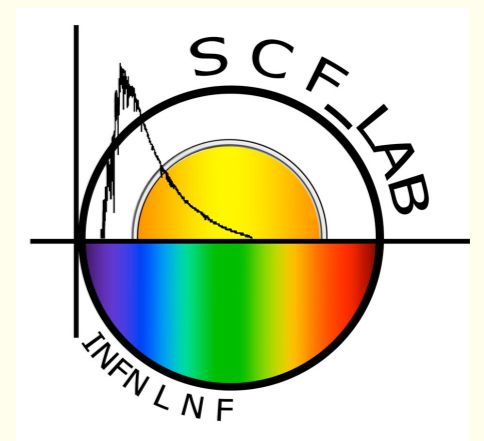


SCF_Lab

Satellite/Lunar/GNSS

laser ranging/altimetry and cube/microsat

Characterization Facilities Laboratory



Optimization and characterization of laser retroreflectors at the SCF_Lab

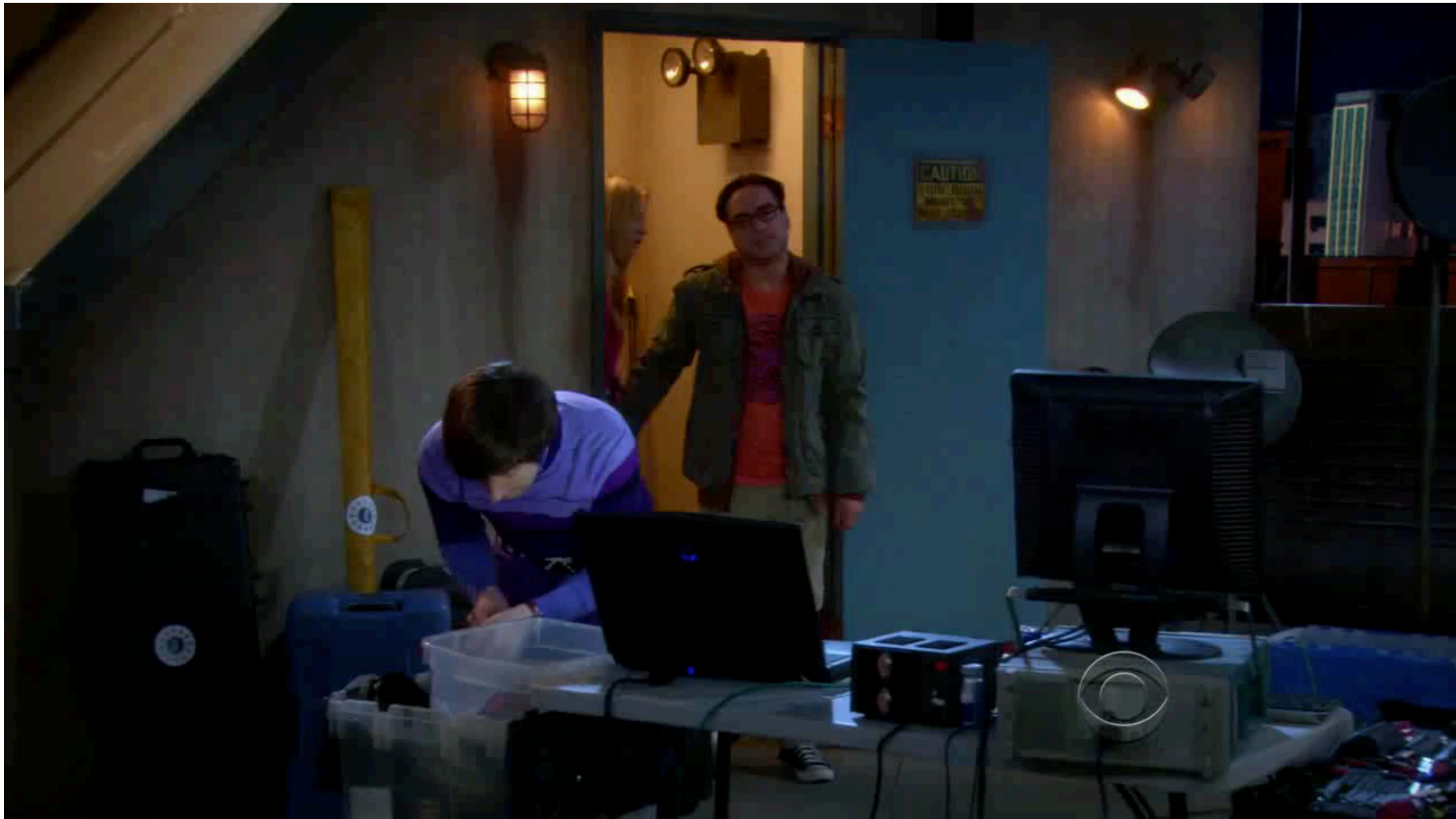
Alessandro Boni

Laboratori Nazionali di Frascati (LNF)-INFN, Frascati, Italy

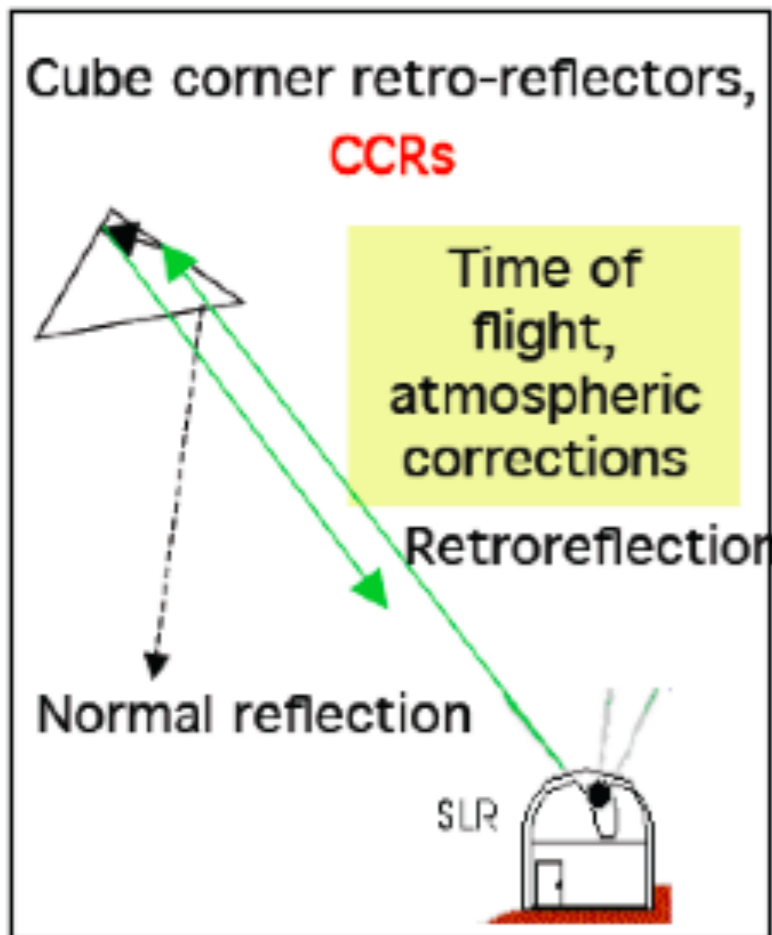
Frascati, 20/10/2015

- Introduction to Satellite and Lunar Laser Ranging
- SLR applications to Earth observation missions
- SLR applications to Global Navigation Satellite System (GNSS)
- Laser Ranging to the Moon
- The SCF_Lab and its innovative activity: characterization of laser retroreflectors in representative space conditions
- The SCF_Lab activities

Leonard explaining Laser Ranging



Satellite Laser Ranging (SLR)
Lunar Laser Ranging (LLR)
Time of flight measurements

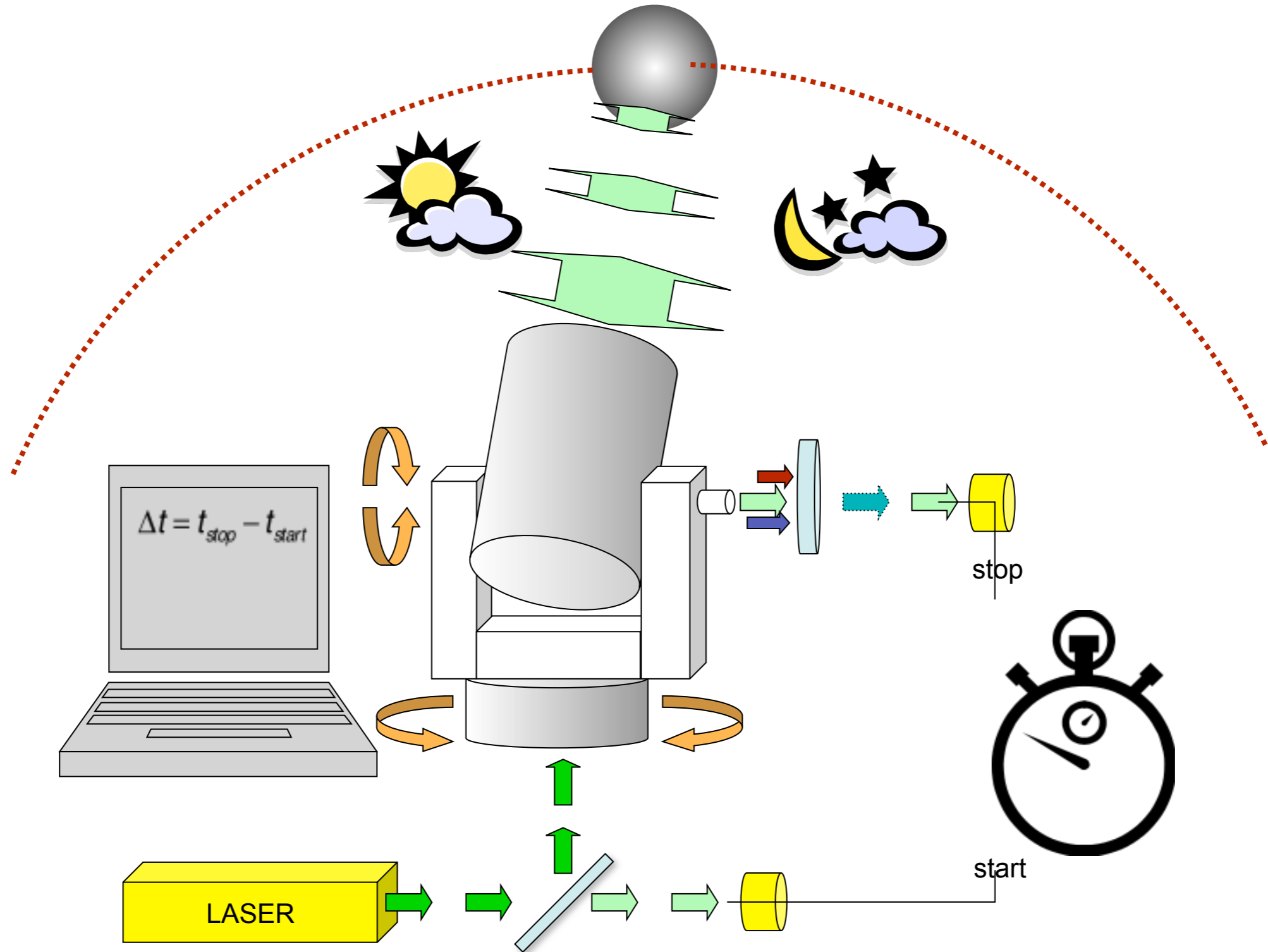


Greenbelt (MD) SLR station

**The most precise and cost effective distance measurement in space
(few millimeters to few centimeters) and (100K€ to M€)**

Laser interferometry much more precise but much more expensive/difficult

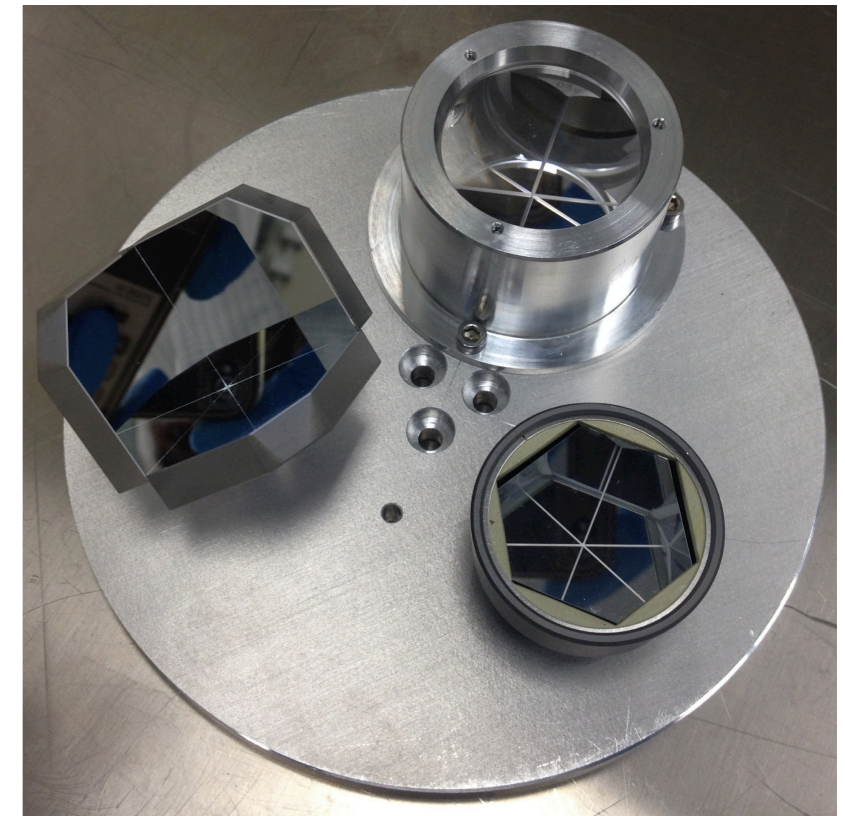
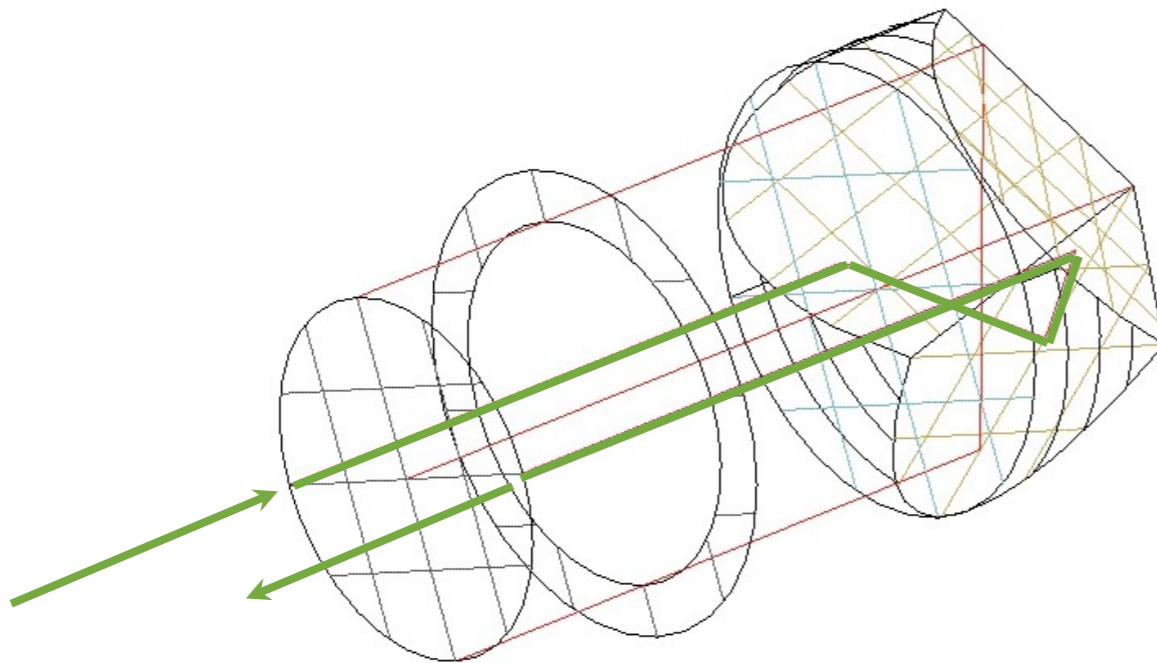
Ranging technique



Corner Cube Reflector

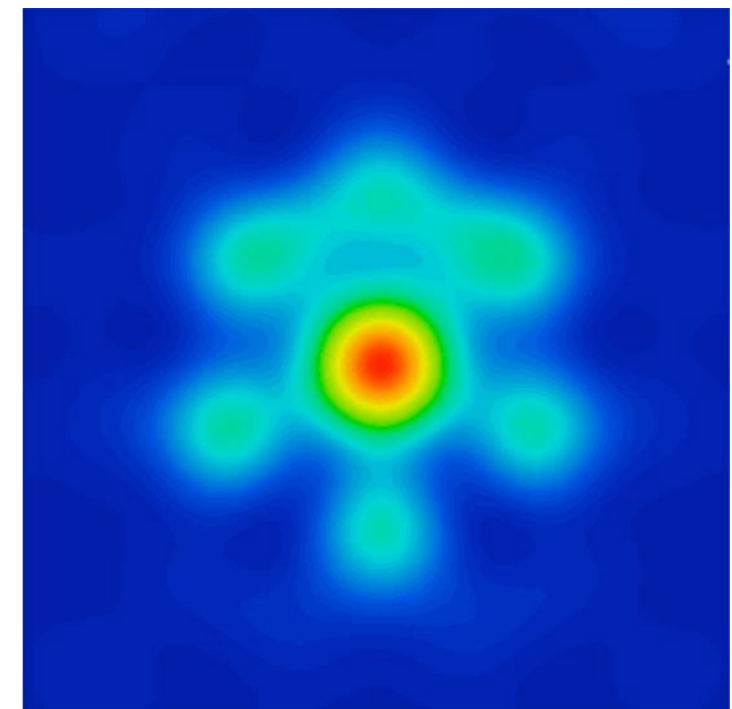
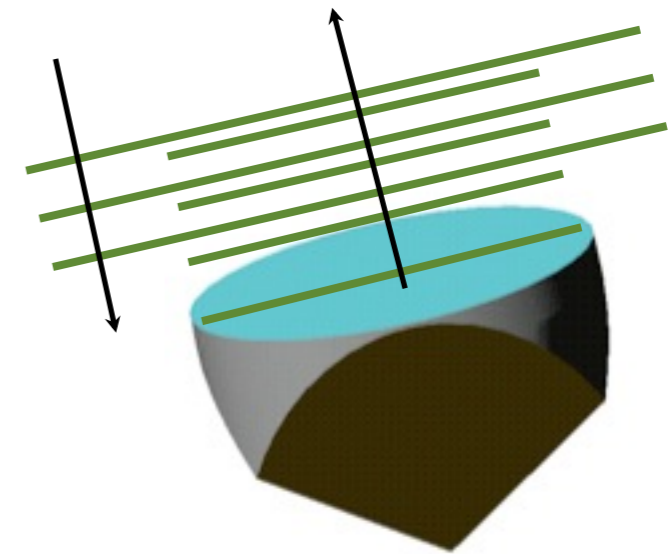
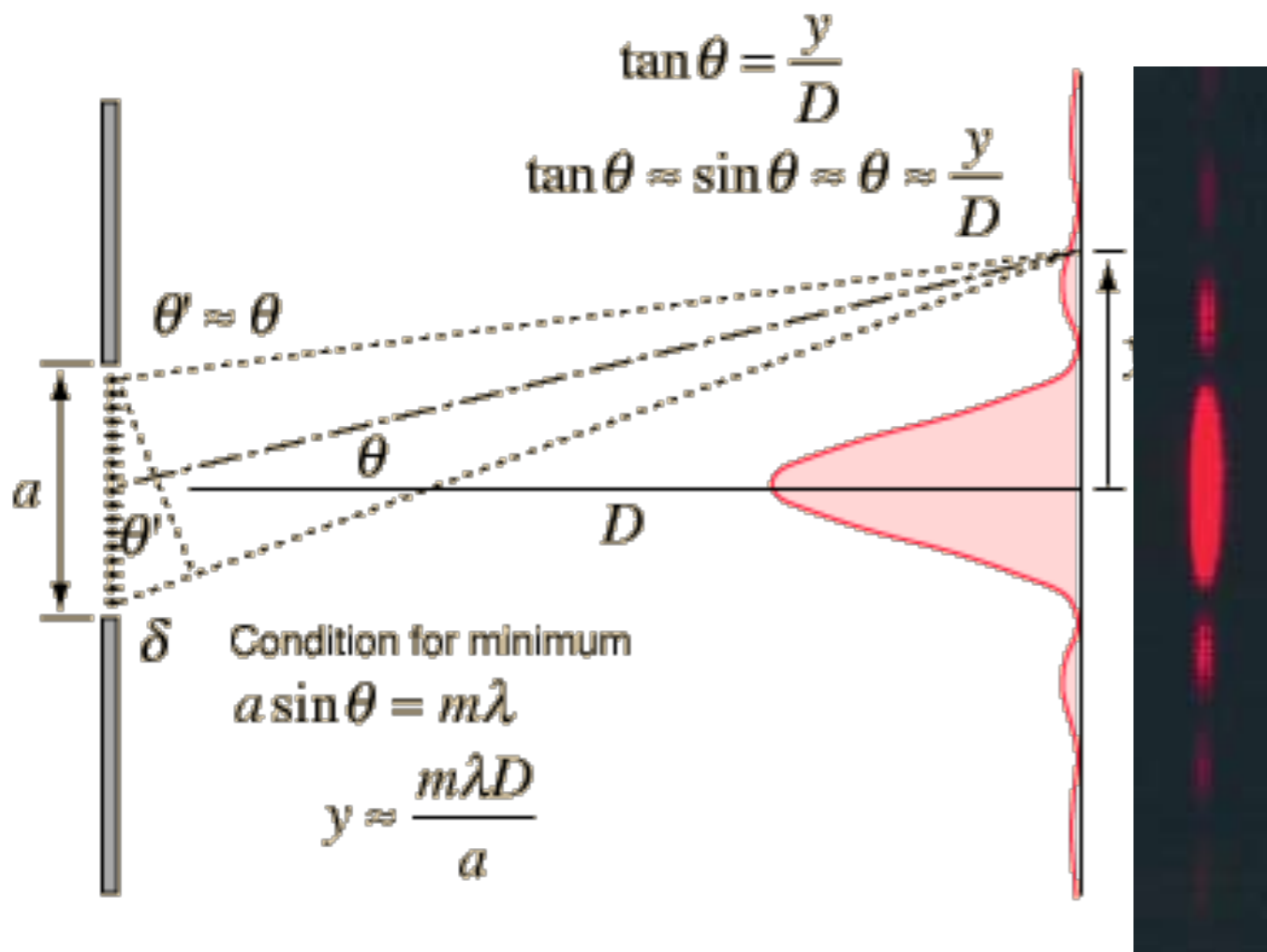


- A CCR is a prism, usually made of Fused Silica, whose vertex is a corner of a cube. Each one of the back faces has an angle of 90° with another.
- A ray entering the CCR is retroreflected along the same direction.
- A ray entering the CCR, comes out in a point opposite to the origin.



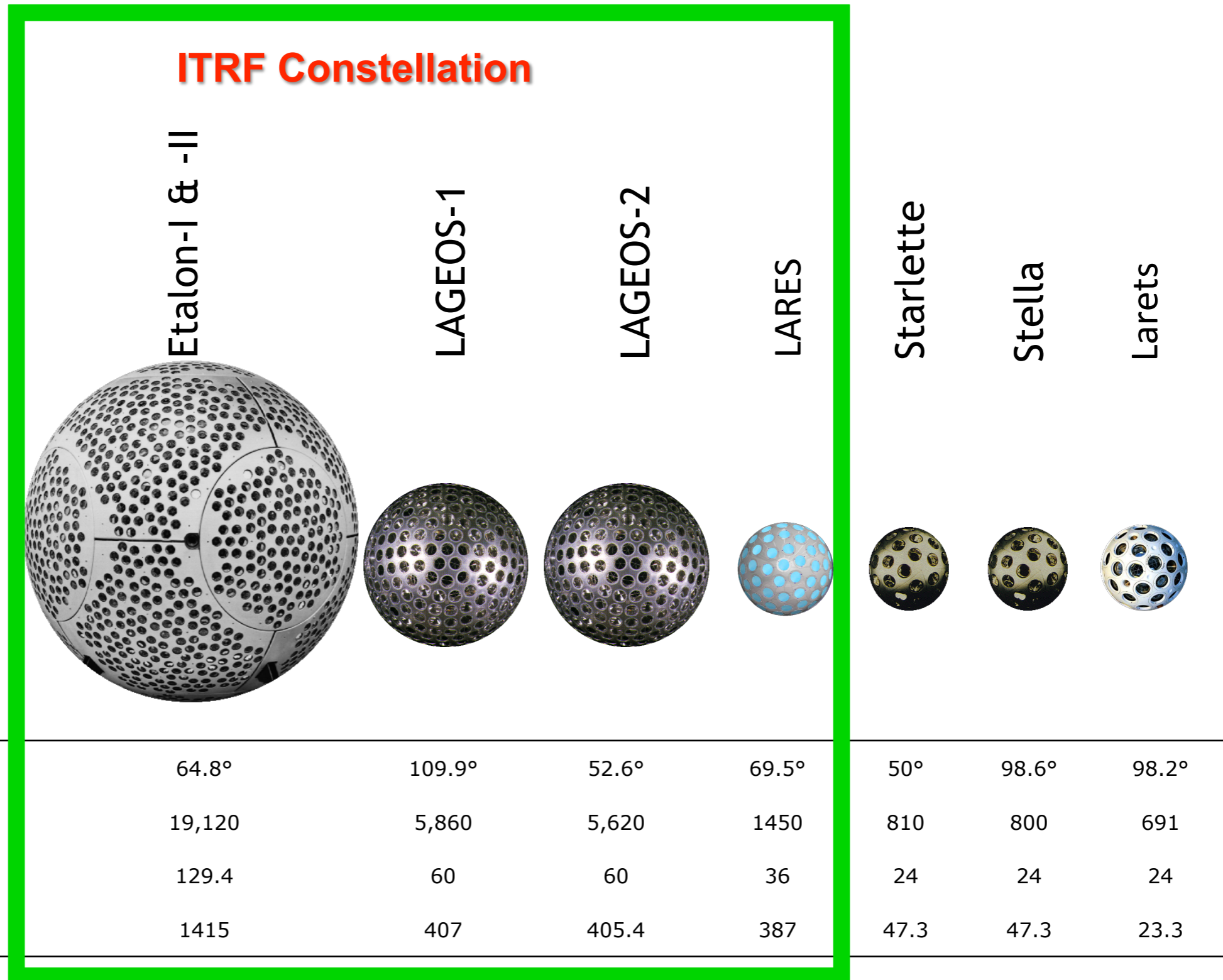
reflection on the three back surfaces through total internal reflection

Diffraction is a phenomenon that occurs when a wave passes through an obstacle or a limited portion of space. On a distant screen a plane wave will result to have a known intensity variation.



Laser Ranging Applications

(Earth geocenter and Inertial frame)



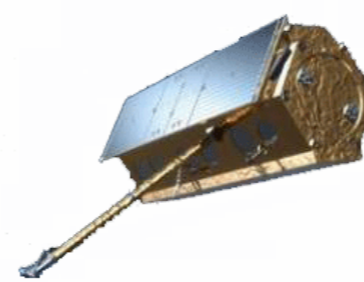
GFO-1



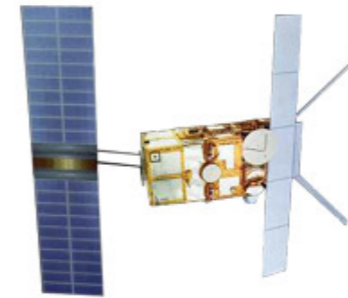
ERS-1



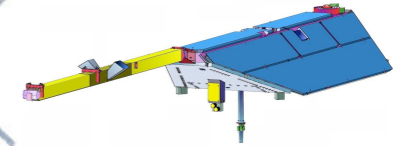
Terra-SAR-X



ERS-2



CHAMP

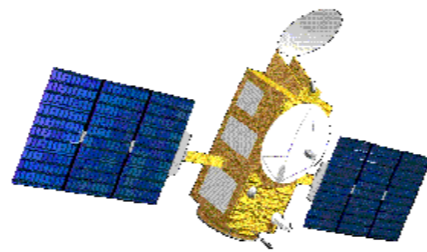


Inclination	108°	98.5°	97.4°	98.5°	87.27°
Perigee ht. (km)	800	780	514	785	474
Mass (kg)	300	2,400	1,230	2,516	400

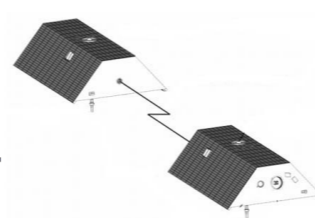
Meteor-3M



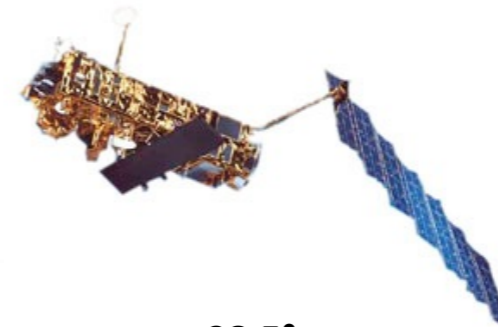
Jason-2



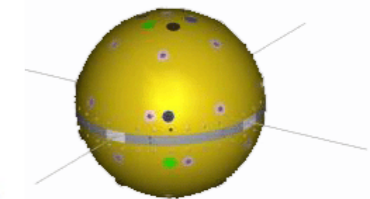
GRACE



Envisat

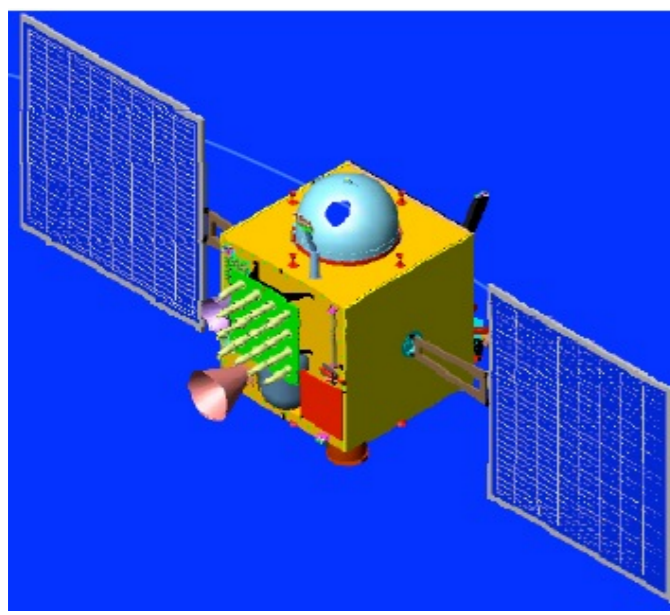


ANDE-RR



Inclination	99.64°	66°	89°	98.5°	51.6°
Perigee ht. (km)	1,012	1,336	450	796	250
Mass (kg)	2,477	500	432/sat.	8,211	50

SLR on GNSS



Indian IRNSS: 7 regional satellites



American GPS: 24 global satellites

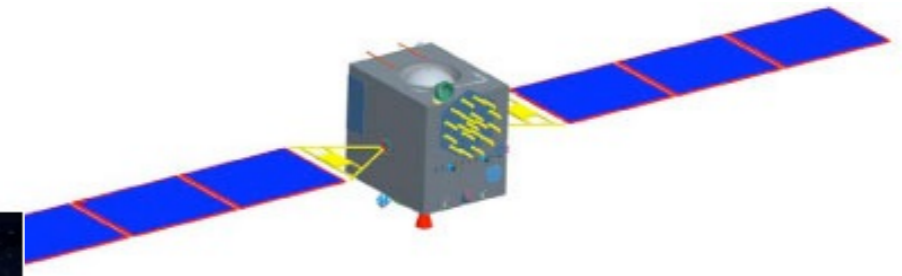
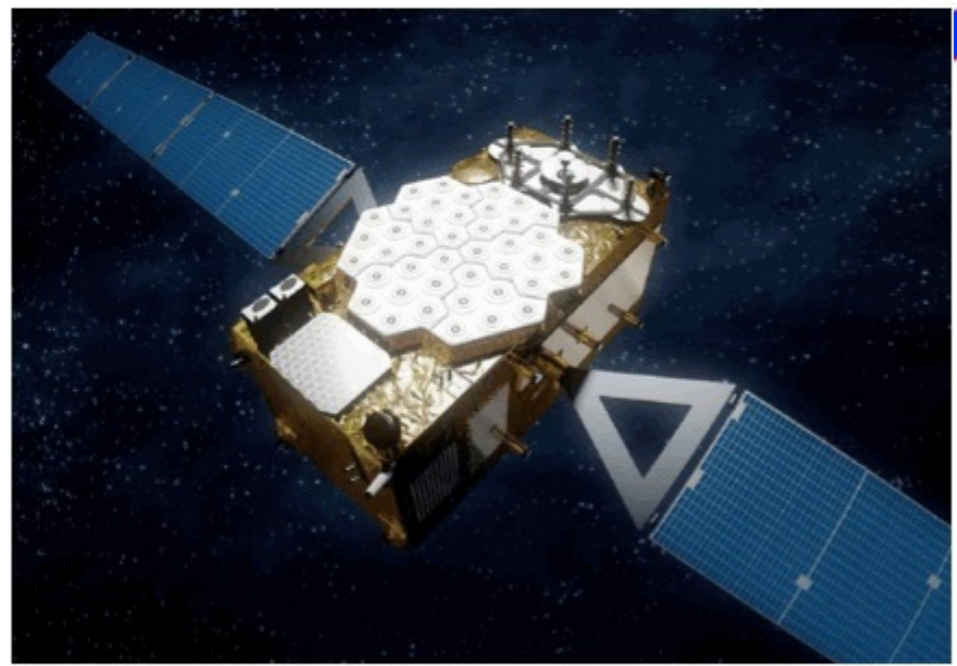


Japanese QZSS: 3 regional satellites



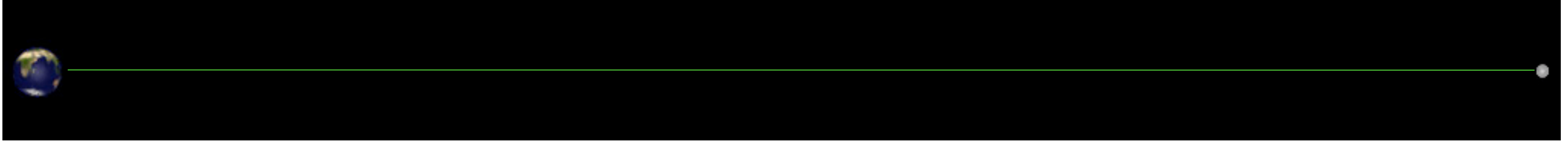
Russian GLONASS: 24 global satellites

European Galileo: 24 global satellites



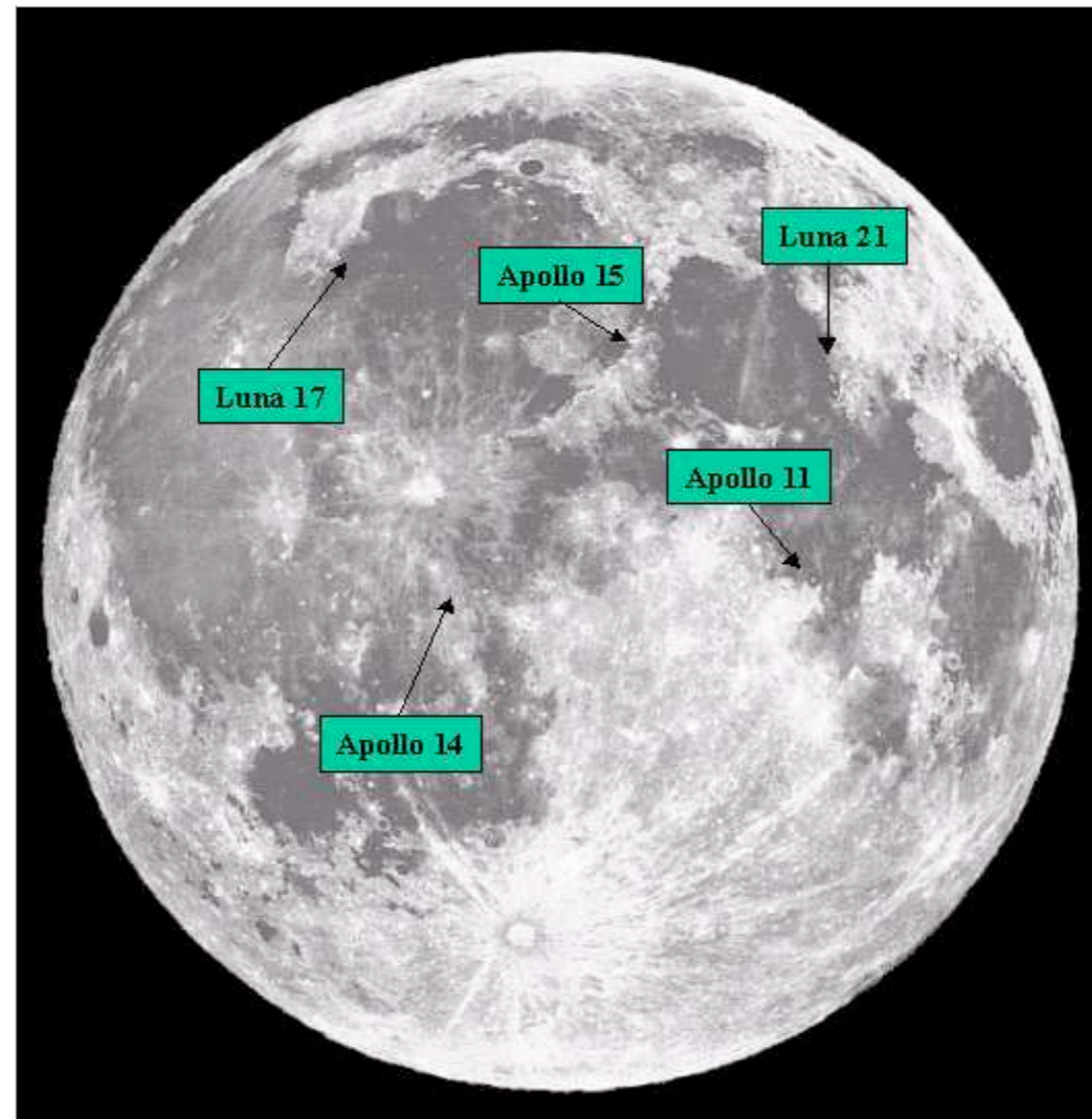
Chinese COMPASS: 30 global and 5 regional satellites

Lunar Laser Ranging: ToF/orbit at $2 \text{ cm} \sim 5 \cdot 10^{-11}$ of Earth-Moon distance

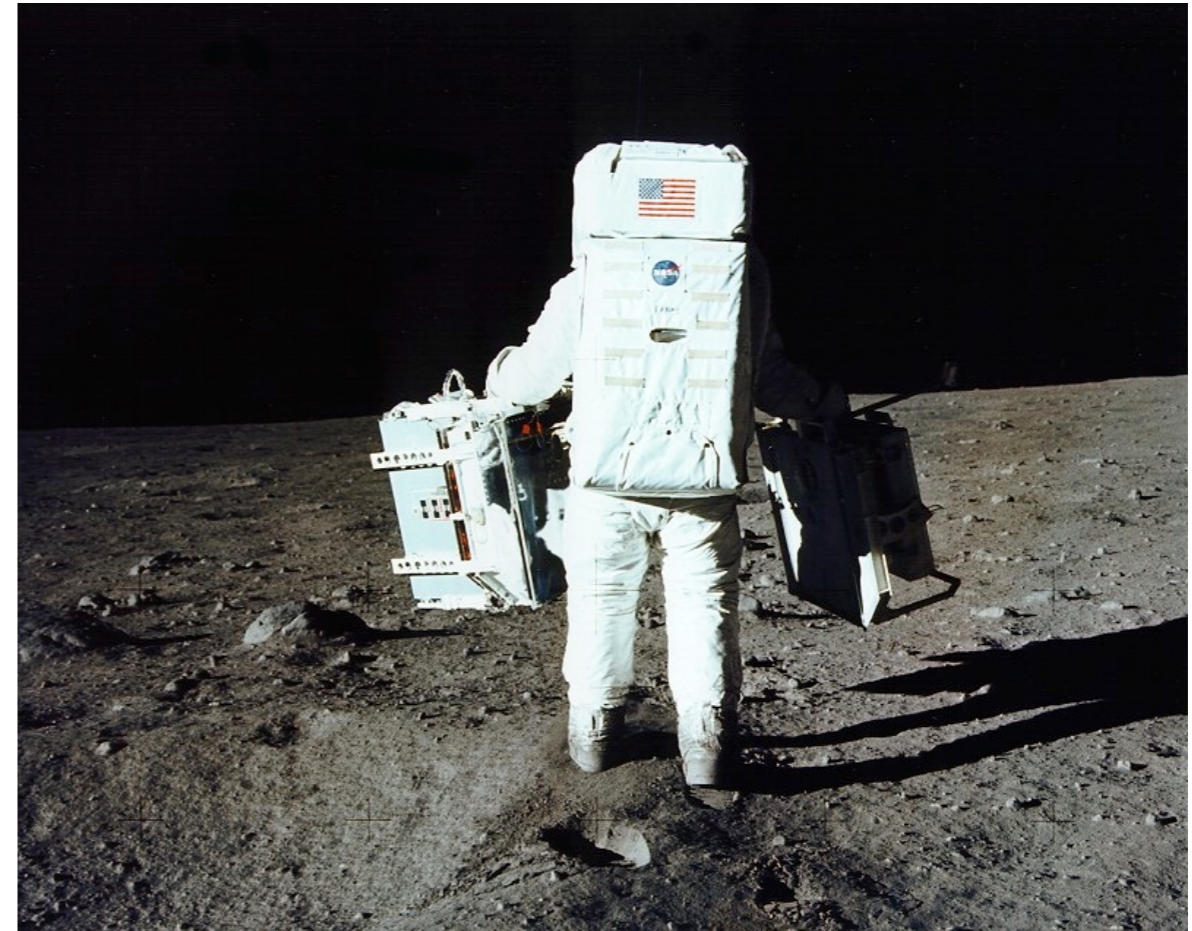
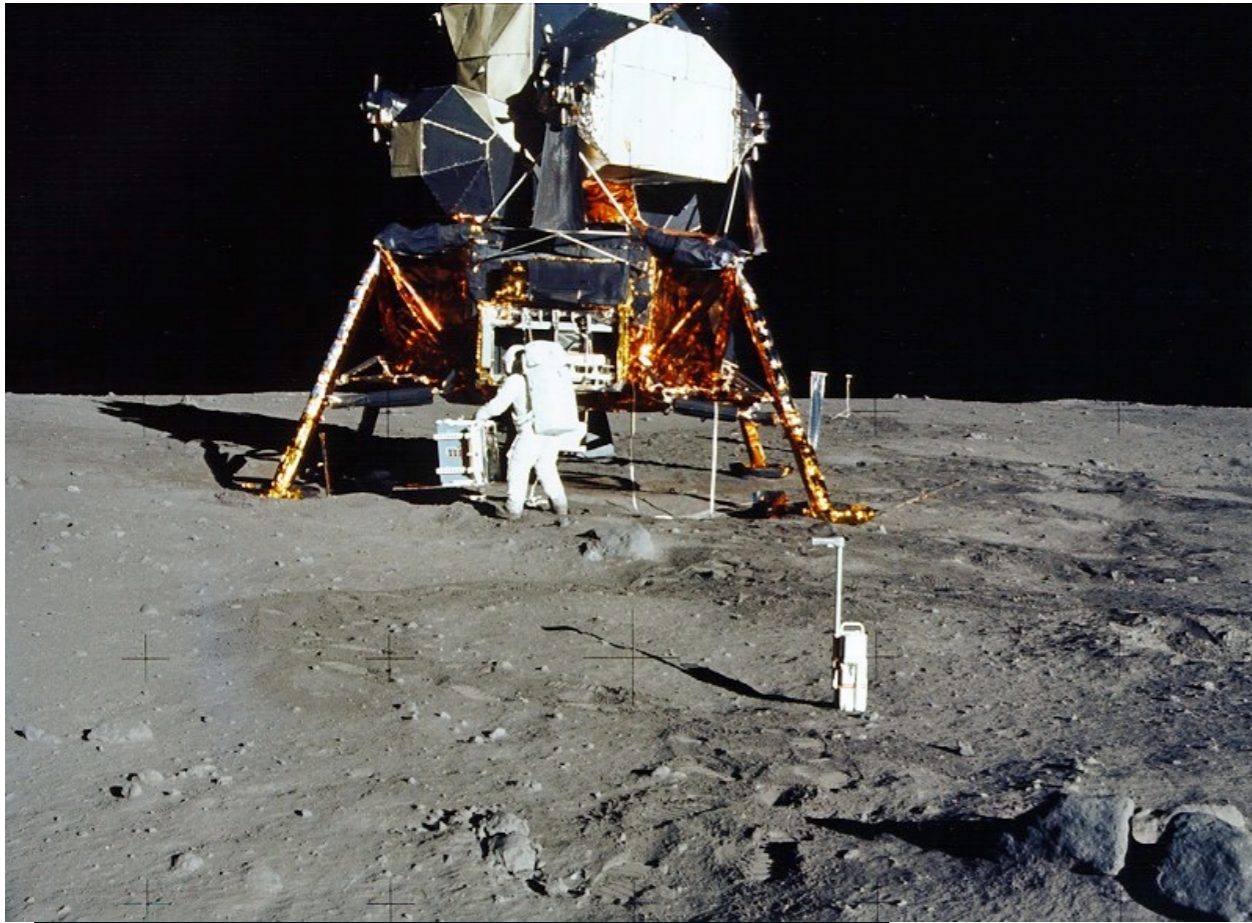


Relative sizes and separation of the Earth–Moon. LLR tof $\sim 2.6 \text{ sec}$ (2-way)

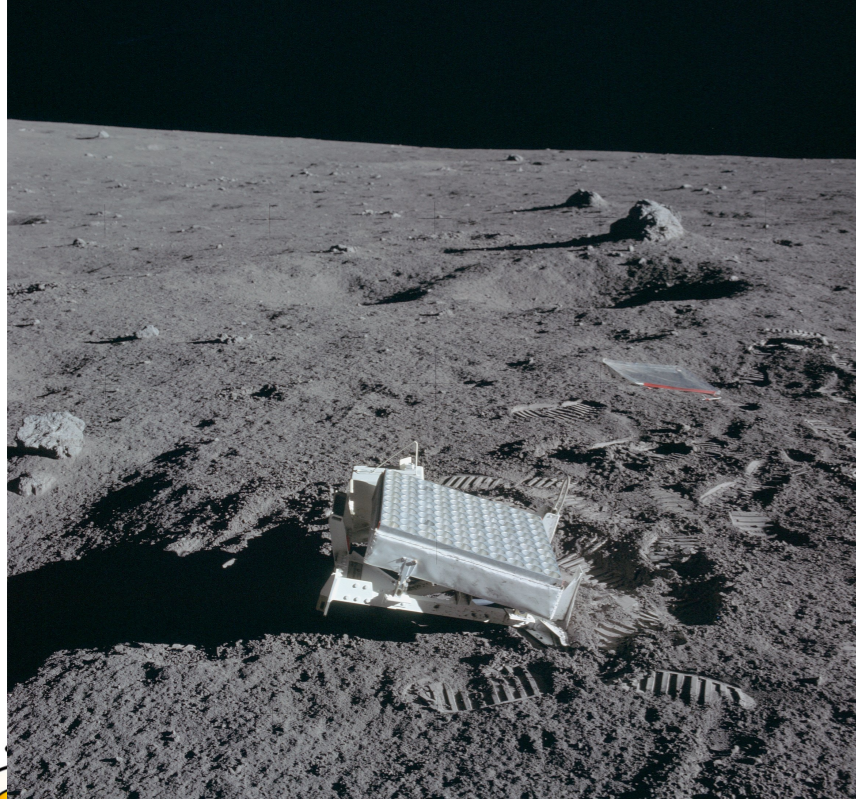
Locations
of 1st Generation
LLR Arrays



Apollo 11 mission

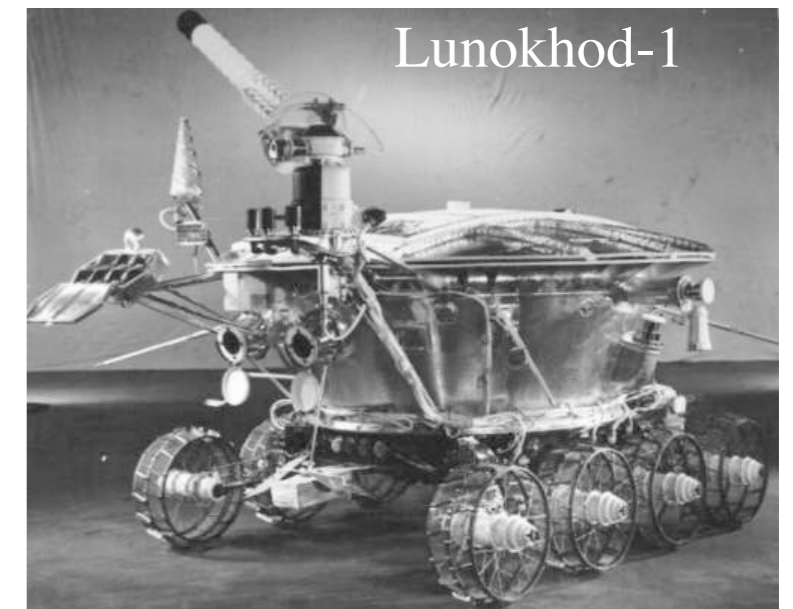
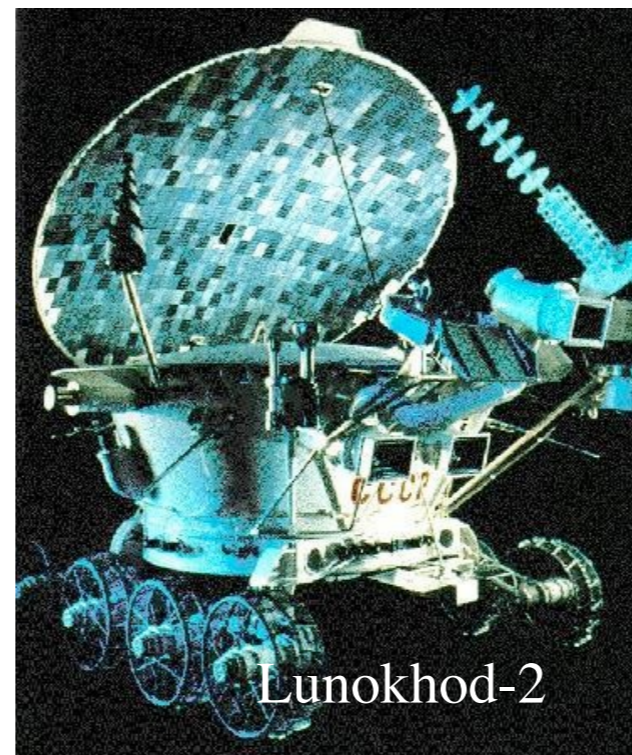
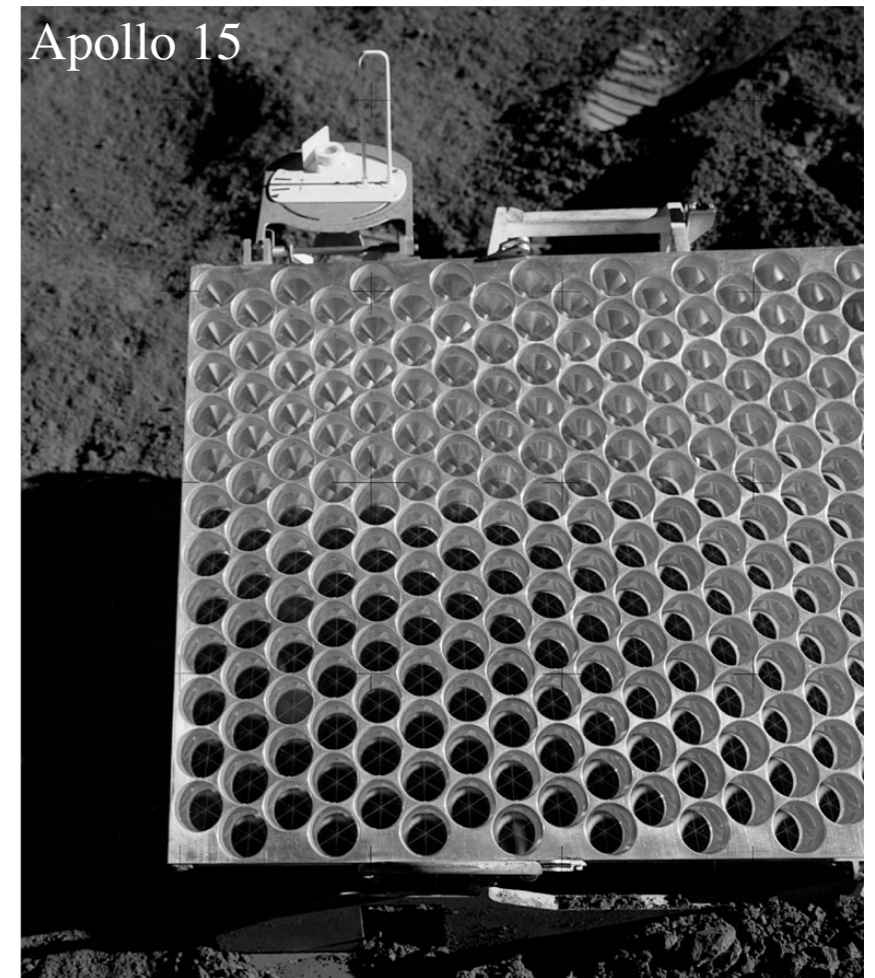
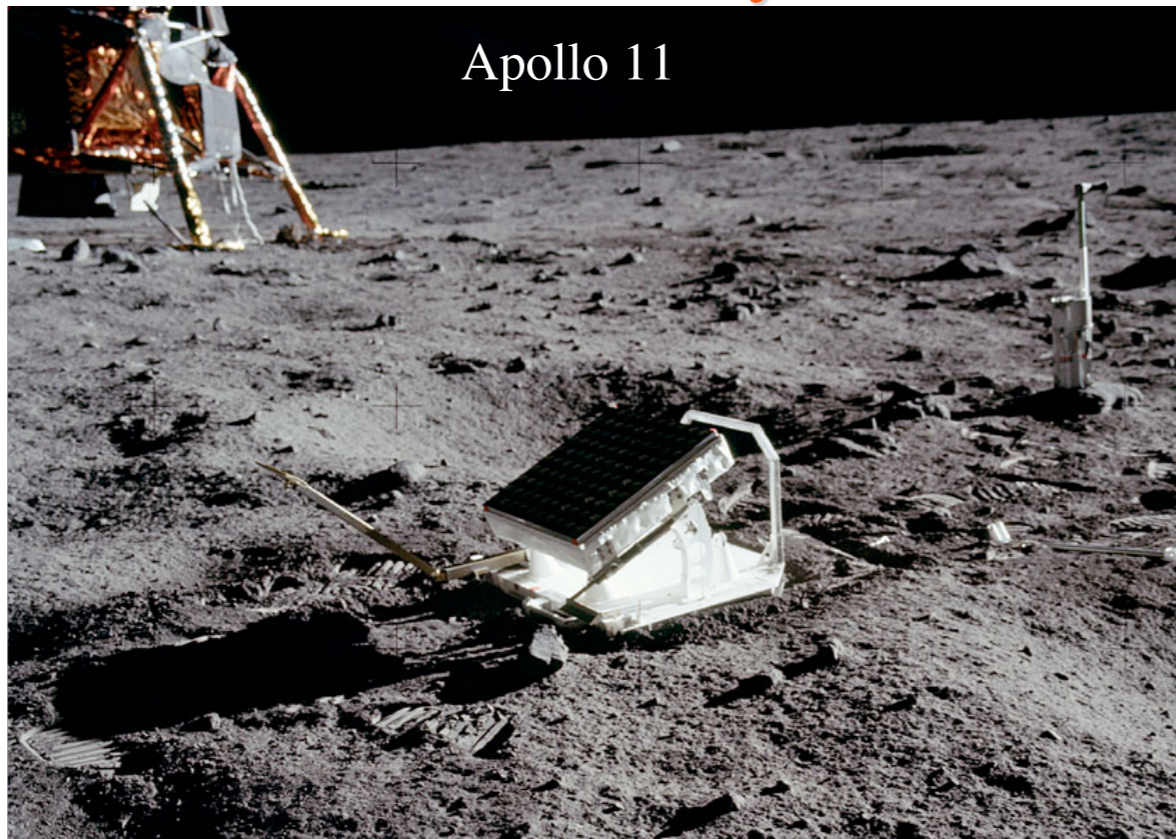


LLR: only Apollo experiment still providing data since July 20, 1969



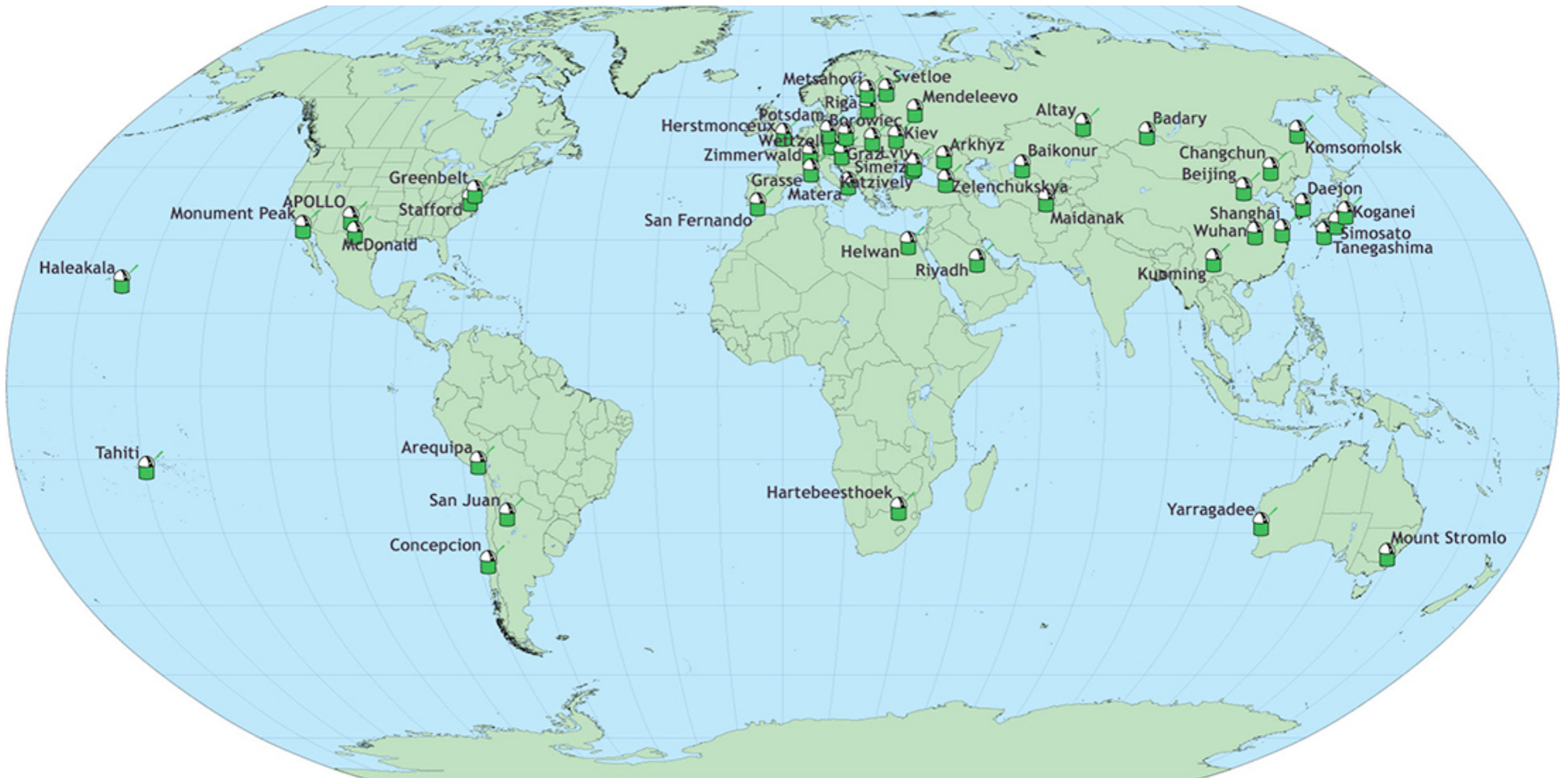
Lunar Laser Ranging arrays

1st Gen. Lunar Reflector Arrays



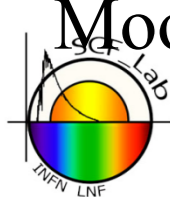
* J. G. Williams, S. G. Turyshev, and D. H. Boggs, PRL 93, 261101 (2004)

Science measurement / Precision test of violation of General Relativity	Apollo/Lunokhod few cm accuracy*	Single Reflectors	
		1 mm	0.1 mm
Parameterized Post-Newtonian (PPN) β	$ \beta-1 < 1.1 \times 10^{-4}$	10^{-5}	10^{-6}
Weak Equivalence Principle (WEP)	$ \Delta a/a < 1.4 \times 10^{-13}$	10^{-14}	10^{-15}
Strong Equivalence Principle (SEP)	$ \eta < 4.4 \times 10^{-4}$	3×10^{-5}	3×10^{-6}
Time Variation of the Gravitational Constant	$ \dot{G}/G < 9 \times 10^{-13} \text{yr}^{-1}$	5×10^{-14}	5×10^{-15}
Inverse Square Law (ISL)	$ \alpha < 3 \times 10^{-11}$	10^{-12}	10^{-13}
Geodetic Precession	$ K_{gp} < 6.4 \times 10^{-3}$	6.4×10^{-4}	6.4×10^{-5}

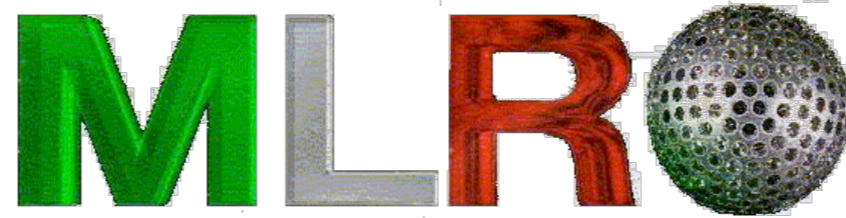


A network of about 40 ground stations routinely track satellites equipped with retroreflectors and give information about their orbit. 4 stations in the world track the

Moon

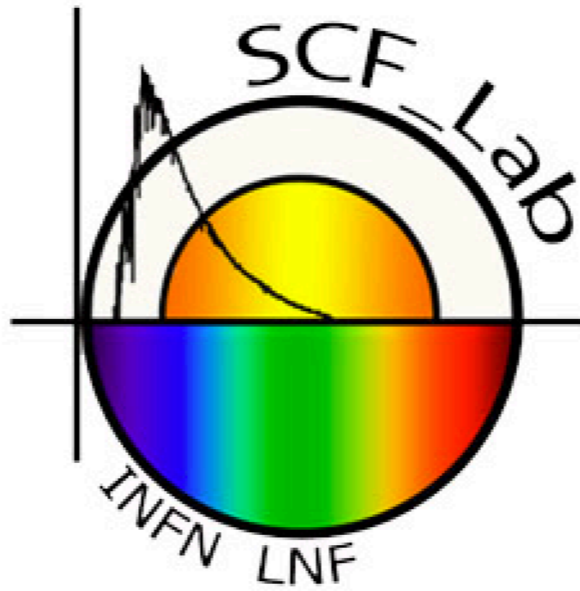


Matera Laser Ranging Station



Slide courtesy of
G. Bianco





The SCF_Lab and its activities

The SCF_Lab

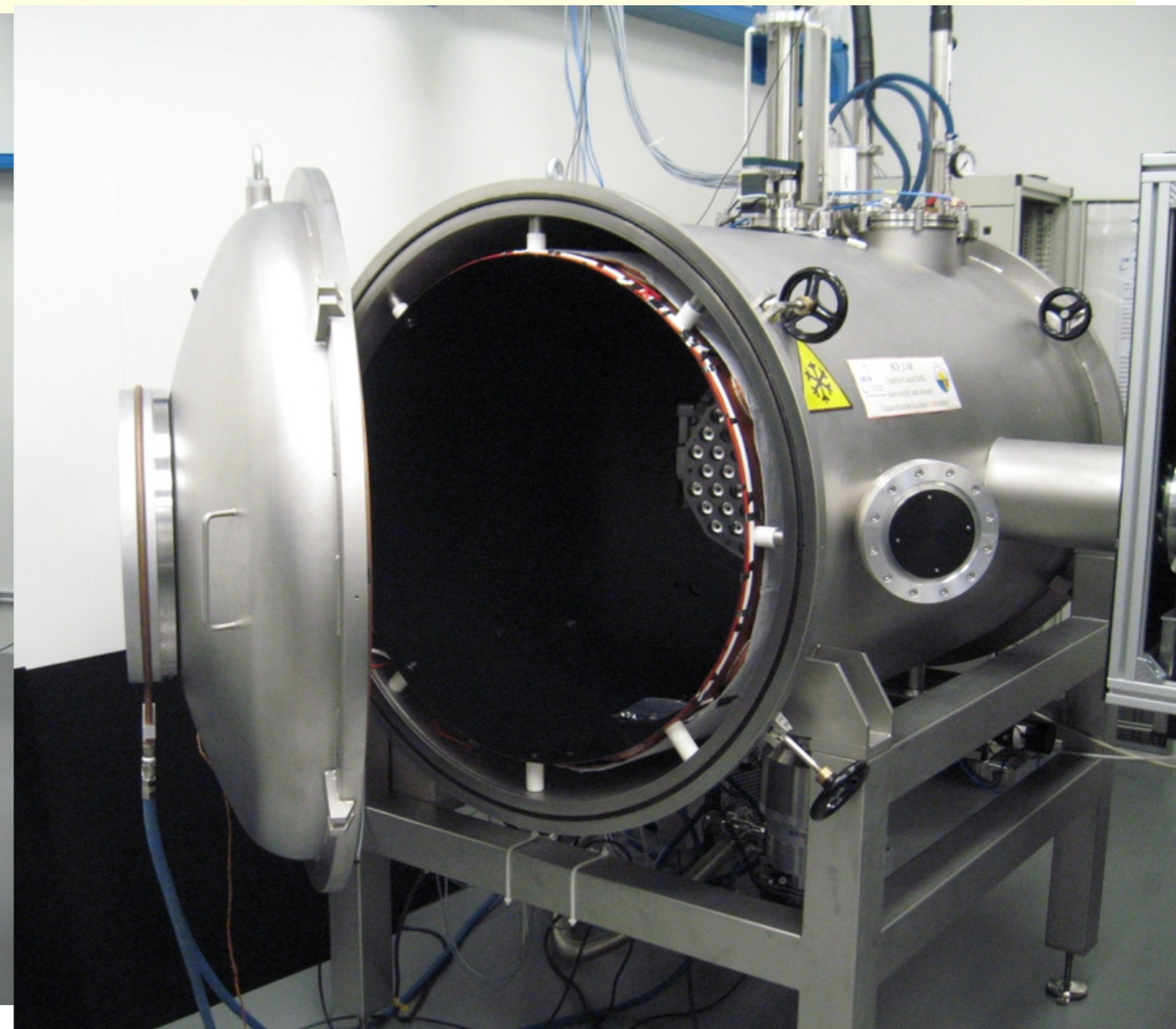
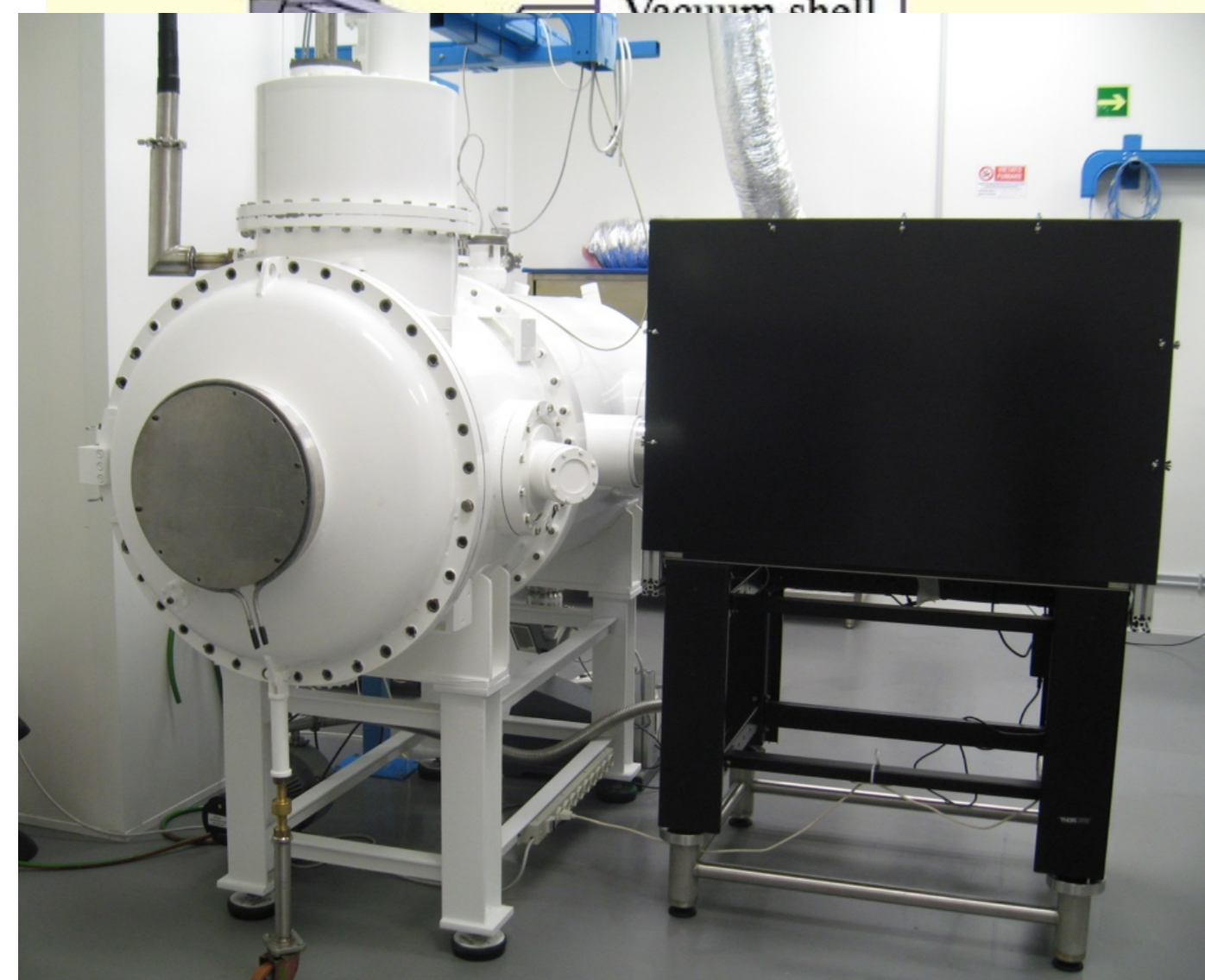


(Satellite/lunar/GNSS laser ranging/altimetry and cube/microsat Characterization Facility)

SCF

Vacuum shell

SCF-G



- environment pressure down to 10^{-6} mbar
- temperature of the chamber at ~ 80 K
- high emissivity cold shield
- solar radiation through a Solar Simulator replicating AM0

Temperature

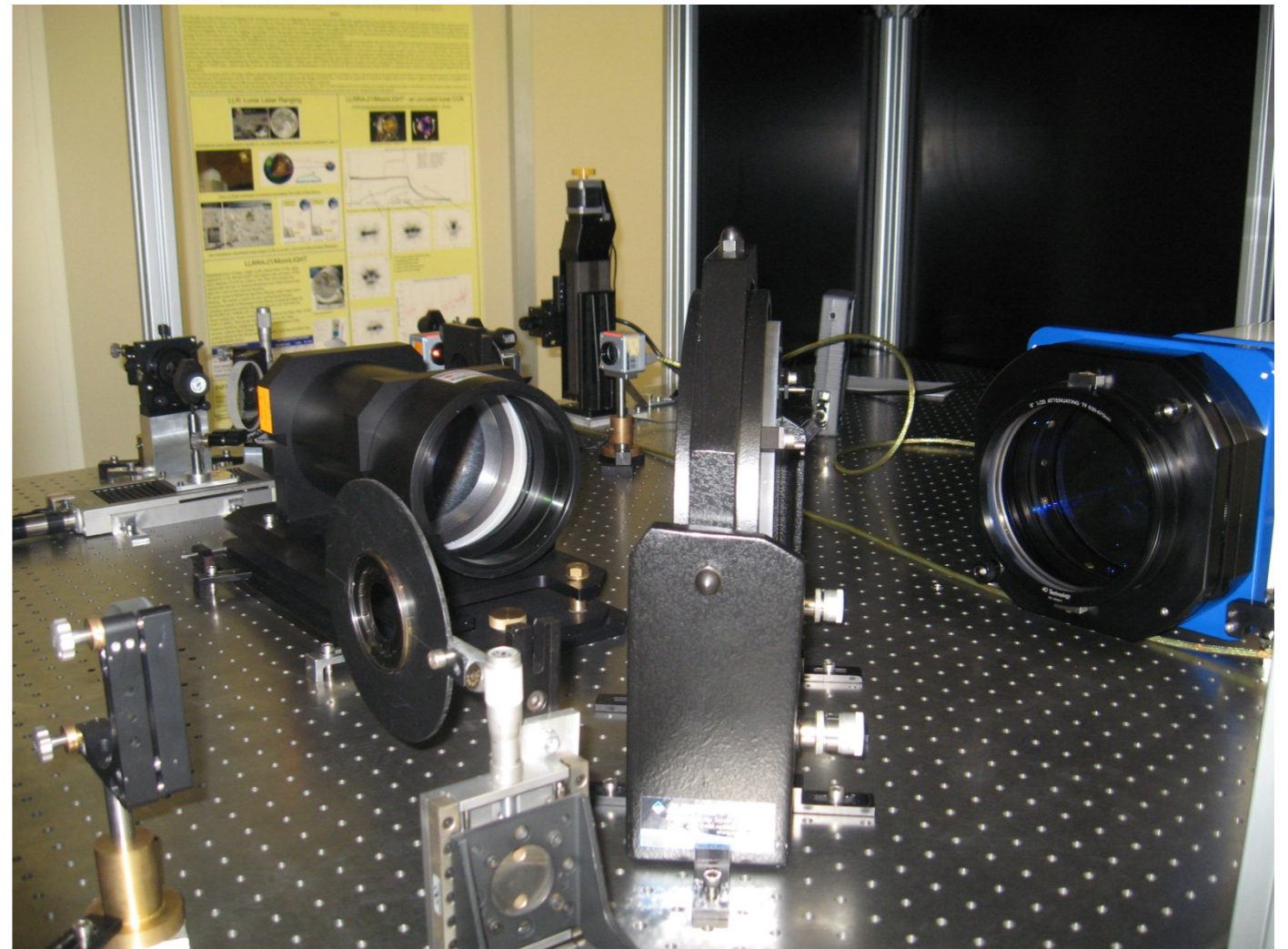


Invasive: platinum RTD probes



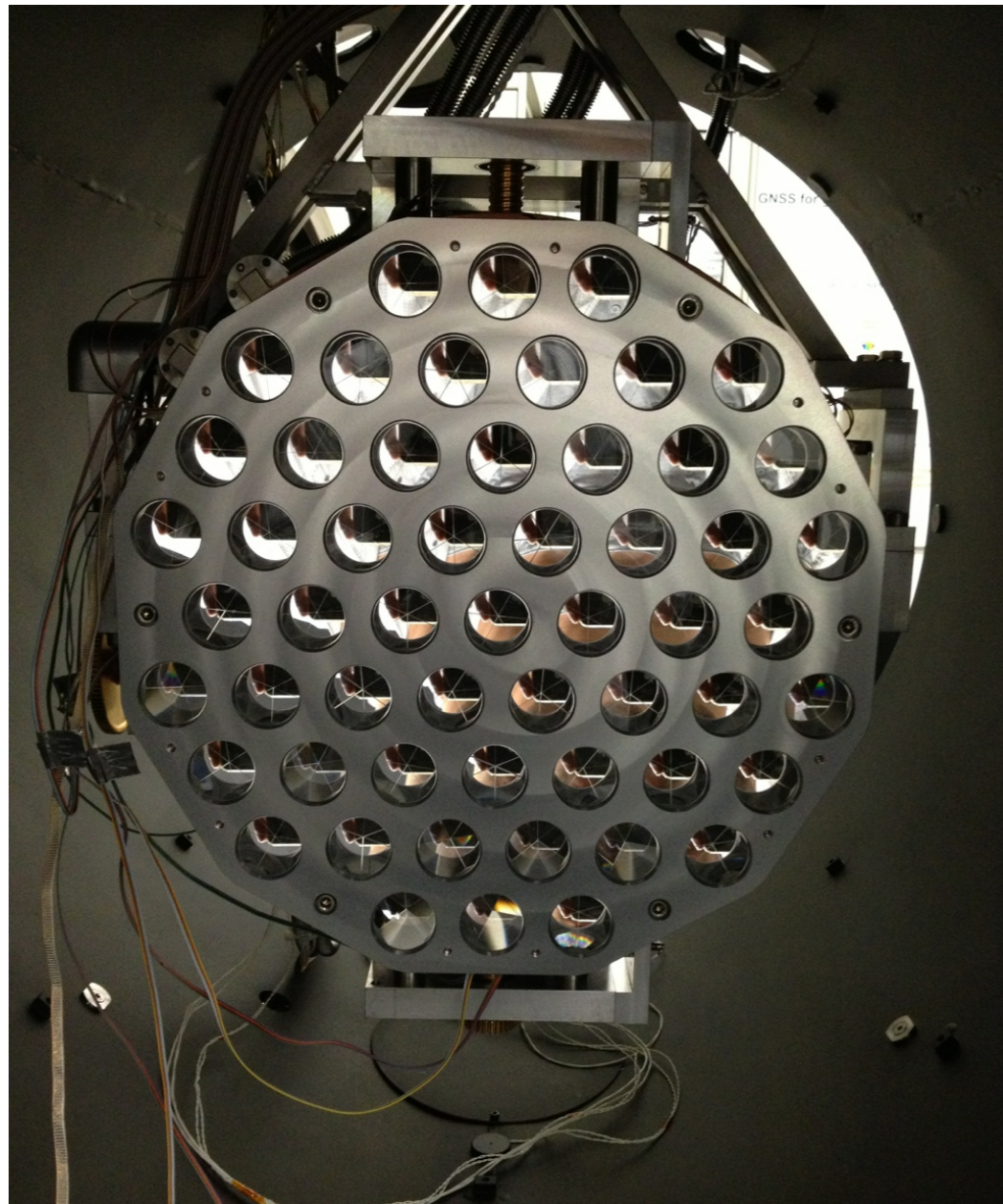
Non-invasive: InfraRed Camera

Optical Response



Optical circuit for FFDP measurement

The purpose of the SCF_Lab measurements is to characterize the whole payload, retroreflectors and their supporting structure under realistic space conditions, in order to determine their compliance to design specification and variation of performance in space.

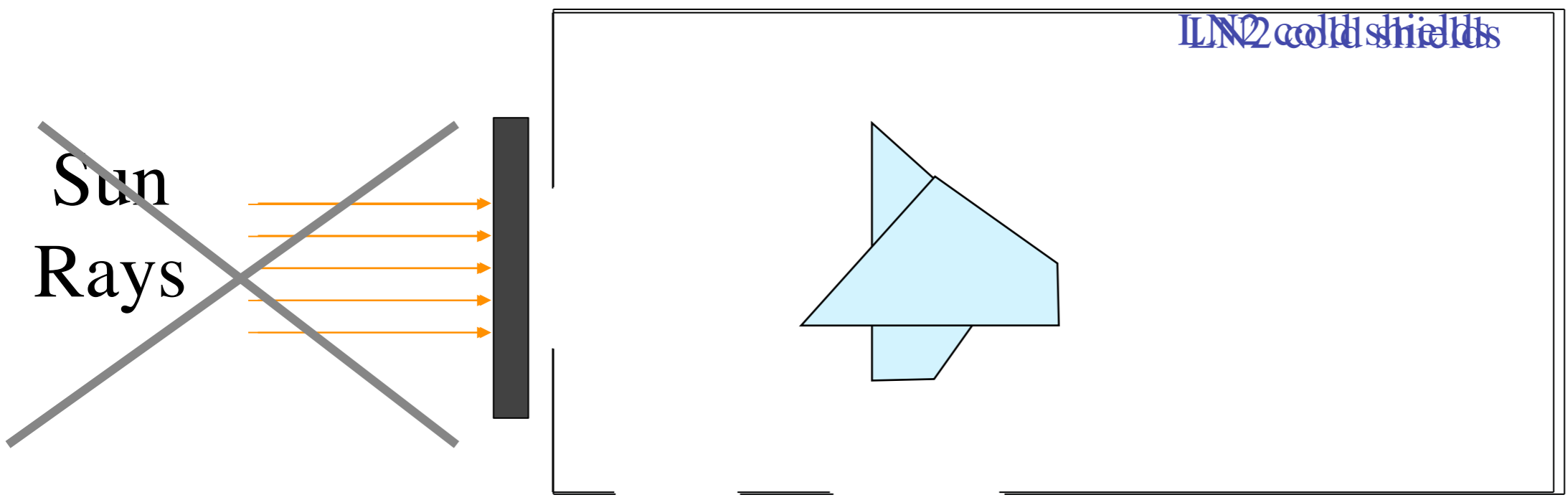


SCF_Lab measurements

- Far Field Diffraction Pattern (FFDP) measurement in Air of CCRs
- SCF-Test
- Simulated orbital measurement

S. Dell'Agnello, et al., **Creation of a New Industry-Standard Space Test of Laser Retroreflectors for GNSS and LAGEOS**, J. Adv. Space Res., DOI: 10.1016/j.asr.2010.10.022

S. Dell'Agnello et al., **ETRUSCO2: an ASI-INFN project of technological development and SCF-Test of GNSS Retroreflector Arrays**, 3rd Int. Colloquium-Galileo Science, Copenhagen, Sept 2011



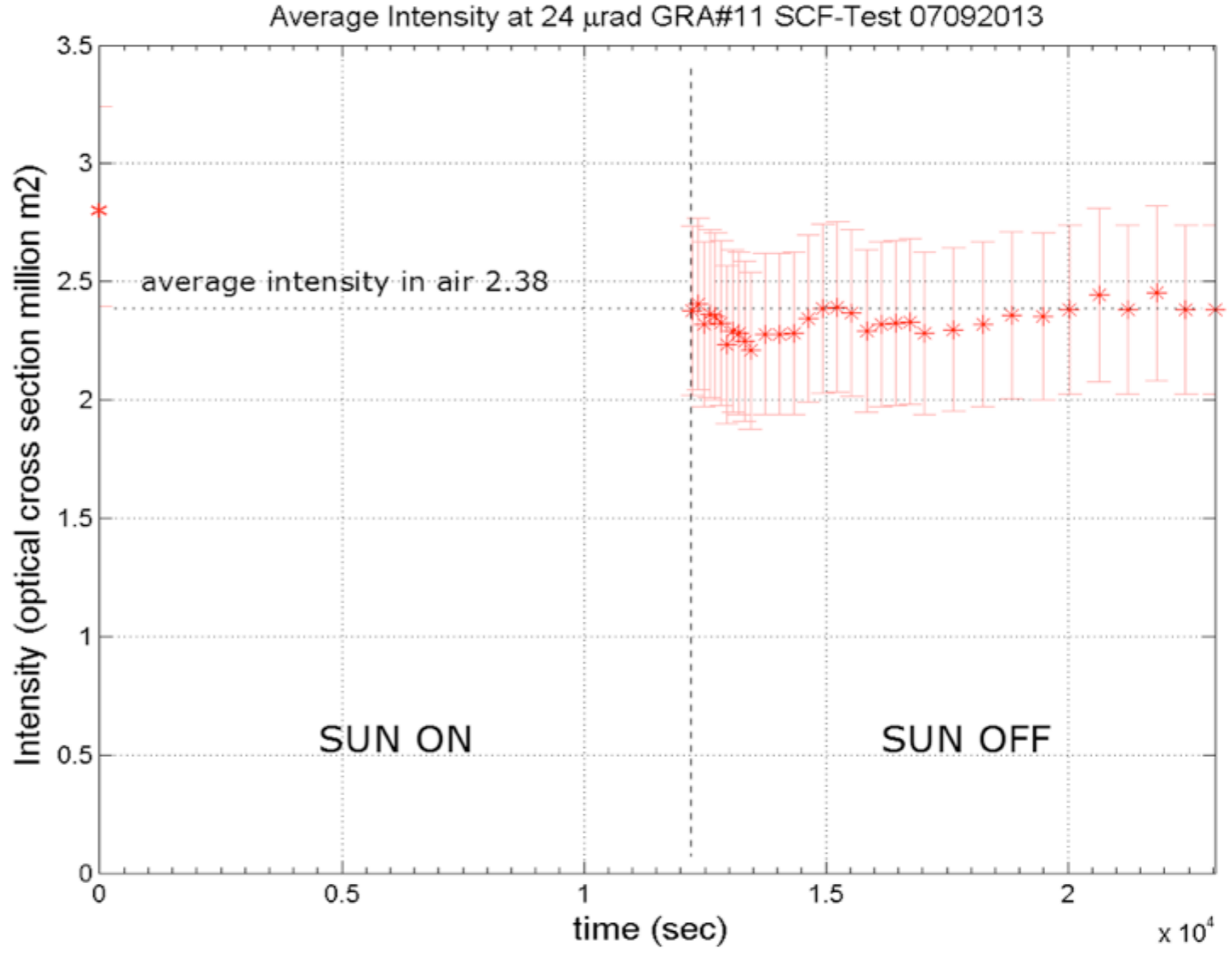
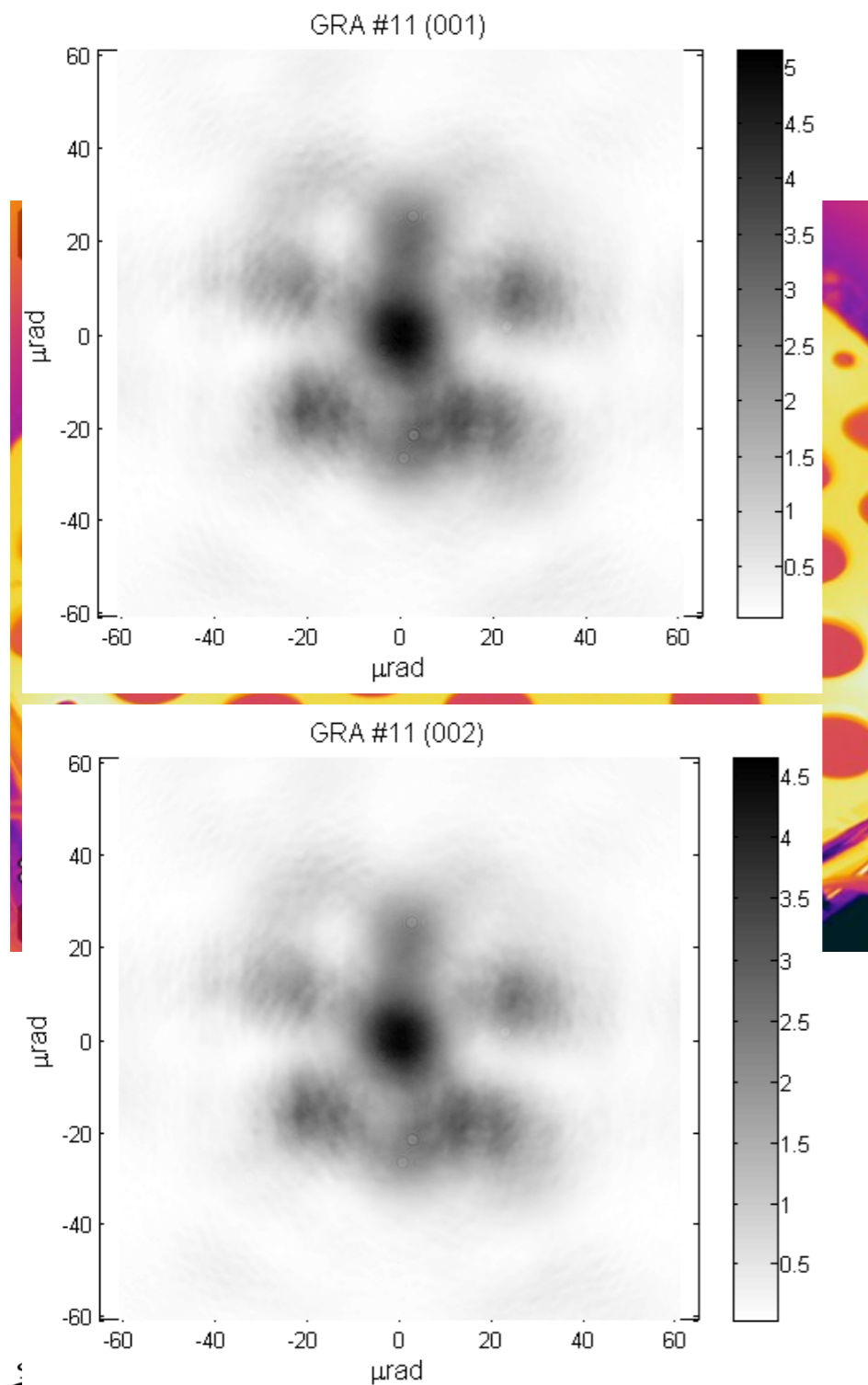
IR camera
Image acquisition



532 nm laser beam
FFDP acquisition

1. **preliminary**: necessary to achieve conditions of equilibrium with space environment
2. **SUN ON**: CCR in front of Solar Simulator for 3 hours. (IR measurements)
3. **SUN OFF**: CCR in front of laser window (IR and FFDP measurements)

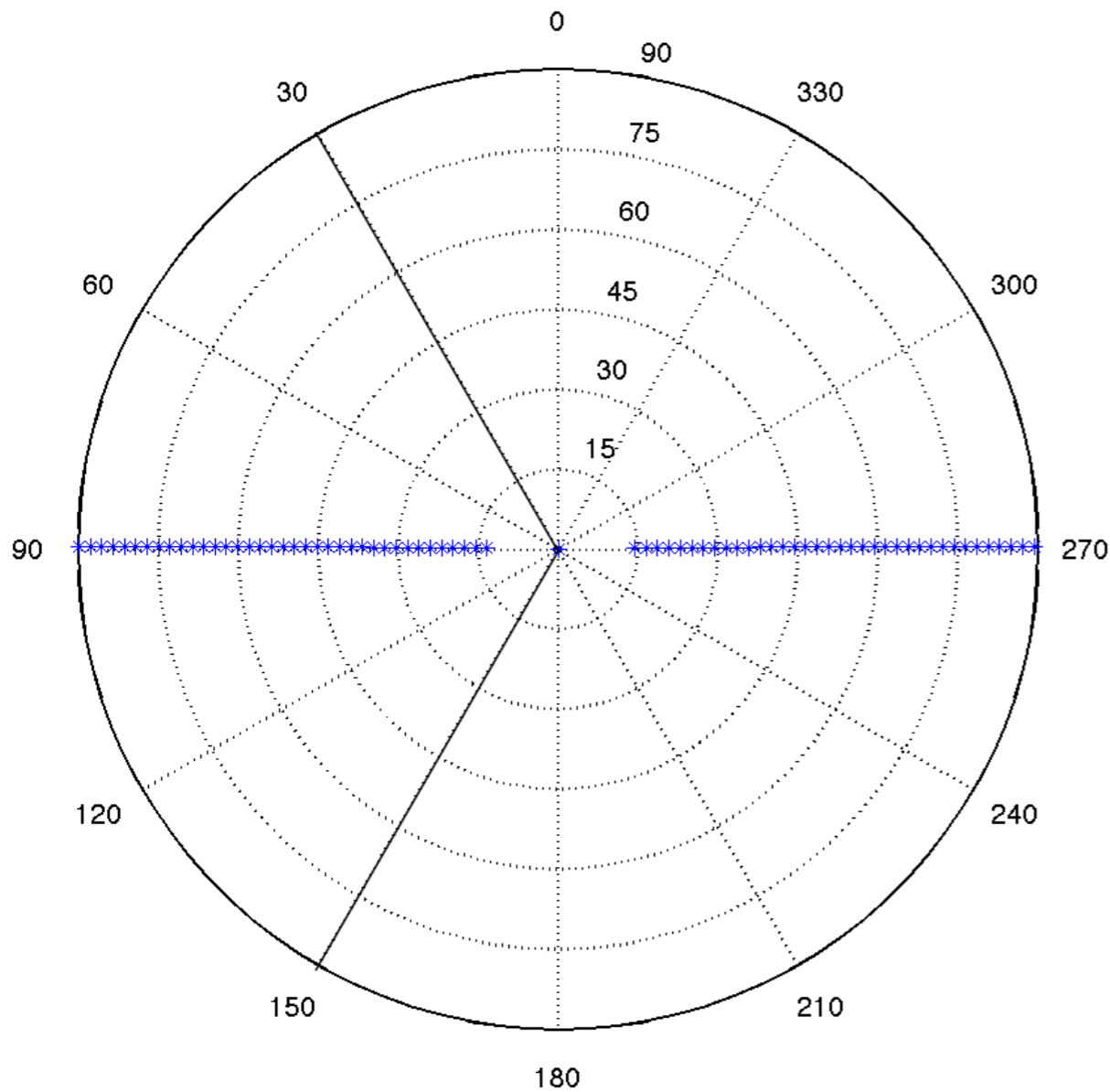
Fluctuation of the optical cross section during SUN ON and SUN OFF phases



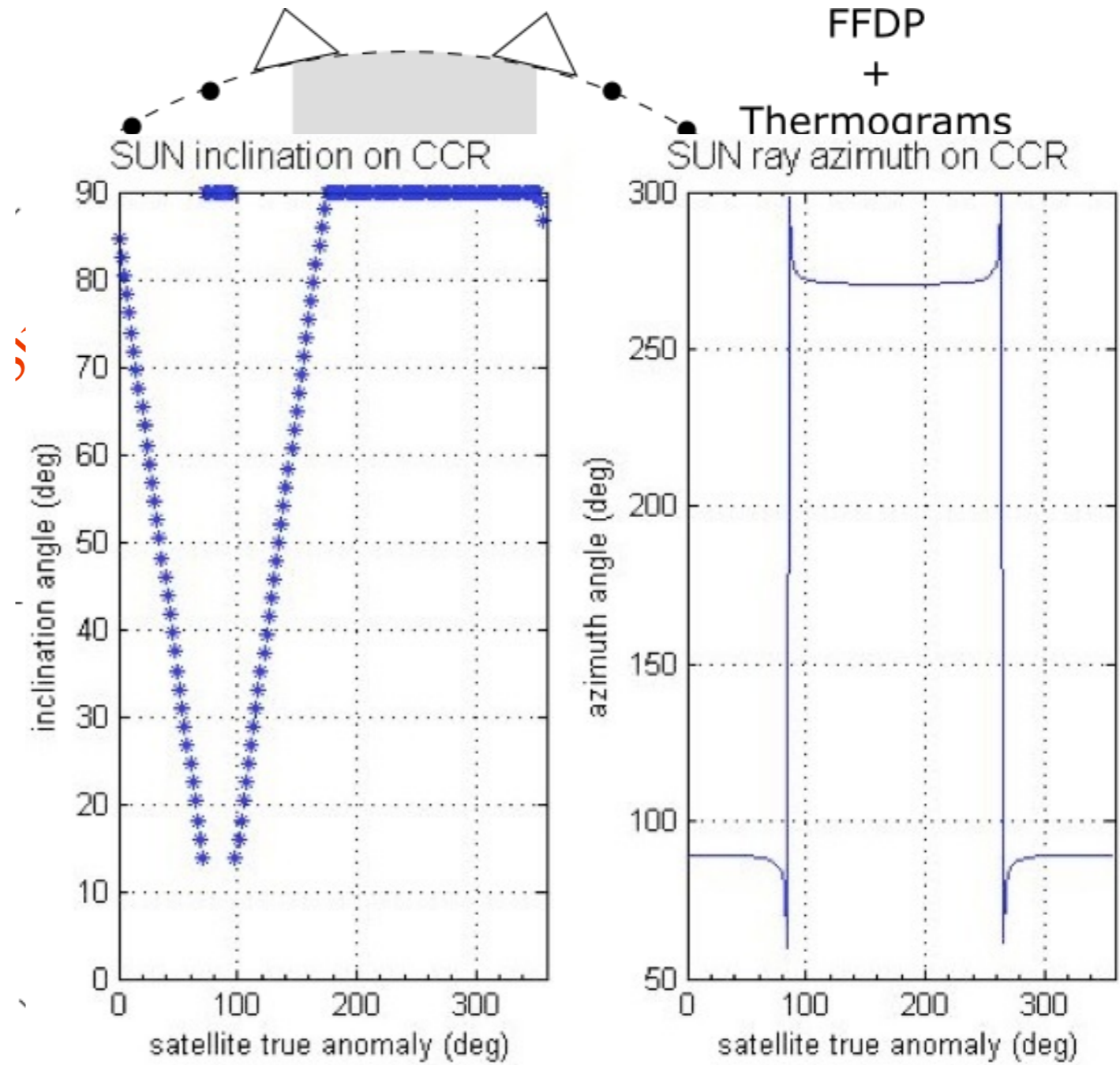
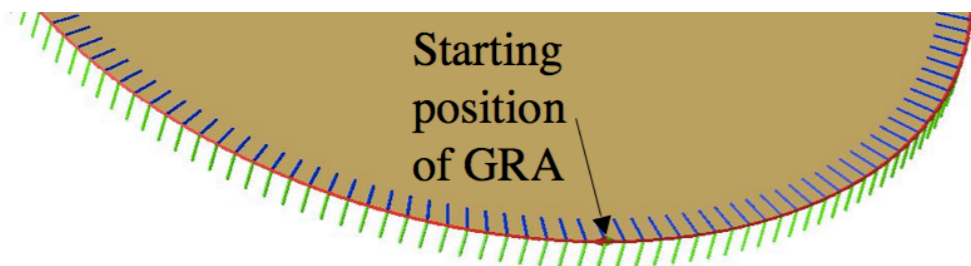
Optical Cross Section variation at the Galileo velocity aberration $\sim 24\mu\text{rad}$
GRA @ 300 K



SCF-testing of GNSS Critical half-Orbit (GCO)

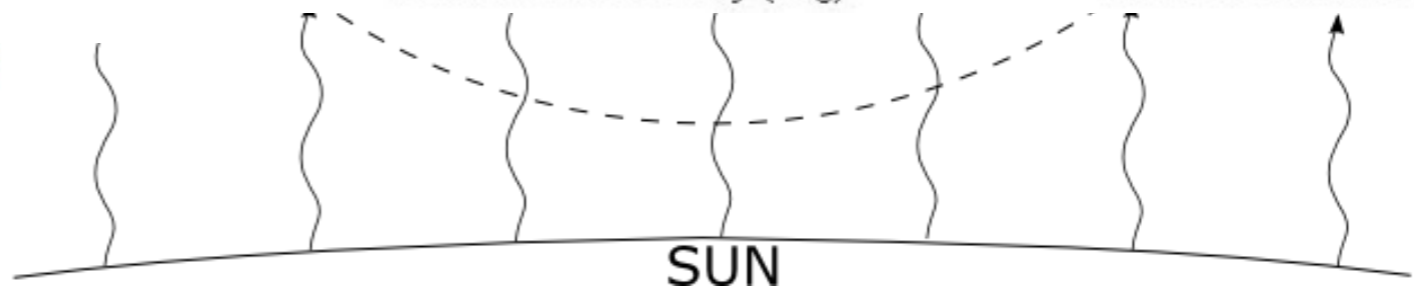


Inclination and azimuth on CCR GCO



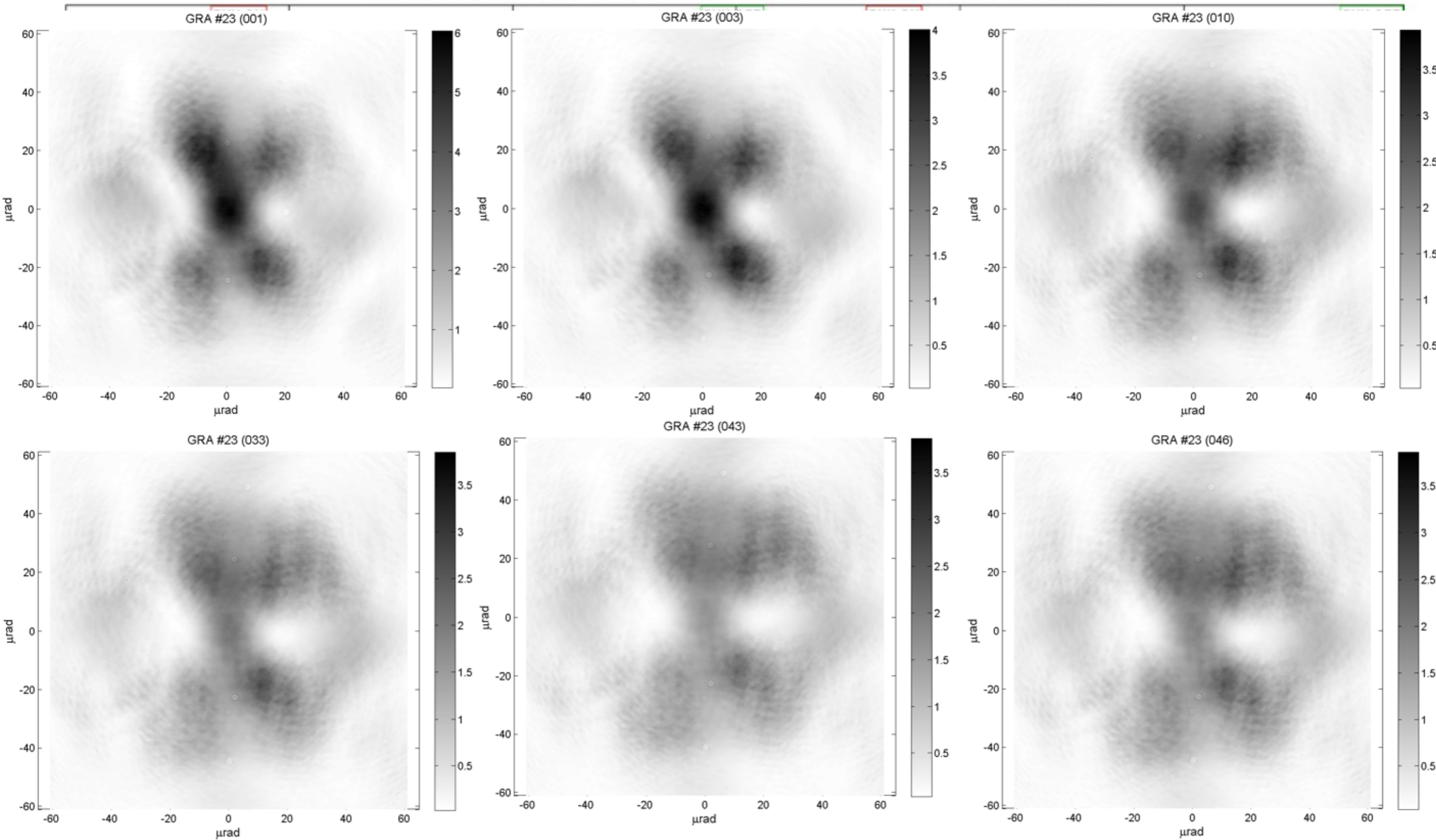
SUN inclination on CCR

FFDP +
Thermoarams
SUN ray azimuth on CCR

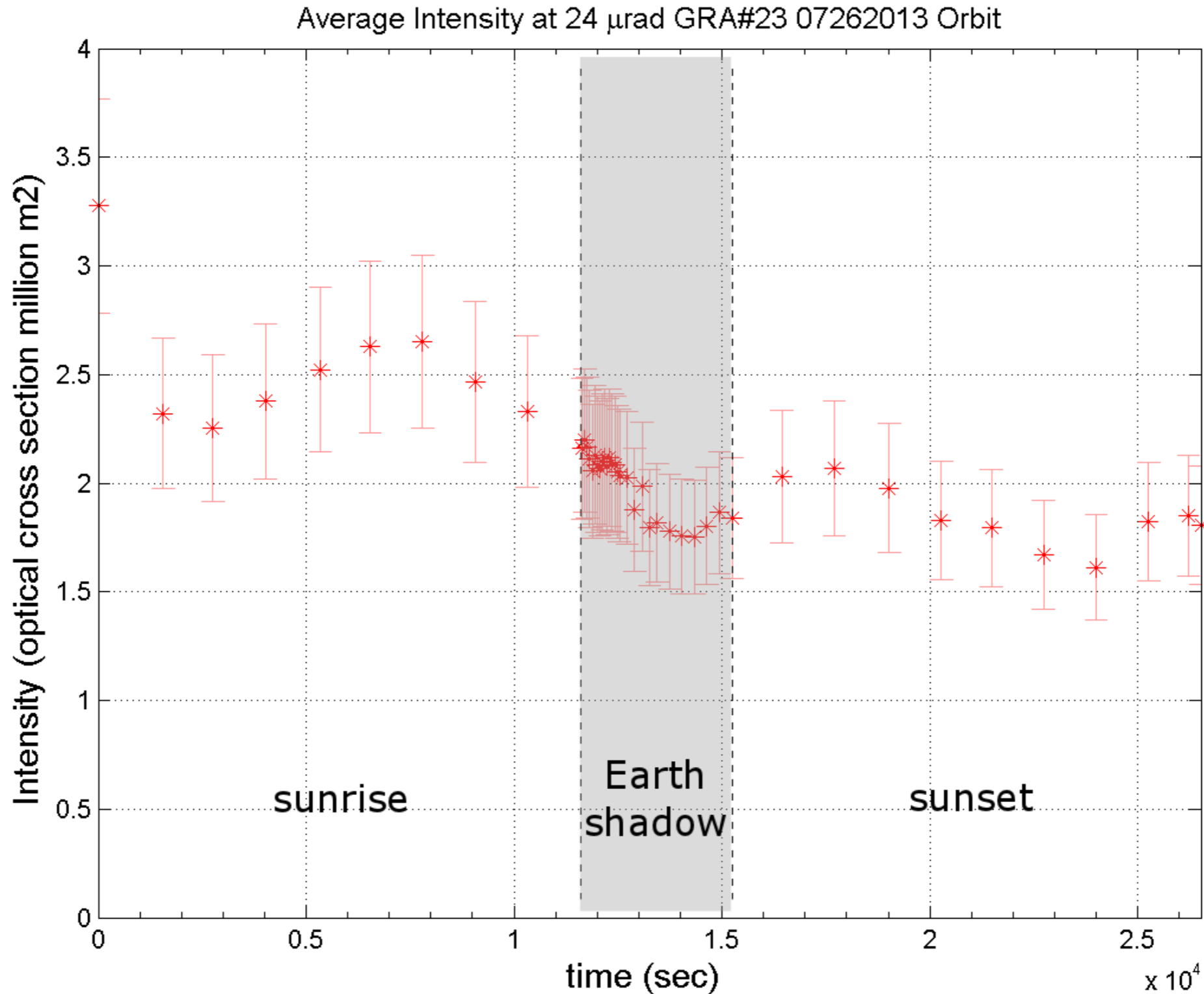


FFDP variation during Orbit

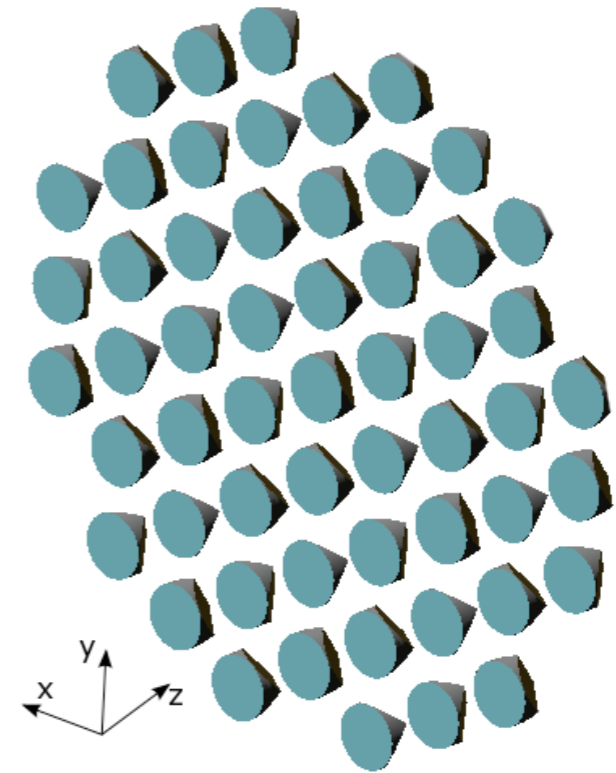
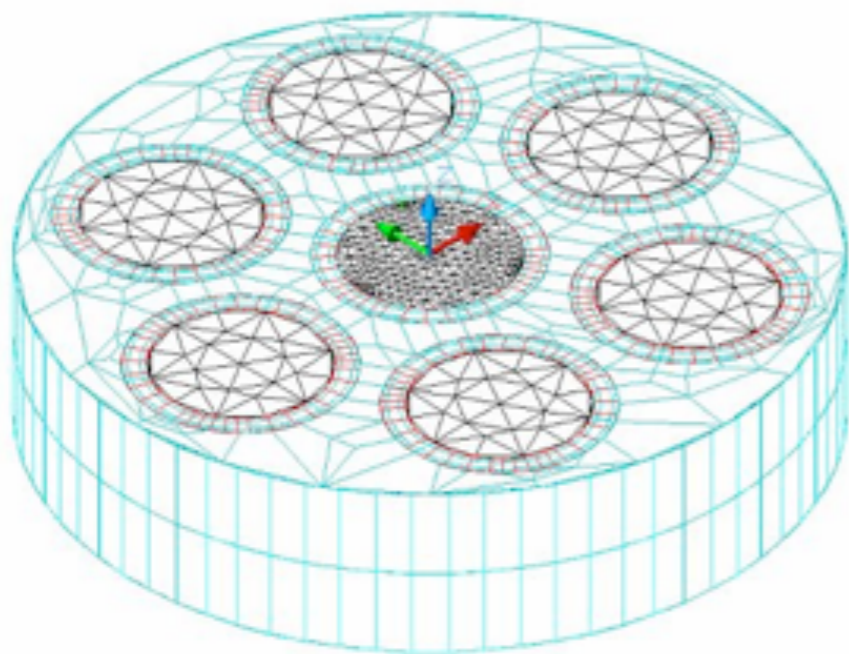
GRA - 23.07.2013 - Orbit - FFDP on CCR1 - Global Probes Temperatures - Prototype



Optical Cross Section variation at the Velocity Aberration of the satellite

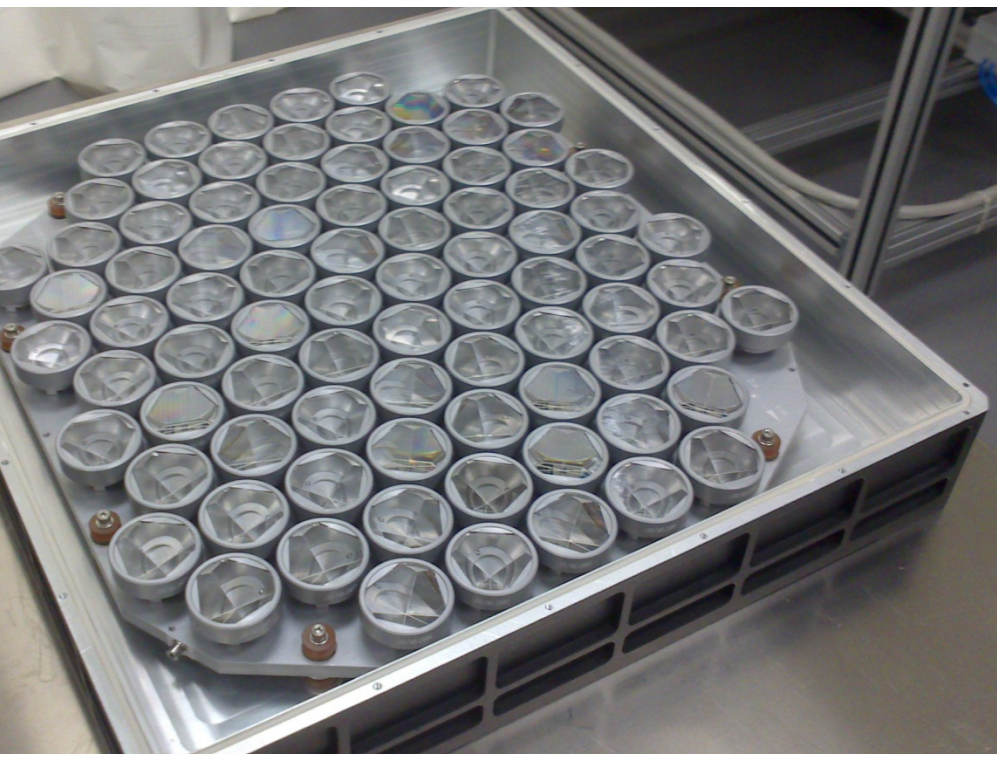
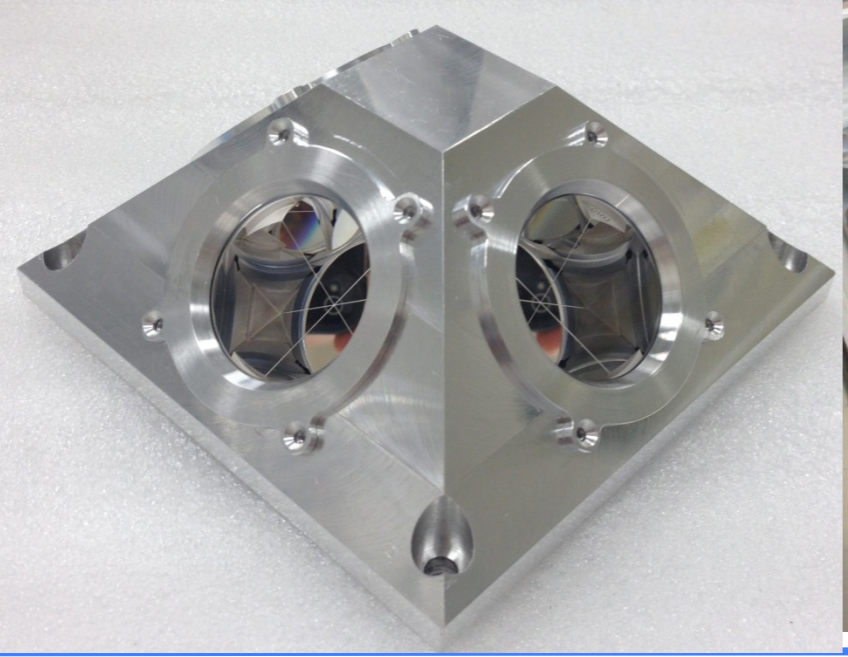
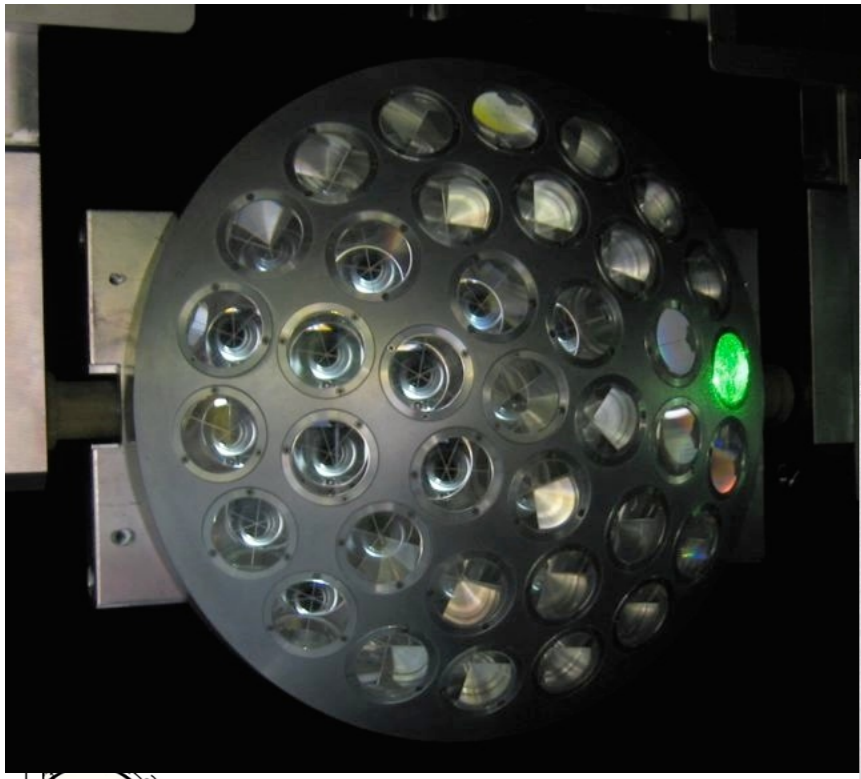
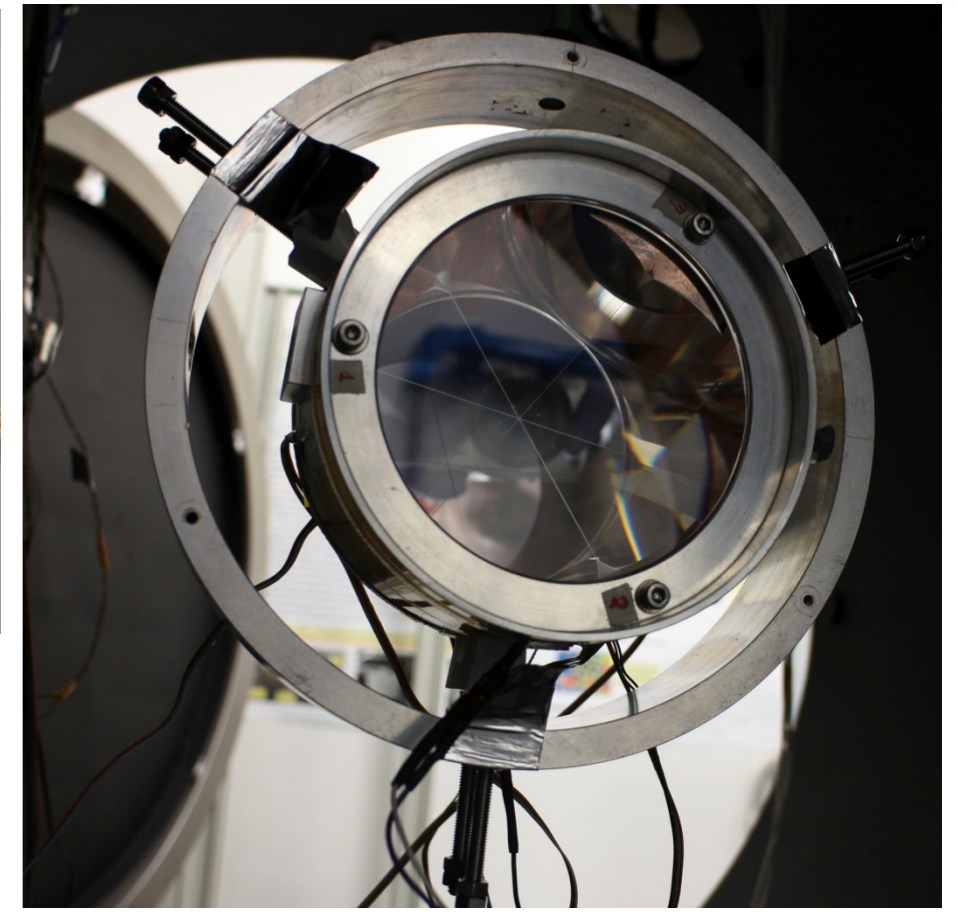
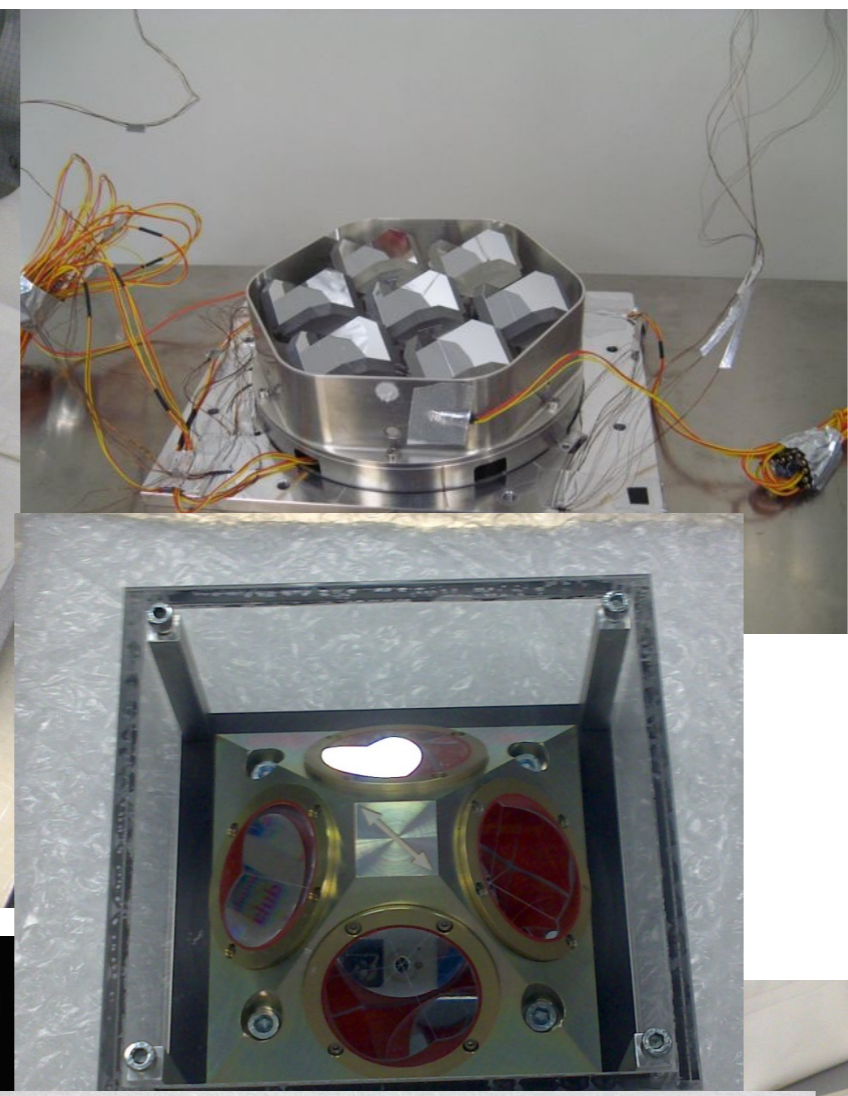
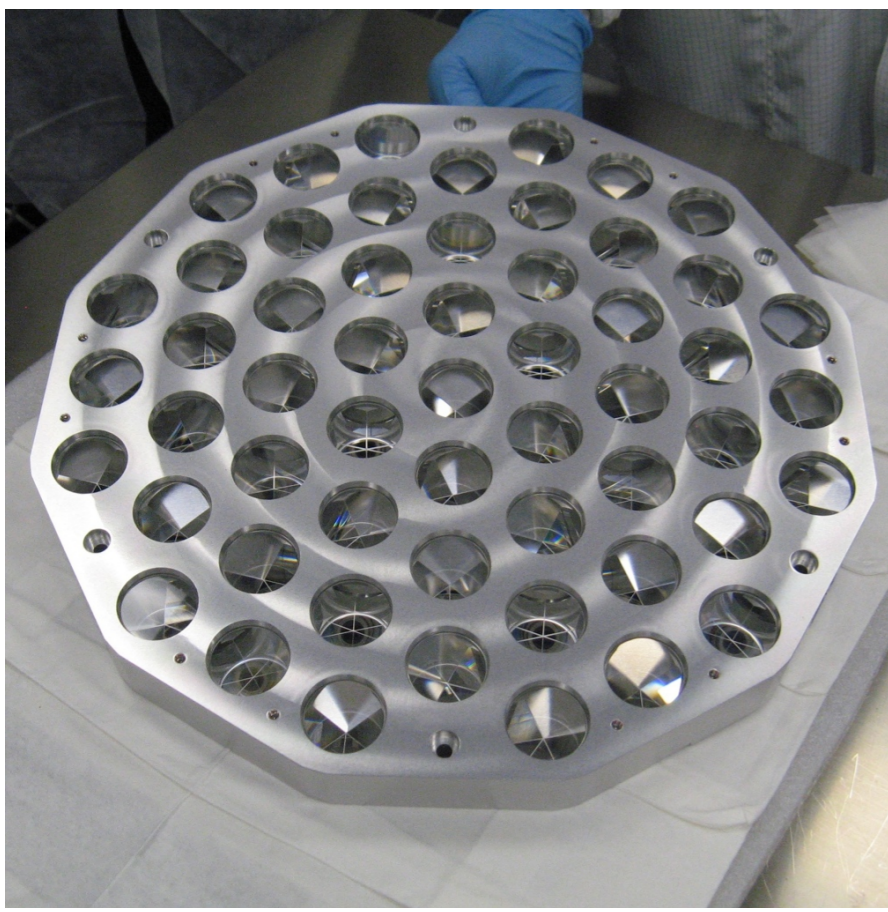


- Optical model of single retroreflectors and payloads
- Structural and thermal studies of retroreflector payloads
- Integrated thermal/optical simulations of retroreflector payloads in characteristic orbits.
- Fine tuning of simulated models with SCF_Lab measurements
- Mechanical drawing of designed payloads



- Collaboration with NASA to characterize **LAGEOS** satellite CCRs.
- Characterized retroreflectors for ASI satellite LARES.
- Characterization of a **Galileo-IOV retroreflector** for ESA.
- **ETRUSCO-2** project for design and characterization of a retroreflector payload for the GNSS (GRA)
- Design and characterization of a retroreflector payload for the **next generation Lunar Laser Ranging**, in collaboration with University of Maryland.
- SCF_Lab measurements of current GNSS retroreflector payloads (**IRNSS** and **Galileo**).
- Design of payloads for next generation **Earth observation satellites**.
- Design of payloads for **laser altimetry** on solar system planets and moons (Mars and Moon)
- Affiliation with **NASA SSERVI Institute** for Solar System Exploration

SCF_Lab projects



Thank you for your attention!

SCF_Lab website: <http://www.lnf.infn.it/esperimenti/etrusco/>