{(\eta+1)}{(N+\mu_N)^3} N(N+1) + \frac{1}{3} \lambda_N^2 \frac{(\eta+1)}{(N+\mu_N)^3} N(N+1) + \frac{1}{3} \lambda_N^2$$

non-autonomous system (due to external scale a = size of grow manifold)

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2nd message

problem of the continuum in QG (GFT,LQG): crucial to connect to macrophysics, difficult and open

we are addressing it

RG is crucial tool, different strategies, many results (renormalizable models, RG flows,...)

QG phase transition (condensation?) could be physical

cosmological interpretation: realization of "Emergent Spacetime" and of "Universe as a Condensate" ideas
Quantum spacetime: the difficult path from microstructure to cosmology

Quantum Gravity problem:

identify microscopic d.o.f. of quantum spacetime and their fundamental dynamics

derive effective (QG-inspired) models for fundamental (quantum) cosmology: explain features of early Universe, obtain testable QG predictions

various models: loop quantum cosmology,

task is daunting (imagine analogue problem in condensed matter theory)

S. Gielen, DO, L. Sindoni, PRL, arXiv:1303.3576 [gr-qc]; JHEP, arXiv:1311.1238 [gr-qc]

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problem 2: extract from fundamental theory an effective macroscopic dynamics for such states

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Quantum GFT condensates

a simple choice of quantum GFT condensate (homogeneous continuum quantum space)

> other constructions possible, depending on how much information the condensate state has to encode (in a coarse-grained form) S. Gielen, DO, L. Sindoni, '13; DO, D. Pranzetti, J. Ryan, L. Sindoni, '15

various procedures for estimating validity of chosen ansatz for vacuum state, e.g. L. Sindoni, arXiv:1408.3095 [gr-qc]

where due to (1) and $[\hat{\varphi}^{\dagger}(g_I), \hat{\varphi}^{\dagger}(h_I)] = 0$ the function ξ can be taken to satisfy $\xi(g_I) = \xi(kg_Ik')$ for all k, k' in SU(2) and $\xi(g_I) = \xi(g_I^{-1})$. ξ is a function on the gauge-

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Fixing a G-invariant inner prod-Assuming that the simplicity constraints have been im- \mathfrak{g} this basis is unique up to the plemented by (6), \mathfrak{G} is a field on $SU(2)^4$ and we require w demand that the *embedded tetra*-9 this additional symmetry under the action of SU(2). It can be imposed on a one-particle state created by S. Gielen, DO, L. Sindoni, the left-invariant vector fields, follow closely procedure used in real BECs $(m) = \mathbf{e}_i(x_m),$ (1) single-particle GFT condensate: $\hat{\sigma} := \int d^4g \ \sigma(g_I) \hat{\varphi}^{\text{FRL, arXiv:1303.3576}}_{\text{JHEP, arXiv:1311.1238 [gr-qc]}};$ (14)

of the physical metric now reads

$$(x_m)(\mathbf{e}_i(x_m), \mathbf{e}_j(x_m)), \qquad (15)$$

metric components in the frame homogeneous metric will be one ents. We can then say that a distetrahedra, specified by the data ith spatial homogeneity if

$$_{j(k)} \quad \forall k, m = 1, \dots, N.$$
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> A second possibility is to use a two-particle operator which automatically has the required gauge invariance:

$$\hat{\xi} := \frac{1}{2} \int d^4g \, d^4h \, \xi(g_I h_I^{-1}) \hat{\varphi}^{\dagger}(g_I) \hat{\varphi}^{\dagger}(h_I), \qquad (18)$$

where due to (1) and $[\hat{\varphi}^{\dagger}(g_I), \hat{\varphi}^{\dagger}(h_I)] = 0$ the function ξ can be taken to satisfy $\xi(g_I) = \xi(kg_Ik')$ for all k, k' in SU(2) and $\xi(q_I) = \xi(q_I^{-1})$. ξ is a function on the gaugeinvariant configuration space of a single tetrahedron.

We then consider two types of candidate states for macroscopic, homogeneous configurations of tetrahedra:

$$|\sigma\rangle := \exp(\hat{\sigma}) |0\rangle, \quad |\xi\rangle := \exp(\hat{\xi}) |0\rangle.$$
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S. Gielen, DO, arXiv:1407.8167 [gr-qc]

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microscopic (single vertex, 1st quantized) variables:

$$\begin{split} [\hat{g}, \hat{B}^i] &= -\mathrm{i}\kappa\tau^i \hat{g} \,, \, [\hat{B}^i, \hat{B}^j] = -\mathrm{i}\kappa\,\epsilon^{ij}{}_k \hat{B}^k \\ g &= \sqrt{1 - \vec{\pi}[g]^2}\,\mathbf{1} - \mathrm{i}\vec{\sigma}\cdot\vec{\pi}[g] \,, \quad |\vec{\pi}[g]| \le 1 \end{split}$$

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$$\hat{b}_a^i = \mathrm{i}\kappa \int (\mathrm{d}g)^4 \,\hat{\varphi}^{\dagger}(g_I) \frac{\mathrm{d}}{\mathrm{d}t} \hat{\varphi} \left(\exp\left(\tau_a^i t\right) g_I \right) \Big|_{t=0}$$

total (non-commutative) flux

satisfying:
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ight] \,\propto\, \widehat{N}$$

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entering effective (semiclassical) cosmological equations via expectation values:

macroscopic geometric conjugate variables are instead:

one extensive, other intensive

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 $(\vec{\Pi}[g_a]^{\mathrm{av.}}) = \langle \hat{\Pi}[g_a] \rangle / \langle \hat{N} \rangle$

"average holonomy"

$$B_a^i = \langle \, \widehat{b_a^i} \, \rangle$$

S. Gielen, DO, arXiv:1407.8167 [gr-qc]
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two immediate (generic) consequences:

1) GFT condensate cosmology gives quantum corrections to cosmological equations akin to LQC ones

 effective cosmological equations will carry a dependence on <N> (purely quantum observable) when expressed in terms of cosmological variables

exact relation between <N> and cosmological variables depends on quantum state

GFT condensates are interesting candidates for physical, geometric vacua of QG theory

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derivation of (quantum) cosmological equations from GFT quantum dynamics very general it rests on:

- continuum homogeneous quantum space (at microscopic scales) ~ GFT condensate
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non-linear quantum cosmology is QG analogue of Gross-Pitaevskii hydrodynamics for BECs consistent with "geometrogenesis" hypothesis and general "macro-from-micro" scenario

"standard" directions (but calculations to be done):

phenomenology as in LQC, here derived from fundamental theory

basis for most LQC phenomenology: modified Friedmann equation plus quantum corrections

$$\left(\frac{\dot{a}}{a}\right)^2 = \left(8\pi G\,\rho/3\right)\,\left(1 - \frac{\rho}{\rho_{\rm crit}}\right)$$

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other source of LQC phenomenology (Bojowald et al.): deformation of diffeomorphism algebra & signature change

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phenomenology as in LQC, here derived from fundamental theory

basis for most LQC phenomenology: modified Friedmann equation plus quantum corrections

$$\left(\frac{\dot{a}}{a}\right)^2 = (8\pi G \,\rho/3) \,\left(1 - \frac{\rho}{\rho_{\rm crit}}\right)$$

• in GFT condensate cosmology, modified gravity equations at effective level basically inevitable

modifications from several ingredients (in addition to involved microscopic dynamics):

"expected from LQC":

- holonomy corrections
- inverse triad corrections
- non-commutativity of fluxes

new, due to embedding into full theory:

- new QG observable N: number of "QG atoms of space"
- non-linear terms in effective cosmological equations
- hydrodynamic character of cosmological dynamics

other source of LQC phenomenology (Bojowald et al.): deformation of diffeomorphism algebra & signature change

diffeos in GFT also expected to be deformed:

A. Baratin, F. Girelli, DO, '11

- simplicial diffeos realised as global quantum group symmetry in topological models
 - expect more surprises at effective cosmological level

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exactly as one would do in a BEC....

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needed for computation of CMB spectrum needed for tests of fate of Lorentz invariance

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define modified FRW metric from expectation values for cosmological variables derived from GFT
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- 2. derive effective dynamics for GFT fluctuations above condensate from full theory
- 3. recast it in standard spacetime-based QFT form using information from background GFT condensate (difficulty is: the formalism naturally gives it in diffeo-invariant variables, spacetime-free form)

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expect deformation of standard QFT:

 holonomization of the connection and non-commutativity of triad variables, both entering definition of basic variables for perturbations (momenta, positions)

> derivation of effective dynamics of perturbations around mean field in topological GFT: non-commutative scalar field theory on non-commutative flat space W. Fairbairn, E. Livine, '07; F. Girelli, E. Livine, DO, '09

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a new avenue for analogue gravity: analogue GFT condensate quantum cosmology in real BECs?

need to simulate mini-superspace, not spacetime BEC hydrodynamics needs to reproduce GFT "classical" equations", not GR equations no problem with lack of diffeo invariance or relativistic symmetry in the lab

Foundational issues: the universe as a (quantum) fluid

- GFT condensates encoded in "collective wave function" identical to Quantum Cosmology one
 - corresponding dynamical equation is non-linear and non-local (on mini-superspace)
- cosmological dynamics is hydrodynamics of fundamental GFT (analogue of GP equation for BEC)
 - still, fundamental dynamics is (more or less) standard quantum mechanics for QG d.o.f.s (although big interpretational issues (e.g. concerning probabilities, unitarity, etc) remain)

calls for new interpretation of "quantum cosmology" (see also Bojowald, '15):

- quantum cosmology is not quantum at all; rather, "cosmological hydrodynamics" no probability for "the whole universe", no "Hilbert space of states of the Universe"
- can still use expectation values (average quantities) but in "hydrodynamics" (realistic/statistical) sense
- no problem of "collapse of cosmological wave function" or spontaneous collapse due to non-linearites?

4th message

a new promising direction to extract effective cosmological dynamics (and associated phenomenology)

directly from full QG theory!

Thank you for your attention!