



Search for SUSY in Dijet Events with Novel Data-Driven Background Estimation

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for the CMS Collaboration

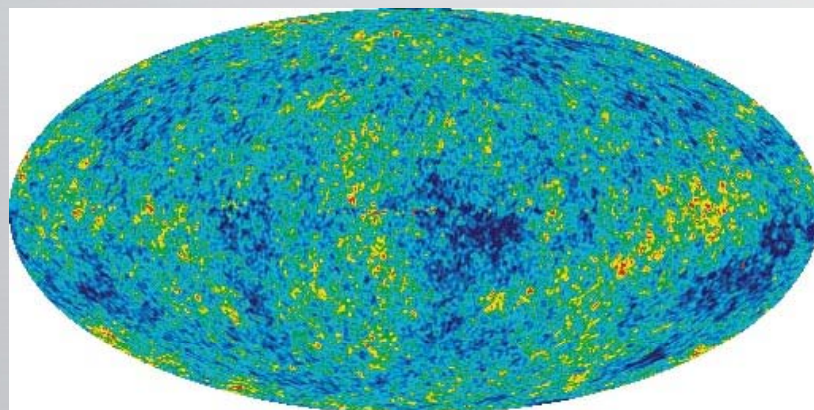
- Motivation for di-jet search
 - Dark Matter at Colliders
- Missing Energy Signatures in Di-Jet Events

 - robust analysis techniques
 - data-driven background estimates
- Conclusions

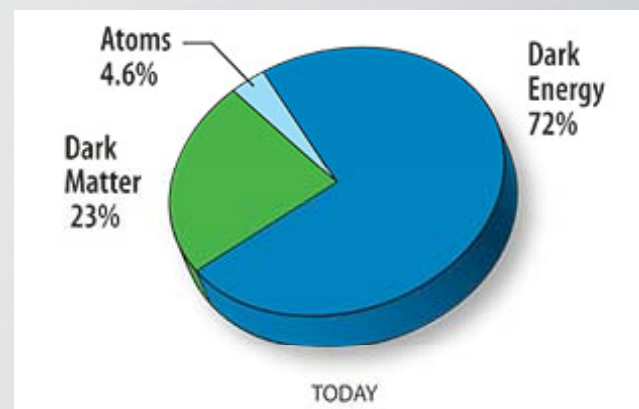
La Thuile 2009 (March 1st -7th, 2009)

A Look at the Energy and Matter Content of the Universe

- Cosmic microwave background gives precise information about dark matter content of the universe
- WMAP 5 year result:



Credit: NASA/WMAP Science Team



- relic dark matter density of the universe
 $\Omega_{\text{DM}} h^2 = 0.110 \pm 0.006$
- Only 5% is made from baryonic matter, 23% from unknown “dark matter”
- Attractive explanation for Dark Matter:
 - new weakly interacting particle



Can we produce Dark Matter particles at Colliders?



- Dark Matter candidate is a weakly interacting massive particle (WIMP)
- Many New Physics Models provide viable dark matter candidates, e.g.
 - R-parity conserving Supersymmetry
 - minimal super gravity mSugra → neutralino is WIMP
 - Gauge mediated SUSY → gravitino is WIMP (too light)
 - Universal Extra Dimensions
 - Warped Extra Dimensions
 - Little Higgs Models
 - Technicolor Models
- **Production of WIMP's in cascade decays of heavy new particles**
 - WIMP's escape the detector and remain undetected
 - **Leads to a missing energy signature**

- from decay of charginos and neutralinos

- # La Thuile 2009



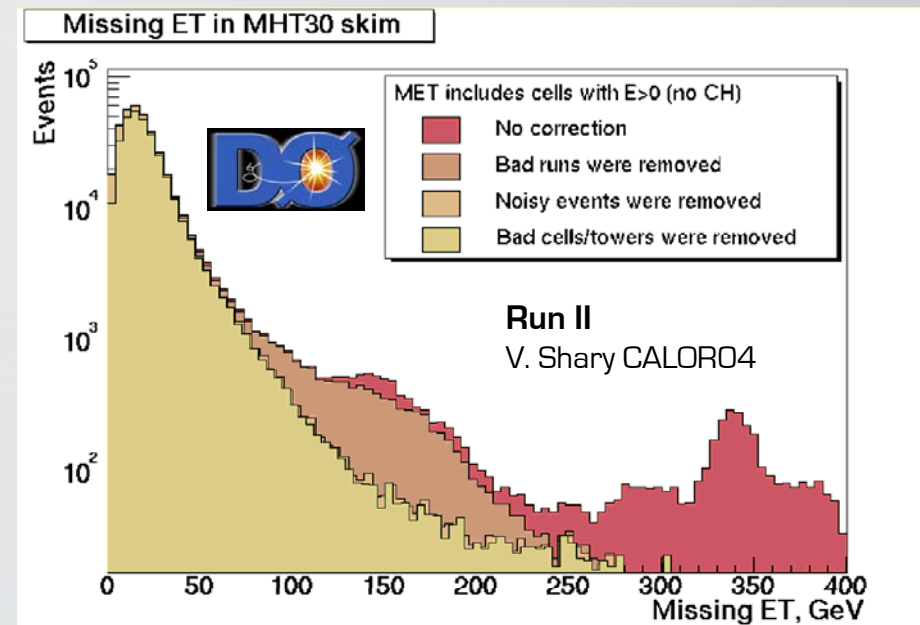
- from decay of the N's

Model examples are Extra dimensions, Little Higgs, Technicolour, etc

Missing Energy Measurement

- “Traditional” approach:
 - Calculate missing energy as negative vectorial sum of all calorimeter deposits
 - Susceptible to mis-measurements from, e.g.
 - Calorimetric noise (hot cells)
 - Cosmic rays
 - Beam-gas interactions
 - Beam-halo events
 - Difficult to understand in the early days of data taking
- Need for robust measurement techniques

Missing Energy from Tevatron during several cleanup stages:



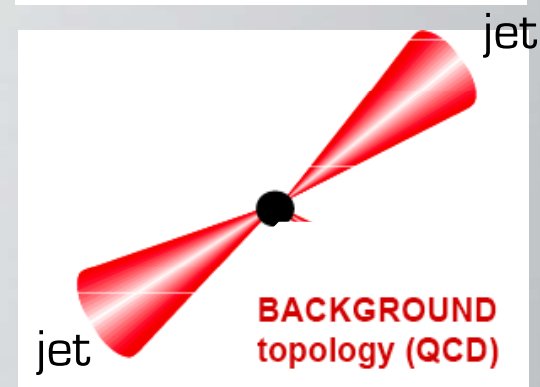
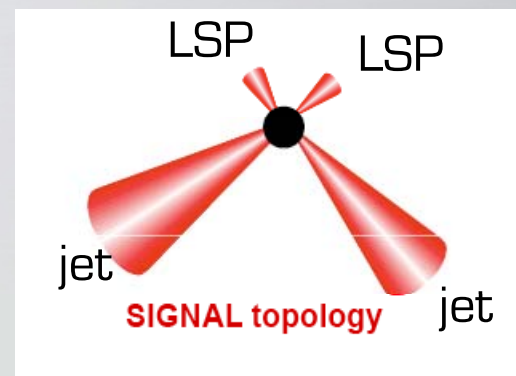
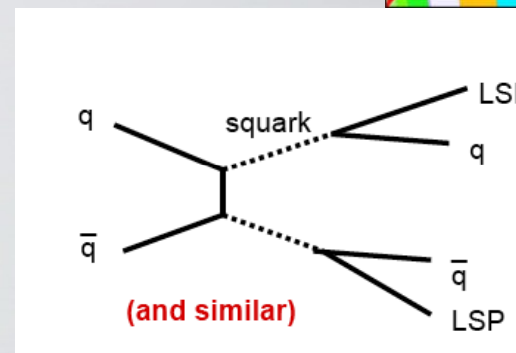
IDEA:

infer missing energy from well measured objects by applying transverse energy/momentum conservation

Di-jet Analysis

- New CMS study: [PAS-SUS-08/005](#)
 - CMS PTDR II focused on inclusive SUSY searches with ≥ 3 jets
- Motivated in addition by recent paper by
 - L. Randall, D.Tucker-Smith ([Phys.Rev.Lett.101:221803,2008](#))
- Idea:
 - Squarks pair produced and directly decaying to quarks and neutralinos
- Event topology
 - Only two jets + missing energy
- Background:
 - QCD dijet events
 - No real missing momentum
 - $Z \rightarrow \nu\nu$ events
 - Irreducible background due to real missing E_T
 - $W \rightarrow l\nu$
 - Leads to missing E_T when lepton not reconstructed or out of acceptance

Transverse momentum conservation
 Jets back-to-back in ϕ
 E_T of jets equal in magnitude



- Main variables of interest

- HT: Scalar sum of Jet p_T 's:

- $HT = p_T^{\text{Jet1}} + p_T^{\text{Jet2}}$

- MHT: Jet based missing E_T

- $MHT = -(\vec{p}_T^{\text{Jet1}} + \vec{p}_T^{\text{Jet2}})$

- MT: Transverse Mass

- $$M_T = \sqrt{(\sum_i E_{T_i})^2 - (\sum_i p_{x_i})^2 - (\sum_i p_{y_i})^2}$$

$$= \sqrt{HT^2 - MHT^2}$$

- but also p_T of a possible 3rd jet

- $\Delta\phi$ between the jets

- α (α_T) from 2 leading jets

- Trigger

- di-jet trigger (2 jets with $p_T > 150$ GeV)

- Preselection:

- Jet Selection

- 2 jets with $p_T > 50$ GeV, $F_{em} < 0.9$
 - 3rd jet veto: $p_T < 50$ GeV
 - $\Delta\phi(MHT, \text{jet}_{1,2,3}) > 0.3$ rad
 - $|\eta_{j1}| < 2.5$

- Lepton veto's:

- no e, μ with $pt > 10$ GeV

- Full Selection

- $HT > 500$ GeV

- α (α_T) > 0.55

- $[\Delta\phi < 2\pi/3]$

Accounting for
finite resolution
(not optimised)

Discriminating Variables

- Exploit kinematics of the event
 - Define new variable α (Randall – Tucker-Smith):

$$\alpha = \frac{E_{Tj2}}{M_{j1j2}} = \frac{E_{Tj2}}{\sqrt{2E_1E_2(1-\cos\theta)}}$$

- Can be at most 0.5 for QCD, $\alpha < 0.5$
 - $\alpha > 0.5$ implies missing momentum
- And transverse α_T :

$$\alpha_T = \frac{E_{Tj2}}{M_{Tj1j2}} = \frac{\sqrt{E_{Tj2}/E_{Tj1}}}{\sqrt{2(1-\cos\Delta\varphi)}}$$

- Exploits that for QCD jets need to be back-to-back and of equal magnitude
 - For QCD dijets $\alpha = 0.5$

Analysis does not rely
on calorimetric MET,
MHT inferred from 2 jets

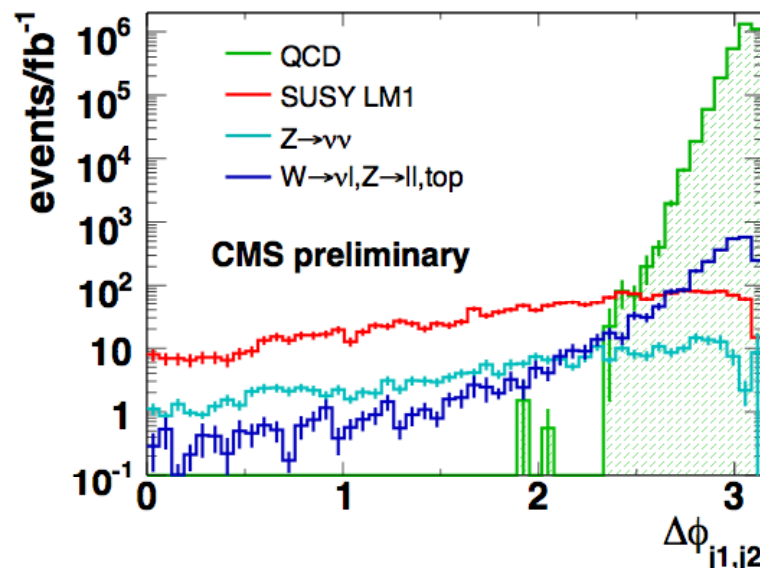
⇒ well suited for
early data

Discriminating Variables (II)

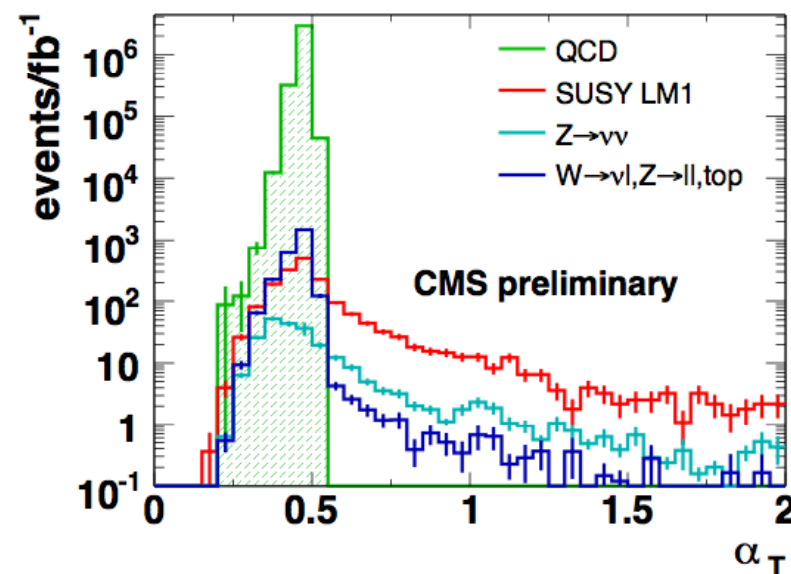
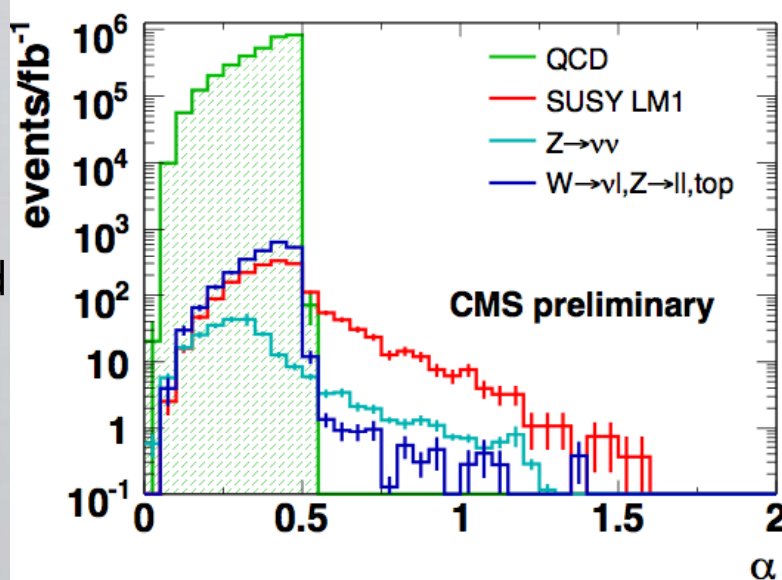
- Cuts applied:
 - Preselection & $HT > 500$ GeV

1 fb^{-1}

QCD peaking
at $\Delta\phi = \pi$,
but exhibiting tail



Sharp drop of
QCD background
for $\alpha [\alpha_T] > 0.5$



- $Z \rightarrow \nu\nu$ is main background
- W and other Z decays small

Signal & Background yields

- Expected event yields for 1fb^{-1}

Selection cut	QCD	$t\bar{t}, W, Z$	$Z \rightarrow \nu\bar{\nu}$	LM1
Trigger	1.1×10^8	147892	1807	25772
Preselection	3.4×10^7	9820	878	2408
HT > 500 GeV	3.2×10^6	2404	243	1784
$\alpha > 0.55$	0	7.2	19.7	227.6
$\alpha_T > 0.55$	0	19.9	58.2	439.6
$\Delta\phi_{j1,j2} < 2\pi/3$	0	18.7	57.2	432.4

=> Signal/Background = 5.6

Reminder: desired topology is 2 squarks decaying to 2 squarks and 2 neutralinos (LSPs)

Sample	Events	$\tilde{q} \tilde{q}$ (invisible)	$\tilde{q} \tilde{q}$ (other)	$\tilde{q} \tilde{g}$	$\tilde{g} \tilde{g}$	other
LM1	432	39%	22%	34%	3%	1%
LM2	132	46%	33%	18%	0%	2%
LM3	138	69%	17%	12%	0%	2%
LM4	195	49%	10%	36%	3%	1%

- Variation of jet energy scale and resolution
 - 10% gaussian smearing of jet p_T 's and of 0.1 rad of ϕ measurement
 - Scaling of jet energy by $\pm 5\%$
 - Scaling of jet energy by $\pm 3\%$ for endcap/forward ($|\eta| > 1.4$)
- Stable S/B for all variations



Data-Driven Background Studies



- LHC data in explores a new energy regime
 - Monte Carlo simulations should not be taken at face value
 - develop data-driven techniques
 - identify data control samples
- Two main sources of background:
- QCD
 - Seems to be under control but huge cross-section
 - MC uncertainties due to higher order QCD effects
- $Z \rightarrow \nu\nu + \text{Jets}$
 - represents an irreducible background
 - two jets + real missing E_T
 - Ideally study $Z \rightarrow \mu\mu$ events but not enough statistics in the early days
 - Other control samples:
 - W + Jets
 - Photon + Jets as shown in CMS-AN 36/2008

Central Production of Heavy Objects

- Idea: define **signal enriched and depleted regions** by splitting data sample in events with first jet in barrel and forward region

➤ SUSY jets are more central

➤ Use ratio of events

$$R_a = \alpha_T > 0.55 / \alpha_T < 0.55 \text{ in}$$

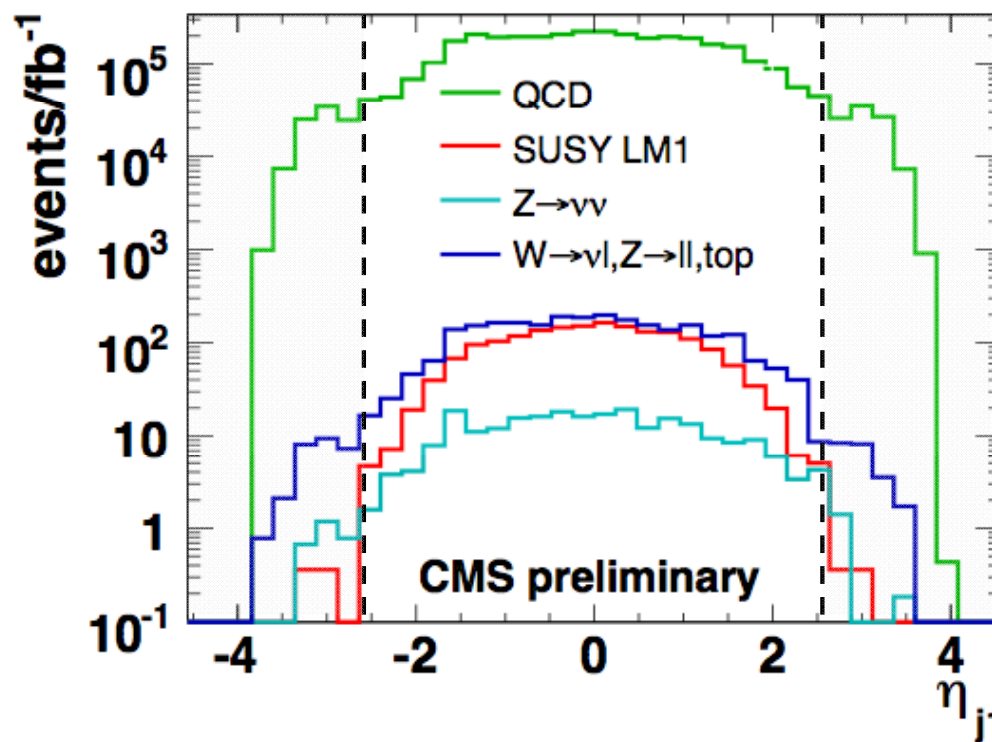
(signal depleted) forward η region to predict background in (signal enriched) barrel region.

$ \eta $		D	C
	2.5	B	A
		0.55 α	

See also: Background Modeling in New Physics Searches
Using Forward Events at LHC.

V. Pavlunin, D. Stuart, Phys.Rev.D78:035012,2008.

Pre-selection (no η cut) + $HT > 500$ GeV



$R = C/D$: assumed to be constant over η
and nearly signal free

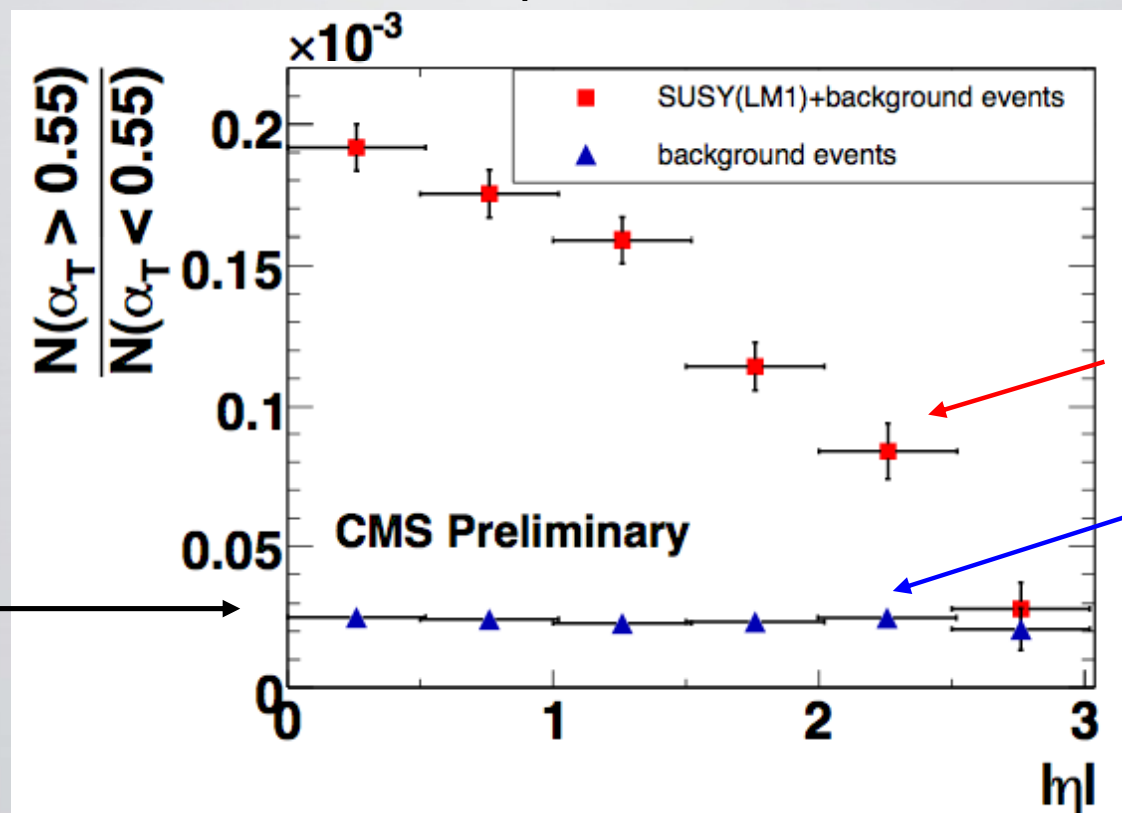
also: constant for all background
contributions individually

Then, background in A can be obtained as:

$$A = B * R$$

η Dependence of Matrix Method

Pre-selection (no $|\eta|$ cut) + HT > 500 GeV



Verified that
flat for all Bkgd
contributions
(see AN-08-071)

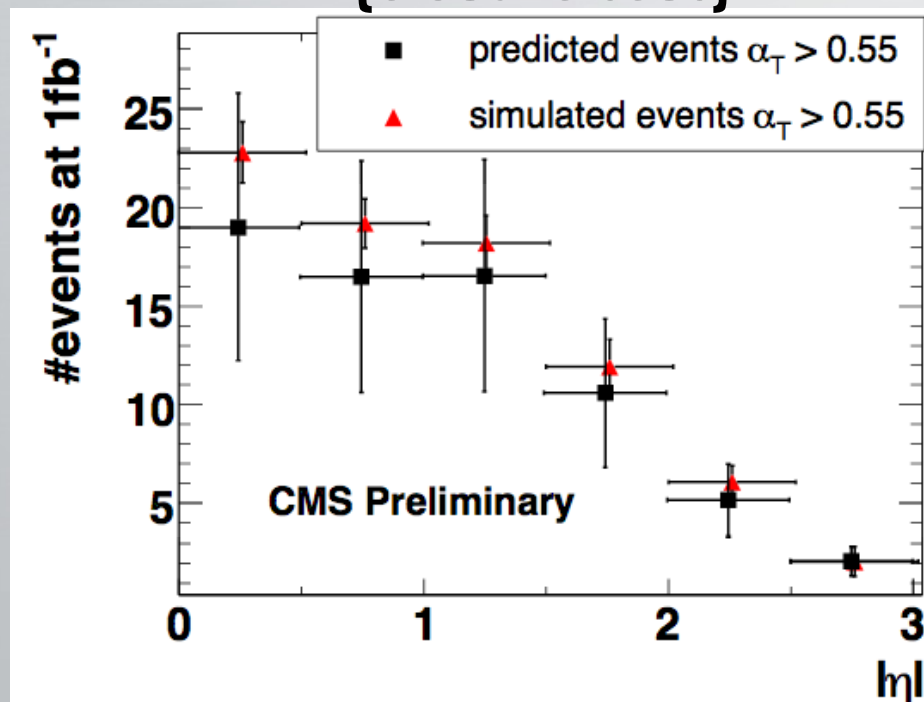
- R_α flat for background as function of $|\eta_{j1}|$
- α_T and $|\eta_{j1}|$ can be used for ABCD-matrix method

★ Measure $R\alpha_T$ in $2.5 < |\eta| < 3.0$ region.

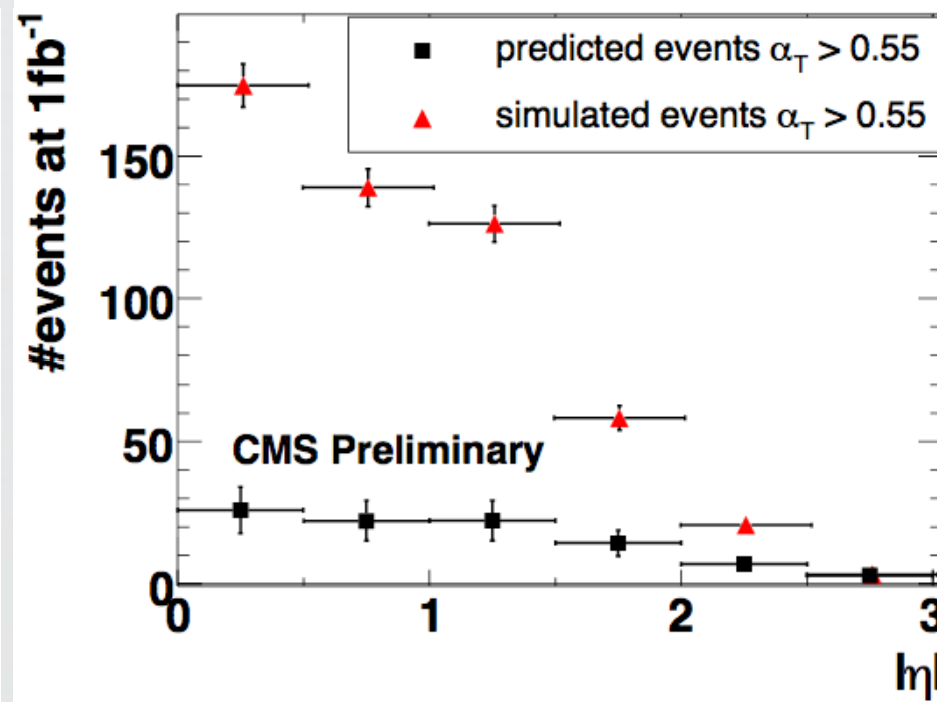
η Dependence of Matrix Method

without SUSY
(closure test)

with LM1 SUSY
“contamination”



simulated: 77 ± 3 (8 @ 1fb^{-1})
predicted: 68 ± 24 (42 @ 1fb^{-1})



simulated: 517 ± 13 (22 @ 1fb^{-1})
predicted: 91 ± 30 (51 @ 1fb^{-1})

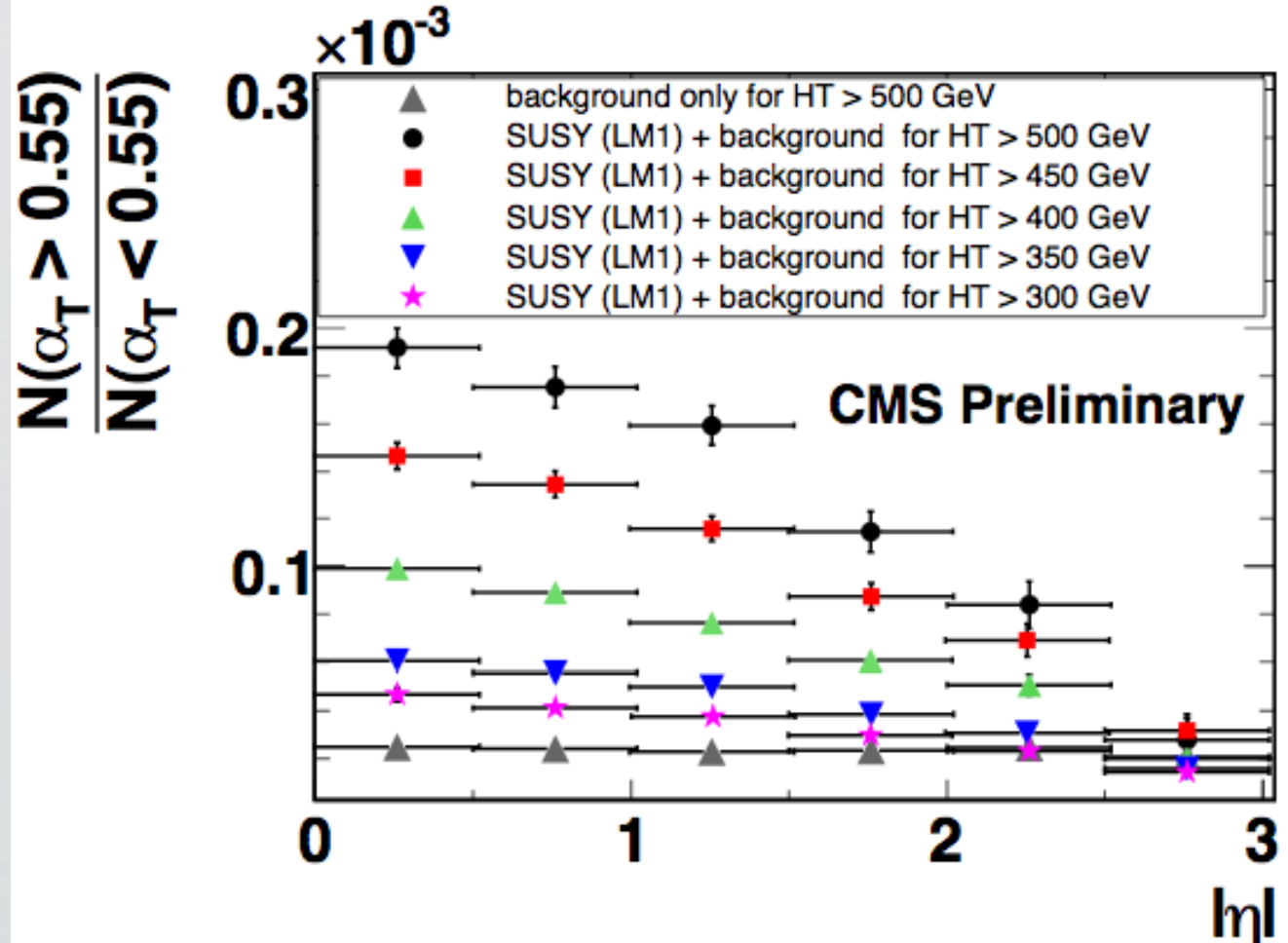
- Predicted BKGD agrees well with simulated BKGD
- SUSY LM1 leads to **significant** excess in signal region
- Method also verified with systematic variations

Variation of HT cut

Idea:

Increase background to check that R_{α_T} is flat in $|\eta|_1$ when signal sufficiently diluted

- Loosen HT cut to decrease signal to background ratio.



- As HT loosened $|\eta|_1$ dependence gets flatter

=> Clear indication that at HT > 500 GeV signal is present

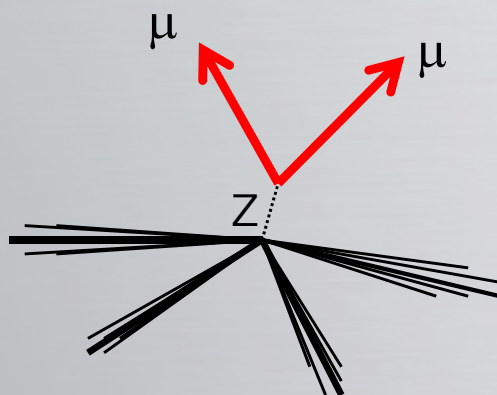
Data Driven Background Estimations

An illustrative example: $Z \rightarrow \nu\nu + \text{jets}$

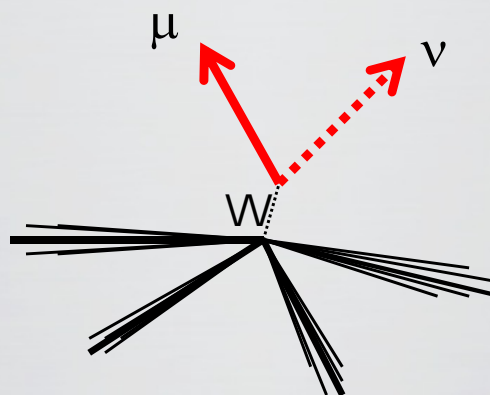
Irreducible background for $\text{Jets} + E_t^{\text{mis}}$ search

Data driven strategy:

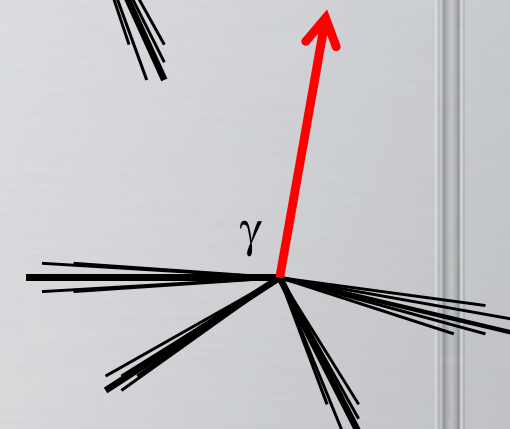
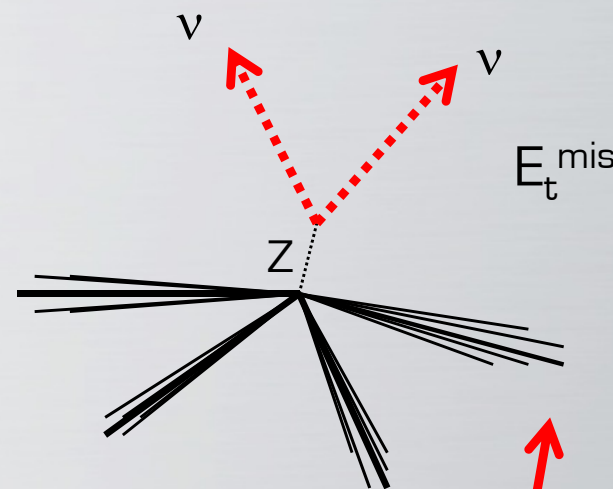
- define control samples and understand their strength and weaknesses:



$Z \rightarrow \ell\ell + \text{jets}$



$W \rightarrow \ell\nu + \text{jets}$



$\gamma + \text{jets}$

Strength:

- very clean, easy to select

Weakness:

- low statistic: factor 6 suppressed wrt. to $Z \rightarrow \nu\nu$

Strength:

- larger statistic

Weakness:

- not so clean, SM and signal contamination

Strength:

- large stat, clean for high E_γ

Weakness:

- not clean for $E_\gamma < 100$ GeV,
- possible theo. issues for normalization (u. investigation)

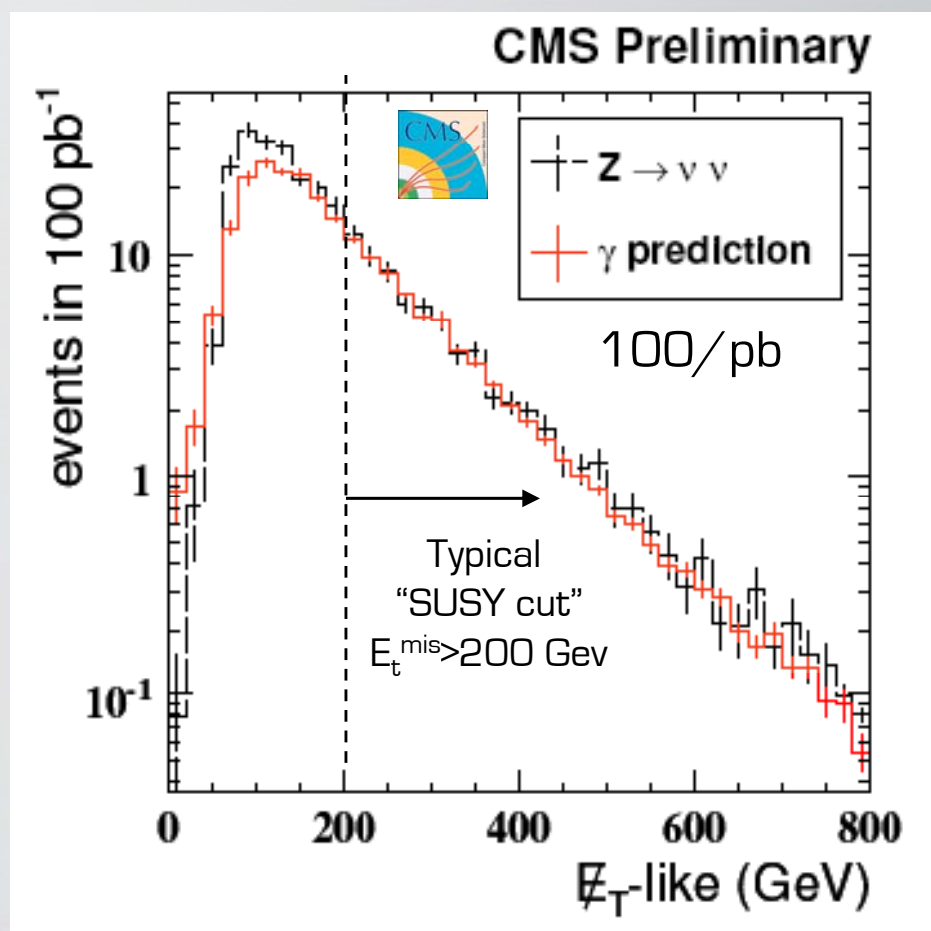
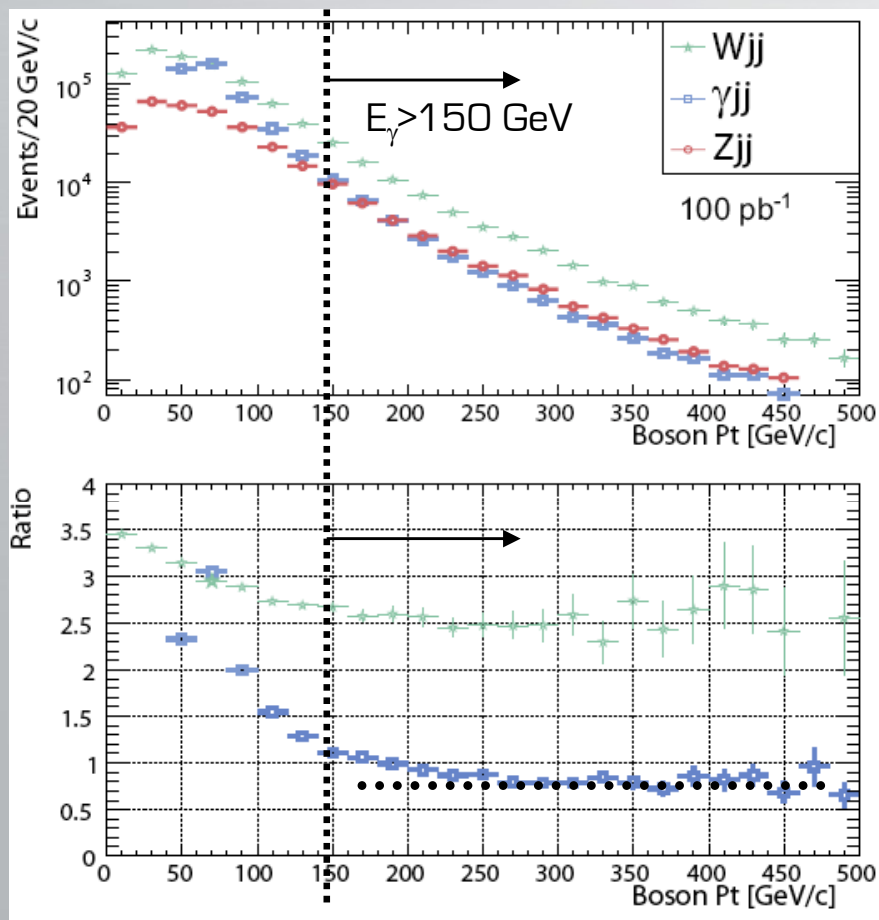
γ +jets: Estimate Z to invisible

γ +jets selection & properties:

- $E_\gamma > 150$ GeV
- clean sample: $S/B > 20$
- ratio $\sigma[Z+\text{jet}]/\sigma[\gamma+\text{jet}]$ constant

γ +jets: Strategy:

- remove γ from the event:
→ γ becomes E_T^{mis}
- take $\sigma[Z+\text{jet}]/\sigma[\gamma+\text{jet}]$ for $E_\gamma > 200$ GeV from MC or measure in data





Conclusions

- Good prospects to find Dark Matter Candidate at LHC
- Di-jet analysis promising, exploiting particular event topology
 - α (α_T) and $\Delta\phi$ very powerful
 - Shown results do not rely on calorimetric MET
- Data-driven background determinations have been developed
 - Subtraction of all backgrounds using matrix method **CMS PAS SUS-08-005**
 - checks on real data in place
 - $Z \rightarrow \nu\nu$ can be obtained from γ + jets **CMS PAS SUS-08-002**
- Extension to calorimetric MET independent multi-jet analyses under study
- Benchmark points (e.g. LM1) could be observed in dataset of $\sim 100\text{pb}^{-1}$
 - Assuming detector performance is understood
- Eagerly awaiting first collision data in fall of this year
 - Exciting times are ahead!

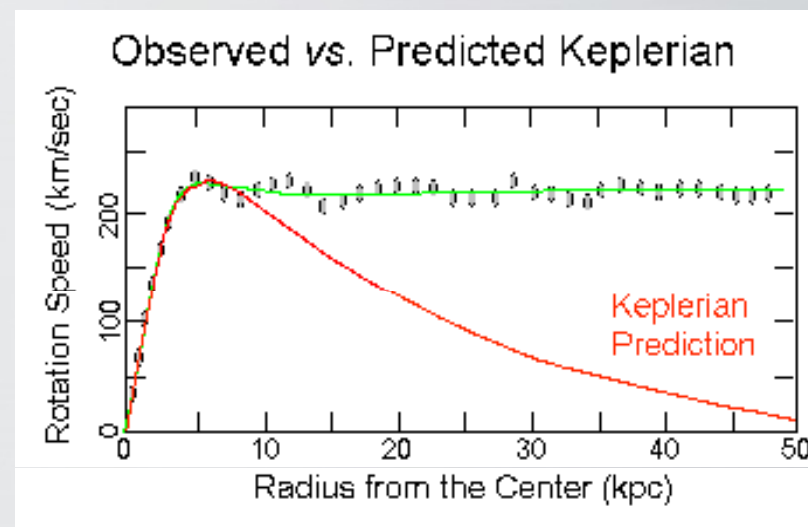


BACKUP

Experimental Evidence for Dark Matter

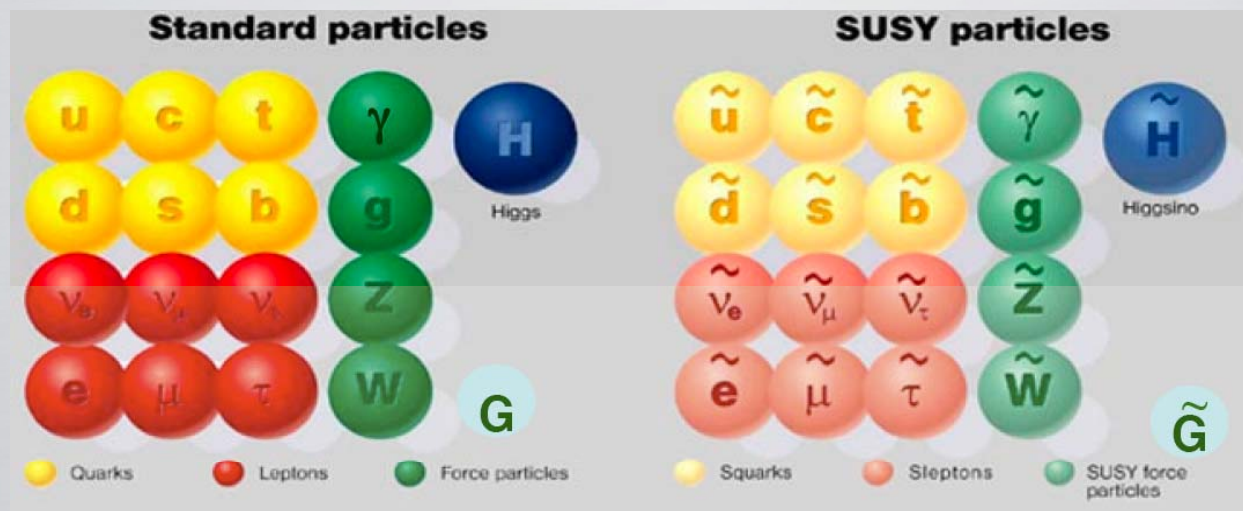
- Zwicky 1933
 - rotation frequencies of galaxies
 - high rotation speed at large radii suggests matter far from the center of the galaxy that is not emitting light
 - Dark matter within the galactic halo

- Bullet cluster
 - collision of two galaxy clusters
 - mass distribution shown in blue
 - determined with gravitational lensing
 - hot gas distribution in red
 - Most of the mass does not interact, only visible matter (gas) is slowed down

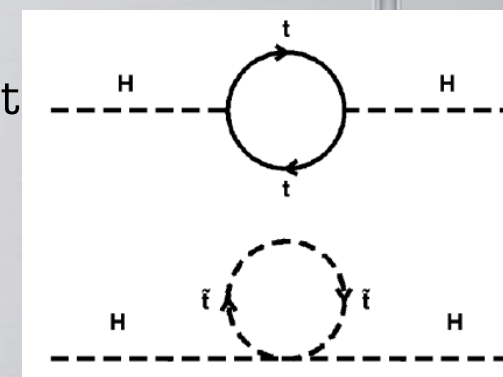


Credit: X-ray: NASA/CXC/CfA/M.Markevitch et al.; Optical: NASA/STScI; Magellan/U.Arizona/D.Clowe et al.; Lensing Map: NASA/STScI; ESO WFI; Magellan/U.Arizona/D.Clowe et al.

Brief intro to SUSY

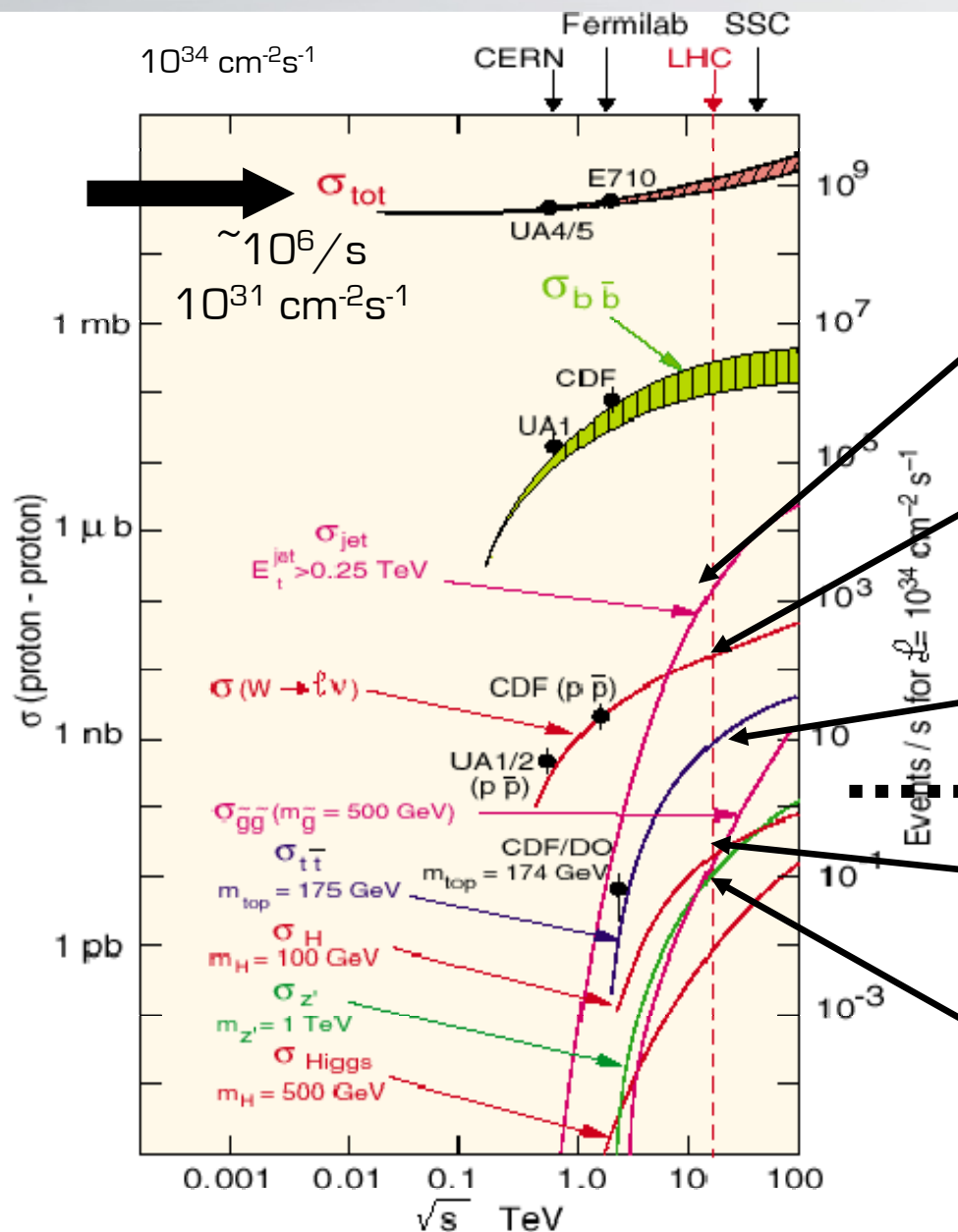


- SUSY partner for every SM particle (with $\frac{1}{2}$ unit of spin different)
 - spin 0 Sfermions (squark, sleptons)
 - spin $\frac{1}{2}$ Gauginos (chargino, neutralino)
- SUSY mass scale expected to be $\sim 1\text{TeV}$ in order to:
 - Solve hierarchy problem (stabilize Higgs mass to radiative correct)
 - Allow unification of strong and electroweak forces
 - Provide sensible dark matter candidate (R-parity)
 - Naturalises scalar (Higgs) sector of SM
- Downside of SUSY
 - Large parts of parameter space ruled out already
 - Many parameters

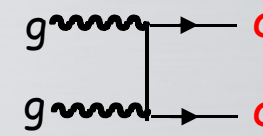


What do we expect to see at LHC?

many



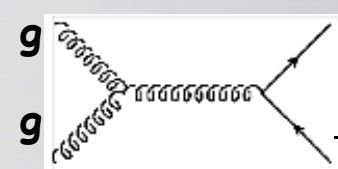
High- p_T QCD jets



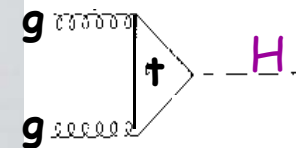
W, Z



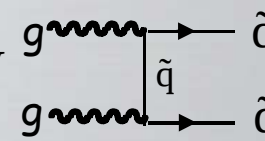
Top-Top



Higgs
 $m_H \approx 100 \text{ GeV}$



$\tilde{q}, \tilde{g}$ pairs, $m \sim 1 \text{ TeV}$



few

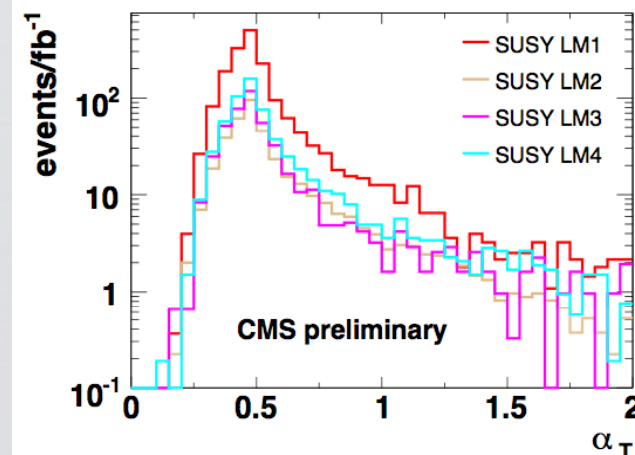
A closer look at SUSY yields

- CMS SUSY benchmark points

Sample	m_0 (GeV)	$m_{1/2}$ (GeV)	A_0	$\tan\beta$	$\text{sign}(\mu)$	σ NLO (pb)	(LO) (pb)	lightest $\tilde{q}$ (GeV)	$\tilde{\chi}_1^0$ (GeV)
LM1	60	250	0	10	+	54.86	(43.28)	410 ($\tilde{t}_1$)	97
LM2	185	350	0	35	+	9.41	(7.27)	582 ($\tilde{t}_1$)	141
LM3	330	240	0	20	+	45.47	(34.20)	446 ($\tilde{t}_1$)	94
LM4	210	285	0	10	+	25.11	(19.43)	483 ($\tilde{t}_1$)	112

- Reminder: desired topology is 2 squarks decaying to squarks and 2 neutralinos (LSPs)

Sample	Events	$\tilde{q} \tilde{q}$ (invisible)	$\tilde{q} \tilde{q}$ (other)	$\tilde{q} \tilde{g}$	$\tilde{g} \tilde{g}$	other
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LM4	195	49%	10%	36%	3%	1%



For comparison:

QCD: 0

$Z \rightarrow \nu\nu$: 57

W/Z: 19

Total: 76

- Dominated by squark-squark, but not only:

- Squark - gluino contribution, where gluino decays to squark+quark
- In LM1: small mass difference between gluino and squark $\Rightarrow$ low p_T 3rd jet

Production process	$p_T^{j3} < 30 \text{ GeV}$	$p_T^{j3} < 50 \text{ GeV}$	$p_T^{j3} < 70 \text{ GeV}$
$\tilde{q} \tilde{q}$	80%	61%	51%
$\tilde{q} \tilde{g}$	18%	34%	44%
$\tilde{g} \tilde{g}$	1%	3%	5%

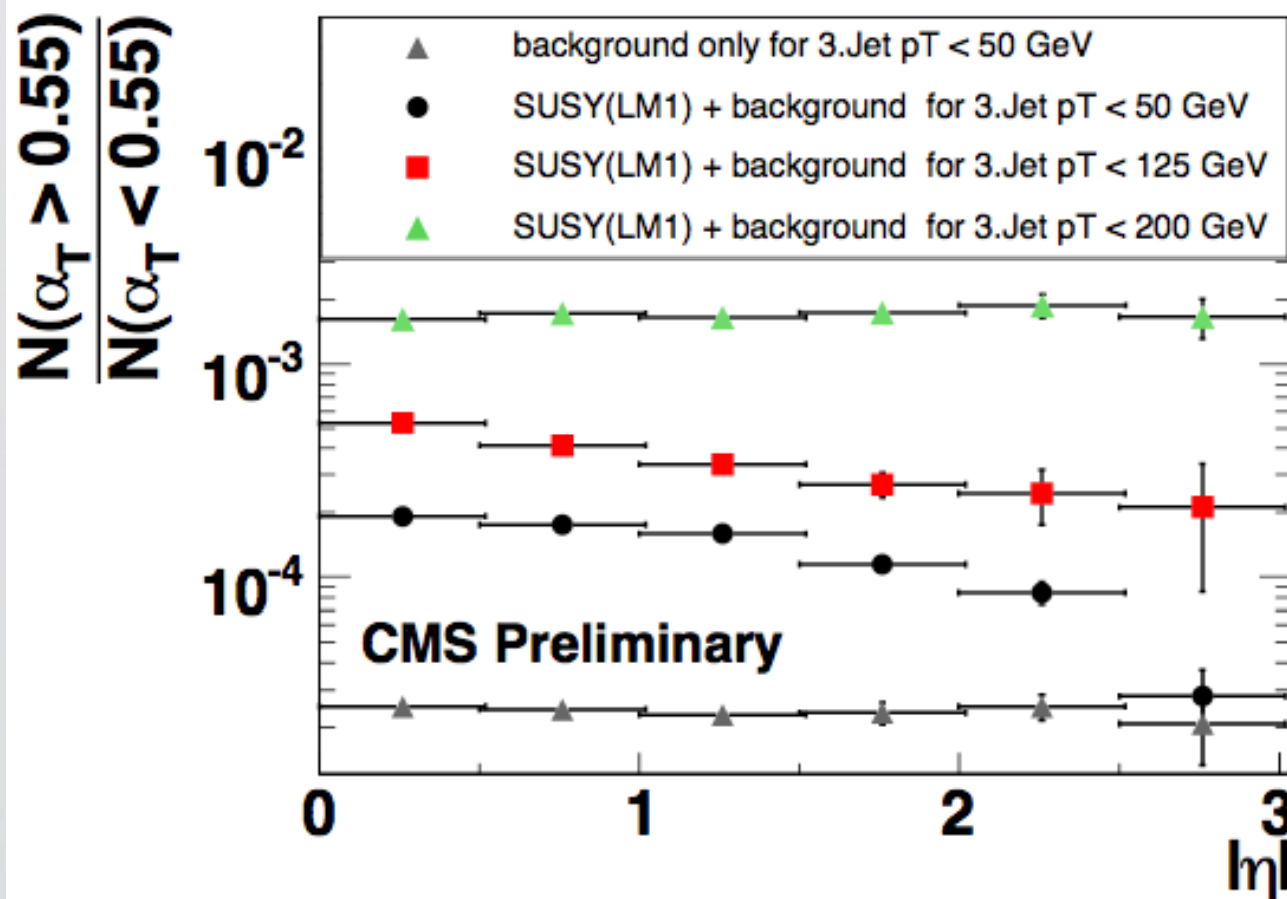
- Indeed observe increase in squark-gluino contribution when relaxing 3rd jet veto

Variation of 3rd jet p_T

Idea:

dilute signal by increasing background contribution
Loosen cut on 3rd jet p_T
to create missing E_T
 $\Rightarrow$ tail in $\alpha(\alpha_T)$

Test if $R\alpha_T$ is stable
Slope should be observed
when signal contribution
becomes sizable



$\Rightarrow$ Slope is observed for hard enough jet veto