l atest results from NFMM-3 and status of SuperNEMO

Alberto Remoto remoto@in2p3.fr

Laboratoire d'Annecy-le-vieux de Physique des Particules







uperne

collaboration

Monday, 23 March 15

What is on the menu?

- Searching 0vββ decay
- Looking for $\beta\beta$ events with a tracker & calorimeter
- Recent results from NEMO-3
- Toward the new generation with SuperNEMO



1936 — E. Majorana proposes a real wave to describe massive & electrically neutral fermion

1957 — B. Pontecorvo set the basis for the neutrino flavour oscillation

Last 20 years: Huge effort to confirm neutrino oscillation and measure parameters





We know a lot about neutrino oscillation but not everything yet: mass hierarchy, $\delta_{CP\!\!,}\,\theta_{23}$ octant

Is the neutrino a Majorana particle?

We know that neutrino is:

- 1) fermion
- 2) electrically neutral



It could be a Majorana particle

3) massive

Search for 0vββ decay is the only practical way to test Dirac/Majorana nature of neutrinos

If Majorana particle exists, there are interesting implications for particle physics:

- Lepton number violation must occur: $v = v^c \rightarrow |\Delta L| = 2$
- GUT, Leptogenesis model, See-Saw mechanism

Double beta decay in a nutshell

$(A, Z) \to (A, Z+2) + 2e^- + 2\bar{\nu}_e$

- 2nd order process allowed in the SM
- Single β decay forbidden (energy & angular momentum)
- 11 isotopes have been experimentally observed undergoing 2vββ decay



$$(T_{1/2}^{2\nu})^{-1} = G_{2\nu}(Q_{\beta\beta}, Z)|M_{2\nu}|^2$$

Neutrino-less double beta decay in a nutshell

 $(A,Z) \rightarrow (A,Z+2) + 2e^{-1}$

- Process forbidden in the SM
- Half-life strongly suppressed





$$(T_{1/2}^{0\nu})^{-1} = G_{0\nu}(Q_{\beta\beta}, Z)|M_{0\nu}|^2\eta^2$$

Few different mechanisms may induce 0vββ

- Light Majorana neutrino exchange
- Right-handed current (V+A), SUSY, Majoron(s), etc.

Different topology in the final state!

Sensitivity to the light Majorana neutrino



Limits from direct $\boldsymbol{\nu}$ mass measurement or cosmology

 $\Sigma m_{\nu} < 0.17 \text{ eV}$ (Planck 2015)

$$(T_{1/2}^{0\nu})^{-1} = G_{0\nu}(Q_{\beta\beta}, Z)|M_{0\nu}|^2 \frac{\langle m_{\beta\beta} \rangle^2}{m_e^2}$$
$$\langle m_{\beta\beta} \rangle = \left| \sum_i U_{ei}^2 m_i \right|$$

Related to:

- 1) neutrino oscillation
- 2) mass hierarchy
- 3) mass scale

3 searches in practice



Few important aspects...

Low energy process ($Q_{\beta\beta} \leq 5$ MeV):

- Natural radioactivity is an issue (²³⁸U, ²³²Th)
- Cosmic muons are an issue



Distinguish 0v from 2v mode \rightarrow irreducible background

Good detector energy resolution

Few available isotopes \rightarrow multiply experimental efforts



Isotope	Q _{ββ} [keV]	Nat. abund. (enrich.) [%]	
⁴⁸ Ca → ⁴⁸ Tl	4270	0.187 (73)	
⁷⁶ Ge → ⁷⁶ Se	2039	7.8 (86)	
⁸² Se → ⁸² Kr	2995	8.7 (97)	
⁹⁶ Zr → ⁹⁶ Mo	3350	2.8 (57)	
¹⁰⁰ Mo → ¹⁰⁰ Ru	3034	9.6 (99)	
$^{110}\text{Pd} \rightarrow ^{110}\text{Cd}$	2018	7.5	
¹¹⁶ Cd → ¹¹⁶ Sn	2802	7.5 (93)	
¹³⁰ Te → ¹³⁰ Xe	2527	34.5 (90)	
¹³⁶ Xe → ¹³⁶ Ba	2480	8.9 (80)	
¹⁵⁰ Nd → ¹⁵⁰ Sm	3367	5.6 (91)	

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$^{130}\text{Te} \rightarrow ^{130}\text{Xe}$	2527	34.5 (90)	
¹³⁶ Xe → ¹³⁶ Ba	2480	8.9 (80)	
$^{150}Nd \rightarrow ^{150}Sm$	3367	5.6 (91)	

Which technology?



 $(T_{1/2}^{0\nu})^{-1} = G_{0\nu}(Q_{\beta\beta}, Z) (M_{0\nu})^2 \eta^2$

- Contain nuclear structure effects
- Many approximation methods
- Different among isotopes
- Measuring NME in $2\nu\beta\beta$ does not help



Main limitation in interpreting results & comparing among different isotopes

... but also the axial coupling constant

- g_A is know to be quenched in β and ββ decay
- An effective constant is extracted from experimental measurement
- Quenching factor ~0.8 0.5
- g_A quenched in $0\nu\beta\beta$ as much as in $2\nu\beta\beta$?



g_A appear at the 4th power in T_{1/2} calculation \rightarrow it may impact $\langle m_{\beta\beta} \rangle$ sensitivity up to x6–34

What is the status?

HdM (35.5 kg x y) & IGEX, ⁷⁶Ge

• T^{0v}_{1/2} > 1.9 x 10²⁵ y @ 90% C.L.

HdM claim: $(m_v) = 0.32 + - 0.03 \text{ eV}$

Cuoricino (19.75 kg x y): TeO₂ bolometer • 130 Te: T ${}^{0v}{}_{1/2} > 2.8 \times 10^{24}$ y @ 90% C.L. NEMO3 (34.7 kg x y with Mo): track + calo.

• ¹⁰⁰Mo: $T^{0v}_{1/2} > 1.1 \times 10^{24} \text{ y} @ 90\% \text{ C.L.}$

EXO200 (> 95 kg x y): Liquid Xe TPC Kamland-ZEN (190 kg x y): Liquid Scintillator GERDA Phase 1 (>20 kg x y): Ge diodes



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2000

Т

1993

Future projects

5 years time scale:

- M ~ 10 50 kg of ββ isotope
- Background level 10⁻³ cts. /(keV kg y)
- Explore quasi-degenerate region

10 years time scale:

- M ~ 100 kg 1t of ββ isotope
- Background level 10⁻⁴ cts. /(keV kg y)
- Approach Inverse Hierarchy region

Extended R&D: Energy resolution, particle ID, radio-purity

Multi-phase approach: demonstrate scalability and background levels

CUORE, Gerda, Majorana, Lucifer, AMORE, NEXT, COBRA, EXO, SNO+, KamLAND-Zen, CANDLES, SuperNEMO, DCBA, ...

$$T_{1/2}^{0\nu} > \frac{N_A \ln 2}{n_\sigma} \times \frac{\epsilon}{A} \times \sqrt{\frac{M \times t}{B \times \Delta E}}$$







ββ decay experiment combining tracker and calorimetric measurement

Located at the Modane underground laboratory (~4800 m.w.e.)

10 kg of different ββ isotopes taking data from February 2003 to January 2011



sources

60 mg/cm² foils 10 kg of ββ isotopes

tracker

6180 Geiger cells vertex resolution : σ_{xv} ~ 3 mm σ_z ~ 10 mm

calorimeter

1940 optical modules : polystyren scintillators + 3" and 5" PMTs FWHM_E ~ 15% / √E_{MeV} σ_t ~ 250 ps

NEMO-3



Laboratoire Souterain de Modane









- Mainly ¹⁰⁰Mo (7 kg) & ⁸²Se (1 kg) + smaller quantities of others isotopes
- Blank foils to cross-check background measurements (Cu & ^{Nat}Te)



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Full reconstruction of 2e⁻ kinematics: unique!

 Individual e⁻ energy, arrival time, track curvature, emission vertex and angle

Excellent background rejection

• Identification e^- , e^+ , γ , α , Internal/External Low energy resolution: [14 - 17] % / Sqrt(E)



SOUFCES 60 mg/cm² foils

10 kg of ββ isotopes

tracker

6180 Geiger cells vertex resolution : σ_{xv} ~ 3 mm σ_z ~ 10 mm

calorimeter

1940 optical modules : polystyren scintillators + 3" and 5" PMTs FWHM_E ~ 15% / √E_{MeV} σ_t ~ 250 ps

Sensitivity equivalent to best calorimetric experiment













Selection efficiency for $E_{2e} > 2.0$ MeV is 11.3 % (MC)

NEMO-3: energy calibration

Radioactive sources:

- ²⁰⁷Bi: 482 keV and 976 keV conversion electron
- ${}^{90}\text{Sr} {}^{90}\text{Y}$: β -decay end point $Q_{\beta} = 2280 \text{ keV}$
- ²⁰⁷Bi: 1682 keV conversion electron to test energy scale: 99% PMTs Data/MC < 0.2%

Laser inter-calibration system:

- Gain and time survey twice a day PMTs linearity < 1% for E < 4 MeV
- 82% of PMTs stable < 5% over the whole data taking



NEMO-3: backgrounds

Radio-impurities in material, γ from (n,γ) and μ bremsstrahlung

- γ from ²⁰⁸Tl at 2.6 MeV
- (n,γ) up to 10 MeV

²⁰⁸TI (from ²³²Th) and ²¹⁴Bi (from ²³⁸U) contamination in foil source and ²¹⁴Bi from Rn decay in tracker volume

- ²⁰⁸TI Q_{β} at 5 MeV
- ²¹⁴Bi Q_{β} at 3.27 MeV





NEMO-3: external backgrounds

Take advantage of PID capabilities of NEMO-3: e^- , e^+ , γ , α and TOF measurement



NEMO-3: internal backgrounds

Take advantage of PID capabilities of NEMO-3: e⁻, e⁺, γ, α and TOF measurement

1e⁻ (single β emitters), e⁻N γ (²⁰⁸TI) , e⁻ α (²¹⁴Bi) channels channels



²⁰⁸TI activity checked with ²³²U sources \rightarrow 10% systematics w.r.t. HPGe measurement ²¹⁴Bi activity compared in two different channel \rightarrow e-Ny, e-a: 10% systematics



Background checked in 2e Ny (^{208}TI) and 2e α (^{214}Bi) channels

- ²⁰⁸TI: 7 events, 8.8 expected
- 214 Bi Phase 1: 3 events, 6.5 ± 0.4 expected
 - 214 Bi Phase 2: 3 events, 2.9 ± 0.2 expected

NEMO-3: ²²²Rn background

²²²Rn in the gas of the tracking chamber monitored through the 1e1α channel

Strongly suppressed upon flushing Rn-free air into a dedicated tent surrounding the detector





Phase 1: 37.7±0.1 mBq/m³

Phase 2: 6.46±0.02 mBq/m³

NEMO-3: hot spots identification

Vertex reconstruction capabilities: $\sigma_z = 0.6$ cm and $\sigma_r = 1.0$ cm



Able to reduce backgrounds removing activity hot spots from the foils surface

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NEMO-3: ¹⁰⁰Mo $2\nu\beta\beta$ results

- About 700 000 2vββ events
- Detection efficiency = $4.3 \pm 0.7 \%$
- Signal over Background ratio = 76

```
T^{2v}_{1/2} = [7.16 \pm 0.01 \text{ (stat)} \pm 0.54 \text{ (syst)}] \times 10^{18} \text{ y}
```

Consistent with previously published [PRL 95 (2005) 182302]





[Phys. Rev. D. 89.111101 (2014)]

Detailed paper to be published in the following weeks

NEMO-3: ¹⁰⁰Mo $0\nu\beta\beta$ result

- No event excess after 34.3 kg×y exposure
- $T^{0v}_{1/2} > 1.1 \times 10^{24} \text{ y} (90 \% \text{ C.L.}) \rightarrow \langle m_v \rangle < 0.3 0.9 \text{ eV}$



Expected background in [2.8 - 3.2] MeV

2νββ	8.45 ± 0.05
²¹⁴ Bi from radon	5.2 ± 0.5
²⁰⁸ TI internal	3.3 ± 0.3
²¹⁴ Bi internal	1.0 ± 0.1
External	< 0.2
Total	18.0 ± 0.3
Data	15

Total background: 1.3×10⁻³cts / (keV×kg×y)

NEMO-3: ¹⁰⁰Mo $0\nu\beta\beta$ result

• $T^{0v}_{1/2}$ limit set with a modified frequentist analysis

Account for statistical and systematic uncertainties

• Using full information in $E_{Tot} = [2.0; 3.2]$ MeV

Detection efficiency: 11.3 ± 0.8 %

[Phys. Rev. D. 89.111101 (2014)]

Detailed paper to be published in the following weeks

Systematics

2vββ events in window	0.7%
0vββ detection efficinecy	7.0%
²¹⁴ Bi contamination	10.0%
²⁰⁸ TI contamination	10.0%

HP/Np 250	$= \begin{array}{c} 2\beta\chi\chi\\ \mathfrak{y}=^{-7}2\beta^{-2}\mu - 2\beta\chi\chi\end{array}$	2BV 2B	Ον
	$n=5 2\beta\chi$	n=1	
200			
150			
100			
100	$ / / \rangle$		
50			
50	$[//] \land \land \land$	1911 1911 1911 1911 1911 1911 1911 191	
0			
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[N.I.M. A 434 (1999) 435]

and their correlation

Limits at 90%	C.L. in	units	of 10 ²⁴ y
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Process	Stat. Only	Stat. + Syst.	Expected
Mass mechanism	1.1	1.1	1.0 [0.7; 1.4]
RH Current $\langle \lambda \rangle$ (q _{r.h.} – I _{r.h.})	0.7	0.6	0.5 [0.4; 0.8]
RH Current $\langle \eta \rangle$ (q _{I.h.} – I _{r.h.})	1.0	1.0	0.9 [0.6; 1.3]
Majoron (n=1)	0.050	0.044	0.039 [0.027; 0.059]

NEMO-3: other results

Other isotopes: only partial exposure has been published



Analysis of whole statistics ongoing (⁸²Se, ⁴⁸Ca, ⁹⁶Zr, ¹¹⁶Cd, ¹⁵⁰Nd)...stay tuned!

¹⁰⁰Mo 0vββ decay to the ¹⁰⁰Ru excited states [Nuclear Physics A781 (2007) 209-226]

NEMO-3: high energy background

No events in 100 Mo foils after 34.3 kg×y > 3.2 MeV

No events in Cu & Te foils after 13.5 kg×y > 3.1 MeV



Promising background free technique for high Q_{ββ} isotopes ⁴⁸Ca (4.272 MeV), ¹⁵⁰Nd (3.368 MeV) or ⁹⁶Zr (3.350 MeV)

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[Eur. Phys. J. C70: 927-943,2010]

SuperNEMO: toward the new generation

Extrapolate a well known technique:

- 100 kg of ββ emitter in 20 detection module
- Approach Inverted Hierarchy region

	NEMO-3	SuperNEMO
Efficiency	18%	~30%
Isotope	7 kg ¹⁰⁰ Mo	~100 kg ⁸² Se (¹⁵⁰ Nd, ⁴⁸ Ca)
Exposure	35 kg y	~500 kg y
Energy res.	8% @ 3 MeV	4% @ 3 MeV
²⁰⁸ TI (source)	~100 µBq/kg	< 2 µBq/kg
²¹⁴ Bi (source)	~ 300 µBq/kg	< 10 µBq/kg
Rn (in tracker)	5 mBq/m ³	0.15 mBq/m ³
T _{1/2}	10 ²⁴ y	10 ²⁶ y
$\langle m_v \rangle$	0.3 - 0.9 eV	0.04 - 0.1 eV



A challenge in many aspects:

- R&D program in the past years almost completed!
- Next step: Demonstrator module

[Eur. Phys. J. C70: 927-943,2010]

SuperNEMO: the demonstrator module

One SuperNEMO module \rightarrow 7 kg ⁸²Se running ~2.5 y

- To be installed @ LSM (replacing NEMO-3)
- Match SuperNEMO requirements

Reach NEMO-3 (¹⁰⁰Mo) sensitivity in 4.5 months

• Sensitivity: $\langle m_v \rangle \sim 0.20 - 0.40 \text{ eV}$



Schedule:

- Calorimeter & tracker under production
- Installation starting in 2015
- Commissioning & First data by 2016

SuperNEMO: the calorimeter

- 5" and 8" high quantum efficiency PMT directly coupled to a scintillator block with optimised geometry
- Energy resolution: 7.2 % FWHM @ 1 MeV
- Electronics, optical modules, shield, mechanical structure under production







SuperNEMO: the tracker

- 2034 Geiger cells in a Rn-tight chamber surrounded by optical modules for veto
- Drift cells under production whit automatic wiring robot
- Tracker assembled in 1/4 @ MSSL (UK) then moved to LSM for integration
- Commissioning of C0 ongoing at sea level, C1 under construction



SuperNEMO: the source foil

- About 37 foils installed on the source frame in the detector center
- ⁸²Se powder mixed with PVA glue + mylar or nylon mechanical support
- Limits on foil contamination in ²⁰⁸TI (2 μBq/kg) and ²¹⁴Bi (10 μBq/kg) are challenging
- Purification technique under investigation: chemical chromatography, distillation, etc.
- LAPP is in charge for the production of 1/2 of the source for the Demonstrator



SuperNEMO: radio-purity measurements

Detector radio-purity budget:

Materials validation with HPGe detectors (sensitivity ~ mBq)

Source foils:

- HPGe not sensitive enough for SuperNEMO requirement: dedicated setup @ LSC (Canfranc) BiPo
- Detecting delayed $\beta \alpha$ coincidence from Bi Po chain
- First two ⁸²Se foils currently under measurement



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SuperNEMO: Radon measurements

- The Rn gas in the tracker volume was the dominant background in NEMO-3
- Reduce Rn contamination to 0.15 mBq/m³
- Control the Radon emanation of the materials
- Radon purification/absorption with dedicated setup
- Preliminary radon emanation of $C0 = 0.236 \pm 0.035 \text{ mBq/m}^3$ limit is close!



Esher - Ascending and descending (1960)

Summany

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Summary & conclusions

- Unique: allowing direct reconstruction of the 2e⁻
- Full signature of 0vββ events and powerful background rejection
- Background-free technique for high energy Q_{ββ} isotopes

Latest NEMO-3 results

technique

Tracking + Calo.

- Total ¹⁰⁰Mo exposure of 34.3 kg×y shows no event excess
- T^{0v}_{1/2} > 1.1×10²⁴ y $\rightarrow \langle m_v \rangle < 0.3 0.9 \text{ eV} @ (90 \% \text{ C.L.})$
- Other isotopes: re-analysis of full statistics ongoing
- Under construction: commissioning by 2016
- SuperNEMO demonstrator
- Foresee to run for 2.5 years with 7 kg of ⁸²Se
- T^{0v}_{1/2} > $6.5 \times 10^{24} \text{ y} \rightarrow \langle m_v \rangle < 0.20 0.40 \text{ eV} @ (90 \% \text{ C.L.})$
- Future: Full SuperNEMO
- 20 demonstrator-like modules: 100 kg of ⁸²Se for 5 years
- T^{0v}_{1/2} > 1 × 10²⁶ y $\rightarrow \langle m_v \rangle < 0.04 0.10 \text{ eV} @ (90 \% \text{ C.L.})$
- ⁴⁸Ca,¹⁵⁰Nd or ⁹⁶Zr are also possible candidates

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Backup



Which isotope?

	Isotope	Q _{ββ} [keV]	Nat. abund. (enrich.) [%]	G _{0v} [10 ⁻¹⁴ y ⁻¹] ^(*)	T ^{2v} 1/2 [10 ¹⁹ y]	Experiment
	⁴⁸ Ca	4270	0.187 (73)	6	4.2 ^{+2.1} -1.0	NEMO3
	⁷⁶ Ge	2039	7.8 (86)	1	150±10	GERDA
	⁸² Se	2995	8.7 (97)	3	9.0±0.7	NEMO3
	⁹⁶ Zr	3350	2.8 (57)	6	2.0±0.3	NEMO3
11 7/1	¹⁰⁰ Mo	3034	9.6 (99)	4	0.71±0.04	NEMO3
ר מו ק	¹¹⁶ Cd	2802	7.5 (93)	5	3.0±0.2	NEMO3
	¹³⁰ Te	2527	34.5 (90)	4	70±10	NEMO3
	¹³⁶ Xe	2480	8.9 (80)	4	238±14	KamlandZEN
130rCP	¹⁵⁰ Nd	3367	5.6 (91)	19	0.78±0.7	NEMO3

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2

What is the status?

Light Majorana neutrino exchange		Rig e 🔪 cui	Right handedSUSY: neutralino orcurrentgluino exchange			Majoron emission	
					^		
Isotope	Exposure (kg•y)	Half life (10 ²⁵ y) published	、 〈m _v 〉(eV) published	〈λ〉(10 ⁻⁶) published	<η> (10 ⁻⁸) published	λ' ₁₁₁ /f (10 ⁻²) published	〈g _{ee} 〉(10 ⁻⁵) published
¹⁰⁰ Mo ^[1] (NEMO-3)	34.7	0.1	0.33 - 0.87	0.9 - 1.3	0.5 - 0.8	4.4 - 6.0	2 - 5
¹³⁰ Te ^{[2][3]} (CUORICINO)	19.75	0.3	0.31 - 0.71	1.6 - 2.4	0.9 - 5.3		17 - 33
¹³⁶ Xe ^{[4][5]} (KamLAND-Zen)	89.5	1.9	0.14 - 0.34				
¹³⁶ Xe ^[9] (KamLAND-Zen)	109.4 + 89.5	2.6	0.14 - 0.28				
¹³⁶ Xe ^[6] (EXO-200)	99.8	1.1	0.19 - 0.45				
⁷⁶ Ge ^{[7][8]} (GERDA)	21.6	2.1	0.2 - 0.4				3.4 - 8.7
⁷⁶ Ge ^[9] (HdM)	35.5	1.9	0.4	1.1	0.6		8.1

NEMO-3: $2\nu\beta\beta$ of ¹⁰⁰Mo SSD/HSD

If the intermediate nucleus is a $J^{\pi}=1+$ state, the NME could be dominated by GT transitions through this state.

If the SSD hypothesis is confirmed

- $2\nu\beta\beta$ half-life could be determined from single- β and electron capture (EC) measurements.
- simplification in the theoretical description of the intermediate nucleus





The BB source foil

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Foil source design

A. Remoto, D. Duchesneau, J.M. Dubois, A. Jeremie, T. Le Noblet



Se + PVA bulk



The $\beta\beta$ emitter is shaped in thin foil (150–200 um) mixing ⁸²Se powder with PVA glue — very fragile!

An embedded mechanical support is necessary to provide mechanical strength over the foil length

Different designs of the mechanical support are under consideration:

- NEMO-3 like design (ITEP): ⁸²Se+PVA within mylar backing film
- New design (LAPP): ⁸²Se+PVA with light nylon fabric support
- New design (LAPP): standalone ⁸²Se+PVA foil with clean mylar film welded to offer a protective layer

Foil source production

A. Remoto, D. Duchesneau, J.M. Dubois, A. Jeremie, T. Le Noblet

- Foil production protocol is defined. All the tools are ready! Improving the technique with practice...
- All materials to be used in foil production have been defined.
- Radio-purity measurements have been performed (collaboration with LSM, LAL and LSC - Canfranc)



Where we were last year

Where we are now



T. Le Noblet, A. Remoto

- Study ββ0v sensitivity w.r.t. foil design
- Check the foil design doesn't alter physics performance and results
- Generating signal, background and detector response





reco'ed e- tracks



- ε : Signal selection efficiency
- S(b) : Average upper limit on the number of signal events (Feldman & Cousins)

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dN/dE [A.U.]

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- Study ββ0v sensitivity w.r.t. foil design
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$$T_{1/2}^{0
u} > rac{N_A \ln 2}{W} imes rac{\epsilon imes M imes T}{\mathcal{S}(b)}$$

- ϵ : Signal selection efficiency
- S(b) : Average upper limit on the number of signal events (Feldman & Cousins)
- Optimise R.O.I w.r.t. s/b ratio

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dN/dE [A.U.]

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• Optimise R.O.I w.r.t. s/b ratio

T. Le Noblet, A. Remoto

- Study ββ0v sensitivity w.r.t. foil design
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• Optimise R.O.I w.r.t. s/b ratio



- Compare different design of the foil
- Recent radio-purity measurements are taken into account
- Currently limited by the PVA glue (x5 times the limit in ²¹⁴Bi)
- PVA purification procedure under R&D