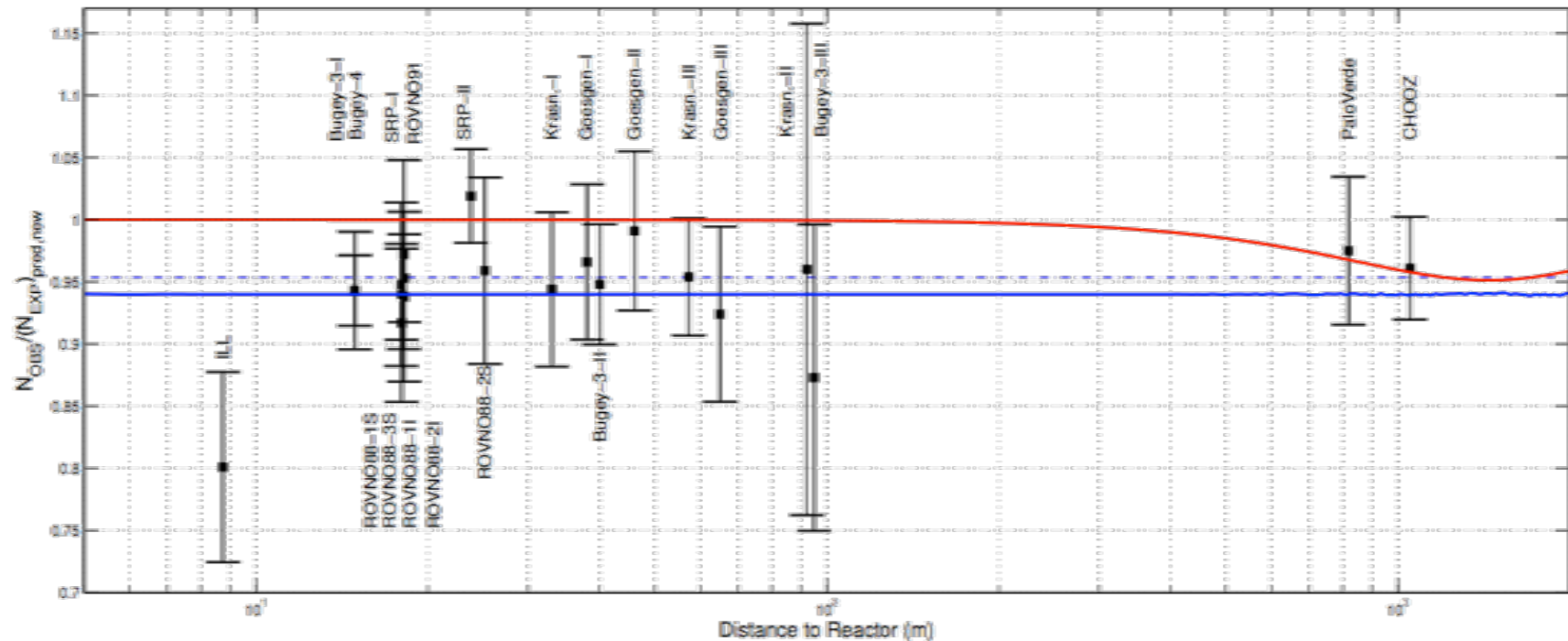


# The Reactor Antineutrino Anomaly and implications



Th. Lasserre (CEA-Saclay, Irfu SPP & APC)

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# New Reactor Antineutrino Spectra

*T.A. Mueller, D. Lhuiller\*, M. Fallot, A. Letourneau, S. Cormon,  
M. Fechner, L. Giot, T. Lasserre, J. Martino, G. Mention, A. Porta, F. Yermia.*

*CEA / Irfu & IN2P3 / Subatech*

arXiv:1101.2663 [hep-ex], submitted to PRC

\* corresponding author

# $\nu$ spectrum emitted by a reactor

The prediction of reactor  $\nu$  spectrum is the dominant source of systematic error for single detector reactor neutrino experiments

**Reactor data**  
Thermal power,  $\delta P_{th} \leq 1\%$

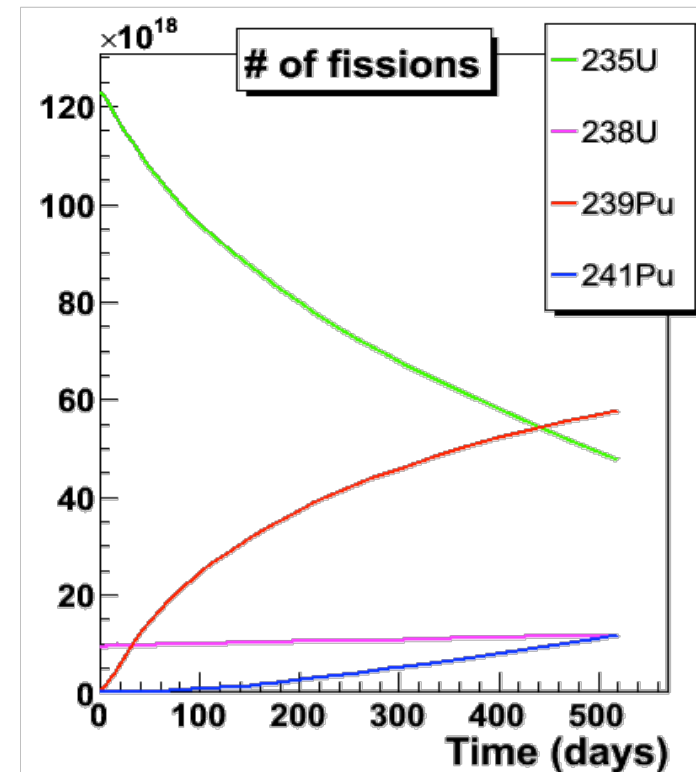
**Reactor evolution codes**  
Fraction of fissions from isotope  $k$ ,  $\delta \alpha_k = \text{few } \%$   
but large anti-correl @ fixed  $P_{th}$

**Nuclear databases**  
E released per fissions of isotope  $k$ ,  
 $\delta E_k \approx 0.3\%$

$$\Phi_\nu(E, t) = \frac{P_{th}(t)}{\sum_k \alpha_k(t) E_k} \times \sum_k \alpha_k(t) S_k(E)$$

$\nu$  spectrum per fission  
**This work !**

$k = {}^{235}\text{U}, {}^{238}\text{U}, {}^{239}\text{Pu}, {}^{241}\text{Pu}$



# The guts of $S_k(E)$

Sum of all fission products' activities

$$S_k(E) = \sum_{fp=1}^{N_{fp}} \mathcal{A}_{fp}(T) \times S_{fp}(E)$$

Sum of all  $\beta$ -branch of each fission product

$$S_{fp}(E) = \sum_{b=1}^{N_b} BR_{fp}^b \times S_{fp}^b(Z_{fp}, A_{fp}, E_{0fp}^b, E)$$

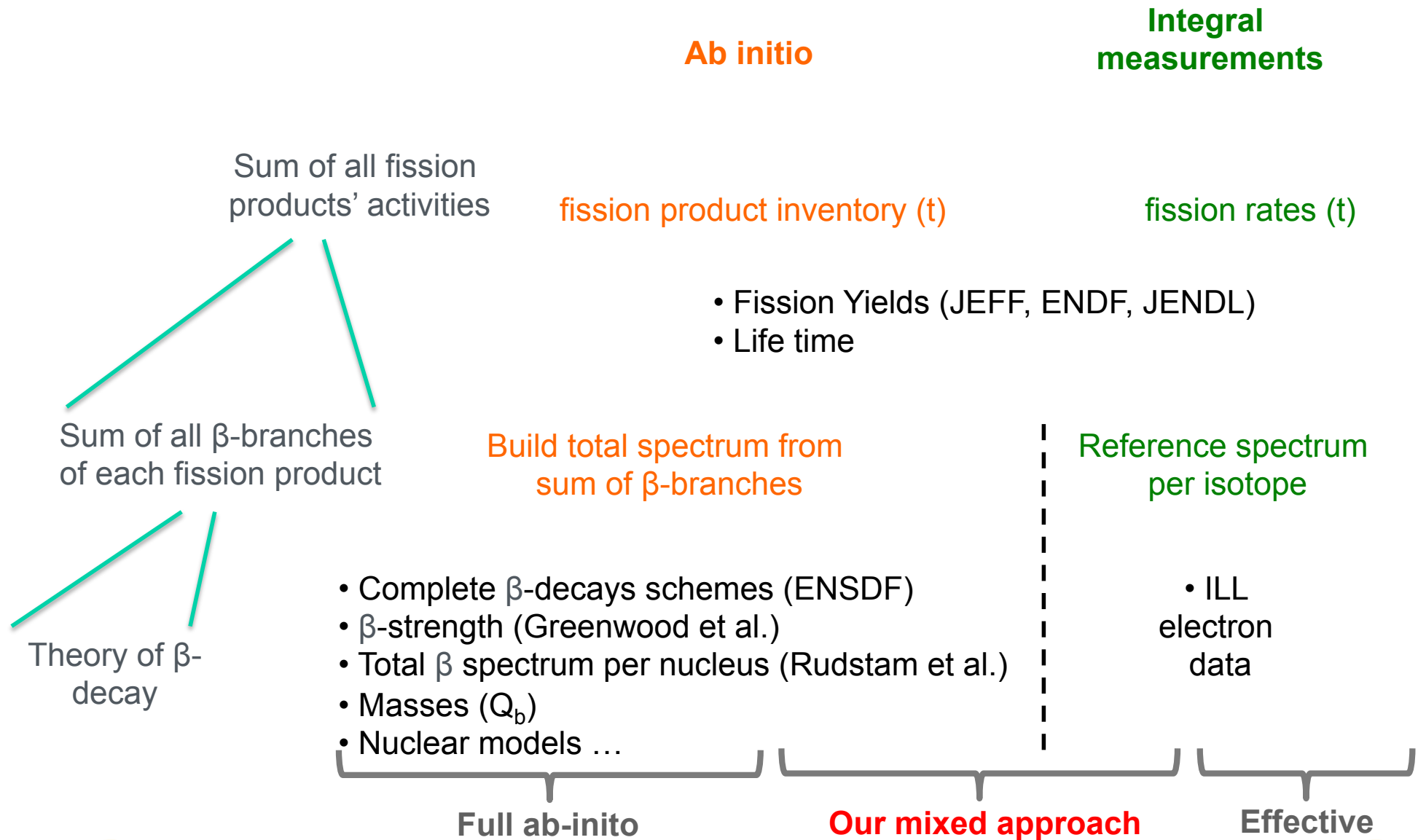
Theory of  $\beta$ -decay

$$S_{fp}^b = \underbrace{K_{fp}^b}_{\text{Norm.}} \times \underbrace{\mathcal{F}(Z_{fp}, A_{fp}, E)}_{\text{Fermi function}} \times \underbrace{pE(E - E_{0fp}^b)^2}_{\text{Phase space}} \\ \times \underbrace{C_{fp}^b(E)}_{\text{Shape factor}} \times \underbrace{\left(1 + \delta_{fp}^b(Z_{fp}, A_{fp}, E)\right)}_{\text{Correction}}$$

$$\delta_{fp}^b(Z_{fp}, A_{fp}, E) = \delta_{QED}(E) + A_C(Z_{fp}, A_{fp}) \times E + A_W \times E$$



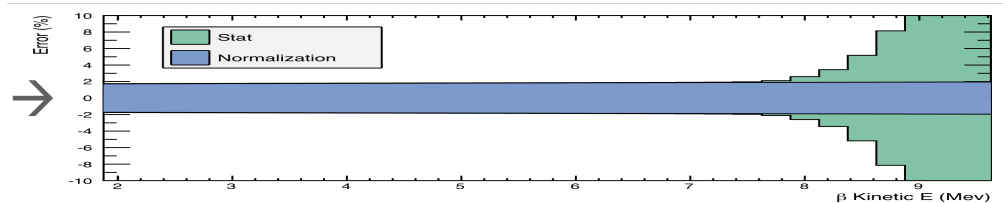
# Complementary approaches to compute the $\nu$ flux



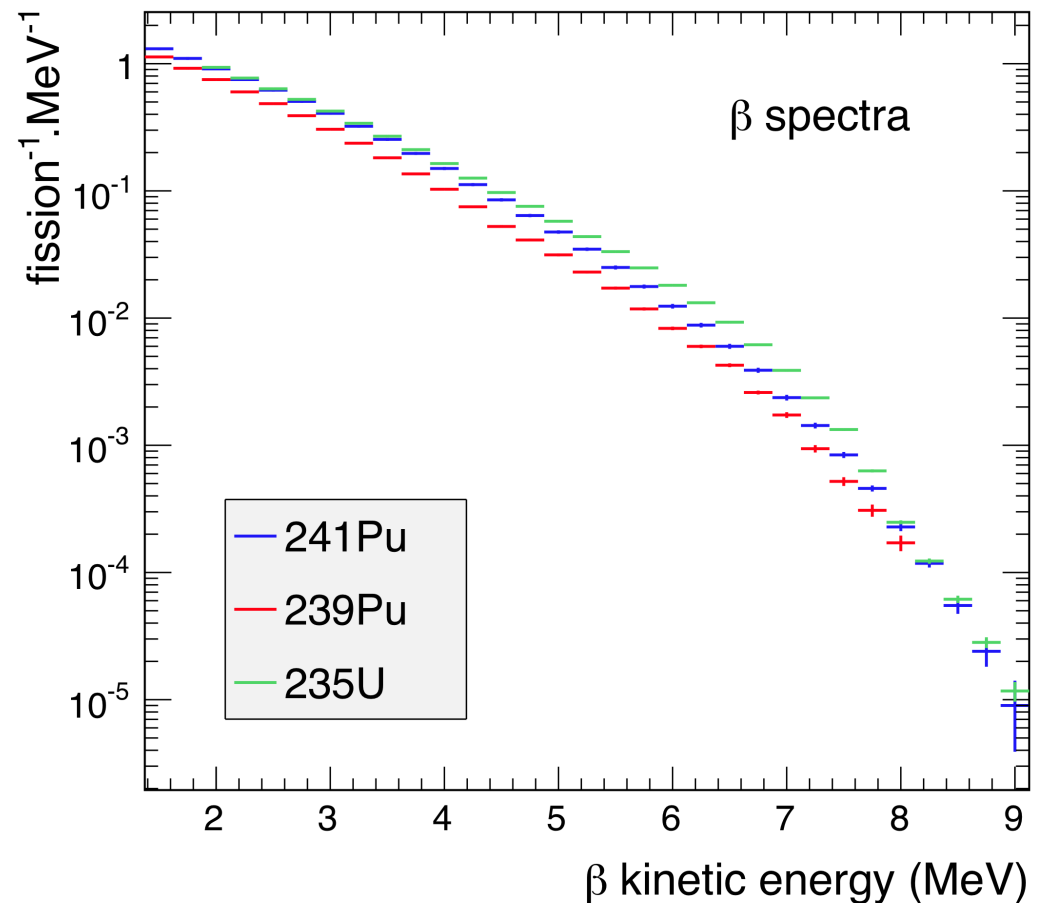
# The ILL electron Data Anchorage

Unique reference to be met by any other measurement or calculation

uncertainty →

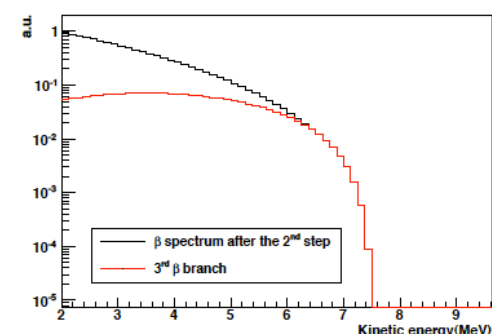
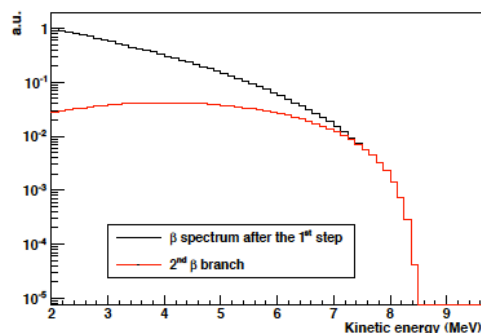
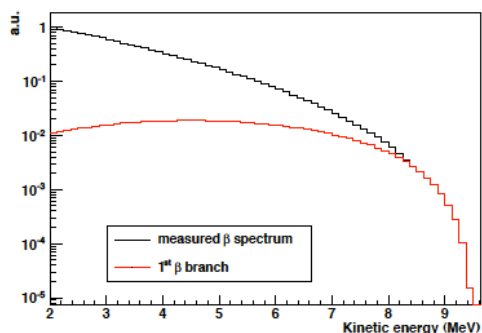


- Accurate  $e^-$  measurements @ ILL' (1980-89):
  - High resolution magn. spectrometer
  - Intense and pure thermal n spectrum from the core
  - Extensive use of reference internal conversion electron lines → Normalization (1.8%)



# ILL data: conversion to $\nu$ spectra

- Fit  $e^-$  spectrum with a sum of 30 effective branches
- Conversion of the effective branches to  $\nu$  spectra



- All theory included in these effective branches but:

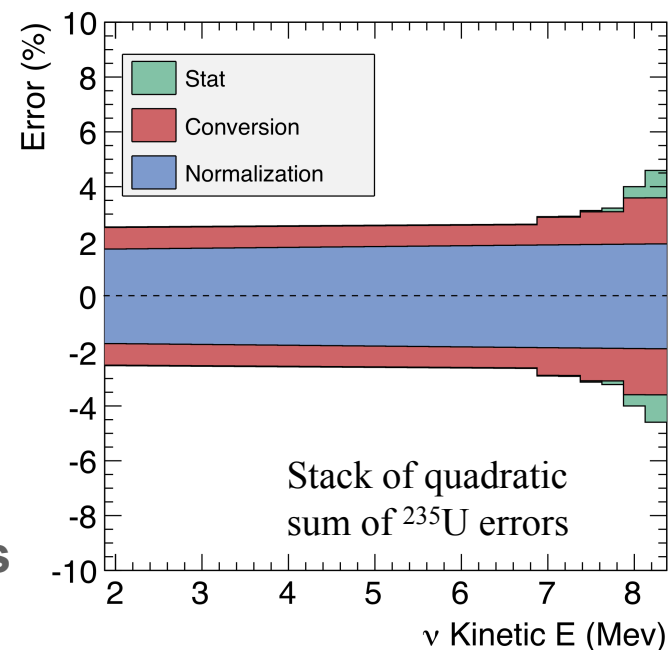
- What  $Z$ ? : Mean fit on nuclear data  $Z=f(E_0)$

$$Z(E_0) \approx 49.5 - 0.7E_0 - 0.09E_0^2, \quad Z \geq 34$$

- What  $A_{CW}$ ? : effective correction on the  $\nu$ -spectra

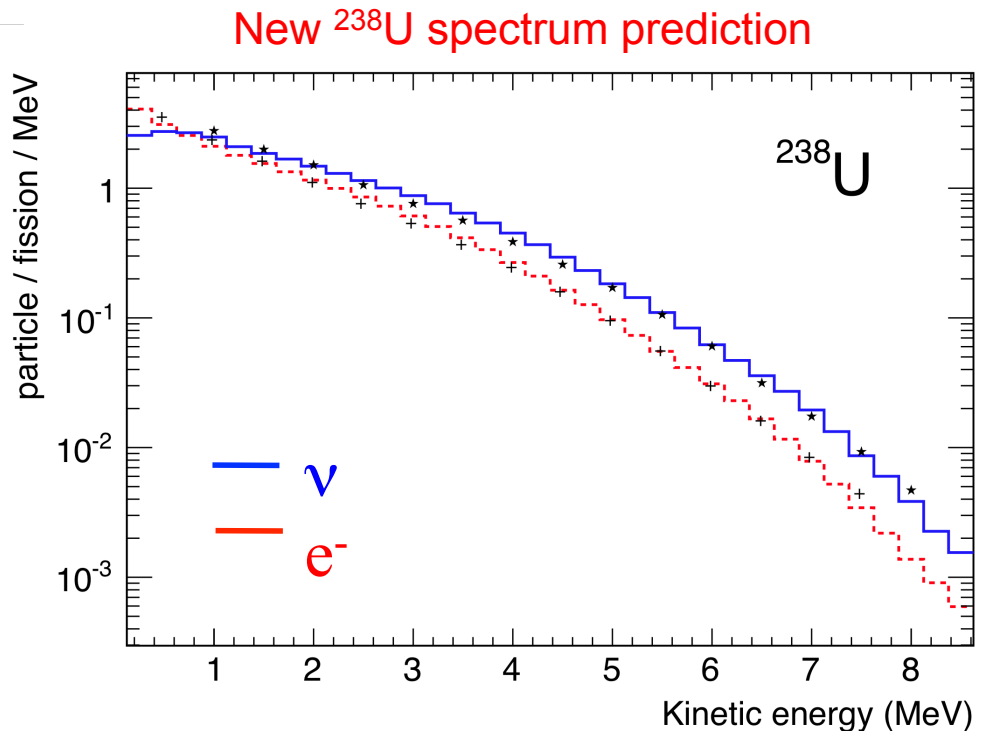
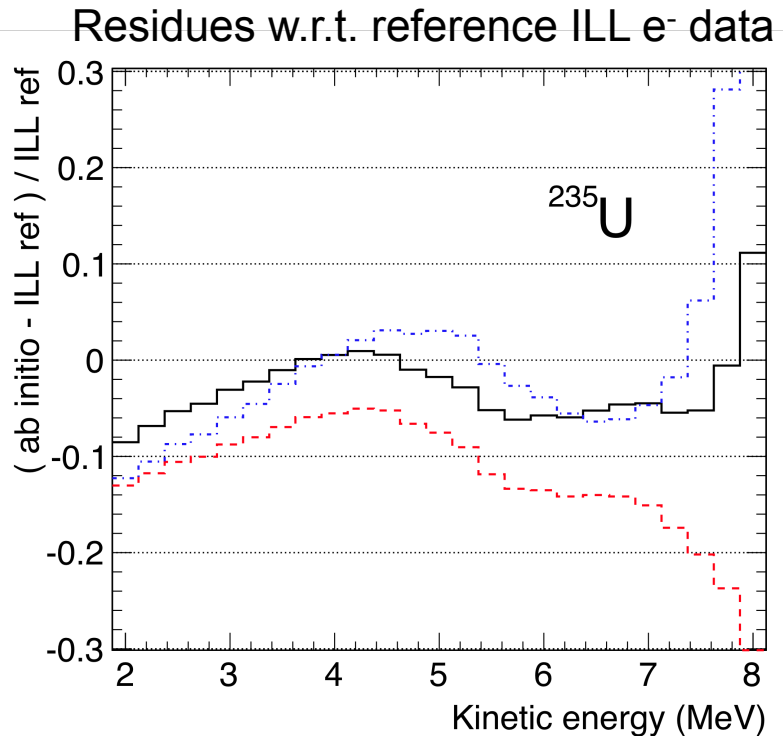
$$DN_n^{C,W}(E_n) \approx 0.65 \times (E_n - 4\text{MeV}) \quad \%$$

- Conversion error from envelop of numerical studies



# The Full *Ab Initio* Attempt (electron data)

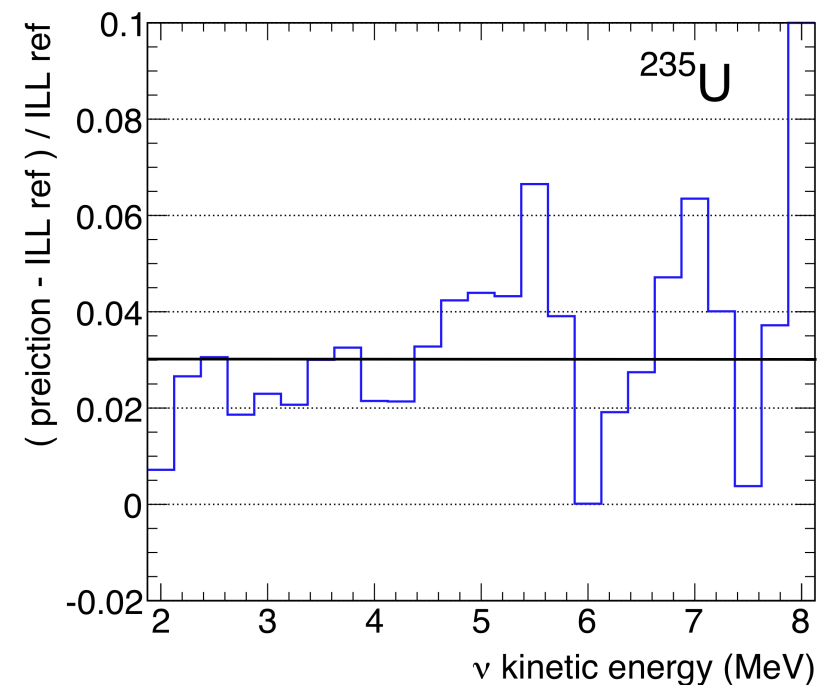
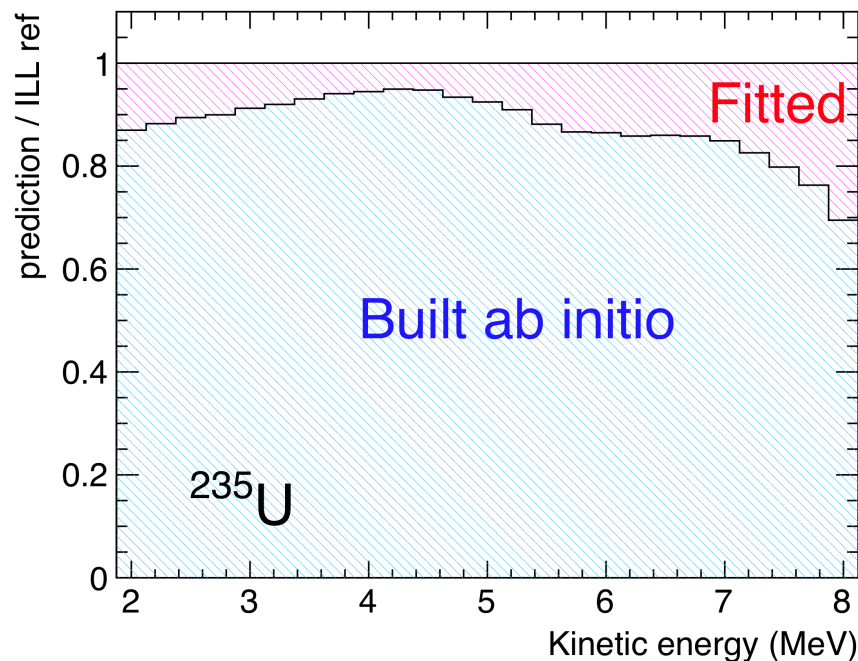
- MURE evolution code: core composition and off equilibrium effects
- BESTIOLE code: build up database of ~800 nuclei and 10000  $\beta$ -branches



- 95+/-5% of the spectrum reproduced but still not meeting required precision
- Useful estimate of  $^{238}\text{U}$  spectrum which couldn't be measured @ ILL
- Measurement at FRMII ongoing (N. Haag & K Schreckenbach)

# The New Mixed Conversion Approach

1. **SAME** ILL e- data Anchorage
2. Ab-Initio: “true” distribution of  $\beta$ -branches reproduces >90% of ILL e- data.
3. Old-procedure: five effective anchorage-branches to the remaining 10%.



- +3% normalization shift with respect to old  $\nu$  spectrum
- Similar result for all isotopes ( $^{235}\text{U}$ ,  $^{239}\text{Pu}$ ,  $^{241}\text{Pu}$ )
- Stringent Test Performed – Origin of the bias identified

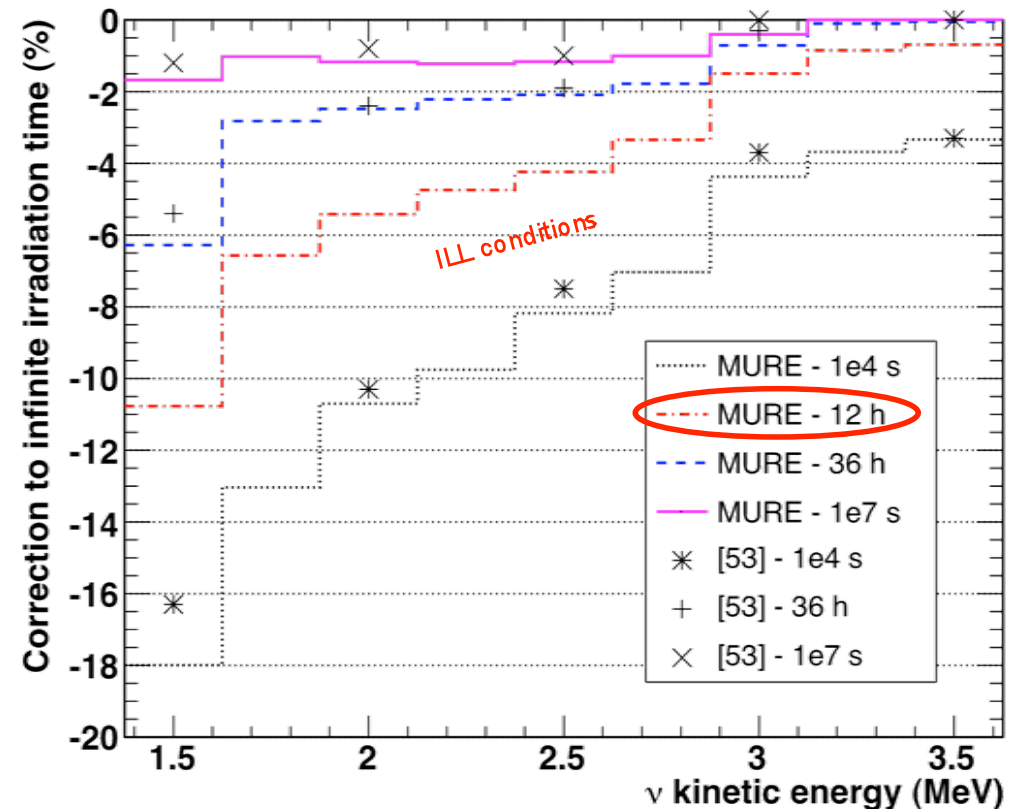
# Off-Equilibrium Effects

MURE evolution code (IN2P3/Subatech)

- ILL electron reference spectra : 12 hours to 1.8 days irradiation time
- Neutrino reactor experiments irradiation time  $\gg$  months
- **BUT 10% of fission products have a  $\beta$ -decay life-time long enough to keep accumulating after several days**
  - need a correction through simulation
  - Not included prior to the CHOOZ experiment

Relative change of  $\nu$  spectrum w.r.t. infinite irradiation time

Correction included by default in our new reference model



# The Reactor Antineutrino Anomaly

→ Updated with 7 additional results

*G. Mention, M. Fechner, T. Lasserre\*,  
M. Cribier, Th. Mueller D. Lhuillier, A. Letourneau,*

*CEA / Ifu*

arXiv:1101.2755 [hep-ex], submitted to PRD

\* corresponding author

# V-A IBD Cross Section

- **Inverse Beta Decay:**  $\bar{\nu}_e + p \rightarrow e^+ + n$
- **Theoretical predictions: our results agree with**
  - Vogel 1984 (Phys Rev D29 p1918). **Fayans 1985 (Sov J Nucl Phys 42)**
  - Vogel-Beacom 1999: “supersedes” Vogel 84 (Phys Prev D60 053003)
  - Strumia-Vissani Phys. Lett. B564 (2003) 42-54

$$\sigma_{V-A}(E_e) = \kappa p_e E_e (1 + \delta_{rec} + \delta_{wm} + \delta_{rad})$$

- **The pre-factor  $\kappa$**  (two pseudo-independent approaches)

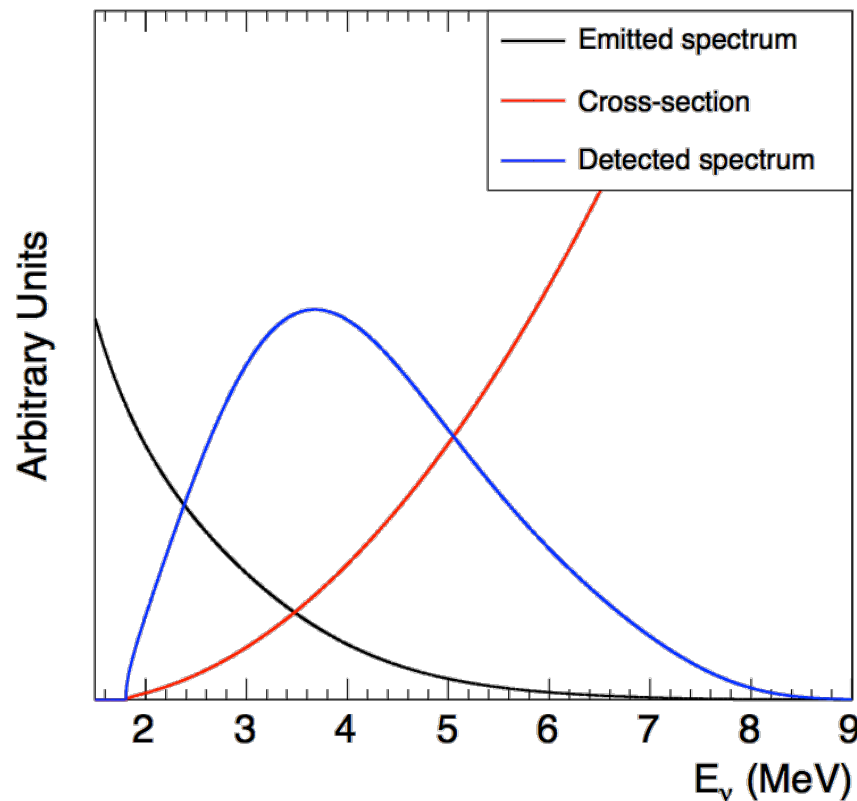
$$\kappa = \frac{G_F^2 \cos^2(\theta_C)}{\pi} (1 + \Delta_{inner}^R) (1 + 3\lambda^2) = \frac{2\pi^2}{m_e^5 f^R \tau_n} \quad \lambda = \left| \frac{g_A}{g_V} \right|$$

- **$\kappa$  ran down over the history**, from  $0.914 \cdot 10^{-42} \text{ cm}^2$  in 1981
  - Vogel-Beacom 1999 :  $\kappa = 0.952 \cdot 10^{-42} \text{ cm}^2$
  - **Our work is based on 2010 PDG  $\tau_n$  :  $\kappa = 0.956 \cdot 10^{-42} \text{ cm}^2$**
  - But we anticipate 2011  $\kappa = 0.961 \cdot 10^{-42} \text{ cm}^2$  ( $\langle \tau_n \rangle$  revision +0.5%)



# Computing the expected rate/spectrum

$$\sigma_f^{pred} = \int_0^\infty S_{tot}(E_\nu) \sigma_{V-A}(E_\nu) dE_\nu = \sum_k f_k \sigma_{f,k}^{pred}$$



## • Bugey-4 Benchmark

- Phys Lett B 338(1994) 383
- $\tau_n = 887.4$  s
- “old” spectra (30 effective branches)
- no off-equilibrium corrections

$10^{-43}$ cm <sup>2</sup> /fission	<sup>235</sup> U	<sup>239</sup> Pu	<sup>241</sup> Pu
BUGEY-4	6.39±1.9%	4.18±2.4%	5.76±2.1%
This work	6.39±1.8%	4.19±2.3%	5.73±1.9%

- Final agreement to better than 0.1% on best known <sup>235</sup>U

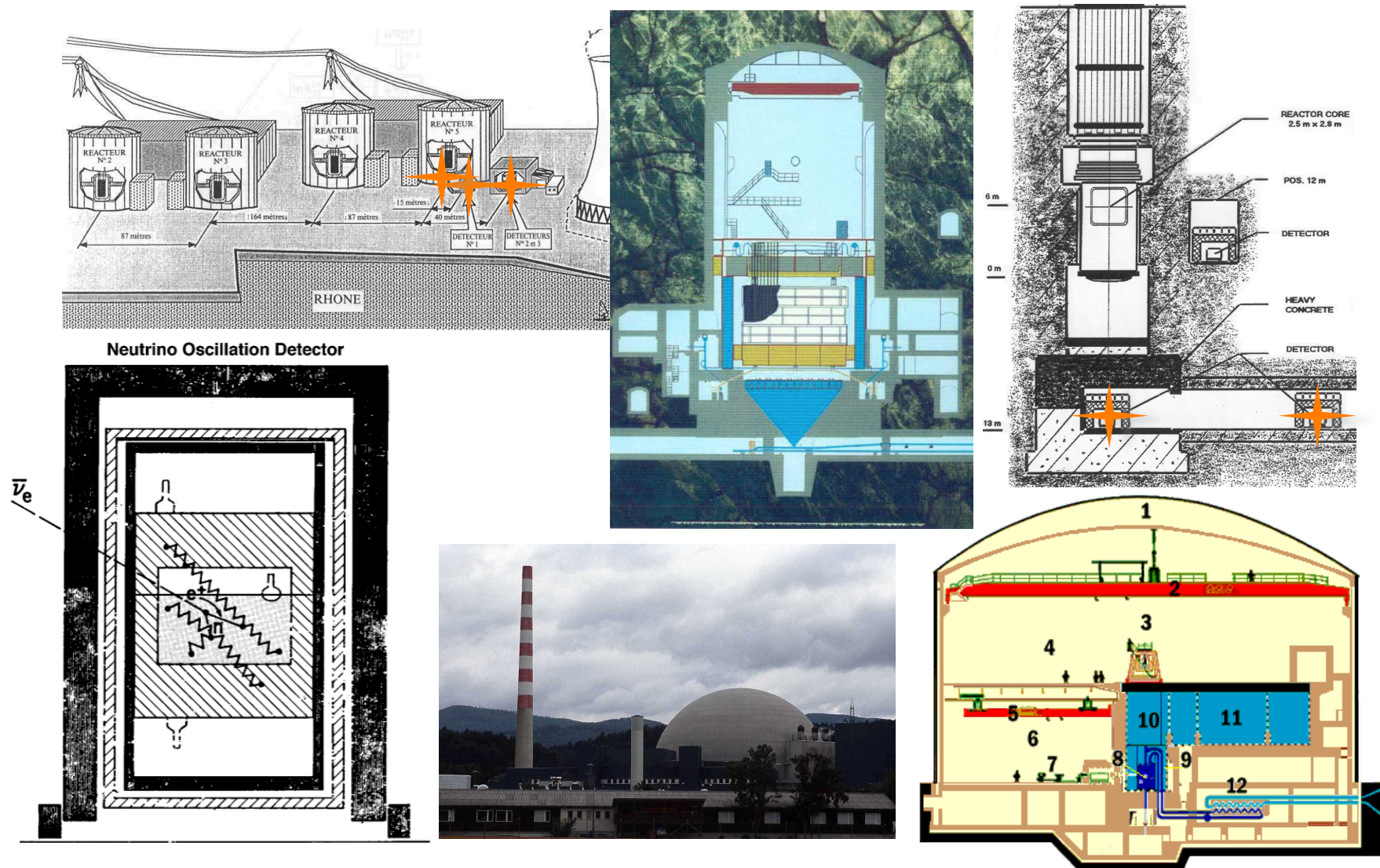
# The New Cross Section Per Fission

- $\nu$ -flux:  $^{235}\text{U}$  +2.5%,  $^{239}\text{Pu}$  +3.1%,  $^{241}\text{Pu}$  +3.7%,  $^{238}\text{U}$  +9.8% ( $\sigma_f^{\text{pred}}$  ↗)
- Off-equilibrium corrections now included ( $\sigma_f^{\text{pred}}$  ↗)
- Neutron lifetime decrease by a few % ( $\sigma_f^{\text{pred}}$  ↗)
- Slight evolution of the phase space factor ( $\sigma_f^{\text{pred}}$  →)
- Slight evolution of the energy per fission per isotope ( $\sigma_f^{\text{pred}}$  →)
- Burnup dependence:  $\sigma_f^{\text{pred}} = \sum_k f_k \sigma_{f,k}^{\text{pred}}$  ( $\sigma_f^{\text{pred}}$  →)

▪ New Results:

	old [3]	new
$\sigma_{f,^{235}\text{U}}^{\text{pred}}$	$6.39 \pm 1.9\%$	$6.61 \pm 2.11\%$
$\sigma_{f,^{239}\text{Pu}}^{\text{pred}}$	$4.19 \pm 2.4\%$	$4.34 \pm 2.45\%$
$\sigma_{f,^{238}\text{U}}^{\text{pred}}$	$9.21 \pm 10\%$	$10.10 \pm 8.15\%$
$\sigma_{f,^{241}\text{Pu}}^{\text{pred}}$	$5.73 \pm 2.1\%$	$5.97 \pm 2.15\%$

# 19 Experimental Results at L<100m



Measured cross sections are taken at their face values



# 19 Experimental Results Revisited (L<100m)

Technology

Baseline

#	result	techno	$\tau_n$ (s)	$^{235}\text{U}$	$^{239}\text{Pu}$	$^{238}\text{U}$	$^{241}\text{Pu}$	old	new	err(%)	corr(%)	L(m)
1	Bugey-4	$^3\text{He}+\text{H}_2\text{O}$	888.7	0.538	0.328	0.078	0.056	0.987	0.943	3.0	3.0	15
2	ROVNO91	$^3\text{He}+\text{H}_2\text{O}$	888.6	0.614	0.274	0.074	0.038	0.985	0.940	3.9	3.0	18
3	Bugey-3-I	$^6\text{Li-LS}$	889	0.538	0.328	0.078	0.056	0.988	0.943	5.0	5.0	15
4	Bugey-3-II	Li-LS	889	0.538	0.328	0.078	0.056	0.994	0.948	5.1	5.0	40
5	Bugey-3-III	Li-LS	889	0.538	0.328	0.078	0.056	0.915	0.873	14.1	5.0	95
6	Goesgen-I	$^3\text{He}+\text{LS}$	897	0.6198	0.274	0.074	0.042	1.018	0.966	6.5	6.0	38
7	Goesgen-II	$^3\text{He}+\text{LS}$	897	0.584	0.298	0.068	0.050	1.045	0.991	6.5	6.0	45
8	Goesgen-II	$^3\text{He}+\text{LS}$	897	0.543	0.329	0.070	0.058	0.975	0.924	7.6	6.0	65
9	ILL	$^3\text{He}+\text{LS}$	889	$\simeq 1$	<0.01	<0.01	<0.01	0.832	0.801	9.5	6.0	9
10	Krasn. I	$^3\text{He}+\text{PE}$	899	$\simeq 1$	<0.01	<0.01	<0.01	1.013	0.944	5.1	4.1	33
11	Krasn. II	$^3\text{He}+\text{PE}$	899	$\simeq 1$	<0.01	<0.01	<0.01	1.031	0.960	20.3	4.1	92
12	Krasn. II	$^3\text{He}+\text{PE}$	899	$\simeq 1$	<0.01	<0.01	<0.01	0.989	0.954	4.1	4.1	57
13	SRP I	Gd-LS	887	$\simeq 1$	<0.01	<0.01	<0.01	0.987	0.953	3.7	3.7	18
14	SRP II	Gd-LS	887	$\simeq 1$	<0.01	<0.01	<0.01	1.055	1.019	3.8	3.7	24
15	ROVNO88-1I	$^3\text{He}+\text{PE}$	898.8	0.607	0.277	0.074	0.042	0.969	0.917	6.9	6.9	18
16	ROVNO88-2I	$^3\text{He}+\text{PE}$	898.8	0.603	0.276	0.076	0.045	1.001	0.948	6.9	6.9	18
17	ROVNO88-1S	Gd-LS	898.8	0.606	0.277	0.074	0.043	1.026	0.972	7.8	7.8	18
18	ROVNO88-2S	Gd-LS	898.8	0.557	0.313	0.076	0.054	1.013	0.959	7.8	7.8	25
19	ROVNO88-3S	Gd-LS	898.8	0.606	0.274	0.074	0.046	0.990	0.938	7.2	7.2	18

# 19 Experimental Results Revisited (L<100m)

## Neutron lifetime

#	result	techno	$\tau_n$ (s)	$^{235}\text{U}$	$^{239}\text{Pu}$	$^{238}\text{U}$	$^{241}\text{Pu}$	old	new	err(%)	corr(%)	L(m)
1	Bugey-4	$^3\text{He}+\text{H}_2\text{O}$	888.7	0.538	0.328	0.078	0.056	0.987	0.943	3.0	3.0	15
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# 19 Experimental Results Revisited (L<100m)

## Averaged Fuel Composition

#	result	techno	$\tau_n$ (s)	$^{235}\text{U}$	$^{239}\text{Pu}$	$^{238}\text{U}$	$^{241}\text{Pu}$	old	new	err(%)	corr(%)	L(m)
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# 19 Experimental Results Revisited (L<100m)

OBSERVED/PREDICTED ratios: OLD & NEW (this work)

#	result	techno	$\tau_n$ (s)	$^{235}\text{U}$	$^{239}\text{Pu}$	$^{238}\text{U}$	$^{241}\text{Pu}$	old	new	err(%)	corr(%)	L(m)
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7	Goesgen-II	$^3\text{He}+\text{LS}$	897	0.584	0.298	0.068	0.050	1.045	0.991	6.5	6.0	45
8	Goesgen-II	$^3\text{He}+\text{LS}$	897	0.543	0.329	0.070	0.058	0.975	0.924	7.6	6.0	65
9	ILL	$^3\text{He}+\text{LS}$	889	$\simeq 1$	$<0.01$	$<0.01$	$<0.01$	0.832	0.801	9.5	6.0	9
10	Krasn. I	$^3\text{He}+\text{PE}$	899	$\simeq 1$	$<0.01$	$<0.01$	$<0.01$	1.013	0.944	5.1	4.1	33
11	Krasn. II	$^3\text{He}+\text{PE}$	899	$\simeq 1$	$<0.01$	$<0.01$	$<0.01$	1.031	0.960	20.3	4.1	92
12	Krasn. II	$^3\text{He}+\text{PE}$	899	$\simeq 1$	$<0.01$	$<0.01$	$<0.01$	0.989	0.954	4.1	4.1	57
13	SRP I	Gd-LS	887	$\simeq 1$	$<0.01$	$<0.01$	$<0.01$	0.987	0.953	3.7	3.7	18
14	SRP II	Gd-LS	887	$\simeq 1$	$<0.01$	$<0.01$	$<0.01$	1.055	1.019	3.8	3.7	24
15	ROVNO88-1I	$^3\text{He}+\text{PE}$	898.8	0.607	0.277	0.074	0.042	0.969	0.917	6.9	6.9	18
16	ROVNO88-2I	$^3\text{He}+\text{PE}$	898.8	0.603	0.276	0.076	0.045	1.001	0.948	6.9	6.9	18
17	ROVNO88-1S	Gd-LS	898.8	0.606	0.277	0.074	0.043	1.026	0.972	7.8	7.8	18
18	ROVNO88-2S	Gd-LS	898.8	0.557	0.313	0.076	0.054	1.013	0.959	7.8	7.8	25
19	ROVNO88-3S	Gd-LS	898.8	0.606	0.274	0.074	0.046	0.990	0.938	7.2	7.2	18



# 19 Experimental Results Revisited (L<100m)

Total Errors Exp.+v-Spectra (%) & Correlated errors (%)

#	result	techno	$\tau_n$ (s)	$^{235}\text{U}$	$^{239}\text{Pu}$	$^{238}\text{U}$	$^{241}\text{Pu}$	old	new	err(%)	corr(%)	L(m)
1	Bugey-4	$^3\text{He}+\text{H}_2\text{O}$	888.7	0.538	0.328	0.078	0.056	0.987	0.943	3.0	3.0	15
2	ROVNO91	$^3\text{He}+\text{H}_2\text{O}$	888.6	0.614	0.274	0.074	0.038	0.985	0.940	3.9	3.0	18
3	Bugey-3-I	$^6\text{Li-LS}$	889	0.538	0.328	0.078	0.056	0.988	0.943	5.0	5.0	15
4	Bugey-3-II	Li-LS	889	0.538	0.328	0.078	0.056	0.994	0.948	5.1	5.0	40
5	Bugey-3-III	Li-LS	889	0.538	0.328	0.078	0.056	0.915	0.873	14.1	5.0	95
6	Goesgen-I	$^3\text{He}+\text{LS}$	897	0.6198	0.274	0.074	0.042	1.018	0.966	6.5	6.0	38
7	Goesgen-II	$^3\text{He}+\text{LS}$	897	0.584	0.298	0.068	0.050	1.045	0.991	6.5	6.0	45
8	Goesgen-II	$^3\text{He}+\text{LS}$	897	0.543	0.329	0.070	0.058	0.975	0.924	7.6	6.0	65
9	ILL	$^3\text{He}+\text{LS}$	889	$\simeq 1$	$<0.01$	$<0.01$	$<0.01$	0.832	0.801	9.5	6.0	9
10	Krasn. I	$^3\text{He}+\text{PE}$	899	$\simeq 1$	$<0.01$	$<0.01$	$<0.01$	1.013	0.944	5.1	4.1	33
11	Krasn. II	$^3\text{He}+\text{PE}$	899	$\simeq 1$	$<0.01$	$<0.01$	$<0.01$	1.031	0.960	20.3	4.1	92
12	Krasn. II	$^3\text{He}+\text{PE}$	899	$\simeq 1$	$<0.01$	$<0.01$	$<0.01$	0.989	0.954	4.1	4.1	57
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17	ROVNO88-1S	Gd-LS	898.8	0.606	0.277	0.074	0.043	1.026	0.972	7.8	7.8	18
18	ROVNO88-2S	Gd-LS	898.8	0.557	0.313	0.076	0.054	1.013	0.959	7.8	7.8	25
19	ROVNO88-3S	Gd-LS	898.8	0.606	0.274	0.074	0.046	0.990	0.938	7.2	7.2	18



# Error Budget & Correlations

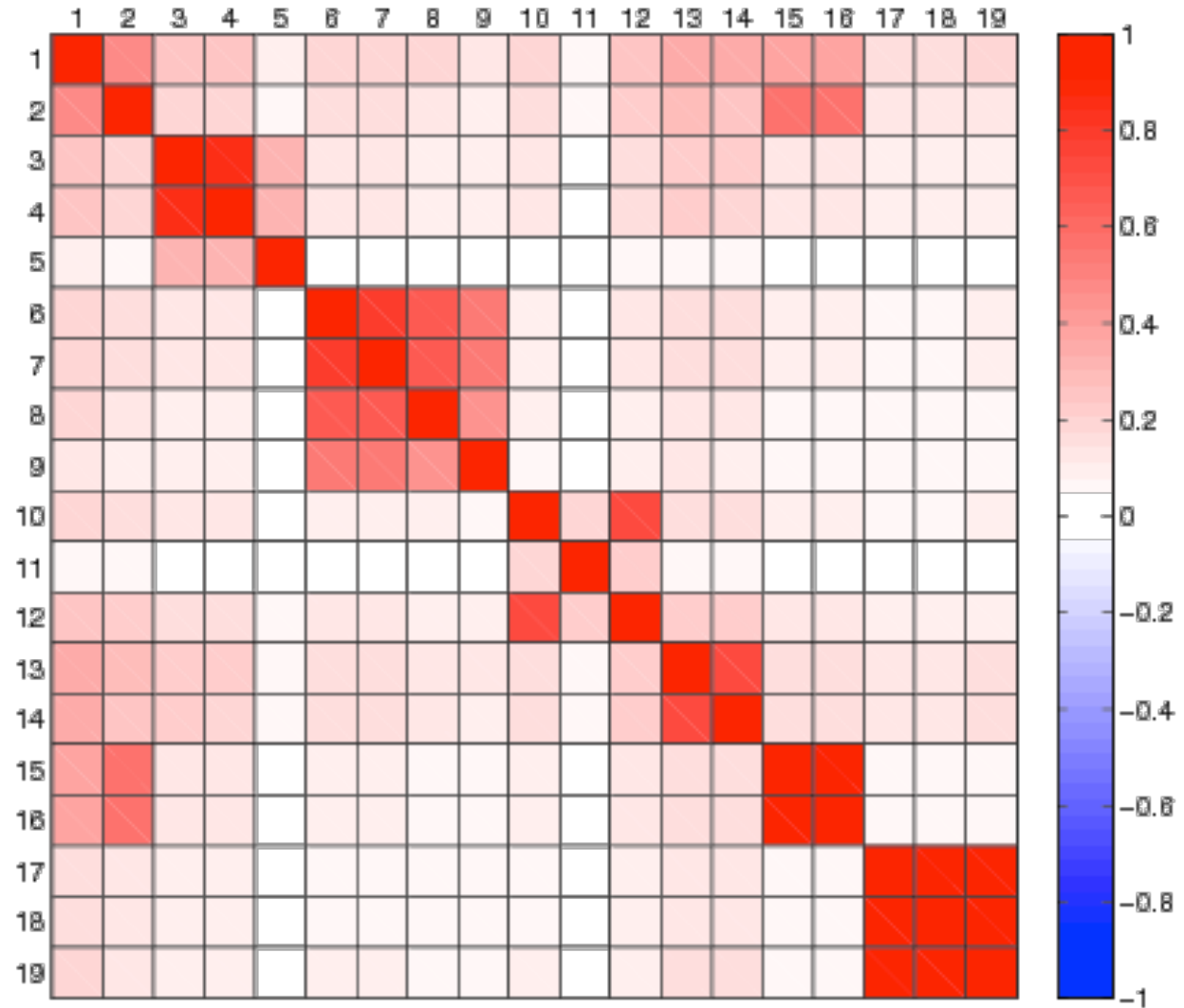
- Our guiding principles: Be conservative - Be stable numerically (SRP case)
- **Reactor Antineutrino Sources**
  - 2% systematic on  $\nu$ -flux 100% correlated over ALL measurements
    - Different from 2.7% published on Arxiv:1101.2755
    - 2% corresponds to the normalization error on the ILL e- data
- **Detector: Non-flux systematic error correlations across measurements:**
  - Same experiment with same technology: 100% correlated
  - ILL shares 6% correlated error with Goesgen although detector slightly different. Rest of ILL error is uncorrelated.
  - Rovno88 integral measurements 100% corr. with Rovno 91 despite detector upgrade, but not with Rovno88 LS data
  - Rovno91 integral meas. 100% correlated with Bugey-4
  - Rovno88 integral meas. 50% correlated with Bugey-4

# Experimental correlation matrix

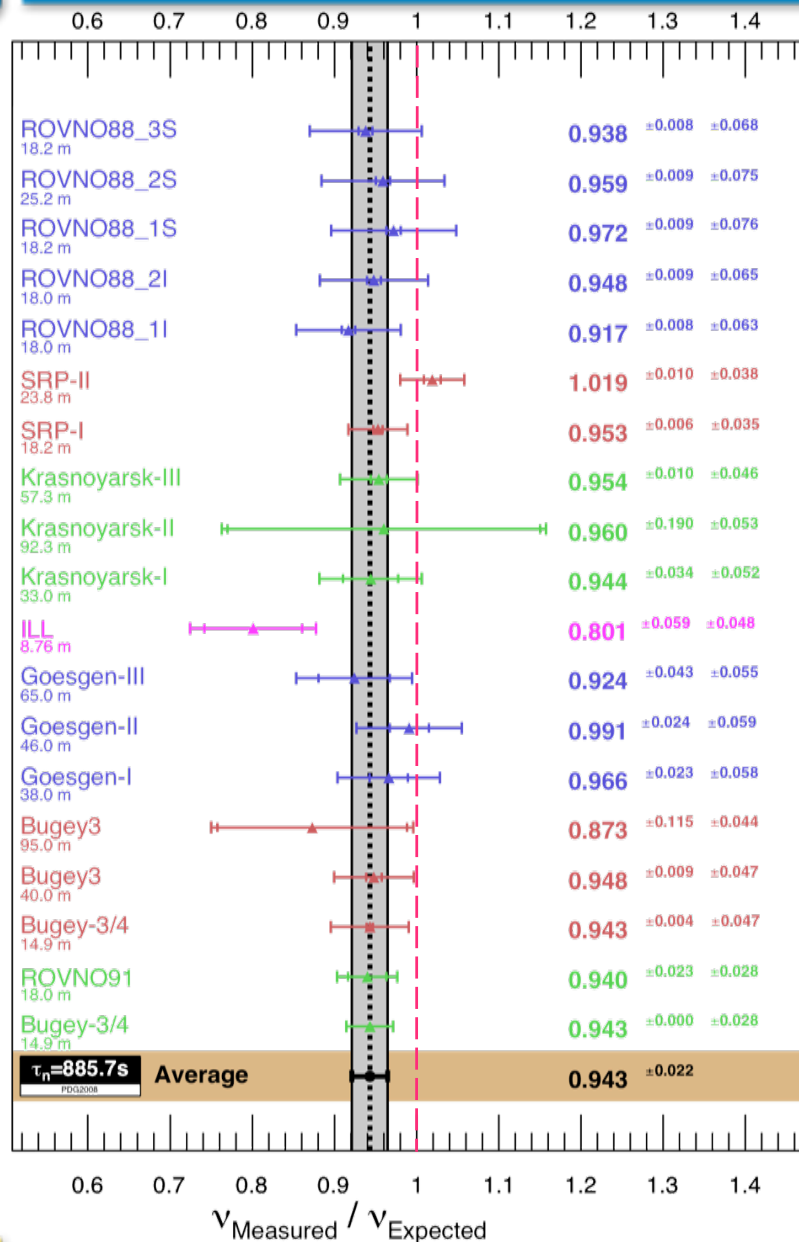
$\vec{R}$

$W$

- Bugey-4 15m
- Rovno91 18m
- Bugey-3 15m
- Bugey-3 40m
- Bugey-3 92m
- Goesgen 38m
- Goesgen 45m
- Goesgen 65m
- ILL 9m
- Krasno 33m
- Krasno 92m
- Krasno 57m
- SRP I 18m
- SRP II 25m
- Rovno88 1I 18m
- Rovno88 2I 18 m
- Rovno88 1S 18m
- Rovno88 2S 25m
- Rovno88 3S 18m



# The reactor antineutrino anomaly

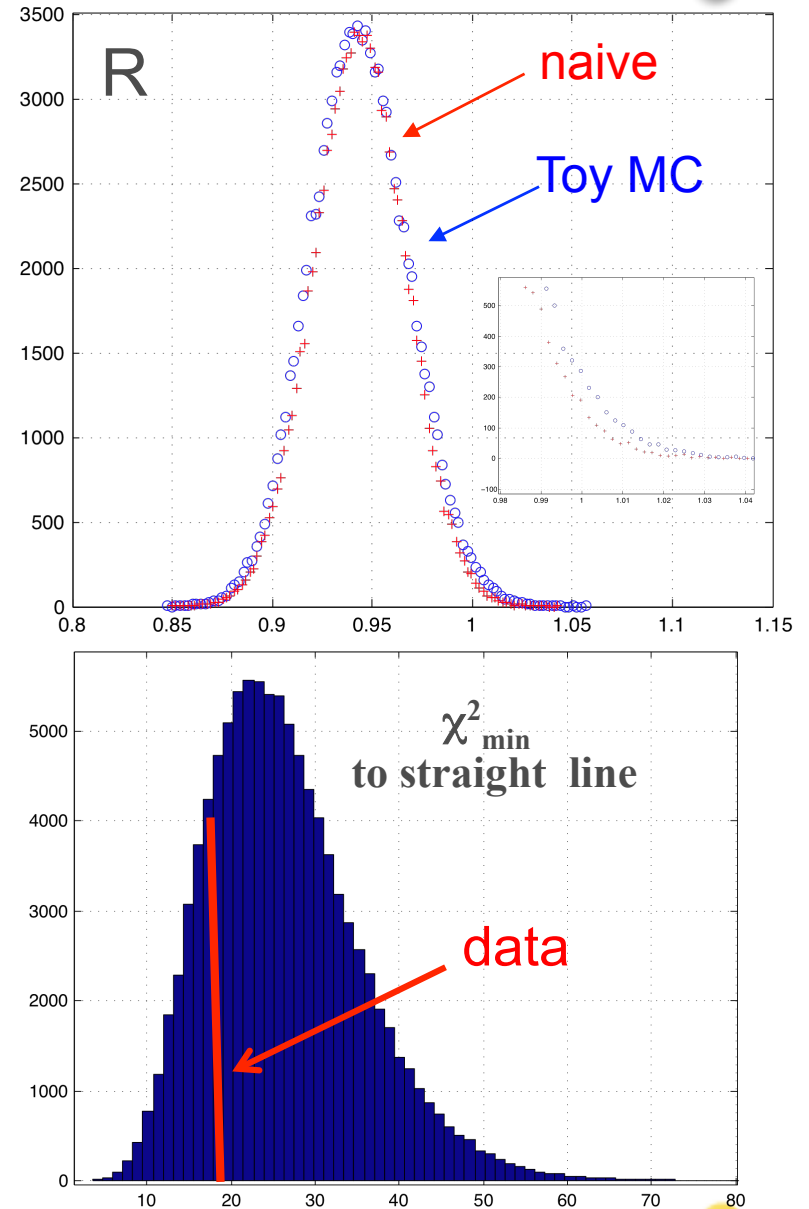


$$\chi^2 = \left( r - \vec{R} \right)^T W^{-1} \left( r - \vec{R} \right)$$

- **Best fit :  $\mu = 0.943$**
- **Uncertainty : 0.023**
- $\chi^2 = 19.6/19$
- Deviation from unity
  - Naïve Gaussian : 99.3% C.L.
  - Toy MC: 98.6% C.L. ( $10^6$  trials)
- No hidden covariance
  - 18% of Toy MC have  $\chi^2_{\min} < 19.6$

# Are the ratios normally distributed?

- Our data points are ratios of Gaussians:
  - **Numerator:** measurement, Gaussian with stat & syst error, partially correlated
  - **Denominator:** common prediction, assumed to have Gaussian fluctuation of 2%
- Toy MC with correlated denominator with 2% fluctuation  $\rightarrow 10^6$  events
  - Estimate weighted average R of 19 random points with correlations around 0.943.
  - P-value for ( $R \geq 1$ ) : 1.4% ( $2.2\sigma$ ) compared to naive Gaussian  $2.4\sigma$ .
  - Our contours are reweighted by  $(2.2/2.4)^2$  to take this slight non-normality into account
- Hidden Covariance
  - $\chi^2_{\min}$  of data to straight line in the 18% quantile  $\rightarrow$  Data not incompatible with fluctuations

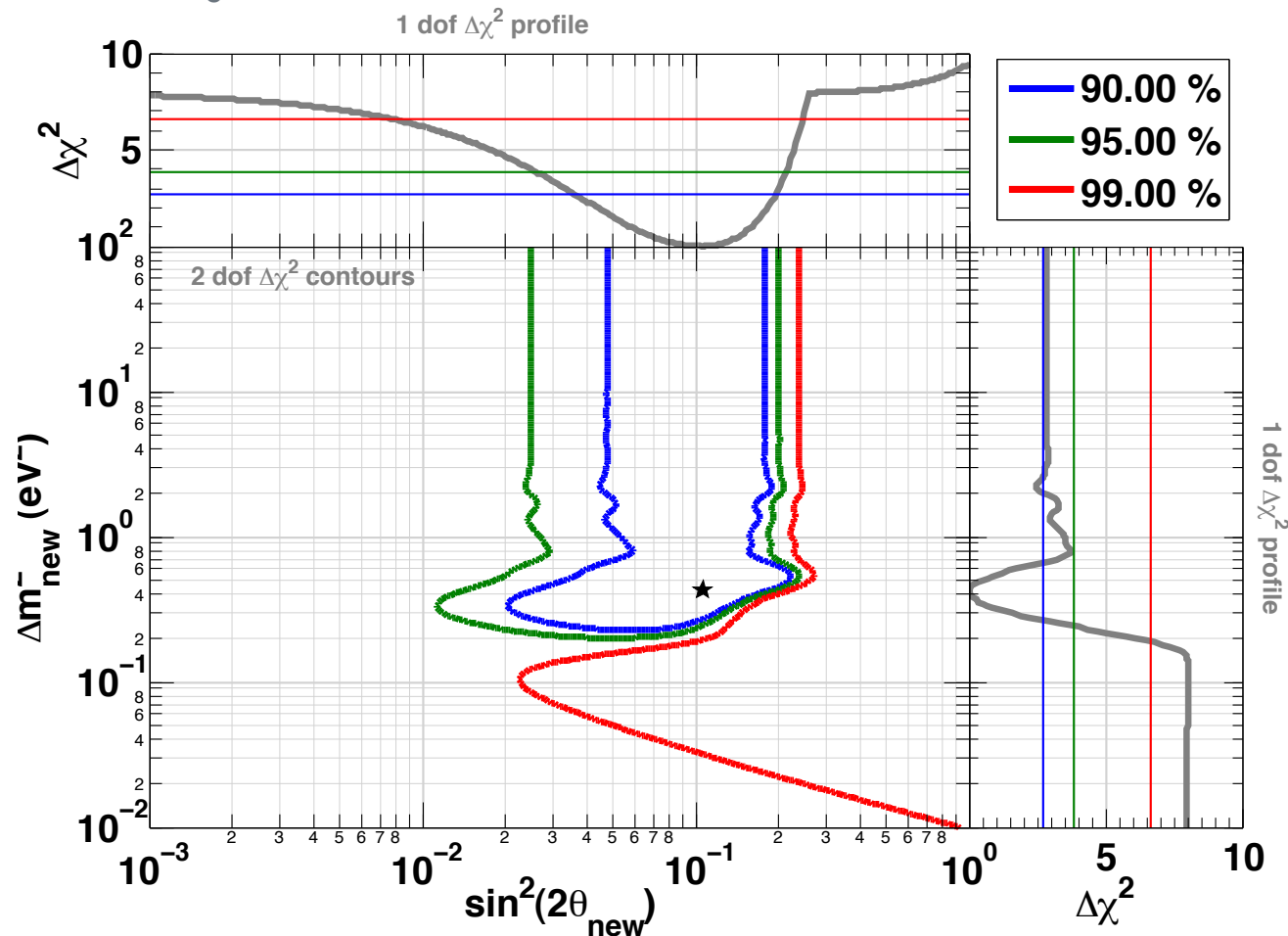


# The reactor rate anomaly

- 18/19 short baseline experiments  $< 100\text{m}$  from a reactor observed a deficit of anti- $\nu_e$  compared to the new prediction
- The effect is statistically significant at more 98.6%
- Effect partly due to re-evaluation of cross-section parameters, especially updated neutron lifetime, accounting for off equ. effect
- **At least three alternatives:**
  - Our conversion calculations are wrong. Anchorage at the ILL electron data is unchanged w.r old prediction
  - Bias in all short-baseline experiments near reactors : unlikely...
  - New physics at short baselines, explaining a deficit of anti- $\nu_e$  :
    - **Oscillation towards a 4<sup>th</sup>, sterile  $\nu$  ?**
    - **a 4<sup>th</sup> oscillation mode with  $\theta_{\text{new}}$  and  $\Delta m^2_{\text{new}}$**

# The 4<sup>th</sup> neutrino hypothesis

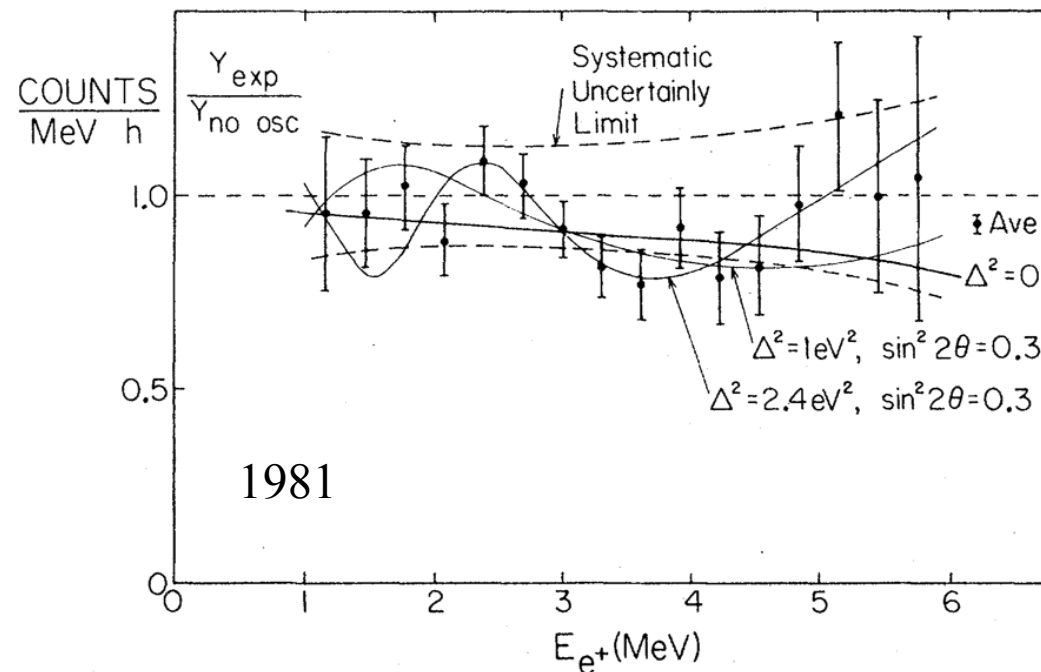
- Combine all rate measurements, no spectral-shape information
- Fit to anti- $\nu_e$  disappearance hypothesis



- Absence of oscillations disfavored at 98.6% C.L.

# The 1981 ILL measurement

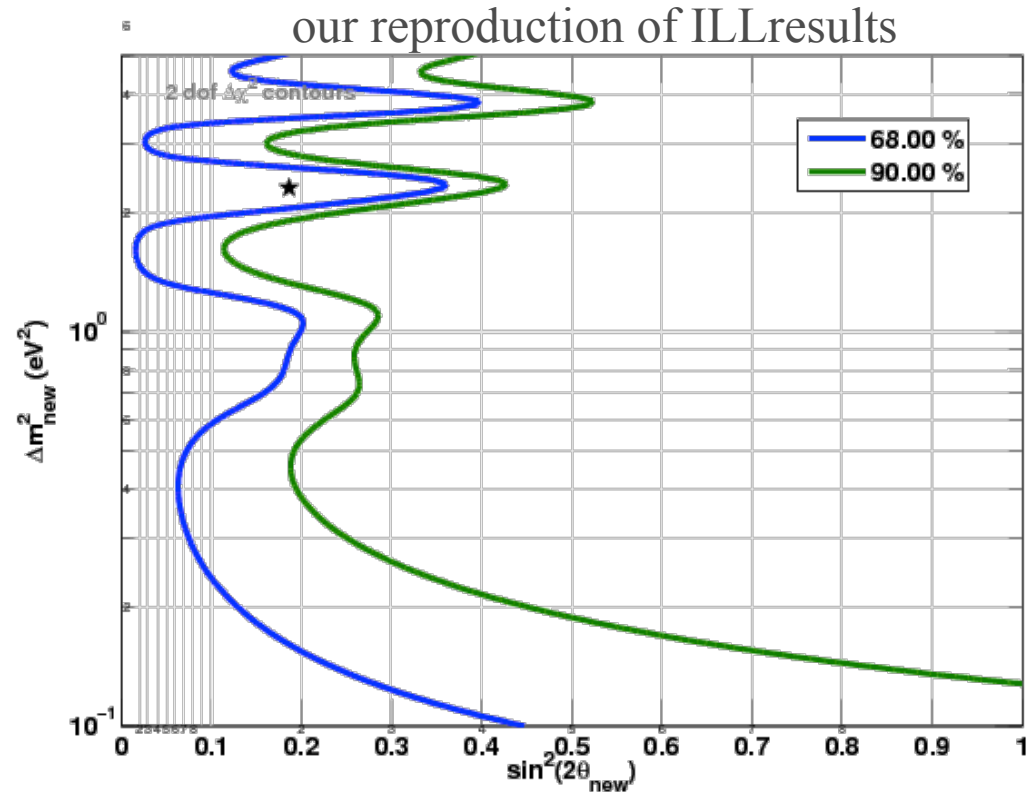
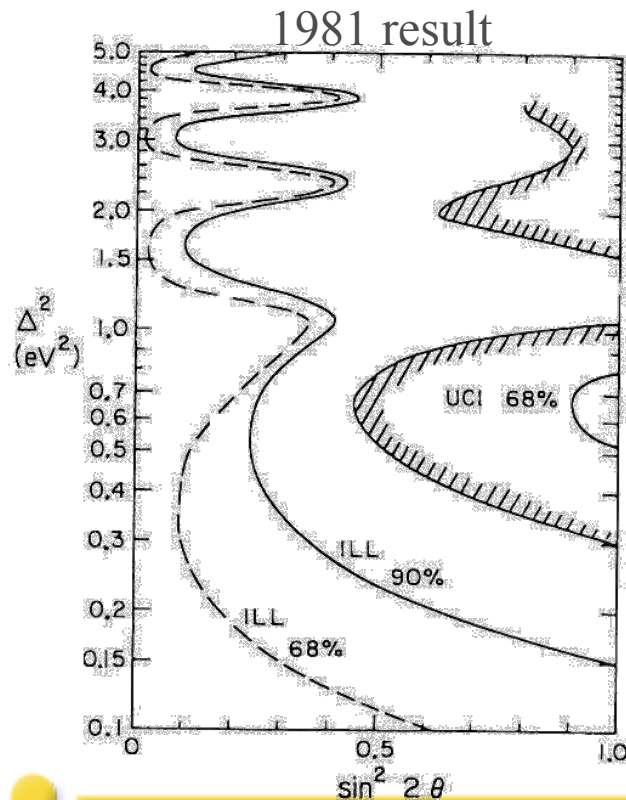
- Reactor at ILL with **almost pure  $^{235}\text{U}$** , with **compact core**
- Detector **8.76(?) m** from core. Any bias?
- Reanalysis in 1995 by part of the collaboration to account for overestimation of flux at ILL reactor by 10%... Affects the rate only



- Large errors, but a striking pattern is seen by eye ?

# Our ILL re-analysis (reproduce no-oscillation claim)

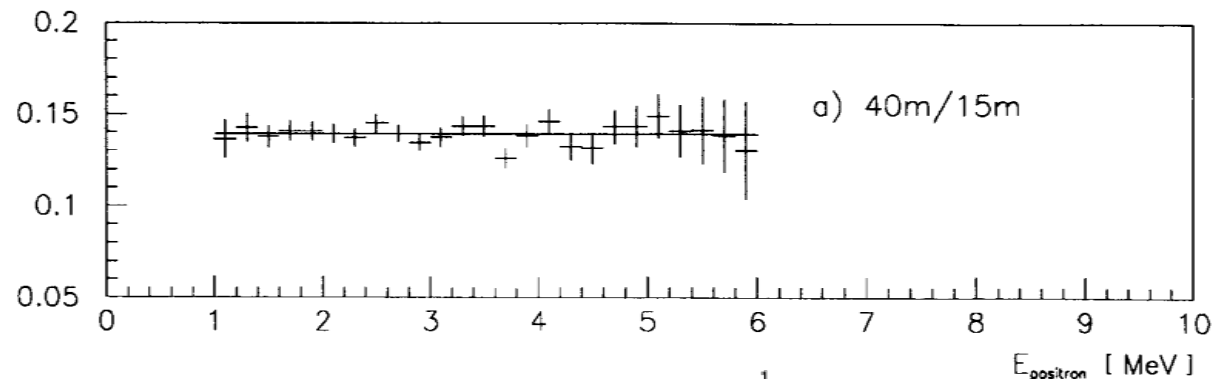
- 1981: Try to reproduce published contour
- 1995: Reproduce claim that global fit disfavors oscillation at  $2\sigma$
- **How ?** We add uncorrelated systematic in each bin until it's large enough
  - **Needed error : 11%, uncorrelated, in each bin.**





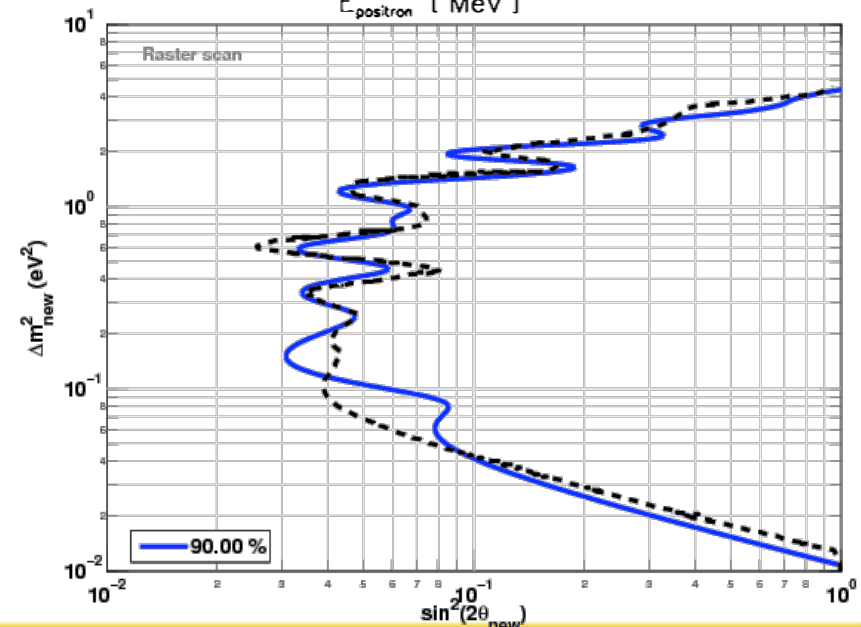
# Spectral shape analysis of Bugey-3

- Bugey-3 spectral measurements at 15 m, 40 m, 90 m
  - Best constraint from high statistics R=15m/40m ratio
  - Very robust since it does not rely on reactor spectra

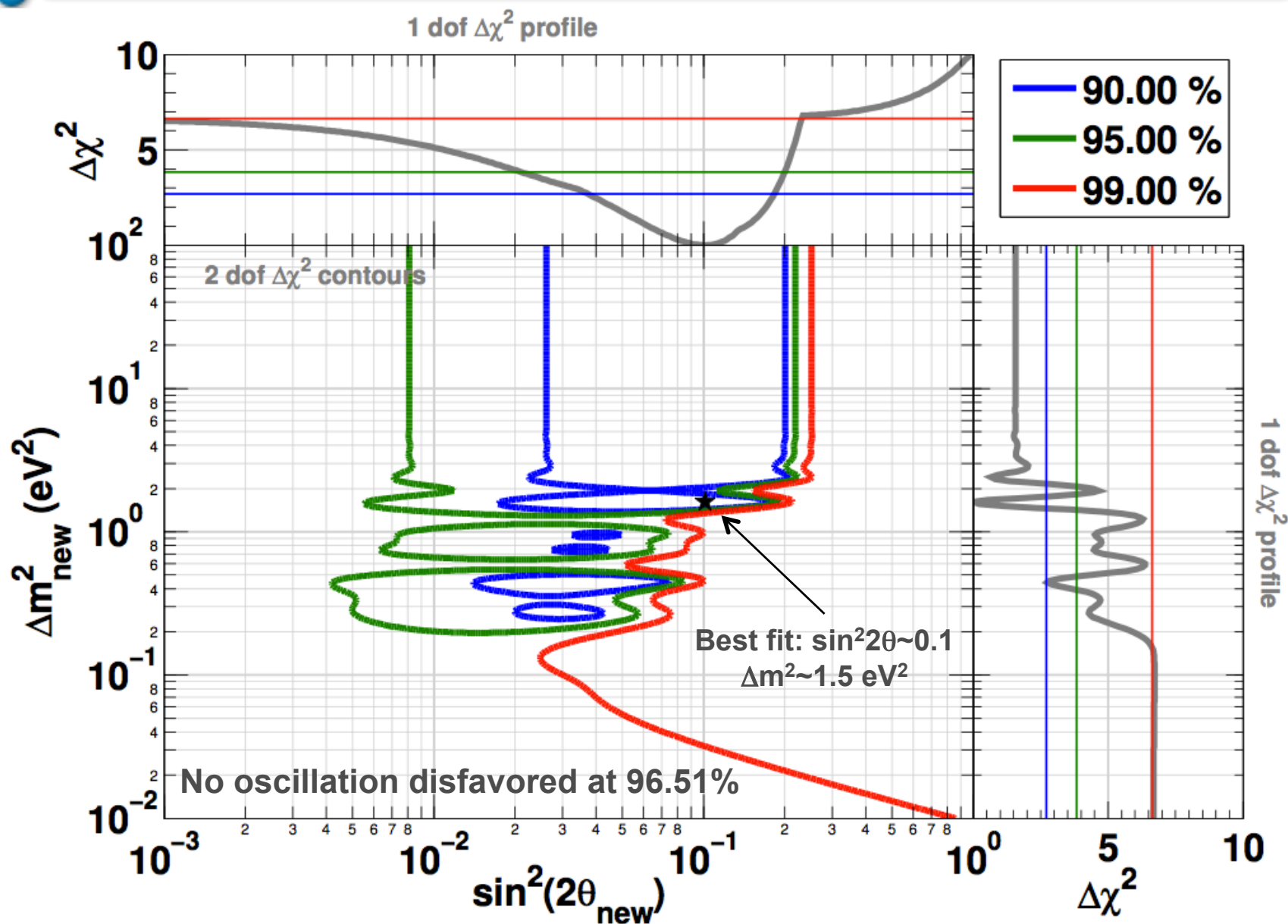


$$\chi^2 = \sum_{i=1}^{N=25} \left( \frac{(1+a)R_{th}^i - R_{obs}^i}{\sigma_i} \right)^2 + \left( \frac{a}{\sigma_a} \right)^2$$

- Reproduction of the collaboration's raster-scan analysis
- Use of a global-scan in combined analysis



# Combined Reactor Rate+Shape contours





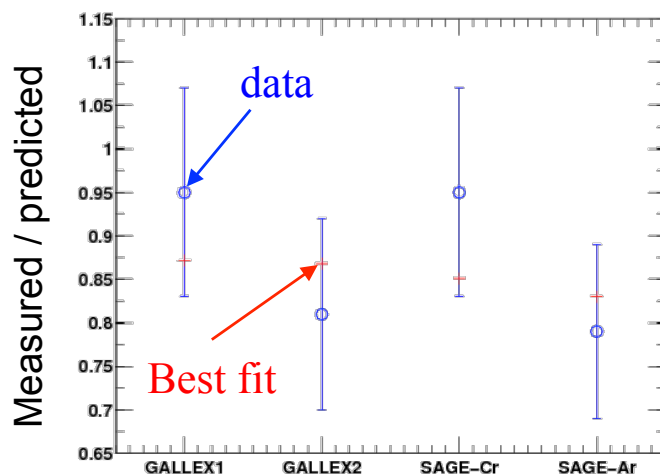
# The Gallium Neutrino Anomaly

Based on PRD82 053005 (2010)

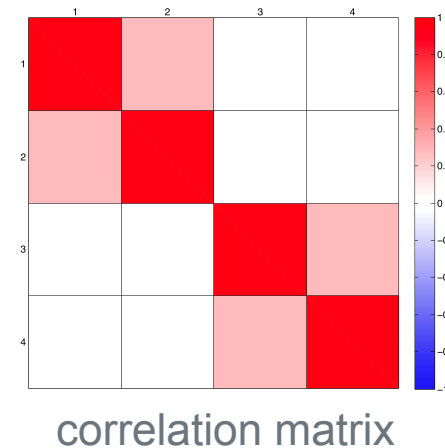
See C. Giunti's talk

# The Gallium anomaly

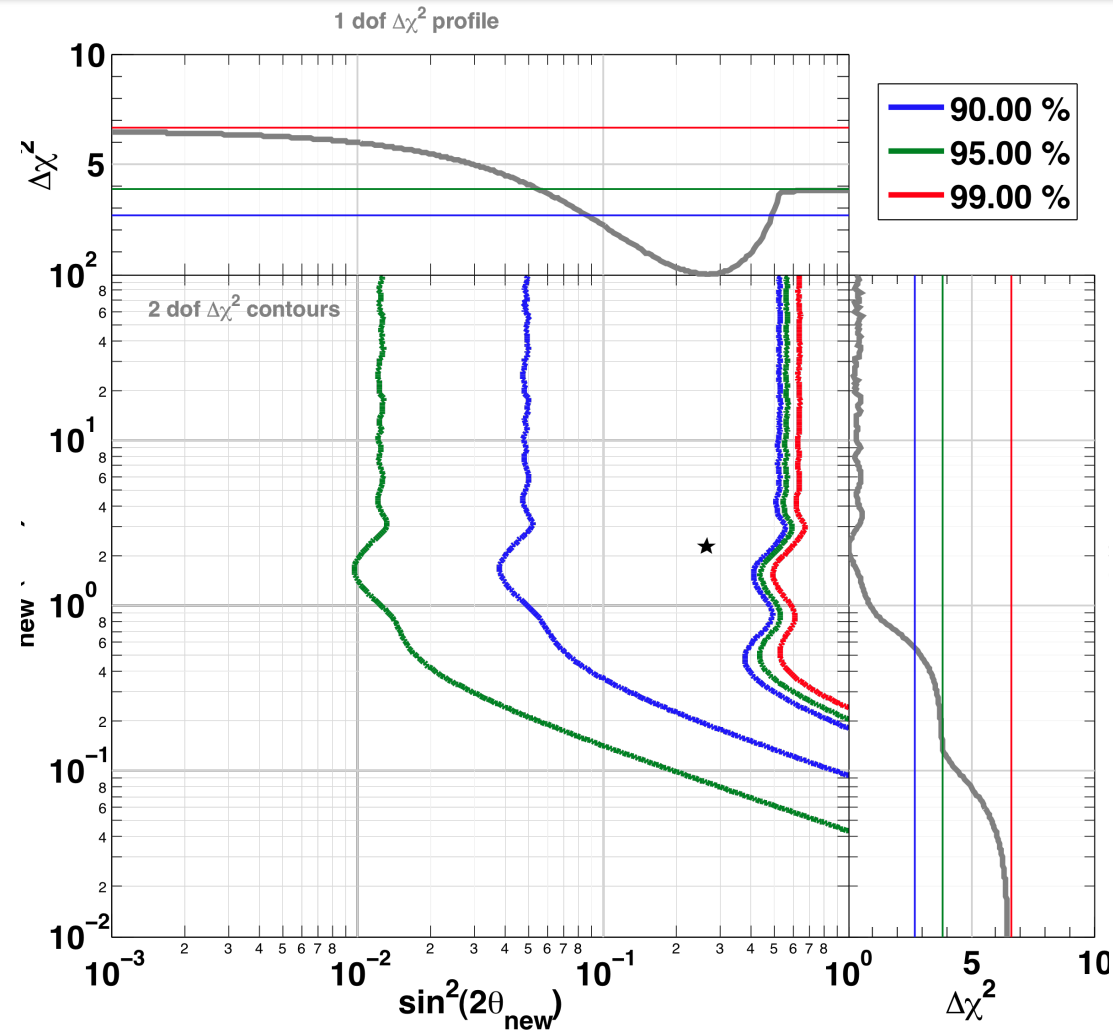
- 4 calibration runs with intense MCi neutrino sources:
  - 2 runs at Gallex with a  $^{51}\text{Cr}$  source (750 keV  $\nu_e$  emitter)
  - 1 run at SAGE with a  $^{51}\text{Cr}$  source
  - 1 run at SAGE with a  $^{37}\text{Ar}$  source (810 keV  $\nu_e$  emitter)
  - All observed a deficit of neutrino interactions compared to the expected activity. Hint of oscillation ?
- Our analysis for Gallex & Sage:
  - Monte Carlo computing mean path lengths of neutrinos in Gallium tanks
  - **NEW** : Correlate the 2 Gallex runs together & the 2 SAGE runs together



- Gallex-I
- Gallex-II
- Sage-Cr
- Sage-Ar

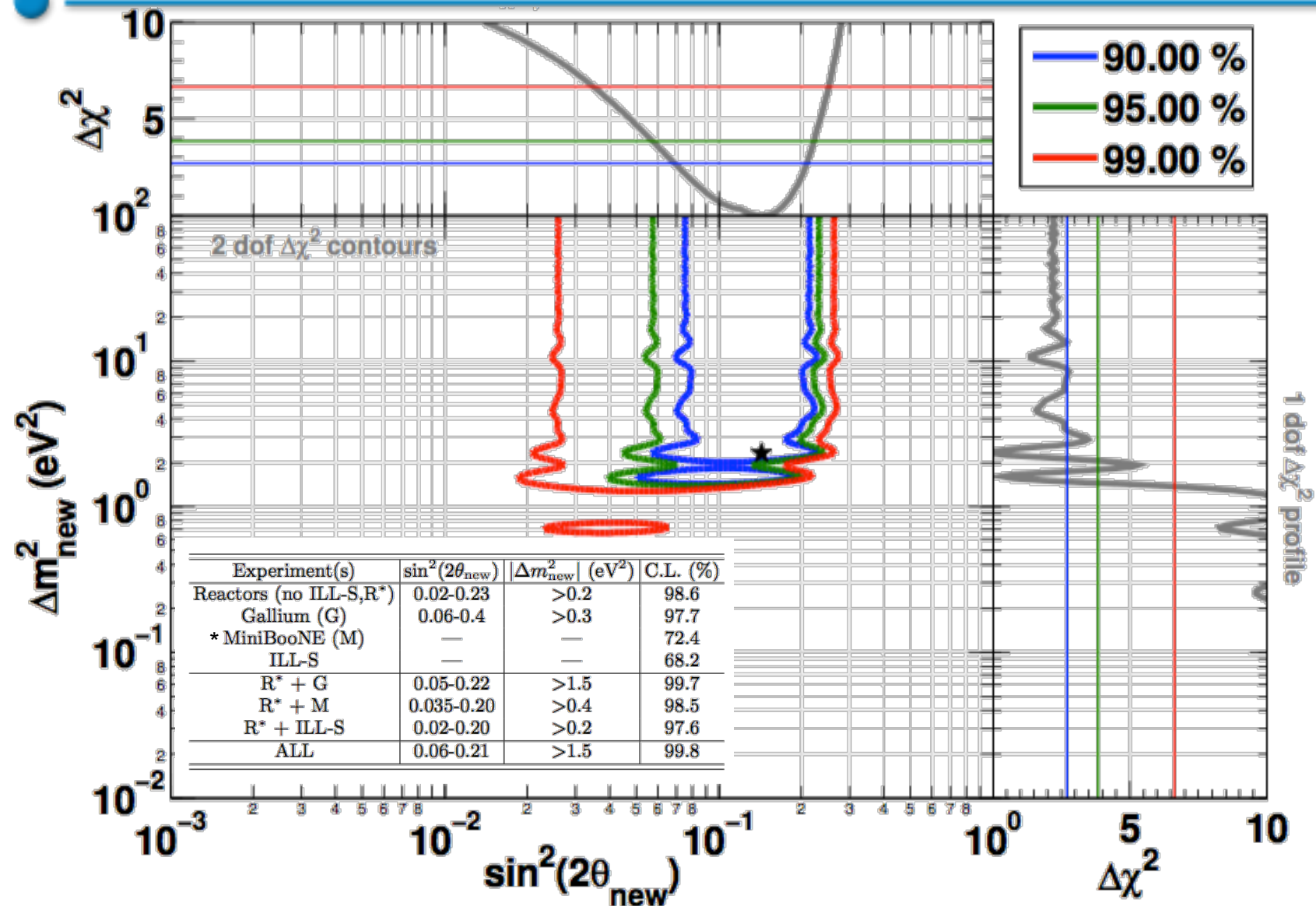


# The Gallium anomaly



- Effect reported in C. Giunti & M. Laveder in PRD82 053005 (2010)
- Significance reduced by additional correlations in our analysis
- No-oscillation hypothesis disfavored at 97.7% C.L.

# Putting it all together: reactor rates + shape + Gallium + (MB)



**The no-oscillation hypothesis is disfavored at 99.8% CL**



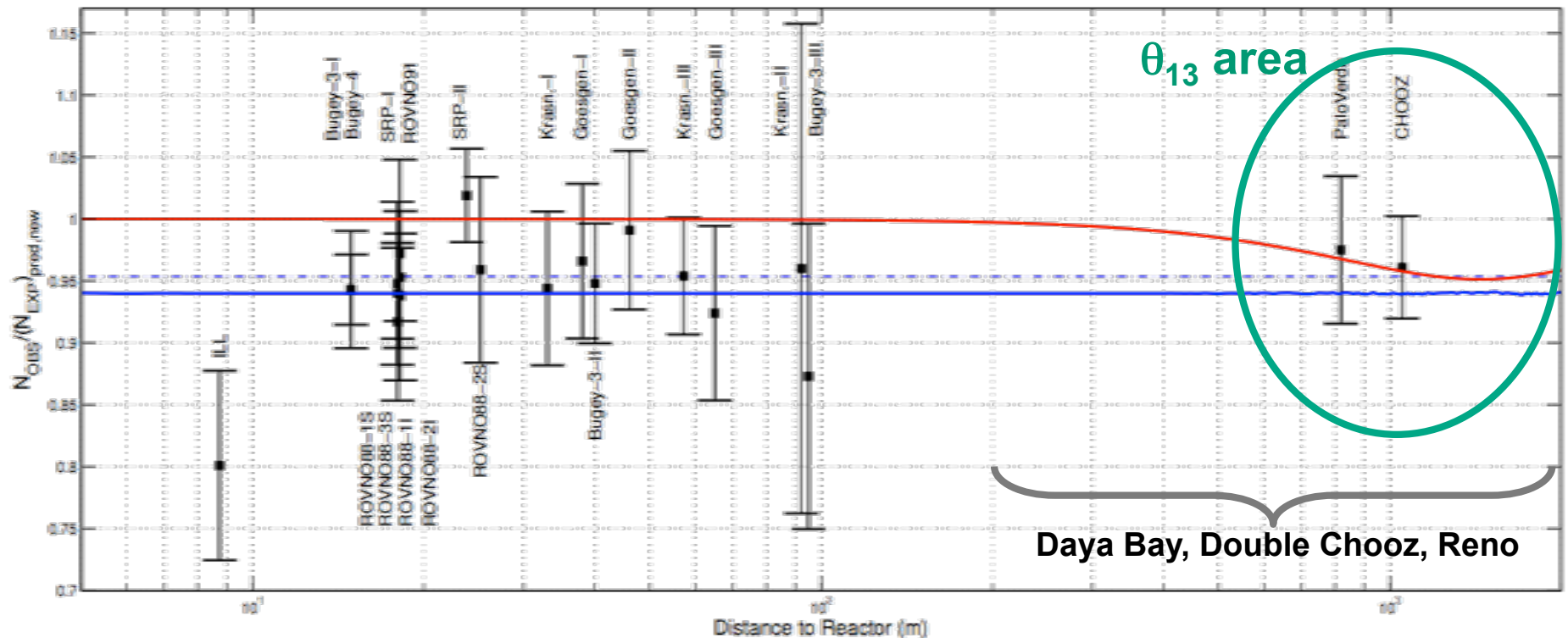
# Implication for $\theta_{13}$

# Implication for $\theta_{13}$ at 1-2 km baselines

- The choice of normalization is crucial for reactor experiments looking for  $\theta_{13}$  without near detector

$\sigma_f^{\text{pred,new}}$  : new prediction of the antineutrino fluxes

$\sigma_f^{\text{ano}}$  : experimental cross section (best fitted mean averaged)

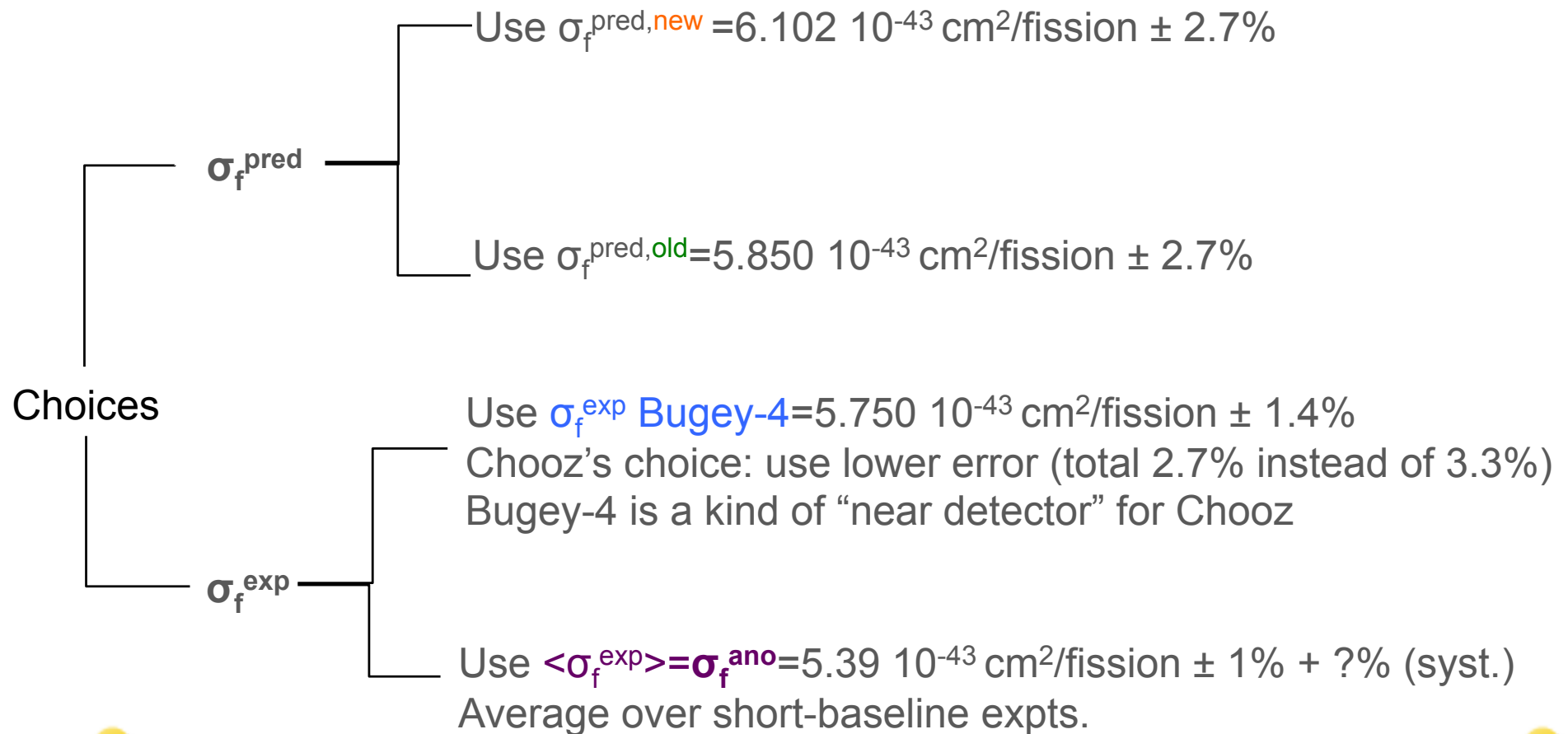


- A deficit observed at 1-2 km can either be induced by  $\theta_{13}$  induced oscillation BUT also by other explanations (experimental, new  $\phi$ , ...)



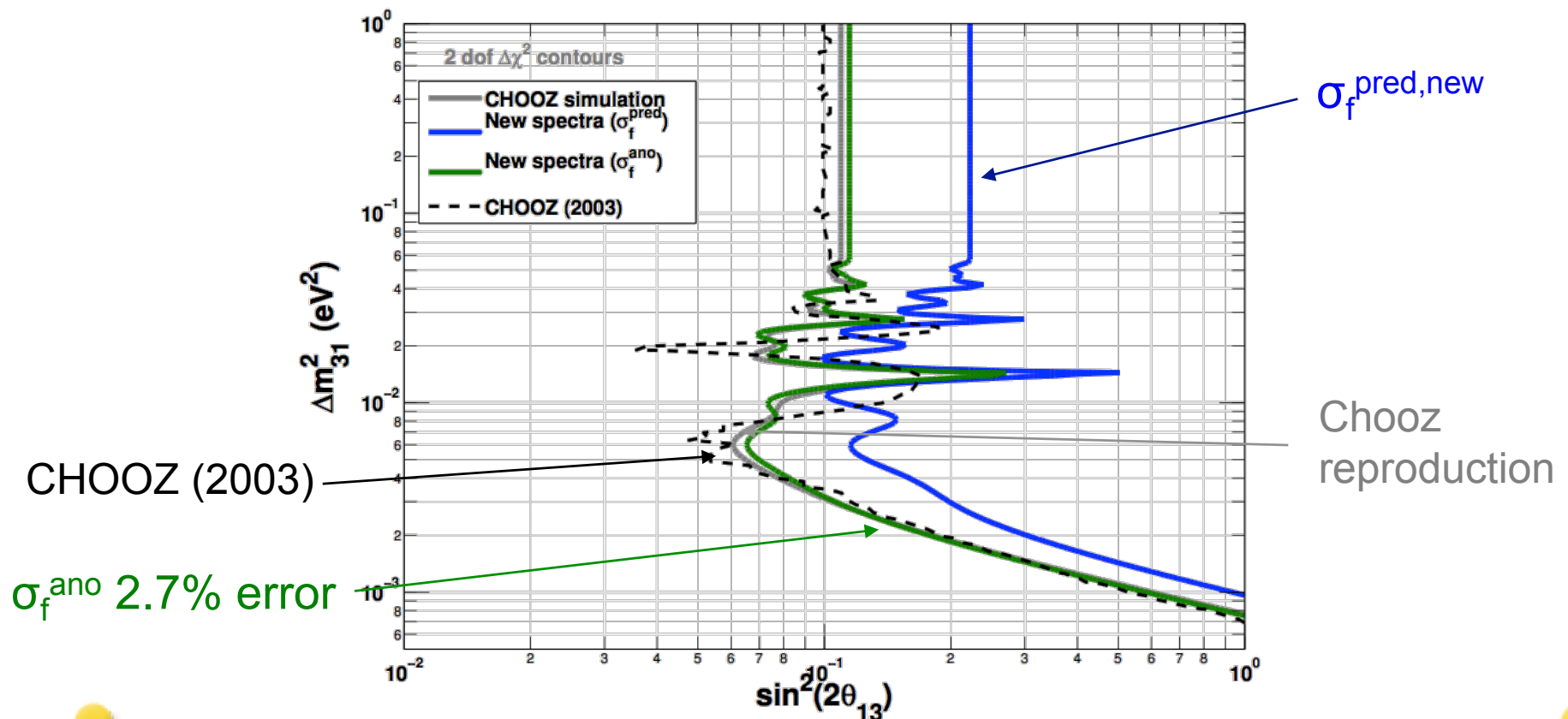
# The Normalization Dilemma

- Experiments with baselines > 500 m
- How do you normalize the expected flux, knowing the fuel composition?
- **If near + far detector, not an issue anymore**



# CHOOZ reanalysis

- The choice of  $\sigma_f$  changes the limit on  $\theta_{13}$
- Chooz original choice was  $\sigma_f^{\text{exp}}$  from Bugey-4 with low error
- If  $\sigma_f^{\text{pred,new}}$  is used, limit is worse by factor of 2
- If  $\sigma_f^{\text{ano}}$  is used with 2.7%, we obtain the original limit  
 → But which error should we associate to  $\sigma_f^{\text{ano}}$  (burnup up error?)



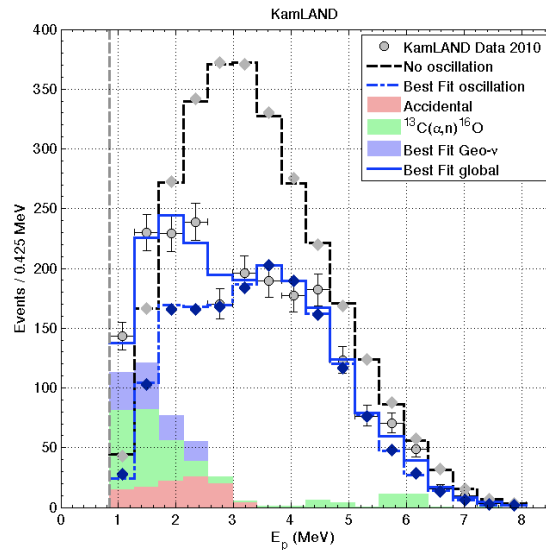
# Reanalysis of KamLAND's 2010 results

arXiv:1009.4771v2 [hep-ex]

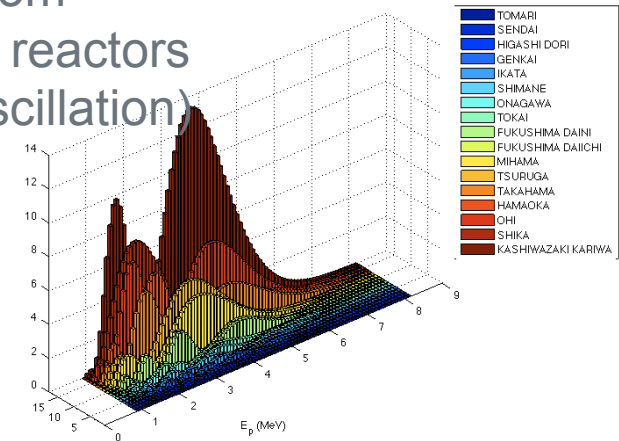
## Systematics

	Detector-related (%)	Reactor-related (%)
$\Delta m_{21}^2$	Energy scale 1.8 / 1.8	$\bar{\nu}_e$ -spectra [31] 0.6 / 0.6
Rate	Fiducial volume 1.8 / 2.5	$\bar{\nu}_e$ -spectra 2.4 / 2.4
	Energy scale 1.1 / 1.3	Reactor power 2.1 / 2.1
	$L_{cut}(E_p)$ eff. 0.7 / 0.8	Fuel composition 1.0 / 1.0
	Cross section 0.2 / 0.2	Long-lived nuclei 0.3 / 0.4
	Total 2.3 / 3.0	Total 3.3 / 3.4

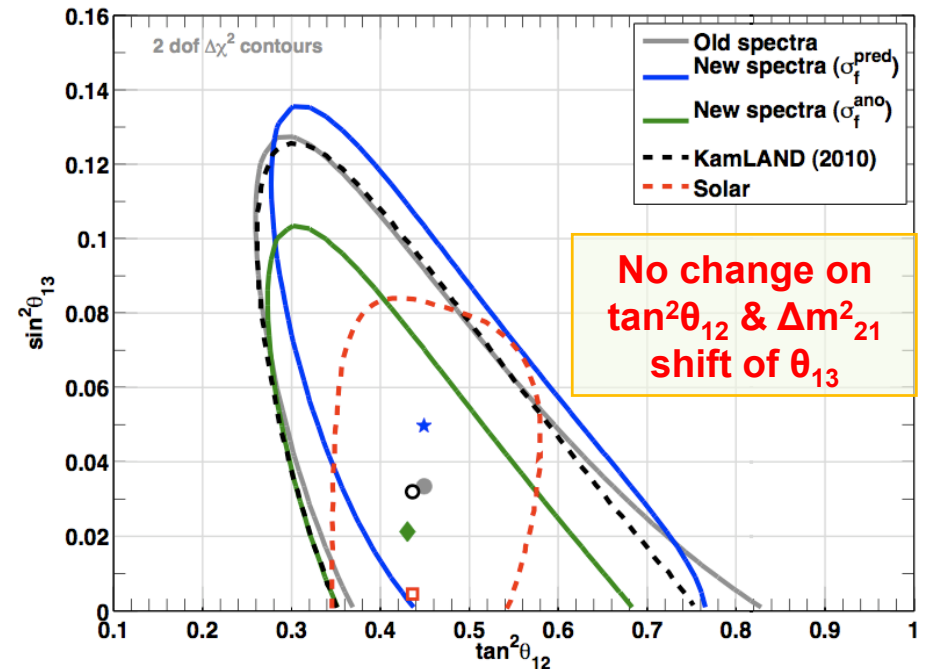
Reproduced KamLAND spectra within 1% in [1-6] MeV range



Spectra from Japanese reactors (with  $\nu_e$  oscillation)



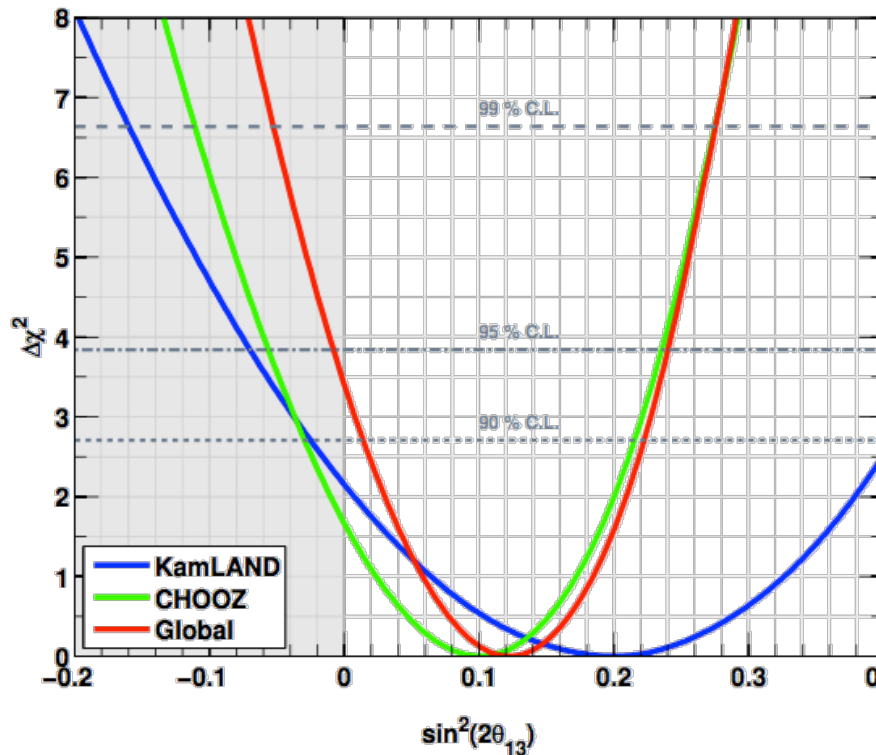
With new spectra predictions



# CHOOZ and KamLAND combined limit on $\theta_{13}$

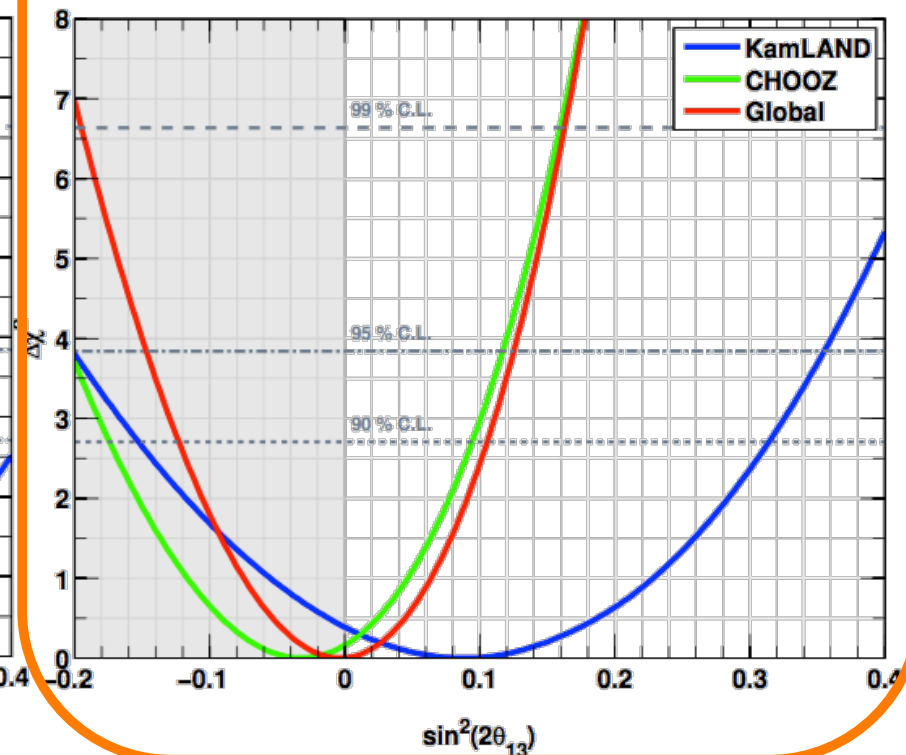
Normalization with  $\sigma_f^{\text{pred,new}}$

3-v framework & 2.7% uncertainty



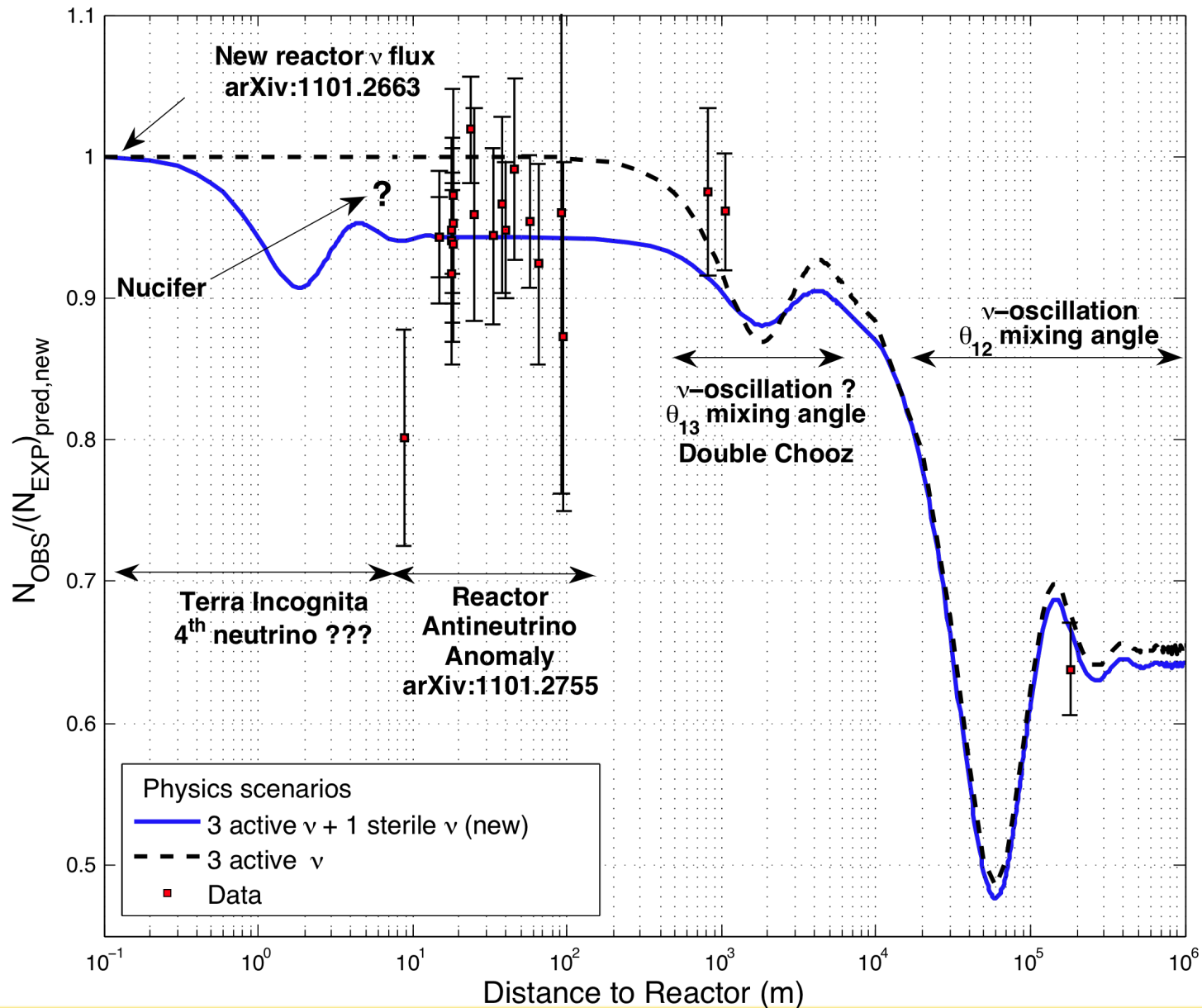
Normalization using  $\sigma_f^{\text{ano}}$

3-v framework & 2.7% uncertainty



- **Our interpretation** (different from Arxiv:1103:0734 for KamLAND- $\sigma_f^{\text{pred,new}}$ , T. Schwetz's talk)
  - No hint on  $\theta_{13} > 0$  from reactor experiments :  $\sin^2(2\theta_{13}) < 0.11$  (90% C.L., 1dof)
  - CHOOZ 90 % CL limit stays identical to Eur. Phys. J. C27, 331-374 (2003)
  - Multi-detector experiments are not affected

# Need for new experimental inputs !



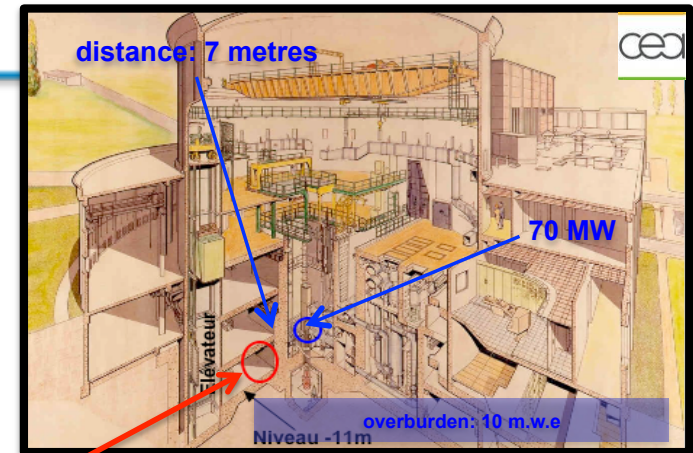
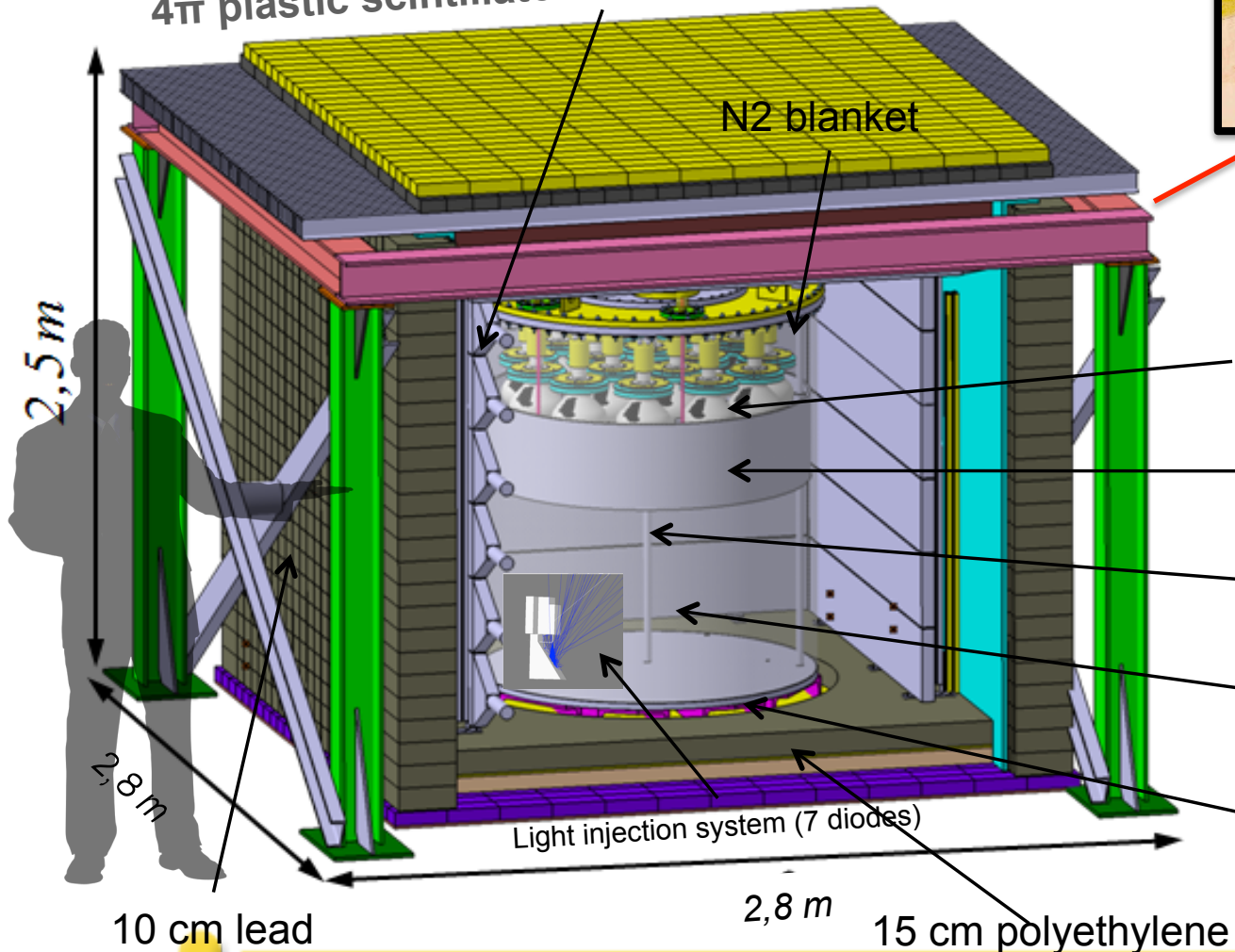
# Nucifer

**First goal: Non Proliferation**

Thermal Power Measurement

Fuel Composition Measurement U/Pu

**4 $\pi$  plastic scintillator Muon Veto (30 PMTs)**



Osiris research reactor  
CEA-Saclay (600 v/d)  
CEA – IN2P3 coll.

16 x 8' PMTs low background

25 cm acrylics buffer

Calibration pipe

Target: 0.85 m<sup>3</sup> Gd-LS (0.5%)

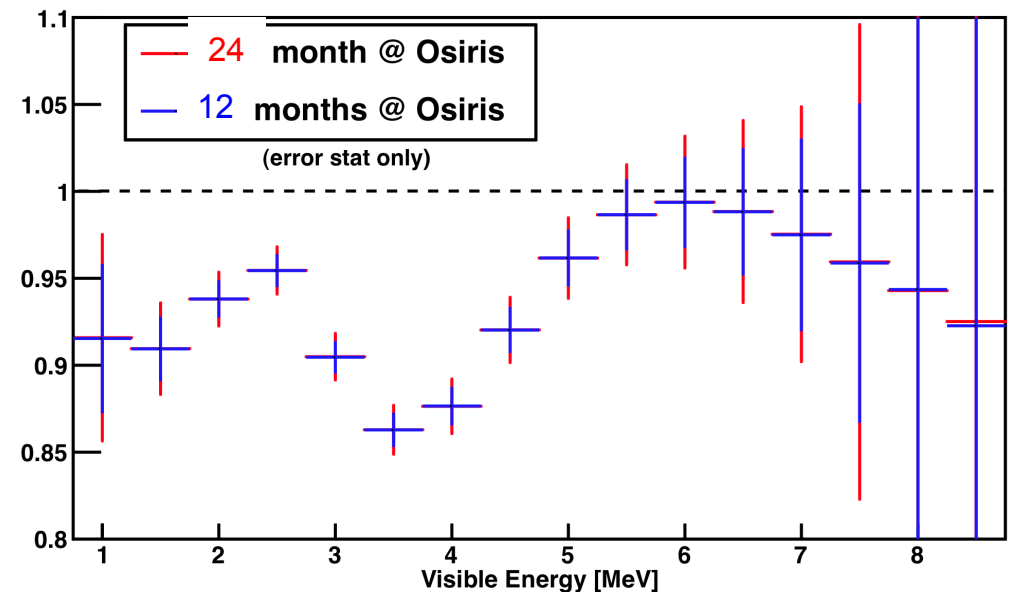
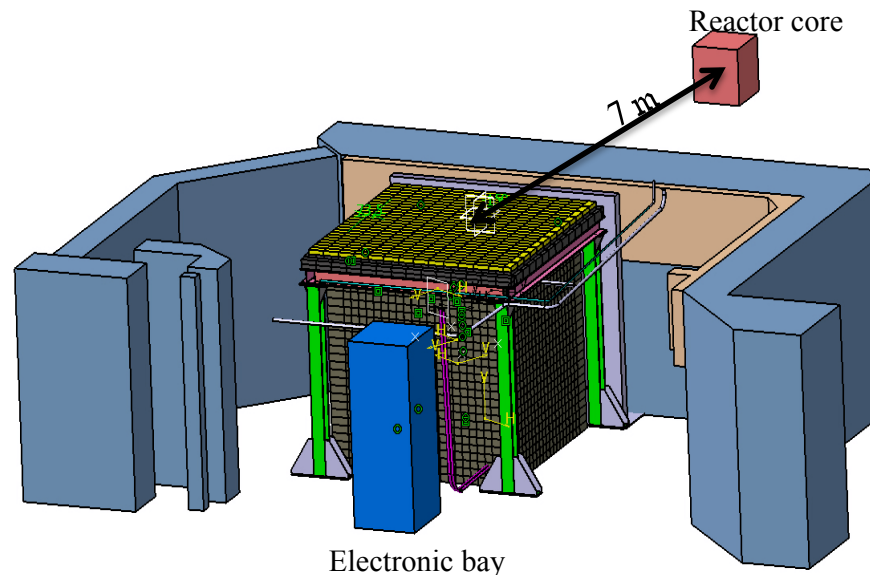
Stainless steel double  
containment vessel coated with  
white Teflon coating inside





# NUCIFER in Saclay

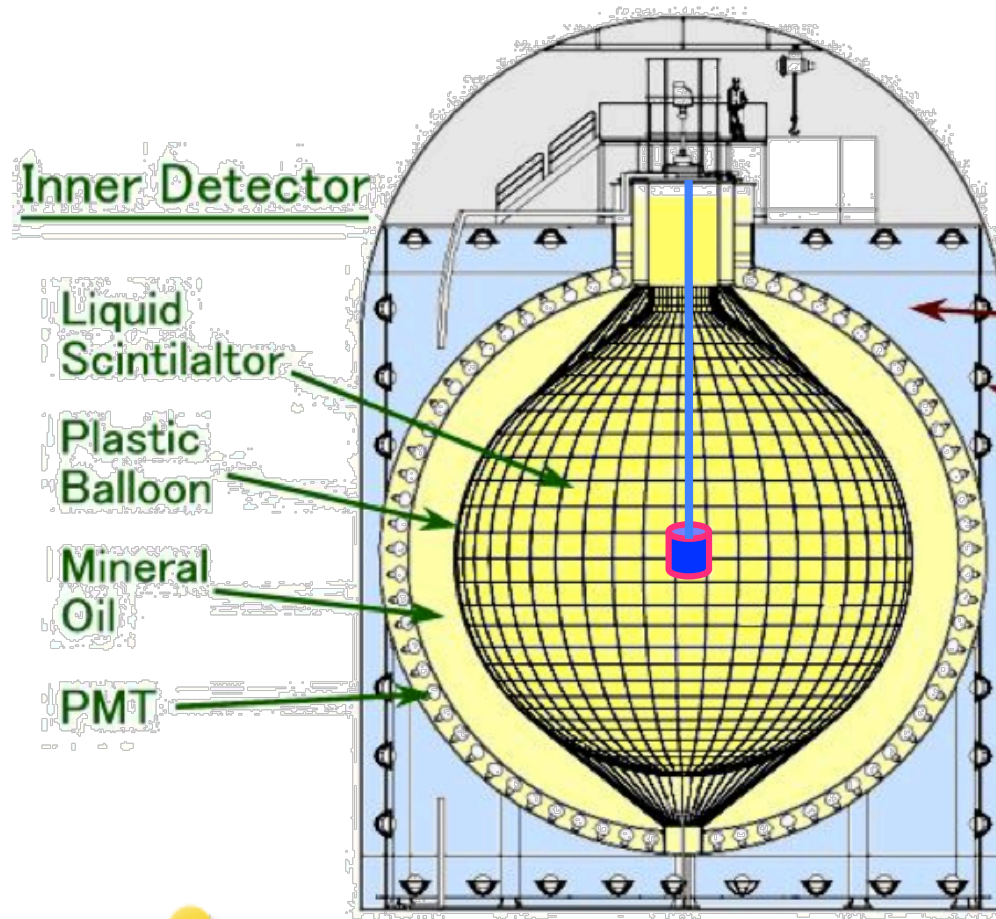
- Osiris-Saclay: Core Size: 57x57x60 cm
- Nucifer Detector Size : 1.2x0.7m (850l)
- Baseline distribution
  - $\langle L \rangle = 7.0$  m,  $\sigma = 0.3$  m  $\rightarrow$   $eV^2$  oscillations are not washed out
- Folding Nucifer Geant4 Monte Carlo detector response
- $\Delta m^2 = 2.4$   $eV^2$  &  $\sin^2(2\theta) = 0.15$
- No backgrounds. Thus to be taken with a grain of salt ...



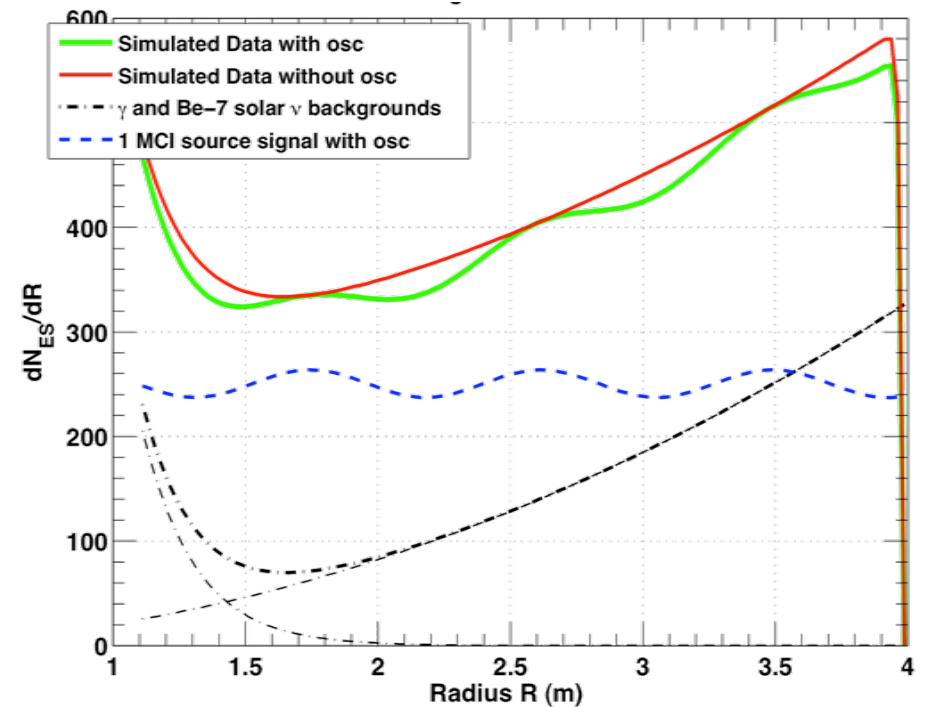
- Such pattern could not be seen at Bugey-3 (extended core & 14 m baselin

# Mci $^{51}\text{Cr}/^{37}\text{Ar}$ Experiment Concept

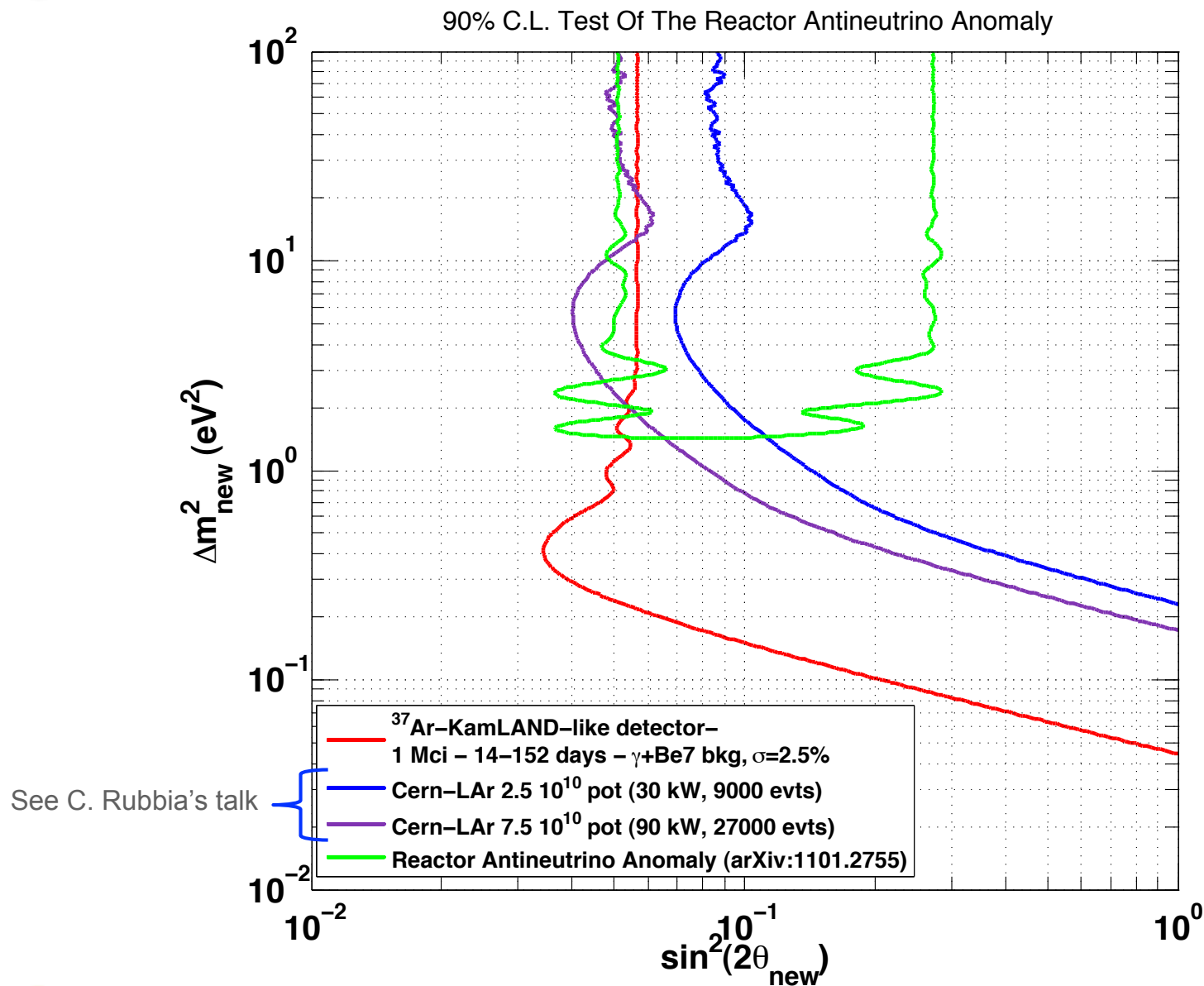
- A strong 1 Mci  $\nu$  source in the middle of a large LS detector
- Elastic scattering on electrons (few 1000 evts, 150 days,  $>250$  keV)
- A good resolution in position (20 cm)



Real oscillation pattern VS radius  
(preliminary)



# Promising experimental prospect testing the RAA!



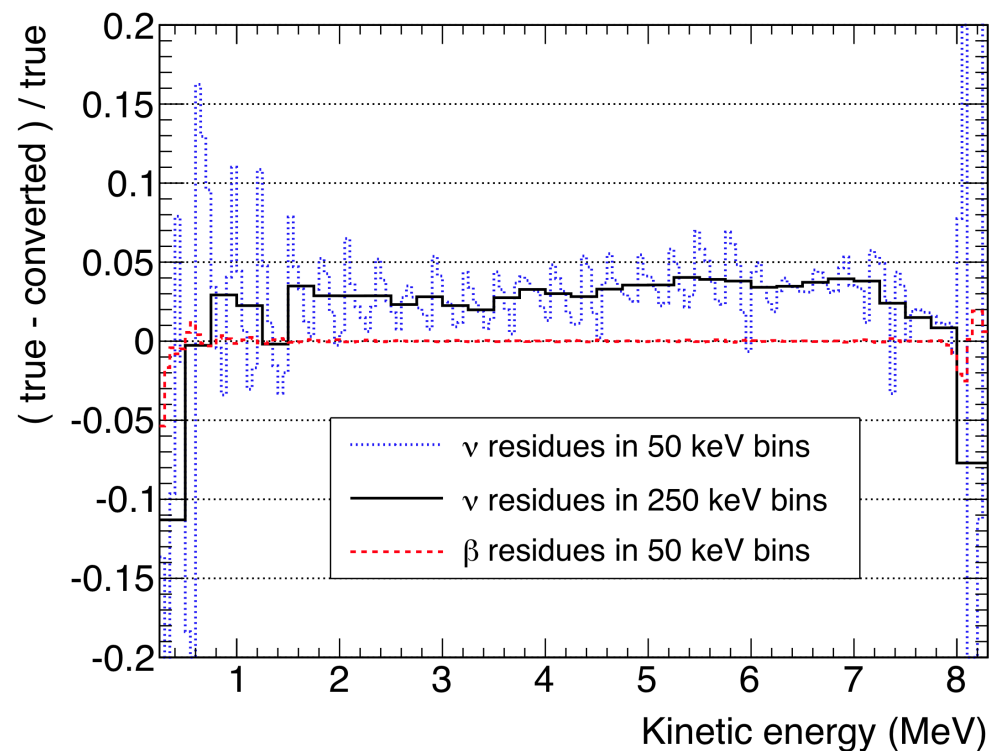


# Backup

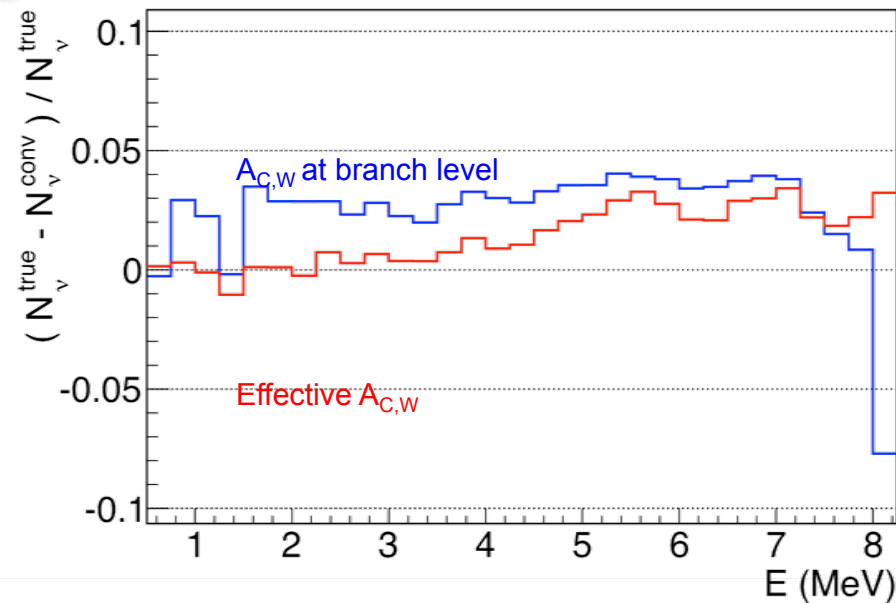
# Consistency Check

## Stringent test

1. Define “true”  $e^-$  and  $\nu$  spectra from reduced set of well-known ENSDF branches.
2. Apply exact same ILL conversion procedure to true  $e^-$  spectrum.
3. Compare converted  $\nu$  spectrum to the true one.
4. ‘old’ effective method bias is confirmed

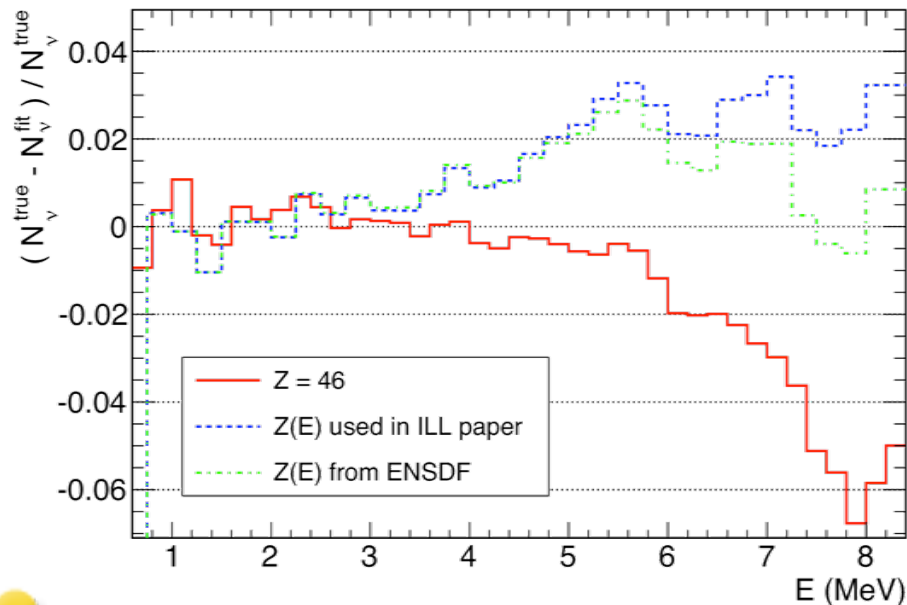


# Origin of the 3% shift



- **$E < 4$  MeV:** deviation from effective linear  $A_{C,W}$  correction of ILL data

$$\Delta N_v^{C,W}(E_v) \approx 0.65 \times (E_v - 4 \text{ MeV}) \quad \%$$



- **$E > 4$  MeV:** mean fit of  $Z(E_0)$  doesn't take into account the very large dispersion of  $Z$  around the mean curve

$$Z(E_0) \approx 49.5 - 0.7E_0 - 0.09E_0^2, \quad Z \geq 34$$