



ICHEP2012 ●
Melbourne

36th International Conference on High Energy Physics

4 – 11 July 2012

Melbourne Convention and Exhibition Centre

-- a cherry picking summary --





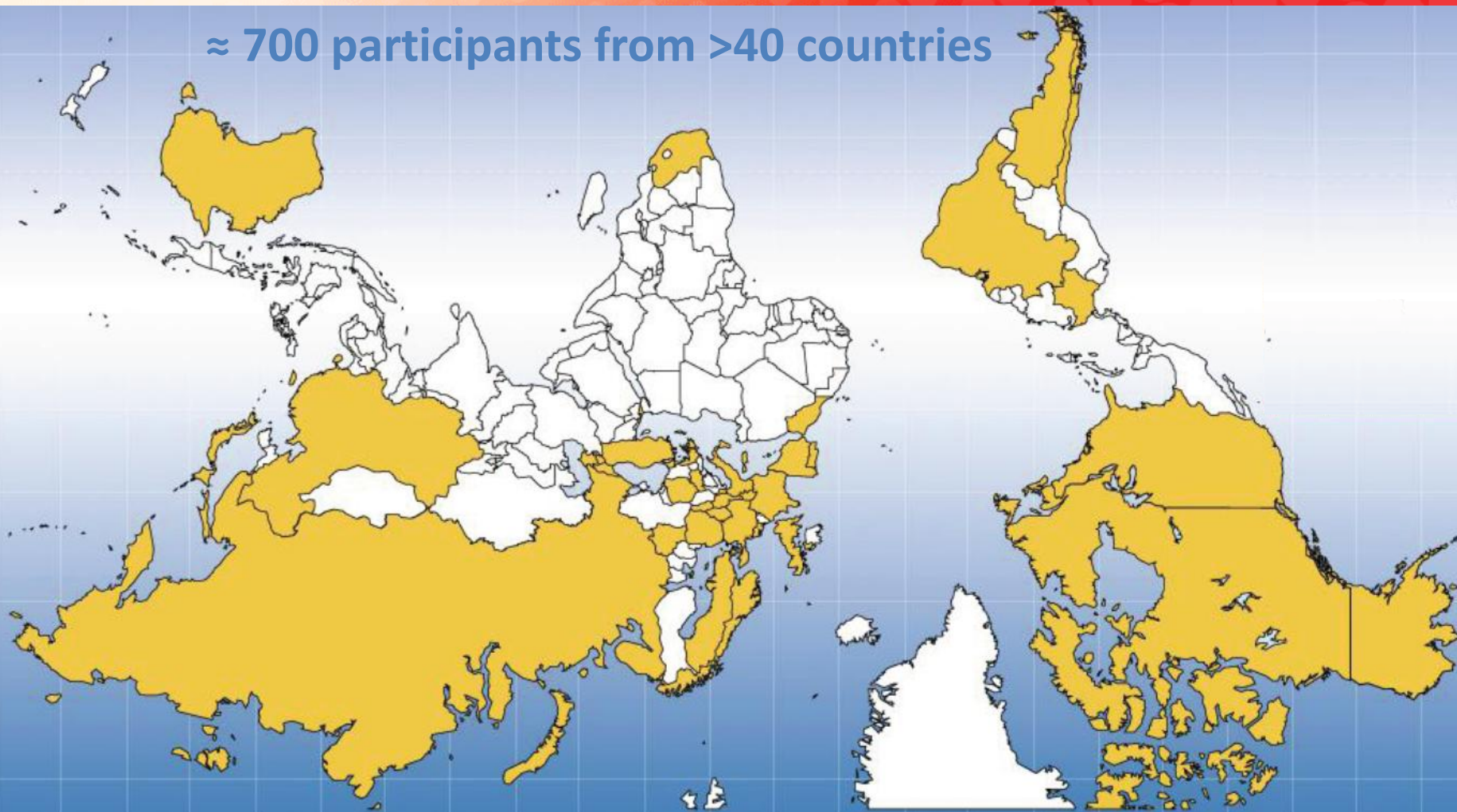
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≈ 700 participants from >40 countries





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➤ 3 days of parallel sessions:
grouped in 15 topics

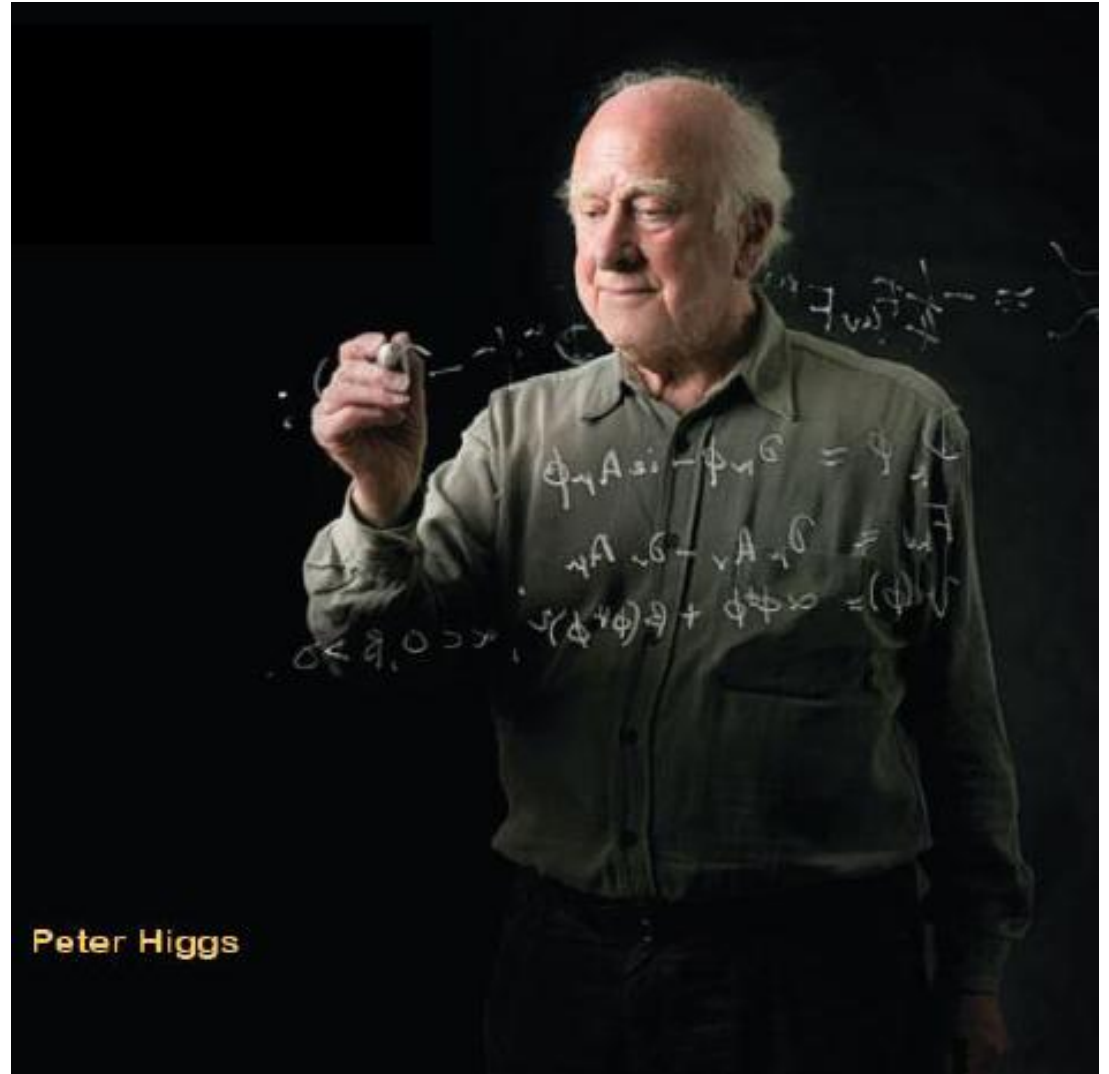
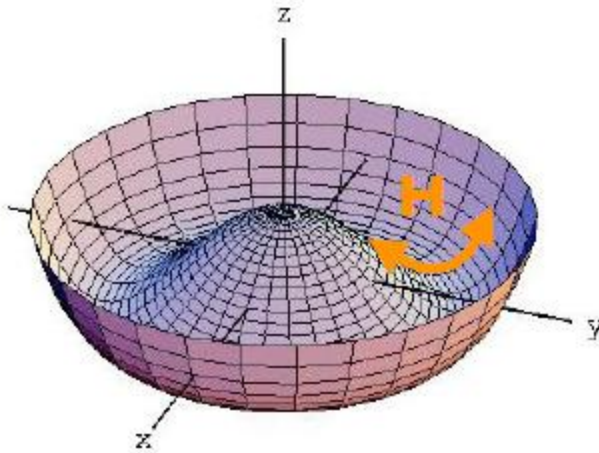
[www.ichep2012.com.au/Program]

➤ 3 days of planary talks

Parallel session convenors

Session	Convenors
1. The Standard Model and EW Symmetry Breaking- Higgs Searches	Sara Diglio Toyoko Orimoto Albert De Roeck Marumi Kado Eric James Pietro Slavich
2. Beyond the Standard Model - SUSY	Joanne Hewitt Iacopo Vivarelli Frederic Ronga
3. BSM - non- SUSY, Exotics	Kenji Hamano Shahram Rahatlou Christophe Grojean
4. Top Quark Physics	Pamela Ferrari Roberto Chierici Stefano Frixione
5. B Physics	Jure Zupan Phillip Urquijo Timothy Gershon
6. QCD, Jets, Parton Distributions	Voica Radescu Andreas Vogt
7. CP Violation, CKM, Rare Decays, Meson Spectroscopy	Chrisoph Schwannda Valerie Gibson
8. Neutrinos	Jelena Maric Kevin McFarland
9. Heavy Ion Collisions	Raimond Snellings Carlos Salgado
10. Lattice QCD	Ross Young Johnathon Flynn
11. Particle Astrophysics and Cosmology	Pat Scott Jodi Cooley Jason Kumar
12. Formal Theory Developments	Gary Shiu Emilian Dudas
13. Detectors and Computing for HEP	Ted Tiehui Liu Su Dong Sunanda Bannerjee Kenneth Bloom
14. Future Accelerators	Mark Boland
15. Education and Outreach	Steven Goldfarb David Barney

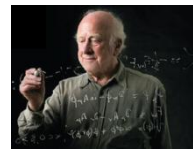
THE highlight





Press Office

THE highlight



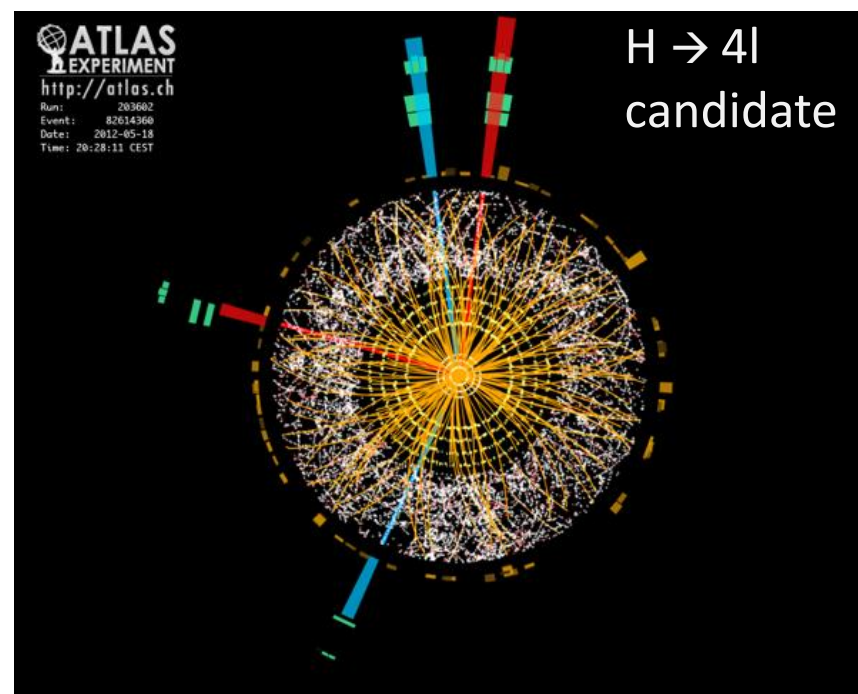
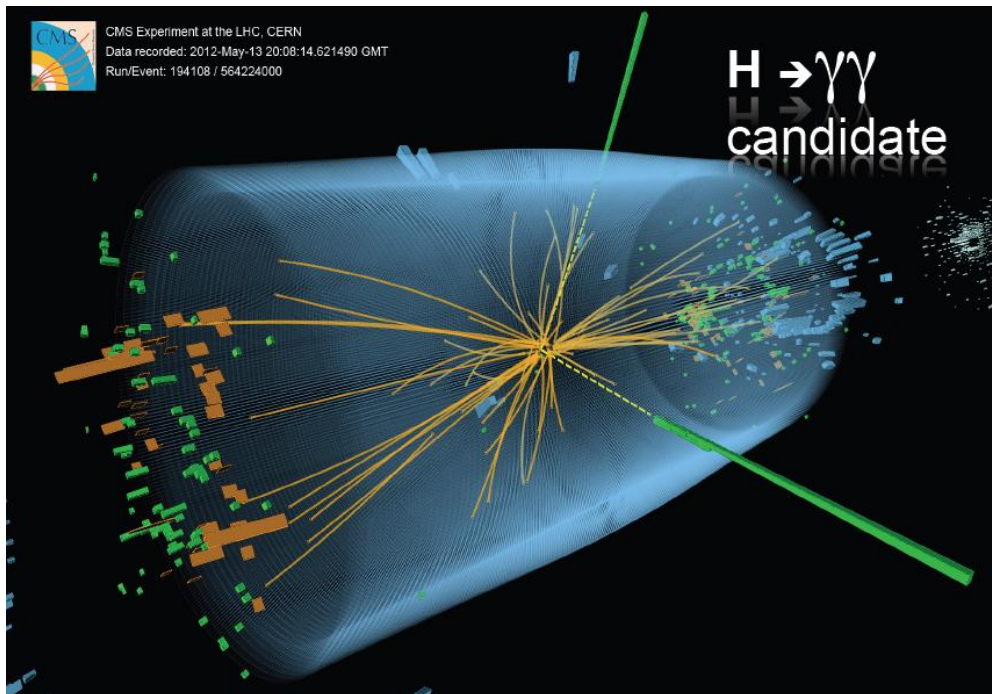
CERN experiments observe particle consistent with long-sought Higgs boson

PR17.12
04.07.2012

Geneva, 4 July 2012. At a seminar held at CERN¹ today as a curtain raiser to the year's major particle physics conference, ICHEP2012 in Melbourne, the ATLAS and CMS experiments presented their latest preliminary results in the search for the long sought Higgs particle. Both experiments observe a new particle in the mass region around 125-126 GeV.

<http://www.atlas.ch/news/2012/latest-results-from-higgs-search.html>

<http://cms.web.cern.ch/news/observation-new-particle-mass-125-gev>



ICHEP topics

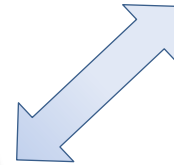
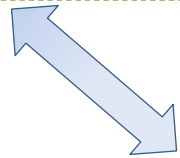
*HERA
legacy*

base line: QCD

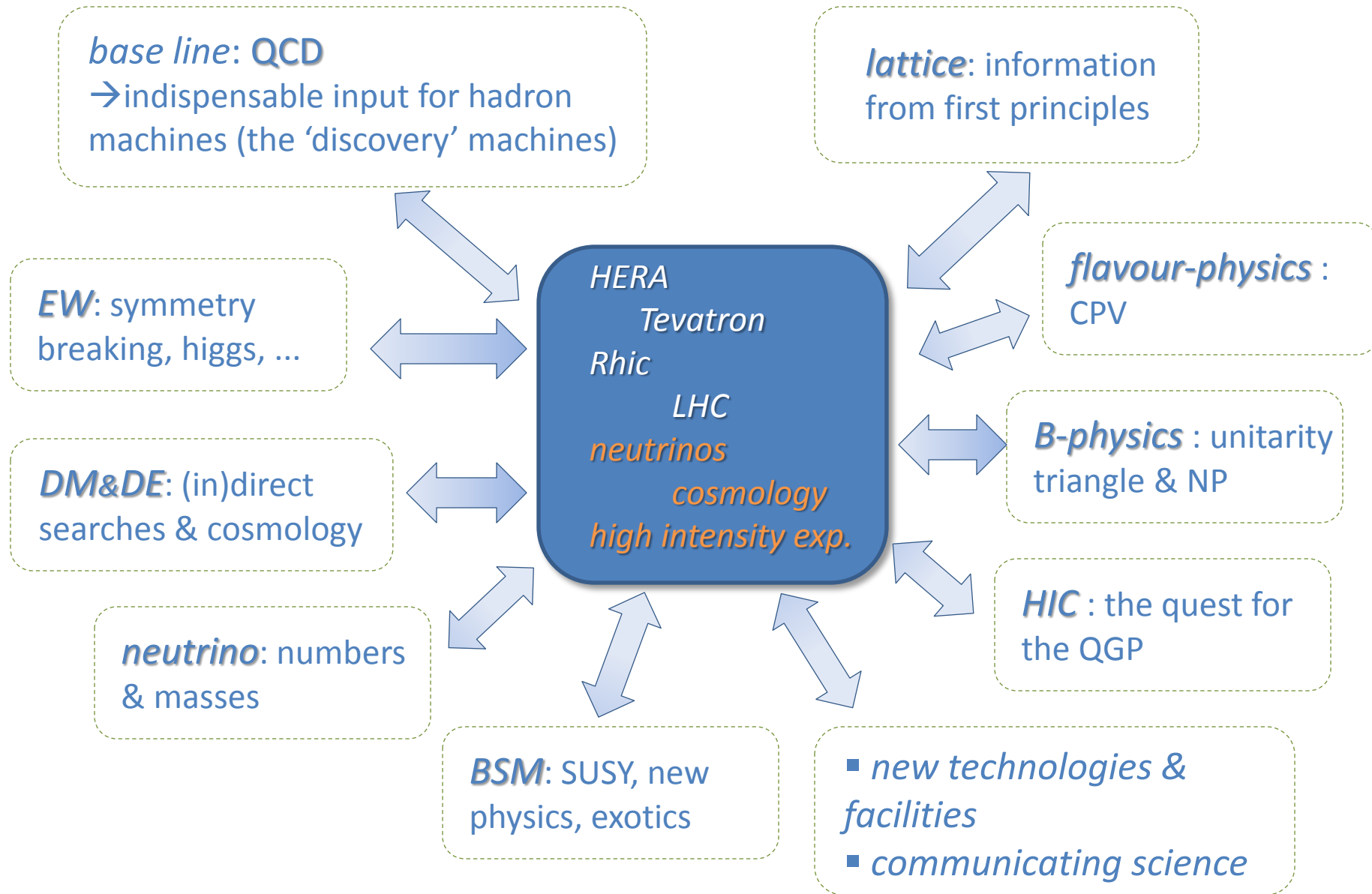
→ indispensable input for hadron machines (the 'discovery' machines)

lattice: information from first principles

*HERA
Tevatron
Rhic
LHC
neutrinos
cosmology
high intensity exp.*

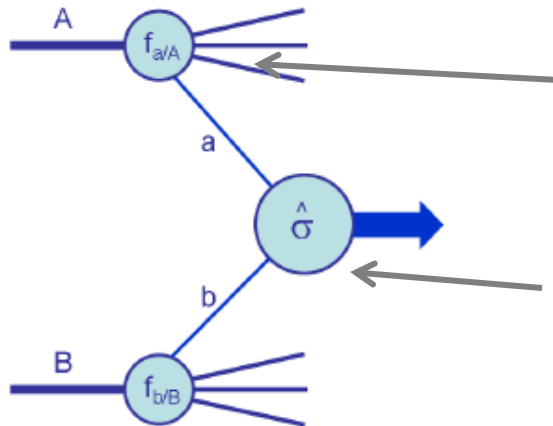


ICHEP topics



pQCD

-- the baseline --



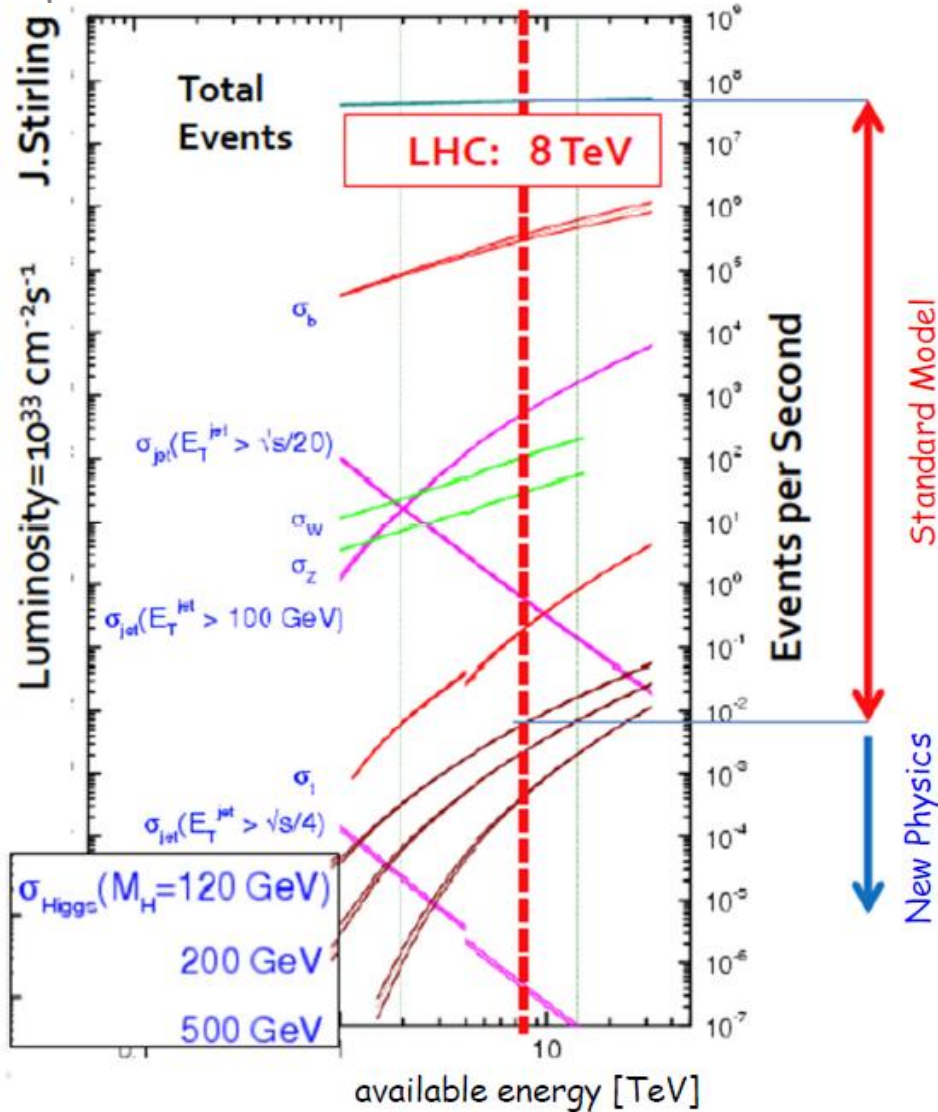
determination of parton distribution functions

calculation of hard scattering matrix elements

expansion in strong coupling α_s

QCD – why do we still care (perhaps more than ever)

production rates at hadron colliders



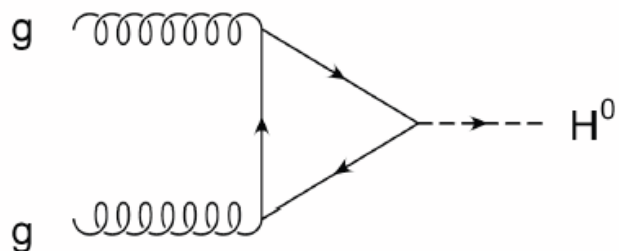
hadron colliders inevitably
have to deal with QCD

discovering the Higgs or
some New Physics requires
a sophisticated **quantitative**
understanding of QCD



Higgs and QCD

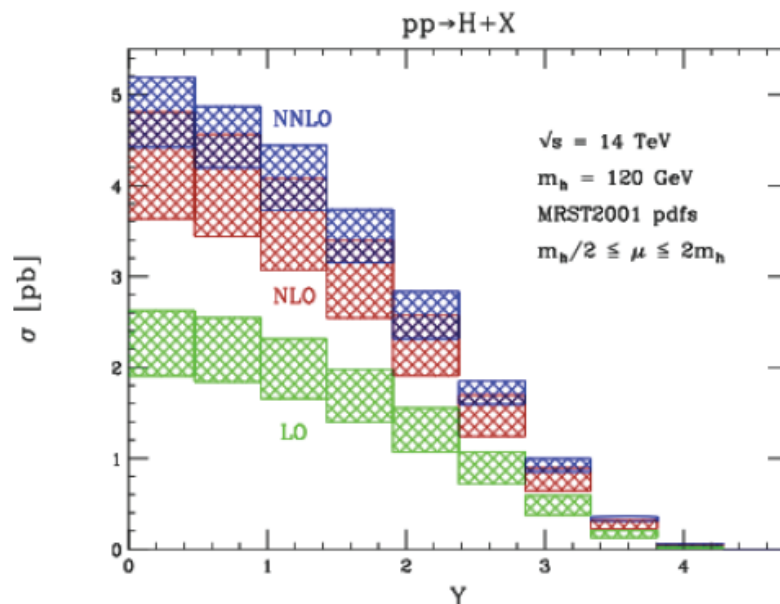
- ◆ Despite EW role a QCD problem



- ◆ Search strategies require good understanding of QCD issues
 - ◆ boosted jets and substructure
 - ◆ jet vetoes and resummation techniques
- ◆ Much study: “Handbook of LHC Higgs cross sections”

Challenging perturbative expansion, NNLO required

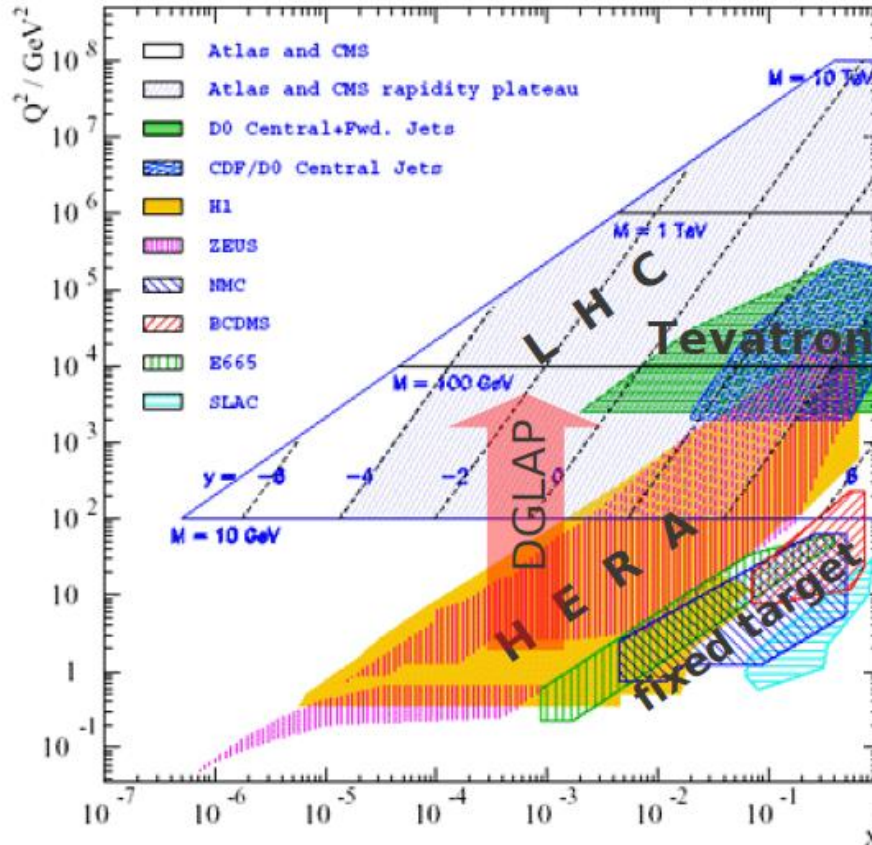
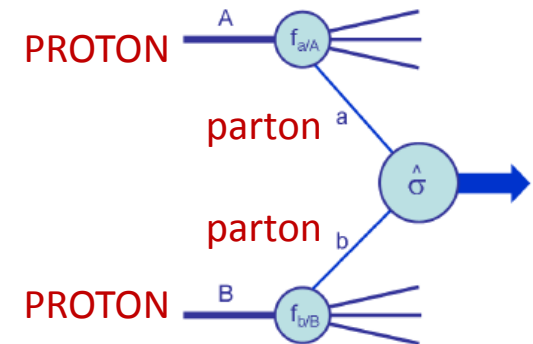
Anastasiou, Melnikov, Petriello



Dittmaier, Mariotti, Passarino, Tanaka

Parton distribution functions

- ◆ Parton content of the proton: life-blood of hadron collider physics
- ◆ Systematic exploration of proton structure at NNLO



reaching for precision

$$P_{ns}^{(0)}(x) = C_F(2p_{qq}(x) + 3\delta(1-x))$$

$$P_{ps}^{(0)}(x) = 0$$

$$P_{qs}^{(0)}(x) = 2n_f p_{qg}(x)$$

$$P_{gq}^{(0)}(x) = 2C_F p_{gq}(x)$$

$$P_{gg}^{(0)}(x) = C_A(4p_{gg}(x) + \frac{11}{3}\delta(1-x)) - \frac{2}{3}n_f\delta(1-x)$$

LO: 1973

Curci, Furmanski, Petronzio;
Floratos et al., ...

$$P_{ns}^{(1)+}(x) = 4C_A C_F \left(p_{qq}(x) \left[\frac{67}{18} - \zeta_2 + \frac{11}{6}H_0 + H_{0,0} \right] + p_{qq}(-x) \left[\zeta_2 + 2H_{-1,0} - H_{0,0} \right] \right. \\ \left. + \frac{14}{3}(1-x) + \delta(1-x) \left[\frac{17}{24} + \frac{11}{3}\zeta_2 - 3\zeta_3 \right] \right) - 4C_F n_f \left(p_{qg}(x) \left[\frac{5}{9} + \frac{1}{3}H_0 \right] + \frac{2}{3}(1-x) \right. \\ \left. + \delta(1-x) \left[\frac{1}{12} + \frac{2}{3}\zeta_2 \right] \right) + 4C_F^2 \left(2p_{qg}(x) \left[H_{1,0} - \frac{3}{4}H_0 + H_2 \right] - 2p_{qg}(-x) \left[\zeta_2 + 2H_{-1,0} \right. \right. \\ \left. \left. - H_{0,0} \right] - (1-x) \left[1 - \frac{3}{2}H_0 \right] - H_0 - (1+x)H_{0,0} + \delta(1-x) \left[\frac{3}{8} - 3\zeta_2 + 6\zeta_3 \right] \right)$$

$$P_{ns}^{(1)-}(x) = P_{ns}^{(1)+}(x) + 16C_F \left(C_F - \frac{C_A}{2} \right) \left(p_{qq}(-x) \left[\zeta_2 + 2H_{-1,0} - H_{0,0} \right] - 2(1-x) \right. \\ \left. - (1+x)H_0 \right)$$

$$P_{ps}^{(1)}(x) = 4C_F n_f \left(\frac{20}{9} \frac{1}{x} - 2 + 6x - 4H_0 + x^2 \left[\frac{8}{3}H_0 - \frac{56}{9} \right] + (1+x) \left[5H_0 - 2H_{0,0} \right] \right)$$

$$P_{qs}^{(1)}(x) = 4C_A n_f \left(\frac{20}{9} \frac{1}{x} - 2 + 25x - 2p_{qg}(-x)H_{-1,0} - 2p_{qg}(x)H_{1,1} + x^2 \left[\frac{44}{3}H_0 - \frac{218}{9} \right] \right. \\ \left. + 4(1-x) \left[H_{0,0} - 2H_0 + xH_1 \right] - 4\zeta_2 x - 6H_{0,0} + 9H_0 \right) + 4C_F n_f \left(2p_{qg}(x) \left[H_{1,0} + H_{1,1} + H_2 \right. \right. \\ \left. \left. - \zeta_2 \right] + 4x^2 \left[H_0 + H_{0,0} + \frac{5}{2} \right] + 2(1-x) \left[H_0 + H_{0,0} - 2xH_1 + \frac{29}{4} \right] - \frac{15}{2} - H_{0,0} - \frac{1}{2}H_0 \right)$$

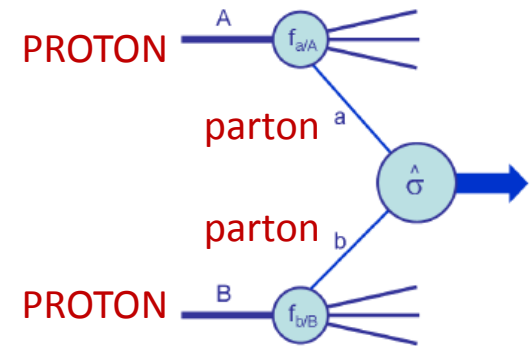
$$P_{gq}^{(1)}(x) = 4C_A C_F \left(\frac{1}{x} + 2p_{gq}(x) \left[H_{1,0} + H_{1,1} + H_2 - \frac{11}{6}H_1 \right] - x^2 \left[\frac{8}{3}H_0 - \frac{44}{9} \right] + 4\zeta_2 - 2 \right. \\ \left. - 7H_0 + 2H_{0,0} - 2H_1 x + (1+x) \left[2H_{0,0} - 5H_0 + \frac{37}{9} \right] - 2p_{gq}(-x)H_{-1,0} \right) - 4C_F n_f \left(\frac{2}{3}x \right. \\ \left. - p_{gq}(x) \left[\frac{2}{3}H_1 - \frac{10}{9} \right] \right) + 4C_F^2 \left(p_{gq}(x) \left[3H_1 - 2H_{1,1} \right] + (1+x) \left[H_{0,0} - \frac{7}{2} + \frac{7}{2}H_0 \right] - 3H_{0,0} \right. \\ \left. + 1 - \frac{3}{2}H_0 + 2H_1 x \right)$$

$$P_{gg}^{(1)}(x) = 4C_A n_f \left(1-x - \frac{10}{9}p_{gq}(x) - \frac{13}{9} \left(\frac{1}{x} - x^2 \right) - \frac{2}{3}(1+x)H_0 - \frac{2}{3}\delta(1-x) \right) + 4C_A^2 \left(27 \right. \\ \left. + (1+x) \left[\frac{11}{3}H_0 + 8H_{0,0} - \frac{27}{2} \right] + 2p_{gq}(-x) \left[H_{0,0} - 2H_{-1,0} - \zeta_2 \right] - \frac{67}{9} \left(\frac{1}{x} - x^2 \right) - 12H_0 \right. \\ \left. - \frac{44}{3}x^2 H_0 + 2p_{gq}(x) \left[\frac{67}{18} - \zeta_2 + H_{0,0} + 2H_{1,0} + 2H_2 \right] + \delta(1-x) \left[\frac{8}{3} + 3\zeta_3 \right] \right) + 4C_F n_f \left(2H_0 \right. \\ \left. + \frac{2}{3} \frac{1}{x} + \frac{10}{3}x^2 - 12 + (1+x) \left[4 - 5H_0 - 2H_{0,0} \right] - \frac{1}{2}\delta(1-x) \right)$$

NLO: 1980

Parton distribution functions

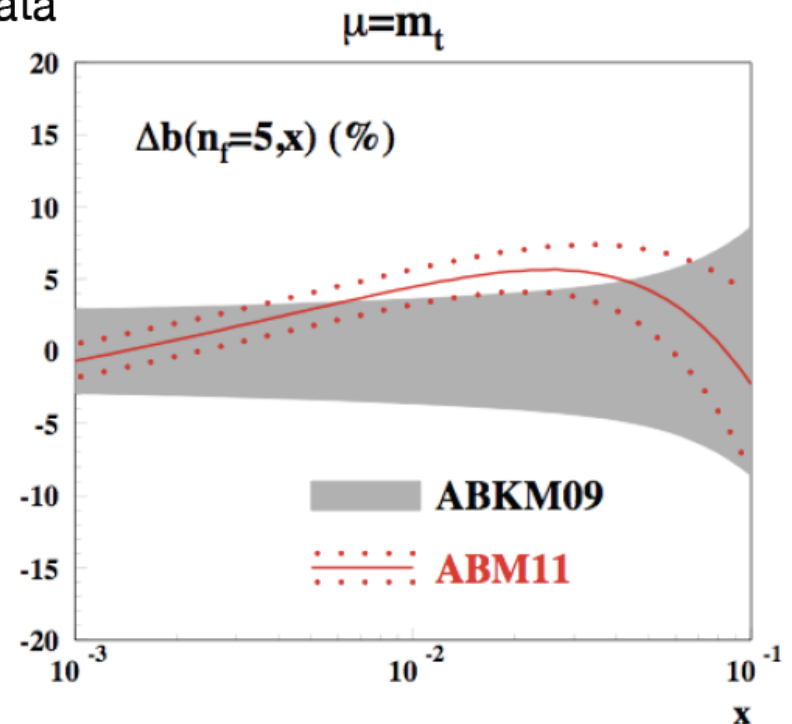
- ◆ Parton content of the proton: life-blood of hadron collider physics
- ◆ Systematic exploration of proton structure at NNLO
- ◆ New this year: ABM11 Alekhin, Blumlein, Moch



- ◆ fit to DIS and fixed-target Drell-Yan data
 - ◆ improved treatment of heavy quarks in DIS, running $\overline{\text{MS}}$ mass
- Alekhin, Moch

b-quark pdf uncertainty

- ◆ much-reduced in ABM11
- ◆ impact on many LHC cross sections, e.g. single top, charged Higgs



Higgs production at 125 GeV

<https://twiki.cern.ch/twiki/bin/view/LHCPhysics/CrossSections>

- ◆ Model testing requires assessment of theoretical uncertainties
 - ◆ uncertainties from **scale variation** and **PDF+strong coupling**

σ (8 TeV)

uncertainty

NNLL QCD
+NLO EW

$gg \rightarrow H$

19.5 pb

14.7%

VBF

1.56 pb

2.9%

NNLO QCD
+NLO EW

WH

0.70 pb

3.9%

ZH

0.39 pb

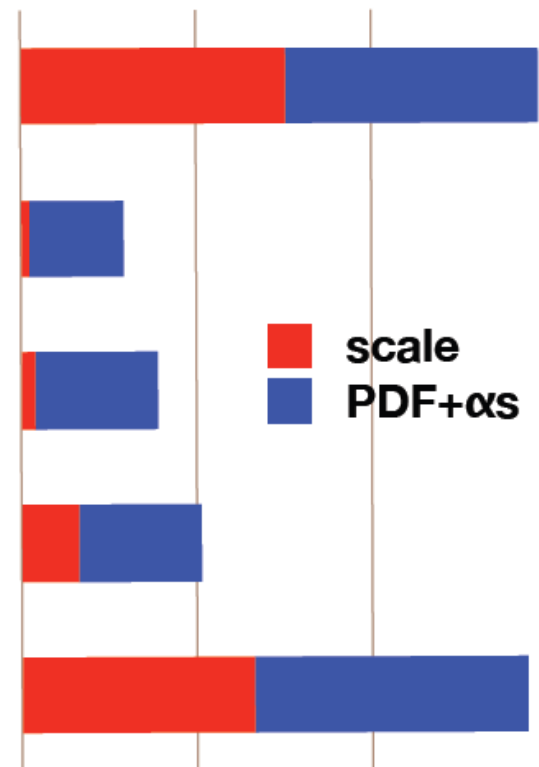
5.1%

NLO QCD

ttH

0.13 pb

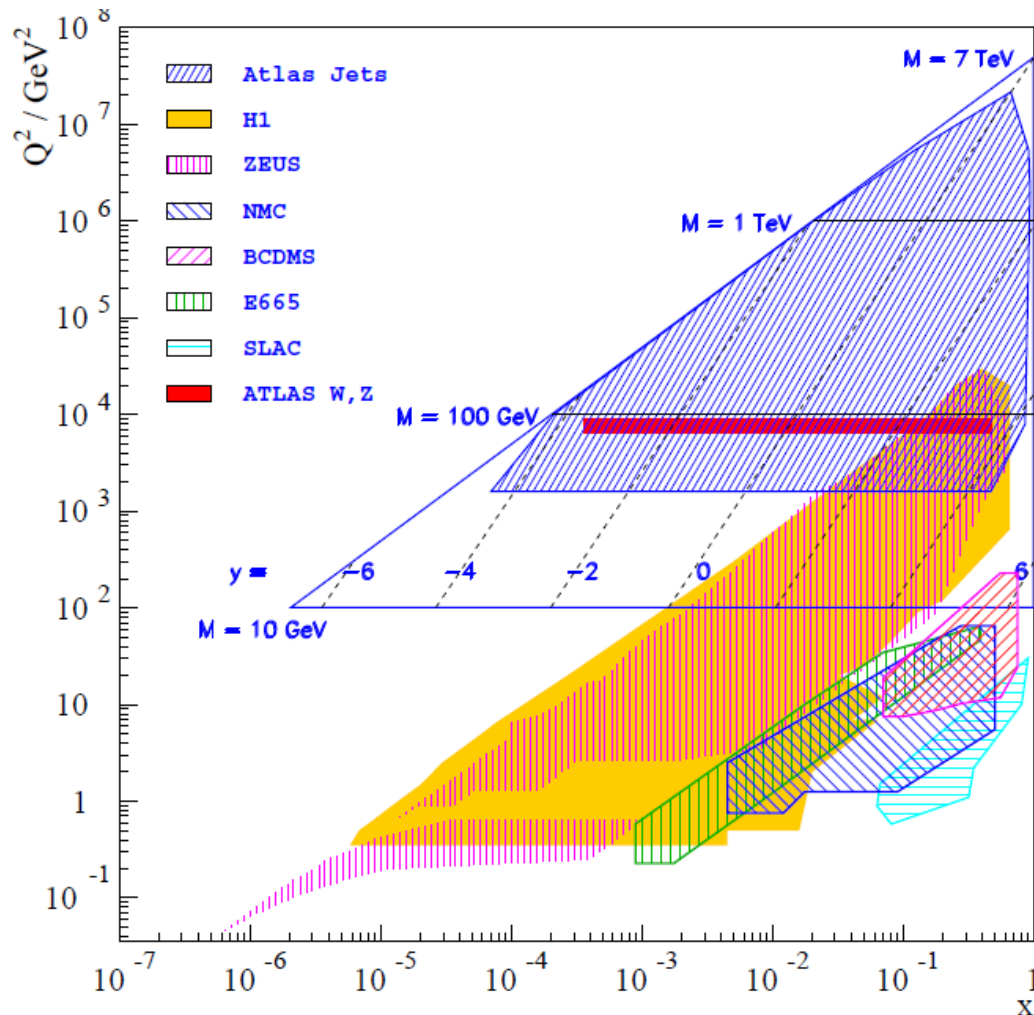
14.4%



cross sections @LHC

$$\sim x_1 f_1(x_1, \mu) x_2 f_2(x_2, \mu) \hat{\sigma}(x_1, x_2, \mu)$$

→ PDFs have on average > 1.3 citation per LHC paper (estimated based on study of G. Salam, La Thuile 2012)



- LHC x-range well covered by HERA data
- QCD: *evolution*
- strange distribution: constraints from neutrino data (SIDIS, LHC)
- gluon distribution: from *evolution* + constraints from jet prod. F_L from HERA

PDF sets

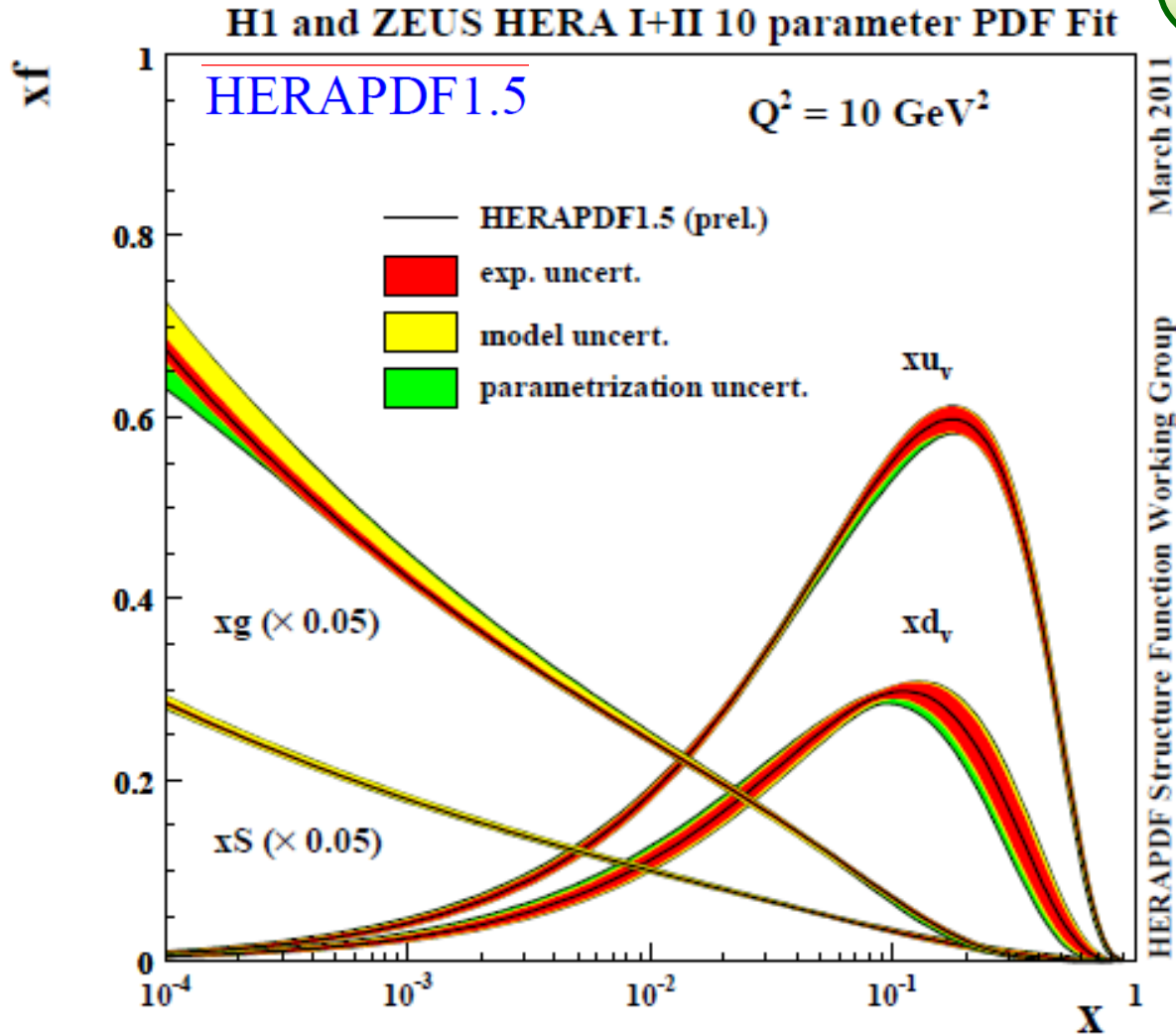
	MSTW08	CTEQ6.6/CT10	NNPDF2.1/2.3	HERAPDF1.0/1.5	ABKM09/ABM11	GJR08/JR09
Evolution Order	LO NLO NNLO	LO NLO NNLO(prel)	LO NLO NNLO	— NLO NNLO	— NLO NNLO	— NLO NNLO
HF Scheme	RT-GMVF	ACOT-GMVF	FONLL-GMVF	RT-GMVF (*)	BMSN-FFNS	FFNS
α_S NLO	0.120	0.118(f)	0.1191(b)	0.1176(f)	0.118	0.1135
α_S NNLO	0.1171	0.118(f)	0.1174(b)	0.1176(f)	0.1135	0.1124
HERA DIS	not up-to-date	+	+	+/prelim.	partial	+
Fixed target DIS	+	+	+	-	+	+
DY	+	+	+	-	+	+
Tevatron W,Z	some	some	some	-	some	some
Tevatron jets	some	+	+	-	some	some
LHC	-	-	<i>W, Z+jets (NNPDF2.3)</i>	-	-	-

NNPDFs [Ball et.al.]

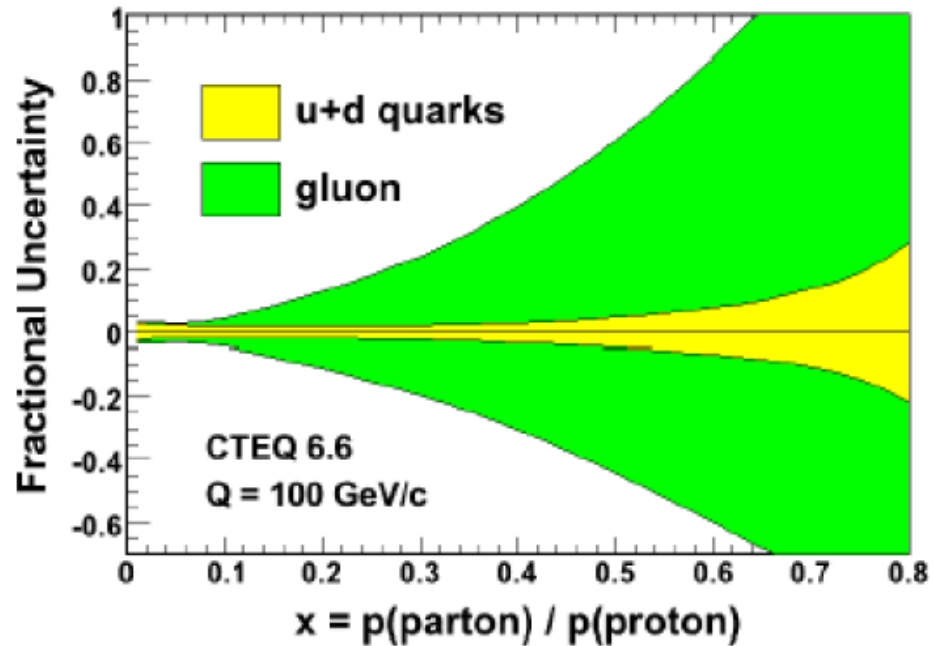
free from uncertainty due to
underlying parametrisation

PDF sets: just one example

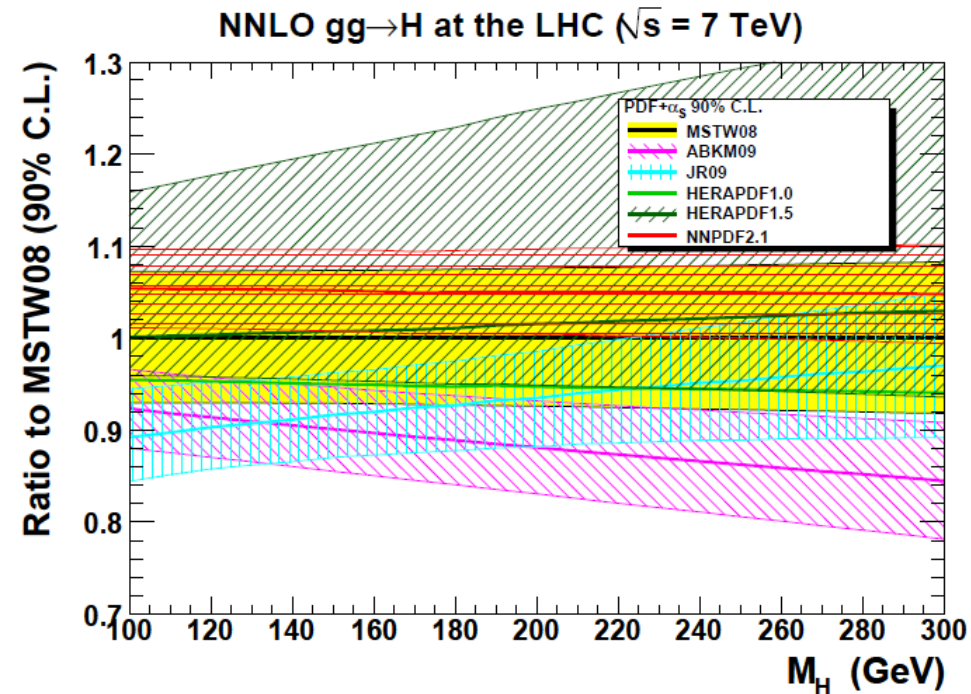
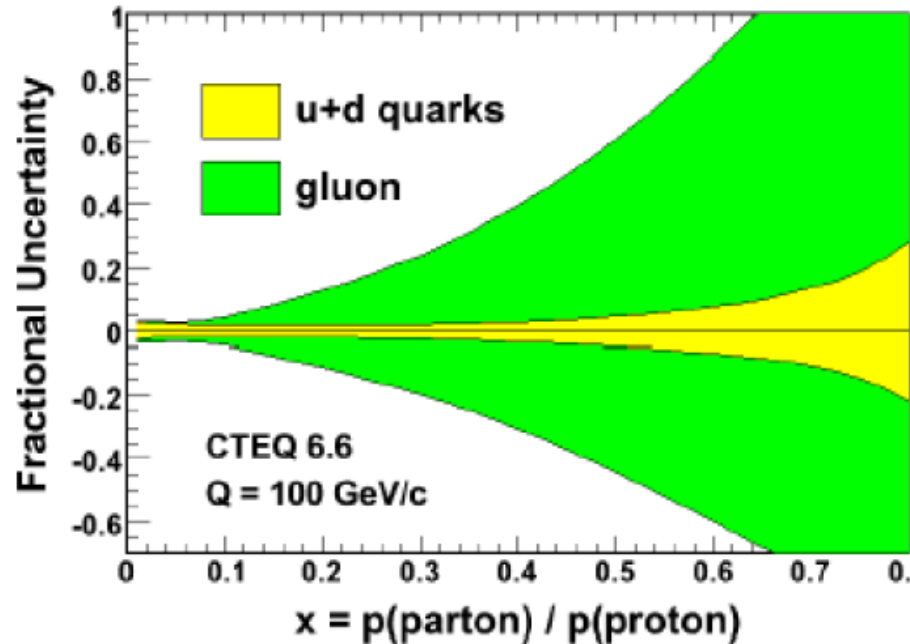
HERA legacy



PDF sets: typical uncertainties

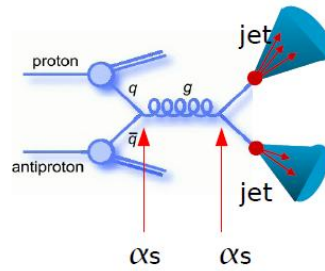


PDF sets: typical uncertainties & cross section estimates



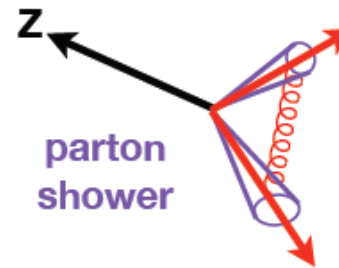
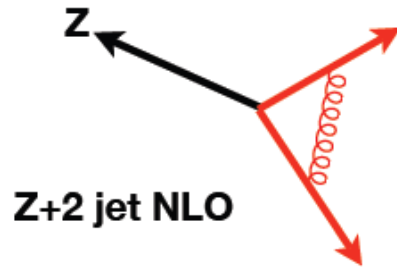
typically 10% uncertainty due to choice of PDF set

pQCD: jets



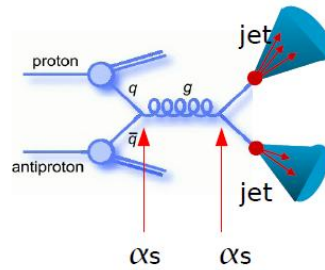
[John Campbell/Dmitry Bandurin]

◆ Refining “well-known” calculations, e.g. NLO QCD



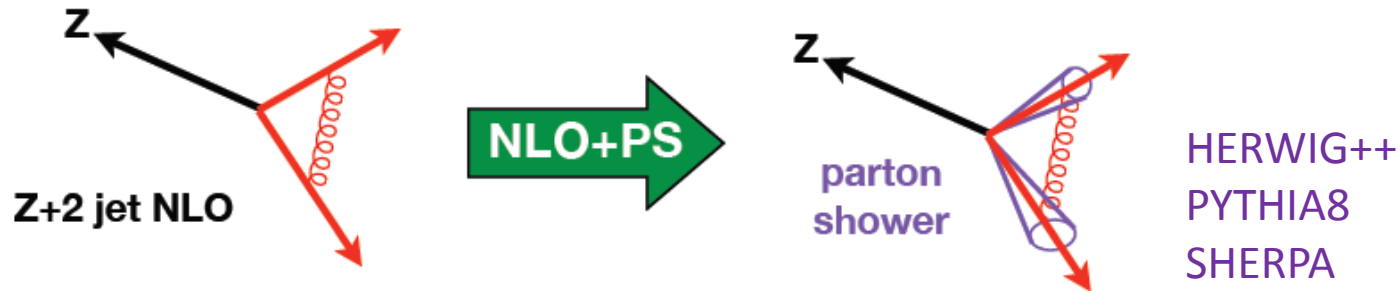
HERWIG++
PYTHIA8
SHERPA

pQCD: jets



[John Campbell/Dmitry Bandurin]

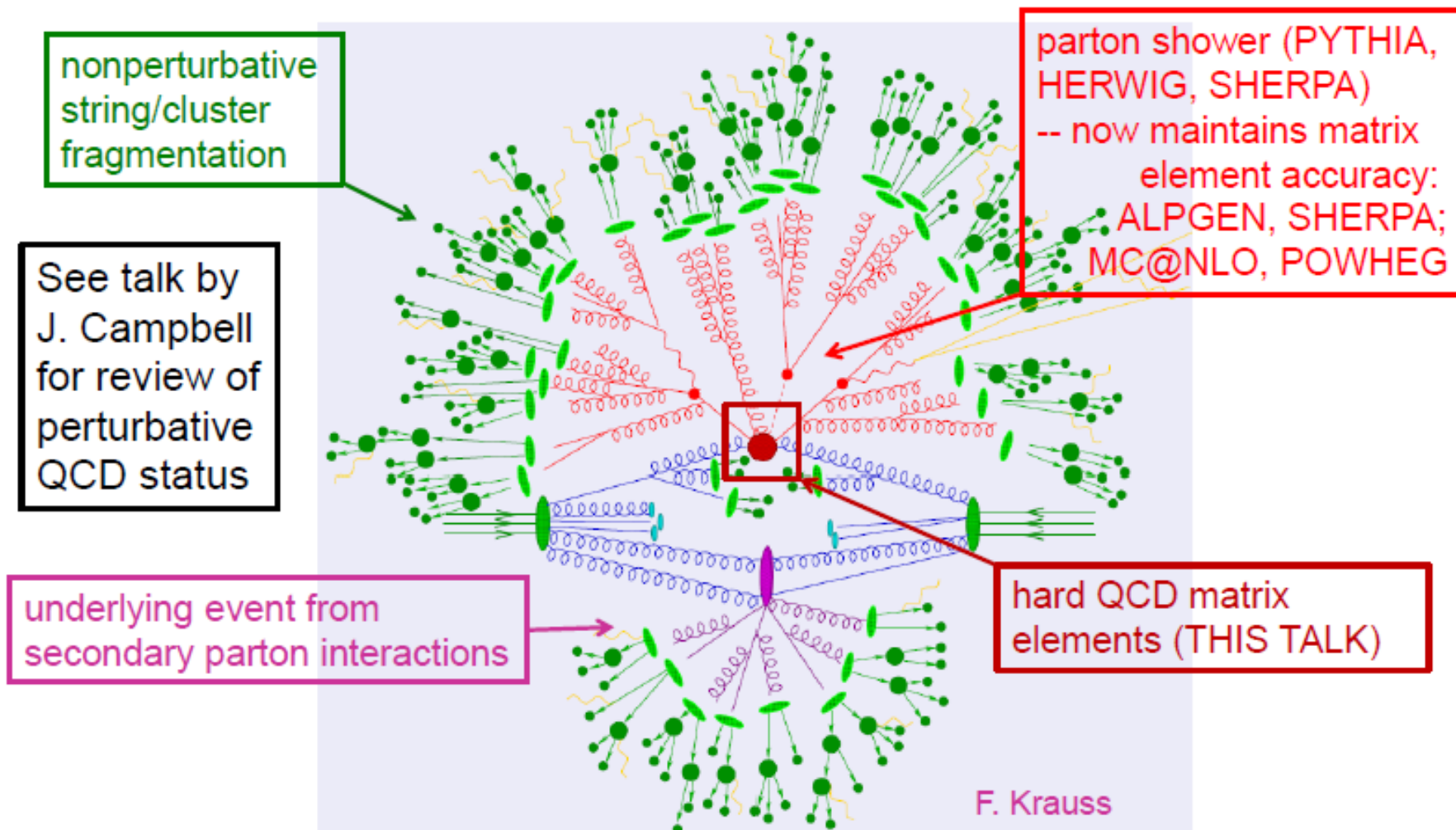
◆ Refining “well-known” calculations, e.g. NLO QCD



- **Jet results:** Precision measurement of fundamental observables.
=> sensitivity to PDF sets, strongest constraint on gluon PDF, extraction of α_s and test of its running up to 400 GeV, detailed studies of the effect of different jet algorithms, study of jet substructure, limits on many NP models.
- **Z/W results:** extensive tests of pQCD and MC models; in most cases, a triumph of NLO and ME-PS MC predictions.

with these tools in hands...

“Typical” hadron collider event





CMS

Event
CMS Experiment at LHC, CERN
Data recorded: Mon May 28 01:16:20 2012 CERN
Run/Event: 195098 / 35488128
Lumi section: 65
Orbit/Crossing: 16992111 / 2295

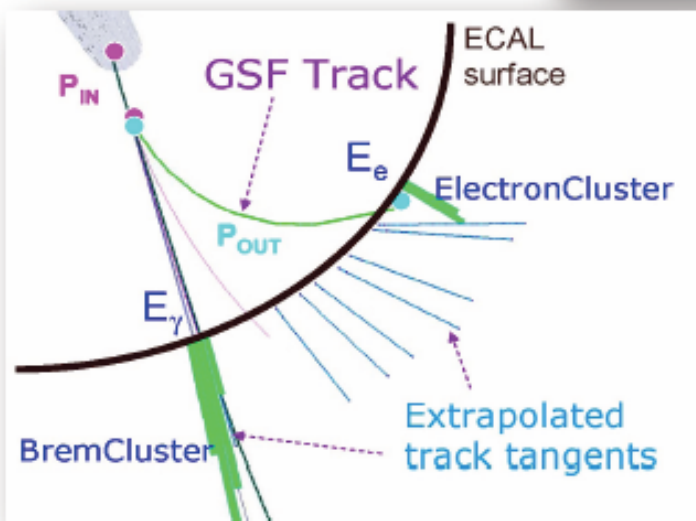
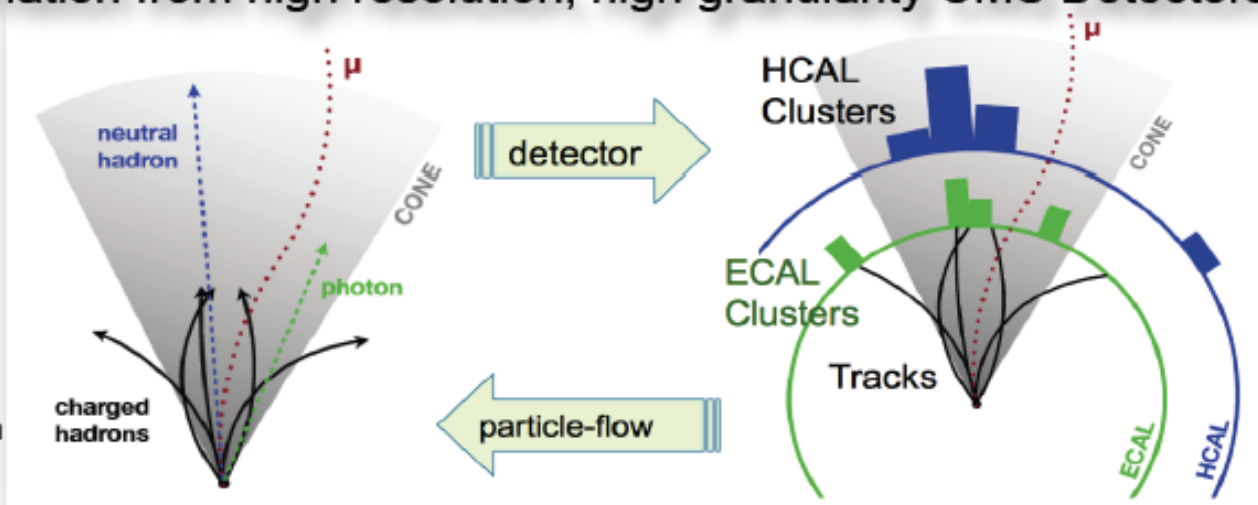
Raw $\Sigma E_T \sim 2$ TeV
14 jets with $E_T > 40$ GeV
Estimated PU ~ 50



Innovation: Global Event Description

Optimal use of information from high resolution, high granularity CMS Detectors

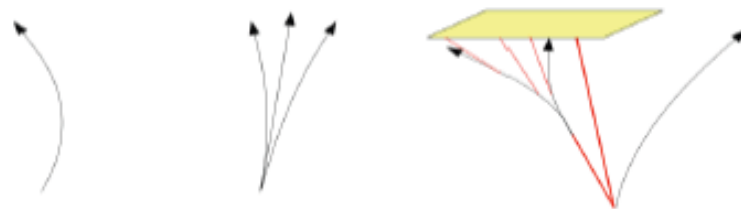
- Lists reconstructed particles
 - e, μ, γ , charged and neutral hadrons
 - Used like "generated particles"
 - Building blocks for jets, taus, missing E_T , isolation
 - PU particle identification



$\tau \rightarrow \pi \nu$

$\tau \rightarrow a_1 \nu$

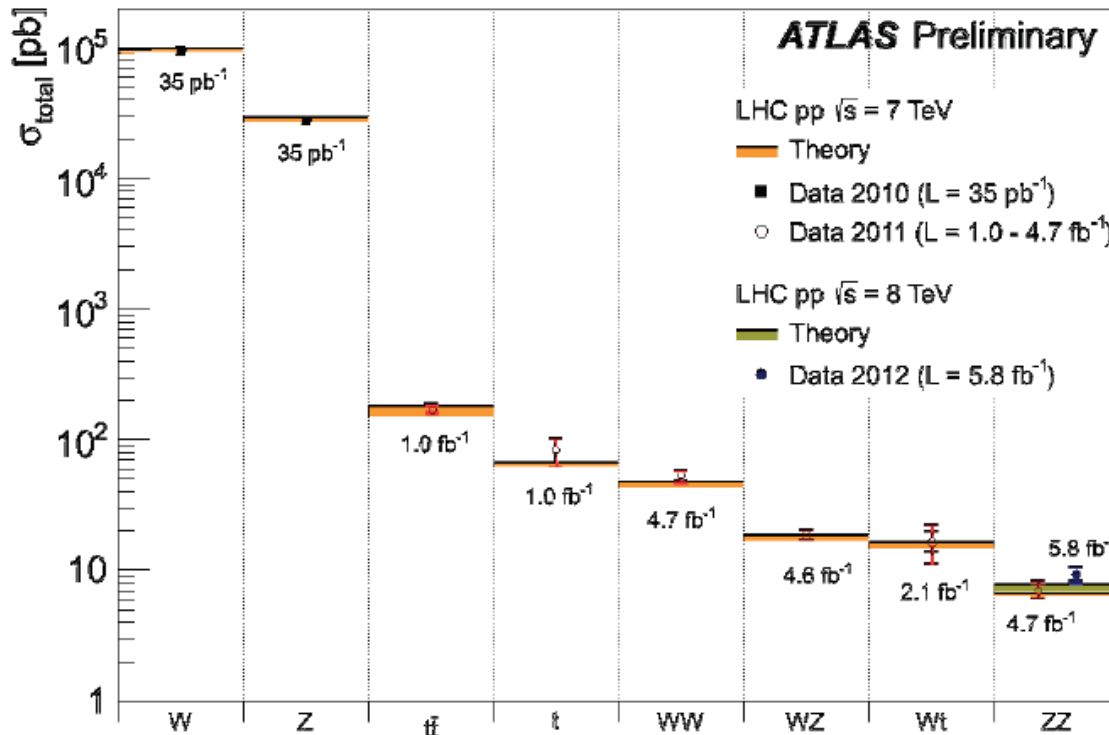
$\tau \rightarrow \rho \nu$



- Sophisticated algorithms
 - Examples: e/γ and hadronic τ

standard candles @LHC

- Foundations for searches - measurements of W, Z, diboson and top prodⁿ:

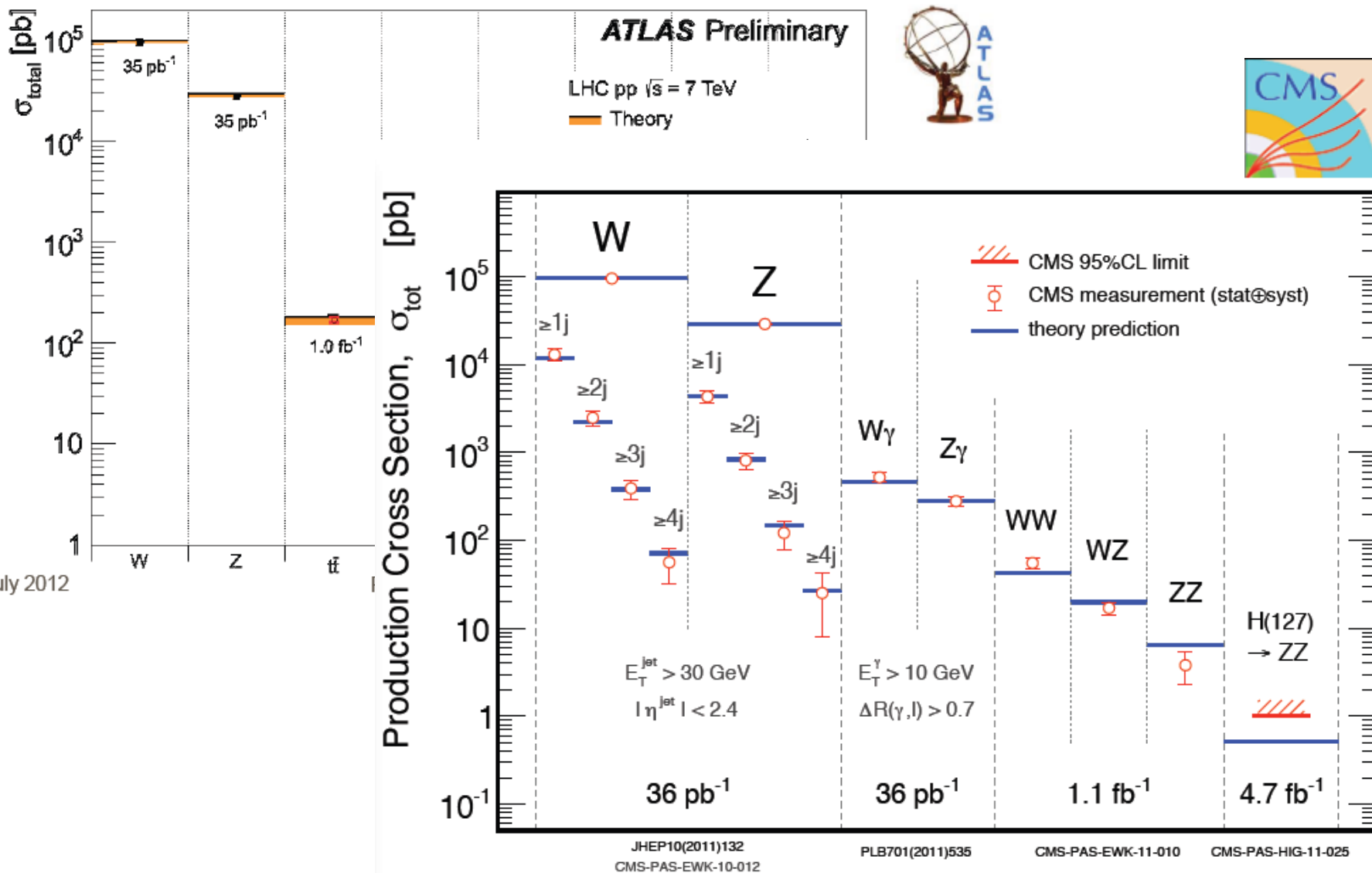


9th July 2012

Richard Hawking

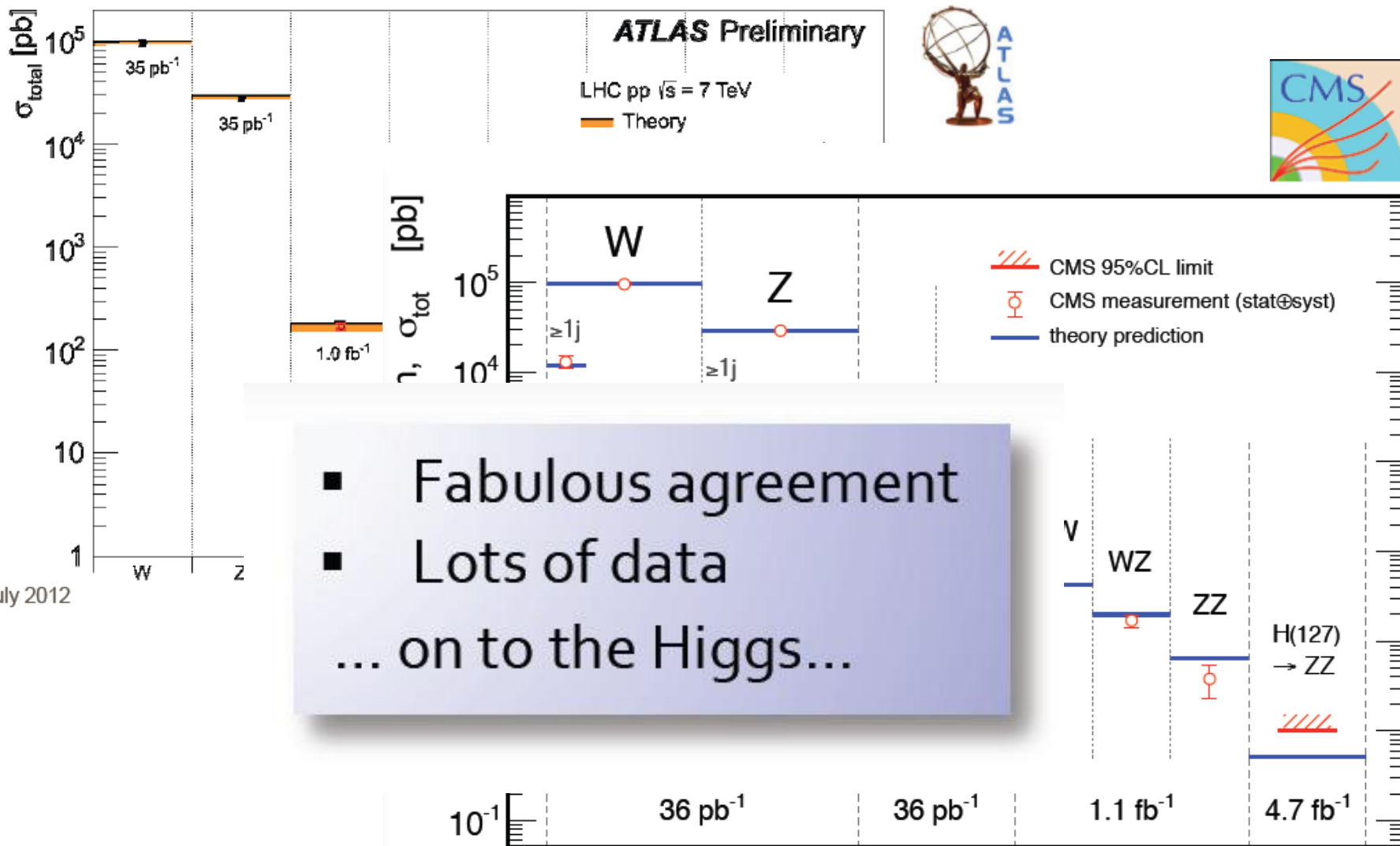
standard candles @LHC

- Foundations for searches - measurements of W, Z, diboson and top prodⁿ:



standard candles @LHC

- Foundations for searches - measurements of W, Z, diboson and top prodⁿ:



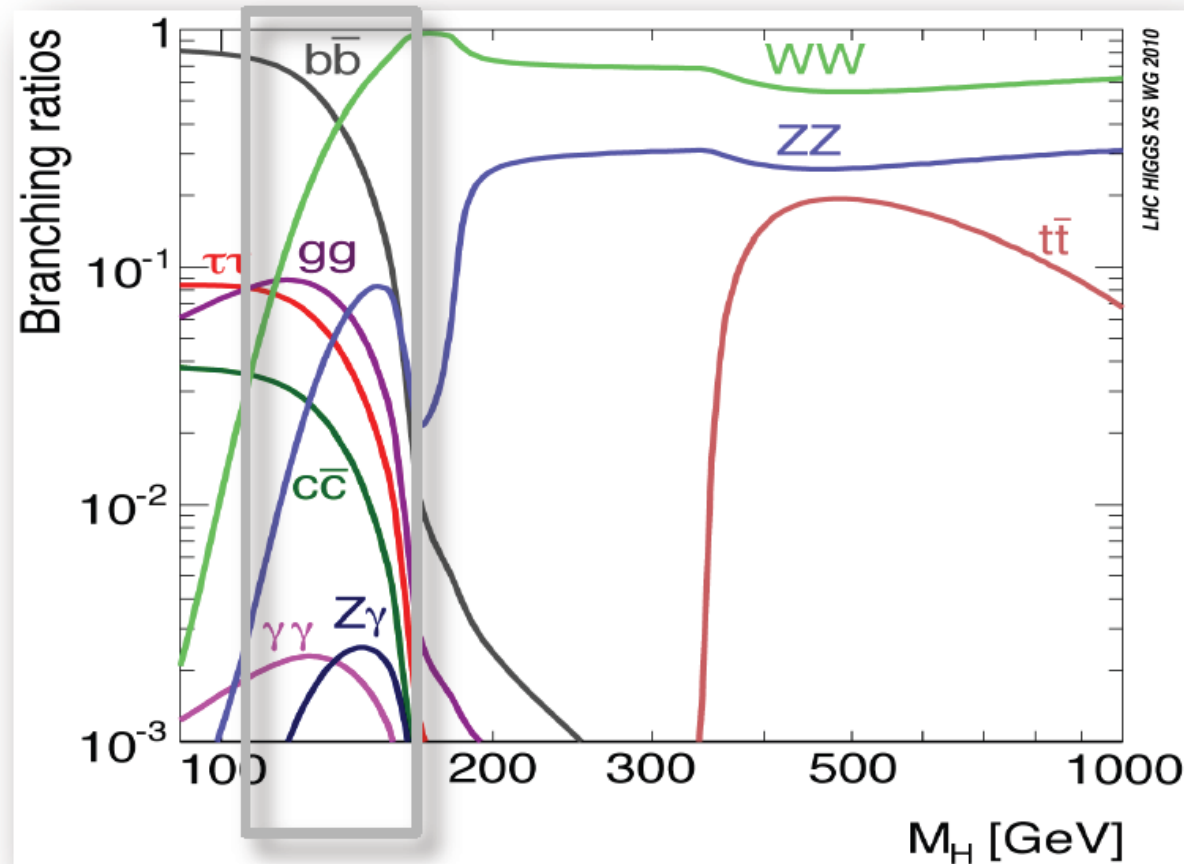
higgs search @LHC



➤ low mass region is left for the higgs:

by 2011:

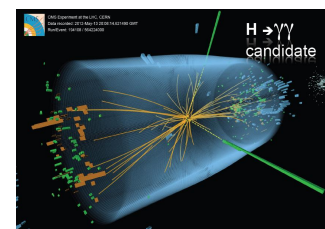
Decay	m_H range (GeV)
$H \rightarrow \gamma\gamma$	110-150 (*)
$H \rightarrow ZZ(*) \rightarrow 4l$	110-600 (*)
$H \rightarrow ZZ \rightarrow llqq$	200-280-600
$H \rightarrow ZZ \rightarrow ll\nu\nu$	200-300-600
$H \rightarrow WW(*) \rightarrow l\nu l\nu$	110-200-300-600
$H \rightarrow WW \rightarrow l\nu qq$	300-600
$H \rightarrow \tau\tau \rightarrow ll, l\tau_h, \tau_h\tau_h$	110-150
$VH \rightarrow l\nu bb, llbb, \nu bb$	110-130



- Low mass region is very rich but also very challenging:
 main decay modes (bb , $\tau\tau$) are hard to identify in the huge background

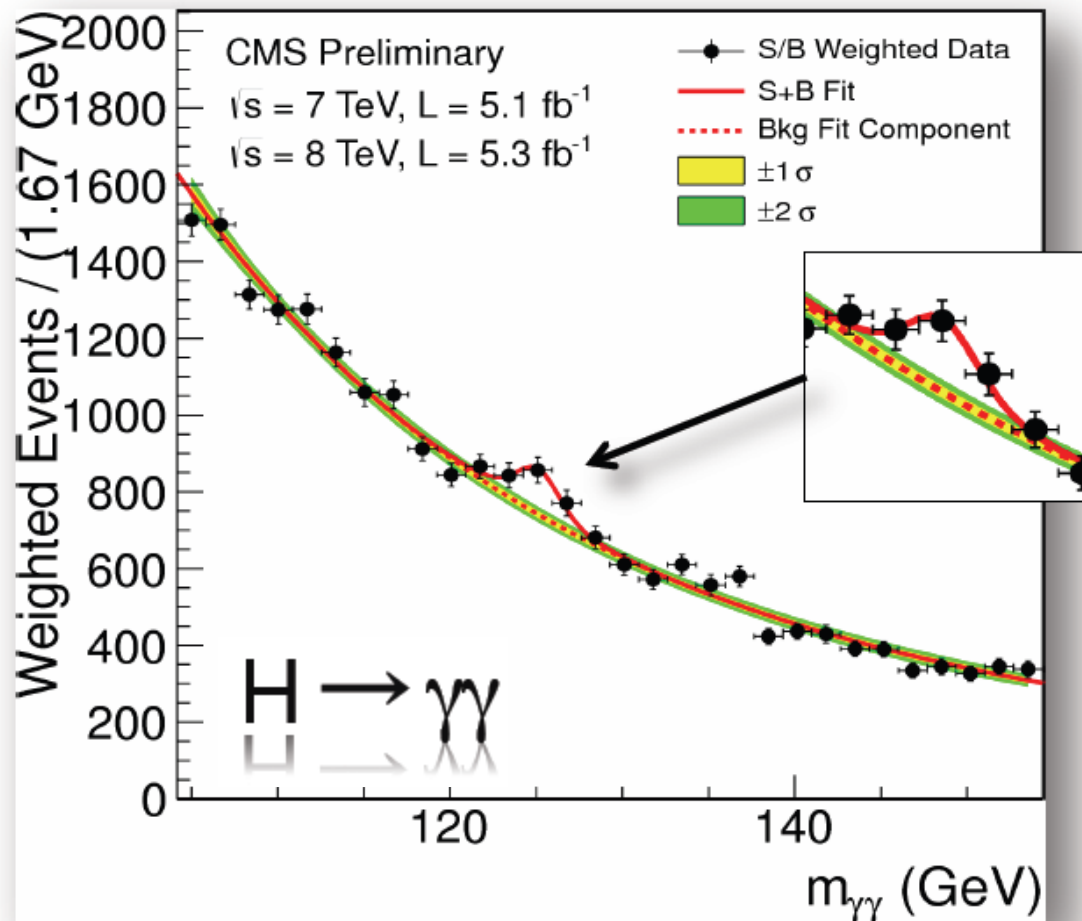
- Very good mass resolution (1%): $H \rightarrow \gamma\gamma$ and $H \rightarrow ZZ \rightarrow 4l$

higgs search @LHC

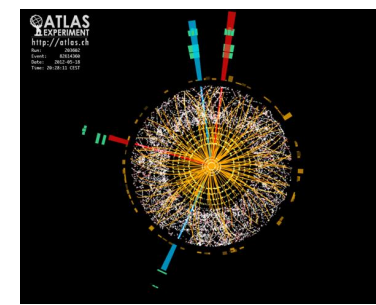


S/B Weighted Mass Distribution

- Sum of mass distributions for each event class, weighted by S/B
 - B is integral of background model over a constant signal fraction interval

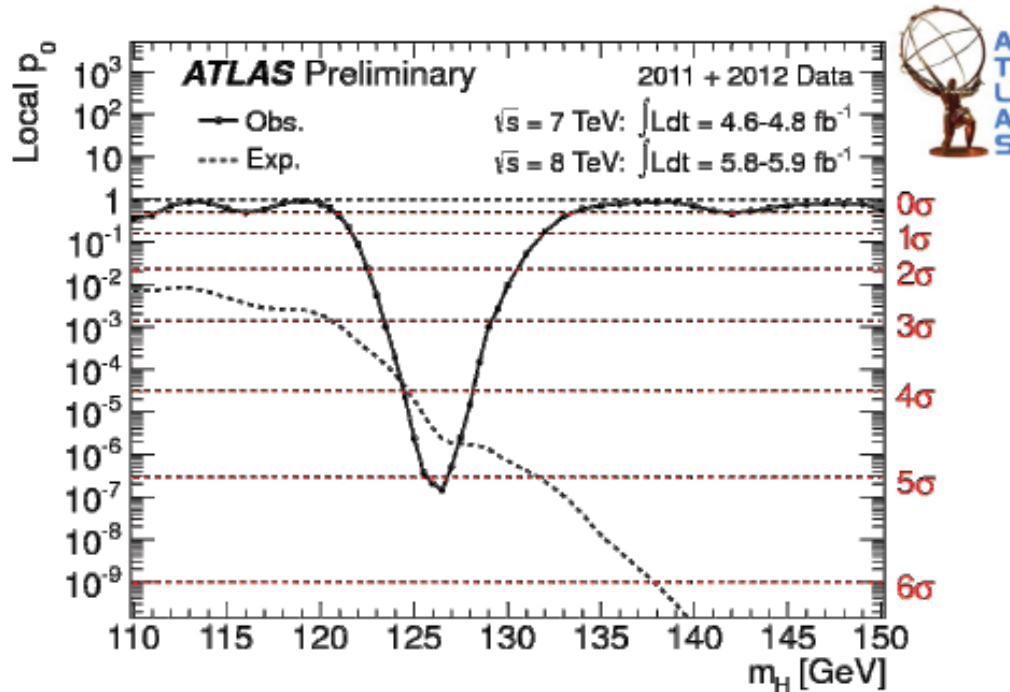


higgs search @LHC

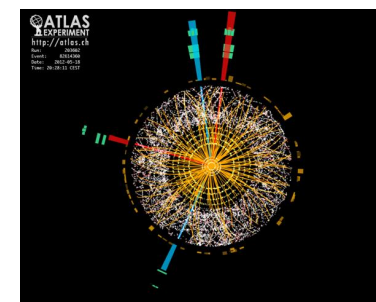


$H \rightarrow ZZ(*) \rightarrow 4l$ ($l = e, \mu$): the golden channel

- A great performing channel in the whole mass range ...
 - Clean signature: narrow peak, low background
 - Background: irreducible $ZZ(*)$; reducible Z +jets, $t\bar{t}$, WZ
- But extremely demanding
 - Requires highest possible efficiencies (lepton Reco/ID/Isolation).

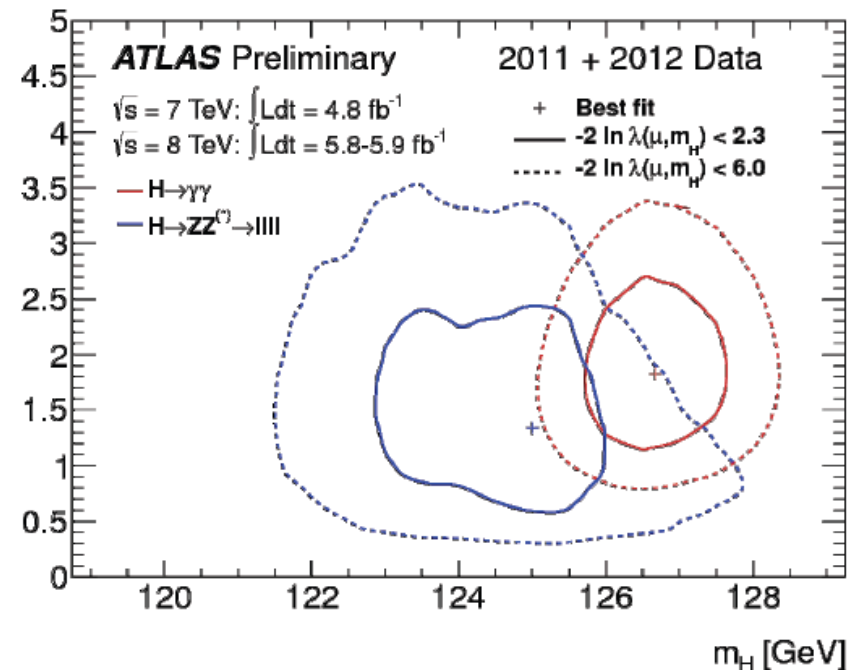
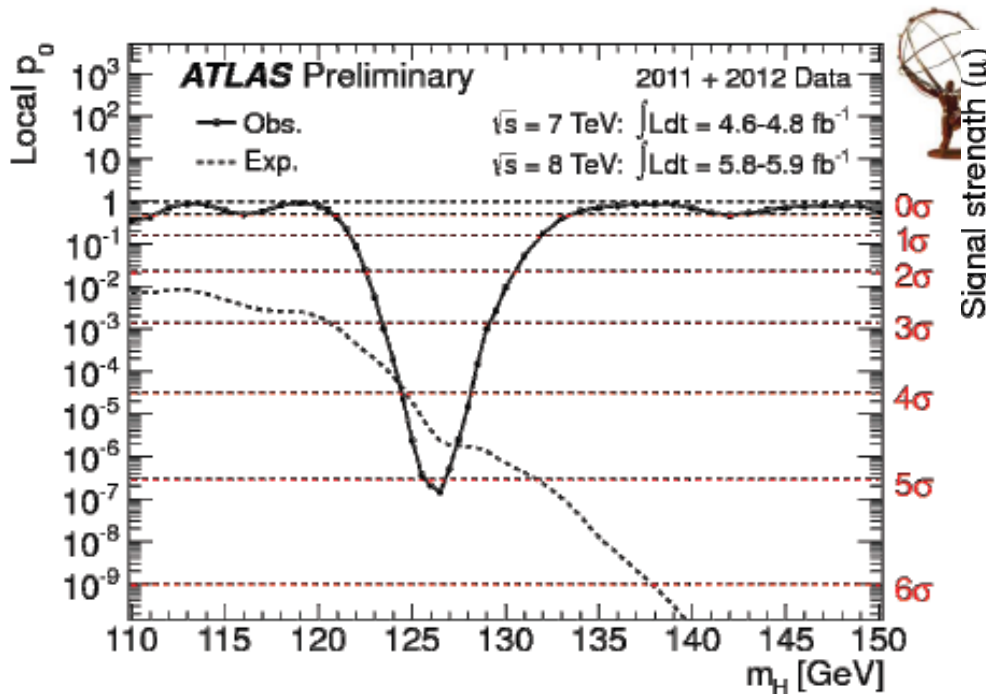


higgs search @LHC



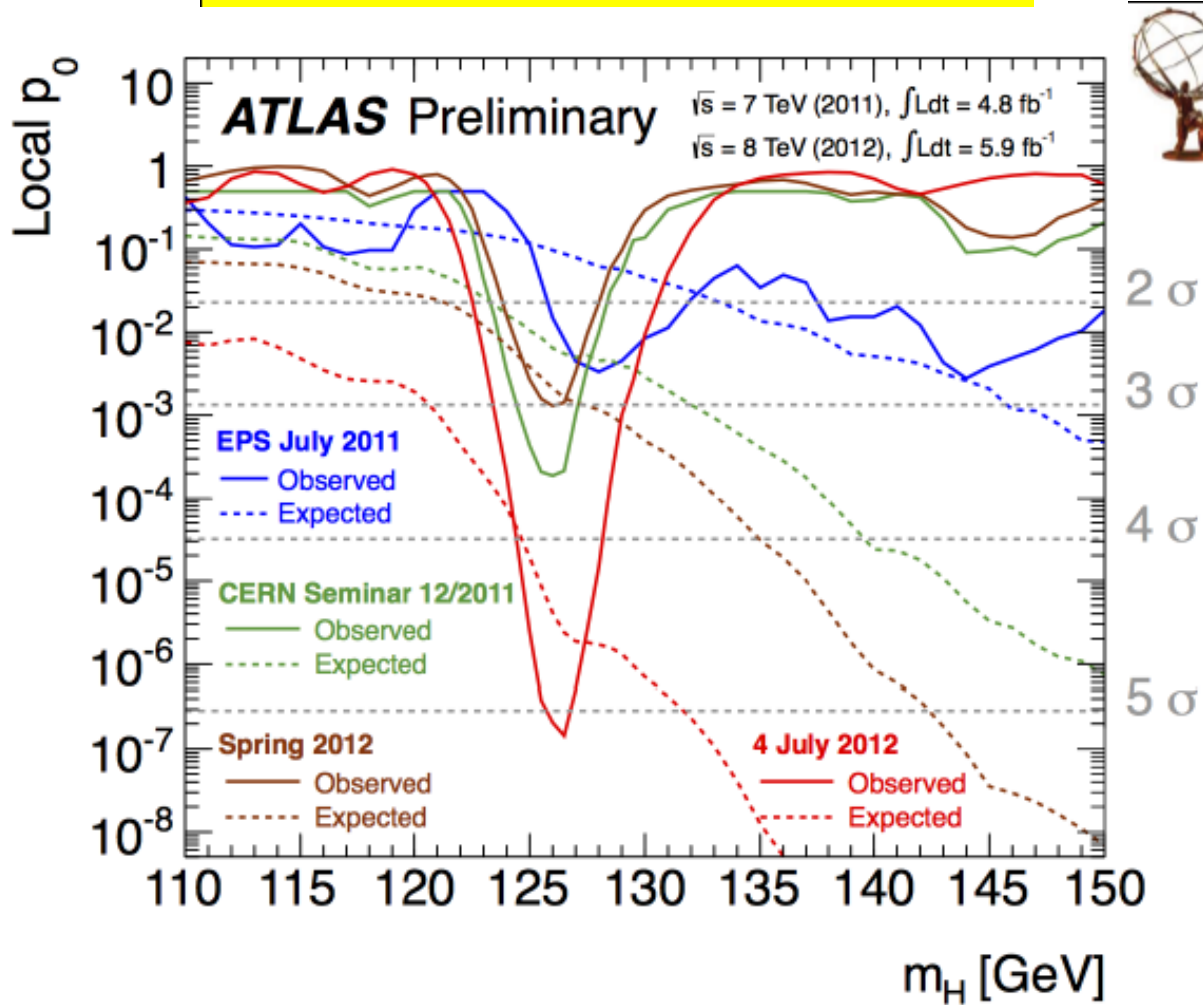
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higgs search @LHC

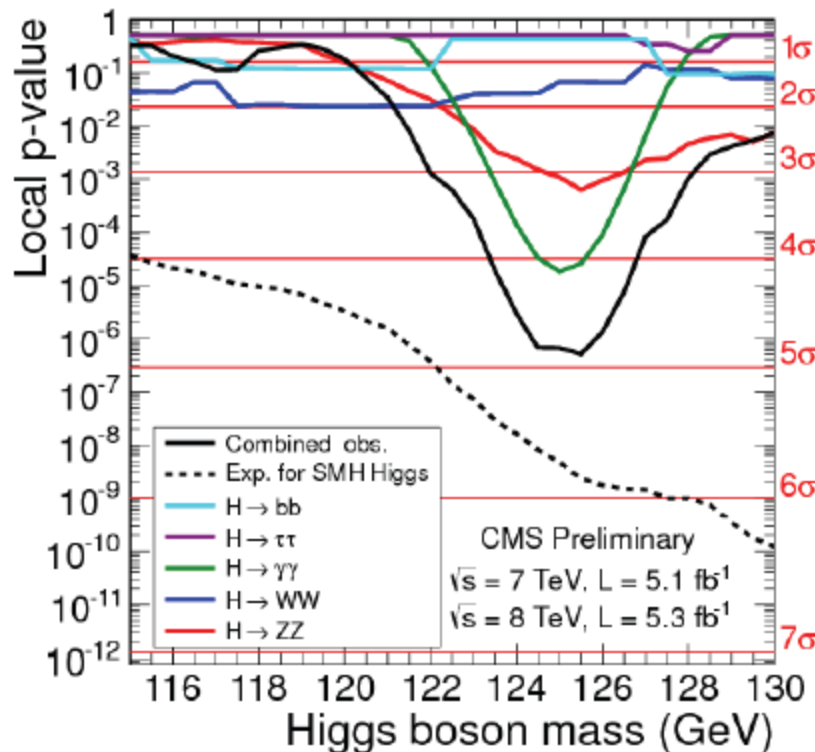
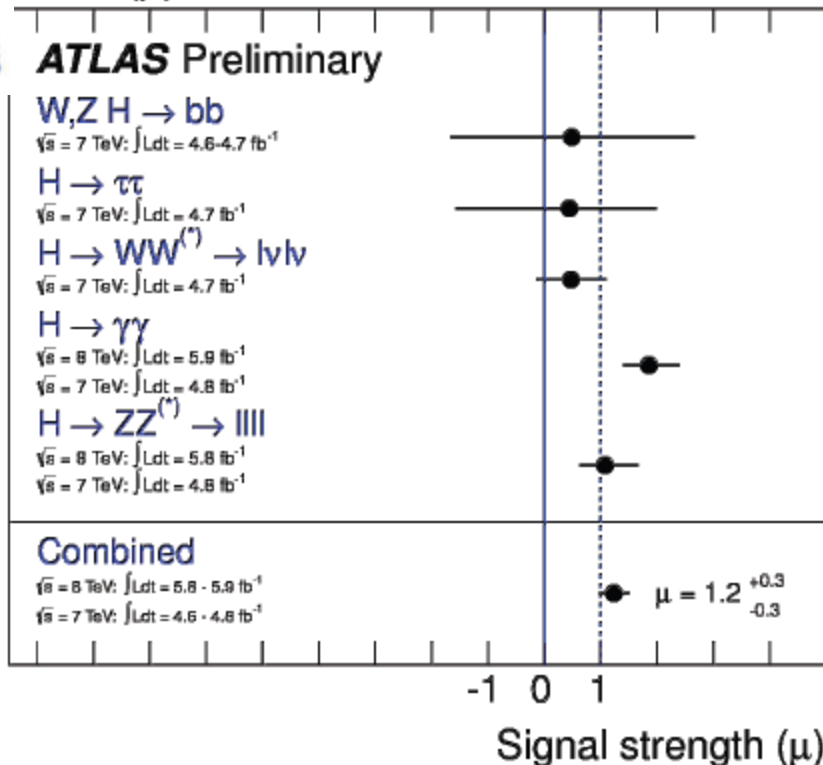
Evolution of the excess with time



all channels

higgs search @LHC

All together



Combined search sees excess with local significance of **5.0 σ** at $m_H = 126.5 \text{ GeV}$

Combination	Expected (σ)	Observed (σ)
$\gamma\gamma + ZZ \rightarrow 4l$	4.7	5.0
$\gamma\gamma + ZZ \rightarrow 4l + WW$	5.2	5.1
All	5.9	4.9

Conclusion

We have observed a new boson
with a mass of

$$125.3 \pm 0.4 \text{ (stat)} \pm 0.5 \text{ (syst)}$$

at significance level of 5σ

Conclusion

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3. As next step, need to know:

[Ricardo Barbieri]

- ⇒ Its quantum numbers: $J^{PC} = 0^{++}$, gauge q.n.s
- ⇒ The strength of its interactions with all other particles and with itself
- ⇒ Is it alone or accompanied?
- ⇒ Is it “elementary” or “composite”?
- ⇒ Is it “natural”?

is it a higgs ?

Disclaimer: We have gotten **more data on EWSB in one single day** than in more than **40 years**
↳ **Not enough time to digest it!**

What makes the Higgs special?

**Not just about finding the
the condensate responsible
for giving masses**



[Alex Pomarol]

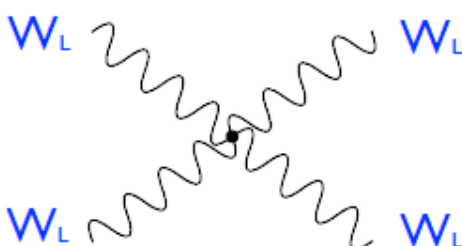
What makes the Higgs special?

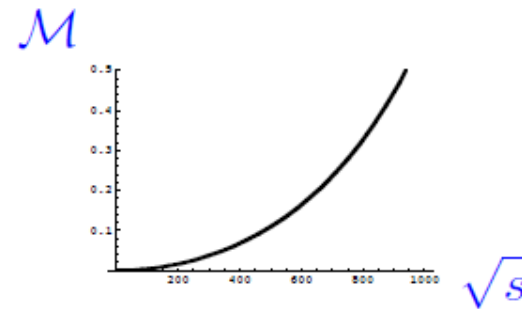
[Alex Pomarol]

**Not just about finding the
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for giving masses**

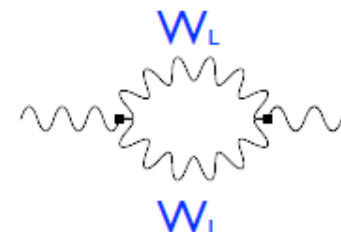


Without a Higgs, the states W_L, Z_L **spoil** the nice
calculability power of gauge theories

a)  $\propto \frac{s}{v^2}$



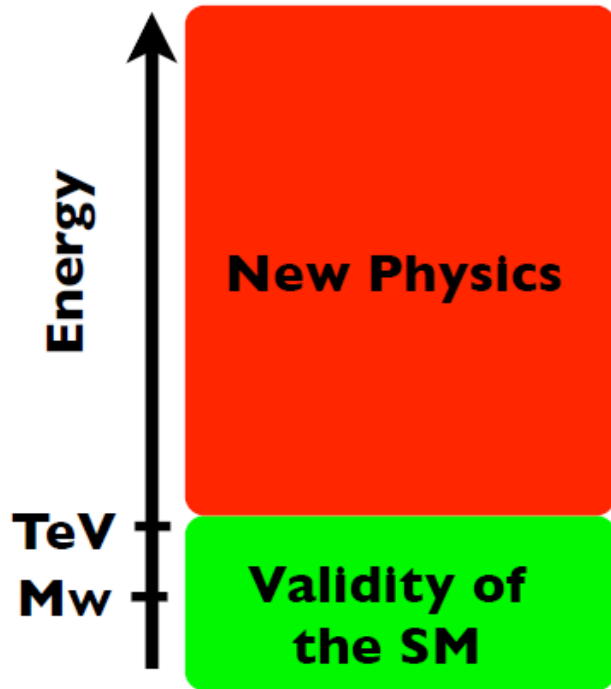
Unitarity is lost at high-energies

b) 

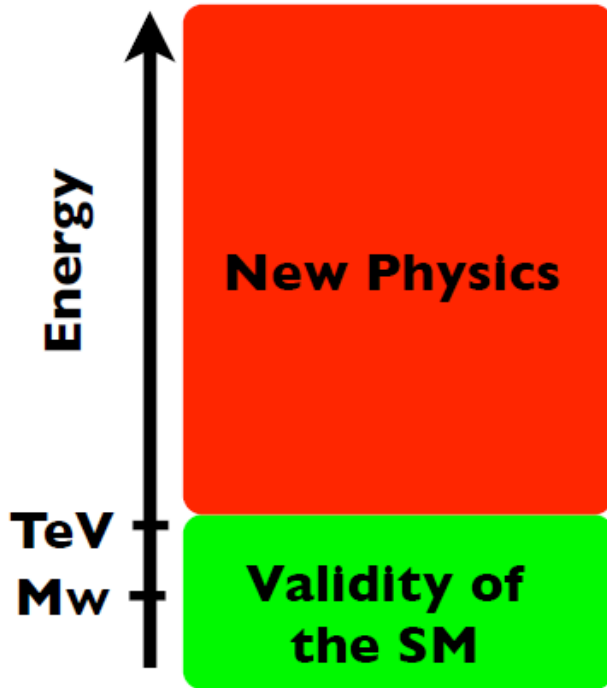
Loops are not finite!

Do not allow for precision calculations

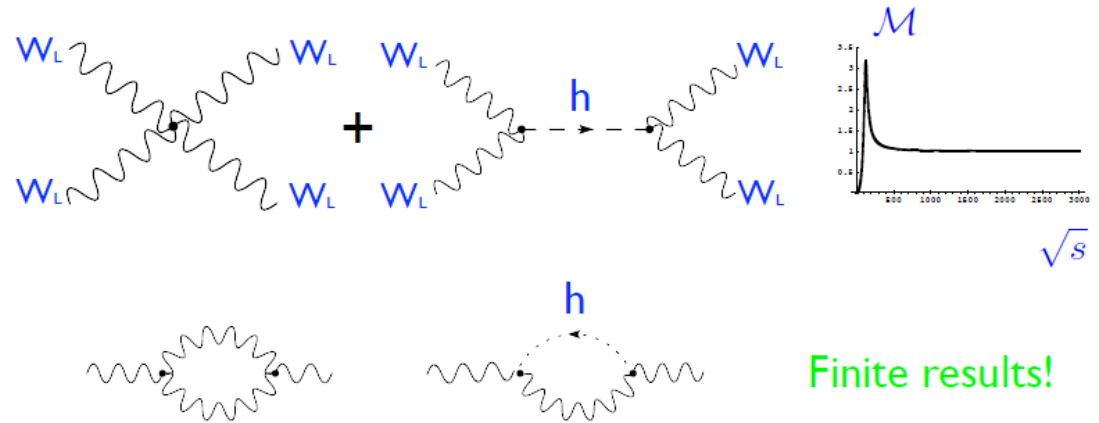
Without a Higgs...



Without a Higgs...

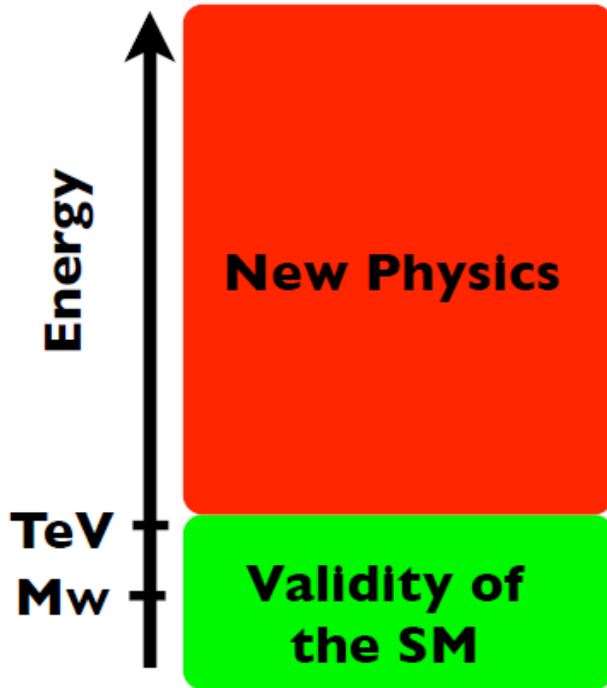


With the Higgs **calculability** is recovered:



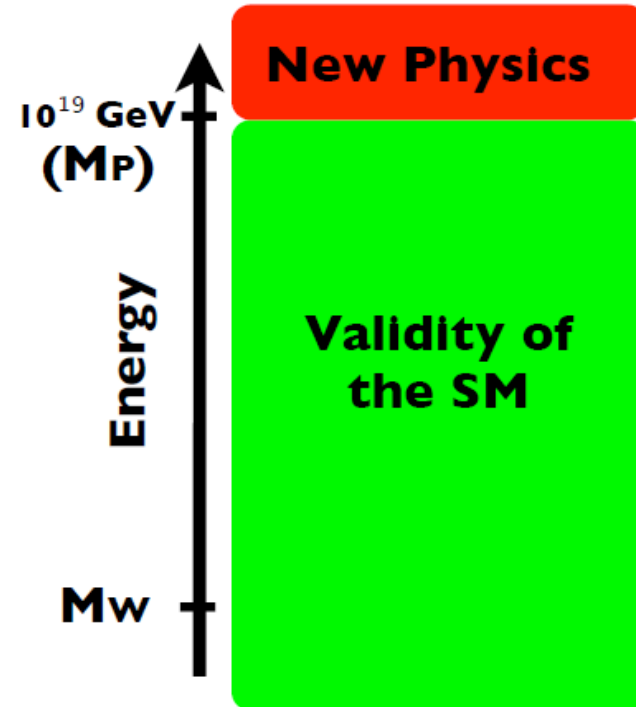
Back to the prediction era!

Without a Higgs...



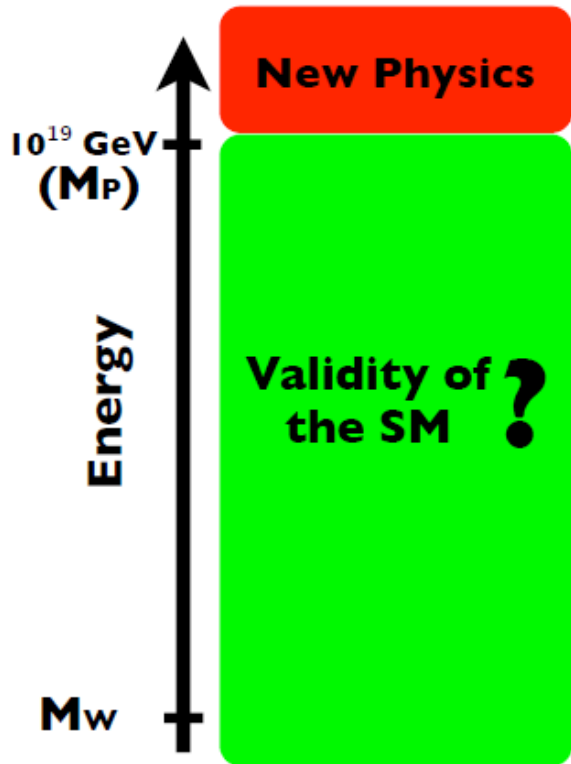
With a Higgs

(100 GeV < mh < 170 GeV)



Although consistent, we think (and hope) the SM is not the full story

[Alex Pomarol]



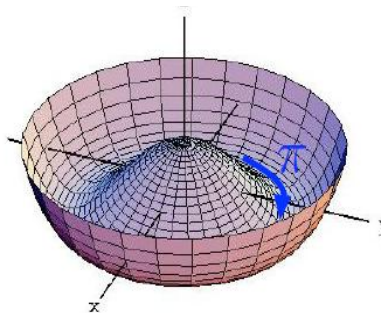
Not understandable
the origin of such a
small EW scale
as compared to the
Planck scale

**Possibilities that theorists envisage
to tackle this problem:**

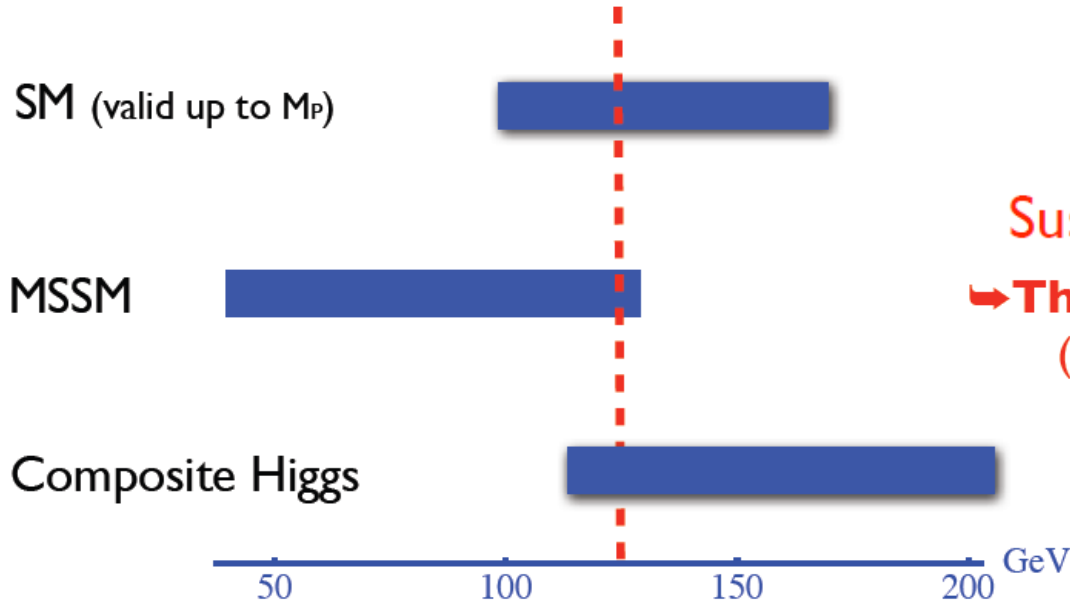
- 1) Keep the Higgs elementary, but protect it by symmetries: **Supersymmetry**
- 2) The Higgs is not elementary: **Composite Higgs**

➔ Both imply **changes** in the Higgs sector

Pseudo-Goldstone bosons (PGB)



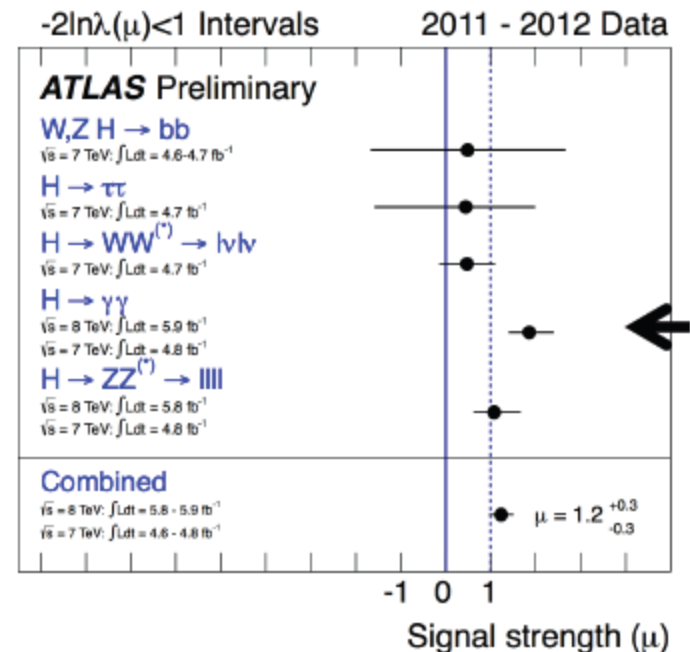
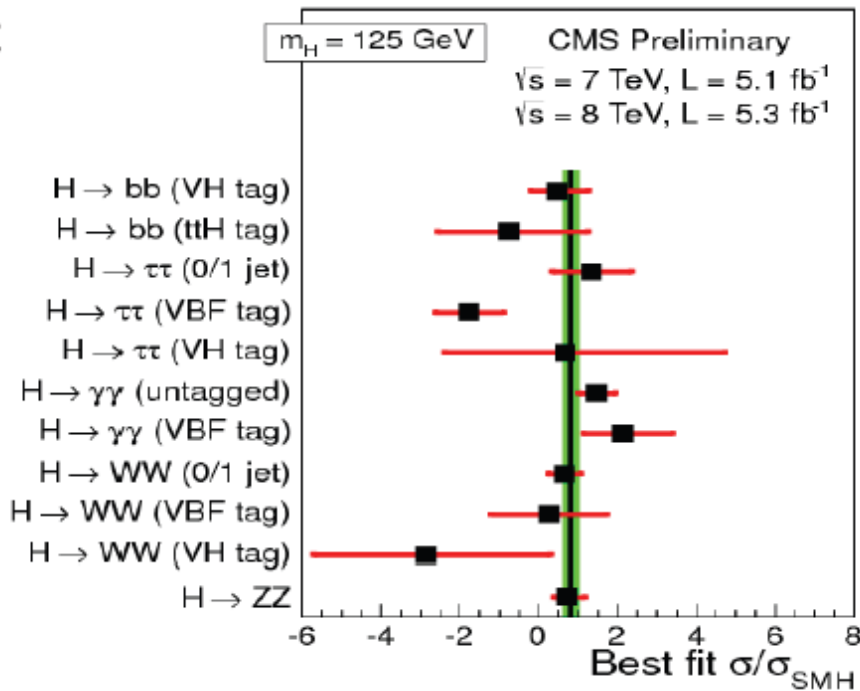
Higgs mass range



Susy must be “badly” broken!

→ **The MSSM is becoming unnatural**
(>99% parameter space excluded)

What the Higgs couplings tells us?



Not significant deviations from a SM Higgs

(The **more** natural the Higgs sector is,
the **more** we expect
deviations from the SM Higgs couplings)

A new era has begun ...

"Sit down before fact as **a little child**,
be prepared to **give up** every preconceived notion,
follow humbly wherever and to whatever abysses nature leads,
or you shall learn nothing"

Thomas Henry Huxley

ν highlight

Reactor $\bar{\nu}_e$ (D-Chooz, Daya-Bay, Reno)

- *At the intensity frontier:*

Large mixing angle

θ_{13} around 9°

neutrinos are massive

and There is Physics Beyond SM

The New Minimal Standard Model

Neutrino Mixing

In a 3- ν framework

$$\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = \begin{pmatrix} U_{e1} & U_{e2} & U_{e3} \\ U_{\mu1} & U_{\mu2} & U_{\mu3} \\ U_{\tau1} & U_{\tau2} & U_{\tau3} \end{pmatrix} \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}$$



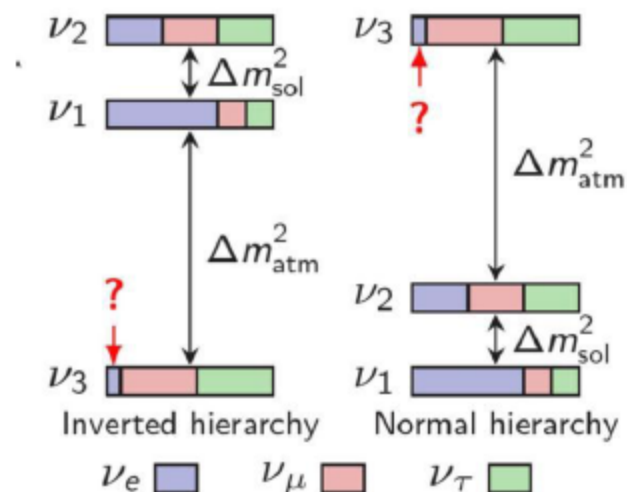
$$U = \begin{pmatrix} 1 & 0 & 0 \\ 0 & c_{23} & s_{23} \\ 0 & -s_{23} & c_{23} \end{pmatrix} \begin{pmatrix} c_{13} & 0 & s_{13} \\ 0 & e^{-i\delta} & 0 \\ -s_{13} & 0 & c_{13} \end{pmatrix} \begin{pmatrix} c_{12} & s_{12} & 0 \\ -s_{12} & c_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix} \begin{pmatrix} e^{i\rho} & 0 & 0 \\ 0 & e^{i\sigma} & 0 \\ 0 & 0 & 1 \end{pmatrix}$$

$\theta_{23} \sim 45^\circ$
Atmospheric
Accelerator

$\theta_{13} = ?$
Reactor
Accelerator

$\theta_{12} \sim 34^\circ$
Solar
Reactor

$0\nu\beta\beta$



The New Minimal Standard Model

Neutrino Mixing

- ◆ Sensitive to
 - ❖ $CPV \propto \sin \delta \cdot s_{12} \cdot s_{23} \cdot s_{13}$
 - ❖ sign of Δm_{13} (Mass hierarchy) thru Matter effect

Existence of ν_e appearance

→ Non-zero (reasonable size) θ_{13}

→ CPV term can exist

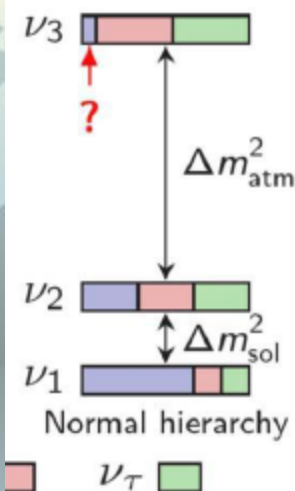
Chance to detect CPV in the future!

$\theta_{23} \sim 45^\circ$
Atmospheric
Accelerator

$\theta_{13} = ?$
Reactor
Accelerator

$\theta_{12} \sim 34^\circ$
Solar
Reactor

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$$U = \begin{pmatrix} e^{i\rho} & 0 & 0 \\ 0 & e^{i\sigma} & 0 \\ 0 & 0 & 1 \end{pmatrix}$$

- ❖ ν_e appearance
- ❖ ν_μ disappearance
- ❖ ν_τ appearance
- ❖ Speed of neutrino

(Accelerator-based) Long baseline neutrino experiments

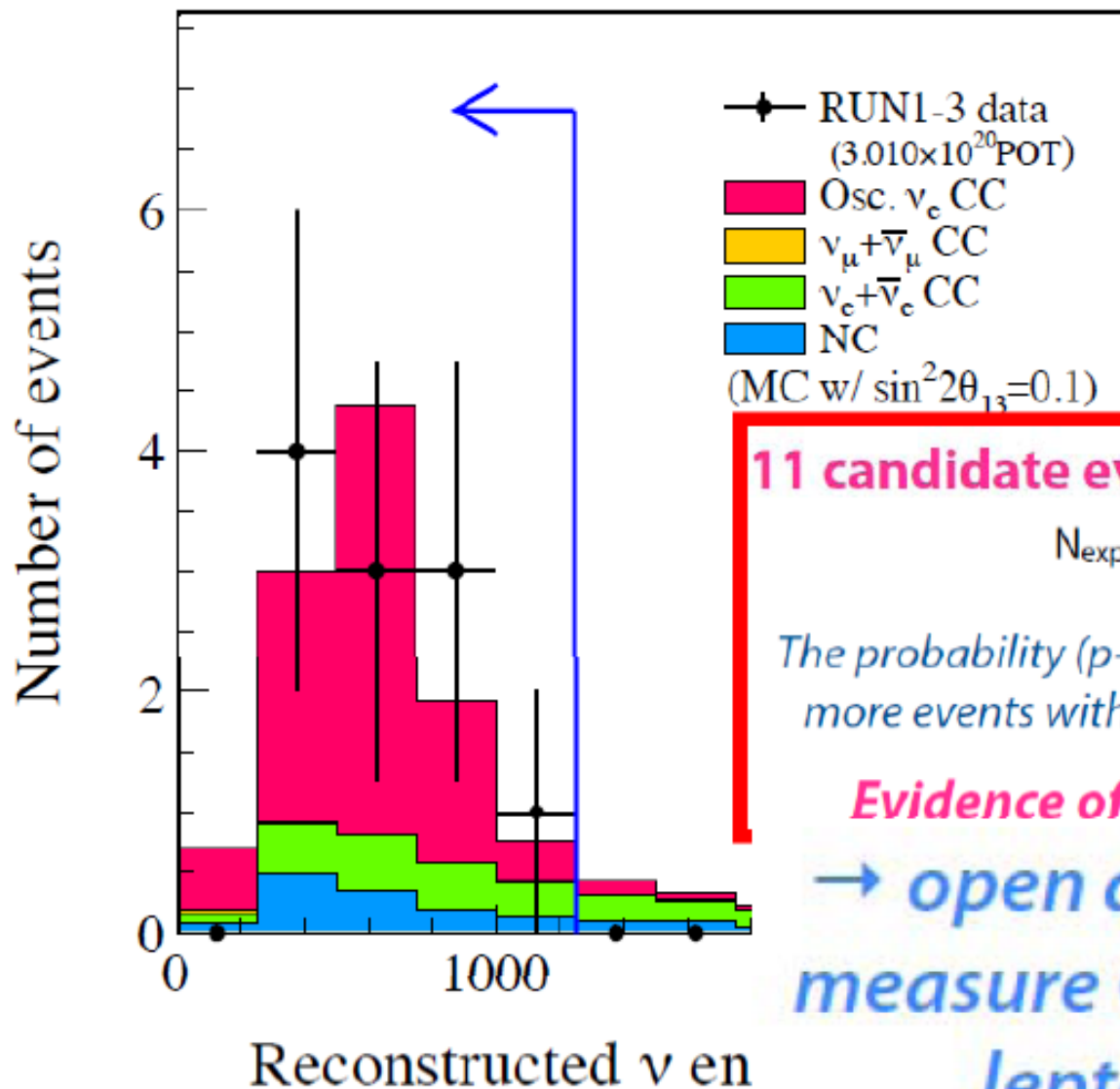
- ❖ Exhibit first violation of standard “SM”: Non-zero mass
- ❖ Surprisingly large mixing unlike quarks
- ◆ Yet unknowns & puzzles (>40yrs behind quark)
 - ❖ Flavor mixing
 - ◆ Standard 3 x 3 PMNS mixing picture is correct?
 - ◆ All three flavors participate mixing? (θ_{13} ?)
 - ◆ Why so different from quark mixing?
 - ◆ CP is violated?
 - ❖ Mass
 - ◆ Absolute mass
 - ◆ “Unexplained lightness of the existence”
 - ◆ Mass ordering (hierarchy)
 - ❖ Any additional neutrino??

more results from MINOS ($\nu-\mu$ disappearance) & OPERA ($\nu-\tau$ appearance)

K.Sakashita

T2K: 11 candidate events

ν_e appearance



11 candidate events are observed

$$N_{\text{exp}} = 3.22 \pm 0.43 \text{ for } \sin^2 2\theta_{13} = 0$$

The probability (*p*-value) to observe 11 or more events with $\theta_{13} = 0$ is 0.08% (3.2σ)

Evidence of ν_e appearance

→ open a possibility to measure CP violation in lepton sector

Precision Measurement at Reactors

[JunCao]



8 proposals, most in 2003 (3 on-going)

- **Fundamental parameter**
- **Gateway to ν -CPV and Mass Hierachy measurements**
- **Less expensive**

Precision Measurement at Reactors

April 8, 2012:

$$\sin^2 2\theta_{13} = 0.113 \pm 0.013 (\text{Stat}) \pm 0.019 (\text{Syst}),$$

results on Jun.4, 2012,

4.9 σ for non-zero θ_{13}

$$\sin^2 2\theta_{13} = 0.109 \pm 0.030 (\text{Stat}) \pm 0.025 (\text{Syst}),$$

3.1 σ for non-zero θ_{13}

Double Chooz, France

RENO, Korea

Diablo Canyon, USA

Daya Bay, China

Angra

Jun.4, 2012, with 139 day data

$$\sin^2 2\theta_{13} = 0.089 \pm 0.010 (\text{stat}) \pm 0.005 (\text{syst})$$

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$$\sin^2 2\theta_{13} = 0.089 \pm 0.010 (\text{stat}) \pm 0.005 (\text{syst})$$

7.7σ for non-zero θ_{13}

- ◆ Daya Bay experiment discovered the new oscillation and proved θ_{13} is quite large.

⇒ We can measure the MH and CPV in our lifetime!

- Why are neutrinos so light?

The Origin of Neutrino Mass

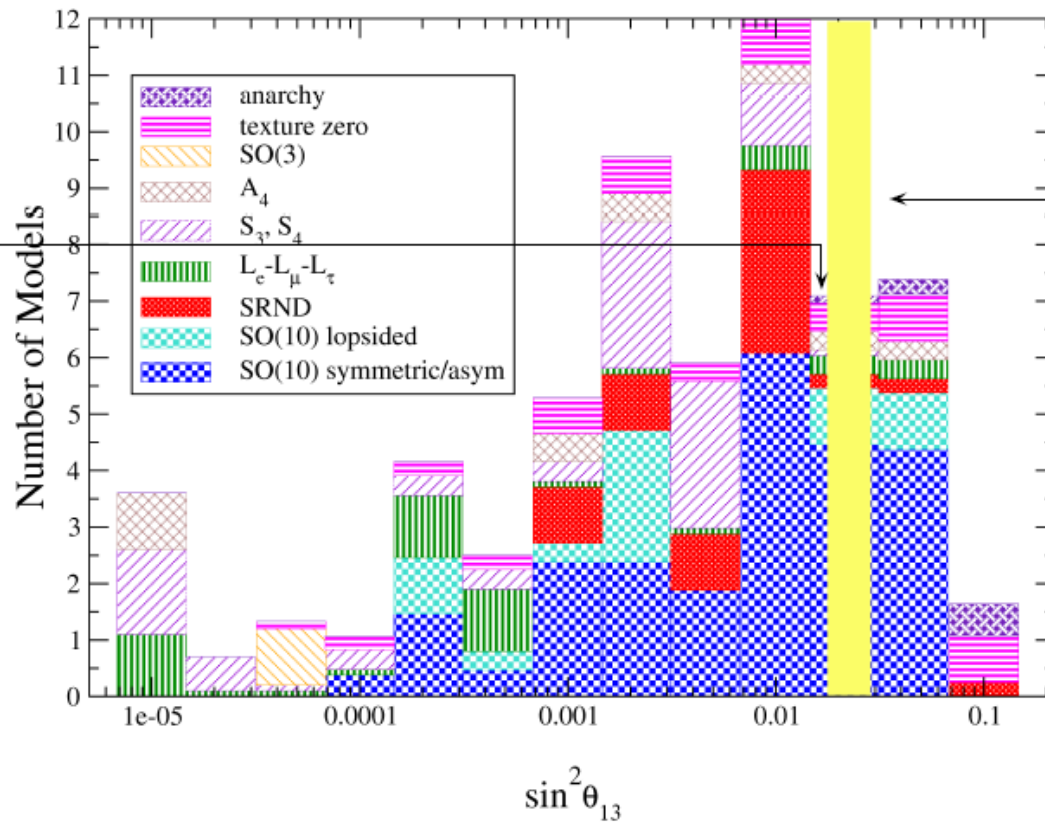
- Why are lepton mixing so different from quark's?

The Flavour Puzzle

- Why are neutrinos so light?

The Origin of Neutrino Mass

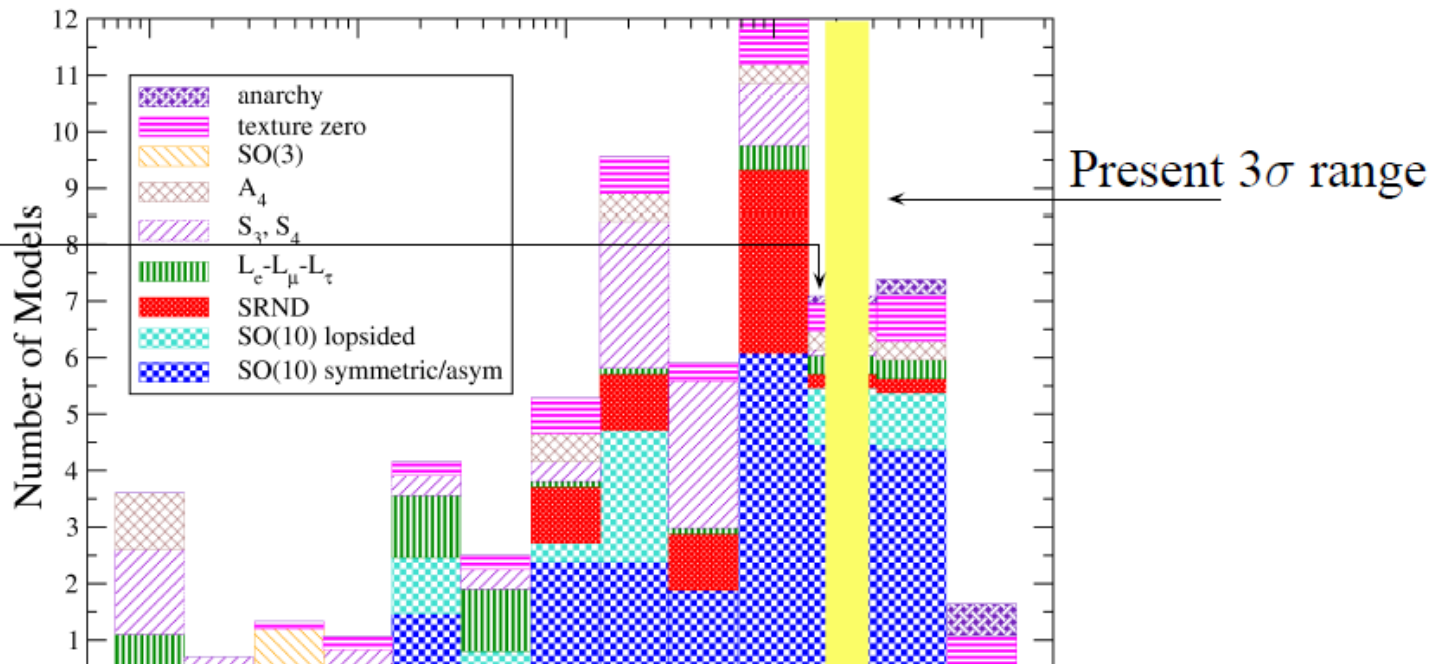
Survey of 63 ν mass models in 2006



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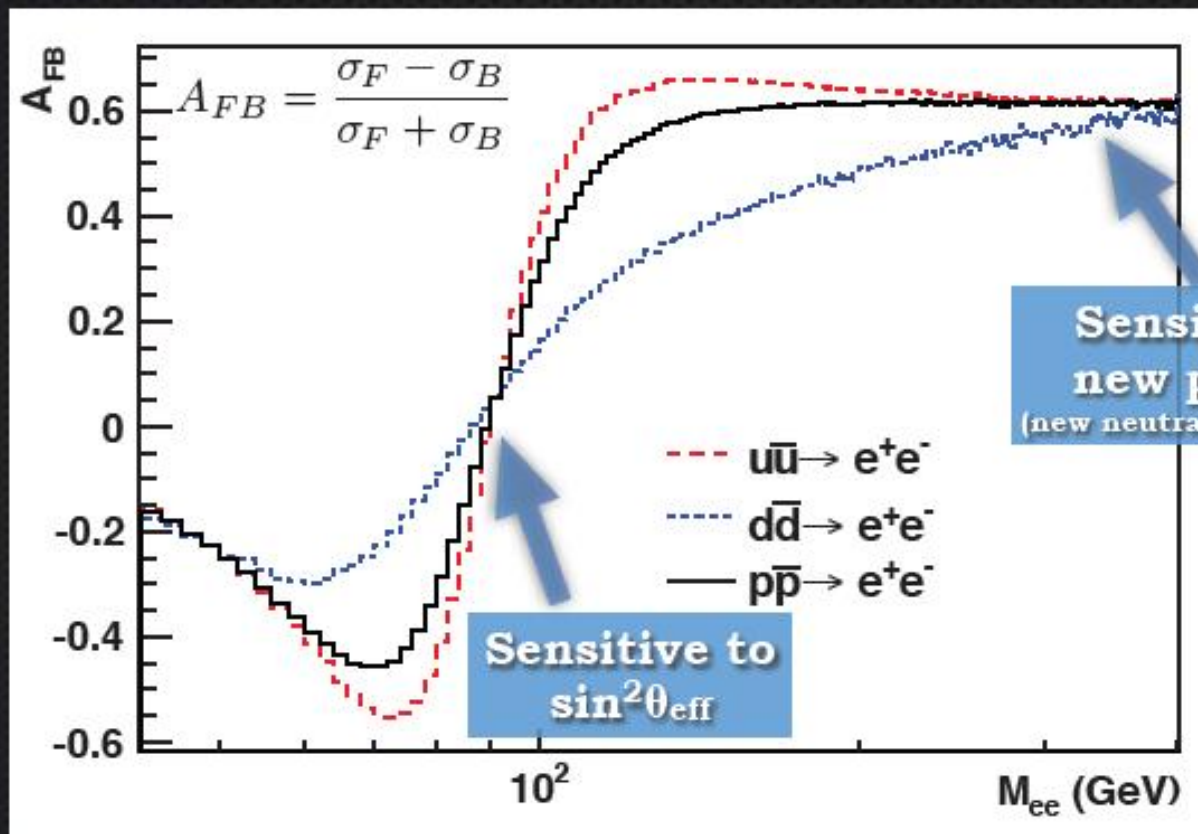
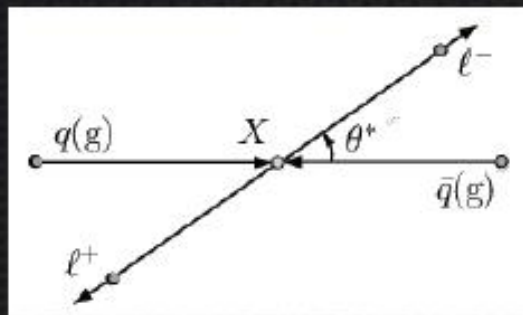
ν masses are BSM physics effects to be put together with *all other NP effects*: from charged LFV, Collider signals, Cosmo-astroparticle... to establish the Next Standard Model

new physics

- **Large Hadron Collider program well underway towards precision physics with W and Z bosons**
 - Stable ground for new physics searches
 - SM physics offers spin-offs into discoveries

Z forward-backward asymmetry

at Hadron Colliders

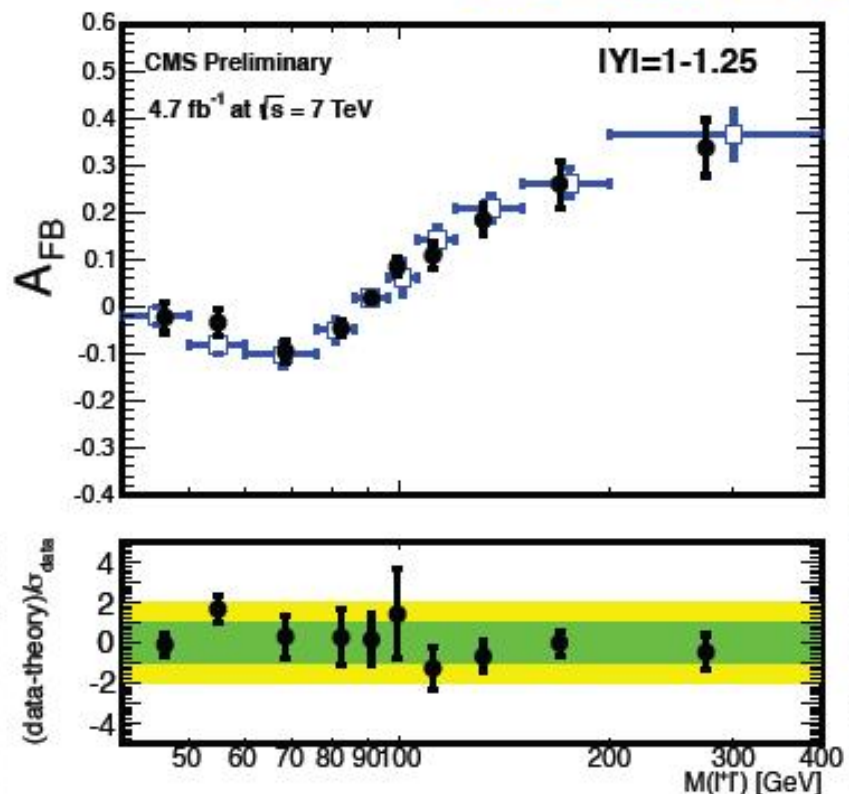
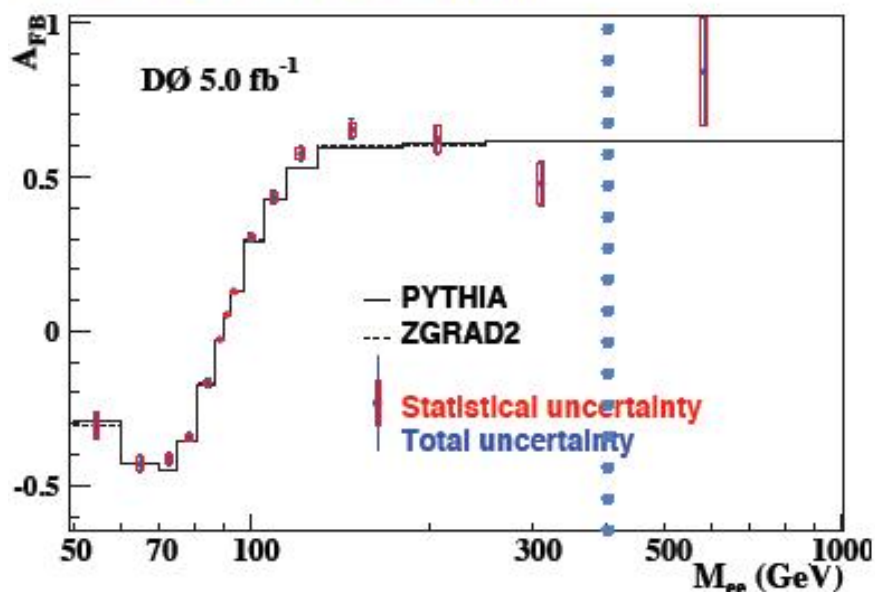


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Phys. Rev. D 84, 012007 (2011)

CMS PAS EWK-11-004

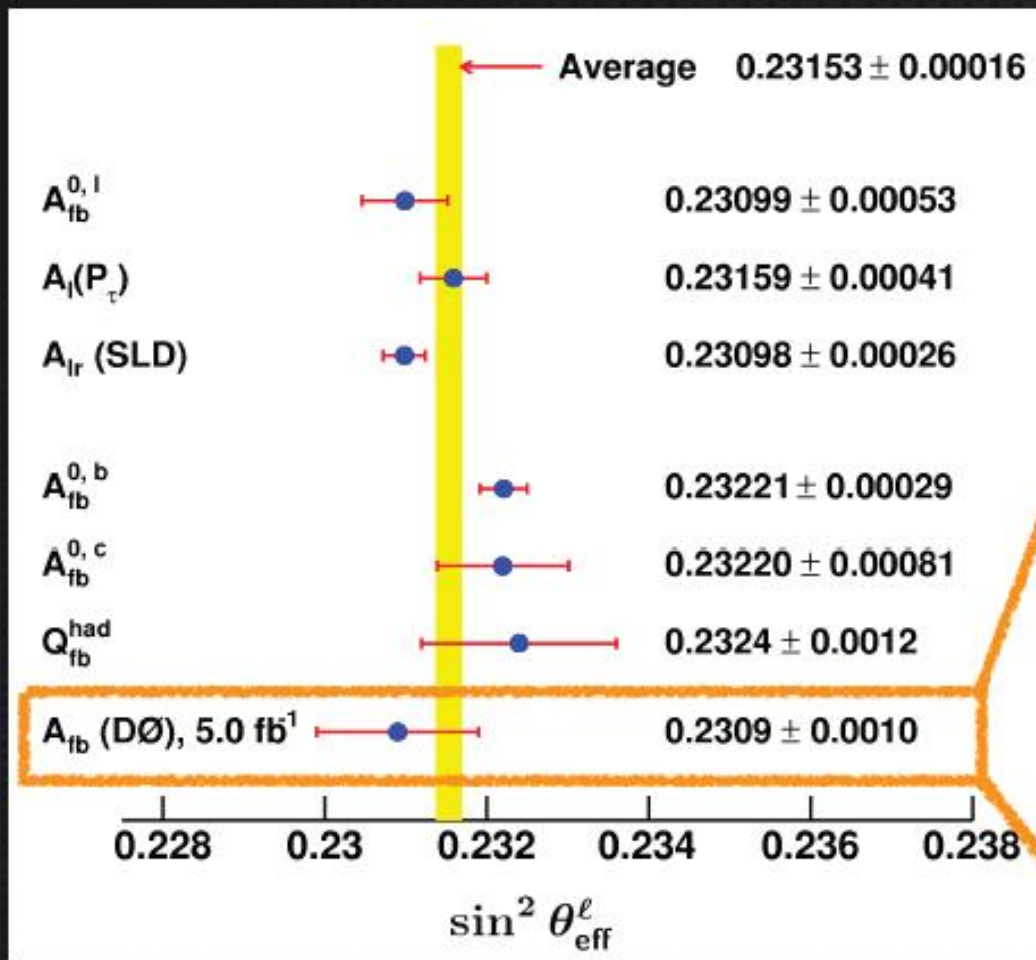


Unfolded AFB agrees well with theoretical predictions

No evidence for new physics at high-mass

Effective Weak Mixing Angle

DØ



$0.2309 \pm 0.0008 \text{ (stat)} \pm 0.0006 \text{ (syst)}$

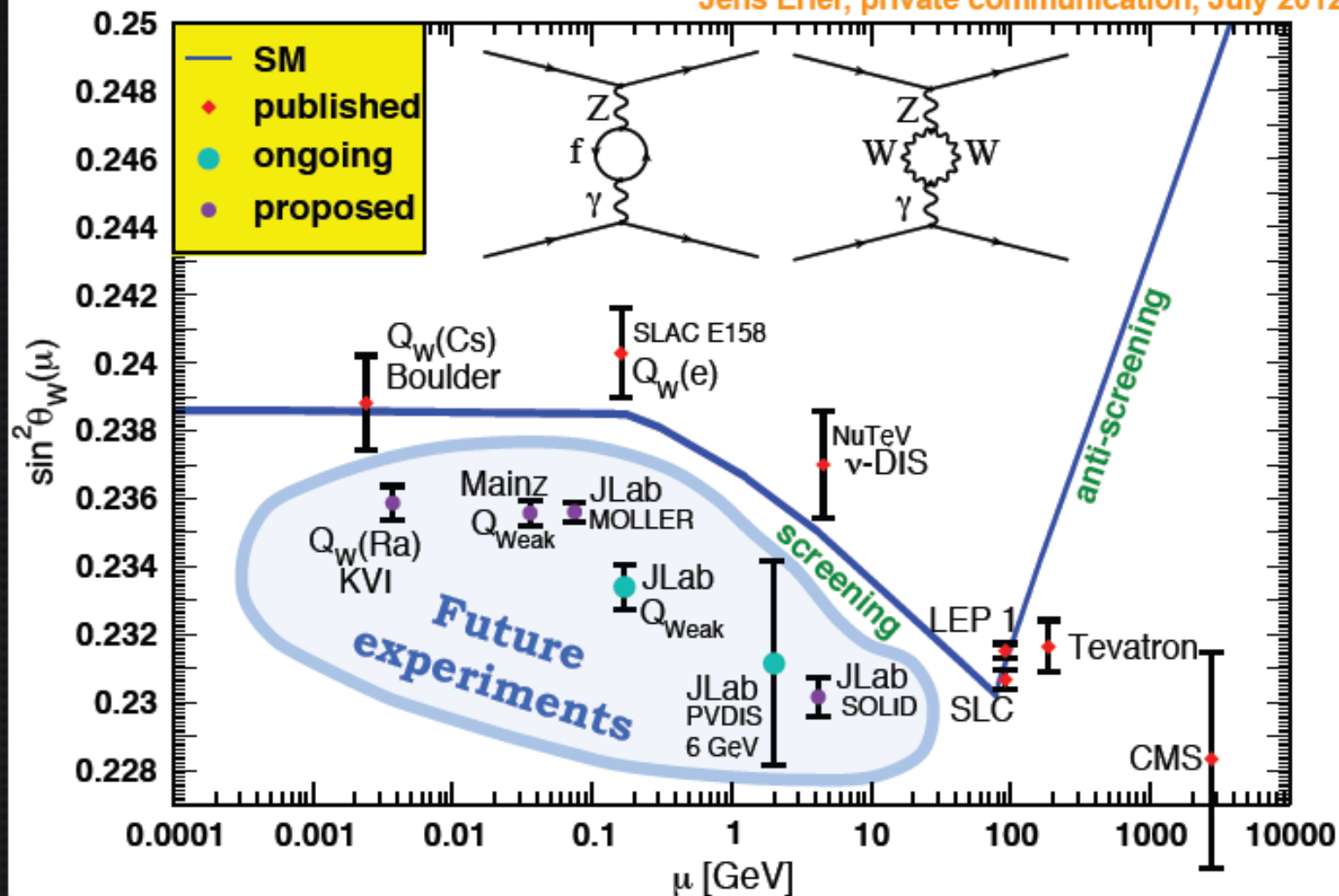
**Most precise
measurement from
Z to light-quark coupling**

**Statistical uncertainty
still dominant**

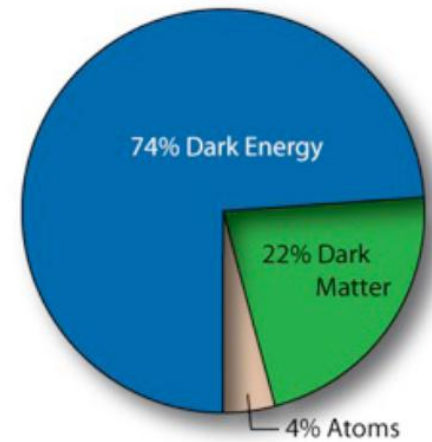
**Dominant systematic
uncertainty
PDF uncertainty
(0.00048)**

Weak mixing angle: scale dependence

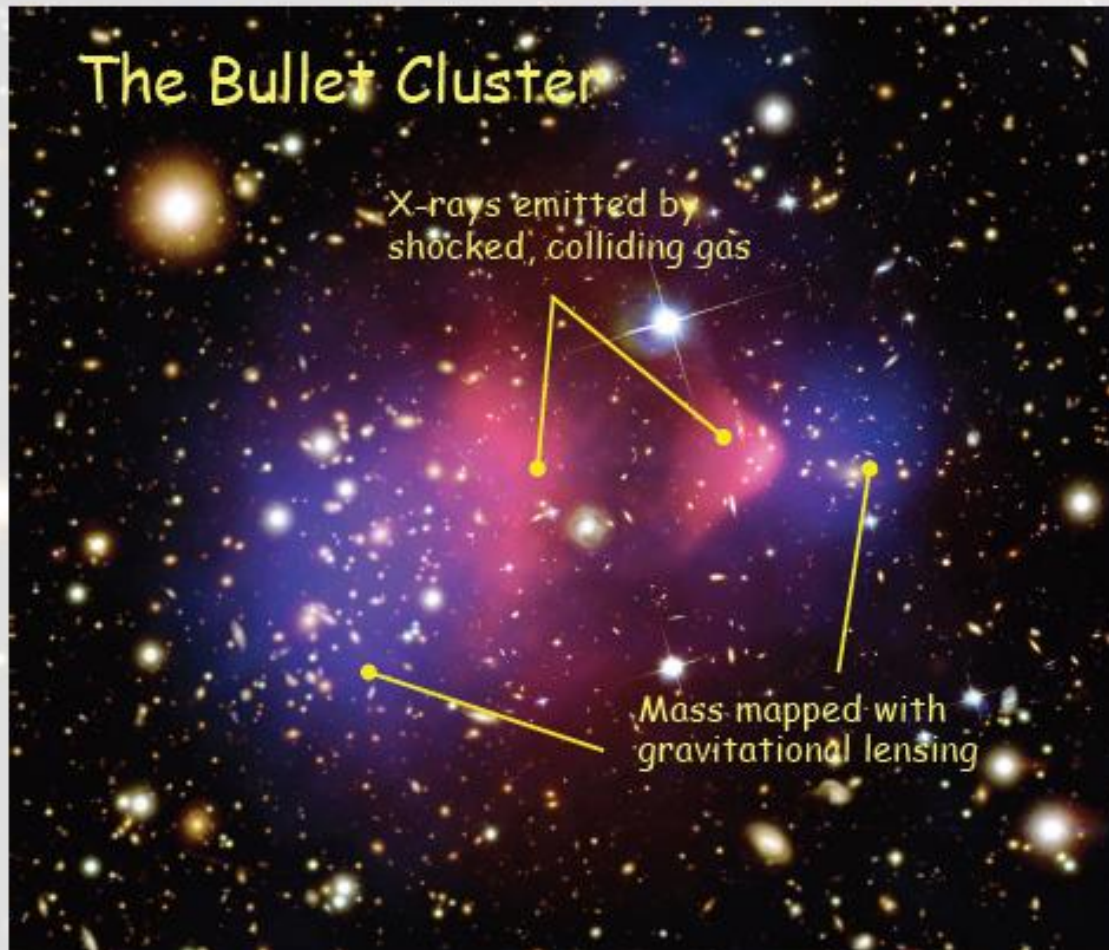
Jens Erler, private communication, July 2012



Scale dependence due to radiative effects



The Dark Matter Problem



What we know:

It's stable, cold, gravitationally interacting, non-baryonic, interacts little with itself (or not at all), composes ~85% of matter in the Universe...

But:

No particle in the Standard Model fits !

Very weakly interacting GeV-TeV particles do...

How to detect WIMPs

[Neal Weiner]

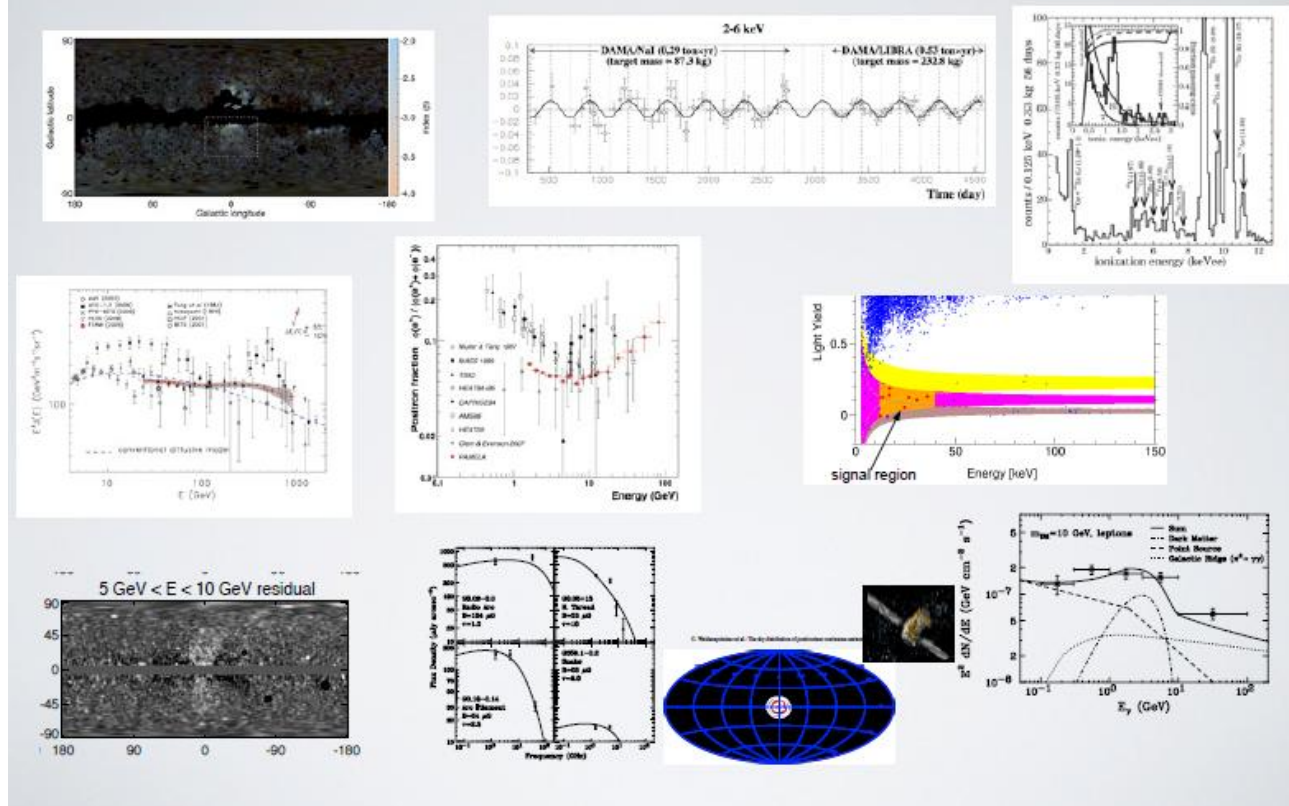
Our ideas of what dark matter is gives us
ideas on how to find it

How to detect WIMPs

[Neal Weiner]

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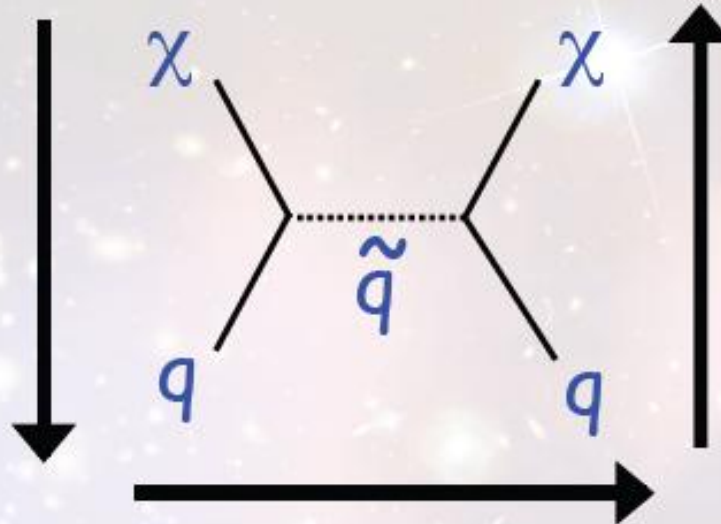
- Unlike the Higgs DM has been discovered many times



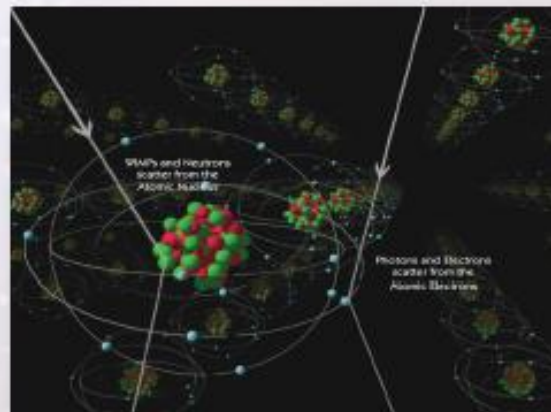
How to detect WIMPs



*Relic
annihilation in
the cosmos*
**INDIRECT
DETECTION**



LHC
*man-made COLLIDER
production*



M. Attisha

*Relic WIMP-
nucleon
elastic
scattering*
**DIRECT
DETECTION**

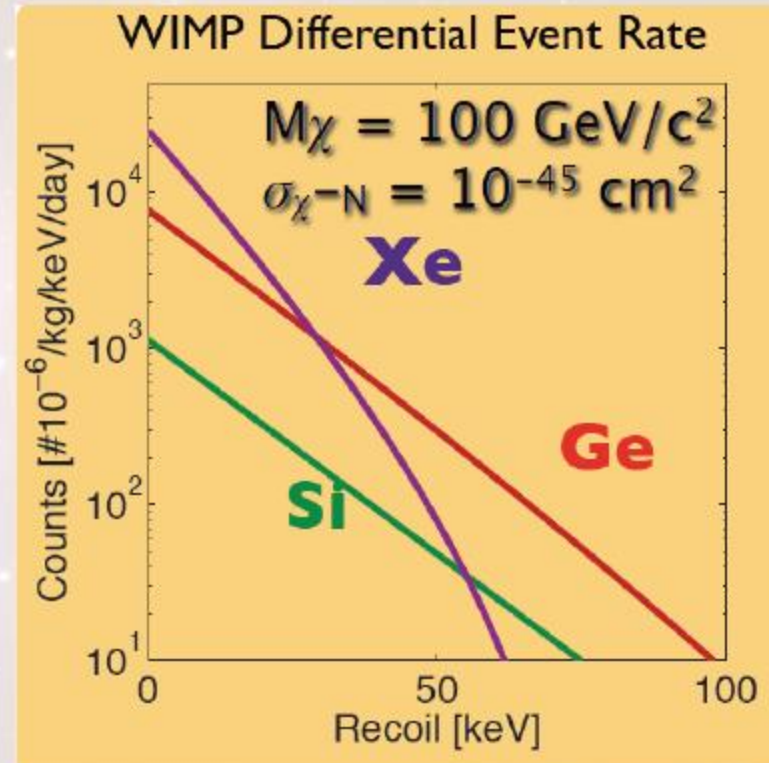
direct searches:

Expected signal:

- nuclear recoil (from elastic scattering of WIMPs)
- featureless exponential
- rates $\ll 0.1$ events /kg/day

Challenges:

- low energy thresholds (≤ 10 keV)
- mitigation of natural radioactive background (by factors $> 10^7$)
- long exposures, underground operation



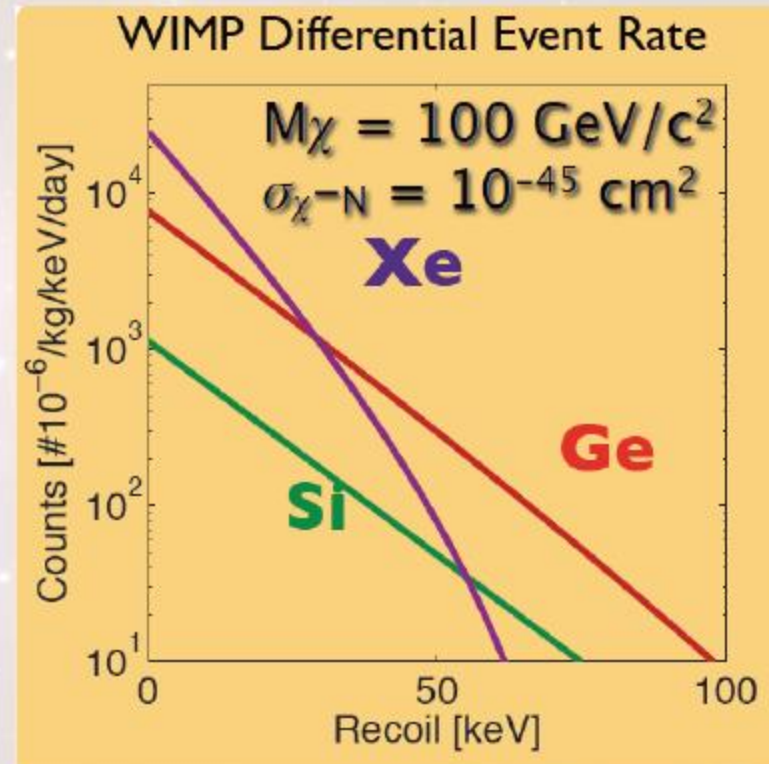
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Strangeness and Dark Matter

- Significant uncertainty coming from nucleon “sigma” terms

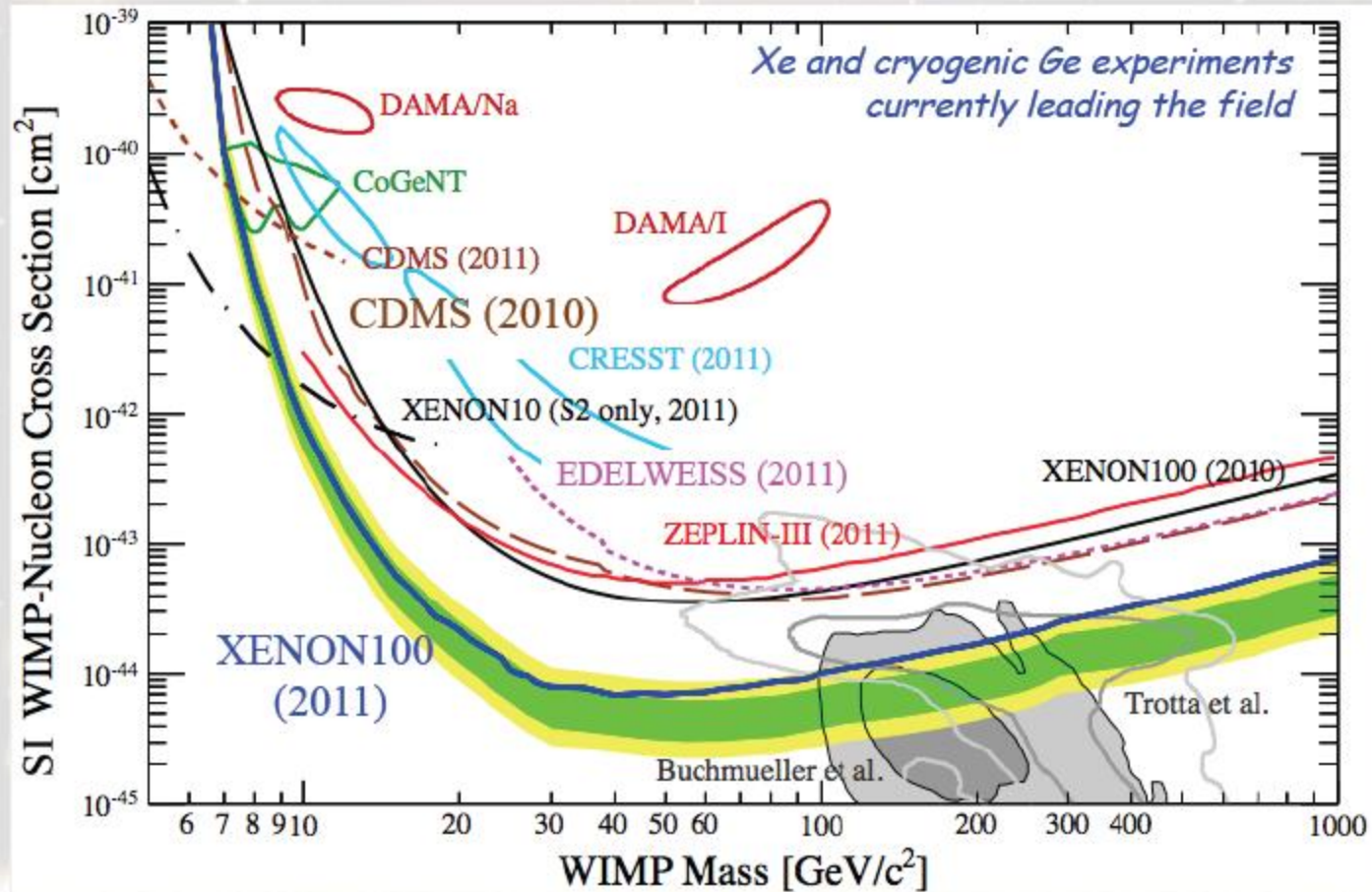
$$\sigma_q = m_q \langle N | \bar{q}q | N \rangle$$

lattice info

[James Zanotti]

direct searches:

Spin-Independent Landscape



direct searches:

Running experiments and those soon to be commissioned are about to explore one of the most interesting theoretical regions

Now that we think we've found the Higgs, will dark matter be the next great discovery of particle physics?

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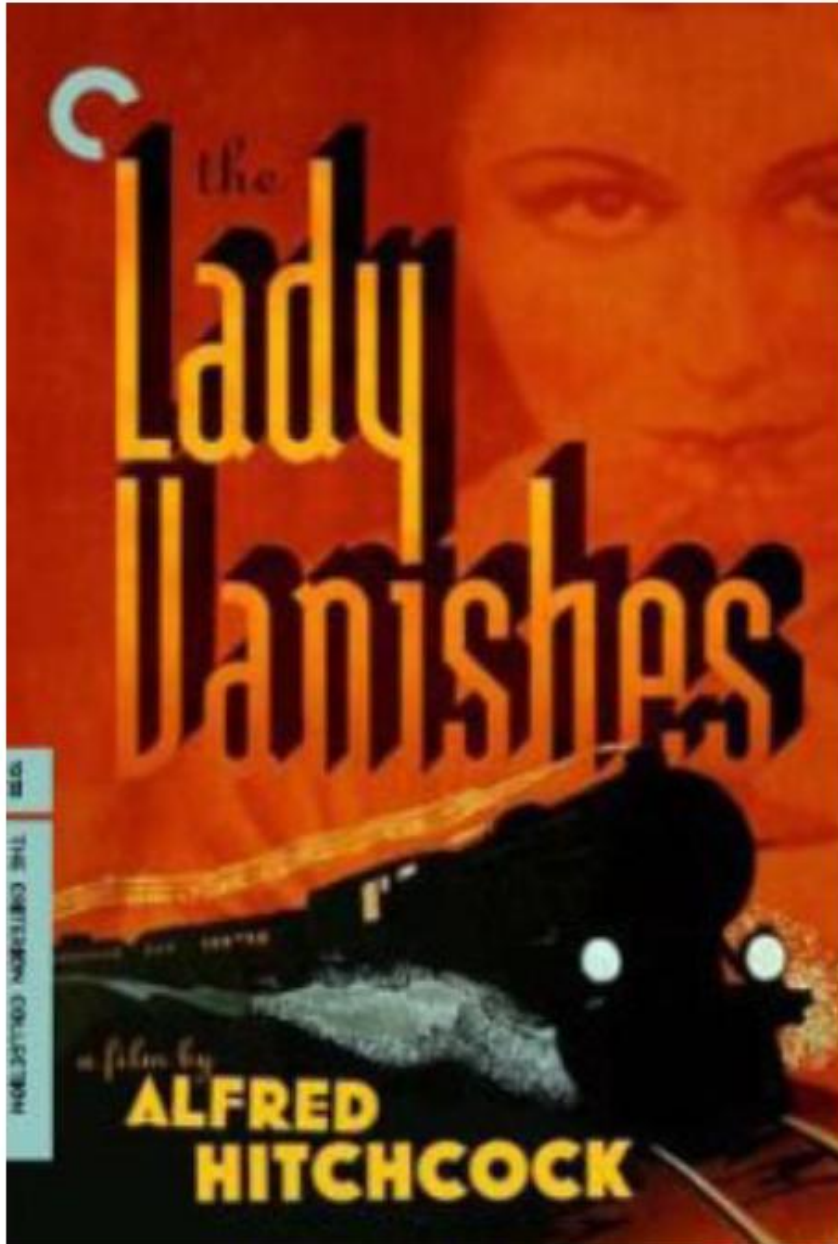
THE HIGGS AND DM IN SUSY

In SUSY the
Higgs is light
for no good
reason!

"X" may
keep other
things light,
too

[Neal Weiner]

and finally also a word about
SUSY



SUSY has been expected for a long time, but no trace has been found so far...

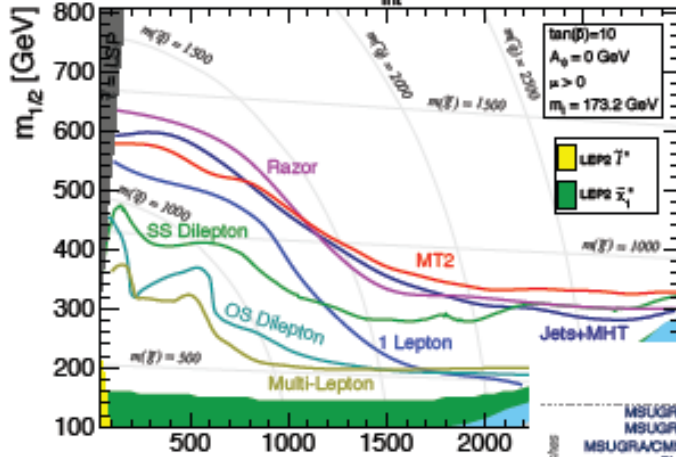
Like the plot of the excellent movie “The Lady Vanishes” (Alfred Hitchcock 1938).

A lady is seen, then disappears on a train:

- is she imaginary?
- has she been kidnapped and hidden?
- is she in disguise?
- is she dead?



CMS Preliminary $L_{int} = 4.98 \text{ fb}^{-1}, \sqrt{s} = 7 \text{ TeV}$



Both experiments provide nice search summaries



Larger versions in backup slides

ATLAS SUSY Searches* - 95% CL Lower Limits (Status: ICHEP 2012)

Search Category	Search Name	Lower Limit [GeV]	Notes
Inclusive searches	MSUGRA/CMSM: 0 lep + $\tilde{\nu}_\tau + E_{miss}$	1.26 TeV	$\tilde{g} \rightarrow \tilde{g} + \tilde{g}$
	MSUGRA/CMSM: 0 lep + multijets + E_{miss}	980 GeV	$\tilde{g} \rightarrow \tilde{g} + g$
	Pheno model: 0 lep + $\tilde{\nu}_\tau + E_{miss}$	1.3 TeV	$\tilde{g} \rightarrow \tilde{g} + \tilde{g}$ ($m_{\tilde{g}} < 2 \text{ TeV}, \mu > 0$)
	Pheno model: 0 lep + $\tilde{\nu}_\tau + E_{miss}$	980 GeV	$\tilde{g} \rightarrow \tilde{g} + \tilde{g}$ ($m_{\tilde{g}} < 2 \text{ TeV}, \mu > 0$)
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2nd gen squarks gluino-mediated	$\tilde{g} \rightarrow b\bar{b}$ (virtual): 0 lep + 1.2 b- $\tilde{\nu}_\tau + E_{miss}$	1.26 TeV	$\tilde{g} \rightarrow \tilde{g} + \tilde{g}$
	$\tilde{g} \rightarrow b\bar{b}$ (virtual): 0 lep + 3 b- $\tilde{\nu}_\tau + E_{miss}$	1.26 TeV	$\tilde{g} \rightarrow \tilde{g} + \tilde{g}$
	$\tilde{g} \rightarrow b\bar{b}$ (real): 0 lep + 3 b- $\tilde{\nu}_\tau + E_{miss}$	1.26 TeV	$\tilde{g} \rightarrow \tilde{g} + \tilde{g}$
	$\tilde{g} \rightarrow t\bar{t}$ (virtual): 1 lep + 1.2 b- $\tilde{\nu}_\tau + E_{miss}$	1.26 TeV	$\tilde{g} \rightarrow \tilde{g} + \tilde{g}$
	$\tilde{g} \rightarrow t\bar{t}$ (virtual): 2 lep (SS) + $\tilde{\nu}_\tau + E_{miss}$	1.26 TeV	$\tilde{g} \rightarrow \tilde{g} + \tilde{g}$
	$\tilde{g} \rightarrow t\bar{t}$ (virtual): 0 lep + multijets + E_{miss}	1.26 TeV	$\tilde{g} \rightarrow \tilde{g} + \tilde{g}$
	$\tilde{g} \rightarrow t\bar{t}$ (real): 0 lep + 3 b- $\tilde{\nu}_\tau + E_{miss}$	1.26 TeV	$\tilde{g} \rightarrow \tilde{g} + \tilde{g}$
	$\tilde{g} \rightarrow t\bar{t}$ (real): 0 lep + 3 b- $\tilde{\nu}_\tau + E_{miss}$	1.26 TeV	$\tilde{g} \rightarrow \tilde{g} + \tilde{g}$
	$\tilde{g} \rightarrow t\bar{t}$ (real): 0 lep + 3 b- $\tilde{\nu}_\tau + E_{miss}$	1.26 TeV	$\tilde{g} \rightarrow \tilde{g} + \tilde{g}$
	$\tilde{g} \rightarrow t\bar{t}$ (real): 0 lep + 3 b- $\tilde{\nu}_\tau + E_{miss}$	1.26 TeV	$\tilde{g} \rightarrow \tilde{g} + \tilde{g}$
2nd gen squarks direct production	$\tilde{g} \rightarrow b\bar{b}$: 0 lep + 2 b- $\tilde{\nu}_\tau + E_{miss}$	980 GeV	$\tilde{g} \rightarrow \tilde{g} + \tilde{g}$
	$\tilde{g} \rightarrow t\bar{t}$ (very light): 1.2 lep + b- $\tilde{\nu}_\tau + E_{miss}$	1.26 TeV	$\tilde{g} \rightarrow \tilde{g} + \tilde{g}$
	$\tilde{g} \rightarrow t\bar{t}$ (heavy): 1.2 lep + b- $\tilde{\nu}_\tau + E_{miss}$	1.26 TeV	$\tilde{g} \rightarrow \tilde{g} + \tilde{g}$
	$\tilde{g} \rightarrow t\bar{t}$ (heavy): 1.2 lep + b- $\tilde{\nu}_\tau + E_{miss}$	1.26 TeV	$\tilde{g} \rightarrow \tilde{g} + \tilde{g}$
	$\tilde{g} \rightarrow t\bar{t}$ (heavy): 1.2 lep + b- $\tilde{\nu}_\tau + E_{miss}$	1.26 TeV	$\tilde{g} \rightarrow \tilde{g} + \tilde{g}$
	$\tilde{g} \rightarrow t\bar{t}$ (heavy): 1.2 lep + b- $\tilde{\nu}_\tau + E_{miss}$	1.26 TeV	$\tilde{g} \rightarrow \tilde{g} + \tilde{g}$
	$\tilde{g} \rightarrow t\bar{t}$ (heavy): 1.2 lep + b- $\tilde{\nu}_\tau + E_{miss}$	1.26 TeV	$\tilde{g} \rightarrow \tilde{g} + \tilde{g}$
	$\tilde{g} \rightarrow t\bar{t}$ (heavy): 1.2 lep + b- $\tilde{\nu}_\tau + E_{miss}$	1.26 TeV	$\tilde{g} \rightarrow \tilde{g} + \tilde{g}$
	$\tilde{g} \rightarrow t\bar{t}$ (heavy): 1.2 lep + b- $\tilde{\nu}_\tau + E_{miss}$	1.26 TeV	$\tilde{g} \rightarrow \tilde{g} + \tilde{g}$
	$\tilde{g} \rightarrow t\bar{t}$ (heavy): 1.2 lep + b- $\tilde{\nu}_\tau + E_{miss}$	1.26 TeV	$\tilde{g} \rightarrow \tilde{g} + \tilde{g}$
RPV direct	$\tilde{g} \rightarrow b\bar{b}$: 2 lep + $\tilde{\nu}_\tau + E_{miss}$	980 GeV	$\tilde{g} \rightarrow \tilde{g} + \tilde{g}$
	$\tilde{g} \rightarrow t\bar{t}$: 2 lep + $\tilde{\nu}_\tau + E_{miss}$	980 GeV	$\tilde{g} \rightarrow \tilde{g} + \tilde{g}$
	$\tilde{g} \rightarrow t\bar{t}$: 3 lep + $\tilde{\nu}_\tau + E_{miss}$	980 GeV	$\tilde{g} \rightarrow \tilde{g} + \tilde{g}$
	$\tilde{g} \rightarrow t\bar{t}$: 3 lep + $\tilde{\nu}_\tau + E_{miss}$	980 GeV	$\tilde{g} \rightarrow \tilde{g} + \tilde{g}$
	$\tilde{g} \rightarrow t\bar{t}$: 3 lep + $\tilde{\nu}_\tau + E_{miss}$	980 GeV	$\tilde{g} \rightarrow \tilde{g} + \tilde{g}$
	$\tilde{g} \rightarrow t\bar{t}$: 3 lep + $\tilde{\nu}_\tau + E_{miss}$	980 GeV	$\tilde{g} \rightarrow \tilde{g} + \tilde{g}$
	$\tilde{g} \rightarrow t\bar{t}$: 3 lep + $\tilde{\nu}_\tau + E_{miss}$	980 GeV	$\tilde{g} \rightarrow \tilde{g} + \tilde{g}$
	$\tilde{g} \rightarrow t\bar{t}$: 3 lep + $\tilde{\nu}_\tau + E_{miss}$	980 GeV	$\tilde{g} \rightarrow \tilde{g} + \tilde{g}$
	$\tilde{g} \rightarrow t\bar{t}$: 3 lep + $\tilde{\nu}_\tau + E_{miss}$	980 GeV	$\tilde{g} \rightarrow \tilde{g} + \tilde{g}$
	$\tilde{g} \rightarrow t\bar{t}$: 3 lep + $\tilde{\nu}_\tau + E_{miss}$	980 GeV	$\tilde{g} \rightarrow \tilde{g} + \tilde{g}$
Long-lived particles	Stable \tilde{R} -hadrons: Full detector	980 GeV	$\tilde{g} \rightarrow \tilde{g} + \tilde{g}$
	Stable \tilde{B} -hadrons: Full detector	980 GeV	$\tilde{g} \rightarrow \tilde{g} + \tilde{g}$
	Stable \tilde{T} -hadrons: Full detector	980 GeV	$\tilde{g} \rightarrow \tilde{g} + \tilde{g}$
	Metastable \tilde{R} -hadrons: Full det. only	980 GeV	$\tilde{g} \rightarrow \tilde{g} + \tilde{g}$
	GMSB: stable $\tilde{\chi}_1^0$	980 GeV	$\tilde{g} \rightarrow \tilde{g} + \tilde{g}$
	RPV: high-mass $\tilde{\chi}_1^0$	980 GeV	$\tilde{g} \rightarrow \tilde{g} + \tilde{g}$
	Stable \tilde{R} -hadrons: Full detector	980 GeV	$\tilde{g} \rightarrow \tilde{g} + \tilde{g}$
	Stable \tilde{B} -hadrons: Full detector	980 GeV	$\tilde{g} \rightarrow \tilde{g} + \tilde{g}$
	Stable \tilde{T} -hadrons: Full detector	980 GeV	$\tilde{g} \rightarrow \tilde{g} + \tilde{g}$
	Metastable \tilde{R} -hadrons: Full det. only	980 GeV	$\tilde{g} \rightarrow \tilde{g} + \tilde{g}$
RPV	Bilinear RPV: 1 lep + $\tilde{\nu}_\tau + E_{miss}$	980 GeV	$\tilde{g} \rightarrow \tilde{g} + \tilde{g}$
	BC1 RPV: 4 lep + E_{miss}	980 GeV	$\tilde{g} \rightarrow \tilde{g} + \tilde{g}$
	Hyperscalar mass (gluons): 4 lep + E_{miss}	980 GeV	$\tilde{g} \rightarrow \tilde{g} + \tilde{g}$
	Spin dep. WIMP interaction: monojet + E_{miss}	980 GeV	$\tilde{g} \rightarrow \tilde{g} + \tilde{g}$
	Spin indep. WIMP interaction: monojet + E_{miss}	980 GeV	$\tilde{g} \rightarrow \tilde{g} + \tilde{g}$
	Spin dep. WIMP interaction: monojet + E_{miss}	980 GeV	$\tilde{g} \rightarrow \tilde{g} + \tilde{g}$
	Spin indep. WIMP interaction: monojet + E_{miss}	980 GeV	$\tilde{g} \rightarrow \tilde{g} + \tilde{g}$
	Spin dep. WIMP interaction: monojet + E_{miss}	980 GeV	$\tilde{g} \rightarrow \tilde{g} + \tilde{g}$
	Spin indep. WIMP interaction: monojet + E_{miss}	980 GeV	$\tilde{g} \rightarrow \tilde{g} + \tilde{g}$
	Spin dep. WIMP interaction: monojet + E_{miss}	980 GeV	$\tilde{g} \rightarrow \tilde{g} + \tilde{g}$

*Only a selection of the available mass limits on new states or phenomena shown

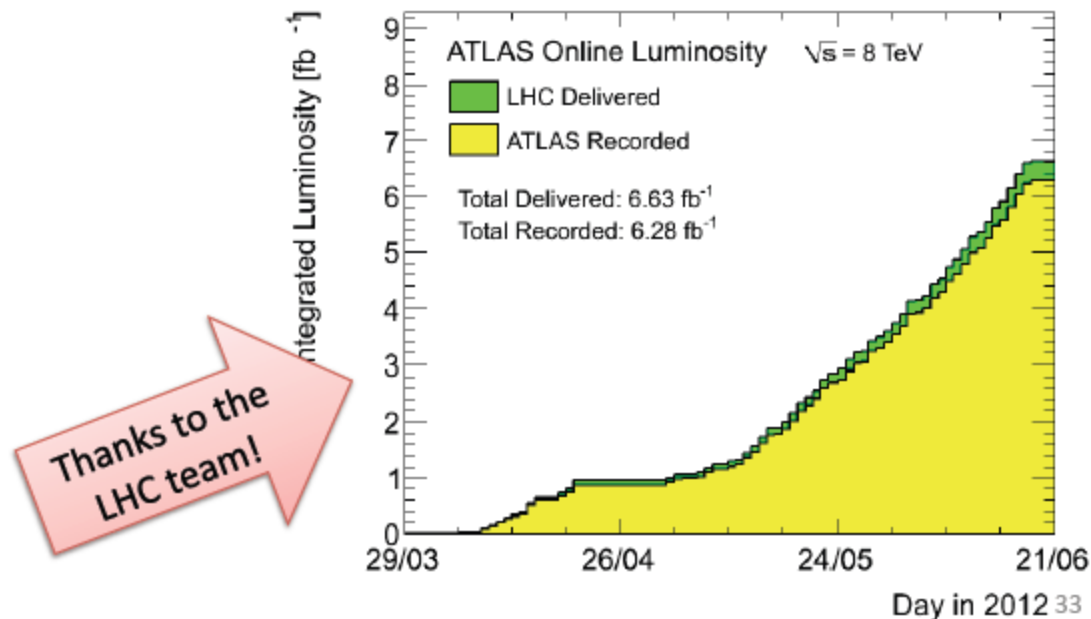
Is SUSY Dead?

- The searches leave little room for SUSY inside the reach of the existing data.
- But interpretations within SUSY models rely on many simplifying assumptions, and so care must be taken when making use of the limit plots

Is SUSY Dead?

- The searches leave little room for SUSY inside the reach of the existing data.
- But interpretations within SUSY models rely on many simplifying assumptions, and so care must be taken when making use of the limit plots

Maybe a happy ending....?



Is SUSY Dead?

The lady is found alive and well in the final scene...



I apologize for not having covered your favourite topic

lattice: information
from first principles

B-physics : unitarity
triangle & NP

no time to cover any detail of...

flavour-physics :
CPV

HIC : the quest for
the QGP

please consult: www.ichep2012.com.au/Program/



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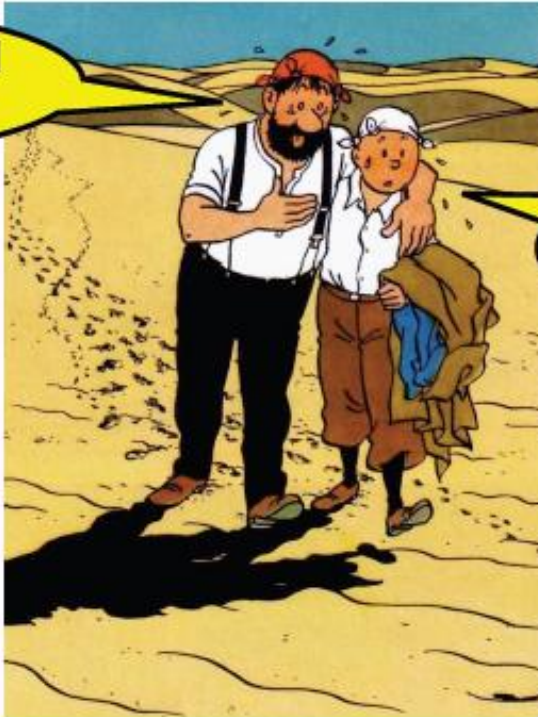
... in summary

4th of July 2012

[Alex Pomarol]

A very stirring day for the EWSB practitioners

**We've been more than 40 years
of mainly wandering in the desert...**



Do you think we'll
find a Higgs?

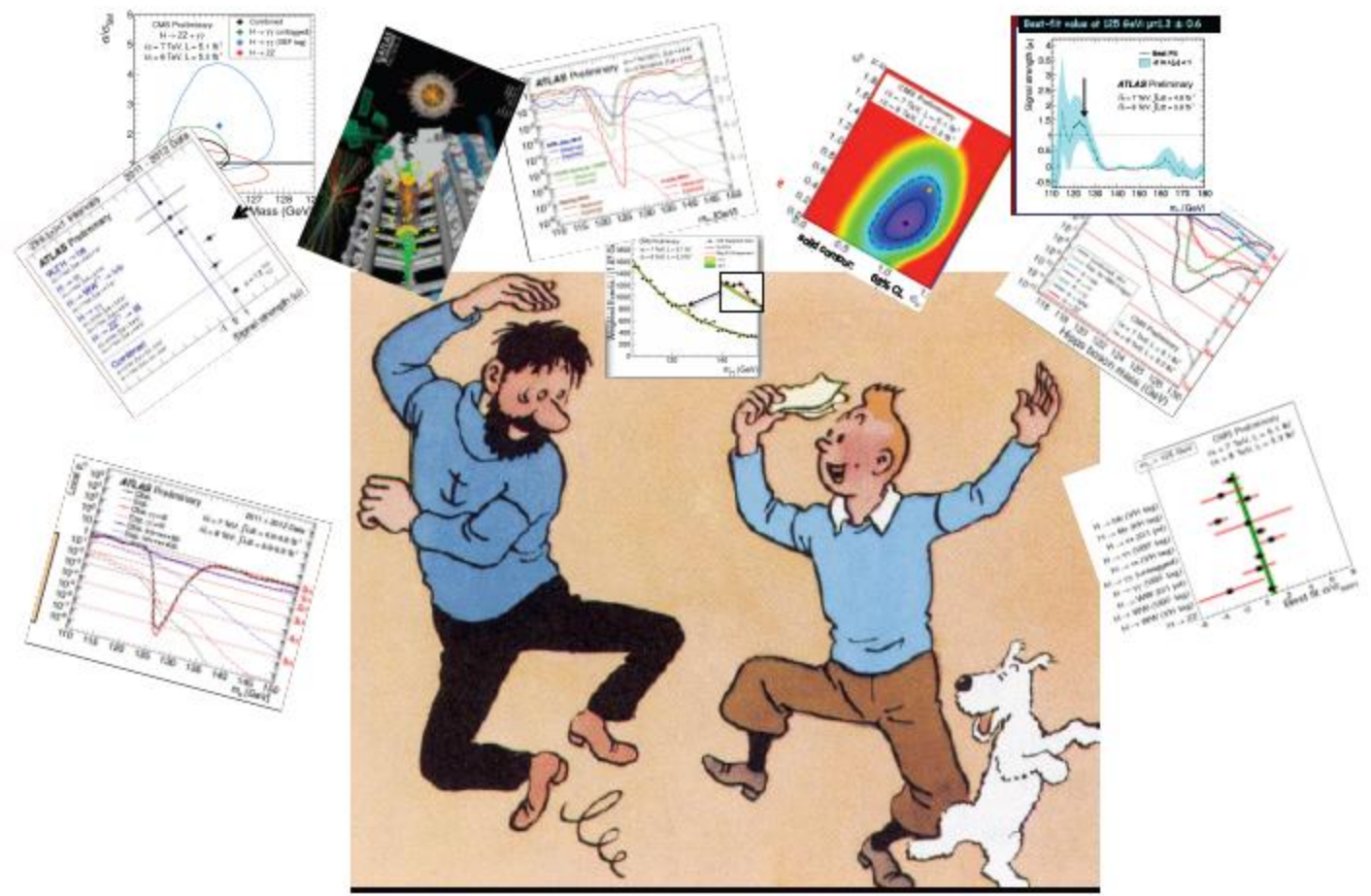
A cartoon illustration of two men walking through a desert. The man on the left is taller, has a beard, and is wearing a red cap, a white shirt with suspenders, and black pants. The man on the right is shorter, wearing a white cap, a white shirt, and brown pants, and is carrying a brown jacket. They are walking on a sandy path with shadows cast on the ground.

I don't know,
this looks Higgsless

... in summary

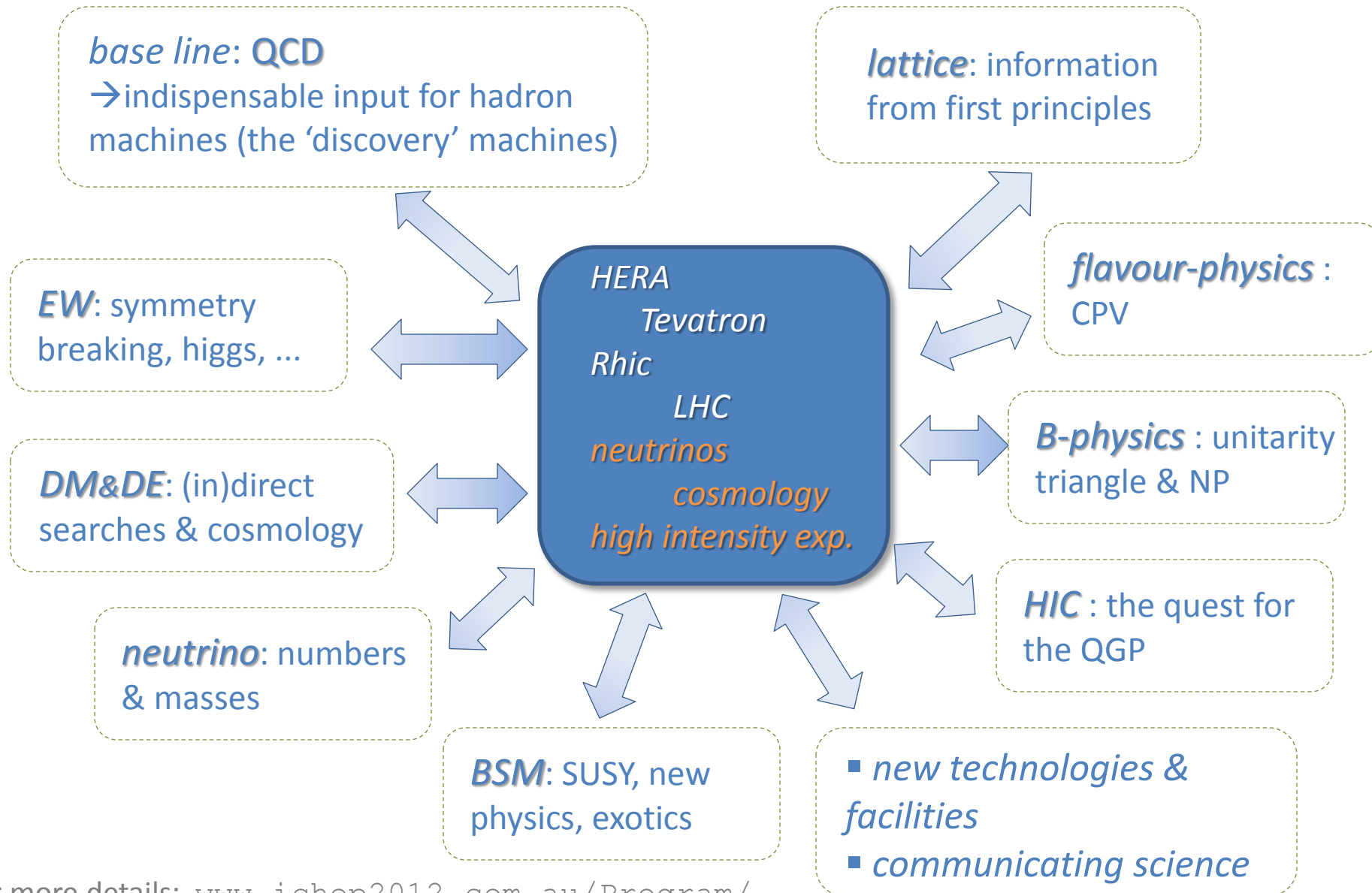
[Alex Pomarol]

... and finally plenty of new relevant data has begun to fall over us!





exciting news from the high energy & high intensity frontiers





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very stimulating conference
& amazing new results
still many ways to go ...

[Maringka Baker: Pukara]





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