FDIRC tests in the SLAC cosmic ray telescope

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# **FDIRC**

### • Successfully tested in the SLAC test beam:

- J. Benitez, I. Bedajanek, D.W.G.S. Leith, G. Mazaheri, B.N. Ratcliff, K. Suzuki, J. Schwiening, J. Uher, and J. Va'vra, "Development of a Focusing DIRC," IEEE Nucl. Sci., Vol. 3, pp. 1550-1556, Conference Record, October 2006, and SLAC-PUB-12236.
- J. Va'vra, J. Benitez, D.W.G.S. Leith, G. Mazaheri, B. Ratcliff, J. Schwiening, and K. Suzuki, "The Focusing DIRC the first RICH detector to correct the chromatic error by timing, and the development of a new TOF detector concept," Presented at Vienna conference on Instrumentation, February 19, 2007, Vienna, Austria, September 2007, SLAC-PUB-12803.
- J. Benitez, D.W.G.S. Leith, G. Mazaheri, B.N. Ratcliff, J. Schwiening, J. Va'vra, L. Ruckman, G. Varner, "Status of fast forcusing DIRC (fDIRC)," Nucl. Inst. & Meth., A595(2008)104-107.
- J. Va'vra, "Simulation of the Focusing DIRC Optics with Mathematica," Presented at IEEE in Dresden in 2008, SLAC-PUB-13464. October, 2008; IEEE conference records.

### • FDIRC main points:

- Its "pixel imaging quality" is the same as that of BaBar DIRC the same  $\sigma_{\theta}$
- However, it is ~10x faster that BaBar DIRC =>
  - => can correct chromatic error by timing,
  - => less sensitive to background.
- Its "detector canvas" size is ~2x smaller than a non-focusing DIRC with similar detector concept.

#### FDIRC tests in the SLAC cosmic ray

### **Focusing DIRC Prototype Optics**



#### • Radiator:

– 1.7 cm thick, 3.5 cm wide, 3.7 m long fused silica bar (spares from BABAR DIRC).

### • Optical expansion region:

- filled with a mineral oil to match the fused silica refraction index (KamLand oil).
- include optical fiber for the electronics calibration (PiLas laser diode).
- Focusing optics:
  - a spherical mirror with 49 cm focal length focuses photons onto a detector plane.

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FDIRC tests in the SLAC cosmic ray



telescope



### SLAC cosmic ray telescope - our "test beam" for the next 1-2 years





- ~ 4 feet of iron (an old TPC magnet) => ~ 1.6 GeV muon energy cutoff
- Tracking resolution: ~1 mrad.
- Status: taking data with the 1-st Hawaii electronics package; 6 more in March.

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## Test setup in the cosmic ray telescope



## **Images in the SLAC cosmic ray telescope**

Side view:



- Will start with a nominal position:  $\theta_{dip} \sim 90^{\circ}$  & position # 3.
- Need to shift the detector plane to reach smaller dip angles (~1.8 cm/10<sup>o</sup> dip angle).
  - With some modification we could reach, perhaps,  $\theta_{dip} \sim 75^{\circ}$ .

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# **New readout configuration**



### **Cosmic ray telescope counters:**

#### 2) Linear scale now:



- As one goes to lower stack counter number of TDC zeroes increases, as one would expect.

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# **Monitoring of FDIRC in the telescope**



#### See forward and backward part of the Cherenkov ring in time • domain, the PiLas laser monitoring pulses and the BLAB2 trigger

time (ns)

# The new Hawaii electronics



## **FDIRC electronics chain**

Gary Varner, Larry Ruckman, Kurtis Nishima, and Andrew Wong, Nucl.Instrum.Meth. A591,534-545,2008



#### **Status:**

- One H-8500 is running in the telescope.
- Will have six more ready in March.



# **BLAB ASIC further studies**



BLAB1 chip Nucl.Instrum.Meth. A591,534-545,2008, and arXiv:0805.2225





- Comparable timing performance to best CFD + HPTDC methods
- MUCH lower power, no need for huge cable plant!
- Using full samples significantly reduces the impact of noise
- Photodetector limited



## **Summary of BLAB2 ASIC parameters**

G. Varner

Parameter	Value
Samples	6 rows & 1024 samples
Triggering mode	Trigger on individual pixel
Analog BW of the present BLAB2 chip	~ 0.85 GHz
BLAB2 chip input impedance	<b>30-80</b> $\Omega$ (adjustable)
Number of MaPMT pixels / BLAB2 ASIC	16
Number of BLAB2 ASICS / MaPMT (final vs. prototype)	2 - 4
Dynamic range	1mV / 1V
Cross-talk	< 0.1%
BLAB2 wafeform sampling speed	1 - 10 GSa/sec
On chip ADC	1 GHz Wilkinson
Number of Wilkinson conversions in parallel	32
ADC resolution	10 bits
ADC conversion time for 10 bits	1 µs
Number of words / event	32 - 512
Read time for 16 channels (1 BLAB2 chip) / event	<b>16 μs</b>
Sustained readout speed	50 kHz
12 μs latency accomplished by	Self-trigger & analog or digital storage
Cost per chanel	< \$10 in volume

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# Details

### Waveform sampling principle: Switched Capacitor Array

G. Varner, Larry Ruckmann and A. Wong, arXiv:0802.2278v1 [physics.ins-det], submitted to NIM, 2008



### **Principle of wafeform sampling in BLAB chip:**

- Fast signal waveform is terminated in 50  $\Omega$  on the chip
- BLAB2 ASIC: 17ch & 6 rows & 1024 samples (102k storage cells)
- Each row can be independently addressed to initiate a storage cycle
- When analog switch is closed the instantaneous signal is stored on a 14 fF capacitor
- ADC conversion is done via the Wilkinson method:

Comparators done inside BLAB chip & high speed encoding done in FPGA

- 32 samples are converted in parallel
- 10-bits corresponds to a conversion time of 1  $\mu$ s in the current scheme.
- Up to 10 GSa/s (100ps sampling interval <=> ~5 samples on the leading edge); the present limit: 2.5GSa/s.
- 12 µs latency: (a) TOF: self-trigger and digital temporary storage

(b) FDIRC: analog storage

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# Highly Integrated Readout

### Buffered LABRADOR

TABLE II: BLAB2 ASIC Specifications.

Item	Value
Photodetector Input Channels	16
Linear sampling arrays/channel	2 6
Storage cells/linear array	512 10
Sampling speed (Giga-samples/s)	2.0 - 10.0
Outputs (Wilkinson)	32





### **BLAB2 ASIC**



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### **BLAB2** lectronics for the next beam or cosmic ray tests at SLAC

Gary Varner & Larry Ruckman



- Prototype boards with BLAB1 chip now exists; BLAB2 chip is almost ready and will be sumbitted to a foundry on June 2.
- To instrument all 64 pixels one needs 4 BLAB chips. That is what we want to have in the prototype in the fall. However, the final DIRC will very likely use 2 chips per MaPMT (2 pads ganged together, 32 pixels/MaPMT).
- Worry about a minituarization later.

# **Readout Requirements**



4x fiber pairs/FINESSE (= 1 iTOP) module 2x dual Shark DSP/FINESSE



4 FINESSE/COPPER (2k chan) 8 COPPER Total 16k total system channels

Even less if use GaAs (4-channel PMTs)

8x COPPER-sized cards for trigger collection/Trigger fan-out

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# Data volume estimate





Assume: 100kHz charged track hits on each bar

~32 p.e./track (1% of 100ns windows) 30kHz trigger rate Each PMT pair sees  $\langle 8 \rangle$  hits 240k hits/s Each hit = 64samples \* 8bits = 512bits  $\rightarrow$ ~125Mbits/s (link is 1.2Gb/s ~ x10 margin)

### **BlackFin DSP**

Pedestal subtract
Feature extract → T, Q
(tentatively allow up to 4x hits in 100ns)
Time = 2Bytes, Q = 2Bytes

8 hit chan \* 1 hit typ \* 4By = 32By/link 256 By/iTOP counter 1kB/event/COPPER

8kByte/event iTOP

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### **SLAC cosmic ray telescope**



#### Energy

### Trigger rates:

- T1\*T2 ~ 5-6 Hz
- T1\*T2\*S1 ~ 1-2 Hz
- T1\*T2\* Quartz\_counter ~ 5-6 k/day
- T1\*T2\*S1\*Quartz\_counter ~ 2-3 k/day

### New DAQ:

- μPCI ceate: μPCI CPU master, fiber optics transievers talk to boards on H-8500 MaPMTs
- Communicate with the rest of CAMAC via the CC-USB controllers
- The new DAQ system created by Gary Varner & Kurtis Nishima
- Status: cosmic ray telescope is running, together with the PiLas laser diode.